

Introduction to the “Beyond the Standard Model” session

JJC 2014
Dec. 11th 2014

Samuel Calvet



Outline

- ◆ Why do we need “Beyond the Standard Model” (BSM) theories ?
- ◆ BSM theories on the market : their predictions/particles
 - SuperSYmetry, Extra-dimensions, compositness, ...



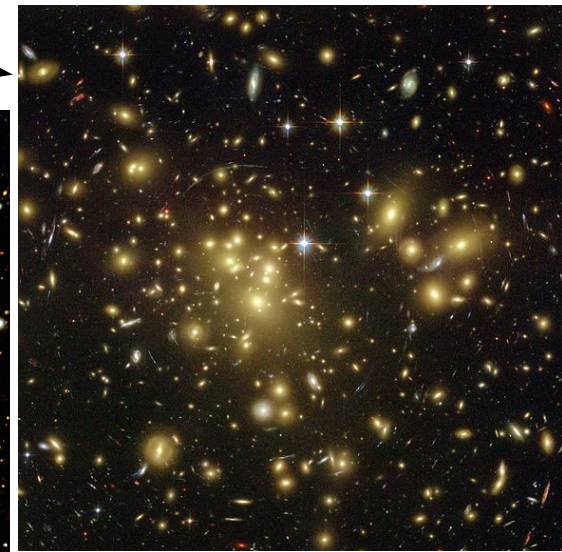
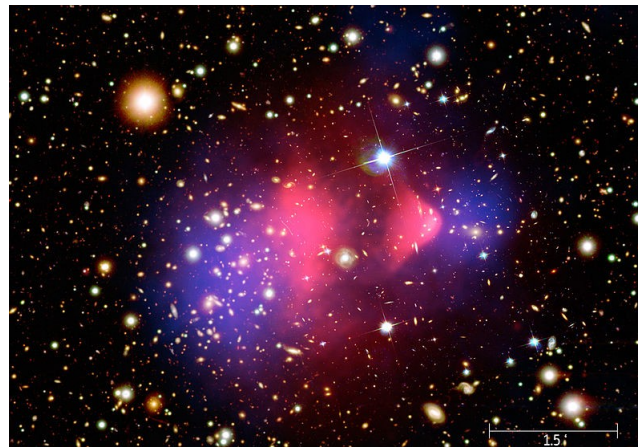
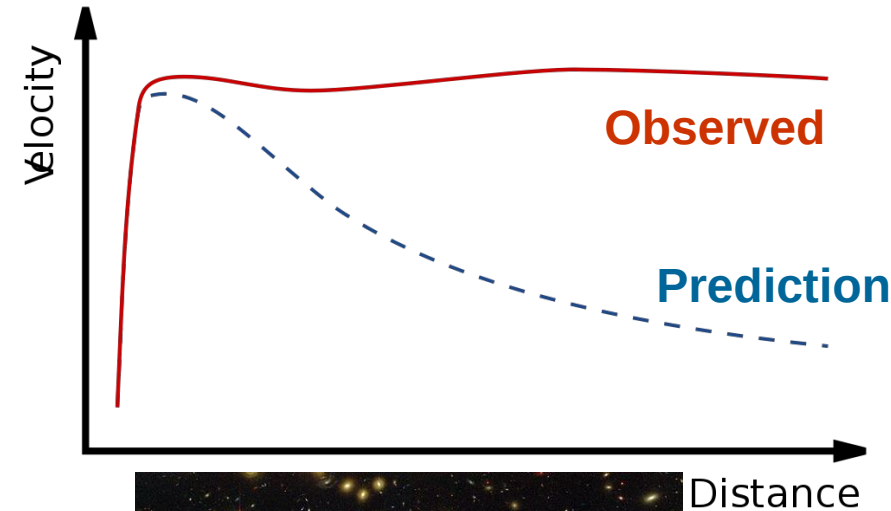
Why do we need BSM theories ?

2 reasons (among others) that drive searches at LHC ?

Dark matter

◆ Astrophysical observations in contradiction w/ the theories

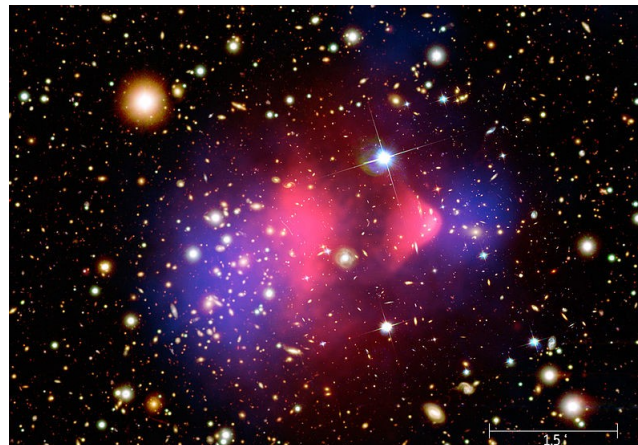
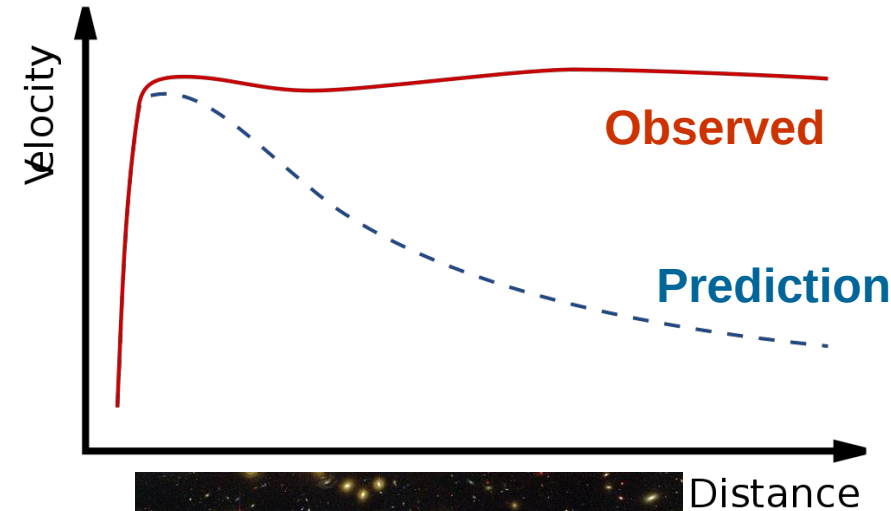
- Rotation of galaxies needs “extra mass”
- Cluster of galaxies:
 - Same velocity issue
 - Gravitational lensing
 - Bullet cluster
 -



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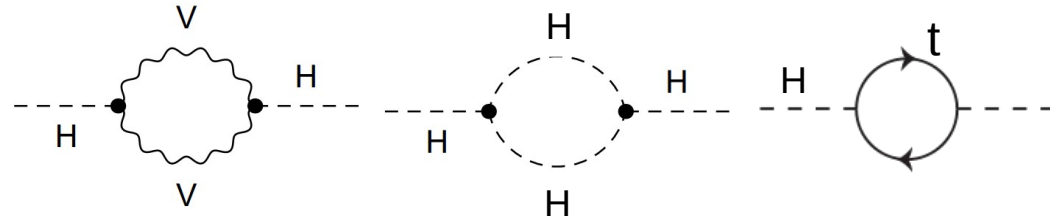
→ Need of neutral particles weakly interacting (dark matter)

Naturalness

- ◆ Higgs mass modified by quantum corrections

$$m_{h_{SM}}^2 = m_0^2 + \frac{3}{16\pi^2 v_{SM}^2} (2m_W^2 + m_Z^2 + m_{h_{SM}}^2 - 4m_t^2) \Lambda^2$$

Λ : scale of new physics



- ◆ If:

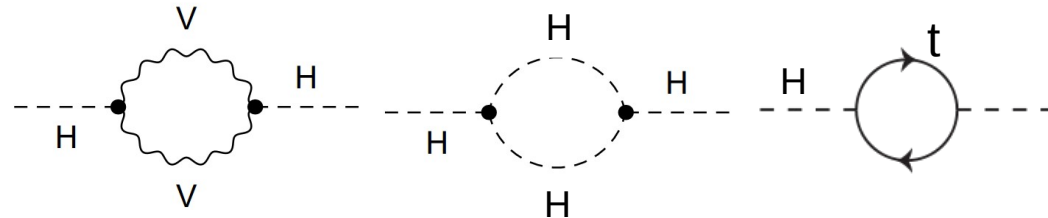
- Λ is large (up to Planck scale to include the gravity ?)
 - No ultra precise cancellation of terms (\rightarrow fine tuning)
- \rightarrow Large m_h

Naturalness

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$$m_{h_{\text{SM}}}^2 = m_0^2 + \frac{3}{16\pi^2 v_{\text{SM}}^2} (2m_W^2 + m_Z^2 + m_{h_{\text{SM}}}^2 - 4m_t^2) \Lambda^2$$

Λ : scale of new physic



- ◆ If:
 - Λ is large (up to Planck scale to include the gravity ?)
 - No ultra precise cancellation of terms (\rightarrow fine tuning) \rightarrow Large m_h

- ◆ Observation : $m_h \sim 125\text{GeV} \rightarrow$ light !

Not natural !

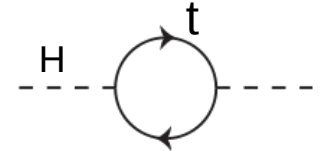
Solving naturalness issue

- ◆ Main ideas to solve the naturalness issue:

- No ultra precise cancellation of terms (\rightarrow fine tuning)
- Λ can be large (up to Planck scale to include the gravity ?)

- New symmetry

- Each correction balanced by another (new) one
- Protects m_H
 - \rightarrow supersymmetry



Solving naturalness issue

◆ Main ideas to solve the naturalness issue:

- No ultra precise cancellation of terms (→ fine tuning)
- Λ can be large (up to Planck scale to include the gravity ?)

- New symmetry
- New spatial dimensions
 - Bring the Planck scale to lower value
→ Λ is small

Solving naturalness issue

- ◆ Main ideas to solve the naturalness issue:

- No ultra precise cancellation of terms (→ fine tuning)
- Λ can be large (up to Planck scale to include the gravity ?)

- New symmetry
- New spatial dimensions
- Higgs boson is not the SM one
 - Higgs is a composite particle at scale Λ
 - naturalness issue disappears

Solving naturalness issue

◆ Main ideas to solve the naturalness issue:

- No ultra precise cancellation of terms (→ fine tuning)
- Λ can be large (up to Planck scale to include the gravity ?)

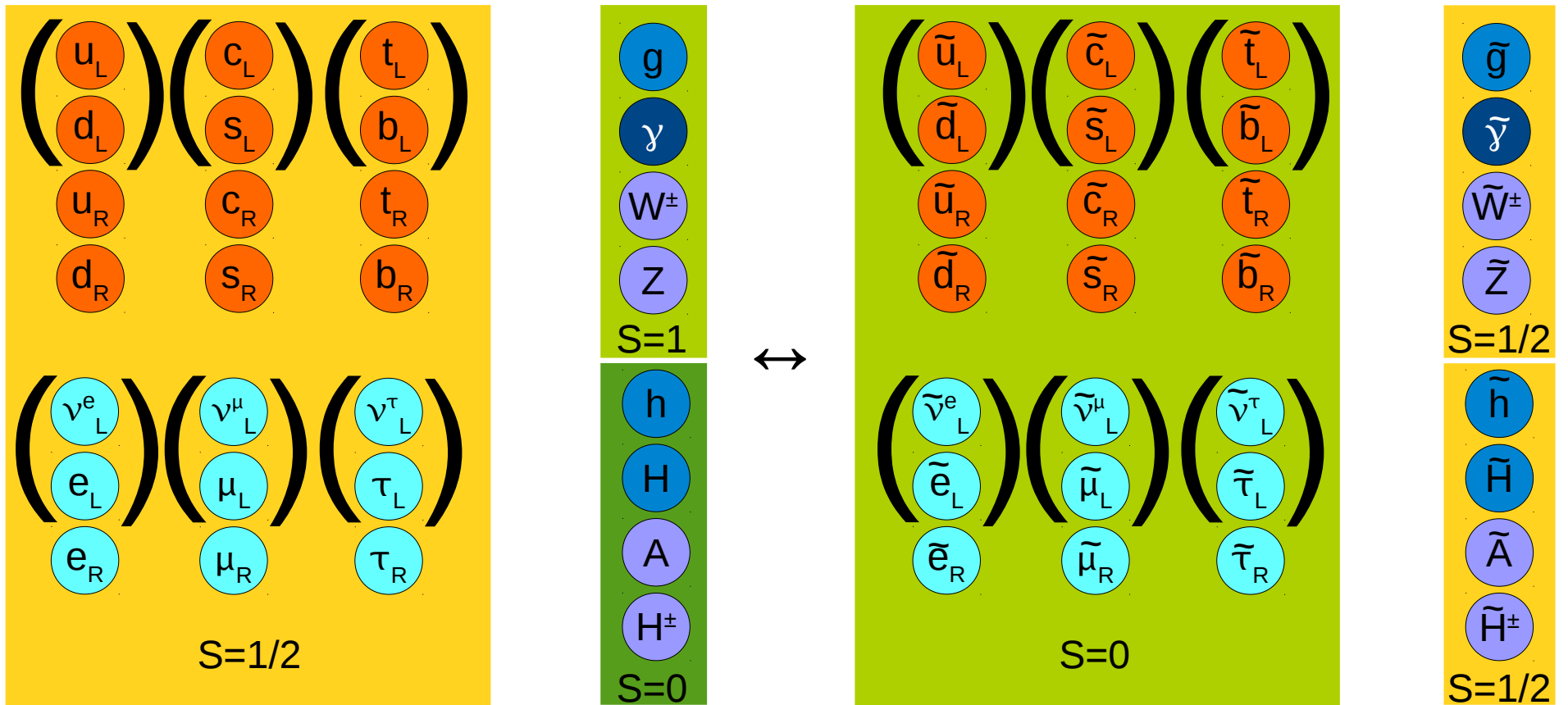
- New symmetry
- New spatial dimensions
- Higgs boson is not the SM one

