

Updated SM predictions of $B_q \rightarrow \ell^+ \ell^-$ at NLO EW and NNLO QCD

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– arXiv:1311.0903 –

Outline

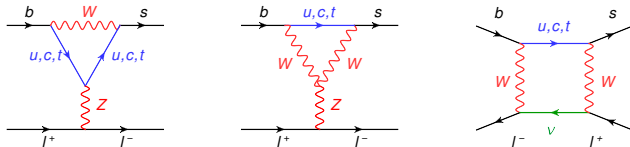
I) NNLO QCD corrections to C_A

II) Updated $Br(B_q \rightarrow \ell^+ \ell^-)$ @ NNLO QCD + NLO EW

!!! Throughout only Standard Model (=SM) is considered

Motivation to study $B_q \rightarrow \ell^+ \ell^-$

- test of the SM at loop-level (FCNC decay) \Rightarrow **loop-suppressed**



\Rightarrow @ LO in QED, hadronic uncertainty only from $B_{d,s}$ decay constant

- in addition **helicity suppressed**
 \Rightarrow sensitive to non-SM (pseudo-) scalar interactions
- important B -decay @ LHCb, CMS & ATLAS \Rightarrow **first measurements from 2013:**

$$\overline{B}(B_s \rightarrow \mu^+ \mu^-) = (2.9 \pm 0.7) \times 10^{-9}$$

$$\overline{B}(B_d \rightarrow \mu^+ \mu^-) = (3.6_{-1.4}^{+1.6}) \times 10^{-10}$$

LHCb (3/fb) + CMS (25/fb)

[LHCb arXiv:1307.5024]

[CMS arXiv:1307.5025]

- exp. prospects for $B_s \rightarrow \mu^+ \mu^-$:

@ LHCb with 50 fb^{-1} : $\sim 0.15 \times 10^{-9} = 5\%$ error of SM (only stat. err) [LHCb arXiv:1208.3355]

@ CMS with 100 fb^{-1} : 15% error of SM

[Kai-Feng Chen, KEK Flavor Factory WS, 2014]

Higher order QCD and EW
corrections to $B_q \rightarrow l^+ l^-$

B -Meson decays are a Multi-scale problem . . .

. . . with hierarchical interaction scales

electroweak IA

\gg

ext. mom'a in B restframe

\gg

QCD-bound state effects

$$M_W \approx 80 \text{ GeV}$$

$$M_Z \approx 91 \text{ GeV}$$

$$M_B \approx 5 \text{ GeV}$$

$$\Lambda_{\text{QCD}} \approx 0.5 \text{ GeV}$$

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OPE = expansion in

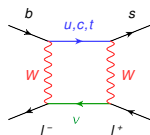
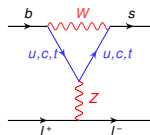
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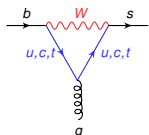
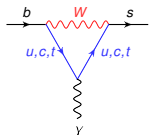
$$M_B \approx 5 \text{ GeV}$$

$$M_B^2/M_W^2 \sim 0.004$$

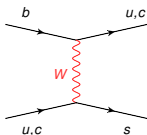
$$b \rightarrow s \ell^+ \ell^-$$



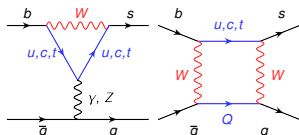
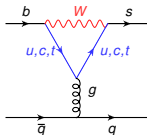
$$b \rightarrow s \gamma \text{ \& } b \rightarrow s g$$



$$b \rightarrow s (u\bar{u}, c\bar{c})$$



$$b \rightarrow s q \bar{q} (q = u, d, s, c, b)$$



B-Meson decays are a Multi-scale problem ...

