

Overview of PandaX-4T experiment

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On behalf of the PandaX Collaboration

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PANDA X
PARTICLE AND ASTROPHYSICAL XENON TPC



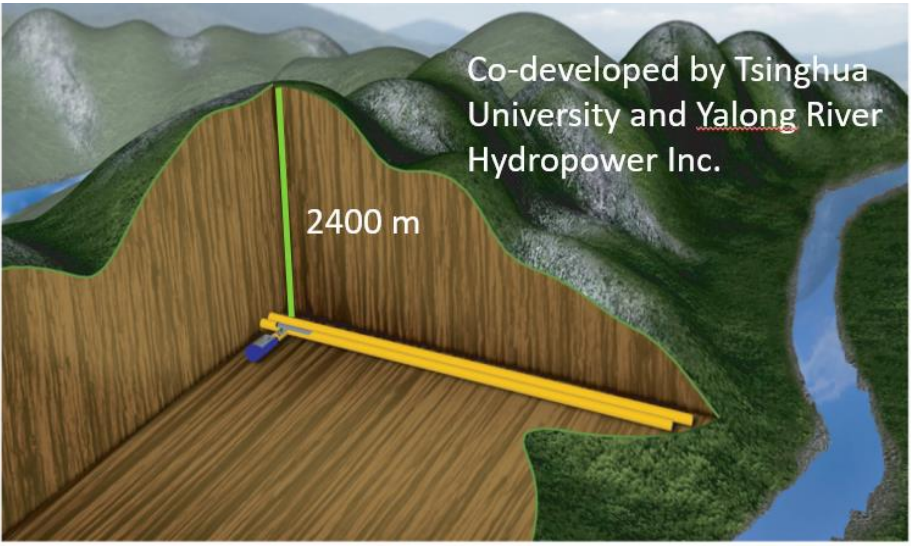
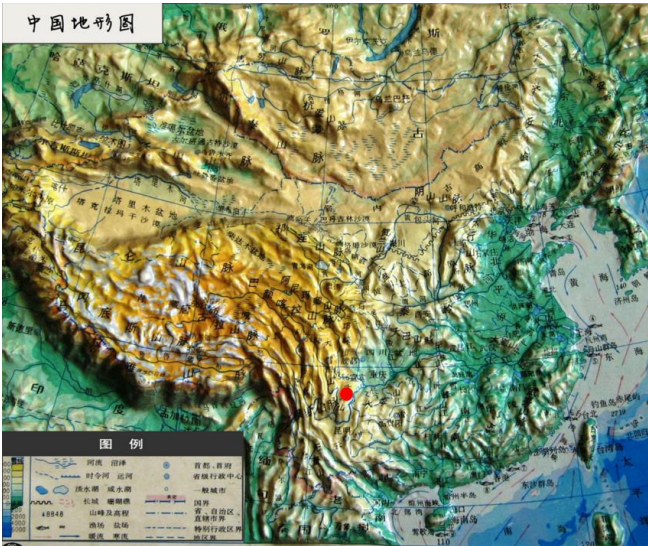
PandaX Collaboration



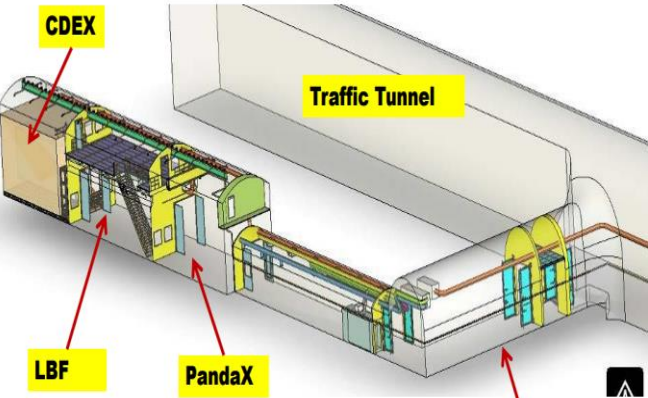
PANDA X



China JinPing Underground Laboratory – CJPL

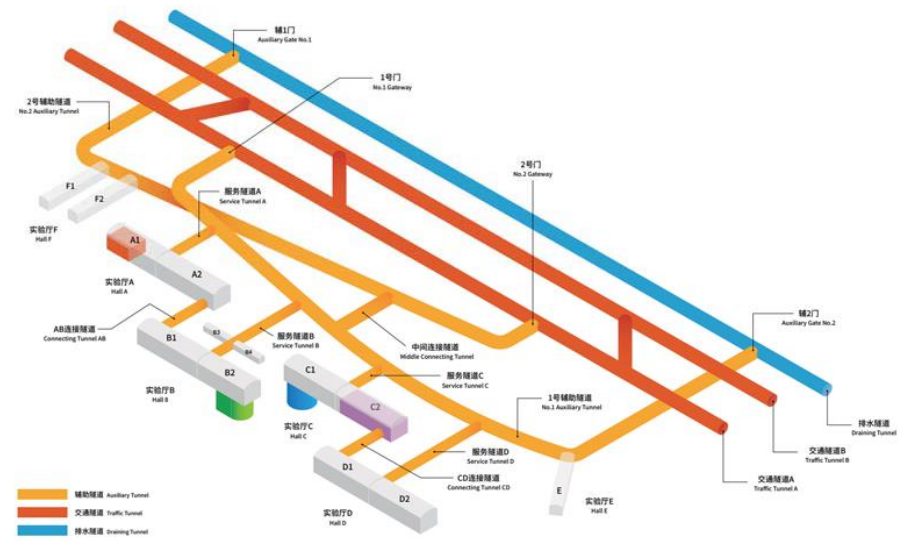


- Deepest (6800 m.w.e)
- Horizontal access
- Muon rate: 1 count/week/m²
- From CJPL-I to CJPL-II



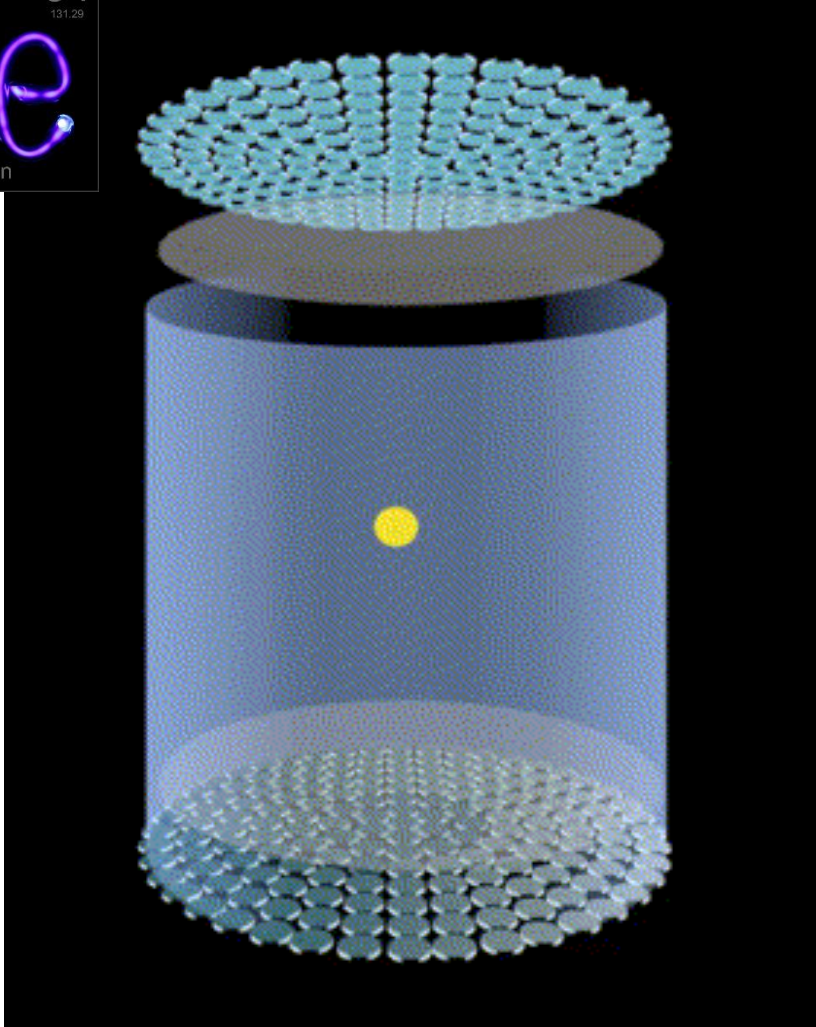
■ Total space: 4000 m³
■ Main Lab Space: 6.5(W) x 6.5(H) x 42(L)

2023/6/22

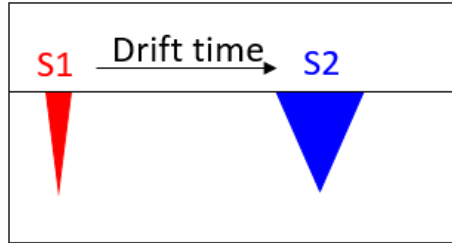


XeSAT2023

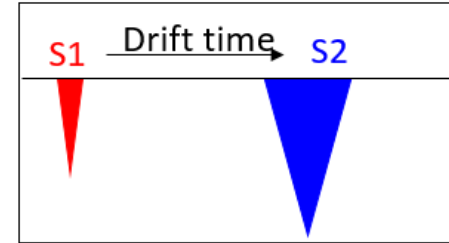
Dual Phase Liquid Xenon TPC



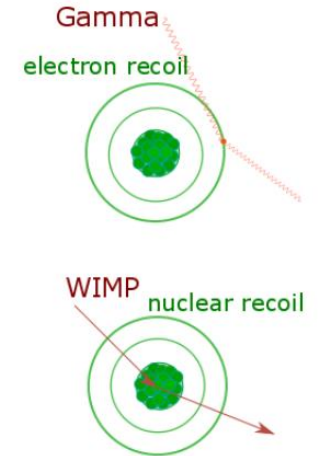
Dark matter: nuclear recoil (NR)



γ background: electron recoil (ER)



$$(S2/S1)_{NR} \ll (S2/S1)_{ER}$$



- Purity liquid xenon target, high light & charge yield;
- Good ER/NR discrimination by S2/S1 ratio;
- 3D reconstruction rejects external background;

PandaX Roadmap

PANDA X Particle and Astrophysical Xenon Experiments

Collaboration formed



2009.3

2014.5-10



PandaX-I, 120 kg operation

PandaX-II, 580 kg operation



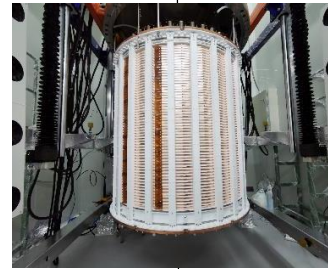
2016.7-2019.7

2019.8



PandaX-4T moved to CJPL-II

PandaX-4T Commissioning



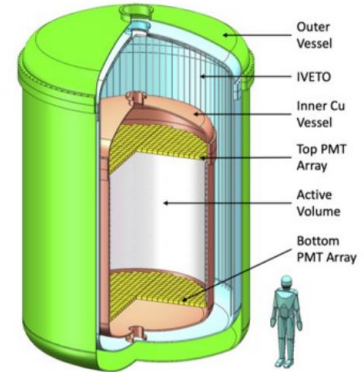
2020.11-2021.5

2021.11



PandaX-4T Run1

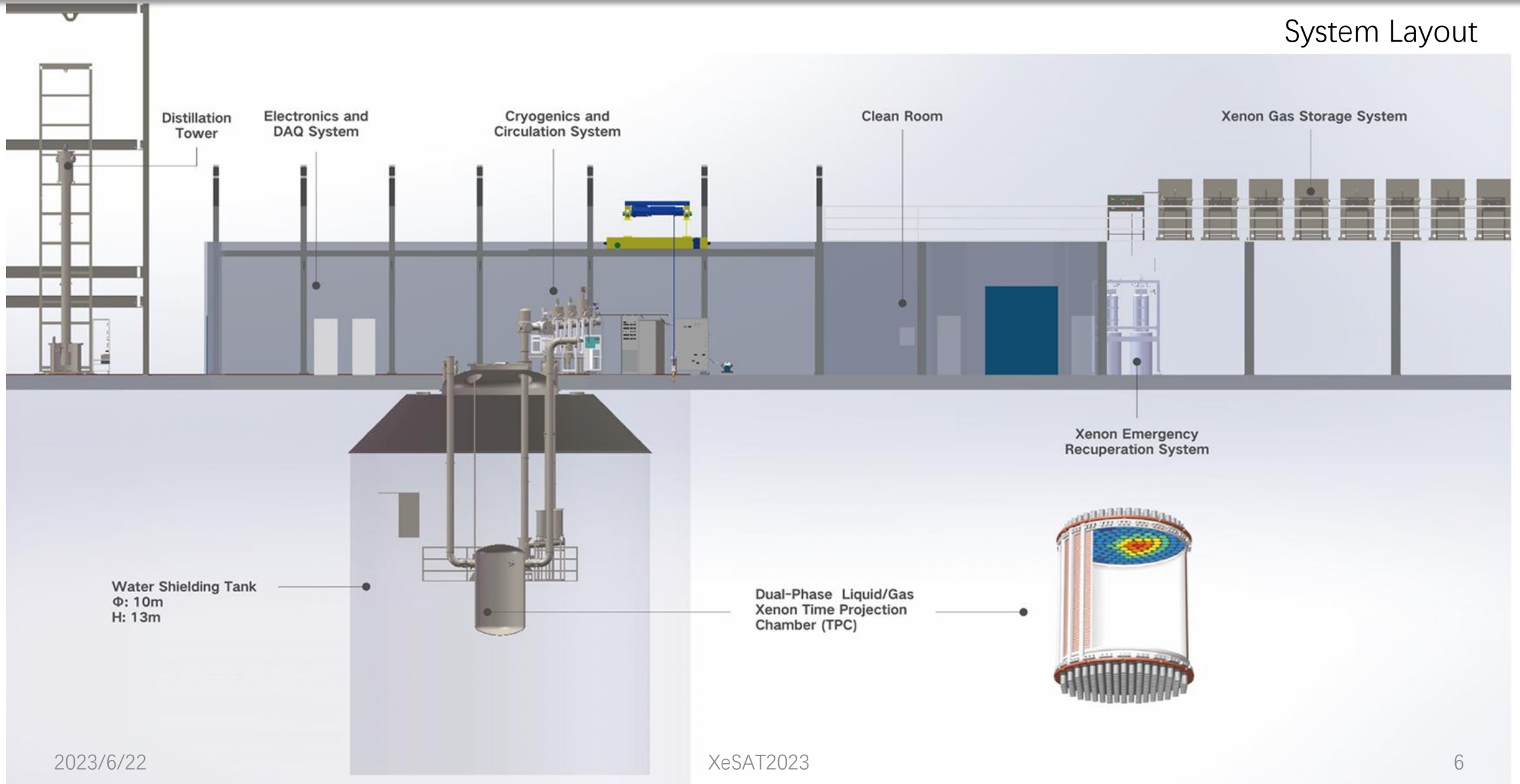
Ongoing



PandaX-xT

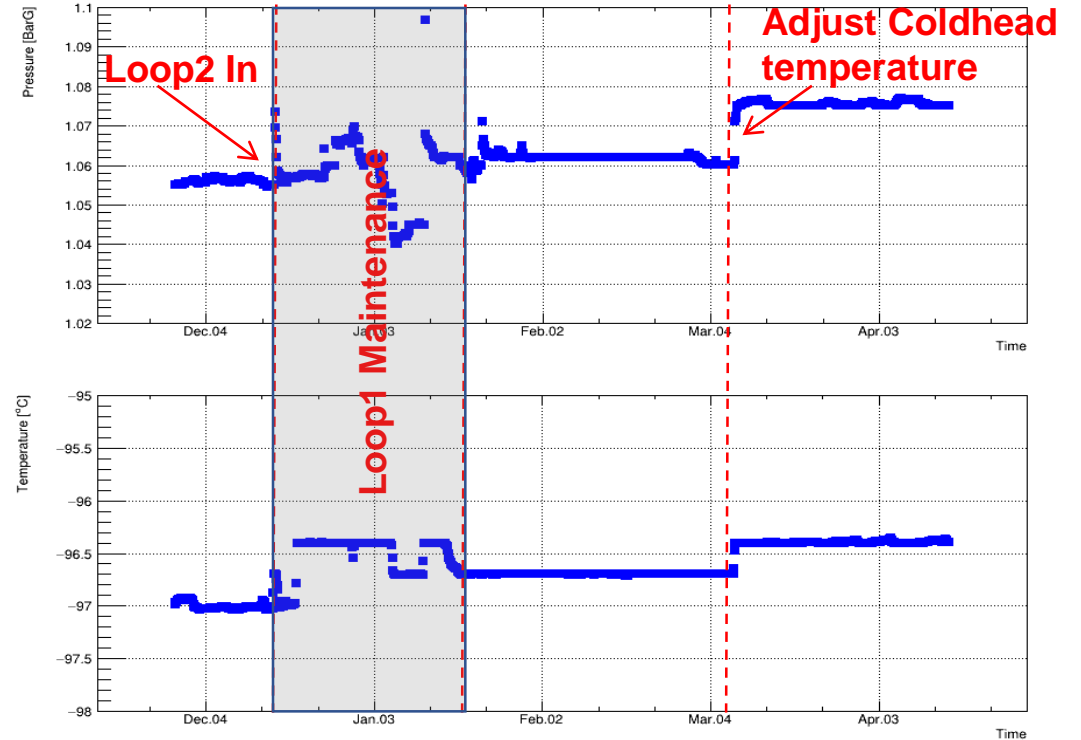
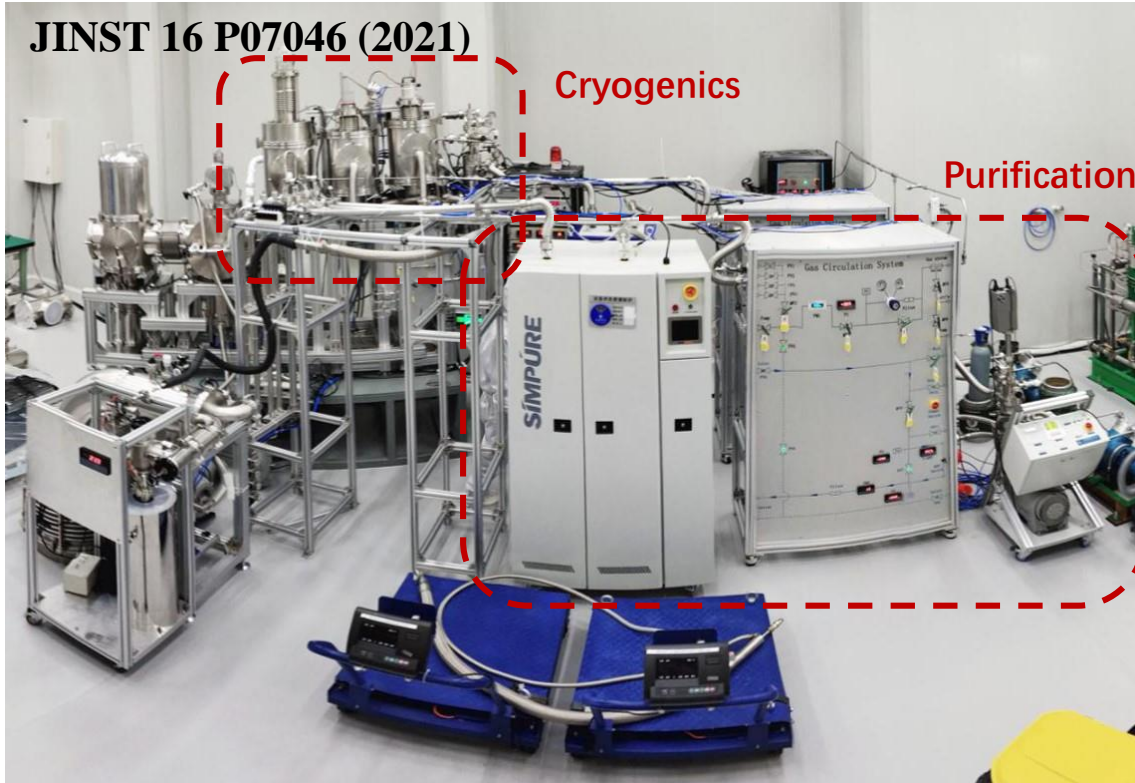
PandaX-4T Overview

System Layout



Cryogenics System

- Pressure and temperature stability within 0.5% and 0.1K .

