

Second International Summer School of the GDR PH-QCD

"Correlations between partons in nucleons"



Multidimensional pictures of the nucleon (1/3)

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Outline

Lecture 1

- $\boldsymbol{\cdot} \textbf{ Introduction}$
- $\boldsymbol{\cdot}$ Tour in phase space
- Galileo vs Lorentz
- $\boldsymbol{\cdot}$ Photon point of view

Structure of matter



Elementary particles





Graviton ?

Degrees of freedom



« Resolution »

Relevant degrees of freedom depend on typical energy scale



Nucleon pictures





3 non-relativistic heavy quarks







Quantum Mechanics + Special Relativity = Quantum Field Theory

Realistic



Indefinite # of relativistic light quarks and gluons

Goal : understanding the nucleon internal structure

Why?



At the energy frontier



Goal : understanding the nucleon internal structure



Goal : understanding the nucleon internal structure





Classical Mechanics



Statistical Mechanics



Position-space density

$$\rho(r) = \int \mathrm{d}p \,\rho(p,r)$$

Momentum-space density

$$\rho(p) = \int \mathrm{d}r \, \rho(p,r)$$

Phase-space average $\overline{O} = \int \mathrm{d}p \, \mathrm{d}r \, O(p,r) \, \rho(p,r)$

[Gibbs (1902)]

Quantum Mechanics



Position-space density $|\psi(r)|^2 = \int \mathrm{d}p \, \rho_W(p,r)$

Momentum-space density $\frac{1}{2\pi} |\varphi(p)|^2 = \int \mathrm{d}r \,\rho_W(p,r)$

Phase-space average
$$\begin{split} \langle \widehat{O} \rangle &= \int \mathrm{d}r \, \psi^*(r) \, \widehat{O}(-i\frac{\partial}{\partial r}, r) \, \psi(r) \\ &= \int \frac{\mathrm{d}p}{2\pi} \, \varphi^*(p) \, \widehat{O}(p, i\frac{\partial}{\partial p}) \, \varphi(p) \\ &= \int \mathrm{d}p \, \mathrm{d}r \, O(p, r) \, \rho_W(p, r) \end{split}$$

> [Wigner (1932)] [Moyal (1949)]

Quantum Mechanics



$$\psi(r) = \int \frac{\mathrm{d}p}{2\pi} e^{ipr} \varphi(p)$$

$$\rho_W(p,r) = \int \frac{\mathrm{d}z}{2\pi} e^{-ipz} \psi^* \left(r - \frac{z}{2}\right) \psi\left(r + \frac{z}{2}\right)$$

$$= \int \frac{\mathrm{d}\Delta}{(2\pi)^2} e^{-i\Delta r} \varphi^* \left(p + \frac{\Delta}{2}\right) \varphi\left(p - \frac{\Delta}{2}\right)$$

$$p \rho_W(p,r) = \int \frac{\mathrm{d}z}{2\pi} p \, e^{-ipz} \, \psi^*(r - \frac{z}{2}) \, \psi(r + \frac{z}{2})$$

$$= \int \frac{\mathrm{d}z}{2\pi} \left[\left(i \frac{\partial}{\partial z} \right) e^{-ipz} \right] \psi^*(r - \frac{z}{2}) \, \psi(r + \frac{z}{2})$$

$$= \int \frac{\mathrm{d}z}{2\pi} \, e^{-ipz} \left(-i \frac{\partial}{\partial z} \right) \left[\psi^*(r - \frac{z}{2}) \, \psi(r + \frac{z}{2}) \right]$$

$$= \int \frac{\mathrm{d}z}{2\pi} \, e^{-ipz} \, \psi^*(r - \frac{z}{2}) \left(-\frac{i}{2} \frac{\overleftarrow{\partial}}{\partial r} \right) \psi(r + \frac{z}{2})$$

$$\overset{\text{Hermitian}}{\text{(symmetric)}} \equiv -\frac{i}{2} \left(\frac{\overrightarrow{\partial}}{\partial r} - \frac{\overleftarrow{\partial}}{\partial r} \right)$$

[Wigner (1932)] [Moyal (1949)]

Wigner distributions have applications in:



- Quantum chemistry
- Quantum molecular dynamics
- Quantum information
- Quantum optics
- Classical optics
- Signal analysis
- Image processing
- Quark-gluon plasma

• ...



Heisenberg's uncertainty relations

 $\Delta p \Delta r \ge \frac{\hbar}{2}$

Harmonic oscillator



In quantum optics, Wigner distributions are « measured » using homodyne tomography



[Lvovski *et al.* (2001)] [Bimbard *et al.* (2014)]



Idea : measuring projections of Wigner distributions from different directions



Binocular vision in phase space !

