



Lrfu - CEA Saclay

Institut de recherche
sur les lois fondamentales
de l'Univers

DE LA RECHERCHE À L'INDUSTRIE

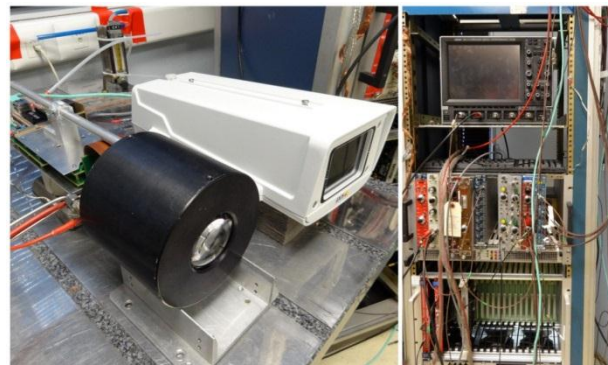
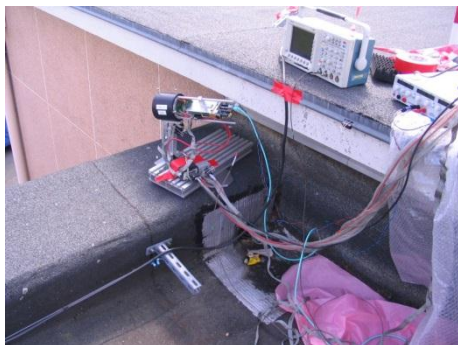


Instrumentation Days on gaseous detectors

25-26 June 2014
IPNO, Orsay

*Feedback from Micromegas operation
without gas renovation
(FORFIRE/Piggyback)*

Thomas Papaevangelou
CEA Saclay



The FORFIRE project



The FORFIRE project

Objective: the development of an outdoor fire detection system, using an innovative solar blind camera based on the technology of photosensitive gas and solid state detectors



R&D investment

CEA	475.380
ITAV	443.250
UOA	80.977
Total	999.607

Provide a fire detection system capability for:

- ✓ highly reliable
- ✓ cost effective
- ✓ early detection
- ✓ accurate localization.

Beneficiary Number	Beneficiary name	Beneficiary short name	Country	Date enter project	Date exit project
1 – SME coordinator	Irish Precision Optics	IPO	Ireland	Month 1	Month 24
2 – SME	HERON Technologies	HERON	France	Month 1	Month 24
3 – SME	OPTOEL	OPTOEL	Rumania	Month 1	Month 24
4 – SME	PINDIATEC	PINDIATEC	Spain	Month 1	Month 24
5 – End User	Forest Research Institute	FRI	Poland	Month 1	Month 24
6 – RTD	CEA	CEA	France	Month 1	Month 24
7 – RTD	ITAV	ITAV	Spain	Month 1	Month 24
8 – RTD	University of Athens	UOA	Greece	Month 1	Month 24

VUV area of the electromagnetic spectrum ($190\text{nm} < \lambda < 240\text{nm}$).

on the Earth surface only fire flames emit in this spectral range avoiding potential cross interferences from other wave sources including the Sun.

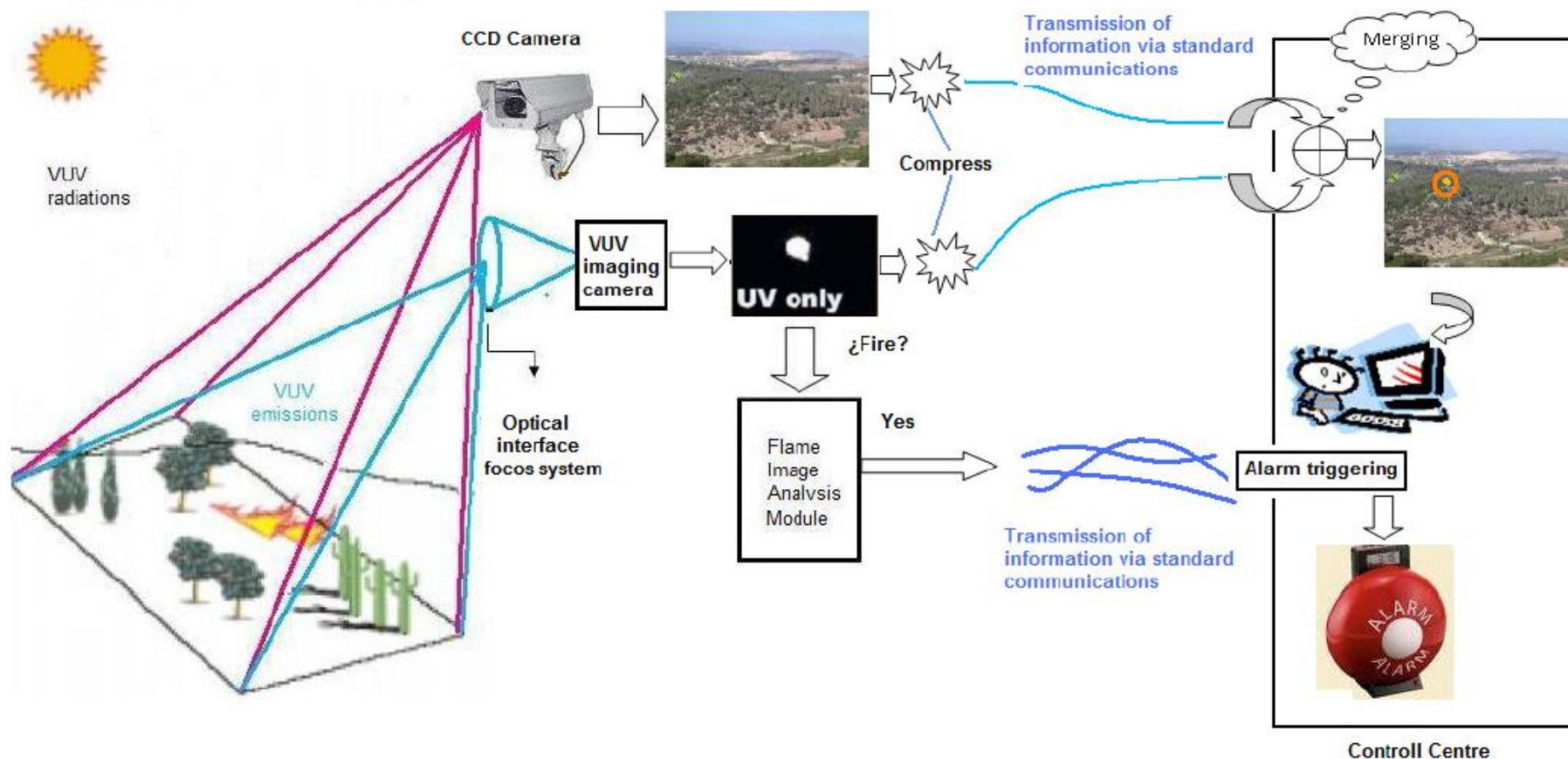
The project received a European subsidy of **1.1 million €**

within the framework of the FP7 program- Technology transfer to *Small and Medium Enterprises*.

Fire detection principle



- Detection of UV light
- Detector Network



False alarms



Where is the fire ?



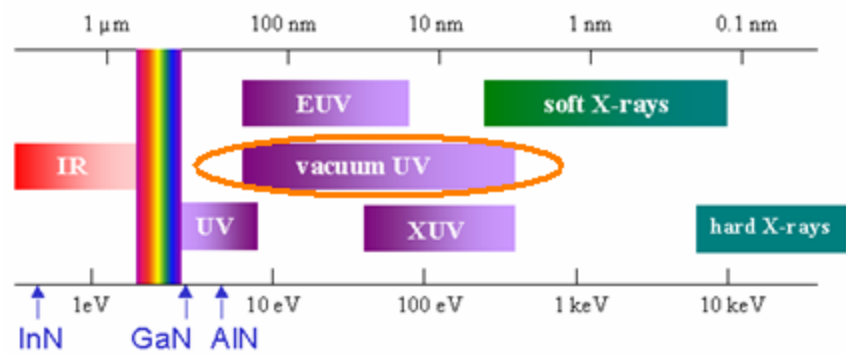
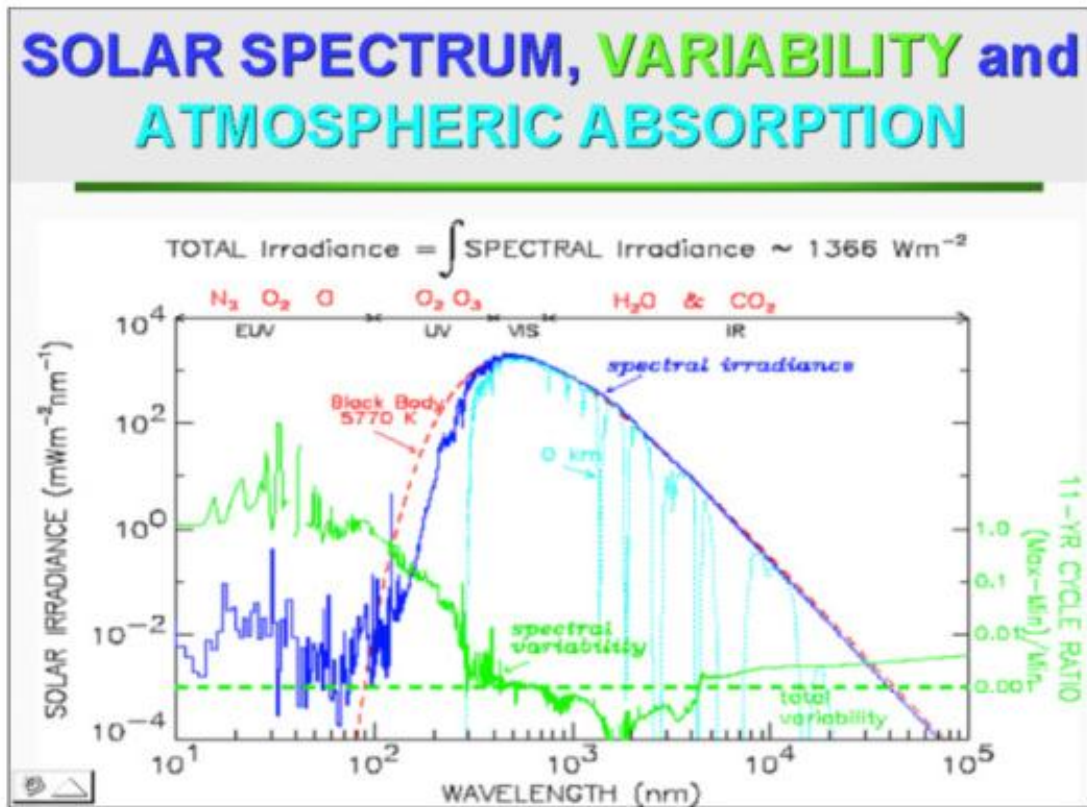
Where is the fire ?