Have to appear at the TeV scale
to be efficient

Supersymmetry

Supersymmetry

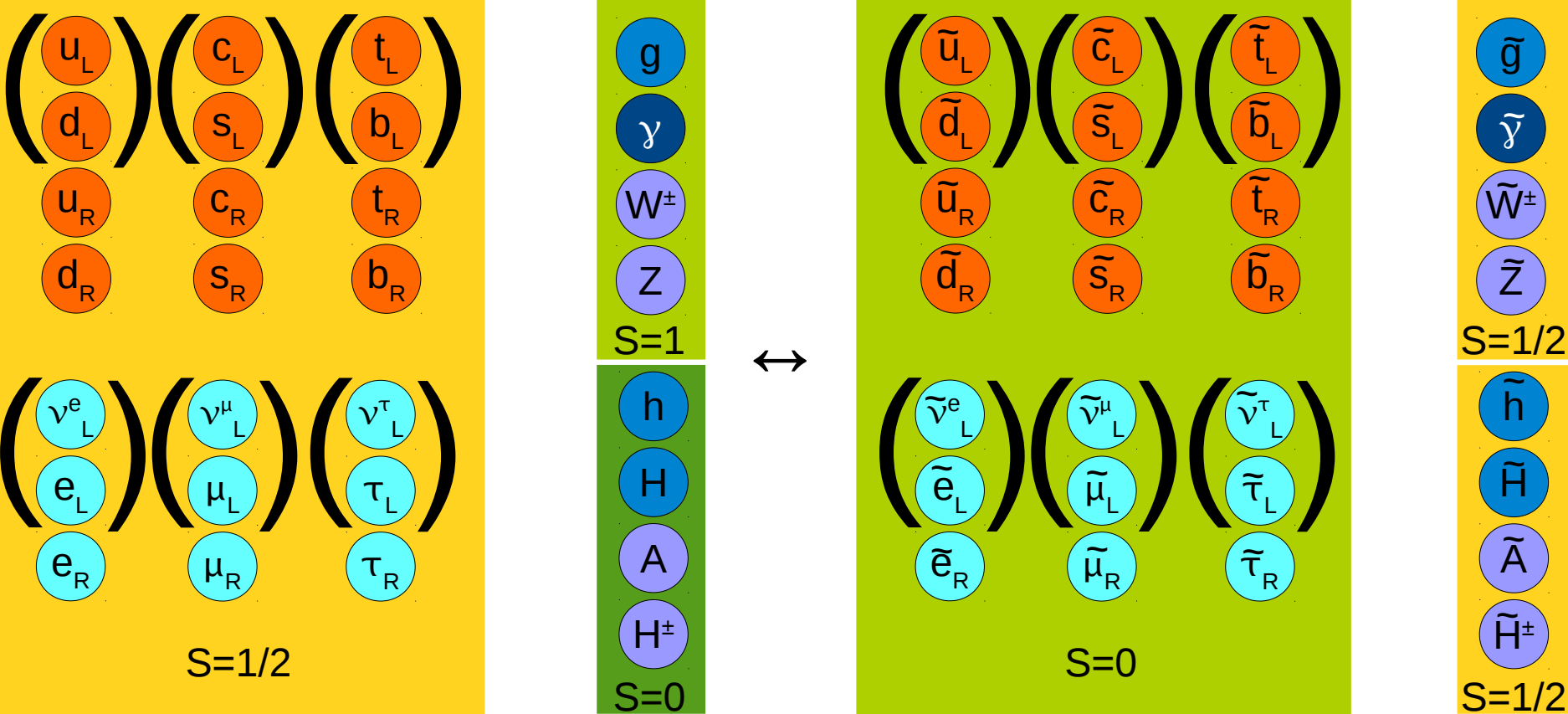
◆ Add new symmetry: fermion \leftrightarrow boson



Extended Higgs sector

Supersymmetry

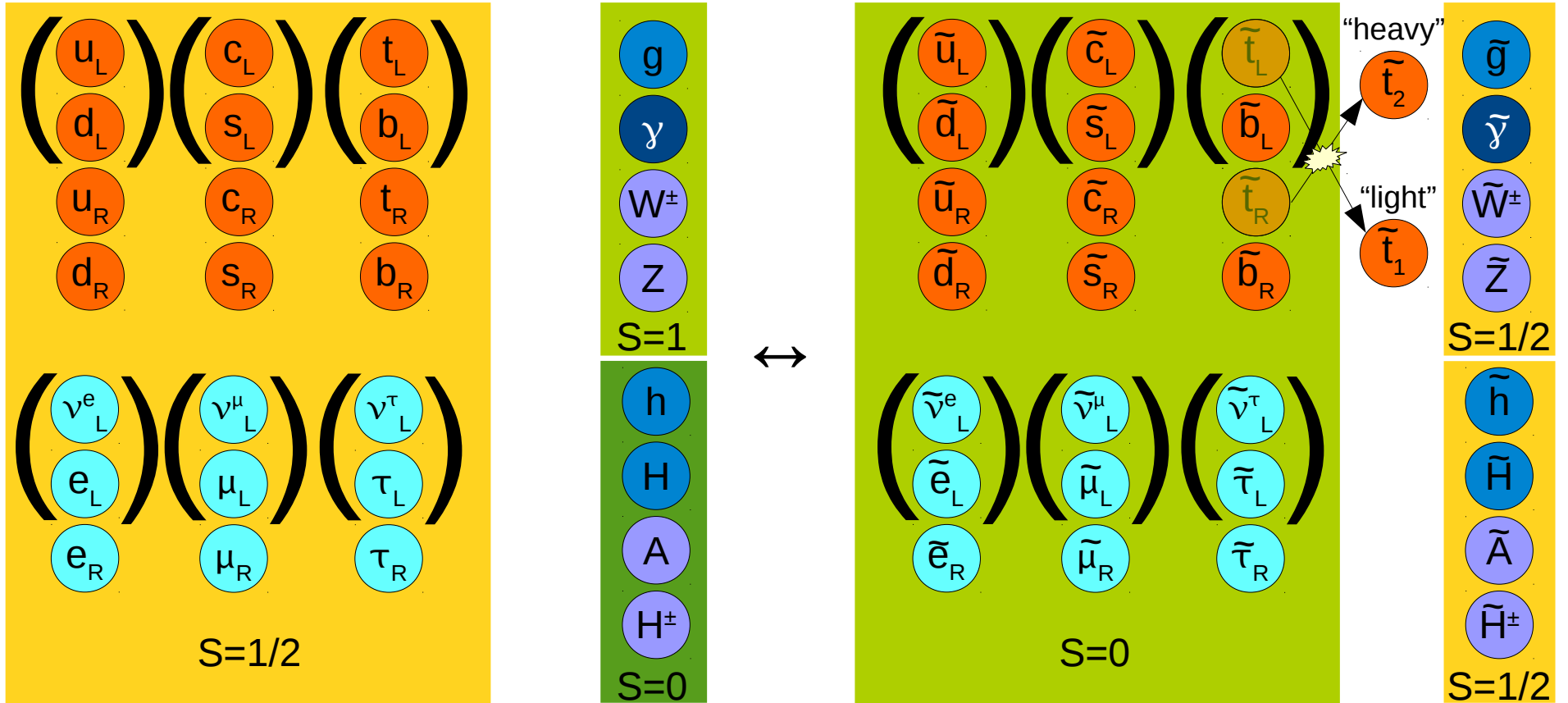
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 → cancellation of radiative corrections



Extended Higgs sector

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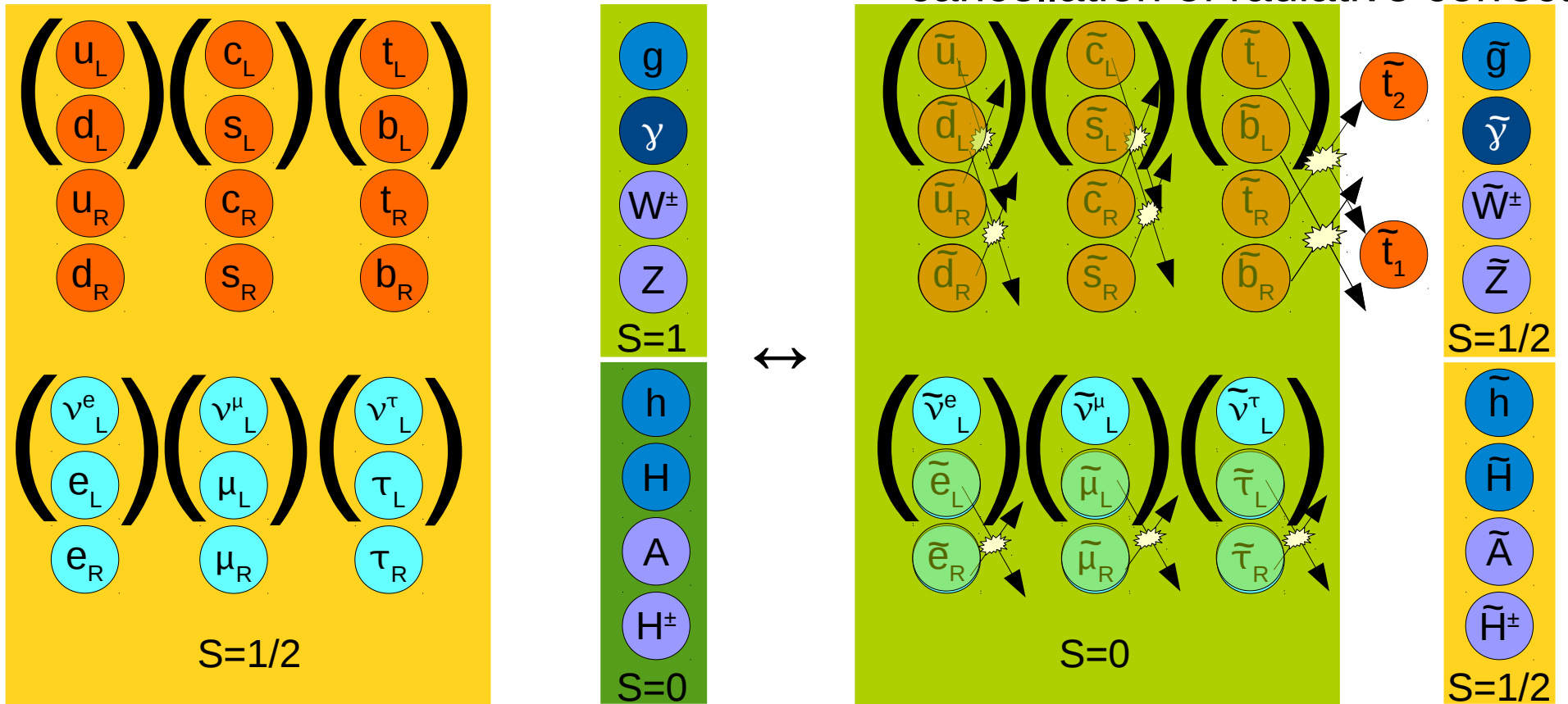
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- ◆ Add a pinch of mixing

