

# Component separation

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# Objective

- We must make sure that the CMB polarization sensitivity is limited by **the instrumental noise** and ***NOT*** by **astrophysical confusion**
- We have to fight ***at least*** two foregrounds
  - Synchrotron
  - Dust
- The more sensitive the mission, the more likely it is that we must deal with additional components

# Validation of detected B-modes

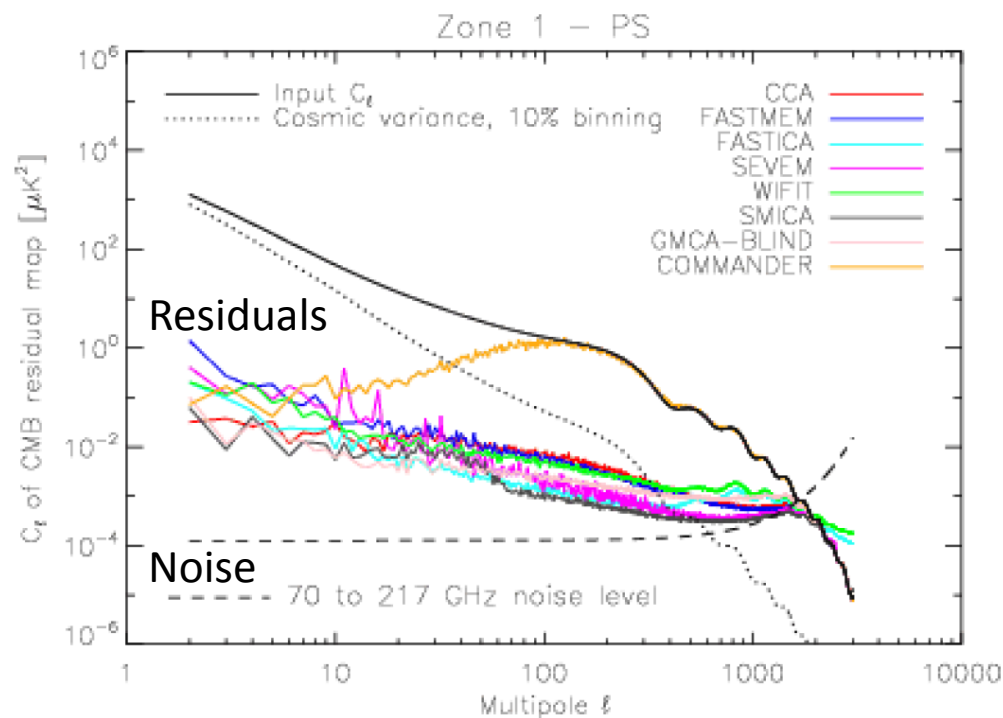
- The only way to guarantee that any excess of power that we see is CMB emission is to check that its color is right
- Hence we need several CMB channels
- For a sensitive CMB B-mode mission we need at least 6 channels to monitor foregrounds, and at least 3 CMB channels (to get more than 1 cross-spectrum).

All FG could be polarised up to a fraction of a percent !

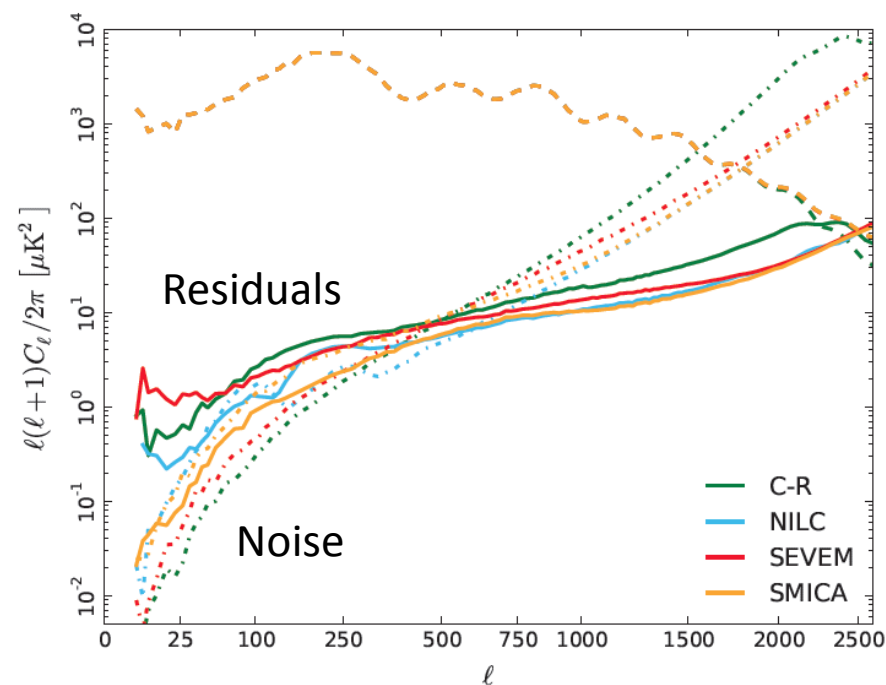
# More than 9 channels is better

- With 9 channels, component separation is a limitation to Planck performance.

