#### Theoretical aspects of Beyond the Standard Model physics



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# Experimental hot subjects (a very long quest in just few lines)

- Discovery and properties of the Higgs Boson
- The top quark
- Electroweak precision tests
- Flavour measurements
- Neutrino oscillations
- Dark matter content of the universe

Standard Model quite successful on the first 4, way beyond expectations!

# Hints for BSM

we have data to explain beyond SM

- neutrinos and flavour hierarchies in the quark/ lepton sectors
- dark matter (and dark energy if you dare...)

data to search/fulfill (precision tests, small deviations)

- Higgs boson mass, couplings
- top and flavour data
- EW precision data, etc.

# Guidelines for BSM physics

#### • Top-down

- Fundamental symmetries 😳
- Unification of the couplings ③
- Unitarity 😳
- Renormalisability 🙂
- Hierarchies, naturalness, fine-tuning?? etc...

• Bottom-up

- Dark matter as a WIMP
- Electroweak precision tests
- Flavour observables
- Observation of new particles

#### Usual list, but realistic model-building not obvious

# New Physics @ the TeV scale?

- Electroweak symmetry breaks at energies 100 GeV-1 TeV
- Weakly Interacting Massive Particle needs a mass
- - 100 GeV-1 TeV to fit observed Dark Matter density
- WL WL scattering unitarised at energies 1 TeV, by just the Higgs?
- Hierarchy: fine-tuning or Higgs mass must be stabilized by a scale 1 TeV

This makes a few strong hints to go beyond the SM at the TeV scale but...

Constraints from higher dim operators Unless symmetries/cancellations:

- Baryon Number Violation
- Lepton Number Violation
- Flavour Violation
- CP Violation
- Precision Electroweak
- Contact Operators

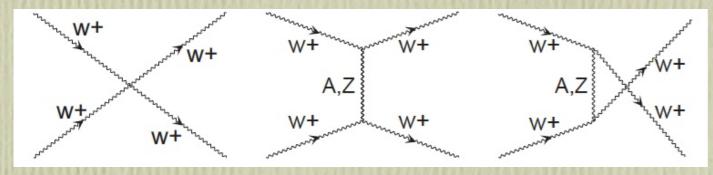


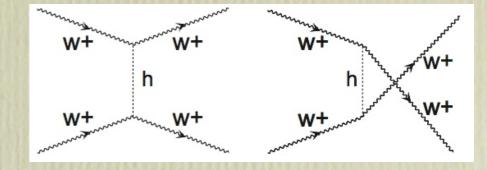
### Some ideas behind models

- WW scattering unitarity
- naturalness and vacuum stability
- matter parities
- extra symmetries (space-time, global, gauge)
- compositeness & Technicolor
- extra dimensions

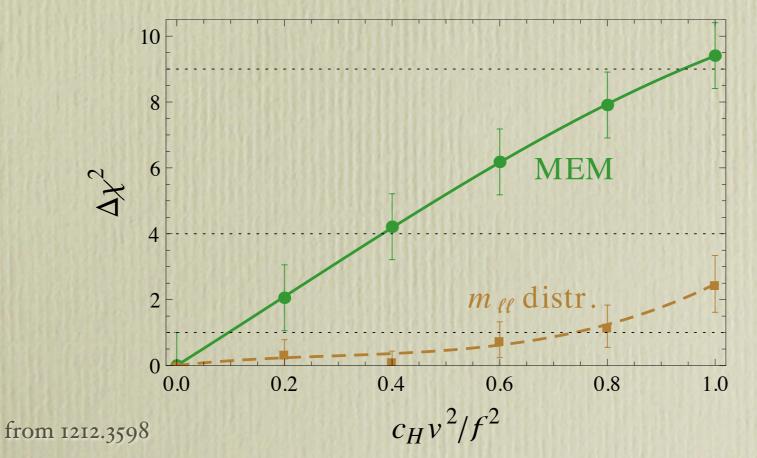
### WW and unitarity

• Higgs boson exchange crucial (see Lee, Quigg, Thacker 1977)