Supersymmetry

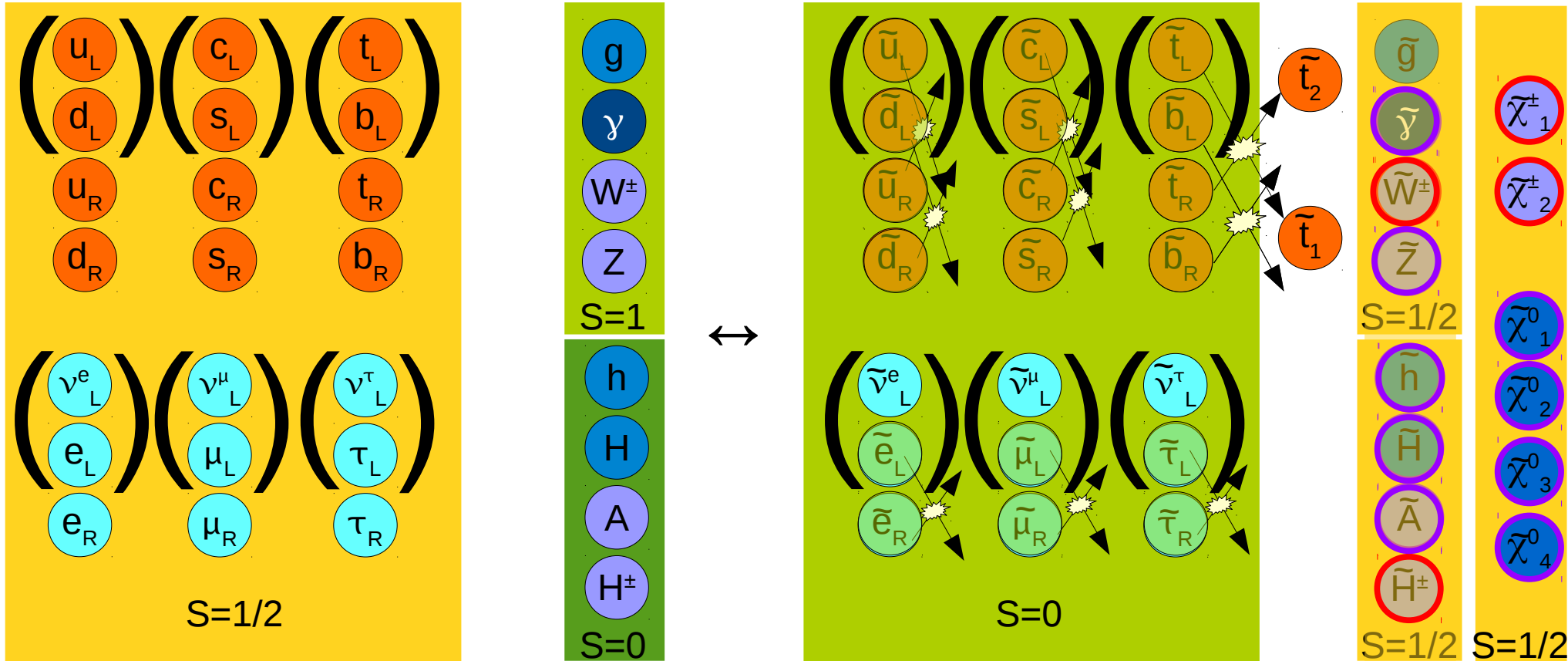
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Supersymmetry

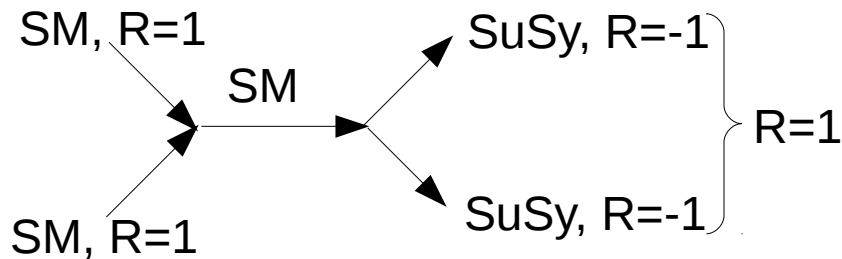
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 → cancellation of radiative corrections



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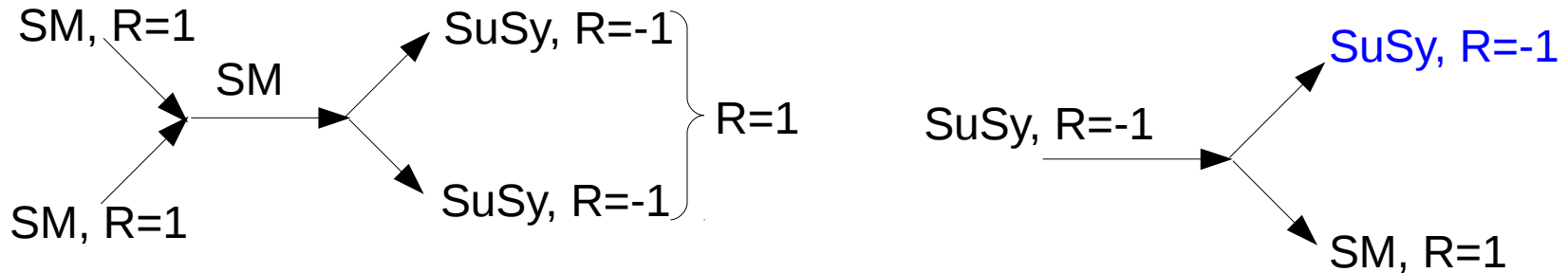
Supersymmetry: R-parity

- ◆ $R = (-1)^{L+3B+2J}$ L/B: leptonic/baryoni number, J:spin
 - +1 for SM particles
 - -1 for SuSy particles
- ◆ R-parity:
 - Assumed prefect or weakly violated (to preserve proton lifetime)
 - SuSy particle **produced by pair**



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- ◆ R-parity:
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 - SuSy particle **produced by pair**
 - Lightest SuSy particle (**LSP**) is stable → **Dark matter candidate**

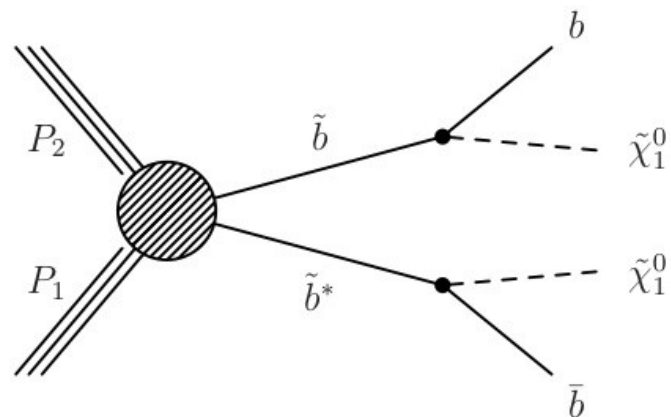


Breaking SuSy

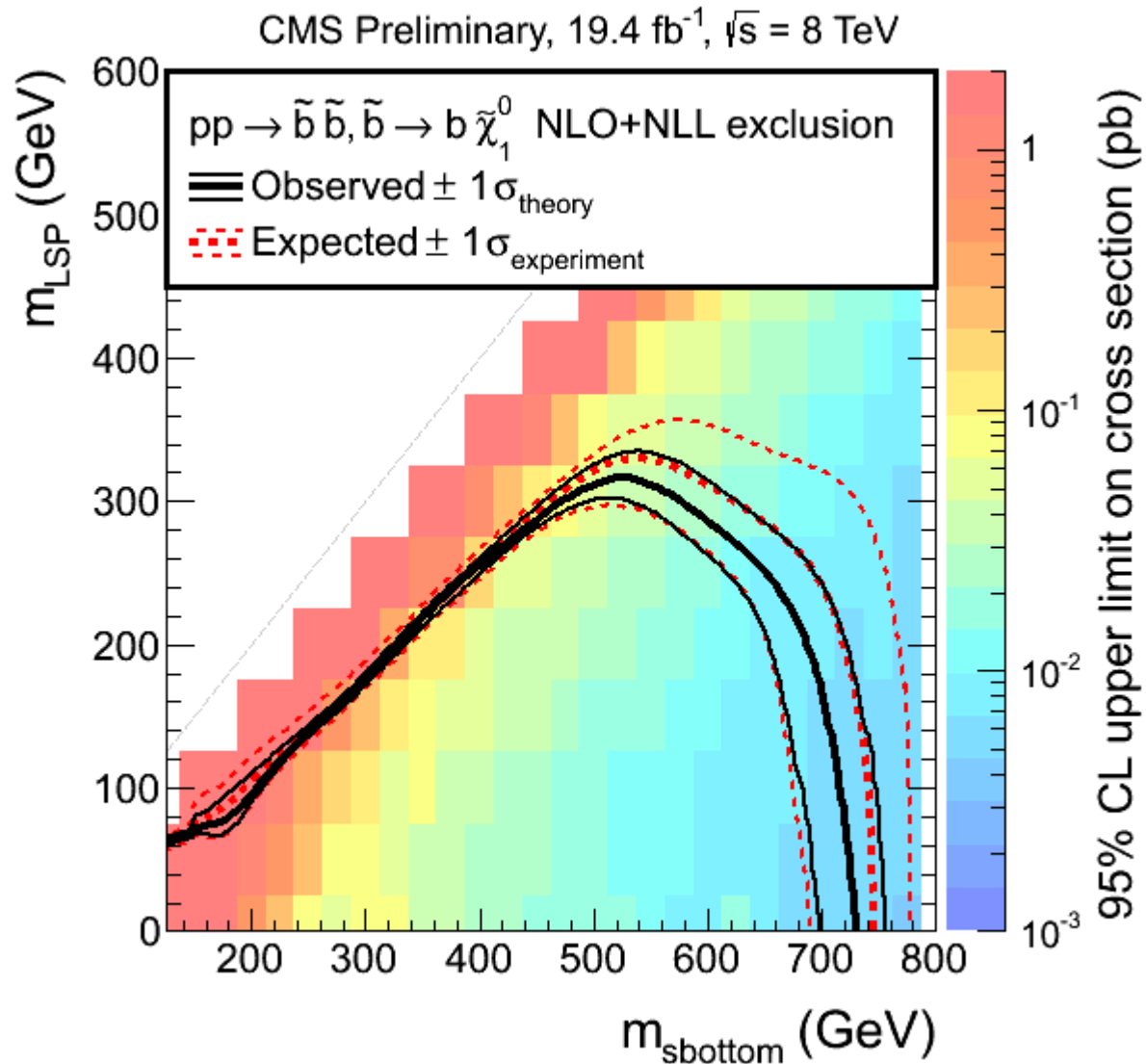
- ◆ Super-partners not yet observed → heavier than SM partners
- ◆ SuSy has to be **broken**
 - By an unknown mechanism
 - Introduces **many free parameters**
 - **Phenomenological assumptions** to express results
(mSUGRA, MSSM, nMSSM, pMSSM, ...)
- or
- **Simplified models:**
 - branching ratio=100%
 - decouple other sparticles

A rich phenomenology

- ◆ A simple example

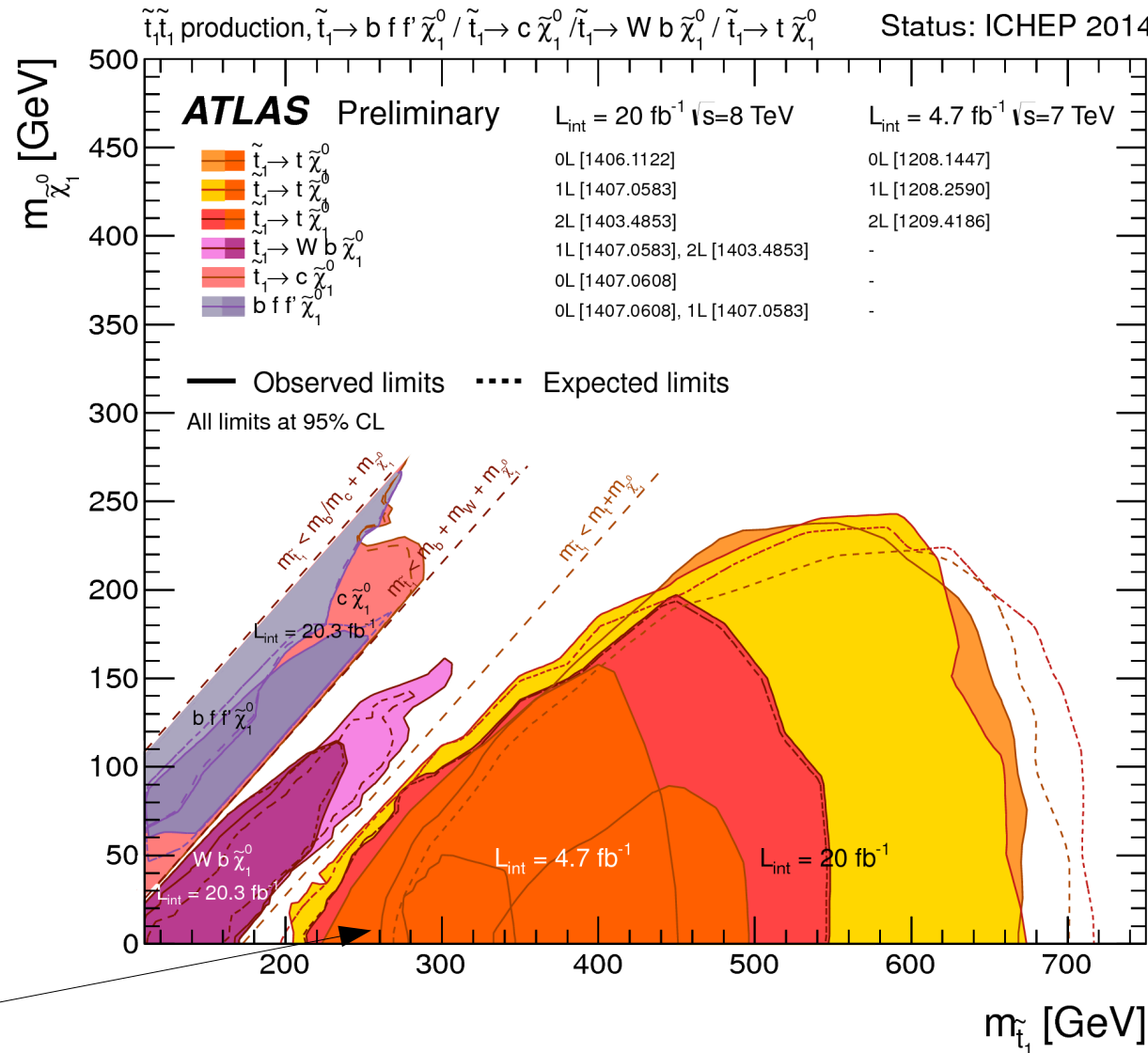
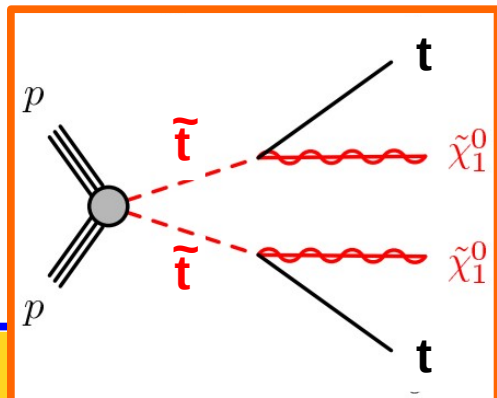


