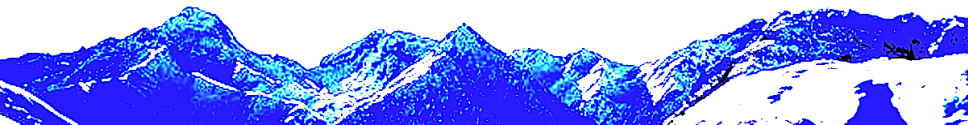


Moriond 2014

Gravitino DM and Leptogenesis after the first LHC results

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University



Introduction

- R -parity conserving Supersymmetry well motivated (DM, stability of EW scale, gauge unification)

- SUSY does not address baryon asymmetry

$$\frac{n_B}{n_\gamma} \simeq \frac{n_B - n_{\bar{B}}}{n_\gamma} \simeq 6 \times 10^{-10}$$

- Attractive: Thermal Leptogenesis [Fukugita, Yanagida '86]

- Leptogenesis requires: $T_R \gtrsim 10^9 \text{ GeV}$

[Davidson, Ibarra '02; Buchmüller, Di Bari, and Plümacher '04]

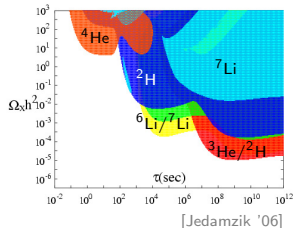
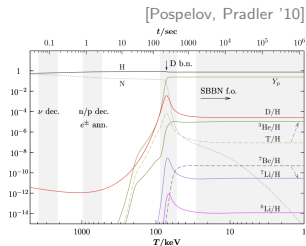
- Production of \tilde{G} in the early universe $\Omega_{\tilde{G}} \sim \frac{T_R}{m_{\tilde{G}}}$

[Bolz, Buchmüller, Plümacher '98; Bolz, Brandenburg, Buchmüller '00; Pradler, Steffen '07]

- Leptogenesis and Supersymmetry: Gravitino Problem

The Gravitino Problem

- Neutralino LSP ($m_{\tilde{G}} \sim m_{\text{soft}}$):
 - Decays $\tilde{G} \rightarrow \text{LSP} + X$ suppressed, \tilde{G} long-lived
 - Decay during/after BBN distorts predictions
 - $\Omega_{\tilde{G}} \sim \frac{T_R}{m_{\tilde{G}}} \rightarrow$ no Leptogenesis!



The Gravitino Problem

Gravitino LSP:

Gravitino stable

NLSP $\rightarrow \tilde{G} + X$ suppressed:

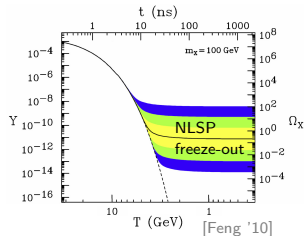
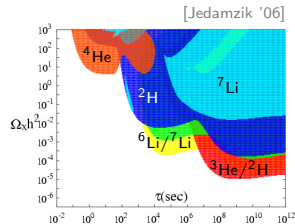
$$\tau_{\text{NLSP}} \sim \frac{M_{\text{Pl}}^2 m_{\tilde{G}}^2}{m_{\tilde{\tau}_1}^5}$$

Gravitino DM: $\Omega_{\tilde{G}} \stackrel{!}{=} \Omega_{\text{DM}}$

$$\Omega_{\tilde{G}} \sim \frac{T_{\text{R}}}{m_{\tilde{G}}} + (\text{non-therm. contr.})$$

If NLSP \in MSSM: freeze-out!

Consider yield $Y = n/s \sim \Omega/m$



The Gravitino Problem

□ Gravitino LSP:

□ Gravitino stable

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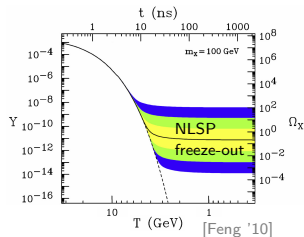
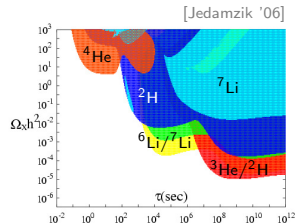
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□ If NLSP \in MSSM: freeze-out!

□ Consider yield $Y = n/s \sim \Omega/m$

\rightarrow Implications from LHC 7/8
and other recent experiments?



