

# Mueller-Navelet jets at the LHC with optimal renormalization

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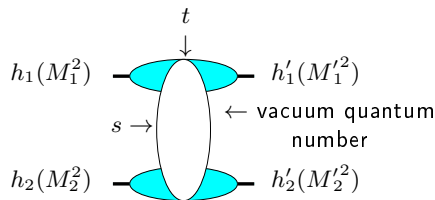
in collaboration with

L. Szymanowski (NCBJ Warsaw), S. Wallon (UPMC & LPT Orsay)

B. D, L. Szymanowski, S. Wallon, JHEP **1305** (2013) 096 [arXiv:1302.7012]

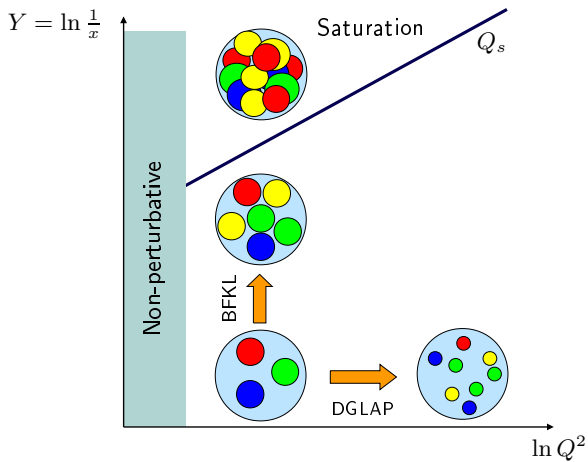
B. D, L. Szymanowski, S. Wallon, arXiv:1309.3229 (to appear in PRL)

- One of the important longstanding theoretical questions raised by QCD is its behaviour in the perturbative **Regge** limit  $s \gg -t$
- Based on theoretical grounds, one should identify and test suitable observables in order to test this peculiar dynamics



hard scales:  $M_1^2, M_2^2 \gg \Lambda_{QCD}^2$  or  $M_1'^2, M_2'^2 \gg \Lambda_{QCD}^2$  or  $t \gg \Lambda_{QCD}^2$   
 where the  $t$ -channel exchanged state is the so-called **hard Pomeron**

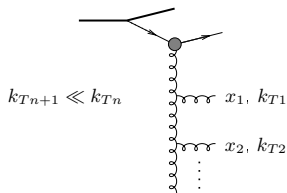
# The different regimes of QCD



Small values of  $\alpha_S$  (perturbation theory applies due to hard scales) can be compensated by large logarithmic enhancements.

⇒ resummation of  $\sum_n (\alpha_S \ln A)^n$  series

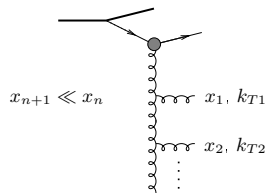
## DGLAP



strong ordering in  $k_T$

$$\sum (\alpha_S \ln \frac{Q^2}{\mu^2})^n$$

## BFKL



strong ordering in  $x$

$$\sum (\alpha_S \ln \frac{s}{s_0})^n$$

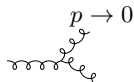
When  $\sqrt{s}$  becomes very large, it is expected that a BFKL description is needed to get accurate predictions

## What kind of observables?

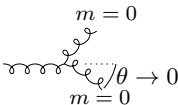
- perturbation theory should be applicable:

selecting external or internal probes with transverse sizes  $\ll 1/\Lambda_{QCD}$  or by choosing large  $t$  in order to provide the hard scale

- governed by the *soft* perturbative dynamics of QCD



and *not* by its *collinear* dynamics



$\Rightarrow$  select semi-hard processes with  $s \gg p_{T,i}^2 \gg \Lambda_{QCD}^2$  where  $p_{T,i}^2$  are typical transverse scale, **all of the same order**

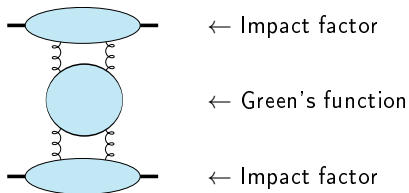
## QCD in the perturbative Regge limit

The amplitude can be written as:

$$\mathcal{A} = \underbrace{\text{Diagram 1}}_{\sim s} + \left( \underbrace{\text{Diagram 2}}_{\sim s (\alpha_s \ln s)} + \underbrace{\text{Diagram 3}}_{\sim s (\alpha_s \ln s)} + \dots \right) + \left( \underbrace{\text{Diagram 4}}_{\sim s (\alpha_s \ln s)^2} + \dots \right) + \dots$$

The diagrams are represented by light blue shapes: two horizontal ovals for the first term, a horizontal oval and a vertical circle for the second and third terms, and two horizontal ovals for the fourth term. Wavy lines connect these shapes to represent gluon exchanges.

