

MORIOND ELECTROWEAK

March 18th 2014

Light Neutralino Dark Matter > 24 GeV from LHC

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ULB



based on collaborations with J. Lindert, T. Ota, Y. Takanishi

Introduction

Searches for EW-interacting SUSY particle at the LHC started to go beyond LEP limits

The EW SUSY sector might be lighter than the strong sector but LHC can test it!

Here an example of the potential of EW searches: probing light Neutralino Dark Matter

Assumptions:

- Only MSSM superfields
- R-parity
- Dark Matter thermal relic, standard history of the universe
- Neutralino DM candidate, $\Omega_{\text{DM}} h^2 \leq 0.124$

How light the neutralino is allowed to be after LHC searches at 8 TeV?

Light Neutralino Dark Matter in the MSSM

MSSM neutralinos:

$$\left(\tilde{B}, \tilde{W}_3, \tilde{H}_d^0, \tilde{H}_u^0 \right)$$

MSSM charginos:

$$(\tilde{W}^\pm, \tilde{H}_u^+, \tilde{H}_d^+)$$

$$M_{\tilde{N}} = \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 \end{pmatrix} \quad \mathcal{M}_\pm = \begin{pmatrix} M_2 & \sqrt{2}M_W \sin\beta \\ \sqrt{2}M_W \cos\beta & \mu \end{pmatrix}$$

Mass eigenstates:

$$\tilde{\chi}_i^0 = N_{i1}\tilde{B} + N_{i2}\tilde{W} + N_{i3}\tilde{H}_d + N_{i4}\tilde{H}_u$$

LEP chargino searches: $M_{\tilde{\chi}_1^+} > 94 \text{ GeV}$



$$M_2, \mu \gtrsim 90 \text{ GeV}$$

$$m_{\tilde{\chi}_1^0} \lesssim 30 \text{ GeV}$$



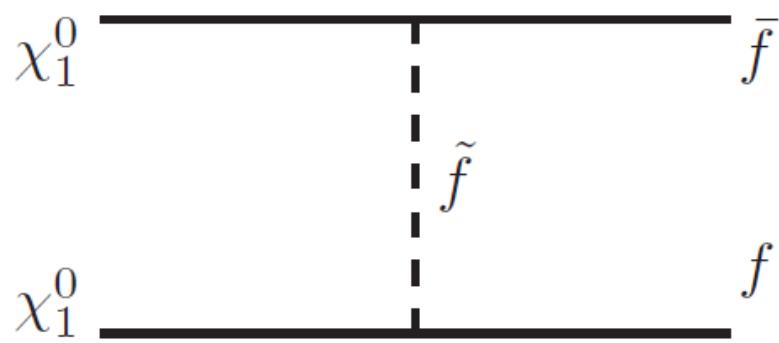
$$M_1 \ll M_2, \mu \iff \tilde{\chi}_1^0 \approx \tilde{B}$$

Relic density (WMAP, Planck): $\Omega_\chi h^2 \sim 0.1$

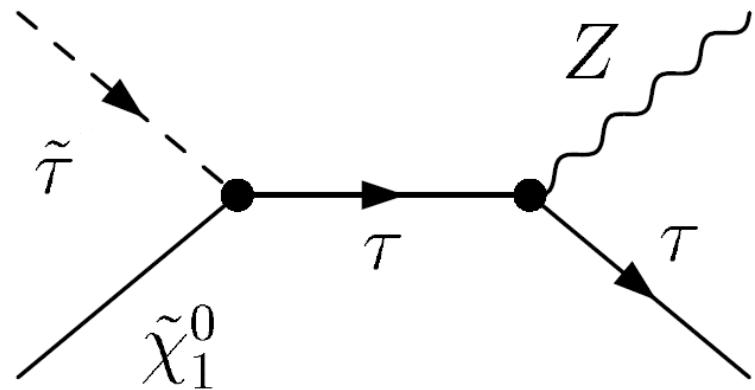
Efficient annihilation required

Annihilation mechanisms

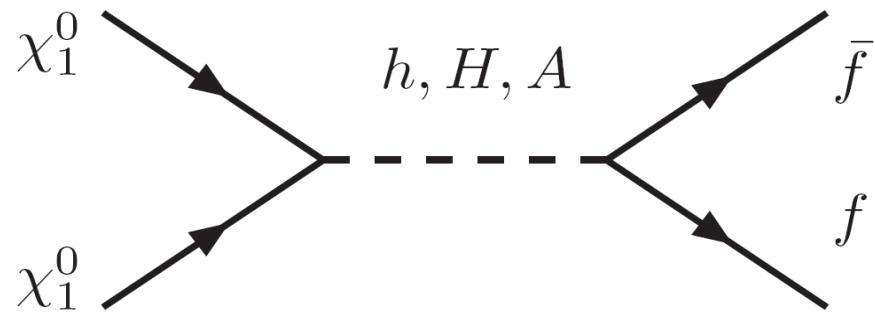
t-channel exchange (e.g. sfermions):