... with hierarchical interaction scales

electroweak IA

» ext. mom'a in B restframe

The extension of Fermi's theory of weak interactions to B decays

$M_W \approx 80 \text{ GeV}$

$M_Z \approx 91 \text{ GeV}$

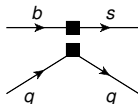
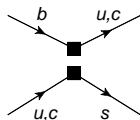
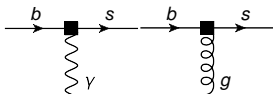
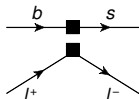
$M_B \approx 5 \text{ GeV}$

semi-leptonic

electro- & chromo-mgn

charged current

QCD & QED -penguin



C_i = **Wilson coefficients**: contains short-dist. pmr's (heavy masses M_t, \dots – CKM factored out) and leading logarithmic QCD-corrections to all orders in α_s

⇒ in SM known up to next-to-next-to-leading order in QCD

O_i = **higher-dim. operators**: flavour-changing coupling of light quarks (+ gluons and photons)

$$\mathcal{L}_{\text{eff}}^{\Delta B=1} \propto V_{CKM} \left[\sum_{V,A} C_i O_i^{\ell\bar{\ell}} + \sum_{7\gamma, 8g} C_i O_i + \text{CC} + (\text{QCD \& QED-peng}) \right] + \mathcal{O} \left(\frac{M_B^2}{M_W^2} \right)$$

“Matching” = Determination of Wilson coefficients

⇒ C_i are determined by requiring equality of full theory (=SM) and effective theory (=EFT) amplitudes order by order in expansion in couplings G_F, α_s (QCD) and α_e (QED)

relevant for $B_s \rightarrow \ell^+ \ell^-$

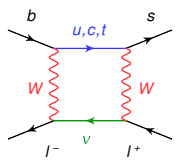
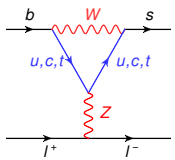
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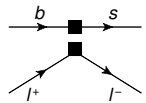
relevant for $B_s \rightarrow \ell^+ \ell^-$

$$G_F \frac{\alpha_e}{s_W^2}$$

LO



≡

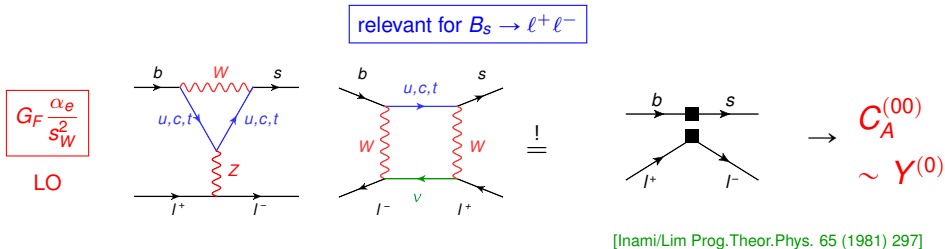


$$\rightarrow C_A^{(00)} \sim Y^{(0)}$$

[Inami/Lim Prog.Theor.Phys. 65 (1981) 297]

“Matching” = Determination of Wilson coefficients

⇒ C_i are determined by requiring equality of full theory (=SM) and effective theory (=EFT) amplitudes order by order in expansion in couplings G_F, α_s (QCD) and α_e (QED)



EFT is “simple” at LO EW (& all orders in QCD), because **only one operator**

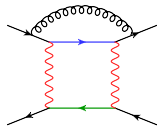
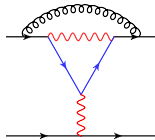
$$\mathcal{O}_A^{\ell\bar{\ell}} = [\bar{s}\gamma_\mu\gamma_5 b][\bar{\ell}\gamma^\mu\gamma_5 \ell] \quad \left(\text{usually known as } \mathcal{O}_{10}^{\ell\bar{\ell}} : C_A = -\frac{C_{10}}{2} \right)$$

contributes to $B_q \rightarrow \ell^+ \ell^- \Rightarrow \mathcal{O}_A^{\ell\bar{\ell}}$ is a **conserved current under QCD** (→ no “Running”)

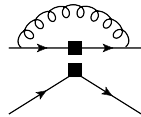
!!! (pseudo-) scalar $\mathcal{O}_{S(P)}^{\ell\bar{\ell}} \propto [\bar{s}\gamma_5 b][\bar{\ell}1(\gamma_5)\ell]$ from Higgs-penguins are suppressed by $M_{B_s}^2/M_W^2$

Adding QCD corrections ...

$$G_F \frac{\alpha_e}{S_W^2} \alpha_s$$



!