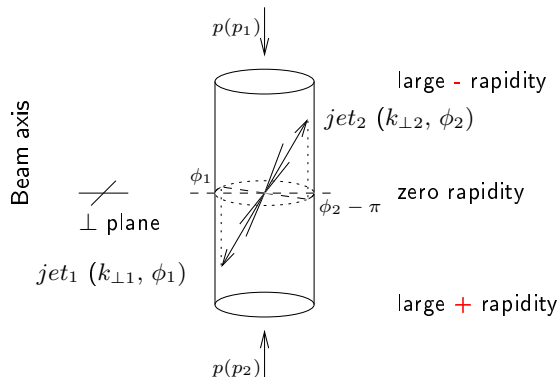
this can be put in the following form :



- Higher order corrections to BFKL kernel are known at NLL order (Lipatov Fadin; Camici, Ciafaloni), now for arbitrary impact parameter  $\alpha_S \sum_n (\alpha_S \ln s)^n$  resummation
- impact factors are known in some cases at NLL
  - $\gamma^* \rightarrow \gamma^*$  at  $t = 0$  (Bartels, Colferai, Gieseke, Kyrieleis, Qiao; Balitski, Chirilli)
  - forward jet production (Bartels, Colferai, Vacca; Caporale, Ivanov, Murdaca, Papa, Perri; Chachamis, Hentschinski, Madrigal, Sabio Vera)
  - inclusive production of a pair of hadrons separated by a large interval of rapidity (Ivanov, Papa)
  - $\gamma_L^* \rightarrow \rho_L$  in the forward limit (Ivanov, Kotsky, Papa)

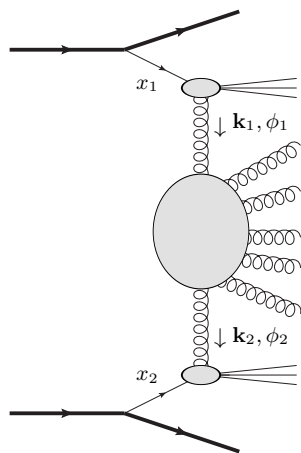
## Mueller-Navelet jets

- Consider two jets (hadrons flying within a narrow cone) **separated by a large rapidity**, i.e. each of them almost fly in the direction of the hadron “close” to it, and with very similar transverse momenta
- in a pure LO collinear treatment, these two jets should be emitted **back to back** at leading order:  $\Delta\phi - \pi = 0$  ( $\Delta\phi = \phi_1 - \phi_2 =$  relative azimuthal angle) and  $k_{\perp 1} = k_{\perp 2}$ . There is no phase space for (untagged) emission between them





## $k_T$ -factorized differential cross section



$$\frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}} = \int d\phi_{J1} d\phi_{J2} \int d^2\mathbf{k}_1 d^2\mathbf{k}_2$$

$$\times \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1)$$

$$\times G(\mathbf{k}_1, \mathbf{k}_2, \hat{s})$$

$$\times \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2)$$

with  $\Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2) = \int dx_2 f(x_2) V(\mathbf{k}_2, x_2)$      $f \equiv \text{PDF}$      $x_J = \frac{|\mathbf{k}_J|}{\sqrt{s}} e^{y_J}$

It is convenient to define the coefficients  $\mathcal{C}_n$  as

$$\begin{aligned} \mathcal{C}_n &\equiv \int d\phi_{J1} d\phi_{J2} \cos(n(\phi_{J1} - \phi_{J2} - \pi)) \\ &\quad \times \int d^2\mathbf{k}_1 d^2\mathbf{k}_2 \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1) G(\mathbf{k}_1, \mathbf{k}_2, \hat{s}) \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2) \end{aligned}$$