Leach et al. (2008)



Planck 2013 results. XII



# Impact of FG emission on $r=T/S$

Betoule et al. (2009)

| case               | $r$   | noise-only   |                             |                          | known foregrounds |                             |                          | SMICA        |                             |                          | $r^{\text{est}}$ | $l_{\text{min}} - l_{\text{max}}$ | $f_{\text{sky}}$ | $D^3$ |
|--------------------|-------|--------------|-----------------------------|--------------------------|-------------------|-----------------------------|--------------------------|--------------|-----------------------------|--------------------------|------------------|-----------------------------------|------------------|-------|
|                    |       | $\sigma_r/r$ | $\sigma_r^{\ell \leq 20}/r$ | $\sigma_r^{\ell > 20}/r$ | $\sigma_r/r$      | $\sigma_r^{\ell \leq 20}/r$ | $\sigma_r^{\ell > 20}/r$ | $\sigma_r/r$ | $\sigma_r^{\ell \leq 20}/r$ | $\sigma_r^{\ell > 20}/r$ |                  |                                   |                  |       |
| PLANCK             | 0.3   | 0.075        | 0.17                        | 0.084                    | 0.1               | 0.2                         | 0.12                     | 0.15         | 0.22                        | 0.2                      | <b>0.26</b>      | 2 - 130                           | 0.95             | 3     |
|                    | 0.1   | 0.17         | 0.25                        | 0.22                     | 0.23              | 0.34                        | 0.32                     | 0.29         | 0.34                        | 0.55                     | 0.086            |                                   |                  |       |
| EPIC-LC            | 0.01  | 0.019        | 0.084                       | 0.019                    | 0.05              | 0.18                        | 0.053                    | 0.079        | 0.18                        | 0.1                      | <b>0.0098</b>    | 2 - 130                           | 0.86             | 4     |
|                    | 0.001 | 0.059        | 0.15                        | 0.064                    | 0.27              | 0.4                         | 0.38                     | 0.37         | 0.43                        | 0.82                     | 0.00088          |                                   |                  |       |
| EPIC-2m            | 0.01  | 0.016        | 0.083                       | 0.016                    | 0.027             | 0.12                        | 0.027                    | 0.032        | 0.11                        | 0.036                    | <b>0.0096</b>    | 2 - 300                           | 0.87             | 4     |
|                    | 0.001 | 0.051        | 0.14                        | 0.055                    | 0.14              | 0.25                        | 0.16                     | 0.16         | 0.24                        | 0.24                     | <b>0.001</b>     |                                   |                  |       |
| EPIC-CS            | 0.01  | 0.017        | 0.084                       | 0.017                    | 0.029             | 0.12                        | 0.03                     | 0.036        | 0.11                        | 0.041                    | <b>0.0096</b>    | 2 - 300                           | 0.87             | 4     |
|                    | 0.001 | 0.058        | 0.15                        | 0.063                    | 0.15              | 0.27                        | 0.19                     | 0.18         | 0.26                        | 0.29                     | <b>0.00098</b>   |                                   |                  |       |
| Ground-based       | 0.1   | 0.083        | —                           | —                        | 0.15              | —                           | —                        | 0.24         | —                           | —                        | <b>0.11</b>      | 50 - 300                          | 0.01             | 2     |
|                    | 0.01  | 0.18         | —                           | —                        | 0.8               | —                           | —                        | 1.6          | —                           | —                        | 0.018            |                                   |                  |       |
| Grnd-based+Planck  | 0.01  | 0.18         | —                           | —                        | 0.51              | —                           | —                        | 0.69         | —                           | —                        | 0.0065           | 50 - 300                          | 0.01             | 2     |
| Deep field mission | 0.001 | 0.082        | —                           | —                        | 0.1               | —                           | —                        | 0.13         | —                           | —                        | <b>0.00092</b>   | 50 - 300                          | 0.01             | 4     |

EPIC-LC    7 channels    34' at 135 GHz     $\sigma_{r=0.001} \times 6$

EPIC-CS    8 channels    3.1' at 150 GHz     $\sigma_{r=0.001} \times 3$

EPIC-2m    9 channels    5' at 150 GHz     $\sigma_{r=0.001} \times 3$

Ground    3 channels    5.5' at 150 GHz     $\sigma_{r=0.01} \times 4-9$     (1% of sky)