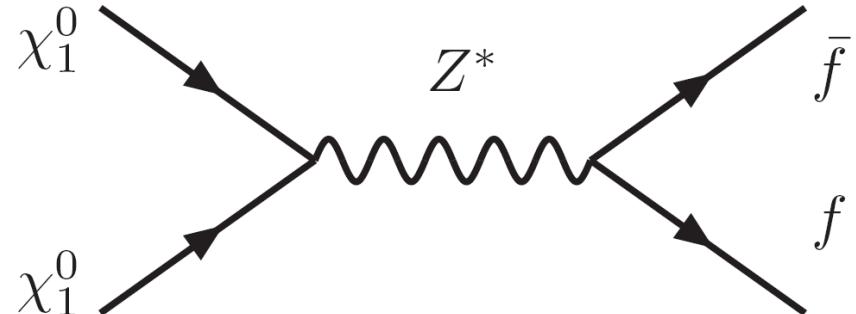
Coannihilations (e.g. stau):



Scalar mediation:



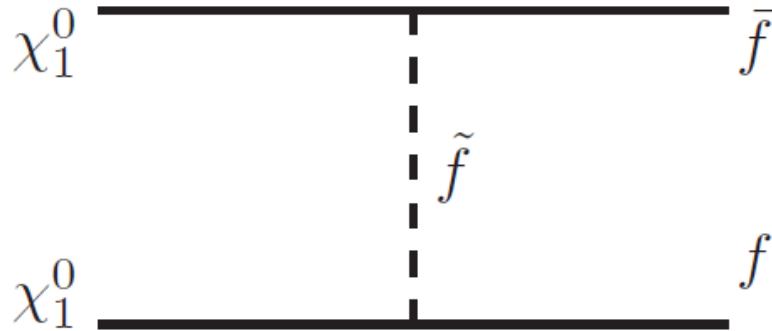
Z mediation:



different regions of the SUSY parameter space are selected

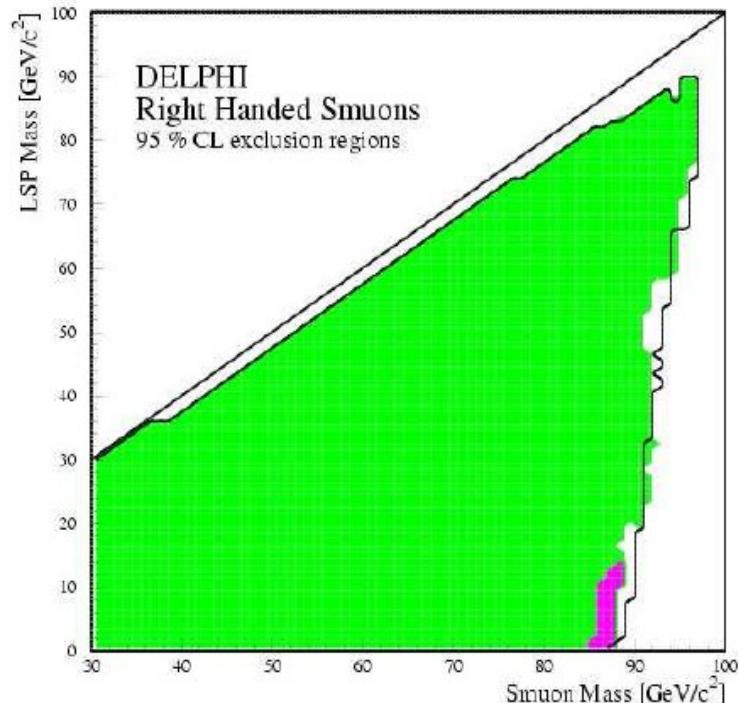
Annihilation mechanisms: sfermion exchange

Sfermion mediated annihilation:



$$\propto \frac{m_{\tilde{\chi}_1^0} m_f}{m_{\tilde{f}}^2}$$

Challenged by LEP bounds, e.g.:



LEP limits: $e^+e^- \rightarrow \tilde{f}\tilde{f}^* \rightarrow f\bar{f} \tilde{\chi}_1^0 \tilde{\chi}_1^0$

Can one evade it with compressed spectra?

