# Introduction to Perturbative QCD

partons, factorization, resummation, and all that

#### Marco Stratmann

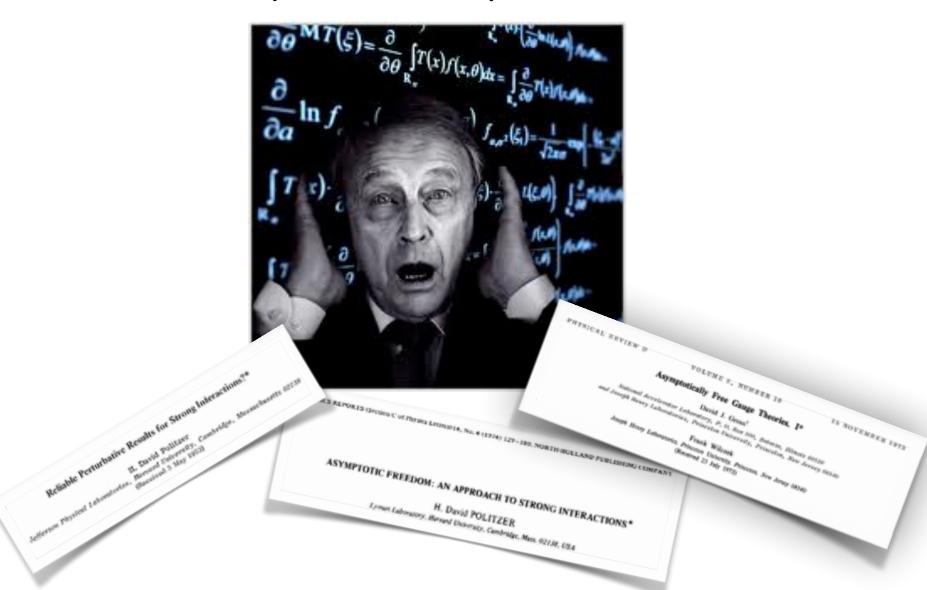


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June 30th/ July 1st, 2014

disclaimer:

pQCD is about 40 years old - impossible to review in 6 hrs



# topics & questions to be addressed

we will mainly concentrate on a few basics and their consequences for phenomenology

- What are the foundations of QCD? keywords: color; SU(3) gauge group; local gauge invariance; Feynman rules
- What are the general features of QCD? keywords: asymptotic freedom; infrared safety; origin of "singularities"
- How to relate QCD to experiment? keywords: partons; factorization; renormalization group eqs. / evolution
- How reliable is a theoretical QCD calculation? keywords: scale dependence; NLO; small-x; all-order resummations
- What is the status of some non-perturbative inputs keywords: global QCD analysis

throughout this will be blended with discussions of some recent results and advanced topics relevant for LHC, RHIC, HERA, COMPASS, EIC, ...

# bibliography – a personal selection

#### textbooks:

- the "pink book" on QCD and Collider Physics by R.K. Ellis, W.J. Stirling, and B.R. Webber
- R.D. Field, Applications of pQCD detailed examples
- Y.V. Kovchegov, E. Levin, QCD at High Energy focus on small x physics
- J. Collins, Foundations of pQCD focus on formal aspects of evolution

#### lecture notes & write-ups:

- D. Soper, Basics of QCD Perturbation Theory, hep-ph/9702203
- Collins, Soper, Sterman, Factorization of Hard Processes in QCD, hep-ph/0409313
- G. Salam, Elements of QCD for Hadron Colliders, arXiv:1011.5131
- Particle Data Group, Review of Particle Physics, pdg.lbl.gov

#### talks & lectures on the web: e.g. by D. Soper; G. Salam; G. Zanderighi; J. Campbell; G. Sterman; ...

- annual CTEQ summer school, tons of material on www.cteq.org
- annual CERN/FNAL Hadron Collider Physics School hcpss.web.cern.ch/hcpss



# tentative outline of the lectures

# <u>Part 1</u>: the foundations

SU(3); color algebra; gauge invariance; QCD Lagrangian; Feynman rules

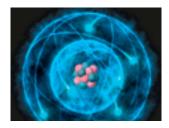
- <u>Part 2</u>: the QCD toolbox asymptotic freedom; infrared safety; the QCD final-state; jets; factorization
- <u>Part 3</u>: inward bound: "femto spectroscopy" QCD initial-state; DIS process; partons; factorization; renormalization group; scales; hadron-hadron collisions

# <u>Part 4</u>: applications:

global analysis of PDFs; scales and theoretical uncertainties; all-order resummations; ...









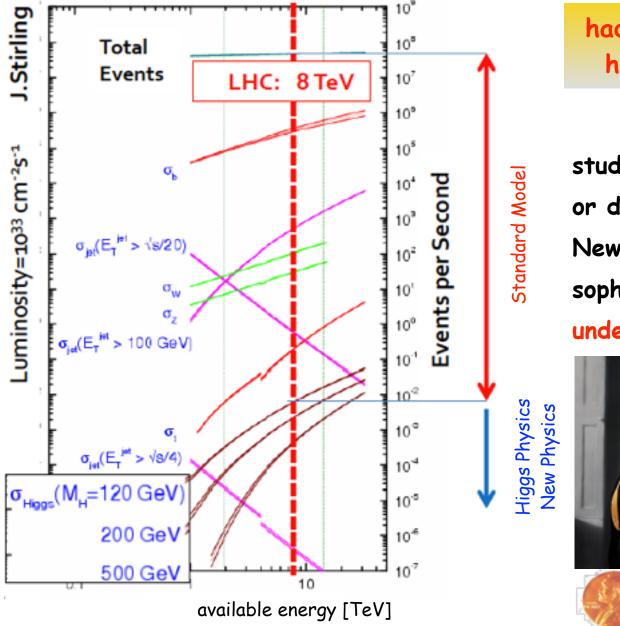


# Part I

shamelessly

excellent och & J. compoent the QCD fundamentals all about color the concept of gauge invariance

#### **QCD** – why do we still care (or perhaps more than ever)



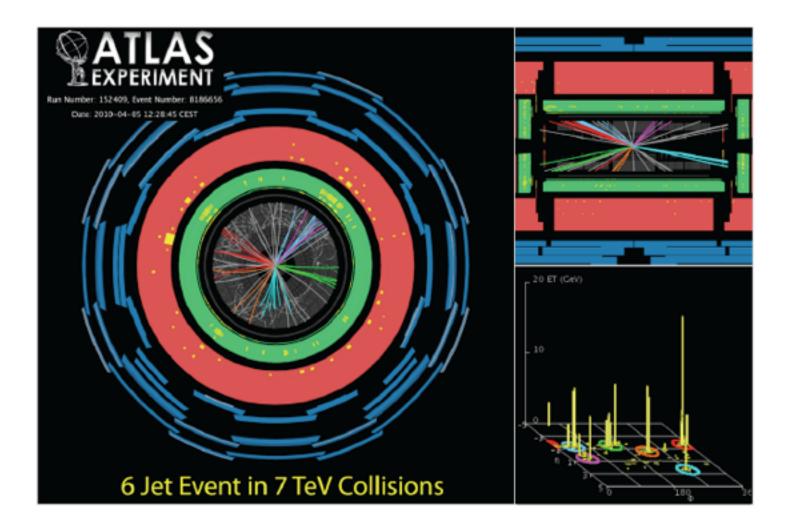
hadron colliders inevitably have to deal with QCD

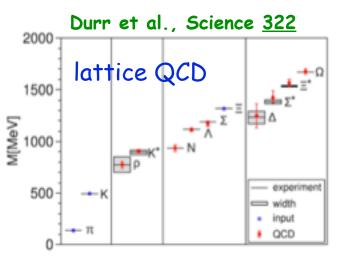
studying the Higgs boson or discovering (perhaps) some New Physics requires a sophisticated quantitative understanding of QCD

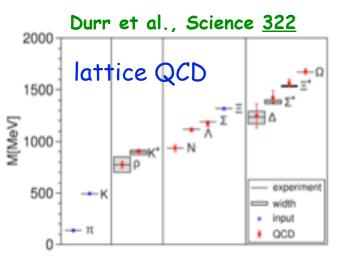


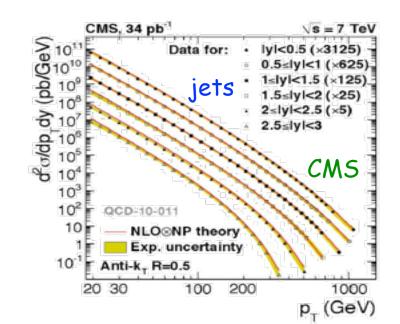
achieving that can be quite a challenge ...

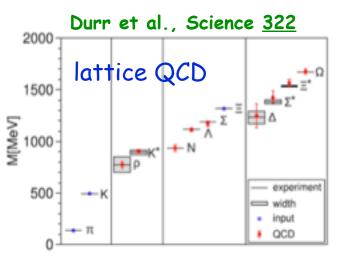
 $\mathcal{L}_{\rm QCD} = -\frac{1}{4} F^A_{\mu\nu} F^{\mu\nu}_A + \sum \bar{q}_i (i D - m)_{ij} q_j$ flavors

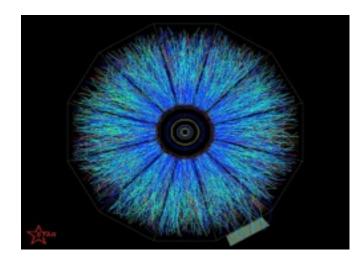




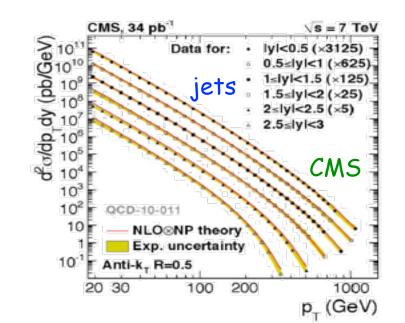




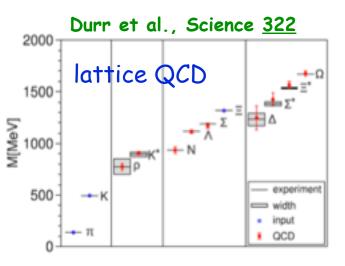


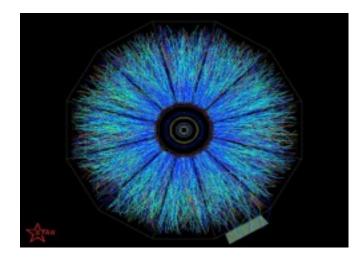


AuAu collision at STAR

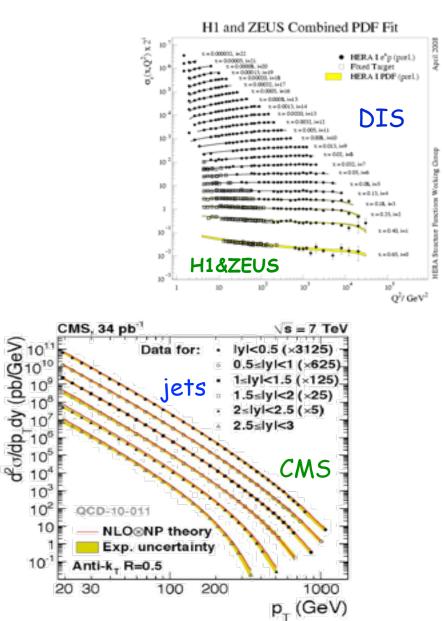


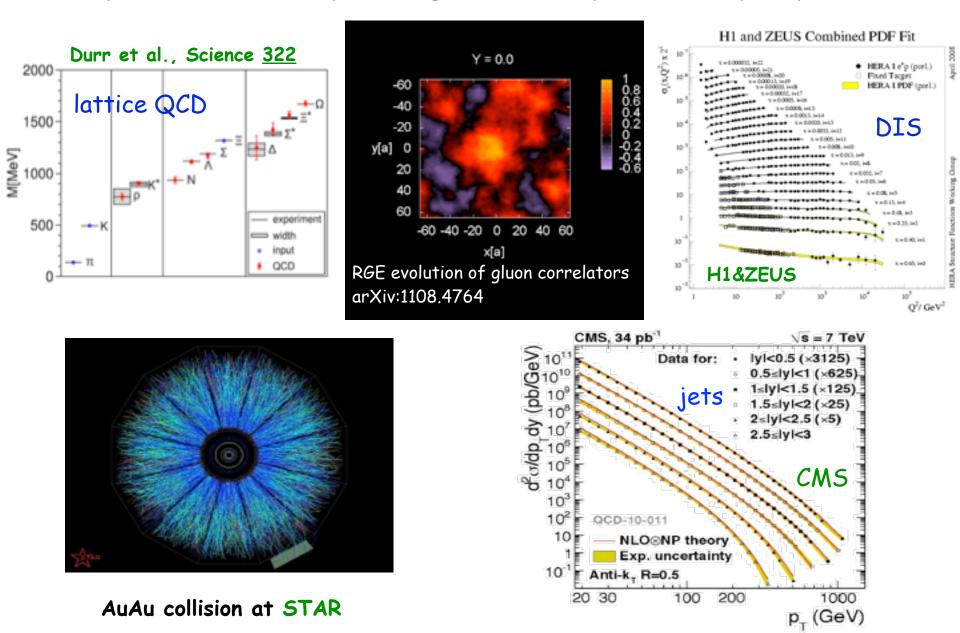
# QCD – the theory of strong interactions a simple QED-like theory, leading to extremely rich & complex phenomena

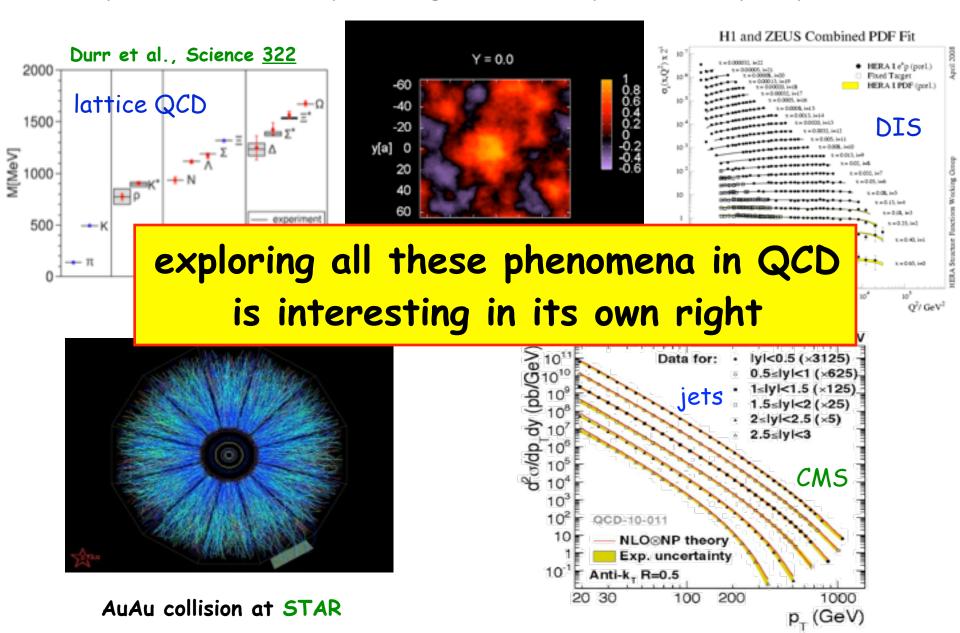




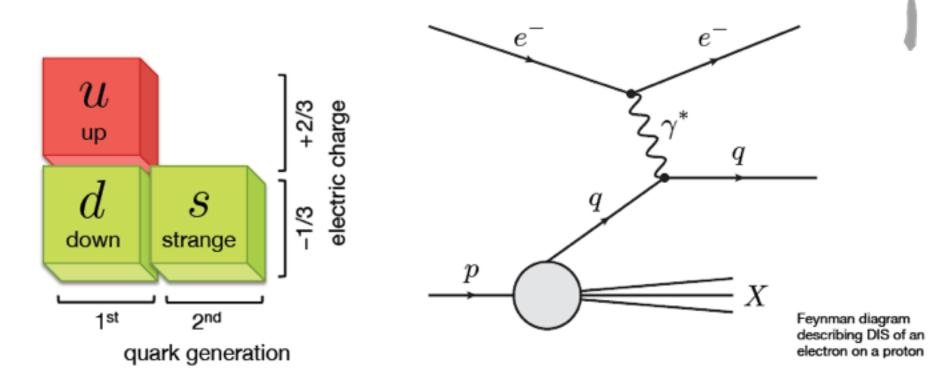
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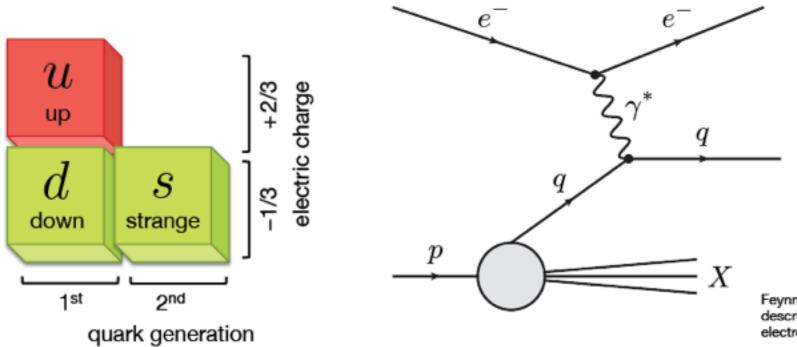


## **QCD** matter sector: Three Quarks for Muster Mark



existence of light quarks validated in deep-inelastic scattering (DIS) experiments carried out at SLAC in 1968

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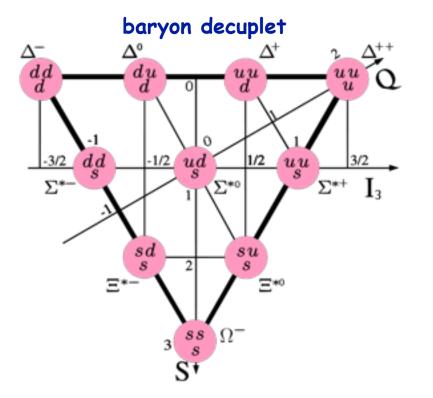
Feynman diagram describing DIS of an electron on a proton

existence of light quarks validated in deep-inelastic scattering (DIS) experiments carried out at SLAC in 1968 strange quarks necessary component in **quark model** to classify the observed slew of mesons/baryons Gell-Mann, Zweig (1964) based on "**Eightfold Way**" (= SU(3)<sub>flavor</sub>) Gell-Mann; Ne'eman (1961)



## quark model: mesons and baryons

categorizes mesons (baryons) in terms of two (three) constituent quarks in **SU(3)**<sub>flavor</sub> **multiplets** = octets and decuplets

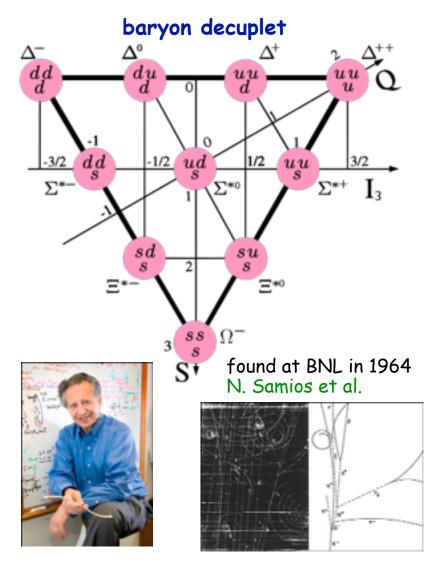


spectrum fully classified by assuming:

- quarks have spin  $\frac{1}{2}$
- quarks have fractional charges (but combine into hadrons with integer charges)

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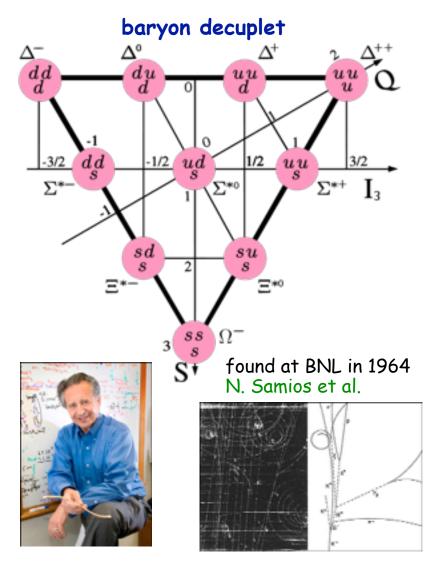
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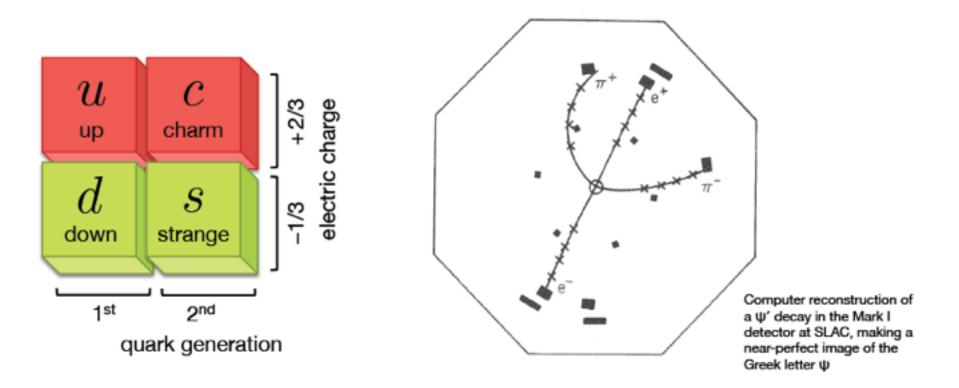
**big success:** prediction of  $\Omega^-$  (sss)

also, first evidence of color

- Δ<sup>++</sup> wave function |uuu> not anti-sym (violates Pauli principle)
- remedy: color quantum number but hadrons remain colorless/color singlets



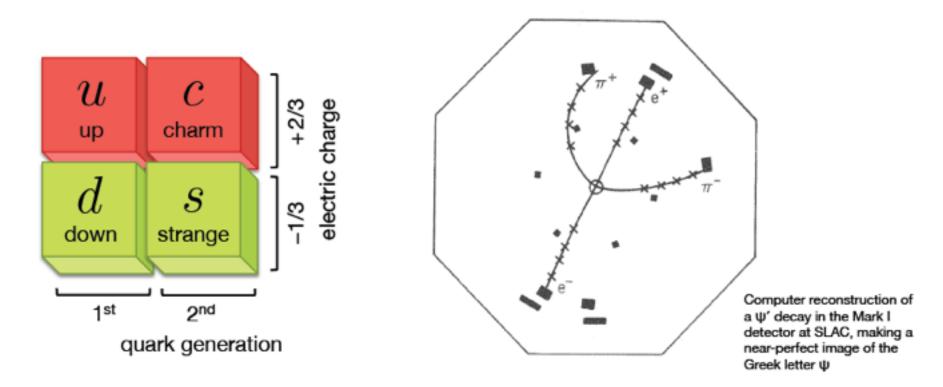
#### **QCD** matter sector: charm



predicted on strong theoretical grounds (suppression of FCNC) "GIM mechanism" in 1970 Glashow, Iliopolus, Maiani



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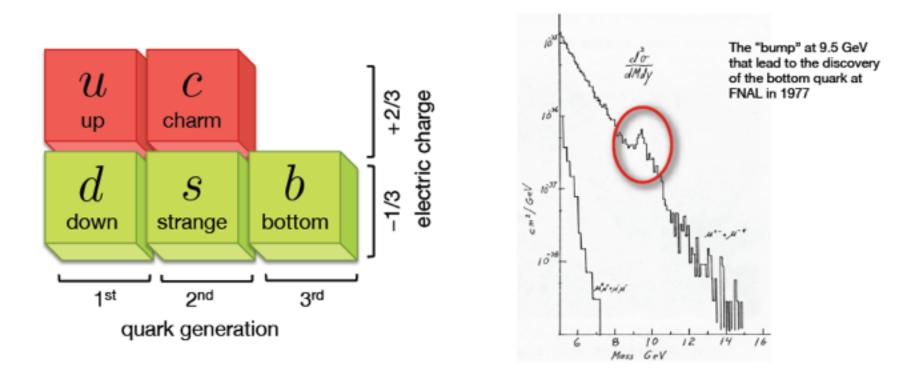


predicted on strong theoretical grounds (suppression of FCNC) "GIM mechanism" in 1970 Glashow, Iliopolus, Maiani

observed during "November revolution" in 1974 both at SLAC (Richter et al.) and BNL (Ting et al.) discovered meson became known as  $J/\Psi$ ; Nobel Prize in 1976



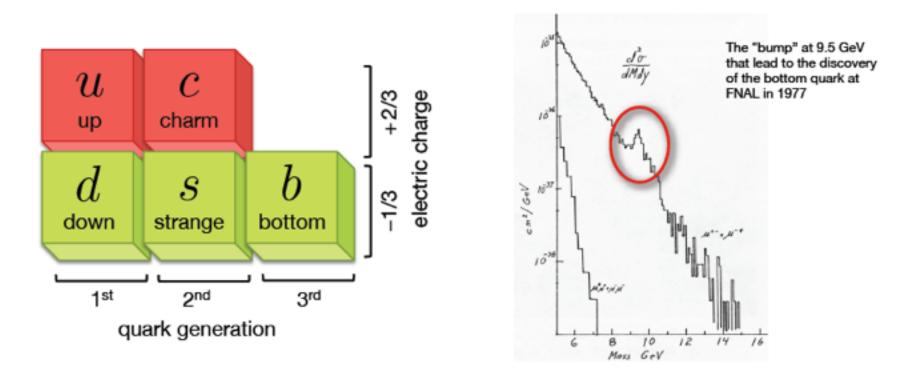
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discovered in 1977 at FNAL (Y meson or "bottomium") Ledermann et al.

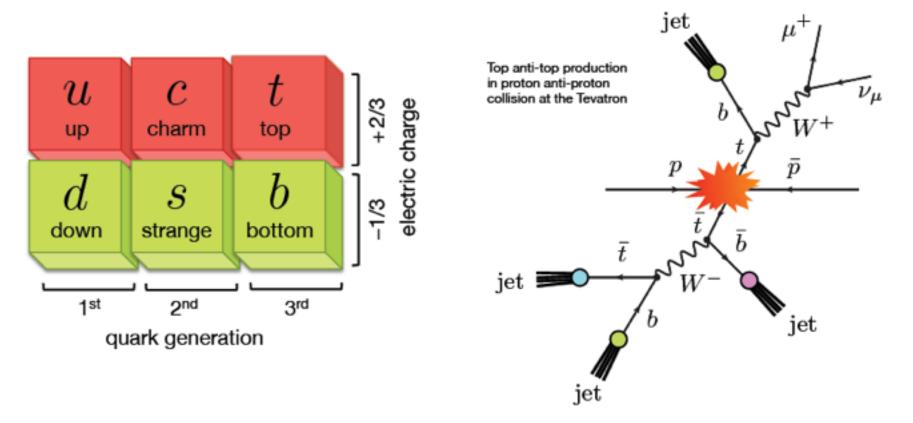
L.L. coined also the term "God particle"



Nobel Prize in 1988 for muon neutrino



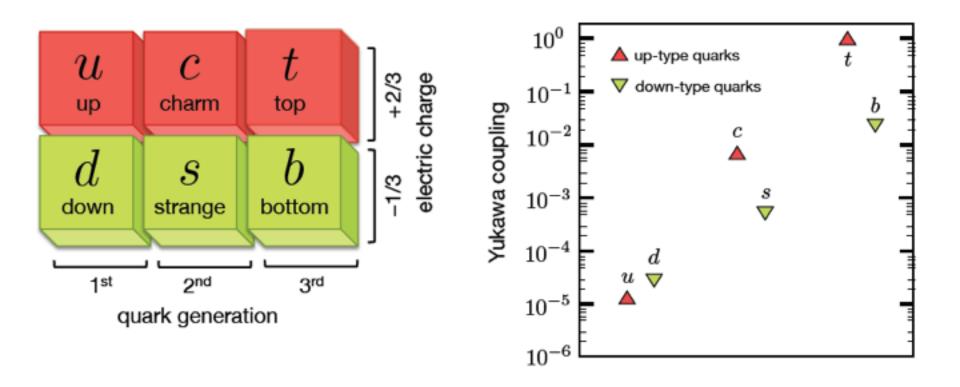
#### **QCD** matter sector: top



by around 1994 electroweak precision fits point towards mass in range 145-185 GeV (vector boson mass and couplings are sensitive to top mass)

eventually discovered in 1995 by CDF and DØ at FNAL (mass nowadays know to about 1 GeV)

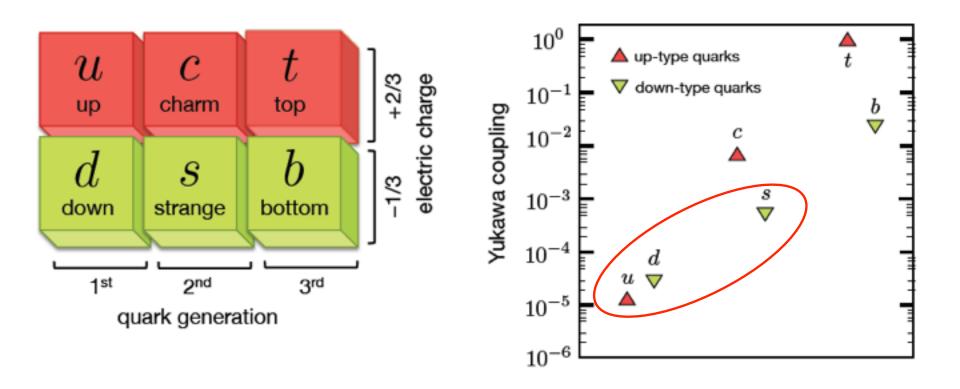
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masses of six quarks range from O(MeV) to about 175 GeV

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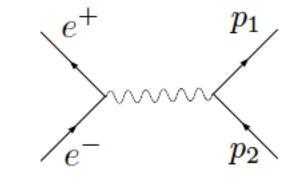


- masses of six quarks range from O(MeV) to about 175 GeV why the masses are split by almost six orders of magnitude remains a big mystery
- masses of u, d, s quarks are lighter than 1 GeV (proton mass) in the limit of vanishing u,d,s masses there is an exact SU(3)<sub>flavor</sub> symmetry

#### • color can be probed directly in e<sup>+</sup>e<sup>-</sup> collisions

#### idea:

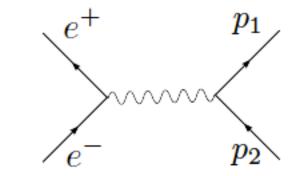
production of fermion pairs (leptons or quarks) through a virtual photon sensitive to electric charge and number of degrees of freedom



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hence, investigate quarks through "R ratio"

$$R \equiv \frac{e^+e^- \to \text{hadrons}}{e^+e^- \to \mu^+\mu^-} \propto N_c \sum_f Q_f^2$$

assumed number of colors of quark

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production of fermion pairs (leptons or quarks) through a virtual photon sensitive to electric charge and number of degrees of freedom  $e^+$   $p_1$  $e^ p_2$ 

> electric charge of quark [in units of e]

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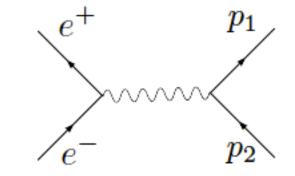
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- in LO described by process  $e^+e^- \to q\bar{q}$ 

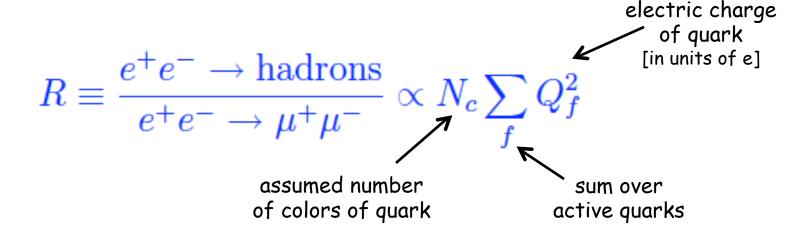
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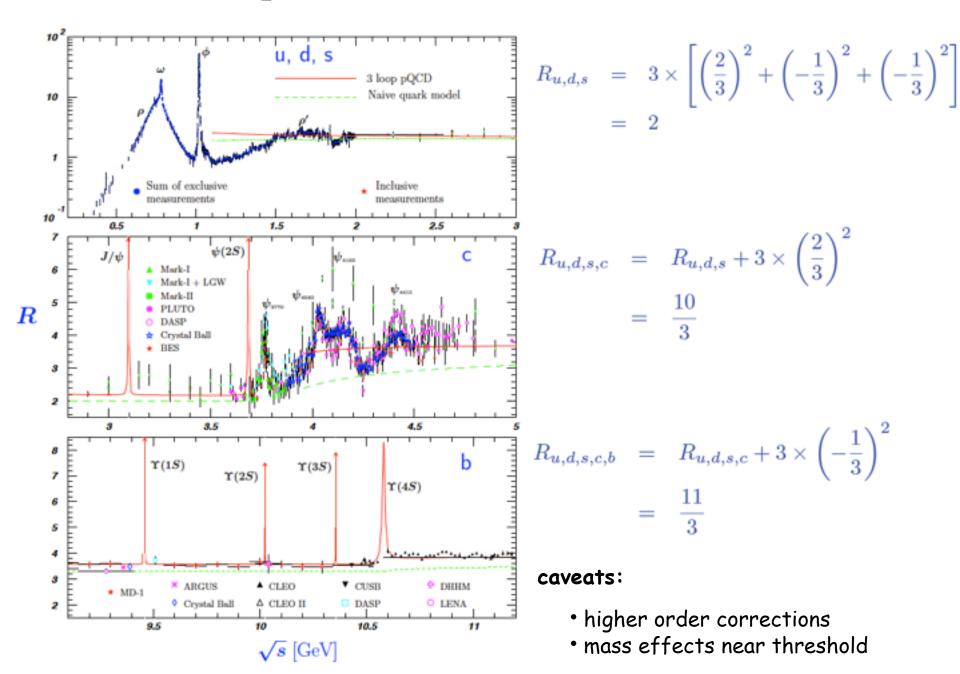


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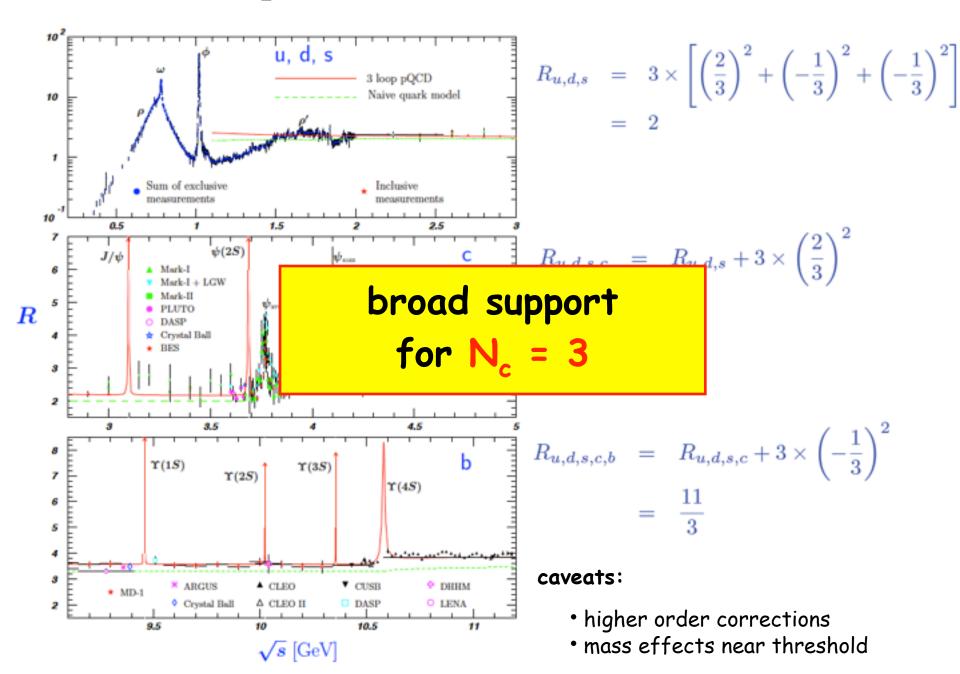


- in LO described by process  $e^+e^- \to q\bar{q}$
- each active quark is produced in one out of  $N_c$  colors above kinematic threshold

#### experimental results for R ratio



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# **QCD** color interactions heuristically

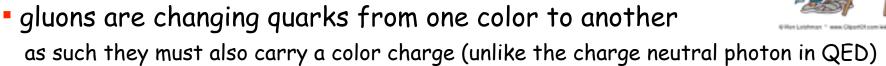
- QCD color quantum number is mediated by the gluon analogous to the photon in QED
- gluons are changing quarks from one color to another
   as such they must also carry a color charge (unlike the charge neutral photon in QED)





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 color charge of each gluon represented by a 3x3 matrix in color space conventional choice: express t<sup>a</sup> (a=1...8) in terms of Gell-Mann matrices

typical color interaction between quarks and gluons

$$\begin{array}{c} (\mathbf{1}, \mathbf{0}, 0) & \left( \begin{array}{cc} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{array} \right) & \left( \begin{array}{c} 0 \\ \mathbf{1} \\ 0 \end{array} \right) \\ \bar{\psi}_i & t_{ij}^1 & \psi_j \end{array}$$

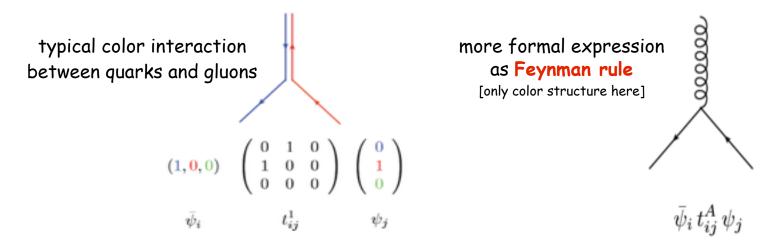


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here: local SU(3) rotations in color space

spin-½ quark fields come as colors triplets (fundamental representation)

$$\Psi = \begin{pmatrix} \bullet \\ \bullet \\ \bullet \end{pmatrix} \xrightarrow{\bullet} \Psi' = \begin{pmatrix} \bullet \\ \bullet \\ \bullet \end{pmatrix}$$

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local SU(3) invariance dictates:

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- 8 massless spin-1 gluons (adjoint representation)
- all interactions between quarks and gluons (covariant derivative)

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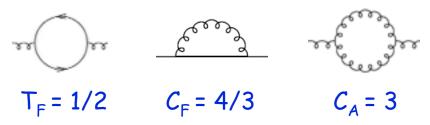
local SU(3) invariance dictates:

#### non-Abelian group structure:

invariants ("color factors") :

$$= \begin{pmatrix} \bullet \\ \bullet \\ \bullet \end{pmatrix} \xrightarrow{\bullet} \Psi' = \begin{pmatrix} \bullet \\ \bullet \\ \bullet \end{pmatrix}$$

- 8 massless spin-1 gluons (adjoint representation)
- all interactions between quarks and gluons (covariant derivative)
- Lie algebra:  $[t_a, t_b] = i f_{abc} t_c$





- choose special unitary group SU(3) as the gauge group for QCD
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  - N x N generic complex matrix has  $N^2$  complex (= 2  $N^2$  real) values



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 $\det(U) = 1$ 

unitary provides N<sup>2</sup> conditions unit determinant ("special"): 1 condition -> SU(N) group has N<sup>2</sup> - 1 generators (-> QCD has 8 gluons)



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• generators are traceless hermitian N x N matrices

 $\mathbf{U} = \mathbf{e}^{\mathbf{i}\theta_{\mathbf{a}}(\mathbf{x})\mathbf{t}^{\mathbf{a}}}$ 



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 $\mathbf{U} = \mathbf{e}^{\mathbf{i}\theta_{\mathbf{a}}(\mathbf{x})\mathbf{t}^{\mathbf{a}}}$ arbitrary parameter element of the group "rotations in color space"
generator
arbitrary
element of the group



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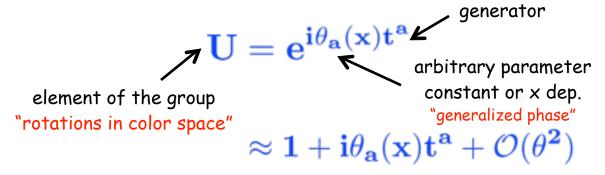
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properties can be studied from infinitesimal transformations



- choose special unitary group SU(3) as the gauge group for QCD
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• why SU(3)?

 $UU^{\dagger} = U^{\dagger}U = 1_{N \times N}$ 

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 $\mathbf{U} = \mathbf{e}^{\mathbf{i}\theta_{\mathbf{a}}(\mathbf{x})\mathbf{t}^{\mathbf{a}}}$   $\mathbf{u} = \mathbf{e}^{\mathbf{i}\theta_{\mathbf{a}}(\mathbf{x})\mathbf{t}^{\mathbf{a}}}$ 

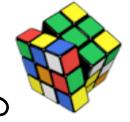
 $\mathbf{a}=\mathbf{1},\mathbf{2},...,\mathbf{N^2}-\mathbf{1}$ 

properties can be studied from infinitesimal transformations

quarks and anti-quarks are different [rules out real SO(3)]

 $\det(U) = 1$ 

• only compact simple Lie group with complex triplet representation



#### color algebra: Fierz identity, Casimir operators

powerful Fierz identity



Entity 
$$\sum_{\mathbf{a}} (\mathbf{t}^{\mathbf{a}})_{\mathbf{ij}} (\mathbf{t}^{\mathbf{a}})_{\mathbf{kl}} = \frac{1}{2} \left( \delta_{\mathbf{il}} \delta_{\mathbf{jk}} - \frac{1}{\mathbf{N}} \delta_{\mathbf{ij}} \delta_{\mathbf{kl}} \right)$$
$$\int_{k}^{i} \frac{\mathbf{a}}{\mathbf{k}} \left[ \frac{\mathbf{a}}{\mathbf{k}} \right]_{k} = \frac{1}{2} \left( \int_{k}^{i} \int_{k}^{i} \frac{1}{N_{c}} \int_{k}^{i} \frac{1}{N_{c}} \int_{k}^{i} \frac{1}{\mathbf{k}} \int_{k}^{i} \frac{1}$$

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• powerful Fierz identity 
$$\sum_{\mathbf{a}} (\mathbf{t}^{\mathbf{a}})_{\mathbf{ij}} (\mathbf{t}^{\mathbf{a}})_{\mathbf{kl}} = \frac{1}{2} \left( \delta_{\mathbf{il}} \delta_{\mathbf{jk}} - \frac{1}{N} \delta_{\mathbf{ij}} \delta_{\mathbf{kl}} \right)$$

• N-1 Casimir operators (commute with all generators; proportional to identity)

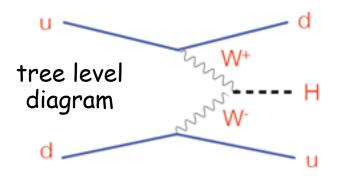
• fundamental representation  

$$i \quad k \quad j = C_F \longrightarrow C_F \longrightarrow C_F = \frac{N^2 - 1}{2N}$$

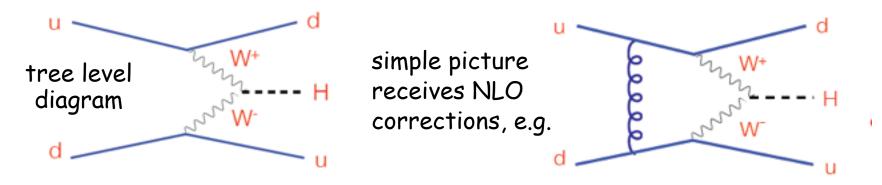
• adjoint representation (defined by  $if_{abc} = 2Tr([t^a, t^b]t^c) \rightarrow 8$  (8x8) matrices)

$$\sum_{\mathbf{a}} \mathbf{f}_{\mathbf{a}\mathbf{c}\mathbf{d}} \mathbf{f}_{\mathbf{b}\mathbf{c}\mathbf{d}} = \mathbf{C}_{\mathbf{A}} \sum_{\mathbf{c}\mathbf{d}} \mathbf{f}_{\mathbf{a}\mathbf{c}\mathbf{d}} \mathbf{f}_{\mathbf{b}\mathbf{c}\mathbf{d}} = \mathbf{C}_{\mathbf{A}} \delta_{\mathbf{a}\mathbf{b}} \qquad \mathbf{C}_{\mathbf{A}} = \mathbf{N}$$

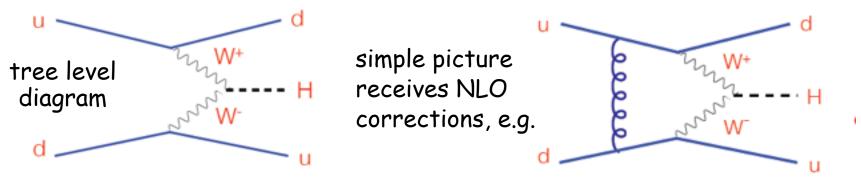
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vanishes when interfered with LO diagram

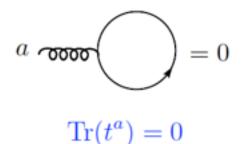
WHY ?

vector boson fusion is an important Higgs search channel at the LHC



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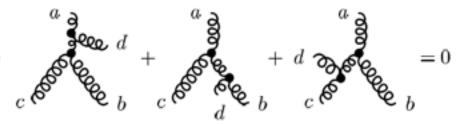


vector boson fusion is an important Higgs search channel at the LHC

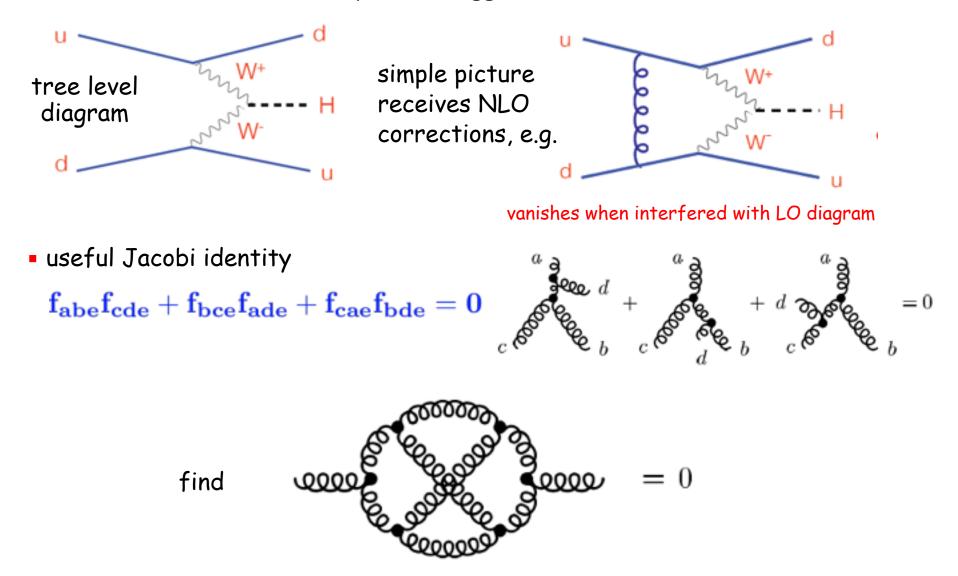


useful Jacobi identity

 $\mathbf{f_{abe}f_{cde}+f_{bce}f_{ade}+f_{cae}f_{bde}=0}$ 



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#### color at work: leading color approximation

example from J. Campbell's lectures

 to simplify large scale QCD calculations, one often works in the leading color approximation

what is it all about?

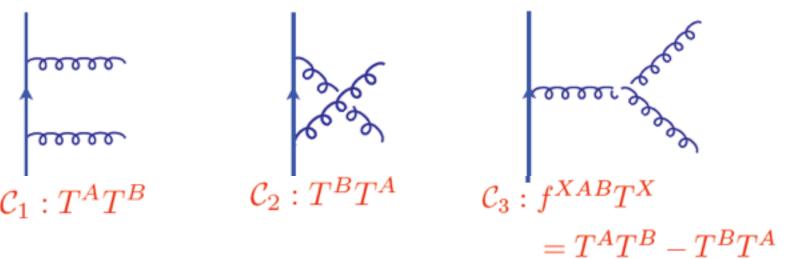
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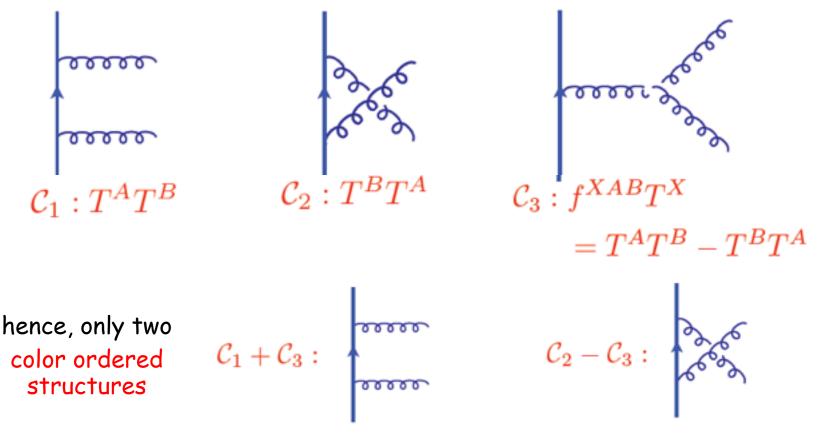
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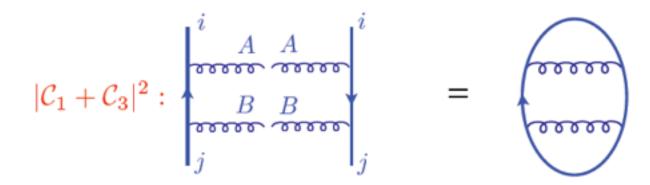
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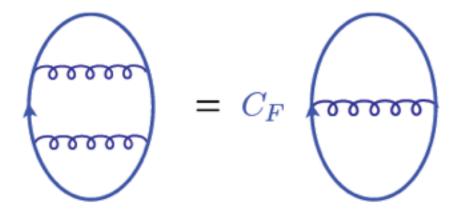


need to square amplitudes to get cross section, e.g.:

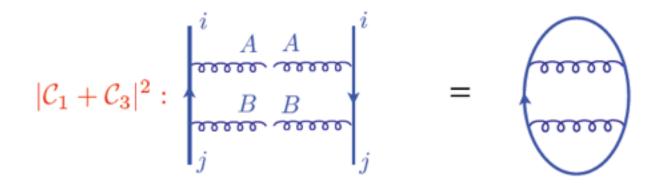
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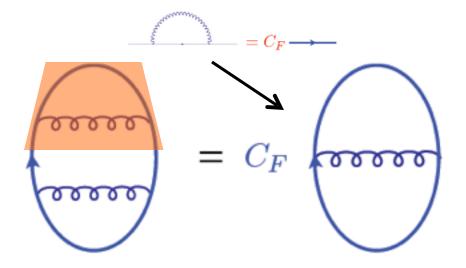
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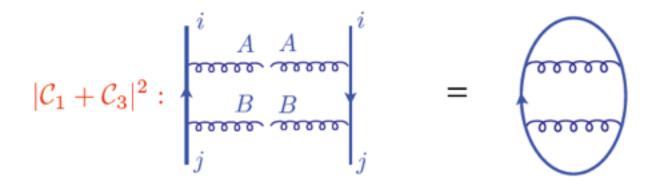
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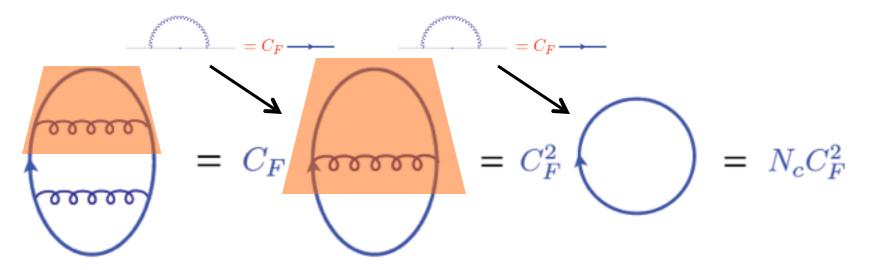
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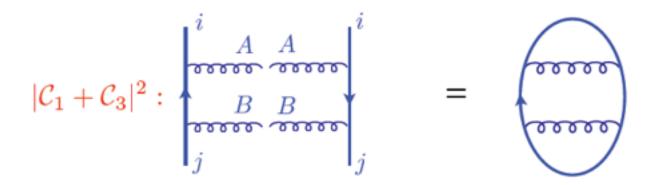
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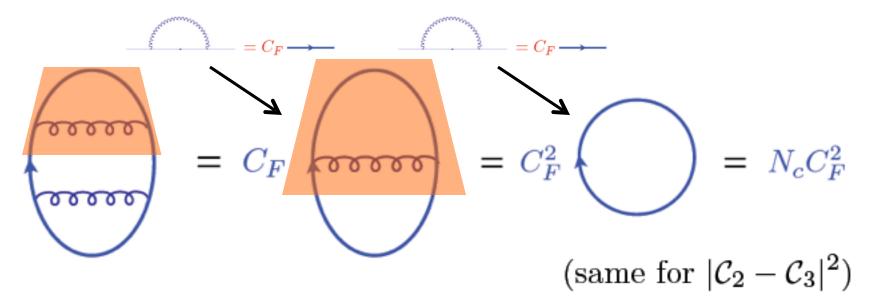
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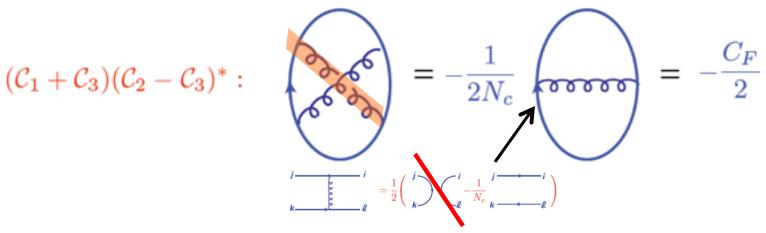
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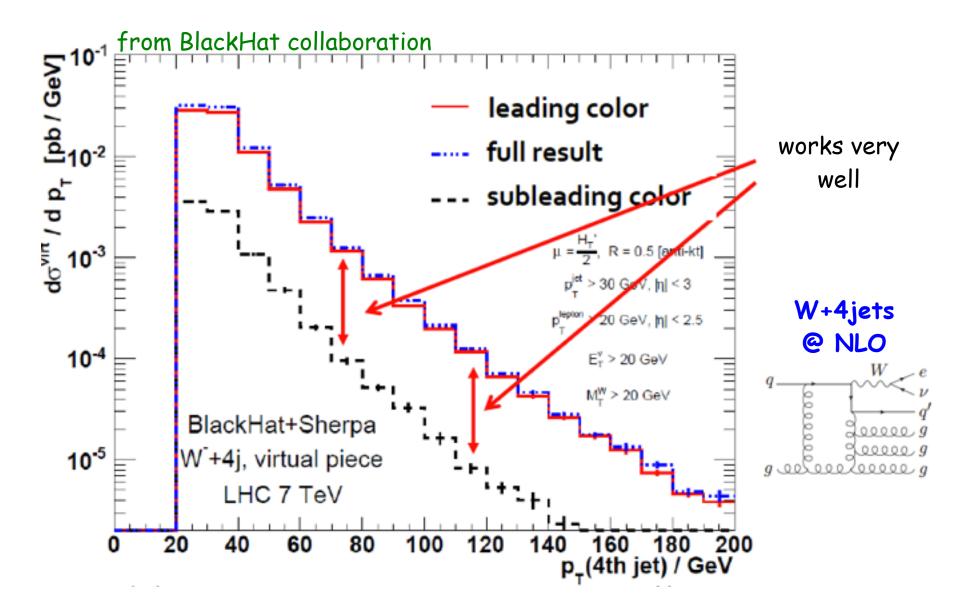
does not contribute: Tr (t<sup>a</sup>) = 0

• combine results for [after some reshuffling, use  $N_c C_F^2 = (N_c^2 C_F - C_F)/2$ ]

$$\frac{N_c^2 C_F}{2} \left( |\mathcal{C}_1 + \mathcal{C}_3|^2 + |\mathcal{C}_2 - \mathcal{C}_3|^2 - \frac{1}{N_c^2} |\mathcal{C}_1 + \mathcal{C}_2|^2 \right)$$

$$\int |\text{leading color} \\ \text{contribution} \\ \frac{1/N_c^2 \text{ suppressed}}{1/N_c^2 \text{ suppressed}} \\ \frac{1}{N_c^2 \text{ suppressed}$$

#### leading color approximation at work



#### experimental support for SU(3)

#### color factors are not just math

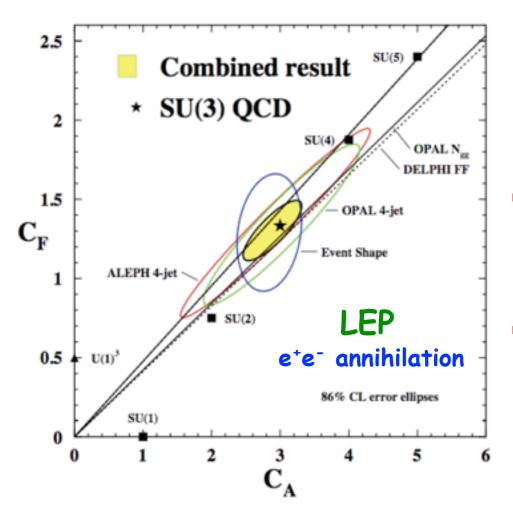
assumed group structure has impact on theoretical predictions

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#### experimental support for SU(3)

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assumed group structure has impact on theoretical predictions





- angular correlations
   between four jets depend
   on  $C_A/C_F$  and  $T_F/C_F$
- sensitivity to non-Abelian three-gluon-vertex
   LO: Ellis, Ross, Terrano

## **QCD** Lagrangian & Feynman rules

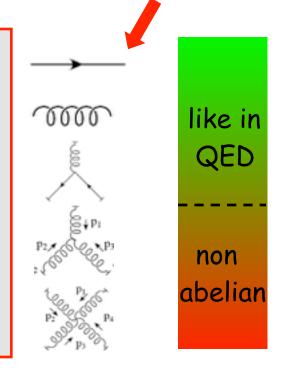
L<sub>QCD</sub> encodes all physics related to strong interactions

for perturbative calculations we simply read off the Feynman rules

 $\mathcal{L}_{QCD} = \bar{\Psi}(i\partial_{\mu}\gamma^{\mu} - m)\Psi$  $- (\partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu})^{2}$  $- g\bar{\Psi}A^{a}_{\mu}T_{a}\gamma^{\mu}\Psi$ 

$$- \frac{1}{2}g(\partial_{\mu}A^{a}_{\nu} - \partial_{\nu}A^{a}_{\mu})f_{abc}A^{\mu b}A^{\nu c}$$

$$- \frac{1}{4}g^2 f_{abc} A^b_\mu A^c_\nu f_{ade} A^{\mu d} A^{\nu e}$$

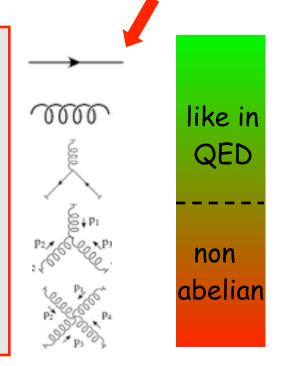


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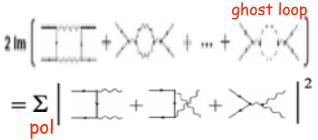
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technical complications due to the gauge-fixing & ghost terms:

**gauge-fixing:** needed to define gluon propagator; breaks gauge-invariance but all physical results are independent of the gauge

<code>ghosts:</code> cancel unphysical degrees of freedom  $\rightarrow$  unitarity



#### recall: gauge invariance in QED

$$\begin{split} \mathcal{L}_{\mathbf{QED}} &= \mathcal{L}_{\mathbf{Dirac}} + \mathcal{L}_{\mathbf{Maxwell}} + \mathcal{L}_{\mathbf{int}} \\ &= \bar{\Psi} (\mathbf{i} \partial \!\!\!/ - \mathbf{m}) \Psi - \frac{1}{4} \mathbf{F}_{\mu\nu} \mathbf{F}^{\mu\nu} - \mathbf{q} \bar{\Psi} \gamma_{\mu} \Psi \mathbf{A}^{\mu} \\ &= \bar{\Psi} (\mathbf{i} \partial \!\!\!/ - \mathbf{m}) \Psi - \frac{1}{4} \mathbf{F}_{\mu\nu} \mathbf{F}^{\mu\nu} \end{split}$$

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invariant under local gauge (phase) transformation

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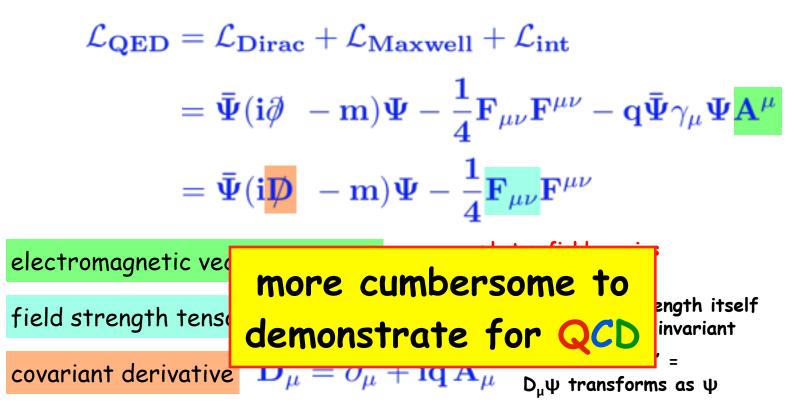
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 Yang and Mills proposed in 1954 that the local "phase rotation" in QED could be generalized to non Abelian groups such as SU(3)



$$\mathcal{L} = -\frac{1}{4} \mathbf{F}_{\mathbf{a}}^{\mu\nu} \mathbf{F}_{\mu\nu}^{\mathbf{a}} + \sum_{\mathbf{f}} \bar{\boldsymbol{\Psi}}_{\mathbf{i}}^{(\mathbf{f})} (\mathbf{i} \boldsymbol{D}_{\mathbf{ij}} - \mathbf{m}_{\mathbf{f}} \delta_{\mathbf{ij}}) \boldsymbol{\Psi}_{\mathbf{j}}^{(\mathbf{f})}$$
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 $a = 1, ..., 8$   $i = 1, 2, 3$ 

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$$\operatorname{QED like but field}_{\text{carries color charge}} \quad \text{non Abelian part gives rise}$$

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$$\operatorname{also in the interaction}_{\text{`covariant derivative''}} (\mathbf{D}_{\mu})_{\mathbf{ij}} = \partial_{\mu} \, \delta_{\mathbf{ij}} + \mathbf{ig_s} \, (\mathbf{t^a})_{\mathbf{ij}} \, \mathbf{A}_{\mu}^{\mathbf{a}}$$

8 generators

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$$\begin{array}{c} \text{QCD interaction is flavor blind} \end{array}$$

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#### more cumbersome to demonstrate

short-hand notation

• redefine quark fields:  $\Psi(\mathbf{x}) \to \Psi'(\mathbf{x}) = e^{i\alpha_{\mathbf{a}}(\mathbf{x})\mathbf{t}^{\mathbf{a}}}\Psi(\mathbf{x}) \equiv \mathbf{U}(\mathbf{x}) \Psi(\mathbf{x})$ likewise  $\bar{\Psi}' = \bar{\Psi}\mathbf{U}^{\dagger}$ 

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• one way is to study explicitly an infinitesimal transformation  $U(x) = 1 + i\alpha_a(x) t^a$ see QCD book by T. Muta for details

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 $\mathbf{D}_{\mu} \Psi \to (\mathbf{D}_{\mu}')_{\mathbf{i}\mathbf{k}} \Psi_{\mathbf{k}}' = \mathbf{U}_{\mathbf{i}\mathbf{j}}(\mathbf{x}) (\mathbf{D}_{\mu})_{\mathbf{j}\mathbf{k}}(\mathbf{x}) \Psi_{\mathbf{k}}(\mathbf{x})$ 

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 sufficient to demonstrate gauge invariance of quark term

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- sufficient to demonstrate gauge invariance of quark term  $\mathcal{L} = -\frac{1}{4} \mathbf{F}_{\mathbf{a}}^{\mu\nu} \mathbf{F}_{\mathbf{a}}^{\mathbf{a}} + \sum_{\mathbf{a}} \bar{\Psi}_{\mathbf{i}}^{(\mathbf{f})} (\mathbf{i} \mathbf{D}_{\mathbf{ij}} \mathbf{m}_{\mathbf{f}} \delta_{\mathbf{ij}}) \Psi_{\mathbf{j}}^{(\mathbf{f})}$

• aside: gauge field transforms as  $t^a A_a \rightarrow t^a A'_a = U t^a A_a U^{-1} + \frac{1}{g_s} (\partial U) U^{-1}$ 

non Abelian part

 invariance of the first term more difficult to show

$$\mathcal{L} = -\frac{1}{4} \mathbf{F}_{\mathbf{a}}^{\mu\nu} \mathbf{F}_{\mu\nu}^{\mathbf{a}} + \sum_{\mathbf{f}} \bar{\boldsymbol{\Psi}}_{\mathbf{i}}^{(\mathbf{f})} (\mathbf{i} \not\!\!{D}_{\mathbf{ij}} - \mathbf{m}_{\mathbf{f}} \, \delta_{\mathbf{ij}}) \boldsymbol{\Psi}_{\mathbf{j}}^{(\mathbf{f})}$$

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$$\mathbf{t_a} \mathbf{F_{\mu 
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to proof: use definition, commutation relation for generators, and consider action on a field

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 $[\mathbf{D}_{\mu},\mathbf{D}_{\nu}]_{\mathbf{i}\mathbf{k}}\Psi_{\mathbf{k}}\rightarrow\mathbf{U}_{\mathbf{i}\mathbf{j}}\,[\mathbf{D}_{\mu},\mathbf{D}_{\nu}]_{\mathbf{j}\mathbf{k}}\Psi_{\mathbf{k}}$ 

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u}]_{\mathbf{j}\mathbf{k}} \Psi_{\mathbf{k}}$ 

• find that field strength is not gauge invariant (unlike in QED)

 $(\mathbf{t_a} \mathbf{F_{\mu\nu}^a})_{ij} \rightarrow \mathbf{U_{ik}} (\mathbf{t_a} \mathbf{F_{\mu\nu}^a})_{kl} \mathbf{U_{lj}^{-1}}$ 

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• easiest to see by first re-writing field strength tensor as

 $\mathbf{t_a} \mathbf{F^a_{\mu
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to proof: use definition, commutation relation for generators, and consider action on a field

• exploit that the commutator transforms as the covariant derivative itself

 $[\mathbf{D}_{\mu},\mathbf{D}_{\nu}]_{\mathbf{i}\mathbf{k}}\Psi_{\mathbf{k}}\rightarrow\mathbf{U}_{\mathbf{i}\mathbf{j}}\,[\mathbf{D}_{\mu},\mathbf{D}_{\nu}]_{\mathbf{j}\mathbf{k}}\Psi_{\mathbf{k}}$ 

• find that field strength is not gauge invariant (unlike in QED)

 $(\mathbf{t_a} \mathbf{F_{\mu\nu}^a})_{ij} \rightarrow \mathbf{U_{ik}} (\mathbf{t_a} \mathbf{F_{\mu\nu}^a})_{kl} \mathbf{U_{lj}^{-1}}$ 

• however, the combination that appears in the Lagrangian is invariant

$$-\frac{1}{4}\mathbf{F}_{\mu\nu}^{\mathbf{a}}\mathbf{F}_{\mathbf{a}}^{\mu\nu} = -\frac{1}{2}\mathbf{Tr}(\mathbf{t}_{\mathbf{a}}\mathbf{F}_{\mu\nu}^{\mathbf{a}}\mathbf{t}^{\mathbf{b}}\mathbf{F}_{\mathbf{b}}^{\mu\nu})$$
  
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Ike in QED, a gluon mass term is prohibited by gauge invariance

- the Lagrangian encodes all the rich physics phenomena of QCD
- in these lectures we are interested in perturbative QCD
   -> how to read off Feynman rules to compute cross sections?