# Counting components

- Polarization
  - Synchrotron: 3
  - Dust: 3+
  - CMB: 1
  - Redundancy: 2

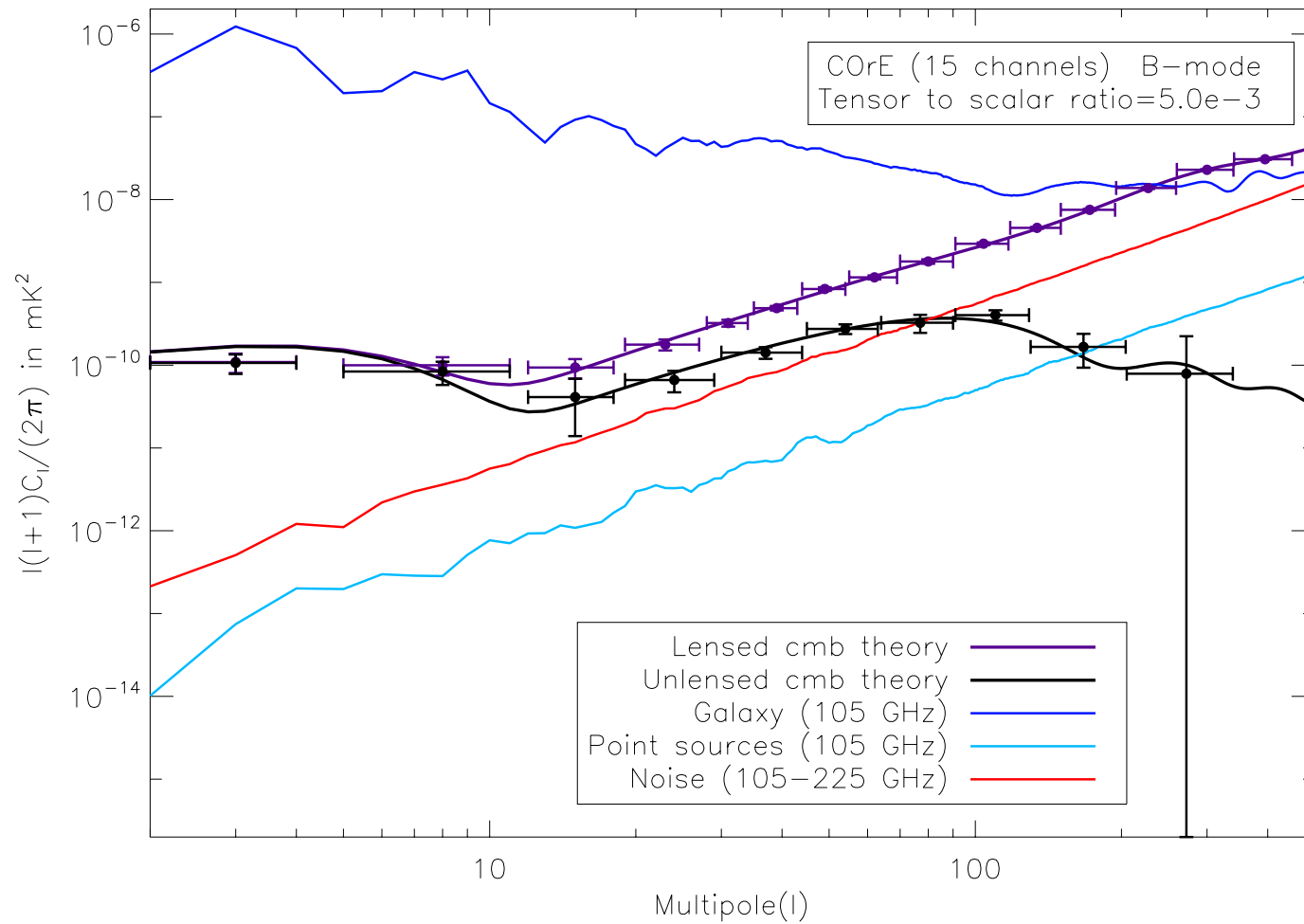
9+ CHANNELS

- Intensity
  - Synchrotron: 3
  - Dust: 3+
  - CMB: 1
  - SZ: 2
  - Free-free: 1-2
  - Spinning dust: 1-3
  - CO: 0-2
  - CIB: 5-10
  - Redundancy: 3