2 b + missing energy



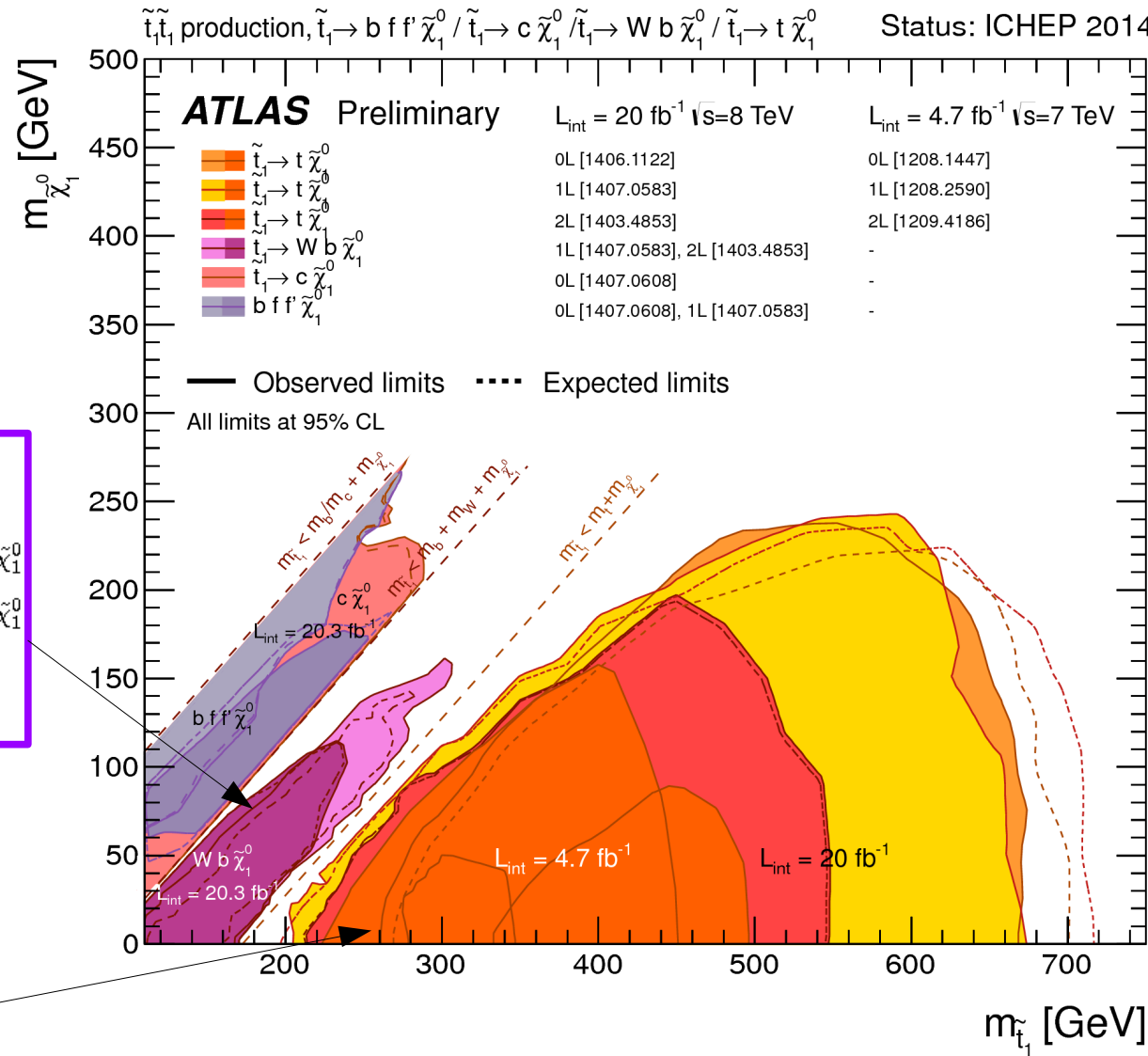
A rich phenomenology

- ◆ A more complex one



A rich phenomenology

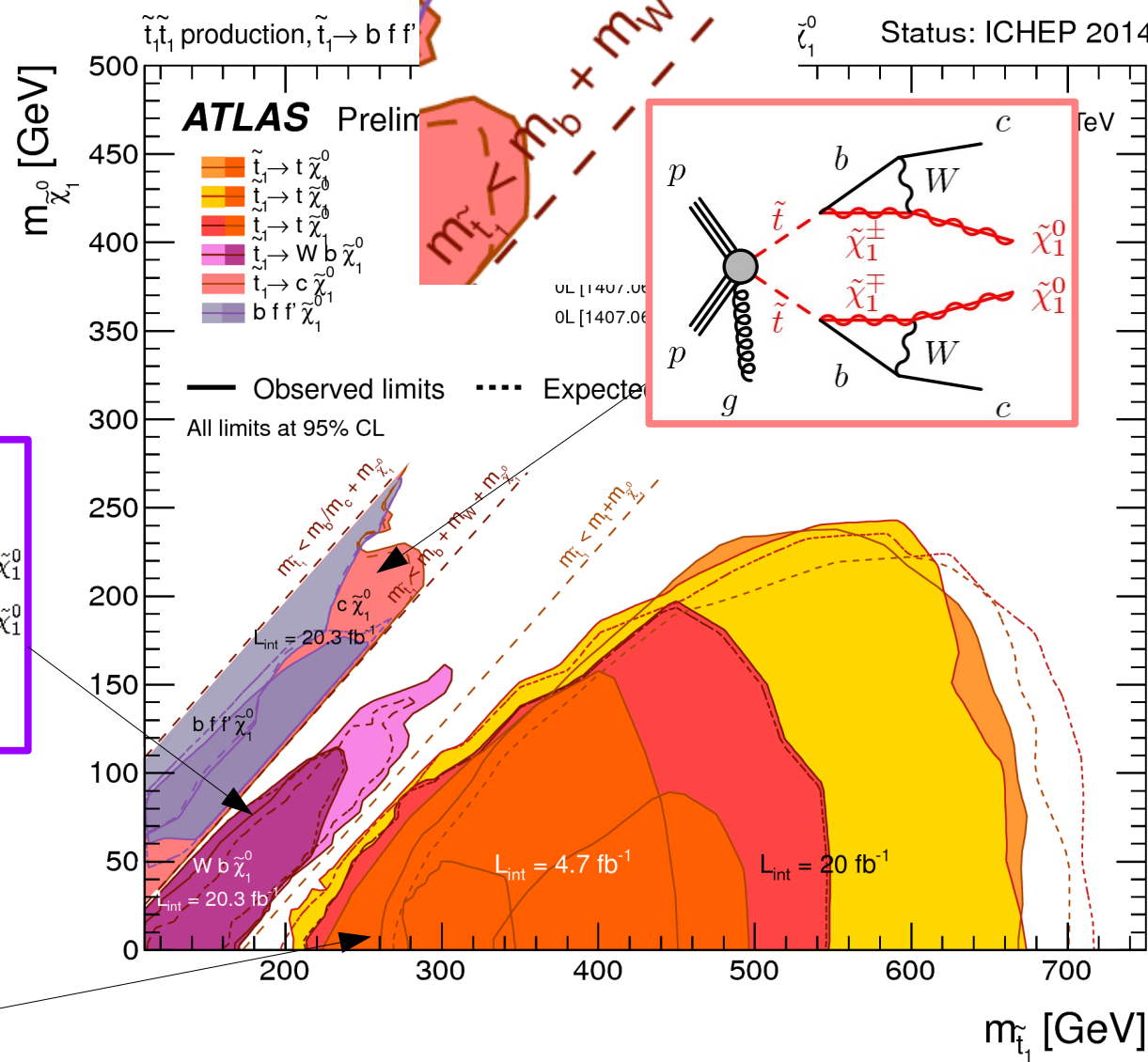
- ◆ A more complex one



$m_{\tilde{t}_1} < m_t + m_{\tilde{\chi}_1^0}$

A rich phenomenology

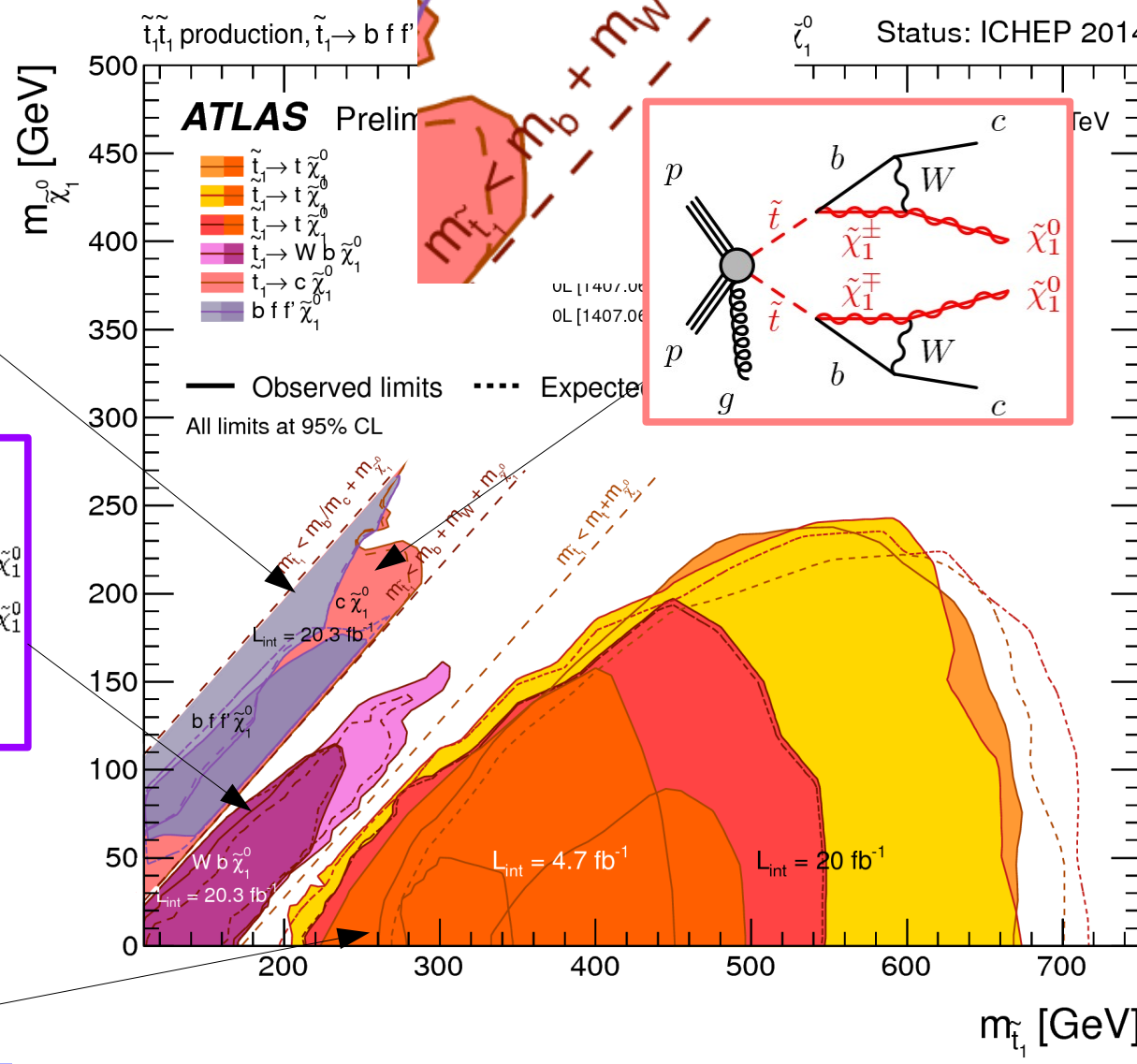
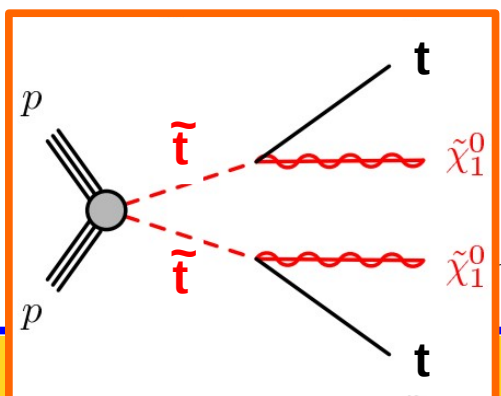
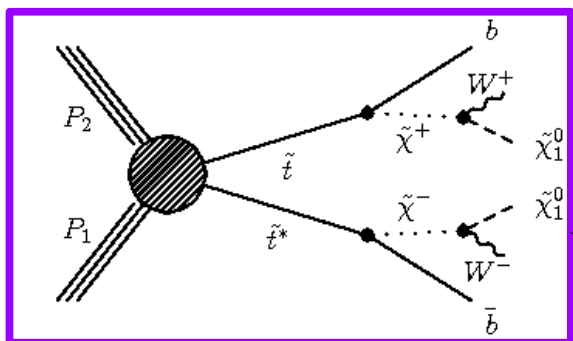
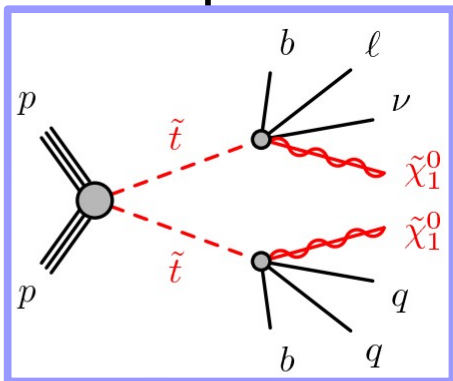
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A rich phenomenology