Solar blind detectors

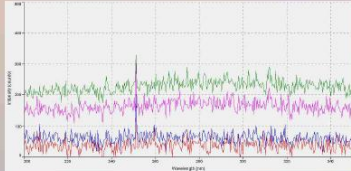
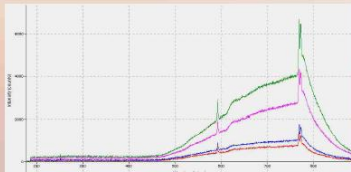
- O₃ layer absorbs solar irradiance below 250–300 nm
- Atmosphere is “transparent” to photons above 200 nm
- Photons between 200–250 nm indicate:
 - ☞ Electrical discharge
 - ☞ Explosion
 - ☞ **Flame**

A detector sensitive in the region 200 – 250 nm and insensitive above is solar blind, producing black/white images



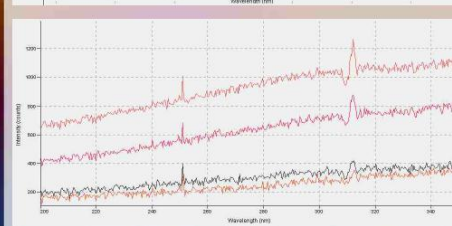
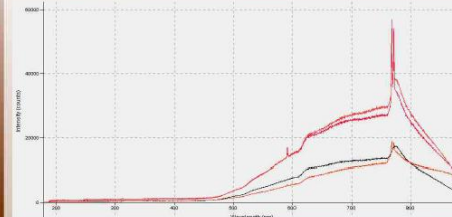
Fire spectra (Athens University)

Wood Fire



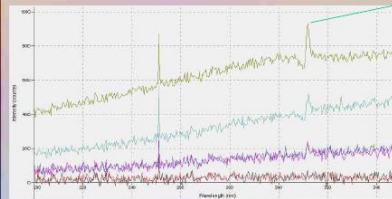
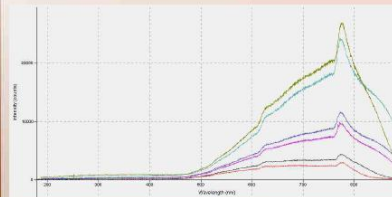
Source – Detector distance: 80 cm Integration Time: 50 ms Scans to Average: 20
 Start of fire-Low intensity_red line High intensity_Green line

Benzene fire



The black line is benzene only fire.
 The red lines are benzene with a piece of cloth inside taken at different intensities.
 The spectrum peak lines seen in the optical part are characteristic to the cloth

Petroleum fire



→ 310 nm OH peak

Petroleum spectra taken at different intensities.
 The fires initiated using a piece of cloth in order to heat the petroleum to the ignition temperature (~100 °C)

A Micromegas for UV is attractive:

- High electron amplification possible (10^6)
 - ➔ Good signal to noise ratio
 - ➔ **High sensitivity**, reaching "**single photon detection**" level.
- Very low power consumption (\ll mW)
- **Intrinsically "solar-blind"**:
the Q.E. of solid photocathodes such as CsI is significant for 200-230 nm and drops by 7 orders of magnitude up to 300 nm
- Very low production cost
- Large scale production possible
- Very fast response ($<1\mu\text{s}$).
- UV imaging possible

 **Challenge: Photocathode and gas aging**

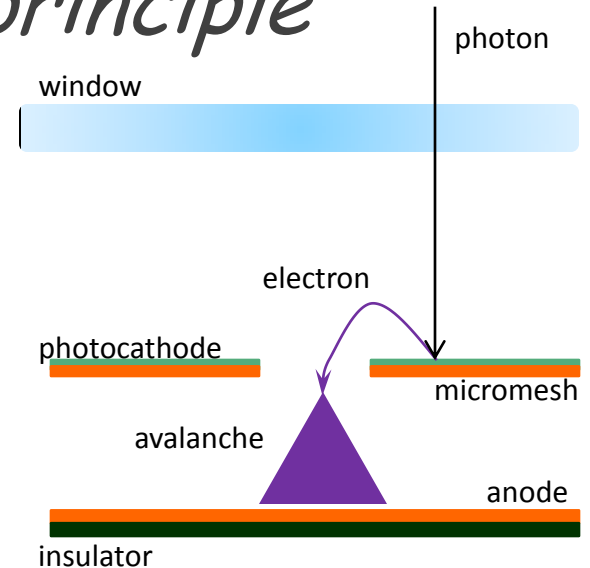
UV photon detection principle

➤ Reflective photocathode:

Photosensitive material is deposited on the top surface of the micromesh.

Photoelectrons extracted by photons will follow the field lines to the amplification region

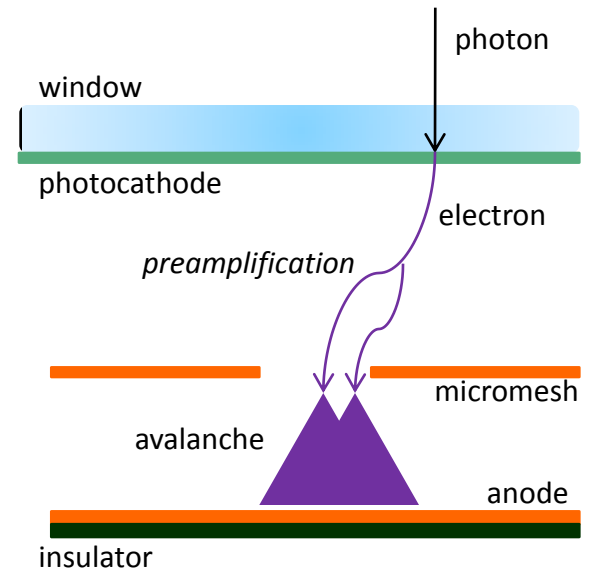
- ✓ The photocathode does not see the avalanche → no ion feedback effect → higher gain (up to 10^6)
- ✓ High electron extraction & collection efficiency
- ✓ Field on photocathode 10^4 V/cm



➤ Semi-transparent photocathode:

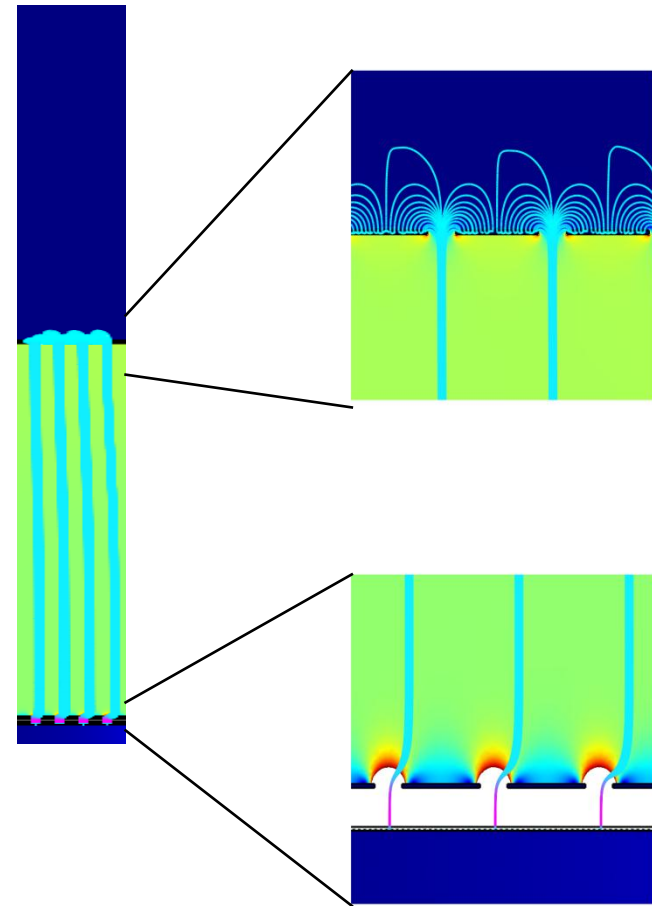
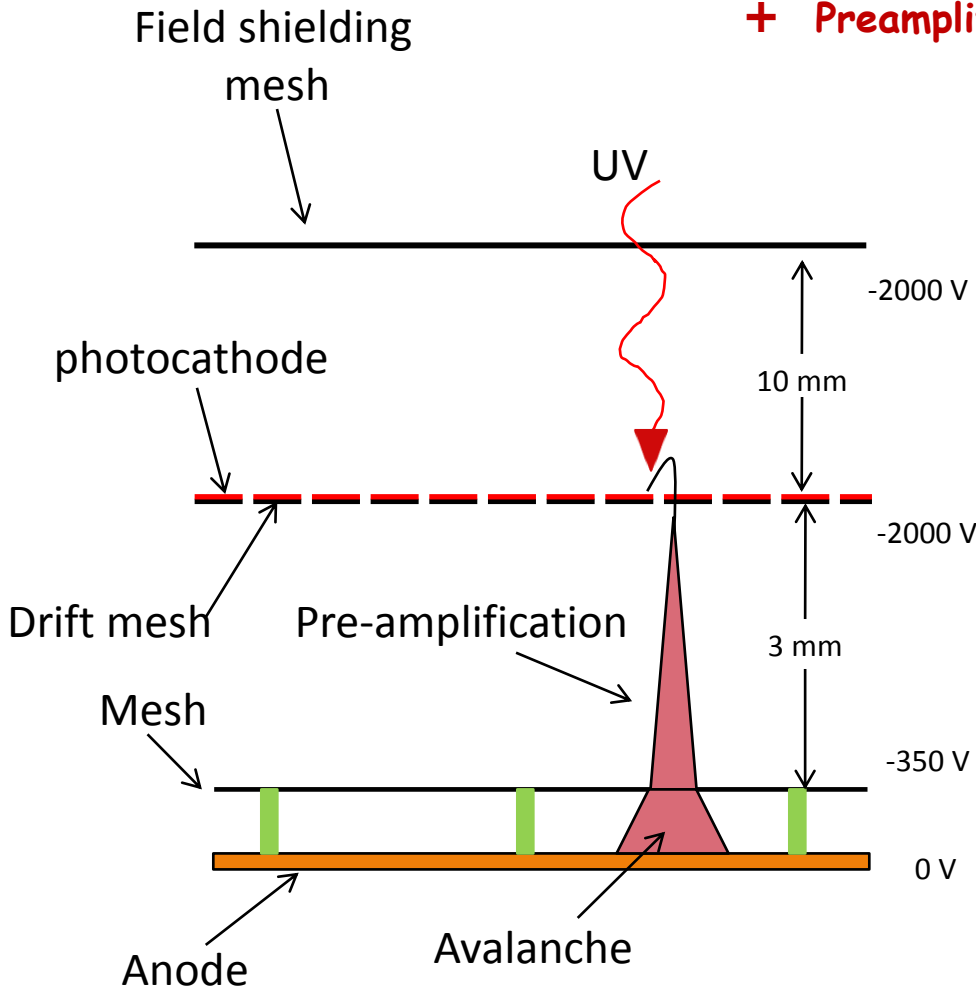
Photosensitive material is deposited on an aluminized quartz window (drift electrode)

- ✓ Extra preamplification stage → better long-term stability
- × Lower photon extraction efficiency (factor 3)
- × Fragility to sparks
- × Ion feedback → gain limitation