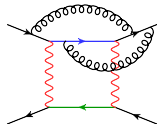
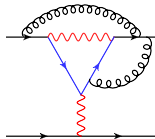
→

$$C_A^{(10)}$$

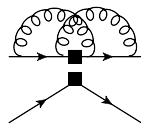
NLO QCD

[Buchalla/Buras NPB 398 (1993) 285, NPB 400 (1993) 225, Misiak/Urban hep-ph/9901278]

$$G_F \frac{\alpha_e}{S_W^2} \alpha_s^2$$



!



→

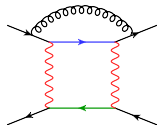
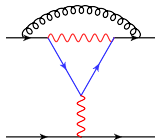
$$C_A^{(20)}$$

NNLO QCD

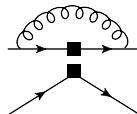
[Hermann/Misiak/Steinhauser arXiv:1311.1347]

Adding QCD corrections ...

$$G_F \frac{\alpha_e}{S_W^2} \alpha_s$$



≡

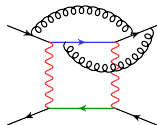
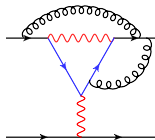


$$\rightarrow C_A^{(10)}$$

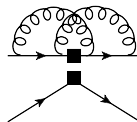
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≡



$$\rightarrow C_A^{(20)}$$

NNLO QCD

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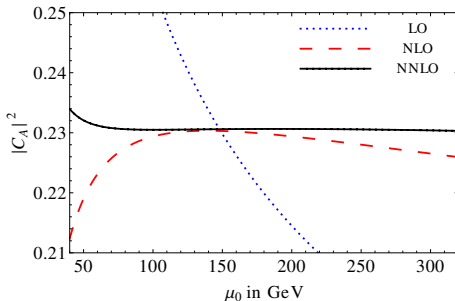
$$C_A = C_A^{(00)} + \frac{\alpha_s}{4\pi} C_A^{(10)} + \left(\frac{\alpha_s}{4\pi}\right)^2 C_A^{(20)}$$

implicite – via $m_t(\mu_0)$ (= $\overline{\text{MS}}$ QCD) – and
 explicite μ_0 dependence

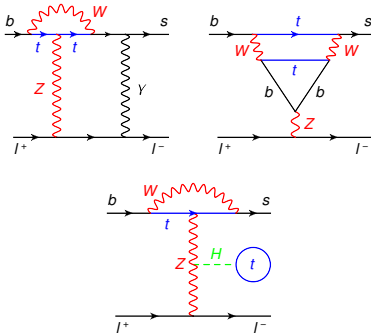
$$\begin{aligned} m_t(m_t) &= 163.5 \text{ GeV} \\ m_t(\mu_0 = 50 \text{ GeV}) &= 180.8 \text{ GeV} \\ m_t(\mu_0 = 300 \text{ GeV}) &= 156.2 \text{ GeV} \end{aligned}$$

⇒ NNLO QCD crss. reduce μ_0 -dep. from

$$1.8\% \text{ at NLO} \rightarrow 0.2\% \text{ at NNLO} \quad \text{on } Br \sim |C_A|^2$$



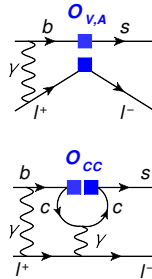
Adding NLO EW corrections ...



$$G_F \frac{\alpha_e}{S_W^2} \alpha_e$$

NLO EW

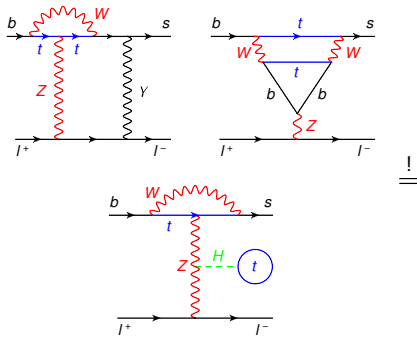
!