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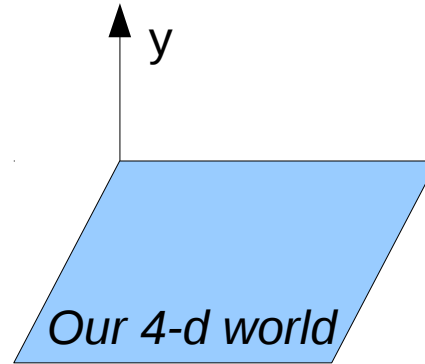
Extra-dimensions

Nice ref: M. Besancon, Moriond EW 2010
Models & signatures of extra dimensions at the LHC

Kaluza–Klein excitations

- ◆ Add a space-dimension (y)

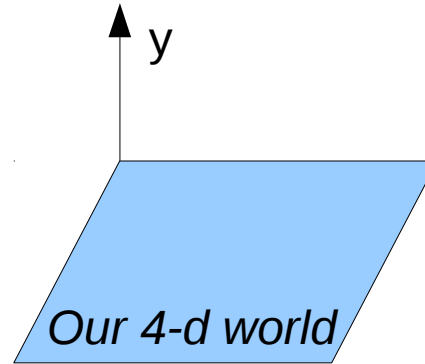
Flat (ie factorisable): $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$
 $\mu, \nu = 0, 1, 2, 3, \dots, D$



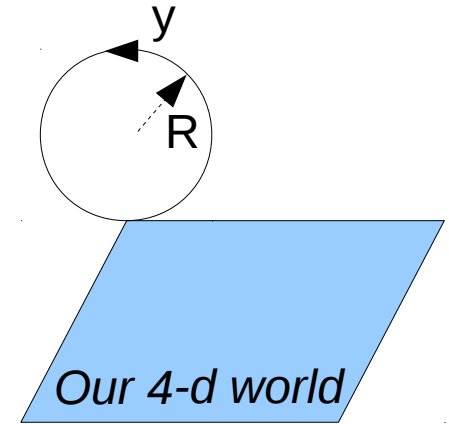
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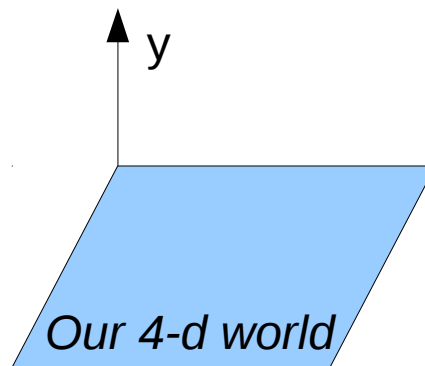
Compactification



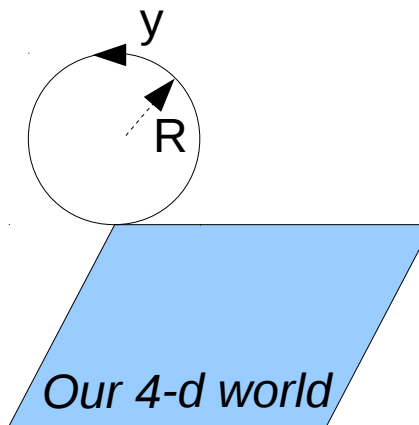
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Compactification



Fourier mode expansion

$$\Phi(x, y) = \sum_k \Phi^{(k)}(x) e^{iky/R}$$

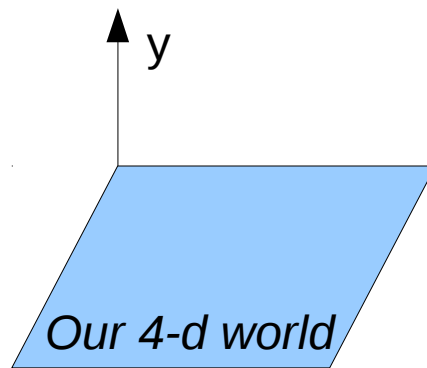
Kaluza–Klein (KK) excitations
 $m_k^2 = m_0^2 + k^2/R^2$

- ◆ Assume some fields can propagate along y
 - Momentum in new dimensions \leftrightarrow mass 4-D (Kaluza–Klein resonance)

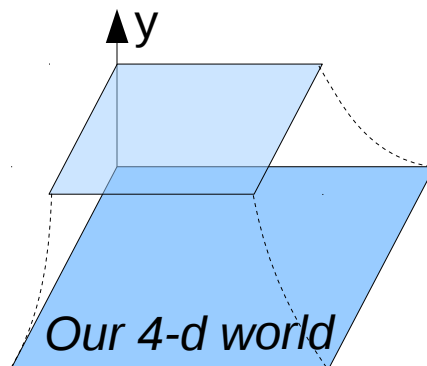
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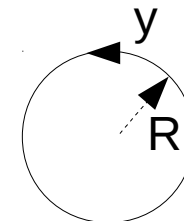
Flat (ie factorisable): $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$
 $\mu, \nu = 0, 1, 2, 3, \dots, D$



Wrapped: $ds^2 = a(y) (g_{\mu\nu} dx^\mu dx^\nu) + dy^2$
 $\mu, \nu = 0, 1, 2, 3$



Compactification



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Kaluza–Klein (KK) excitations

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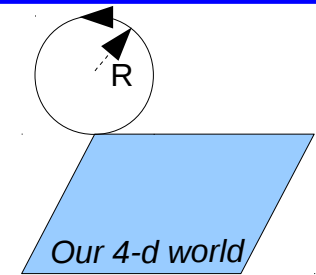
Adding extradimensions...

A lot of room to play...

| | Topology | N extradim | Propagating fields |
|-------------------|----------|------------|------------------------|
| ADD | Flat | ≥ 2 | Graviton |
| TeV ⁻¹ | Flat | ≥ 1 | Vector boson (V) |
| mUED | Flat | ≥ 1 | V, fermions |
| RS1 | Wrapped | 1 | Graviton |
| Bulk RS | Wrapped | 1 | Everything apart Higgs |
| ... | | | |

ADD (Arkani-Hamed, Dimopoulos and Dvali)

- ◆ n compactified flat extra-dimensions
- ◆ Only **gravity propagates in the bulk**
 - Dilution of the gravity



$$M_{Pl}^2 \sim M_D^{2+n} R^n$$

Planck mass in 4+nD

Planck mass in 4D

- ◆ If $M_{Pl} = 1\text{TeV}$ (to address the hierarchy problem):

- ~~$n=1 \rightarrow R=10^{10}\text{km}$~~

- $n=2 \rightarrow R=1\text{mm}$

- $n=3 \rightarrow R=1\text{nm}$

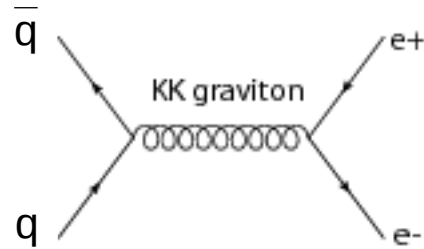
- ...

} Large R , but gravity not yet probed at these scales !

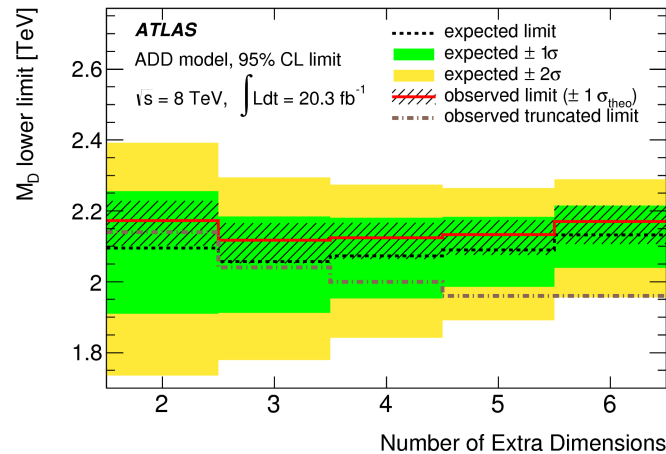
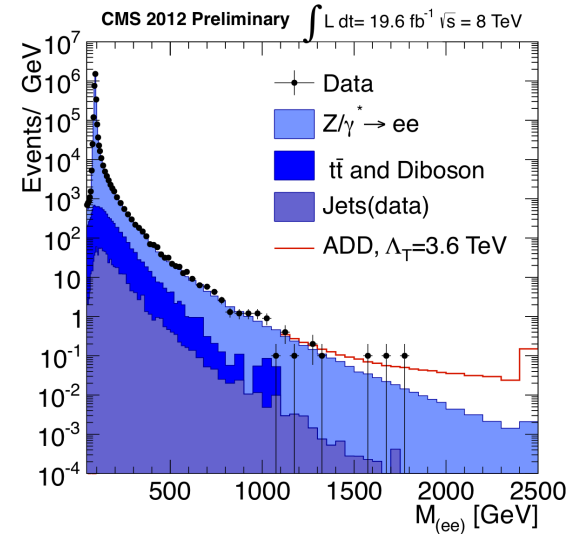
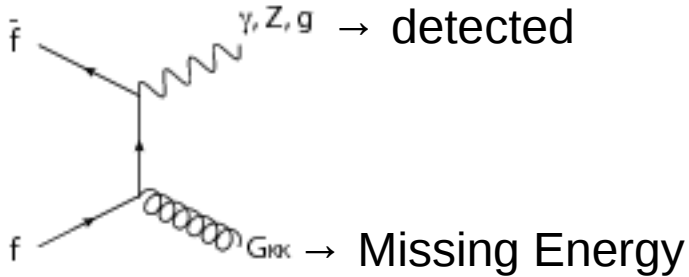
ADD signatures at colliders

◆ “Large” $R \rightarrow$ states close to each other ($\Delta m \sim eV$) \rightarrow continuum

◆ s-channel



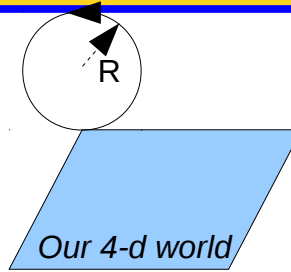
◆ KK graviton in final states



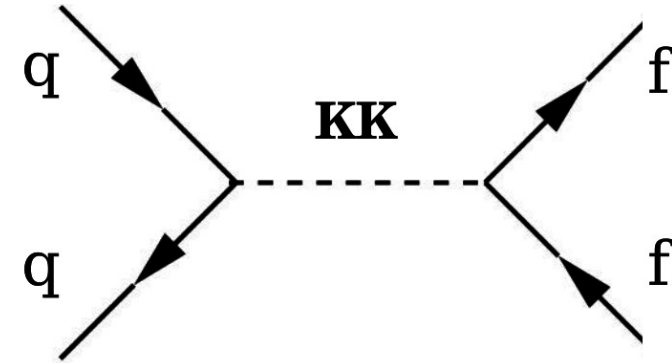
Gamma+MET

TeV⁻¹ / mUED

- 5D flat bulk
 - No gravity included



- TeV⁻¹:**
 - $R = O(\text{TeV}^{-1})$
 - gauge bosons in the bulk $\rightarrow Z', W', g_{\text{KK}}$
 - 1st, 2nd, 3rd, ... modes well separated



Minimal Universal Extra-Dimensions

- All the SM fields in the bulk
- Assume momentum conservation in the bulk
 - \rightarrow KK-parity

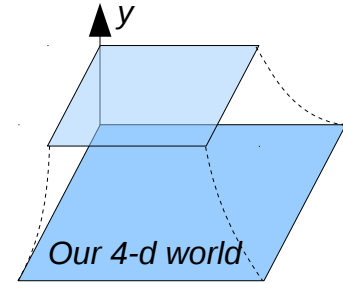
Similar effect than R-parity (pair production of 1st mode, dark matter candidate)

Randal–Sundrum (RS)

- ◆ Wrapped compacted 5D

$$ds^2 = e^{-2kR\theta} (g_{\mu\nu} dx^\mu dx^\nu) + R d\theta^2 \quad (\theta \in [0, \pi], k \sim M_{pl})$$