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### quark and gluon propagators

simple prescription:

• consider free, non-interacting theory  $(g_s = 0)$ 

$$\mathcal{L}_{\mathbf{free}} = \bar{\Psi}_{\mathbf{i}} (\mathbf{i} \partial_{\mu} \gamma^{\mu} - \mathbf{m}) \delta_{\mathbf{ij}} \Psi_{\mathbf{j}} - \frac{1}{4} (\partial_{\mu} \mathbf{A}^{\mathbf{a}}_{\nu} - \partial_{\nu} \mathbf{A}^{\mathbf{a}}_{\mu})^{2}$$

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gluon propagator:

$$\frac{\mathbf{i}}{2}\mathbf{A}_{\mu}(\mathbf{p^{2}}\mathbf{g}^{\mu\nu}-\mathbf{p}^{\mu}\mathbf{p}^{\nu})\mathbf{A}_{\nu}$$

#### inverse does not exist

encounter similar problem in QED problem is freedom of gauge

### gauge fixing and the gluon propagator

solution: add a gauge fixing term to the Lagrangian, e.g.,

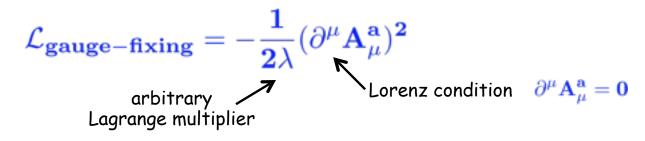
$$\mathcal{L}_{\textbf{gauge-fixing}} = -\frac{1}{2\lambda} (\partial^{\mu} \mathbf{A}^{\mathbf{a}}_{\mu})^{\mathbf{2}}$$

$$\text{arbitrary} \quad \textbf{Lorenz condition} \quad \partial^{\mu} \mathbf{A}^{\mathbf{a}}_{\mu} = \mathbf{0}$$

$$\text{Lagrange multiplier}$$

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leads to extra term such that an inverse now exists

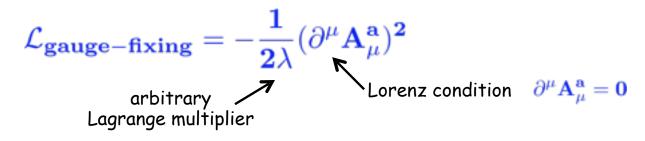
$$\frac{\mathbf{i}}{2\lambda} \mathbf{A}^{\mathbf{a}}_{\mu} \mathbf{p}^{\mu} \mathbf{p}^{\nu} \mathbf{A}^{\mathbf{a}}_{\nu} \qquad \frac{-\mathbf{i}}{\mathbf{p}^{2}} \left( \mathbf{g}_{\mu\nu} - (\mathbf{1} - \lambda) \frac{\mathbf{p}^{\mu} \mathbf{p}^{\nu}}{\mathbf{p}^{2}} \right) \delta^{\mathbf{a}\mathbf{b}}$$

$$A, \mu \qquad \rho \qquad B, \nu$$

• particularly simple choice is Feynman gauge ( $\lambda$ =1)

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$$\overset{\mathsf{A},\mu}{\overbrace{\mathbf{0}},\mathbf{0},\mathbf{0},\mathbf{0},\mathbf{0}} \overset{\mathsf{p}}{\underset{\mathbf{b}}}^{\mathsf{B},\nu}$$

- particularly simple choice is Feynman gauge ( $\lambda$ =1)
- $\bullet$  gauge fixing breaks explicitly gauge invariance though but since  $\lambda$  is arbitrary this leaves us with a powerful check of calculations

any dependence on  $\lambda$  must ultimately cancel in physical observables

# another peculiarity: ghosts

- gauge fixing leads to consistent quantization of QED
- more trouble ahead for non Abelian theories:
  - covariant gauges introduce unphysical longitudinal d.o.f. for the gluon as for a photon only transverse d.o.f. are physical



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    - complex scalar field which obeys Fermi statistics
    - new Feynman rules: propagator and gluon-ghost-ghost coupling
    - eats unphysical degrees of freedom in polarization sum

$$\sum_{\lambda=\pm 1,-1,0} \left| \operatorname{verter}_{\lambda \to \lambda}^{yyyt} \right|^2 = \left| \operatorname{verter}_{\lambda}^{yyyt} \right|^2 = \sum_{\lambda=\pm 1,-1} \left| \operatorname{verter}_{\lambda \to \lambda}^{yyyt} \right|^2$$



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- alternatively one can choose a non-covariant (axial) gauge  $\mathcal{L}_{axial} = -\frac{1}{\xi} (\mathbf{n}^{\mu} \mathbf{A}^{a}_{\mu})^{2}$ 
  - at the expense of a more complicated gluon propagator

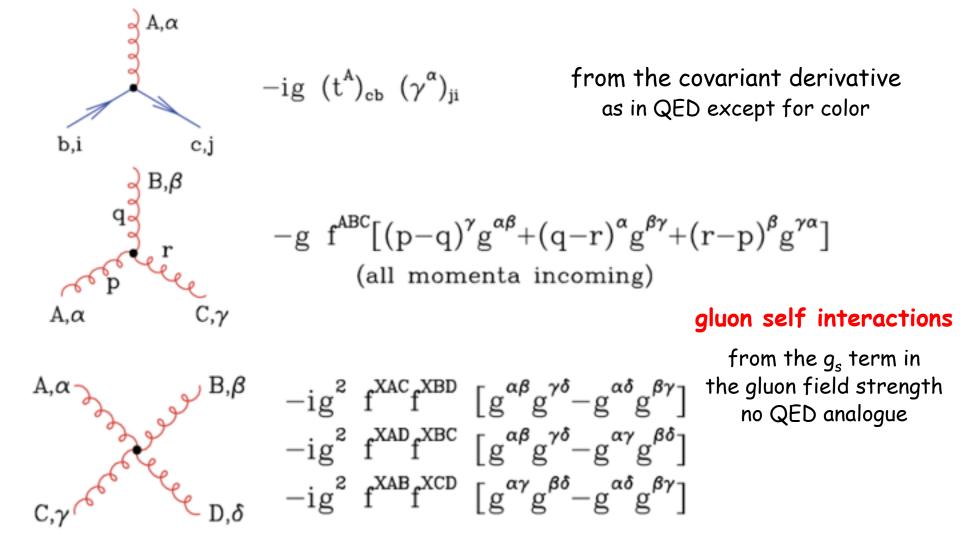
$$\frac{\mathbf{i}}{\mathbf{p^2}} \left( -\mathbf{g}_{\mu\nu} + \frac{\mathbf{n}_{\mu}\mathbf{p}_{\nu} + \mathbf{n}_{\nu}\mathbf{p}_{\mu}}{\mathbf{n}\cdot\mathbf{p}} + \frac{(\mathbf{n^2} + \xi\mathbf{p^2})\,\mathbf{p}_{\mu}\mathbf{p}_{\nu}}{(\mathbf{n}\cdot\mathbf{p})^2} \right) \delta_{\mathbf{ab}}$$



arbitrary direction

### using the QCD Lagrangian: interactions

- interactions between quarks and gluons can be simply read off from the terms in the Lagrangian containing  $g_s$ 



### take home message for part I THE FOUNDATIONS





# QCD is based on a simple Lagrangian but has a rich phenomenology



### QCD is based on the non Abelian gauge group SU(3)

- number of colors and group structure can be tested experimentally
- concept of local gauge invariance dictates interactions
- similarities to QED, yet profound differences (and more to come)
- color leads to self-interactions between "force carrying" gluons
- perturbation theory can be based on a short list of Feynman rules



### color algebra decouples and can be performed separately

- color factors can be expressed in terms of two Casimirs:  $C_A$  and  $C_F$
- powerful pictorial methods; possibility of "leading color approximation"



# Part II the QCD toolbox asymptotic freedom, IR safety, QCD final state, factorization

the gauge principle is elegant and powerful but any theory must ultimately stand (or fall) by its success (or failure)

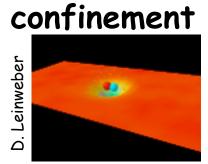
QCD is the theory of **strong** interactions

- how can we make use of **perturbative** methods?

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non-perturbative structure of hadrons e.g. through lattice QCD

## asymptotic freedom

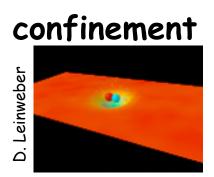
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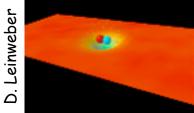
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QCD is the theory of **strong** interactions

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# confinement



non-perturbative structure of hadrons

e.g. through lattice QCD



## asymptotic freedom

hard scattering cross sections and renormalization group

with perturbative methods

interplay

probing hadronic structure with weakly interacting quanta of asymptotic freedom



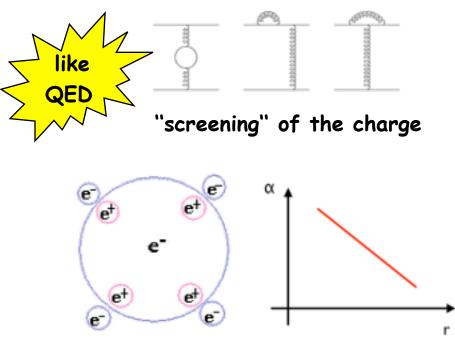


Gross, Wilczek; Politzer ('73/'74) Nobel prize 2004





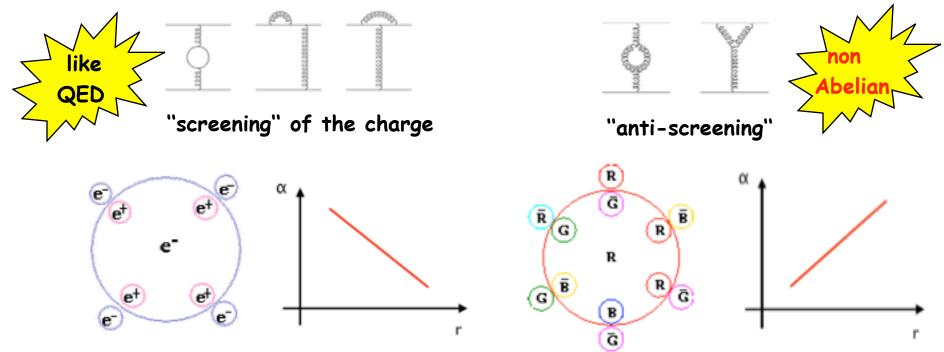
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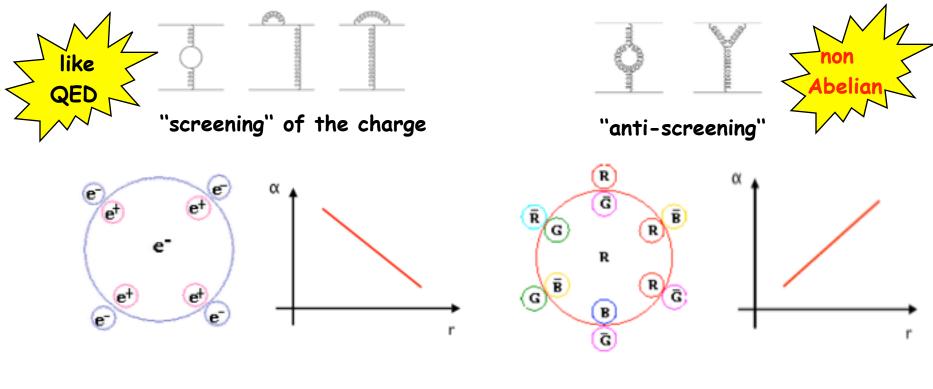
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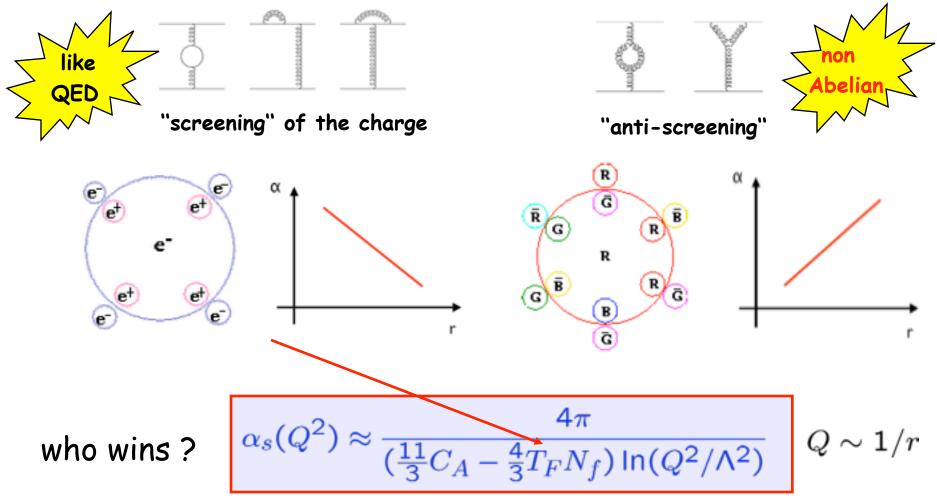


$$\alpha_s(Q^2) \approx \frac{4\pi}{\left(\frac{11}{3}C_A - \frac{4}{3}T_F N_f\right) \ln(Q^2/\Lambda^2)} \quad Q \sim 1/r$$





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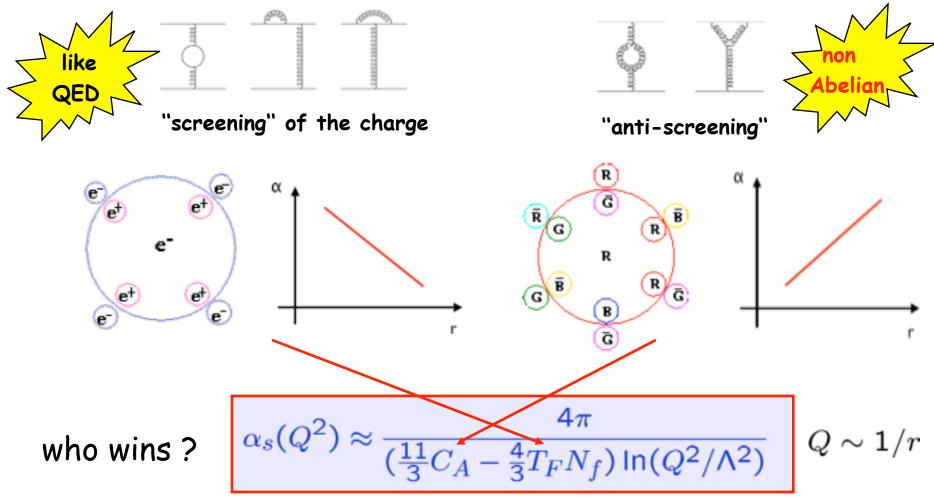






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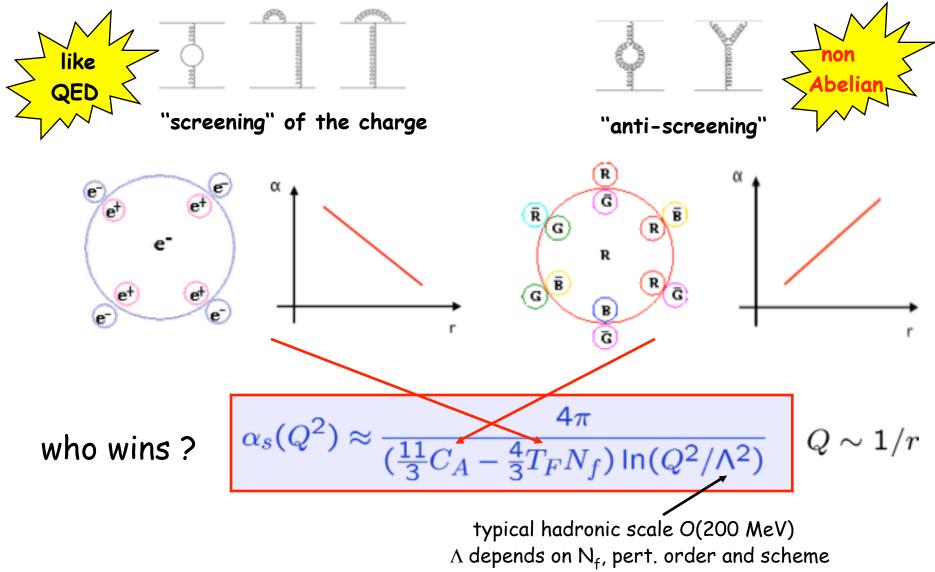


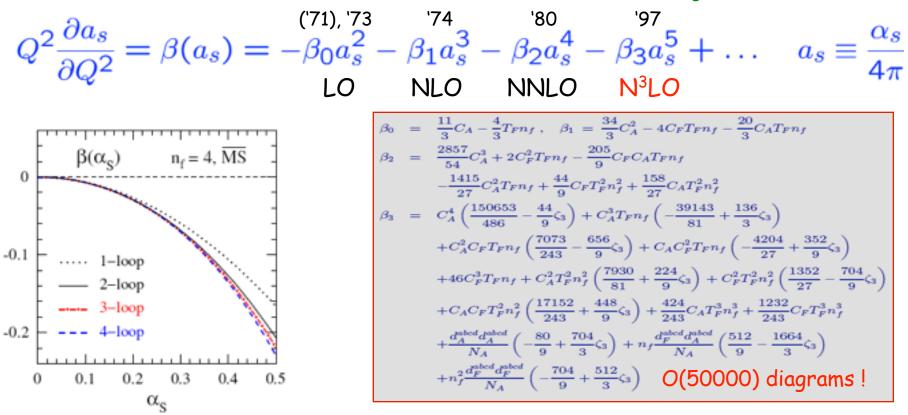


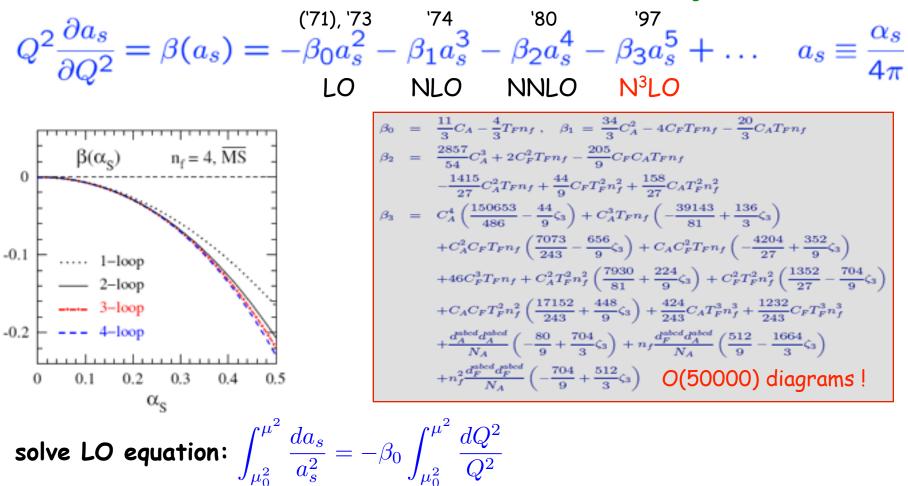


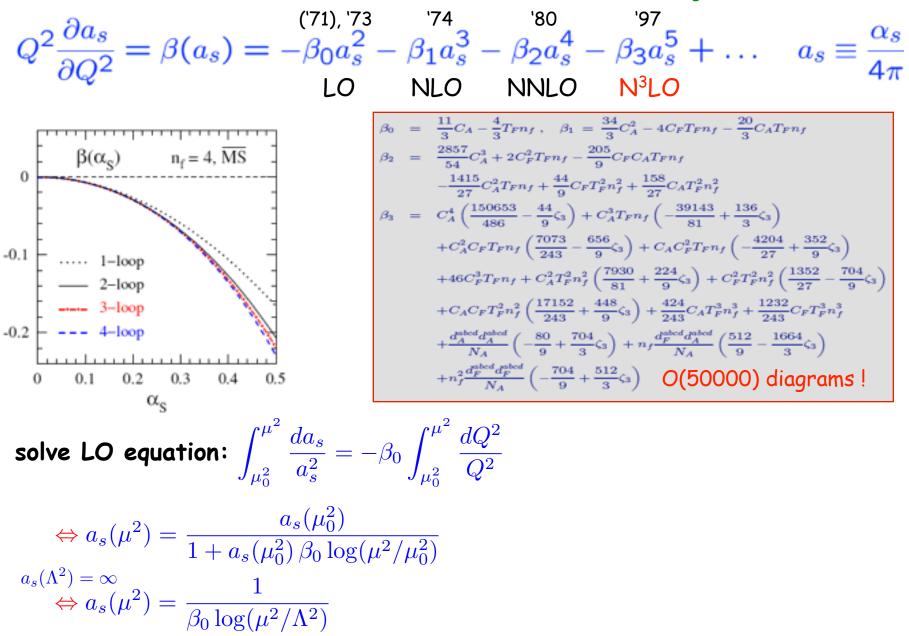
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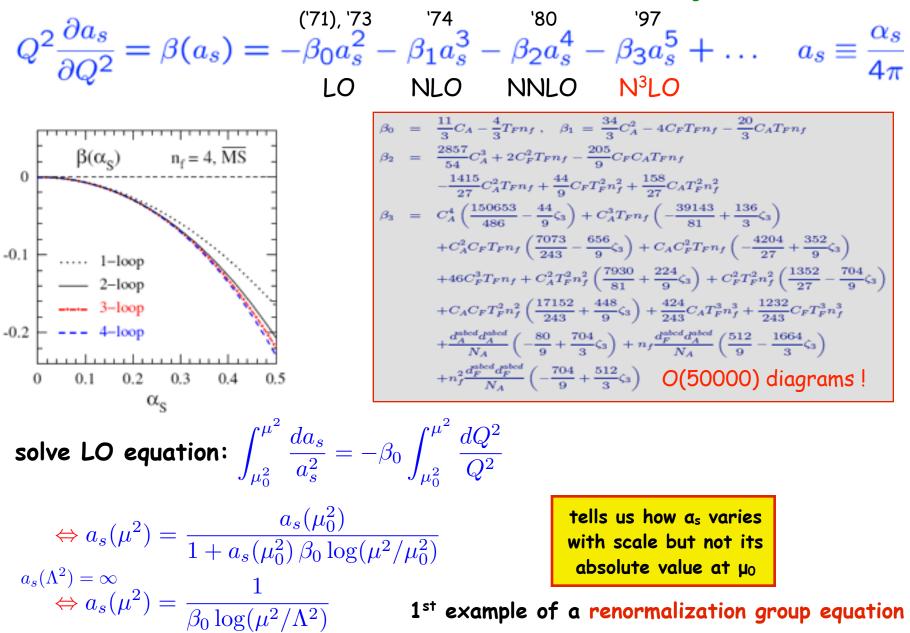












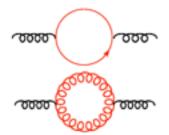
## some further observations

#### recap

$$\beta = -\alpha_s^2(\mu) \sum_i b_i \alpha_s^i(\mu)$$

$$b_0 = \frac{11N_c - 4n_f T_R}{12\pi}$$

- negative contribution to b<sub>0</sub> due to
- $\bullet$  positive contribution to  $b_0$  due to
- positive contribution larger -> b<sub>0</sub> > 0 (-> overall: negative beta function)



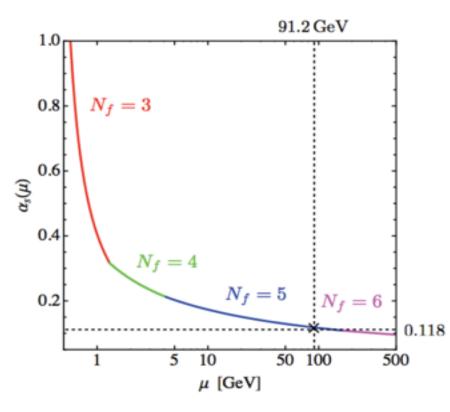
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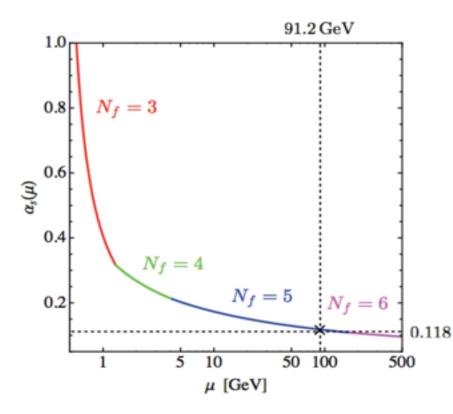
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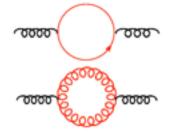
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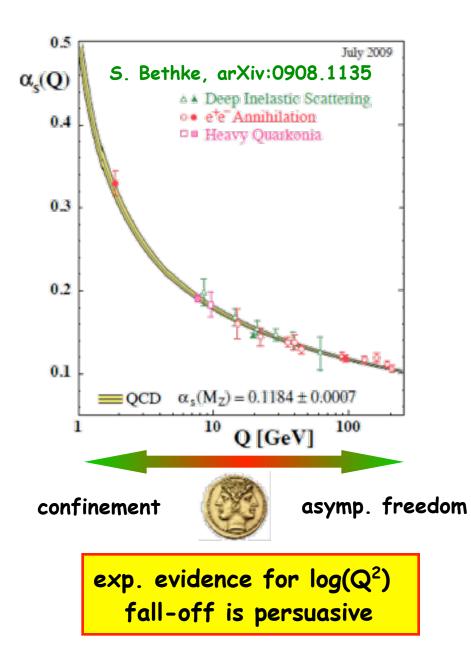
- coupling depends on number of active flavors (need matching a thresholds)
- can read off QED beta function (T<sub>R</sub> coefficient) (only one flavor)

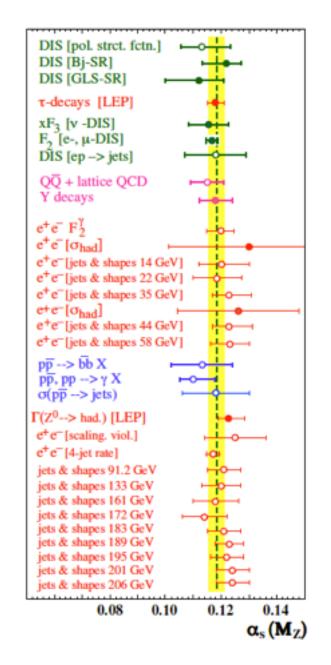
$$\beta_{\rm QED} = \frac{1}{3\pi}\alpha^2 + \dots$$

 $b_0$  negative  $\rightarrow$  overall: positive beta fct.



### consistent picture from many observables





**upshot**: a strongly interacting theory at long-distance can become weakly interacting at short-distance

Is this enough to explain the success of the parton model and pQCD?

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- factorization

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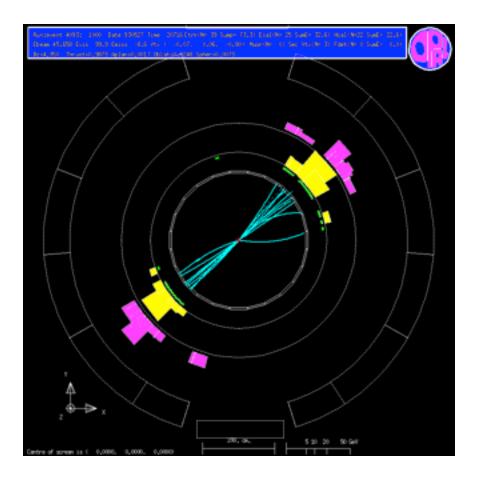
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let's study electron-positron annihilation to see what this is all about ...

## e<sup>+</sup>e<sup>-</sup> annihilation: the **QCD** guinea pig

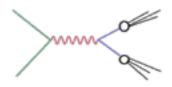
most of the hadronic events at CERN-LEP had two back-to-back jets

1989-2000

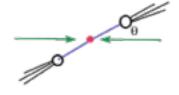


**jet**: pencil-like collection of hadrons

• jets resemble features of underlying 2->2 hard process  $e^+e^- \rightarrow q\bar{q}$ 



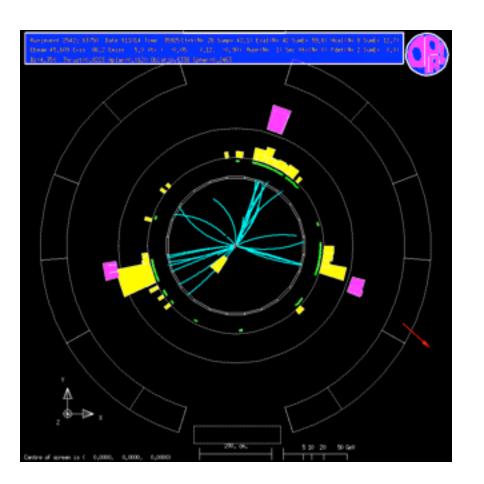
 angular distribution of jet axis w.r.t. beam axis as predicted for spin-<sup>1</sup>/<sub>2</sub> quarks



jets play major role in hadron-hadron collisions at TeVatron, RHIC, LHC

## e<sup>+</sup>e<sup>-</sup> annihilation: three-jet events

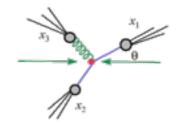
### about 10% of the events had a third jet



first discovered at DESY-PETRA in 1979

 jets resemble features of underlying 2->3 hard process  $e^+e^- \rightarrow q\bar{q}g$ 

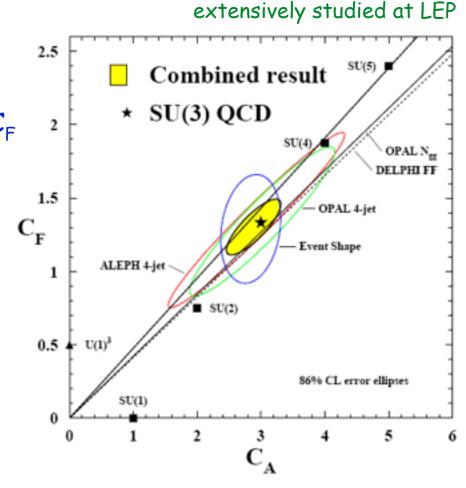
- 10% rate consistent with  $\alpha_{\rm s} \simeq$  0.1 (determination of  $\alpha_{\rm s}$ )
- angular distribution of jets w.r.t. beam axis as expected for spin-1 gluons



### e<sup>+</sup>e<sup>-</sup> annihilation: four-jet events

some events even had a fourth jet

- angular correlations between four jets depend on  $C_A/C_F$  and  $T_F/C_F$
- sensitivity to non-Abelian
   three-gluon-vertex
   LO: Ellis, Ross, Terrano
- crucial test of QCD when combined with results for event shapes (thrust, etc.)



## e<sup>+</sup>e<sup>-</sup> annihilation: four-jet events

some events even had a fourth jet extensively studied at LEP 2.5SU(5) **Combined** result angular correlations between SU(3) QCD four jets depend on  $C_A/C_F$  and  $T_F/C_F$ 2 OPAL N., DELPHI FF sensitivity to non-Abelian 1.5 **OPAL 4-jet** three-gluon-vertex C<sub>F</sub> LO: Ellis, Ross, Terrano Event Shape ALEPH 4-jet 1 crucial test of QCD when 0.5 k U(1)3 combined with results for 86% CL error ellipses SU(1) event shapes (thrust, etc.) 0 2 6

> e<sup>+</sup>e<sup>-</sup> experiments played a vital role in establishing QCD as the correct theory of strong interactions and SU(3) as the underlying gauge group

# recipe for quantitative calculations



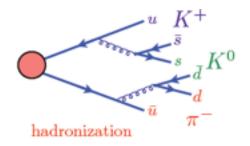
- (1) identify the final-state of interest and draw all relevant Feynman diagrams
- (2) use SU(3) algebra to take care of QCD color factors
- (3) compute the rest of the diagram using "Diracology" traces of gamma matrices, spinors, ...
- (4) to turn squared matrix elements into a cross section we need to
  - account for the available phase space (momentum d.o.f. in final-state)
  - integrate out not observed d.o.f.
  - normalize by incoming flux

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but wait ... experiments do not see free quarks and gluons

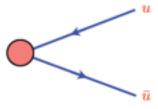


# recipe for quantitative calculations

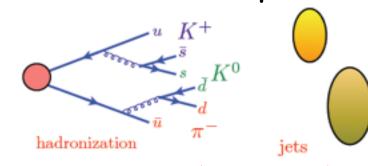


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energetic partons



cleanest observables in QCD

will find that most "stuff" is observed in the directions of produced quarks & gluons parton-hadron duality

## bunch of automated LO tools

- LO estimates of cross sections are practically a solved problem
- many useful fully automated tools available (limitations for high multiplicities)

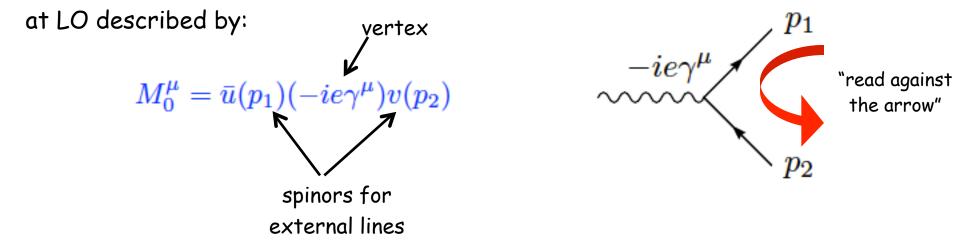
| ALPGEN   | M. L. Mangano et al.<br>http://alpgen.web.cern.ch/alpgen/                                   |
|----------|---------------------------------------------------------------------------------------------|
| AMEGIC++ | F. Krauss et al.<br>http://projects.hepforge.org/sherpa/dokuwiki/doku.php                   |
| CompHEP  | E. Boos et al.<br>http://comphep.sinp.msu.ru/                                               |
| HELAC    | C. Papadopoulos, M. Worek<br>http://helac-phegas.web.cern.ch/helac-phegas/helac-phegas.html |
| Madgraph | F. Maltoni, T. Stelzer<br>http://madgraph.hep.uiuc.edu/                                     |

let's have a closer look at the R-ratio already encountered in Part I

$$R \equiv \frac{e^+e^- \to \text{hadrons}}{e^+e^- \to \mu^+\mu^-} \propto N_c \sum_f Q_f^2$$

let's have a closer look at the R-ratio already encountered in Part I

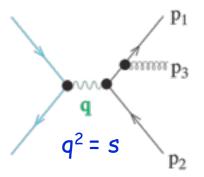
$$R \equiv \frac{e^+e^- \to \text{hadrons}}{e^+e^- \to \mu^+\mu^-} \propto N_c \sum_f Q_f^2$$



### exploring the QCD final-state: $e^+e^- \rightarrow 3$ partons

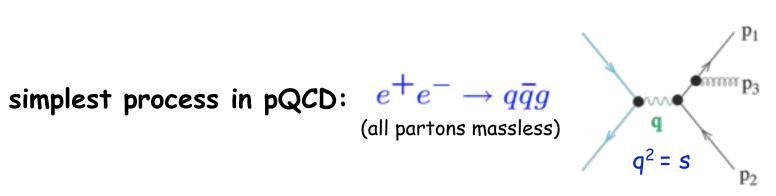
simplest process in pQCD:

$$e^+e^- \rightarrow q\bar{q}g$$



(all partons massless)

### exploring the QCD final-state: $e^+e^- \rightarrow 3$ partons



some kinematics first:

• energy fractions & conservation:  $x_i \equiv \frac{2p_i \cdot q}{s} = \frac{E_i}{\sqrt{s/2}}$   $\sum x_i = \frac{2(\sum p_i) \cdot q}{s} = 2$ 

### exploring the QCD final-state: $e^+e^- \rightarrow 3$ partons

simplest process in pQCD:  $e^+e^- \rightarrow q\bar{q}g$ (all partons massless)  $q^2 = s$ 

some kinematics first:

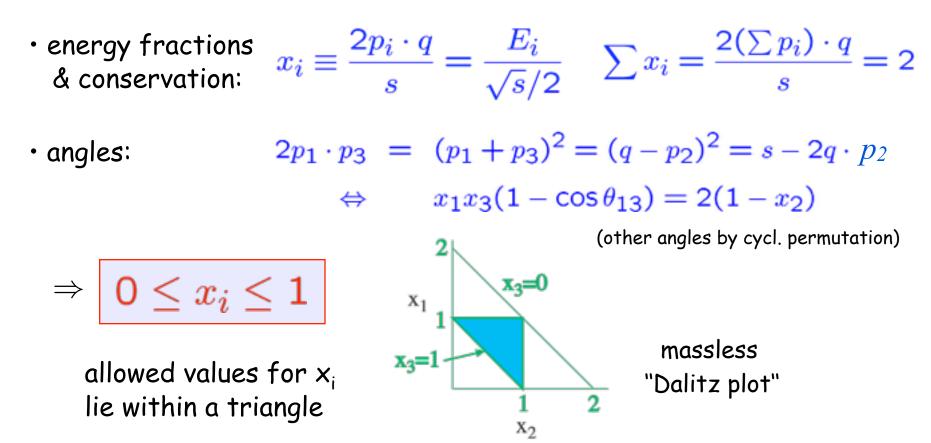
 $\begin{array}{ll} \cdot \text{ energy fractions} \\ \& \text{ conservation:} \end{array} \quad x_i \equiv \frac{2p_i \cdot q}{s} = \frac{E_i}{\sqrt{s/2}} \quad \sum x_i = \frac{2(\sum p_i) \cdot q}{s} = 2 \\ \cdot \text{ angles:} \qquad 2p_1 \cdot p_3 \ = \ (p_1 + p_3)^2 = (q - p_2)^2 = s - 2q \cdot p_2 \\ \Leftrightarrow \qquad x_1 x_3 (1 - \cos \theta_{13}) = 2(1 - x_2) \end{array}$ 

(other angles by cycl. permutation)

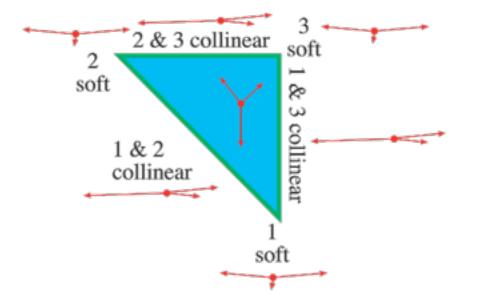
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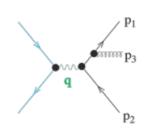
simplest process in pQCD:  $e^+e^- \rightarrow q\bar{q}g$ (all partons massless)

some kinematics first:



at the boundaries of phase space we encounter **special kinematic configurations**:





• "edges": two partons collinear

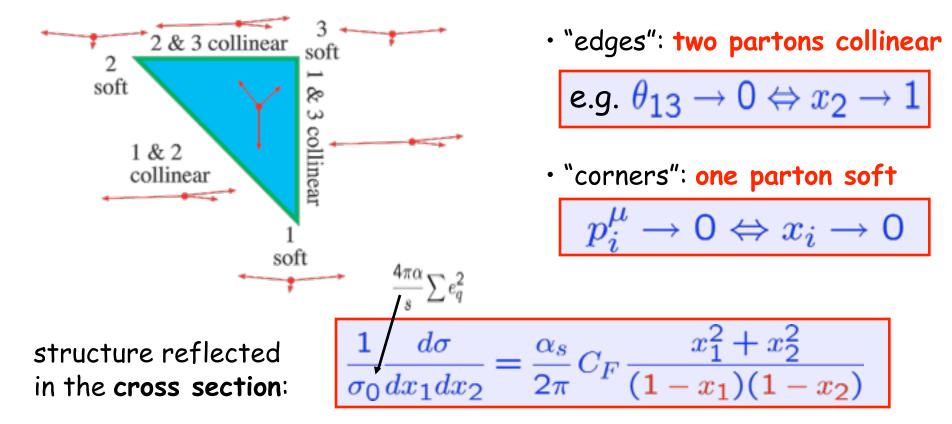
e.g. 
$$\theta_{13} \rightarrow 0 \Leftrightarrow x_2 \rightarrow 1$$

"corners": one parton soft

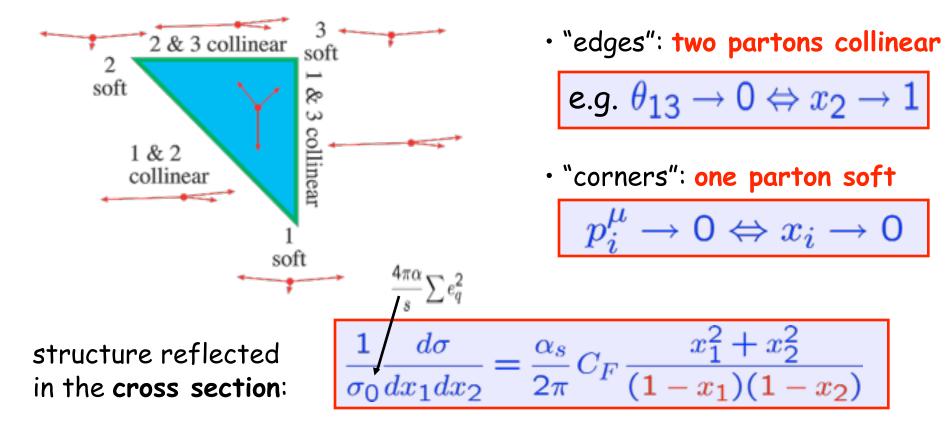
$$p_i^\mu o \mathsf{0} \Leftrightarrow x_i o \mathsf{0}$$

q

at the boundaries of phase space we encounter **special kinematic configurations**:



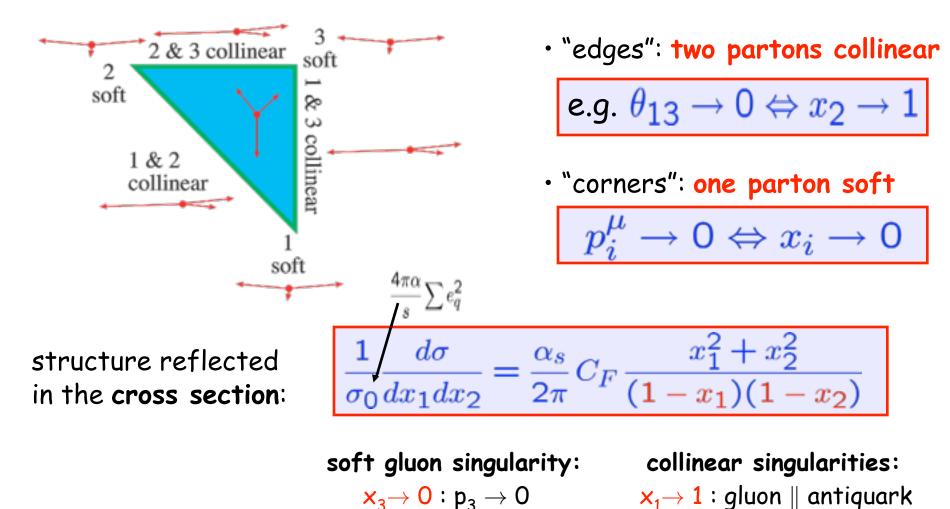
at the boundaries of phase space we encounter **special kinematic configurations**:



 $\begin{array}{l} \textbf{collinear singularities:} \\ \textbf{x}_1 \rightarrow \textbf{1}: \texttt{gluon} \parallel \texttt{antiquark} \\ \textbf{x}_2 \rightarrow \textbf{1}: \texttt{gluon} \parallel \texttt{quark} \end{array}$ 

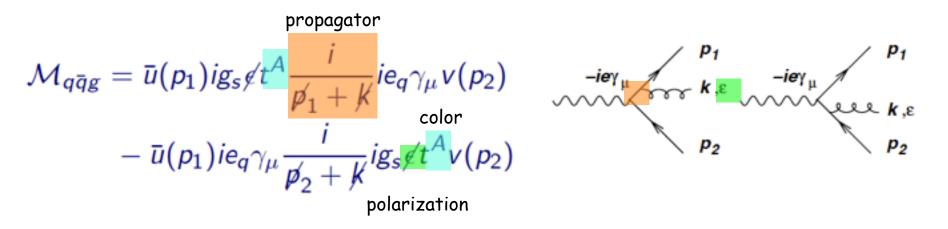
q

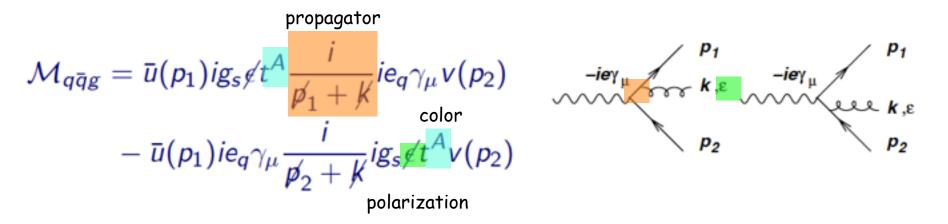
at the boundaries of phase space we encounter **special kinematic configurations**:



 $\leftrightarrow \mathbf{x_1} \rightarrow \mathbf{1} \& \mathbf{x_2} \rightarrow \mathbf{1}$ 

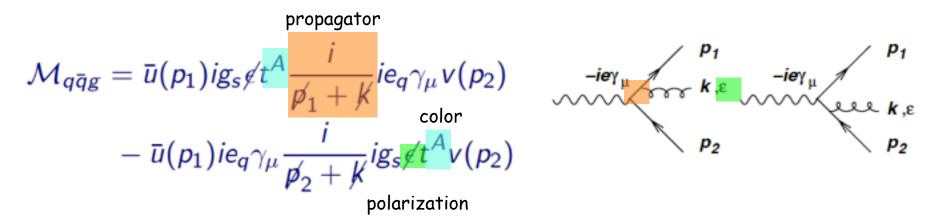
 $x_2 \rightarrow 1$ : gluon || quark





make gluon soft  $k \ll p_{1,2}$  and square the amplitude

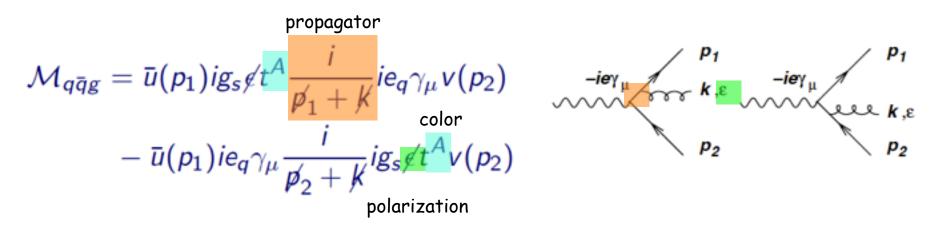
$$|M_{q\bar{q}g}^{2}| \simeq \sum_{A,\text{pol}} \left| \bar{u}(p_{1})ie_{q}\gamma_{\mu}t^{A}v(p_{2}) g_{s} \left( \frac{p_{1}.\epsilon}{p_{1}.k} - \frac{p_{2}.\epsilon}{p_{2}.k} \right) \right|^{2} \frac{e_{F}}{color factor}$$
$$= -|M_{q\bar{q}}^{2}|C_{F}g_{s}^{2} \left( \frac{p_{1}}{p_{1}.k} - \frac{p_{2}}{p_{2}.k} \right)^{2} = |M_{q\bar{q}}^{2}|C_{F}g_{s}^{2} \frac{2p_{1}.p_{2}}{(p_{1}.k)(p_{2}.k)}$$



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**Eikonal factor** 



make gluon soft  $k \ll p_{1,2}$  and square the amplitude

$$|M_{q\bar{q}g}^{2}| \simeq \sum_{A,pol} \left| \bar{u}(p_{1})ie_{q}\gamma_{\mu}t^{A}v(p_{2}) g_{s} \left( \frac{p_{1}.\epsilon}{p_{1}.k} - \frac{p_{2}.\epsilon}{p_{2}.k} \right) \right|^{2} \qquad \text{sum over gluon polarizations}$$

$$= -|M_{q\bar{q}}^{2}|C_{F}g_{s}^{2} \left( \frac{p_{1}}{p_{1}.k} - \frac{p_{2}}{p_{2}.k} \right)^{2} = |M_{q\bar{q}}^{2}|C_{F}g_{s}^{2} \frac{2p_{1}.p_{2}}{(p_{1}.k)(p_{2}.k)}$$
Include phase space for gluon
$$d\Phi_{q\bar{q}g}|M_{q\bar{q}g}^{2}| \simeq \left( d\Phi_{q\bar{q}}|M_{q\bar{q}}^{2}| \right) \xrightarrow{d^{3}\vec{k}}{2E(2\pi)^{3}} C_{F}g_{s}^{2} \frac{2p_{1}.p_{2}}{(p_{1}.k)(p_{2}.k)}$$
note: color will in general not factorize in soft limit

$$d\Phi_{q\bar{q}g}|M_{q\bar{q}g}^2| \simeq (d\Phi_{q\bar{q}}|M_{q\bar{q}}^2|) \frac{d^3\bar{k}}{2E(2\pi)^3} C_F g_s^2 \frac{2p_1 \cdot p_2}{(p_1 \cdot k)(p_2 \cdot k)}$$

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soft emission factor dS

$$EdE \, d\cos\theta \, \frac{d\phi}{2\pi} \cdot \frac{2\alpha_{\rm s}C_F}{\pi} \frac{2p_1.p_2}{(2p_1.k)(2p_2.k)} \begin{array}{l} \theta \equiv \theta_{p_1k} \\ \phi = {\rm azimuth} \end{array}$$

$$d\Phi_{q\bar{q}g}|M_{q\bar{q}g}^{2}| \simeq (d\Phi_{q\bar{q}}|M_{q\bar{q}}^{2}|) \frac{d^{3}\vec{k}}{2E(2\pi)^{3}} C_{F}g_{s}^{2} \frac{2p_{1}.p_{2}}{(p_{1}.k)(p_{2}.k)}$$
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$$\phi = \text{azimuth}$$

$$express \text{ in } = \frac{1}{E^{2}(1-\cos^{2}\theta)}$$

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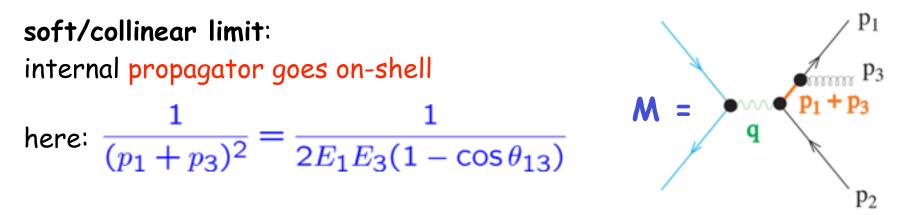
end up with

$$d\mathcal{S} = \frac{2\alpha_{\rm s}C_F}{\pi} \, \frac{dE}{E} \frac{d\theta}{\sin\theta} \, \frac{d\phi}{2\pi}$$

• It diverges for  $E \rightarrow 0$  — infrared (or soft) divergence

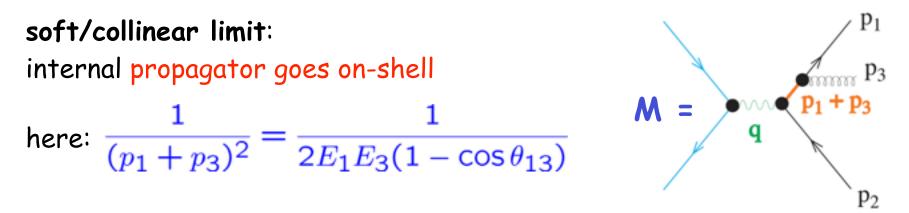
▶ It diverges for  $\theta \rightarrow 0$  and  $\theta \rightarrow \pi$  — collinear divergence

## general nature of these singularities

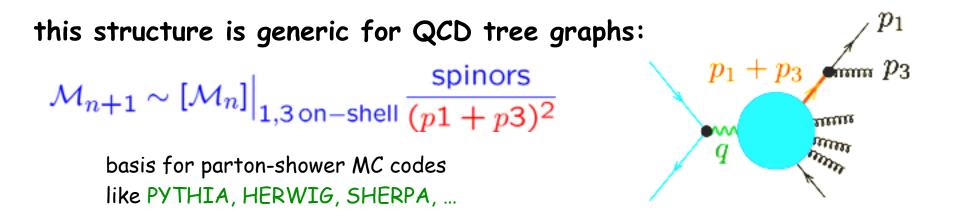


note: "soft quarks" (here  $E_1 \rightarrow 0$ ) never lead to singularities (canceled by numerator)

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## Do we observe a breakdown of pQCD already here?

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the lesson is:

whenever the 2->(n+1) kinematics collapses to an effective 2->n parton kinematics due to

- the emission of a soft gluon
- a collinear splitting of a parton into two partons

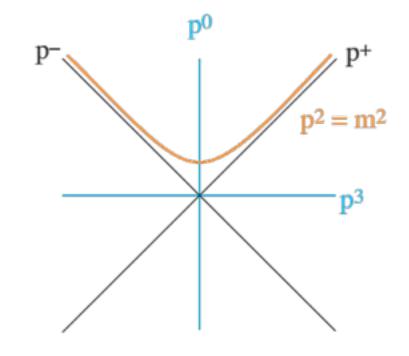
we have to be much more careful and work a bit harder!

this applies to all pQCD calculations

## towards a space-time picture of the singularities

#### interlude: light-cone coordinates

$$p^{\pm} \equiv (p^{0} \pm p^{3})/\sqrt{2}$$
$$p^{2} = 2p^{+}p^{-} - \vec{p}_{T}^{2}$$
$$p^{-} = (p_{T}^{2} + m^{2})/2p^{+}$$

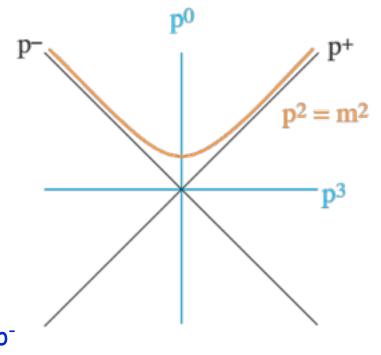


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particle with large momentum in +p<sup>3</sup> direction has large p<sup>+</sup> and small p<sup>-</sup>



## towards a space-time picture of the singularities

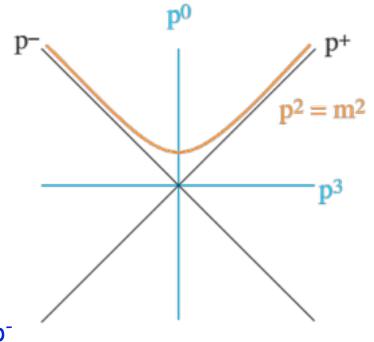
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Fourier transform momentum space  $\xrightarrow{e^{ip \cdot x}}$  coordinate space  $p \cdot x = p^+ x^- + p^- x^+ - \vec{p}_T \cdot \vec{x}_T$ 

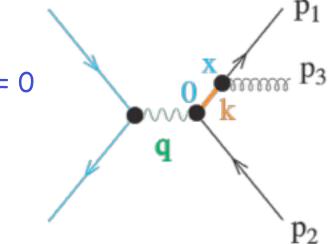
-->  $x^-$  is conjugate to  $p^+$  and  $x^+$  is conjugate to  $p^-$ 



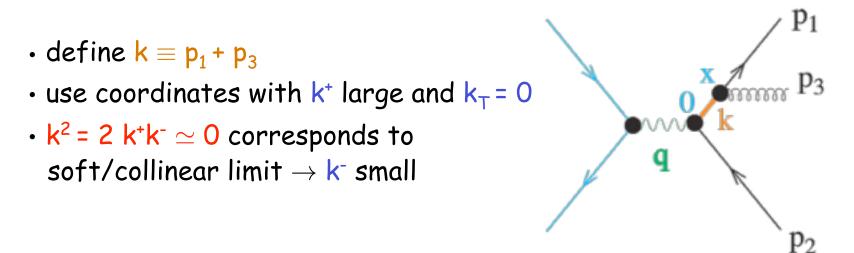
What does this imply for our propagator going on-shell?

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- define  $\mathbf{k} \equiv \mathbf{p}_1 + \mathbf{p}_3$
- use coordinates with  $k^+$  large and  $k_T = 0$
- $k^2 = 2 k^+ k^- \simeq 0$  corresponds to soft/collinear limit  $\rightarrow k^-$  small

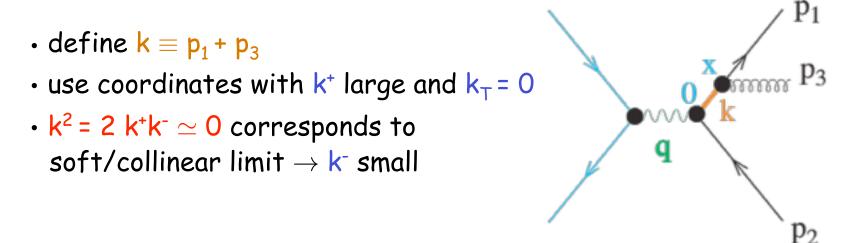


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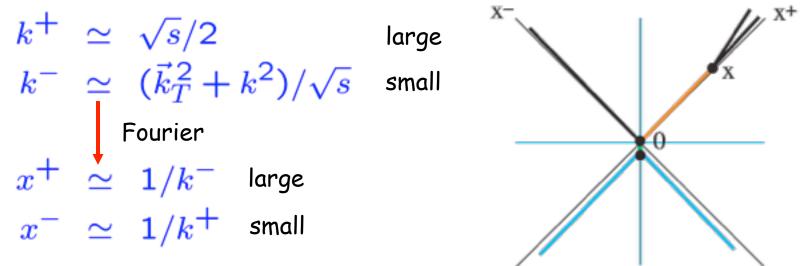


How far does the internal on-shell parton travel in space-time?

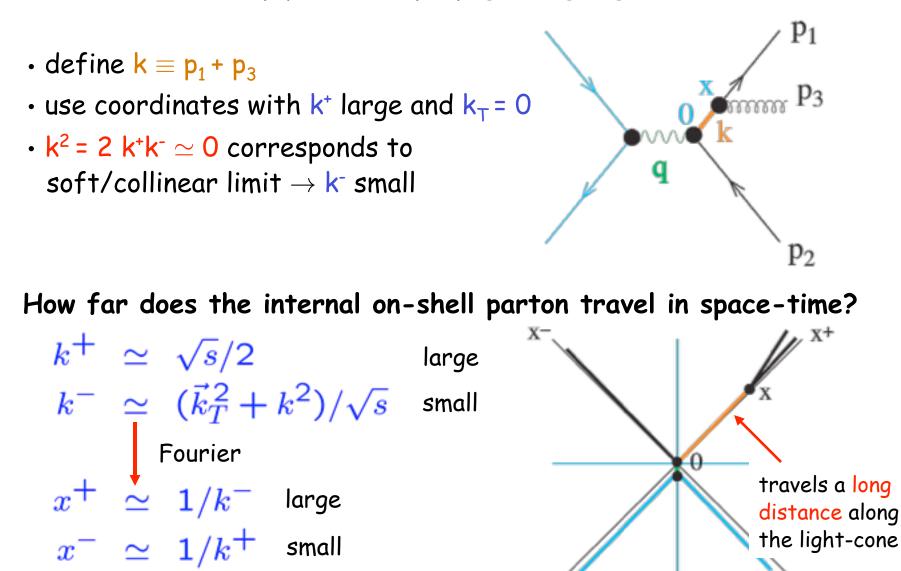
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**SO** ..... What to do with the long-distance physics associated with these soft/collinear singularities? Is there any hope that we can predict some reliable numbers to compare with experiment?

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**SO** ..... What to do with the long-distance physics associated with these soft/collinear singularities? Is there any hope that we can predict some reliable numbers to compare with experiment?

> to answer this, we have to formulate the concept of infrared safety

# infrared-safe observables

#### formal definition of infrared safety:

Kunszt, Soper; ...

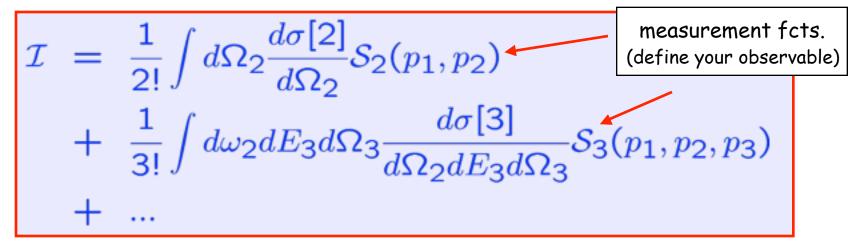
study inclusive observables which do not distinguish between (n+1) partons and n partons in the soft/collinear (=degenerate) limit, i.e., are insensitive to what happens at long-distance

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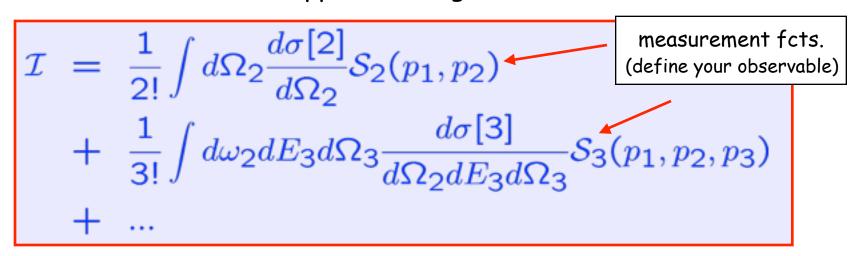


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infrared safe iff [for  $\lambda=0$  (soft) and  $0 < \lambda < 1$  (collinear)]

 $S_{n+1}(p_1,\ldots,(1-\lambda)p_n,\lambda p_n)=S_n(p_1,\ldots,p_n)$ 

# physics behind formal IR safety requirement

cannot resolve soft and collinear partons experimentally

→ intuitively reasonable that a theoretical calculation can be infrared safe as long as it is insensitive to long-distance physics (not a priori guaranteed though)

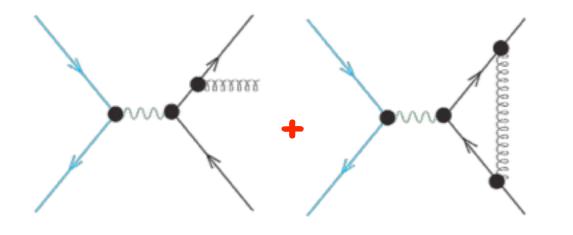
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at a level of a pQCD calculation (e.g.  $e^+e^-$  at  $O(\alpha_s)$ , i.e., n=2)  $S_{n+1}(p_1, \ldots, (1 - \lambda)p_n, \lambda p_n) = S_n(p_1, \ldots, p_n)$ 

 $\rightarrow$  singularities of real gluon emission and virtual corrections cancel in the sum





extension of famous theorems by Kinoshita-Lee-Nauenberg and Bloch-Nordsieck



#### example I: total cross section $e^+e^- \rightarrow hadrons$

### simplest case:

$$\mathcal{S}_n(p_1,\ldots,p_n)=1$$

fully inclusive quantity  $\leftrightarrow$  we don't care what happens at long-distance

- the produced partons will all hadronize with probability one
- we do not observe a specific type of hadron (i.e. sum over a complete set of states)
- we sum over all degenerate kinematic regions

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$$\mathcal{S}_n(p_1,\ldots,p_n)=1$$

fully inclusive quantity  $\leftrightarrow$  we don't care what happens at long-distance

- the produced partons will all hadronize with probability one
- we do not observe a specific type of hadron (i.e. sum over a complete set of states)
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# infrared safe by definition

R ratio:  

$$R = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)} = N_c \sum e_q^2 (1 + \Delta_{\text{QCD}})$$
need to add up real and virtual corrections at a given  $O(a_s)$ 

#### example I: total cross section $e^+e^- \rightarrow hadrons$

### simplest case:

not IR safe:

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multiplicity of gluons or 1-gluon cross section

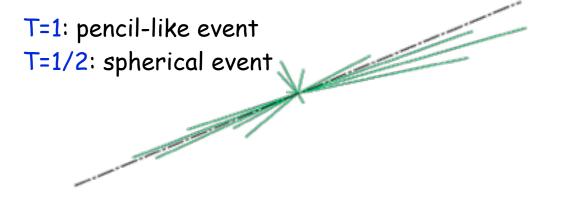
# example II: thrust distribution

#### **somewhat less trivial:** $d\sigma/dT$ (measure of the "event shape")

$$S_n(p_1,\ldots,p_n) = \delta(T - T_n(p_1,\ldots,p_n))$$
  
$$T_n(p_1,\ldots,p_n) \equiv \max|_{\vec{n}} \frac{\sum_{i=1}^n |\vec{p}_i \cdot \vec{n}|}{\sum_{i=1}^n |\vec{p}_i|}$$

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T=1: pencil-like event T=1/2: spherical event why infrared safe? • contributions from soft particles with  $\vec{p_i} \rightarrow 0$  drop out • a collinear splitting does not change the thrust:  $|(1 - \lambda)\vec{p_i} \cdot \vec{n}| + |\lambda \vec{p_i} \cdot \vec{n}| = |\vec{p_i} \cdot \vec{n}|$  $|(1 - \lambda)\vec{p_i}| + |\lambda \vec{p_i}| = |\vec{p_i}|$ 

#### example III: event shape variables

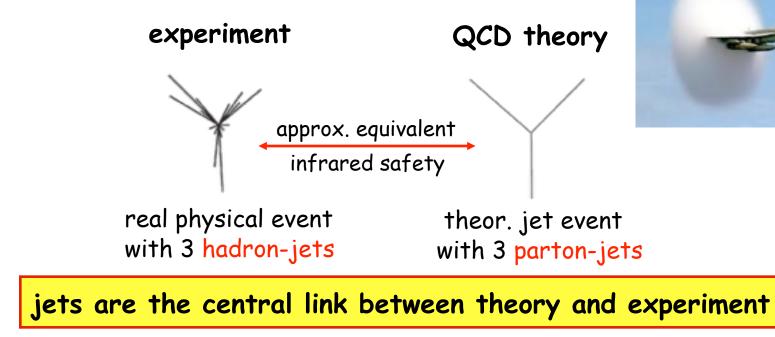
there is a long list of similar infrared safe observables:

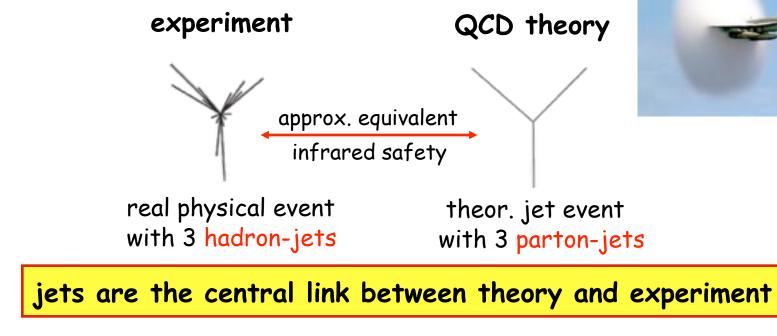
event-shapes: fertile ground for comparison between theory and experiment

- validity of pQCD calculations
- many ways to test SU(3) (color factors)
- spin of quarks and gluons
- measurements of  $\boldsymbol{\alpha}_{\!s}$

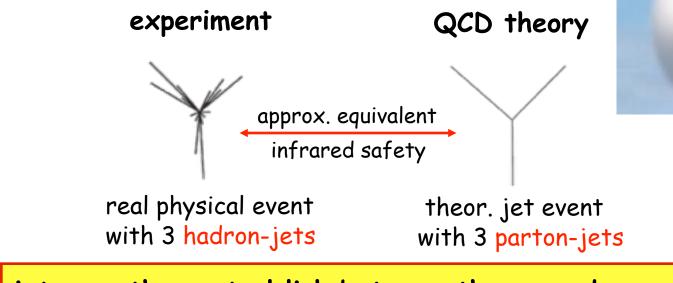
|                               |                                                                                                                                                                                                                        | Typical Value for: |                              |           |                                     |
|-------------------------------|------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|--------------------|------------------------------|-----------|-------------------------------------|
| Name of<br>Observable         | Definition                                                                                                                                                                                                             | Ĵ                  | $\dot{\boldsymbol{\lambda}}$ | ₩         | QCD<br>calculation                  |
| Thrust                        | $\mathbf{T} = \max_{\vec{n}} \left( \frac{\boldsymbol{\Sigma}_i   \vec{p}_i \vec{n}  }{\boldsymbol{\Sigma}_i   \vec{p}_i  } \right)$                                                                                   | 1                  | ≥2/3                         | ≥1/2      | $(resummed) O(\alpha_s^2)$          |
| Thrust major                  | Like T, however $T_{mij}$ and $\vec{n}_{mij}$ in<br>plane $\perp \vec{n}_T$                                                                                                                                            | 0                  | ≤1/3                         | ≤1/√2.    | $O(\alpha_s^2)$                     |
| Thrust minor                  | Like T, however $T_{min}$ and $n_{min}$ in<br>direction $\perp$ to $\vec{n}_T$ and $\vec{n}_{maj}$                                                                                                                     | 0                  | 0                            | ≤1/2      | $O(\alpha_s^2)$                     |
| Oblateness                    | O = T <sub>maj</sub> - T <sub>min</sub>                                                                                                                                                                                | 0                  | ≤1/3                         | 0         | $O(\alpha_s^2)$                     |
| Sphericity                    | $\begin{split} &S = 1.5 (Q_1 + Q_2); Q_1 \leq \leq Q_3 \text{ are} \\ &Eigenvalues of  S^{\alpha\beta} = \frac{\sum_i p_i^{\alpha} p_j^{\beta}}{\sum_i p_i^2} \end{split}$                                             | 0                  | ≤3/4                         | ≤1        | none<br>(aot infrared<br>safe)      |
| Aplanarity                    | A=1.5 Q1                                                                                                                                                                                                               | 0                  | 0                            | ≤1/2      | none (not<br>infrared safe)         |
| Jet (Hemis-<br>phere) masses  | $M_{4}^{2} = (\sum_{i} E_{i}^{2} - \sum_{i} \vec{p}_{i}^{2})_{i \in S_{4}}$ $(S_{5}: \text{Hemispheres } \perp \text{to } \vec{n}_{T})$ $M_{H}^{2} = \max(M_{4}^{2}, M_{-}^{2})$ $M_{D}^{2} =  M_{4}^{2} - M_{-}^{2} $ | 0                  | ≤1/3<br>≤1/3                 | ≤1/2<br>0 | (resummed)<br>O( $\alpha_{s}^{2}$ ) |
| Jet broadening                | $B_{\pm} = \frac{\sum_{i \in S_{\pm}}  \vec{p}_i \times \vec{n}_T }{2 \sum_i  \vec{p}_i }; B_T = B_+ + B$ $B_w = \max(B_+, B)$                                                                                         | 0                  |                              |           | (resummed)                          |
| Energy-Energy<br>Correlations | $EEO(\chi) = \sum_{\text{events}} \sum_{i,j} \frac{E_i E_j}{E_{vis}^2} \int_{x^*}^{x \frac{\delta y}{2}} \widetilde{\delta}(\chi \cdot \chi_{ij})$                                                                     | ľ                  |                              |           | $(resummed)  O(\alpha_s^2)$         |
| Asymmetry of<br>EEC           | $AEEC(\chi) = EEC(\pi-\chi) - EEC(\chi)$                                                                                                                                                                               | ľ                  | π20 π/                       | 20 π/2    | $O(\alpha_s^2)$                     |
| Differential<br>2-jet rate    | $D_2(y) = \frac{R_2(y-\Delta y) - R_2(y)}{\Delta y}$                                                                                                                                                                   |                    |                              |           | (resummed)<br>O( $\alpha_s^2$ )     |

#### taken from S. Bethke, hep-ex/0001023



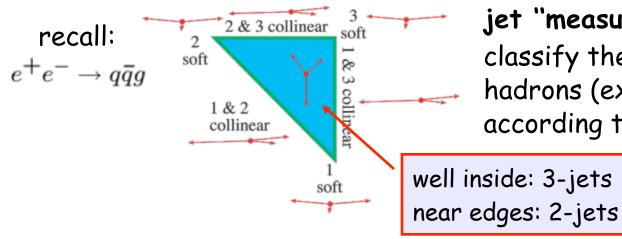


But what is a jet exactly?

