

Developing the SuperNEMO $\beta\beta$ source

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The neutrino-less double beta decay

$2\nu\beta\beta$ decay: $(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e$

- 2nd order process **allowed** in the SM

$0\nu\beta\beta$ decay: $(A, Z) \rightarrow (A, Z + 2) + 2e^-$

- process **forbidden** in the SM

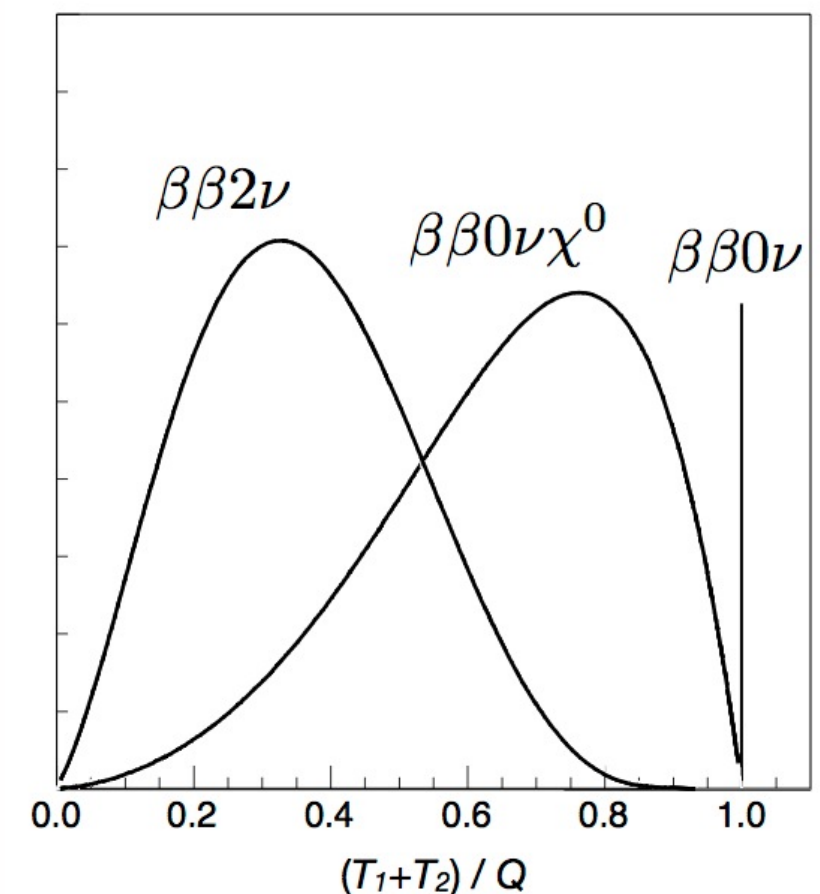
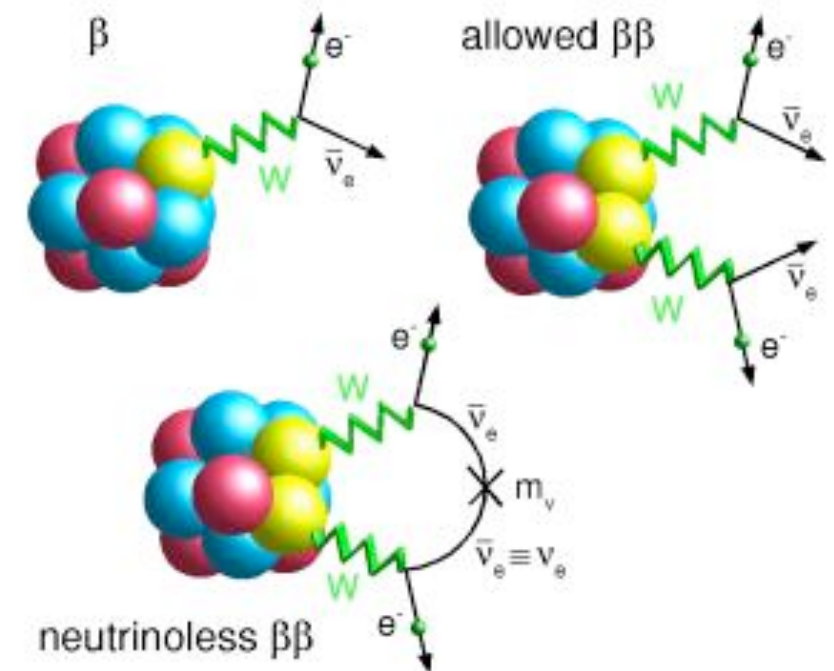
Interesting implication for particle physics:

- Lepton number violation must occur
- GUT, Leptogenesis model, See-Saw mechanism

$$(T_{1/2}^{0\nu})^{-1} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \eta^2$$

- **Light Majorana neutrino exchange**
- Right-handed current (V+A), SUSY, 1 Majoron, etc.

Different event topology in the final state



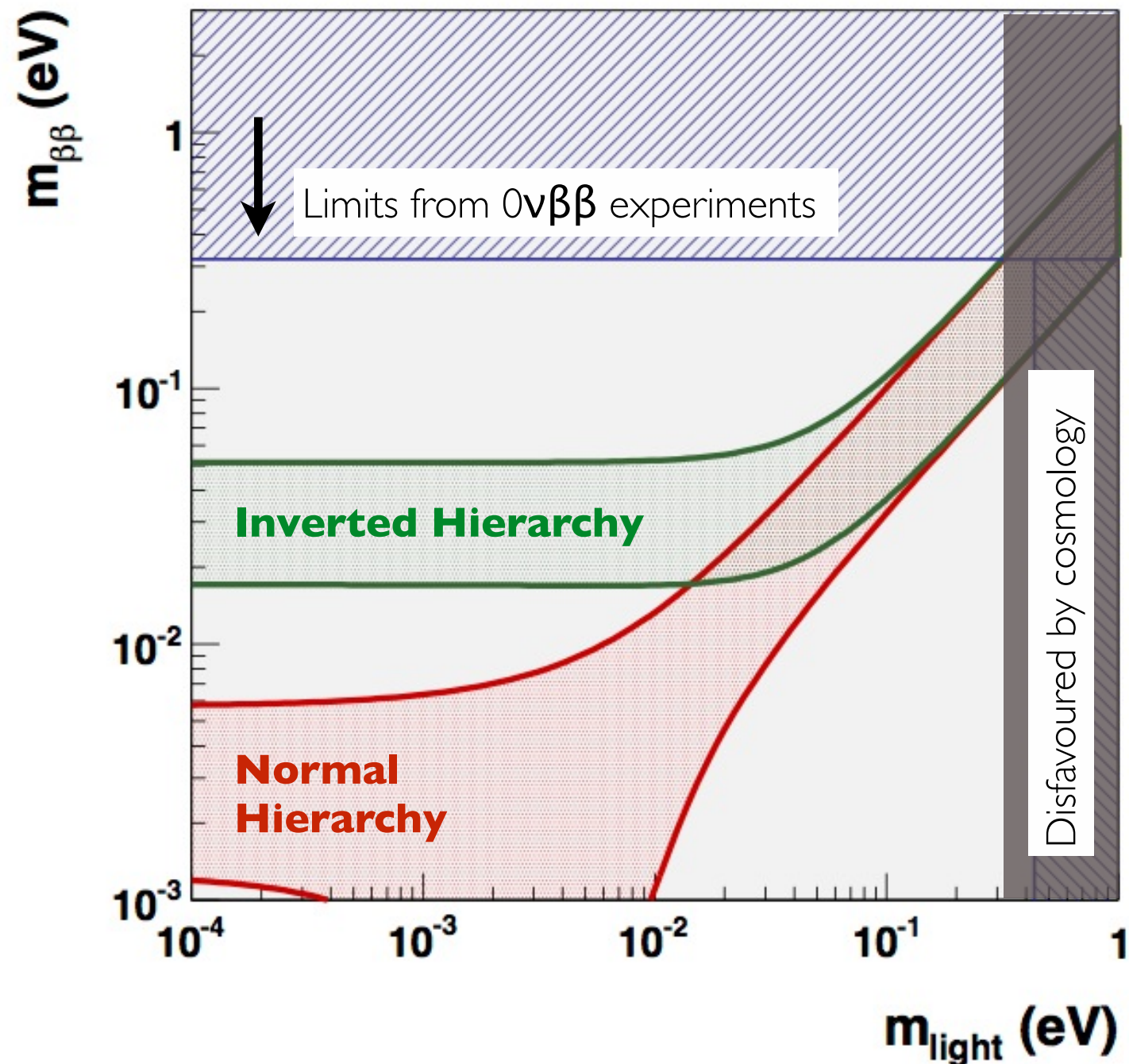
Sensitivity on neutrino mass scale

$$\left(T_{1/2}^{0\nu}\right)^{-1} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{ee} \rangle^2}{m_e^2}$$

Effective mass term: $\langle m_{ee} \rangle = \left| \sum_i U_{ei}^2 m_i \right|$

Related to neutrino mixing matrix

Sensitive to the neutrino mass scale (m_{light})



What's the status?

1993 - 2000:

- **HdM** (~11 kg) & **IGEX** (~2 kg), ^{76}Ge
 - $T_{1/2}^{0\nu} > 1.9 \cdot 10^{25} \text{ y}$ @ 90% C.L.
- **HdM claim**: $\langle m_{ee} \rangle = 0.32 \pm 0.03 \text{ eV}$

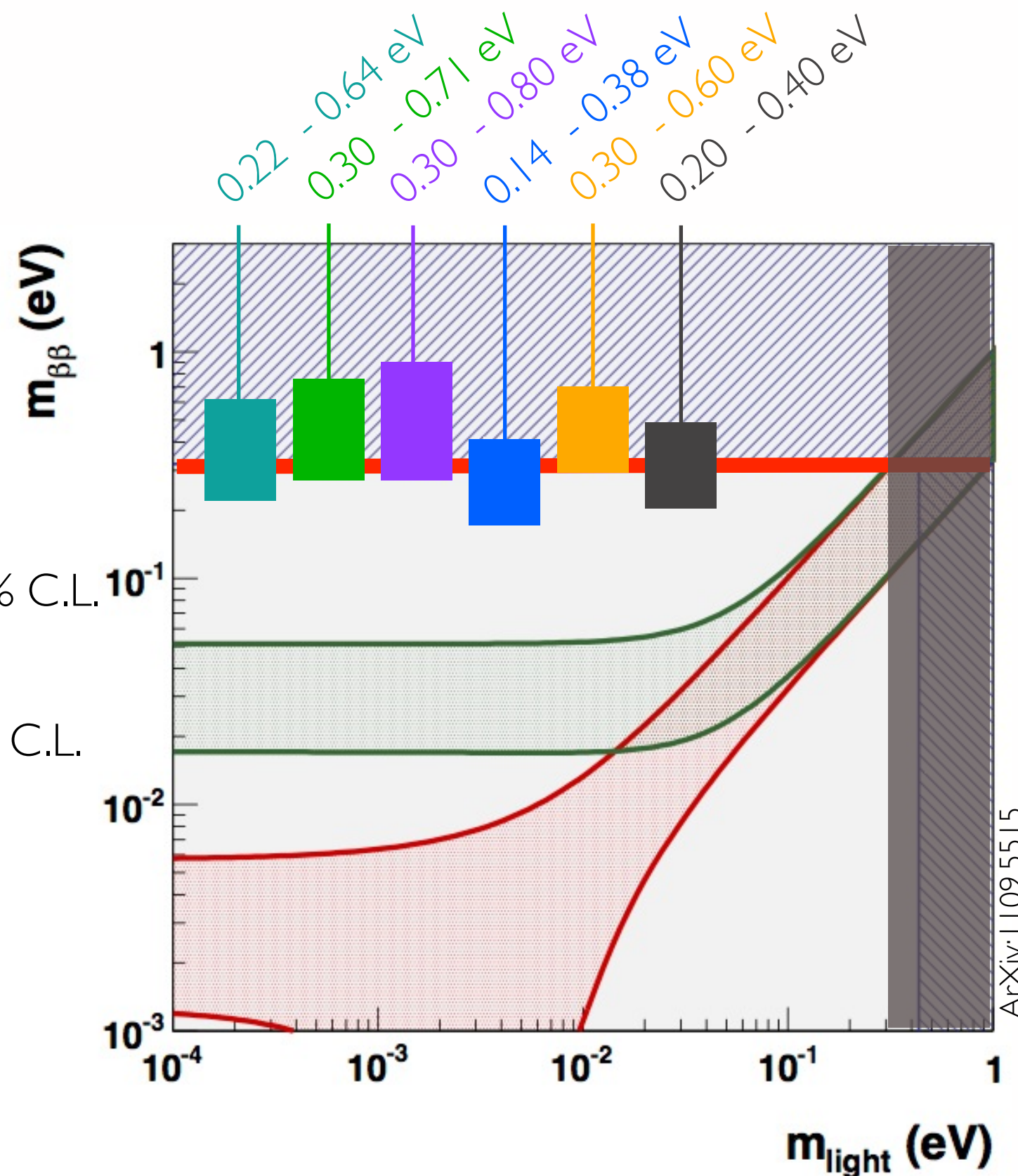
2000 - 2010:

- **Cuoricino**: TeO_2 bolometric detector
 - ~11 kg ^{130}Te : $T_{1/2}^{0\nu} > 2.8 \cdot 10^{24} \text{ y}$ @ 90% C.L.
- **NEMO3**: Tracko-Calo, 7 different isotopes
 - ~7 kg ^{100}Mo : $T_{1/2}^{0\nu} > 1.1 \cdot 10^{24} \text{ y}$ @ 90% C.L.