Wrap factor

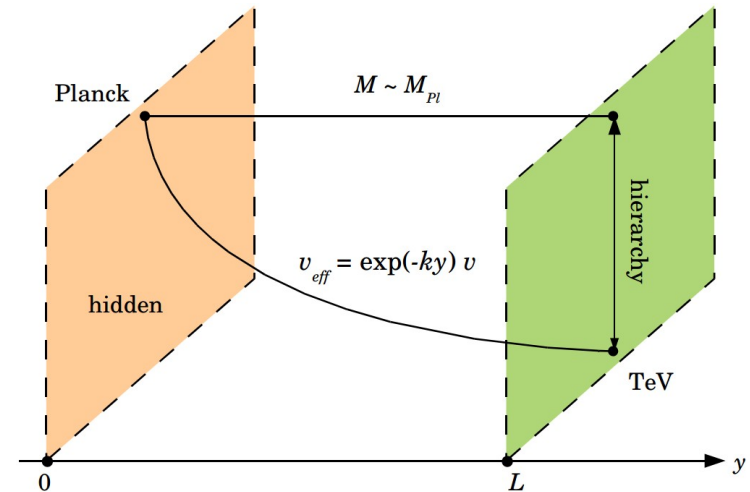


- ◆ Wrap generates **vev ~ O(TeV) on a brane** from **Planck scale on another brane**

→ Fix hierarchy problem

- ◆ Graviton near Planck brane

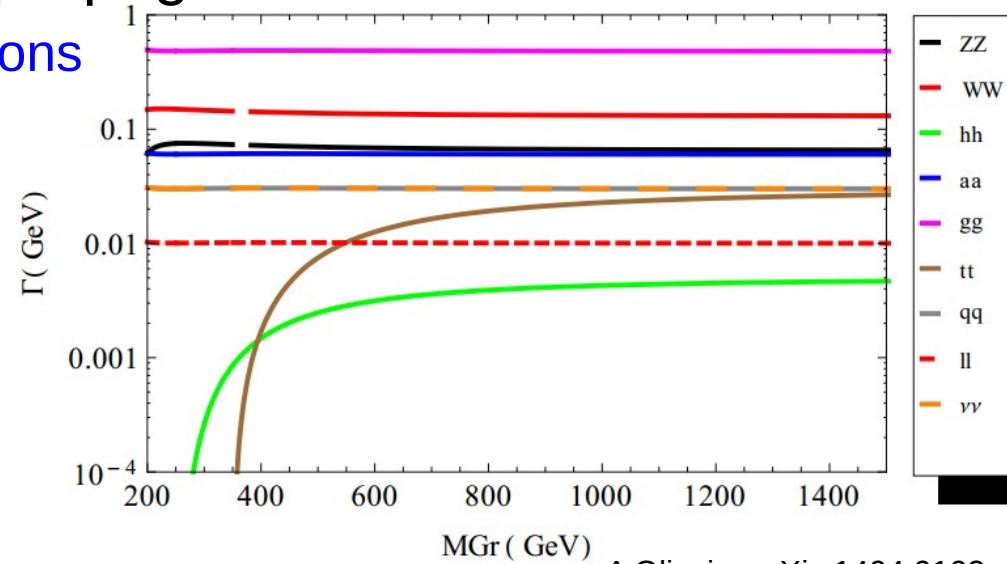
- G_{KK} near TeV brane



Randal–Sundrum (RS)

Minimal RS (RS1): Only gravity propagates in bulk

- Well separated ($\Delta m \sim \text{TeV}$) KK gravitons
- Democratic couplings



A Oliveira arXiv:1404.0102

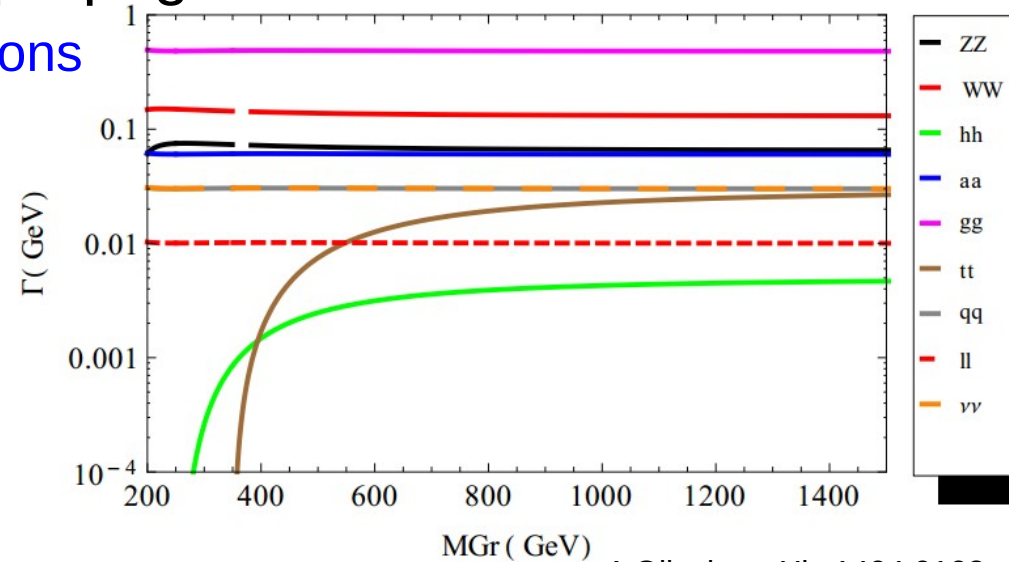
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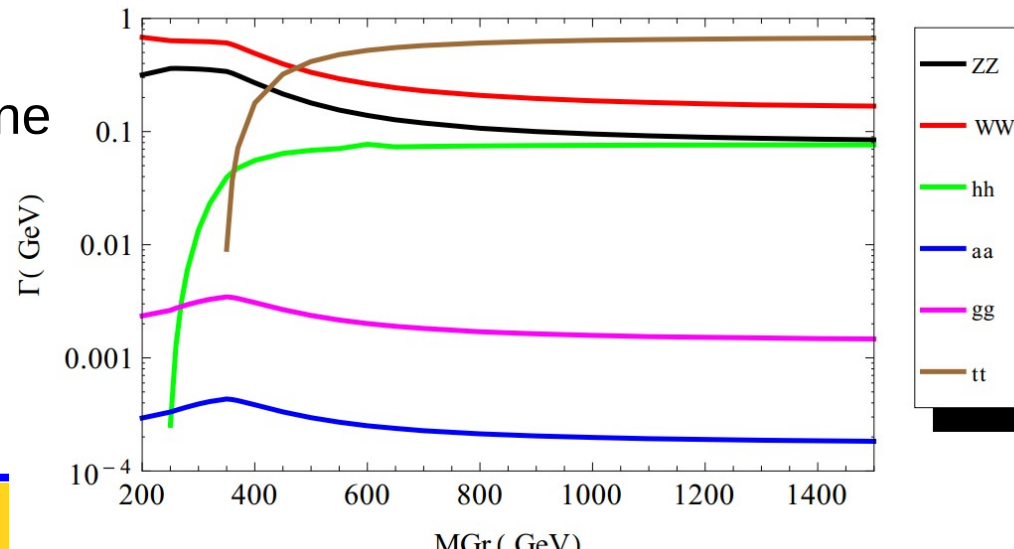
- Well separated ($\Delta m \sim \text{TeV}$) KK gravitons
- Democratic couplings

Bulk RS:

- Higgs localized on/near TeV brane
- Other fields in bulk
 - KK resonances
- 1st and 2nd generations : Planck brane
- 3rd generation: TeV brane
 - \sim large Yukawa
- Prediction of Vector-like quark (see next slides)



A Oliveira arXiv:1404.0102



Compositness

Compositeness

- ◆ Assume **new Strong sector** (QCD-like) at TeV-scale
 - $SO(5) \rightarrow SO(4) + 4$ Goldstone bosons (=Higgs doublet)
→ Higgs boson is composite (like pion in QCD)
- ◆ Consequences:
 - Fix naturalness issue
 - Lagrangian modified as:

$$\mathcal{L} = \frac{M_V^2}{2} V_\mu^2 \left(1 + 2a \frac{h}{v} + b \frac{h^2}{v^2} \right) - m_f \bar{\psi}_L \psi_R \left(1 + c \frac{h}{v} \right) + \dots$$

$$a = \sqrt{1 - \frac{v^2}{f^2}} \quad b = 1 - \frac{2v^2}{f^2} \quad c = \sqrt{1 - \frac{v^2}{f^2}}$$

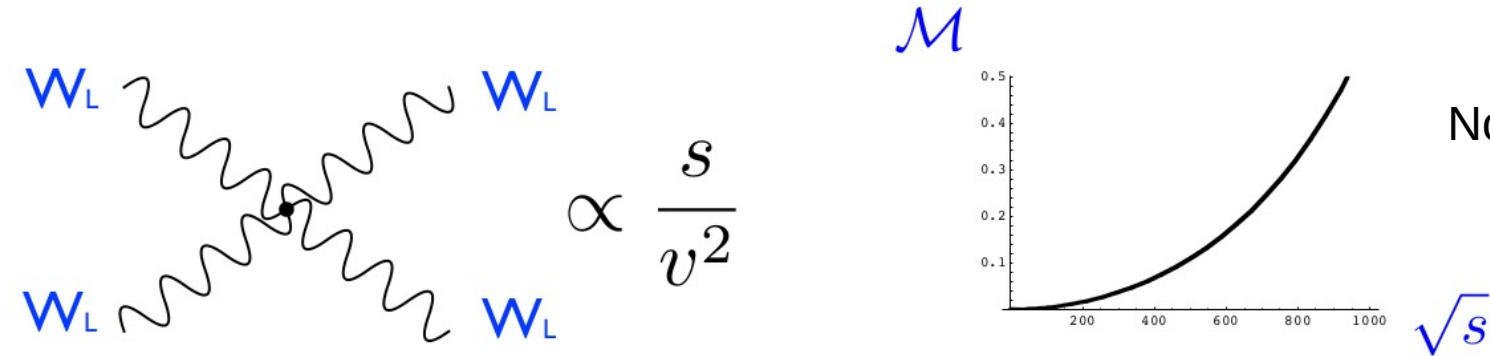
(SM: $a=b=c=1$)

Scale related to the composite-scale

Stolen to A. Pomarol

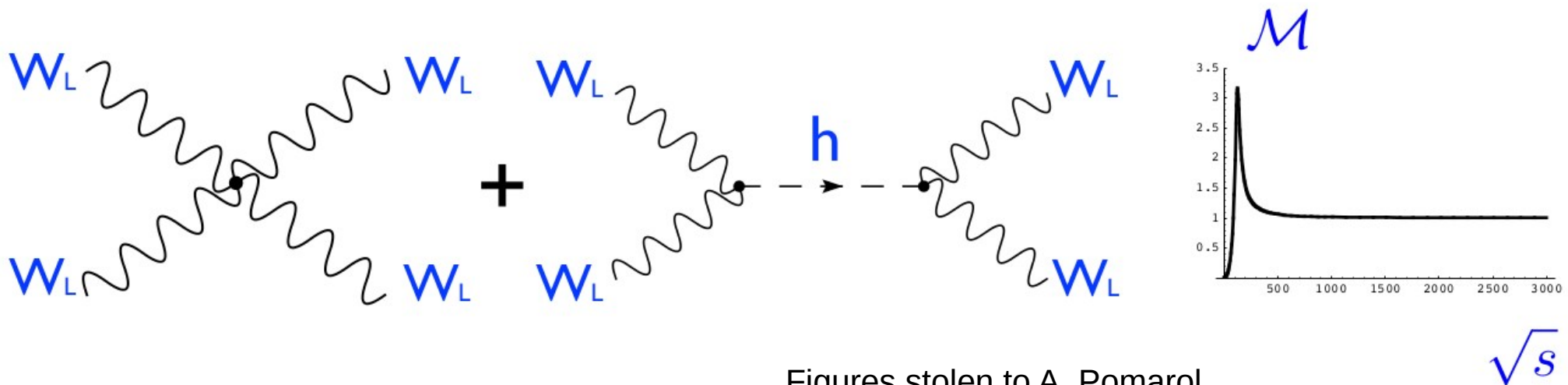
Compositeness: consequence

- SM without Higgs



No unitarity at high energy

- SM with Higgs: unitarity recovered

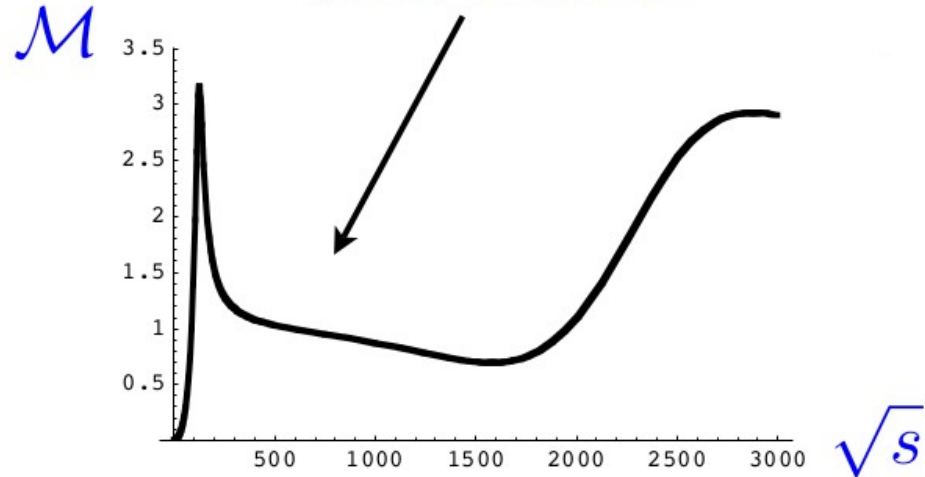


Figures stolen to A. Pomarol

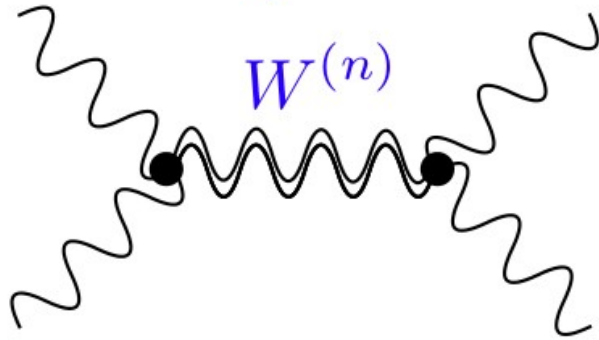
Compositeness: consequence

- ◆ If the Higgs is (partially) composite:

partly unitarize!