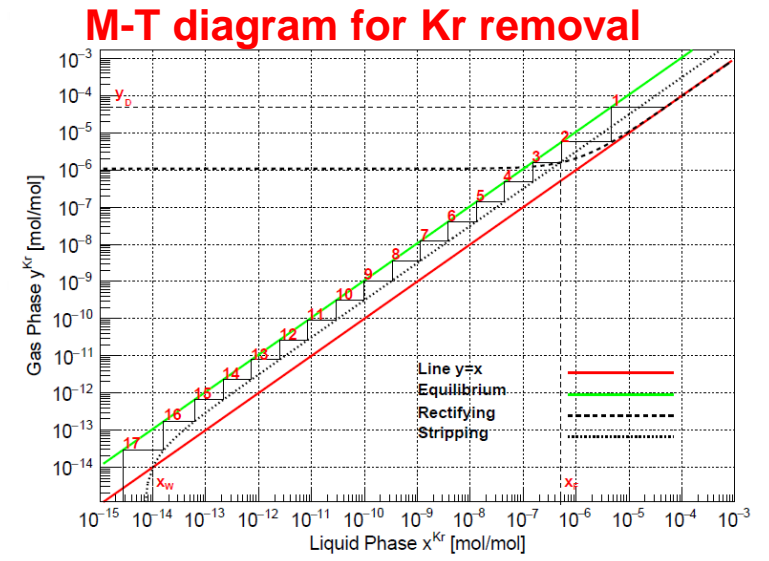


Parameters	Heating load (No purification)	Maximum Cooling Power	Filling/Recuperation flow rate	Outer Vacuum
Value	~50 W	~580 W	~1 ton/day	<2E-4 Pa

Distillation System

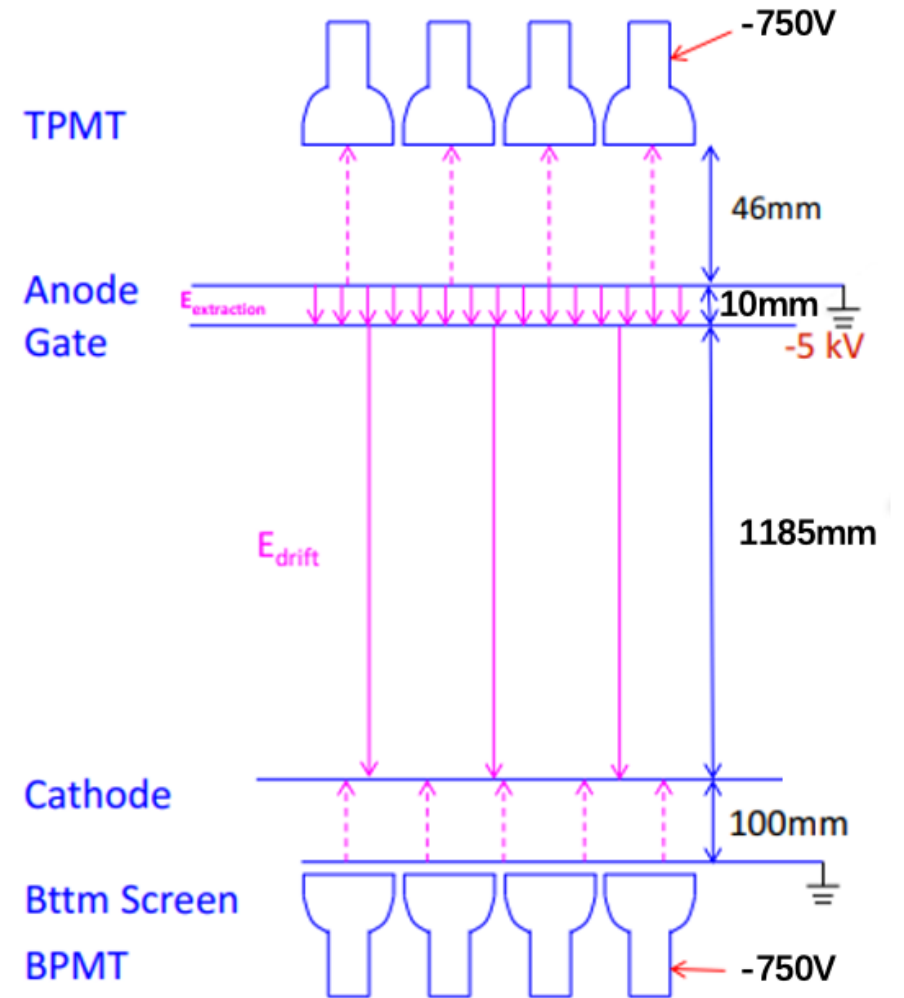
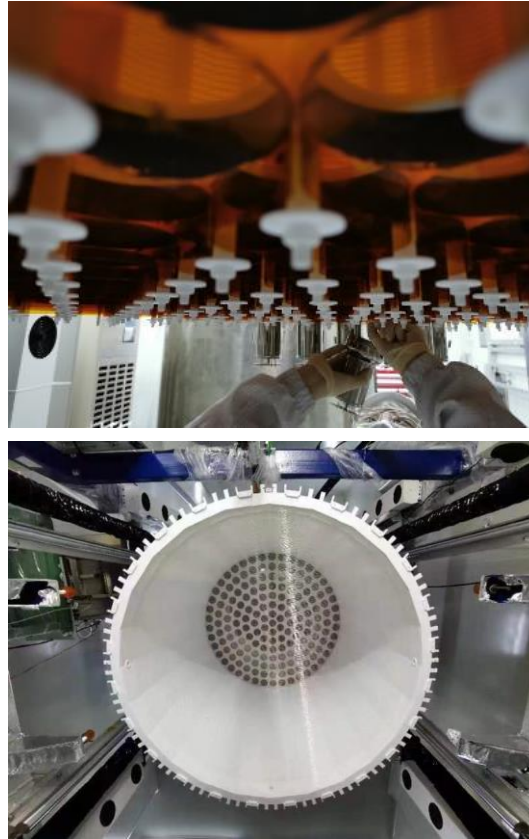


Structured packing (inside)



- Distillation method for the LXe intrinsic background Kr & Rn removal;
- $\sim 10^6$ reduction factor for Kr removal with 10 kg/h;
- Reversed operation mode working for Rn removal;

TPC Conditions



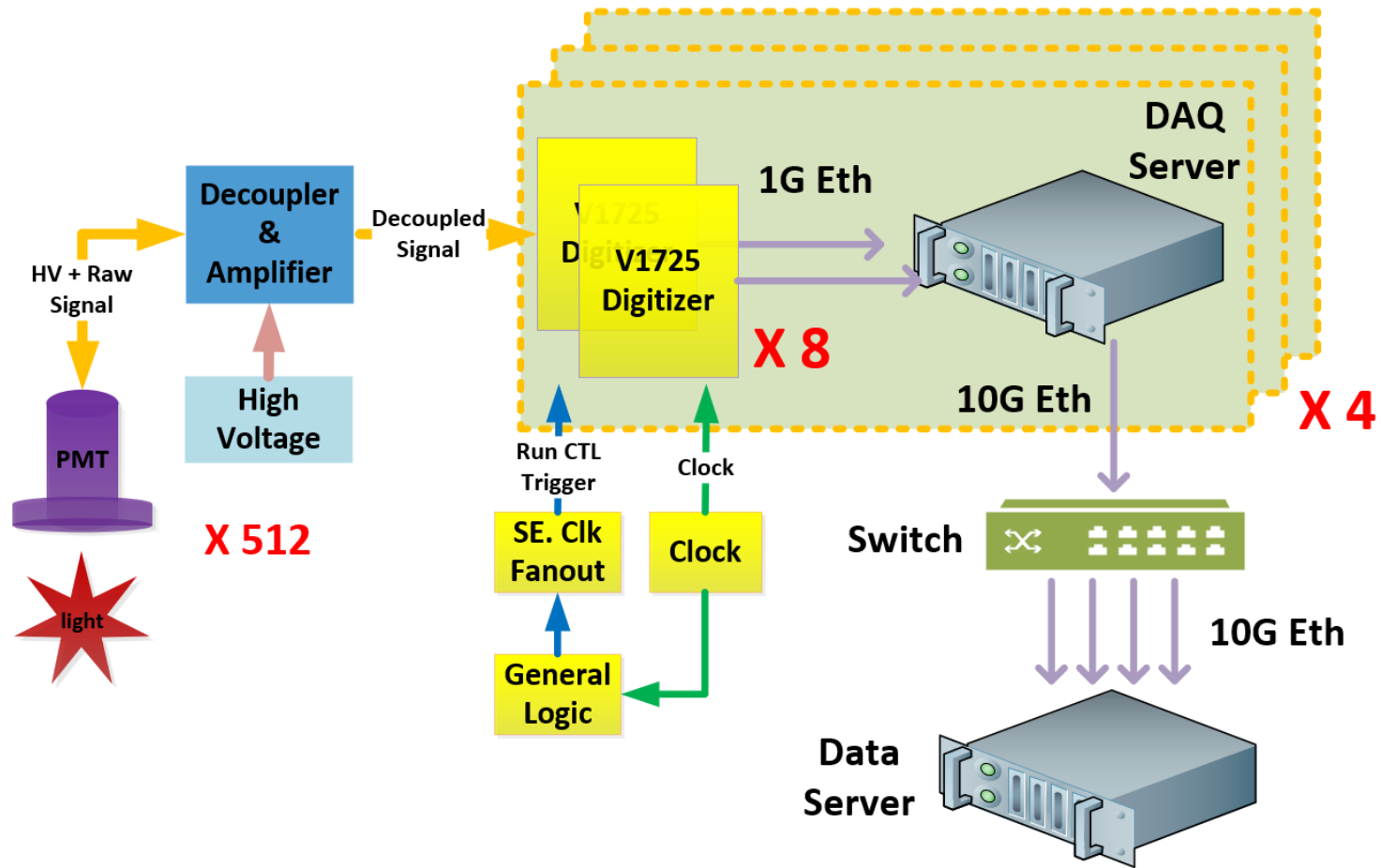
	Set1	Set2	Set3	Set4	Set5
Gate(kV)	-4.9	-5	-5	-5	-5
Cathode (kV)	-20	-18.6	-18	-16	-16

PMT Arrays



- 169 top + 199 bottom R11410-23 3-inch PMTs, with the average gain of 5.5×10^6 ;
- LED calibration every week, monitor gain stability;

Electronics



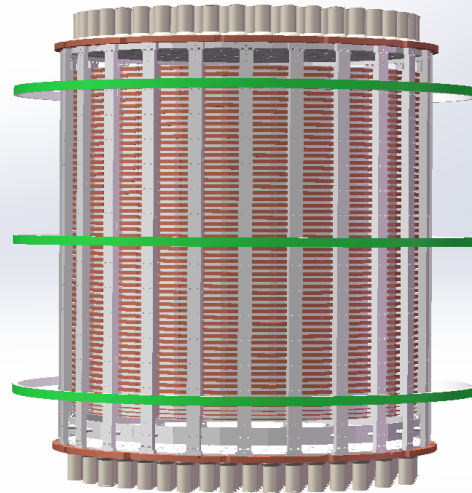
- From PMT to DAQ server;
- CERN V1725 Digitizer, 250 MS/s;
- Self-trigger mode: read out pulses above 20 ADC ($\sim 1/3$ PE);

Calibration System

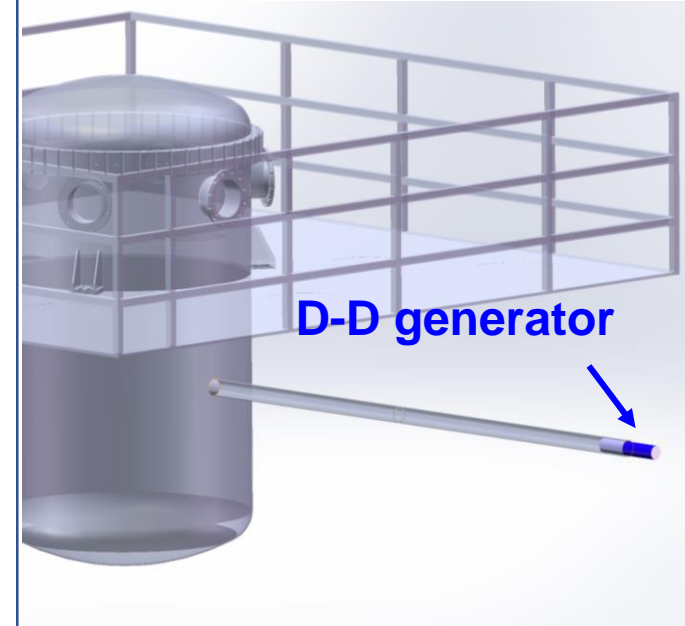
Gaseous source injection panel



Calibration tubes



D-D generator



Calibration source

$^{83m}\text{Kr}/^{220}\text{Rn}$

$^{241}\text{Am}-\text{Be}$

D-D neutron

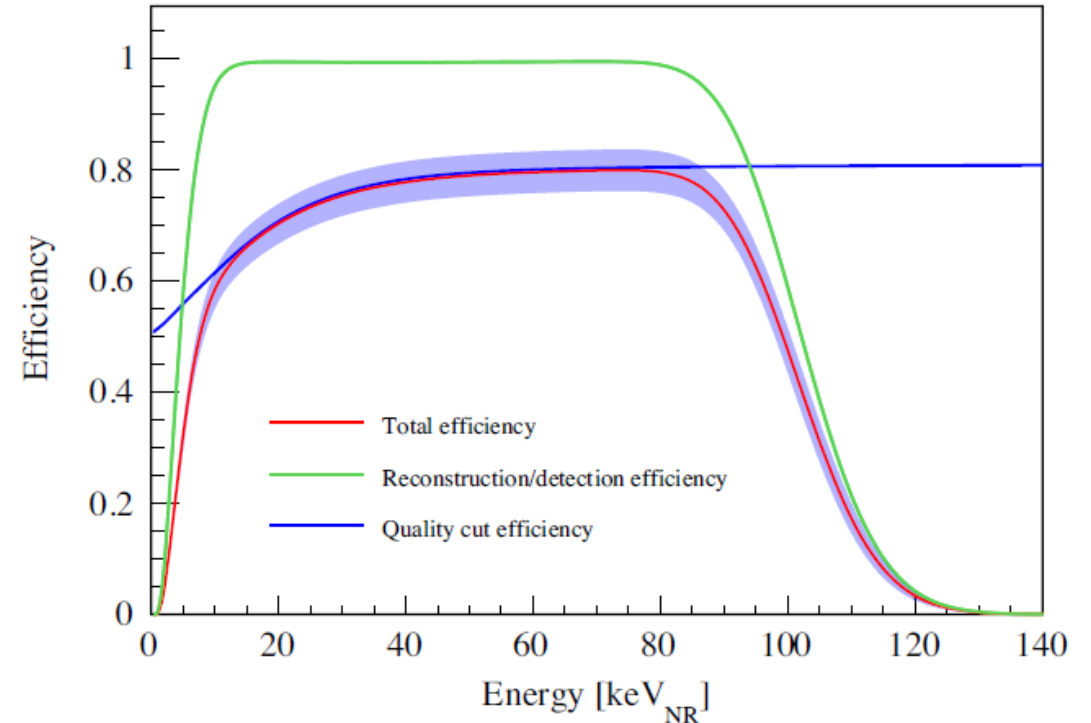
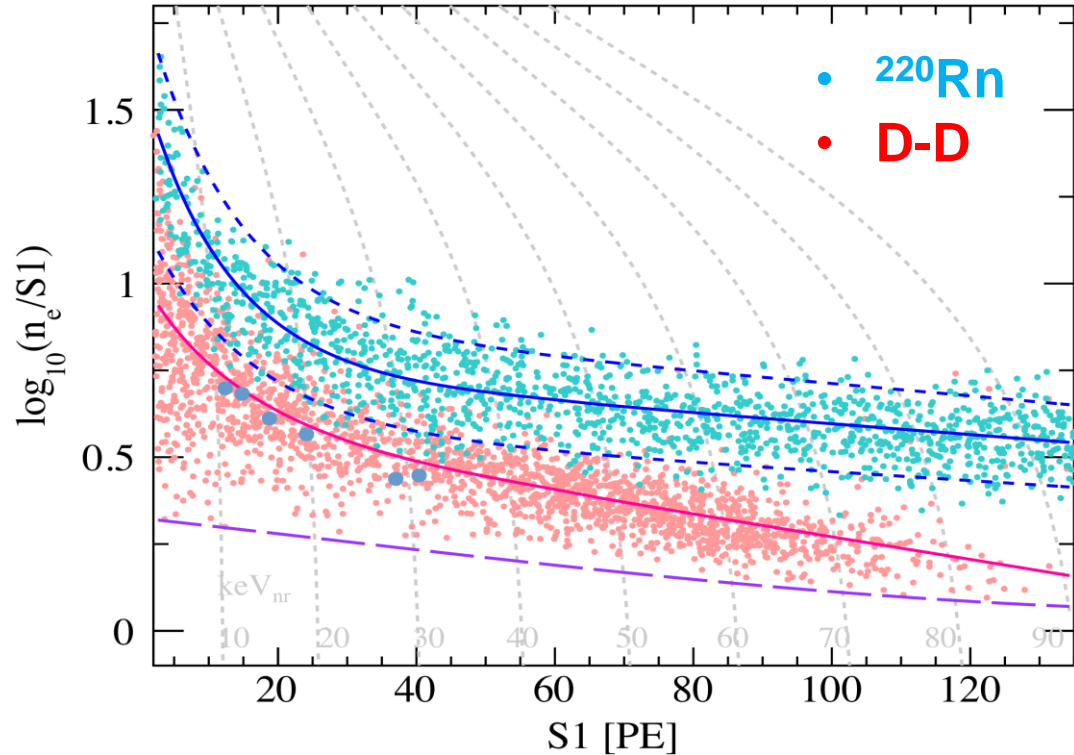
Position