Since 2011: new generation

10 - 100 kg, R&D for future scaling

- **EXO200** (^{136}Xe): Liquid TPC
- **Kamland-ZEN** (^{136}Xe): Liquid Scintillator
- **GERDA Phase I** (^{76}Ge): Ge diodes



Future projects

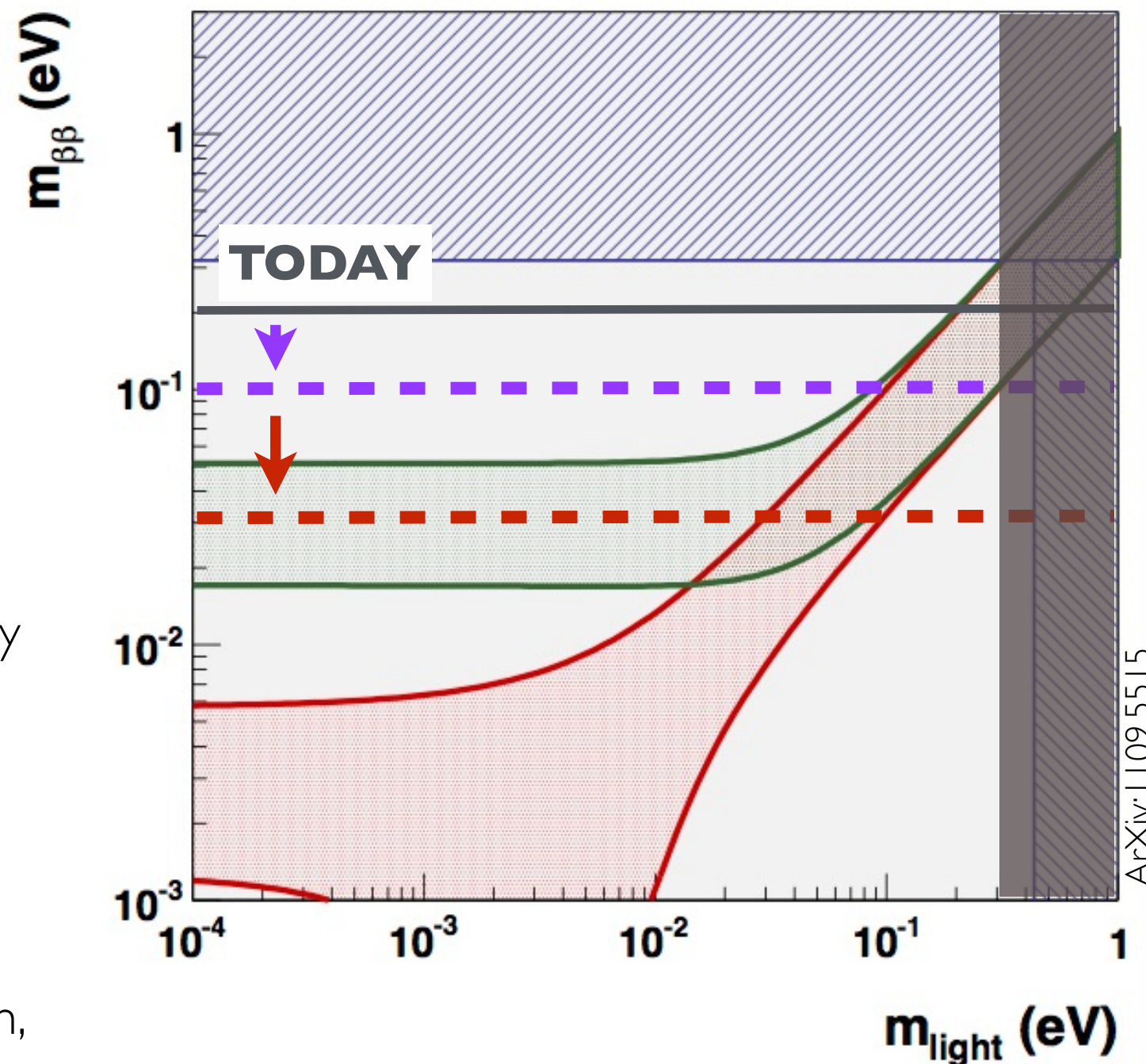
5 years time scale:

- $M \sim 10 - 50$ kg of $\beta\beta$ isotope
- Background level 10^{-3} cts./ (keV kg y)
- Explore quasi-degenerate region

10 years time scale:

- $M \sim 100$ kg - 1t of $\beta\beta$ isotope
- Background level 10^{-4} cts./ (keV kg y)
- Approach Inverse Hierarchy region
- Multi-phase approach: demonstrate scalability to higher mass and background reduction

CUORE, Gerda, Majorana, Lucifer, AMORE, NEXT, COBRA, EXO, SNO+, KamLAND-Zen, Candels, **SuperNEMO**, MOON, DCBA, ...



NEMO3 and the tracko-calorimetry technique

Running @ LSM 2003 - 2011

Wire chamber

Plastic scint.



PMTs

$\beta\beta$ source

Detector $\neq \beta\beta$ source

- **Multiple isotopes** at the same times
- 7 isotopes in NEMO-3

Full reconstruction of $2e^-$ kinematics \rightarrow **unique!**

- Individual e^- energy, arrival time, track curvature in magnetic field, emission vertex and tracks angle

Excellent background rejection

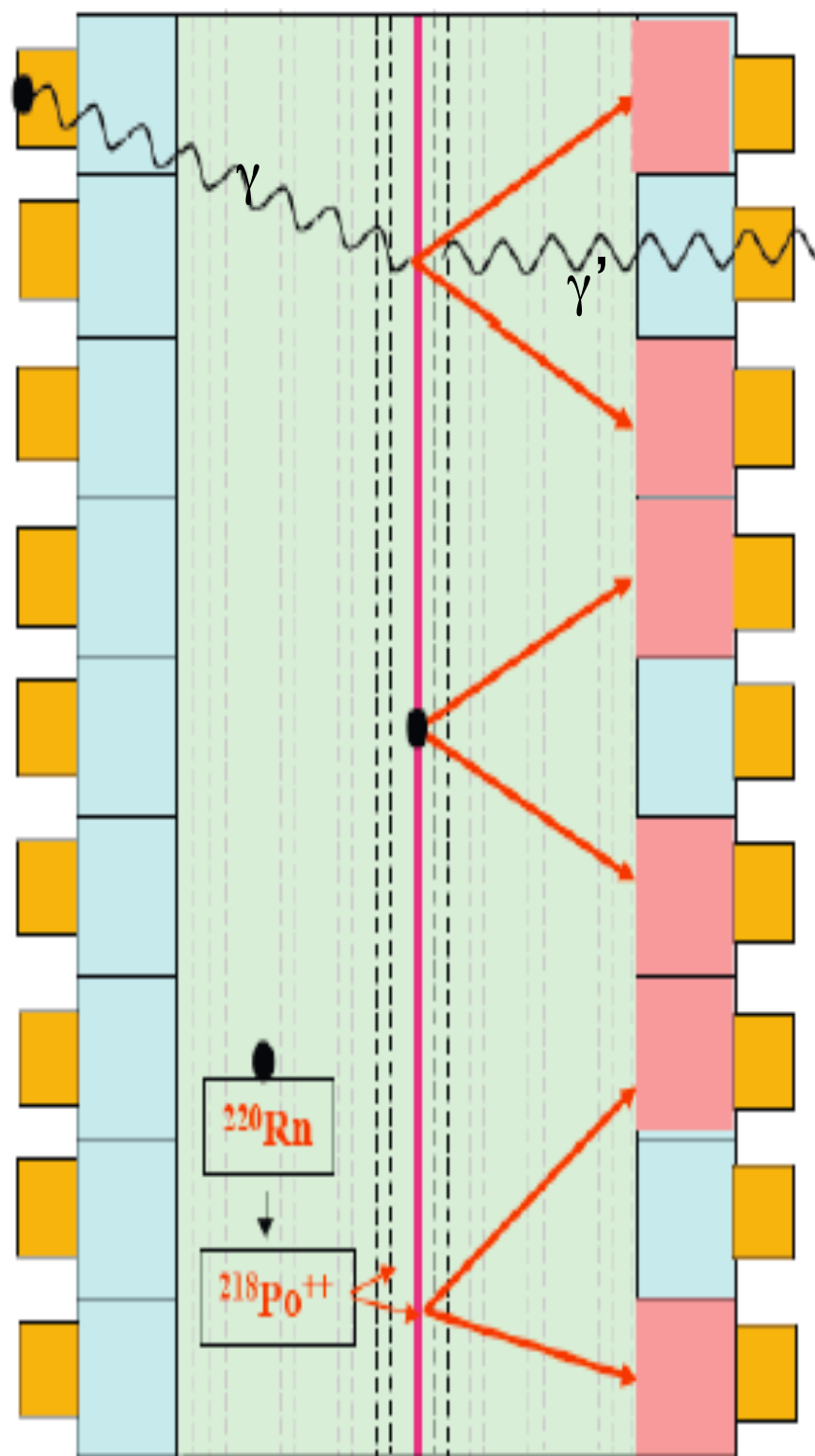
- Identification e^- , e^+ , γ , α

Low energy resolution: $[14 - 17] \% / \sqrt{E}$

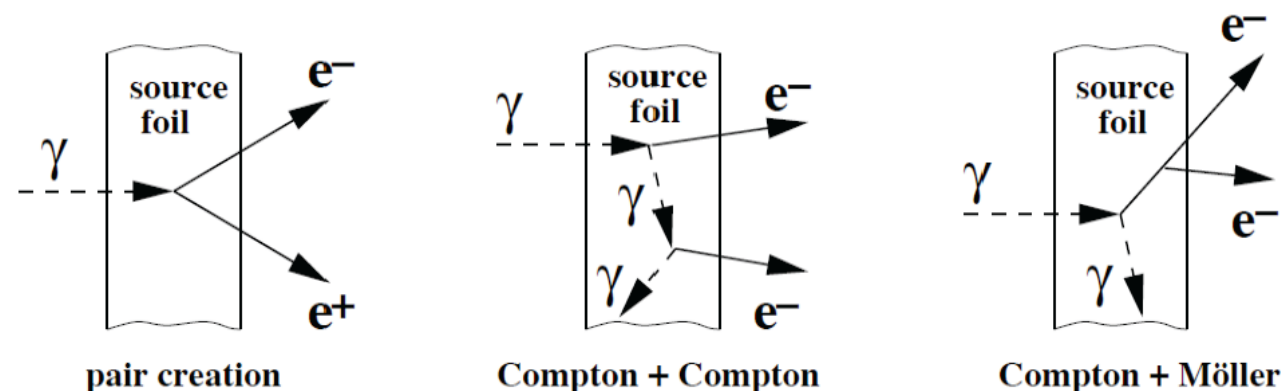
Equivalent to best calorimetric experiment

Discrimination of physics mechanism beyond $\beta\beta 0\nu$!

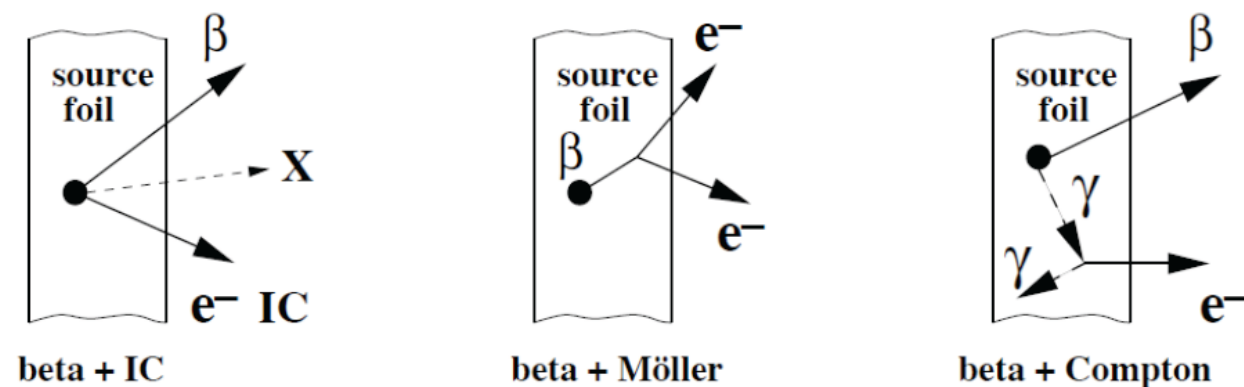
Backgrounds



External background: Radio-impurities in material, γ from (n, γ) and μ bremsstrahlung



Internal background: $\beta\beta 2\nu$ tail, ^{214}Bi (from ^{238}U) and ^{208}Tl (from ^{232}Th) contamination in foil source

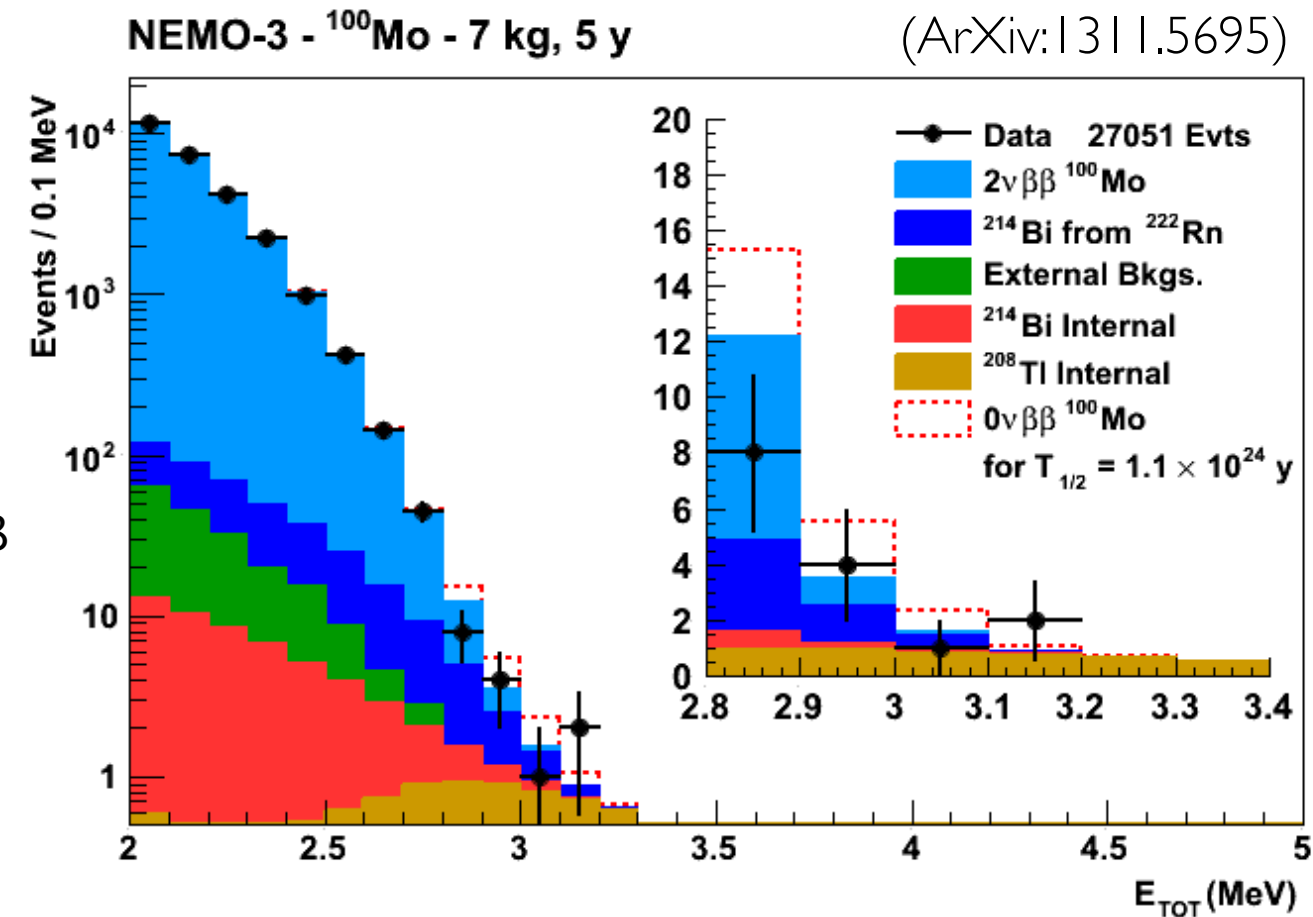


Radon background: ^{214}Bi from Rn decay in tracker volume

Direct measurement of background in different channel combining calo + tracker infos

Latest NEMO-3 results

- Full ^{100}Mo exposure 7 kg, 5 y
- $T_{1/2}^{0\nu} > 1.1 \times 10^{24} \text{ y}$ @ 90% C.L.
- Background level $\sim 0.02 \text{ cts. / (keV kg y)}$ @ $Q_{\beta\beta}$
- No background event $> 3.2 \text{ MeV}$