→ Need extra states:



$$M_{W^{(n)}} \simeq \frac{2 \text{ TeV}}{\sqrt{1 - a^2}}$$

Figures stolen to A. Pomarol

- ◆ Also predicts **Vector-Like Quarks (VLQ)**

Vector-Like Quark (VLQ)

- ◆ Predicted by many theories
 - Extra-dimension, Compositeness, Grand Unified Theory, ...
- ◆ Vector-like ?

Vector-Like Quark (VLQ)

- ♦ Predicted by many theories
 - Extra-dimension, Compositeness, Grand Unified Theory, ...
- ♦ Vector-like ?

$$\mathcal{L}_W = \frac{g}{\sqrt{2}} \left(J^{\mu+} W_\mu^+ + J^{\mu-} W_\mu^- \right) \quad \text{Charged current Lagrangian}$$

- SM chiral quarks: ONLY left-handed charged currents

$$J^{\mu+} = J_L^{\mu+} + J_R^{\mu+} \quad \text{with} \quad \begin{cases} J_L^{\mu+} = \bar{u}_L \gamma^\mu d_L = \bar{u} \gamma^\mu (1 - \gamma^5) d = V - A \\ J_R^{\mu+} = 0 \end{cases}$$

- vector-like quarks: BOTH left-handed and right-handed charged currents

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L. Panizzi

Vector-Like Quark (VLQ)

- ♦ Predicted by many theories
 - Extra-dimension, Compositeness, Grand Unified Theory, ...
- ♦ Vector-like ?

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$$J^{\mu+} = J_L^{\mu+} + J_R^{\mu+} = \bar{u}_L \gamma^\mu d_L + \bar{u}_R \gamma^\mu d_R = \bar{u} \gamma^\mu d = V$$

L. Panizzi

$$\mathcal{L}_M = -M \bar{\psi} \psi \quad \text{Gauge invariant mass term without the Higgs}$$

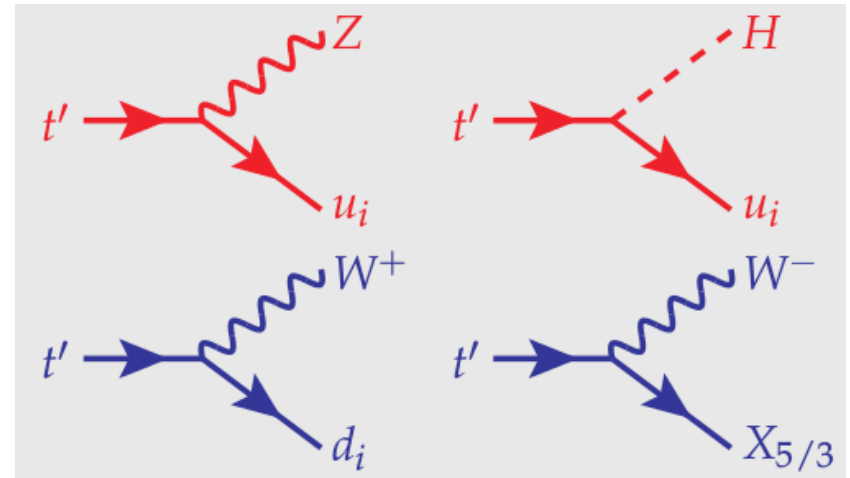
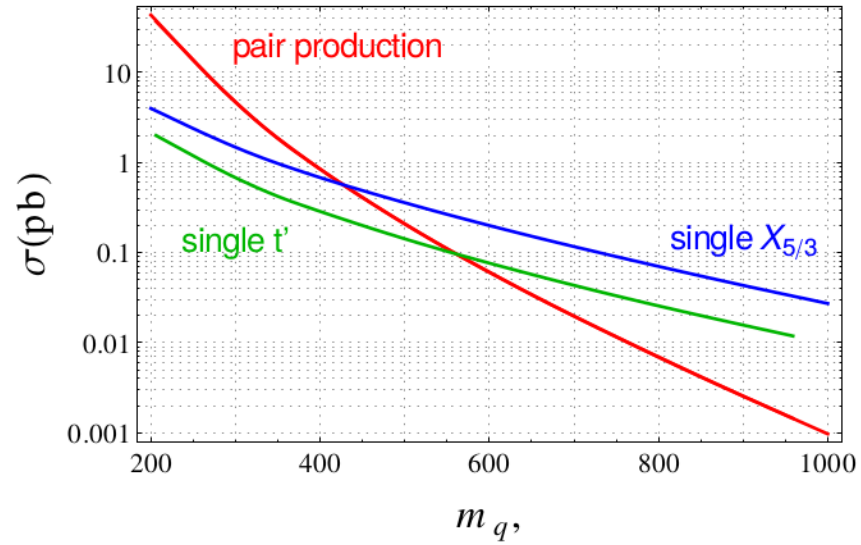
VLQ

- VLQ mix with SM quarks
- ... and interact through Yukawa interactions
- Various incarnation possible:

| | SM | Singlets | Doublets | Triplets |
|-----------------|--|---|--|---|
| | $\begin{pmatrix} u \\ d \end{pmatrix} \begin{pmatrix} c \\ s \end{pmatrix} \begin{pmatrix} t \\ b \end{pmatrix}$ | $\begin{pmatrix} t' \\ b' \end{pmatrix}$ | $\begin{pmatrix} X \\ t' \end{pmatrix} \begin{pmatrix} t' \\ b' \end{pmatrix} \begin{pmatrix} b' \\ Y \end{pmatrix}$ | $\begin{pmatrix} X \\ t' \\ b' \end{pmatrix} \begin{pmatrix} t' \\ b' \\ Y \end{pmatrix}$ |
| $SU(2)_L$ | 2 and 1 | 1 | 2 | 3 |
| $U(1)_Y$ | $q_L = 1/6$ $u_R = 2/3$ $d_R = -1/3$ | $2/3$ $-1/3$ | $7/6$ $1/6$ $-5/6$ | $2/3$ $-1/3$ |
| \mathcal{L}_Y | $-\frac{y_u^i v}{\sqrt{2}} \bar{u}_L^i u_R^i$ $-\frac{y_d^i v}{\sqrt{2}} \bar{d}_L^i V_{CKM}^{i,j} d_R^j$ | $-\frac{\lambda_u^i v}{\sqrt{2}} \bar{u}_L^i U_R$ $-\frac{\lambda_d^i v}{\sqrt{2}} \bar{d}_L^i D_R$ | $-\frac{\lambda_u^i v}{\sqrt{2}} U_L u_R^i$ $-\frac{\lambda_d^i v}{\sqrt{2}} D_L d_R^i$ | $-\frac{\lambda_i v}{\sqrt{2}} \bar{u}_L^i U_R$ $-\lambda_i v \bar{d}_L^i D_R$ |

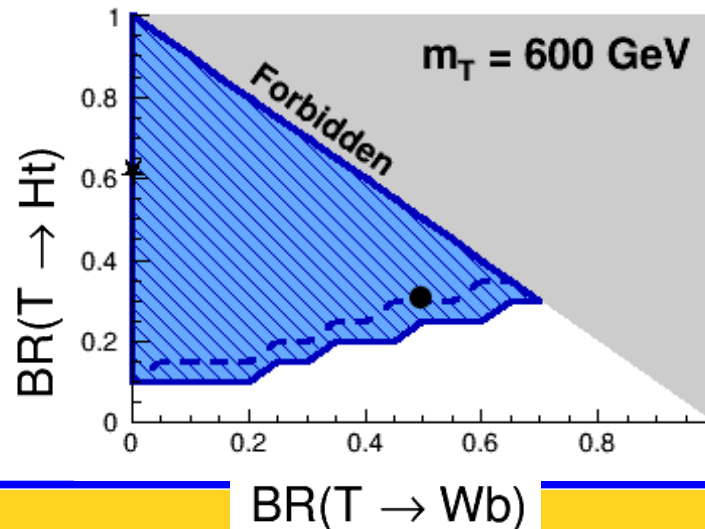
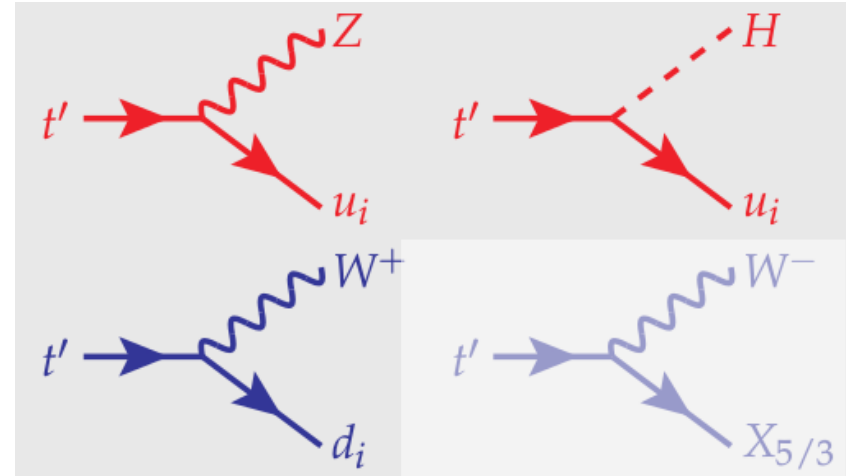
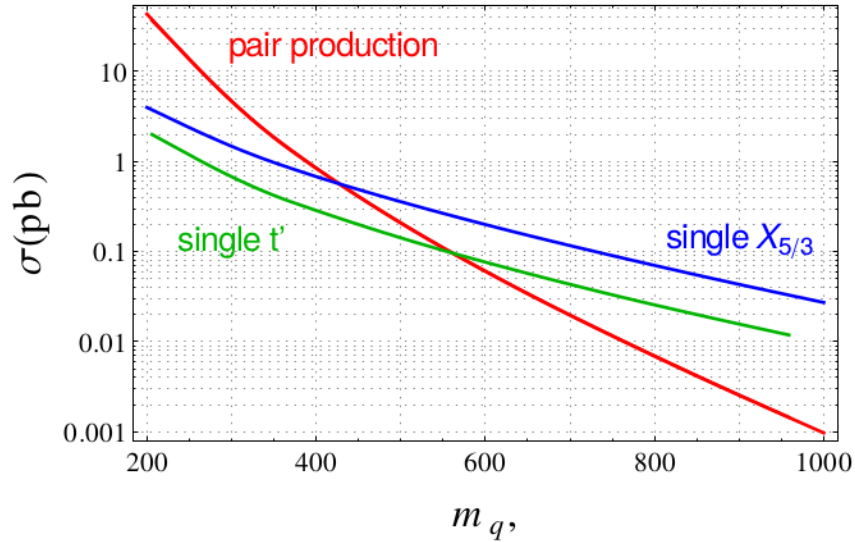
Searching for VLQ

◆ Production / decay



Searching for VLQ

Production / decay



ATLAS Preliminary
Status: ICHEP 2014

$\sqrt{s} = 8$ TeV, $\int L dt = 14.3$ & 20.3 fb $^{-1}$

--- 95% CL exp. excl. — 95% CL obs. excl.