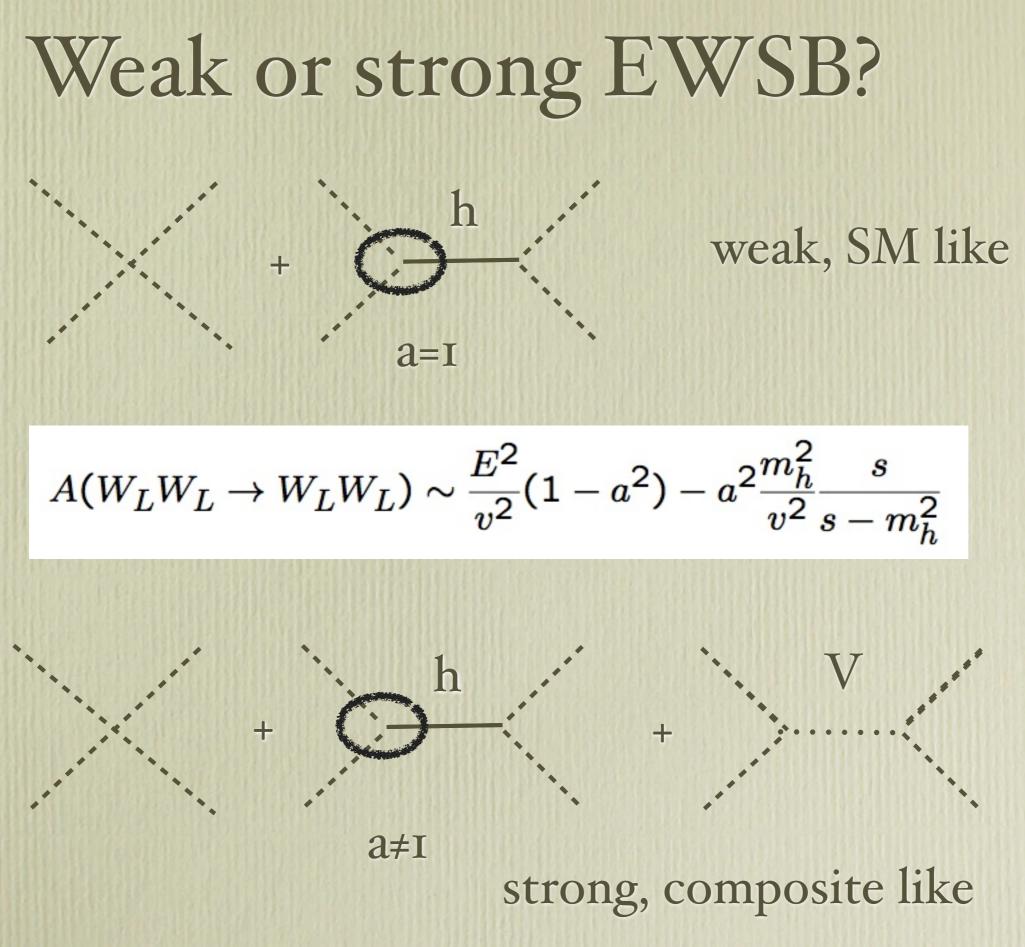




• Now still useful for discriminating models if accurately measured at LHC 14 TeV, example testing 2HDM or effective composite SILH models:



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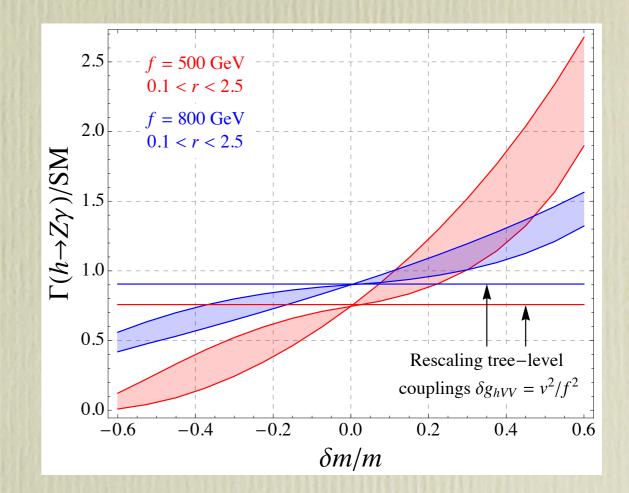


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#### Possible tests weak vs strong

- VV → VV, limited sensitivity at LHC
- $h \rightarrow Z \gamma$ , large effect may be possible
- VV → hh, at a e<sup>+</sup>e<sup>-</sup> collider

see 1305.5251, 1308.2676, 1309.7038



#### Natural theories

$$m^2 = m_0^2 \left( 1 + a(\lambda, g) \log \frac{\Lambda^2}{m_0^2} \right) + b(\lambda, g) \Lambda^2$$

- natural if  $b(\lambda,g)=0$  by a symmetry
- can be natural if  $\Lambda$  is a physical cut-off (ex. compositness)
- quasi-natural if  $b(\lambda,g)=0$  perturbatively (ex. at oneloop in Little Higgs for top contribution)
- tuned: any special value you like, even mo=0 Λ=0 (classically conformal)

#### SM and naturalness

 Naturalness as UV sensitivity of h mass in the SM effective theory (not a principle, just reasonable)

$$\delta M_H^2 \sim \frac{3G_F}{4\pi^2\sqrt{2}} \Lambda^2 (4m_t^2 - 2m_W^2 - m_Z^2 - m_H^2)$$

implies

 $\Lambda\sim 500~GeV$ 

• no new physics (yet) implies it is wrong? Maybe but it works in other cases...

new physics: charm quark mass at 1.3 GeV

# More general question... if fundamental scalar sector

 $V = \text{constant} + M_H^2 |H|^2 + \lambda |H|^4$ 

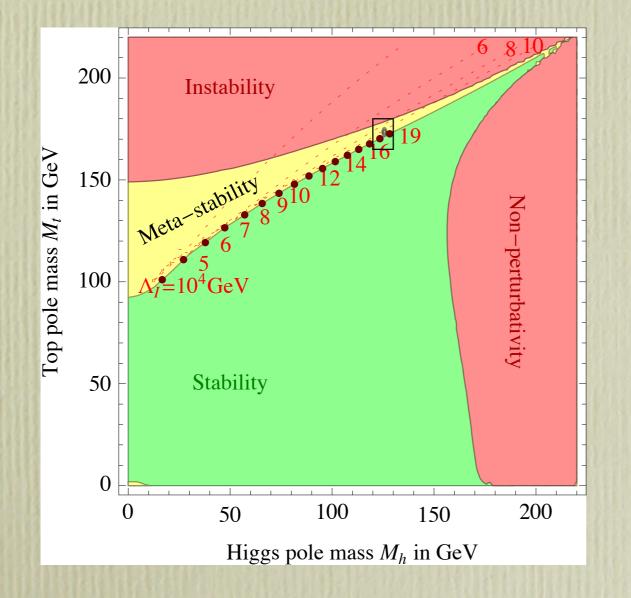
Cosmological constant

Naturalness

Stability

### Stability/criticality

• measurement of m<sub>t</sub>, m<sub>h</sub> allows to "wildly" extrapolate up to 15 orders of magnitude:



Why close to instability?