- $n = 0 \implies$  differential cross-section

$$\mathcal{C}_0 = \frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}}$$

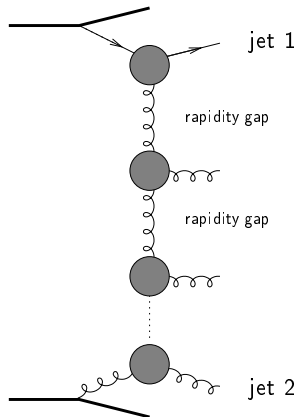
- $n > 0 \implies$  azimuthal decorrelation

$$\frac{\mathcal{C}_n}{\mathcal{C}_0} = \langle \cos(n(\phi_{J,1} - \phi_{J,2} - \pi)) \rangle \equiv \langle \cos(n\varphi) \rangle$$

- sum over  $n \implies$  azimuthal distribution

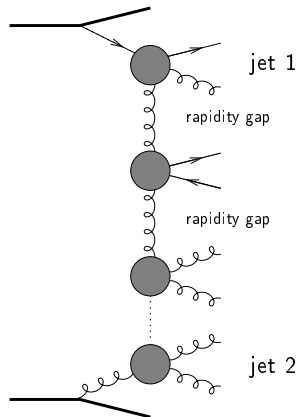
$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}$$

LL BFKL



$$\sum (\alpha_s \ln s)^n$$

NLL BFKL



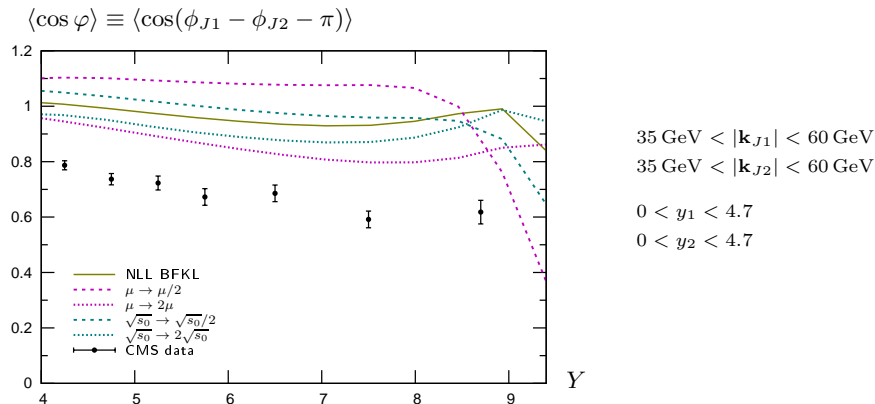
$$\sum (\alpha_s \ln s)^n + \alpha_s \sum (\alpha_s \ln s)^n$$

## Results for a symmetric configuration

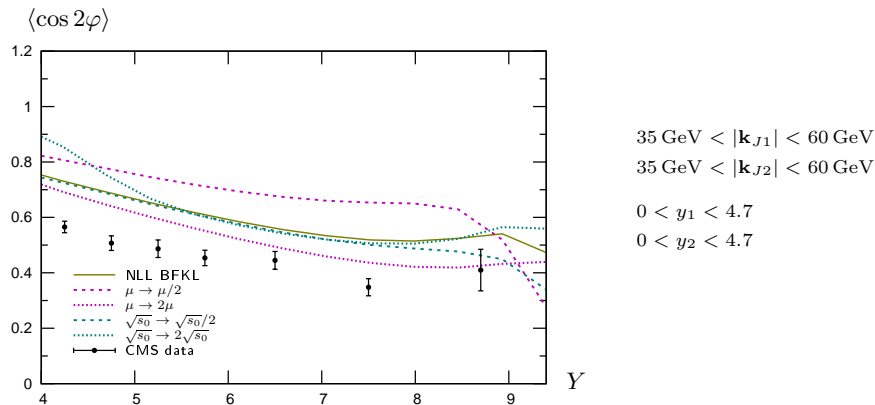
In the following we show results for

- $\sqrt{s} = 7 \text{ TeV}$
- $35 \text{ GeV} < |\mathbf{k}_{J1}|, |\mathbf{k}_{J2}| < 60 \text{ GeV}$
- $0 < y_1, y_2 < 4.7$