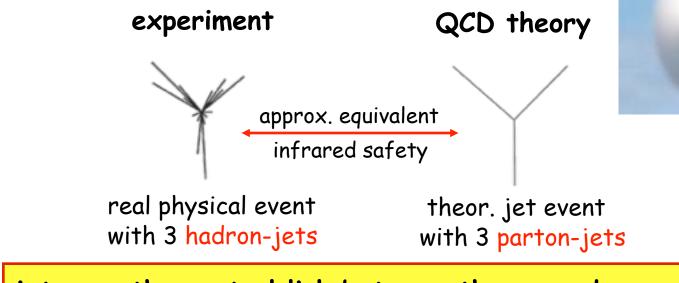


jets are the central link between theory and experiment

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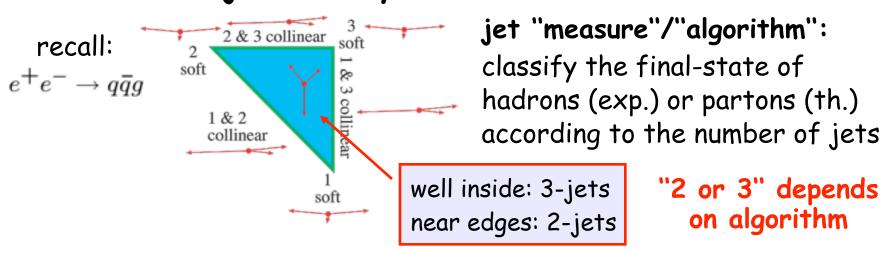


jet "measure"/"algorithm": classify the final-state of hadrons (exp.) or partons (th.) according to the number of jets



jets are the central link between theory and experiment

### But what is a jet exactly?

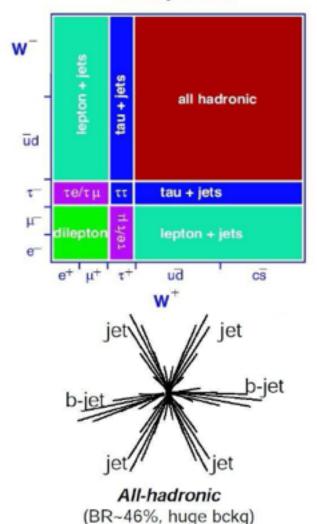


# jets – the central link between theory and experiment

#### input to almost all analyses at the LHC:

BSM & Higgs searches, top physics, PDF analyses, MC validation, ...

tt decay modes

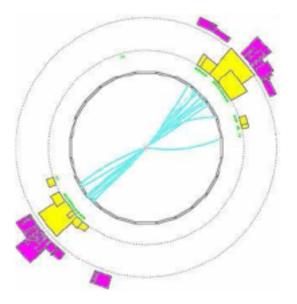


#### heavy objects have multi-jet final-states

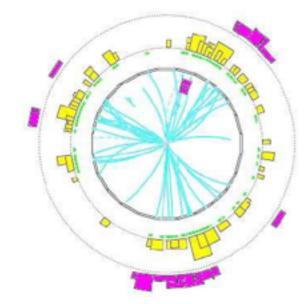
- 10<sup>7</sup> top-antitop pairs for 10 fb<sup>-1</sup>!
- vast number of QCD multi-jets:

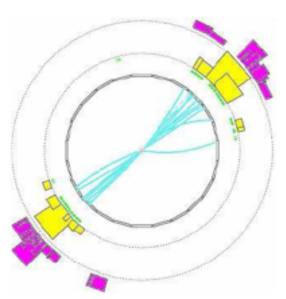
| # jets | $\#$ events for 10 fb $^{-1}$ |
|--------|-------------------------------|
| З      | $9 \cdot 10^8$                |
| 4      | $7 \cdot 10^7$                |
| 5      | $6 \cdot 10^6$                |
| 6      | $3\cdot 10^5$                 |
| 7      | $2 \cdot 10^4$                |
| 8      | $2 \cdot 10^3$                |

tree level estimates: Draggiotis, Kleiss, Papadopoulos p<sub>T</sub>(jet) > 60 GeV, θ<sub>ij</sub> > 30deg., |y<sub>ij</sub>| < 3



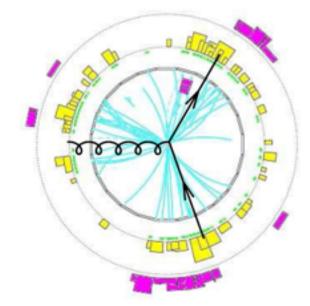
```
clearly (?) a 2-jet event
```

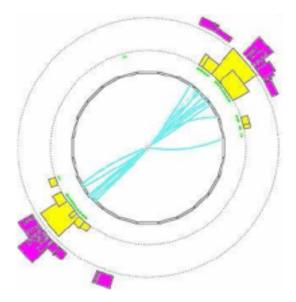




how many jets do you count?

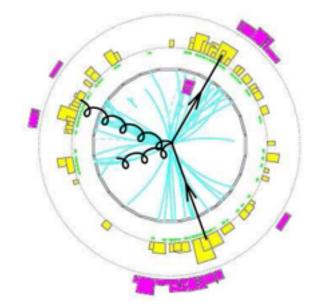
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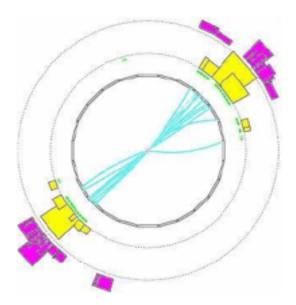




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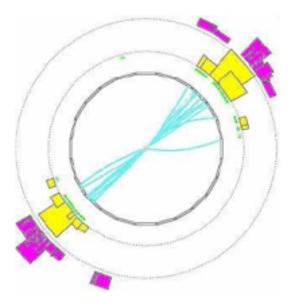
clearly (?) a 2-jet event

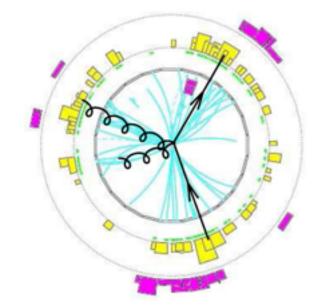




how many jets do you count?

clearly (?) a 2-jet event





clearly (?) a 2-jet event

how many jets do you count?

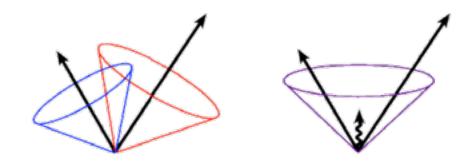
the "best" jet definition does not exist - construction is unavoidably ambiguous basically two issues:

- which particles/partons get put together in a jet  $\rightarrow$  jet algorithm
- how to combine their momenta  $\rightarrow$  recombination scheme

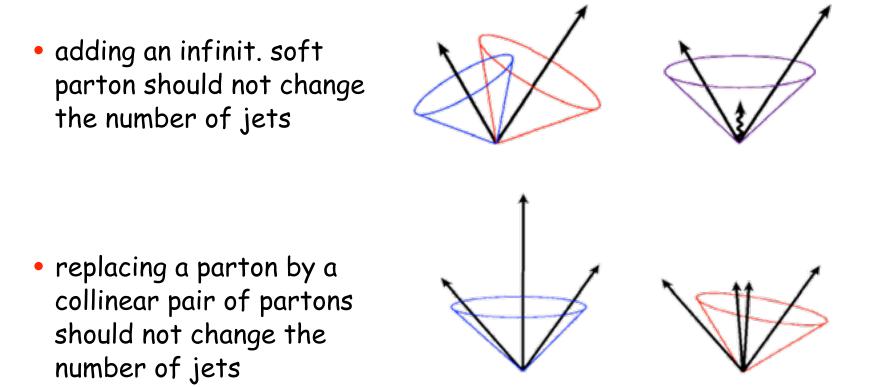
projection to jets should be resilient to QCD & detector effects

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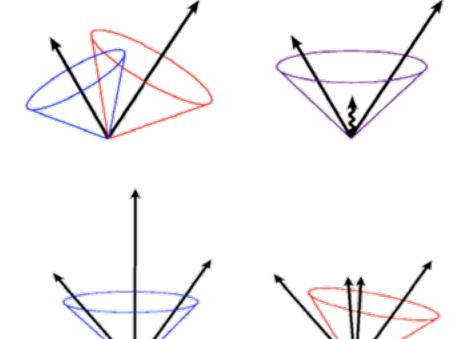


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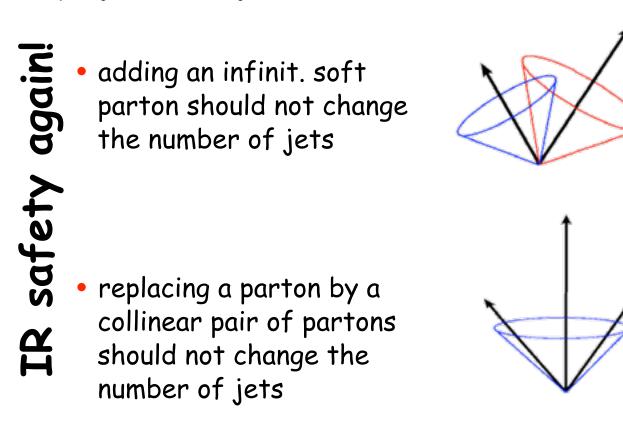


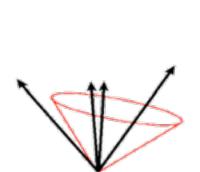
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IR safety again! adding an infinit. soft parton should not change the number of jets replacing a parton by a collinear pair of partons should not change the number of jets



projection to jets should be resilient to QCD & detector effects

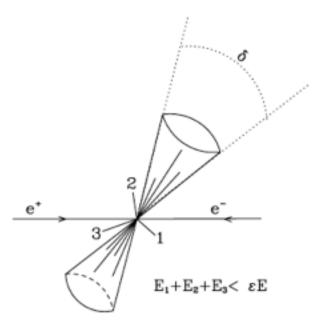




(anti-)  $k_T$  algorithms are the method of choice these days

Cacciari, Salam, Soyez (FastJet tool)

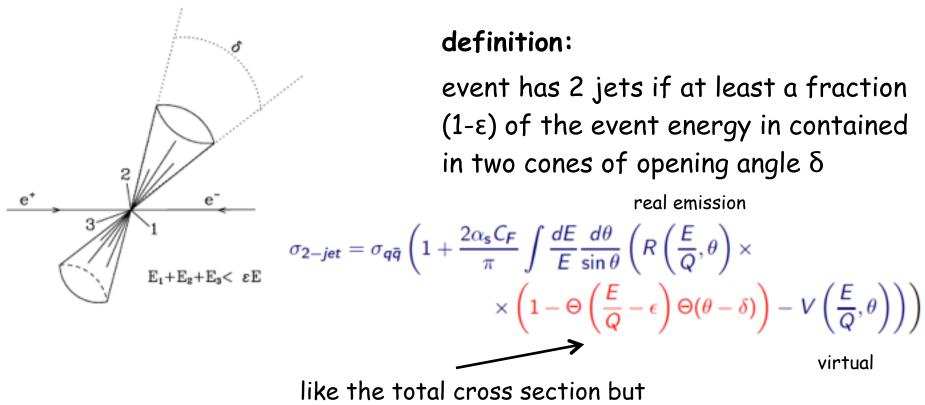
# 1st jet definition: Sterman and Weinberg



#### definition:

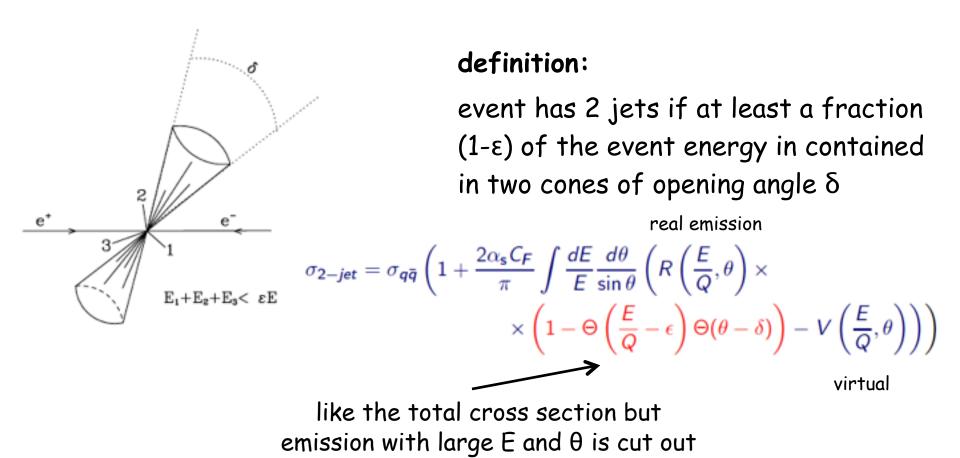
event has 2 jets if at least a fraction (1- $\epsilon$ ) of the event energy in contained in two cones of opening angle  $\delta$ 

# 1st jet definition: Sterman and Weinberg



emission with large E and  $\theta$  is cut out

# 1st jet definition: Sterman and Weinberg



find:

$$\sigma = \sigma_0 \left( \mathbf{1} + \frac{\mathbf{2}\alpha_s \mathbf{C_F}}{\pi} \ln \epsilon \ln \delta^2 \right)$$

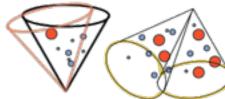
 if ε and/or δ become too small the results makes no sense (spoils KLN cancellation)

# classes of jet algorithms

there are many algorithms to choose from! basically two classes: "k<sub>T</sub>-type" or "cone"

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cone type

long. boost invariant cone size

 $R = \sqrt{\Delta \eta^2 + \Delta \phi^2}$ 

#### top-down approach

cluster particles according to their distance in coordinate space

put cones along dominant direction of energy flow

potential problems with IR safety

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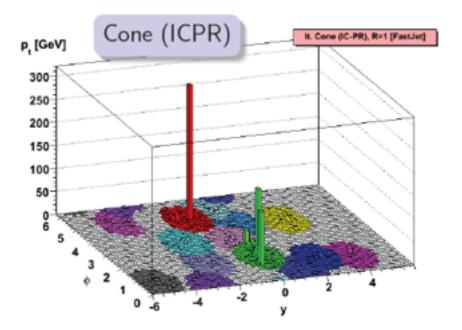
#### bottom-up approach

 $\mathbf{k}_{\mathrm{T}}$  type

cluster particles according to their distance in momentum space

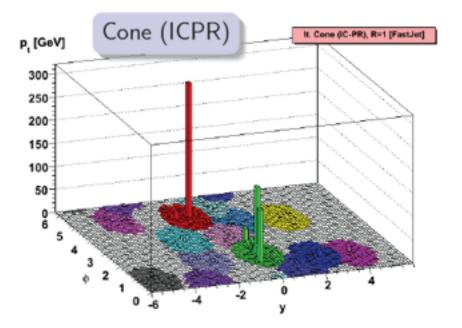
Cone jet

undo branchings occured in the perturbative QCD evolution: e.g., pair particles with the smallest relative k<sub>T</sub>



most cone algorithms produce circular jets in y-Φ plane loved by experimentalists

taken from G. Salam

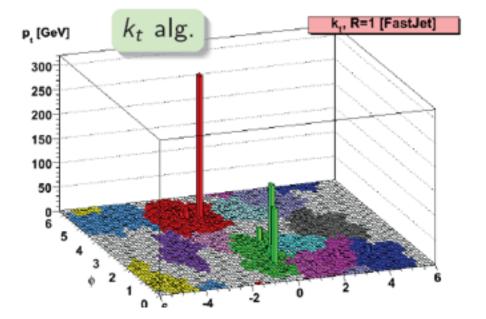


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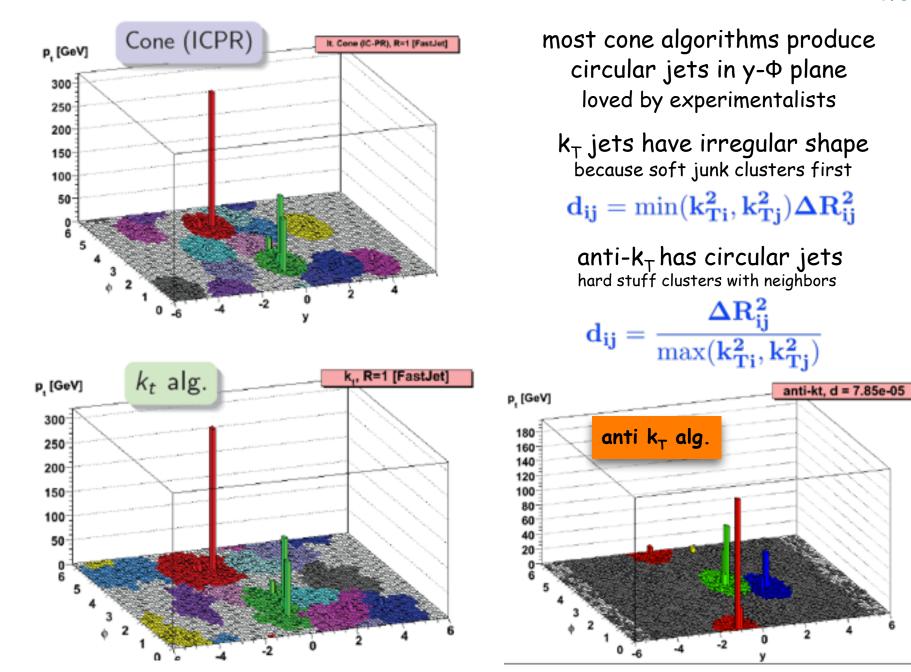
taken from G. Salam

k<sub>T</sub> jets have irregular shape because soft junk clusters first

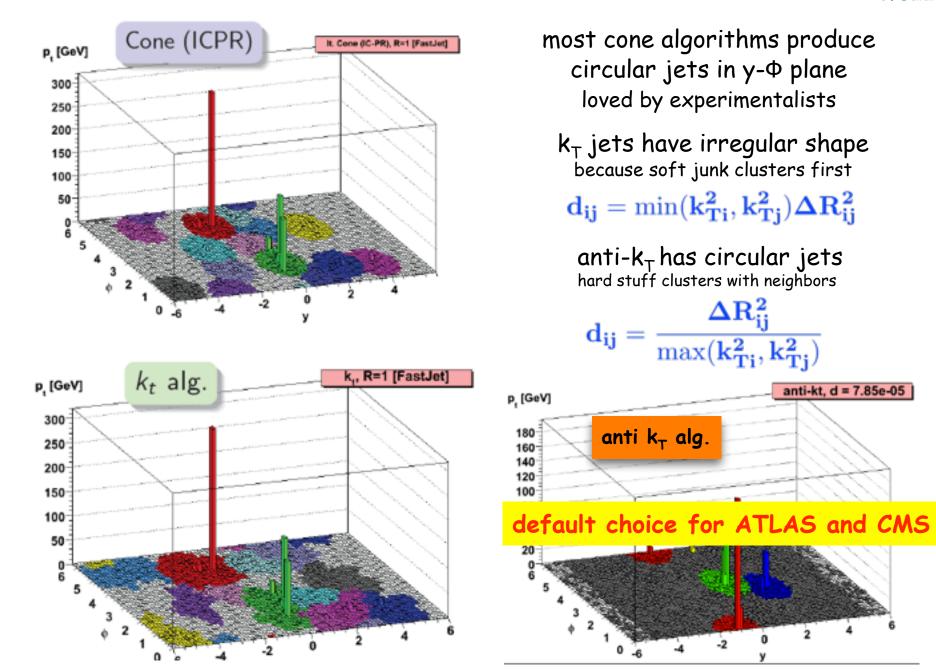
 $\mathbf{d_{ij}} = \min(\mathbf{k_{Ti}^2}, \mathbf{k_{Tj}^2}) \Delta \mathbf{R_{ij}^2}$ 









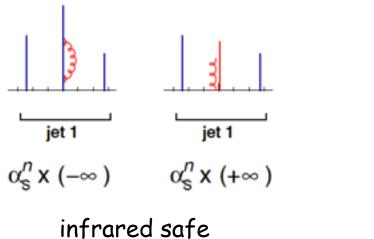


# jets – final remarks

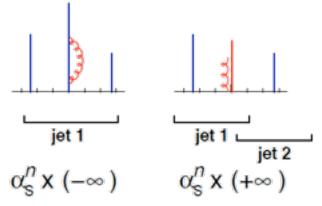
- n-jet vs. (n+1)-jet rate depends on algorithm  $\rightarrow$  have to choose the same jet definition in exp. and theory
- have to be careful when comparing between different experiments or experiment and theory (often different jet algorithms!)
- many widely used jet definitions are NOT IR safe! extensive study by Salam, Soyez, JHEP 0705:086,2007

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- have to be careful when comparing between different experiments or experiment and theory (often different jet algorithms!)
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   extensive study by Salam, Soyez, JHEP 0705:086,2007
- use of non IR safe definition invalidates pQCD approach

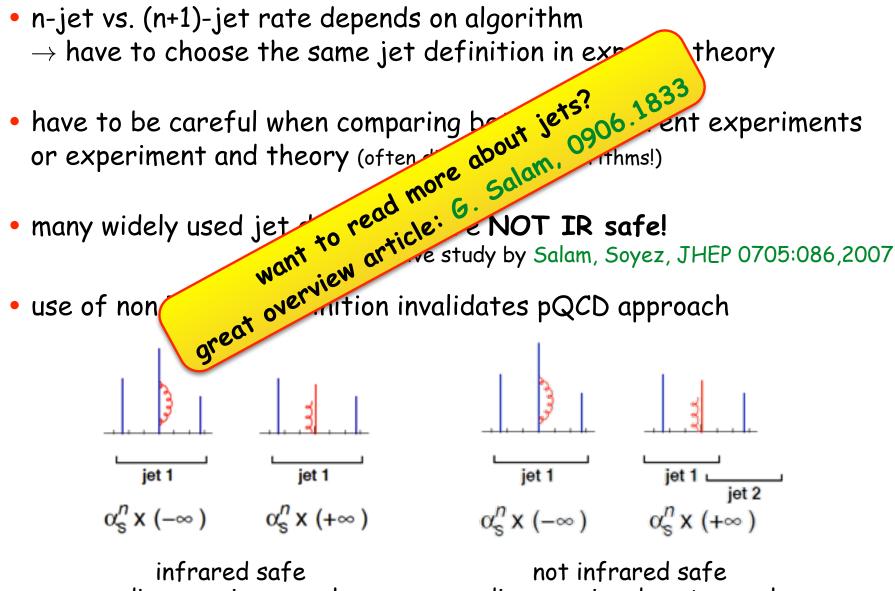


divergencies cancel



not infrared safe divergencies do not cancel

# jets – final remarks



divergencies cancel

divergencies do not cancel

# latest achievement: $e^+e^- \rightarrow 3$ jets at NNLO

Gehrmann-De Ridder, Gehrmann, Glover, Heinrich; Weinzierl

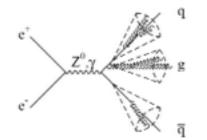
e Z<sup>o</sup> y mainten g

up to 7 jets in NLO !! leading color approx Becker et al., 1111.1733

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- numerous IR singularities to identify and cancel

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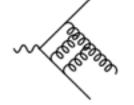
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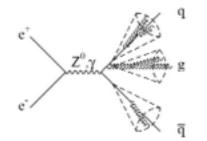
2-loop matrix elements (3 partons)



explicit IR poles from loop integrals

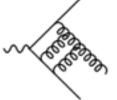
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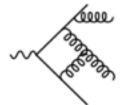
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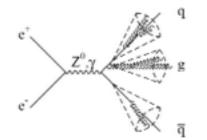
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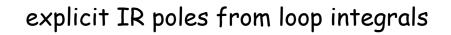
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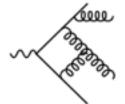
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1-loop matrix elements (4 partons)



explicit IR poles from loop integrals implicit IR poles from 1-unresolved radiation soft, collinear

tree level matrix elements (5 partons)

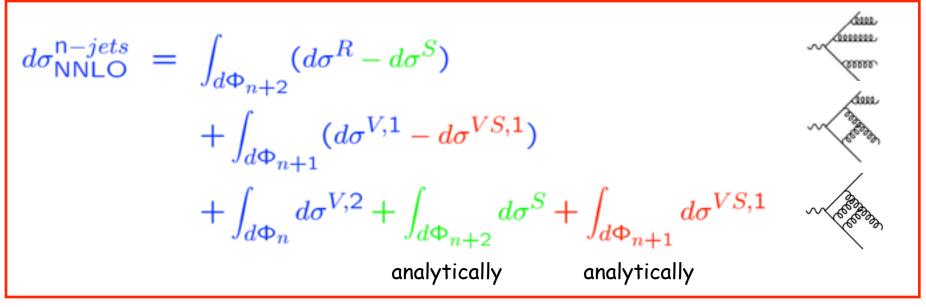
implicit IR poles from 2-unresolved radiation double soft, soft/collinear, double single collinear, triple collinear

#### structure of NNLO cross section

- complicated phase space (d $\Phi$ ) integrations done with numerical (MC) methods
- different strategies for IR cancellations, most common: subtraction method

tricky issue: find NNLO subtraction functions which

- approximate cross section in all singular limits
- are sufficiently simple to be integrated analytically



each line above is free of IR poles and numerically finite; implemented in EERAD3 code 1402.4140

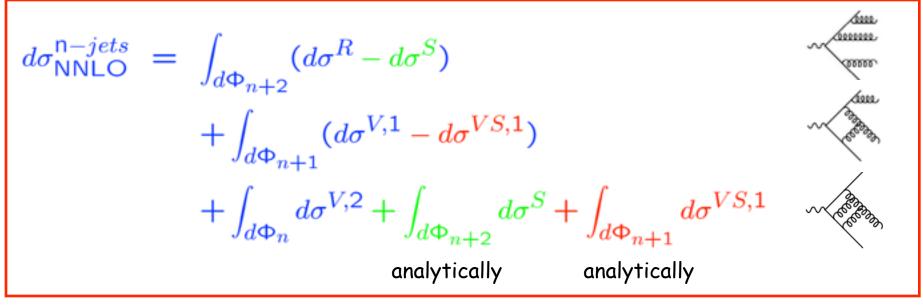
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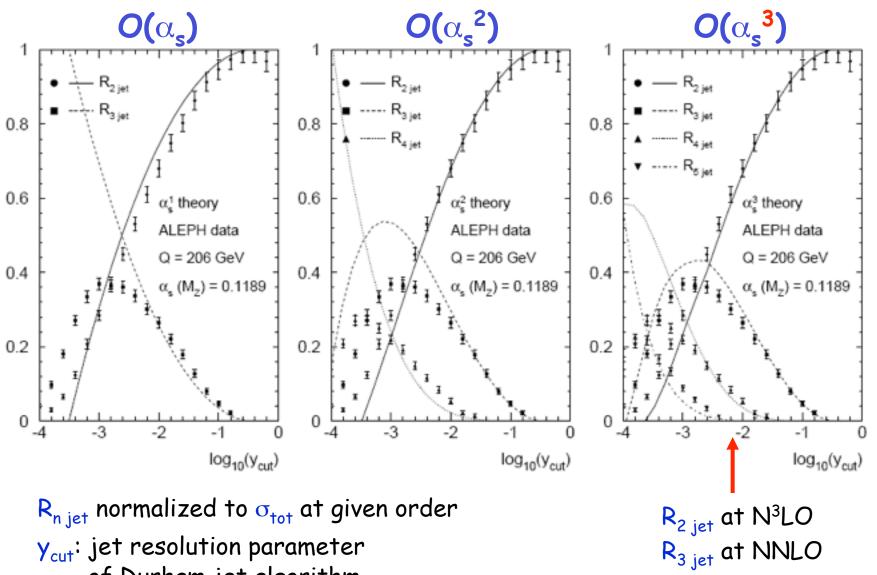


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crucial step towards full NNLO corrections for 2 ightarrow 2 QCD processes

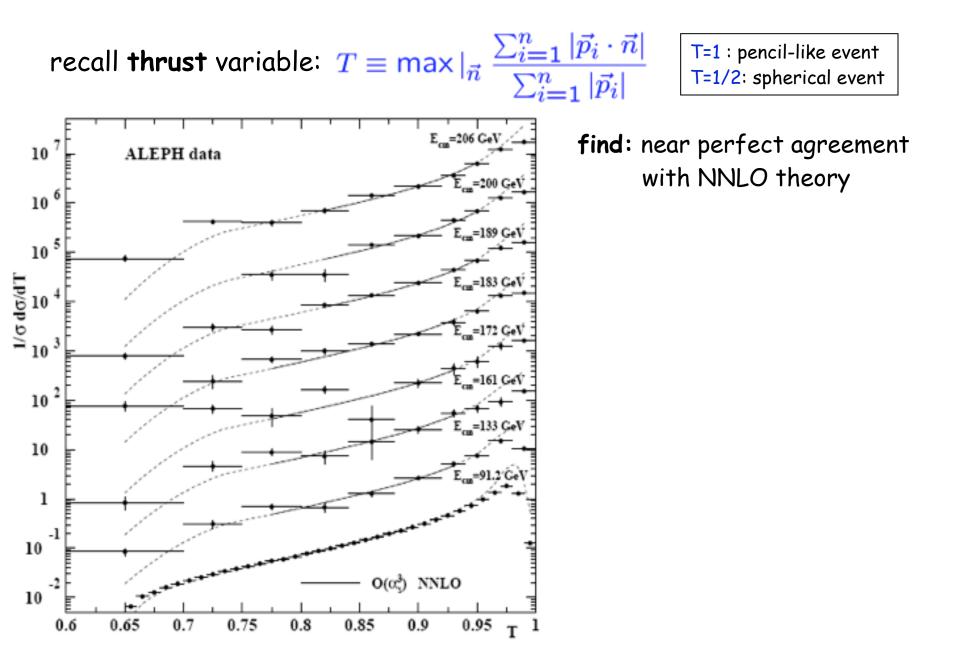
impact on e<sup>+</sup>e<sup>-</sup> jet rates

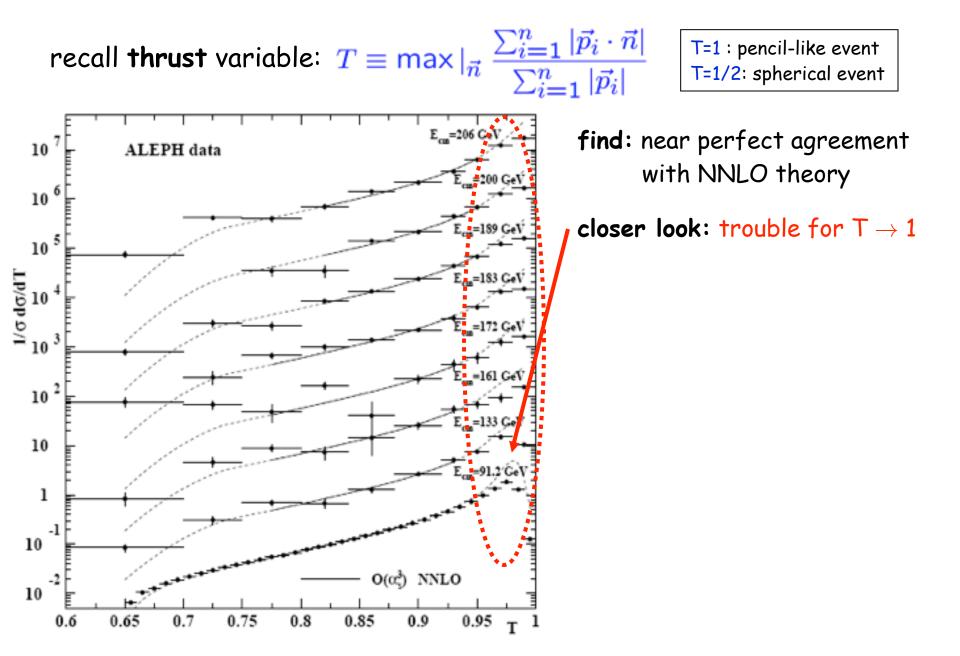
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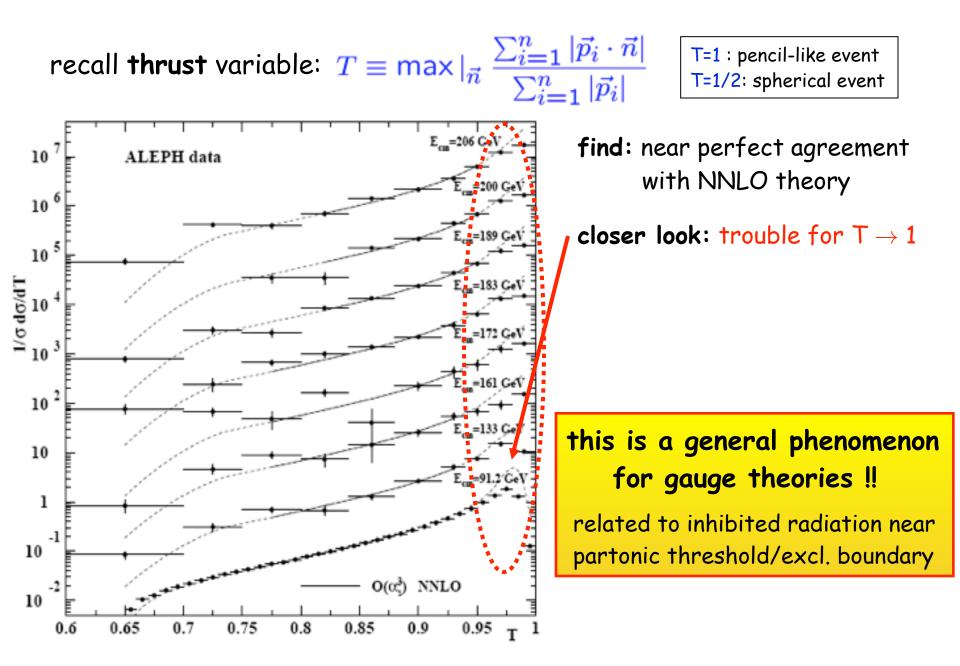


of Durham jet algorithm

 $R_{4,jet}$  at NLO R<sub>5 jet</sub> at LO







#### What goes wrong for thrust?

- T=1 corresponds to 2-parton final state (just two back-to-back jets)
- . if  $T{\rightarrow}1$  only soft/collinear gluons can be emitted ("inhibited radiation") in events with an extra gluon
- IR singularities cancel between real emissions and loop corrections but leave large logarithms behind in each order of α<sub>s</sub>
   here: (α<sub>s</sub> ln<sup>2</sup> [1-T])<sup>n</sup> → spoil convergence of pQCD series even if a<sub>s</sub><<1</li>

Can this be cured?

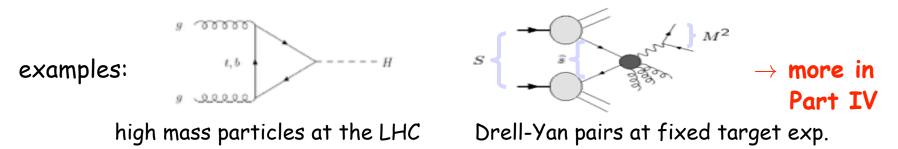
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- T=1 corresponds to 2-parton final state (just two back-to-back jets)
- . if  $T{\rightarrow}1$  only soft/collinear gluons can be emitted ("inhibited radiation") in events with an extra gluon
- IR singularities cancel between real emissions and loop corrections but leave large logarithms behind in each order of α<sub>s</sub>
   here: (α<sub>s</sub> ln<sup>2</sup> [1-T])<sup>n</sup> → spoil convergence of pQCD series even if a<sub>s</sub><<1</li>

#### Can this be cured?

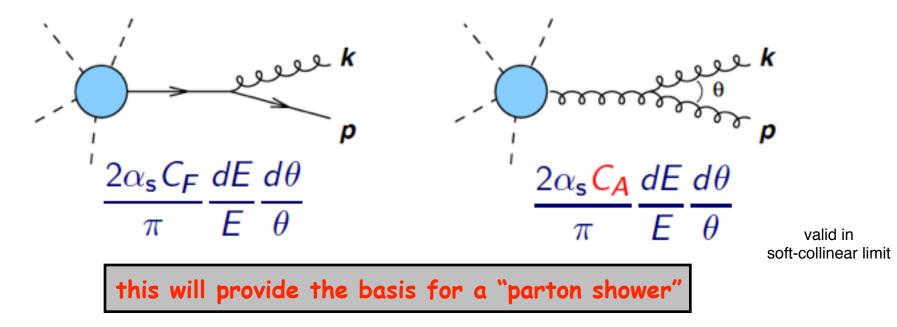
Yes! re-organize pQCD series to resum large logs to all orders Sterman; Catani, Trentadue; Laenen, Oderda, Sterman; Catani et al.; Sterman, Vogelsang; Kidonakis, Owens; ...

#### of great phenomenological relevance in hadronic processes



#### recap: idea behind parton shower MC programs

- we have seen that emission of soft/collinear partons is favored
- we know exactly how and when it occurs (process-independent)



 main idea: seek for an approx. result such that soft/collinear enhanced terms are included to all orders emissions are probabilistic (as needed to set up an event generator)

the possible way to proceed is to ask

"what is the probability of NOT radiating a gluon above a certain scale  $k_{\rm T}$ ?"

$$P(\text{no emission above } k_t) \sim 1 - \frac{2\alpha_s C_F}{\pi} \int \frac{dE}{E} \int \frac{d\theta}{\theta} \Theta(E\theta - k_t)$$

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$$= \text{generalized to all orders by exponentiation (Sudakov exponent)}$$

$$\Delta(k_t, Q) \simeq \exp\left[-\frac{2\alpha_s C_F}{\pi} \int^Q \frac{dE}{E} \int^{\pi/2} \frac{d\theta}{\theta} \Theta(E\theta - k_t)\right] \qquad \text{bounded between } 0 \text{ and } 1 \text{ (probability)}$$

$$(\text{here: some simplifying assumptions !!)}$$

$$= \text{probability distribution for gluon emission given by} \quad \frac{dP}{dk_{t1}} = \frac{d\Delta(k_{t1}, Q)}{dk_{t1}}$$

the possible way to proceed is to ask

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• used in MC to generate subsequent ordered branchings, e.g.,  $k_{T1} > k_{T2} > ...$ 

stop at some small cut-off scale and then use some model to produce hadrons

### some popular parton shower programs

| PYTHIA   | T. Sjöstrand et al.                                   |   |
|----------|-------------------------------------------------------|---|
|          | http://home.thep.lu.se/~torbjorn/Pythia.html          |   |
| HERWIG   | G. Corcella et al.                                    |   |
|          | http://hepwww.rl.ac.uk/theory/seymour/herwig/         | , |
| HERWIG++ | S. Gieseke et al.                                     |   |
|          | http://projects.hepforge.org/herwig/                  | 4 |
| SHERPA   | F. Krauss et al.                                      | - |
|          | http://projects.hepforge.org/sherpa/dokuwiki/doku.php |   |
| ISAJET   | H. Baer et al.                                        |   |
|          | http://www.nhn.ou.edu/~isajet/                        |   |

- can fail in high-multiplicity events or when large-angle emissions are relevant
- do better than fixed order calculations at lowish scales
- matching with NLO matrix elements well advanced: MC@NLO, POWHEG, ...

# summary so far

pQCD cannot give all the answers but it does cover a lot of ground despite the "long-distance problem"

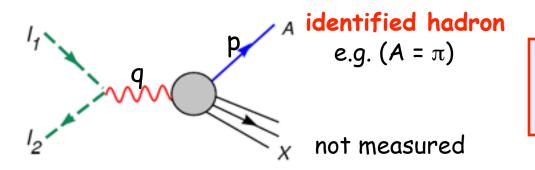
# summary so far

pQCD cannot give all the answers but it does cover a lot of ground despite the "long-distance problem"

the concept of factorization will allow us to compute cross sections for a much wider class of processes than considered so far (involving hadrons in the initial and/or final state) LHC, RHIC, COMPASS, ..., EIC, ...

#### identified hadrons: a new "long distance problem"

consider the one-particle inclusive cross section:

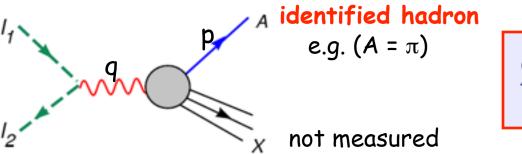


$$\frac{d\sigma(e^+e^- \to \pi + X)}{dE_\pi}$$

## not infrared safe by itself!

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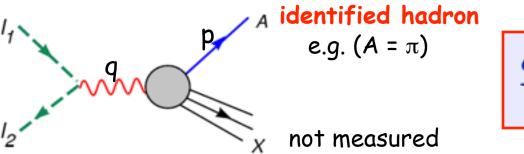
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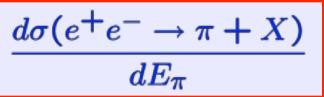
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problem: sensitivity to long-distance physics related to particle emission along with identified/observed hadrons (leads to uncanceled singularities -> meaningless)

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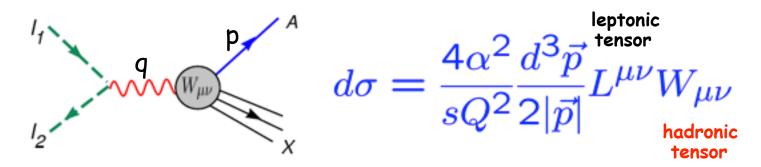
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general feature of QCD processes with observed (=identified) hadrons in the initial and/or final state

#### factorization

**strategy**: try to factorize the physical observable into a calculable infrared safe and a non-calculable but universal piece

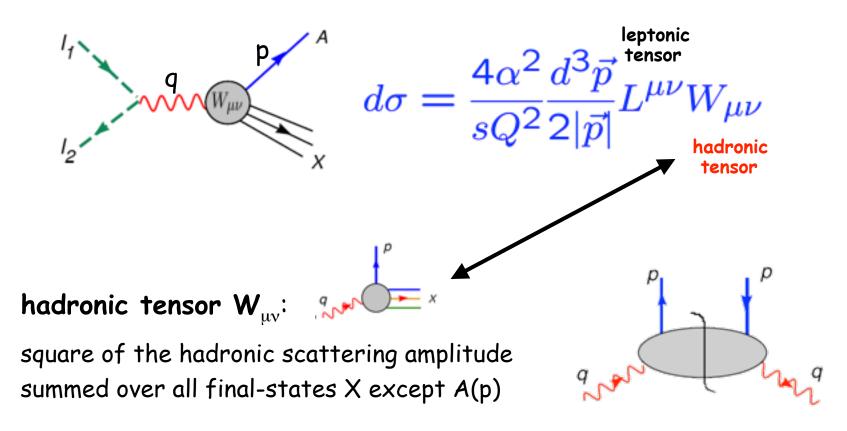
how does it work?



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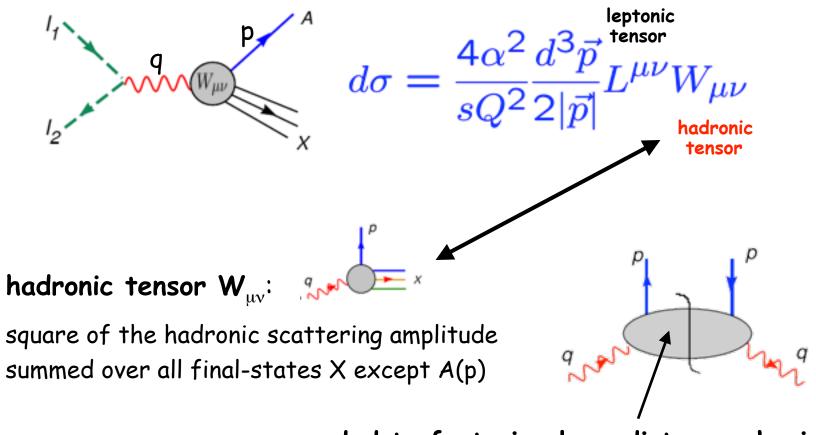
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needed to factorize long-distance physics

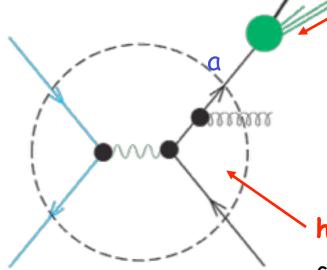
### concept of factorization - pictorial sketch

**factorization** = isolating and absorbing infrared singularities accompanying observed hadrons

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pictorial sketch:



# $\succ$ fragmentation functions $D_a^{ar{h}}$

contains all **long-distance** interactions hence not calculable but universal

#### physical interpretation:

probability to find a hadron carrying a certain momentum of parent parton

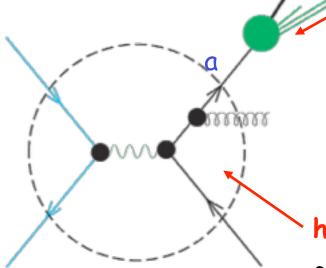
## 

contains only **short-distance** physics amenable to pQCD calculations

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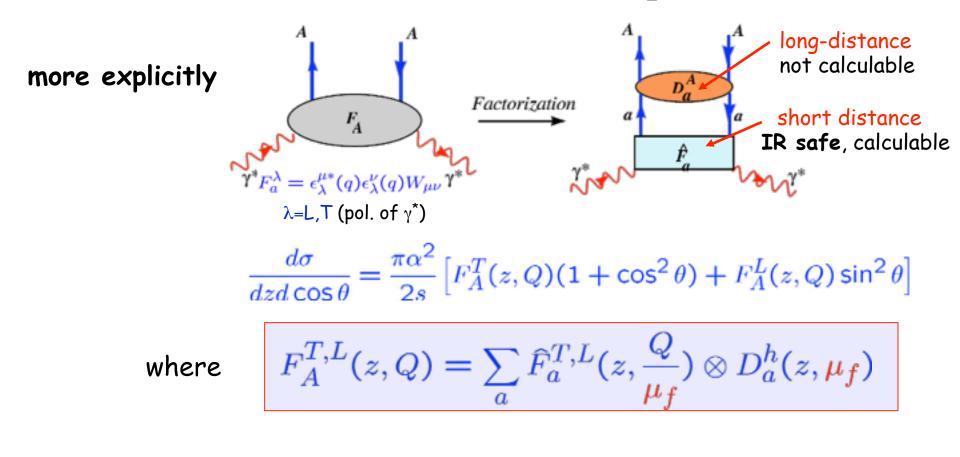
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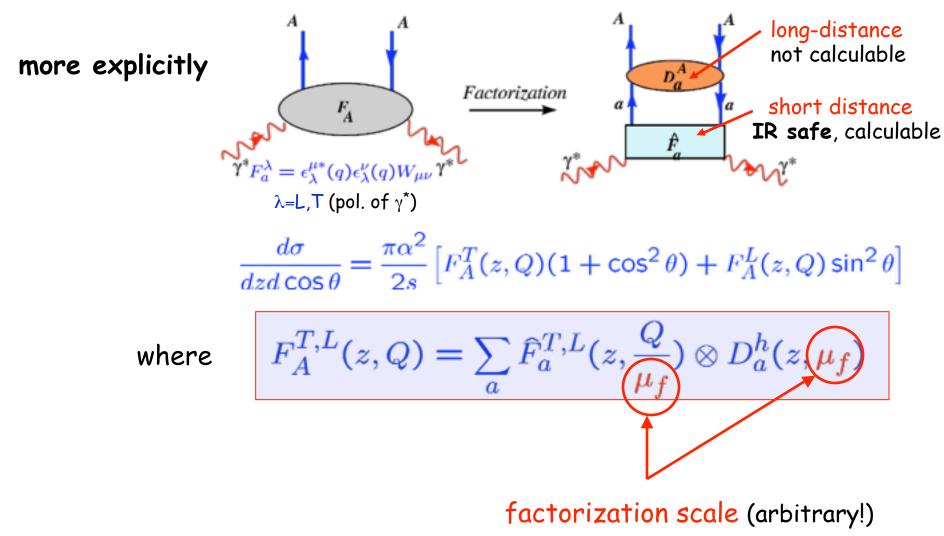
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## hard scattering $\widehat{\widehat{F}}_{a}$

contains only **short-distance** physics amenable to pQCD calculations

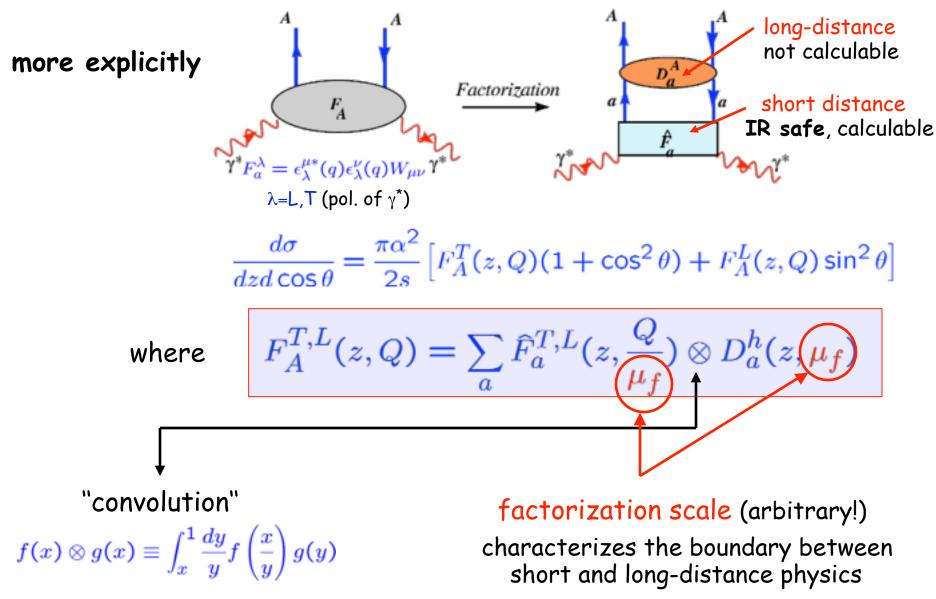
aside: fragmentation fcts. play an important role in learning about nucleon (spin) structure from semi-inclusive DIS data by COMPASS & HERMES or from hadron production at RHIC



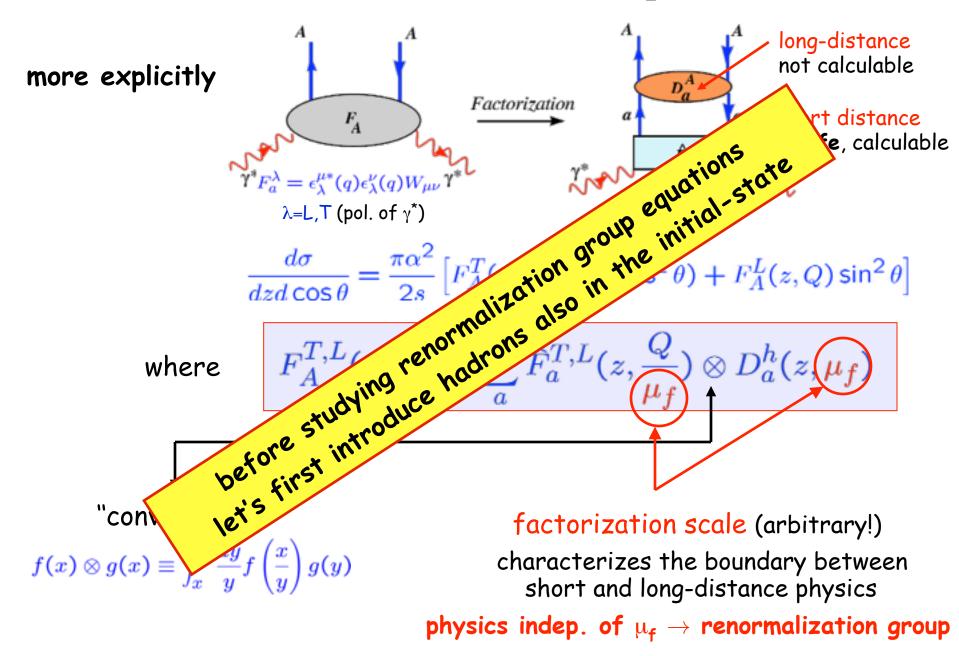


characterizes the boundary between short and long-distance physics

physics indep. of  $\mu_{\text{f}} \rightarrow$  renormalization group



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#### take home message for part II THE QCD TOOLBOX



QCD is a non-Abelian gauge theory: gluons are self-interacting
 –> asymptotic freedom (large Q), confinement (small Q)

- QCD calculations are singular when any two partons become collinear or a gluon becomes soft; basis for parton shower MCs
- choose infrared/collinear safe observables for comparison between experiment and perturbative QCD
- jets (= cluster of partons): best link between theory and exp.; needs a proper IR safe jet definition in theory and experiment
- infrared cancellation leaves large logarithms behind which become important in certain regions of phase-space —> all-order resummations
- factorization allows to deal with hadronic processes introduces arbitrary scale —> leads to RGEs





early microscopes

the World's most powerful microscopes

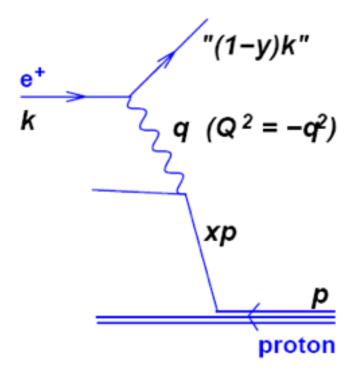
# Part III

inward bound: "femto-spectroscopy"

QCD initial state, partons, DIS, factorization, renormalization group, hadron-hadron collisions

#### partons in the initial state: the DIS process

start with the simplest process: deep-inelastic scattering



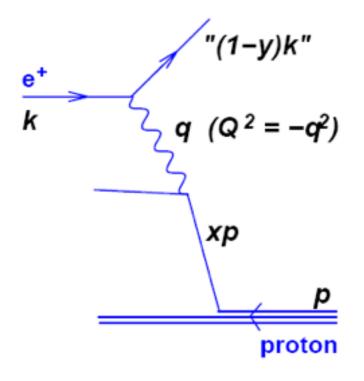
relevant kinematics:

$$x = \frac{Q^2}{2p \cdot q}$$
  $y = \frac{p \cdot q}{p \cdot k}$   $Q^2 = xys$ 

- Q<sup>2</sup>: photon virtuality  $\leftrightarrow$  resolution r~1/Q at which the proton is probed
- x: long. momentum fraction of struck parton in the proton
- y: momentum fraction lost by electron in the proton rest frame

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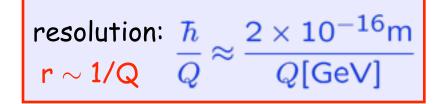


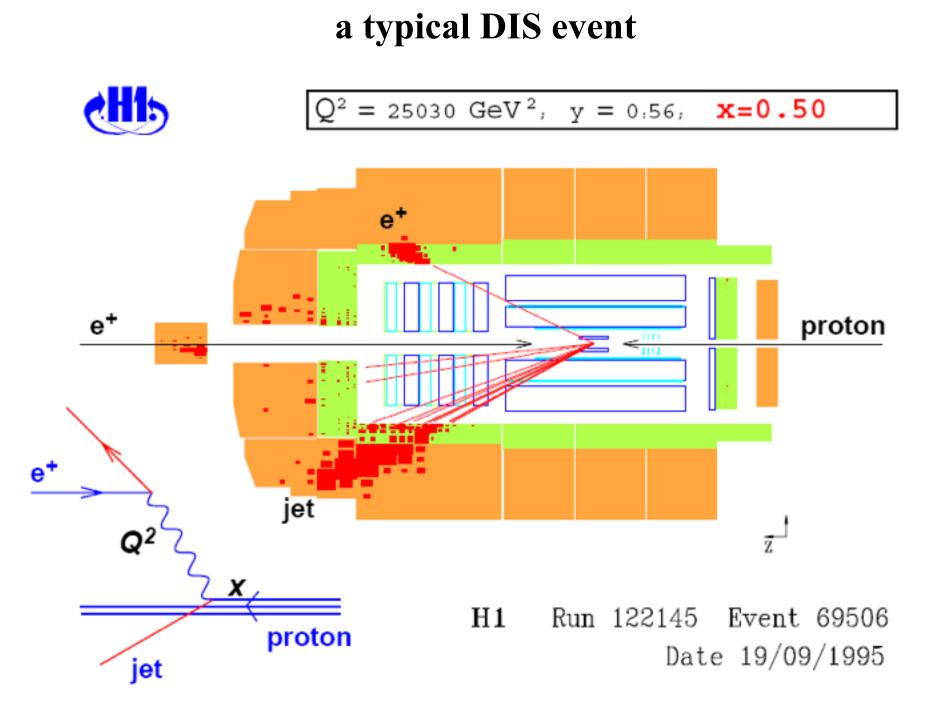
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"deep-inelastic": 
$$Q^2 \gg 1 \text{ GeV}^2$$
  
"scaling limit":  $Q^2 \rightarrow \infty$ , x fixed



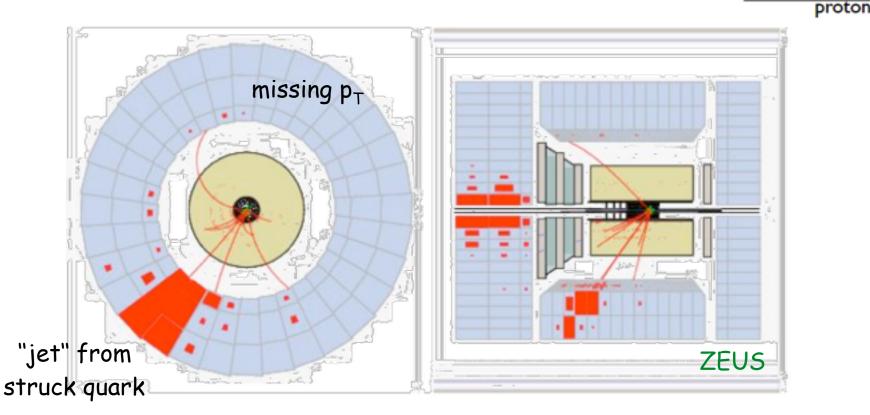


#### a charged current DIS event

 $\nu_{\mu}$ 

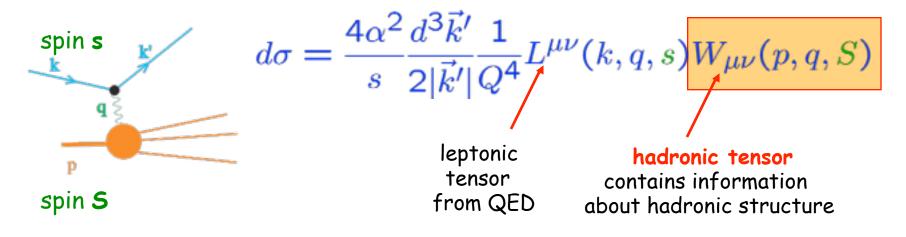
 $W^+$ 

a charged current event with W-boson-exchange (the electron turns into a neutrino which is "invisible")

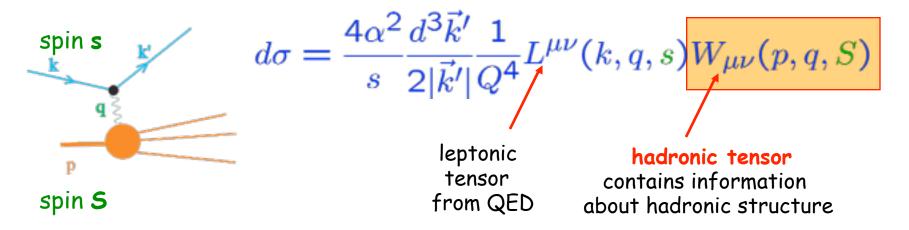


for simplicity we will restrict ourselves to photon exchange though

electroweak theory tells us how the virtual vector boson (here  $\gamma^*$ ) couples:



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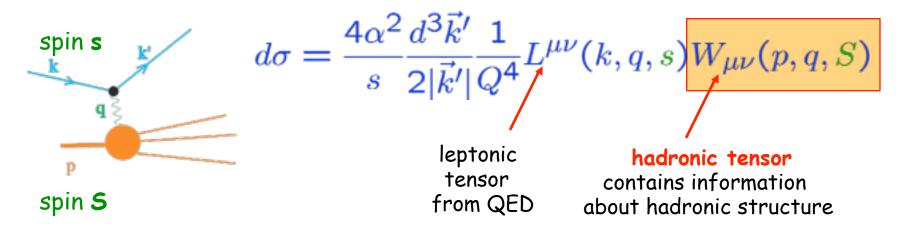
parity & Lorentz inv., hermiticity  $W^{\mu\nu} = W^{\mu\nu}$ , current conservation  $q_{\mu}W^{\mu\nu} = 0$  dictate:  $W^{\mu\nu}(P, q, S) = \frac{1}{2} \int d^4z \, e^{iq \cdot z} \langle P, S | J_{\nu}(z) J_{\nu}(0) | P, S \rangle$ 

$$(I,q,S) = \frac{1}{4\pi} \int u z e^{-i\varphi} \langle I, S | J_{\mu}(z) J_{\nu}(0) | I, S \rangle$$

$$= \left( -g^{\mu\nu} + \frac{q^{\mu}q^{\nu}}{q^2} \right) F_1(x,Q^2) + \left( P^{\mu} - \frac{P \cdot q}{q^2} q^{\mu} \right) \left( P^{\nu} - \frac{P \cdot q}{q^2} q^{\nu} \right) F_2(x,Q^2)$$

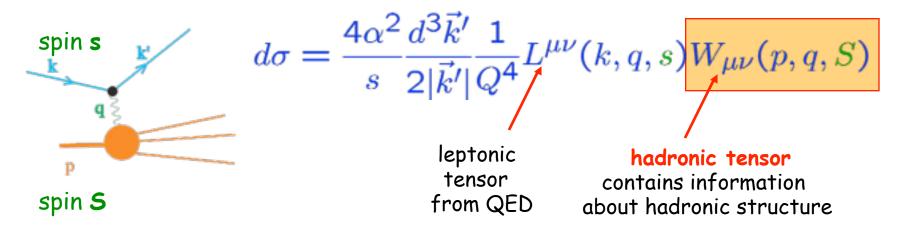
$$+ i M \varepsilon^{\mu\nu\rho\sigma} q_{\rho} \left[ \frac{S_{\sigma}}{P \cdot q} g_1(x,Q^2) + \frac{S_{\sigma}(P \cdot q) - P_{\sigma}(S \cdot q)}{(P \cdot q)^2} g_2(x,Q^2) \right]$$

electroweak theory tells us how the virtual vector boson (here  $\gamma^*$ ) couples:



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$$\begin{split} \mathcal{W}^{\mu\nu}(P,q,S) &= \frac{1}{4\pi} \int d^4 z \, \mathrm{e}^{iq \cdot z} \, \langle P,S | \, J_\mu(z) \, J_\nu(0) \, | P,S \rangle \\ &= \left( -g^{\mu\nu} + \frac{q^\mu q^\nu}{q^2} \right) F_1(x,Q^2) + \left( P^\mu - \frac{P \cdot q}{q^2} q^\mu \right) \left( P^\nu - \frac{P \cdot q}{q^2} q^\nu \right) F_2(x,Q^2) \\ &+ i \, M \, \varepsilon^{\mu\nu\rho\sigma} q_\rho \Biggl[ \frac{S_\sigma}{P \cdot q} g_1(x,Q^2) + \frac{S_\sigma(P \cdot q) - P_\sigma(S \cdot q)}{(P \cdot q)^2} g_2(x,Q^2) \Biggr] \end{split}$$

pol. structure fcts.  $g_{1,2}$  - measure W(P,q,S) - W(P,q,-S) !