→ $C_A^{(01)}$

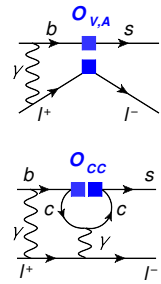
[Bobeth/Gorbahn/Stamou arXiv:1311.1348]

Adding NLO EW corrections ...



$$G_F \frac{\alpha_e}{S_W^2} \alpha_e$$

NLO EW



$$\rightarrow C_A^{(01)}$$

[Bobeth/Gorbahn/Stamou arXiv:1311.1348]

See talk by Martin Gorbahn

Including QED-Log's + combination with NNLO QCD

⇒ Log-enhanced QED corrections known

[Bobeth/Gambino/Gorbahn/Haisch hep-ph/0312090,
Huber/Lunghi/Misiak/Wyler hep-ph/0512066]

Choose OS-2 as default scheme

Solution of RGE

$$C_A(\mu_b) = \sum_i [U(\mu_b, \mu_0)]_{A,i} C_i(\mu_0)$$

→ $U(\mu_b, \mu_0)$ = evolution operator

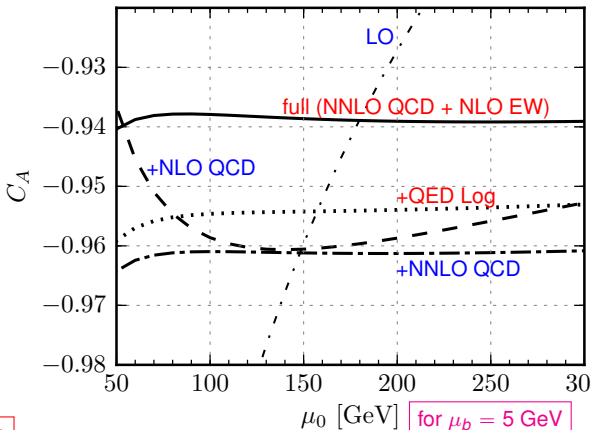
→ $C_i(\mu_0)$ = Wilson coefficients at high scale

!!! $C_A(\mu_b)$ μ_0 -independent

⇒ full EW corr's reduce Br by 4% compared to NNLO QCD

Estimate of higher order uncertainties

- 1) μ_0 variation between $[m_t/2, 2m_t]$ → about 0.2% (QCD) + 0.2% (EW) on $|C_A(\mu_b)|^2$
- 2) additional EW scheme dep. from diff. of OS-2 and HY scheme → about 0.2%

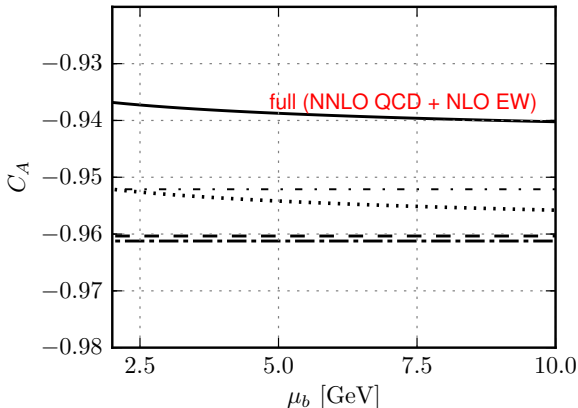


Residual μ_b dependence of C_A ...