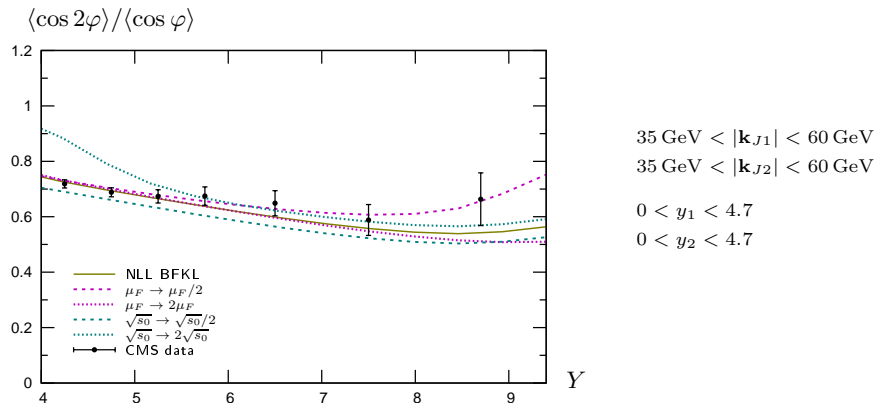
These cuts allow us to compare our predictions with the first experimental data on azimuthal correlations of Mueller-Navelet jets from the LHC presented by the CMS collaboration (CMS-PAS-FSQ-12-002)

Azimuthal correlation  $\langle \cos \varphi \rangle$ 

- NLL BFKL predicts a too small decorrelation
- The NLL BFKL calculation is still rather dependent on the scales, especially the renormalization / factorization scale

Azimuthal correlation  $\langle \cos 2\varphi \rangle$ 

- The agreement with data is a little better for  $\langle \cos 2\varphi \rangle$  but still not very good
- This observable is also very sensitive to the scales

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

- This observable is more stable with respect to the scales than the previous ones
- The agreement with data is good across the full  $Y$  range

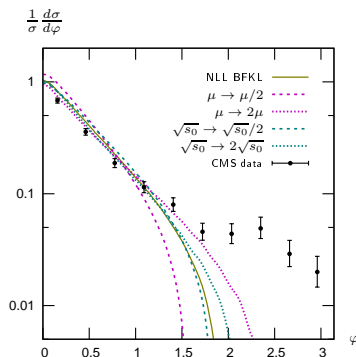
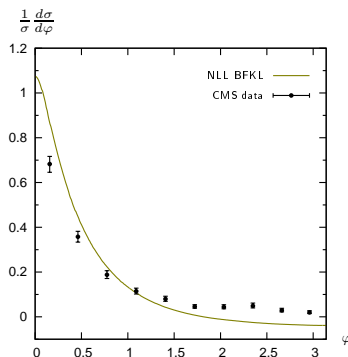
### Azimuthal distribution

The azimuthal distribution  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  has also been measured by the CMS collaboration. It can be written as

$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}$$



## Azimuthal distribution: comparison to CMS data

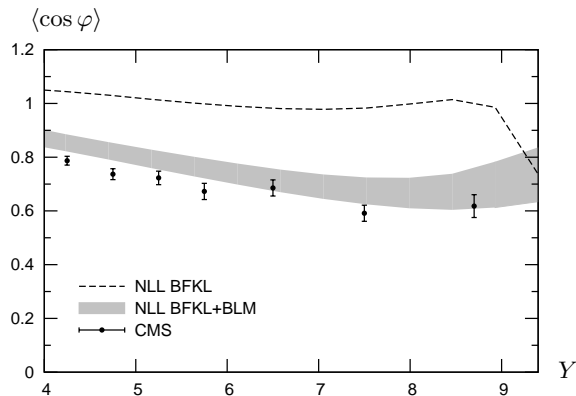


- Our calculation predicts a too large value of  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  for  $\varphi \lesssim \frac{\pi}{2}$  and a too small value for  $\varphi \gtrsim \frac{\pi}{2}$
- For large values of  $\varphi$ , the distribution even becomes negative