# SLAC-MIT experiment of 1969

two unexpected results:

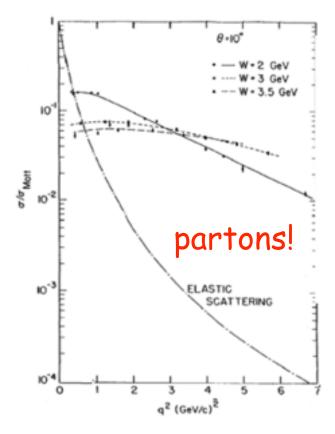




# SLAC-MIT experiment of 1969



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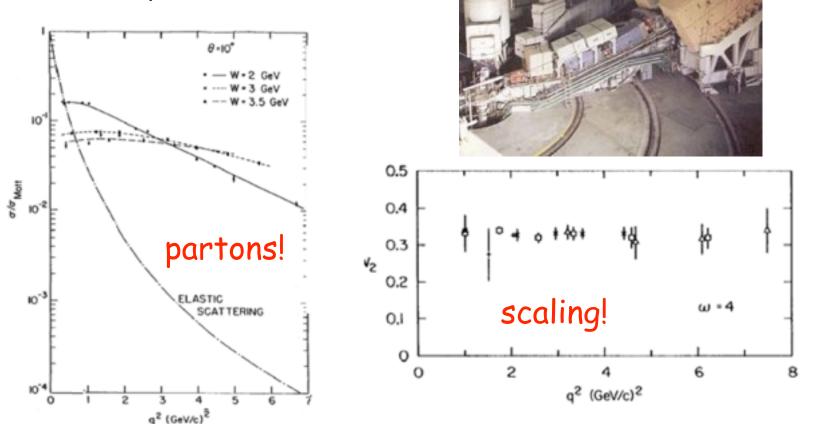




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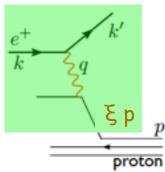


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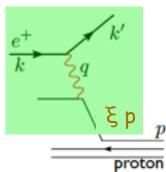
birth of the pre-QCD parton model

let's do a quick calculation: consider electron-quark scattering



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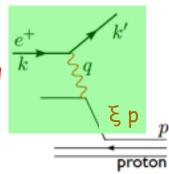
find 
$$\overline{\sum} |\mathcal{M}|^2 = 2e_q^2 e^4 \frac{\hat{s}^2 + \hat{u}^2}{\hat{t}^2}$$



let's do a quick calculation: consider electron-quark scattering

 $\begin{array}{ll} \mbox{find} & \overline{\sum} |\mathcal{M}|^2 = 2 e_q^2 e^4 \, \frac{\hat{s}^2 + \hat{u}^2}{\hat{t}^2} & \mbox{with the usual} & \hat{s} = (k + p_q)^2 \\ \mbox{Mandelstam's} & \hat{t} = (k - k')^2 \end{array}$ 

 $\hat{\mathbf{u}} = (\mathbf{p}_{\mathbf{q}} - \mathbf{k}')^2$ 



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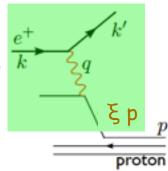
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with the usual Mandelstam's

$$\hat{\mathbf{s}} = (\mathbf{k} + \mathbf{p}_{\mathbf{q}})^{2}$$

$$\hat{\mathbf{t}} = (\mathbf{k} - \mathbf{k}')^{2}$$

$$\hat{\mathbf{u}} = (\mathbf{p}_{\mathbf{q}} - \mathbf{k}')^{2}$$



next: express by usual DIS variables

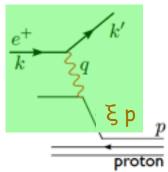
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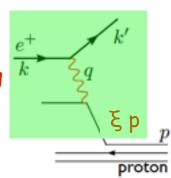
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and use the massless 2->2 cross section

 $\frac{\mathrm{d}\sigma}{\mathrm{d}t} = \frac{1}{16\pi\hat{s}^2} \overline{\sum} |\mathcal{M}|^2$ 



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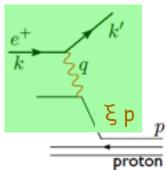
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$$\mathbf{\hat{t}} = (\mathbf{k} + \mathbf{p_q})$$
  
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 $\frac{\mathrm{d}\sigma}{\mathrm{d}t} = \frac{1}{16\pi\hat{s}^2} \sum |\mathcal{M}|^2$ 

to obtain

 $\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^2} = \frac{2\pi\alpha^2 \mathbf{e}_{\mathbf{q}}^2}{\Omega^4} [1 + (1 - \mathbf{y})^2]$ 

Mandelstam'

let's do a quick calculation: consider electron-quark scattering

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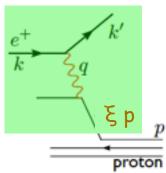
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 $\frac{\mathrm{d}\sigma}{\mathrm{d}t} = \frac{1}{16\pi\hat{s}^2} \overline{\sum} |\mathcal{M}|^2$ to obtain

next: use on-mass shell constraint

 $\mathbf{p}_{\mathbf{q}}^{\prime 2} = (\mathbf{p}_{\mathbf{q}} + \mathbf{q})^2 = \mathbf{q}^2 + 2\mathbf{p}_{\mathbf{q}} \cdot \mathbf{q}$ 

with the usual 
$$\hat{\mathbf{s}} = (\mathbf{k} + \mathbf{p}_{\mathbf{q}})^2$$
  
Mandelstam's  $\hat{\mathbf{t}} = (\mathbf{k} - \mathbf{k}')^2$   
 $\hat{\mathbf{u}} = (\mathbf{p}_{\mathbf{q}} - \mathbf{k}')^2$ 



$$\begin{aligned} \hat{\mathbf{s}} &= \xi \mathbf{Q}^2 / (\mathbf{x}\mathbf{y}) = \xi \, \mathbf{s} \\ \hat{\mathbf{t}} &= \mathbf{q}^2 = -\mathbf{Q}^2 \\ \hat{\mathbf{u}} &= \hat{\mathbf{s}} \, (\mathbf{y} - \mathbf{1}) \end{aligned}$$

$$\frac{\mathrm{d}\sigma}{\mathrm{d}\mathbf{Q}^2} = \frac{2\pi\alpha^2 \mathbf{e}_{\mathbf{q}}^2}{\mathbf{Q}^4} [\mathbf{1} + (\mathbf{1} - \mathbf{y})^2]$$

 $= -2\mathbf{p} \cdot \mathbf{q} (\mathbf{x} - \xi) = \mathbf{0}$ 

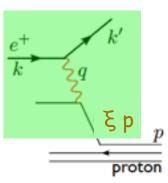
this implies that  $\xi$  is equal to Bjorken x

let's do a quick calculation: consider electron-quark scattering

find 
$$\overline{\sum} |\mathcal{M}|^2 = 2e_q^2 e^4 \frac{\hat{s}^2 + \hat{u}^2}{\hat{t}^2}$$

Mandelstam's

with the usual  $\hat{\mathbf{s}} = (\mathbf{k} + \mathbf{p}_{\mathbf{q}})^2$  $\hat{\mathbf{t}} = (\mathbf{k} - \mathbf{k}')^2$  $\hat{\mathbf{u}} = (\mathbf{p}_{\mathbf{u}} - \mathbf{k}')^2$ 



next: express by usual DIS variables

 $x = \frac{Q^2}{2n \cdot q}$   $y = \frac{p \cdot q}{n \cdot k}$   $Q^2 = xys$ find

 $\hat{\mathbf{s}} = \xi \mathbf{Q}^2 / (\mathbf{x}\mathbf{y}) = \xi \mathbf{s}$  $\hat{\mathbf{t}} = \mathbf{q}^2 = -\mathbf{Q}^2$  $\hat{\mathbf{u}} = \hat{\mathbf{s}} (\mathbf{y} - \mathbf{1})$ 

and use the massless 2->2 cross section

 $\frac{\mathrm{d}\sigma}{\mathrm{d}t} = \frac{1}{16\pi\hat{s}^2} \overline{\sum} |\mathcal{M}|^2$ to obtain  $\frac{\mathrm{d}\sigma}{\mathrm{d}\Omega^2} = \frac{2\pi\alpha^2 \mathbf{e}_{\mathbf{q}}^2}{\Omega^4} [1 + (1 - \mathbf{y})^2]$ 

next: use on-mass shell constraint

$$\mathbf{p_q'^2} = (\mathbf{p_q} + \mathbf{q})^2 = \mathbf{q^2} + 2\mathbf{p_q} \cdot \mathbf{q}$$

$$= -\mathbf{2p} \cdot \mathbf{q} \left( \mathbf{x} - \boldsymbol{\xi} \right) = \mathbf{0}$$

this implies that  $\xi$  is equal to Bjorken x

to obtain 
$$\frac{\mathrm{d}\sigma}{\mathrm{d}x\mathrm{d}\mathbf{Q}^2} = \frac{4\pi\alpha^2}{\mathbf{Q}^4} [\mathbf{1} + (\mathbf{1} - \mathbf{y})^2] \frac{1}{2} \mathbf{e}_{\mathbf{q}}^2 \delta(\mathbf{x} - \xi)$$

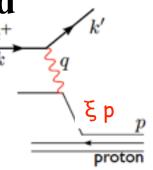
#### DIS in the naïve parton model cont'd

compare our result

$$\frac{\mathrm{d}\sigma}{\mathrm{d}\mathrm{x}\mathrm{d}\mathrm{Q}^2} = \frac{4\pi\alpha^2}{\mathrm{Q}^4} [\mathbf{1} + (\mathbf{1} - \mathbf{y})^2] \frac{1}{2} \mathbf{e}_{\mathbf{q}}^2 \delta(\mathbf{x} - \xi)$$

to what one obtains with the hadronic tensor (on the guark level)

$$\frac{\mathbf{d}^2\sigma}{\mathbf{d}\mathbf{x}\mathbf{d}\mathbf{Q}^2} = \frac{4\pi\alpha^2}{\mathbf{Q}^4} \left[ [\mathbf{1} + (\mathbf{1} - \mathbf{y})^2]\mathbf{F_1}(\mathbf{x}) + \frac{(\mathbf{1} - \mathbf{y})}{\mathbf{x}}(\mathbf{F_2}(\mathbf{x}) - \mathbf{2}\mathbf{x}\mathbf{F_1}(\mathbf{x})) \right]$$



# DIS in the naïve parton model cont'd

ξp

proton

compare our result

$$\frac{\mathrm{d}\sigma}{\mathrm{d}\mathrm{x}\mathrm{d}\mathrm{Q}^2} = \frac{4\pi\alpha^2}{\mathrm{Q}^4} [1 + (1-\mathrm{y})^2] \frac{1}{2} \mathrm{e}_\mathrm{q}^2 \delta(\mathrm{x}-\xi)$$

to what one obtains with the hadronic tensor (on the guark level)

$$\frac{d^2\sigma}{dxdQ^2} = \frac{4\pi\alpha^2}{Q^4} \left[ \left[ 1 + (1-y)^2 \right] F_1(x) + \frac{(1-y)}{x} (F_2(x) - 2xF_1(x)) \right]$$
  
and read off  
$$F_2 = 2xF_1 = xe_q^2 \,\delta(x-\xi) \frac{\text{Callan Gross relation}}{\text{reflects spin 1/2 nature of quarks}}$$

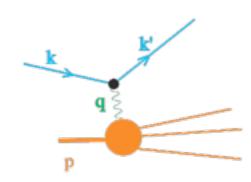
# **DIS** in the naïve parton model cont'd compare our result $\frac{\mathrm{d}\sigma}{\mathrm{d}\mathrm{x}\mathrm{d}\mathrm{Q}^2} = \frac{4\pi\alpha^2}{\mathrm{Q}^4} [1 + (1 - \mathrm{y})^2] \frac{1}{2} \mathrm{e}_\mathrm{q}^2 \delta(\mathrm{x} - \xi)$ protor to what one obtains with the hadronic tensor (on the guark level) $\frac{\mathrm{d}^2\sigma}{\mathrm{d}\mathbf{x}\mathrm{d}\mathbf{O}^2} = \frac{4\pi\alpha^2}{\mathbf{O}^4} \left[ [\mathbf{1} + (\mathbf{1} - \mathbf{y})^2] \mathbf{F_1}(\mathbf{x}) + \frac{(\mathbf{1} - \mathbf{y})}{\mathbf{x}} (\mathbf{F_2}(\mathbf{x}) - \mathbf{2}\mathbf{x}\mathbf{F_1}(\mathbf{x})) \right]$ $\mathbf{F_2} = \mathbf{2xF_1} = \mathbf{xe_q^2} \, \delta(\mathbf{x} - \xi) \qquad \begin{array}{c} \textbf{Callan Gross relation} \\ \textbf{reflects spin 1/2 nature of quarks} \end{array}$ and read off

proton structure functions then obtained by weighting the quark str. fct. with the parton distribution functions (probability to find a quark with momentum  $\xi$ )

$$\begin{split} \mathbf{F_2} &= 2\mathbf{x} \mathbf{F_1} = \sum_{\mathbf{q},\mathbf{q}'} \int_0^1 d\xi \, \mathbf{q}(\xi) \, \mathbf{x} \mathbf{e}_{\mathbf{q}}^2 \, \delta(\mathbf{x} - \xi) \\ &= \sum_{\mathbf{q},\mathbf{q}'} \mathbf{e}_{\mathbf{q}}^2 \, \mathbf{x} \, \mathbf{q}(\mathbf{x}) \\ &= \sum_{\mathbf{q},\mathbf{q}'} \mathbf{e}_{\mathbf{q}}^2 \, \mathbf{x} \, \mathbf{q}(\mathbf{x}) \end{split} \end{split} \\ \end{split}$$

# space-time picture of DIS

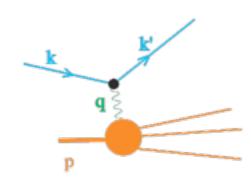
this can be best understood in a reference frame where the proton moves very fast and Q>>m<sub>h</sub> is big

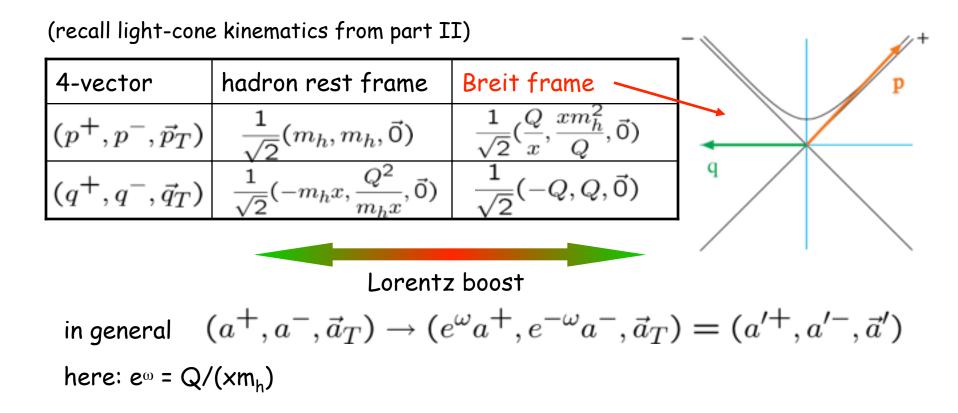


| (recall light-cone      | - +                                                      |                                                            |   |
|-------------------------|----------------------------------------------------------|------------------------------------------------------------|---|
| 4-vector                | hadron rest frame                                        | Breit frame 🔍                                              | р |
| $(p^+, p^-, \vec{p}_T)$ | $\frac{1}{\sqrt{2}}(m_h, m_h, \vec{0})$                  | $\frac{1}{\sqrt{2}}(\frac{Q}{x},\frac{xm_h^2}{Q},\vec{0})$ |   |
| $(q^+,q^-,\vec{q}_T)$   | $\frac{1}{\sqrt{2}}(-m_h x, \frac{Q^2}{m_h x}, \vec{0})$ | $rac{1}{\sqrt{2}}(-Q,Q,ec{0})$                            | q |
|                         |                                                          |                                                            |   |

# space-time picture of DIS

this can be best understood in a reference frame where the proton moves very fast and  $Q \gg m_h$  is big

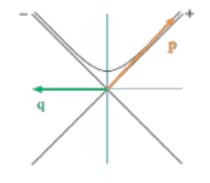


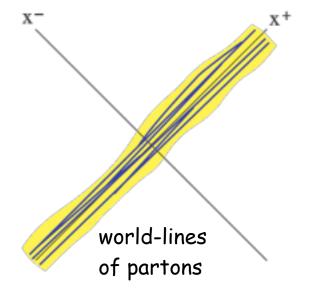


#### space-time picture of DIS – cont'd

simple estimate for typical time-scale of interactions among the partons inside a fast-moving hadron:

rest frame: 
$$\Delta x^+ \sim \Delta x^- \sim \frac{1}{m}$$
  
Breit frame:  $\Delta x^+ \sim \frac{1}{m} \frac{Q}{m} = \frac{Q}{m^2}$  large  
 $\Delta x^- \sim \frac{1}{m} \frac{m}{Q} = \frac{1}{Q}$  small



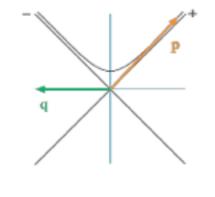


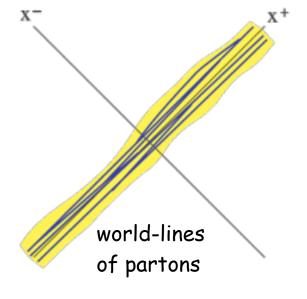
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interactions between partons are spread out inside a fast moving hadron





#### space-time picture of DIS – cont'd

simple estimate for typical time-scale of interactions among the partons inside a fast-moving hadron:

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$$\Delta x^{+} \sim \Delta x^{-} \sim \frac{1}{m}$$
  
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 $\Delta x^{-} \sim \frac{1}{m} \frac{m}{Q} = \frac{1}{Q}$  small

interactions between partons are spread out inside a fast moving hadron xworld-lines of partons

How does this compare with the time-scale of the hard scattering?

Breit frame:

proton moves very fast and  $Q \gg m_h$  is big

$$(p^+, p^-, \vec{p}_T) = \frac{1}{\sqrt{2}} (\frac{Q}{x}, \frac{xm_h^2}{Q}, \vec{0}) \quad (q^+, q^-, \vec{q}_T) = \frac{1}{\sqrt{2}} (-Q, Q, \vec{0})$$

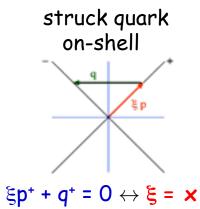
struck quark on-shell q $\xi p^+ + q^+ = 0 \leftrightarrow \xi = x$ 

#### Breit frame:

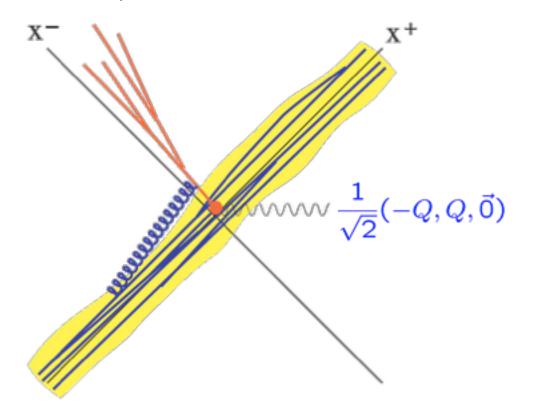
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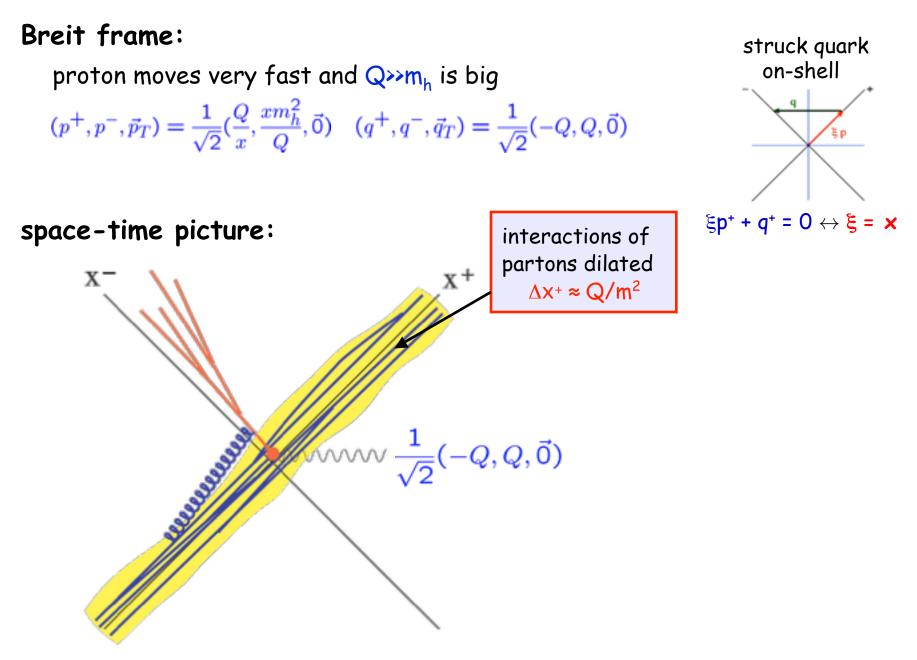
$$(p^+, p^-, \vec{p}_T) = \frac{1}{\sqrt{2}} (\frac{Q}{x}, \frac{xm_h^2}{Q}, \vec{0}) \quad (q^+, q^-, \vec{q}_T) = \frac{1}{\sqrt{2}} (-Q, Q, \vec{0})$$

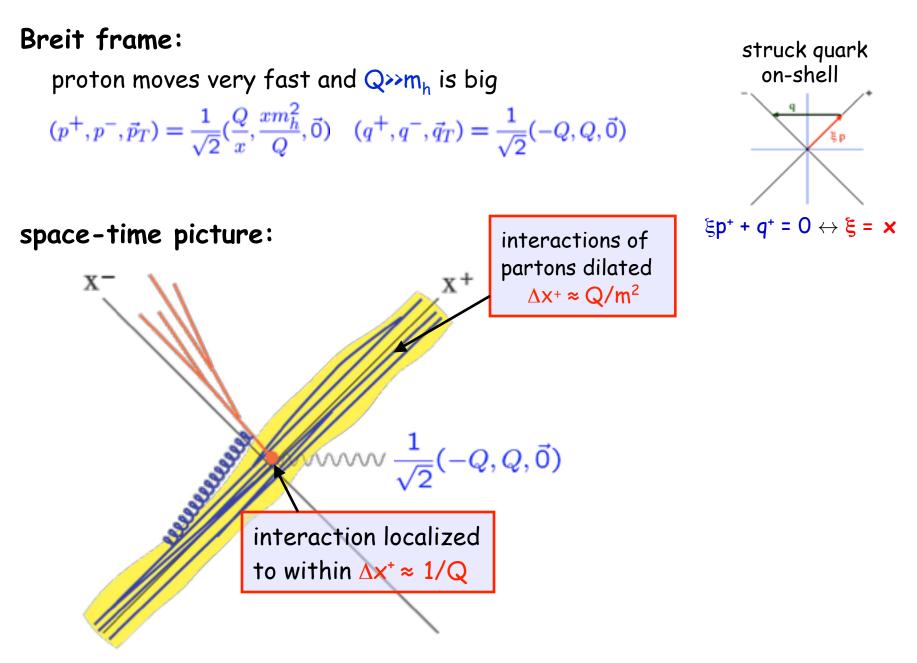
Feynman; Bjorken, Paschos

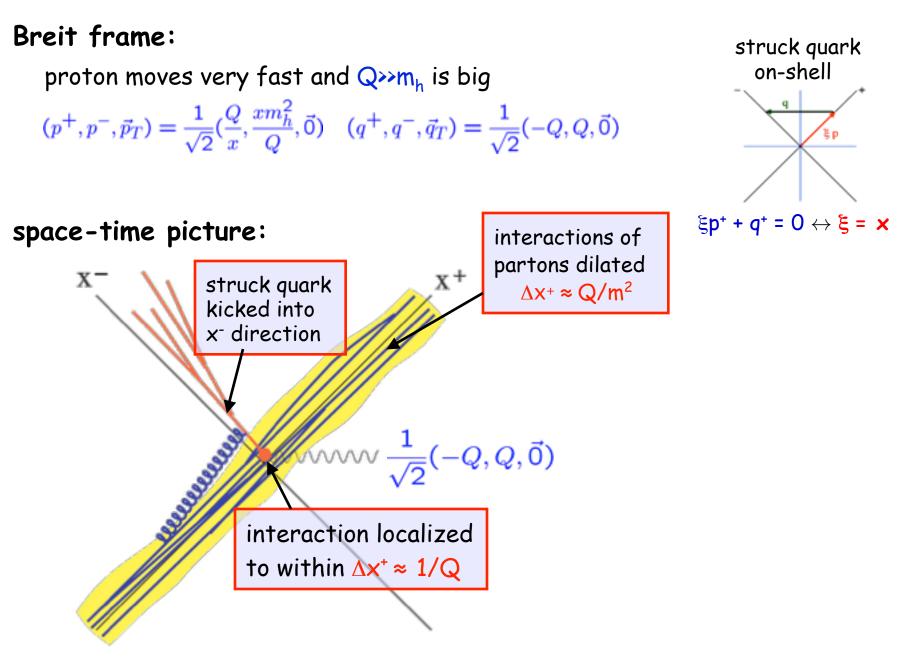


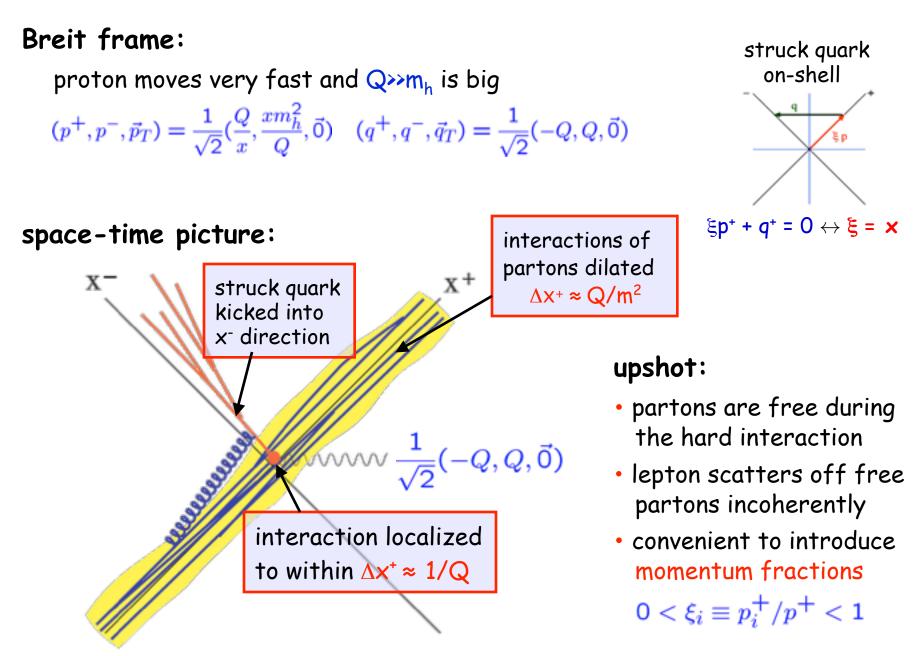
space-time picture:

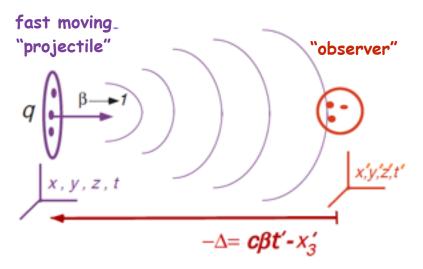








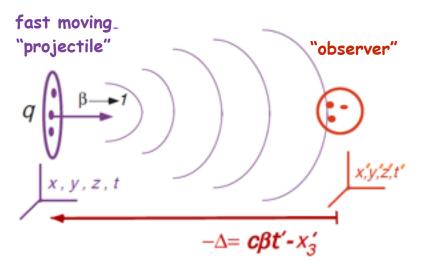




adapted from G. Sterman's lectures

accelerated charges produce classical radiation QFT assembles field from infinite # of soft quanta

Lorentz transformation  $\mathbf{x_3} = \gamma(eta \mathbf{ct'} - \mathbf{x'_3}) \equiv -\gamma \mathbf{\Delta}$ 

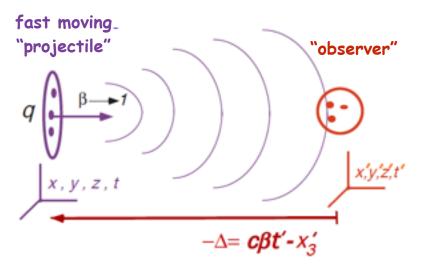


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| field                           | x-frame                                   | x' frame                                                                 | Lorentz<br>factor |
|---------------------------------|-------------------------------------------|--------------------------------------------------------------------------|-------------------|
| scalar field $\phi(\mathbf{x})$ | $\frac{\mathbf{q}}{ \mathbf{\tilde{x}} }$ | $\frac{\mathbf{q}}{(\mathbf{x_T^2} + \gamma^2 \mathbf{\Delta^2})^{1/2}}$ | $1/\gamma$        |
|                                 |                                           |                                                                          |                   |
|                                 |                                           |                                                                          |                   |

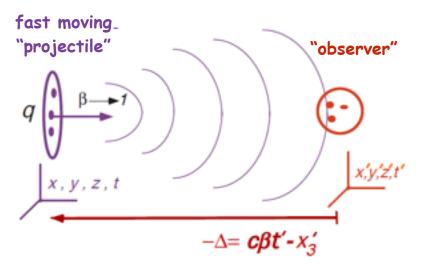


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| field                                  | x-frame                                   | x' frame                                                                        | Lorentz<br>factor |
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| scalar field $\phi(\mathbf{x})$        | $\frac{\mathbf{q}}{ \mathbf{\tilde{x}} }$ | $\frac{\mathbf{q}}{(\mathbf{x_T^2} + \gamma^2 \mathbf{\Delta^2})^{1/2}}$        | $1/\gamma$        |
| gauge field $\mathbf{A_0}(\mathbf{x})$ | $\frac{\mathbf{q}}{ \mathbf{\tilde{x}} }$ | $\frac{-\mathbf{q}\gamma}{(\mathbf{x_T^2} + \gamma^2 \mathbf{\Delta}^2)^{1/2}}$ | $\gamma^{0}$      |
|                                        |                                           |                                                                                 |                   |

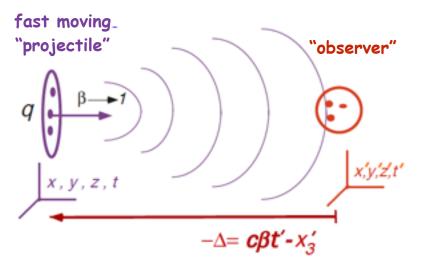


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| field                                       | x-frame                                       | x' frame                                                                                                 | Lorentz<br>factor |
|---------------------------------------------|-----------------------------------------------|----------------------------------------------------------------------------------------------------------|-------------------|
| scalar field $\phi(\mathbf{x})$             | $\frac{\mathbf{q}}{ \mathbf{\tilde{x}} }$     | $\frac{\mathbf{q}}{(\mathbf{x_T^2} + \gamma^2 \mathbf{\Delta^2})^{1/2}}$                                 | $1/\gamma$        |
| gauge field $\mathbf{A_0}(\mathbf{x})$      | $\frac{\mathbf{q}}{ \mathbf{\tilde{x}} }$     | $\frac{-\mathbf{q}\gamma}{(\mathbf{x_T^2}+\gamma^2\mathbf{\Delta^2})^{1/2}}$                             | $\gamma^{0}$      |
| "field strength" $\mathbf{E_3}(\mathbf{x})$ | $\frac{\mathbf{q}}{ \tilde{\mathbf{x}} ^{2}}$ | $\frac{-\mathbf{q}\gamma\mathbf{\Delta}}{(\mathbf{x_T^2}+\gamma^{2}\mathbf{\Delta}^{2})^{\mathbf{3/2}}}$ | $1/\gamma^{2}$    |

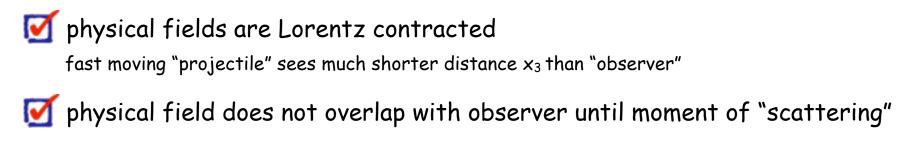


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accelerated charges produce classical radiation QFT assembles field from infinite # of soft quanta

Lorentz transformation  $\mathbf{x_3} = \gamma(\beta \mathbf{ct'} - \mathbf{x'_3}) \equiv -\gamma \mathbf{\Delta}$ 

# upshot



 $\widecheck{ extsf{M}}$  corrections (= "advanced effects") power suppressed  $\propto (1-eta)$ 

Much the same reasoning for final-state



### sum rules and isospin

for the quark distributions in a proton there are several sum rules to obey

$$\int_0^1 dx \sum_i x f_i^{(p)}(x) = 1$$

$$\begin{split} &\int_{0}^{1} dx \left( f_{u}^{(p)}(x) - f_{\bar{u}}^{(p)}(x) \right) = 2 \\ &\int_{0}^{1} dx \left( f_{d}^{(p)}(x) - f_{\bar{d}}^{(p)}(x) \right) = 1 \\ &\int_{0}^{1} dx \left( f_{s}^{(p)}(x) - f_{\bar{s}}^{(p)}(x) \right) = 0 \end{split}$$

#### momentum sum rule

quarks share proton momentum

flavor sum rules conservation of quantum numbers

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#### momentum sum rule

quarks share proton momentum

flavor sum rules conservation of quantum numbers

isospin symmetry relates a neutron to a proton (just u and d interchanged)

$$F_2^n(x) = x\left(\frac{1}{9}d_n(x) + \frac{4}{9}u_n(x)\right) = x\left(\frac{4}{9}d_p(x) + \frac{1}{9}u_p(x)\right)$$

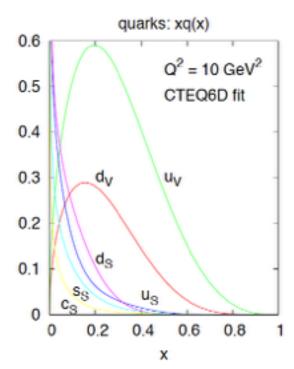
• measuring both allows to determine u<sup>p</sup> and d<sup>p</sup> separately

• note: CC DIS couples to weak charges and separates quarks and antiquarks

$$\int_0^1 dx \sum_i x f_i^{(p)}(x) = 1$$

| Uv             | 0.267 |
|----------------|-------|
| d,             | 0.111 |
| U <sub>s</sub> | 0.066 |
| d,             | 0.053 |
| <b>S</b> 5     | 0.033 |
| Cc             | 0.016 |
| total          | 0.546 |

half of the momentum is missing

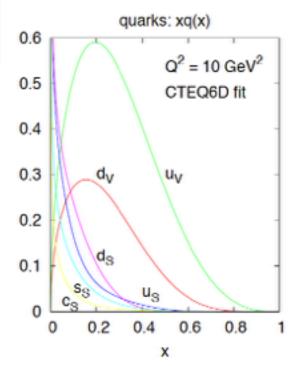


$$\int_0^1 dx \sum_i x f_i^{(p)}(x) = 1$$

| 0.267 |
|-------|
| 0.111 |
| 0.066 |
| 0.053 |
| 0.033 |
| 0.016 |
| 0.546 |
|       |

half of the momentum is missing

gluons!



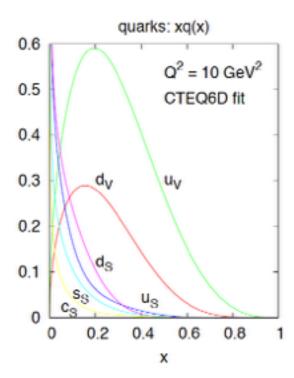
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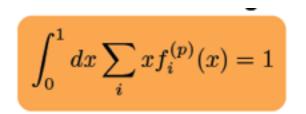
| 0.267 |
|-------|
| 0.111 |
| 0.066 |
| 0.053 |
| 0.033 |
| 0.016 |
| 0.546 |
|       |



gluons!

but they don't carry electric/weak charge how can they couple?





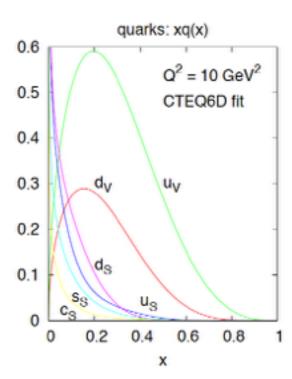
| 0.267 |
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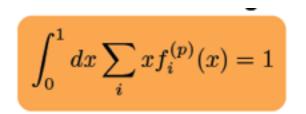


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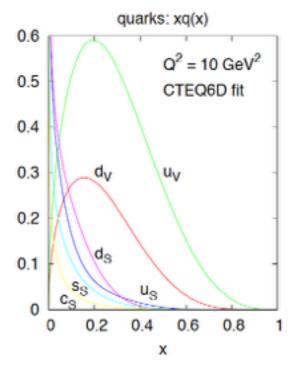


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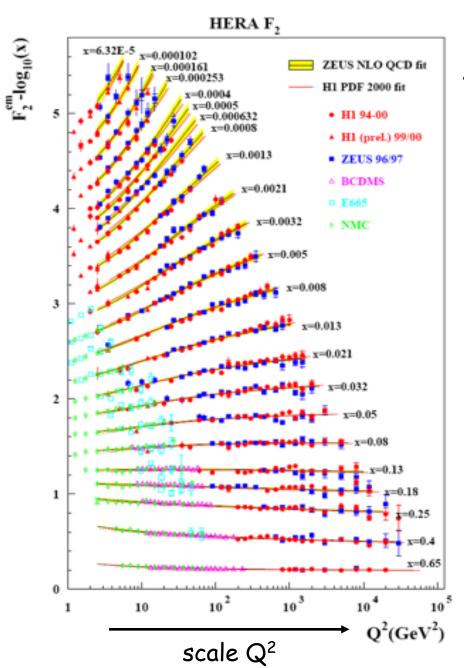
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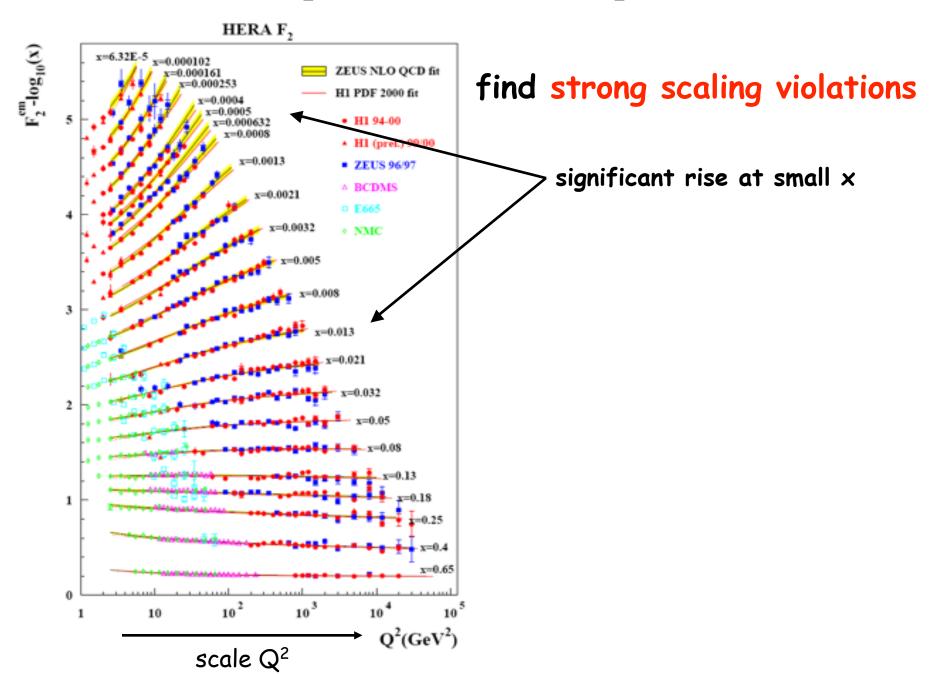


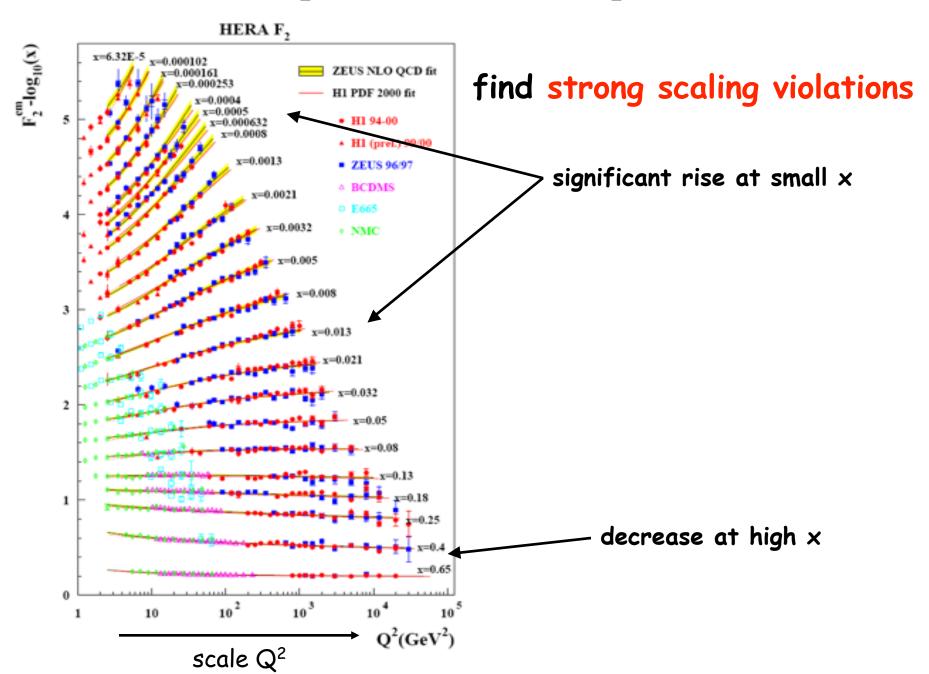
-> we need to discuss QCD radiative corrections to the naïve picture

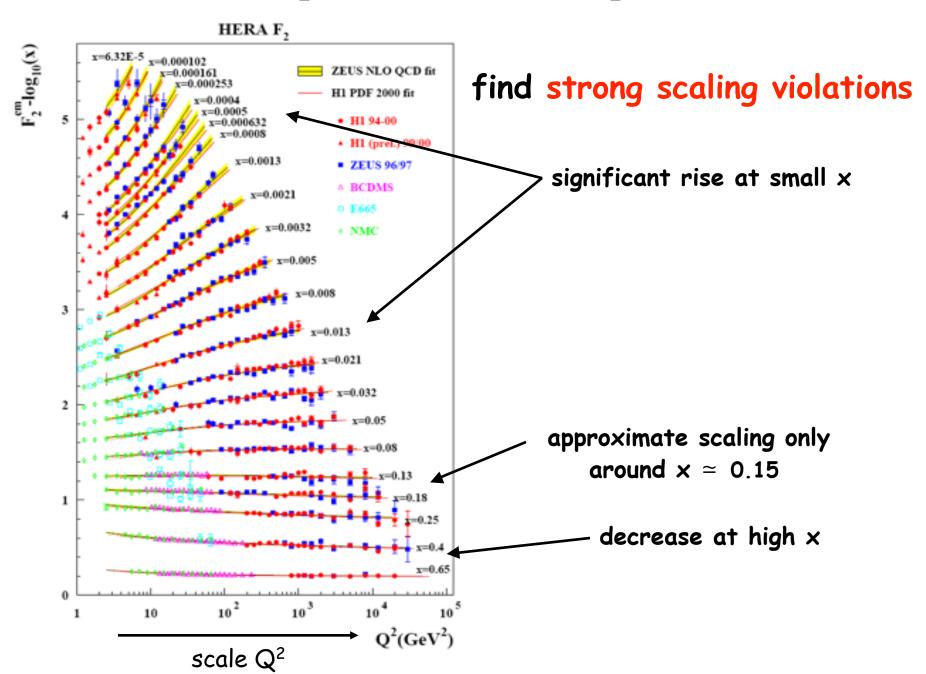
gluons will enter the game and everything will become scale dependent



find strong scaling violations

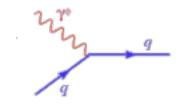






# **DIS in the QCD improved parton model**

we got a long way (parton model) without invoking QCD

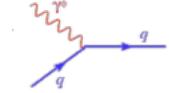


now we have to study QCD dynamics in DIS

- this leads to similar problems already encountered in  $e^+e^-$ 

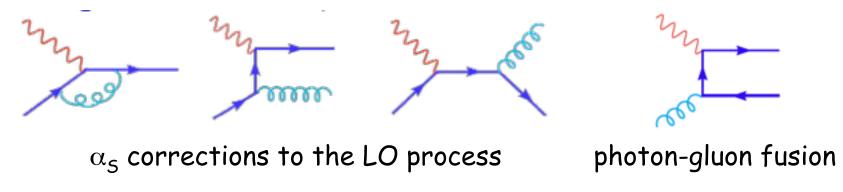
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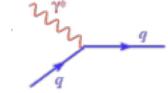
now we have to study QCD dynamics in DIS - this leads to similar problems already encountered in e<sup>+</sup>e<sup>-</sup>

let's try to compute the  $O(\alpha_s)$  QCD corrections to the naive picture



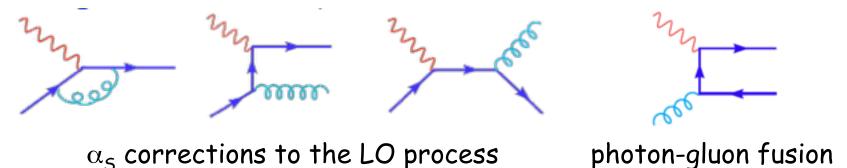
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now we have to study QCD dynamics in DIS - this leads to similar problems already encountered in e<sup>+</sup>e<sup>-</sup>

let's try to compute the  $O(\alpha_s)$  QCD corrections to the naive picture



caveat: have to expect divergencies (recall 2<sup>nd</sup> part) related to soft/collinear emission or from loops

we cannot calculate with infinities  $\rightarrow$  introduce a "regulator" and remove it in the end

regulating divergencies is the  $1^{st}$  step in higher order calculations

standard regulators in QCD calculations:

dimensional regularization

change dimension of space-time to  $4\text{-}2\epsilon$   $\rightarrow$  calculations (integrals) rather involved;

works in general, i.e., to all orders



<u>issues</u>:  $\gamma_5$  (spin, e.-w. couplings), SUSY, helicity violation

regulating divergencies is the 1<sup>st</sup> step in higher order calculations

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change dimension of space-time to  $4-2\epsilon$  $\rightarrow$  calculations (integrals) rather involved; works in general, i.e., to all orders

issues:  $\gamma_5$  (spin, e.-w. couplings), SUSY, helicity violation

# small quark/gluon mass

intuitive and transparent; stick to four dimensions <u>issues</u>: does not work beyond NLO

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only if we have done everything consistently, including factorization, we can safely remove the regulator and can compare to experiment

$$\begin{aligned} \frac{d^2\hat{\sigma}}{dxdQ^2}\Big|_{F_2} &\equiv \hat{F}_2^g \\ &= \sum_q e_q^2 x \left[ 0 + \frac{\alpha_s(\mu_r)}{4\pi} \left[ P_{qg}(x) \ln \frac{Q^2}{m_q^2} + C_2^g(x) \right] \right] \quad \stackrel{\mathcal{W}}{\underset{\sigma}{\longrightarrow}} \end{aligned}$$

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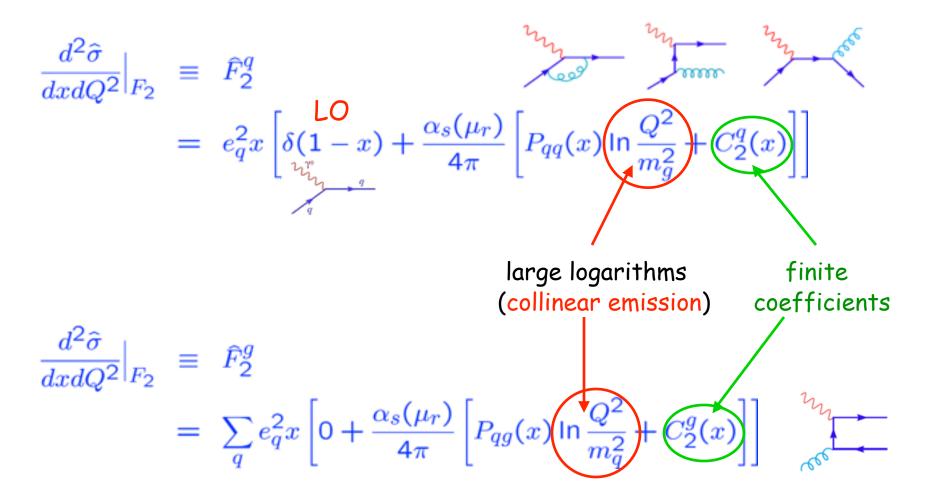
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using small (artificial) quark/gluon masses as regulator we obtain:



to see what happens to the logs we have to convolute our results with the PDFs

for the quark part we obtain:

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$$F_2(x,Q^2) = x \sum_{a=q,\bar{q}} e_q^2 \Big[ f_{a,0}(x) + \frac{\alpha_s}{2\pi} \int_x^1 \frac{d\xi}{\xi} \\ f_{a,0}(x) \Big[ P_{qq} \left( \frac{x}{\xi} \right) \ln \frac{Q^2}{m_q^2} + C_2^q \left( \frac{x}{\xi} \right) \Big] \Big]$$
from from

similarly for the gluonic part

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f<sub>a,0</sub>(x): unmeasurable "bare" (= infinite) parton densities; need to be re-defined (= renormalized) to make them physical

at order  $\alpha_s$ : (can be generalized to all orders)

$$f_a(x,\mu_f^2) \equiv f_{a,0}(x) + \frac{\alpha_s}{2\pi} \int_x^1 \frac{d\xi}{\xi} f_{a,0}(\xi) P_{qq}\left(\frac{x}{\xi}\right) \ln\left(\frac{\mu_f^2}{m_g^2}\right) + z_{qq}$$

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absorbs all long-distance singularities at a factorization scale  $\mu_f$  into  $f_{a,0}$ 

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physical/renormalized densities: not calculable in pQCD but universal

putting everything together, keeping only terms up to  $\alpha_s$ :

$$F_{2}(x,Q^{2}) = x \sum_{a=q,\bar{q}} e_{q}^{2} \int_{x}^{1} \frac{d\xi}{\xi} f_{a}(\xi,\mu_{f}^{2}) \\ \left[ \delta(1-\frac{x}{\xi}) + \frac{\alpha_{s}(\mu_{r})}{2\pi} \left[ P_{qq}\left(\frac{x}{\xi}\right) \ln \frac{Q^{2}}{\mu_{f}^{2}} + (C_{2}^{q} - z_{qq})\left(\frac{x}{\xi}\right) \right] \right]$$

short-distance "Wilson coefficient"

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both, pdf's and the short-dist. coefficient depend on  $\mu_f$  (choice of  $\mu_f$ : shifting terms between long- and short-distance parts)

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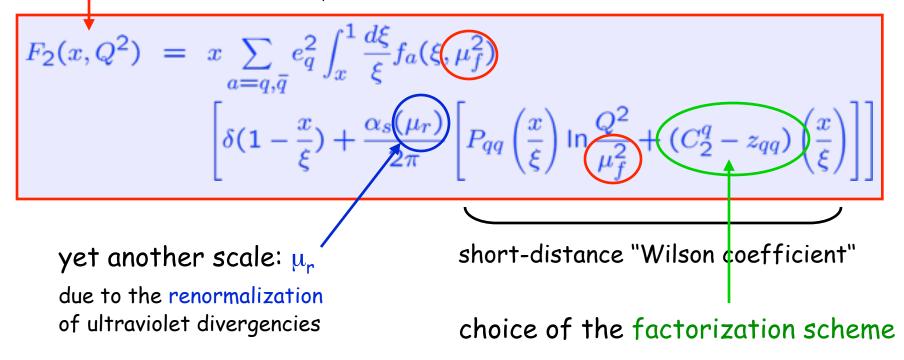
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#### choice of the factorization scheme

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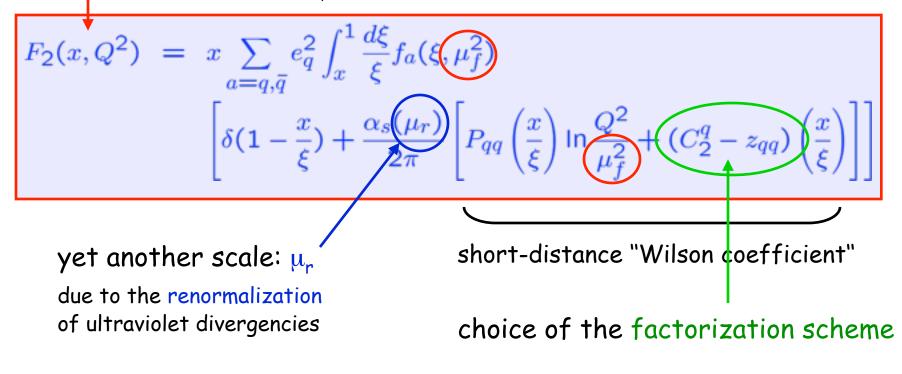
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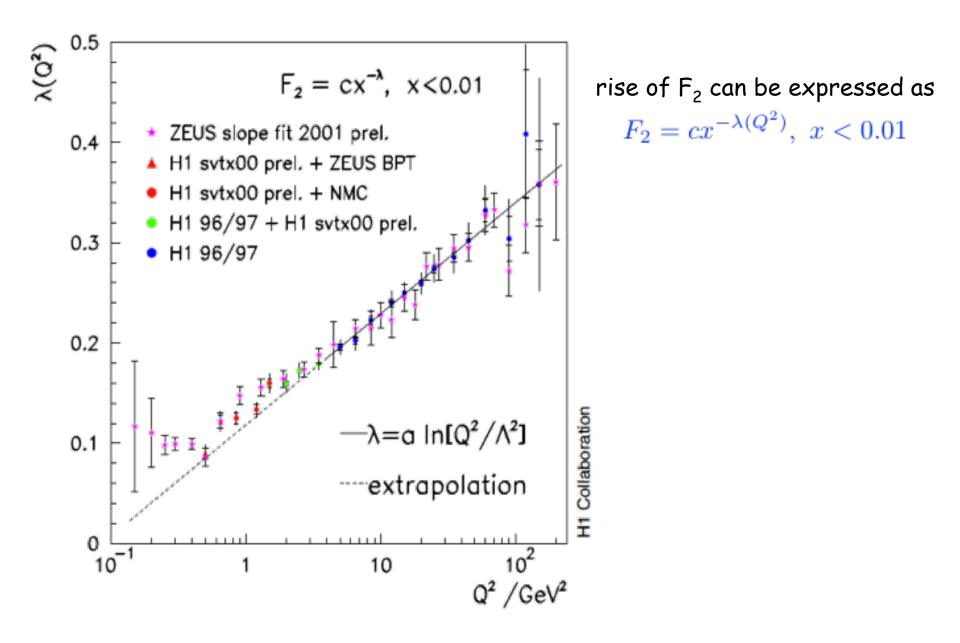


#### this result is readily extended to hadron-hadron collisions

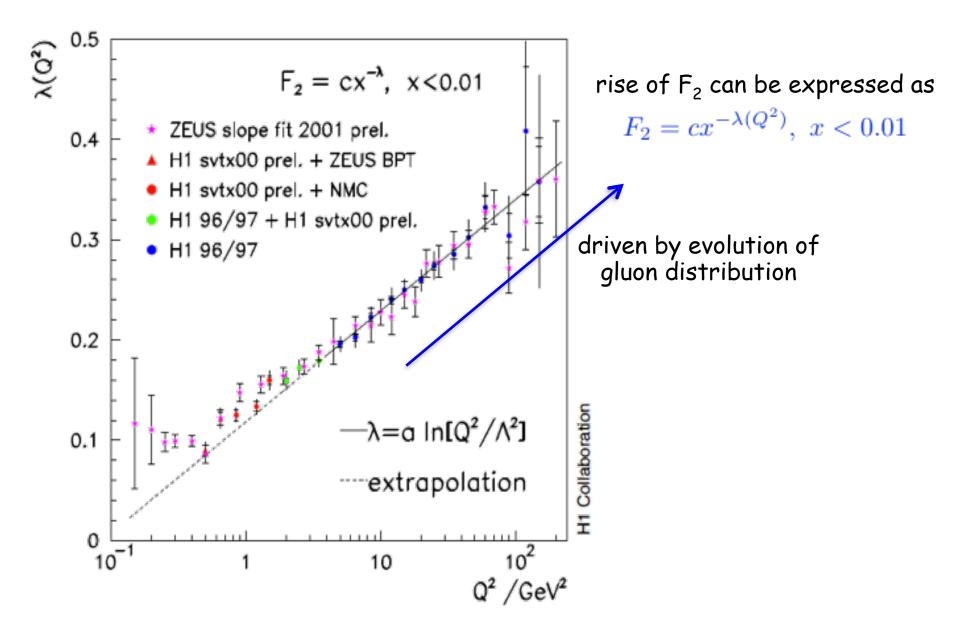
### lesson: theorists are not afraid of infinities



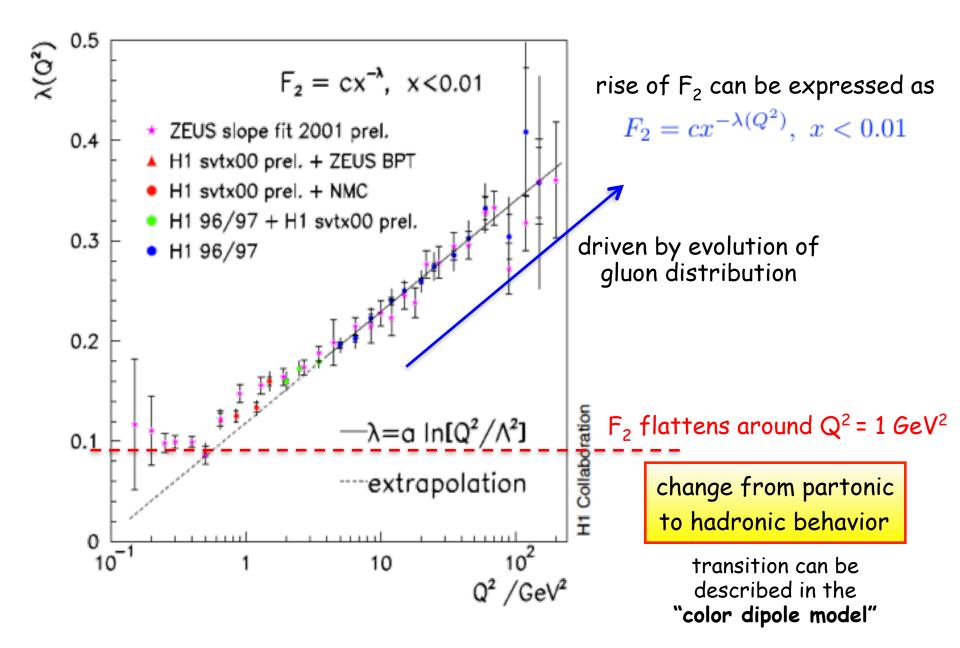
### HERA's legacy: rise of $F_2$ vs $Q^2$



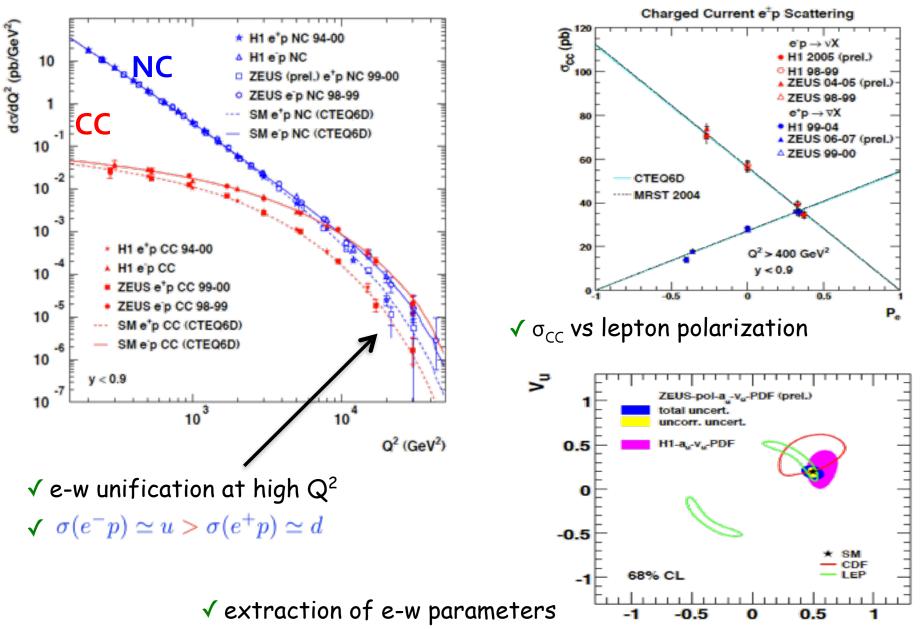
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#### NC & CC DIS: test of e-w theory !



au

### universal PDFs $\rightarrow$ key to predictive power of pQCD

once PDFs are extracted from one set of experiments, e.g. DIS, we can use them to **predict cross sections** in, say, hadron-hadron collisions

parton densities are **universal** 

 $\rightarrow$  there must be a process-independent precise definition

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small print: we need to specify a common factorization scheme for short- and long-distance physics (= choice of z<sub>ii</sub> in our result for F<sub>2</sub>)

> standard choice: modified minimal subtraction (MS) scheme (closely linked to dim. regularization; used in all PDF fits)

less often used: **DIS scheme** = "maximal" subtraction where all  $O(\alpha_s)$  corrections in DIS are absorbed into PDFs (nice for DIS but a bit awkward for other processes)

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Bardeen, Buras, classic (but old-fashioned) definition of PDFs through their Duke, Muta Mellin moments in Wilson-Zimmermann's operator product expansion (OPE)

**more physical formulation** in Bjorken-x space:

matrix elements of bi-local operators on the light-cone

Curci, Furmanski, Petronzio; Collins, Soper see, e.g., D. Soper, hep-lat/9609018

for quarks: (similar for gluons; easy to include spin  $\gamma^* \rightarrow \gamma^* \gamma_5$ )

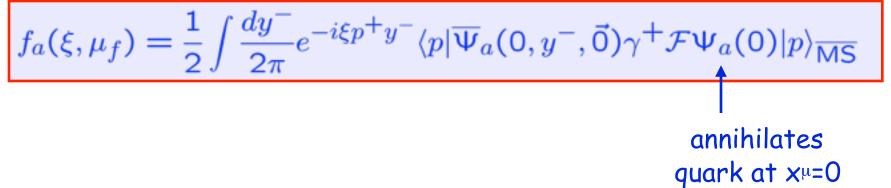
 $f_a(\xi,\mu_f) = \frac{1}{2} \int \frac{dy^-}{2\pi} e^{-i\xi p^+ y^-} \langle p | \overline{\Psi}_a(0,y^-,\vec{0}) \gamma^+ \mathcal{F} \Psi_a(0) | p \rangle_{\overline{\mathsf{MS}}}$ 

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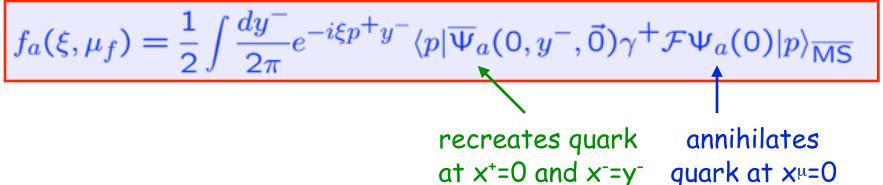


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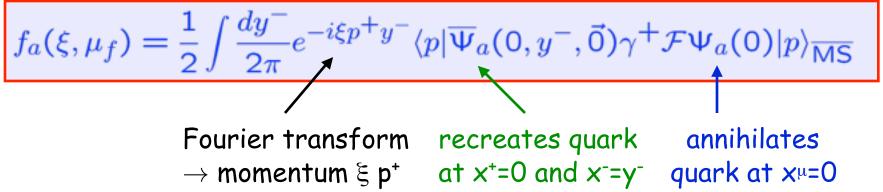


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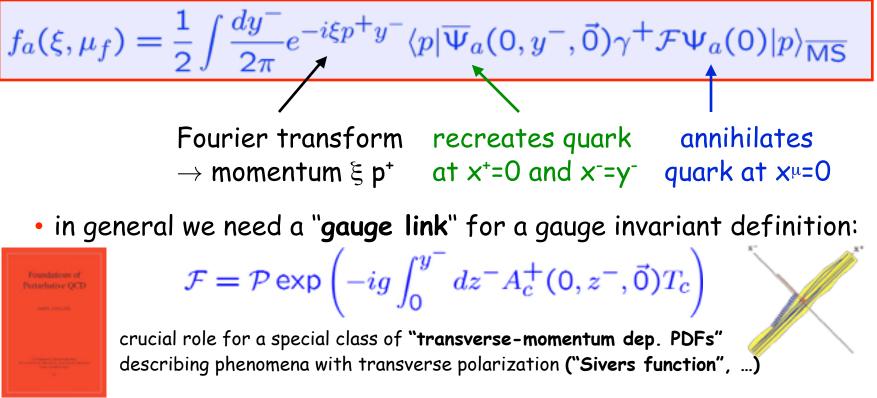
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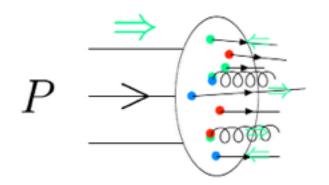
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- interpretation as "number operator" only in "A<sup>+</sup>= 0 gauge"
- turn into local operators (  $\rightarrow$  lattice QCD) if taking moments  $\int_0^1 d\xi \ \xi^n$

# pictorial representation of PDFs

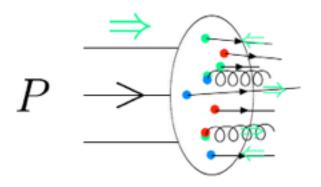
suppose we could take a snapshot of a nucleon with positive helicity



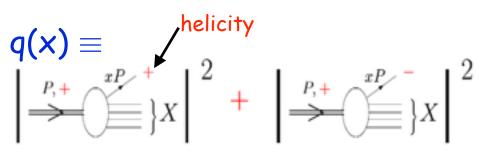
**question:** how many constituents (quark, anti-quarks, gluons) have momenta between xP and (x+dx)P and how many have the same/opposite helicity?

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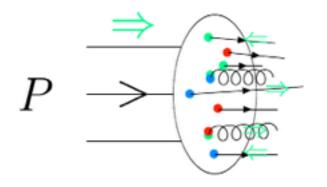
 $g(\mathbf{x}) = \left| \underbrace{\stackrel{P,+}{\longrightarrow}}_{0} \underbrace{\stackrel{xP_{0}}{\longrightarrow}}_{0} \right|^{2} + \left| \underbrace{\stackrel{P,+}{\longrightarrow}}_{0} \underbrace{\stackrel{xP_{0}}{\longrightarrow}}_{0} \right|^{2}$ 

### unpolarized PDFs

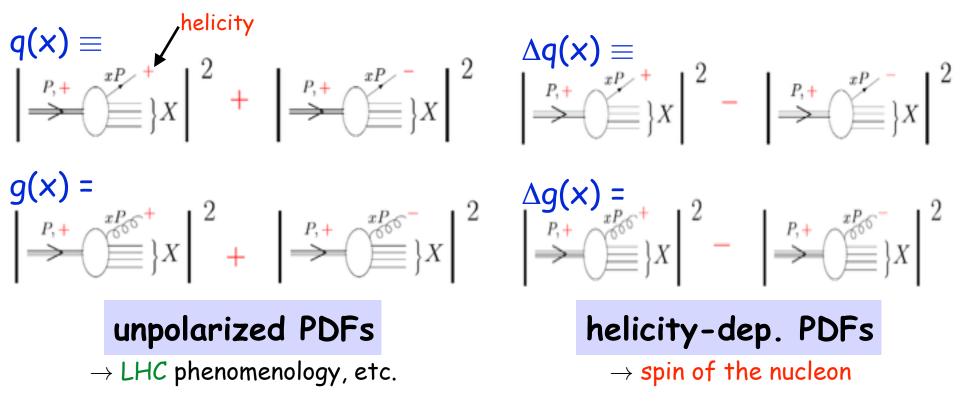
 $\rightarrow$  LHC phenomenology, etc.