- not a nonsense as true even at low scale
- maybe not tuned, due to a model?

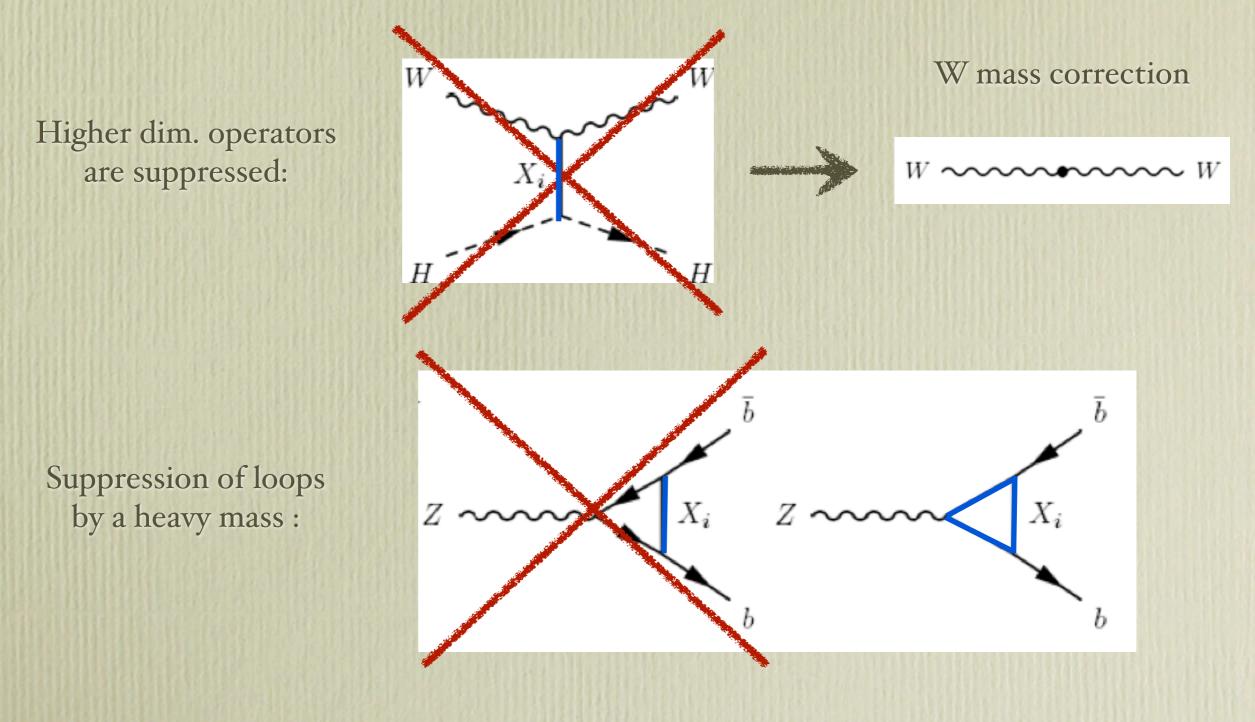
## Matter parities

- Many models share a common ingredient: a Matter parity, M-parity, for the new particles X<sub>A</sub>, X<sub>B</sub> ... or more generally a new global symmetry
- Often known as R-, T-, KK- parity
- Not fundamental, but rather ad hoc
- decay by pairs if initial particles are SM (as M-neutral)
- cannot be resonantly produced
- the lightest M-parity particle is stable (LMP)
- MET (large) at colliders
- once  $X_i$  produced (may cascade) decay to  $X_{LMP}$

A well known example is SUSY phenomenology

#### Matter parities and quantum corrections

• corrections to electroweak precision measurements are typically small, just ad-hoc parity or something more fundamental?



#### Typical BSM models under test

- Supersymmetry
- Technicolor/Composite
- Extra dimensions

# Supersymmetry

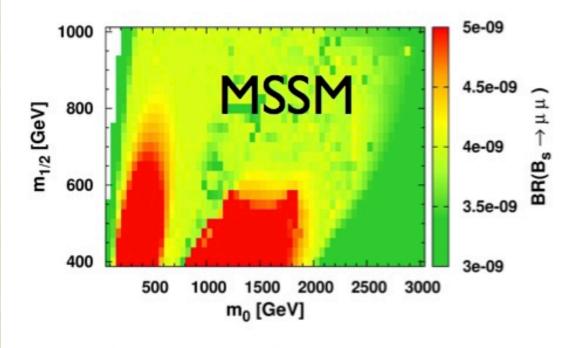
- hierarchy problem
- unification of gauge couplings
- dark matter candidate
- unification with gravity
- essential ingredient in string/brane