Injected from gas panel

Calibration tubes

Beam pipe

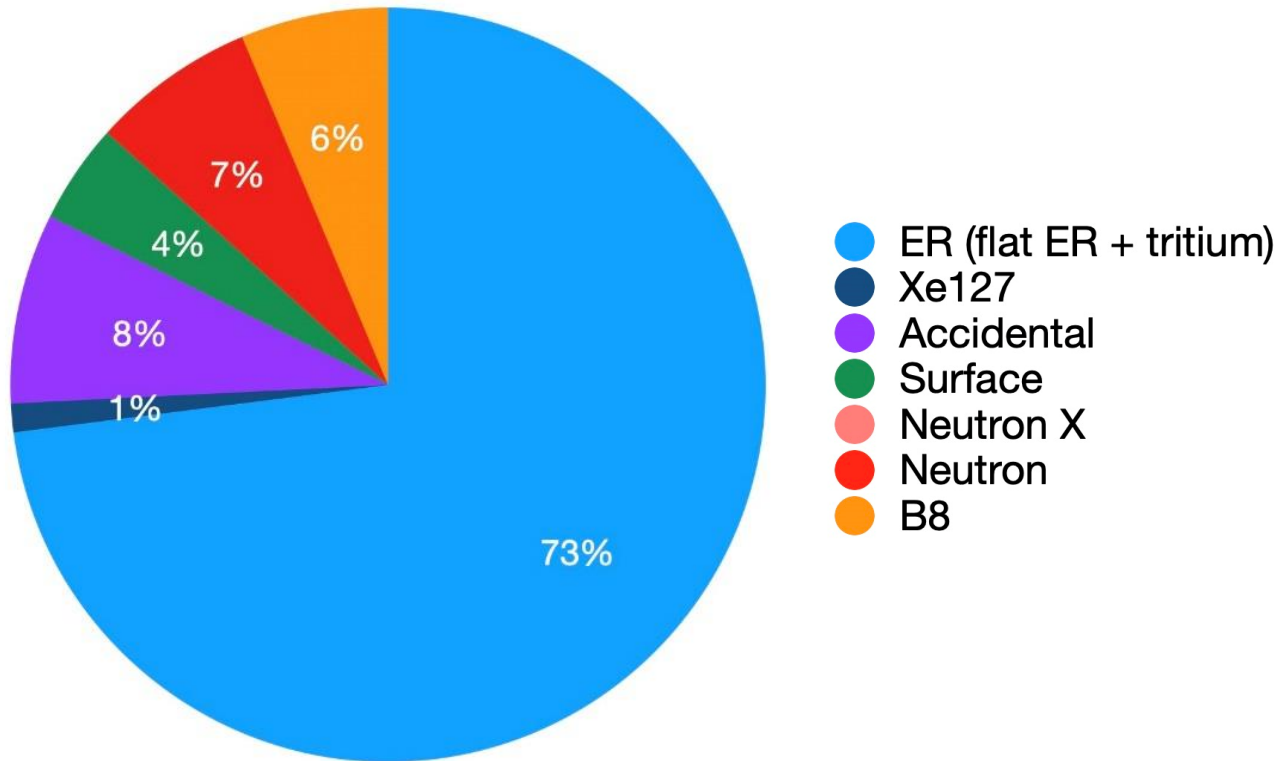
Detector Response



- ER leak ratio (below NR median curve) is $0.43\% \pm 0.18\%$;
- Efficiencies separately determined from ER or NR calibration data are all consistent;

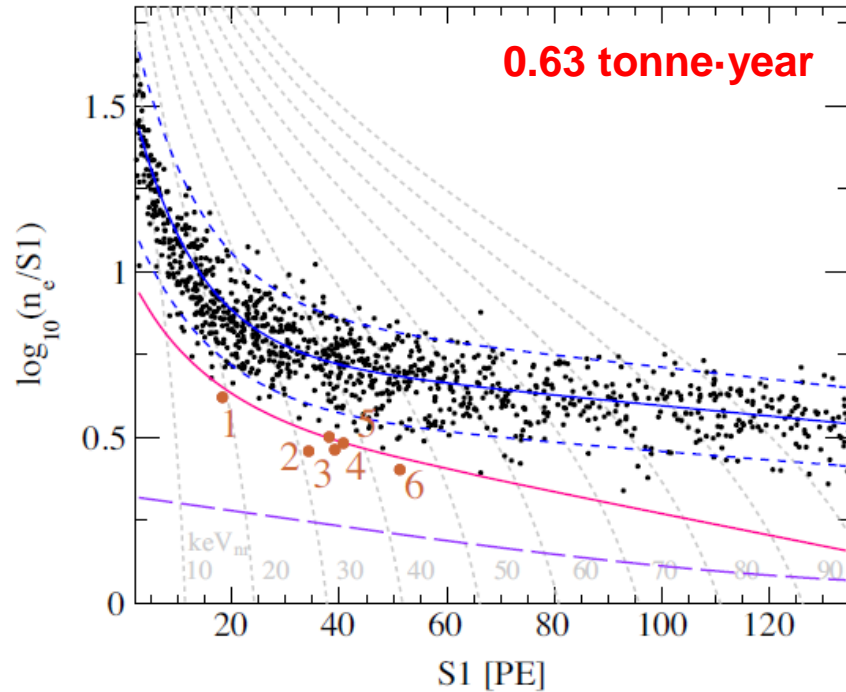
Backgrounds

Expected below-NR-median
events: 9.8 (0.6) evts

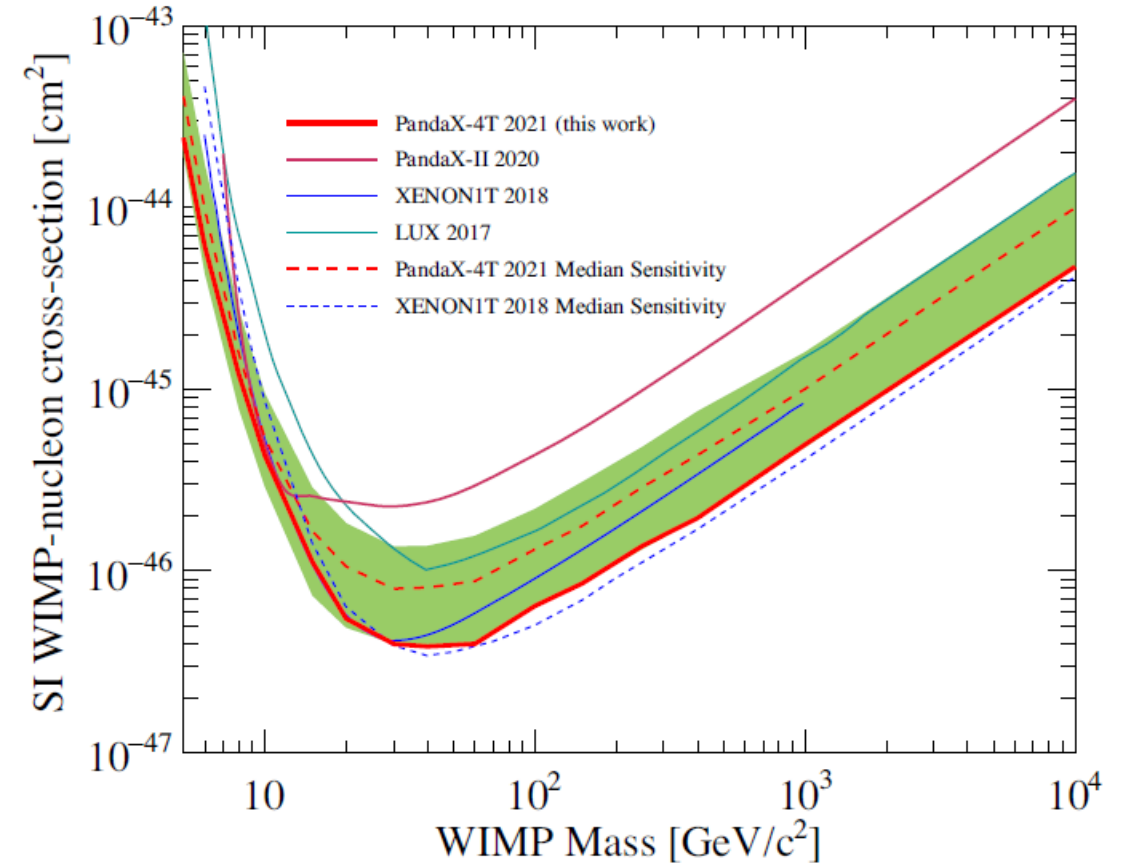


- ER (Rn+Kr+Material+Tritium) background dominated in the selection region;
- Background per unit target is improved from PandaX-II by 4 times (<10 keV);

PandaX-4T first commissioning Result - WIMPs

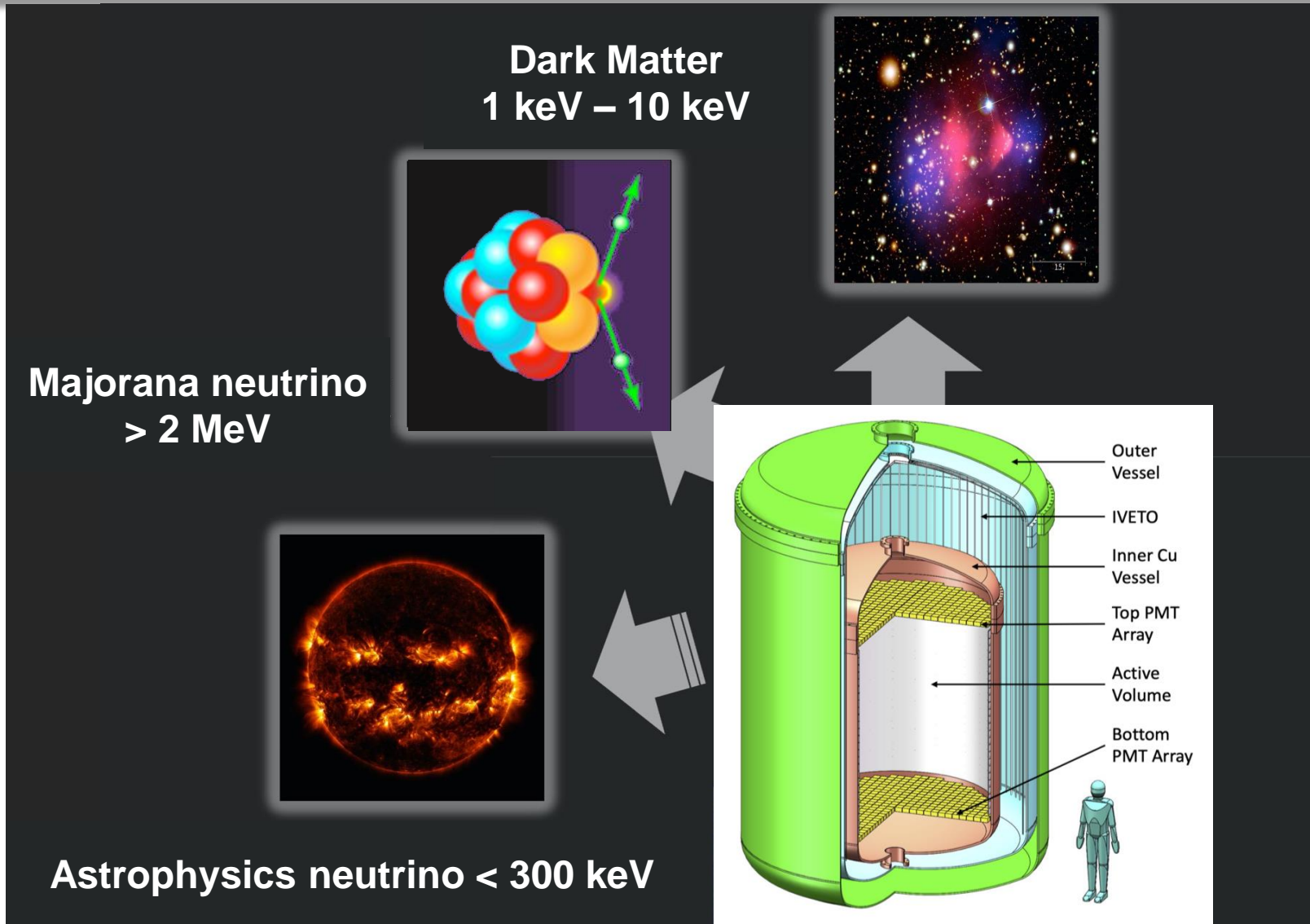


Y. Meng et al. PRL 127, 261802 (2021)



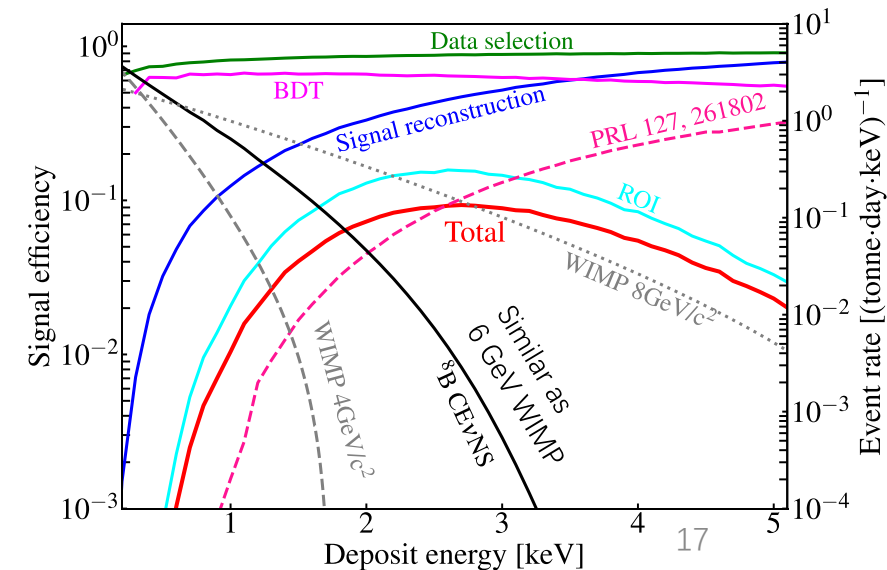
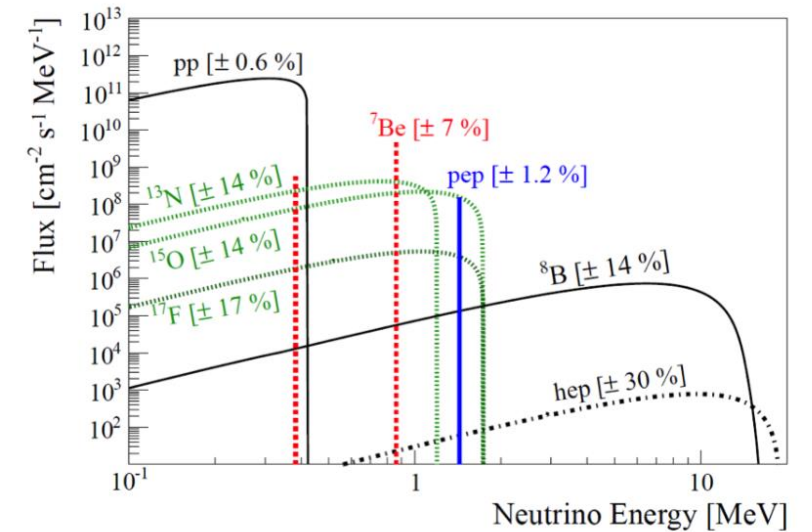
- 1058 candidates (expected 1054 ± 39), 6 below NR median curve (expected 9.8 ± 0.6);
- Sensitivity improved from PandaX-II final analysis by 2.9 times ($30 \text{ GeV}/c^2$);