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these singularities cancel for infrared safe observables or can be systematically removed (factorization) by "hiding" them in some non-perturbative parton distribution or fragmentation functions

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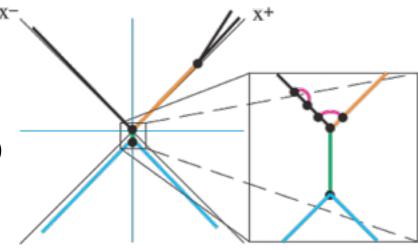
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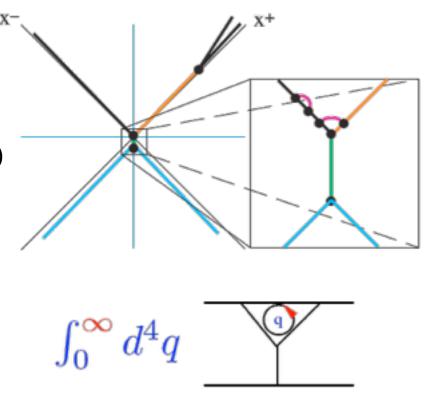
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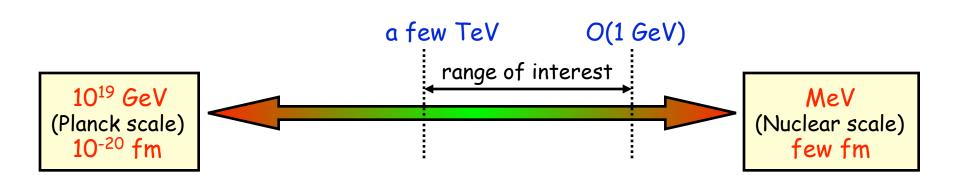
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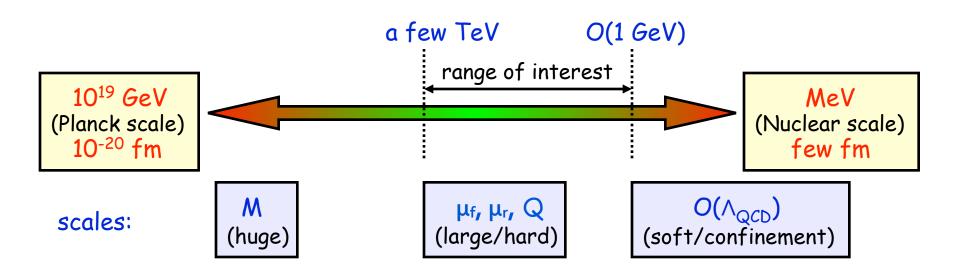
again, we need a suitable regulator for divergent loop integrations:

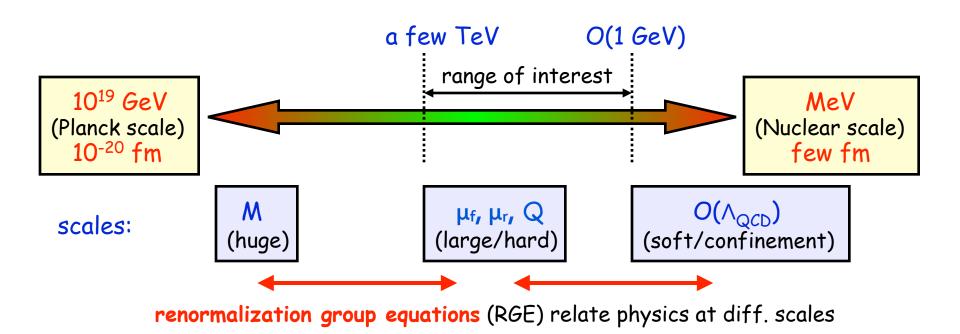
UV cut-off vs. dim. regularization intuitive; involved; not beyond NLO works to all orders

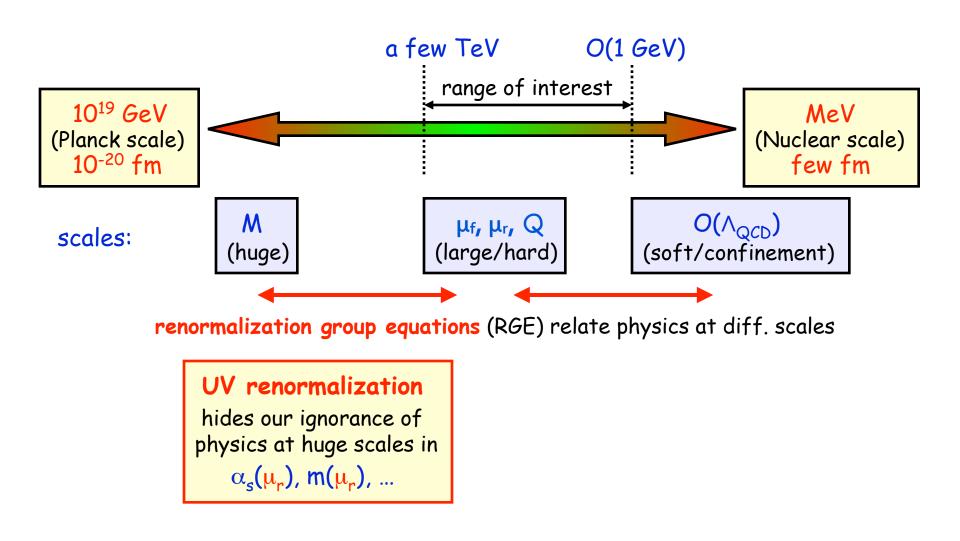


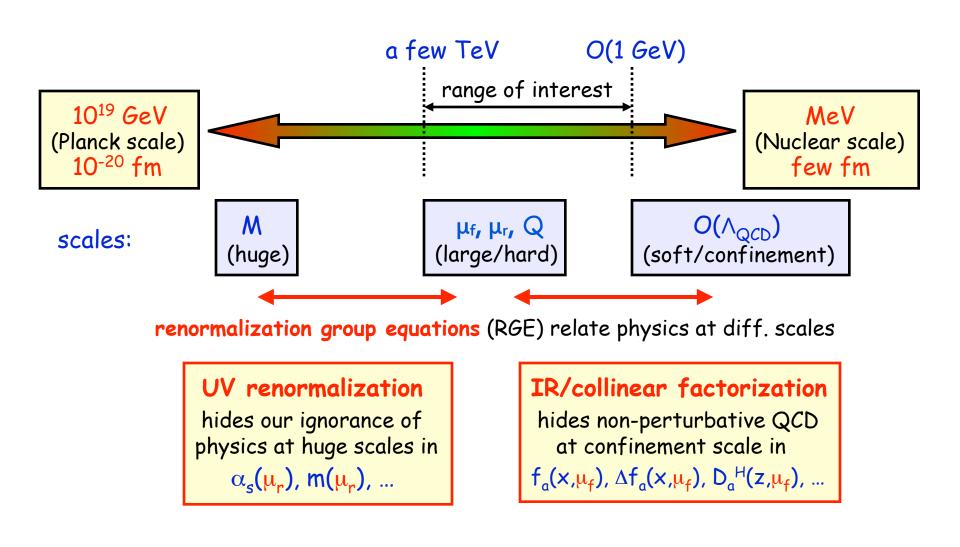














we use  $\alpha_s$  (and  $f_a$ ,  $D_c^H$ ) to absorb UV (IR) divergencies

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all we need is a reference measurement at some scale  $\mu_0$ 

#### scale evolution of $\alpha_s$ and parton densities

simplest example of RGE: running coupling  $\alpha_s$  derived from  $\frac{d\sigma}{d \ln \mu_r} = 0$  $\rightarrow \frac{\text{recall}}{\text{part II}} \frac{da_s}{d \ln \mu^2} = -\beta_0 a_s^2 - \beta_1 a_s^3 - \beta_2 a_s^4 - \beta_3 a_s^5 + \dots \quad a_s \equiv \frac{\alpha_s}{4\pi}$ 

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scale dependence of PDFs: more complicated

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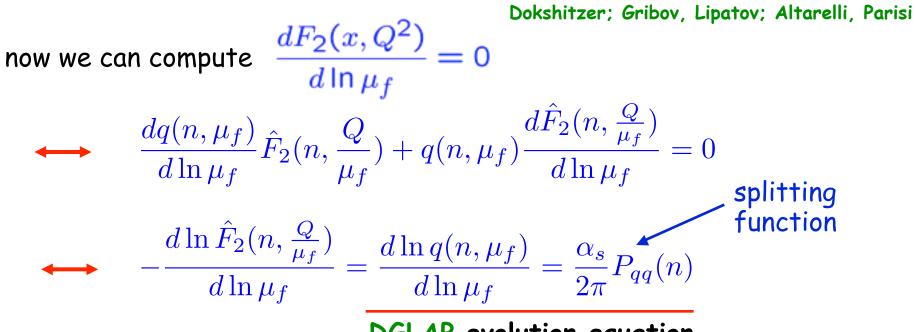
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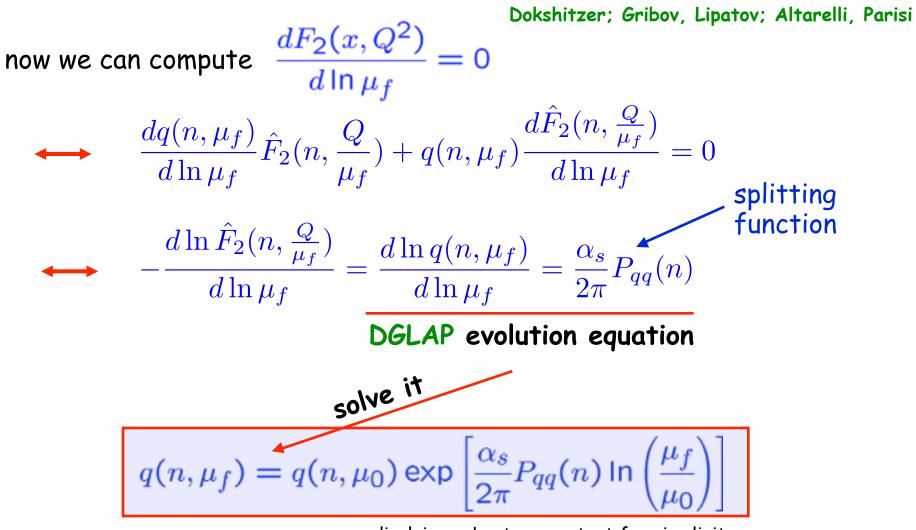
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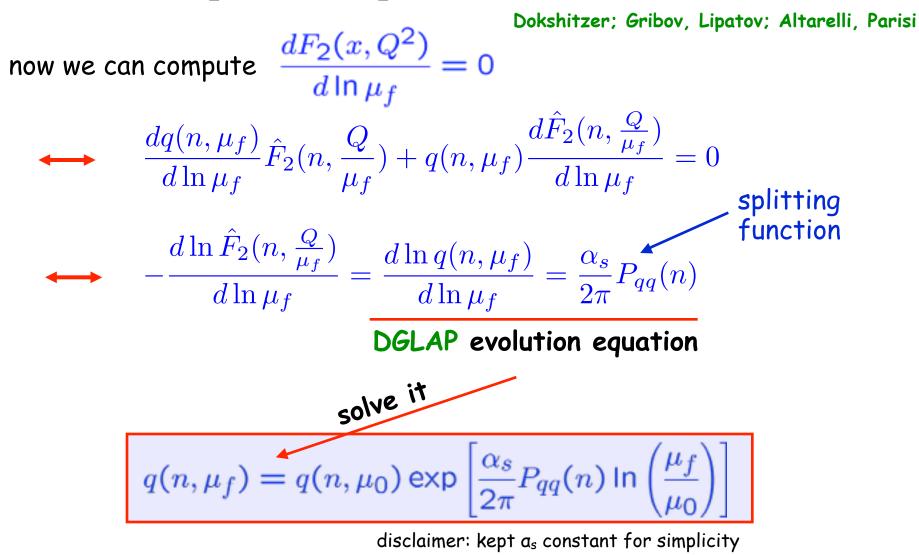
now we can compute  $\frac{dF_2(x,Q^2)}{d\ln \mu_f} = 0$   $\longrightarrow \quad \frac{dq(n,\mu_f)}{d\ln \mu_f} \hat{F}_2(n,\frac{Q}{\mu_f}) + q(n,\mu_f) \frac{d\hat{F}_2(n,\frac{Q}{\mu_f})}{d\ln \mu_f} = 0$ 



**DGLAP** evolution equation



disclaimer: kept as constant for simplicity



ightarrow once we know the PDFs at a scale  $\mu_0$  we can predict them at  $\mu$  >  $\mu_0$ 

physical interpretation of the evolution eqs.:

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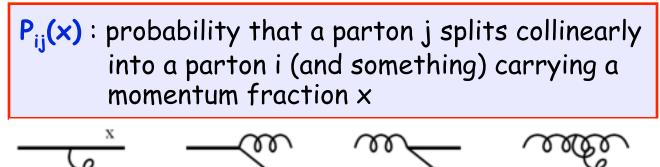
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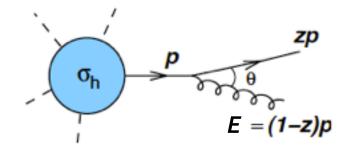
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- the splitting functions  $P_{ij}(n)$  or  $P_{ij}(x)$  multiplying the log's are universal and calculable in pQCD order by order in  $\alpha_s$
- the physical meaning of the splitting functions is easy:



# factorization recap: final-state vs initial-state

recall what we learned for final-state radiation

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 $\int_{1}^{1} \frac{p}{\sigma_{h}} \frac{zp}{\theta}$  E = (1-z)p



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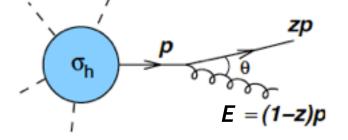
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where we have used

$$\mathbf{E} = (\mathbf{1} - \mathbf{z})\mathbf{p}$$
$$\mathbf{k}_{\mathbf{T}} = \mathbf{E}\sin\theta \simeq \mathbf{E}\theta$$

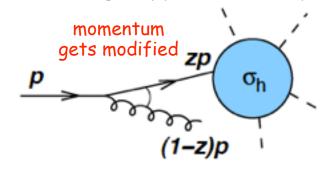
**KLN:** if we avoid distinguishing quark and collinear quark-gluon final-states (like for jets) divergencies cancel against virtual corrections

$$\int_{\mathbf{\sigma}_{h}} \frac{\mathbf{p}}{\mathbf{r}_{h}} \frac{\mathbf{p}}{\mathbf{r}_{h}} \frac{\mathbf{p}}{\mathbf{r}_{h}} \frac{\mathbf{p}}{\mathbf{r}_{h}} \frac{\mathbf{p}}{\mathbf{r}_{h}} \frac{\mathbf{p}}{1-z} \frac{dk_{t}^{2}}{k_{t}^{2}}$$



initial-state radiation: crucial difference - hard scattering happens after splitting

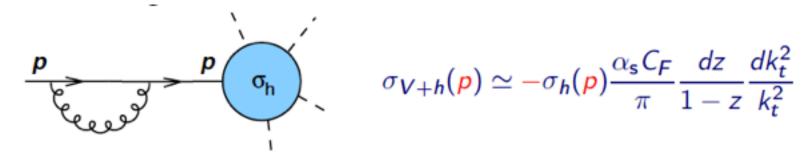
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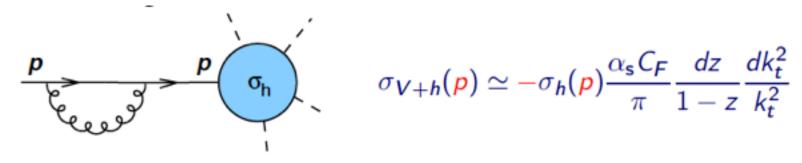


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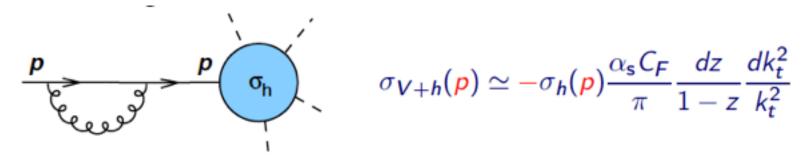
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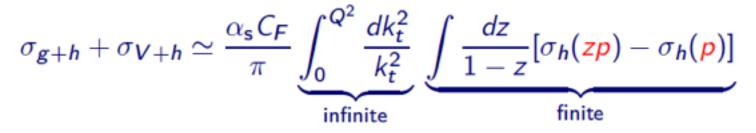
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leads to uncanceled collinear singularity

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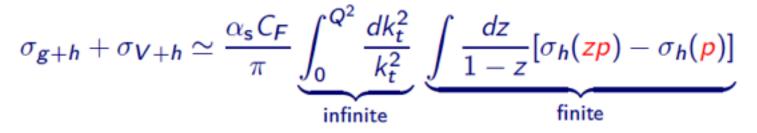
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# factorization revisited: collinear singularity



- z=1: soft divergence cancels (KLN) as  $\sigma_{\mathbf{h}}(\mathbf{zp}) \sigma_{\mathbf{h}}(\mathbf{p}) \rightarrow \mathbf{0}$
- arbitrary z:  $\sigma_{\mathbf{h}}(\mathbf{zp}) \sigma_{\mathbf{h}}(\mathbf{p}) \neq \mathbf{0}$  but z integration is finite
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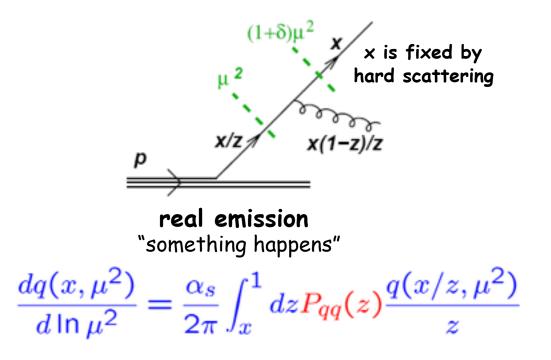
$$\sigma_{g+h} + \sigma_{V+h} \simeq \frac{\alpha_{s} C_{F}}{\pi} \underbrace{\int_{0}^{Q^{2}} \frac{dk_{t}^{2}}{k_{t}^{2}}}_{\text{infinite}} \underbrace{\int \frac{dz}{1-z} [\sigma_{h}(zp) - \sigma_{h}(p)]}_{\text{finite}}$$

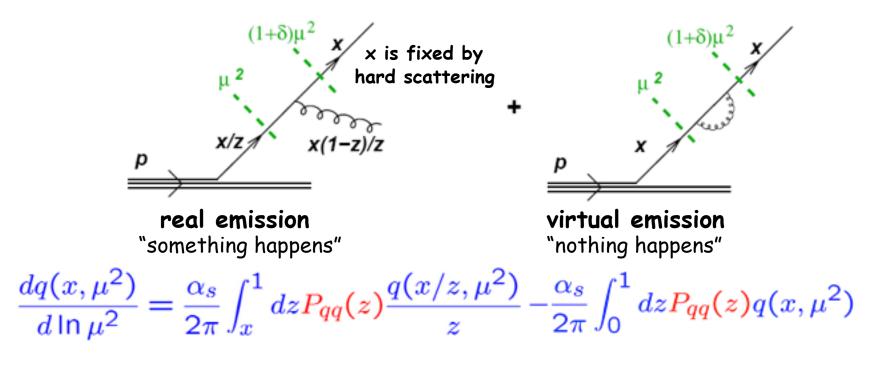
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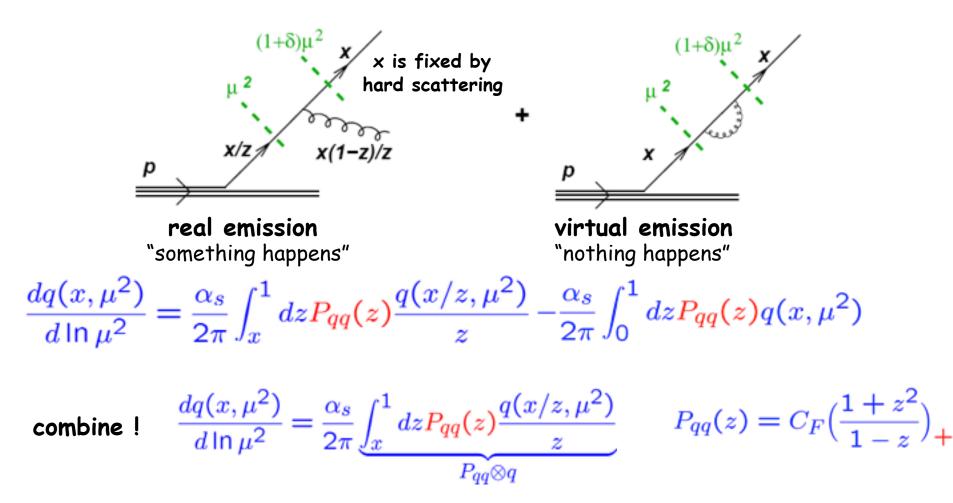
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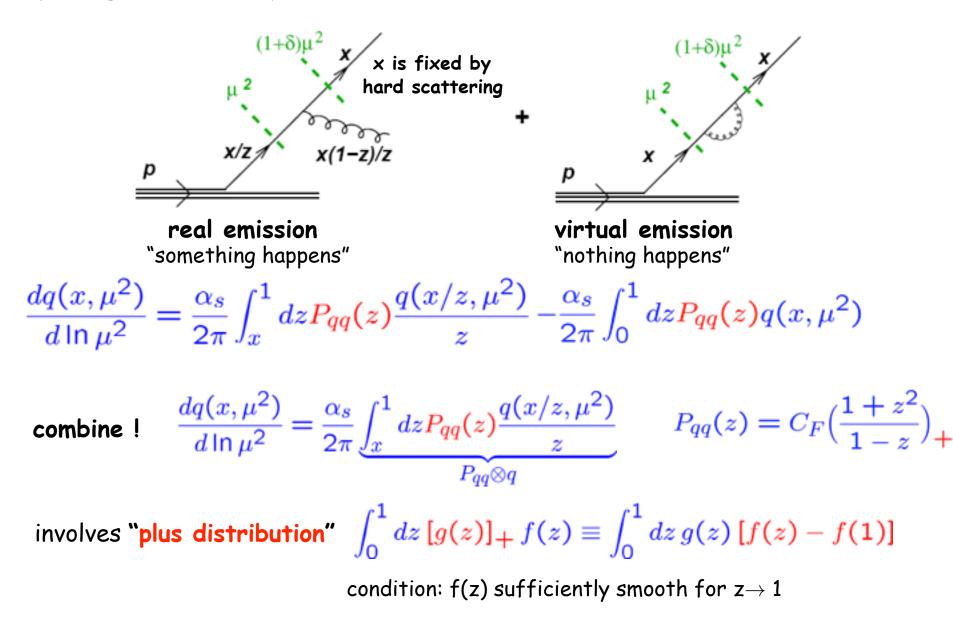
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# factorization = collinear "cut-off" • absorb divergent small $k_T$ region in non-perturbative PDFs $\sigma_1 \simeq \frac{\alpha_s C_F}{\pi} \underbrace{\int_{\mu^2}^{Q^2} \frac{dk_t^2}{k_t^2}}_{\text{finite (large)}} \underbrace{\int \frac{dx \, dz}{1-z} \left[\sigma_h(zxp) - \sigma_h(xp)\right] q(x,\mu^2)}_{\text{finite}}$









# properties of LO splitting functions

in general, quarks and gluons can split into quarks and gluons -> 4 functions

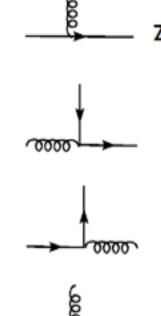
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$$P_{qg}^{(0)} = P_{\bar{q}g}^{(0)} = T_R \left( z^2 + (1-z) \right)$$



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# reaching for precision

 $\begin{array}{rcl} P_{\rm EE}^{(0)}(x) &=& C_F \left( 2 p_{\rm qq}(x) + 3 \delta(1-x) \right) \\ P_{\rm PE}^{(0)}(x) &=& 0 \\ P_{\rm QE}^{(0)}(x) &=& 2 n_f p_{\rm QE}(x) \\ P_{\rm EQ}^{(0)}(x) &=& 2 C_F p_{\rm EQ}(x) \\ P_{\rm EE}^{(0)}(x) &=& C_A \left( 4 p_{\rm EE}(x) + \frac{11}{3} \delta(1-x) \right) - \frac{2}{3} n_f \delta(1-x) \end{array}$ 

LO: 1973

## reaching for precision

 $\begin{array}{rcl} P_{\rm gs}^{(0)}(x) &=& C_F(2p_{\rm qq}(x) + 3\delta(1-x)\,) \\ P_{\rm ps}^{(0)}(x) &=& 0 \\ P_{\rm qg}^{(0)}(x) &=& 2n_f p_{\rm qg}(x) \\ P_{\rm gq}^{(0)}(x) &=& 2C_F p_{\rm gq}(x) \\ P_{\rm gg}^{(0)}(x) &=& 2C_F p_{\rm gq}(x) \\ P_{\rm gg}^{(0)}(x) &=& C_d \left(4p_{\rm gg}(x) + \frac{11}{3}\delta(1-x)\right) - \frac{2}{3}n_f \delta(1-x) \end{array}$ 

# LO: 1973

Curci, Furmanski, Petronzio; Floratos et al., ...

$$\begin{split} P_{\rm ns}^{(1)+}(x) &= 4C_{\rm s}C_{\rm F}\left(p_{\rm eq}(x)\left[\frac{67}{18}-\zeta_2+\frac{11}{6}H_0+H_{0,0}\right]+p_{\rm eq}(-x)\left[\zeta_2+2H_{-1,0}-H_{0,0}\right]\right.\\ &+\frac{14}{3}(1-x)+\delta(1-x)\left[\frac{17}{24}+\frac{11}{3}\zeta_2-3\zeta_3\right]\right)-4C_{\rm F}n_f\left(p_{\rm eq}(x)\left[\frac{5}{9}+\frac{1}{3}H_0\right]+\frac{2}{3}(1-x)\right.\\ &+\delta(1-x)\left[\frac{1}{12}+\frac{2}{3}\zeta_2\right]\right)+4C_{\rm F}^2\left(2p_{\rm eq}(x)\left[H_{1,0}-\frac{3}{4}H_0+H_2\right]-2p_{\rm eq}(-x)\left[\zeta_2+2H_{-1,0}-H_{0,0}\right]-(1-x)\left[1-\frac{3}{2}H_0\right]-H_0-(1+x)H_{0,0}+\delta(1-x)\left[\frac{3}{8}-3\zeta_2+6\zeta_3\right]\right) \end{split}$$

$$\begin{aligned} P_{\rm ns}^{(1)-}(x) &= P_{\rm ns}^{(1)+}(x)+16C_{\rm F}\left(C_{\rm F}-\frac{C_{\rm A}}{2}\right)\left(p_{\rm eq}(-x)\left[\zeta_2+2H_{-1,0}-H_{0,0}\right]-2(1-x)-(1+x)H_0\right) \end{split}$$

$$\begin{aligned} P_{\rm ps}^{(1)}(x) &= 4C_{\rm F}n_f\left(\frac{20}{9}\frac{1}{x}-2+6x-4H_0+x^2\left[\frac{8}{3}H_0-\frac{56}{9}\right]+(1+x)\left[5H_0-2H_{0,0}\right]\right) \end{aligned}$$

$$\begin{split} P_{gq}^{(1)}(x) &= 4C_{d}C_{F}\left(\frac{1}{x} + 2p_{gq}(x)\left[H_{1,0} + H_{1,1} + H_{2} - \frac{11}{6}H_{1}\right] - x^{2}\left[\frac{8}{3}H_{0} - \frac{44}{9}\right] + 4\zeta_{2} - 2 \\ -7H_{0} + 2H_{0,0} - 2H_{1}x + (1+x)\left[2H_{0,0} - 5H_{0} + \frac{37}{9}\right] - 2p_{gq}(-x)H_{-1,0}\right) - 4C_{F}n_{f}\left(\frac{2}{3}x - p_{gq}(x)\left[\frac{2}{3}H_{1} - \frac{10}{9}\right]\right) + 4C_{F}^{2}\left(p_{gq}(x)\left[3H_{1} - 2H_{1,1}\right] + (1+x)\left[H_{0,0} - \frac{7}{2} + \frac{7}{2}H_{0}\right] - 3H_{0,0} \\ +1 - \frac{3}{2}H_{0} + 2H_{1}x\right) \end{split}$$

 $+4(1-x)[H_{0,0}-2H_0+xH_1]-4\zeta_2x-6H_{0,0}+9H_0]+4C_Fn_f(2p_{qg}(x)[H_{1,0}+H_{1,1}+H_2])$ 

 $-\zeta_{2}$  + 4x<sup>2</sup> [H<sub>0</sub> + H<sub>0,0</sub> +  $\frac{5}{2}$ ] + 2(1 - x) [H<sub>0</sub> + H<sub>0,0</sub> - 2xH<sub>1</sub> +  $\frac{29}{4}$ ] -  $\frac{15}{2}$  - H<sub>0,0</sub> -  $\frac{1}{2}$ H<sub>0</sub>)

$$\begin{split} P_{gg}^{(1)}(x) &= 4C_{a}n_{f}\left(1-x-\frac{10}{9}p_{gg}(x)-\frac{13}{9}\left(\frac{1}{x}-x^{2}\right)-\frac{2}{3}(1+x)H_{0}-\frac{2}{3}\delta(1-x)\right)+4C_{a}^{-2}\left(27\right)\\ &+(1+x)\left[\frac{11}{3}H_{0}+8H_{0,0}-\frac{27}{2}\right]+2p_{gg}(-x)\left[H_{0,0}-2H_{-1,0}-\zeta_{2}\right]-\frac{67}{9}\left(\frac{1}{x}-x^{2}\right)-12H_{0}\\ &-\frac{44}{3}x^{2}H_{0}+2p_{gg}(x)\left[\frac{67}{18}-\zeta_{2}+H_{0,0}+2H_{1,0}+2H_{2}\right]+\delta(1-x)\left[\frac{8}{3}+3\zeta_{3}\right]\right)+4C_{F}n_{f}\left(2H_{0}+\frac{2}{3}\frac{1}{x}+\frac{10}{3}x^{2}-12+(1+x)\left[4-5H_{0}-2H_{0,0}\right]-\frac{1}{2}\delta(1-x)\right] \\ &+\frac{2}{3}\frac{1}{x}+\frac{10}{3}x^{2}-12+(1+x)\left[4-5H_{0}-2H_{0,0}\right]-\frac{1}{2}\delta(1-x)\right] \end{split}$$

# **P**<sub>ij</sub> **@** NNLO: a landmark calculation

10000 diagrams, 10<sup>5</sup> integrals, 10 man years, and several CPU years later:

# **P**<sub>ij</sub> **@ NNLO: a landmark calculation**

#### 10000 diagrams, 10<sup>5</sup> integrals, 10 man years, and several CPU years later:

 $\delta_{\mu}^{(0)}(z) = (m_{\mu}^{\prime})_{\mu\nu}^{\prime} \left[ \frac{1}{2} (z + z^{2}) \left[ \frac{1}{2} (z + z^{2}) + \frac{1}{2} (z + z^{2}) (z - 2) + (z - 2) \right] \right]$  $-H_{1,1}\left[+\frac{1}{2}\frac{1}{2}-r^{2}\left(\frac{14}{2}r_{1}+H_{1}+H_{2}+\frac{1}{2}H_{1}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{1}+\frac{2}{2}H_{1}+\frac{2}{2}H_{1}+\frac{2}{2}H_{1}+\frac{2}{2}H_{1}+\frac{2}{2}H_{1}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_{2}+\frac{2}{2}H_$  $\mathbf{M}_{1,1,2} + \mathbf{M}_{2,1,2} + \mathbf{M}_{2,1,2} + (1 - \alpha) \left[ \frac{2 \pi i}{1} \mathbf{H}_{2} + \frac{1 \pi i}{1} + \frac{2 \pi i}{10} \mathbf{H}_{2,2} - \frac{1 \pi}{1} \mathbf{H}_{-1,2,2} + \mathbf{M}_{2,2,2,2} \right]$ - No - Morthar Ray - Roy - New - New - New - New - $\frac{1}{2} [t_{12} + [t_{123} + [t_{123}]] + (1 + \pi) \Big[ \frac{1}{12} [t_{12} + \frac{31}{2} t_{12} + \frac{11}{12} t_{12} + \frac{11}{12} t_{12} - \frac{11}{12} t_{12}$  $\frac{1}{2} \sum_{i=1}^{n} \sum_{j=1}^{n} \sum_{i=1}^{n} \sum_{j=1}^$ (Kaar Shorthigh (RainKoust Shorthon (N)) (Martho  $+ 7 \eta_{(1,1)} - \frac{1}{2} \zeta_{2}^{2} + 4 \theta_{-} ( \chi + 4 \theta_{0} \zeta_{2} - \frac{12}{3} \theta_{0} - \frac{2 \theta_{0}}{12} \theta_{0} - \frac{2 (\theta_{0} - 2 ) \theta_{0}}{12} - \frac{10}{12} \theta_{0} -$  $\frac{11}{12} \sigma_{0} r_{0} + \frac{11}{12} r_{0} + \frac{1}{2} \sigma_{0} r_{0} + \frac{1}{2} \sigma_{0}^{2} \frac{12}{12} \sigma_{0} r_{0} + \frac{12}{12} \sigma_{0} + \frac{11}{12} + \frac{11}{12} \sigma_{0} + \frac{1}{2} \sigma_{0}$ 5- mp - m - m - m - ] - m - + ( [m - 2 - m - p - ] / m - p - 2  $\frac{1}{2}(1+1)\left[\frac{1}{2}(t_{1}-\frac{1}{2}(t_{1}+t_{1})+2t_{1})-2t_{1}+2t_{1}(t_{1}+\frac{12}{2}(t_{1})+2t_{1}(t_{1})\right]+2t_{1}(t_{1}^{2})t_{1}\left[\frac{12}{2}(t_{1}^{2})\right]$ テルコールコー 茶れー茶・子に一子れった。 アコールコールシーパー  $\left[\frac{1}{2} |\mathbf{R}_{0,0}| + \frac{1}{2} \left[\frac{1}{2} - 2 \right] \left[\frac{1}{2} |\mathbf{R}_{0,0} - \frac{1}{2} |\mathbf{R}_{0,0} - \mathbf{R}_{0,0} + \frac{1}{2} + \frac{1}{2} |\mathbf{R}_{0,0} + \mathbf{R}_{0,0} - \mathbf{R}_{0,0} - \mathbf{R}_{0,0} \right]$ 17-18 Martin - 200 - 201 - 201 - 201 - 201 - 10, 100 - 1942 + 1944 - 49444 + 1944 - 1944 - 1944 - 1944 - 1944 - 1944 - 1944

$$\begin{split} & \frac{1}{2} \left[ \hat{V}_{11} \right] = 1 + C_{11} \left[ \hat{V}_{21} \right] + \frac{1}{2} \hat{V}_{11} \left[ \hat{V}_{21} \right] + \hat{V}_{21} \left[ \hat{V}$$

 $2R_{22} - \frac{12}{3}R_{22} - 2R_{-1/2} - \frac{11}{3}R_{2/2} + \frac{12}{3}R_{2/2} + \frac{12}{3}R_{2} - \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3}R_{2} +$ - 30 - 300 - 300 - 300 - 300 - 40<sup>2</sup> - 300 - 80 - 300 - 800 - 800 -1994 - Washington - H. S. 1996 - 1994 - 199 - 1994 - 1994 - 1994 - 1994  $+ \begin{bmatrix} \mathbf{R}_{1,1,2} - \frac{\mathbf{R}_{1,2}}{2} \mathbf{R}_{2,1} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1,2} \end{bmatrix} + \begin{bmatrix} \mathbf{R}_{1,1} - \mathbf{R}_{1,1,2} + \mathbf$  $-481 + \frac{100}{10} (s_1 + \frac{100}{10} (s_1) + s_{20}(s_1) \frac{1}{2} (s_1 + \frac{100000}{10000} + \frac{1}{2} (s_{20} + \frac{100}{10} + s_{20} + \frac{100}{10} (s_{20} + \frac{1$ دِرْنِ 10 هَانُ 10 هَانُ 10 هِانُ ( - يُور د ( يُوَال د رويا 10 د رواله د رواله د رواله د رواله د  $+\frac{100}{10}$  (1) + (10) +  $\frac{17}{2}$  (2) +  $\frac{17}{2}$  (1) + (10) (2) -  $\frac{10}{2}$  (1) -  $\frac{10}{2}$  (1) (2) - 47 (2) - 17 (2)  $\begin{array}{c} -2 H & \leq \gamma + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 2 H + 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$-36_{1}\ell_{2}+\frac{10}{12}\eta_{1,1}+\frac{10}{7}\eta_{1,1}+\frac{20}{7}\eta_{1}-\eta_{2}\eta_{2}+\eta_{3}\eta_{3}+(1+\eta_{1}^{2}\eta_{1}^{2}-\eta_{1})\eta_{3}+\frac{20}{7}\eta_{1}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}$  $-\frac{1}{2}$ -The Theorem - Theorem - Theorem - South - State - Theorem 186.0.0 1 Januar 1800 - Jan 1 Januar 1800 - Januar 1800 - 1800 - 1800 Co-mo-in-in-in-in-in-in-in-in- $+ \frac{1}{2} \mathbf{R}_{1} + \frac{1}{2} \mathbf{R}_{1,1} - 2 \mathbf{R}_{1,1} + \frac{1}{2} \mathbf{R}_{-1,-1,1} + \frac{1}{2} \mathbf{R}_{2,1,2} - \mathbf{R}_{2,1,1} - \mathbf{R}_{2,1,1} - \mathbf{R}_{2,1,1} + \frac{1}{2} \mathbf{R}_{2,1,2} - \mathbf{R}_{2,1,1} - \mathbf{R}_{2,1,1} + \frac{1}{2} \mathbf{R}_{2,1,2} - \mathbf{R}_{2,1,1} - \mathbf{R$ 

-mar [lig]-agent[]-mar [max-]mar [mar ]mar -Mg - Mg - 60 g + 60 g - g - 10 gg + Mg / + 80 y + 10 g - 50 gg Recordence Reg Roy Roy Rev Thus  $-221(j) + \frac{1}{2} + 2^{2} \left[ \frac{1}{2} + \frac{1}{2} + \frac{1}{2} + 2^{2} \left[ \frac{1}{2} + \frac{1}{2} + 2^{2} \left[ \frac{1}{2} +$  $(B_{1,1,2}+\frac{114}{3}B_{1,2}+\frac{101}{3}B_{1,2}+\frac{10}{3}C_{2}^{-1}+\frac{101}{3}C_{2}^{-1}+\frac{101}{3}B_{1}-\frac{5}{3}B_{1,2}+\frac{11}{3}B_{1,2}-\frac{10}{3}B_{1,2}-\frac{10}{3}B_{1,2}$  $\frac{110}{14} \theta_1 - \frac{12}{16} \theta_2 - 2\theta_{10} - \frac{100}{16} \theta_2 - \frac{100}{16} \theta_{100} - \frac{100}{16} \theta_1 + \frac{100}{16} \theta_1 + 2\theta_{100}$ 1314 National Works House Made Marcare Made Many May يكروالا جرين الله - (يكر الاتحادث وجري الله جريزواله - يرالف والف بري الله.  $(0, \frac{1}{2} - 20_{12}) + 2\delta_{22} + 2\delta_{22} - 20_{12} - 20_{12} + \frac{1}{2}\delta_{12} + \frac{1}{2}\delta_{12} + \frac{1}{2}\delta_{22} + \frac{1}{2}\delta_{22} + \frac{1}{2}\delta_{22}$ march marches march - march - march  $-\frac{10}{7}\zeta^2-\frac{10}{12}\zeta_2-\frac{10}{7}\zeta_2-\frac{10}{7}\zeta_2-\frac{10}{7}\zeta_2+\frac{10}{7}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{12}\zeta_2+\frac{10}{1$  $+\frac{1}{2}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}$  $+2\delta \zeta_{2}q_{1}^{2}\left[\frac{1}{2}g_{2}(a)\left[B_{1}-B_{2}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B$ Star Star (an See See Star Store Sector) - Stor (Mar  $-\frac{H}{2M} + \frac{1}{2} n_{1} h_{2} \left( p_{0}(z) \left[ H_{1,1} + \frac{H}{4} H_{1,1,2} - \frac{H}{2} n_{1,2} + \frac{T}{2} n_{1,2} + \frac{H}{2} n_{1,1,2} + \frac{H}{2} n_{1,1,2} + \frac{H}{2} n_{1,2} - \frac{H}{2} n_{1,2} \right)$ my Car Car We Ca. Was at Gauss me an  $-\frac{10}{10}R_{12} - \frac{10}{10}R_{12} - \frac{1}{10} - \frac{1}{10}R_{12} - \frac{10}{10}R_{12} -$ #Tell = Theo = The lot = The \$4 + Theo = The \$2 = Theory = Theory = Theory = Theory 24.11-24.11 + A. ( x) = 1.12 - 21.11 - 21.11 + 21.11 + 21.12 - 21.11 Washington - Jug Kantham Man Jum 「いいいういいないぞれ、「いい」、「いい」でないのとい

 $+ \frac{1}{(p_{0})^{(1)}} [ R_{12} - R_{12} - R_{12} + R_{22} + \frac{1}{(2} R_{12} + 2R_{12} - \frac{1}{(2} R_{1} + \frac{1}{(2} R_{12} - \frac{1}$  $+20_{11}+\left[20_{11}+\left[\frac{10}{2}h_{0}\right]+\left[\frac{10}{2}h_{0}+c\right]\left[\frac{10}{2}h_{0}-\frac{10}{2}+R_{11}\right]+\left[\frac{10}{2}+c\right]\left[20_{11}m-2h-\frac{1001}{24}\right]$ to man har frankra that the provide the  $\frac{1}{2}\left(1-\frac{10}{2}\left(1-\frac{10}{2}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\left(1-\frac{10}{2}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{2}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)\right)-\frac{10}{20}\left(1-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{10}{20}\left(1-\frac{10}{20}\right)-\frac{1$  $+\frac{12\pi i}{2\pi}+2\pi \frac{1}{2}-2\pi \frac{1}{2}\pi \frac{1}{2}\pi$ Phys. Rep. Berry Mark 1994 (1994) - Mark 1994 (1994) - Mark  $-2h_{1,2}+2h_{2,2}-\frac{1}{2}h_{1,2}-\frac{1}{2}h_{1,2}+\frac{1}{2}h_{1,2}+\frac{1}{22}h_{1}-\frac{1}{22}h_{2}+\frac{1}{22}h_{2}+h_{2}(-1)|H|<\infty$ یکی 10–روں یا 18–روں 18–روں 10–<sup>1</sup> <mark>کی</mark>ا – روا<del>ز د</del>یار 18–روں یا 19–ر -H (100) + H - A (H) (10 + H(1) - 100/2 + 100/2 + H(2 - H/2 + H(1) + 10/2 + H(1))  $(0_1+0_{1,1})+(0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}+0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,1})_{1}+(0_{1,1}-0_{1,$ - Marcold Marco + 11+2 - 1492 + 1949 - 11-12 - 12-2942 + 29412 + 19422 - 2942 - 11-2942 tru- Fran - Frank)

$$\begin{split} & \frac{1}{2} \left[ (x_1 - x_2)_{1 \leq i} \left\{ \left| \frac{1}{2} (x_1 + x_2)_{1 \leq i} - \frac{1}{2} (x_1 - \frac{1}{2} (x_2 - \frac{1}{2}$$

 $|\mathbf{x}_{1} + [\mathbf{x}_{1}]_{2} + |\mathbf{x}_{1}|_{2} - \frac{12}{10}|\mathbf{x}_{2} + \frac{12}{10}|\mathbf{x}_{2} - \frac{12}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} + \frac{10000}{10}|\mathbf{x}_{1}$ 10 mar 10 mar 100 mar 100 mar 10 mar 10 mar 10 mar 10 (N.c.o) - N.o. 100.01 (Rep. 20.01 100.00 These Mar Share Shake Shake Share Share Share Share  $+118_{12}+857_{12}-110_{12}=200_{12}(1+20_{12})-1005_{12}-1005_{12}-1000_{12}-100_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{12}-1000_{1$  $= \frac{1}{2} \frac{1}{2} \frac{1}{2} \frac{1}{2} \frac{1}{2} - \frac{1}{2} \frac{1}{2} \frac{1}{2} + \frac{1}{2} \frac{1}{2} \frac{1}{2} + \frac{1}{2} \frac{1$ 
$$\begin{split} & -\frac{1}{2} \mathcal{C}_{0,1} - \frac{1}{2} \mathcal{C}_{0,1} - \frac{1}{2} \mathcal{C}_{0,1} + \frac{1}{2} \mathcal{C}_{0,1} + \frac{1}{2} \mathcal{C}_{0,1} \mathcal{C}_{0,1} + \frac{1}{2} \mathcal{C}_{0,1} - \frac{1}{2} \mathcal{C}_{0,1} + \mathcal{C}_{0,1,1} + \mathcal{C}_{0,1,2} \\ & -\mathcal{C}_{0,1} - \mathcal{C}_{0,1,1} + \frac{1}{2} \mathcal{C}_{1,1} - \mathcal{C}_{1,1} + (1 - 1) \left[ \frac{1}{12} \mathcal{C}_{1,1} - \frac{4 \mathcal{C}_{0,1}}{2 \mathcal{C}_{0,2}} - \frac{\mathcal{C}_{0,1,1}}{2 \mathcal{C}_{0,1,1}} - \frac{\mathcal{C}_{0,1,1}}{2 \mathcal{C}_{0,1,1}} + \frac{\mathcal{C}_{0,1,1}}{2 \mathcal{C$$
 $+ \left[ \frac{1}{2} \left[ \left[ \frac{1}{2} \left[ \frac$ SLor-Nort Northan Northan Northant Northant  $+\frac{e^{2}}{2}\eta_{1}-\frac{e^{2}}{2}\eta_{2}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{2}-\frac{e^{2}}{2}\eta_{2}-\frac{e^{2}}{2}\eta_{2}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{2}$  $+ \frac{1}{2} + \frac{1}{2} \mathbf{a}_{i} - \frac{1}{2} \mathbf{a}_{i} \mathbf{a}_{i} - \frac{1}{2} \mathbf{a}_{i} \mathbf{a}_{i} + \frac{1}{2} \mathbf{a}_{i+1} + \frac{1}{2} \mathbf{a}_{i+1} + \frac{1}{2} \mathbf{a}_{i}^{2} + \mathbf{a}_{i+1}^{2} - \mathbf{a}_{i} \mathbf{a}_{i} \mathbf{a}_{i} + \frac{1}{2} \mathbf{a}_{i}^{2} + \mathbf{a}_{i+1}^{2} - \mathbf{a}_{i} \mathbf{a}_{$ 1184 - Mura - Murger Mouse - Murger - Musari Rosser (Musari Musa Marillo Hard Rose Rose Rose Rose Rose Rose Rear Rear Red (Record - Rear) 127-18 Nov-Nov-No. -Mu-Mu-Mulan, -Mushur Mushik-4-4-24-24-24-24  $[[\mathbf{R}_{i}+]]\mathbf{R}_{i,i}+[[\mathbf{R}_{i,jk}+]]\mathbf{R}_{i}+[[\mathbf{R}_{i}+]]\mathbf{R}_{i,j}-[[\mathbf{R}_{i}+]]\mathbf{R}_{i,j}+[[\mathbf{R}_{i,j}+]]\mathbf{R}_{i,j}+[[\mathbf{R}_{i,j}+]]\mathbf{R}_{i,j}]$  $+ \frac{1}{10} \frac{1}{10} \frac{1}{10} + \frac{10}{10} \frac{1}{10} \frac{1}{$ 

 $+128_{0.047} - \frac{305}{1.6} + \frac{47}{1.84} + \frac{3}{1.84} + \frac{47}{1.6} + \frac{47}{1.6} + \frac{47}{1.6} + \frac{3}{1.6} + \frac{10}{1.6} +$  $\frac{10}{10}\mathbf{R}_{10} + \frac{10}{10}\mathbf{R}_{10} - \mathbf{R}_{10} - \mathbf{R}_{10} + \left(1 + \alpha\right) \frac{10}{10}\mathbf{R}_{10} - \frac{90}{10}\mathbf{R}_{10} + \frac{100}{10}\mathbf{R}_{1} + \frac{100}{100}\mathbf{R}_{1}$  $\begin{array}{c} -m_{1,1} + \frac{m_{1}}{m_{1}} m_{1,1} + \frac{m_{1}}{m_{1}} + \frac{m_$  $-88_{-1,00}\left[-\frac{24}{248}(1-r)\right]+10^{2}\mu_{0}^{2}\left(\frac{10}{34}r_{0}-\frac{1}{24}r_{0}r_{0}-\frac{1}{24}r_{0}r_{0}(1+\frac{10}{34}r_{0}-r^{2})\left[\frac{1}{2}-7r_{0}\right]$  $+[1+\alpha]\left[\frac{1}{\alpha}\mathbf{R}_{1}-\frac{1}{\alpha}\mathbf{R}_{2}\right]+\left[(1+\alpha)\left[\mathbf{p}+\frac{1}{\alpha}\mathbf{R}_{2}-\frac{1}{\alpha}\mathbf{R}_{2}-\mathbf{R}_{1}\right]+\frac{1}{\alpha}\mathbf{R}_{2}-\mathbf{R}_{1}\right]$ 100/16/2 - 10 - 10- 10- 100 - 1110 - 20-1 - 10-10 TO  $-\frac{29}{24} + 4\zeta_2 - 11 + \zeta_2 - \frac{10}{2} \delta_{12} - \frac{10}{2} \delta_{12} - \frac{2}{2} \delta_{12} - 8_{232} - \frac{2}{2} \delta_{12} - 8_{13} + \frac{10}{2} \delta_{12} (-1) \Big[ \zeta_2$  $+28_{-1,0}+\frac{7}{12}(r_{0}r_{1}^{\prime}+8q_{1}^{\prime})+\frac{1}{2}(\frac{1}{2}-r^{2})\left[r_{1}-r_{0}r_{2}-\frac{13}{2}r_{1}+\frac{7+9}{12}-18(r_{2}+\frac{29}{12}r_{1})\right]$ -[m\_1-m\_2]+1]+2][[m-[0+]m\_2+0+0+m\_2+-[m\_2+  $-(\theta_{1,2_{1},2_{2},2}+(\theta_{1,2_{1},2_{1},2_{1}}-\frac{2}{2}\theta_{1})_{2}+\frac{477}{2}\theta_{1}+(\theta_{1,2}+\frac{2}{2}\theta_{1,2_{2}}]+(1+\varepsilon)\Big[\frac{100}{10}\theta_{1}+\frac{11}{2}\theta_{1,2}\theta_{2}$  $(M_{1,1} \cdot M_{1,1} \cdot [H_{1,1} \cdot [H_{1,1} \cdot [H_{1,1} \cdot [H_{1}] \cdot [L_{1} \cdot M_{1,1} \cdot M_{1} \cdot [H_{1} \cdot H_{1,1}]])]$  $+ \frac{1}{2} e^2 + \overline{m}_{-1,0} + e^{\left\lceil \frac{1}{2} \right\rceil} \overline{m}_{0,0} + \frac{1}{2} \overline$  $- \delta(1 - \pi) \Big[ \frac{2\pi}{100} + \frac{1}{2} f_{2} + \frac{1}{10} f_{2}^{2} + \frac{1}{10} f_{2}^{2} + \frac{1}{10} f_{2}^{2} \Big] + 100 f_{2}^{2} \Big[ e^{2} (100 + \mu + 100 \mu_{2}^{2} - \frac{100}{10} \theta_{0,\mu} + 100 \mu_{1}^{2} - \frac{100}{10} \theta_{0,\mu} + \frac{100}{10} \theta_{0,\mu} \Big] + 100 f_{2}^{2} \Big] + 100 f_{2}^{2} \Big[ e^{2} (100 + \mu + 100 \mu_{1}^{2} - \frac{100}{10} \theta_{0,\mu} + \frac{100}{10} \theta_{0,\mu} + \frac{100}{10} \theta_{0,\mu} \Big] + 100 f_{2}^{2} \Big$  $-44 (t_{1,1,2}-\frac{110}{2} (t_1-\frac{11}{2} (t_{1,2}+\frac{11}{2} (t_1+\frac{110}{2} (t_1))+3 (t_2) (t_1) (\frac{110}{2} (t_1-\frac{11}{2} (t_1-\frac{11}{2} (t_1'+\frac{11}{2} (t_1))))))$ - Martin La + Martin - Ma 00-40 - 00-40 - 00-40 - 10-40 - 00-40 - 00-40 - 00-44 - 00-1 Mar 1 May 1 May 1 Mar 1  $+60_{1,2}\left[-1\right]_{0}\left[-1\right]_{0}\left[\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\right]\frac{1}{2}\right]\right]}\right]$ الله من المنظل المن الله من ال المنطقة المنظلة المنظلة المنطقة - This - Theorem + (1-+) [2007 - Theor - This - Theorem - Theorem 
$$\begin{split} & H_{M,M} = \frac{1}{2} H_{M,M} - \frac{1}{2} H_{M,M} = H_{M,M,M} + H_{M,M,M} + H_{M,M,M} + H_{M,M} + \frac{1}{2} H_{M,M} + H_$$

 $\boldsymbol{S}_{\boldsymbol{k}}^{(2)}[\boldsymbol{k}] = \log_{\boldsymbol{k}} \boldsymbol{g}_{\boldsymbol{k}} \cdot \boldsymbol{g}_{\boldsymbol{k}} \left[ \left[ \boldsymbol{k} \right] \left[ \frac{1}{2} \boldsymbol{R}_{\boldsymbol{k}} - \frac{1}{2} + \boldsymbol{R}_{\boldsymbol{k}} \cdot \boldsymbol{R}_{\boldsymbol{k} + \boldsymbol{k}} - \boldsymbol{R}_{\boldsymbol{k} + \boldsymbol{k}} + \frac{1}{2} \boldsymbol{R}_{\boldsymbol{k}} \cdot \boldsymbol{R}_{\boldsymbol{k} + \boldsymbol{k}} \right]$  $= \sum_{k=1}^{n} ([a_{k,1} + a_{k,1} + \frac{a_{k,1}}{a_{k,1}} + \frac{a_{k$ • [1] • • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] • (1] - Ja- Jaw - Jaw - Mar (a. - Jac) - M- Ja- Jaw m  $\frac{1}{2} R_{12} + \frac{11}{12} R_{2} + \frac{11}{12} R_{22} - R_{13} - \frac{1}{12} R_{23} + \frac{11}{12} R_{-12} + R_{2} + 2R_{1} + \frac{1}{12} R_{-12} + \frac{11}{12} R_{1}$  $+ \frac{10}{10} R_{12} + R_{13} - 2R_{13} - \frac{1}{3} R_{13} + \frac{10}{3} R_{13}^2 + R_{1,12} + R_{13}^2 + \frac{1}{3} R_{1,13} - 2R_{13}^2 + \frac{1}{3} R_{13} + \frac{$ But has be much the music has be ين»- يُرُو(د- (د) إين 10+ي رد ر 10+ي رد H - يود ر 10 - يود ر 10  $+\frac{11}{10}R_{12} - \frac{3\pi i}{42}R_{12} - \frac{4\pi i}{10}R_{12} - \frac{11}{10}R_{12} - R_{13} - 2R_{13} - \frac{4\pi^{2}}{42}R_{1} - 2R_{23} - \frac{11}{10}R_{13}$ \$1- for + Bolo - or 6 for - or 1. Bolo - for - Elle  $\frac{411}{100} - \frac{101}{2} \zeta_{2} - \frac{100}{10} \zeta_{2} + \frac{1}{2} \zeta_{2} \zeta_{2} - \frac{81}{2} \delta_{2} + \frac{20}{2} \delta_{2} - \frac{101}{2} \zeta_{2} - 111 \zeta_{2} + 21 \zeta_{2}$  $-2H_{-1,-1,0} + \frac{1}{10}H_{0,0} + \frac{1}{10}H_{0,0,0} - \frac{1}{10}H_{0,0} - \frac{1}{10}H_{0,0} - \frac{10}{10}H_{0,0} - \frac{10}{10}H_{0,0} - \frac{10}{10}H_{0,0}$ 

The first final and while in the first first  $+\mathbf{R}_{n,0,0} - \frac{20}{100}\mathbf{R}_{n,0} + \frac{2}{3}\mathbf{R}_{n,0,0} + \mathbf{R}_{n,0} + 2 \cdot \mathbf{z} + \frac{2}{3}\mathbf{R}_{0} + \frac{20}{10}\mathbf{R}_{0} - \frac{20}{10}\mathbf{R}_{0,0} - \mathbf{R}_{n,0}$ - How 
$$\begin{split} &+(1+1)\frac{12}{2}h_{1}f_{2}-\frac{16}{2}h_{1}+\frac{12}{2}h_{2}+\frac{16}{20}h_{1}+\frac{129}{20}h_{1}-\frac{129}{20}f_{2}+\frac{1}{2}(1+1)f_{2}-\frac{16}{2}f_{2}^{2}+\frac{129}{2}h_{1}f_{2}\\ &+(2)h_{1}f_{2}-(2)h_{2}^{2}-(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^$$
 $\frac{10}{2} R_{1} - \frac{10}{2} R_{1} \left[ k_{1} \right] - 2 R_{10} - \frac{4}{2} R_{100} + 2 R_{10} + \frac{6}{2} R_{10} - 2 R_{1} - 2 R_{10} + \frac{10}{2} k_{10} + \frac{10}{2} k_{$ 100 No 100 100<sup>2</sup> 100 - Multi-1Mar 1Mar 1Maas 181 - 40  $-i_{2}i_{2}+\frac{1}{2}i_{2}+\frac{1}{2}i_{2}+\frac{1}{2}i_{2}^{2}+\frac{1}{2}i_{2}^{2}-2i_{2}^{2}\Big]+22i_{2}\mu_{2}^{2}\Big[\frac{1}{2}i_{1}^{2}\Big[\frac{1}{2}i_{2}+2i_{1}-i_{2}+2i_{2}\mu_{2}-i_{1}^{2}\Big]+\frac{1}{2}i_{2}$  $-\frac{1}{2}(1-\frac{1}{2}(h_{0}-\frac{1}{2}(h_{0}+1)+\frac{1}{2}(\frac{1}{2}-h^{2})\left[\frac{1}{2}(h_{0}-2h_{0}-h_{0}-\frac{1}{2})(1-h^{2})\left[\frac{1}{2}(h_{0}+\frac{1}{2}(h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}$  $\left[ H_{1,1} - 2 H_{2,1} \right] + \frac{11}{14\pi} H(1-\alpha) + 10 H_{2,1}^{-1} h_{2} \left[ \frac{\alpha}{2} d^2 \left[ \frac{\alpha + 1}{14\pi} + \frac{11}{2} H_{2,1} + \frac{1}{2} H_{2,2} - H_{2,2} - \frac{\alpha}{2} + \frac{1}{2} H_{2,2} \right] \right]$  $- \pi_{12} + \frac{1}{2} \pi_{12} + \frac{1}{2} \pi_{14} + \pi_{14} - \pi_{14} - \frac{1}{2} \pi_{14} + \frac{1}{2} \frac{1}{4} - \pi_{14} - \frac{1}{2} \pi_{14} - 2 (f_{1} - 2 (g_{1} + 2 (g_{2} + g_{2} + g_$  $+\frac{10}{10}\mathbf{K}_{0}-\frac{100}{10}+\mathbf{IH}_{0,0}+\frac{100}{10}\mathbf{x}_{0,0}^{2}-\mathbf{IH}_{0,0}+\mathbf{K}_{0,0}-\frac{10}{10}\mathbf{K}_{0}+\frac{10}{10}\mathbf{K}_{0}+\mathbf{IH}_{0,0}+\mathbf{X}_{0,0}^{2}-\mathbf{K}_{0,0}^{2}$  $-\frac{24}{3} f_{0,0} - 4 f_{0,0} + \frac{24}{3} f_{0,0} - 4 f_{0,0} - 4 f_{0,0} - 10 f_0 + 2 f_0^2 + 10 g_0^2 - 10 f_0 - 4 f_0 g_0^2$  $(M_{1,1,1,2}+\frac{12}{7}M_{1,2,2}+M_{1,1,2,2}+2M_{2,2,2}-2M_{2,2,2}-4M_{2,2}+2M_{2,2,2}+2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2}-2M_{2,2} (2n_{11} + n_{12} + 2n_{12} + \frac{1}{2}4(1 + a))$ .

Moch, Vermaseren, Vogt 2004

# **P**<sub>ij</sub> **@ NNLO: a landmark calculation**

#### 10000 diagrams, 10<sup>5</sup> integrals, 10 man years, and several CPU years later:

 $\delta_{\mu}^{(0)}(z) = (m_{\mu}^{\prime})_{\mu\nu}^{\prime} \left[ \frac{1}{2} (z + z^{2}) \left[ \frac{1}{2} (z + z^{2}) + \frac{1}{2} (z + z^{2}) (z - 2) + (z - 2) \right] \right]$  $-H_{1,1}\left[+\frac{1}{2}\frac{1}{2}-r^{2}\left(\frac{14}{2}r_{1}+H_{1}+H_{2}+\frac{1}{2}H_{1}-\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+H_{1}r_{2}-\frac{1}{2}H_{1}\right]$  $\mathbf{M}_{1,1,2} + \mathbf{M}_{2,1,2} + \mathbf{M}_{2,1,2} + (1 - \alpha) \left[ \frac{2 \pi i}{1} \mathbf{H}_{2} + \frac{1 \pi i}{1} + \frac{2 \pi i}{10} \mathbf{H}_{2,2} - \frac{1 \pi}{1} \mathbf{H}_{-1,2,2} + \mathbf{M}_{2,2,2,2} \right]$ - No - Morthar Ray - Roy - New - New - New - New - $\frac{1}{2} (t_1 + (t_{1,1} + (t_{1,1})) + (1 + 4) \left( \frac{1}{12} (t_1 + \frac{31}{2} (t_1 + \frac{31}{2} (t_1 + \frac{11}{12} (t_1 + \frac{11}{12} (t_1 - \frac{112}{22} (t_1 + \frac{11$  $\frac{1}{2} n_{2,1} + \frac{14}{2} n_{1,2} + n_{2,1} + \frac{12}{2} n_{2,2} + \frac{12}{2} n_{2,2} + \frac{14}{2} n_{2,2} + \frac{$ Rus-Burthis Burthout Burtho R. Hurtho  $+ \frac{1}{2} (z_{12} z_{23} - \frac{1}{2} z_{2}^{2} + 4 \overline{z} + (z_{1} + 4 \overline{z} ) z_{2} - \frac{12}{3} \overline{z} (z_{1} - \frac{24}{3} z_{1} - \frac{24}{3} z_{2} - \frac{24}{3} z_{1} - \frac{24}{3} z_{1} - \frac{24}{3} z_{1} - \frac{24}{3} z_{1} - \overline{z} z_{2} - \frac{24}{3} z_{1} - \overline{z} z_{2} - \frac{24}{3} z_{1} - \overline{z} z_{2} - \overline{z} - \overline{z} z_{2} - \overline{z} z_{2} - \overline{$  $\frac{11}{12} \sigma_{0} r_{0} + \frac{11}{12} r_{0} + \frac{1}{2} \sigma_{0} r_{0} + \frac{1}{2} \sigma_{0}^{2} \frac{12}{12} \sigma_{0} r_{0} + \frac{2}{12} \sigma_{0} r_{0} + \frac{11}{12} + \frac{11}{12} r_{0} + \frac{1}{2} \sigma_{0}$  $(g_{1} - 2k_{2} + 2k_{1} - 2k_{2} - 2k_{1} + 2 - 2k_{1$ (m) + (i + 1)(m + (i + 1) + 1) + (m + 1)(m +  $\frac{1}{2}(1+1)\left[\frac{1}{2}(t_{1}-\frac{1}{2}(t_{1}+t_{1})+2t_{1})-2t_{1}+2t_{1}(t_{1}+\frac{12}{2}(t_{1})+2t_{1}(t_{1})\right]+2t_{1}(t_{1}^{2})t_{1}\left[\frac{12}{2}(t_{1}^{2})\right]$ テルコールコー 茶れー茶・子に一子れった。 アコールコールシーパー 1 - Tau - T- Ca - Ta - Tau - Tau - Tau - Tau  $\left[\frac{1}{2} |\mathbf{R}_{0,0}| + \frac{1}{2} \left[\frac{1}{2} - 2 \right] \left[\frac{1}{2} |\mathbf{R}_{0,0} - \frac{1}{2} |\mathbf{R}_{0,0} - \mathbf{R}_{0,0} + \frac{1}{2} + \frac{1}{2} |\mathbf{R}_{0,0} + \mathbf{R}_{0,0} - \mathbf{R}_{0,0} - \mathbf{R}_{0,0} \right]$ 17-18 Martin - 200 - 201 - 201 - 201 - 201 - 10, 100 - 1942 + 1944 - 49444 + 1944 - 1944 - 1944 - 1944 - 1944 - 1944 - 1944

$$\begin{split} & f_{1}^{(0)}(z) = 1 + 0^{1} f_{1}^{(0)} f_{1}^{(0)} (\frac{1}{2} \frac{1}{2} h_{1}^{(0)} + \frac{1}{2} h_{1}^{(0)} +$$

 $2R_{22} - \frac{12}{3}R_{22} - 2R_{-1/2} - \frac{11}{3}R_{2/2} + \frac{12}{3}R_{2/2} + \frac{12}{3}R_{2} - \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3}R_{2} +$ - 30 - 300 - 300 - 300 - 300 - 40<sup>2</sup> - 300 - 80 - 300 - 800 - 800 -1. May 1.  $+ \begin{bmatrix} \mathbf{R}_{1,1,2} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,2} + 2 \mathbf{R}_{1,1} + \mathbf{R}_{1,1,2} + \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \begin{bmatrix} \mathbf{R}_{1,1} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \\ - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \begin{bmatrix} \mathbf{R}_{1,1} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \\ - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \begin{bmatrix} \mathbf{R}_{1,1} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \\ - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \begin{bmatrix} \mathbf{R}_{1,1} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \\ - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \begin{bmatrix} \mathbf{R}_{1,1} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \\ - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \begin{bmatrix} \mathbf{R}_{1,1} - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \\ - \frac{\mathbf{M}}{2} \mathbf{R}_{1,1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \mathbf{L}_{1,1}^{-1} \end{bmatrix} + 2 \mathbf{L}_{1,1}^{-1} \mathbf{$  $m_{1} + \frac{m_{1}}{2} + 2m_{1} + \frac{m_{1}}{2} + \frac{m_{1}}{2}$  $+46(-_{12}+\frac{10}{12}(r_{1})+p_{10}(r)\Big(\frac{1}{2}(r_{1})+\frac{10000}{2000}+\frac{1}{2}(r_{10}+\frac{10}{4}(r_{1}))+\frac{10}{2}(r_{10})+\frac{10}{2}(r_{10})$ دِرْنِ 10 هَارُو 10 هَارُ الْحَالَ ( - أيور د ( ( 10 هـ روالله - ور 10 هـ روالله - ور 10 هـ ور 10 ه  $+\frac{100}{10}(1_{12}+1_{12})_{2}+\frac{10}{2}(1_{1}+1_{12})_{2}+11_{12})_{2}+\frac{10}{2}(1_{12}-\frac{10}{2}(1_{12}-\frac{10}{2}(1_{12}-1_{12})_{2}-1_{12})_{2})_{2}$  $\begin{array}{c} \displaystyle \frac{1}{2} \sum_{i \in U} \left[ \frac{1}{2} \sum_{i \in U} \left( \frac{1}{2} \sum_{i \in U} \left($  $-36_{1}\ell_{2}+\frac{10}{12}\eta_{1,1}+\frac{10}{7}\eta_{1,1}+\frac{20}{7}\eta_{1}-\eta_{2}\eta_{2}+\eta_{3}\eta_{3}+(1+\eta_{1}^{2}\eta_{1}^{2}-\eta_{1})\eta_{3}+\frac{20}{7}\eta_{1}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{1}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}\eta_{3}+\frac{10}{7}\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}-\eta_{3}$  $-\frac{1}{2}$ - The state - The - The - The - The - Sel - Sel - Sec -186.0.0 1 Januar 1800 - Jan 1 Januar 1800 - Januar 1800 - 1800 - 1800 Co-mo-in-in-in-in-in-in-in-in- $+ \frac{1}{2} \mathbf{R}_{1} + \frac{1}{2} \mathbf{R}_{1,1} - 2 \mathbf{R}_{1,1} + \frac{1}{2} \mathbf{R}_{-1,-1,1} + \frac{1}{2} \mathbf{R}_{2,1,2} - \mathbf{R}_{2,1,1} - \mathbf{R}_{2,1,1} - \mathbf{R}_{2,1,1} + \frac{1}{2} \mathbf{R}_{2,1,2} - \mathbf{R}_{2,1,1} - \mathbf{R}_{2,1,1} + \frac{1}{2} \mathbf{R}_{2,1,2} - \mathbf{R}_{2,1,1} - \mathbf{R$ 

-mar [lig]-agent[]-mar [max-]mar [mar ]mar ين 10 من ي 10 بار الله راي 10 باري 10 باري (10 مي ي 10 مي ي 10 مي 10 م Record Road Road Road - Road - Road - Thomas  $-221(j) + \frac{1}{2} + 2^{2} \left[ \frac{1}{2} + \frac{1}{2} + \frac{1}{2} + 2^{2} \left[ \frac{1}{2} + \frac{1}{2} + 2^{2} \left[ \frac{1}{2} +$  $-\frac{1}{2}(t_{1,1}) + 2 - 4 \left[ 4 t_{1,1} + 4 t_{1,1} - \frac{1}{2}(t_{1,1} - 2t_{2,1} - 2t_{1,1} + 10)(t_1 - 4t_2)(t_1 - \frac{1}{2}) \right]$  $(B_{1,1,2}+\frac{114}{3}B_{1,2}+\frac{101}{3}B_{1,2}+\frac{10}{3}C_{2}^{-1}+\frac{101}{3}C_{2}^{-1}+\frac{101}{3}B_{1}-\frac{5}{3}B_{1,2}+\frac{11}{3}B_{1,2}-\frac{10}{3}B_{1,2}-\frac{10}{3}B_{1,2}$  $\frac{110}{14} \theta_1 - \frac{12}{16} \theta_2 - 2\theta_{10} - \frac{100}{16} \theta_2 - \frac{100}{16} \theta_{100} - \frac{100}{16} \theta_1 + \frac{100}{16} \theta_1 + 2\theta_{100}$ 1314 New York House Block March May May May يَسْرِينَا هَا إِنَّا اللَّهُ - إِنَّا اللَّهُ اللَّهِ عَلَيْهِ اللَّهِ عَلَيْهِ عَلَيْهِ عَلَيْهِ عَلَيْهِ عَ  $(0, \zeta + 20, \omega + \frac{1}{2}(0, \omega + \frac$ march mar mar mar mar mar mar mar mar  $\frac{1}{2} + \frac{1}{2} \frac{1}{2} \frac{1}{2} - \frac{1}{2} \frac{1}{2} \frac{1}{2} + \frac{1}{2} \frac{1}{2} \frac{1}{2} \frac{1}{2} - \frac{1}{2} \frac{1}$  $+\frac{1}{2}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}{10}e^{it}+\frac{1}$  $+2\delta \zeta_{2}q_{1}^{2}\left[\frac{1}{2}g_{2}(a)\left[B_{1}-B_{2}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B_{1}(a-B$ Star Star (an See See Star Store Sector) - Stor al Mars  $-\frac{H}{2M} + \frac{1}{2} n_{1} h_{2} \left( p_{0}(z) \left[ H_{1,1} + \frac{H}{4} H_{1,1,2} - \frac{H}{2} n_{1,2} + \frac{T}{2} n_{1,2} + \frac{H}{2} n_{1,1,2} + \frac{H}{2} n_{1,1,2} + \frac{H}{2} n_{1,2} - \frac{H}{2} n_{1,2} \right)$ my Car Car Why Car War M. Carrow M. Bur  $-\frac{10}{10}R_{12} - \frac{10}{10}R_{12} - \frac{1}{10} - \frac{1}{10}R_{12} - \frac{10}{10}R_{12} -$ #Tell = Theo = The lot = The \$4 + Theo = The \$2 = Theory = Theory = Theory = Theory 24.11-24.11 + A. ( x) = 1.12 - 21.11 - 21.11 + 21.11 + 21.12 - 21.11 West Rep - R 「いいいういいないぞれ、「いい」、「いい」でないのとい