#### MSSM

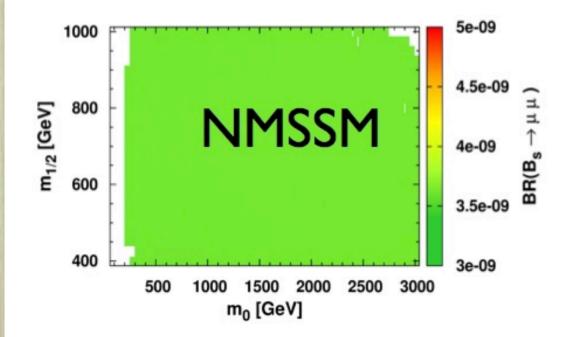
• MSSM has tuning (direct search limit and H mass & couplings constraints)

$$\begin{split} m_{H}^{2} &= m_{Z}^{2}\cos^{2}2\beta + \Delta m_{H}^{2} \\ \Delta m_{H}^{2} &= \frac{3m_{t}^{4}}{4\pi^{2}v^{2}} \left[ \log \frac{M_{S}^{2}}{m_{t}^{2}} \left( 1 - \frac{X_{t}^{2}}{12M_{S}^{2}} \right) \right] \geq 87 \text{ GeV} \\ X_{t} &= A_{t} - \mu \cot \beta \quad M_{S}^{2} = m_{\tilde{t}_{1}} m_{\tilde{t}_{2}} \\ \swarrow \end{split}$$
maximal mixing heavy stops, fine tuning O(1%)

### more general low scale SUSY



$$B_s \to s\gamma, \ B_s \to \mu^+\mu^-, \ B_s \to \tau\nu$$



- not limited to artificially very constrained models (mSUGRA, mGMSB, mAMSB, etc.)
- example NMSSM in much better shape than MSSM
- however  $\lambda$ -SUSY (large coupling  $\lambda$  SH<sub>u</sub>H<sub>d</sub> case of NSSM) partially ruled out, see 1310.0459

#### from D.Kazakov Morion 13

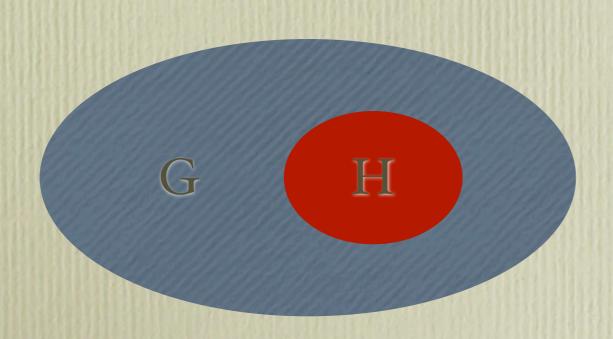
# Compositness/Technicolor

- "Composite" models used today in the effective theory meaning
- "Technicolor" used typically for fundamental fermions forming bound states
- naively the S parameter  $\approx 4\pi (v/m_{\varrho})^2$  implies  $m_{\varrho} \approx 3 \text{ TeV}$
- The lighter the composites, more they affect Higgs couplings

#### Composite effective theories

• The Higgs can be a Goldstone boson (massless in the symmetry limit) of a global symmetry

coset G/H → # Goldstones = dim[G]-dim[H]
 SM ∈ H



Georgi, Kaplan 1980 Arkani-Hamed, Cohen,Georgi 2001 Little Higgs, deconstruction

#### Composite effective theories

8				
G	$\mathcal{H}$	C	$N_G$	$\mathbf{r}_{\mathcal{H}} = \mathbf{r}_{\mathrm{SU}(2) \times \mathrm{SU}(2)} \left( \mathbf{r}_{\mathrm{SU}(2) \times \mathrm{U}(1)}  ight)$
SO(5)	SO(4)	~	4	${f 4}=({f 2},{f 2})$
$SU(3) \times U(1)$	$SU(2) \times U(1)$		5	$2_{\pm 1/2} + 1_0$
SU(4)	Sp(4)	~	5	${f 5}=({f 1},{f 1})+({f 2},{f 2})$
SU(4)	$[\mathrm{SU}(2)]^2 \times \mathrm{U}(1)$	√*	8	$(2,2)_{\pm 2} = 2 \cdot (2,2)$
SO(7)	SO(6)	~	6	${f 6}=2\cdot ({f 1},{f 1})+({f 2},{f 2})$
SO(7)	$G_2$	√*	7	${f 7}=({f 1},{f 3})+({f 2},{f 2})$
SO(7)	$SO(5) \times U(1)$	√*	10	$\mathbf{10_0} = (3, 1) + (1, 3) + (2, 2)$
SO(7)	$[SU(2)]^{3}$	√*	12	$({f 2},{f 2},{f 3})=3\cdot ({f 2},{f 2})$
Sp(6)	$\operatorname{Sp}(4) \times \operatorname{SU}(2)$	~	8	$({f 4},{f 2})=2\cdot ({f 2},{f 2})$
SU(5)	$SU(4) \times U(1)$	√*	8	${f 4}_{-5}+ar{f 4}_{+f 5}=2\cdot ({f 2},{f 2})$
SU(5)	SO(5)	√*	14	${f 14}=({f 3},{f 3})+({f 2},{f 2})+({f 1},{f 1})$
SO(8)	SO(7)	~	7	${f 7}=3\cdot ({f 1},{f 1})+({f 2},{f 2})$
SO(9)	SO(8)	~	8	${f 8}=2\cdot ({f 2},{f 2})$
SO(9)	$SO(5) \times SO(4)$	√*	20	$({f 5},{f 4})=({f 2},{f 2})+({f 1}+{f 3},{f 1}+{f 3})$
$[SU(3)]^2$	SU(3)		8	$8 = \mathbf{1_0} + \mathbf{2_{\pm 1/2}} + \mathbf{3_0}$
$[SO(5)]^2$	SO(5)	√*	10	${f 10}=({f 1},{f 3})+({f 3},{f 1})+({f 2},{f 2})$
$SU(4) \times U(1)$	$SU(3) \times U(1)$		7	$3_{-1/3} + \mathbf{\bar{3}}_{+1/3} + 1_0 = 3 \cdot 1_0 + 2_{\pm 1/2}$
SU(6)	Sp(6)	√*	14	${f 14}=2\cdot ({f 2},{f 2})+({f 1},{f 3})+3\cdot ({f 1},{f 1})$
$[SO(6)]^2$	SO(6)	√*	15	$15 = (1, 1) + 2 \cdot (2, 2) + (3, 1) + (1, 3)$