If $m_{\tilde{\chi}_1^0} \lesssim 30 \text{ GeV}$ this is not possible

because Z width bounds:

$$Z \rightarrow \tilde{f}\tilde{f}^* \xrightarrow{\text{blue arrow}} m_{\tilde{f}} \gtrsim 40 \text{ GeV}$$

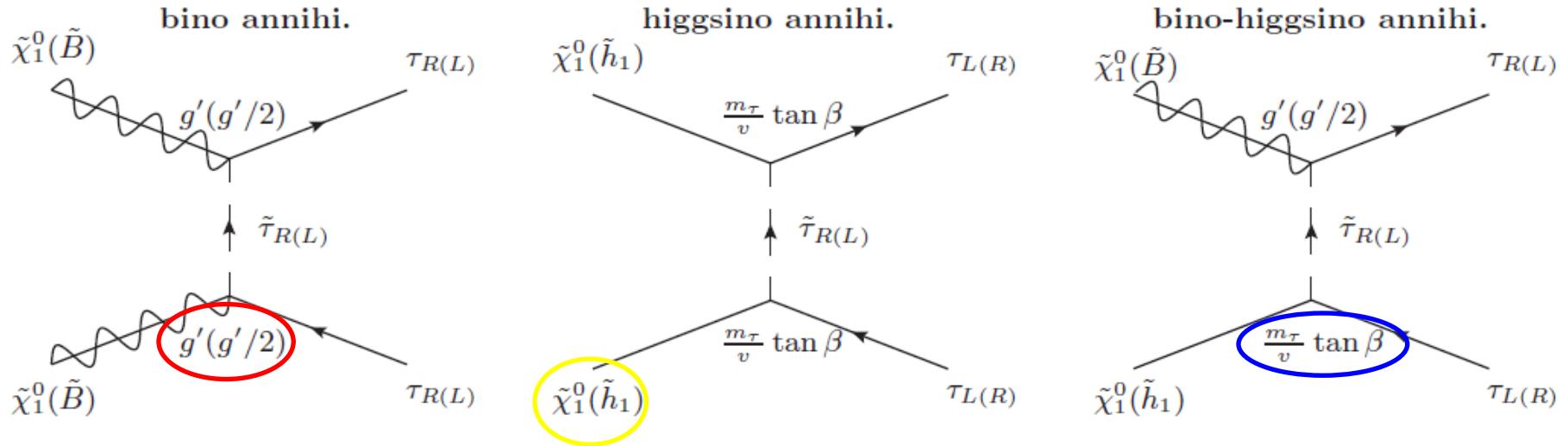
(Possible exception: light sbottom with tuned LR mixing) **Arbey Battaglia Mahmoudi '13**

Stau-mediated annihilation

Light neutralino DM possible in presence of a light stau

Albornoz Bélanger Boehm '11

Grothaus Lindner Takanishi '12



- RH stau much more efficient (cross-section 16x larger than LH one)
- Sizeable higgsino component: small μ
- Yukawa interactions: large $\tan\beta$

Relic density essentially controlled by 4 parameters only:

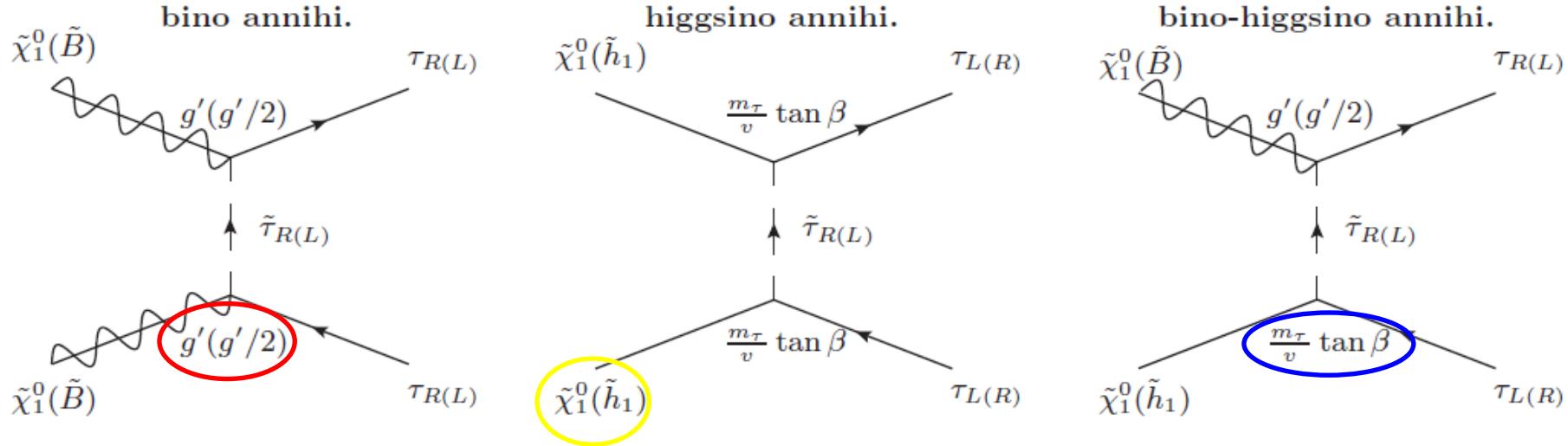
$$M_1, m_{\tilde{\tau}_R}, \mu, \tan \beta$$