Find the hidden 3D picture



Quantum Field Theory

Covariant Wigner operator

$$\widehat{W}(k,r) = \int \frac{\mathrm{d}^4 z}{(2\pi)^4} \, e^{ik \cdot z} \, \phi^*(r - \frac{z}{2}) \, \phi(r + \frac{z}{2}) \qquad \text{Time ordering ?}$$
Scalar fields

Equal-time Wigner operator

$$\begin{split} \widehat{W}(\vec{k},\vec{r},t) &= \int \mathrm{d}k^0 \, \widehat{W}(k,r) \\ &= \int \frac{\mathrm{d}^3 z}{(2\pi)^3} \, e^{-i\vec{k}\cdot\vec{z}} \, \phi^*(\vec{r}-\frac{\vec{z}}{2},t) \, \phi(\vec{r}+\frac{\vec{z}}{2},t) \end{split}$$

$$\int \frac{\mathrm{d}k^0}{2\pi} \, e^{ik^0 z^0} = \delta(z^0)$$

Phase-space/Wigner distribution

$$\rho_W(\vec{k},\vec{r},t;\Psi) = \langle \Psi | \widehat{W}(\vec{k},\vec{r},t) | \Psi \rangle$$

[Carruthers, Zachariasen (1976)] [Ochs, Heinz (1997)]





Phase-space compromise

$$\rho_W(\vec{k}, \vec{r}, t; \vec{P}, \vec{R}) = \int \frac{\mathrm{d}^3 \Delta}{(2\pi)^3} e^{i\vec{\Delta}\cdot\vec{R}} \langle \vec{P} + \frac{\vec{\Delta}}{2} |\widehat{W}(\vec{k}, \vec{r}, t)| \vec{P} - \frac{\vec{\Delta}}{2} \rangle$$
$$= \int \mathrm{d}^3 Z \, e^{i\vec{P}\cdot\vec{Z}} \, \langle \vec{R} - \frac{\vec{Z}}{2} |\widehat{W}(\vec{k}, \vec{r}, t)| \vec{R} + \frac{\vec{Z}}{2} \rangle$$

Intrinsic phase-space/Wigner distribution

$$\rho_W(\vec{k}, \vec{r}, t; \vec{0}, \vec{0}) = \int \frac{\mathrm{d}^3 \Delta}{(2\pi)^3} \left\langle \frac{\vec{\Delta}}{2} | \widehat{W}(\vec{k}, \vec{r}, t) | - \frac{\vec{\Delta}}{2} \right\rangle$$



Identified with intrinsic variables

> [Ji (2003)] [Belitsky, Ji, Yuan (2004)]

Localized state in momentum space $|\Psi\rangle = |\vec{P}\rangle$ Commomentum
momentum $\vec{P} = \sum_{i} \vec{k}_{i}$ in position space $|\Psi\rangle = |\vec{R}\rangle$ $\sum_{i} m_{i} \vec{r}_{i}$ $= \int \frac{\mathrm{d}^{3}P}{(2\pi)^{3}} e^{-i\vec{P}\cdot\vec{R}} |\vec{P}\rangle$ $\vec{R} = \frac{\sum_{i} m_{i}\vec{r}_{i}}{\sum_{i} m_{i}}$

Phase-space compromise

$$\rho_W(\vec{k}, \vec{r}, t; \vec{P}, \vec{R}) = \int \frac{\mathrm{d}^3 \Delta}{(2\pi)^3} e^{i\vec{\Delta}\cdot\vec{R}} \langle \vec{P} + \frac{\vec{\Delta}}{2} |\widehat{W}(\vec{k}, \vec{r}, t)| \vec{P} - \frac{\vec{\Delta}}{2} \rangle$$
$$= \int \mathrm{d}^3 Z \, e^{i\vec{P}\cdot\vec{Z}} \langle \vec{R} - \frac{\vec{Z}}{2} |\widehat{W}(\vec{k}, \vec{r}, t)| \vec{R} + \frac{\vec{Z}}{2} \rangle$$

Intrinsic phase-space/Wigner distribution



Time translation $\widehat{O}(t) = e^{i\widehat{H}t}\,\widehat{O}(0)\,e^{-i\widehat{H}t}$

[Ji (2003)] [Belitsky, Ji, Yuan (2004)]

Galilean symmetry

All is fine as long as space-time symmetry is Galilean

$$\begin{split} [J^{i}, J^{j}] &= i\epsilon^{ijk}J^{k} & [B^{i}, H] = -iP^{i} & E \mapsto E + \vec{P} \cdot \delta \vec{v} \\ [J^{i}, B^{j}] &= i\epsilon^{ijk}B^{k} & [B^{i}, P^{j}] = -i\delta^{ij}M & \vec{P} \mapsto \vec{P} + M\delta \vec{v} \\ [B^{i}, B^{j}] &= 0 & [B^{i}, M] = [J^{i}, M] = 0 \end{split}$$