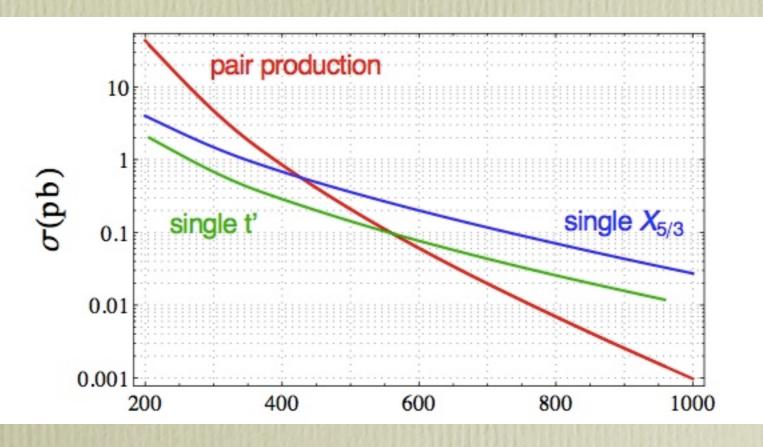
from 1401.2457 Csaki et al.

#### Compositness and flavour

- Extra vector-like quark multiplets often present (even exotic ones)
- Partial compositness (elementary/composite talk through mixings)
- Anarchic scenario (no hierarchies, SM ones generated by mixings of strongly composite top and elementary light quarks)
- MFV (light quarks strongly composite)

#### Vector-like quarks

- Unique window to test models (Xdim, composite, Little Higgs, SUSY)
- Reach at LHC substantial and only partially exploited
- Mixings with all the 3 SM generation important (production/decay)
- Single production dominant with present mass bound at LHC (~700 GeV)



#### see talk by L.Panizzi

# Mixing structure

- Key assumption: new fermions interact with the SM fermions via Yukawa interactions
- The Q-numbers of the new fermions under the weak  $SU(2)_L \times U(1)_Y$  gauge group are limited by interaction with the SM Higgs doublet and one of the SM fermions
- Possible Q-numbers :
- 1 SM-like singlet
- 3 doublets : 1 with SM hypercharge Y, the others Y+/-1
- 2 triplets with Y+/-1
- if more than one VL multiplet, inter-multiplet interactions and more general structure

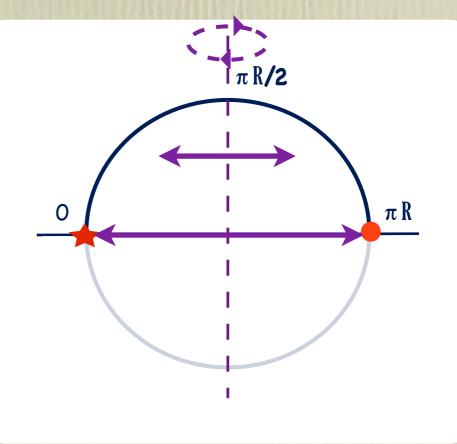
#### Extra dimensional models

- Typical Scenarios:
  - Large (flat) \ Warped Extra Dimensions
  - only gravity in the bulk, all SM fields in the bulk
- Issues you may explain (or describe geometrically) :
  - Weak scale stability: Gauge-Higgs unification
  - Fermion mass hierarchy, neutrino masses
  - Gauge symmetry breaking, strongly interacting conformal sector
  - Higgs composite models/technicolor

But : non-renormalizable, often large arbitrariness (localized interactions)

# Selecting Xdim models

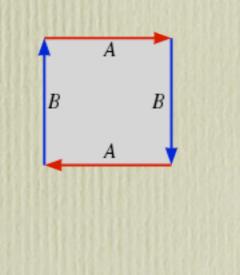
- Many models but a personal view:
  - promote (dark-)Matter parity to a fundamental orbifold symmetry
  - "natural" chiral fermions in the spectrum (no Z<sub>N</sub>-type quotient tricks)
  - no fixed points (source of arbitrary localized interactions)

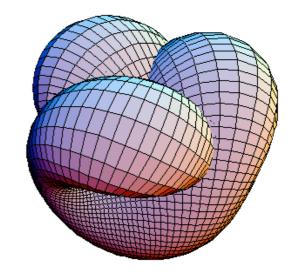


standard orbifold S<sup>1</sup>/Z<sub>2</sub>, KKparity ad hoc, fixed points are not the same point!