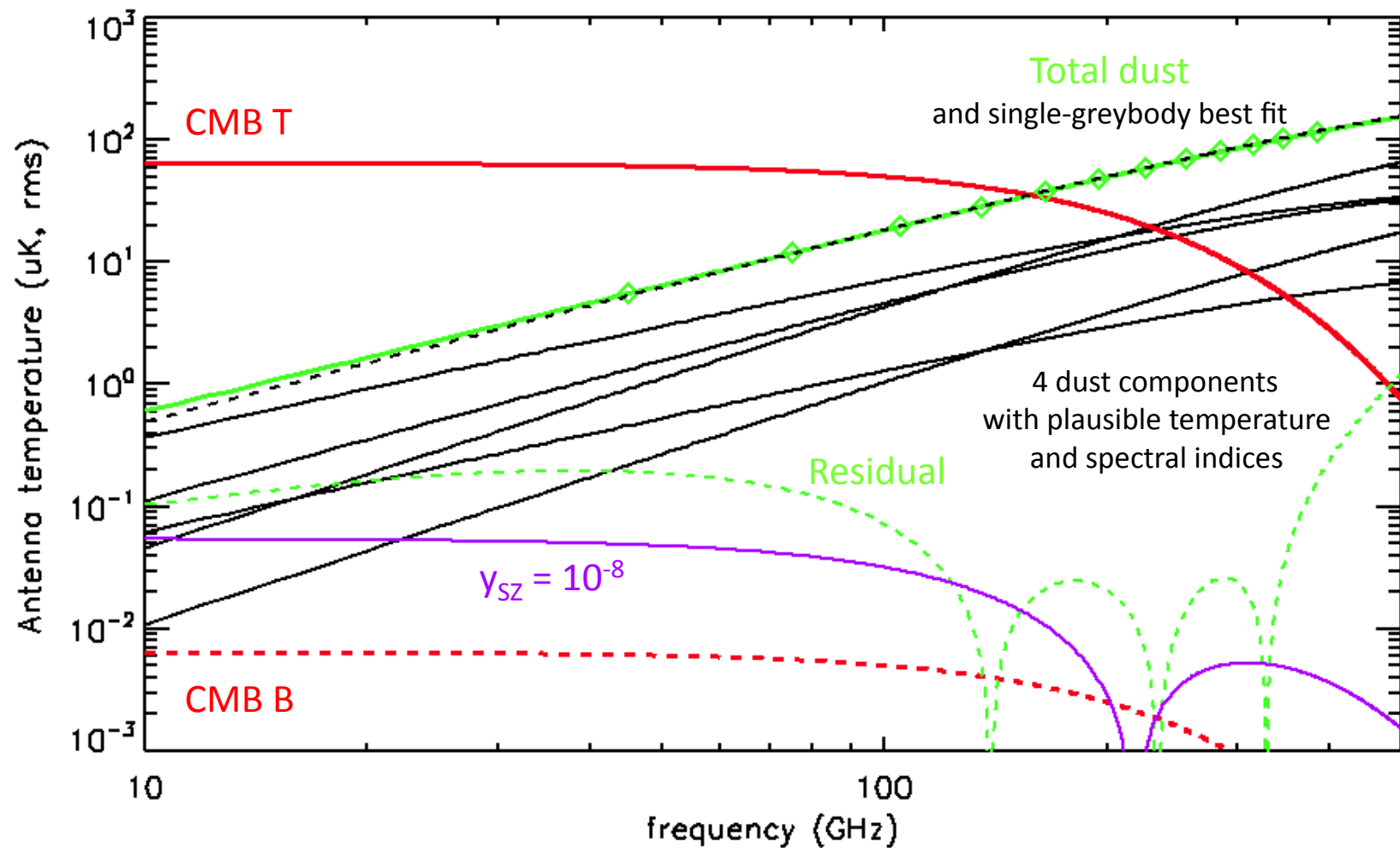
18+ CHANNELS

All FG could be polarised up to a fraction of a percent !