The FORFIRE Micromegas concept

- Our choice:  Reflective photocathode
+ Photocathode separated from detector
+ Preamplification



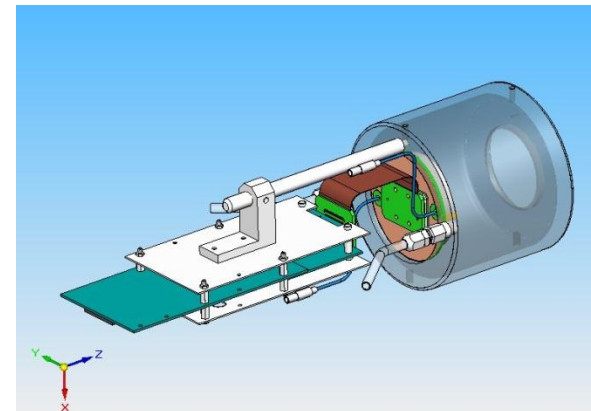
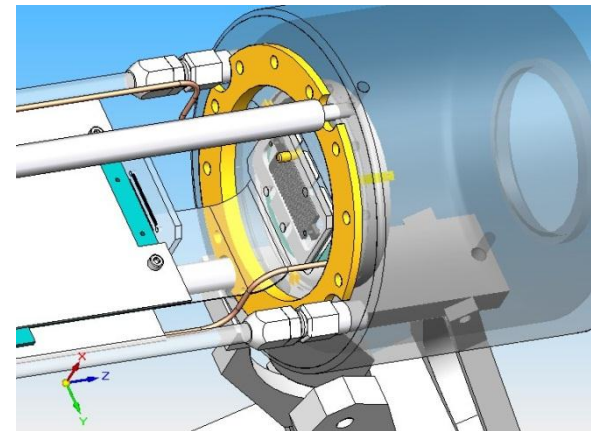
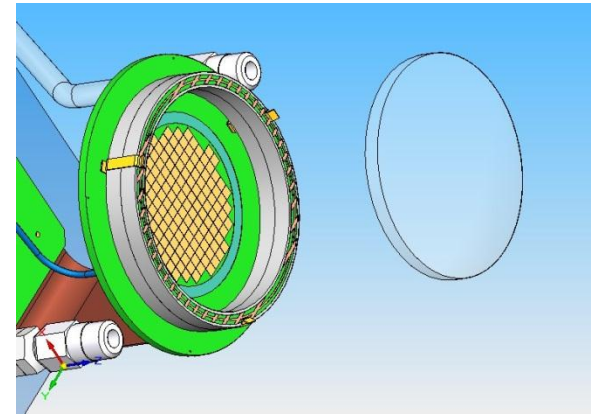
Advantages of the new concept

*Combining the advantages of the two modes,
while suppressing the disadvantages!!!*

- **Reflective photocathode**
 - ➔ **High electron extraction efficiency**
- **Preamplification**
- **No ion feedback**
 - ➔ **Very high total gain ($\gg 10^7$)**
 - ➔ **Stability in sealed mode**
 - ➔ **Exceptional signal to noise ratio**
- **Photocathode separated from detector**
 - ➔ **Easy fabrication/handling of CsI**

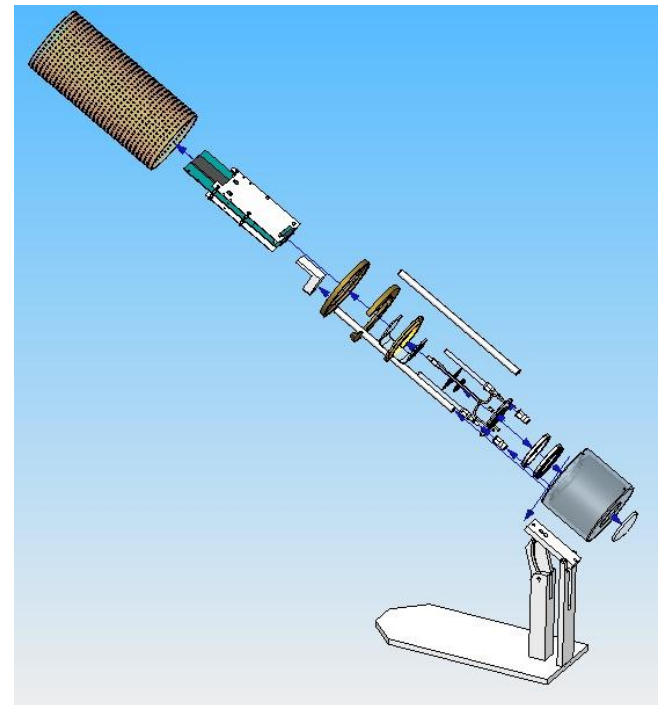
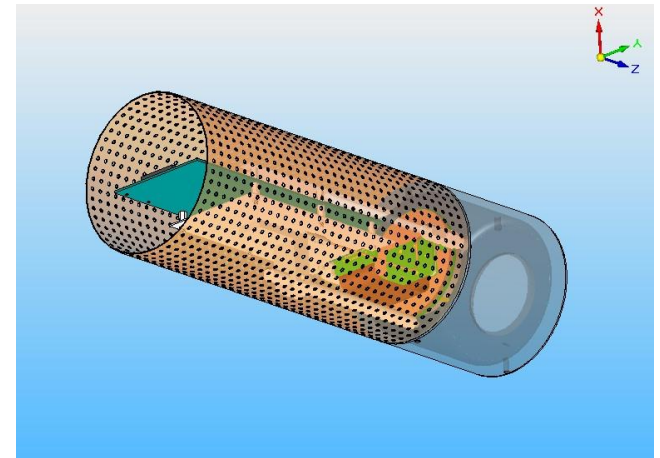
The FORFIRE detector

- Bulk Micromegas
 - 144 pixels
 - Gas: 90% Ne + 10% Ethane
 - High gain
 - Good electron extraction efficiency
 - Photocathode = drift electrode
-
- ✓ High purity Ne gas (6.0)
 - ✓ Metallic tubes and components
 - ✓ Quartz lens glued with glue appropriate for ultra vacuum
 - ✓ Anodized aluminum chamber
 - ✓ Detector pumped and heated up before sealing
-
- ✗ Gas tube feedthroughs (Stubli) leaky → glued
 - ✗ PCB outgassing (?)
 - ✗ O-ring to seal the PCB



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Photocathode

Selected Material: **CsI deposited on Ni mesh.**
The photocathode is the drift electrode of the detector

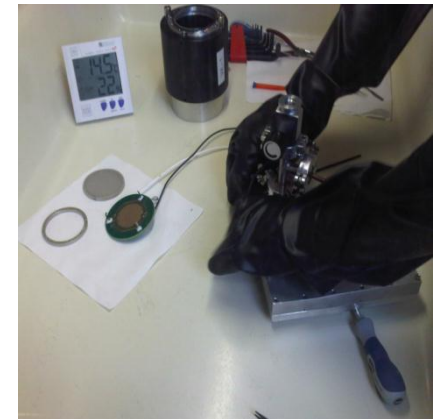
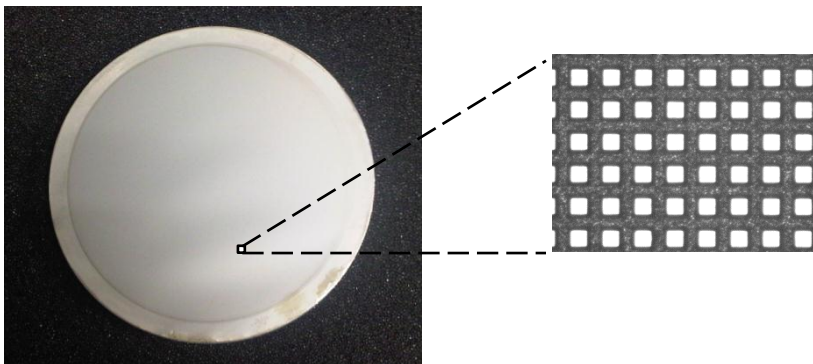
All photocathodes used were fabricated at CERN with CsI evaporation.

Evaporation facility at CEA: ready to start!

Aging:

- Photocathodes produced Jan / Apr / May / Oct 2011
- Measurements done Feb / May / Jul / Sep / Nov 2011

Similar behavior



General behavior

High gain ($\gg 10^5$) in a single stage

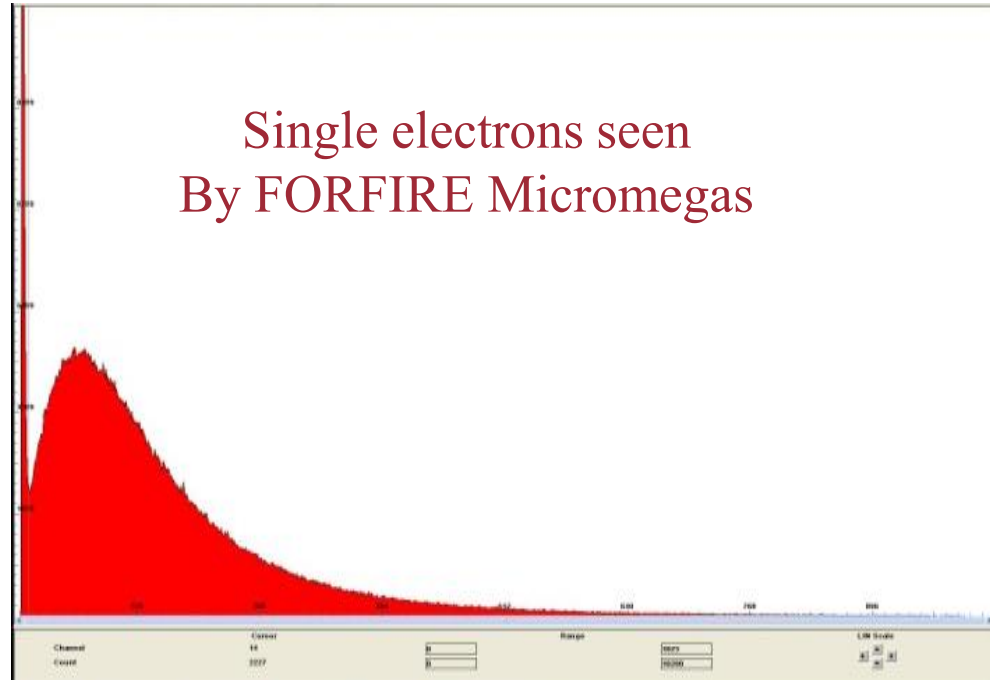
☞ Single electron

No ion feedback

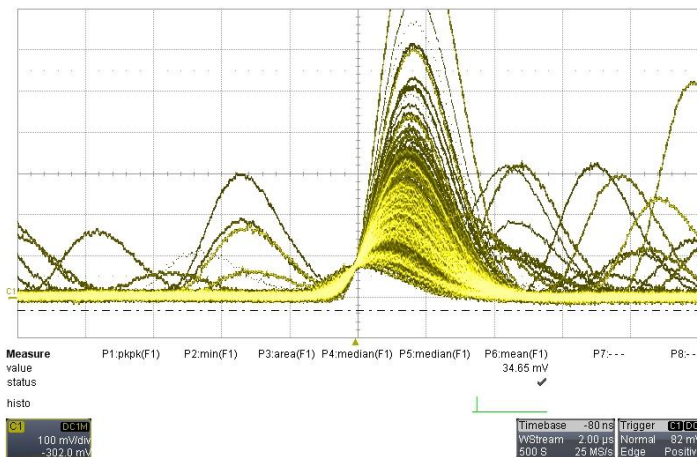
Sealed

Preamplification (>100)