#### Orbifolds without fixed points

- One flat X-dim does not have candidates with chiral fermions and no fixed points
- 2 flat X-dim has the 17 wallpaper groups in the plane
  - Only 3 have no fixed points (Torus, Klein bottle, Real Projective Plane)
  - Only 1 has moreover chiral fermions (Real Projective Plane)
  - Isometries of X<sup>n</sup>: Noether theorem imposes selection rules

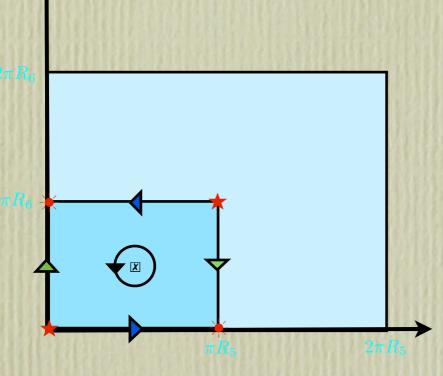


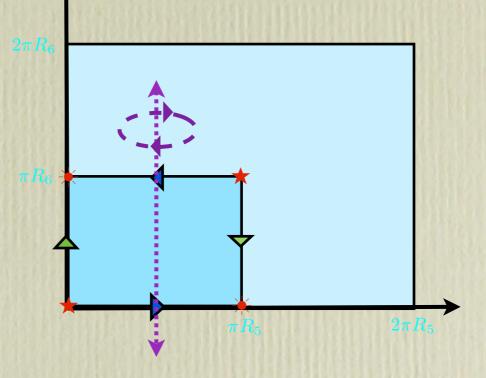


for details 0907.4993 1104.3800, 1210.0384

#### Selection rules in flat RPP

kk-modes (m,n)





 $p_{kk} = (-I)^{(m+n)}$ exact symmetry  $p_{kk} = (-I)^{(m)}$ violated by localised interactions

pictures courtesy of G.Cacciapaglia

# Beyond flat geometries

- RPP can be obtained from the sphere, but positive curvature implies massive eigenvalues for the Dirac operator. Extra gauge field can compensate the connection...but not than nice!
- Negative curvature more interesting (hyperbolic orbifolds M4 x H<sup>d</sup>/ $\Gamma$ ):
  - massless fermionic modes, large mass gap with KK modes
  - M<sub>p</sub> TeV exponential hierarchy
  - stability of the extra space (rigidity, only radion stabilization)
  - standard Friedman-Roberson-Walker cosmology

#### see N.Deutschmann's talk

### Bump hunt at LHC

- LHC at 7-8 TeV measured the Higgs boson (remember the no-loose "theorem" either Higgs or something else?)
- LHC at 13 TeV, more uncertain discovery of new particles, but well in the TeV range for testing extensions of the EW sector, putting bounds, challenging more naturalness