Position operator can be defined

Lorentz symmetry

But in relativity, space-time symmetry is Lorentzian

$$[J^{i}, J^{j}] = i\epsilon^{ijk}J^{k}$$
$$[J^{i}, K^{j}] = i\epsilon^{ijk}K^{k}$$
$$[K^{i}, K^{j}] = -i\epsilon^{ijk}J^{k}$$

$$J^{i} = \frac{1}{2} \epsilon^{ijk} M^{jk}$$
$$K^{i} = M^{0i}$$

$$[K^{i}, P^{0}] = -iP^{i} \qquad E \mapsto E + \vec{P} \cdot \delta \vec{v}$$
$$[K^{i}, P^{j}] = -i\delta^{ij}P^{0} \qquad \vec{P} \mapsto \vec{P} + E\delta \vec{v}$$

Position operator is ill-defined !



No separation of CoM and internal coordinates

 $[K^i, P^0] = -iP^i$



Further issues :

Lorentz contraction



Creation/annihilation of pairs





Spoils (quasi-) probabilistic interpretation

Forms of dynamics

Space-time foliation

Light-front components



Instant-form dynamics



Light-front form dynamics



[Dirac (1949)]

Ordinary point of view



Photon point of view

$$x^+ = \frac{1}{\sqrt{2}}(t+z)$$



Initial frame



Boosted frame



(Quasi) infinite-momentum frame



Light-front operators

$$\begin{split} K^{1}_{\perp} &= M^{+1} & J^{1}_{\perp} &= M^{-1} \\ &= \frac{1}{\sqrt{2}}(K^{1} + J^{2}) &= \frac{1}{\sqrt{2}}(K^{1} - J^{2}) \\ K^{2}_{\perp} &= M^{+2} & J^{2}_{\perp} &= M^{-2} \\ &= \frac{1}{\sqrt{2}}(K^{2} - J^{1}) &= \frac{1}{\sqrt{2}}(K^{2} + J^{1}) \end{split}$$

Transverse space-time symmetry is Galilean

$$\begin{split} [J^3, J^i_{\perp}] &= i\epsilon^{3ij}J^j_{\perp} \\ [J^3, K^i_{\perp}] &= i\epsilon^{3ij}K^j_{\perp} \\ [K^i_{\perp}, K^j_{\perp}] &= 0 \end{split}$$

$$[K_{\perp}^{i}, P^{-}] = -iP_{\perp}^{i}$$
$$[K_{\perp}^{i}, P_{\perp}^{j}] = -i\delta_{\perp}^{ij}P^{+}$$
$$[K_{\perp}^{i}, P^{+}] = [J^{3}, P^{+}] = 0$$

Transverse position operator can be defined !

Longitudinal momentum plays the role of mass in the transverse plane $P^+ \sim M$

Transverse boost $p'^+ = p^+$ $\vec{p}'_\perp = \vec{p}_\perp + p^+ \vec{v}_\perp$

[Kogut, Soper (1970)]

Quasi-probabilistic interpretation

What about the further issues with Special Relativity ?





Non-relativistic phase space

Localized state in momentum space
$$|\Psi\rangle = |\vec{P}\rangle$$
Commomentum $\vec{P} = \sum_{i} \vec{k}_{i}$ in position space $|\Psi\rangle = |\vec{R}\rangle$ $\mathbf{Commomentum}$ $\vec{R} = \frac{\sum_{i} m_{i} \vec{r}_{i}}{\sum_{i} m_{i}}$ $= \int \frac{\mathrm{d}^{3}P}{(2\pi)^{3}} e^{-i\vec{P}\cdot\vec{R}} |\vec{P}\rangle$ $\vec{R} = \frac{\sum_{i} m_{i} \vec{r}_{i}}{\sum_{i} m_{i}}$

Equal-time Wigner operator

$$\widehat{W}(\vec{k},\vec{r},t) = \int \mathrm{d}k^0 \,\widehat{W}(k,r)$$

Intrinsic phase-space/Wigner distribution

$$\rho_W(\vec{k}, \vec{r}, \vec{k}, \vec{0}, \vec{0}) = \int \frac{\mathrm{d}^3 \Delta}{(2\pi)^3} \left\langle \frac{\vec{\Delta}}{2} | \widehat{W}(\vec{k}, \vec{r}, t) | - \frac{\vec{\Delta}}{2} \right\rangle$$

Identified with intrinsic variables

Time translation $\widehat{O}(t) = e^{i\widehat{H}t}\,\widehat{O}(0)\,e^{-i\widehat{H}t}$

[Ji (2003)] [Belitsky, Ji, Yuan (2004)]