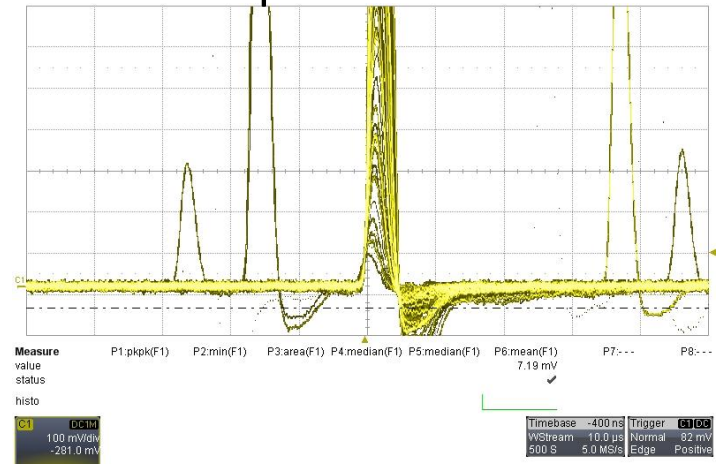
- ✓ Stable operation, far from sparking limit
- ✓ Huge gain (even causing trouble using standard electronics!)



Normal mode



Preamplification mode



Performance of FORFIRE prototype



Q.E & Sensitivity

Photodiode Q.E.(200nm)≈70%

FORFIRE Q.E.(200nm)≈1%

However

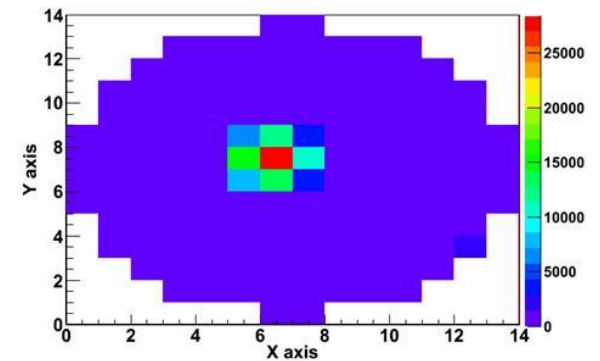
Photodiode minimum signal $\approx 5 \text{ pA} \rightarrow$

Photodiode minimum sensitivity \approx

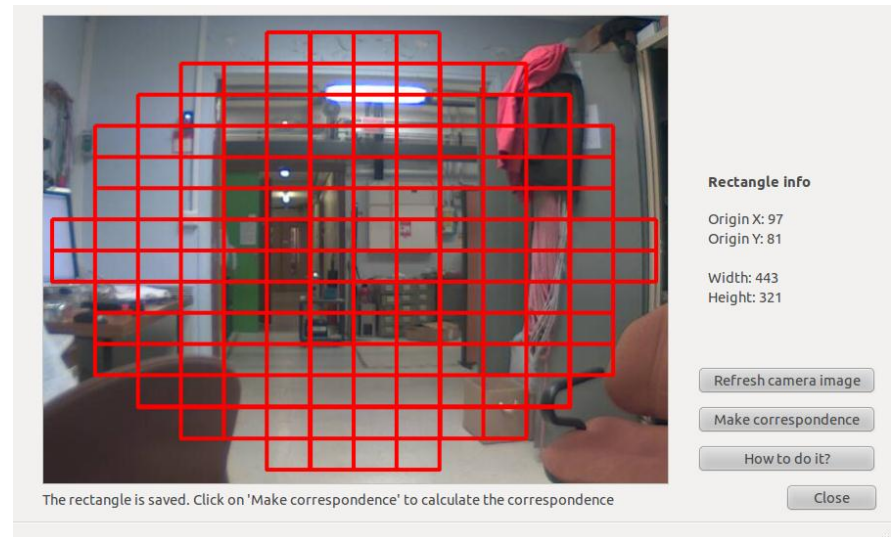
5×10^7 photons

FORFIRE minimum sensitivity \approx

5×10^2 photons



Imaging - alarms



See the correspondence grid on the camera image and note the origin $(x,y) = (97, 81)$, width = 443 and height = 321

Use these coordinates in the FORFIRE software to calibrate both images.

OK

- To mark the square window for merging the camera image with that of the detector.
- Changing the software values in `ImakeT2K.cpp` and recompiling
- Gives the gui the correct dimensions to the pixel mapping image and the IP camera

Imaging - alarms

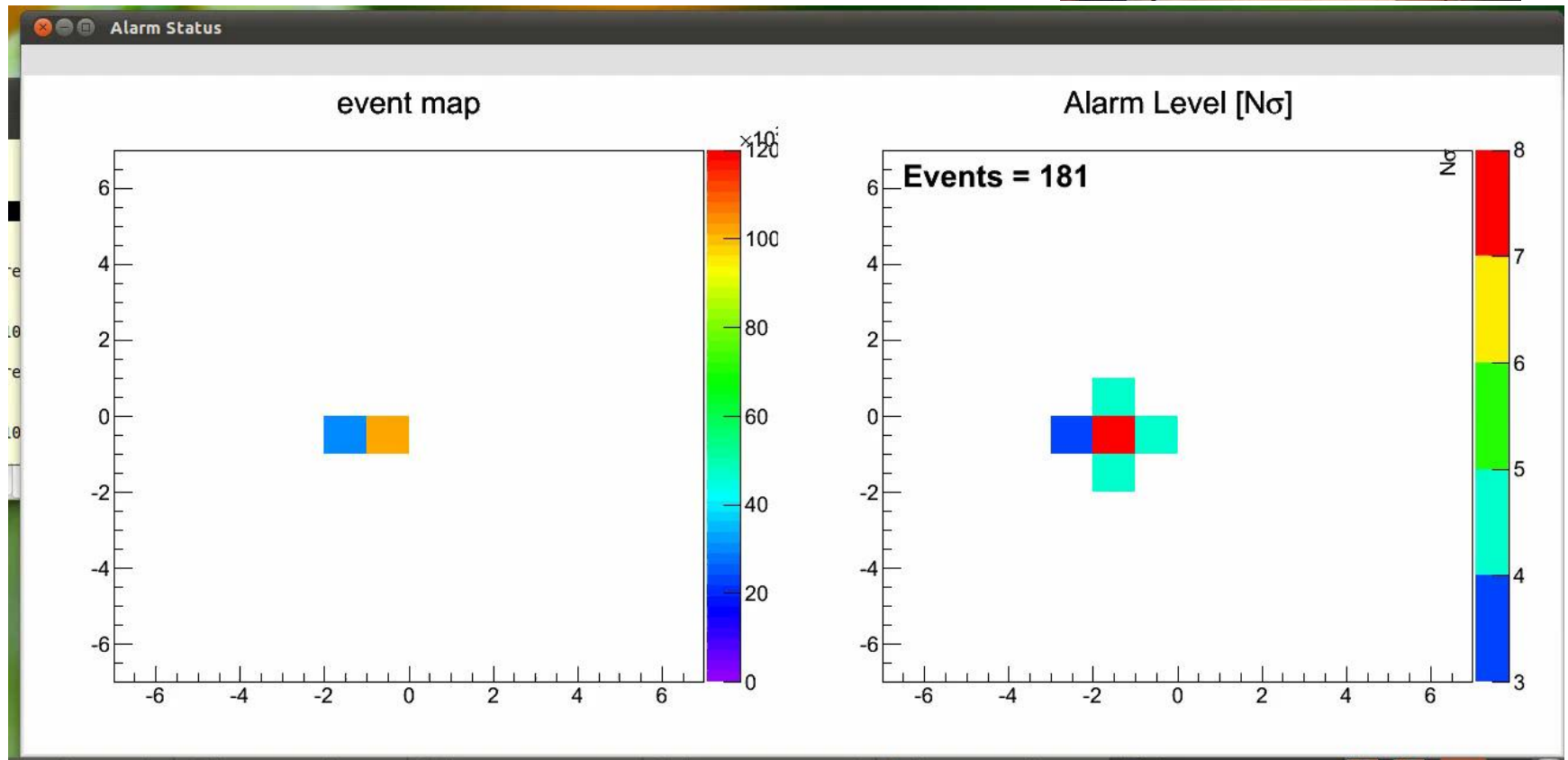
The image displays two side-by-side screenshots of the 'DAQ INTERFACE' software, showing the process of triggering an alarm based on imaging data.

Left Screenshot: The interface shows 'Acquisition from DAQ file' with parameters: Asic no. 1, Channel no. 18, Events to read: 2, Delay between events: 0.1. Under 'Online HW Acquisition', 'Change Working Directory' and 'Read from HW device' are visible. A 'Delay between events: 0.005' is set. A checkbox 'Write HW events into a set of files?' is unchecked. The 'No. event:' is 1253, 'Max. value:' is 298, and 'Max. position:' is [7, 7]. The 'Channel calculation:' is 'Max cell value', 'Window size (secs):' is 60, 'Threshold channel:' is 1400, and 'Threshold image:' is 40. The 'Tracking algorithms and merging' checkbox is checked. A 'Watch Camera' button is at the bottom. A log window on the right shows a sequence of 'Image captured.' and 'Waiting...' messages, followed by '<< End of acquisition core.' and '>> Acquisition cleaned up...'. A red 'ALARM' label is visible at the bottom right.