## Example of limits @LHC

#### ATLAS SUSY Searches\* - 95% CL Lower Limits

Status: SUSY 2013

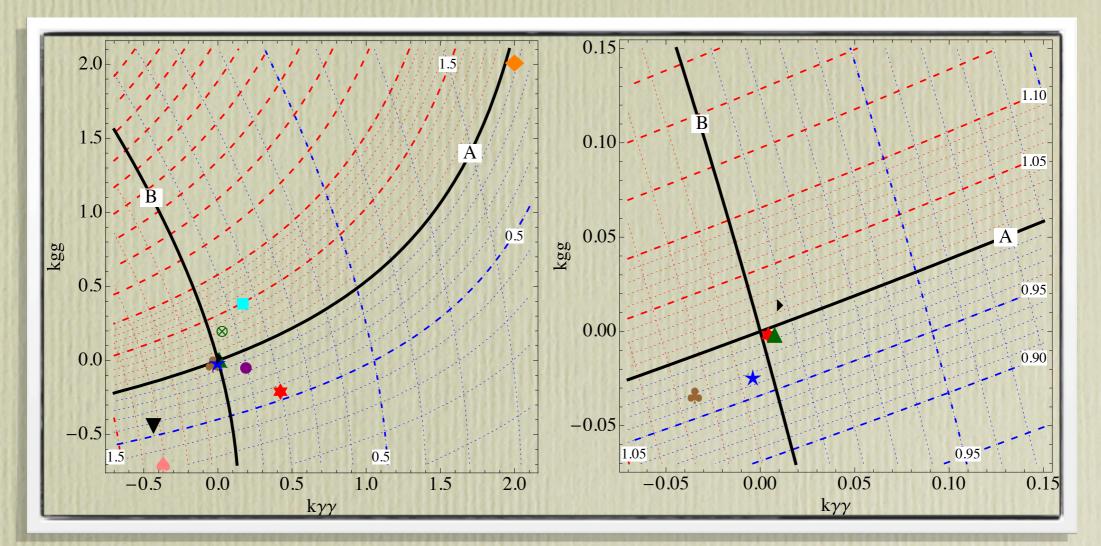
010	llus. 3031 2013					$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$	$\sqrt{s} = 7, 8 \text{ lev}$
	Model	e, μ, τ, γ	Jets	E <sup>miss</sup> T	∫£ dt[fb	-	Reference
Inclusive Searches	$ \begin{array}{l} \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \overline{q}\overline{q},\overline{q} \rightarrow q \overline{\chi}_{1}^{0} \\ \overline{g}\overline{g},\overline{g} \rightarrow q \overline{q} \overline{\chi}_{1}^{0} \\ \overline{g}\overline{g},\overline{g} \rightarrow q \overline{q} \overline{\chi}_{1}^{0} \\ \overline{g}\overline{g},\overline{g} \rightarrow q q \overline{\chi}_{1}^{0} \\ q q W^{\pm} \overline{\chi}_{1}^{0} \\ \overline{g}\overline{g},\overline{g} \rightarrow q q (\ell/(\nu/\nu)\nu) \overline{\chi}_{1}^{0} \\ \overline{g}\text{MSB}(\overline{\ell}\text{NLSP}) \\ \text{GMSB}(\overline{\ell}\text{NLSP}) \\ \text{GGM}(\text{bino}\text{NLSP}) \\ \text{GGM}(\text{bino}\text{NLSP}) \\ \text{GGM}(\text{higgsino-bino}\text{NLSP}) \\ \text{GGM}(\text{higgsino-bino}\text{NLSP}) \\ \text{GGM}(\text{higgsino}\text{NLSP}) \\ \text{Gravitino}\text{LSP} \end{array} $	$\begin{array}{c} 0 \\ 1 \ e, \mu \\ 0 \\ 0 \\ 0 \\ 1 \ e, \mu \\ 2 \ e, \mu \\ 2 \ e, \mu \\ 1 - 2 \ \tau \\ 2 \ \gamma \\ 1 \ e, \mu + \gamma \\ \gamma \\ 2 \ e, \mu \left( Z \right) \\ 0 \end{array}$	2-6 jets 3-6 jets 7-10 jets 2-6 jets 2-6 jets 3-6 jets 0-3 jets 0-2 jets 0-3 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 20.3	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	ATLAS-CONF-2013-047 ATLAS-CONF-2013-062 1308.1841 ATLAS-CONF-2013-047 ATLAS-CONF-2013-047 ATLAS-CONF-2013-062 ATLAS-CONF-2013-069 1208.4688 ATLAS-CONF-2013-026 1209.0753 ATLAS-CONF-2012-144 1211.1167 ATLAS-CONF-2012-152 ATLAS-CONF-2012-147
3 <sup>rd</sup> gen. ĝ med.	$\begin{array}{l} \tilde{g} \rightarrow b \bar{b} \tilde{\chi}_{1}^{0} \\ \tilde{g} \rightarrow t \bar{t} \tilde{\chi}_{1}^{0} \\ \tilde{g} \rightarrow t \bar{t} \tilde{\chi}_{1}^{1} \\ \tilde{g} \rightarrow b \bar{t} \tilde{\chi}_{1}^{1} \end{array}$	0 0 0-1 e,μ 0-1 e,μ	3 b 7-10 jets 3 b 3 b	Yes Yes Yes Yes	20.1 20.3 20.1 20.1	Ř         1.2 TeV         m( $\tilde{k}_{1}^{0}$ )<600 GeV           Ř         1.1 TeV         m( $\tilde{k}_{1}^{0}$ )<350 GeV           Ř         1.34 TeV         m( $\tilde{k}_{1}^{0}$ )<400 GeV           Ř         1.3 TeV         m( $\tilde{k}_{1}^{0}$ )<300 GeV	ATLAS-CONF-2013-061 1308.1841 ATLAS-CONF-2013-061 ATLAS-CONF-2013-061
3rd gen. squarks direct production	$ \begin{array}{l} \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \rightarrow b\tilde{\chi}_{1}^{0} \\ \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}(\text{light}), \tilde{t}_{1} \rightarrow b\tilde{\chi}_{1}^{1} \\ \tilde{t}_{1}\tilde{t}_{2}(\text{light}), \tilde{t}_{1} \rightarrow Wb\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{3}(\text{medium}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}(\text{medium}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}(\text{nedium}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}(\text{heavy}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{3}(\text{heavy}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{2}(\text{heavy}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{3}(\text{heavy}), \tilde{t}_{1} \rightarrow t\tilde{\chi}_{1}^{0} \\ \tilde{t}_{2}\tilde{t}_{2}\tilde{t}_{2}, \tilde{t}_{2} \rightarrow \tilde{t}_{1} + Z \end{array} $	$\begin{array}{c} 0\\ 2\ e,\mu\ ({\rm SS})\\ 1\text{-}2\ e,\mu\\ 2\ e,\mu\\ 2\ e,\mu\\ 0\\ 1\ e,\mu\\ 0\\ 3\ e,\mu\ (Z) \end{array}$	2 b 0-3 b 1-2 b 0-2 jets 2 jets 2 b 1 b 2 b ono-jet/c-t: 1 b 1 b	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.1 20.7 4.7 20.3 20.3 20.1 20.7 20.5 20.3 20.7 20.7	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	1308.2631 ATLAS-CONF-2013-007 1208.4305, 1209.2102 ATLAS-CONF-2013-048 ATLAS-CONF-2013-048 1308.2631 ATLAS-CONF-2013-037 ATLAS-CONF-2013-024 ATLAS-CONF-2013-025 ATLAS-CONF-2013-025