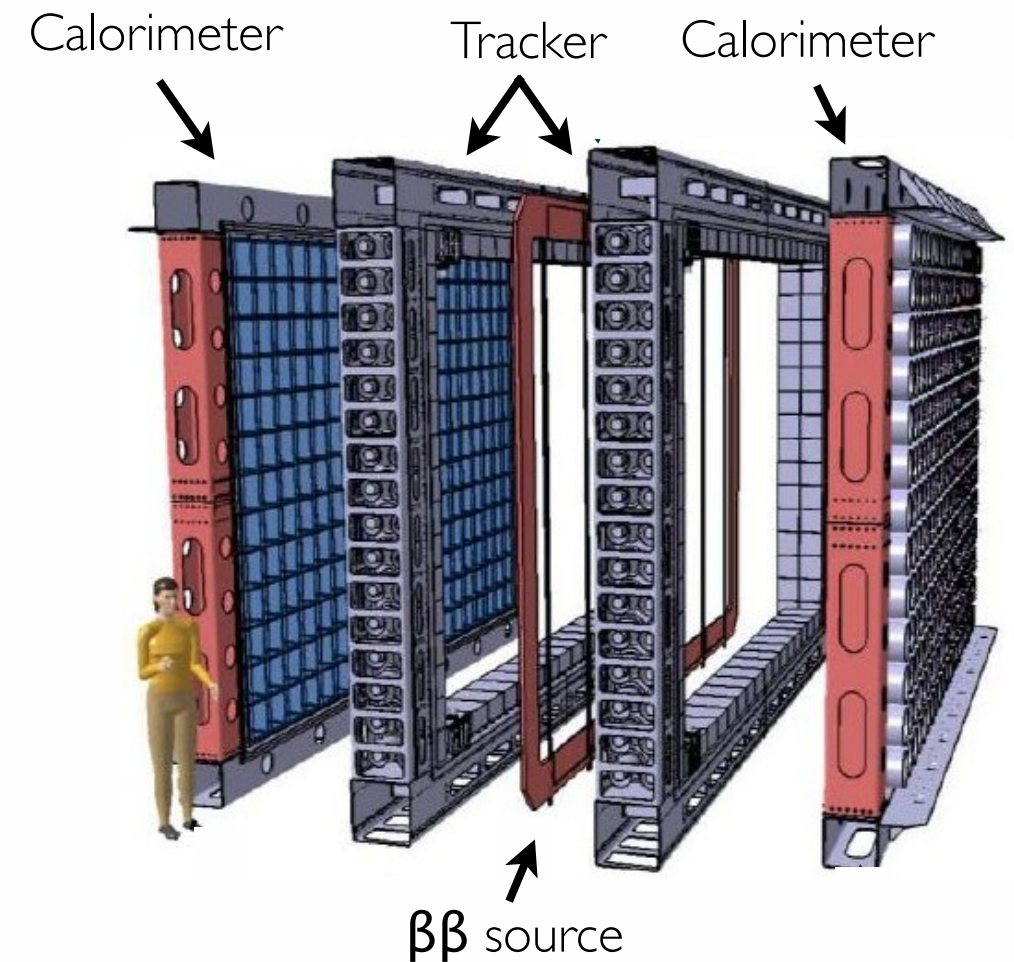
Potential background free technique for high energy $Q_{\beta\beta}$ isotopes (^{48}Ca , ^{150}Nd , ^{96}Zr)

The next step: SuperNEMO

Extrapolate a well known technique:

- 100 kg of $\beta\beta$ emitter in 20 detection module
- Approach IH region

	NEMO-3	SuperNEMO
Efficiency	18%	~30%
Isotope	7 kg ^{100}Mo	~100 kg ^{82}Se (^{150}Nd , ^{48}Ca)
Exposure	35 kg y	~500 kg y
Energy res.	8% @ 3 MeV	4% @ 3 MeV
^{208}Tl (source)	~100 $\mu\text{Bq/kg}$	< 2 $\mu\text{Bq/kg}$
^{214}Bi (source)	~ 300 $\mu\text{Bq/kg}$	< 10 $\mu\text{Bq/kg}$
Rn (in tracker)	5 mBq/m ³	0.15 mBq/m ³
$T_{1/2}$	10^{24} y	10^{26} y
$\langle m_{ee} \rangle$	0.31 - 0.79 eV	0.04 - 0.1 eV



- A **challenge** under **many** aspects
- Extended R&D program in the past years
 - Almost completed!
- Next step → **Demonstrator module**

The demonstrator module

One SuperNEMO module \rightarrow 7 kg ^{82}Se running ~ 2.5 y

- To be installed @ LSM (replacing NEMO-3)

Match SuperNEMO requirements

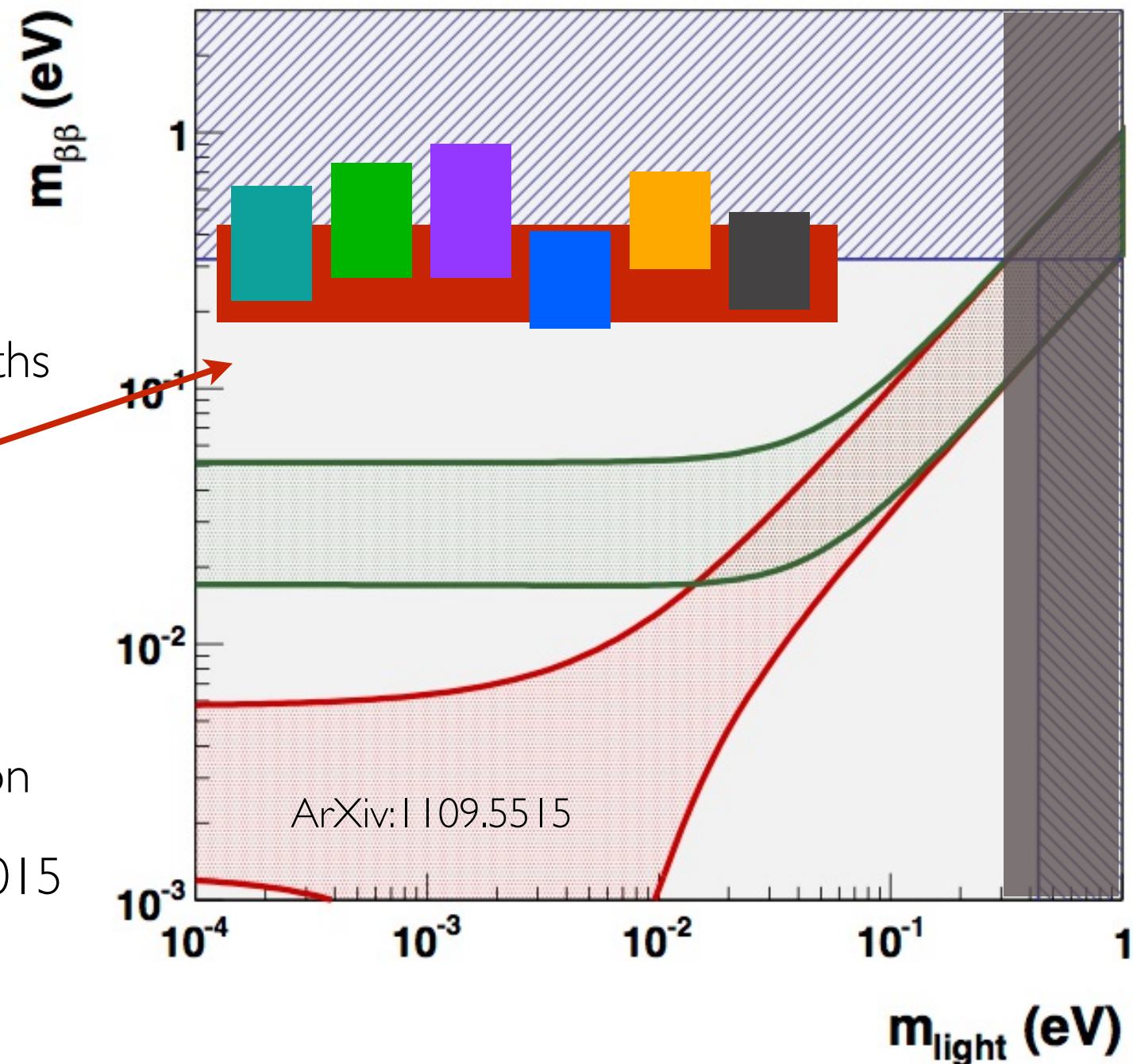
- Background level $\sim 10^{-4}$ cts./(keV kg y)
- Background free at high energy (^{150}Nd)

Reach NEMO-3 (^{100}Mo) sensitivity in 4.5 months

- Sensitivity: $\langle m_{ee} \rangle \sim \mathbf{0.20 - 0.40 \text{ eV}}$
- Test HdM claim with ^{82}Se

Schedule:

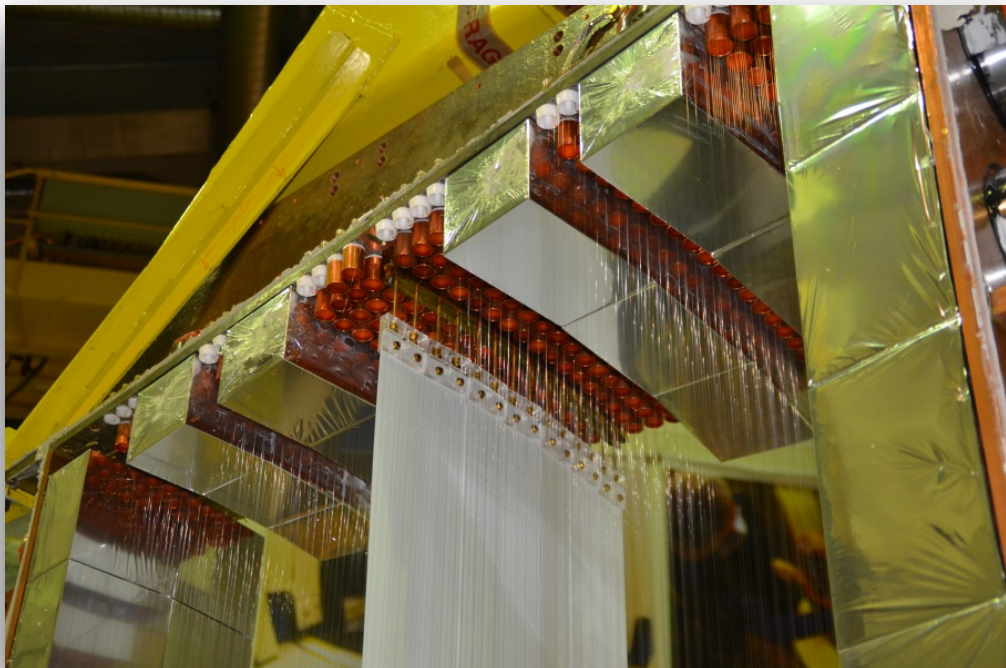
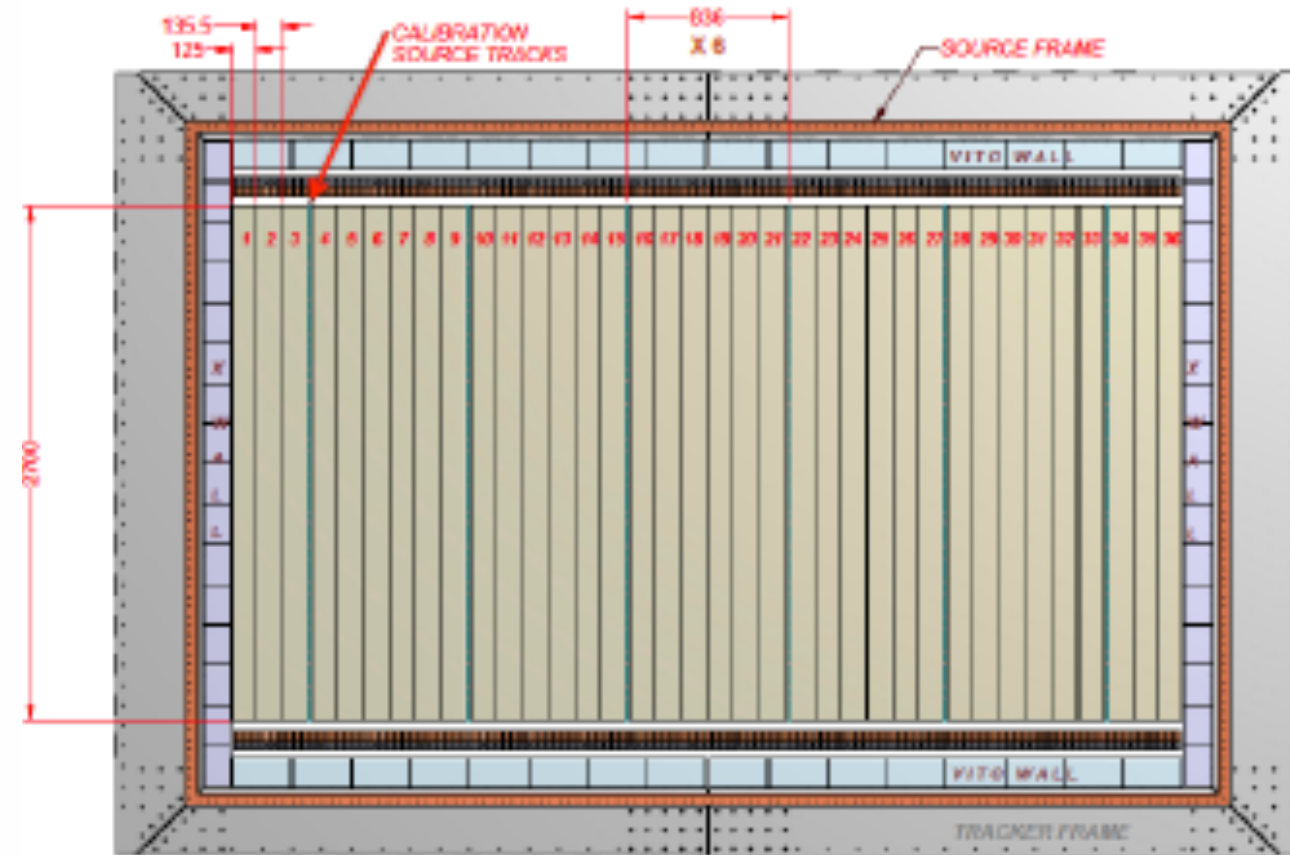
- Calorimeter and tracker under production
- Installation & commissioning beginning 2015
- First **physics data in 2015!**



Developing the foil source @ LAPP

- 36 strips 3 m long, ~200 μm thick ($50 \text{ mg}/\text{cm}^2$)
- Strong material radio-purity constrain:

^{208}Tl	$< 2 \text{ } \mu\text{Bq}/\text{kg}$
^{214}Bi	$< 10 \text{ } \mu\text{Bq}/\text{kg}$



Similar to NEMO-3:

- Enriched Se powder: $> 99 \% ^{82}\text{Se}$
- PVA glue: 5 % - 10 % of Se mass
- **Mechanical support...**
 - Three solutions under consideration

LAPP recently join the collaboration and start R&D activity on the foil source (March 2013)

Developing the foil source @ LAPP

R&D with PVA glue (July - September 2013)

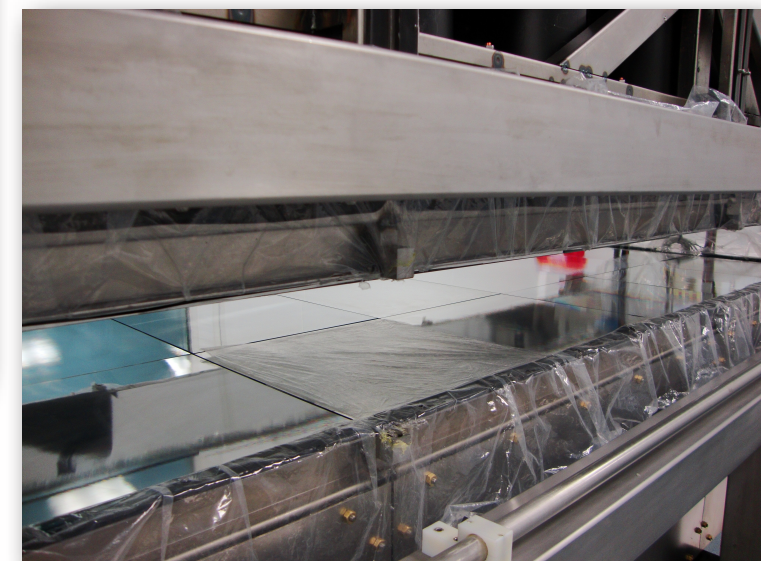
- Defining and training thin foil production technique