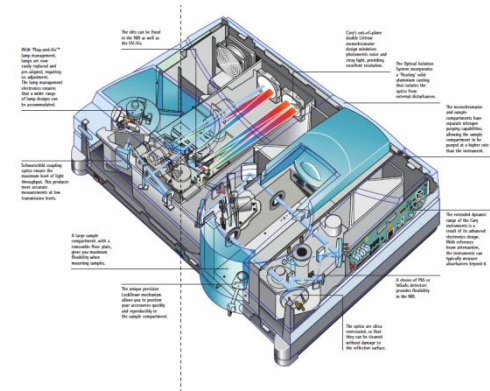
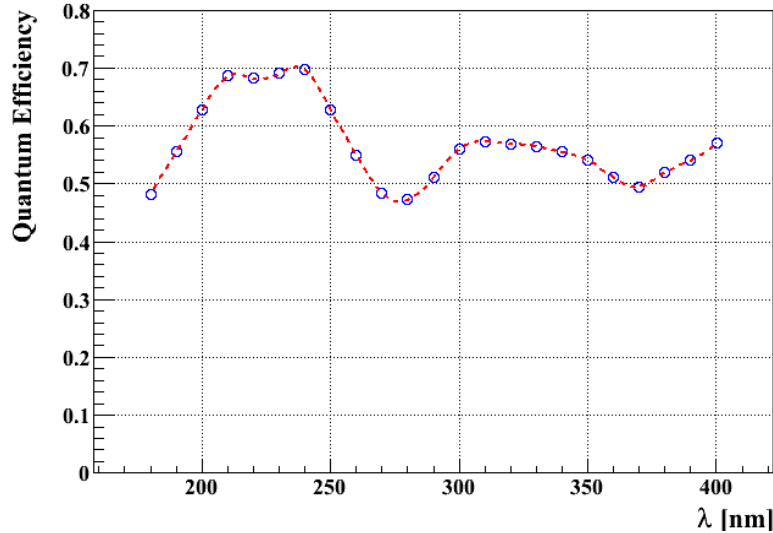
Right Screenshot: The interface shows 'Acquisition from Camera Viewer'. The 'Camera Viewer' window displays a live video feed of a room with a red square overlaid on the floor, corresponding to the alarm location. The 'No. event:' is 1253, 'Max. value:' is 298, and 'Max. position:' is [7, 7]. The 'Tracking algorithms and merging' checkbox is checked. A 'Watch Camera' button is at the bottom. A log window on the right shows a sequence of 'Image captured.' and 'Waiting...' messages, followed by '<< End of acquisition core.' and '>> Acquisition cleaned up...'. A red 'ALARM' label is visible at the bottom right.

- Click on Tracking Algorithms and Merging to raise an image on alarm event
- Imaging software defines the position from the pixel map onto the camera image to pin-point the location of the triggered alarm.

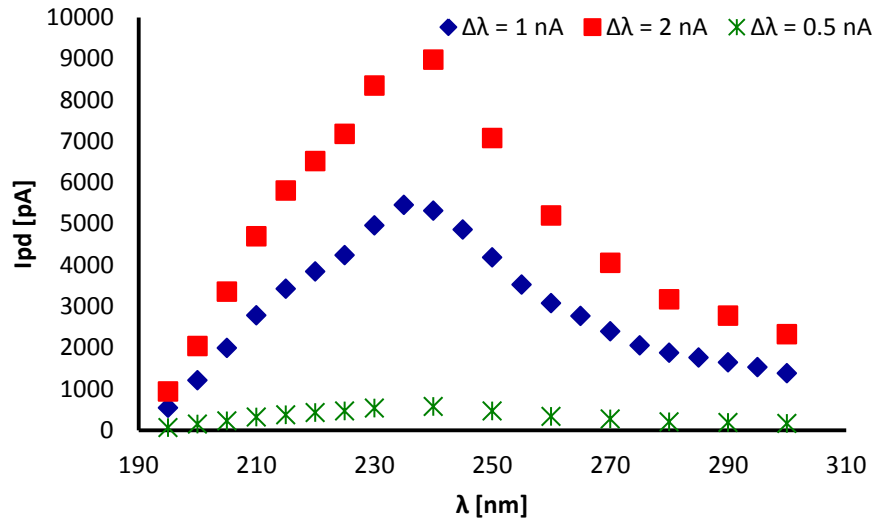
Recording a candle's flame



Quantum Efficiency of the FORFIRE detector

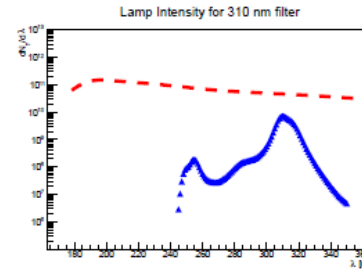
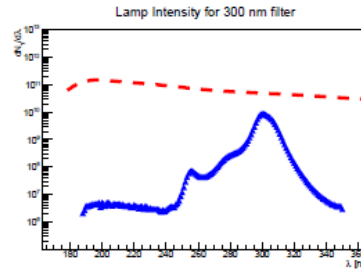
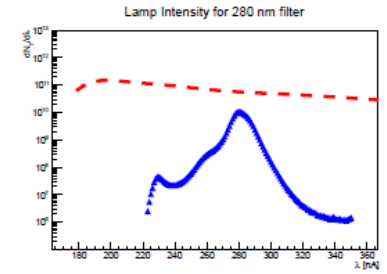
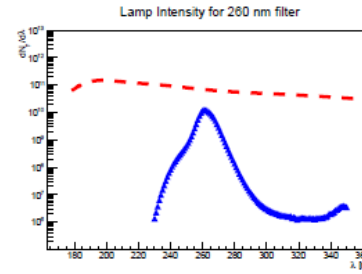
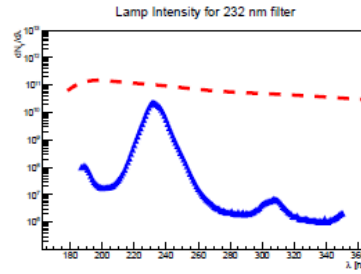
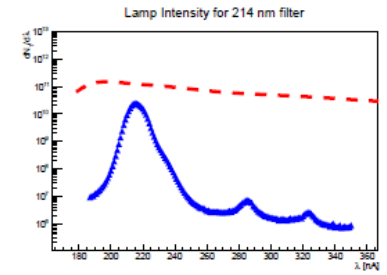
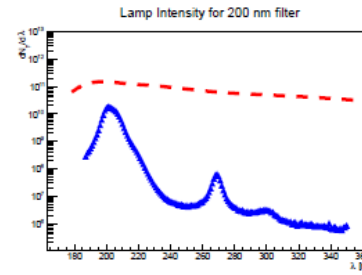
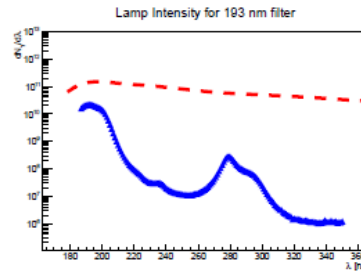
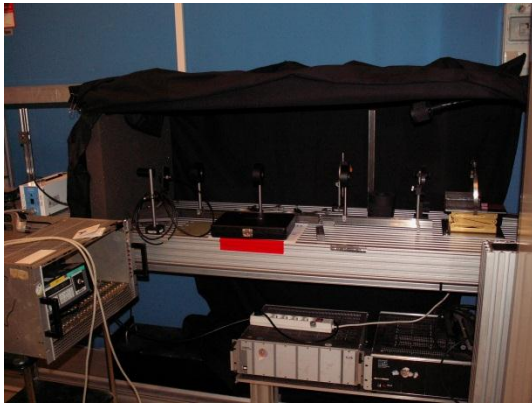


The measurements were performed using the Varian 5000 spectrometer. The photon flux as a function of the wavelength was measured using a calibrated PD222AUV photodiode



Spectrometer flux for several bandwidth acceptances. Measurements repeated within a period of 1 month revealed stability of the order of few %.

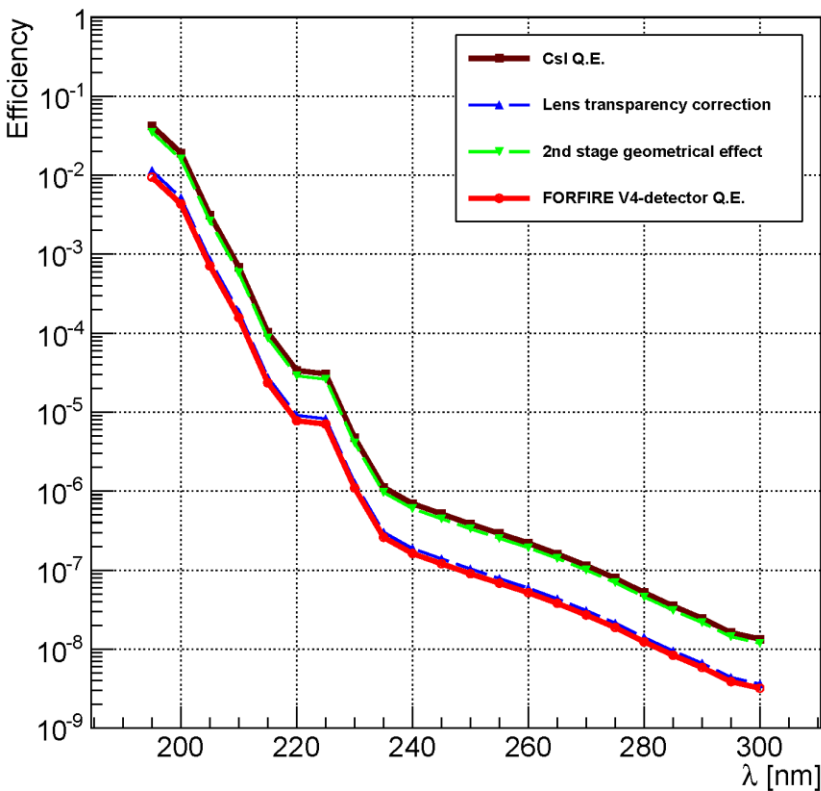
UV filters + deuterium lamp



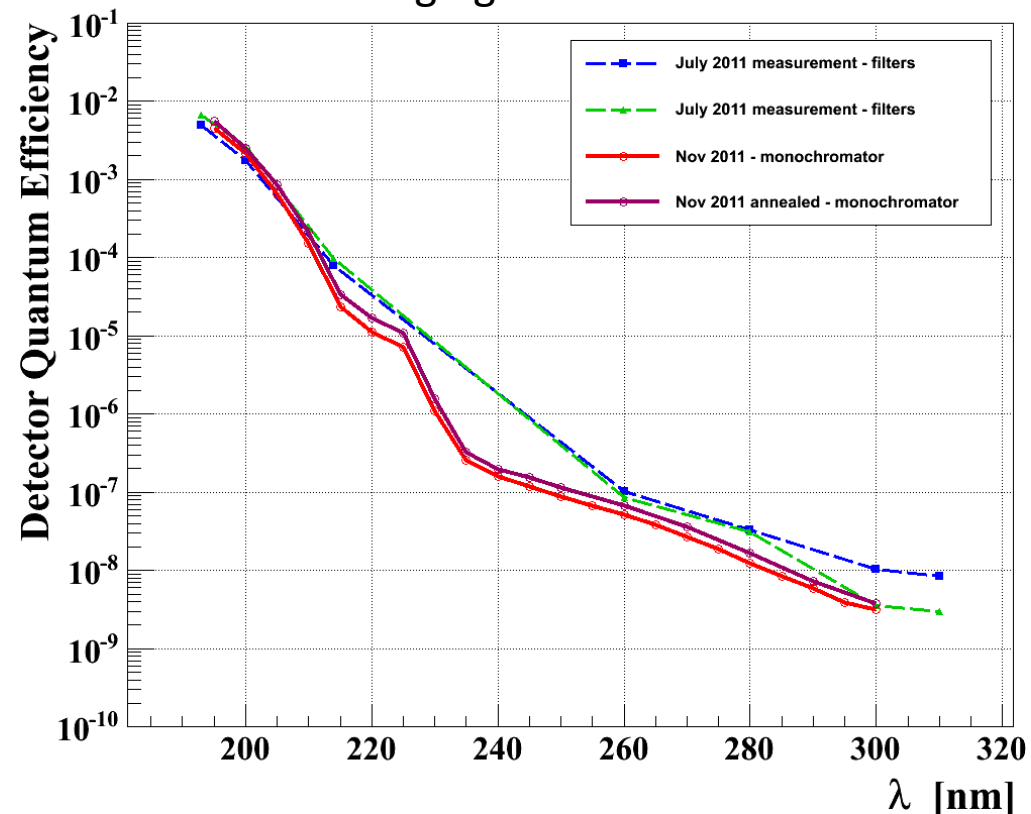
Integrated γ flux at 50 cm	
193 nm filter:	$3.202e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
200 nm filter:	$1.953e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
214 nm filter:	$2.673e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
232 nm filter:	$2.440e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
260 nm filter:	$1.394e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
280 nm filter:	$1.391e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
300 nm filter:	$1.139e+11 \gamma \text{ s}^{-1} \text{ cm}^{-2}$
310 nm filter:	$9.430e+10 \gamma \text{ s}^{-1} \text{ cm}^{-2}$