PandaX: Multi-physics Goals



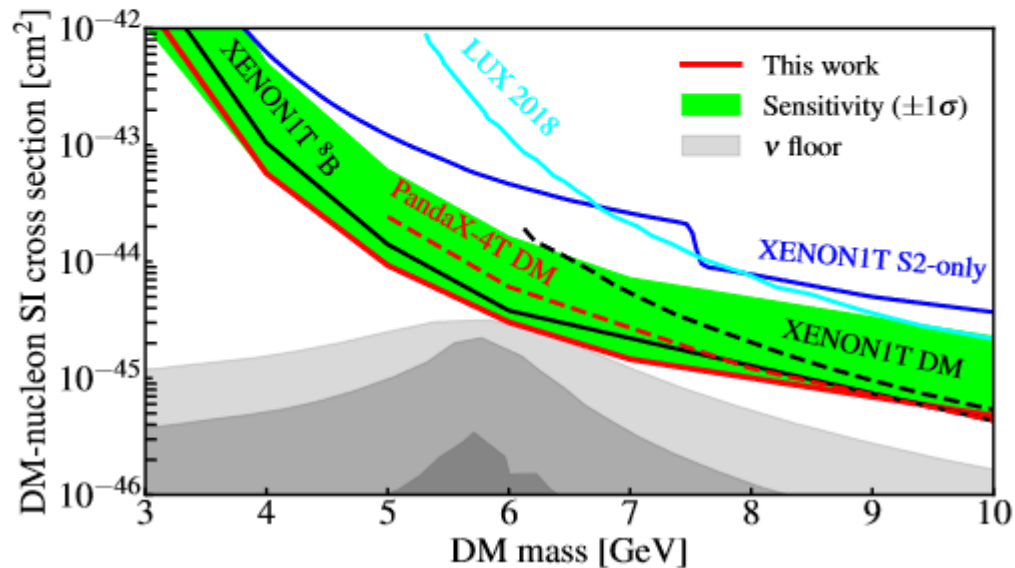
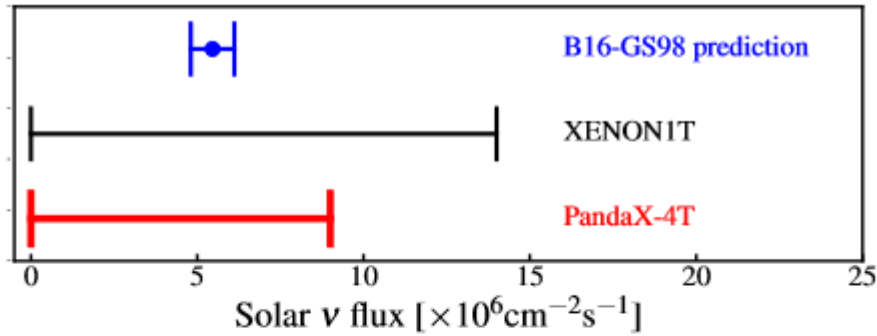
PandaX-4T Search for ^8B CEvNS

- To enhance sensitivity on ^8B (like a 6 GeV WIMP), need to lower selection threshold ($S1 \downarrow$, $S2 \downarrow$)
- Major challenge: Accidental background (AC, non-physical S1 and S2 randomly paired)
- Blind analysis: 0.48 ton-year data, excluding data with increase of noises rate (micro-discharge)



Constraints on ^8B neutrino

W. Ma et al. PRL 130, 021802 (2023)



- A multi-variate (BDT) algorithm trained to suppress AC background

– Some downward fluctuation

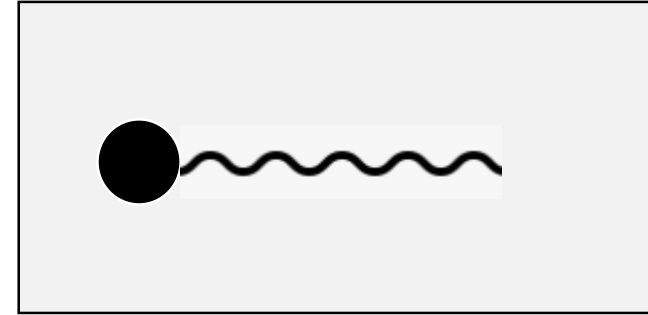
ROI (BDT applied)

ER+NR+AC	8B	Total prediction	Unblind data
1.46	1.42	2.88	1
0.04	0.29	0.33	0

- Leading constraint on ^8B neutrino flux through CEvNS
- Assuming nominal ^8B background, set strongest constraints on light WIMP of 3 -10 GeV

Luminance of Dark Matter

- Residual weak EM properties
- Coupling with photons



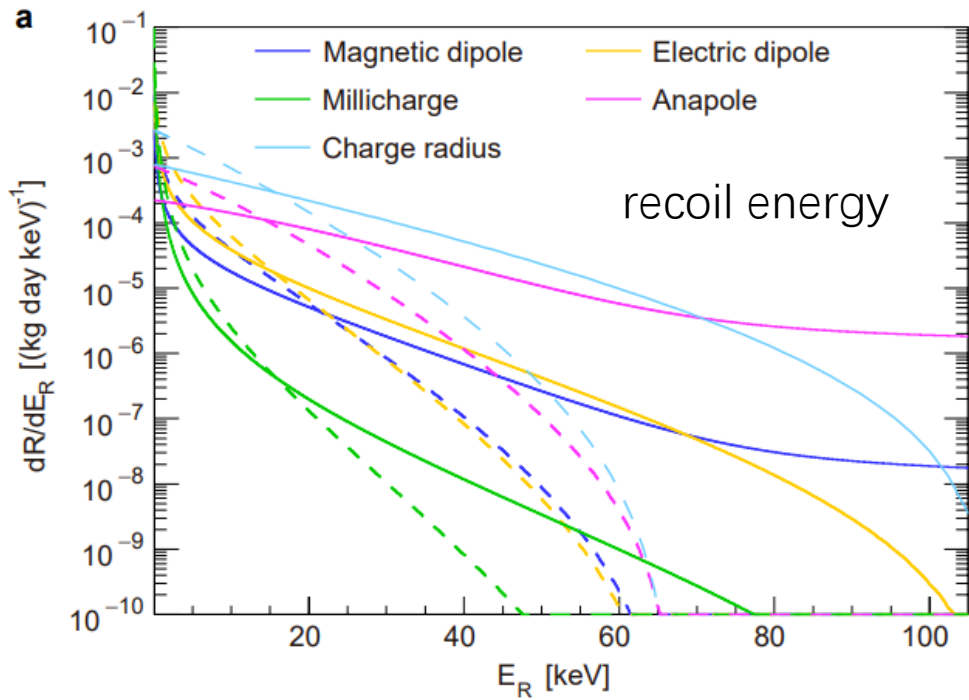
$$\mathcal{L} = \underbrace{Qe\bar{\chi}\gamma^\mu\chi A_\mu}_{\text{millicharge}} + \underbrace{\frac{\mu_\chi}{2}\bar{\chi}\sigma^{\mu\nu}\chi F_{\mu\nu}}_{\text{magnetic dipole}} + \underbrace{i\frac{d_\chi}{2}\bar{\chi}\sigma^{\mu\nu}\gamma^5\chi F_{\mu\nu}}_{\text{electric dipole}} + \underbrace{b_\chi\bar{\chi}\gamma^\mu\chi\partial^\nu F_{\mu\nu}}_{\text{charge radius}} + \underbrace{a_\chi\bar{\chi}\gamma^\mu\gamma^5\chi\partial^\nu F_{\mu\nu}}_{\text{anapole}}$$

tree-level

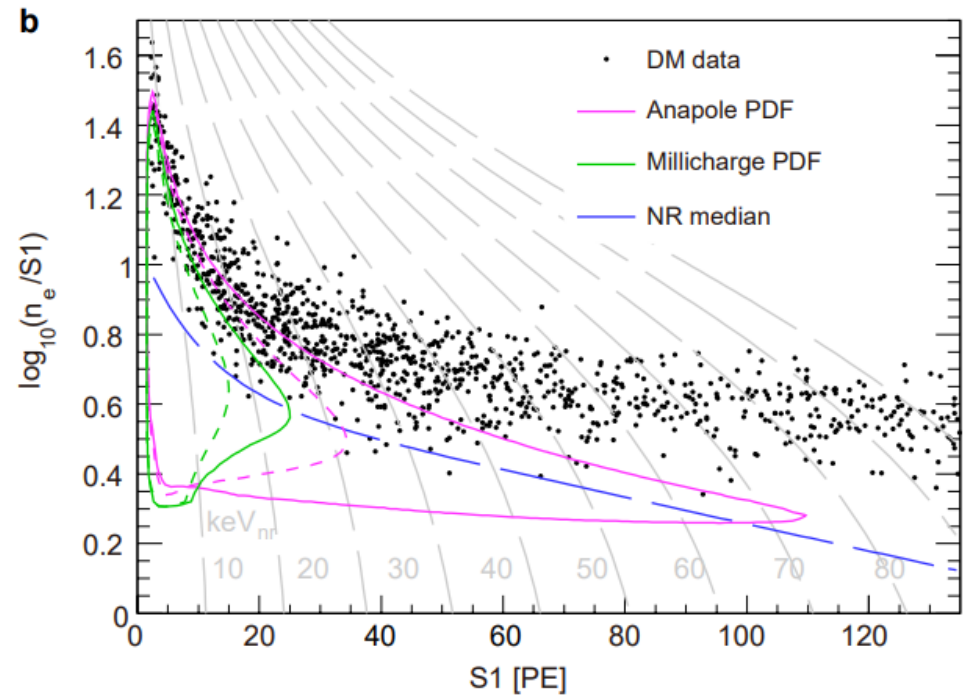
higher-order loop-level

Photon-mediated Interaction

- Nuclear recoil signature
 - Charge radius: similar to SI
 - Millicharge: $1/E_R^2$
 - MD and ED: $1/E_R$



30 and 100 GeV/c² DM



Results from Xenon Recoil Data

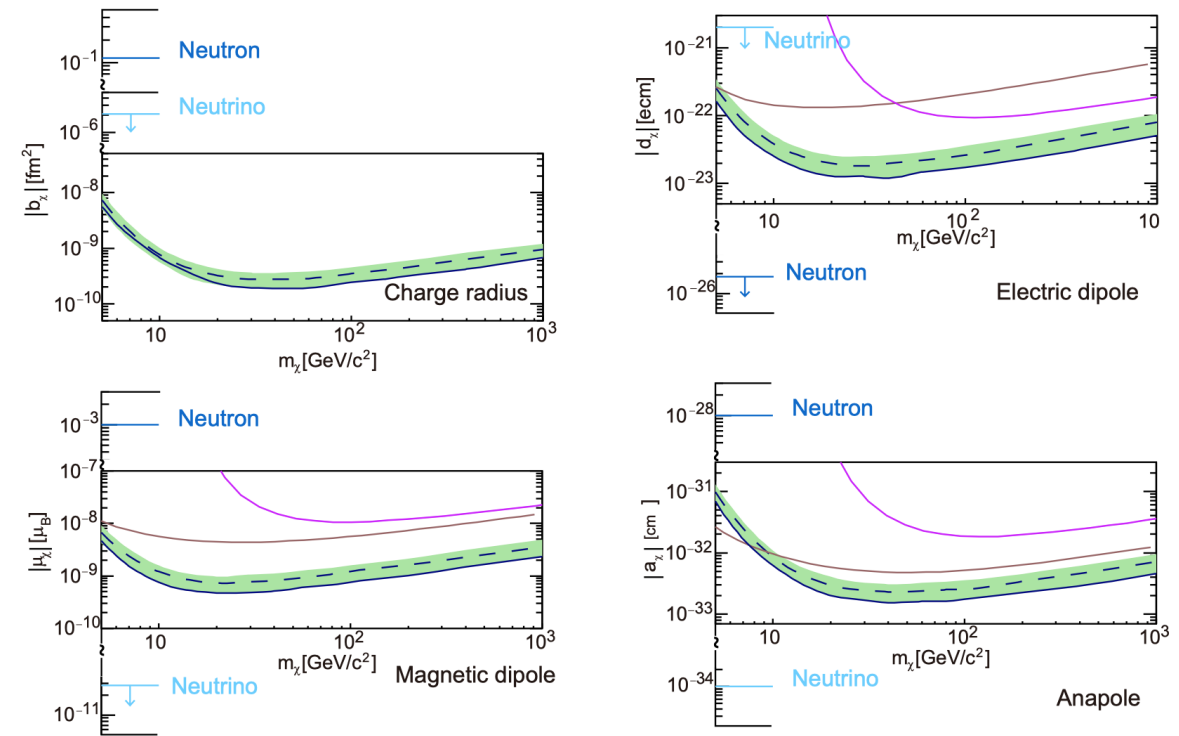
- First experimental constraints on DM charge radius
 - 4 orders of magnitude smaller than neutrino
- Other EM properties
 - up to 3 – 10 times improvement

Table 1 | Comparison of electromagnetic properties

	dark matter	neutrino	neutron
Charge radius (fm ²)	$<1.9 \times 10^{-10}$	$[-2.1, 3.3] \times 10^{-6}^*$	-0.1155^*
Millicharge (e)	$<2.6 \times 10^{-11}$	$<4 \times 10^{-35}^*$	$(-2 \pm 8) \times 10^{-22}^*$
Magnetic dipole (μ_B)	$<4.8 \times 10^{-10}$	$<2.8 \times 10^{-11}^*$	$-1 \times 10^{-3}^*$
Electric dipole (ecm)	$<1.2 \times 10^{-23}$	$<2 \times 10^{-21}^\dagger$	$<1.8 \times 10^{-26}^*$
Anapole (cm ²)	$<1.6 \times 10^{-33}$	$\sim 10^{-34}^\ddagger$	$\sim 10^{-28}^\S$

* Datas are taken from PDG [33]
 † Taken from [32]
 ‡ Taken from [34]
 § Taken from [35]

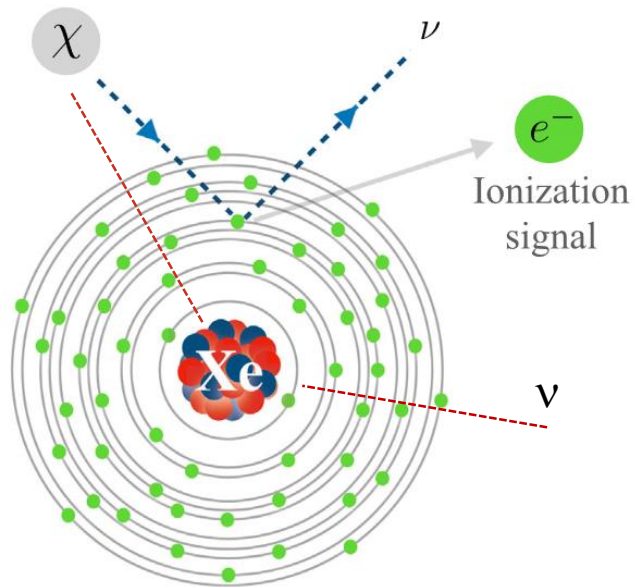
X. Ning et al. Nature (2023)



— PandaX-4T — PICO-60 — DEAP-3600

DM- ν Conversion

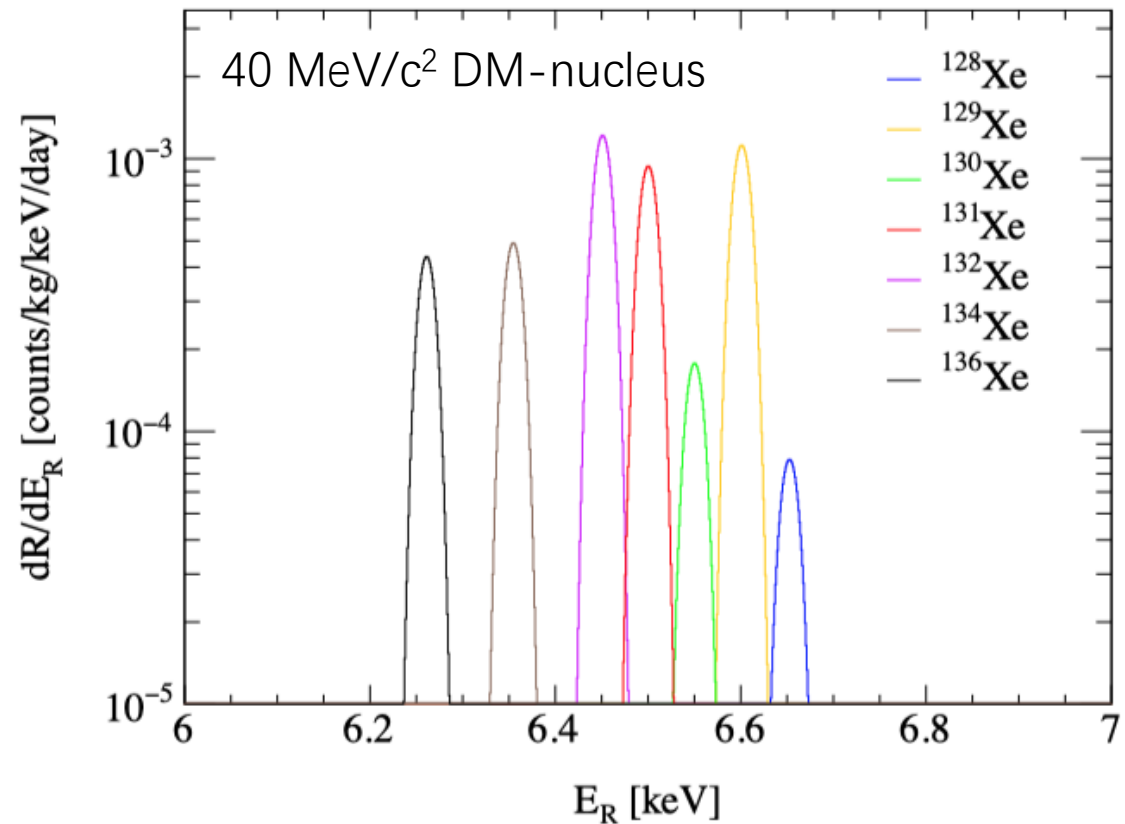
$$E_R = \frac{m_\chi^2}{2(m_T + m_\chi)}$$