# COrE proposal

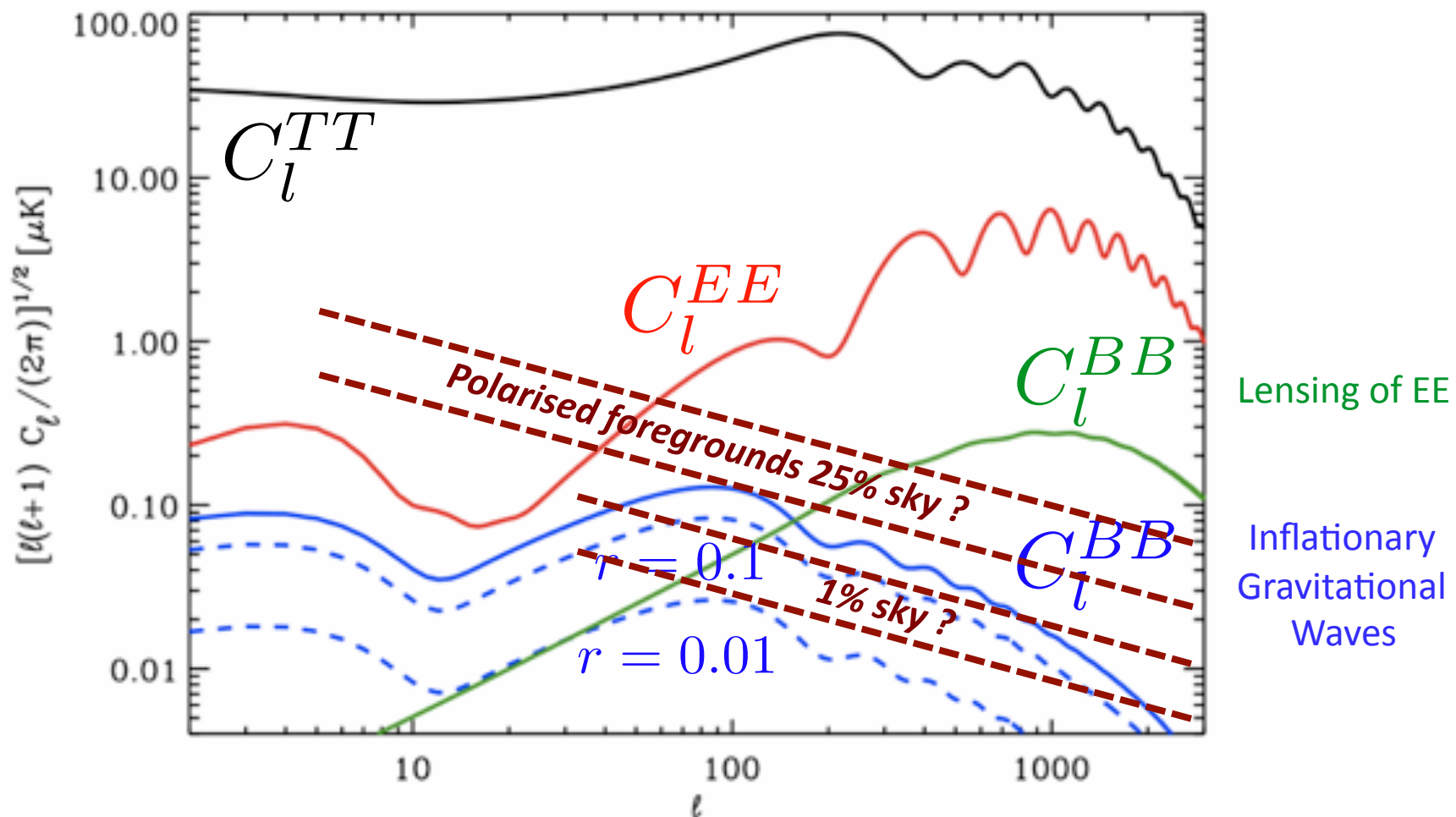


# Beware fitting emission laws