- Implications for possible NLSP yields

Based on J. Kersten, B. Panes, T. Robens, JH: 1310.2825, acc. by JHEP

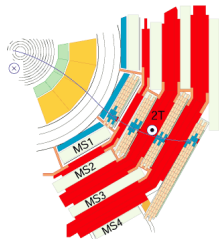
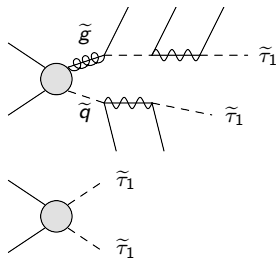
- Implications for the reheating temperature

Based on JH: 1310.6352 sub. to JCAP

Choice of NLSP

Consider NLSP = $\tilde{\tau}_1$:

- ❑ Direct SUSY searches = HSCP searches (no MET!)
- ❑ Clean signature \rightarrow high sensitivity
- ❑ Robust mass bounds possible
- ❑ E.g. $m_{\tilde{\tau}_1} \gtrsim 340 \text{ GeV}$ [CMS 2013]
- ❑ Great prospects for LHC
13/14



pMSSM Monte Carlo scan

- Monte Carlo scan over 17-dim. pMSSM

$$A_t, A_b, A_\tau; \mu, \tan \beta, m_A; M_1, M_2, M_3; \theta_{\tilde{\tau}}, m_{\tilde{\tau}_1}; \theta_{\tilde{t}}, m_{\tilde{t}_1}, m_{\tilde{b}_1}; \\ m_{\tilde{L}_{1,2}}, m_{\tilde{e}_{1,2}}, m_{\tilde{Q}_{1,2}} = m_{\tilde{u}_{1,2}} = m_{\tilde{d}_{1,2}}$$

- Interpret Higgs discovery in MSSM:

$$123 \text{ GeV} < m_{h/H} < 128 \text{ GeV}$$

[ATLAS, CMS '12]

- Used tools:

- Spectrum, Higgs decays and precision observables: SUSPECT, FEYNHIGGS
 - Decay tables: SDECAY, WHIZARD
 - Cross sections: Fast XS estimation based on PROSPINO and NLL FAST, WHIZARD
 - Stau yield and flavor observables: MICROMEGAS
- No 'blind' scan – systematically scan co-annihilation regions, resonances, larger mixings

pMSSM scan – constraints

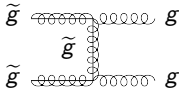
- ❑ Interpretation of the HSCP search at the 7 and 8 TeV LHC
 - ❑ Consider all SUSY xs (also $pp \rightarrow h/H \rightarrow \tilde{\tau}_1 \tilde{\tau}_1$)
 - ❑ Estimated $\sigma_{\text{limit}}^{\text{obs}}$ for each point from [CMS Collaboration '13]
- ❑ MSSM Higgs searches at the LHC, Tevatron and LEP via HIGGSBOUNDS 4.0.0
- ❑ Flavor and precision observables
 - ❑ $M_W = 80.385 \pm 0.060 \text{ GeV} @ 95\% \text{ C.L. (Exp.+Theo. error)}$
[TEW Group '12; Bechtle, Heinemeyer, Stål, Stefaniak, Weiglein Zeune '12]
 - ❑ $\text{BR}(B \rightarrow X_s \gamma) = (3.43 \pm 0.56) \times 10^{-4} @ 95\% \text{ C.L.}$
[Heavy Flavor Averaging Group '12]
 - ❑ $\text{BR}(B_s^0 \rightarrow \mu^+ \mu^-) = (3.2_{-2.1}^{+3.2}) \times 10^{-9} @ 95\% \text{ C.L.}$
[LHCb Collaboration '12]
- ❑ Constraints from vacuum (meta-)stability (CCB)
 - ❑ Constraints on $|\mu \tan \beta|$ [Kitahara, Yoshinaga '13]
 - ❑ Constraints on A_τ, A_b, A_t [Casas, Lleyda, Muñoz '96]



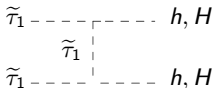
Implications for low stau yields, $Y < 10^{-14}$

Scan exceptional regions [Griest, Seckel '91, Ratz, Schmidt-Hoberg, Winkler '08; Pradler, Steffen '08]

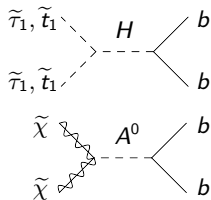
- Co-annihilation with colored sparticles, e.g. $m_{\tilde{g}} \simeq m_{\tilde{\tau}_1}$:



- Enhanced stau-Higgs couplings, large mixing, large $\tan \beta$



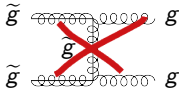
- H/A -resonance, $m_H, m_A \simeq 2m_{\tilde{\tau}_1}$