Performance of FORFIRE prototype

Q. E

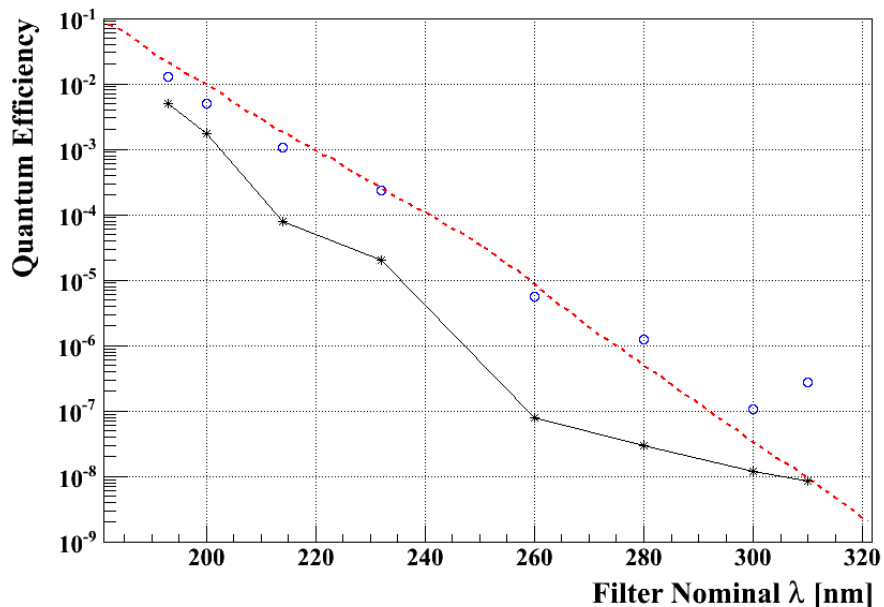


Aging measurements

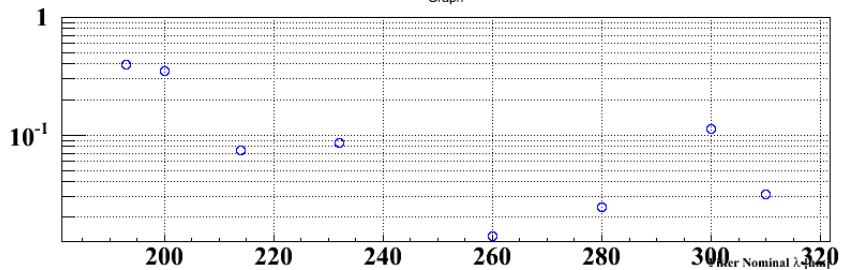


Performance of FORFIRE prototype

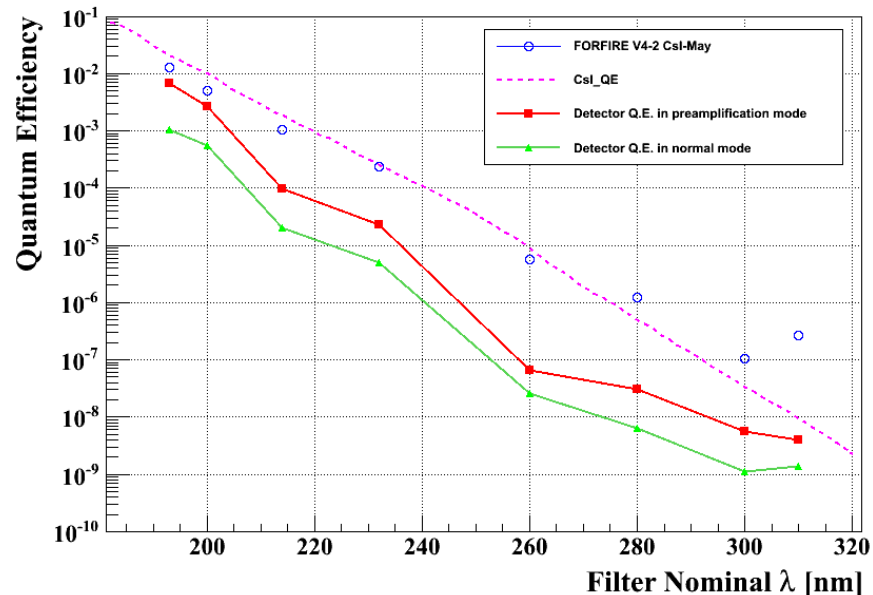
V4-2 May Csl measured July 2011



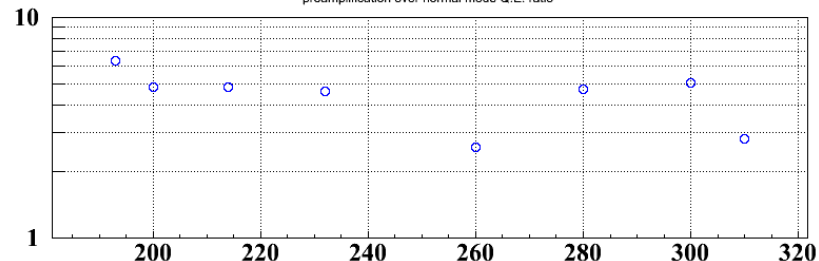
Graph



FORFIRE V4-2 Csl-May



preamplification over normal mode Q.E. ratio



Feedback from the FORFIRE detector

- ✓ Operation in sealed mode possible using standard vacuum technics (baking, gluing etc)
- ✓ Ne - Ethane mixtures have good behavior
- ✓ Stable performance for several weeks
- ✓ Best aging measurement:
 - Detector filled in April
 - QE measurements in May
 - Shipped to Athens for field measurements (June)
 - Q.E. measurements in July → similar performance
- No systematic tests done for gas aging → concentrate on photocathode aging
- Most of problems due to the design: techniques for laboratory prototypes (o-rings feedthroughs etc) are not appropriate for sealed mode operation → Industrial methods should be used (glass, Coca-Cola cans...)
- Passing voltages & signals through the PCB helps a lot → Piggyback



The Piggyback Micromegas



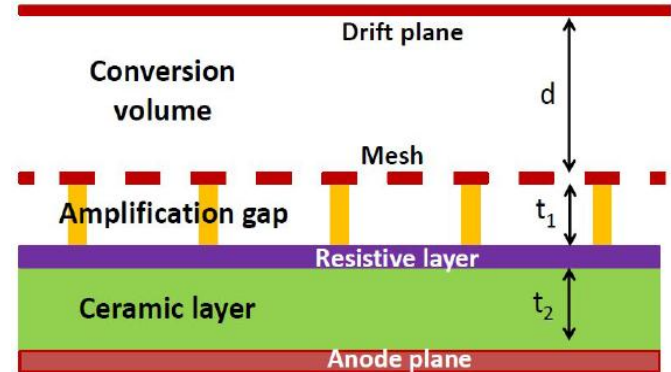
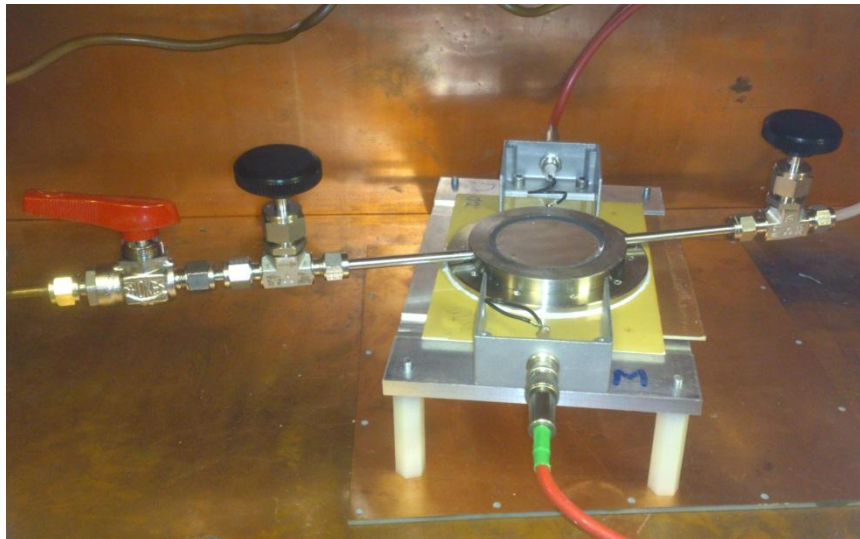
The "piggyback" concept

Micromegas on a ceramic

Resistive layer

Readout through capacitance coupling

- **Detector completely decoupled from readout electronics!**
- **Readout without fit-throughs**
- ✓ Spark protection
- ✓ Appropriate for sealed operation
- ✓ Window can be grounded

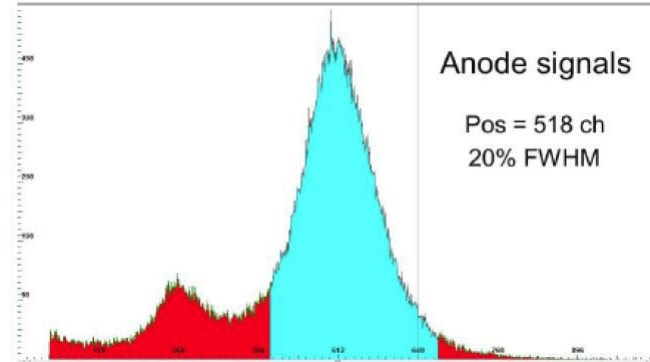
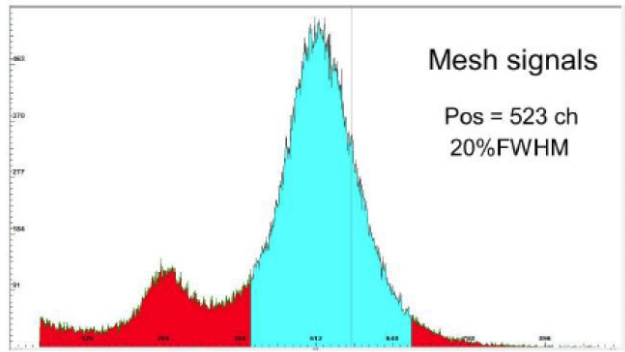