EW direct	$ \begin{array}{l} \tilde{\ell}_{\mathbb{L},\mathbb{R}}\tilde{\ell}_{\mathbb{L},\mathbb{R}},\tilde{\ell} \rightarrow \ell\tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-},\tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell}\nu(\ell\tilde{\nu}) \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-},\tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau}\nu(\tau\tilde{\nu}) \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{\nu}\nu\tilde{\ell}_{\nu}\ell(\tilde{\nu}\nu),\ell\tilde{\nu}\tilde{\ell}_{\nu}\ell(\tilde{\nu}\nu) \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{2}^{0} \rightarrow W\tilde{\chi}_{1}^{0}Z\tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+}\tilde{\chi}_{2}^{0} \rightarrow W\tilde{\chi}_{1}^{0}h\tilde{\chi}_{1}^{0} \end{array} $	2 e,μ 2 e,μ 2 τ 3 e,μ 3 e,μ 1 e,μ	0 0 - 0 2 b	Yes Yes Yes Yes Yes Yes	20.3 20.3 20.7 20.7 20.7 20.7 20.3	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	ATLAS-CONF-2013-049 ATLAS-CONF-2013-049 ATLAS-CONF-2013-028 ATLAS-CONF-2013-035 ATLAS-CONF-2013-035 ATLAS-CONF-2013-093
Long-lived particles	$\begin{array}{l} \text{Direct} \tilde{\chi}_1^+ \tilde{\chi}_1^- \operatorname{prod., long-lived} \tilde{\chi}_1^\pm \\ \text{Stable, stopped} \; \tilde{g} \; \text{R-hadron} \\ \text{GMSB, stable} \; \tilde{\tau}, \tilde{\chi}_1^0 {\rightarrow} \tilde{\tau}(\tilde{e}, \tilde{\mu}) {+} \tau ( \\ \text{GMSB,} \; \tilde{\chi}_1^0 {\rightarrow} \gamma \tilde{G}, \operatorname{long-lived} \tilde{\chi}_1^0 \\ \tilde{q} \; \tilde{q}, \; \tilde{\chi}_1^0 {\rightarrow} q q \mu \; (\text{RPV}) \end{array}$	Disapp. trk 0 e, μ) 1-2 μ 2 γ 1 μ, displ. vtx	1 jet 1-5 jets - -	Yes Yes Yes -	20.3 22.9 15.9 4.7 20.3	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	ATLAS-CONF-2013-069 ATLAS-CONF-2013-057 ATLAS-CONF-2013-058 1304.6310 ATLAS-CONF-2013-092
RPV	$ \begin{array}{l} LFV \ pp \rightarrow \tilde{v}_{\tau} + X, \ \tilde{v}_{\tau} \rightarrow e + \mu \\ LFV \ pp \rightarrow \tilde{v}_{\tau} + X, \ \tilde{v}_{\tau} \rightarrow e(\mu) + \tau \\ Bilinear \ RPV \ CMSSM \\ \tilde{\chi}_1^+ \tilde{\chi}_1^-, \ \tilde{\chi}_1^+ \rightarrow W \tilde{\chi}_1^0, \ \tilde{\chi}_1^0 \rightarrow ee \tilde{v}_{\mu}, e \mu \tilde{v} \\ \tilde{\chi}_1^+ \tilde{\chi}_1, \ \tilde{\chi}_1^+ \rightarrow W \tilde{\chi}_1^0, \ \tilde{\chi}_1^0 \rightarrow \tau \tau \tilde{v}_e, e \tau \tilde{v} \\ \tilde{g} \rightarrow q q \\ \tilde{g} \rightarrow \tilde{t}_1 t, \ \tilde{t}_1 \rightarrow bs \end{array} $	$\begin{array}{c} 2 \ e, \mu \\ 1 \ e, \mu + \tau \\ 1 \ e, \mu \\ e \\ \tau \\ 3 \ e, \mu + \tau \\ 0 \\ 2 \ e, \mu \ (\text{SS}) \end{array}$	- 7 jets - - 6-7 jets 0-3 b	- Yes Yes - Yes	4.6 4.7 20.7 20.7 20.3 20.7	\$\vec{v}_r\$         1.61 TeV $\lambda_{311}^2$ =0.10, $\lambda_{132}$ =0.05           \$\vec{v}_r\$         1.1 TeV $\lambda_{311}^2$ =0.10, $\lambda_{12233}$ =0.05           \$\vec{v}_r\$         1.2 TeV         m(\$\vec{v}_1\$)=m(\$\vec{v}_r\$), \$c_{1,25}\$           \$\vec{v}_r\$         760 GeV         m(\$\vec{v}_1\$)=300 GeV, \$\lambda_{121}\$>0           \$\vec{v}_r\$         350 GeV         m(\$\vec{v}_1\$)=300 GeV, \$\lambda_{121}\$>0           \$\vec{v}_r\$         916 GeV         BR(t)=BR(c)=0%           \$\vec{v}_r\$         880 GeV         BR(t)=BR(c)=0%	1212.1272 1212.1272 ATLAS-CONF-2012-140 ATLAS-CONF-2013-036 ATLAS-CONF-2013-036 ATLAS-CONF-2013-091 ATLAS-CONF-2013-007
Other		$ \begin{array}{c} 0\\2 e, \mu (SS)\\0\\ \sqrt{s} = 8 \text{ TeV}\\ \text{artial data} \end{array} $	4 jets 1 <i>b</i> mono-jet √s = 3 full o	Yes Yes 8 TeV data	4.6 14.3 10.5	sgluon         100-287 GeV         incl. limit from 1110.2693           sgluon         800 GeV         m(¿)<80 GeV, limit of <687 GeV for D8	1210.4826 ATLAS-CONF-2013-051 ATLAS-CONF-2012-147

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1\sigma theoretical signal cross section uncertainty.

#### $\mathbf{ATLAS} \text{ Preliminary}$ $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \quad \sqrt{s} = 7, 8 \text{ TeV}$

- I TeV but read the fine prints!!! NOT general bounds

#### Large BSM effects only in few cases



#### from hep-ph 0901.0927 (in unsuspicious times...)

	4 <sup>th</sup>	*	Littlest Higgs	•	Warped GHU Space	$\star$	Flat GHU
•	SUSY gold		Color Octet		Flat BH with Flavour		UED Model
	SLH		Lee Wick SM	٨	Warped BH with Flavour		

### Sort of conclusion: BSM hints?

- naive/oversimplified models are under serious attack
- Hints may hide in details... Experimental bounds should be kept as general as possible (we may overlook interesting physics)
- model building has new challenges and opportunities