J. Dror et al, PRL 124, 181301 (2020)

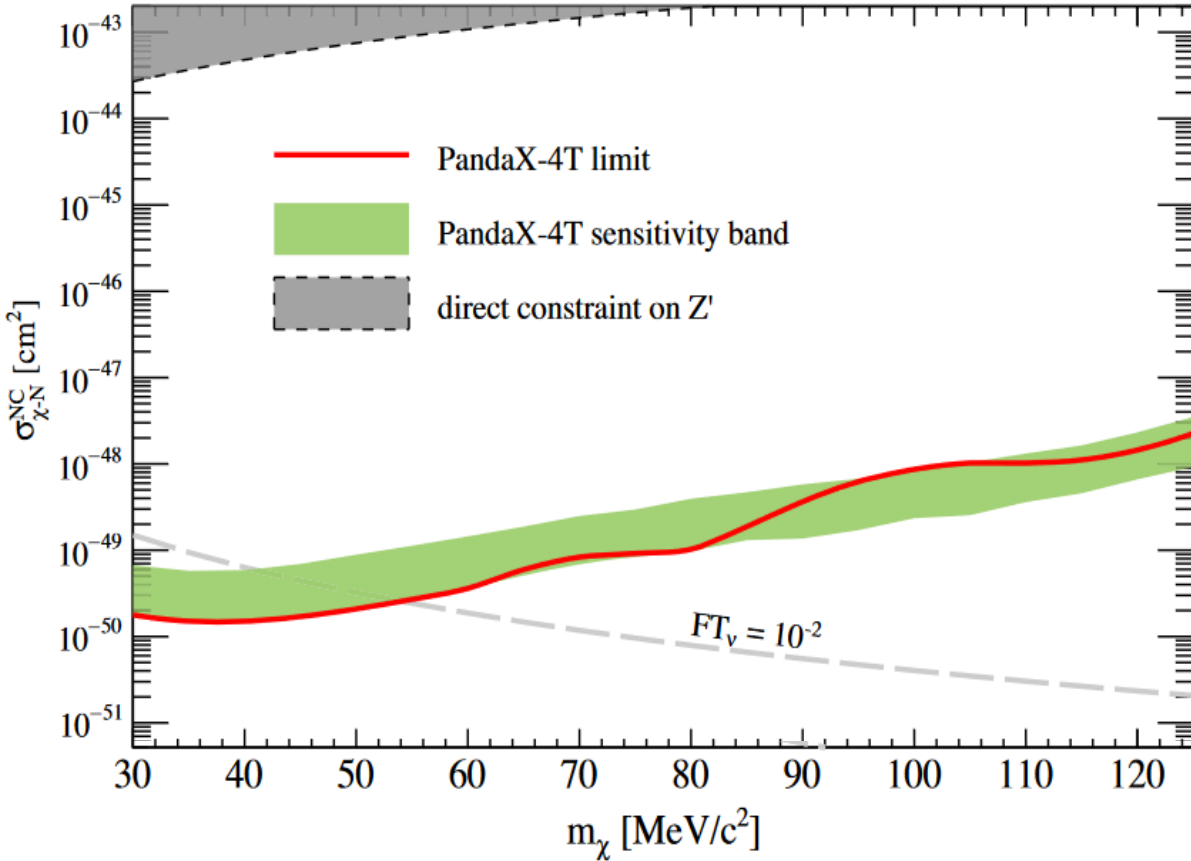
Characteristic mono-energetic signal:

- (a) Xe-nucleus targets, $m_\chi = 40 \text{ MeV}$, $E_R \sim 6.5 \text{ keV}$
- (b) Electrons targets, $m_\chi = 40 \text{ keV}$, $E_R \sim 1.5 \text{ keV}$



Mono-energetic Signal Search

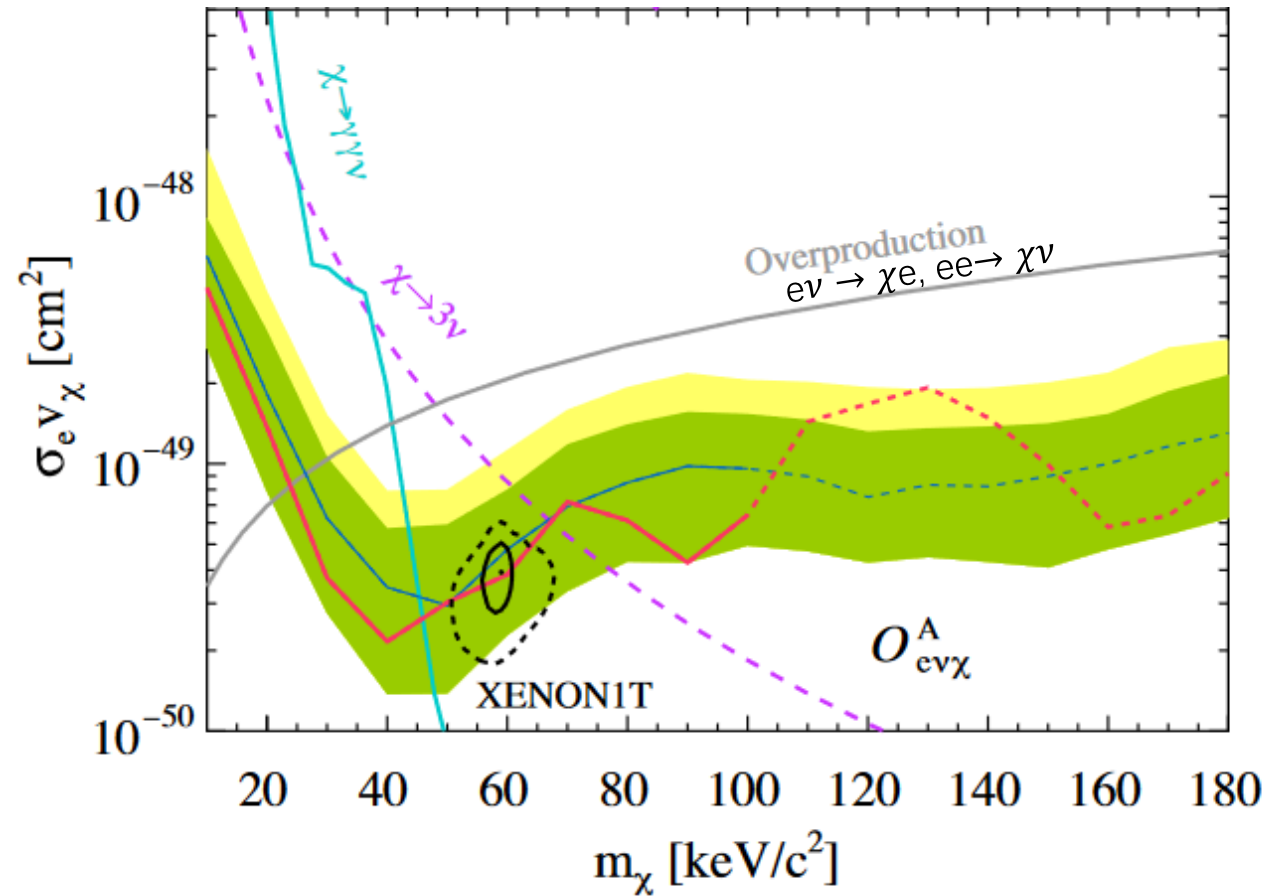
$$\chi + \text{Xe} \rightarrow \nu + \text{Xe}$$



L. Gu et al. PRL 129, 161803 (2022), Editors' Suggestion

2023/6/22

$$\chi + e \rightarrow \nu + e$$

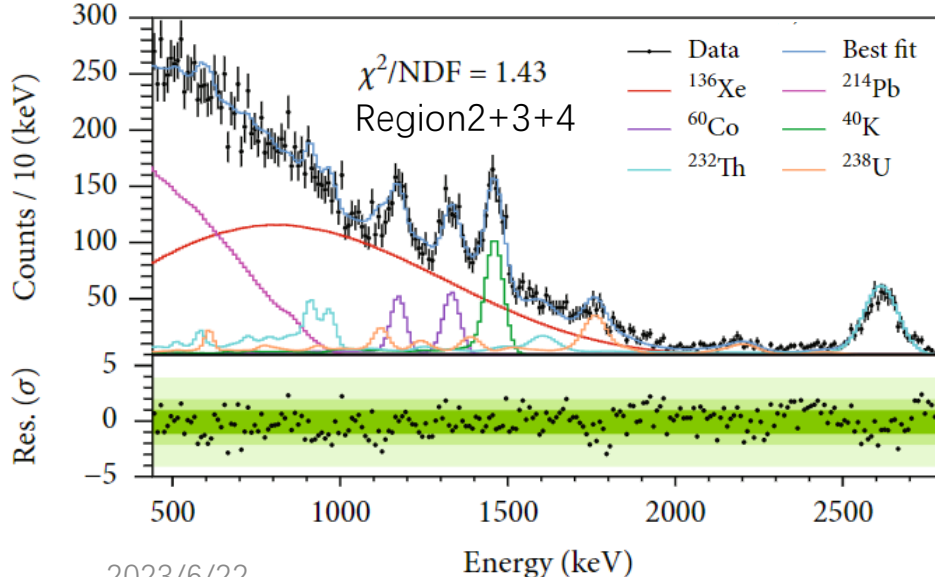
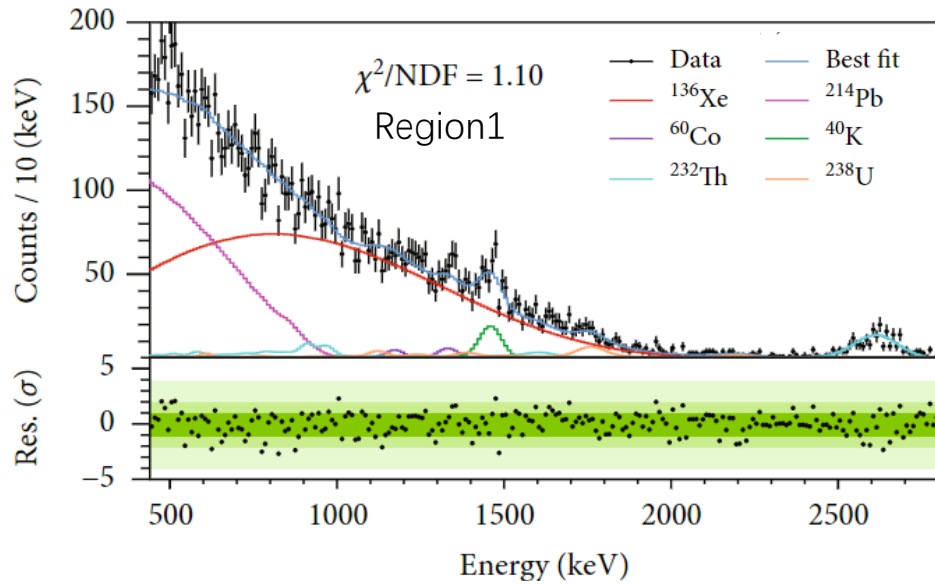


D. Zhang et al. PRL 129, 161804 (2022), Editors' Suggestion

XeSAT2023

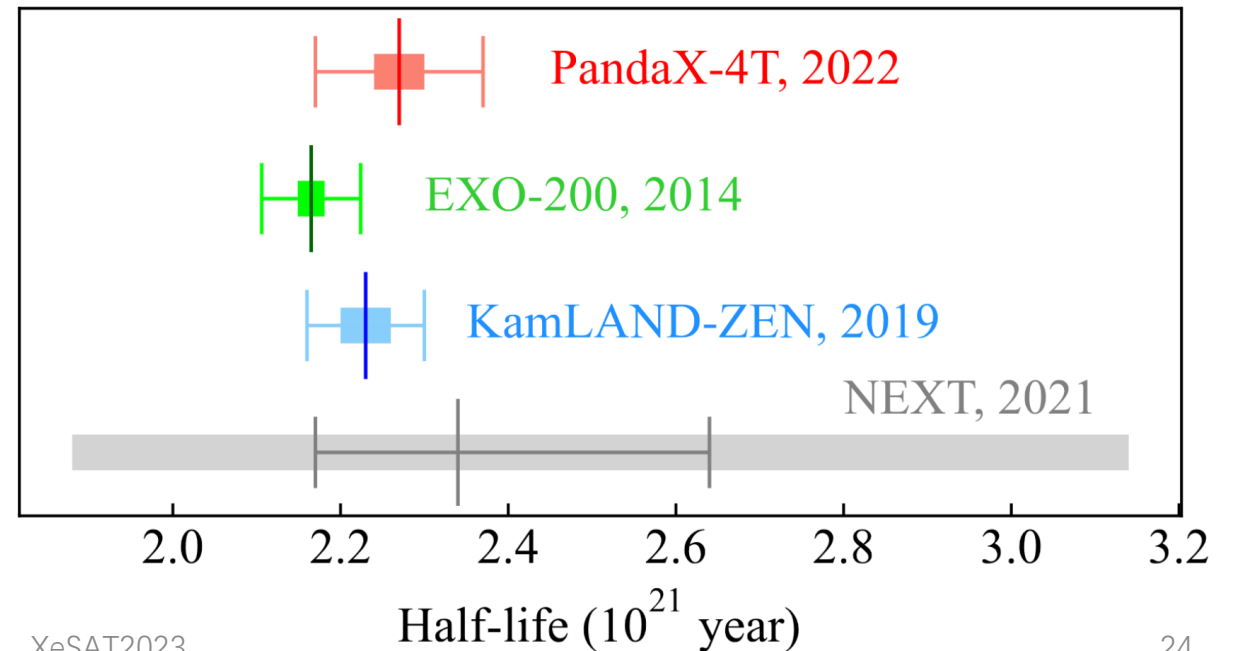
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^{136}Xe $2\nu\text{DBD}$ Half-life Measurement



- PandaX-4T: $^{136}\text{Xe}/\text{Xe}$ 8.86%, 59.6kg in FV, up to 2.8MeV
- First natural xenon measurement with a dark matter detector
 2.27 ± 0.03 (stat) ± 0.10 (syst) $\times 10^{21}$ years
- Consistent with ^{136}Xe -enrich experiments.

Research Vol 2022, 9798721 (2022)



After Commissioning

- Tritium identified in commissioning data
- Offline xenon distillation
- 1st physics run (Run1)
 - **Data still under blind analysis**
- CJPL-II B2 hall construction
- Detector upgrade

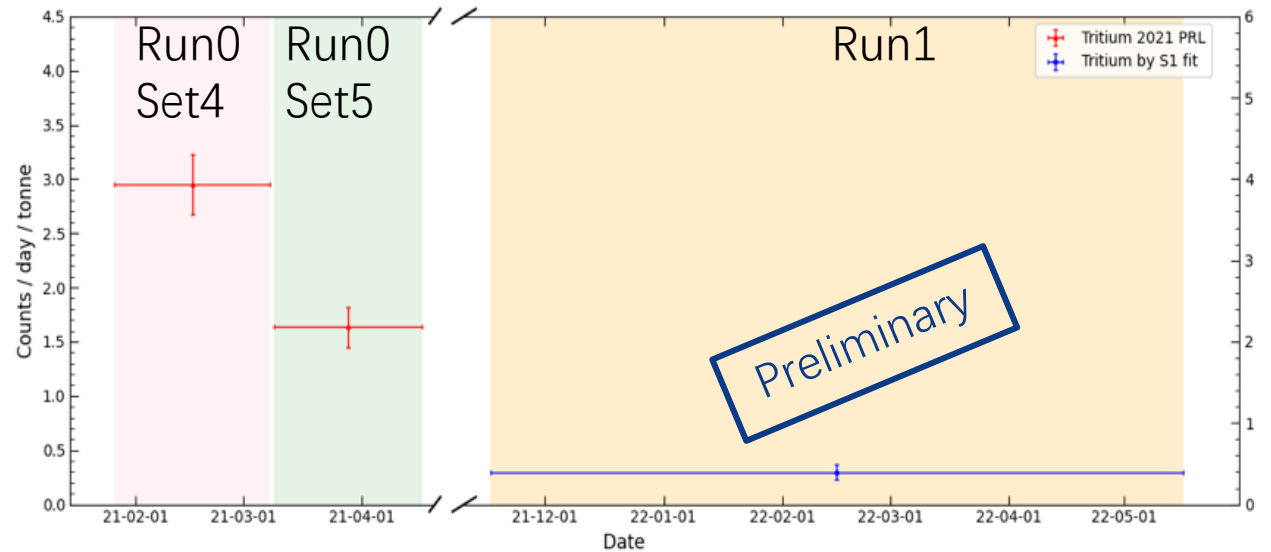
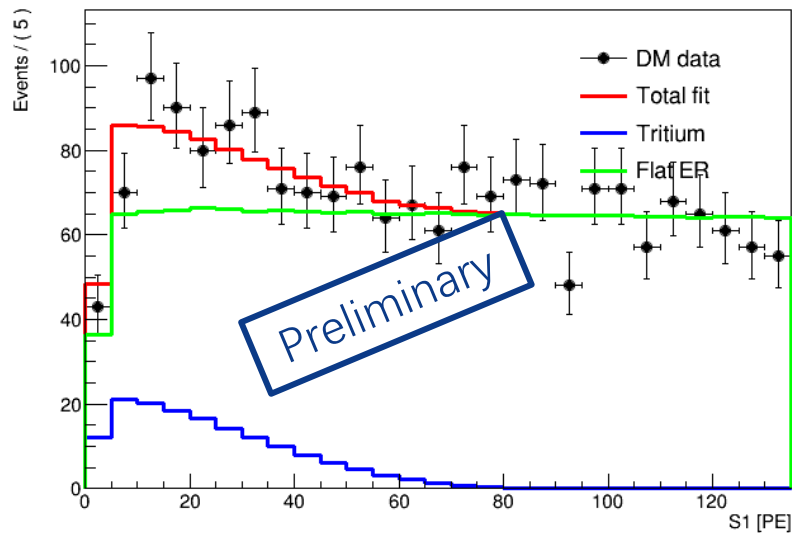