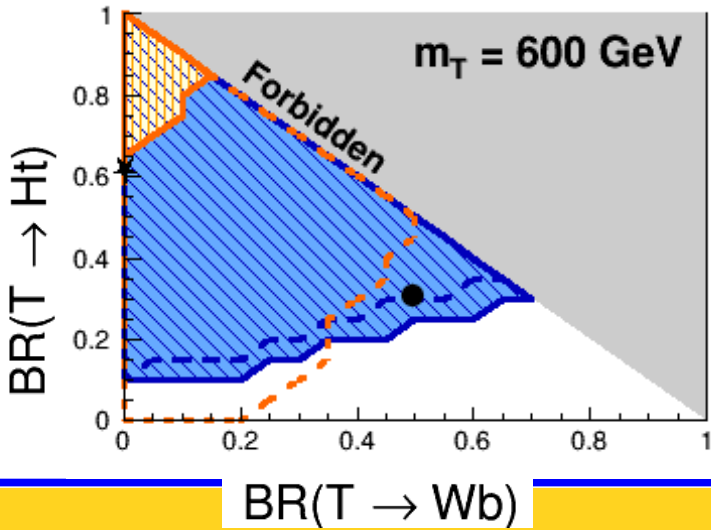
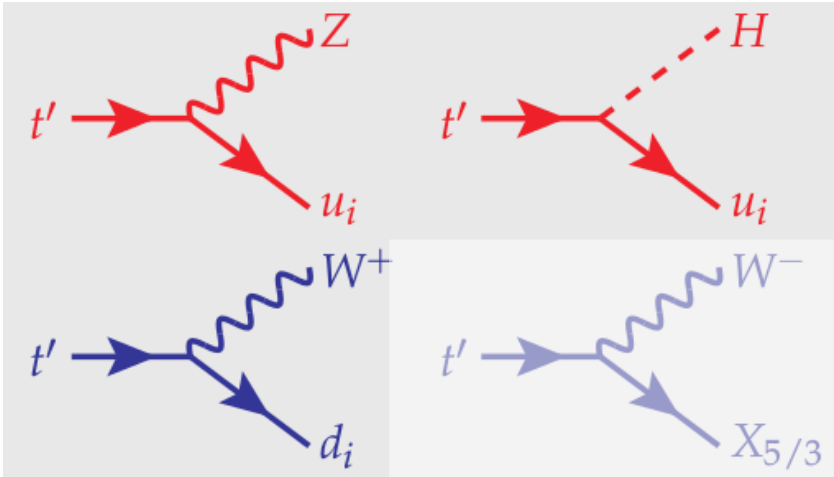
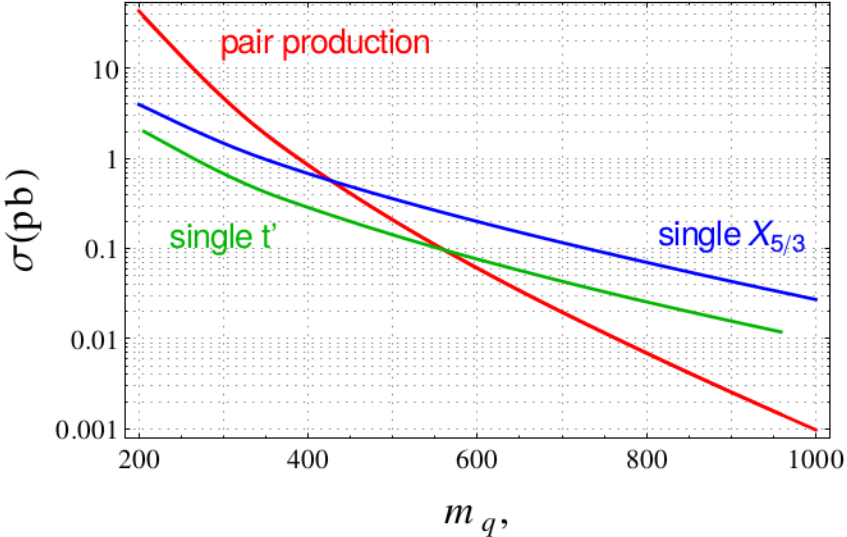


Ht+X

[ATLAS-CONF-2013-018]

Searching for VLQ

Production / decay



ATLAS Preliminary

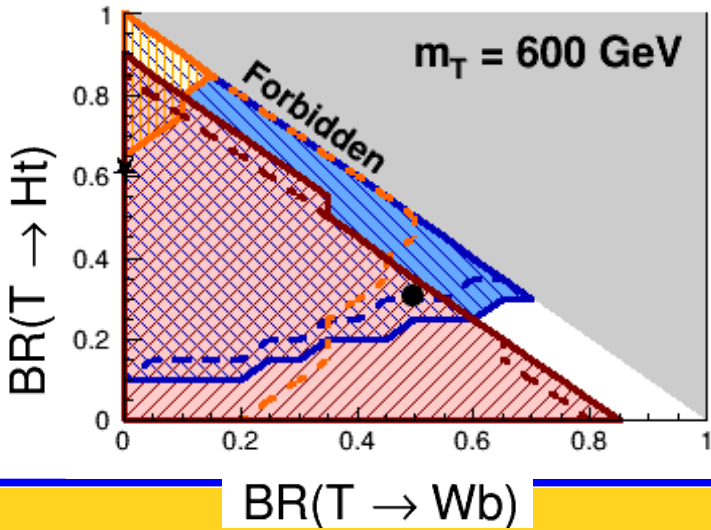
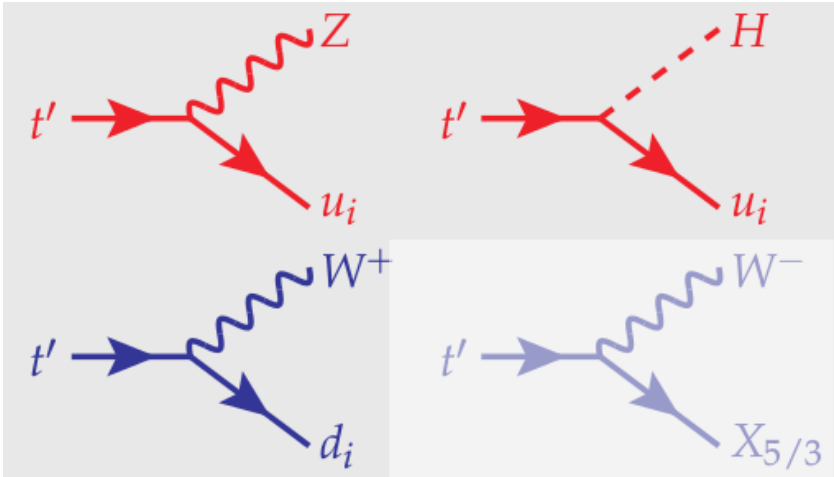
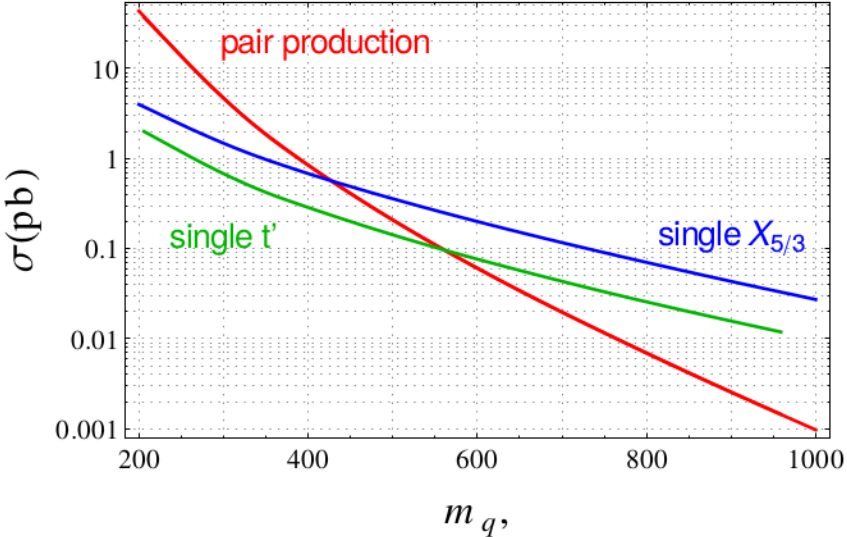
Status: ICHEP 2014

$\sqrt{s} = 8$ TeV, $\int L dt = 14.3$ & 20.3 fb $^{-1}$

- 95% CL exp. excl. — 95% CL obs. excl.
- Ht+X [ATLAS-CONF-2013-018]
- Same-Sign II [ATLAS-CONF-2013-051]

Searching for VLQ

Production / decay



ATLAS Preliminary

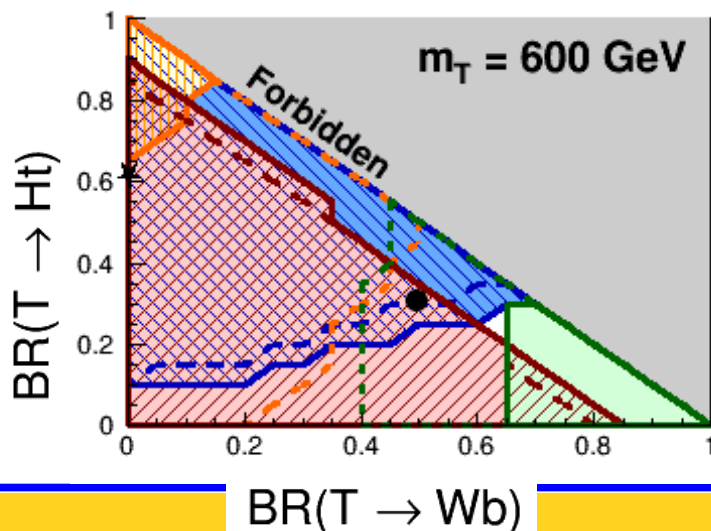
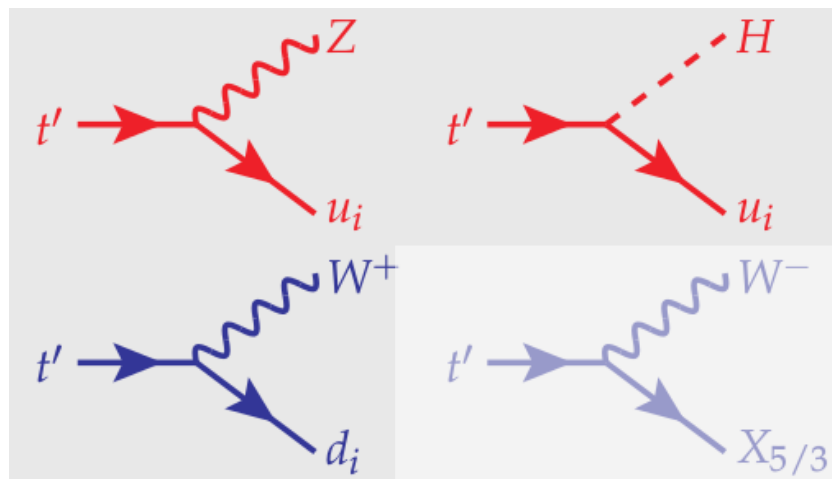
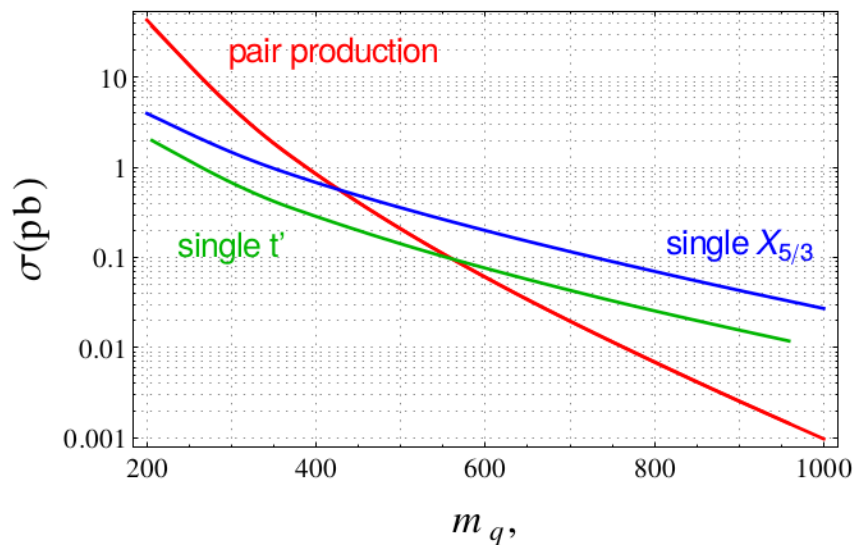
Status: ICHEP 2014

$\sqrt{s} = 8$ TeV, $\int L dt = 14.3$ & 20.3 fb $^{-1}$

- 95% CL exp. excl.
- 95% CL obs. excl.
- $Ht+X$ [ATLAS-CONF-2013-018]
- Same-Sign II [ATLAS-CONF-2013-051]
- $Zb/t+X$ [ATLAS-CONF-2014-036]

Searching for VLQ

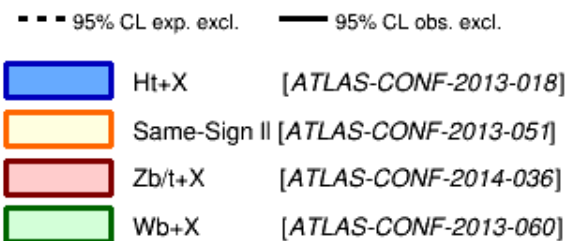
Production / decay



ATLAS Preliminary

Status: ICHEP 2014

$\sqrt{s} = 8$ TeV, $\int L dt = 14.3$ & 20.3 fb $^{-1}$



Conclusion

- ◆ Lots of models on the market
... trying to address
 - dark matter
 - naturalness
- ◆ Of course, I could not present
 - many other models
 - many other problems (ν mass, ...)

QUE CHOISIR Hors-Série
Beyond SM

N° 125 - JANVIER 2012
www.quechoisir.org
09697 - Abonnements - Paris - France

> Votre Higgs à l'abri!
RIEN NE VA PLUS...
LES MEILLEURS PLACEMENTS EN TEMPS DE CRISE p.4

Interdit bancaire
Comment s'en sortir p.46

COMPARATIF p.34
Garanties accidents de la vie contre individuelles accident

> Bien assurer vos objets d'art et de valeur p.49

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> Chercher un emploi, combien ça coûte p.54

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L 12234 - 125 - P 4,00 € - 160

Backup

Top-quark mass

- ◆ Important parameter of the SM
 - (meta-)stability of the Higgs field ?