... gives measure of uncertainties due to lacking virtual QED corrections to matrix elements of operators:

$\mathcal{O}_{A,V,cc}$

0.3% residual
 μ_b -dependence
on $|C_A(\mu_b)|^2$



Updated $Br(B_q \rightarrow \ell^+ \ell^-)$ at
NNLO QCD + NLO EW

Branching ratio $B_q \rightarrow \ell^+ \ell^-$

Up to now discussed (NNLO QCD + NLO EW) corrections to $C_A(\mu_b)$

\Rightarrow need to calculate (matrix elements)² to obtain Br including NLO QED

$$Br \sim |\langle \ell \bar{\ell} | \mathcal{H}_{\text{eff}} | \bar{B} \rangle|^2 \sim |C_A(\mu_b)|^2 \langle 0 | \bar{s} \gamma_\mu \gamma_5 b | \bar{B} \rangle^2 + \mathcal{O}(\alpha_e)$$

1) Hadronic matrix element (to all orders in QCD and LO in QED):

$$\langle 0 | \bar{q} \gamma^\mu \gamma_5 b | \bar{B}_q(p_B) \rangle \equiv i f_{B_q} p_B^\mu$$

$$f_{B_d} = (190.5 \pm 4.2) \text{ MeV}$$

$$f_{B_s} = (227.7 \pm 4.5) \text{ MeV}$$

[FLAG (lattice average) 2013]

2) Account for B_s -mixing: fully time-integrated CP-averaged Br

[De Bruyn et al. arXiv:1204.1737]

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!!! for consistency, should include $\mathcal{O}(\alpha_e)$ corrections:

A) bremsstrahlung

B) virtual corrections

can be sizeable depending on photon energy cuts in experiment

expected to be small $\alpha_e/\pi \sim 1/400 \sim 0.25\%$, but non-perturbative → no theoretical framework

$$\overline{B}(B_s \rightarrow \mu^+ \mu^-) = (3.65 \pm 0.23) \times 10^{-9}$$

$$\overline{B}(B_d \rightarrow \mu^+ \mu^-) = (1.06 \pm 0.09) \times 10^{-10}$$

$$\overline{B}(B_s \rightarrow e^+ e^-) = (8.54 \pm 0.55) \times 10^{-14}, \quad \overline{B}(B_s \rightarrow \tau^+ \tau^-) = (7.73 \pm 0.49) \times 10^{-7},$$

$$\overline{B}(B_d \rightarrow e^+ e^-) = (2.48 \pm 0.21) \times 10^{-15}, \quad \overline{B}(B_d \rightarrow \tau^+ \tau^-) = (2.22 \pm 0.19) \times 10^{-8}$$

[CB/Gorbahn/Hermann/Misiak/Stamou/Steinhauser arXiv:1311.0903]

Error budget	f_{B_q}	CKM	τ_H^q	M_t	α_s	non-param.	other param.	Σ
$\overline{B}_{S\mu}$	4.0%	4.3%	1.3%	1.6%	0.1%	1.5%	< 0.1%	6.4%
$\overline{B}_{d\mu}$	4.5%	6.9%	0.5%	1.6%	0.1%	1.5%	< 0.1%	8.5%

⇒ CKM: use here $|V_{cb}|_{\text{incl}} = 0.0424 \pm 0.0009$
rather than $|V_{cb}|_{\text{excl}} = 0.0396 \pm 0.0009$

[Gambino/Schwanda arXiv:1307.4551]

as used as input by [UTfit post-EPS13]

⇒ non-parametric uncertainties

- 0.3% from $\mathcal{O}(\alpha_{em})$ corrections from $\mu_b \in [m_b/2, 2m_b]$
- $2 \times 0.2\%$ from $\mathcal{O}(\alpha_s^3, \alpha_{em}^2, \alpha_s \alpha_{em})$ matching corrections from $\mu_0 \in [m_t/2, 2m_t]$
- 0.3% from top-mass conversion from on-shell to $\overline{\text{MS}}$ scheme
- 0.5% further uncertainties (power corrections $\mathcal{O}(m_b^2/M_W^2), \dots$)

Another precise SM prediction @ (NLO EW + NNLO QCD)