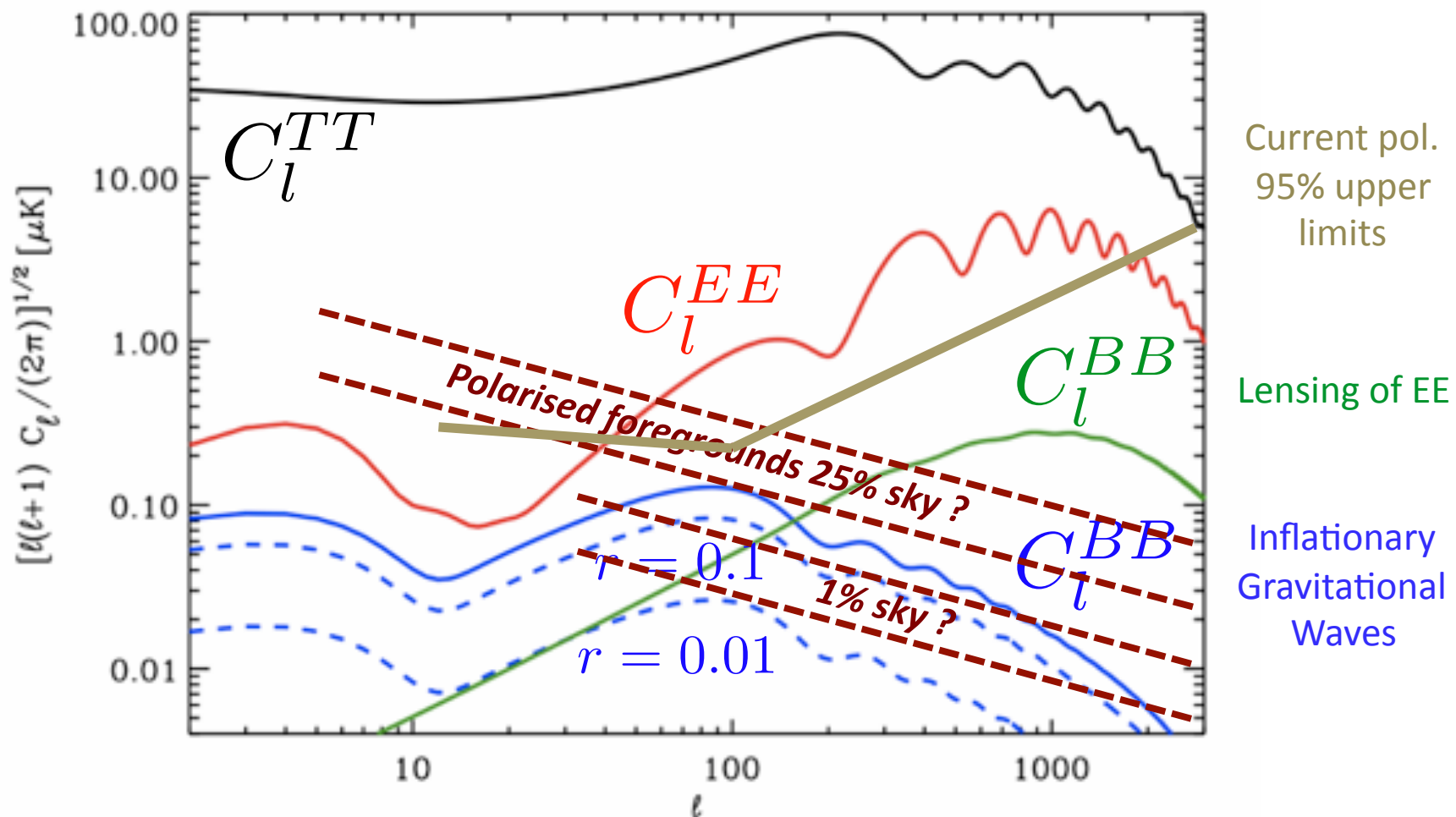




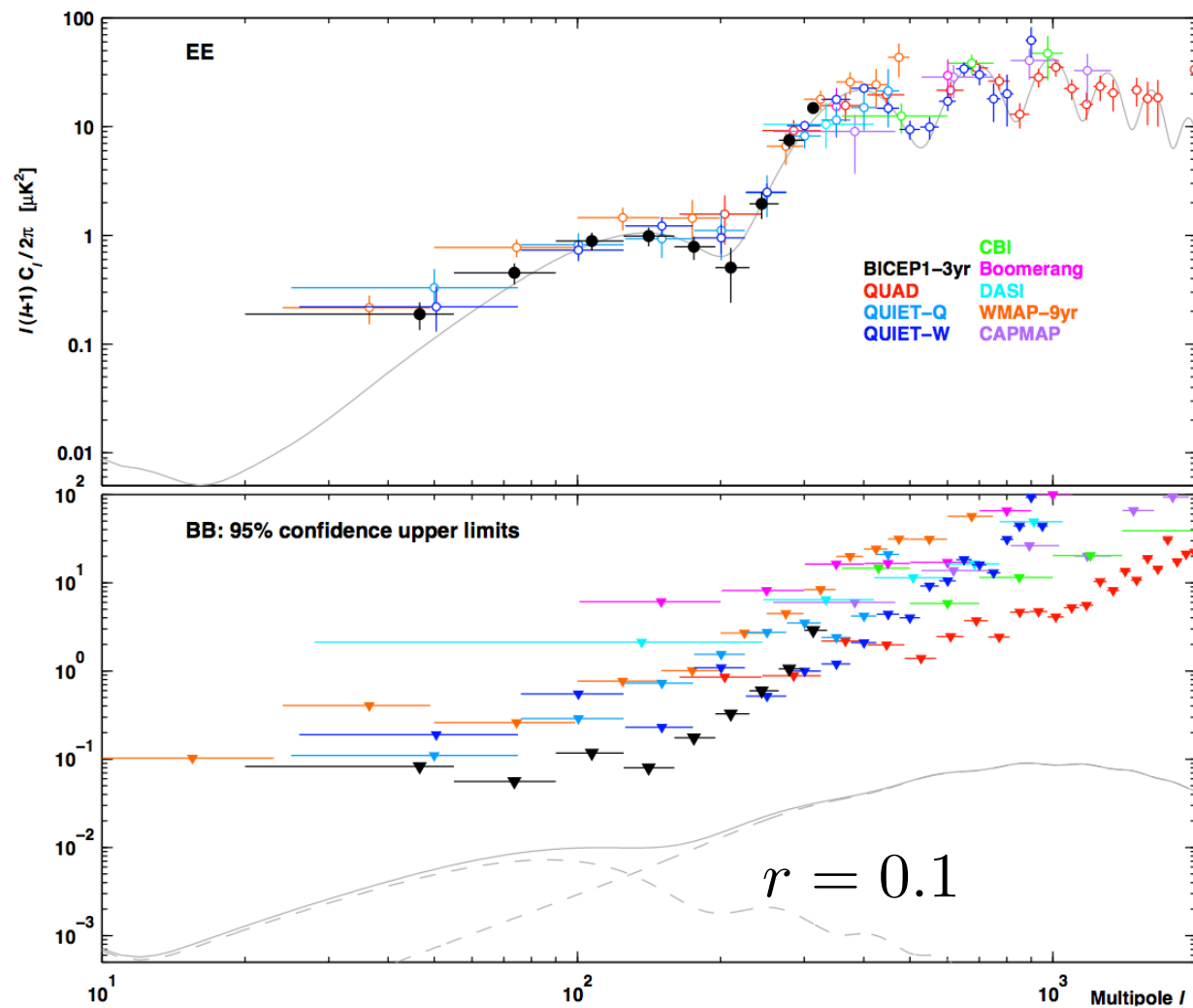
# Foregrounds (150 GHz)