Relativistic phase space

Localized state in momentum space
$$|\Psi\rangle = |P^+, \vec{P}_{\perp}\rangle$$

in position space $|\Psi\rangle = |P^+, \vec{R}_{\perp}\rangle$
 $= \int \frac{d^2 P_{\perp}}{(2\pi)^2} e^{-i\vec{P}_{\perp} \cdot \vec{R}_{\perp}} |P^+, \vec{P}_{\perp}\rangle$
[Soper (1977)]

[Burkardt (2000)] [Burkardt (2003)]

Equal light-front time Wigner operator

$$\widehat{W}(k^{+}, \vec{k}_{\perp}, r^{-}, \vec{r}_{\perp}, r^{+}) = \int dk^{-} \widehat{W}(k, r)$$
[Ji (2003)]
[Belitsky, Ji, Yuan (2004)]

Intrinsic relativistic phase-space/Wigner distribution

$$\rho_{W}(x,\vec{k}_{\perp},\vec{b}_{\perp}) = \frac{P^{+}}{\langle P^{+}|P^{+}\rangle} \int \mathrm{d}r^{-} \frac{\mathrm{d}^{2}\Delta_{\perp}}{(2\pi)^{2}} \langle P^{+},\frac{\vec{\Delta}_{\perp}}{2}|\widehat{W}(xP^{+},\vec{k}_{\perp},r^{-},\vec{b}_{\perp},r^{+})|P^{+},-\frac{\vec{\Delta}_{\perp}}{2}\rangle$$

$$x = \frac{k^{+}}{P^{+}} = \frac{1}{2} \int \frac{\mathrm{d}^{2}\Delta_{\perp}}{(2\pi)^{2}} \langle P^{+},\frac{\vec{\Delta}_{\perp}}{2}|\widehat{W}(xP^{+},\vec{k}_{\perp},0,\vec{b}_{\perp},0)|P^{+},-\frac{\vec{\Delta}_{\perp}}{2}\rangle$$
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Boost invariant !

$$\int \mathrm{d}k^+ f(k^+) = \int \mathrm{d}x \, P^+ f(xP^+)$$

Space-time translation

 $\widehat{O}(r) = e^{i\widehat{P}\cdot r}\,\widehat{O}(0)\,e^{-i\widehat{P}\cdot r}$ Normalization $\langle P^+|P^+\rangle = 2P^+ 2\pi \,\delta(0)$

[C.L., Pasquini (2011)]

Relativistic phase space

Our intuition is instant form and not light-front form

<u>NB</u>: $\rho_W(x, \vec{k}_{\perp}, \vec{b}_{\perp})$ is invariant under light-front boosts

It can be thought as instant form phase-space/Wigner distribution in IMF !



[C.L., Pasquini (2011)]

Link with parton correlators

Phase-space/Wigner distributions are Fourier transforms

General parton correlator

$$W(x,\xi,\vec{k}_{\perp},\vec{\Delta}_{\perp}) = \frac{1}{2} \int \frac{\mathrm{d}z^{-} \,\mathrm{d}^{2} z_{\perp}}{(2\pi)^{3}} \,e^{ik\cdot z} \,\langle p' | \phi^{*}(-\frac{z}{2}) \,\phi(\frac{z}{2}) | p \rangle \big|_{z^{+}=0}$$



Link with parton correlators





Leading twist

$$\gamma^{+} \sim \delta_{\lambda'\lambda}$$
$$\gamma^{+}\gamma_{5} \sim (\tau_{3})_{\lambda'\lambda}$$
$$i\sigma^{j+}\gamma_{5} \sim (\tau_{i})_{\lambda'\lambda}$$

Link with parton correlators



Adding spin and color to the picture

$$W_{\Lambda'\Lambda}^{[\Gamma]}(x,\xi,\vec{k}_{\perp},\vec{\Delta}_{\perp};\mathcal{W}) = \frac{1}{2} \int \frac{\mathrm{d}z^{-} \,\mathrm{d}^{2}z_{\perp}}{(2\pi)^{3}} \,e^{ik\cdot z} \,\langle p',\Lambda'|\overline{\psi}(-\frac{z}{2})\,\Gamma\,\mathcal{W}_{-\frac{z}{2}\frac{z}{2}}\,\psi(\frac{z}{2})|p,\Lambda\rangle\big|_{z^{+}=0}$$

Wilson line

$$\mathcal{W}_{ba} = \mathcal{P}\left[e^{ig\int_a^b \mathrm{d}x^- A^+}\right]$$

Gauge transformation

$$\psi(x) \mapsto U(x)\psi(x)$$
$$\mathcal{W}_{yx} \mapsto U(y)\mathcal{W}_{yx}U^{-1}(x)$$

Very very very complicated object not directly measurable Let's look for simpler measurable correlators





Lecture 1

- Understanding nucleon internal structure is essential
- Concept of phase space can be generalized to QM and QFT
- Relativistic effects force us to abandon 1D in phase space