Stau-mediated annihilation

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Light neutralino DM necessarily implies:

light stau, light higgsino-like neutralinos and charginos

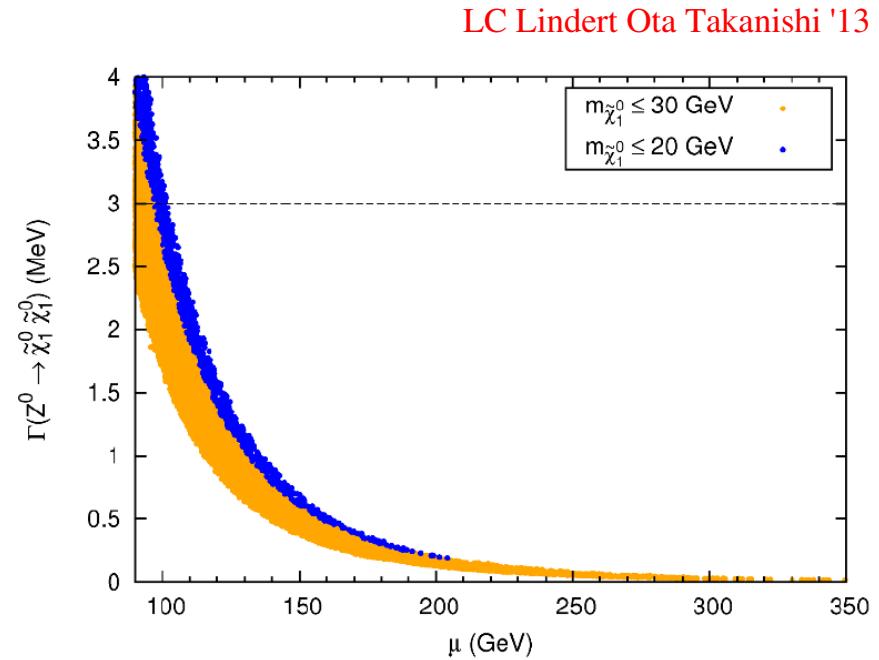
Parameters scan and constraints

$$\begin{aligned}
 10 \text{ GeV} &\leq M_1 \leq 45 \text{ GeV}, & 65 \text{ GeV} &\leq m_{\tilde{\tau}_R} \leq 200 \text{ GeV}, \\
 90 \text{ GeV} &\leq \mu \leq 400 \text{ GeV}, & 5 &\leq \tan \beta \leq 60. \\
 m_{\tilde{f}} = M_3 = m_A &= 2 \text{ TeV}, & M_2 &= 500 \text{ GeV}, & A_t &= 1.5 \times m_{\tilde{f}}.
 \end{aligned}$$