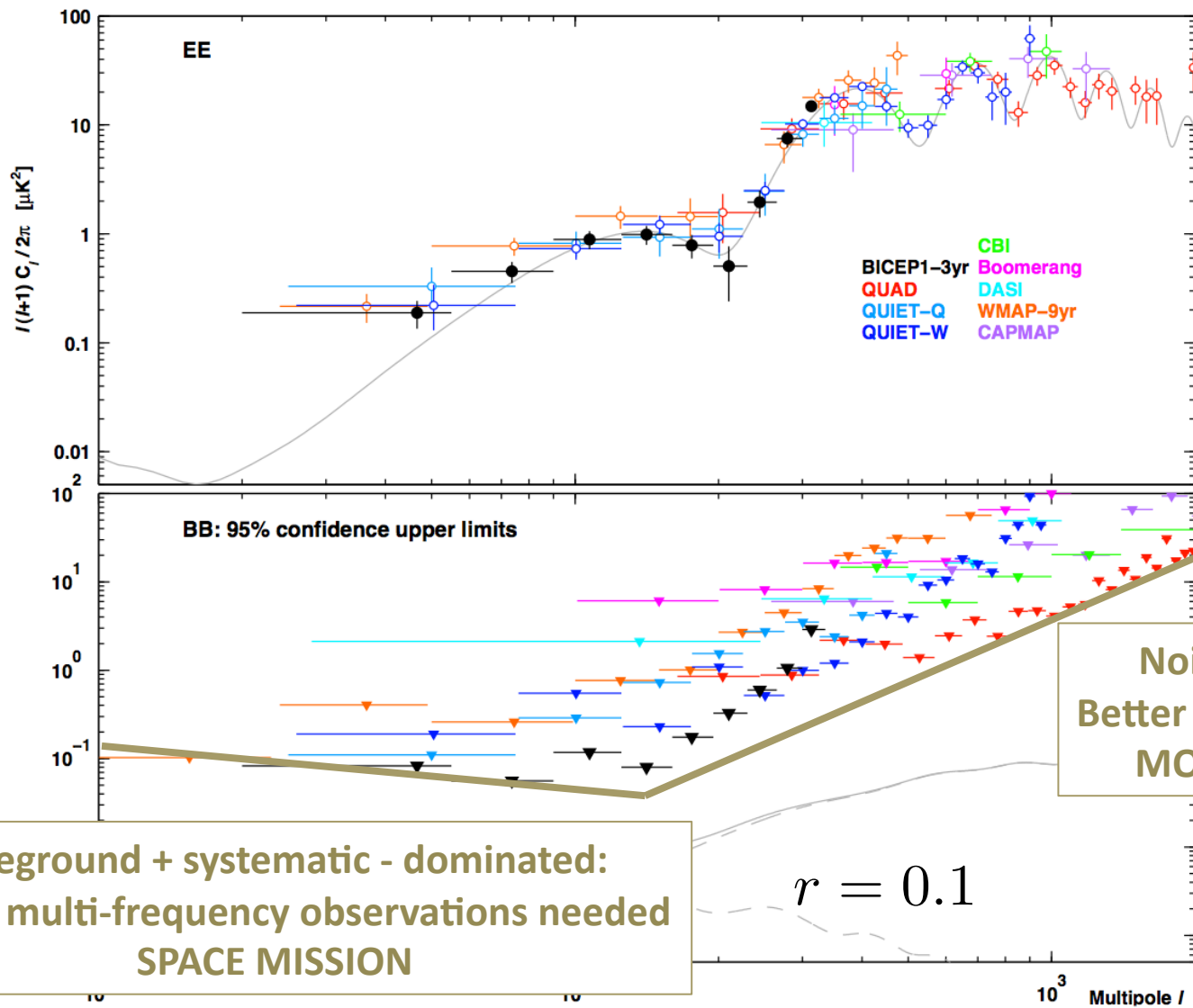
# Foregrounds (150 GHz)



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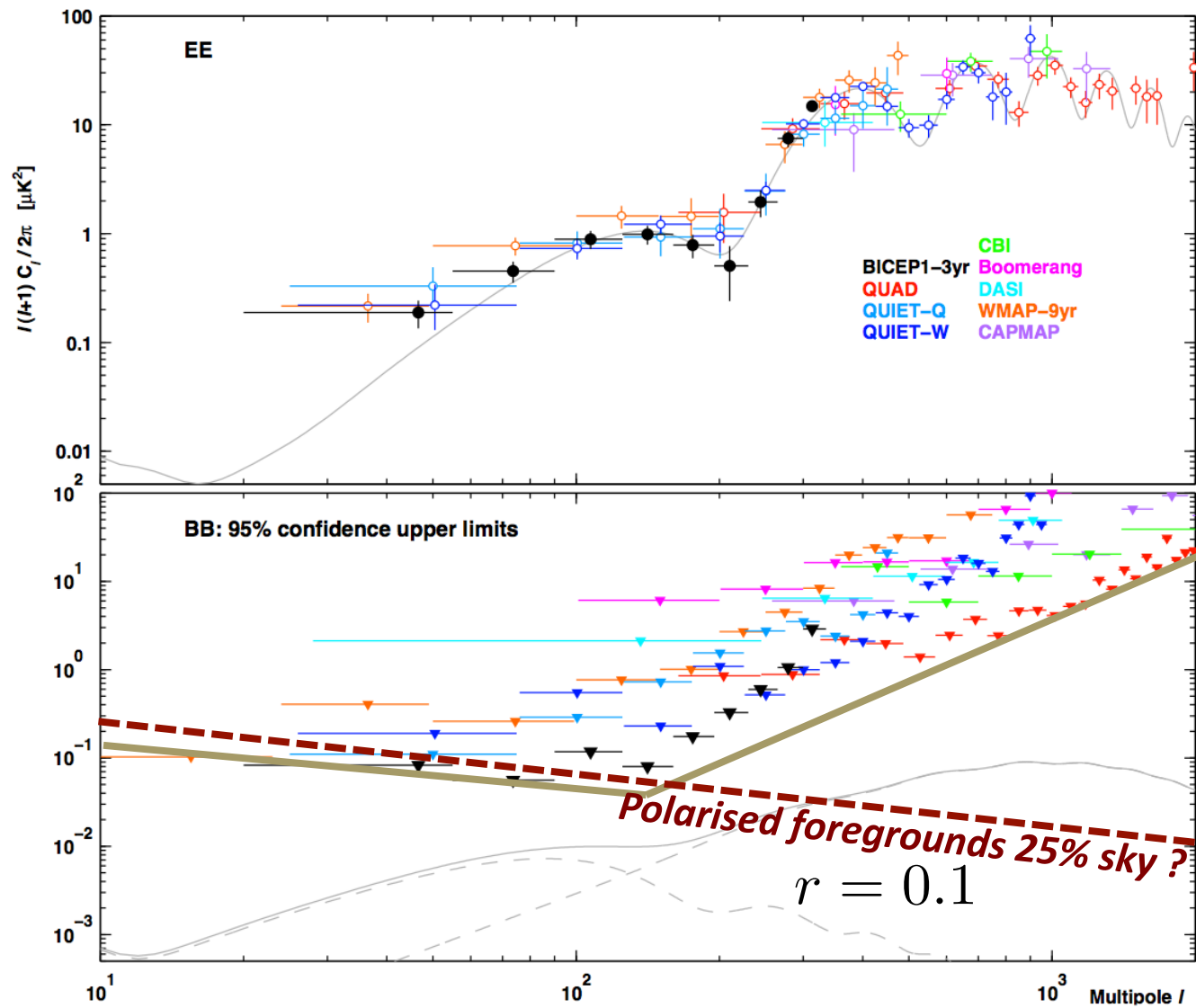
# Current measurements