Radio-purity measurement

- HGe (LSM): ~ 1.5 kg PVA powder
- BiPo (LSC): 20 thin foil \rightarrow 30x30 cm 200 μm thick



	M [g]	T [d]	$A(^{208}\text{Tl})$ [$\mu\text{Bq/kg}$]	$A(^{214}\text{Bi})$ [$\mu\text{Bq/kg}$]
Ge	1485	26	< 120	< 350
BiPo	210	71	< 32	no yet
Limit	5 (10) % M_{Se}	—	40 (20)	200 (100)

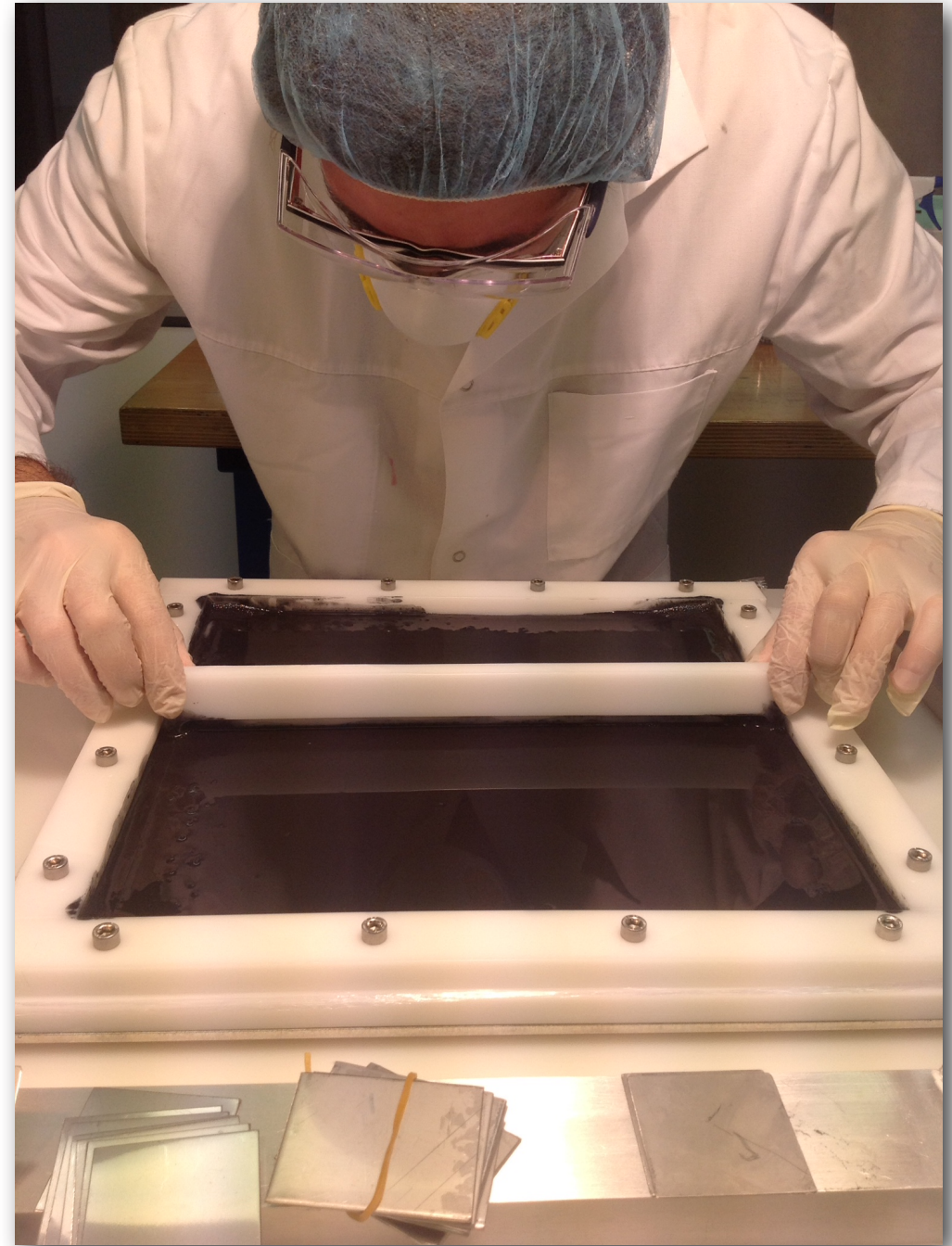
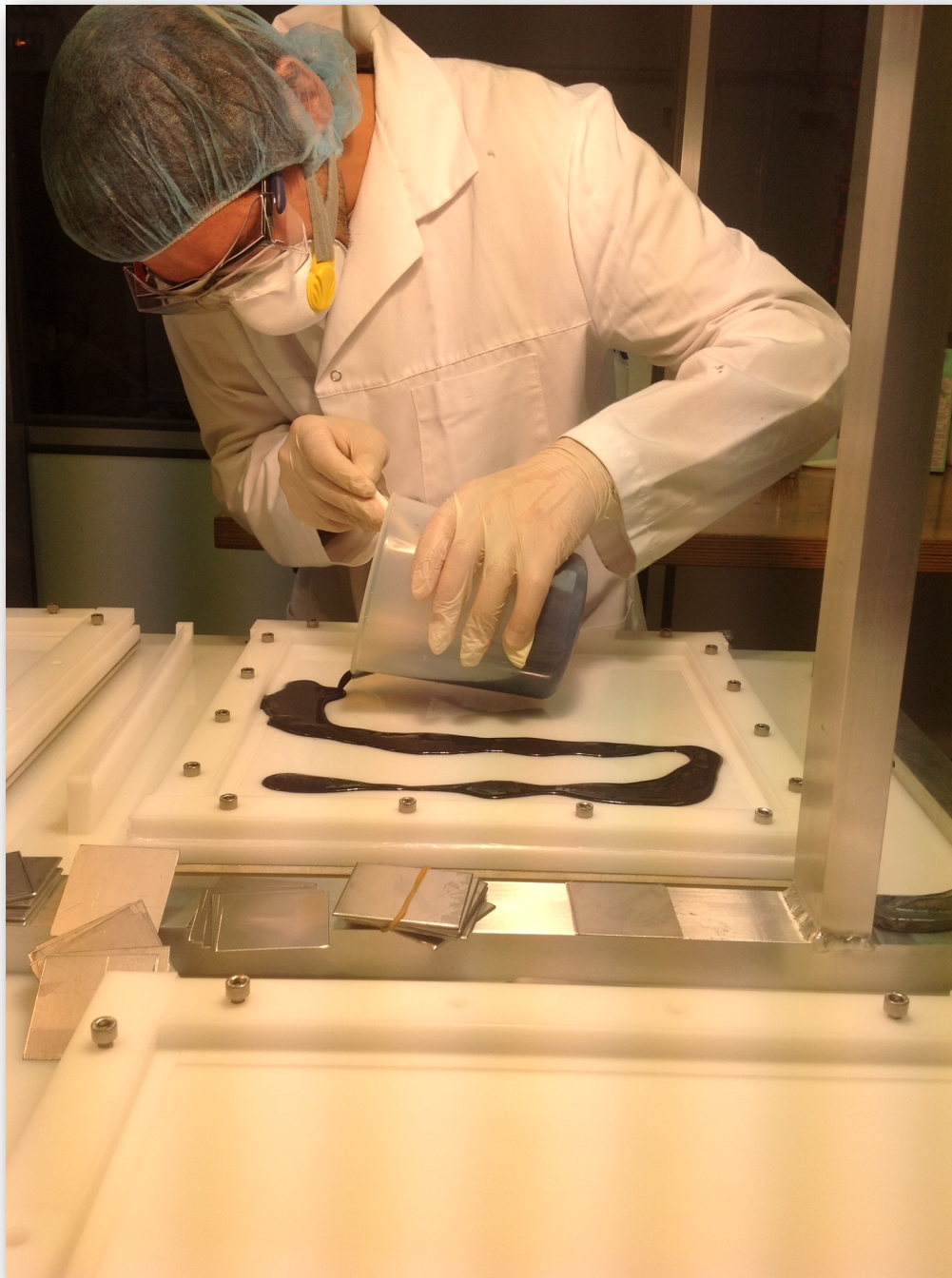


PVA radio-purity is good enough!

Developing the foil source @ LAPP

R&D with Se powder (October 2013)

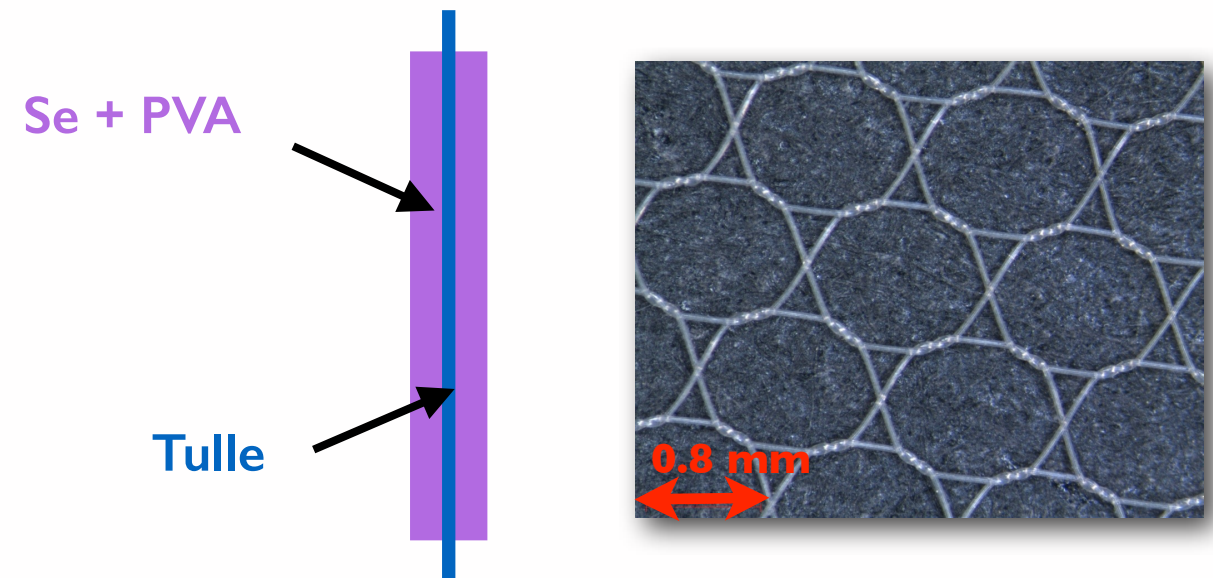
Defining working procedure with Se, approach main problematics



Developing the foil source @ LAPP

R&D for mechanical support (November 2013 - January 2014)

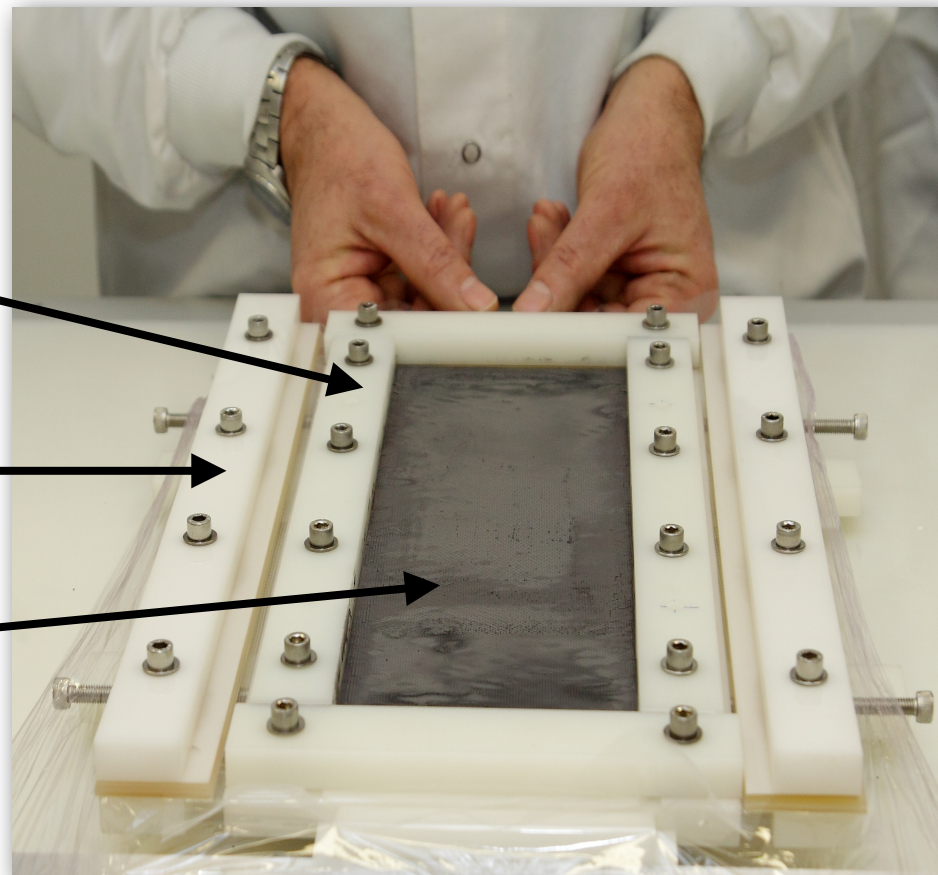
- Fine mesh fabric (Tulle) as central backbone
- New idea proposed by LAPP:
 - Flexible and resistant foil
 - Small support mass $\sim 1\%$ of Se mass
 - High level of radio-purity is expected
- Preparing samples for radio-purity measurement



Derlin support

Tulle tensioning bars

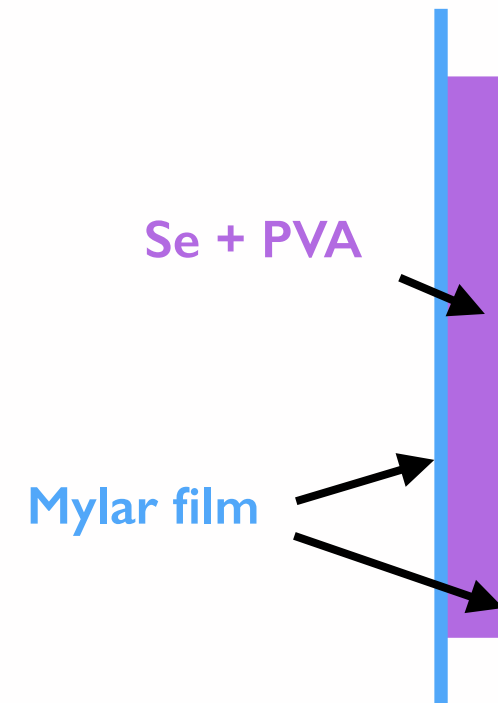
Se + Tulle foil
(30 cm x 12 cm)