Ratio of $\overline{B}(B_q \rightarrow \ell^+ \ell^-)$ and B_q -mass difference ΔM_{B_q}

$$\kappa_{q\ell} \equiv \frac{\overline{B}(B_q \rightarrow \ell^+ \ell^-)}{\Delta M_{B_q}} \frac{\Gamma_H^q}{(G_F M_W m_\ell)^2 \beta_{q\ell}} \stackrel{\text{(SM)}}{=} \frac{3 |C_A(\mu_b)|^2}{\pi^3 C_{LL}(\mu_b) B_{B_q}(\mu_b)}$$

NLO EW corr. to ΔM_q [Gambino/Kwiatkowski/Pott hep-ph/9810400]

- decay constant f_{B_q} and CKM cancel in ratio
- BUT depends on
 - $\Rightarrow B_{B_q}(\mu_b)$ bag factor of $\Delta B = 2$ hadronic matrix element
- depends on perturbatively calculable
 - $\Rightarrow C_{LL}(\mu_b)$ Wilson coefficient of $\Delta B = 2$ operator

SM prediction with comparable theory uncertainty to $\overline{B}(B_q \rightarrow \ell^+ \ell^-)$

\Rightarrow dominated by uncertainties of lattice predictions of $B_{B_q}(\mu_b)$

$$\kappa_{sl} = (1.26 \pm 0.07)\%$$

$$\kappa_{dl} = (1.32 \pm 0.12)\%$$

Conclusion

SM prediction of $\overline{B}(B_s \rightarrow \mu^+ \mu^-)$ has $\lesssim 7\%$ theory uncertainty

\Rightarrow mainly f_{B_s} (4.0%), V_{cb} (4.3%), non-parametric (1.5%)

included are

- NNLO QCD crrs. reduce μ_0 -dep. of $m_t(\mu_0)$ from 1.8% at NLO \rightarrow 0.2% at NNLO
- NLO EW crrs. enhanced by: $\frac{1}{\sin^2 \theta_W}$, $\frac{m_t^2}{M_W^2}$, $\ln^2 \frac{m_b^2}{M_W^2}$
reduce LO-scheme dependence from $\gtrsim 8\% \rightarrow 0.6\% \lesssim$ at NLO
- the size of NLO EW crrs. is about $\sim (3 \dots 5)\%$ (depending on μ_0)
- photon bremsstrahlung corrections due to PHOTOS

Lacking: Virtual QED corrections to matrix elements $\sim \alpha_e/\pi$

\Rightarrow expected to be small from μ_b -variation

\Rightarrow HOWEVER, nonperturbative physics involved below scales $\mu_b \sim m_b$

Backup Slides

Impact of NLO EW crrs. at μ_0 . . . neglecting operator mixing

⇒ α_e always $\overline{\text{MS}}$ and m_t always $\overline{\text{MS}}$ w.r.t QCD

⇒ 3 different EW renormalization schemes:

OS) masses and ($s_W^2 = 1 - M_W^2/M_Z^2$) on-shell renormalized

$\overline{\text{MS}}$) masses and s_W^2 in minimal subtraction

HY) masses on-shell but s_W^2 in minimal subtraction

⇒ 2 normalizations of \mathcal{L}_{eff}

[Misiak arXiv:1112.5978]

$$1) \quad \frac{4G_F}{\sqrt{2}} \frac{\alpha_e}{s_W^2} \left[C_A^{(00)} + \frac{\alpha_e(\mu_0)}{4\pi} C_A^{(01)}(\mu_0) \right] \quad 2) \quad \frac{G_F^2 M_W^2}{\pi^2} \left[C_A^{(00)} + \frac{\alpha_e(\mu_0)}{4\pi} C_A^{(01)}(\mu_0) \right]$$

Impact of NLO EW crrs. at $\mu_0 \dots$ neglecting operator mixing

$\Rightarrow \alpha_e$ always $\overline{\text{MS}}$ and m_t always $\overline{\text{MS}}$ w.r.t QCD

\Rightarrow 3 different EW renormalization schemes:

OS) masses and ($s_W^2 = 1 - M_W^2/M_Z^2$) on-shell renormalized

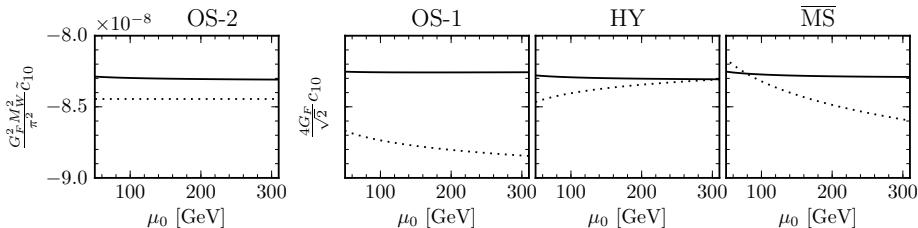
$\overline{\text{MS}}$) masses and s_W^2 in minimal subtraction

HY) masses on-shell but s_W^2 in minimal subtraction

\Rightarrow 2 normalizations of \mathcal{L}_{eff}

[Misiak arXiv:1112.5978]

$$1) \frac{4G_F}{\sqrt{2}} \frac{\alpha_e}{s_W^2} \left[C_A^{(00)} + \frac{\alpha_e(\mu_0)}{4\pi} C_A^{(01)}(\mu_0) \right] \quad 2) \frac{G_F^2 M_W^2}{\pi^2} \left[C_A^{(00)} + \frac{\alpha_e(\mu_0)}{4\pi} C_A^{(01)}(\mu_0) \right]$$

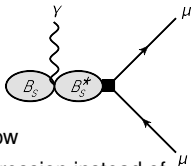


scheme dependence @ LO: $[-8.9, -8.2] \rightarrow \pm 8\%$ at Br
 reduces @ NLO: $[-8.31, -8.25] \rightarrow \pm 0.8\%$ at Br

LO = dotted
 NLO = solid

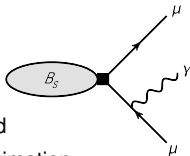
Bremsstrahlung: initial and final state

Initial State Radiation



- tiny in signal window
- phase-space suppression instead of helicity suppression
- can be avoided with cuts

Final State Radiation



- helicity suppressed
- soft-photon approximation
- extrapolated from signal window over all $m_{\mu^+\mu^-}^2$ via PHOTOS by LHCb and CMS

ISR

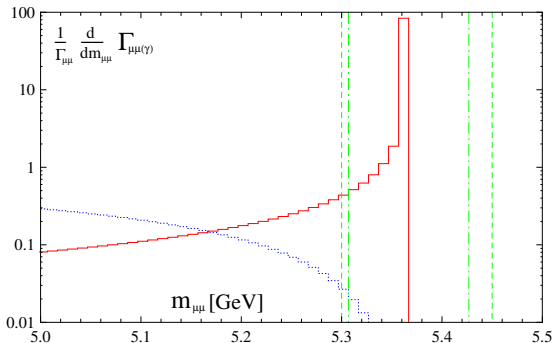
[Aditya/Healey/Petrov arXiv:1212.4166]

FSR

[Buras et al. arXiv:1208.0934]

experimental signal windows (LHCb, CMS)

[LHCb arXiv:1307.5024,
CMS arXiv:1307.5025]



What should theorists include in predictions of Br ?

- no need to include ISR \rightarrow currently removed by cuts
 - !!! it should be counted as background in MC
 - ??? should not be included in PHOTOS in experimental analysis
(especially when experimental accuracy increases)
- FSR should be included due to cuts in experimental analysis
BUT already accounted for by LHCb and CMS using PHOTOS
(extrapolation along red curve to zero on previous slide)
 \Rightarrow this corresponds to the limit where:

$$\text{photon-inclusive } Br = \text{non-radiative } Br$$

[Buras/Girrbach/Guadagnoli/Isidori arXiv:1208.0934]

- virtual corrections \rightarrow requires nonperturbative method
 \Rightarrow NOT included \rightarrow assign 0.3% uncertainty from μ_b -variation