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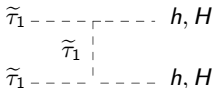
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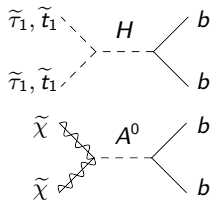


Constrained by HSCP bounds

- Enhanced stau-Higgs couplings, large mixing, large $\tan \beta$



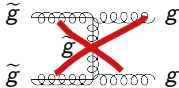
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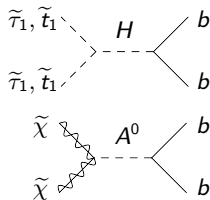
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Constrained by FP bounds
MSSM Higgs searches
CCB bounds

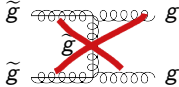
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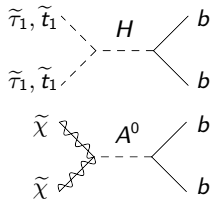
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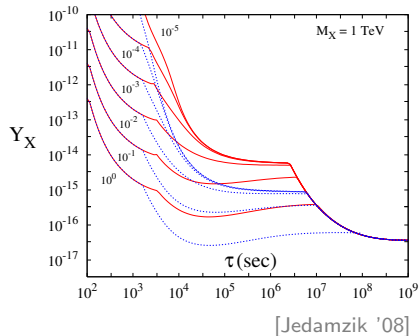
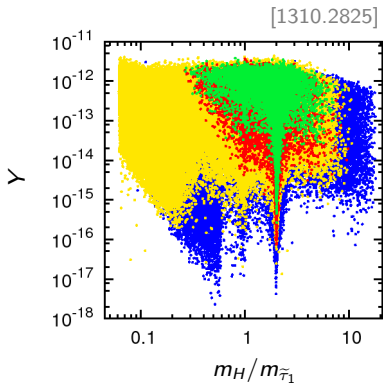
Constrained by FP bounds
MSSM Higgs searches
CCB bounds

- H/A -resonance, $m_H, m_A \simeq 2m_{\tilde{\tau}_1}$



Resonant stau annihilation or
resonant stop/EWino co-annihilation
provides $Y < 10^{-14}$

Implications for the stau yield



■ no constraints
■ passed HSCP search

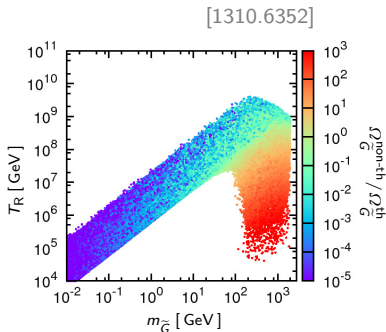
■ (additionally) passed FP+HB
■ (additionally) passed CCB bounds

Implications for T_R from Ω_{DM}

- Extend param. space by $m_{\tilde{G}}$
- Gravitino DM:
 $\Omega_{\tilde{G}} h^2 \stackrel{!}{=} \Omega_{DM} h^2 = 0.119$ [Planck '13]
 $\Omega_{\tilde{G}} = \Omega_{\tilde{G}}^{th} + \Omega_{\tilde{G}}^{non-th}$
- Non-thermal production:
 $\Omega_{\tilde{G}}^{non-th} \sim Y_{\tilde{\tau}_1} m_{\tilde{G}}$
- Thermal production:

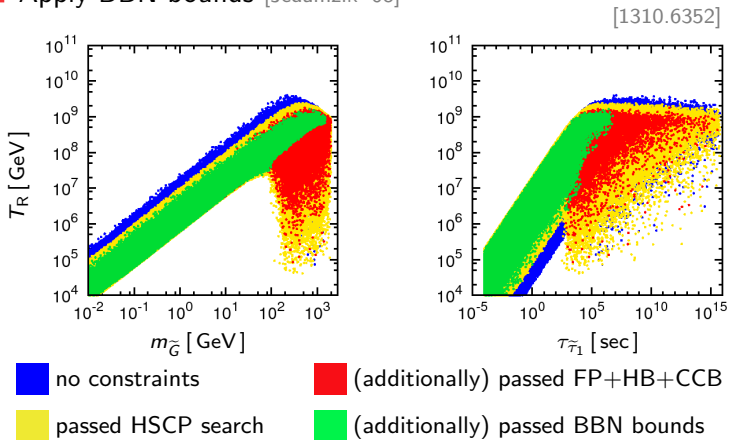
$$\Omega_{\tilde{G}}^{th} \sim \sum_{i=1}^3 C_i \left(1 + \frac{M_i^2}{3m_{\tilde{G}}^2} \right) m_{\tilde{G}} T_R$$