Alternative solution under consideration

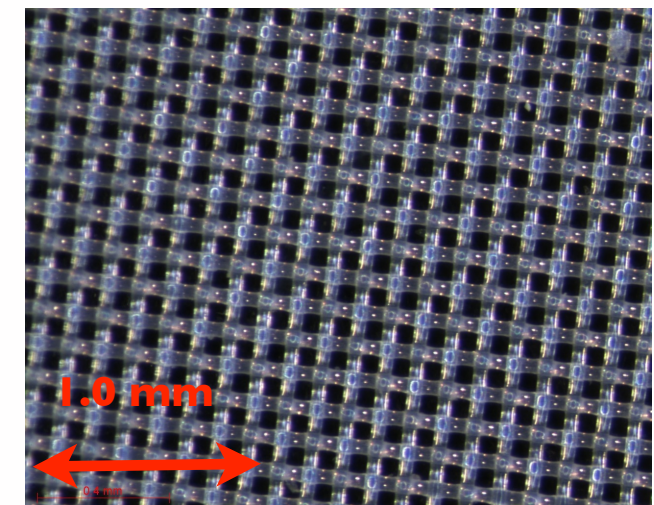
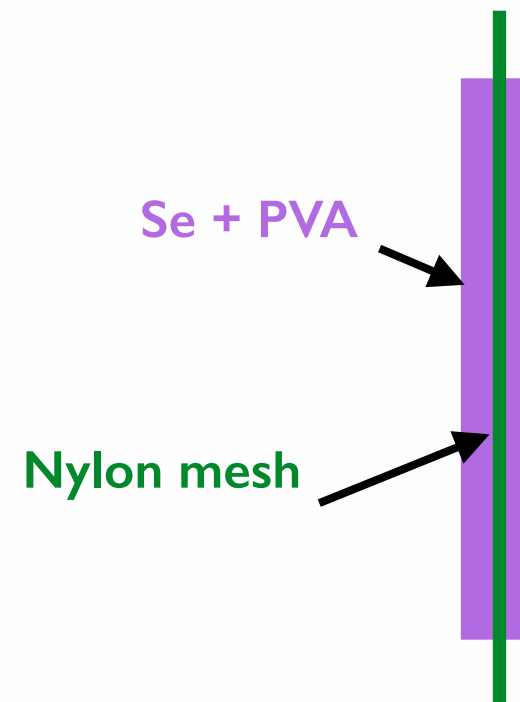
Mylar backing film

- Solution adopted in NEMO-3
- Resistant and flexible foil
- Higher support mass $\sim 10\%$ of Se mass
- Issue with radio-purity [$A(^{208}\text{Tl}) \sim 7\ \mu\text{Bq/kg}$]



Nylon mesh

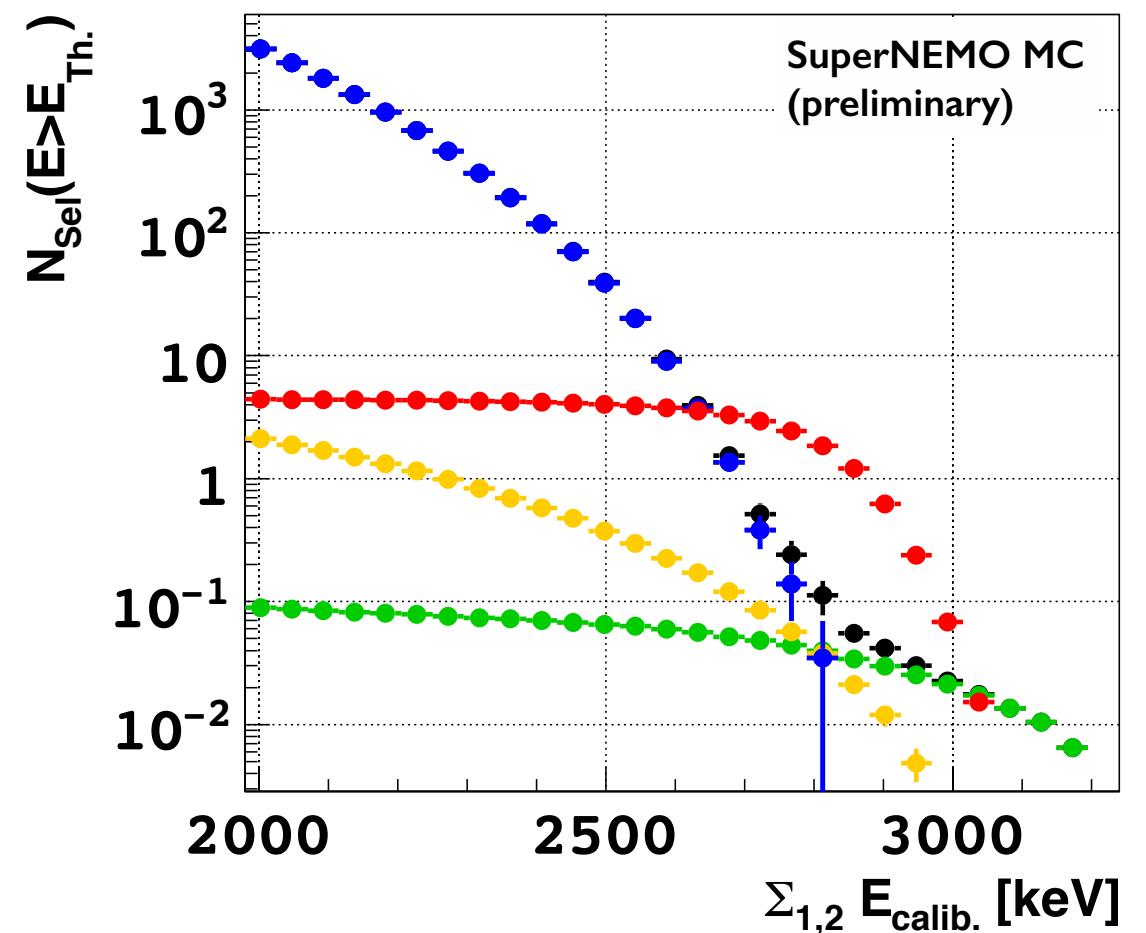
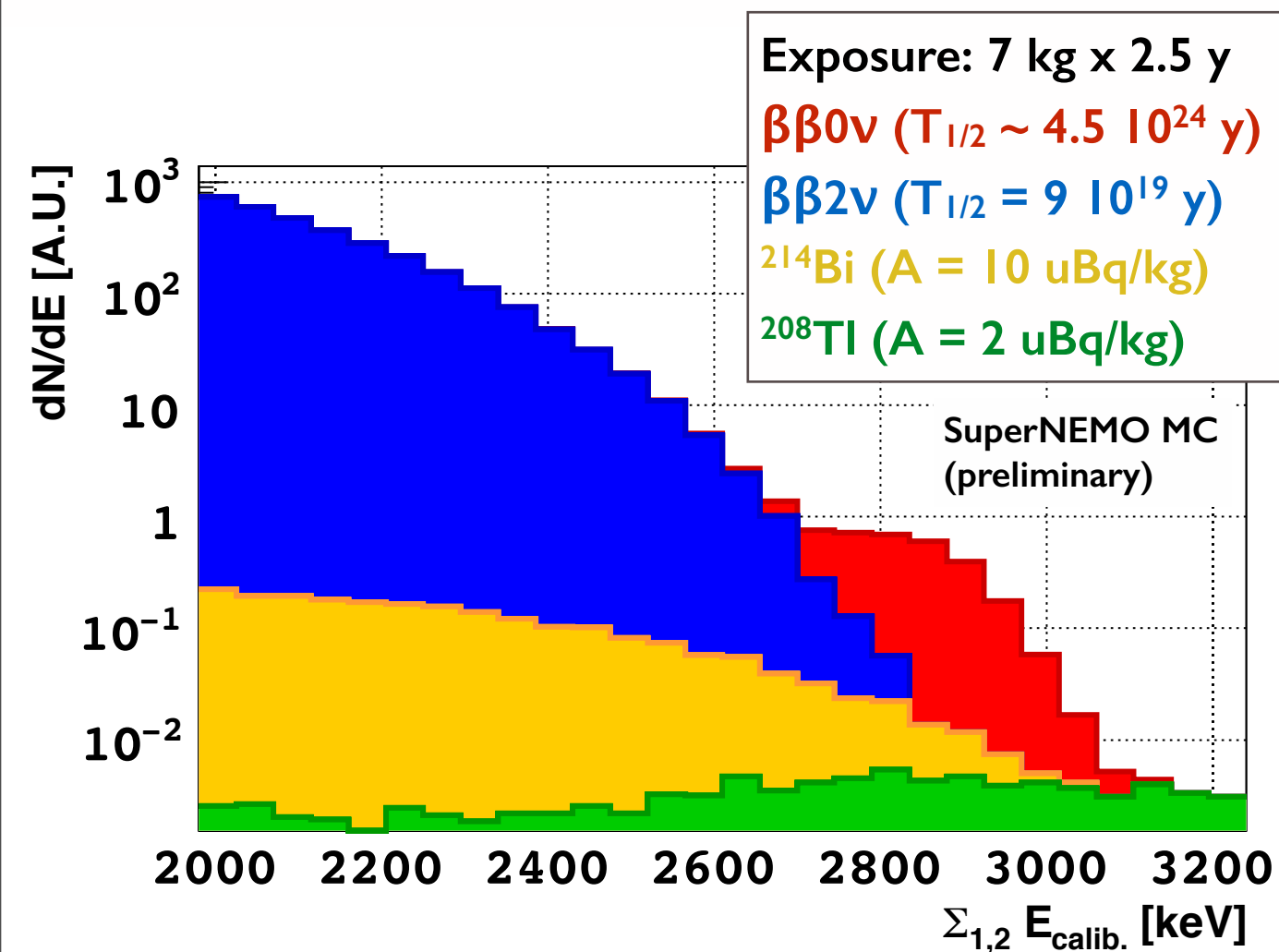
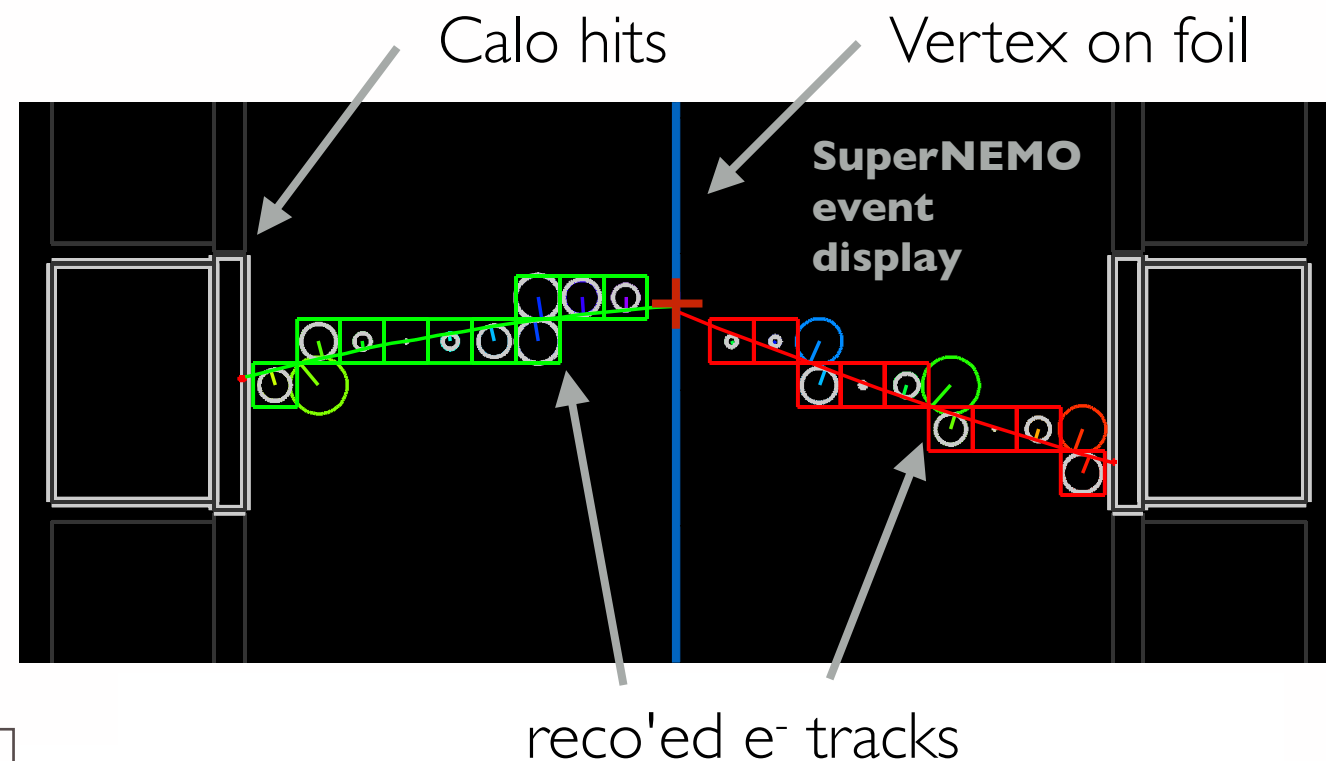
- Used for screen printing and liquid filtering
- Similar to Tulle but more dense
- Higher support mass $\sim 50\%$ of Se mass
- Preparing for radio-purity measurement



Final decision will depend from radio-purity, feasibility and performances

$\beta\beta 0\nu$ Sensitivity study

- Study $\beta\beta 0\nu$ sensitivity w.r.t. foil design
- Check the foil design doesn't alter physics performance and results
- Generating signal and background
- Simple event selection: 2 calo hit associated to 2 negative tracks + Vertex on foil

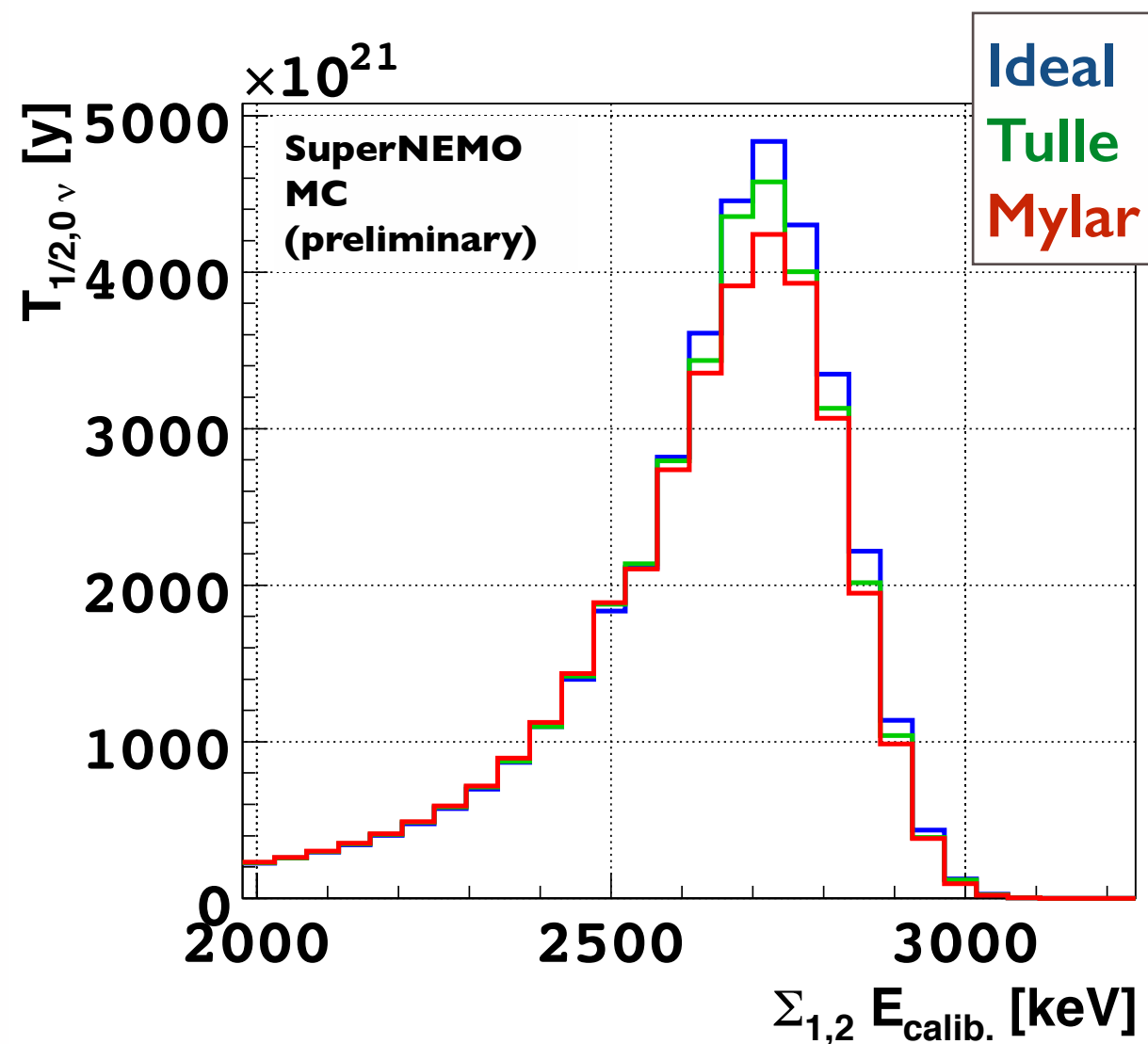


$\beta\beta 0\nu$ Sensitivity study

In hypothesis there is no signal, which the best limit on $\beta\beta 0\nu$ half-life?



$$T_{1/2}^{0\nu} > \frac{\log 2N_A}{W} \times \epsilon^{0\nu} \frac{M \times T}{N_{\text{EXC}}}.$$



- Consider 0ν and 2ν only for the moment
- Study performed as function of E_{cut}

	$T_{1/2}$ [10^{24} y]
Ideal	4.83 (1)
Tulle	4.68 (0.969)
Mylar	4.57 (0.946)

There is a slight preference for the “**tulle**” design...