Noise-dominated:  
Better sensitivity needed  
**MORE DETECTORS**

Foreground + systematic - dominated:  
Better, multi-frequency observations needed  
**SPACE MISSION**

# Current measurements

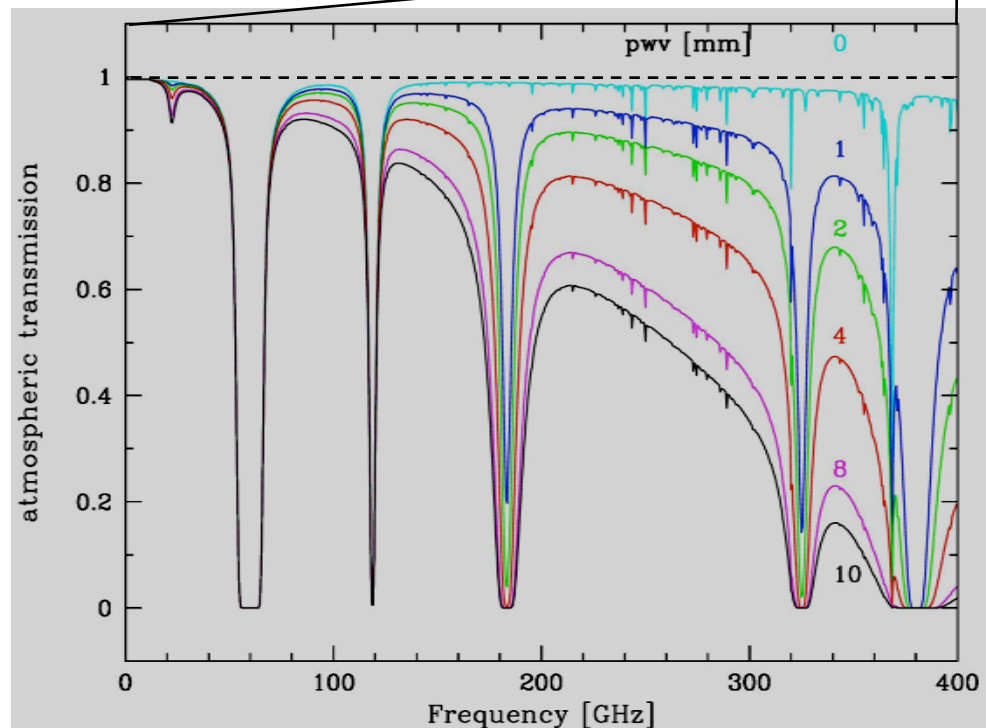
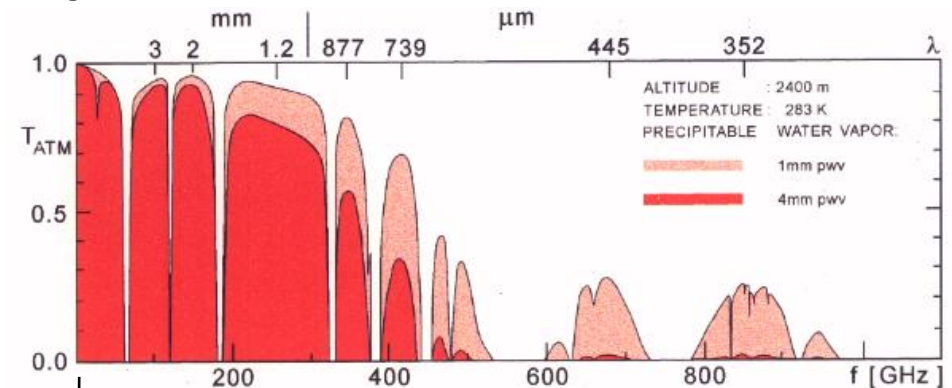


# Atmosphere

Atmospheric transmission  
and emission

Systematics

Complete survey



At 150 GHz a Planck detector is  $\approx 10 \times$  more sensitive than ground-based detectors at ATACAMA or South Pole in good observing conditions.

Factor of 100-400 in observing time.

The situation is much worse for ground-based observations at higher frequency.

# Conclusion

- Plan **at least 10 channels**, preferably more
- **COrE**: 15 channels, but not enough on the low frequency end
- **PRISM**: 20 channels in the frequency range useful for the CMB, + line monitors
- Proposed mission baseline: **15-20 frequency channels with optimised number of detectors**
- None of this is possible from the ground !