[Bolz, Buchmüller, Plümacher '98/'00; Pradler, Steffen '07]



Implications for T_R from BBN

- Compute $\tau_{\tilde{\tau}_1}$, hadronic BRs $\Gamma(\tilde{\tau}_1 \rightarrow \tilde{G}\tau q\bar{q})$, $\Gamma(\tilde{\tau}_1 \rightarrow \tilde{G}\nu_\tau q\bar{q}')$
- Apply BBN bounds [Jedamzik '08]



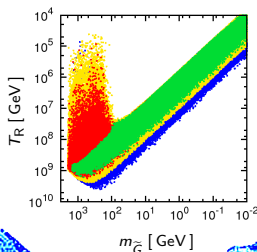
Implications for T_R

Points with $T_R > 10^9$ GeV point to very particular corner:

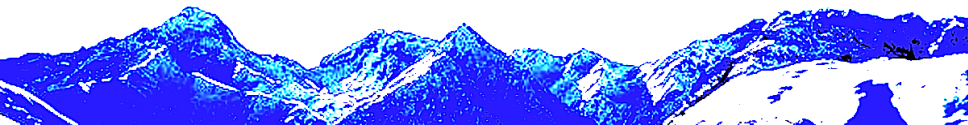
- ❑ Large gravitino masses: $300 \text{ GeV} < m_{\tilde{G}} < (m_{\tilde{\tau}_1} - 200 \text{ GeV})$
- ❑ Life-times: $10^4 \text{ sec} < \tau_{\tilde{\tau}_1} < 10^7 \text{ sec}$
- ❑ Exceptionally small yields $Y \lesssim 10^{-14}$
- ❑ All in H/A resonant region (w EWino/stop or w/o co-ann.)
- ❑ Gaugino masses just above exclusion limits, in particular M_2 (running)
- ❑ Small stau-gaugino splittings, $m_{\tilde{\tau}_1} > 800 \text{ GeV}$
- ❑ 1st and 2nd generation sfermions arbitrary heavy

Conclusions

- ❑ SUSY+Leptogenesis → Gravitino Problem
- ❑ Gravitino-stau scenario: high LHC sensitivity
- ❑ Performed bottom-up approach (no high scale model)
- ❑ Points with $T_R > 10^9$ GeV are allowed
- ❑ Point to very particular corner in parameter space with interesting LHC pheno – testable @ LHC 13/14
- ❑ Thermal Leptogenesis:



Thank you for your attention!



Backup – Prospects for LHC 14

Important channels for points with $T_R > 10^9$ GeV:

- EWino & gluino production
- Resonant production of staus

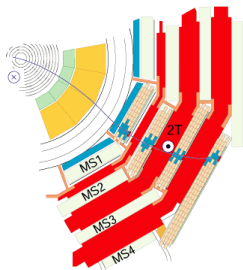
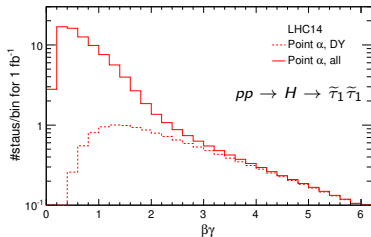
→ Extremely slowly moving HSCPs

- Challenging for trigger

[Kersten, JH; '12]

- Many stopped staus

[Lindert, Steffen, Trenkel; '11]



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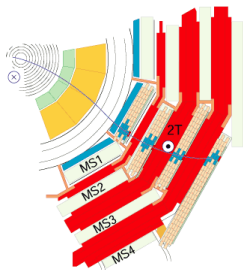
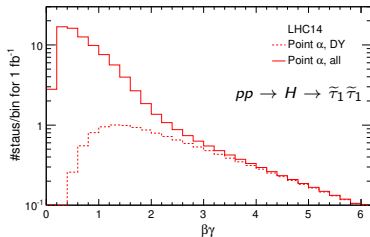
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Backup – Prospects for LHC 14: Stopped Staus

Determination of $m_{\tilde{G}}$ from
 $\tilde{\tau}_1 \rightarrow \tilde{G}\tau$ in detector

$$E_{\tau} \simeq \frac{m_{\tilde{\tau}_1}}{2} \left(1 - \frac{m_{\tilde{G}}^2}{m_{\tilde{\tau}_1}^2} \right)$$

$m_{\tilde{\tau}_1}, m_{\tilde{G}}, \tau_{\tilde{\tau}_1} \rightarrow$ Probe Supergravity

[Feng, Rajaraman, Takayama '04; W. Buchmüller, Hamaguchi, M. Ratz, and T. Yanagida '04]