What next?

Foil R&D @ LAPP:

- Measure radio-purity of tulle → Is it good enough?
- Finalise production protocols → start testing on final dimensions (3 m)
- Test radio-purity on final foil design in BiPo (natural & enriched Se)

Foil design optimisation studies:

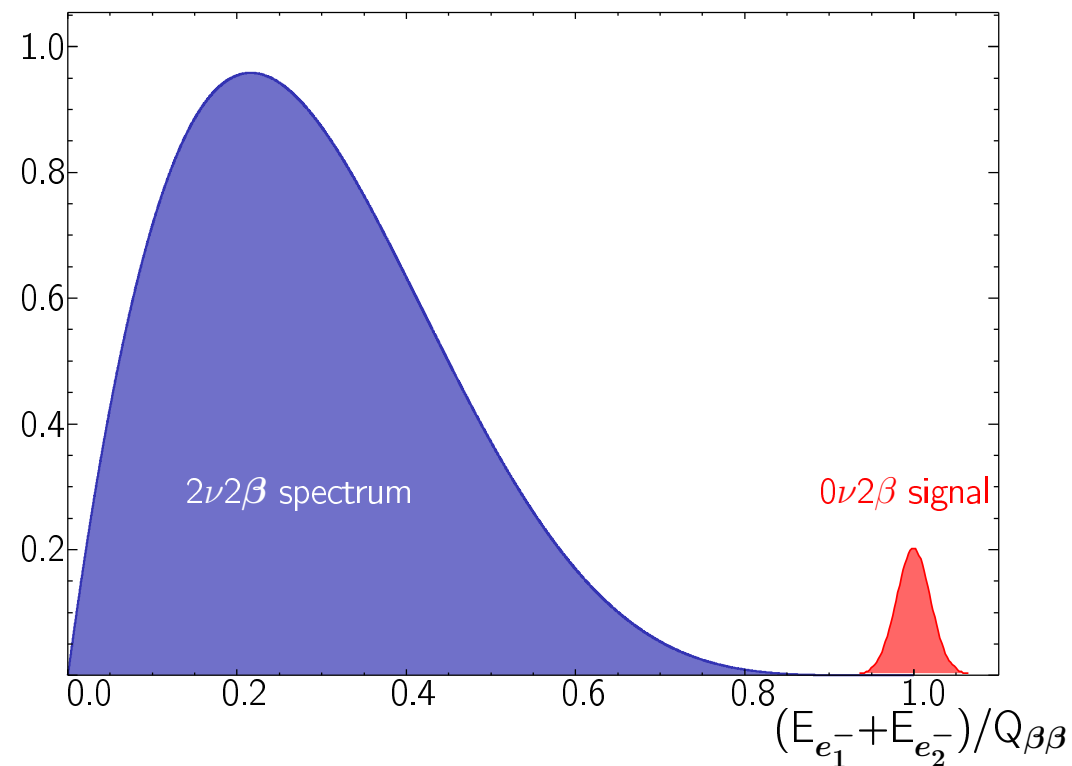
- Account for internal backgrounds and material radio-purity level
- Setup a complete sensitivity study → Master 2 project from March
- Provide the best foil design choice w.r.t. detector performance

Prospectives

- Wide interest on $0\nu\beta\beta$ searches \rightarrow answer a fundamental question
- SuperNEMO aims to IH region with a Tracko-Calo detector \rightarrow Unique!
- The Demonstrator module is the first step \rightarrow Match SuperNEMO requirements
- Sensitivity of $\sim 0.20 - 0.40$ eV in 2 years
- Tracker & Calorimeter under production
- **Installation & commissioning in early 2015**
- **Data analysis shortly after \rightarrow Results in 2017**

Backup slides

Searching for $0\nu\beta\beta$ process



Measure the 2 e^- energy spectrum

- $2\nu\beta\beta$ signature \rightarrow Broad spectrum
- $0\nu\beta\beta$ signal signature \rightarrow **Peak** @ $Q_{\beta\beta}$

If no signal \rightarrow set a **limit** on half life

$$T_{1/2}^{0\nu} > \frac{N_A \ln 2}{n_\sigma} \times \frac{\epsilon}{A} \times \sqrt{\frac{M \times t}{B \times \Delta E}}$$

Detection efficiency $\rightarrow \epsilon$
 $\beta\beta$ emitter mass $\rightarrow M$
 Exposure time $\rightarrow t$
 Excluded events at a given C.L. $\rightarrow n_\sigma$
 Atomic mass $\rightarrow A$
 Bkg. index $\rightarrow B$
 E res. @ $Q_{\beta\beta}$ $\rightarrow \Delta E$

Choosing the $\beta\beta$ isotopes

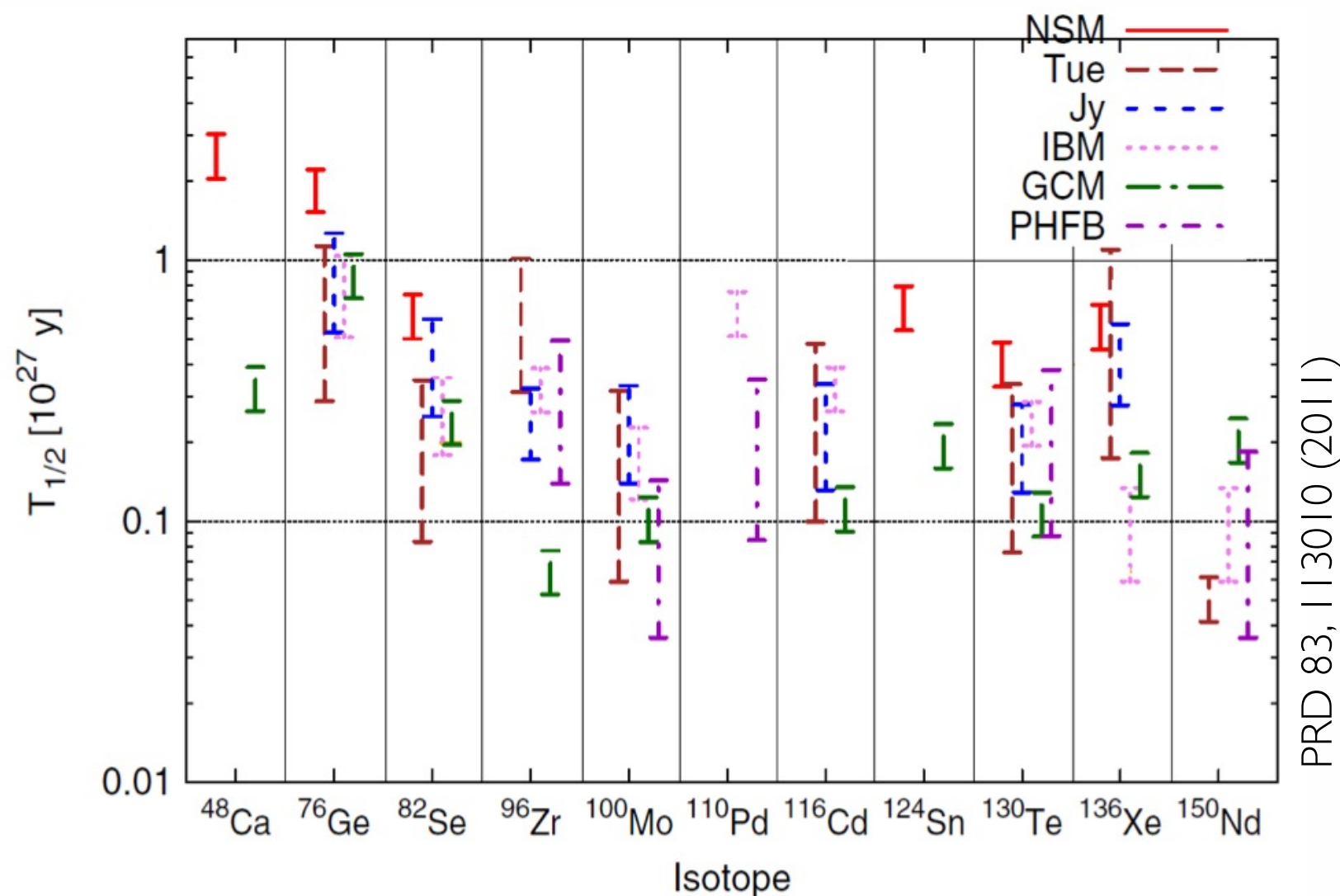
Isotopes enrichment and $T_{1/2}^{2\nu}$ from respective experiment

Isotope	$Q_{\beta\beta}$ [keV]	Nat. abund. (enrich.) [%]	$G_{0\nu}$ [10^{-14} y^{-1}](*)	$T_{1/2}^{2\nu}$ [10^{19} y]	Experiment
^{48}Ca	4270	0.187 (73)	6	$4.2^{+2.1}_{-1.0}$	NEMO3
^{76}Ge	2039	7.8 (86)	1	150 ± 10	HM
^{82}Se	2995	8.7 (97)	3	9.0 ± 0.7	NEMO3
^{96}Zr	3350	2.8 (57)	6	2.0 ± 0.3	NEMO3
^{100}Mo	3034	9.6 (99)	4	0.71 ± 0.04	NEMO3
^{116}Cd	2802	7.5 (93)	5	3.0 ± 0.2	NEMO3
^{130}Te	2527	34.5 (90)	4	70 ± 10	NEMO3
^{136}Xe	2480	8.9 (80)	4	238 ± 14	KamlandZEN
^{150}Nd	3367	5.6 (91)	19	0.78 ± 0.7	NEMO3

Nuclear Matrix Element

Contain nuclear structure effects → only **approximative** theoretical calculation

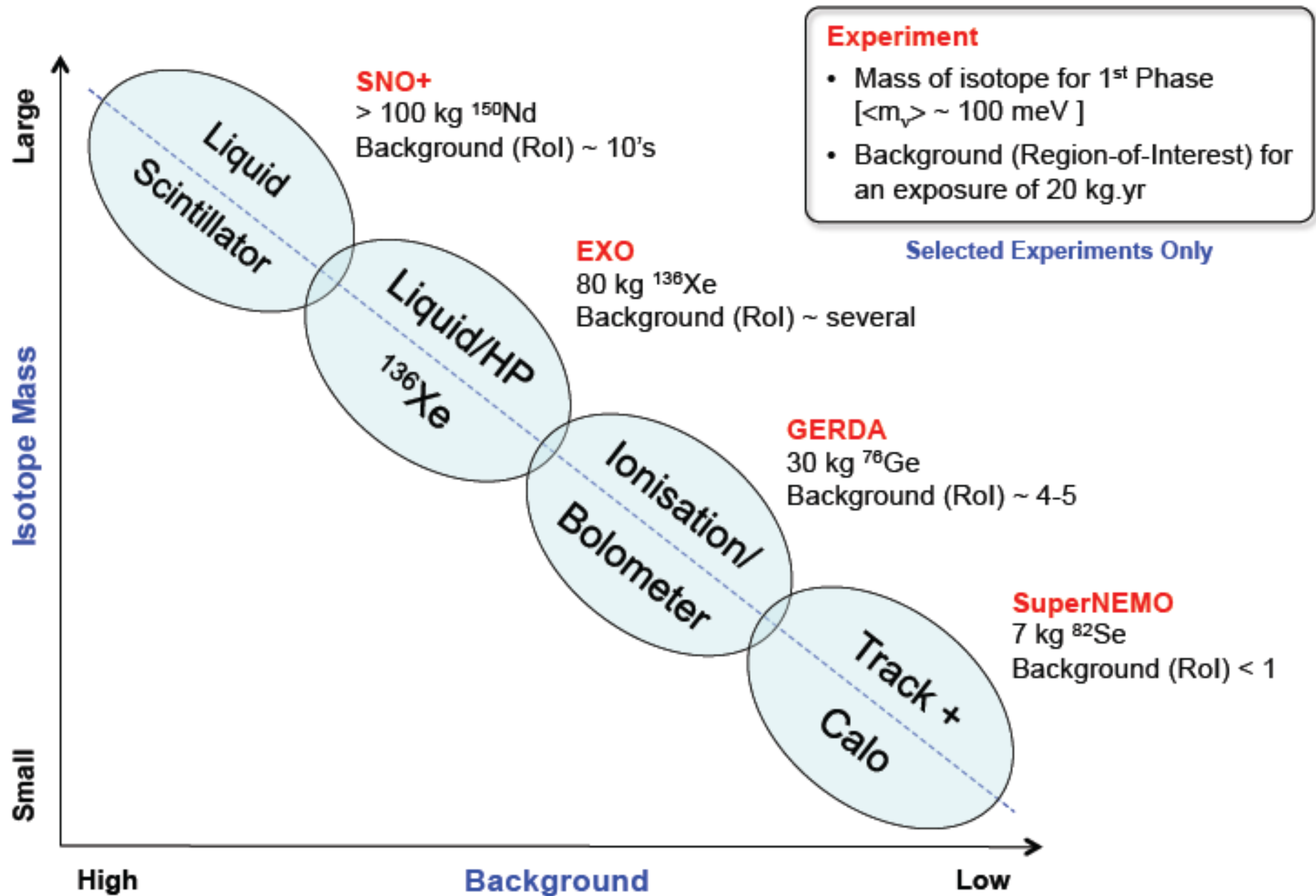
Required $T_{1/2}^{0\nu}$ sensibility for $\langle m_{ee} \rangle \sim 0.05$ eV (IH)



- Many approximation methods
- Variation up to factor ~ 10
- Different among isotopes
- Up to factor 10 on required mass! (^{150}Nd , ^{100}Mo w.r.t. ^{76}Ge)