Commissioning (Run 0)	Calibration	Distillation	Physics Run (Run 1)	Calibration	Detector Upgrade
2020/11/28 – 2021/04/16	2021/04/17 – 2021/06/09		2021/11/15 – 2022/05/15	2022/05/16 – 2022/07/08	

Tritium Removal

- Preliminary estimation of tritium level (keep S2 blinded)
- Extensive tritium measures planned for the next run (Run 2)

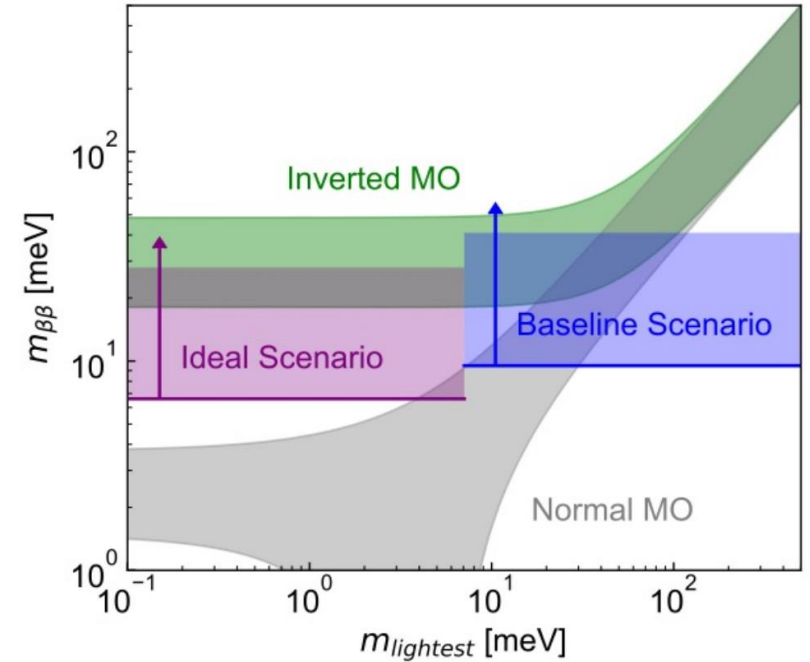
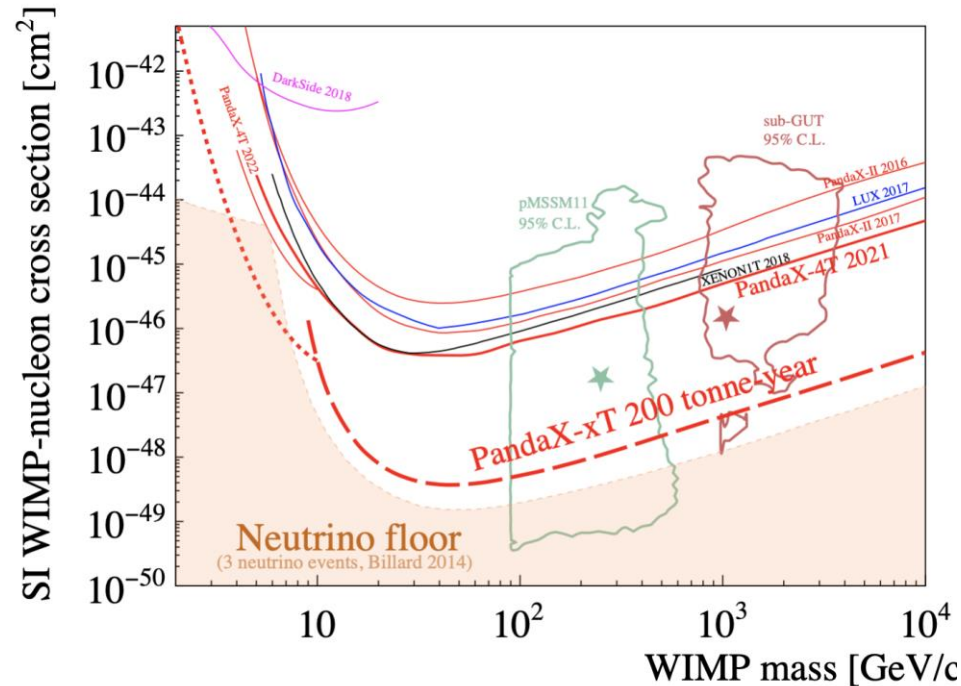
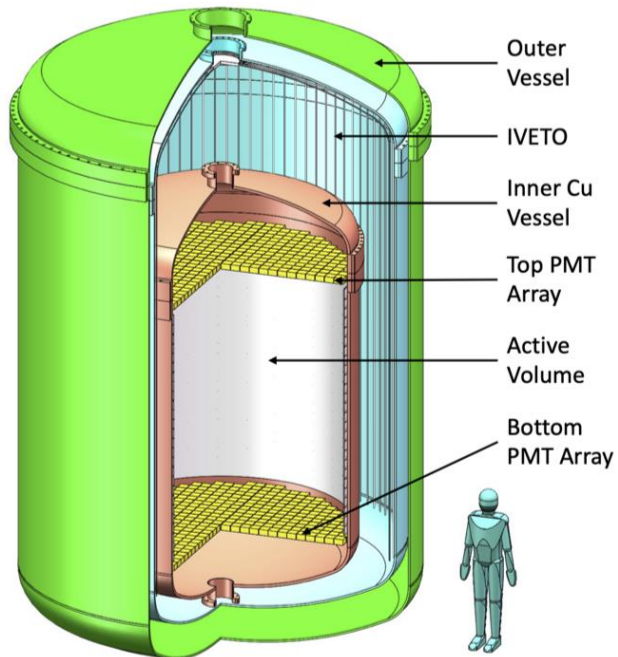
Period	Run0 Set 4	Run0 Set 5	Run1
Tritium Counts/day/tonne	3.0 ± 0.3	1.6 ± 0.2	0.4 ± 0.1



Future plan: PandaX-xT

- “Ultimate” liquid xenon experiment

- With >30 tonne sensitive volume
- Letter-of-interest sent to Chinese funding agency
- Decisive test on WIMP and key test on Dirac/Majorana neutrino



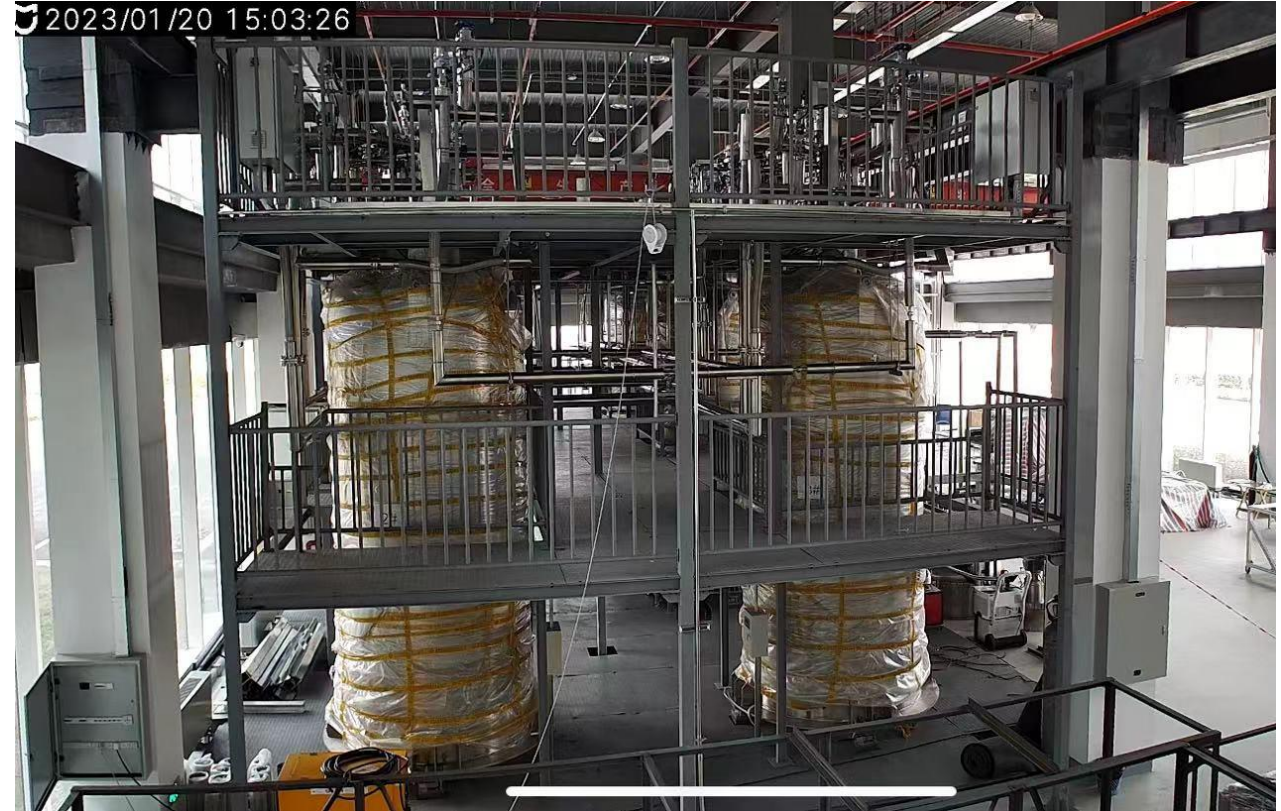
PandaX-xT R&D: Electrode and Xenon Storage

TPC: Large size high light transmission electrode



2023/6/22

Large xenon storage system



XeSAT2023

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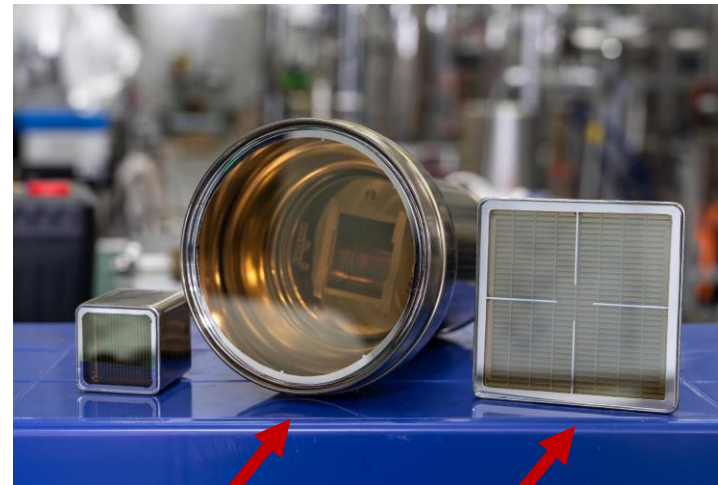
PandaX-xT R&D: 2-inch PMTs and New Electronics

PMT:

- R12699 2-inch PMTs with 4 independent anode readout;
- Better time response for better waveform building;
- Lower radioactivity;

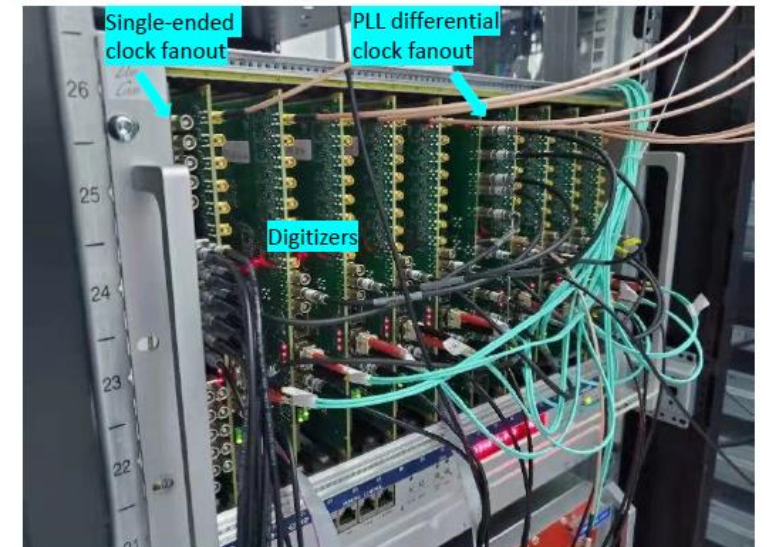
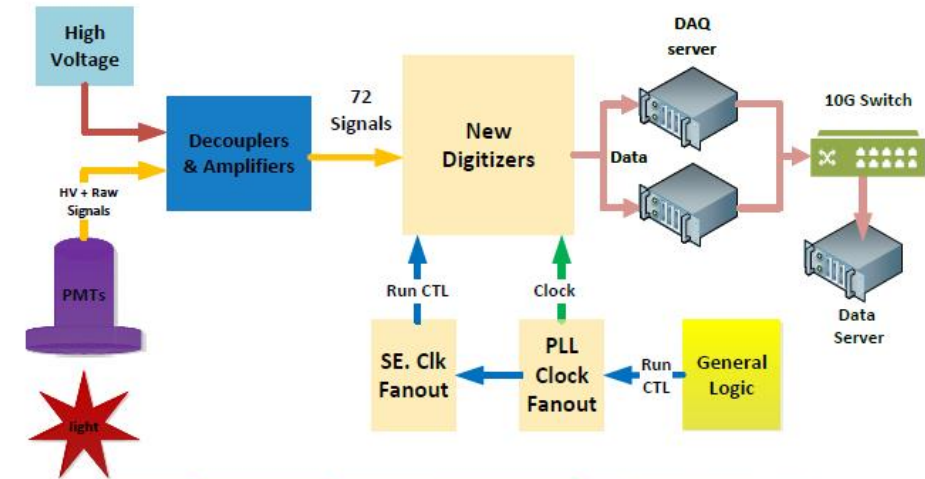
Electronics:

- Higher sampling rate;
- Accept out-trigger mode;

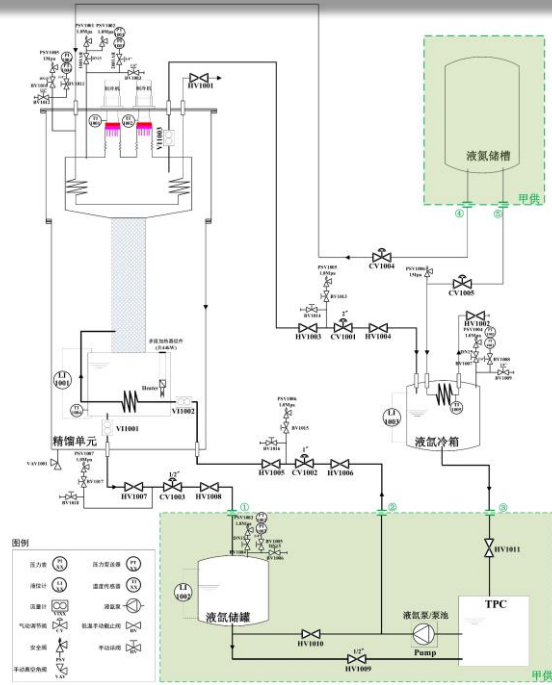
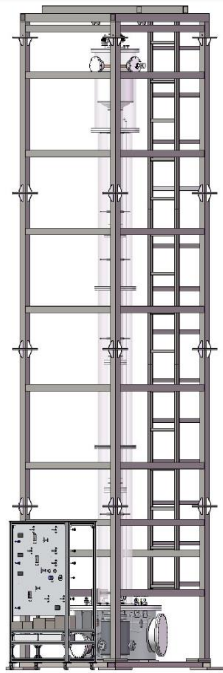


		R11410	R12699 (4ch/piece)
Time Response [ns]	Rise Time	5.5	1.2
	Transit Time	46	5.9
Radioactivity [mBq/pc]	Co-60	< 2.34	< 0.07
	Th-232	< 7.82	< 0.40
	U-238	< 56.48	0.47 ± 0.11

Custom-developed: 14-bit, 500MS/s



PandaX-xT R&D: Distillation and Purification



		PandaX-4T	Upgraded
Flow rate [kg/h]	Kr	10	30
	Rn	56.5	856
Reduction factor	Kr	10^6	10^8
	Rn	2.2	4.4

- Low Outgassing

- High flow rate

Summary

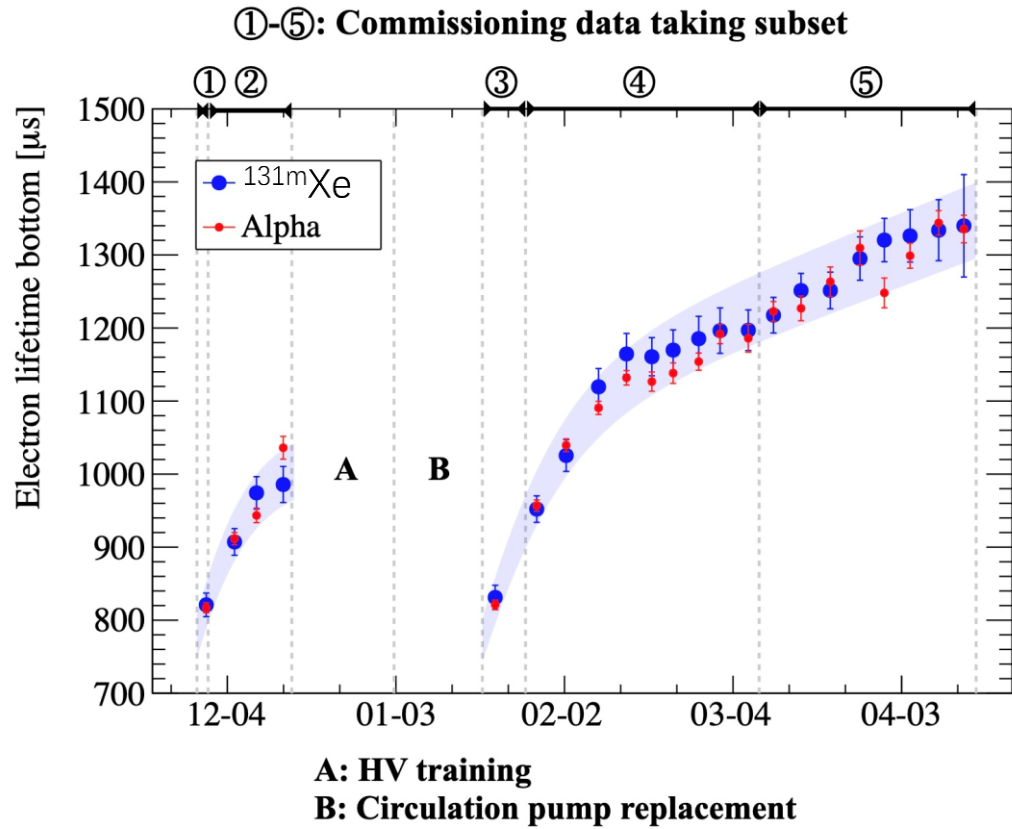
- **PandaX-4T is one of the new generation multi-tonne xenon experiments**
- **Intense searches for various types of physics, including DMs and neutrinos**
- **Expecting more interesting results from PandaX**
- **Highly welcome new collaborators!**

Thank You!