[arXiv:1208.6525](https://arxiv.org/abs/1208.6525) [physics.ins-det]

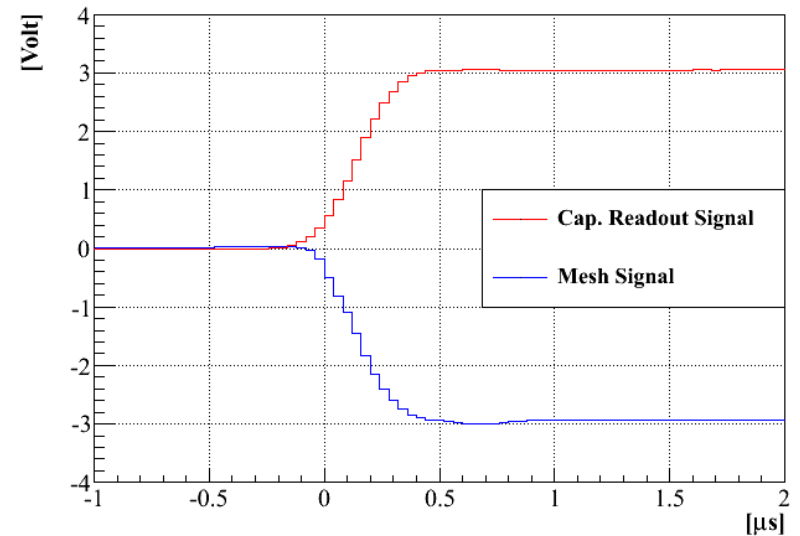
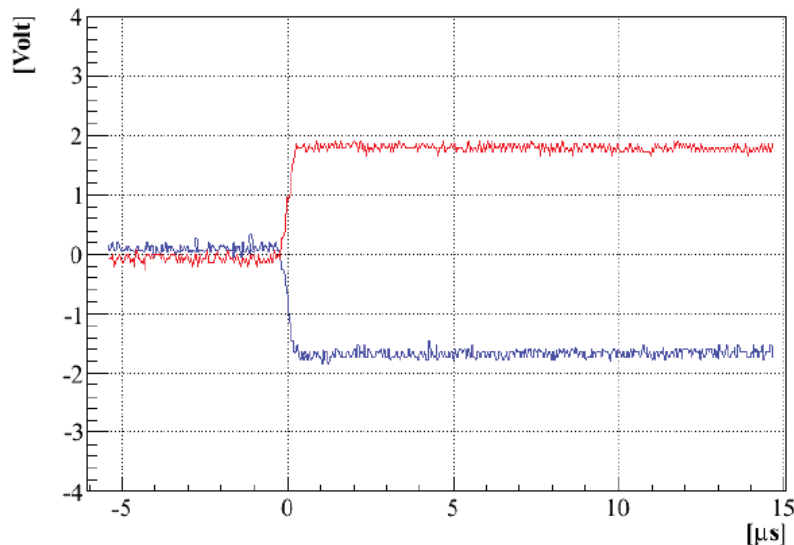


The "piggyback" concept

Checking possible looses by the ceramic layer: signal entirely transmitted

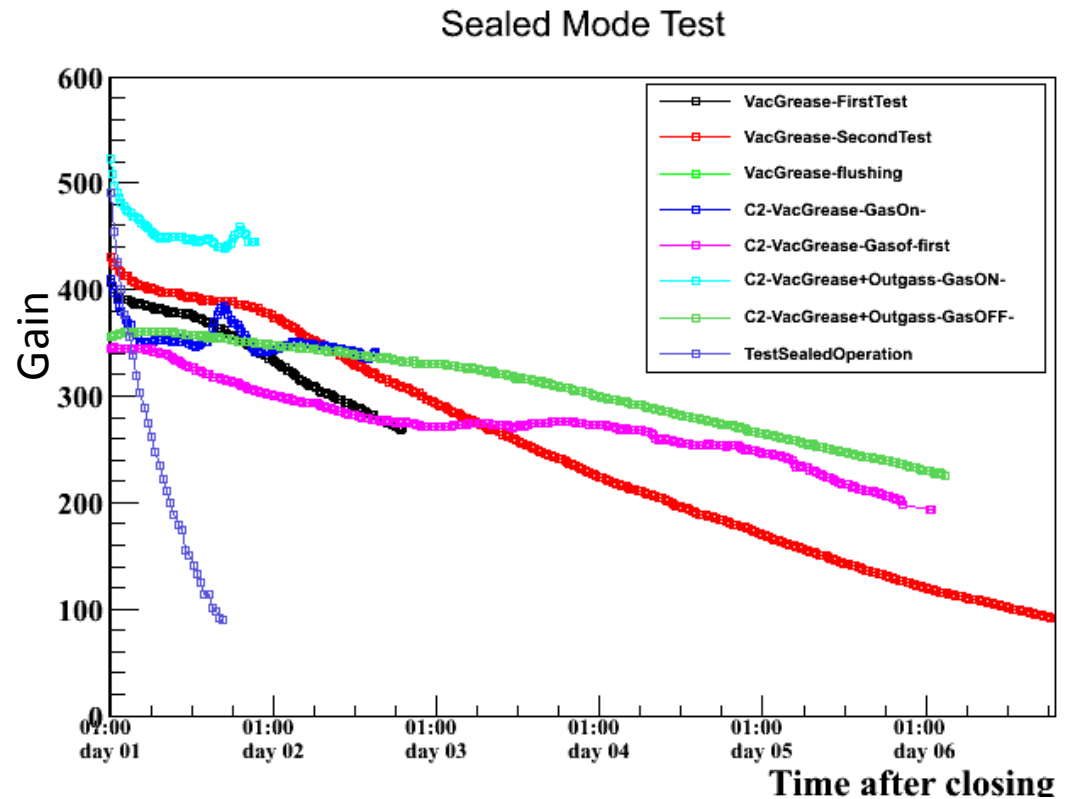


Test with a ^{252}Cf (fission fragments signals) reading simultaneously mesh and anode



First attempts for a "sealed" Piggyback

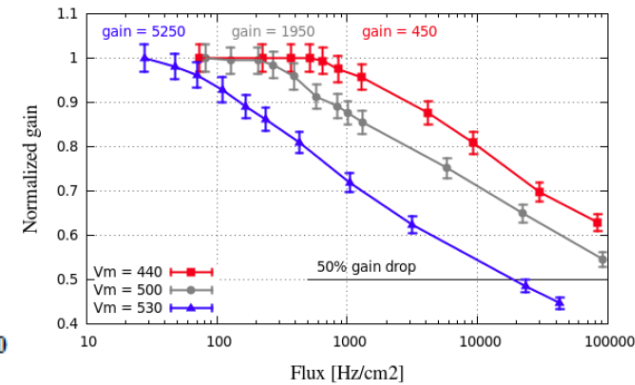
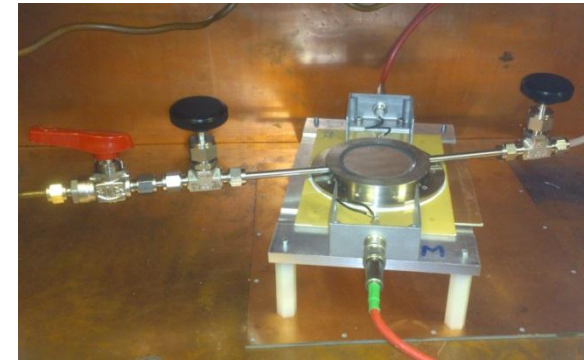
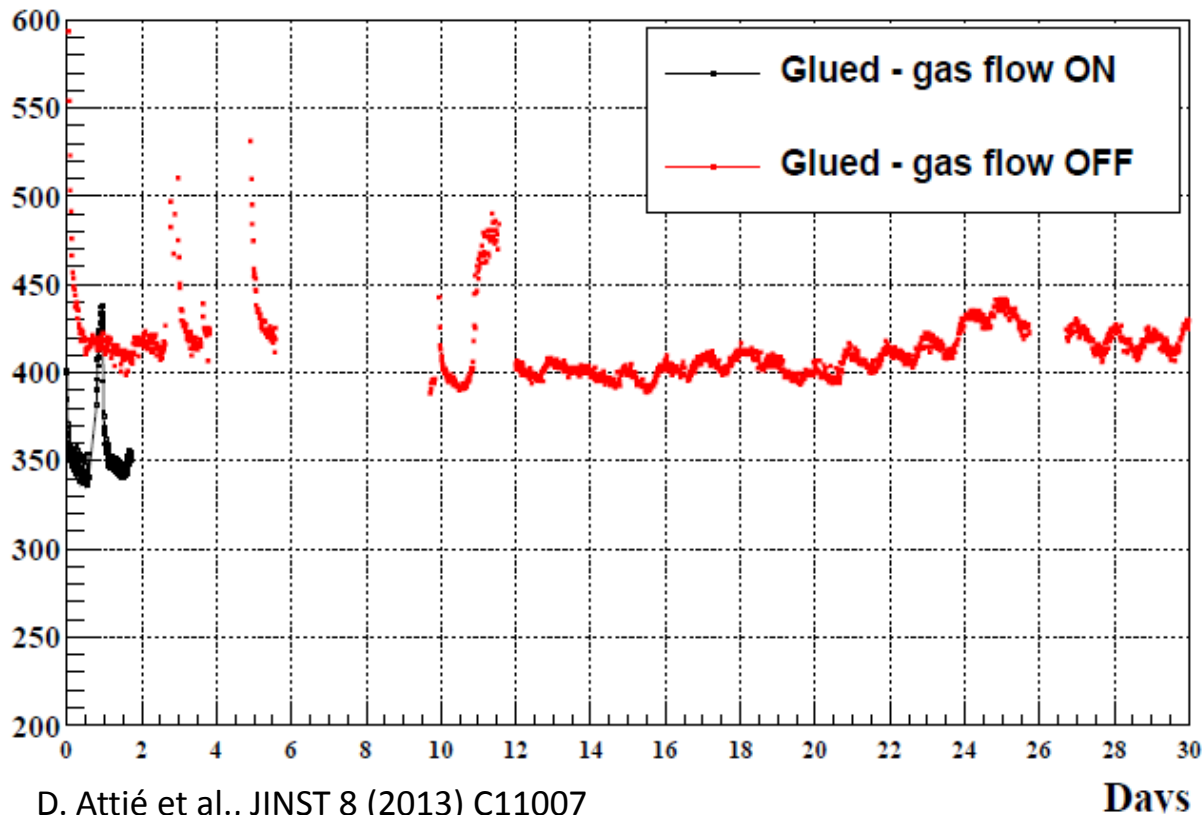
- Piggyback closed with a metallic chamber + Al window
 - ➔ Use standard o-rings
 - ➔ Improve with standard vacuum grease
 - ➔ Improve by pumping for few hours
- Close the chamber with ultra vacuum glue + outgas by heating and pumping
 - ➔ Wrong manipulation of the valves blew up the window. Tests just restarted...
- Design of new chamber (by an expert...)
- *Another approach: Microbulk*
 - ✓ Can heat up to $> 300\text{ }^{\circ}\text{C}$
 - ✓ More robust



Is it due to leaks from the detector or material outgassing?