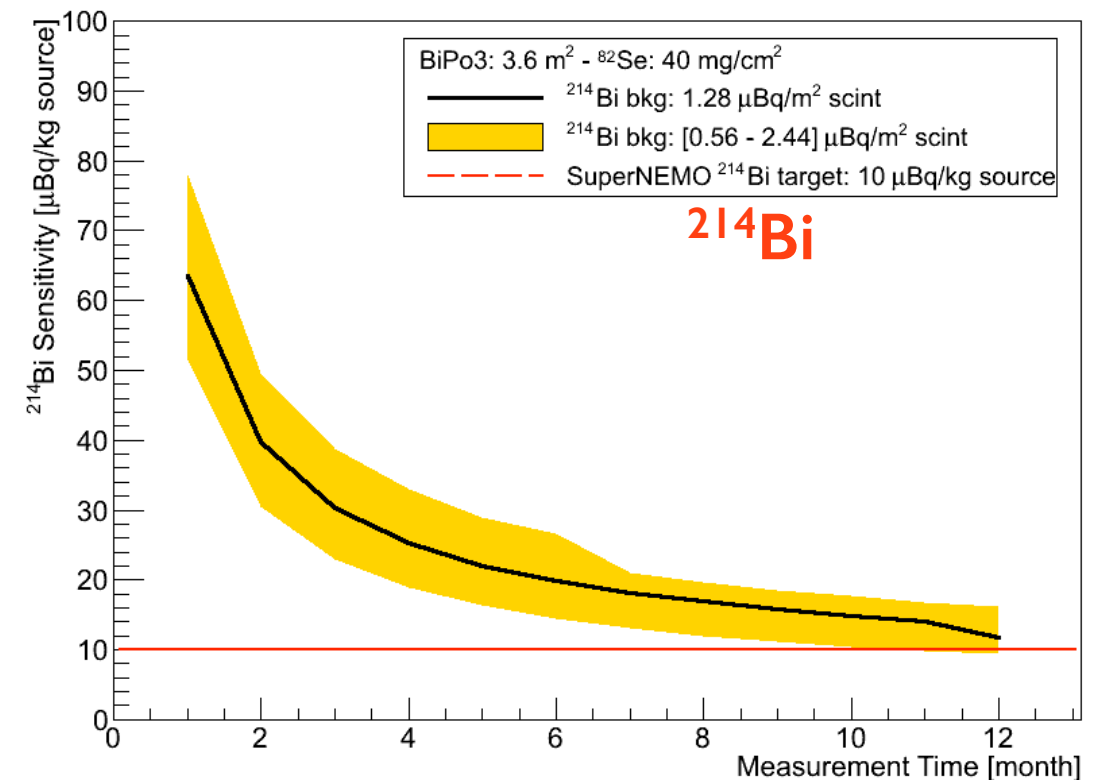
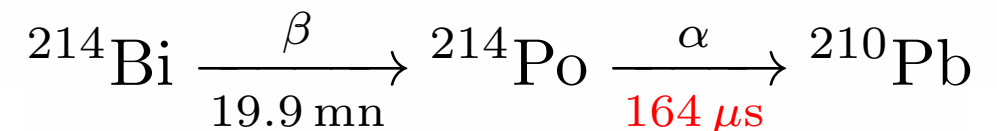
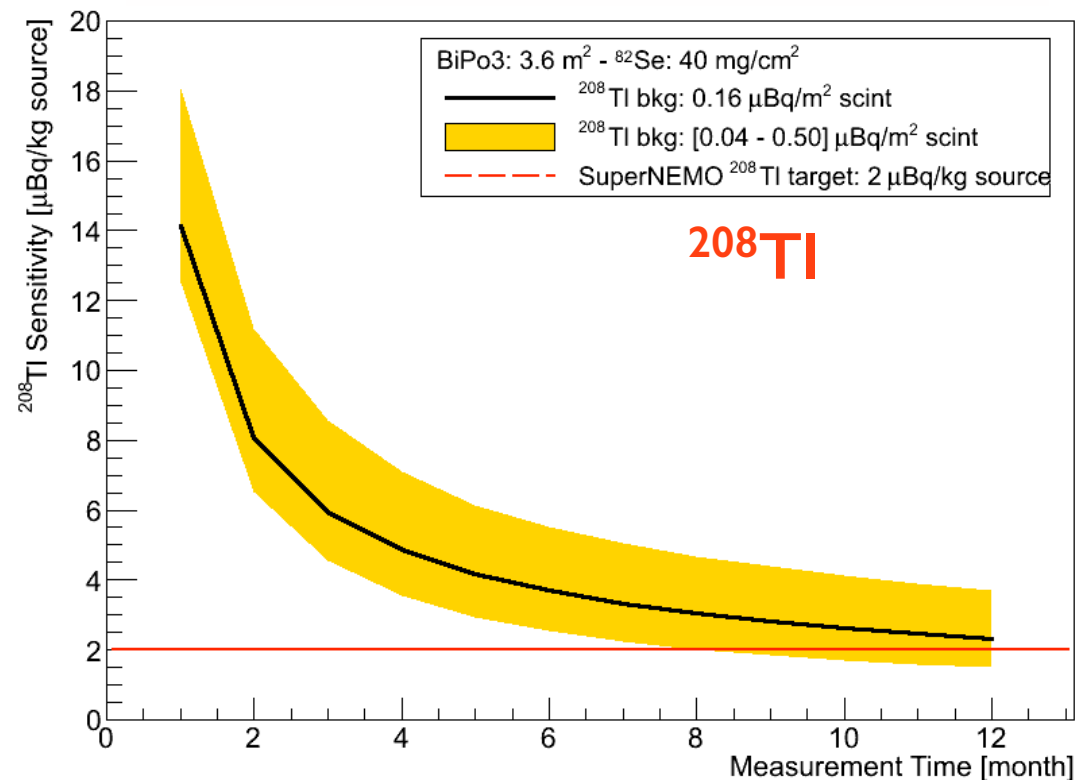
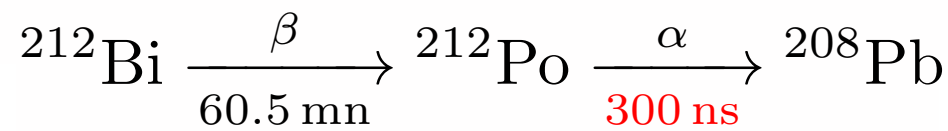
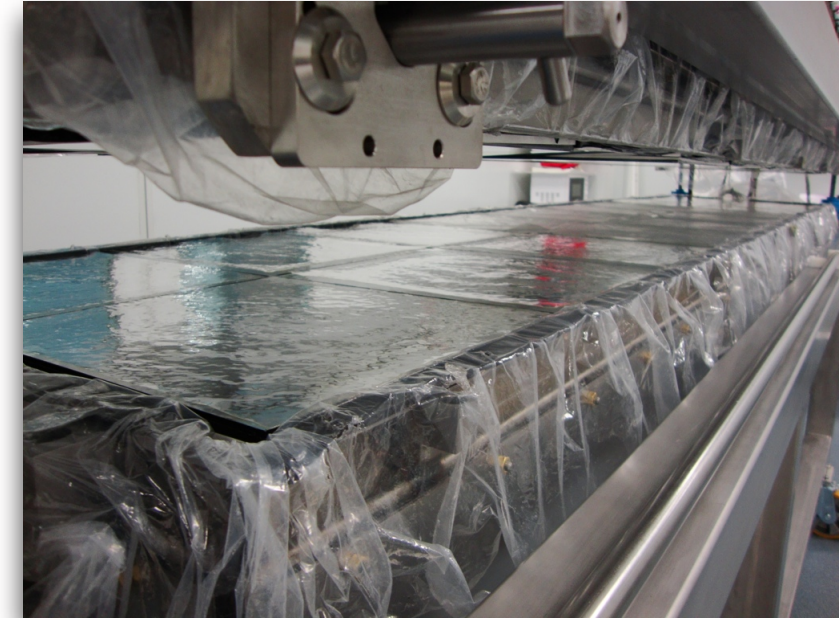
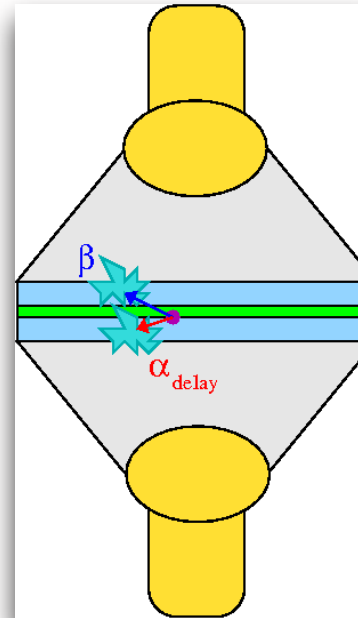
Main **limitation** in interpreting result & comparing among different isotopes

Comparison of Current Techniques



Radio-purity measurement with BiPo

- HGe sensitivity is not enough for SuperNEMO
 - ~ 1 mBq/kg in 1 month measurement
- Dedicated detector measuring the Bi-Po chain
 - Up and running since July 2012 in Canfranc
 - ~ 10 uBq/kg in 2 months measurement (^{208}Tl)



SuperNEMO demonstrator: status

Most R&D completed

Radio purity measurements of materials ongoing

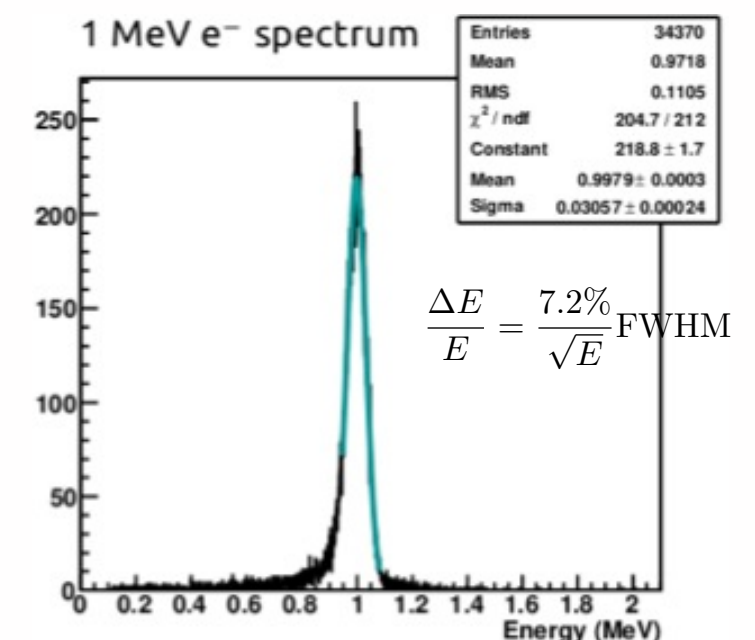
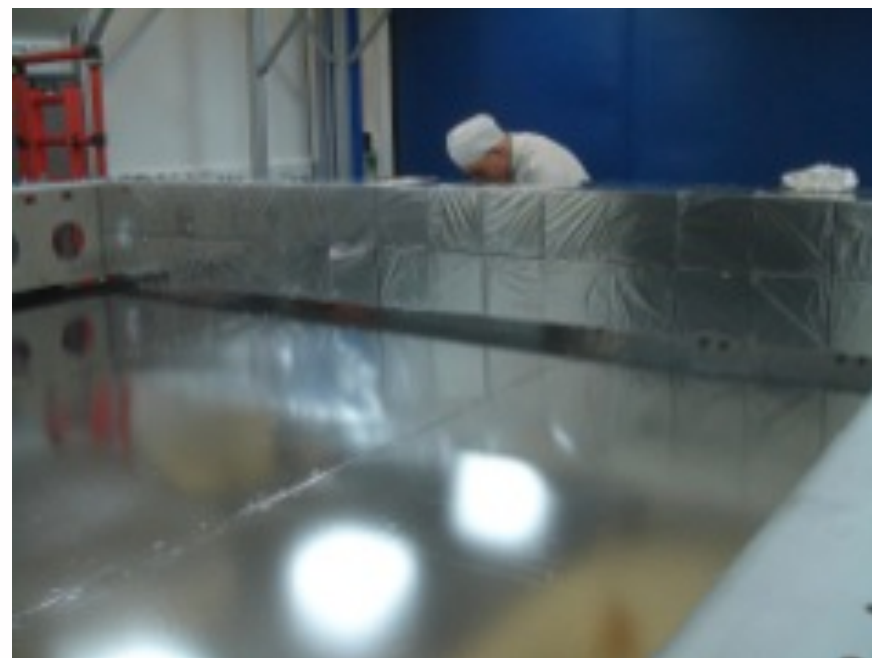
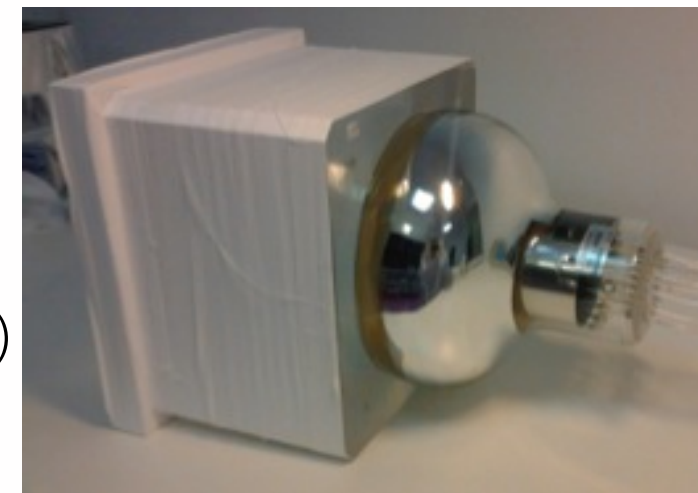
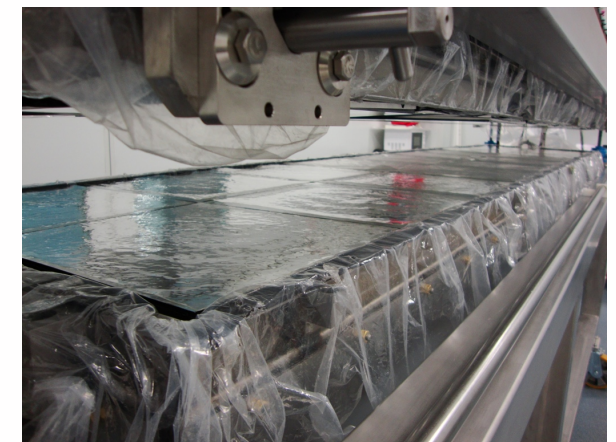
- HPGe, Radon emanation chamber, BiPo

Calorimeter: main wall under construction

- Scintillator block under production
- 8" Hamamatsu PMT by February 2014
- Electronic (FE digitiser & trigger board) under production

Tracker: Under construction.