 $+ \frac{1}{(p_{0})^{(1)}} [R_{12} - R_{12} - R_{12} + R_{22} + \frac{1}{(2} R_{12} + 2R_{12} - \frac{1}{(2} R_{1} + \frac{1}{(2} R_{12} - \frac{1}{$  $+20_{11}+\left[20_{11}+\left[\frac{10}{2}h_{0}\right]+\left[\frac{10}{2}h_{0}+c\right]\left[\frac{10}{2}h_{0}-\frac{10}{2}+R_{11}\right]+\left[\frac{10}{2}+c\right]\left[20_{11}m-2h-\frac{1001}{24}\right]$  $\frac{1}{2} \left[ \left( \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \right) - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \right) - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \right) - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \right) - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \left( \frac{1}{2} - \frac{1}{2} \right) - \frac{1}{2} \left( \frac{1}{2} -$  $+\frac{12\pi i}{2\pi}+2\pi \frac{1}{2}-2\pi \frac{1}{2}\pi \frac{1}{2}\pi$ Phys. Rep. Berry Mark 1994 (1994) - Mark 1994 (1994) - Mark  $-2h_{1,2}+2h_{2,2}-\frac{1}{2}h_{1,2}-\frac{1}{2}h_{1,2}+\frac{1}{2}h_{1,2}+\frac{1}{22}h_{1}-\frac{1}{22}h_{2}+\frac{1}{22}h_{2}+h_{2}(-1)|H|<\infty$ یکی 10–روں یا 18–روں 18–روں 10–<sup>1</sup> <mark>کی</mark>ا – روا<del>ز د</del>یار 18–روں یا 19–ر -H (100) + H - A (H) (10 + H(1) - 100/2 + 100/2 + H(2 - H/2 + H(1) + 10/2 + H(1)) - Marcolline - May - Marcolling - The Inc. - The - May - May + 110, + 140, + 10, 12, - 11, - 210, + 210, 1, + 110, 1, - 210, - 210, tru- Fran - Frank)

$$\begin{split} & \frac{1}{2} \left[ (x_1 - x_2)_{1 \leq i} \left\{ \left| \frac{1}{2} (x_1 + x_2)_{1 \leq i} - \frac{1}{2} (x_1 - \frac{1}{2} (x_2 - \frac{1}{2}$$

anar 1949 - Nafer [[16 - 1633] - D - al] Masso - 1946 - [[4 - [[16 a  $|\mathbf{x}_{1} + [\mathbf{x}_{1}]_{2} + |\mathbf{x}_{1}|_{2} - \frac{12}{10}|\mathbf{x}_{2} + \frac{12}{10}|\mathbf{x}_{2} - \frac{12}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} + \frac{1000}{10}|\mathbf{x}_{1} - \frac{100}{10}|\mathbf{x}_{1} + \frac{1000}{10}|\mathbf{x}_{1} - \frac{100}{10}|\mathbf{x}_{1} - \frac{100}{10}|\mathbf{x}_{1}$  $+\frac{126}{124} + \frac{27}{7} (t_{11}) + \frac{31}{7} (t_{12}) + (26)_{12} + \frac{12}{12} (t_{12}) + \frac{27}{7} (t_{12}) - \frac{47}{12} (t_{12}) + \frac{2}{7} (t_{12}) - \frac{12}{12} (t_{12}) + \frac{12}{$ 10 mar 10 mar 100 mar 100 mar 10 mar 10 mar 10 mar 10 (Marco - May - May - 100 and - 100 and - 100 and - 100 and Thur Mar Thur The Star Thur The Star  $+121 \pm 10 - \frac{10}{10}(1_{12} - \frac{1}{2}(1_{12} - \frac{1}{2}(1_{12})_{2} + 1)_{12} - \frac{10}{10}(1_{12} + \frac{10}{10}(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10}(1_{12} - \frac{10}{10}(1_{12} - \frac{10}{10}(1_{12} - \frac{10}{10}(1_{12} - \frac{10}{10}(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10}(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})(1_{12} - \frac{10}{10})$ 111 - 187 R. - 198 R. - 208 G + R. S. - 1007 R. - 100 - 100 Lat - 21 R.  $n_{10} - \frac{11}{2}n_{10} - \frac{1}{2}n_{10} + \frac{$ 
$$\begin{split} & -\frac{1}{2} \mathcal{C}_{0,1} - \frac{1}{2} \mathcal{C}_{0,1} - \frac{1}{2} \mathcal{C}_{0,1} + \frac{1}{2} \mathcal{C}_{0,1} + \frac{1}{2} \mathcal{C}_{0,1} \mathcal{C}_{0,1} + \frac{1}{2} \mathcal{C}_{0,1} - \frac{1}{2} \mathcal{C}_{0,1} + \mathcal{C}_{0,1,1} + \mathcal{C}_{0,1,2} \\ & -\mathcal{C}_{0,1} - \mathcal{C}_{0,1,1} + \frac{1}{2} \mathcal{C}_{1,1} - \mathcal{C}_{1,1} + (1 - 1) \left[ \frac{1}{12} \mathcal{C}_{1,1} - \frac{4 \mathcal{C}_{0,1}}{2 \mathcal{C}_{0,2}} - \frac{\mathcal{C}_{0,1,1}}{2 \mathcal{C}_{0,1,1}} - \frac{\mathcal{C}_{0,1,1}}{2 \mathcal{C}_{0,1,1}} + \frac{\mathcal{C}_{0,1,1}}{2 \mathcal{C$$
 $+ \left[ \frac{1}{2} \bar{m}_{1} + \left[ \frac{1}{2} \bar{m}_{1} + \left[ \frac{1}{2} \bar{m}_{2} \right] - \left[ 1 + \alpha \right] \left[ \frac{1}{2} \frac{1}{2} \bar{m}_{1} + \frac{1}{2} \frac{1}{2} \bar{m}_{2} \right] \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{2}^{-1} \frac{1}{2} \left[ m_{1} + m_{2} \right] + 10 C_{$ SLor-Nort Northan Northan Northant Northant  $+\frac{e^{2}}{2}\eta_{1}-\frac{e^{2}}{2}\eta_{2}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{2}-\frac{e^{2}}{2}\eta_{2}-\frac{e^{2}}{2}\eta_{2}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{1}+\frac{e^{2}}{2}\eta_{2}$  $+ \frac{1}{2} + \frac{1}{2} \mathbf{a}_{i} - \frac{1}{2} \mathbf{a}_{i} \mathbf{a}_{i} - \frac{1}{2} \mathbf{a}_{i} \mathbf{a}_{i} + \frac{1}{2} \mathbf{a}_{i+1} + \frac{1}{2} \mathbf{a}_{i+1} + \frac{1}{2} \mathbf{a}_{i}^{2} + \mathbf{a}_{i+1}^{2} - \mathbf{a}_{i} \mathbf{a}_{i} \mathbf{a}_{i} + \frac{1}{2} \mathbf{a}_{i}^{2} + \mathbf{a}_{i+1}^{2} - \mathbf{a}_{i} \mathbf{a}_{$ 1184 - Mura - Murger Mouse - Murger - Musari Rosser (Musari Musa Marillo Hard Rose Rose Rose Rose Rose Rose Rose Ros + Rog + Room - Rose + 21 - 4 Ros - Ros - Ro  $[[\mathbf{R}_{i}+]]\mathbf{R}_{i,i}+[[\mathbf{R}_{i,jk}+]]\mathbf{R}_{i}+[[\mathbf{R}_{i}+]]\mathbf{R}_{i,j}-[[\mathbf{R}_{i}+]]\mathbf{R}_{i,j}+[[\mathbf{R}_{i,j}+]]\mathbf{R}_{i,j}+[[\mathbf{R}_{i,j}+]]\mathbf{R}_{i,j}]$  $+ \frac{1}{10} \frac{1}{10} + \frac{1}{10} \frac{1}{1$ 

 $+128_{0.047} - \frac{305}{1.6} + \frac{47}{1.84} + \frac{3}{1.84} + \frac{47}{1.6} + \frac{47}{1.6} + -\frac{47}{1.6} + -\frac{36}{1.6} + 108_{-2,-1,2} - 48_{-2,1,2} + 108_{-2,2}$  $\frac{10}{10}\mathbf{R}_{10} + \frac{10}{10}\mathbf{R}_{10} - \mathbf{R}_{10} - \mathbf{R}_{10} + \left[(1+\alpha)\right]^2 \mathbf{R}_{10} - \frac{10}{10}\mathbf{R}_{10} - \frac{10}{10}\mathbf{R}_{10} - \frac{100}{10}\mathbf{R}_{10} - \frac{100}{10}\mathbf{R}_{10}$  $\begin{array}{c} -m_{1,1} + \frac{m_{1}}{m_{1}} m_{1,1} + \frac{m_{1}}{m_{1}} + \frac{m_$  $-88_{-1,00}\left[-\frac{24}{248}(1-r)\right]+10^{2}\mu_{0}^{2}\left(\frac{10}{34}r_{0}-\frac{1}{24}r_{0}r_{0}-\frac{1}{24}r_{0}r_{0}(1+\frac{10}{34}r_{0}-r^{2})\left[\frac{1}{2}-7r_{0}\right]$  $+[1+\alpha]\left[\frac{1}{\alpha}\mathbf{R}_{1}-\frac{1}{\alpha}\mathbf{R}_{2}\right]+\left[(1+\alpha)\left[\mathbf{p}+\frac{1}{\alpha}\mathbf{R}_{2}-\frac{1}{\alpha}\mathbf{R}_{2}-\mathbf{R}_{1}\right]+\frac{1}{\alpha}\mathbf{R}_{2}-\mathbf{R}_{1}\right]$ 100/16/2 - 10 - 10- 10- 100 - 1110 - 20-1 - 10-10 TO  $-\frac{29}{24} + \theta_{1} - 10 + \phi_{1} - \frac{10}{2} \theta_{1} - \frac{2}{2} \theta_{1} - \theta_{1} + \frac{2}{2} \theta_{1} - \theta_{1} + \frac{2}{2} \theta_{1} - \theta_{1} + \frac{2}{2} \theta_{2} - \theta_{1} + \frac{10}{2} \theta_{2} + \frac{10}{2} \theta$  $+28_{-1,0}+\frac{7}{12}(\phi_{2}^{*}+B_{0,0}^{*})+\frac{1}{9}(\frac{1}{2}-\theta^{2})\left[B_{1}-B_{0,1}^{*}-\frac{13}{9}B_{1}+\frac{7+9}{12}+20(f_{2}+\frac{23}{12}\theta_{1})\right]$  $-(\theta_{1,2_{1},2_{2},2}+(\theta_{1,2_{1},2_{1},2_{1}}-\frac{2}{2}\theta_{1})_{2}+\frac{477}{2}\theta_{1}+(\theta_{1,2}+\frac{2}{2}\theta_{1,2_{2}}]+(1+\varepsilon)\Big[\frac{100}{10}\theta_{1}+\frac{11}{2}\theta_{1,2}\theta_{2}$  $(M_{1,1} \cdot M_{1,1} \cdot [H_{1,1} \cdot [H_{1,1} \cdot [H_{1,1} \cdot [H_{1}] \cdot [L_{1} \cdot M_{1,1} \cdot M_{1} \cdot [H_{1} \cdot H_{1,1}]])]$  $+ \left[ \varphi^2 + \Theta_{-1,0} + z \right] \frac{11}{12} \Theta_{0,0} + \left[ \Theta_{0,0} + \left[ \Theta_{0} - \Theta_{0,0,0} + \left[ \Theta_{0,0,0} + \frac{1100}{12} \Theta_{0,0} + \Theta_{0,0} \right] \right] \right]$  $- \delta(1 - \pi) \Big[ \frac{2\pi}{100} + \frac{1}{2} f_{2} + \frac{1}{10} f_{2}^{2} + \frac{1}{10} f_{2}^{2} + \frac{1}{10} f_{2}^{2} \Big] + 100 f_{2}^{2} \Big[ e^{2} (100 + \mu + 100 \mu_{2}^{2} - \frac{100}{10} \theta_{0,\mu} + 100 \mu_{1}^{2} - \frac{100}{10} \theta_{0,\mu} + \frac{100}{10} \theta_{0,\mu} \Big] + 100 f_{2}^{2} \Big] + 100 f_{2}^{2} \Big[ e^{2} (100 + \mu + 100 \mu_{1}^{2} - \frac{100}{10} \theta_{0,\mu} + \frac{100}{10} \theta_{0,\mu} + \frac{100}{10} \theta_{0,\mu} \Big] + 100 f_{2}^{2} \Big$  $-44 (t_{1,1,2}-\frac{110}{2} (t_1-\frac{11}{2} (t_{1,2}+\frac{11}{2} (t_1+\frac{110}{2} (t_1))+3 (t_2) (t_1) (\frac{110}{2} (t_1-\frac{11}{2} (t_1-\frac{11}{2} (t_1'+\frac{11}{2} (t_1))))))$ - Martin La + Martin - Ma Bug + Buss - Bug - Buss + 1Bus - Bug + Buss + Buss + Bu 1 Mar 1 May 1 May 1 Mar  $+60_{1,2}\left[-1\right]_{1}\left[-1\right]_{1}\left[-\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\left[\frac{1}{2}+\frac{11}{2}\right]\frac{1}{2}\right]\right]}\right]$  $-\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+m_{12}\left(-\frac{m_{12}}{m_{12}}\right)\left[\frac{2m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{12}}+\frac{m_{12}}{m_{1$  
$$\begin{split} & H_{0,0,0,1} = \frac{1}{2} K_{0,0,1} - \frac{1}{2} K_{0,1,0,1} + \frac{1}{2} K_{0,1,0,1} + \frac{1}{2} K_{0,1,0,1} + \frac{1}{2} K_{0,1,0} + \frac{1}{2} K_{0,1,0} + \frac{1}{2} K_{0,1,0} + \frac{1}{2} K_{0,1,0} + \frac{1}{2} K_{0,0,1} - \frac{1}{2} K_{0,0,1} -$$

 $A_{n}^{(2)}[t] = M_{n}^{(2)}[t_{n}^{(2)}[\frac{1}{2}B_{n} - \frac{1}{2}+B_{n}^{(2)} - B_{n+1} - B_{n} + B_{n} + \frac{1}{2}B_{n} - B_{n}]$  $= \sum_{k=1}^{n} ([a_{k,1} + a_{k,1} + \frac{a_{k,1}}{a_{k,1}} + \frac{a_{k$ • [1] • • (1] • [1] • (1] • [1] • (1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] • [1] \bullet [1] - Ja- Jaw - Jaw - Mar (a. - Jac) - M- Ja- Jaw m  $\frac{1}{2} R_{12} + \frac{11}{12} R_{2} + \frac{11}{12} R_{22} - R_{13} - \frac{1}{12} R_{23} + \frac{11}{12} R_{-12} + R_{2} + 2R_{1} + \frac{1}{12} R_{-12} + \frac{11}{12} R_{1}$ 12 - 24 - 24 - 24 ] + W/ / (2 ] + 12 - 12 - 24 - 24  $+ \frac{10}{10} R_{12} + R_{13} - 2R_{13} - \frac{1}{3} R_{13} + \frac{10}{3} R_{13}^2 + R_{1,12} + R_{13}^2 + \frac{1}{3} R_{1,13} - 2R_{13}^2 + \frac{1}{3} R_{13} + \frac{$ But has be much the music has be +10 c da = 40 c cu = 10 c ca + 00 c c u + 10 cu | + 0 - 40 da = 10a  $+\frac{11}{10}R_{12} - \frac{3\pi i}{42}R_{12} - \frac{4\pi i}{10}R_{12} - \frac{11}{10}R_{12} - R_{12} - 2R_{12} - \frac{4\pi^2}{42}R_{1} - 2R_{23} - \frac{11}{10}R_{13}$  $\frac{411}{100} - \frac{101}{2} \zeta_{2} - \frac{100}{10} \zeta_{2} + \frac{1}{2} \zeta_{2} \zeta_{2} - \frac{81}{2} \delta_{2} + \frac{20}{2} \delta_{2} - \frac{101}{2} \zeta_{2} - 111 \zeta_{2} + 21 \zeta_{2}$  $-24E_{-2,-1,0} + \frac{1}{10}E_{0,0} + \frac{1}{10}E_{0,0,0} - \frac{1}{10}E_{-1,0} - \frac{1}{10}E_{0,0} - \frac{10}{10}E_{0,0} - \frac{10}{10}E_{0,0}$ 

The first final and while in the first first  $+ \mathbb{E}_{-1, -1, 0} - \frac{101}{100} \mathbb{E}_{-1, 0} + \frac{10}{10} \mathbb{E}_{-1, 0} + \mathbb{E}_{-1, 0} \Big] + (1 - \alpha) \Big[ \frac{10}{10} \mathbb{E}_{0} + \frac{10}{10} \mathbb{E}_{0, 0} - \frac{10}{10} \mathbb{E}_{0, 0} - \mathbb{E}_{0, 0} \Big]$ - How 
$$\begin{split} &+(1+1)\frac{12}{2}h_{1}f_{2}-\frac{16}{2}h_{1}+\frac{12}{2}h_{2}+\frac{16}{20}h_{1}+\frac{129}{20}h_{1}-\frac{129}{20}f_{2}+\frac{1}{2}(1+1)f_{2}-\frac{16}{2}f_{2}^{2}+\frac{129}{2}h_{1}f_{2}\\ &+(2)h_{1}f_{2}-(2)h_{2}^{2}-(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^{2}+(2)h_{2}^$$
 $\frac{10}{2} R_{1} - \frac{10}{2} R_{1} \left[ k_{1} \right] - 2 R_{10} - \frac{4}{2} R_{100} + 2 R_{10} + \frac{6}{2} R_{10} - 2 R_{1} - 2 R_{10} + \frac{10}{2} k_{10} + \frac{10}{2} k_{$ 100 No 100 100<sup>2</sup> 100 - Multi-1Mar 1Mar 1Maas 181 - 40  $-i_{2}i_{2}+\frac{1}{2}i_{2}+\frac{1}{2}i_{2}+\frac{1}{2}i_{2}^{2}+\frac{1}{2}i_{2}^{2}-2i_{2}^{2}\Big]+22i_{2}\mu_{2}^{2}\Big[\frac{1}{2}i_{1}^{2}\Big[\frac{1}{2}i_{2}+2i_{1}-i_{2}+2i_{2}\mu_{2}-i_{1}^{2}\Big]+\frac{1}{2}i_{2}$  $-\frac{1}{2}(1-\frac{1}{2}(h_{0}-\frac{1}{2}(h_{0}+1)+\frac{1}{2}(\frac{1}{2}-h^{2})\left[\frac{1}{2}(h_{0}-2h_{0}-h_{0}-\frac{1}{2})(1-h^{2})\left[\frac{1}{2}(h_{0}+\frac{1}{2}(h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}$ 得一览之后;同口云苍风一致一致一变之之之法之之之。  $\left[ H_{1,1} - 2 H_{2,1} \right] + \frac{11}{14\pi} H(1-\alpha) + 10 H_{2,1}^{-1} h_{2} \left[ \frac{\alpha}{2} d^2 \left[ \frac{\alpha + 1}{14\pi} + \frac{11}{2} H_{2,1} + \frac{1}{2} H_{2,2} - H_{2,2} - \frac{\alpha}{2} + \frac{1}{2} H_{2,2} \right] \right]$  $- \pi_{12} + \frac{1}{2} \pi_{12} + \frac{1}{2} \pi_{14} + \pi_{14} - \pi_{14} - \frac{1}{2} \pi_{14} + \frac{1}{2} \frac{1}{4} - \pi_{14} - \frac{1}{2} \pi_{14} - 2 (f_{1} - 2 (g_{1} + 2 (g_{2} + g_{2} + g_{2} + g_{2} + g_{2} + g_{2}))))))))))))))))))))))$  $+\frac{10}{10}\mathbf{K}_{0}-\frac{100}{10}+\mathbf{IH}_{0,0}+\frac{100}{10}\mathbf{x}_{0,0}^{2}-\mathbf{IH}_{0,0}+\mathbf{K}_{0,0}-\frac{10}{10}\mathbf{K}_{0}+\frac{10}{10}\mathbf{K}_{0}+\mathbf{IH}_{0,0}+\mathbf{X}_{0,0}^{2}-\mathbf{K}_{0,0}^{2}$  $-\frac{24}{3} f_{0,0} - 4 f_{0,0} + \frac{24}{3} f_{0,0} - 4 f_{0,0} - 4 f_{0,0} - 10 f_0 + 2 f_0^2 + 10 g_0^2 - 10 f_0 - 4 f_0 g_0^2$ (BLOOD) - BLOOD BLOOD - BLOOD  $(2n_{11} + n_{12} + 2n_{12} + \frac{1}{2}4(1 + a))$ .

# Moch, Vermaseren, Vogt 2004

#### NNLO the new emerging standard in QCD – essential for precision physics

# **P**<sub>ii</sub> **@ NNLO:** a landmark calculation

#### 10000 diagrams, 10<sup>5</sup> integrals, 10 man years, and several CPU years later:

 $\delta_{\mu}^{(0)}(z) = (m_{\mu}^{\prime})_{\mu\nu}^{\prime} \left[ \frac{1}{2} (z + z^{2}) \left[ \frac{1}{2} (z + z^{2}) + \frac{1}{2} (z + z^{2}) (z - 2) + (z - 2) \right] \right]$  $-H_{1,1}\left[+\frac{1}{2}\frac{1}{2}-r^{2}\left(\frac{14}{2}r_{1}+H_{1}+H_{2}+\frac{1}{2}H_{1}-\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+\frac{2H_{1}}{2H}+H_{1}r_{2}-\frac{1}{2}H_{1}\right]$  $\mathbf{M}_{1,1,2} + 2\mathbf{M}_{1,1,2} + 2\mathbf{M}_{2,1,2} + (1-\alpha) \left[ \frac{2M}{2} \mathbf{H}_{2} + \frac{1+\alpha}{2} + \frac{2M}{2} \mathbf{H}_{2,2} - \frac{1}{2} \mathbf{H}_{-1,2,2} + 2\mathbf{H}_{2,2,2,2} \right]$ - No - Morthar Ray - Roy - New - New - New - New - $\frac{1}{2} [t_{12} + [t_{123} + [t_{123}]] + (1 + \pi) \Big[ \frac{1}{12} [t_{12} + \frac{31}{2} t_{12} + \frac{11}{12} t_{12} + \frac{11}{12} t_{12} + \frac{11}{12} t_{12} - \frac{11}{12} t_{12} + \frac{11}{12} t_{12}$ Rus-Burthip Burthout Burtho R. Hurtho  $+ \frac{1}{2} (z_{12} z_{23} - \frac{1}{2} z_{2}^{2} + 4 \overline{z} + (z_{1} + 4 \overline{z} ) z_{2} - \frac{12}{3} \overline{z} (z_{1} - \frac{24}{3} z_{1} - \frac{24}{3} z_{2} - \frac{24}{3} z_{1} - \frac{24}{3} z_{1} - \frac{24}{3} z_{1} - \frac{24}{3} z_{1} - \overline{z} z_{1} - \frac{24}{3} z_{1} - \overline{z} z_{1} -$  $\frac{11}{2} \theta_{1} f_{2} + \frac{11}{2} f_{2} + \frac{1}{2} \theta_{1,0} + 10 \theta_{0,0,0} + \frac{3}{2} \phi \Big[ \frac{10}{2} \theta_{1,0} + \frac{20}{2} \theta_{1,0} + 10 f_{2} + \frac{11}{2} + \frac{17}{2} \theta_{1} + \frac{1}{2} \theta_{1} + \frac{1}{2} \theta_{1} \Big]$  $\left\{ p_{1} - 2k_{1} \left[ p_{1} - 2k_{1} - 2k_{1}$  $\frac{1}{2}(1+t)\left[\frac{1}{2}(t_{1}-\frac{1}{2}(t_{1}+t_{2}+2t_{2})-2t_{1}+2t_{2}(t_{1}+\frac{1}{2}(t_{1}+2t_{2}))\right]+2t_{1}(t_{1}^{2}t_{1})\left[\frac{1}{2}(t_{1}^{2}-t_{1}^{2})+2t_{1}(t_{1}^{2}-t_{1}^{2})+2t_{1}(t_{1}^{2}-t_{1}^{2})\right]$ テルコールコー 茶れー茶・子に一子れった。 アコールコールシーパー 1 - Tau - T- Ca - Ta - Tau - Tau - Tau - Tau  $+\frac{10}{1}\mathbf{R}_{1,0}\left[+\frac{1}{2}(1-z)\left[\frac{10}{12}\mathbf{R}_{1,0}-\frac{10}{12}\mathbf{R}_{1,1}-\mathbf{X}_{2,1}+\frac{10}{12}+\frac{10}{12}\mathbf{R}_{1,10}+\mathbf{R}_{1,1}-\mathbf{R}_{1,10}-\mathbf{R}_{1,10}\right]$ 17-18 Martin - 200 - 201 - 201 - 201 - 201 - 10, 100 - 1944 + 1944 - 49444 + 1944 - 1944 - 1944 - 1944 - 494

NEW - WARDON WILL WALL + Mars - EAL + Mars -144 - 2411 - 144 - 1944 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194 - 194  $\frac{30^4}{10}R_{1,2}-\frac{10}{7}R_{1,2}-\frac{101}{10}R_{1}+\frac{4^2}{7}R_{2,2}+\frac{8}{7}R_{2,2}+\frac{38}{7}R_{2,2,2}+\frac{105}{10}R_{2}-\frac{1099}{10}+\frac{1099}{10}+\frac{1099}{10}R_{2}$ 1994 + 19, 10 + 19, 14 + 19, 14 + 19, 15 + 19, 111 - 19, 11 - 19, 11 - 19, 11

 $2R_{22} - \frac{12}{3}R_{22} - 2R_{-1/2} - \frac{11}{3}R_{2/2} + \frac{12}{3}R_{2/2} + \frac{12}{3}R_{2} - \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3} + \frac{12R_{2}}{3}R_{2} +$ - 30 - 300 - 300 - 300 - 300 - 40<sup>2</sup> - 300 - 80 - 300 - 800 - 800 -وكروال ويريها ( معروي الله من و الله موكر الله موكر الله والرال الم والرال الم ويوافق مواقعه  $+ \left[ \mathbf{E}_{i,i,j} - \frac{1}{2} \mathbf{E}_{i,j} + 2 \mathbf{E}_{i} + 2 \mathbf{E}_{i,i,j} + \mathbf{E}_{i,j} \right] + 2 \mathbf{E}_{i,j} \left[ \mathbf{E}_{i,j} \left[ \mathbf{E}_{i,j} - \frac{2 \mathbf{E}_{i,j}}{2} - \frac{2 \mathbf{E}_{i,j}}{2} - \frac{2 \mathbf{E}_{i,j}}{2} \right] \right]$  $+46(z_1+\frac{200}{2}(z_1)+y_{20}(z_1)\frac{1}{2}(z_1)z_2+\frac{100000}{2000}+\frac{1}{2}(z_{10}+\frac{10}{2}(z_1)+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_{10})+\frac{10}{2}(z_$ دِرْنِ 10 هَارُو 10 هَارُ الْحَالَ (- "يَوَرَد (يَا 10 هَارَ) - يَوَالَّهُ وَرَيْ 10 هُوَ الْعَالَ ا  $+\frac{100}{10}(1_{12}+1_{12})_{2}+\frac{10}{2}(1_{1}+1_{12})_{2}+11_{12})_{2}+\frac{10}{2}(1_{12}-\frac{10}{2}(1_{12}-\frac{10}{2}(1_{12}-1_{12})_{2}-1_{12})_{2})_{2}$ يون 16 ميرو 16 مين ي 10 مركي 181 ميں ي 101 ميرو ي 101 ميں ي ي 181  $-\frac{36}{2} \left[ \frac{17}{10} \left[ \frac{10}{10} \left[ \frac{10}{10} + \frac{10}{10} \right] \left[ \frac{20}{10} \left[ \frac{10}{10} - \frac{10}{10} \left[ \frac{10}{10} + \frac{10}{10} \left[ \frac{10}{10} + \frac{10}{10} \right] + \frac{10}{10} \left[ \frac{10}{10} \left[ \frac{10}{10} - \frac{10}{10} \right] + \frac{10}{10} \left[ \frac{10}{10} + \frac{10}{10} \right] + \frac{10}{10} \left[ \frac{10}{10} \left[ \frac{10}{10} + \frac{10}{10} \right] + \frac{10}{10} \left[ \frac{10}{10$  $-\frac{1}{2}$   $(x_1 - x_2)$   $-\frac{1}{2}$   $(x_2 - \frac{1}{2})$   $(x_3 - \frac{1}{2})$   $(x_4 - \frac{1}{2})$   $(x_5 - \frac{1}{2})$   $(x_5 - \frac{1}{2})$   $(x_5 - \frac{1}{2})$   $(x_5 - \frac{1}{2})$ - The state - The - The - The - The - Sel - Sel - Sec -186.0.0 1 Januar 1800 - Jan 1 Januar 1800 - Januar 1800 - 1800 - 1800 Co-mo-in-in-in-in-in-in-in-in- $\left[\frac{1}{2}\mathbf{R}_{1}+\frac{1}{2}\mathbf{R}_{2}-2\mathbf{R}_{2}+\frac{1}{2}\mathbf{R}_{2}]_{2}+1\right]\mathbf{R}_{2}\mathbf{u}+\mathbf{R}_{2}\mathbf{u}-\mathbf{R}_{2}\left[\left[+2\mathbf{L}_{2}\mathbf{v}^{2}\right]_{2}^{2}-\frac{1}{2}\right]_{2}$ 

-ma+[ig]-a\_1+([img-]max-]ma+[ma+]ma+]ma - May - May - Milling and May - Million - May - May - Million - May - Million latest progress: A. 2014 Nogt @ Loops & Legs 2014 Nu a in order Record Road Road Road - Road - Road - Thomas est progress' APije finally done (NLO in 95) also helicity 8-1422 + 7 - 8 (19144 + 1914 - 1987 / + 1947 + 1947 - 1974 + 1944 - 1974 HANTIMA MAY MANY MAY MANY MANY + 110, + 140, + 10, 12, - 11, - 210, + 210, 1, + 110, 1, - 210, - 210, - 210, tru- Fran - Frank)

> $f_{\mathbf{k}}^{(1)}(z) = 10^{1} f_{\mathbf{k}}^{(1)} g_{\mathbf{k}}^{(1)} \left\{ f_{\mathbf{k}}^{(1)} + 10_{1} g_{\mathbf{k}} - \frac{10}{2} \eta_{\mathbf{k}}^{(1)} + \frac{10}{2} \eta_{\mathbf{k}}^{(1)} - \frac{10}{2} \eta_{\mathbf{k}}^{(1)} (H_{-1,2}+H_{0,2}-H_{-1,2}-\frac{12}{3}H_{0,2}-\frac{42}{3}H_{0}+\frac{141}{3}H_{0}+\frac{513}{12}-\frac{1}{3}H_{0}-2H_{0,2}-4H_{0,0,2}$

 $+\frac{126}{124} + \frac{27}{7} (t_{11}) + \frac{31}{7} (t_{12}) + (26)_{12} + \frac{12}{12} (t_{12}) + \frac{27}{7} (t_{12}) - \frac{47}{12} (t_{12}) + \frac{2}{7} (t_{12}) - \frac{12}{12} (t_{12}) + \frac{12}{$ 

 $|\mathbf{x}_{1} + [\mathbf{x}_{1}]_{2} + |\mathbf{x}_{1}|_{2} - \frac{12}{10}|\mathbf{x}_{2} + \frac{12}{10}|\mathbf{x}_{2} - \frac{12}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} + \frac{10000}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} + \frac{10000}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} - \frac{10000}{10}|\mathbf{x}_{1} - \frac{12}{10}|\mathbf{x}_{1} - \frac{10000}{10}|\mathbf{x}_{1} - \frac{1000}{10}|\mathbf{x}_{1} - \frac{10000}{10}|\mathbf{x}_{1} - \frac{10000}{10}|\mathbf{x}_{1} - \frac{1000}{10}|\mathbf{x}_{1} - \frac{1000}{$ 

 $+28_{-1,0}+\frac{7}{12}(\phi_{2}^{*}+B_{0,0}^{*})+\frac{1}{9}(\frac{1}{2}-\theta^{2})\left[B_{1}-B_{0,1}^{*}-\frac{13}{9}B_{1}+\frac{7+9}{12}+20(f_{2}+\frac{23}{12}\theta_{1})\right]$ [manma] + [ + 2] [ [m - [ 2 + [m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 + m + 2 +  $-(\theta_{1,2_{1},2_{2},2}+(\theta_{1,2_{1},2_{1},2_{1}}-\frac{2}{2}\theta_{1})_{2}+\frac{477}{2}\theta_{1}+(\theta_{1,2}+\frac{2}{2}\theta_{1,2_{2}}]+(1+\varepsilon)\Big[\frac{100}{10}\theta_{1}+\frac{11}{2}\theta_{1,2}\theta_{2}$  $(M_{1,1} \cdot M_{1,1} \cdot [H_{1,1} \cdot [H_{1,1} \cdot [H_{1,1} \cdot [H_{1}] \cdot [L_{1} \cdot M_{1,1} \cdot M_{1} \cdot [H_{1} \cdot H_{1,1}]])]$  $+ \left[ \varphi^2 + \Theta_{-1,0} + z \right] \frac{11}{12} \Theta_{0,0} + \left[ \Theta_{0,0} + \left[ \Theta_{0} - \Theta_{0,0,0} + \left[ \Theta_{0,0,0} + \frac{1100}{12} \Theta_{0,0} + \Theta_{0,0} \right] \right] \right]$  $- \delta_{1}^{-1} - \pi \left[ \frac{2 \pi i}{2 \pi i} + \frac{1}{2} \zeta_{2}^{-1} + \frac{1}{2} \zeta_{2}^{-1} + \frac{1}{2} \zeta_{2}^{-1} \right] + 10 \varepsilon_{2}^{-1} \left[ e^{2} \left[ 10 \pi i \zeta_{2} + 10 \pi i \zeta_{2} - \frac{10 \pi i}{2} \sigma_{0,2} \right] \right]$  $-44 (t_{1,1,2}-\frac{110}{2} (t_1-\frac{11}{2} (t_{1,2}+\frac{11}{2} (t_1+\frac{110}{2} (t_1))+3 (t_2) (t_1) (\frac{110}{2} (t_1-\frac{11}{2} (t_1-\frac{11}{2} (t_1'+\frac{11}{2} (t_1))))))$ - M. 11 + M. 14 + M. 14 + M. 14 - M. 14 + M. 14 Bug + Buss - Bug - Buss + 1Bus - Bug + Buss + Buss + Bu 1 Mar 1 May 1 May 1 Mar 1  $+ 60_{1,1,0} + \beta_{1,0} (-\epsilon) \left[ \frac{11}{2} \beta_{1}^{-1} - \frac{11}{2} \beta_{1} \beta_{2}^{-1} + 60_{1,0,0} + 100_{1,0,0} - 100_{1,0,0} - \frac{110_{1,0,0}}{2} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0} + 100_{1,0,0}$ - This - Theorem + (1-+) [2007 - Theor - This - Theorem - Theorem

Rose [Kar [Ka] - Micco Hass - Micco Hass | Kar  $+ 4 R_{-1,0,0} + \frac{111}{4} R_1 + \frac{11}{4} Q_2 + \frac{11}{4} R_1 + \frac{11}{2} R_2 + \frac{11}{4} R_1 + \frac{1}{2} R_2 + \frac{1}{2} R_2 Q_2 + \frac{1}{2} R_{1,0} + \frac{1}{2} R_{$  $+\frac{1}{2}\xi_{1}^{2}+\frac{1}{2}R_{1}+\frac{1}{2}R_{1}\xi_{2}+\frac{12}{2}R_{1}g_{2}+\frac{12}{2}R_{2}g_{3}+\frac{12}{2}R_{3}+R_{1}g_{3}+R_{1}g_{3}+R_{1}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_{3}+R_{2}g_$ 30-36-4- But 2.4-8-4-

 $\boldsymbol{S}_{\boldsymbol{k}}^{(2)}[\boldsymbol{k}] = \log_{\boldsymbol{k}} \boldsymbol{g}_{\boldsymbol{k}} \cdot \boldsymbol{g}_{\boldsymbol{k}} \left[ \left[ \boldsymbol{k} \right] \left[ \frac{1}{2} \boldsymbol{R}_{\boldsymbol{k}} - \frac{1}{2} + \boldsymbol{R}_{\boldsymbol{k}} \cdot \boldsymbol{R}_{\boldsymbol{k} + \boldsymbol{k}} - \boldsymbol{R}_{\boldsymbol{k} + \boldsymbol{k}} + \frac{1}{2} \boldsymbol{R}_{\boldsymbol{k}} \cdot \boldsymbol{R}_{\boldsymbol{k} + \boldsymbol{k}} \right]$ ية (1) مدارية (10 من (1 من 1) × 1) م (1 من 11 من 10 م - Ja- Jan - Jan - Man - Jan -  $\frac{1}{2} R_{12} + \frac{11}{12} R_{2} + \frac{11}{12} R_{22} - R_{13} - \frac{1}{12} R_{23} + \frac{11}{12} R_{-12} + R_{2} + 2R_{1} + \frac{1}{12} R_{-12} + \frac{11}{12} R_{1}$ The fair fair fair the fair that the fair fair fair the  $+ \frac{10}{2} R_{12} + R_{13} - 2R_{13} - \frac{1}{2} R_{13} + \frac{1}{2} R_{23} + R_{14} + \frac{1}{2} R_{14} + \frac{1}{2} R_{14} - 2R_{15} + \frac{1}{2} R_{14} + \frac{1}{2} R_{14} - 2R_{15} + \frac{1}{2} R_{14} + \frac{1}$ Rob Karly Mary Rob Mary Ray +10 c da = 40 c cu = 10 c ca + 00 c c u + 10 cu | + 0 - 40 da = 10a  $+\frac{11}{10}R_{12} - \frac{3\pi i}{42}R_{12} - \frac{4\pi i}{10}R_{12} - \frac{11}{10}R_{12} - R_{13} - 2R_{13} - \frac{4\pi^{2}}{42}R_{1} - 2R_{23} - \frac{11}{10}R_{13}$  $\frac{411}{100} - \frac{100}{4} \zeta_2 - \frac{100}{100} \beta_{12} + \frac{1}{2} \beta_{12} \zeta_2 - \frac{10}{2} \beta_{12} + \frac{20}{2} \beta_{12} - \frac{100}{4} \zeta_2 - 100 \zeta_{12} + 20 \zeta_2$  $-2H_{-1,-1,0} + \frac{1}{10}H_{0,0} + \frac{1}{10}H_{0,0,0} - \frac{1}{10}H_{0,0} - \frac{1}{10}H_{0,0} - \frac{10}{10}H_{0,0} - \frac{10}{10}H_{0,0} - \frac{10}{10}H_{0,0}$ 

The first final and while in the first first  $+ \mathbb{E}_{-1, -1, 0} - \frac{101}{100} \mathbb{E}_{-1, 0} + \frac{10}{10} \mathbb{E}_{-1, 0} + \mathbb{E}_{-1, 0} \Big] + (1 - \alpha) \Big[ \frac{10}{10} \mathbb{E}_{0} + \frac{10}{10} \mathbb{E}_{0, 0} - \frac{10}{10} \mathbb{E}_{0, 0} - \mathbb{E}_{0, 0} \Big]$  $+ (1 + 1) \left[ \frac{17}{7} h_0 f_0 - \frac{10}{7} h_0 + \frac{10}{7} h_{0,1} + \frac{100}{77} h_0 - \frac{100}{77} h_0 + \frac{10}{7} h_0 + \frac{10}{7}$ 2364Q - 186Q - 186, july + 188, july + 286, july + 188, july - 48, Q + 188, july + 286,  $\frac{10}{2} \mathbf{R}_{1} - \frac{10}{2} \mathbf{R}_{1} \left[ \mathbf{k}_{1} \right] - 2 \mathbf{R}_{10} - \frac{\mathbf{n}}{2} \mathbf{R}_{100} + 2 \mathbf{R}_{10} + \frac{\mathbf{n}}{2} \mathbf{R}_{10} \left[ \mathbf{k}_{1} - 2 \mathbf{R}_{1} - 2 \mathbf{R}_{10} + \frac{\mathbf{n}}{2} \mathbf{k}_{10} \right]$  $-i_{2}i_{2}+\frac{1}{2}i_{2}+\frac{1}{2}i_{2}+\frac{1}{2}i_{2}^{2}+\frac{1}{2}i_{2}^{2}-2i_{2}^{2}\Big]+22i_{2}\mu_{2}^{2}\Big[\frac{1}{2}i_{1}^{2}\Big[\frac{1}{2}i_{2}+2i_{1}-i_{2}+2i_{2}\mu_{2}-i_{1}^{2}\Big]+\frac{1}{2}i_{2}$  $-\frac{1}{2}(1-\frac{1}{2}(h_{0}-\frac{1}{2}(h_{0}+1)+\frac{1}{2}(\frac{1}{2}-h^{2})\left[\frac{1}{2}(h_{0}-2h_{0}-h_{0}-\frac{1}{2})(1-h^{2})\left[\frac{1}{2}(h_{0}+\frac{1}{2}(h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}-h_{0})+\frac{1}{2}(h_{0}$ 得一览之后;同口云苍风一致一致一变之之之法之之之。  $\left[ H_{1,1} - 2 H_{2,1} \right] + \frac{11}{14\pi} H(1-\alpha) + 10 H_{2,1}^{-1} h_{2} \left[ \frac{\alpha}{2} d^2 \left[ \frac{\alpha + 1}{14\pi} + \frac{11}{2} H_{2,1} + \frac{1}{2} H_{2,2} - H_{2,2} - \frac{\alpha}{2} + \frac{1}{2} H_{2,2} \right] \right]$  $- \pi_{12} + \frac{1}{2} \pi_{122} + \frac{1}{2} \pi_{12} + \pi_{12} - 2\pi_{222} - \frac{1}{2} \pi_{12}^{2} + \frac{1}{2} \frac{1}{2} - \pi_{12}^{2} + \frac{1}{2} \pi_{12}^$  $- 2 (f_{12} - 2 f_{12} + 2 f_{1$  $+\frac{10}{10}\mathbf{K}_{0}-\frac{100}{10}+\mathbf{IH}_{0,0}+\frac{100}{10}\mathbf{x}_{0,0}^{2}-\mathbf{IH}_{0,0}+\mathbf{K}_{0,0}-\frac{10}{10}\mathbf{K}_{0}+\frac{10}{10}\mathbf{K}_{0}+\mathbf{IH}_{0,0}+\mathbf{X}_{0,0}^{2}-\mathbf{K}_{0,0}^{2}$  $-\frac{24}{3} (h_{12}-40)_{12} + \frac{24}{3} (h_{1}-40)_{12} + \frac{1}{3} (h_{12}-1) (h_{1}+1) (h_{1}+1) (h_{1}-1) (h_{1}-40)_{12} (h_{1}-1) (h_{1}-1)$  $(20_{111} + 20_{11} + 20_{11}^2 + \frac{1}{10}(1 + \alpha))$ .

#### Moch, Vermaseren, Vogt 2004

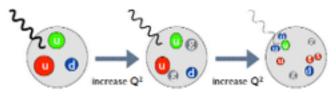
#### NNLO the new emerging standard in QCD – essential for precision physics

# **DGLAP** evolution in full glory

taking quarks and gluons together: coupled integro-differential equations

 $\frac{d}{d\ln\mu} \begin{pmatrix} q(x,\mu) \\ g(x,\mu) \end{pmatrix} = \int_{x}^{1} \frac{dz}{z} \begin{pmatrix} \mathcal{P}_{qq} & \mathcal{P}_{qg} \\ \mathcal{P}_{gq} & \mathcal{P}_{gg} \\ \mathcal{P}_{gq} & \mathcal{P}_{gg} \end{pmatrix}_{(z,\alpha_{s})} \cdot \begin{pmatrix} q(x/z,\mu) \\ g(x/z),\mu \end{pmatrix}$ 

best solved in Mellin moment space: set of ordinary differential eqs.; no closed solution in exp. form beyond LO (commutators of P matrices!)



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taking quarks and gluons together: coupled integro-differential equations

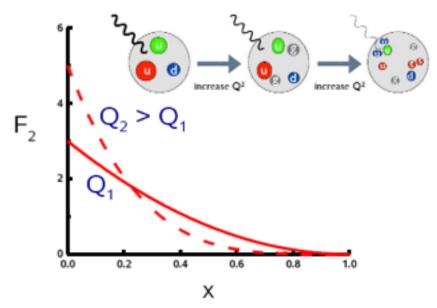
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#### main effect/prediction of evolution:

partons loose energy by evolution!

- large x depletion
- small x increase



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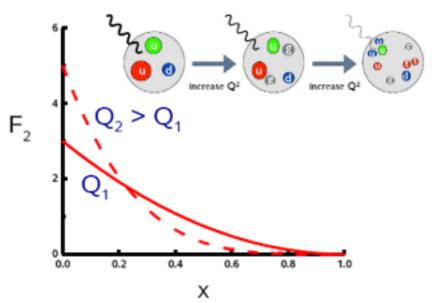
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#### main effect/prediction of evolution:

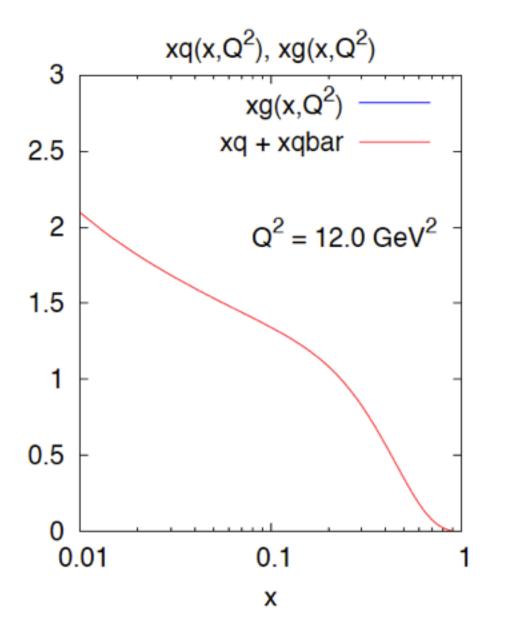
partons loose energy by evolution!

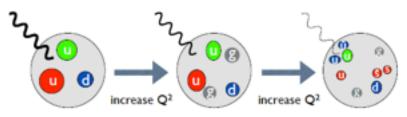
- large x depletion
- small x increase

exactly as observed in experiment huge success of pQCD



### **DGLAP** evolution at work: toy example

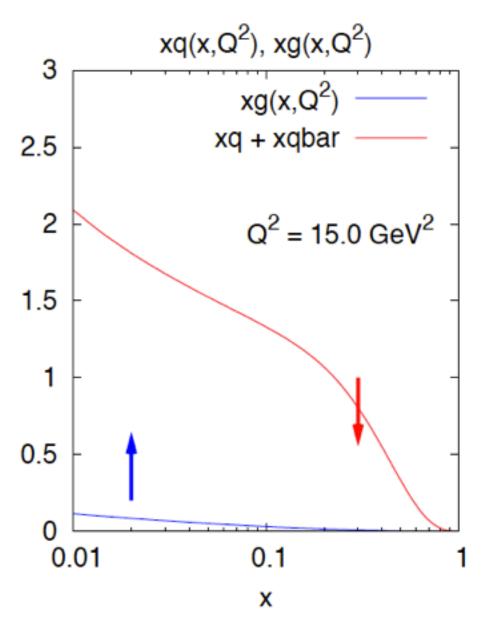


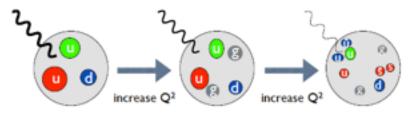


start off from just quarks, no gluons

- quarks reduced at large x
- gluons rise quickly at small x (which, btw, also generates sea quarks)

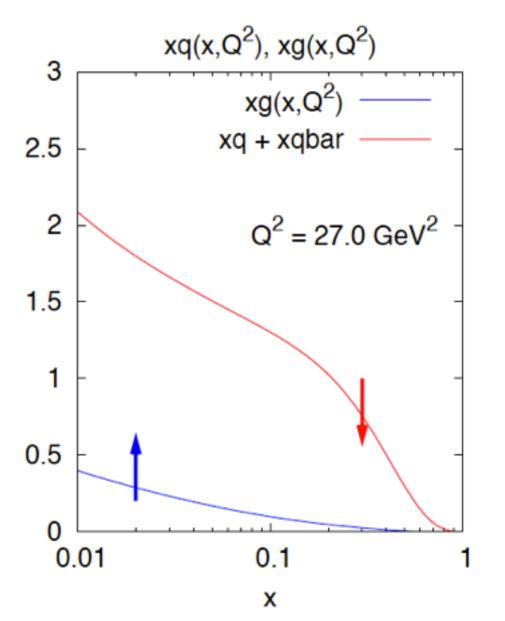
taken from G. Salam





start off from just quarks, no gluons

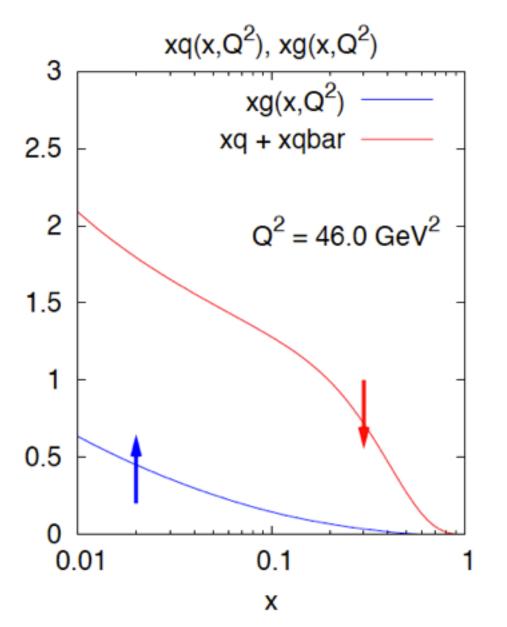
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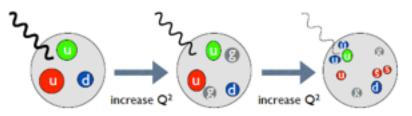


 $\begin{array}{c} & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & & \\ & & & & \\ & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ &$ 

start off from just quarks, no gluons

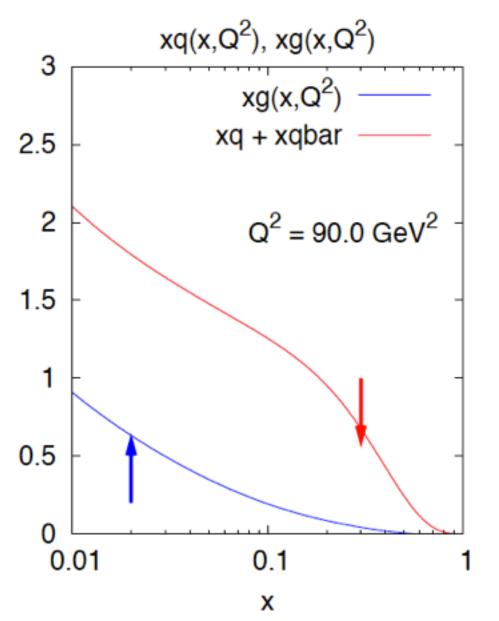
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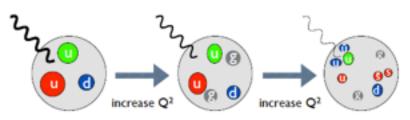




start off from just quarks, no gluons

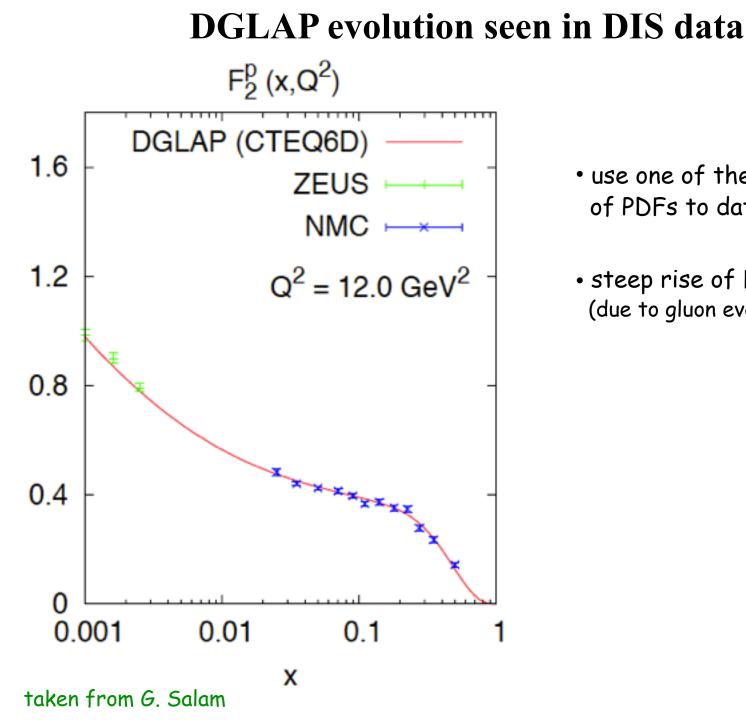
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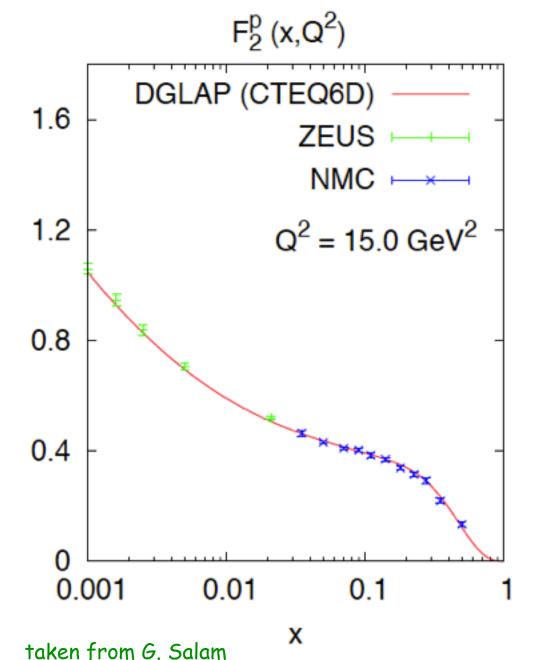
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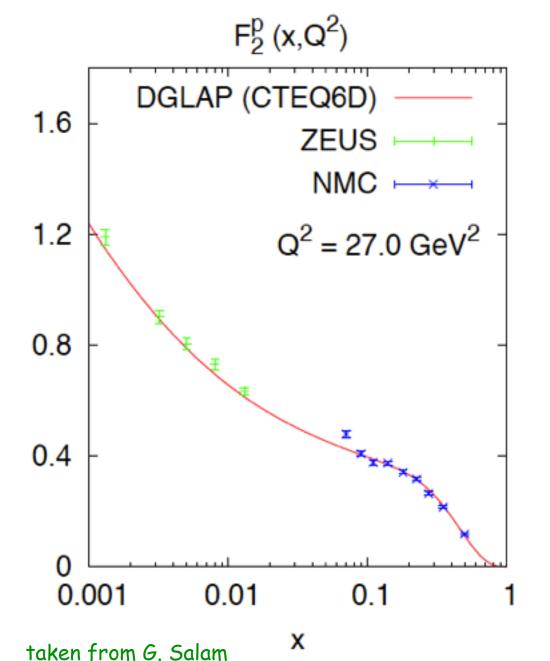


 use one of the global fits of PDFs to data by CTEQ

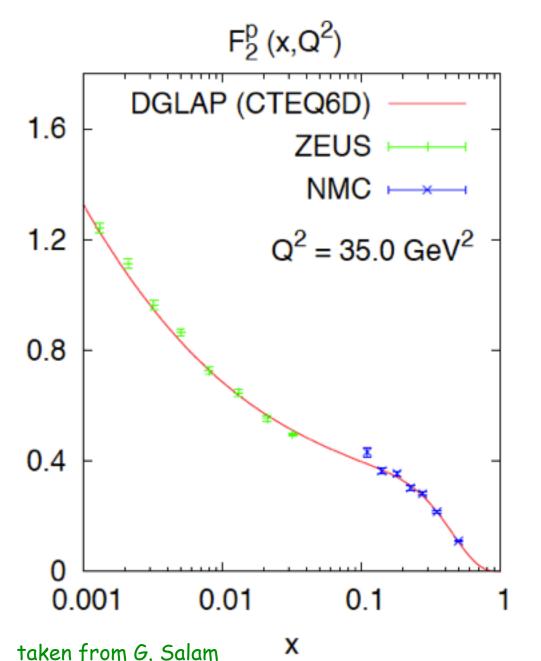
• steep rise of  $F_2$  at small x (due to gluon evolution)



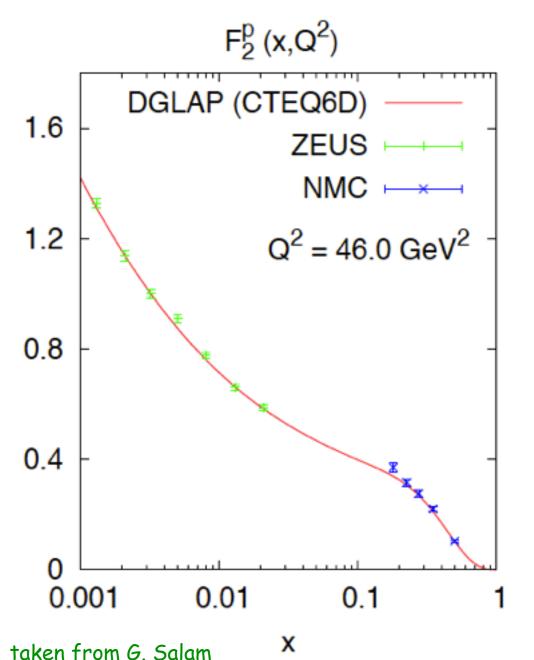
 use one of the global fits of PDFs to data by CTEQ



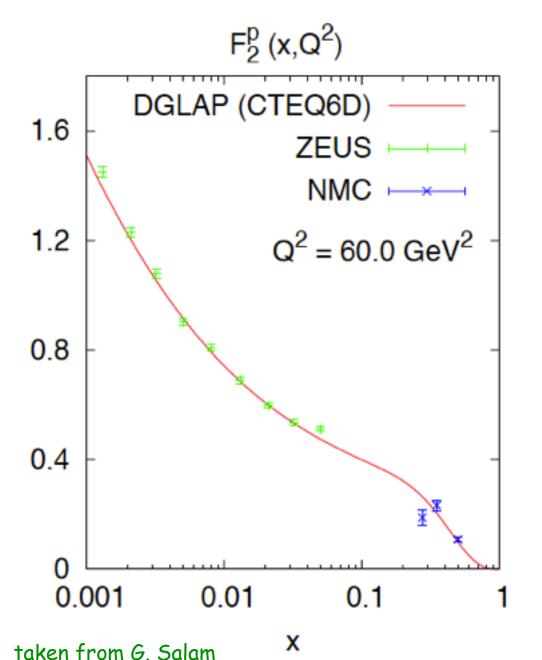
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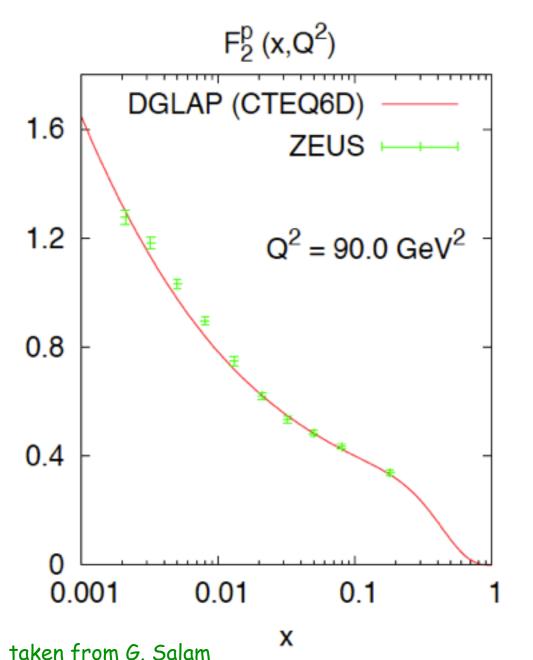
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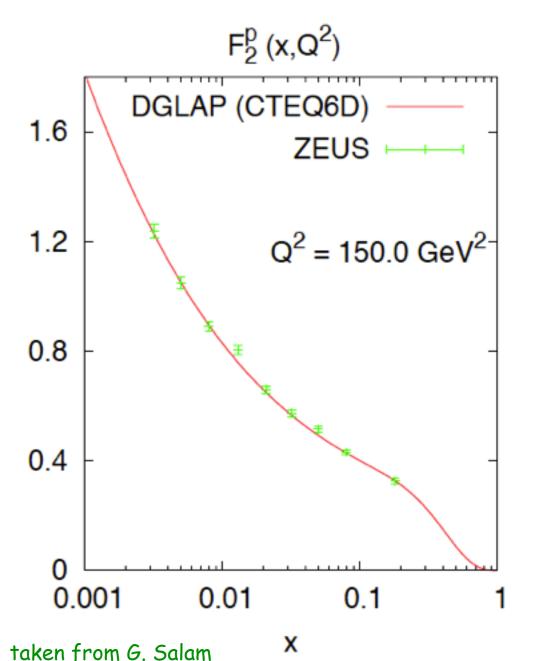
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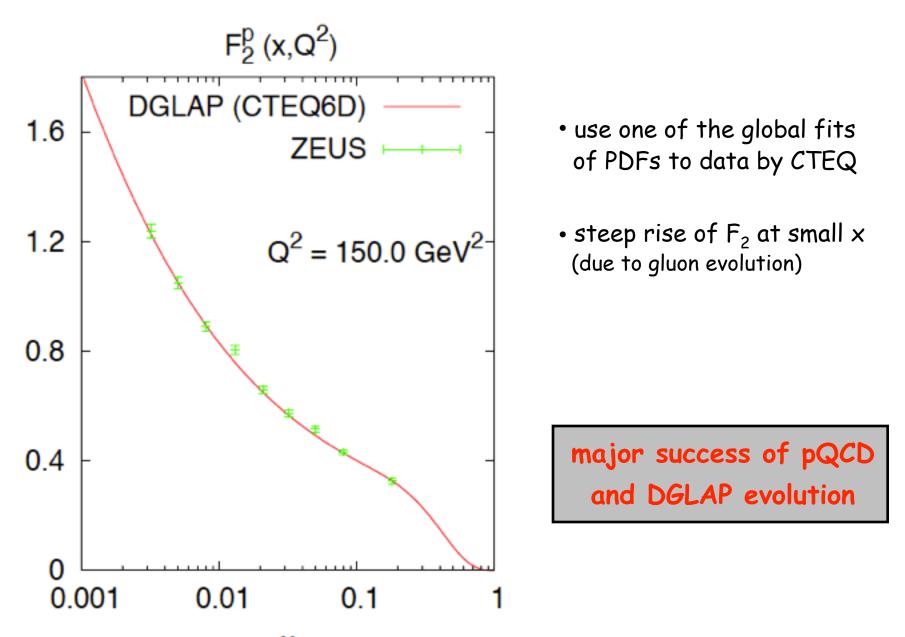
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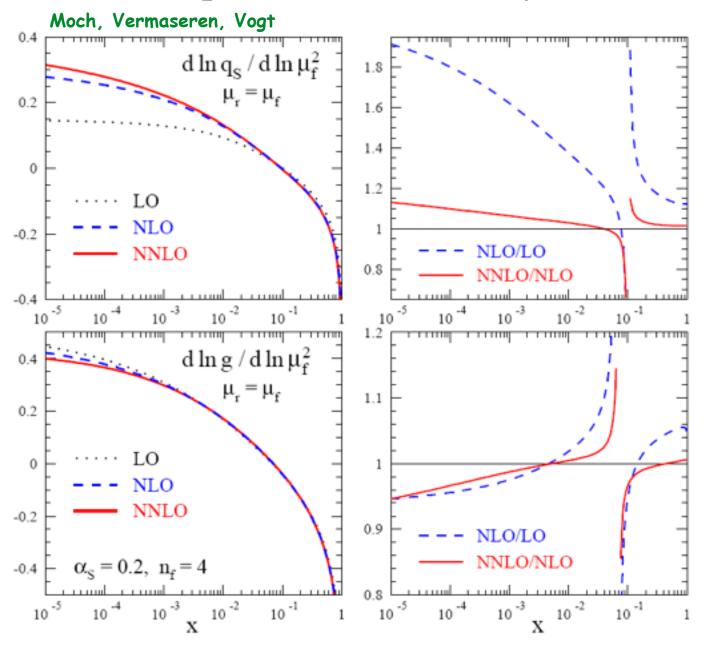
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taken from G. Salam

Х

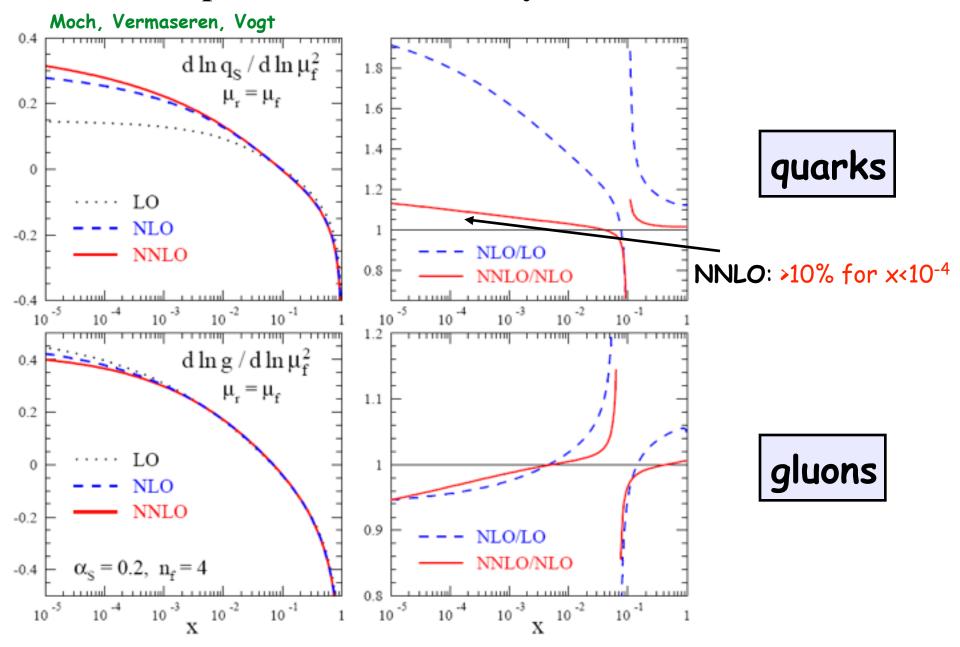
#### perturbative stability of evolution



# quarks



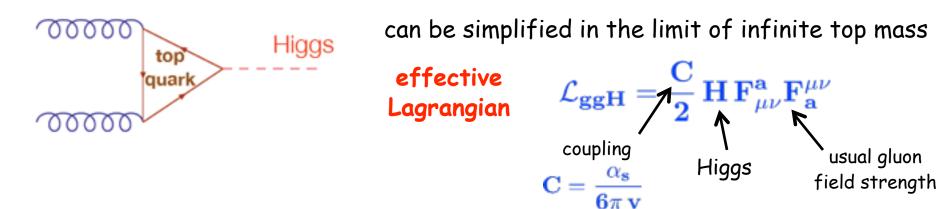
#### perturbative stability of evolution



# aside: universality of splitting fcts

example taken from J. Campbell

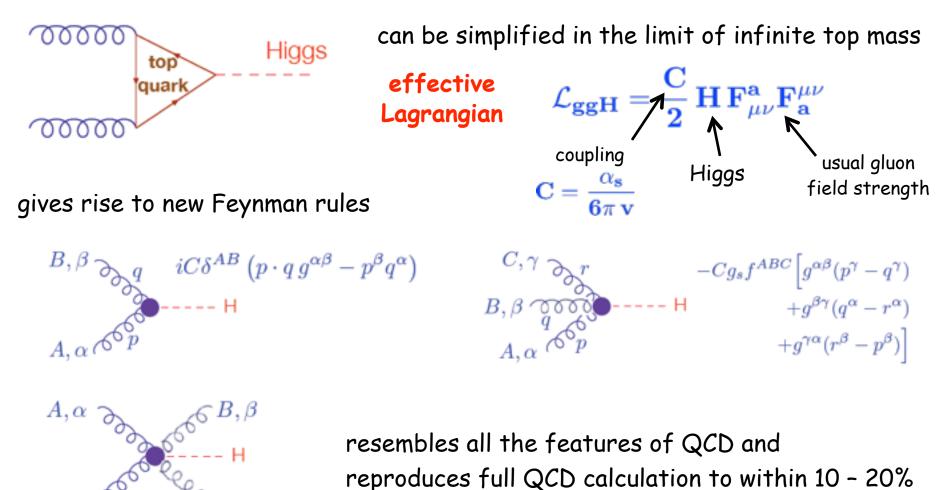
let's look at collinear singularities in a "QCD-ish" effective theory



## aside: universality of splitting fcts

example taken from J. Campbell

let's look at collinear singularities in a "QCD-ish" effective theory



so, what do we encounter in an actual calculation?