Backups

Purification System

NEXT



Ref. the maximum drift time $\sim 840 \mu\text{s}$



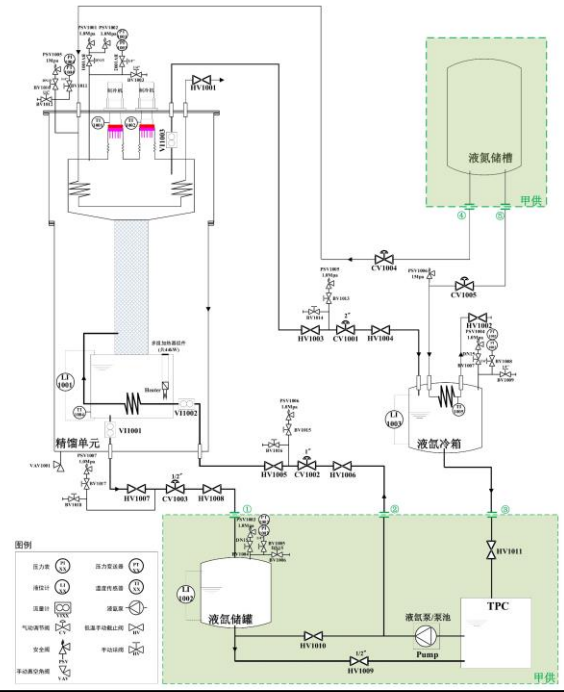
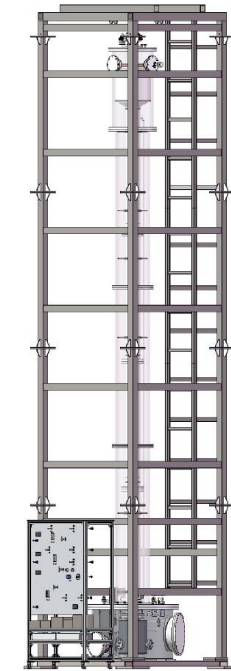
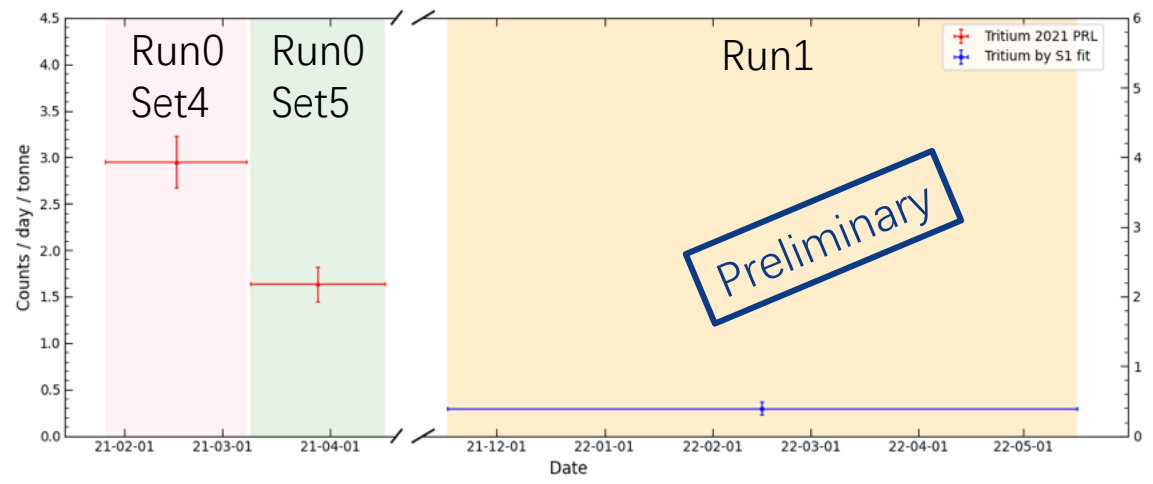
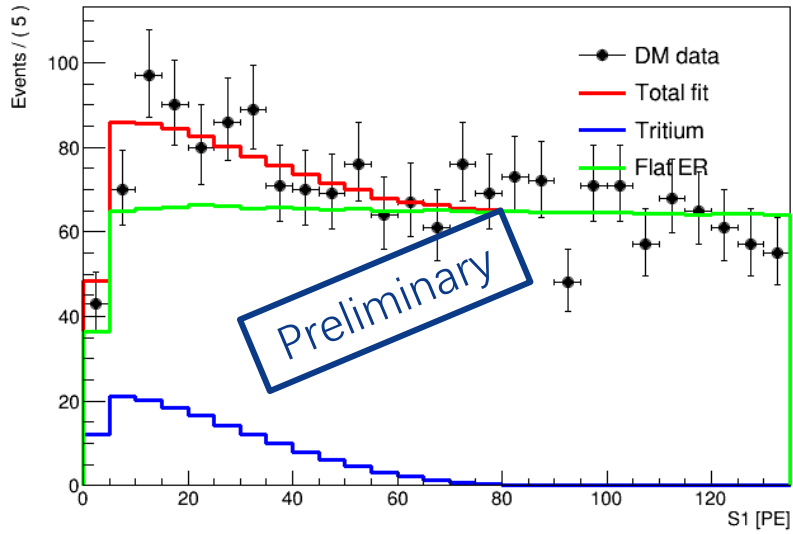
- Low Outgassing



- High flow rate

Distillation System

NEXT

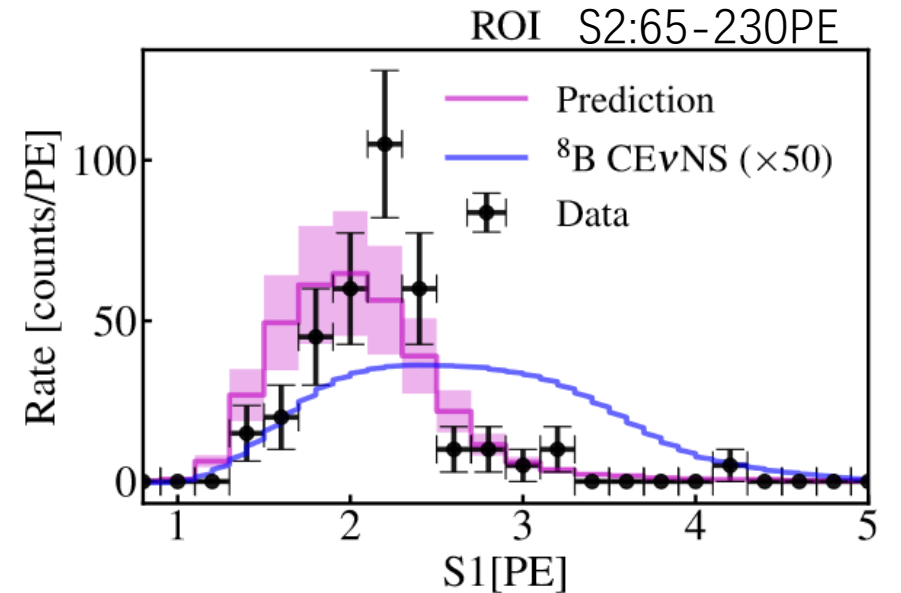
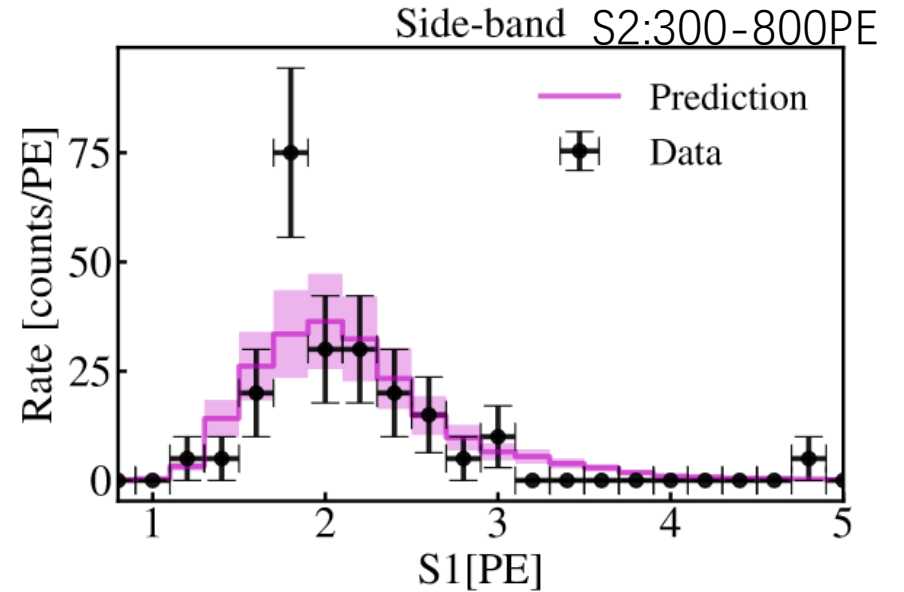


	PandaX-4T	Upgraded
Flow rate [kg/h]	Kr 10 Rn 56.5	30 856
Reduction factor	Kr 10^6 Rn 2.2	10^8 4.4

Control of Accidental Background in ^8B Search

- Use “scrambled” real data to model accidental background
- A multi-variate (BDT) algorithm trained to suppress AC background
- Training/selection is blinded
- postBDT: $N_{\text{obs}}=1$, $N_{\text{bkg}}=1.6$, $N_{\text{sig}}=1.7$

N_{hit}	S2 range [PE]	BDT	ER	NR	Surf	AC	Total BKG	^8B	Obs
2	65-230	pre	0.04	0.10	0.14	62.43	62.71	2.32	59
		post	0.02	0.04	0.03	1.41	1.50	1.42	1
3	65-190	pre	0.01	0.05	0.08	0.79	0.93	0.42	2
		post	0.00	0.02	0.03	0.02	0.07	0.29	0



How dark is dark matter?



Photon-mediated Interaction

- Matched to non-relativistic EFT
 - Nuclear response function [N. Anand, A. Fitzpatrick, W. Haxton PRC 89, 065501 \(2014\)](#)

$$\mathcal{L}_{\text{MC}} = Q\bar{\chi}\gamma^\mu\chi A_\mu \rightarrow \mathcal{O}_{\text{MC}} = eQ Q_{p,n} \frac{\mathcal{O}_1}{q^2},$$

$$\mathcal{L}_{\text{MD}} = \frac{\mu_\chi}{2} \bar{\chi} \sigma^{\mu\nu} \chi F_{\mu\nu}$$

$$\rightarrow \mathcal{O}_{\text{MD}} = e\mu_\chi \left[-Q_{p,n} \left(\frac{\mathcal{O}_1}{2m_\chi} + \frac{2m_{p,n}}{q^2} \mathcal{O}_5 \right) - \frac{g_{p,n}}{m_{p,n}} \left(\mathcal{O}_4 - \frac{m_{p,n}^2}{q^2} \mathcal{O}_6 \right) \right],$$

$$\mathcal{L}_{\text{ED}} = i\frac{d_\chi}{2} \bar{\chi} \sigma^{\mu\nu} \gamma^5 \chi F_{\mu\nu} \rightarrow \mathcal{O}_{\text{ED}} = ed_\chi Q_{p,n} \frac{2m_{p,n}}{q^2} \mathcal{O}_{11},$$

$$\mathcal{L}_{\text{CR}} = b_\chi \bar{\chi} \gamma^\mu \chi \partial^\nu F_{\mu\nu} \rightarrow \mathcal{O}_{\text{CR}} = eb_\chi Q_{p,n} \mathcal{O}_1,$$

$$\mathcal{L}_{\text{A}} = a_\chi \bar{\chi} \gamma^\mu \gamma^5 \chi \partial^\nu F_{\mu\nu} \rightarrow \mathcal{O}_{\text{A}} = ea_\chi (2Q_{p,n} \mathcal{O}_8 - g_{p,n} \mathcal{O}_9),$$

$$\mathcal{O}_1 = 1$$

$$\mathcal{O}_4 = \mathbf{S}_\chi \cdot \mathbf{S}_N$$

$$\mathcal{O}_5 = i\mathbf{S}_\chi \cdot \left(\frac{\mathbf{q}}{m_N} \times \mathbf{v}^\perp \right)$$

$$\mathcal{O}_6 = \left(\mathbf{S}_\chi \cdot \frac{\mathbf{q}}{m_N} \right) \left(\mathbf{S}_N \cdot \frac{\mathbf{q}}{m_N} \right)$$

$$\mathcal{O}_8 = \mathbf{S}_\chi \cdot \mathbf{v}^\perp$$

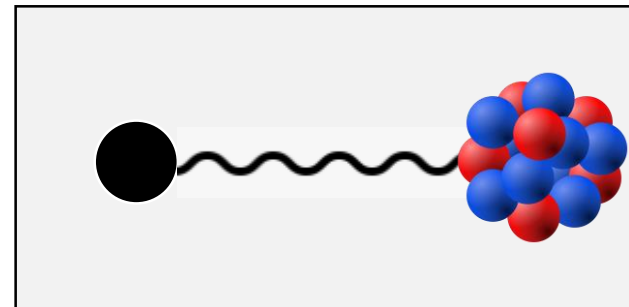
$$\mathcal{O}_9 = i\mathbf{S}_\chi \cdot \left(\mathbf{S}_N \times \frac{\mathbf{q}}{m_N} \right)$$

$$\mathcal{O}_{11} = i\mathbf{S}_\chi \cdot \frac{\mathbf{q}}{m_N}$$

Nuclear response function:

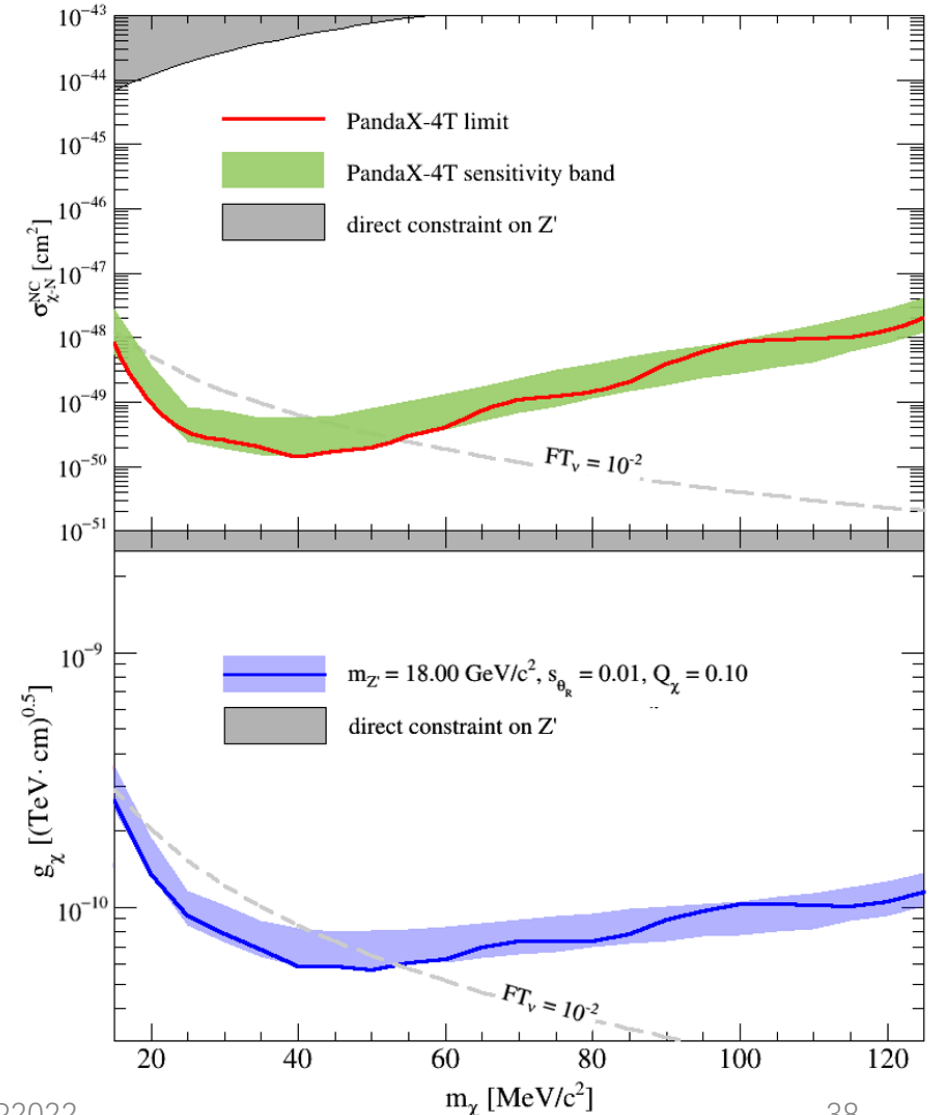
- full-basis shell-model calculations (GCN5082 interaction).
- No truncation (9 billion Slater determinants)!