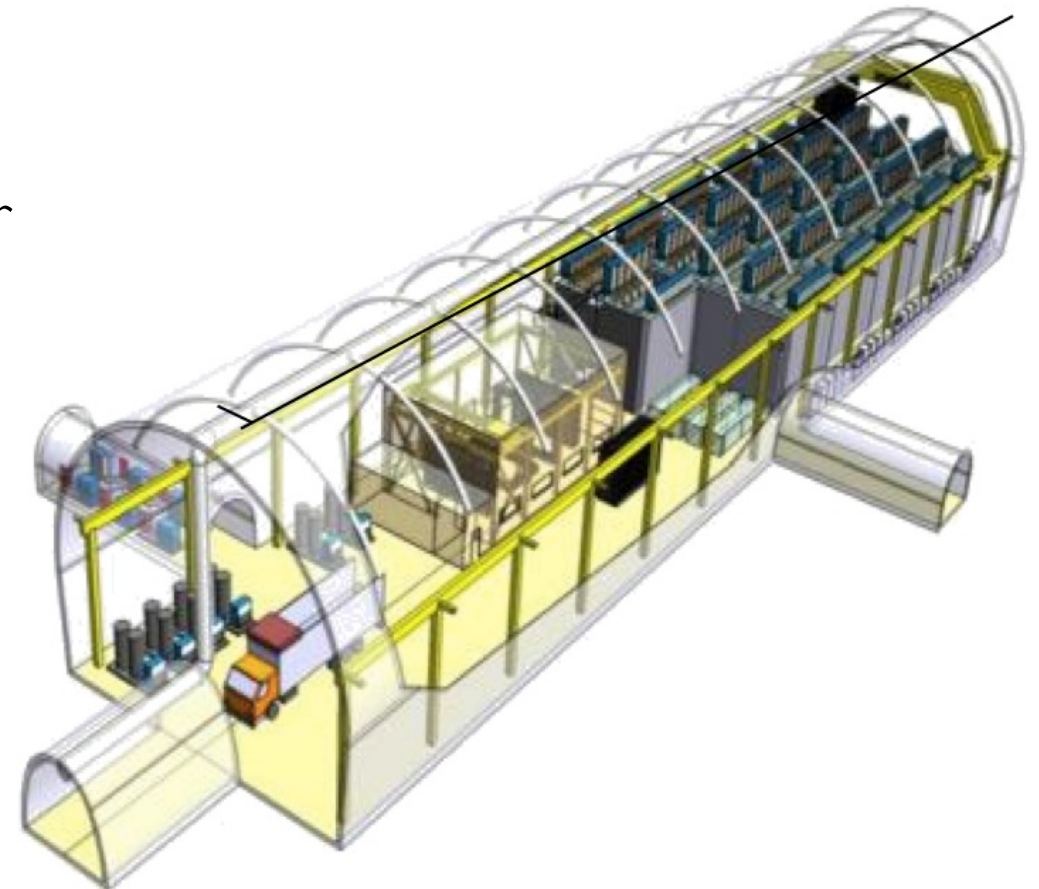
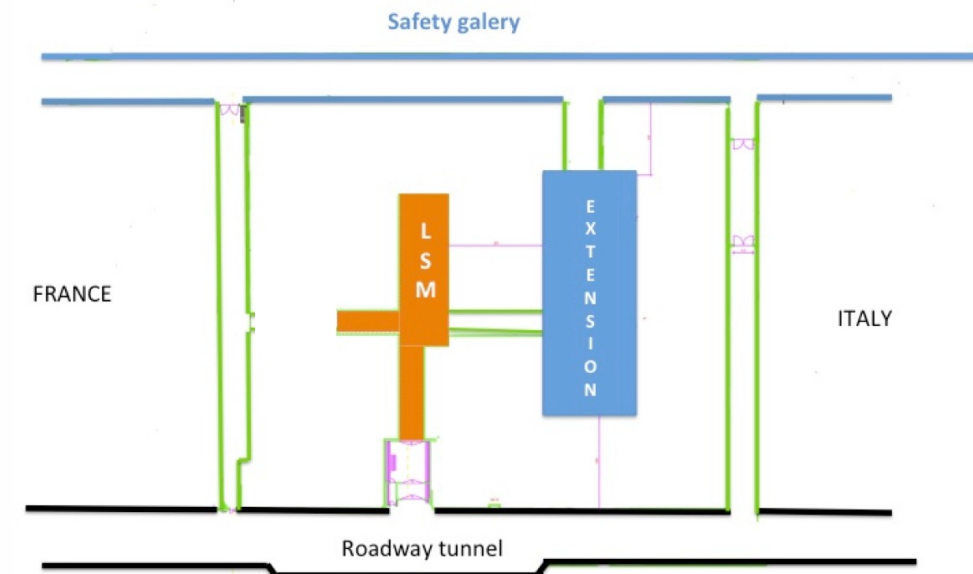
- Commissioning begin 2014 (surface), end 2014 (underground)



LSM Extension: DOMUS

From F. Piquemal (TAUP2013)

- Project accepted by Ministry of Research and programmed by CNRS
- Cost estimated by project supervisor of safety gallery :
- 7 M€ including 20% hazards
- Funding secured from CNRS, Rhône-Alpes region, Savoie department, FEDER
- Technical studies completed
- Negotiations with the civil work company in October
- Digging Spring 2014 or end 2015 depending of the company schedule
- 6 months for excavation, 10 months for outfittings
- Extension in operation 2016 -2017

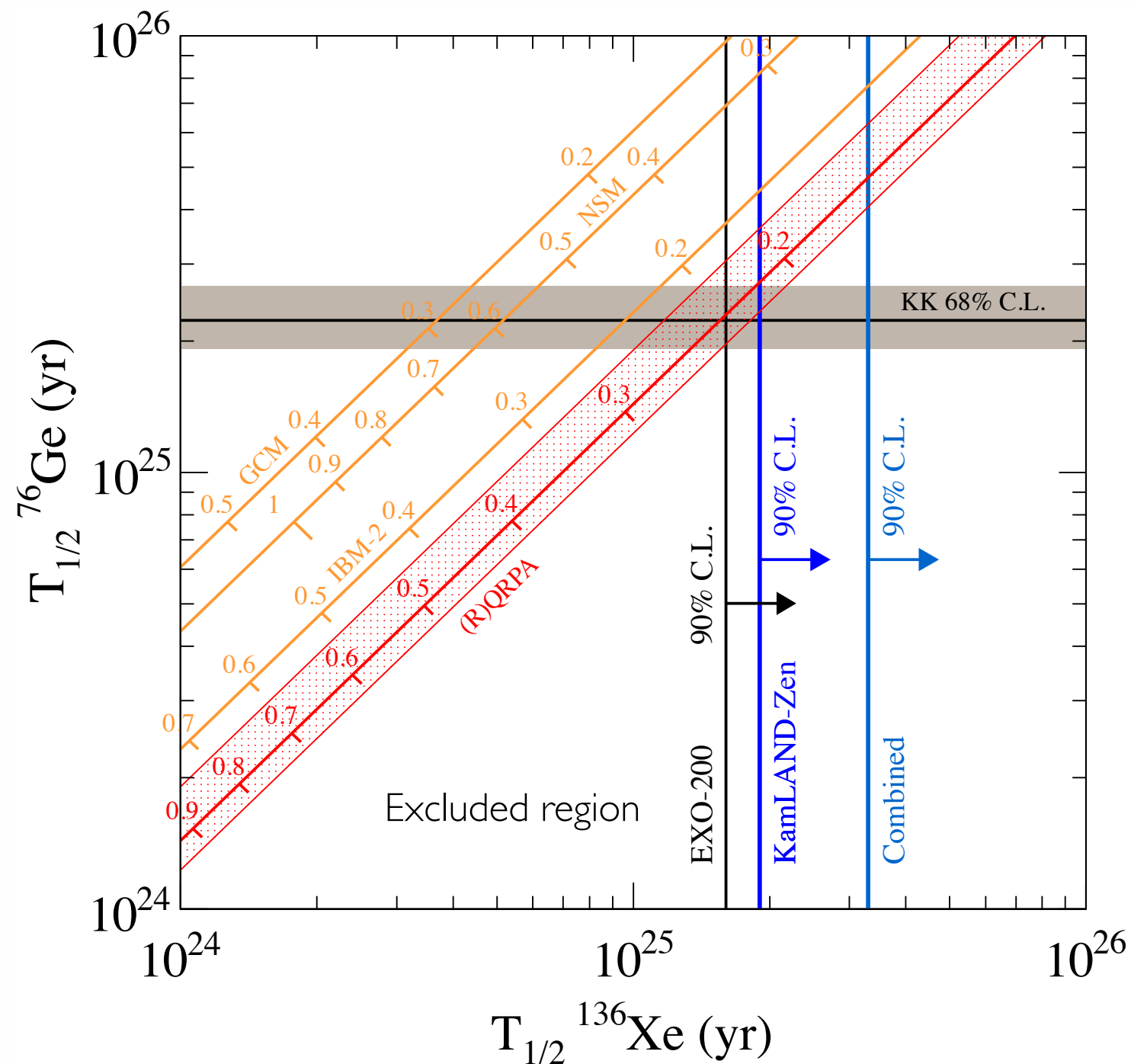


Developing the foil source (2)



Setting up a test area in ISO 5-6 clean room (May - June 2013)

KamLAND-Zen & EXO results



EXO-200

$$T_{1/2,0\nu} > 1.6 \times 10^{25} \text{ y}$$

$$\langle m_{\beta\beta} \rangle < 0.14 - 0.38 \text{ eV} \quad (90\% \text{ C.L.})$$

KamLAND-Zen

$$T_{1/2,0\nu} > 1.9 \times 10^{25} \text{ y}$$

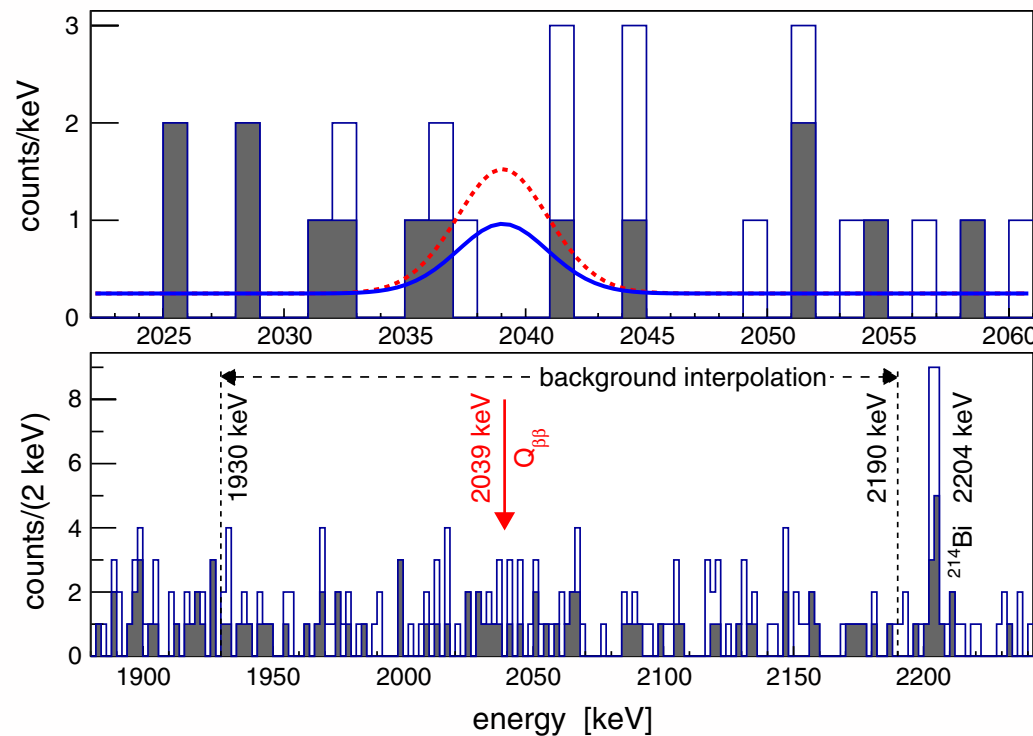
Combined

$$T_{1/2,0\nu} > 3.4 \cdot 10^{25} \text{ yr}$$

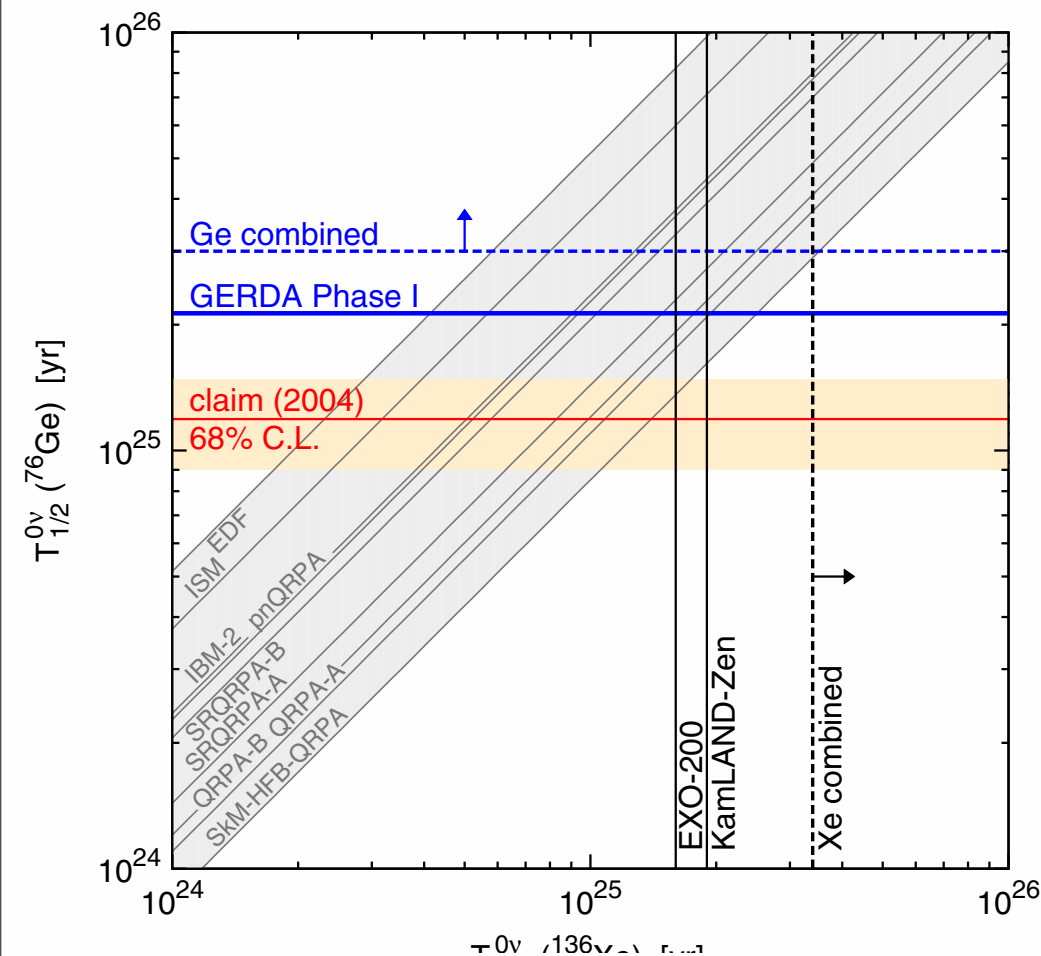
$$\langle m_{\beta\beta} \rangle < 0.12 - 0.25 \text{ eV} \quad (90\% \text{ C.L.})$$

Phys.Rev.Lett. 110:062502 (2013)

Gerda phase I results



- Background index: 0.01 cts/(keV kg yr)
- Exposure: 21.6 kg yr
- No $0\nu\beta\beta$ -signal observed at $Q_{\beta\beta}=2039$ keV
- Background-only hypothesis H_0 strongly favoured
- Claim strongly disfavoured



Limit on half-life:

- GERDA: $T_{1/2,0\nu} > 2.1 \times 10^{25}$ yr (90% C.L.)
- GERDA+IGEX+HdM: $T_{1/2,0\nu} > 3.0 \times 10^{25}$ yr (90% C.L.)

$$\langle m_{\beta\beta} \rangle < 0.2-0.4 \text{ eV}$$

Phys.Rev.Lett. 111:122503 (2013)