## sketch of a calculation in the effective Hgg theory

start with the tree-level diagram (recall: one-loop in full QCD)

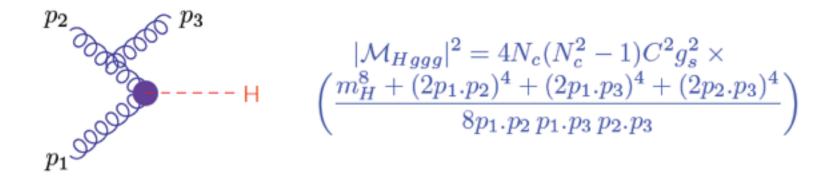
$$p_2$$
  
 $p_1$   
 $M_{Hgg}|^2 = 2(N_c^2 - 1)C^2 m_H^4$ 

## sketch of a calculation in the effective Hgg theory

start with the tree-level diagram (recall: one-loop in full QCD)



then add another gluon



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then add another gluon

 $\begin{array}{c} p_{2} & |\mathcal{M}_{Hggg}|^{2} = 4N_{c}(N_{c}^{2} - 1)C^{2}g_{s}^{2} \times \\ & |\mathcal{M}_{Hgggg}|^{2} = 4N_{c}(N_{c}^{2} - 1)C^{2}g_{s}^{2} \times \\ & \left(\frac{m_{H}^{8} + (2p_{1}.p_{2})^{4} + (2p_{1}.p_{3})^{4} + (2p_{2}.p_{3})^{4}}{8p_{1}.p_{2}p_{1}.p_{3}p_{2}.p_{3}}\right) \\ \text{and evaluate in the collinear limit for } p_{2} \text{ and } p_{3} \\ & 2p_{2} \cdot p_{3} \rightarrow 0 \\ \text{use} \quad 2p_{1} \cdot p_{2} \rightarrow zm_{H}^{2} \\ & 2p_{1} \cdot p_{3} \rightarrow (1 - z)m_{H}^{2} \end{array}$ 

find

$$|\mathcal{M}_{Hggg}|^2 \xrightarrow{\text{coll.}} 4N_c (N_c^2 - 1)C^2 g_s^2 m_H^4 \left(\frac{1 + z^4 + (1 - z)^4}{2z(1 - z)p_2 p_3}\right)$$

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can factor out the LO result  $|\mathcal{M}_{Hgg}|^2 = 2(N_c^2-1)C^2m_H^4$ 

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m  
$$p_1 \qquad p_2 \qquad p_3$$
  
m

similarly, one can obtains  $\mathsf{P}_{\mathsf{qq}}$  from

## factorization in hadron-hadron collisions

What happens when two hadrons collide ?

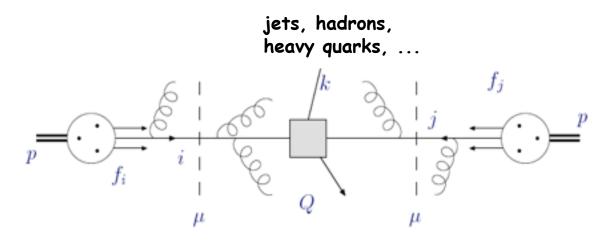


## factorization in hadron-hadron collisions

What happens when two hadrons collide ?



straightforward generalization of the concepts discussed so far:

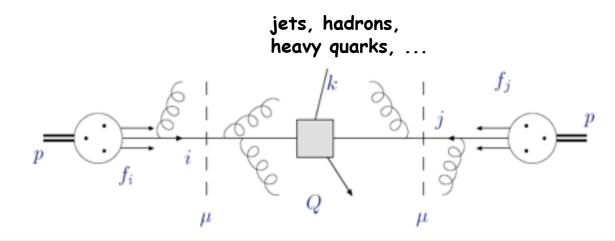


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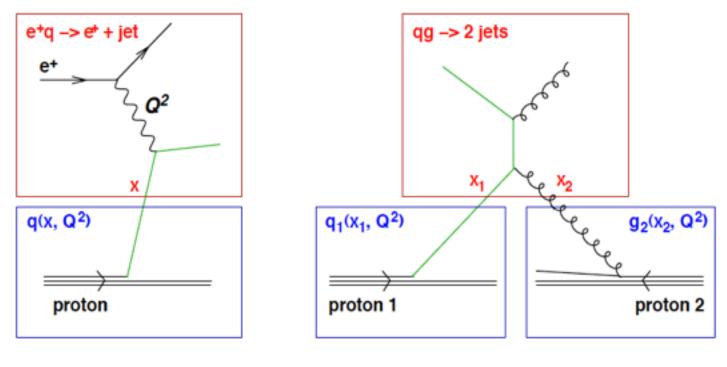
 $\int dx_i dx_j f_i(x_i, \boldsymbol{\mu^2}) f_j(x_j, \boldsymbol{\mu^2}) d\hat{\sigma}_{ij}(\alpha_s(\mu_r), Q^2, \boldsymbol{\mu^2}, x_i, x_j)$ non-perturbative <u>linked</u> hard scattering of but universal PDFs  $\phantom{a}$  by  $\mu$  two partons  $\rightarrow$  pQCD

## factorization at work

key assumption that a cross section factorizes into

- hard (perturbatively calculable) process-dep. partonic subprocesses
- non-perturbative but universal parton distribution functions

has great predictive power and can be challenged experimentally:





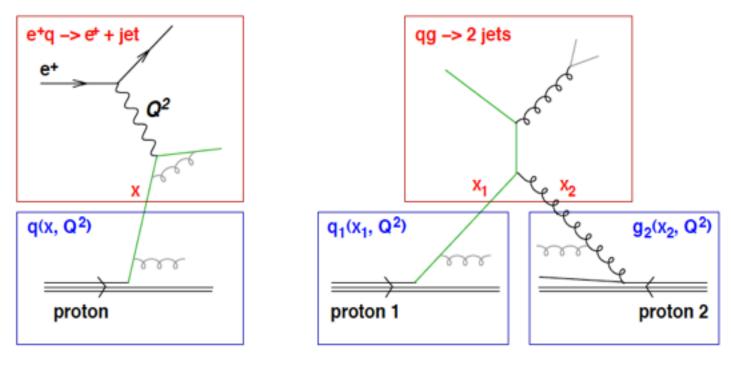


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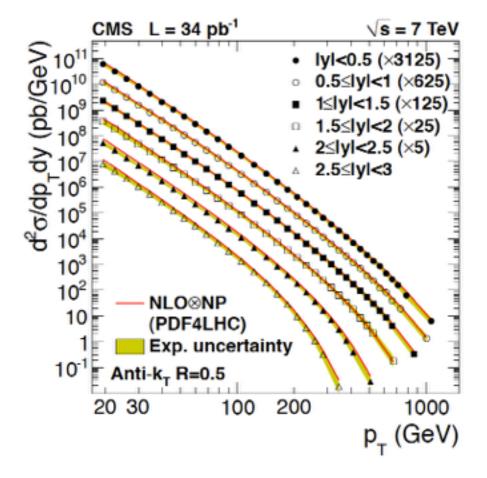
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 $\sigma_{ep} = \sigma_{eq} \otimes q$ 

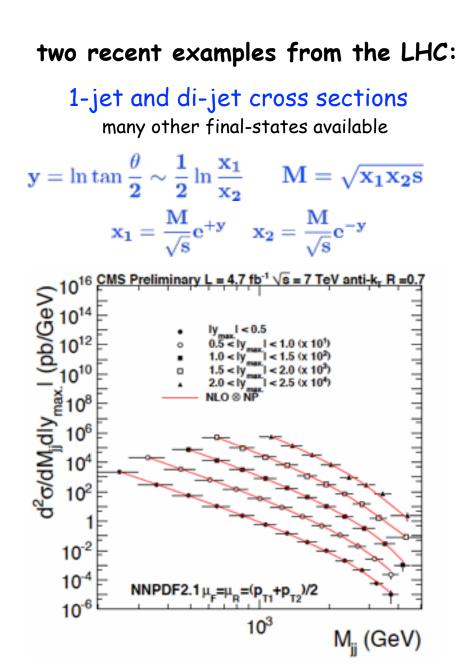


#### factorization: so far a success story



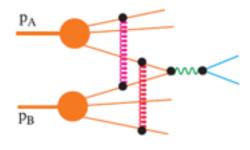
results now start to being used in global fits to constrain PDFs **particularly sensitive to gluons** 

 $\mathbf{g}\mathbf{g} 
ightarrow \mathbf{g}\mathbf{g} \quad \mathbf{g}\mathbf{q} 
ightarrow \mathbf{g}\mathbf{q}$ 



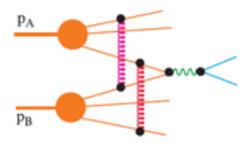
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   such as DIS and Drell-Yan Libby, Sterman; Ellis et al.; Amati et al.; Collins et al.;...



<u>issues</u>: factorization does not hold graph-by-graph; saved by the interplay between graphs, unitarity, causality, and gauge invariance

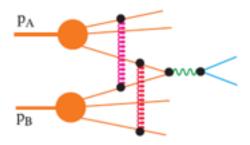
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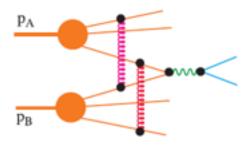


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**recall:** the **renormalizibility** of a non-abelian gauge theory like QCD was demonstrated by 't Hooft and Veltman







recap: salient features of pQCD



#### recap: salient features of pQCD

- strong interactions, yet perturbative methods are applicable
- confined quarks, yet calculations based on free partons can describe large classes of processes

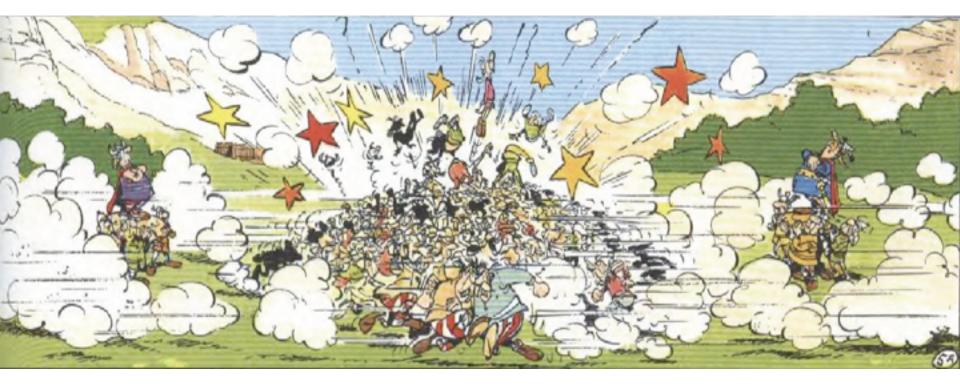


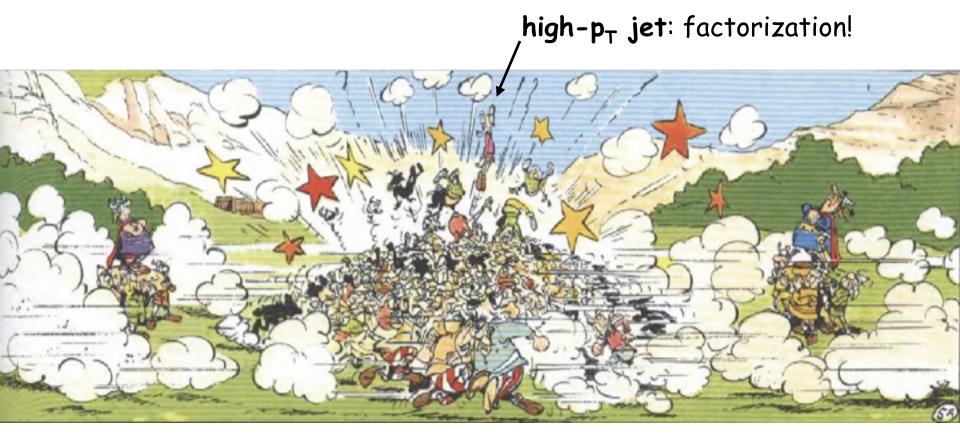
#### recap: salient features of pQCD

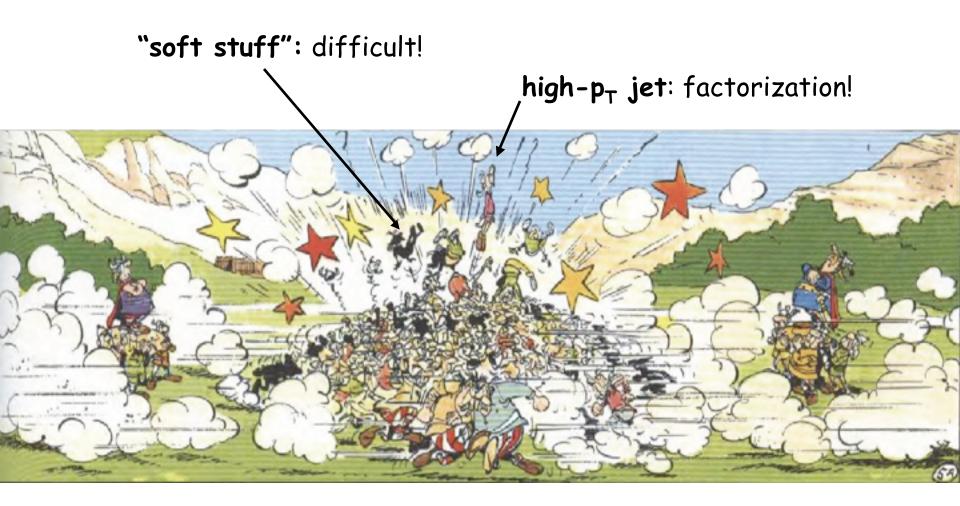
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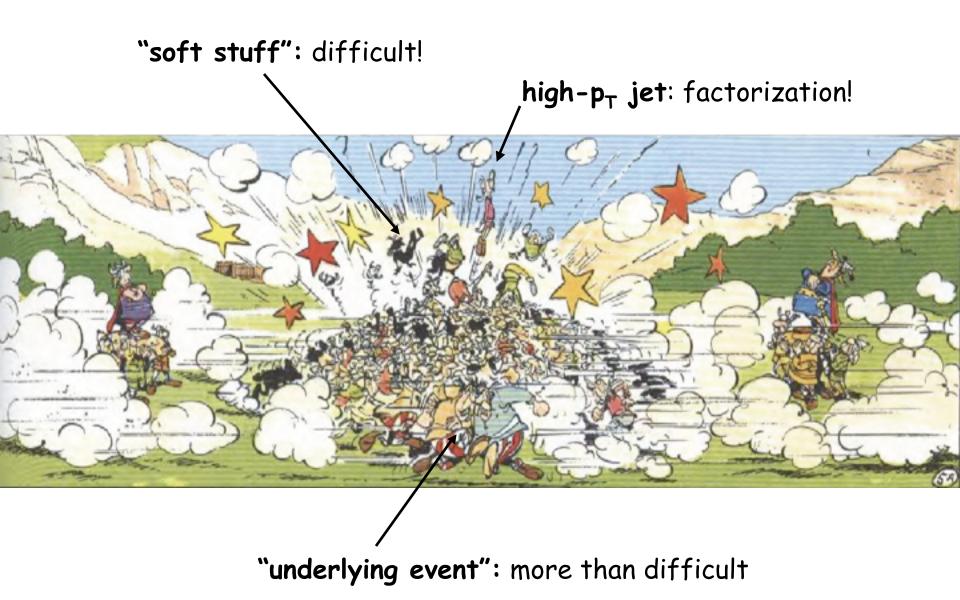
keys to resolve the apparent dilemma:

- asymptotic freedom
- infrared safety
- factorization theorems & renormalizibility

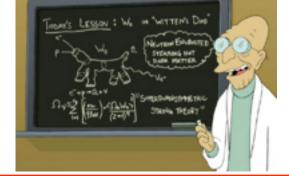








to take home from this part of the lectures INWARD BOUND - FEMTO'SCOPY



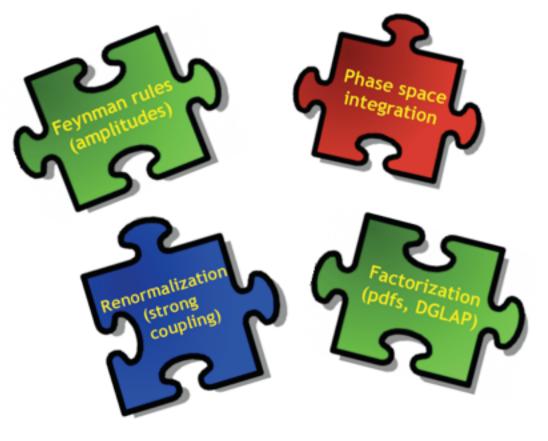
factorization = isolating and absorbing long-distance singularities accompanying identified hadrons into parton densities (initial state) and fragmentation fcts. (final state)

• factorization and renormalization introduce arbitrary scales  $\rightarrow$  powerful concept of renormalization group equations  $\rightarrow \alpha_s$ , PDFs, frag. fcts. depend on energy/resolution

PDFs (and frag. fcts) have definitions as bilocal operators

 $\blacksquare$  hard hadron-hadron interactions factorize as well: <code>ffd\sigma</code>

strict proofs of factorization only for limited class of processes



### Part IV

### some applications & advanced topics

scales and theoretical uncertainties; Drell-Yan process small-x physics; global QCD analysis; resummations

**CERN SppS** [1981  $\rightarrow$  1990] pp collisions 540, 630 GeV

W,Z discovery, jets, ... early successes of QCD





Fermilab TeVatron [1987  $\rightarrow$  2011]

pp collisions 0.63, 1.8, 1.96 TeV

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## **BNL RHIC** $[2000 \rightarrow ...]$ pp collisions up to 500 GeV

the World's first and only **polarized collider** spin dep. phenomena, spin strct. of the nucleon also very versatile heavy ion program





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CERN LHC [operating] pp collisions up to 14(?) TeV a QCD machine, discoveries ? also a PbPb and pPb program

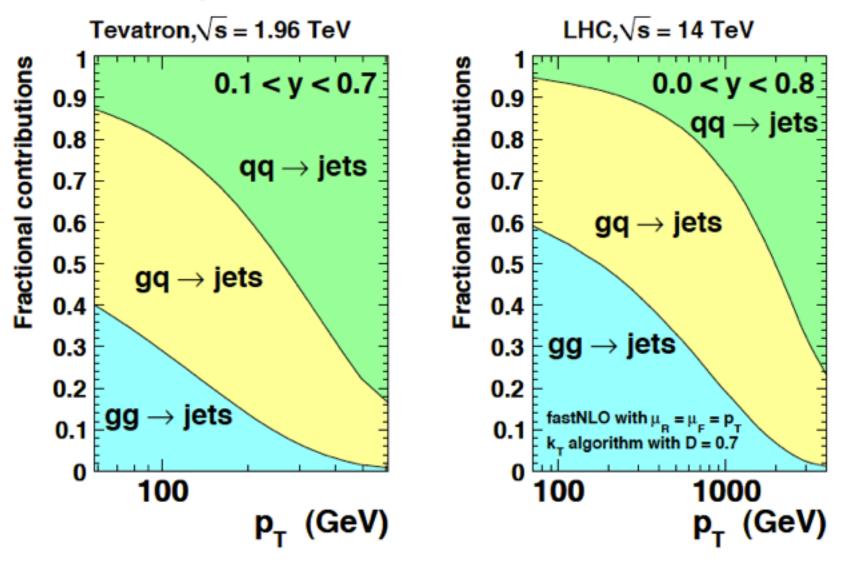
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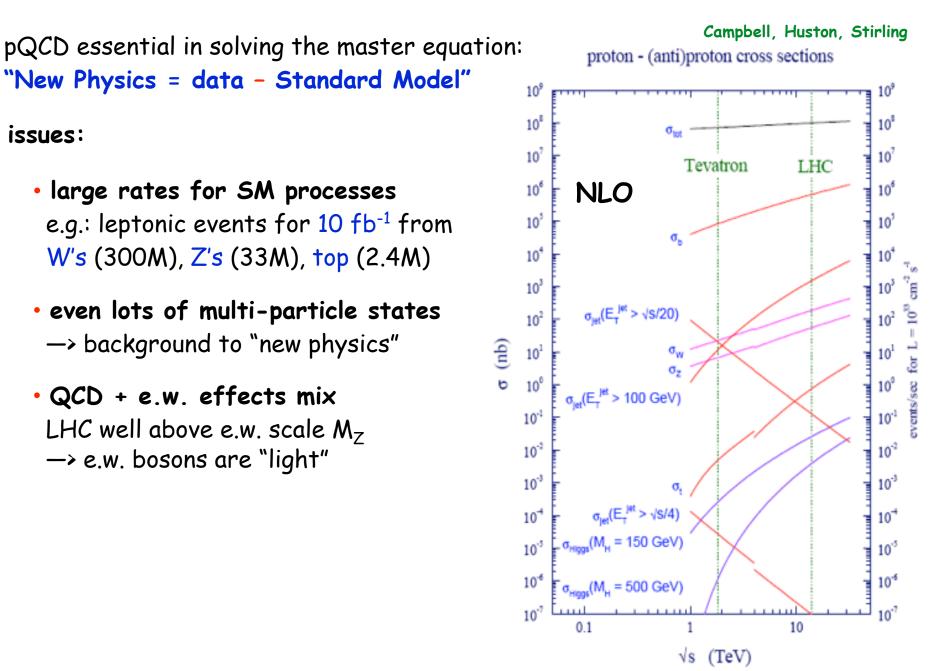
#### jets: which parton processes contribute

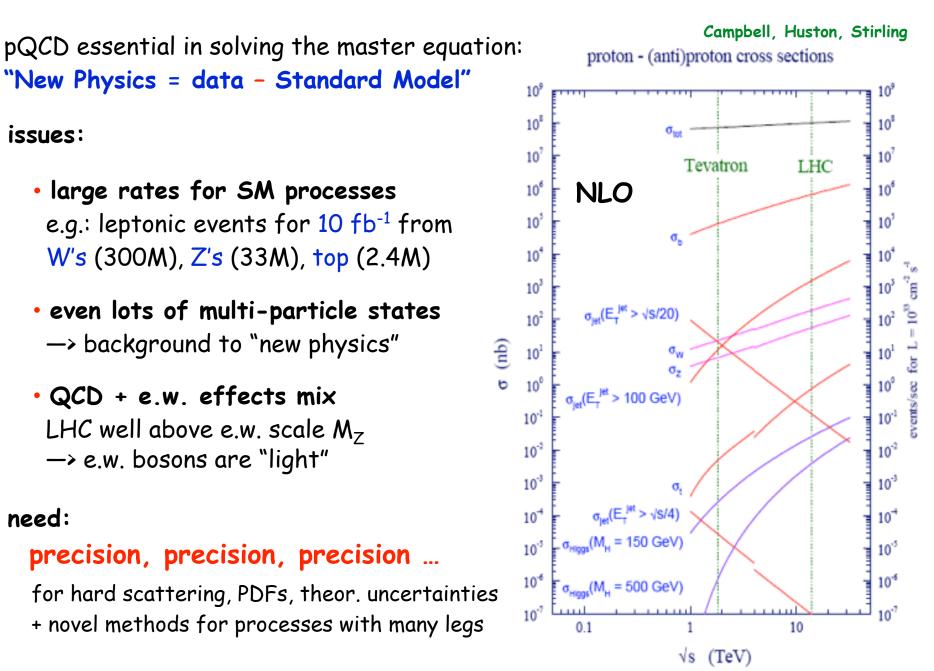
Inclusive jet cross sections with MSTW 2008 NLO PDFs

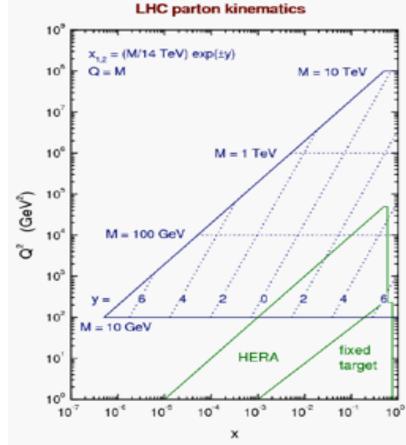


from G. Salam's lectures

hadron colliders are gluon dominated up to rather large  $p_{T}$ 

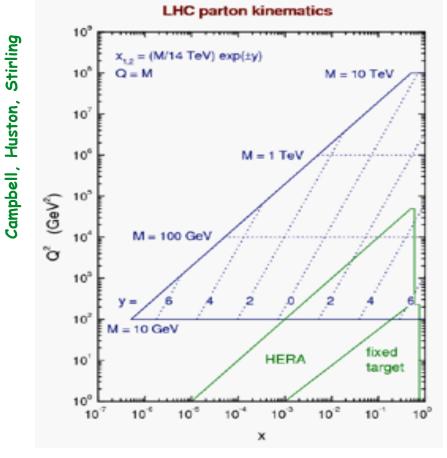






PDFs: vastly extended x,Q<sup>2</sup> landscape

- HERA -> LHC: evolution across up to 3 decades in Q<sup>2</sup>
- M < 100 GeV physics: small x relevant
- TeV scale physics: large x relevant
- large angles/rapidities: extreme x

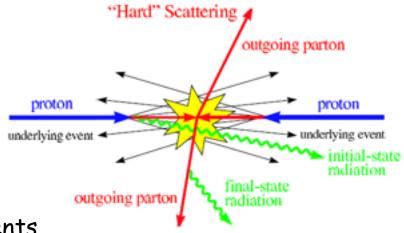


real events at the LHC are very messy:

- possible interactions of spectator partons leading to multiple interactions/underlying events
  - -> relies on event generators (Sherpa, Herwig, ...); state-of-the-art: merge with NLO calculations (MC@NLO, POWHEG, ...)

PDFs: vastly extended x,Q<sup>2</sup> landscape

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### 4-1 the Whys and Hows of NLO Calculations & Beyond

#### why go beyond LO (and even NLO)?

recall factorization theorem for hadronic processes:

$$d\sigma = \sum_{ij} \int dx_i dx_j f_i(x_i, \mu^2) f_j(x_j, \mu^2) d\hat{\sigma}_{ij}(\alpha_s(\mu_r), Q^2, \mu^2, x_i, x_j)$$
  
non-perturbative hard scattering of  
but universal PDFs by  $\mu$  two partons  $\rightarrow$  pQCD

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**caveat:** we work with a perturbative series truncated at LO, NLO, NNLO, ...  $\rightarrow$  at any fixed order N there will be a **residual scale dependence** 

- in our theoretical prediction
- $\rightarrow$  since  $\mu$  is completely arbitrary this limits the precision of our results



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simplest example:  $e^+e^- \rightarrow hadrons$   $\frac{d}{d\ln\mu_r}\sum_{n=1}^N c_n(\mu_r)\alpha_s^n(\mu_r) \sim \mathcal{O}\left(\alpha_s^{N+1}(\mu_r)\right)$ 

applies in general also for  $\mu_f$ 

uncertainty is formally of higher order -> gets smaller if higher orders are known

#### explicit example: scale dependence of $e^+e^- \rightarrow jets$

recall: at NLO we have  $\sigma^{\text{NLO}}(\mu_R) = \sigma_{q\bar{q}} \left(1 + c_1 \alpha_s(\mu_R)\right)$ 

result

NLO coefficient independent of scale

all scale uncertainty from strong coupling

independent of scale

from strong coupling

**recall**: at NLO we have 
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LO  $\Lambda_{\text{LO}} \Lambda_{\text{result}}$  NLO coefficient all scale uncertainty

suppose we want to choose a different scale Q - what do we need to do?

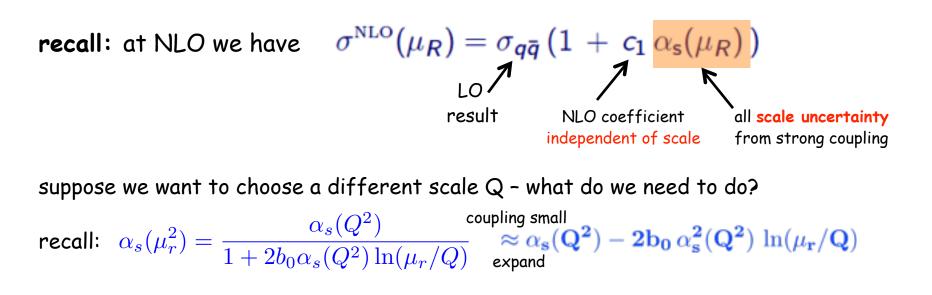
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recall:  $\alpha_s(\mu_r^2) = \frac{\alpha_s(Q^2)}{1 + 2b_0\alpha_s(Q^2)\ln(\mu_r/Q)}$ 



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suppose we want to choose a different scale Q - what do we need to do? recall:  $\alpha_s(\mu_r^2) = \frac{\alpha_s(Q^2)}{1+2b_0\alpha_s(Q^2)\ln(\mu_r/Q)} \approx \alpha_s(\mathbf{Q}^2) - 2\mathbf{b_0}\alpha_s^2(\mathbf{Q}^2)\ln(\mu_r/\mathbf{Q})$ expand

independent of scale

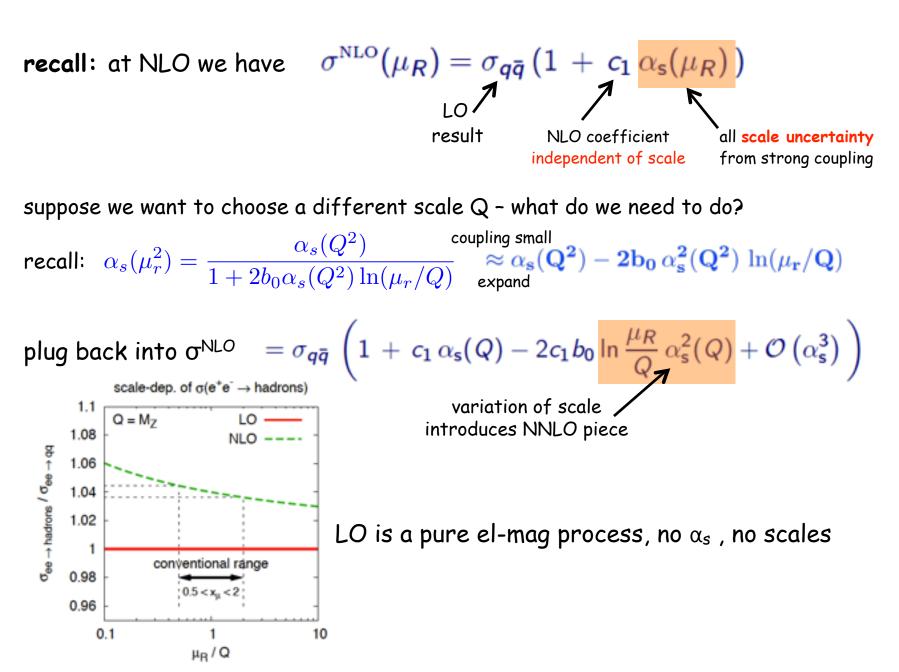
from strong coupling

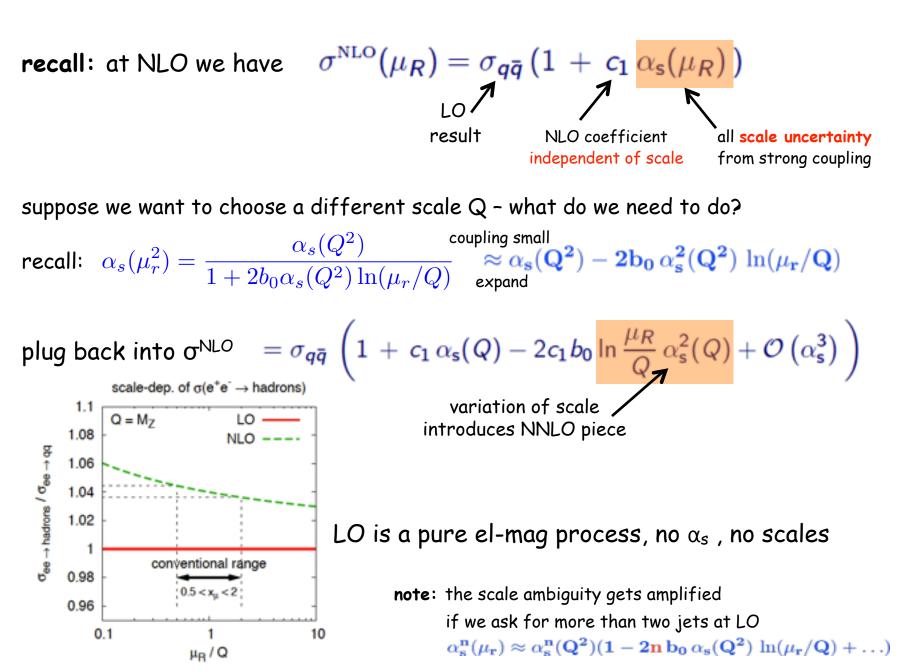
plug back into 
$$\sigma^{\mathsf{NLO}} = \sigma_{q\bar{q}} \left( 1 + c_1 \alpha_{\mathsf{s}}(Q) - 2c_1 b_0 \ln \frac{\mu_R}{Q} \alpha_{\mathsf{s}}^2(Q) + \mathcal{O}\left(\alpha_{\mathsf{s}}^3\right) \right)$$

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$$\sigma^{\text{NLO}}(\mu_R) = \sigma_{q\bar{q}} \left(1 + c_1 \alpha_s(\mu_R)\right)$$

$$\underset{\text{IO}}{\underset{\text{result}}{\underset{\text{independent of scale}}{\underset{\text{NLO coefficient}}{\underset{\text{from strong coupling}}{\underset{\text{from strong coupling}}{\underset{\text{suppose we want to choose a different scale Q - what do we need to do?}}$$

$$\underset{\text{recall:}{\underset{\alpha_s(\mu_r^2)}{\underset{1+2b_0\alpha_s(Q^2)}{\underset{(1+2b_0\alpha_s(Q^2)\ln(\mu_r/Q)}{\underset{\text{expand}}{\underset{\text{expand}}{\underset{\text{expand}}{\underset{\text{odd}}{\underset{\text{expand}}{\underset{\text{odd}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{odd}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{\text{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}}{\underset{suppose}$$





#### explicit example - cont'd

next calculate full NNLO result:

$$\sigma^{\text{NNLO}}(\mu_R) = \sigma_{q\bar{q}} \left[ 1 + c_1 \alpha_s(\mu_R) + c_2(\mu_R) \alpha_s^2(\mu_R) \right]$$
NNLO term starts to
depend on the scale

#### explicit example - cont'd

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in fact  $c_2$  must (and will !) cancel the scale ambiguity found at NLO:

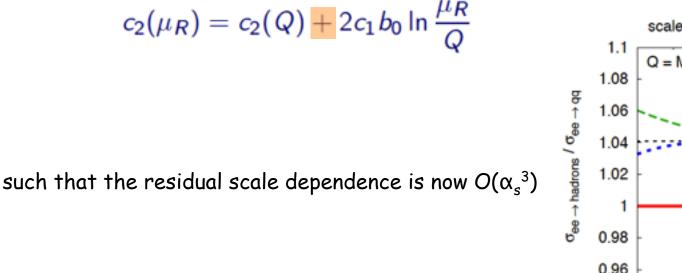
$$c_2(\mu_R) = c_2(Q) + 2c_1b_0 \ln \frac{\mu_R}{Q}$$

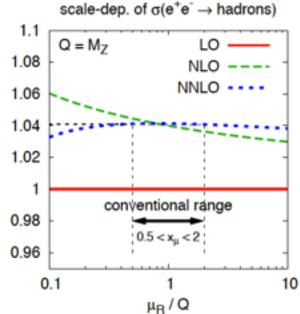
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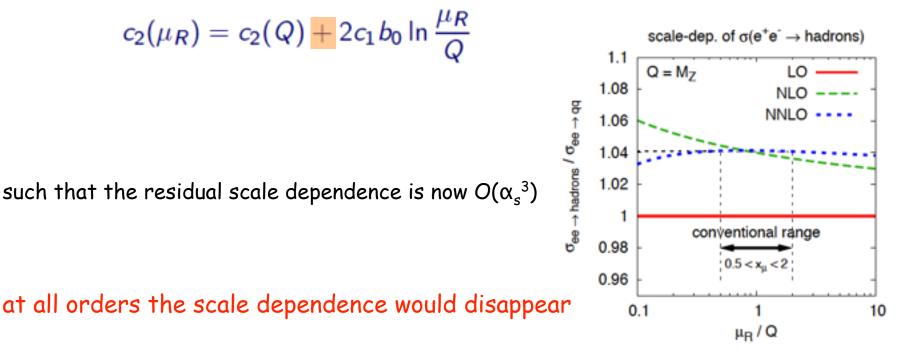


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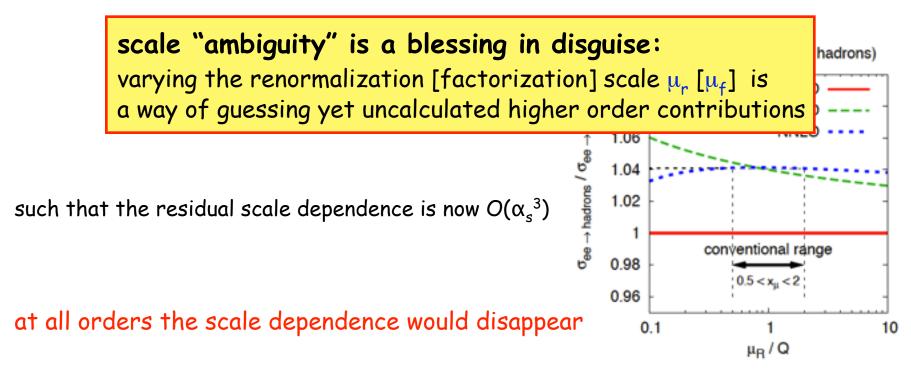


## explicit example - cont'd

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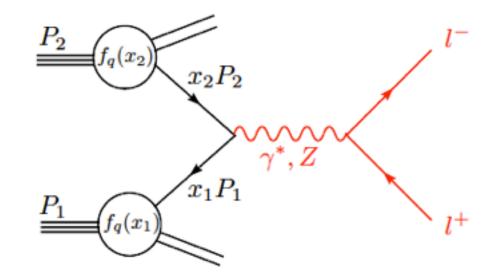
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in fact  $c_2$  must (and will !) cancel the scale ambiguity found at NLO:



## example from hadronic collisions

#### take the "classic" Drell Yan process



- dominated by quarks in the initial-state
- at LO no colored particles in the final-state
- clean experimental signature
- at LO an electromagnetic process (low rate)
- one of the best studied processes (known to NNLO)

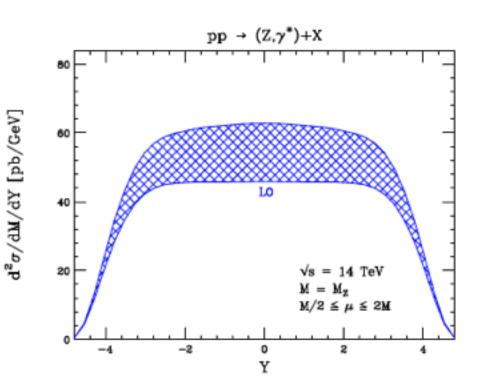
#### as "clean" as it can get at a hadron collider

$$\sigma_{pp \rightarrow Z}^{\text{NLO}} = \sum_{i,j} \int dx_1 dx_2 f_i(x_1, \mu_F^2) f_j(x_2, \mu_F^2) \left[ \hat{\sigma}_{0,ij \rightarrow Z}(x_1, x_2) + \alpha_s(\mu_R) \hat{\sigma}_{1,ij \rightarrow Z}(x_1, x_2, \mu_F) \right]$$

- no  $\alpha_s$  at LO but  $\mu_F$  appears in PDFs
- $\alpha_s$  enters at NLO and hence  $\mu_R$
- NLO terms reduce dep. on  $\mu_F$

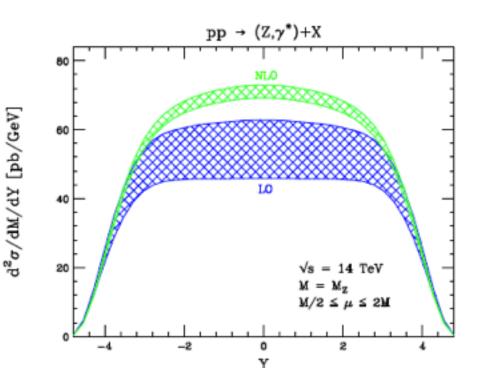
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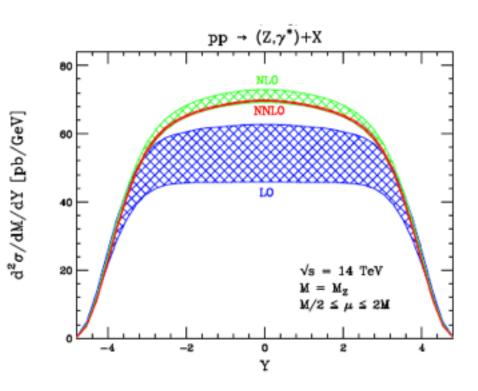
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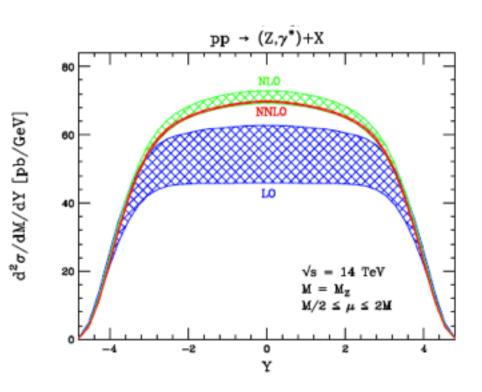
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at NLO:

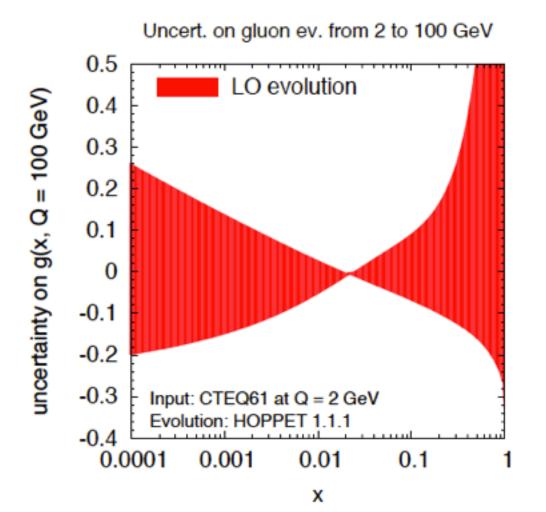
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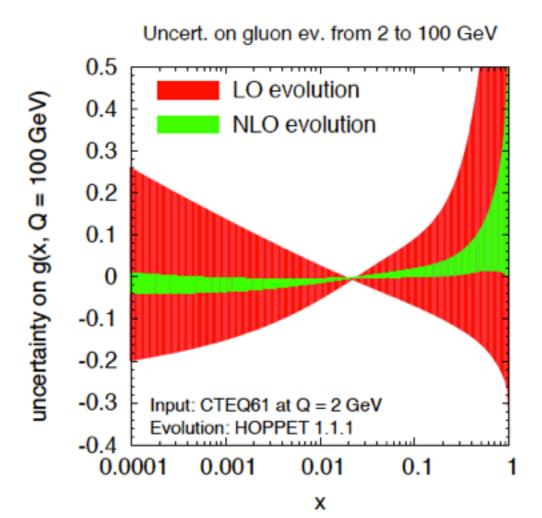
perturbative accuracy of O(percent) achieved

estimate by G. Salam: vary the scale of  $\alpha_s$  in the DGLAP kernel



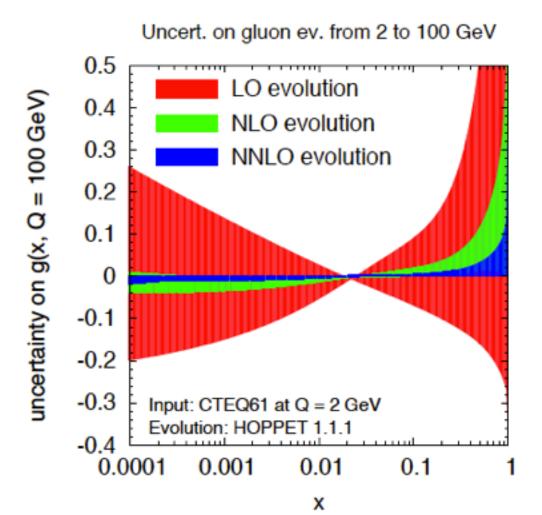


estimate by G. Salam: vary the scale of  $\alpha_s$  in the DGLAP kernel



- about 30% in LO
- down to about 5% in NLO

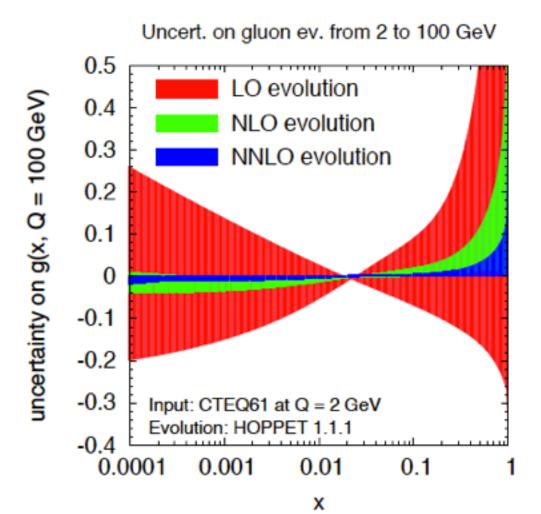
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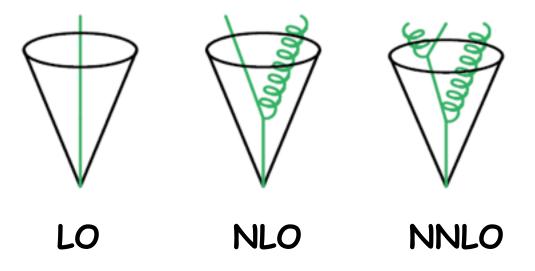


- about 30% in LO
- down to about 5% in NLO
- NNLO brings it down to 2%

which is about the precision of the HERA DIS data

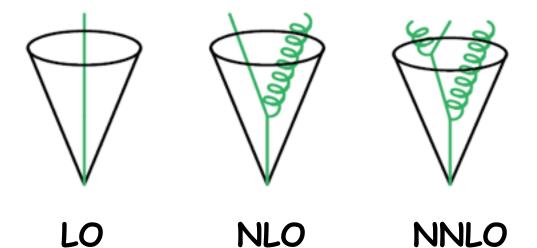
## other motivations for NLO and beyond

• much more realistic final states, e.g., more partons can form a jet

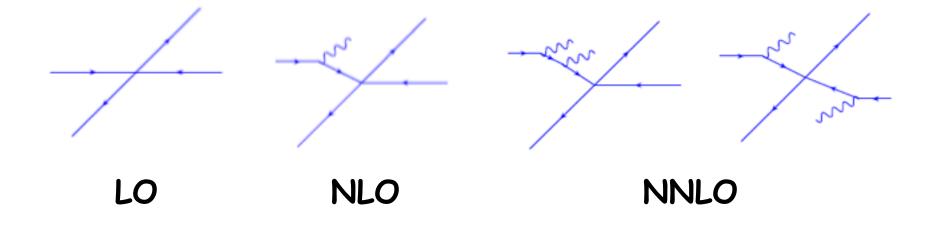


## other motivations for NLO and beyond

• much more realistic final states, e.g., more partons can form a jet



- higher orders generate non-trivial  $k_{\rm T}$  effects/dependence



|   | 2->1 | 2->2 | 2->3 | 2->4 | 2->5 | 2->6 |
|---|------|------|------|------|------|------|
| 1 | LO   |      |      |      |      |      |
| α | NLO  | LO   |      |      |      |      |
| α | NNLO | NLO  | LO   |      |      |      |
| α |      | NNLO | NLO  | LO   |      |      |
| a |      |      | NNLO | NLO  | LO   |      |
| α |      |      |      |      | NLO  | LO   |

|   | 2->1 | 2->2 | 2->3 | 2->4 | 2->5 | 2->6 |
|---|------|------|------|------|------|------|
| 1 | LO   |      |      |      |      |      |
| ۵ | NLO  | LO   |      |      |      |      |
| ۵ | NNLO | NLO  | LO   |      |      |      |
| ۵ |      | NNLO | NLO  | LO   |      |      |
| ۵ |      |      | NNLO | NLO  | LO   |      |
| α |      |      |      |      | NLO  | LO   |

LO matrix elements up to  $2 \rightarrow 8$  and phase space integration (automatically generated); interfaced with parton shower; large  $\mu$  uncertainties though

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NLO NLO all 2  $\rightarrow$  2 SM/MSSM processes; matching with parton shower started some 2  $\rightarrow$  3 results: pp  $\rightarrow$  jjj, Hjj, VVV, ... some 2  $\rightarrow$  4 results: pp  $\rightarrow$  VVjj, Hjjj, ttbb, ttjj,Vjjj, VVbb; also Wjjjj

|   | 2->1 | 2->2 | 2->3 | 2->4 | 2->5 | 2->6 |
|---|------|------|------|------|------|------|
| 1 | LO   |      |      |      |      |      |
| α | NLO  | LO   |      |      |      |      |
| α | NNLO | NLO  | LO   |      |      |      |
| α |      | NNLO | NLO  | LO   |      |      |
| α |      |      | NNLO | NLO  | LO   |      |
| α |      |      |      |      | NLO  | LO   |

LO matrix elements up to  $2 \rightarrow 8$  and phase space integration (automatically) generated); interfaced with parton shower; large  $\mu$  uncertainties though

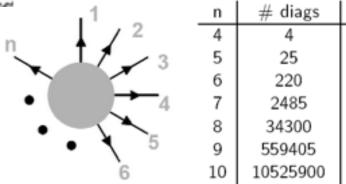
NLO

all  $2 \rightarrow 2$  SM/MSSM processes; matching with parton shower started some 2  $\rightarrow$  3 results: pp  $\rightarrow$  jjj, Hjj, VVV, ... NLO some 2  $\rightarrow$  4 results: pp  $\rightarrow$  VVjj, Hjjj, ttbb, ttjj,Vjjj, VVbb; also Wjjjj

**NNLO** NNLO Drell-Yan type  $2 \rightarrow 1$  processes (total and differential cross sections); splitting functions;  $e^+e^- \rightarrow jjj$ ; progress towards general 2  $\rightarrow$  2 processes including heavy flavor production ( $\sigma_{tot}$  at NNLO done)



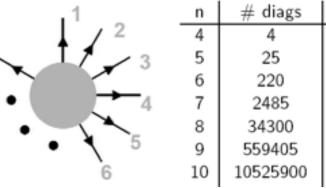
traditional Feynman diagram technique still going strong but becomes very clumsy for high-multiplicity processes:



rapid growth in complexity, but final answers often very simple  $\rightarrow$  new ways to compute amplitudes?



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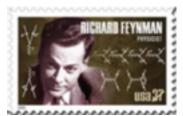


rapid growth in complexity, but final answers often very simple  $\rightarrow$  new ways to compute amplitudes?

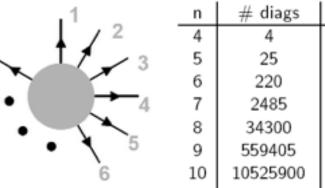
#### ideas:



- use analytical properties of amplitudes (unitarity) as calculational tools
- build amplitudes from simpler amplitudes with fewer legs by recursion
- get "loops from trees"



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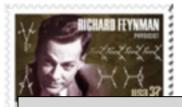


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#### amazing progress in a short time (few years) guided by two principles:

The best way to have a good idea is to have a lot of ideas --- Linus Pauling Those are my principles, and if you don't like them ... well, I have others --- Groucho Marx

#### currently aiming at full automatization at 1-loop level



id

traditional Feynman diagram technique still going strong but becomes very clumsy for high-multiplicity processes:

ble

les?

on

for some ideas, see:

Berends, Giele 1988 - recursion relations (off-shell) Britto, Cachazo, Feng 2004 - recursion relations (on-shell) / unitarity Cachazo, Svrcek, Witten 2004 - MHV amplitudes

Ossola, Pittau, Papadopolous 2006 - NLO loop integrals w/o doing integrals ols

recent report on unitarity method:

Ellis, Kunszt, Melnikov, Zanderighi, arXiv:1105.4319 amazing progress in a snort time (rew years) gaided by two principles.

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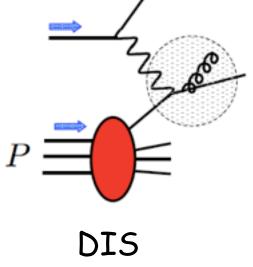
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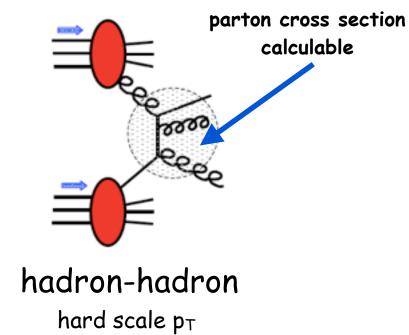
# 4-2 Anatomy of a Global QCD PDF Analysis

## how to determine PDFs from data?

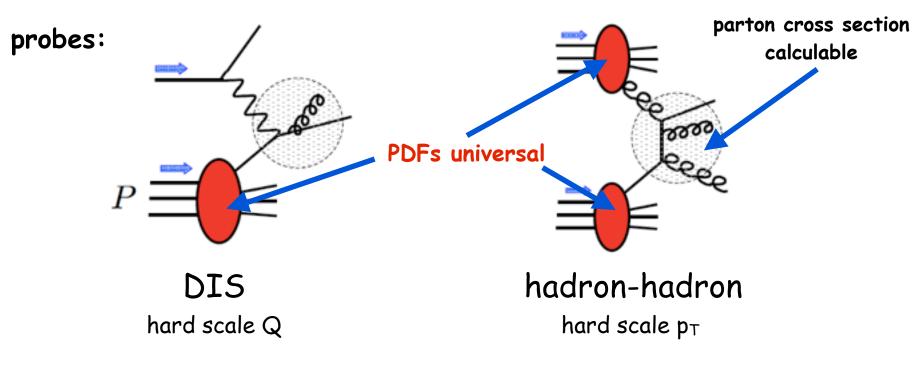
probes:



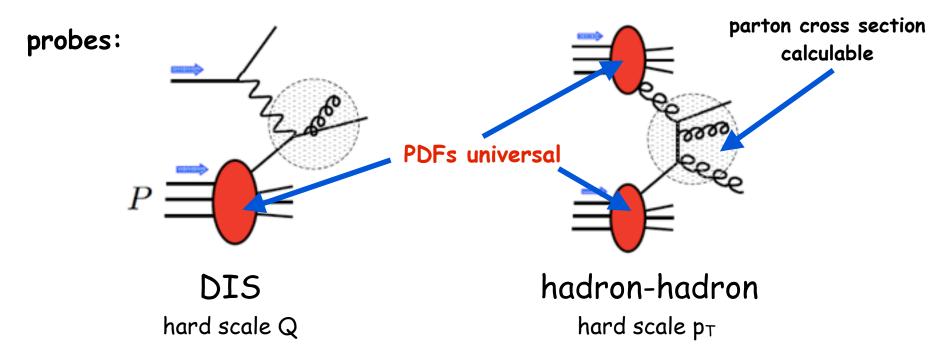
hard scale Q



## how to determine PDFs from data?



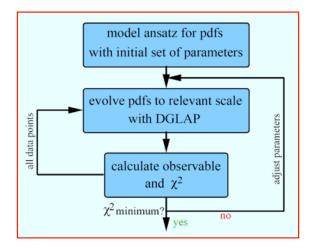
## how to determine PDFs from data?



**task:** extract PDFs and their uncertainties (assume factorization)

- all processes tied together: universality of pdfs & Q<sup>2</sup> evolution
- each reaction provides insights into different aspects and kinematics
- need at least NLO accuracy for quantitative analyses
- information on PDFs "hidden" inside complicated (multi-)convolutions

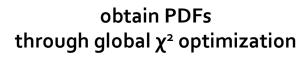
# obtain PDFs through global $\chi^{\rm 2}$ optimization

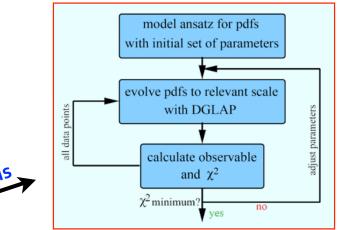


set of **optimum parameters** for *assumed* functional form

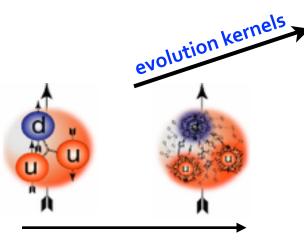
#### computational challenge:

- up to O(20-30) parameters
- many sources of uncertainties
- very time-consuming NLO expressions

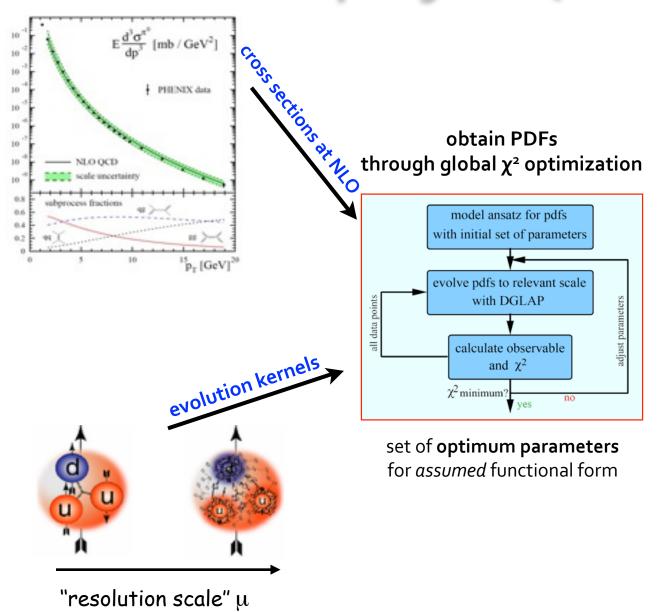


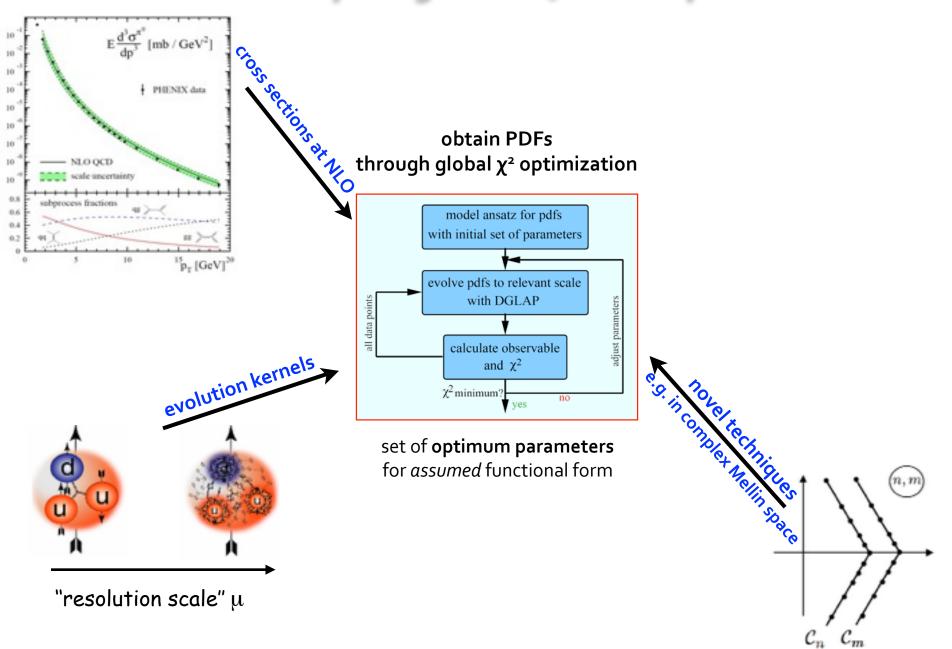


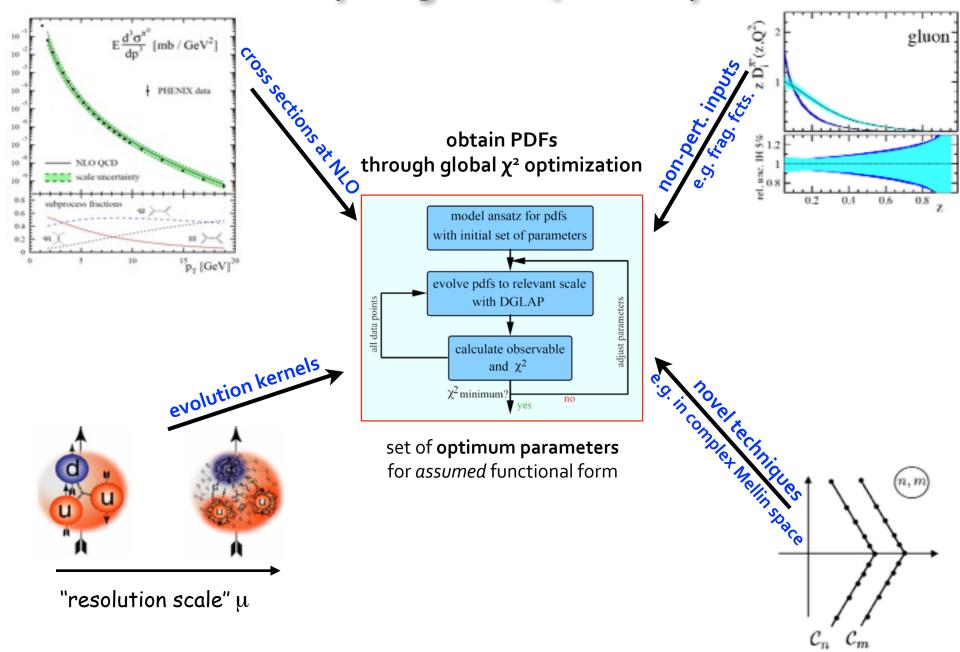
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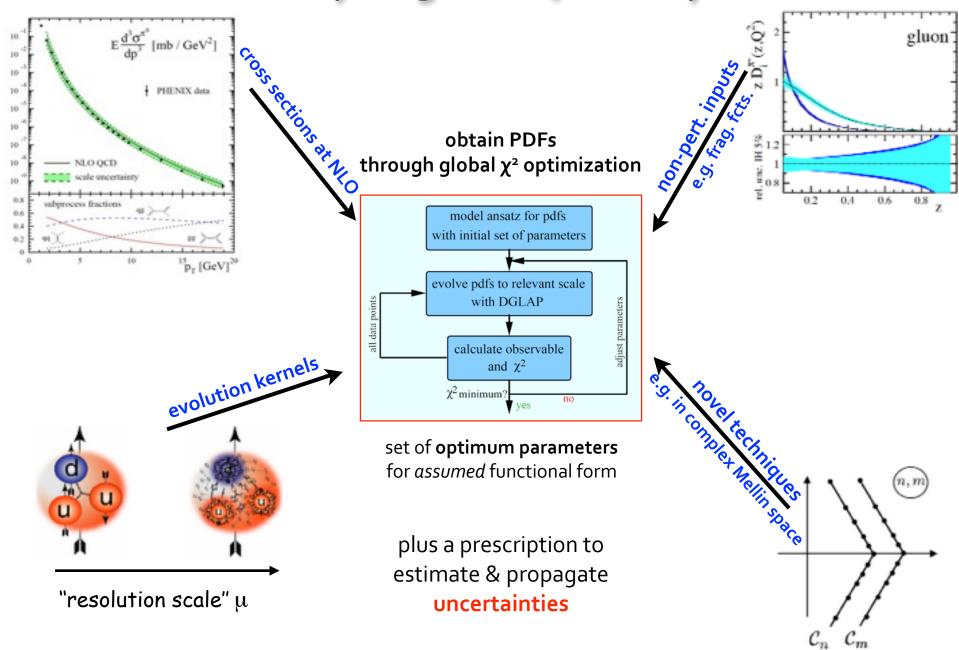


"resolution scale"  $\boldsymbol{\mu}$ 









## global analysis: computational challenge

- one has to deal with O(2800) data points from many processes and experiments
- need to determine O(20-30) parameters describing PDFs at  $\mu_0$
- NLO expressions often very complicated  $\rightarrow$  computing time becomes excessive  $\rightarrow$  develop **sophisticated algorithms & techniques**, e.g., based on Mellin moments Kosower; Vogt; Vogelsang, MS

## global analysis: computational challenge

- one has to deal with O(2800) data points from many processes and experiments
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#### data sets & (x,Q<sup>2</sup>) coverage used in MSTW fit Martin, Stirling, Thorne, Watt, arXiv:0901.0002

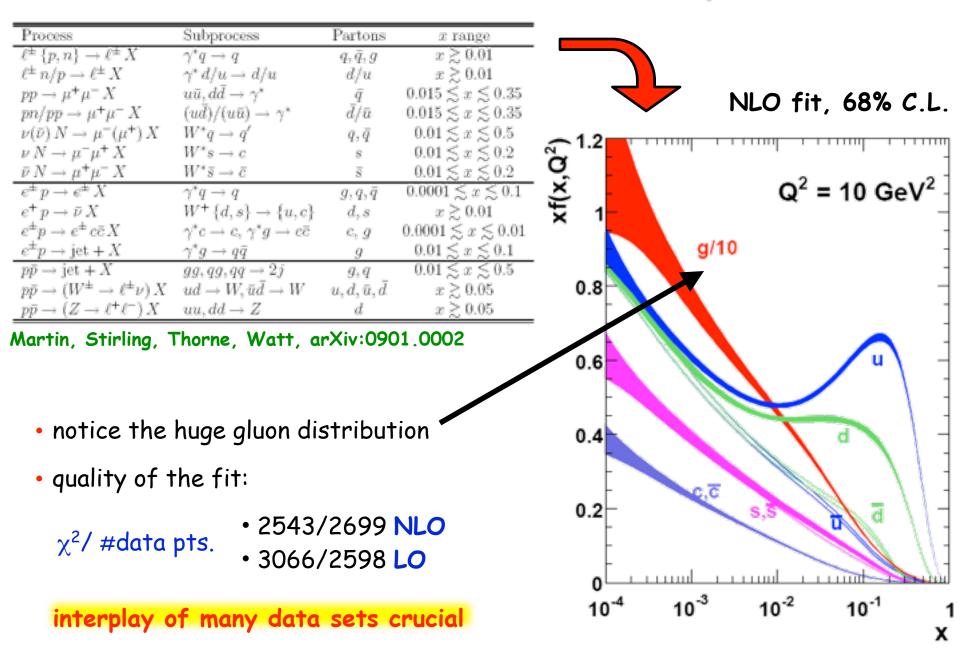
|                                                  |                  |                                     |                    | 2 10 5              | H1        |                                    |          |              |                    |
|--------------------------------------------------|------------------|-------------------------------------|--------------------|---------------------|-----------|------------------------------------|----------|--------------|--------------------|
| Data set                                         | N <sub>pts</sub> | Data set                            | N <sub>pts</sub> . | Q <sup>2</sup> (GeV | ZEUS      |                                    |          | -            | 4                  |
| H1 MB 99 e <sup>∞</sup> ρ NC                     | 8                | BCDMS $\mu p F_2$                   | 163                | ~ F                 |           |                                    |          |              | 7 🥒                |
| H1 MB 97 e <sup>+</sup> p NC                     | 64               | BCDMS µd F2                         | 151                | ° l                 | CDF/D     | Inclusive jets y <sub>1</sub> <0.9 | ,        |              |                    |
| H1 low Q <sup>2</sup> 96–97 e <sup>+</sup> p NC  | 80               | NMC $\mu p F_2$                     | 123                | 10 *                | D0 Ind    | lusive jets n=3                    | _        | $\leftarrow$ |                    |
| H1 high Q <sup>2</sup> 98–99 e <sup>−</sup> p NC | 126              | NMC $\mu d F_2$                     | 123                | Ē                   | 🛄 Fized T | arget Experiments:                 |          |              |                    |
| H1 high Q2 99-00 e+p NC                          | 147              | NMC µn/µp                           | 148                | J.                  | CCER      | NMC, BCDMS,                        | · · · ·  |              |                    |
| ZEUS SVX 95 e <sup>+</sup> p NC                  | -30              | E665 µp F2                          | 53                 | 10 2                |           |                                    |          | 4            |                    |
| ZEUS 96-97 e <sup>+</sup> p NC                   | 144              | E665 µd F2                          | 53                 | F                   | E465, S   | LAC                                |          | /            |                    |
| ZEUS 98-99 e <sup>-</sup> p NC                   | 92               | SLAC ep F2                          | 37                 | - F                 |           |                                    | S.       |              |                    |
| ZEUS 99-00 e <sup>+</sup> p NC                   | 90               | SLAC ed F2                          | 38                 | 10 <sup>2</sup>     |           |                                    | 34       |              |                    |
| H1 99-00 e <sup>+</sup> p CC                     | 28               | NMC/BCDMS/SLAC F                    | 31                 | E                   |           |                                    |          |              |                    |
| ZEUS 99-00 e+p CC                                | - 30             | E866/NuSea pp DY                    | 184                | t t                 |           | ALC: NO.                           |          |              |                    |
| $H1/ZEUS e^{\pm}p F_{2}^{charm}$                 | 83               | E866/NuSea pd/pp DY                 | 15                 | 10                  |           | 3/                                 |          | 1            |                    |
| H1 99-00 e+p incl. jets                          | 24               | NuTeV $\nu N F_2$                   | 53                 |                     |           |                                    |          |              |                    |
| ZEUS 96-97 e <sup>+</sup> p incl. jets           | - 30             | CHORUS WN F                         | 42                 | E                   |           |                                    |          |              |                    |
| ZEUS 98–00 $e^{\pm}p$ incl. jets                 | 30               | NuTeV $\nu N x F_3$                 | 45                 | , [                 |           | <i>/</i> 4 <del></del>             | 1000 a   |              |                    |
| DØ II pp incl. jets                              | 110              | CHORUS #N xF3                       | 33                 | 1                   |           |                                    |          |              |                    |
| CDF II pp incl. jets                             | 76               | CCFR $\nu N \rightarrow \mu \mu X$  | 86                 | F                   |           |                                    |          | ,            |                    |
| CDF II $W \rightarrow l\nu$ asym.                | 22               | NuTeV $\nu N \rightarrow \mu \mu X$ | 84                 | 10 -1               |           |                                    |          |              |                    |
| DØ II $W \rightarrow l\nu$ asym.                 | 10               |                                     | -                  | 10 1                |           |                                    |          |              |                    |
| DØ II Z rap.                                     | 28               | All data sets                       | 2743               | E                   | - M       | <u></u>                            | <u>.</u> | ىتىپ است     |                    |
| CDF II Z rap.                                    | 29               | Red = New w.r.t. MR                 | ST 2006 f          | it.                 | 10 10     | ) 10 "                             | 10 -3    | 10 -         | 10 <sup>-1</sup> 1 |
|                                                  |                  | CARD CARD DIFFE FOR                 |                    |                     |           |                                    |          |              | v                  |