Eugenio Del Nobile [Phys. Rev. D, 98\(12\), Dec 2018](#)
 A.Liam Fitzpatrick et al. [JCAP, 2013\(02\):004004, Feb 2013](#)
 Bradley J. et al. [JHEP 2019\(4\), Apr 2019](#)
 Chiara Arina et al. [Eur. Phys. J. C, 81\(3\):223, 2021](#)
 A.Liam Fitzpatrick et al. [arXiv: 1211.2818](#)
 Fady Bishara et al. [JHEP 11:059, 2017](#)

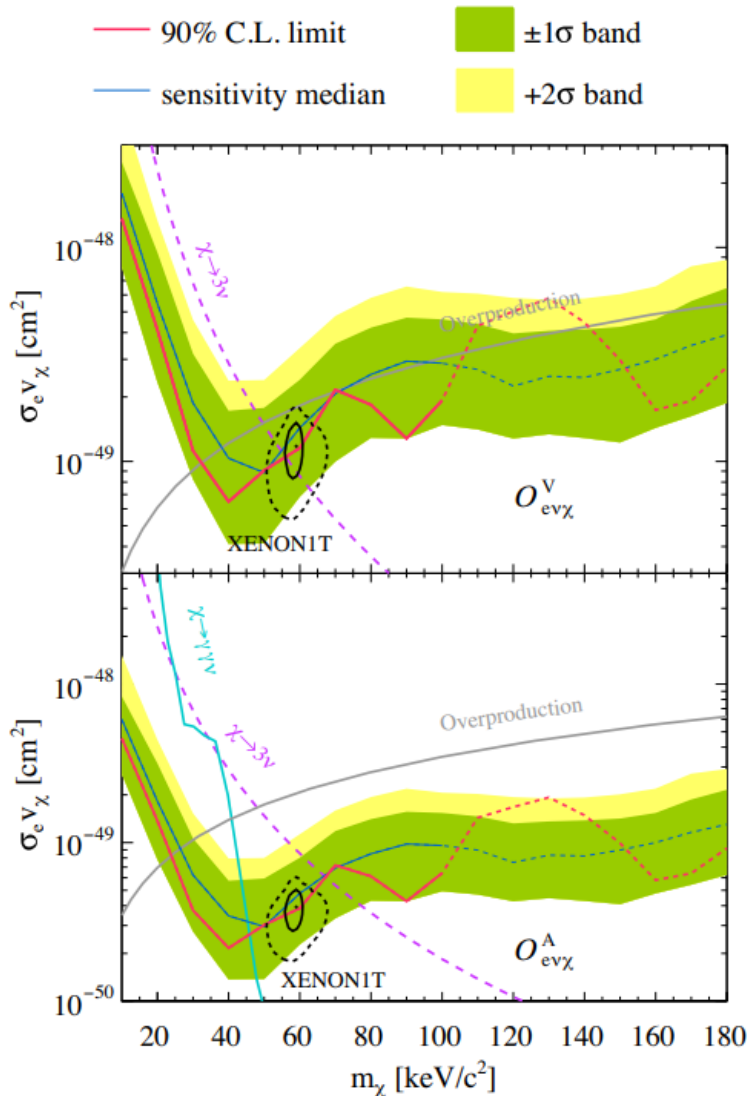


Exclusion Limits of Neutral current FDM Absorption (NR)

- **No significant excess above 1σ . 90% C.L.** upper limit within the $\pm 1\sigma$ sensitivity band (slight downward fluctuation in the DM mass range [40,55] MeV/c², power-constrained to -1σ)
- The strongest limit achieved is **$1.4 \times 10^{-50} \text{ cm}^2$** at a fermionic dark matter mass of **35 MeV/c²**
- Constraints on the coupling g_χ to the order of **$10^{-10} (\text{TeV} \cdot \text{cm})^{0.5}$**



Exclusion Limits of FDM Absorption on Electron

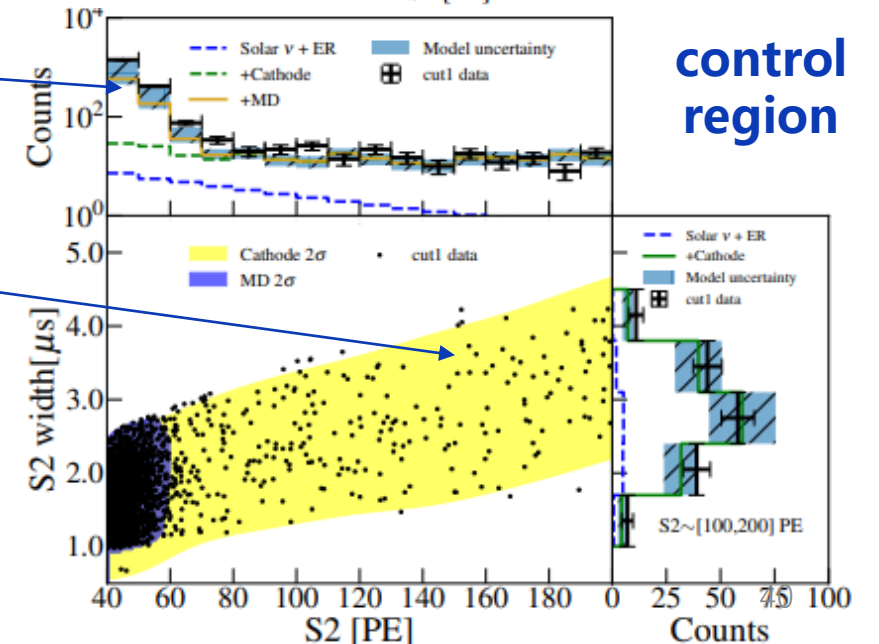
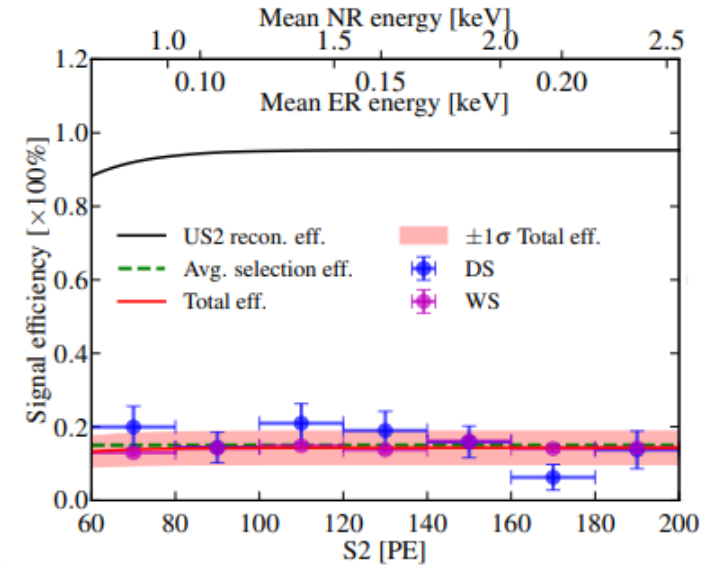


$$\sigma_{e\chi\nu}^A \approx 3\sigma_{e\chi\nu}^V \text{ w.r.t. the same } \Lambda$$

- A general fermionic (sterile neutrino-like) dark matter absorption on electron.
- In **soft** tension with XENON1T's ER excess, consistent with XENONnT latest result.
- A strong sensitivity to **vector** and **axial-vector** mediators, complementary to astrophysical constraints, but much less theoretical uncertainties.
- Competitive constraint in **20-55 keV/c²**.

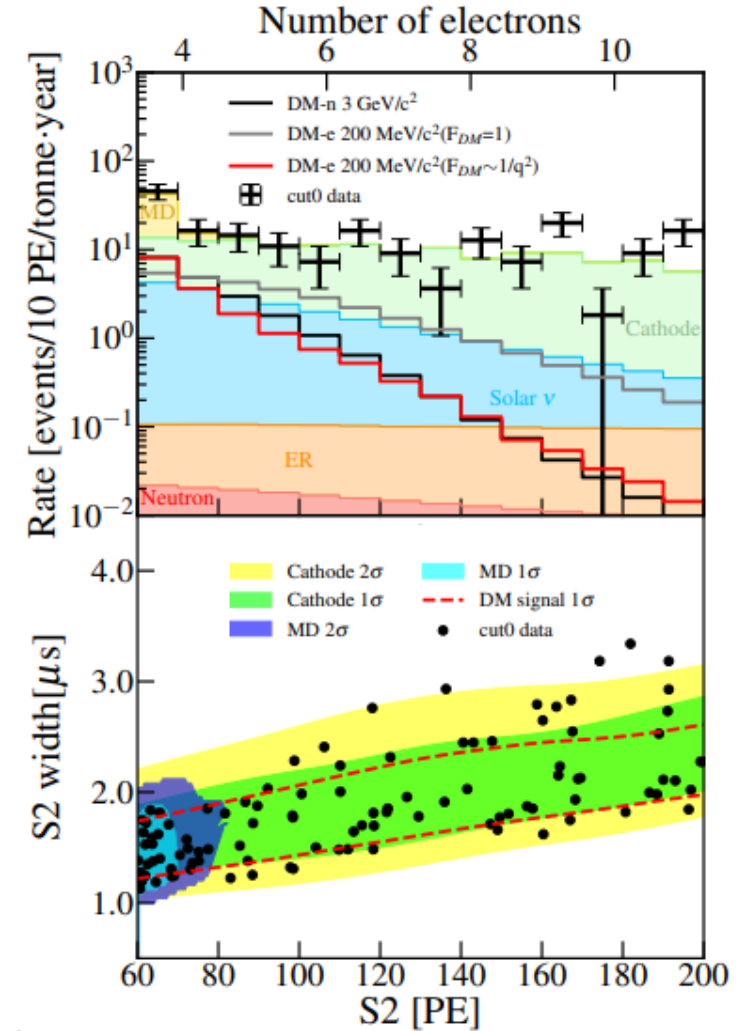
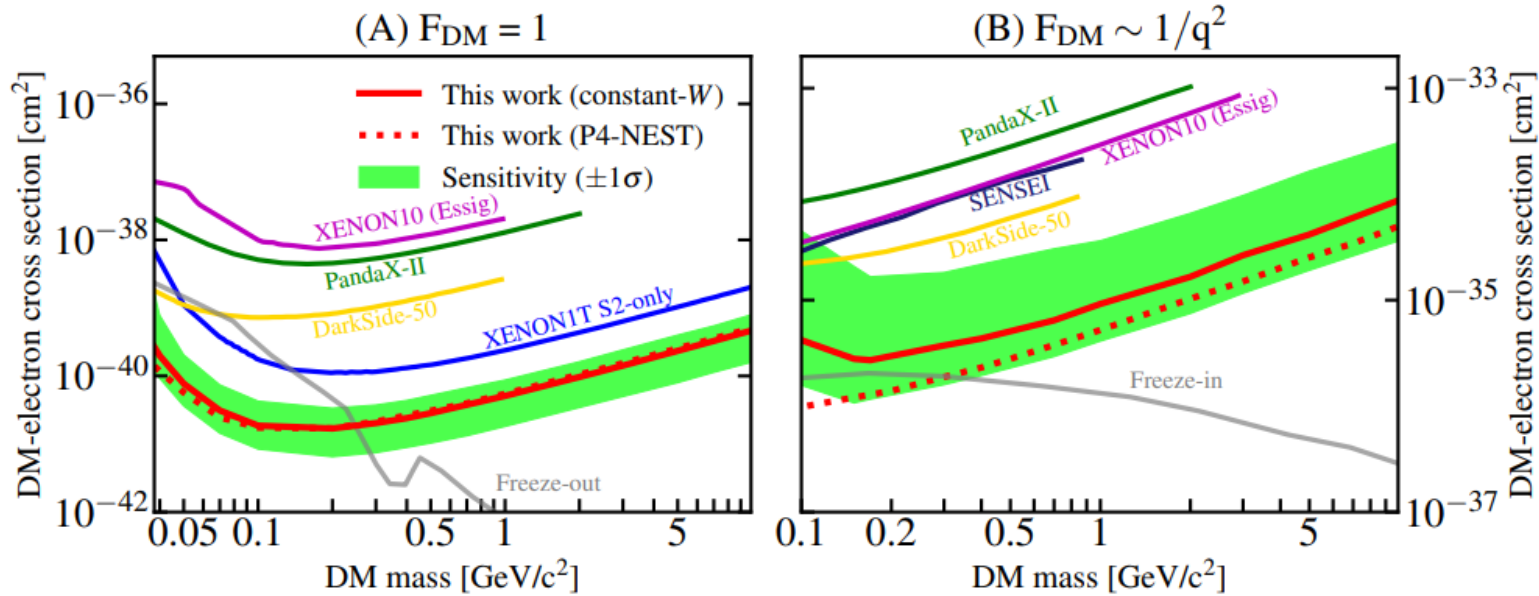
Ionization-only Search

- **Abandon the scintillation signal cut**
 - ROI: S2 [60, 200] PE
 - Threshold down to ~ 100 eV (from ~ 1 keV)
 - Tight quality cuts on the ionization signal
- **Background components**
 - **Micro-discharging (MD)**
 - Small charge, strong run-condition dependence
 - **Cathode activity**
 - Large charge, large pulse-shape width
 - **Data-driven estimation**
 - Validated in control region



Constraints on light dark matter

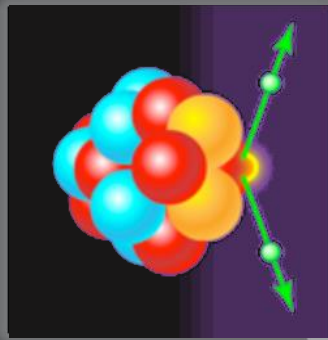
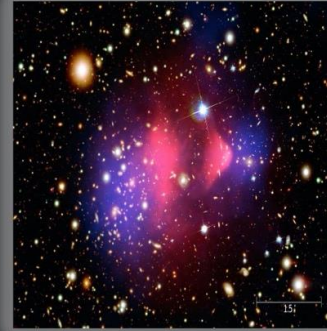
- Blind analysis of 0.55 tonne-year exposure
- Most stringent constraints are derived
 - DM-electron interaction, $2 \times 10^{-41} \text{ cm}^2$



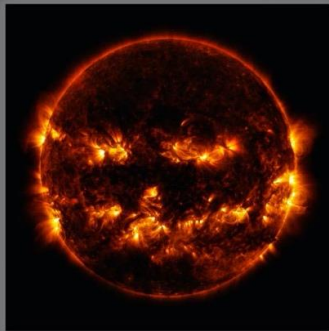
S. Li et al. [arXiv:2212.10067](https://arxiv.org/abs/2212.10067) Accepted by PRL, Editors' Suggestion

Future

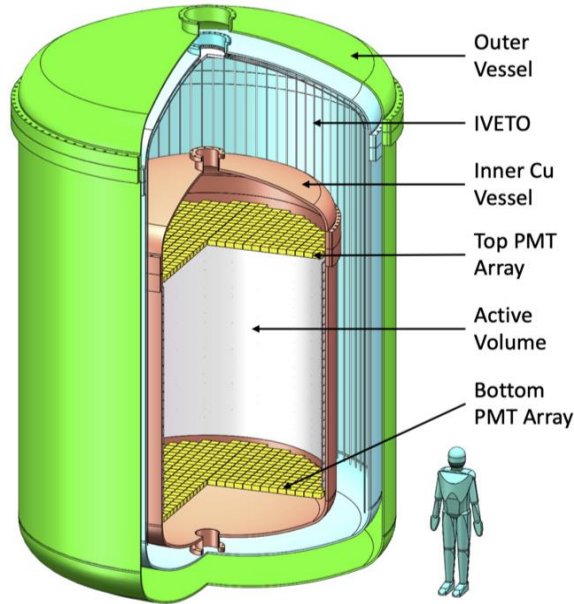
Dark Matter
1 keV – 10 keV



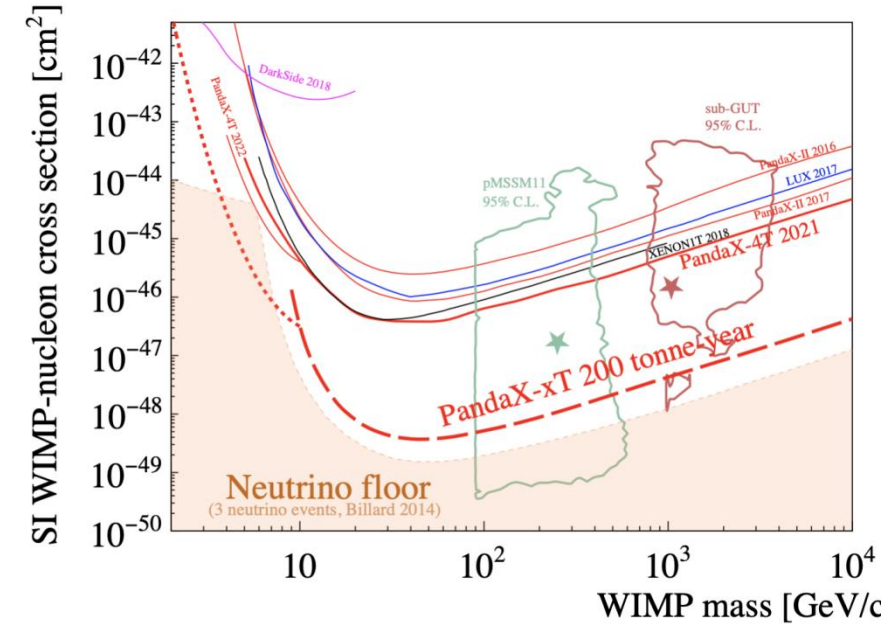
Majorana neutrino
> 2 MeV



Astrophysics neutrino < 300 keV



XeSAT2023



- Multi-physics goals;
- With > 30 tonne liquid xenon in the sensitive volume;
- Decisive test on WIMP and key test on Dirac/Majorana neutrino;
- Stay tuned!