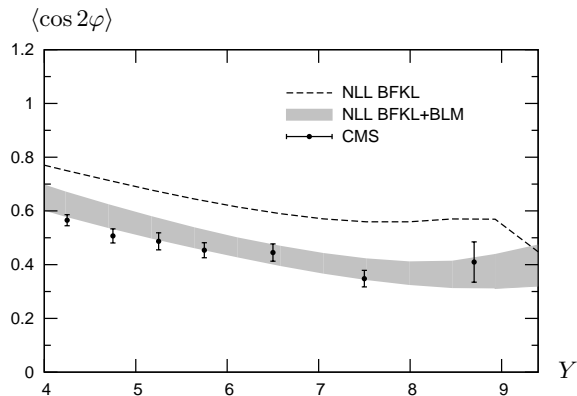
- The agreement of our calculation with the data for  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is good and very stable with respect to the scales
- The agreement for  $\langle \cos n\varphi \rangle$  and  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  is not very good and very sensitive to the choice of the renormalization scale  $\mu_R$
- An all-order calculation would be independent of the choice of  $\mu_R$ . This feature is lost if we truncate the perturbative series  
 $\Rightarrow$  How to choose the renormalization scale?  
 'Natural scale': sometimes the typical momenta in a loop diagram are different from the natural scale of the process

The **Brodsky-Lepage-Mackenzie** (BLM) procedure resums the self-energy corrections to the gluon propagator at one loop into the running coupling. These contributions are formally of higher-order but they are proportional to

$$\beta_0 = \frac{11N_c - 2N_f}{3} \simeq 7.67$$

Azimuthal correlation  $\langle \cos \varphi \rangle$ 
 $35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$ 
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$ 
 $0 < y_1 < 4.7$ 
 $0 < y_2 < 4.7$ 

Using the BLM scale setting, the scale uncertainty is reduced and the agreement with data becomes much better

Azimuthal correlation  $\langle \cos 2\varphi \rangle$ 

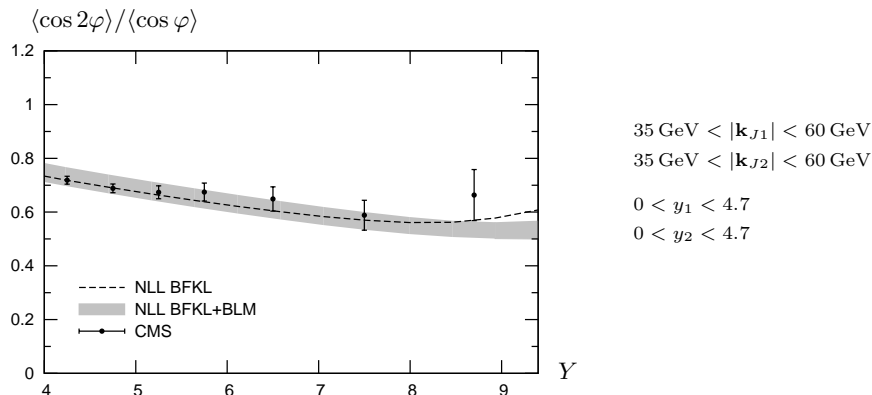
$$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$$

$$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$$

$$0 < y_1 < 4.7$$

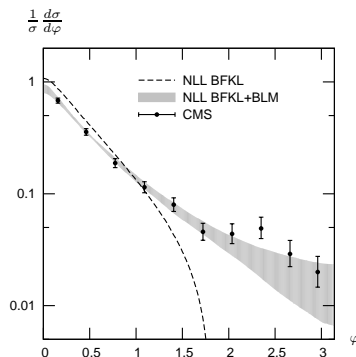
$$0 < y_2 < 4.7$$

Using the BLM scale setting, the scale uncertainty is reduced and the agreement with data becomes much better

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

Because it is much less dependent on the scales, the observable  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is almost not affected by the BLM procedure and is still in very good agreement with the data

## Azimuthal distribution: comparison to CMS data



With the BLM scale setting the azimuthal distribution no longer reaches negative values and is in good agreement with the data across the full  $\varphi$  range.

- We studied Mueller-Navelet jets at full (vertex + Green's function) NLL accuracy and compared our results with the first data from the LHC
- The observables  $\langle \cos n\varphi \rangle$  and  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  are very dependent on the choice of the scales and don't agree very well with data when using a 'natural' scale
- The agreement with CMS data is greatly improved by using the BLM scale fixing procedure
- For the observable  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ :
  - NLL BFKL predictions are much more stable with respect to the scales
  - the data is well described by BFKL
  - in our opinion this would be a good observable to distinguish between BFKL and other scenarios in the future