## which data sets determine which partons

| Process                                                      | Subprocess                                              | Partons                  | x range                           |                                                                                                                |
|--------------------------------------------------------------|---------------------------------------------------------|--------------------------|-----------------------------------|----------------------------------------------------------------------------------------------------------------|
| $\ell^{\pm} \{p, n\} \rightarrow \ell^{\pm} X$               | $\gamma^* q \rightarrow q$                              | $q, \bar{q}, g$          | $x \gtrsim 0.01$                  |                                                                                                                |
| $\ell^{\pm} n/p \rightarrow \ell^{\pm} X$                    | $\gamma^* d/u  ightarrow d/u$                           | d/u                      | $x \gtrsim 0.01$                  |                                                                                                                |
| $pp \rightarrow \mu^+ \mu^- X$                               | $u\bar{u}_{\underline{i}}d\bar{d} \rightarrow \gamma^*$ | $\bar{q}$                | $0.015 \lesssim x \lesssim 0.35$  | NLO fit, 68% C.L.                                                                                              |
| $pn/pp \rightarrow \mu^+\mu^- X$                             | $(u\bar{d})/(u\bar{u}) \rightarrow \gamma^*$            | $\bar{d}/\bar{u}$        | $0.015 \lesssim x \lesssim 0.35$  |                                                                                                                |
| $\nu(\bar{\nu}) N \rightarrow \mu^{-}(\mu^{+}) X$            | $W^*q \rightarrow q'$                                   | $q, \bar{q}$             | $0.01 \lesssim x \lesssim 0.5$    | - 1.2 - · · · · · · · · · · · · · · · · · ·                                                                    |
| $\nu N \rightarrow \mu^- \mu^+ X$                            | $W^*s \rightarrow c$                                    | 8                        | $0.01 \lesssim x \lesssim 0.2$    | $\hat{\mathbf{Q}}^{1.2}_{\mathbf{X}}$<br>$\hat{\mathbf{Q}}^{2}_{\mathbf{X}} = 10 \text{ GeV}^{2}_{\mathbf{X}}$ |
| $\bar{\nu} N \rightarrow \mu^+ \mu^- X$                      | $W^*\bar{s} \rightarrow \bar{c}$                        | ŝ                        | $0.01 \leq x \leq 0.2$            |                                                                                                                |
| $e^{\pm} p \rightarrow e^{\pm} X$                            | $\gamma^* q \rightarrow q$                              | $g, q, \bar{q}$          | $0.0001 \lesssim x \lesssim 0.1$  | $Q^2 = 10 \text{ GeV}^2$                                                                                       |
| $e^+ p \rightarrow \bar{\nu} X$                              | $W^+ \{d, s\} \rightarrow \{u, c\}$                     | $d_2 s$                  | $x \gtrsim 0.01$                  | <b>∑</b> 1 –                                                                                                   |
| $e^{\pm}p \rightarrow e^{\pm}c\bar{c}X$                      | $\gamma^* c \to c, \gamma^* g \to c \bar{c}$            | c, g                     | $0.0001 \lesssim x \lesssim 0.01$ |                                                                                                                |
| $e^{\pm}p \rightarrow \text{jet} + X$                        | $\gamma^* g \rightarrow q \bar{q}$                      | g                        | $0.01 \lesssim x \lesssim 0.1$    | g/10 -                                                                                                         |
| $p\bar{p} \rightarrow jet + X$                               | $gg, qg, qq \rightarrow 2j$                             | g, q                     | $0.01 \lesssim x \lesssim 0.5$    |                                                                                                                |
| $p\bar{p} \rightarrow (W^{\pm} \rightarrow \ell^{\pm}\nu) X$ |                                                         | $u, d, \bar{u}, \bar{d}$ | $x \gtrsim 0.05$                  | 0.8                                                                                                            |
| $p\bar{p} \rightarrow (Z \rightarrow \ell^+ \ell^-) X$       | $uu, dd \rightarrow Z$                                  | d                        | $x \gtrsim 0.05$                  |                                                                                                                |
| Martin, Stirling, 7                                          | Thorne, Watt, a                                         | rXiv:090                 | 01.0002                           |                                                                                                                |
| ,                                                            |                                                         |                          |                                   | 0.6 u -                                                                                                        |
|                                                              |                                                         |                          |                                   |                                                                                                                |
|                                                              |                                                         |                          |                                   |                                                                                                                |
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|                                                              |                                                         |                          |                                   | 0.2 s,s a l                                                                                                    |
|                                                              |                                                         |                          |                                   |                                                                                                                |
|                                                              |                                                         |                          |                                   |                                                                                                                |
|                                                              |                                                         |                          |                                   |                                                                                                                |
|                                                              |                                                         |                          |                                   | 10 <sup>-4</sup> 10 <sup>-3</sup> 10 <sup>-2</sup> 10 <sup>-1</sup> 1                                          |
|                                                              |                                                         |                          |                                   |                                                                                                                |

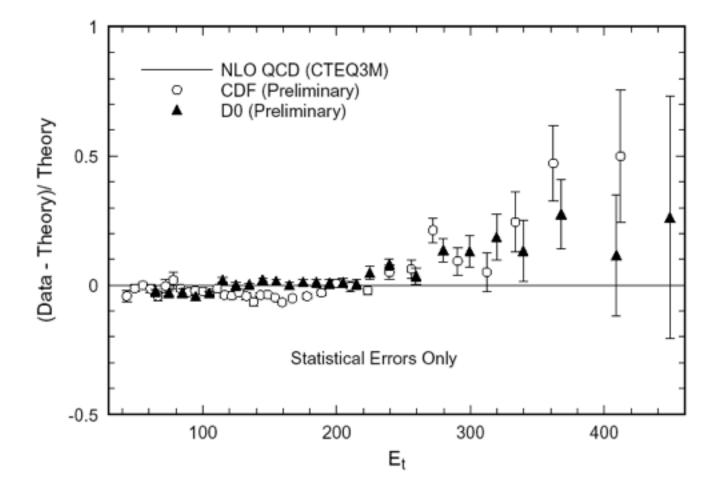
Х

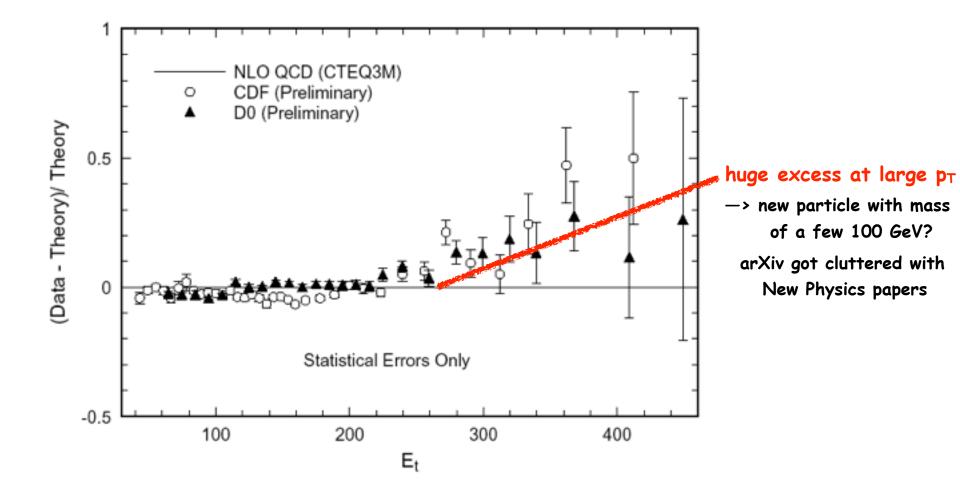
## which data sets determine which partons

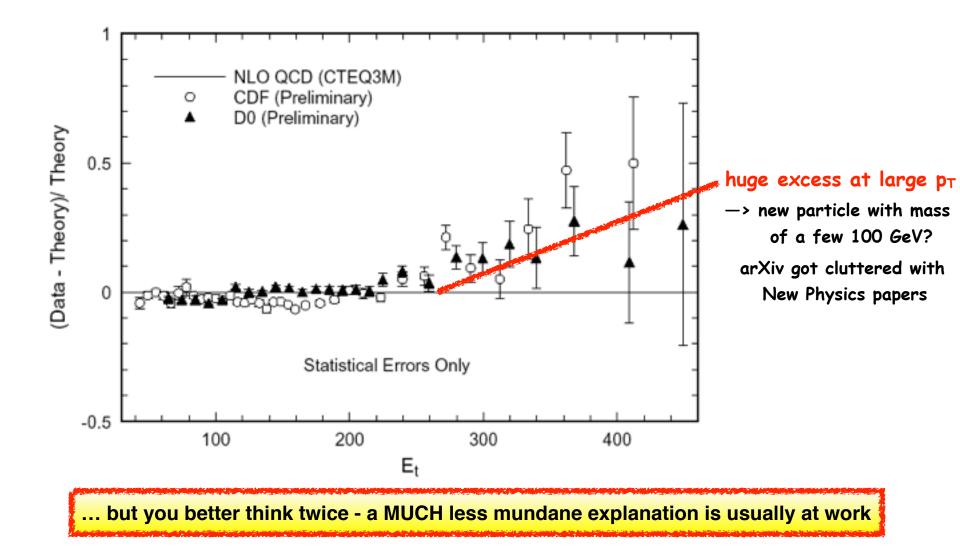


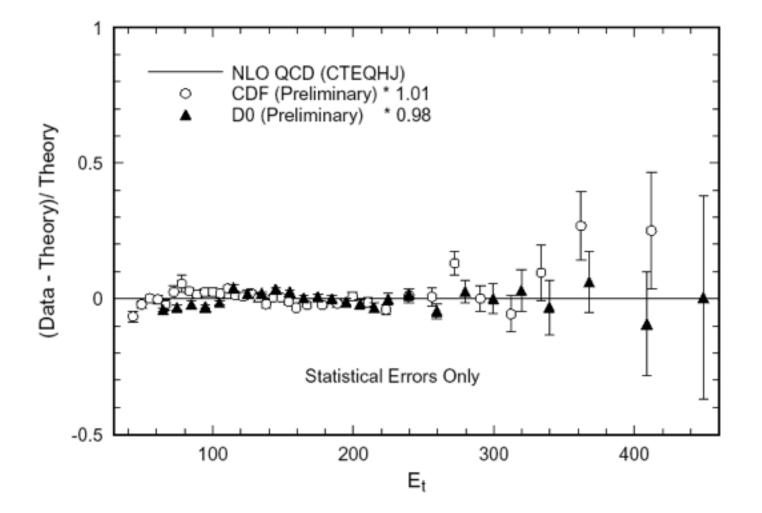


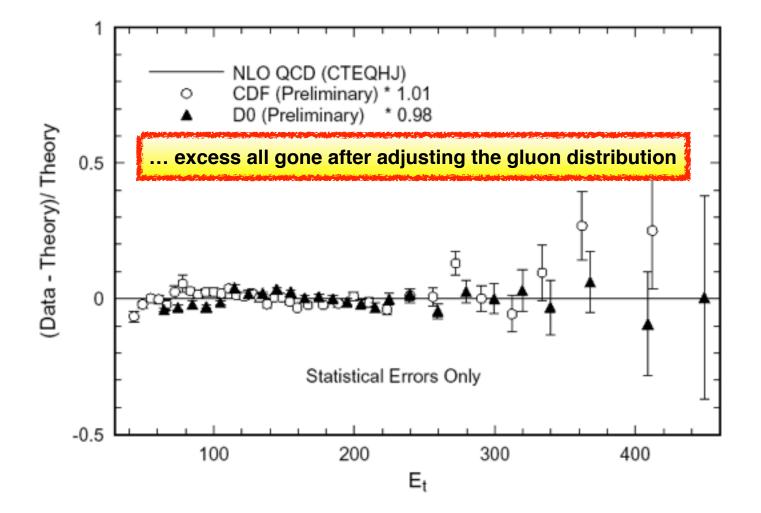
from R.D. Ball











# what's on the market?



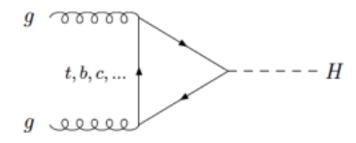
| set          | H.O. | data                        | $\alpha_s(M_Z)@NNLO$ | uncertainty                      | HQ                       | Comments                   |
|--------------|------|-----------------------------|----------------------|----------------------------------|--------------------------|----------------------------|
| MSTW<br>2008 | NNLO | DIS+DY+Jets                 | 0.1171               | Hessian (dynamical<br>tolerance) | GM-VFN<br>(ACOT+TR')     | old HERA DIS               |
| СТІО         | NNLO | DIS+DY+Jets                 | 0.118                | Hessian (dynamical<br>tolerance) | GM-VFN<br>(SACOT-X)      | New HERA<br>DIS            |
| NNPDF        | NNLO | DIS+DY+Jets<br>+LHC         | 0.1174               | Monte Carlo                      | GM-VFN<br>(FONLL)        | New HERA<br>DIS            |
| ABKM         | NNLO | DIS+DY(f.t.)<br>+DY-tT(LHC) | 0.1132               | Hessian                          | FFN<br>BMSN              | New HERA<br>DIS            |
| (G)JR        | NNLO | DIS+DY(f.t.)+<br>some jet   | 0.1124               | Hessian                          | FFN<br>(VFN<br>massless) | valence like<br>input pdfs |
| HERA<br>PDF  | NNLO | only DIS<br>HERA            | 0.1176               | Hessian                          | GM-VFN<br>(ACOT+TR')     | Latest HERA<br>DIS         |

#### compilation by D. de Florian (DIS 2014)

important example: Higgs production through gluon-gluon fusion

PDF uncertainty: look a parton-parton luminosities

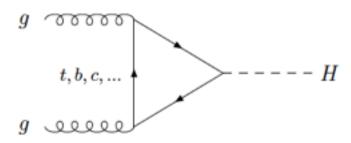
$$\mathcal{L}_{ij}(\tau \equiv M_X^2/S) = \frac{1}{S} \int_{\tau}^{1} \frac{dx}{x} f_i(x, M_X^2) f_j(\tau/x, M_X^2)$$

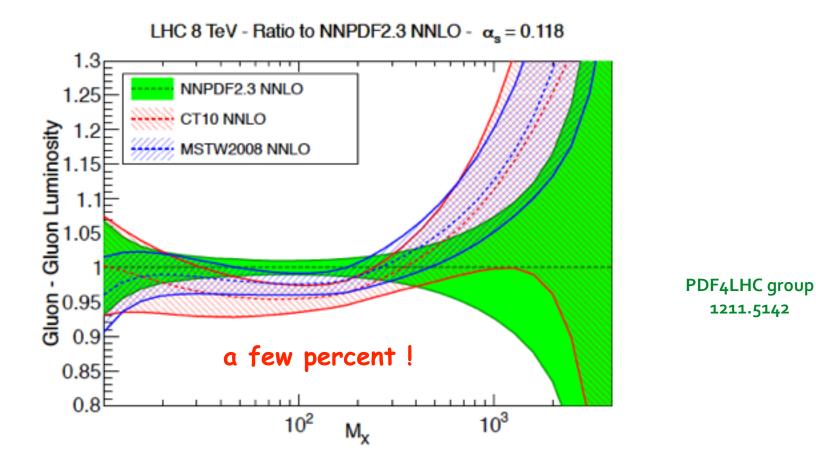


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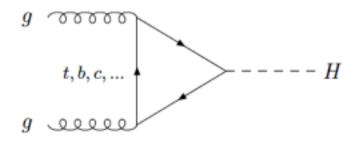




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another culprit is the strong coupling:

 optimum a<sub>s</sub> in global fits varies by about 5% error much larger than for "PDG average"

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current (theory) status for total Higgs cross section:

 $\sigma(\mathbf{M_{H}} = \mathbf{125}\,\mathrm{GeV}) = \mathbf{19.27}^{+7.2\%}_{-7.8\%} \stackrel{\text{PDF \& }\alpha_{s}}{_{-7.8\%}}\mathrm{pb}$ 

de Florian, Grazzini

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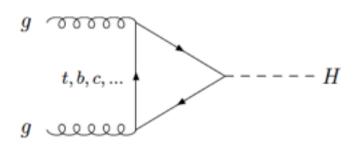
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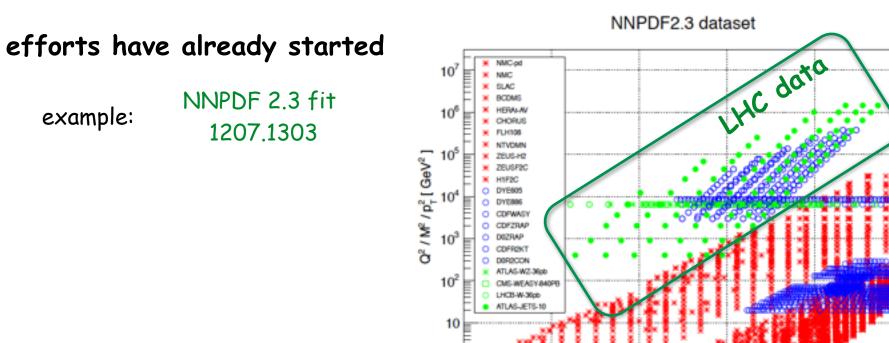
$$\sigma({
m M_{H}}=125\,{
m GeV})=19.27^{+7.2\%}_{-7.8\%}{}^{+7.5\%}_{-6.9\%}{
m pb}$$

de Florian, Grazzini

precise LHC data important for validating and improving PDF and as determinations



## improving PDF's at the LHC



10.5

10<sup>-3</sup>

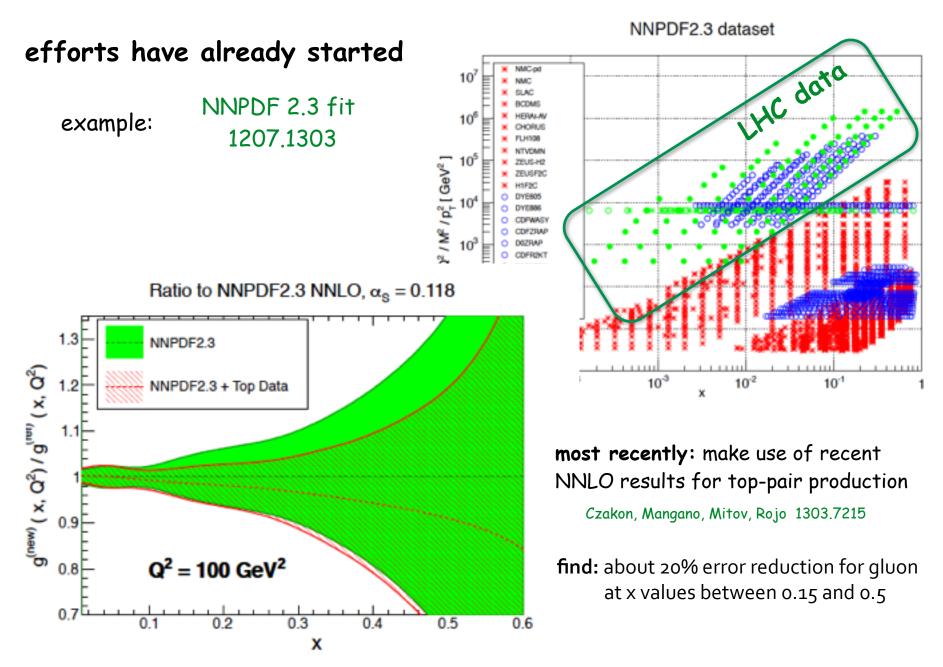
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10-4

10<sup>-2</sup>

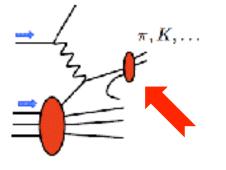
10.1

## improving PDF's at the LHC



## status of fragmentation functions



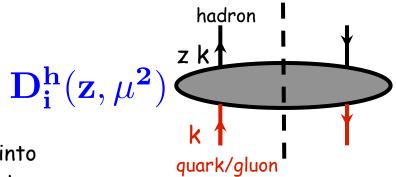


#### recall:

crucial for pQCD interpretation (factorization) of all data with detected (identified) hadrons, e.g., SIDIS (HERMES, COMPASS),  $pp \rightarrow \pi X$  (PHENIX, STAR, ALICE, ...)

### very similar to PDFs:

- non-perturbative but universal
- pQCD predicts scale evolution
- describe the collinear transition of a parton "i" into a massless hadron "h" carrying fractional momentum z



 $\mathbf{D_i^h}(\mathbf{z},\mu)$ 

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no inclusive final-state no local OPE —> no lattice formulation

hadron

guark/gluon

z k1

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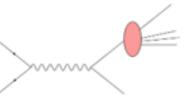
hadron

guark/gluon

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#### also determined from global fits to data:

 key process is e<sup>+</sup>e<sup>-</sup> annihilation to hadrons (plays similar role than DIS for PDFs)



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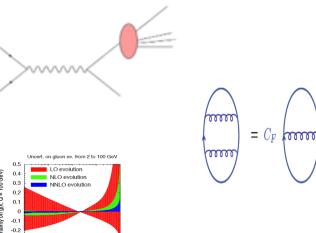
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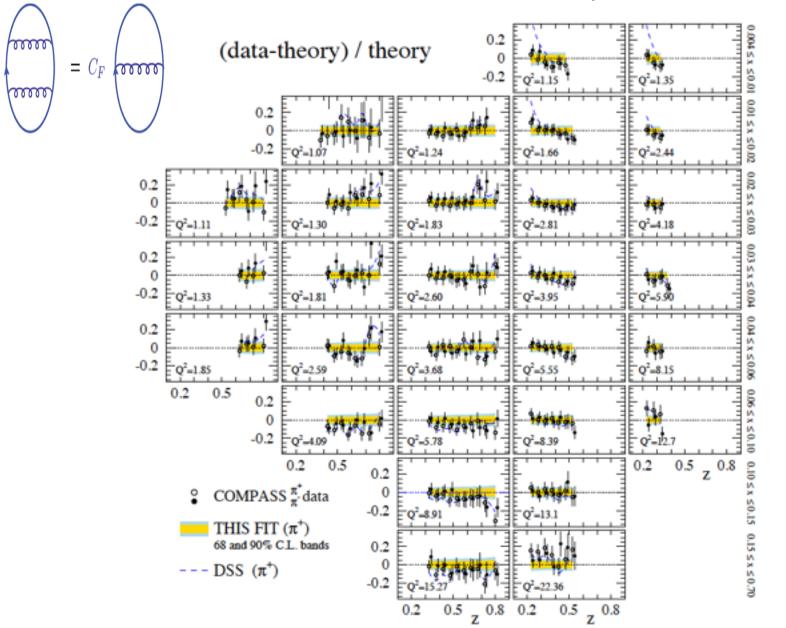
 $D_i^h(z, \mu^2)$ 

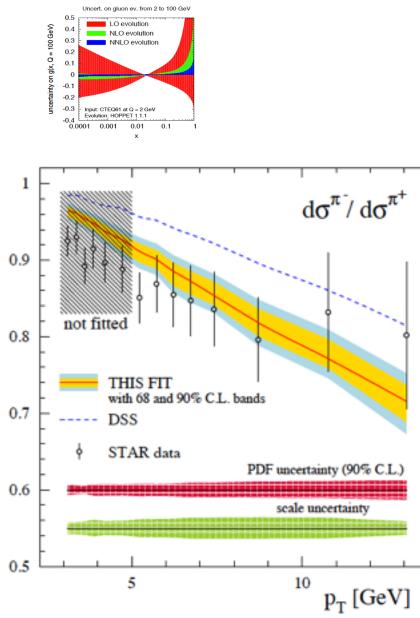
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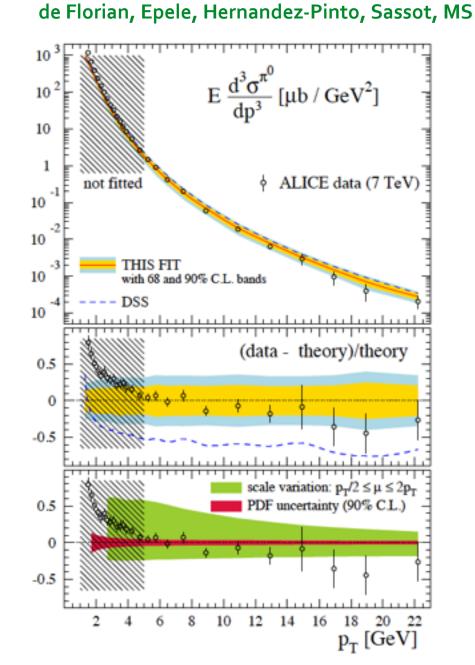
- key process is e<sup>+</sup>e<sup>-</sup> annihilation to hadrons (plays similar role than DIS for PDFs)
- semi-inclusive DIS provides flavor separation
- pp data (RHIC, LHC) important for gluon FF



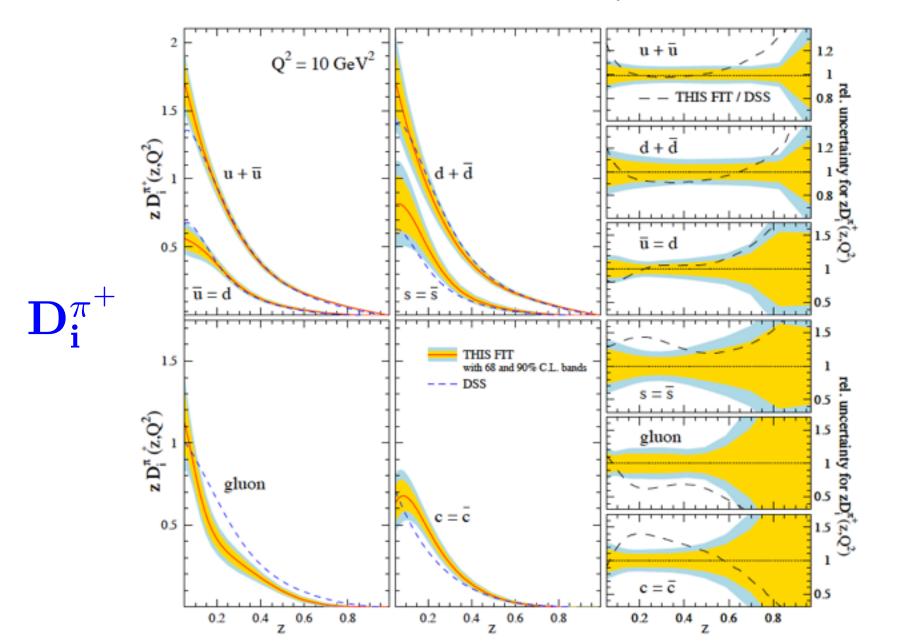
de Florian, Epele, Hernandez-Pinto, Sassot, MS



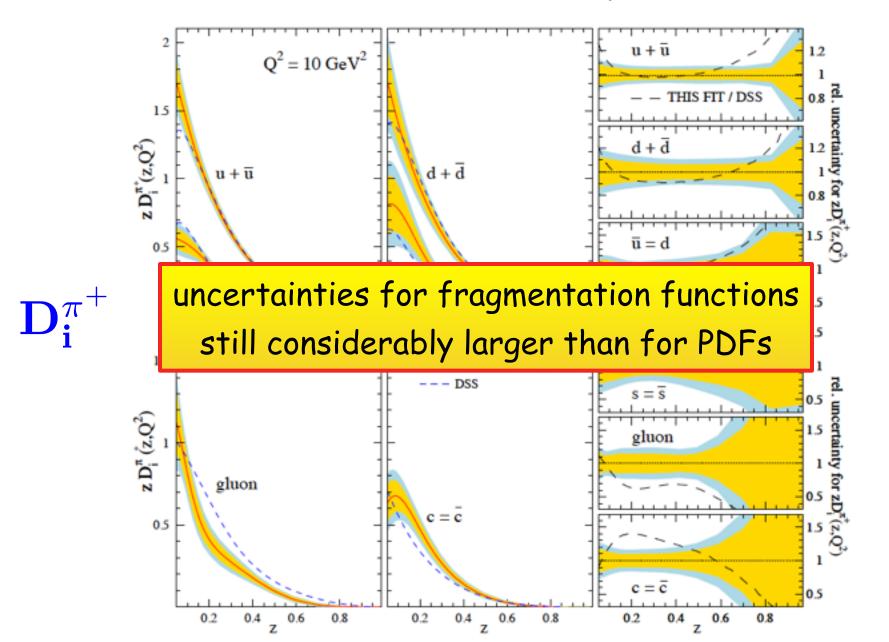


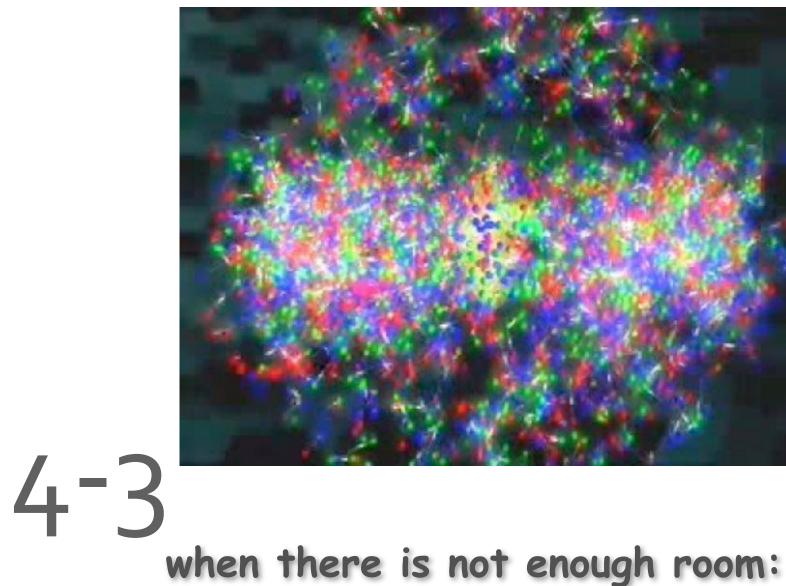


de Florian, Epele, Hernandez-Pinto, Sassot, MS



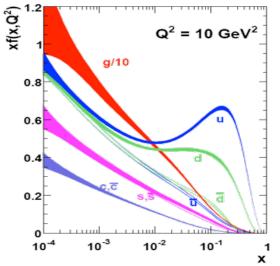
de Florian, Epele, Hernandez-Pinto, Sassot, MS





when there is not enough room: gluons at small x

## what drives the growth of the gluon density

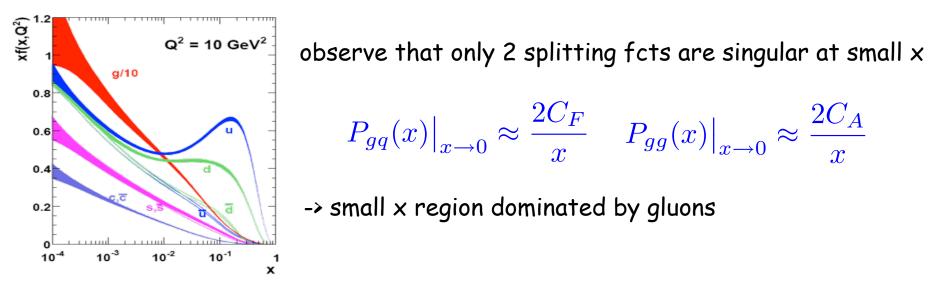


observe that only 2 splitting fcts are singular at small x

$$P_{gq}(x)\big|_{x\to 0} \approx \frac{2C_F}{x} \quad P_{gg}(x)\big|_{x\to 0} \approx \frac{2C_A}{x}$$

-> small x region dominated by gluons

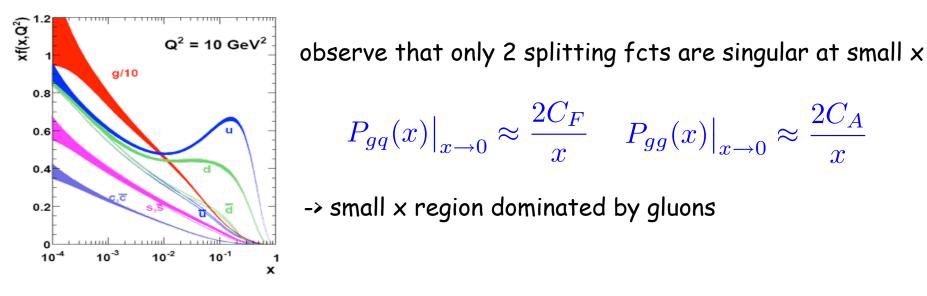
## what drives the growth of the gluon density



• write down "gluon-only" DGLAP equation only valid for small x and large Q<sup>2</sup>

$$\frac{dg(x,\mu^2)}{d\log\mu^2} = \frac{\alpha_s}{2\pi} \int_x^1 \frac{dz}{z} \frac{2C_A}{z} g(x/z,\mu^2)$$

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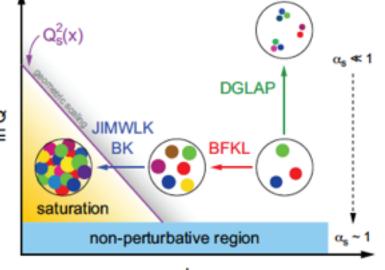
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• for fixed coupling this leads to

"double logarithmic approximation"

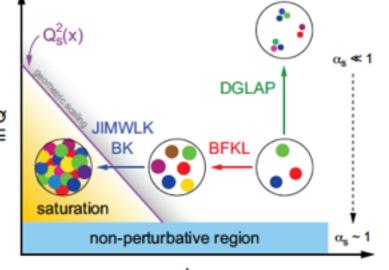
$$xg(x,Q^2) \sim \exp\left(2\sqrt{\frac{\alpha_S C_A}{\pi}\log(1/x)\log(Q^2/Q_0^2)}\right)$$

predicts rise that is faster than  $\log^{a}(1/x)$  but slower than  $(1/x)^{a}$ 



ln x

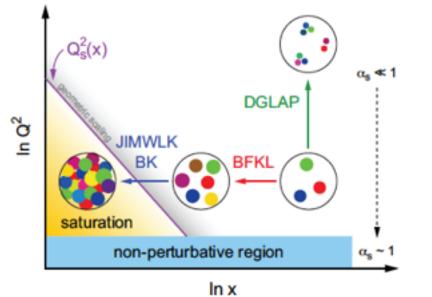
 DGLAP predicts an increase of gluons at small x but proton becomes more dilute as Q<sup>2</sup> increases transverse size of partons ≈ 1/Q



ln x

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but what happens at small x for not so large (fixed)  $Q^2$ ?



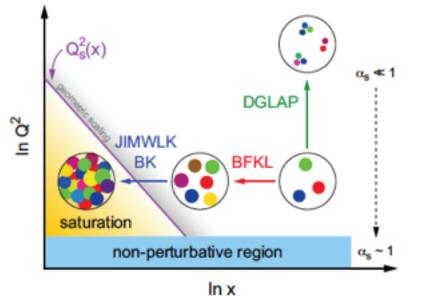
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- aim to resum terms  $\approx \alpha_s \log(1/x)$
- Balitsky-Fadin-Kuraev-Lipatov (BFKL) equation: evolves in x not  $Q^2$
- BFKL predicts a power-like growth  $xg(x,Q^2)\sim (1/x)^{\alpha_P-1}$

much faster than in DGLAP



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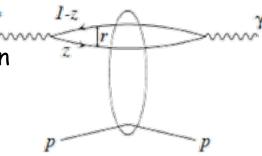
much faster than in DGLAP

### **BIG** problem

- proton quickly fills up with gluons (transverse size now fixed !)
- hadronic cross sections violate In<sup>2</sup>s bound (Froissart-Martin) and grow like a power

make progress by viewing, e.g., DIS from a "different angle"

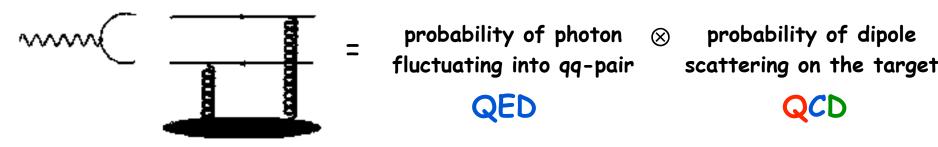
DIS in the **proton rest frame** can be viewed as the photon splitting into a quark-antiquark pair ("**color dipole**") which scatters off the proton (= "slow" gluon field)



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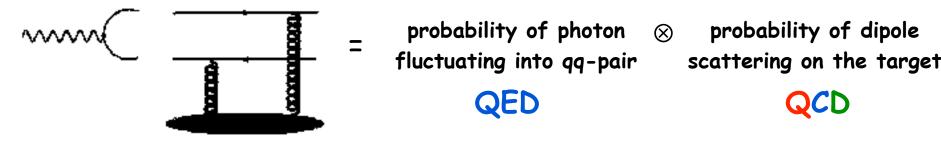
factorization now in terms of



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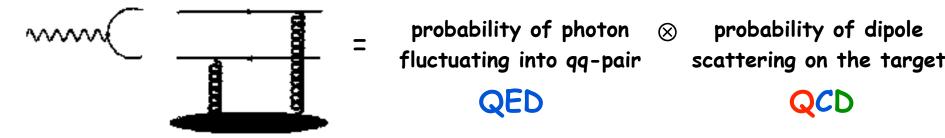


introduces dipole-nucleon scattering amplitude N as fund. building block
energy dependence of N described by Balitsky-Kovchegov equation

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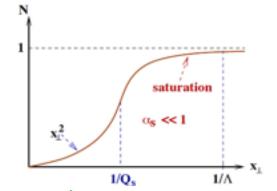
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factorization now in terms of



• introduces dipole-nucleon scattering amplitude N as fund. building block

- energy dependence of N described by Balitsky-Kovchegov equation
  - non-linear -> includes multiple scatterings for unitarization
  - $\ensuremath{\cdot}$  generates saturation scale  $Q_{\ensuremath{\mathsf{s}}}$
  - suited to treat collective phenomena (shadowing, diffration)
  - impact parameter dependence







# when N×LO is not enough: all order resummations

#### when a N<sup>x</sup>LO calculation is not good enough

observation: fixed N×LO order QCD calculations are not necessarily reliable this often happens at low energy fixed-target experiments and can be an issue also at colliders, even the LHC

**reason:** structure of the perturbative series and IR cancellation

at partonic threshold / near exclusive boundary:

- just enough energy to produce, e.g., high- $p_T$  parton
- "inhibited" radiation (general phenomenon for gauge theories)

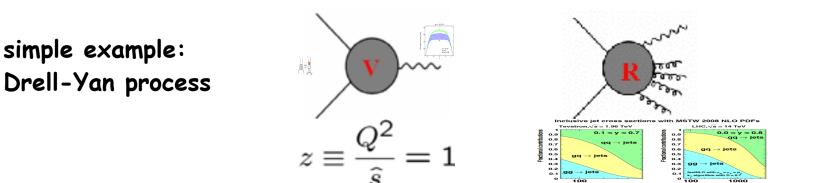
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"imbalance" of real and virtual contributions: IR cancellation leaves large log's

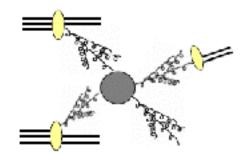
#### all order structure of partonic cross sections

#### let's consider pp scattering:

logarithms related to partonic threshold

Reliable Perturbative Results for Strong Interactions?\*

II. David Politzer lefferson Physical Laboratories, Reveard University, Cambridge, Museuchasetts 62238 (Received 3 May 1953)



general structure of partonic cross sections at the k<sup>th</sup> order:

$$p_T^3 \frac{d\hat{\sigma}_{ab}}{dp_T} = p_T^3 \frac{d\hat{\sigma}_{ab}^{\text{Born}}}{dp_T} \left[ 1 + \underbrace{\mathcal{A}_1 \alpha_s \ln^2 \left(1 - \hat{x}_T^2\right) + \mathcal{B}_1 \alpha_s \ln \left(1 - \hat{x}_T^2\right)}_{\text{NLO}} + \dots + \mathcal{A}_k \alpha_s^k \ln^{2k} \left(1 - \hat{x}_T^2\right) + \dots \right] + \dots \right]$$
``threshold logarithms''

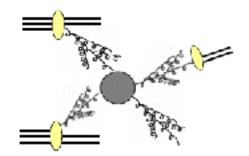
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"threshold logarithms"

where relevant? ... convolution with steeply falling parton luminosity Lab:

$$d\sigma \propto \sum_{a,b} \int_{\tau}^{1} \frac{dz}{z} \mathcal{L}_{ab} \left(\frac{\tau}{z}\right) d\hat{\sigma}_{ab}(z) \qquad z = 1 \text{ emphasized,}$$
in particular as  $\tau \to 1$ 

large at small  $\tau/z$ 

 $\rightarrow$  important for fixed target phenomenology: threshold region more relevant (large  $\tau$ )

#### resummations – how are they done

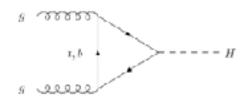
 $\alpha_s^k \ln^{2k}(1-\widehat{x}_T^2)$ 

may spoil perturbative series – unless taken into account to all orders

resummation of such terms has reached a high level of sophistication

Sterman; Catani, Trentadue; Laenen, Oderda, Sterman; Catani et al.; Sterman, Vogelsang; Kidonakis, Owens; ...

- worked out for most processes of interest at least to NLL
- well defined class of higher-order corrections
- often of much phenomenological relevance even for high mass particle production at the LHC



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resummation (= exponentiation) occurs when "right" moments are taken:

Mellin moments for threshold logs

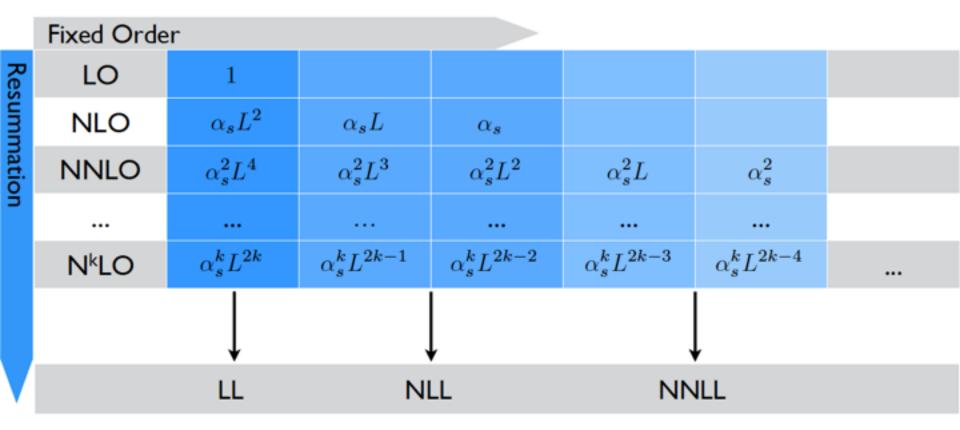


- fixed order calculations needed to determine "coefficients"
- the more orders are known, the more subleading logs can be resummed

#### resummations – terminology

| Fixed Order       |                     |                       |                       |                       |                       |  |
|-------------------|---------------------|-----------------------|-----------------------|-----------------------|-----------------------|--|
| LO                | 1                   |                       |                       |                       |                       |  |
| NLO               | $\alpha_s L^2$      | $\alpha_s L$          | $\alpha_s$            |                       |                       |  |
| NNLO              | $\alpha_s^2 L^4$    | $\alpha_s^2 L^3$      | $\alpha_s^2 L^2$      | $\alpha_s^2 L$        | $\alpha_s^2$          |  |
|                   |                     |                       |                       |                       |                       |  |
| N <sup>k</sup> LO | $\alpha_s^k L^{2k}$ | $\alpha_s^k L^{2k-1}$ | $\alpha_s^k L^{2k-2}$ | $\alpha_s^k L^{2k-3}$ | $\alpha_s^k L^{2k-4}$ |  |

#### resummations – terminology



#### some leading log exponents

(assuming fixed  $\alpha_s$  for simplicity)

color factors for soft gluon radiation matter:

unobserved parton Sudakov "suppression"

$$\exp\left[\frac{\mathbf{C_F}\,\alpha_{\mathbf{s}}}{\pi}\,\ln^2(\mathbf{N})\,-\,\frac{\mathbf{C_F}\,\alpha_{\mathbf{s}}}{\pi}\,\frac{\mathbf{1}}{\mathbf{2}}\,\ln^2(\mathbf{N})\,\right]$$



DIS

#### some leading log exponents

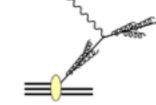
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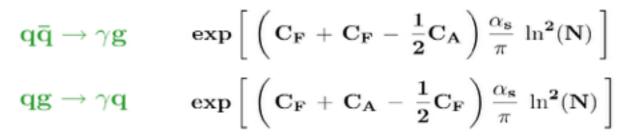
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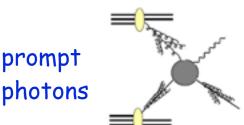
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moderate enhancement, unless  $x_{Bj}$  large



DIS





exponents positive — enhancement

## some leading log exponents

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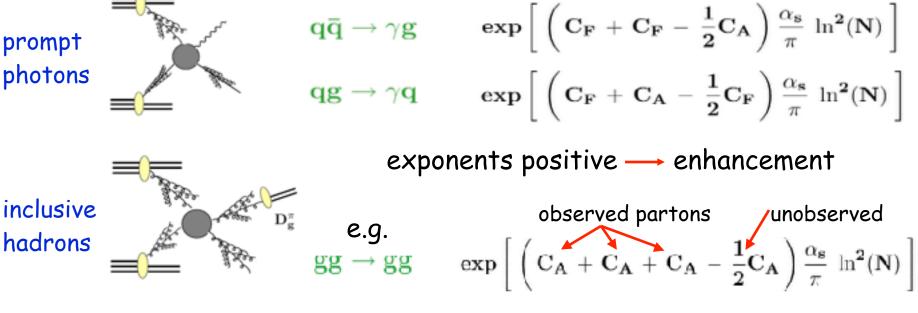
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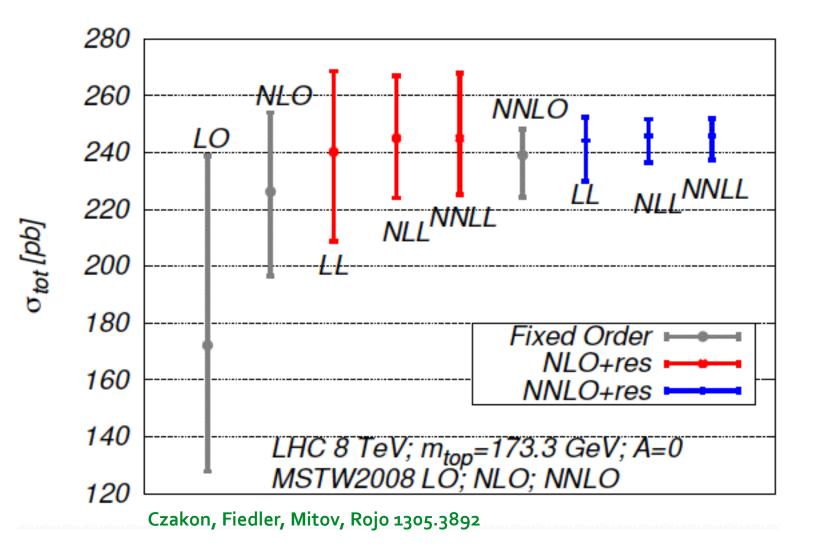
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moderate enhancement, unless  $x_{Bj}$  large



expect much larger enhancement

#### one recent example: top-pair production



#### resummations: window to non-perturbative regime

#### important technical issue:

resummations are sensitive to strong coupling regime

 $\rightarrow$  need some "minimal prescription" to avoid Landau pole (where  $\alpha_s \rightarrow \infty$ ) Catani, Mangano, Nason, Trentadue:

define resummed result such that series is asymptotic w/o factorial growth associated with power corrections [achieved by particular choice of Mellin contour]

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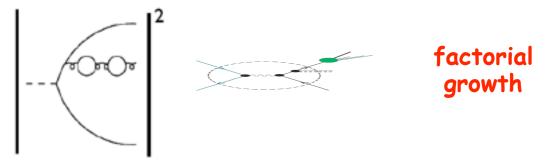
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window to the non-perturbative regime so far little explored

### "convergence" of an asymptotic series

see, "Renormalons" review by M. Beneke, hep-ph/9807443

suppose we keep calculating higher and higher orders

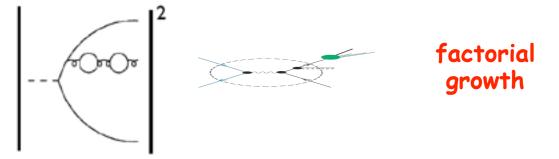


→ **big trouble**: the perturbative series is not convergent but only asymptotic

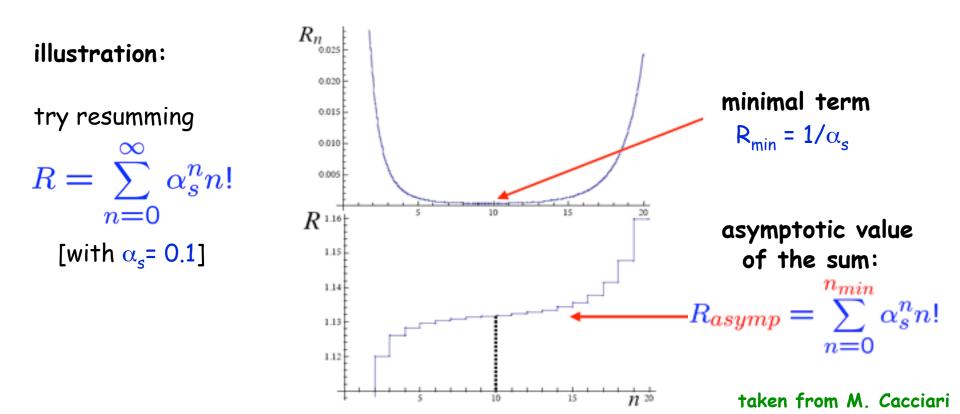
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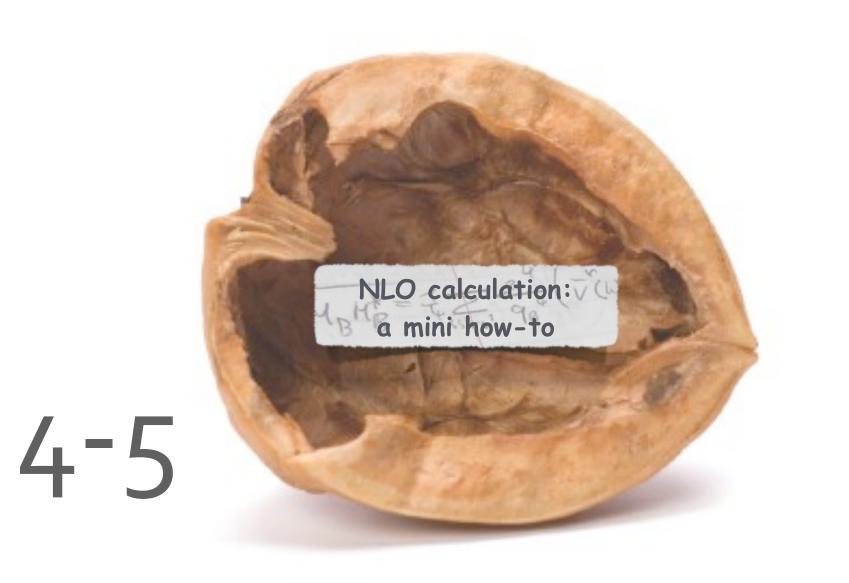
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QCD: NP corrections are power suppressed:

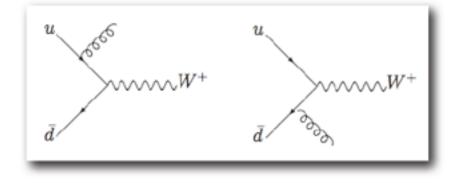


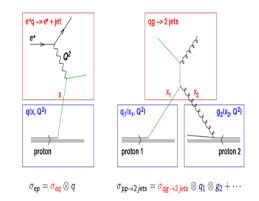
the value of p depends on the process and can sometimes be predicted



#### NLO calculation in a nutshell: Drell-Yan

at NLO we need to compute two contributions:





#### real radiation corrections

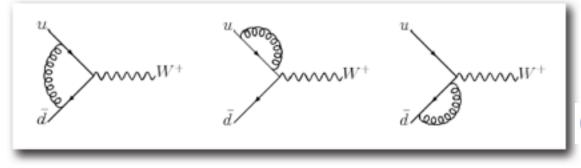
one extra parton in final-state

$$|\mathcal{M}_{W+g}|^2 \sim (g_s)^2$$



only interference with Born contributes at NLO

 $(\mathcal{M}_{W,1-\text{loop}} \times \mathcal{M}_{W,\text{tree}}) \sim g_s^2 \times 1$ 



#### NLO in a nutshell: real radiation

recall: collinear/soft kinematics

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$$d\sigma_{(...)ac} \sim \int |\mathcal{M}_{(...)ac}|^2 E_a^2 dE_a \,\theta_a d\theta_a \sim d\sigma_{(...)b} \left(\frac{\alpha_s}{2\pi}\right) \frac{dt}{t} P_{ab}(z) \, dz$$

since we cannot calculate with infinities we need to regularize them: this time we choose dimensional regularization (i.e. work in d=4-2 e dimensions)

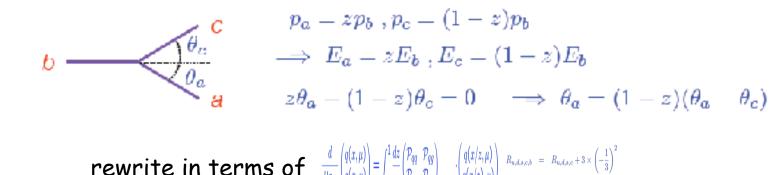
$$E_a^2 dE_a \,\theta_a d\theta_a \to E_a^{2-2\epsilon} dE_a \,\theta_a^{1-2\epsilon} d\theta_a = E_a^2 dE_a \,\theta_a d\theta_a \, z^{-\epsilon} (1-z)^{-\epsilon} \, t^{-\epsilon}$$

and obtain

$$d\sigma_{(\ldots)ac}^{4-2\epsilon} = d\sigma_{(\ldots)b} \left(\frac{\alpha_s}{2\pi}\right) \frac{dt}{t^{1+\epsilon}} P_{ab}(z) z^{-\epsilon} (1-z)^{-\epsilon} dz$$

#### NLO in a nutshell: real radiation

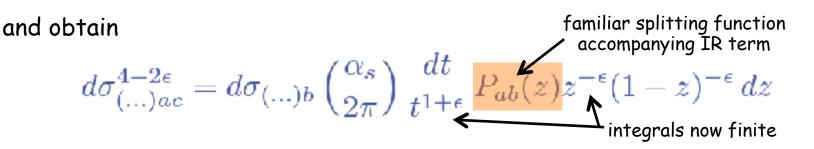
recall: collinear/soft kinematics



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#### NLO in a nutshell: poles

we can now see how the singularities are regularized in d dimensions

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• collinear pole  $\int \frac{dt}{t^{1+\epsilon}} \to \frac{1}{\epsilon}$  recall: such a factor is present in P<sub>qq</sub> (and P<sub>gg</sub>)  
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putting all together one obtains the following (general) structure

$$d\sigma_{W+g} = \left(\frac{2}{\epsilon^2} + \frac{3}{\epsilon} - \frac{2}{\epsilon}P_{qq} + \mathcal{O}(\epsilon^0)\right)d\sigma_{W,\text{tree}} + \text{finite terms}$$

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#### NLO in a nutshell: virtual corrections

only one loop diagram to consider at NLO (selfenergy on massless external lines zero in d dimensions)

obtain for amplitude:

$$\int \frac{d^{4-2\epsilon}\ell}{\ell^2(\ell+p_{\bar{d}})^2(\ell+p_{\bar{d}}+p_u)^2}$$



with some complicated Dirac structure in numerator

 $\mathcal{N} = \left[\bar{u}(p_{\bar{d}})\gamma^{\alpha} \not \!\!\!/ \gamma^{\mu} (\not \!\!\!/ + \not \!\!\!/ _{\bar{d}} + \not \!\!\!/ _{u})\gamma_{\alpha} u(p_{u})\right] V_{\mu}(p_{W})$ 

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inspect denominator:

can shift momenta

$$\ell^2(\ell+p_{\bar{d}})^2(\ell+p_{\bar{d}}+p_u)^2 \longrightarrow \ell^2(\ell-p_{\bar{d}})^2(\ell+p_u)^2$$

- soft singularity for I -> 0
- singularities for I collinear to quark lines

regularize again in d dimensions

can decompose Dirac structure into given set of simpler scalar integrals (Passarino Veltman decomposition)

then:  $\frac{1}{\ell^2 (\ell + p_{\bar{d}})^2 (\ell + p_{\bar{d}} + p_u)^2}$ 

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$$\frac{1}{\ell^2(\ell+p_{\bar{d}})^2(\ell+p_{\bar{d}}+p_u)^2}$$

$$= 2\int_0^1 dx_1 \int_0^1 dx_2 \int_0^1 dx_3 \frac{\delta(x_1+x_2+x_3-1)}{\left[x_1\ell^2+x_2(\ell+p_{\bar{d}})^2+x_3(\ell+p_{\bar{d}}+p_u)^2\right]^3}$$

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Then:  

$$\frac{1}{\ell^{2}(\ell+p_{\bar{d}})^{2}(\ell+p_{\bar{d}}+p_{u})^{2}}$$
need to combine different terms in denominator with help of Feynman parameter integrals  

$$= 2 \int_{0}^{1} dx_{1} \int_{0}^{1} dx_{2} \int_{0}^{1} dx_{3} \frac{\delta(x_{1}+x_{2}+x_{3}-1)}{[x_{1}\ell^{2}+x_{2}(\ell+p_{\bar{d}})^{2}+x_{3}(\ell+p_{\bar{d}}+p_{u})^{2}]^{3}}$$

$$F_{2}^{n}(x) = x \left(\frac{1}{9}d_{n}(x) + \frac{4}{9}u_{n}(x)\right) = x \left(\frac{4}{9}d_{p}(x) + \frac{1}{9}u_{p}(x)\right) \quad \text{where} \quad \frac{d^{2}\bar{a}}{dxd^{2}}|_{F_{2}} = \frac{\ell^{2}}{c_{q}^{2}} \left[\delta(1-x) + \frac{\alpha_{s}(\mu)}{4\pi} \left[P_{M}(x)\ln\frac{Q^{2}}{m_{s}^{2}} + C_{s}^{q}(x)\right]\right]$$
this can be evaluated using: 
$$\int \frac{d^{d}L}{(2\pi)^{d}} \frac{1}{(L^{2}-\Delta)^{n}} = i \frac{(-1)^{n}}{(4\pi)^{d/2}} \frac{\Gamma\left(n-\frac{d}{2}\right)}{\Gamma(n)} \Delta^{d/2-n}$$

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$$\frac{1}{\ell^{2}(\ell+p_{\bar{d}})^{2}(\ell+p_{\bar{d}}+p_{u})^{2}} \qquad \text{need to combine different terms in denominator with help of Feynman parameter integrals} \\ = 2\int_{0}^{1} dx_{1} \int_{0}^{1} dx_{2} \int_{0}^{1} dx_{3} \frac{\delta(x_{1}+x_{2}+x_{3}-1)}{[x_{1}\ell^{2}+x_{2}(\ell+p_{\bar{d}})^{2}+x_{3}(\ell+p_{\bar{d}}+p_{u})^{2}]^{3}} \\ F_{2}^{n}(x) = x\left(\frac{1}{9}d_{n}(x)+\frac{4}{9}u_{n}(x)\right) = x\left(\frac{4}{9}d_{p}(x)+\frac{1}{9}u_{p}(x)\right) \qquad \text{where } \begin{array}{c} \frac{d^{2}\hat{r}}{dxdQ^{2}|_{F_{2}}} = \tilde{r}_{2}^{2} \\ = c_{q}^{2}x\left[\delta(1-x)+\frac{\alpha_{s}(\mu)}{4\pi}\left[P_{qq}(x)\ln\frac{Q^{2}}{m_{q}^{2}}+C_{2}^{q}(x)\right] \end{array}$$

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and one obtains:

#### NLO in a nutshell: final result

once all scalar integrals are computed and put together, find:

$$d\sigma_{W,1-\text{loop}} = \left(-\frac{2}{\epsilon^2} - \frac{3}{\epsilon} + \text{finite}\right) d\sigma_{W,\text{tree}}$$

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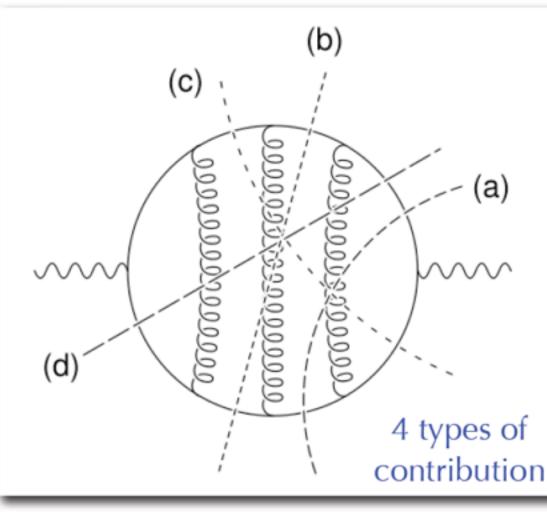
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and one ends up with the finite NLO result (where d->4)

this is one of the simplest loop calculations ! in general it is much more complicated but the general ideas are the same for high multiplicity final states one needs novel methods "beyond" Feynman diagrams

# **NNLO complexity**

one can envision the contributions to a NNLO calculation by considering all possible cuts to a 3-loop diagram:



#### example:

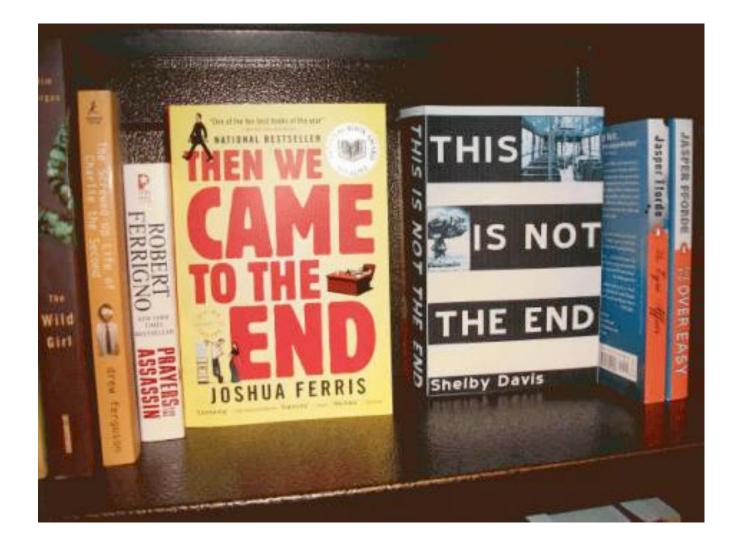
```
3 jet production in e^+e^-
```

(a) two-loop virtual correction

(b) one-loop x one-loop

(c) one-loop x real both with an extra parton
(d) real

with two extra partons



# SUMMARY & OUTLOOK

# QCD: the most perfect gauge theory (so far)

simple  $\mathcal{L}$  but rich & complex phenomenology; few parameters

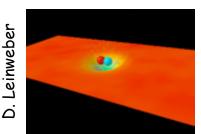
in principle complete up to the Planck scale (issue: CP, axions?)

highly non-trivial ground state responsible for all the structure in the visible universe

emergent phenomena: confinement, chiral symmetry breaking, hadrons



#### confinement

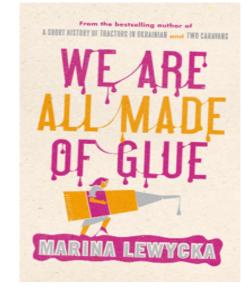


non-perturbative structure of hadrons

e.g. through lattice QCD



interplay between High Energy and Hadron Physics



#### asymptotic freedom

hard scattering cross sections and renormalization group

perturbative methods

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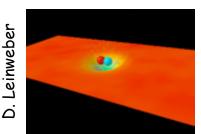
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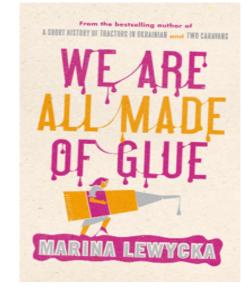


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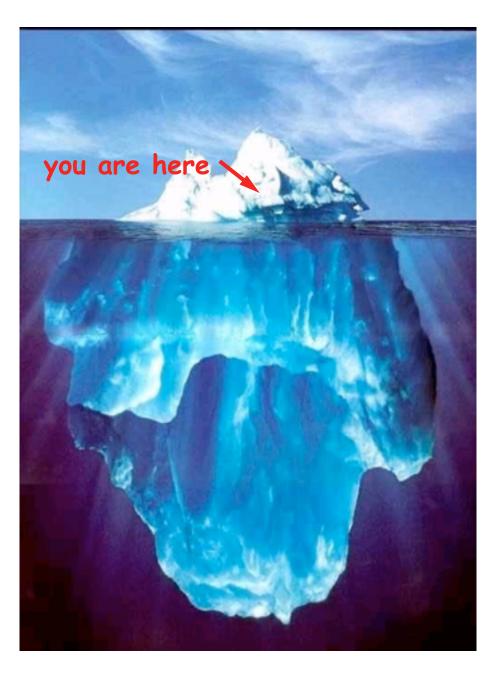
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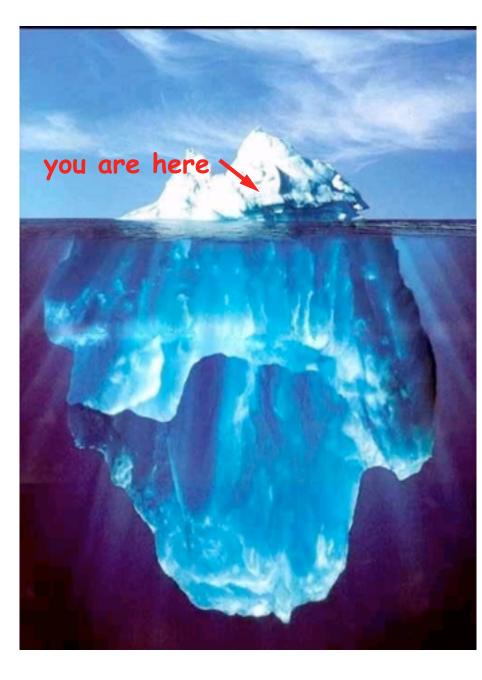
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# we have just explored the tip of the iceberg



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enjoy the other lectures !