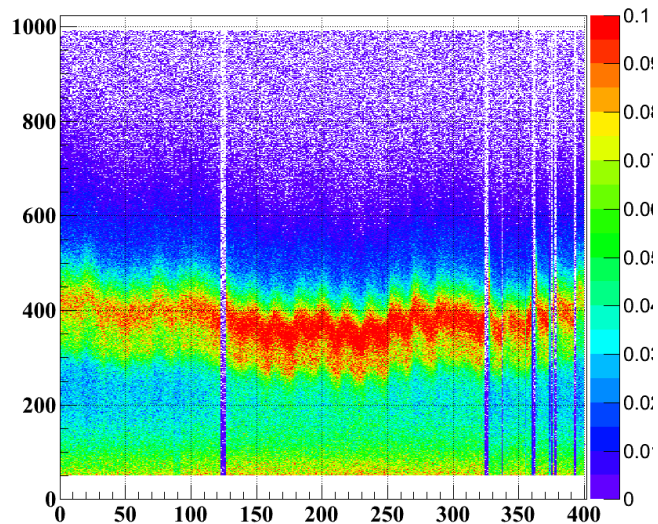
Using lots of Torrseal glue + bake-out

Gain evolution after one month sealed mode!

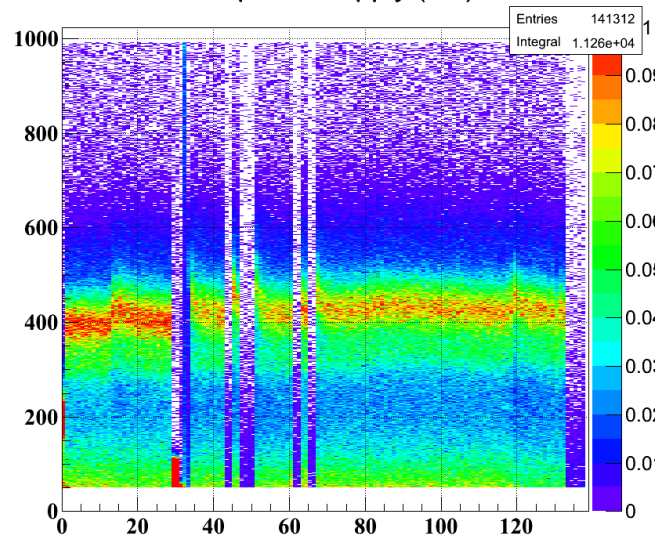


Exploring the variations

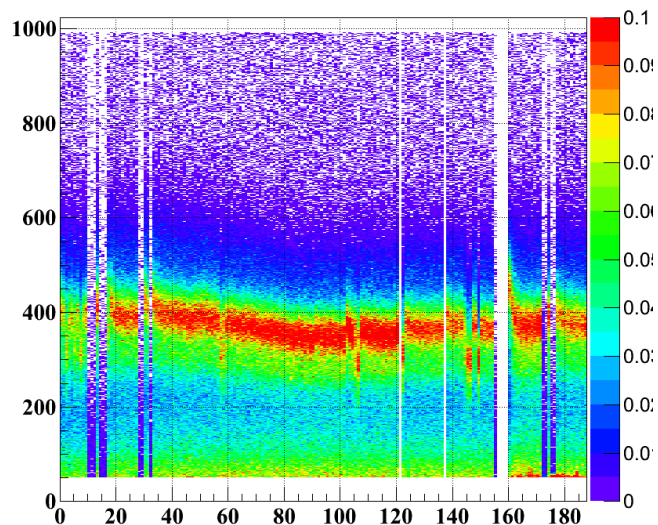
New run - August settings



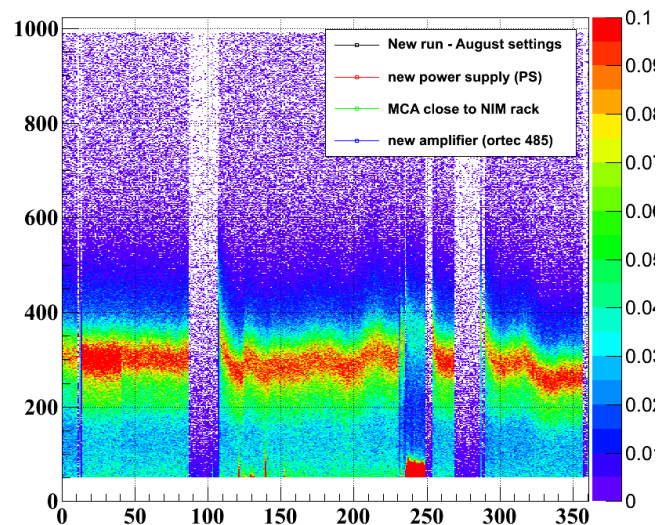
new power supply (PS)



MCA close to NIM rack



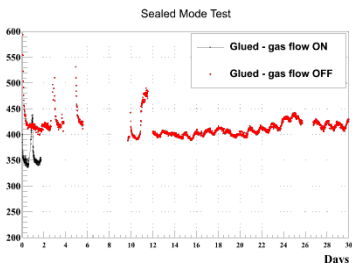
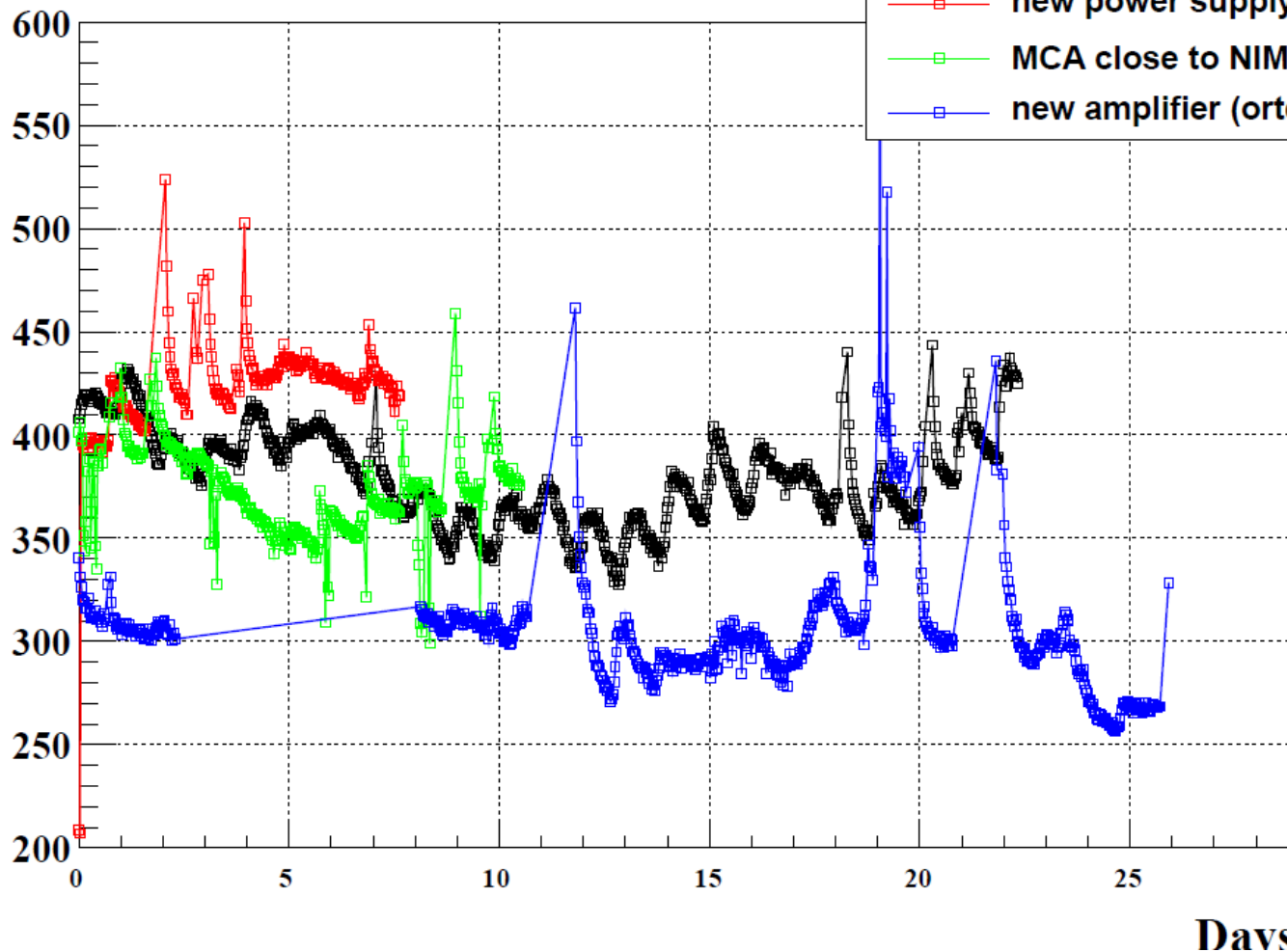
new amplifier (ortec 485)



June - November 2013

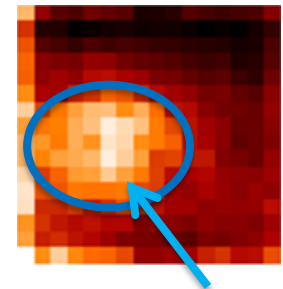
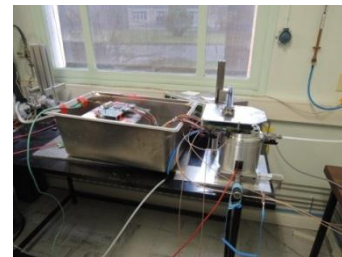
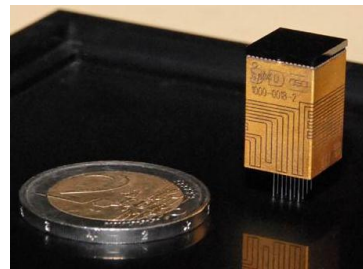
Sealed Mode Test

- New run - August settings
- new power supply (PS)
- MCA close to NIM rack
- new amplifier (ortec 485)



Feedback from the FORFIRE detector

- ✓ Similar techniques with the FORFIRE detector
 - ✓ Same gas system
 - ✓ Ceramic → excellent for signal extraction
 - ✓ **Operation in sealed mode possible for at least 6 months**
 - ✓ Ne - Ethane mixtures better behavior than Ar mixtures
 - ✓ Detector is still sealed since last year (to be tested!)
- Verified that o-rings are not good for sealed mode. **Industrial methods should be used**
 - The thin window is a potential leak source. Best operation with a 25 μm aluminum foil over a \varnothing 2mm hole (+some glue on the hole...)
 - A second attempt was not so successful. Gluing should be done by experts...
 - Tests stopped in order to concentrate to signal readout with a chip



Thank you!