SuSpect,
micrOMEGAs

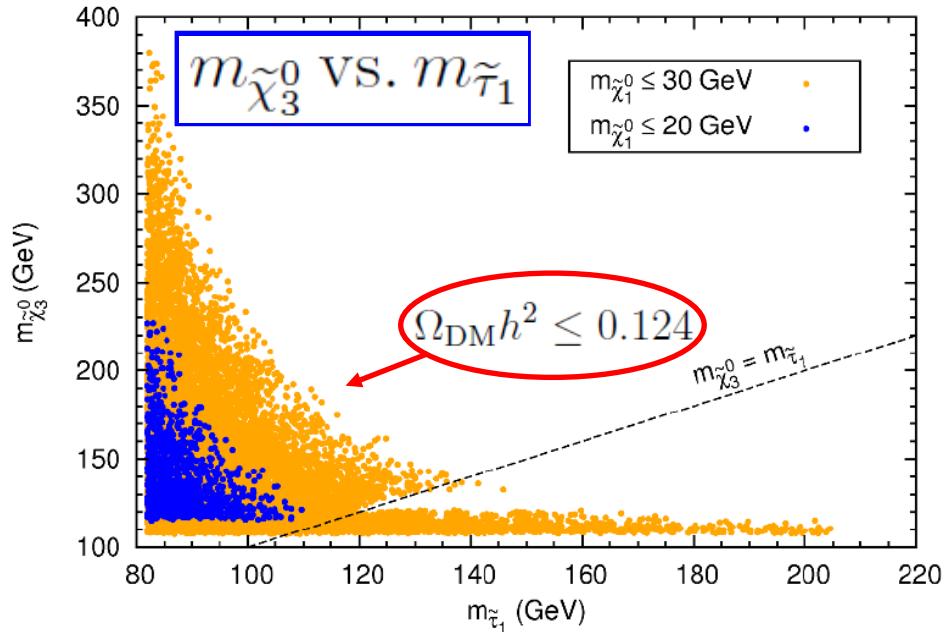
Djouadi et al. '02
Belanger et al. '06

- $m_{\tilde{\chi}_1^0} \leq 30 \text{ GeV}$
- CMB, Planck (3\(\sigma\)):** $\Omega_{\text{DM}} h^2 \leq 0.124$
- LEP2:** $m_{\tilde{\tau}_R} \geq 81.9 \text{ GeV}$, $m_{\tilde{\chi}_1^\pm} \geq 103.5 \text{ GeV}$
- LHC:** limits on charginos depend on smuon/selectron masses and can be evaded
- Flavour:** Ω_{DM} does not depend on heavy Higgs/squark masses, flavour observables do not constrain the DM parameter space
- Z invisible width, LEP:** $\Delta\Gamma_Z^{\text{inv}} < 3 \text{ MeV}$



$$\Gamma(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = \frac{G_F}{\sqrt{2}} \frac{M_Z^3}{12\pi} \left(1 - \frac{4M_{\tilde{\chi}_1^0}^2}{M_Z^2}\right)^{\frac{3}{2}} \left|N_{13}^2 - N_{14}^2\right|^2$$

Parameter space



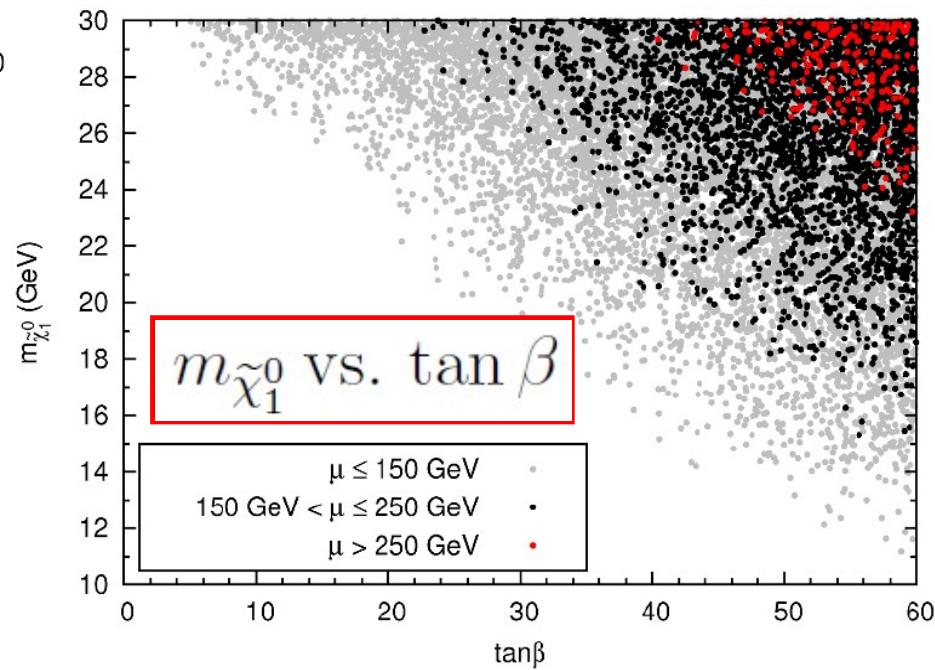
$$m_{\tilde{\tau}_1} \lesssim 210 \text{ GeV}$$

$$m_{\tilde{\chi}_1^\pm} \approx m_{\tilde{\chi}_2^0} \approx m_{\tilde{\chi}_3^0} \lesssim 380 \text{ GeV}$$

LC Lindert Ota Takanishi '13

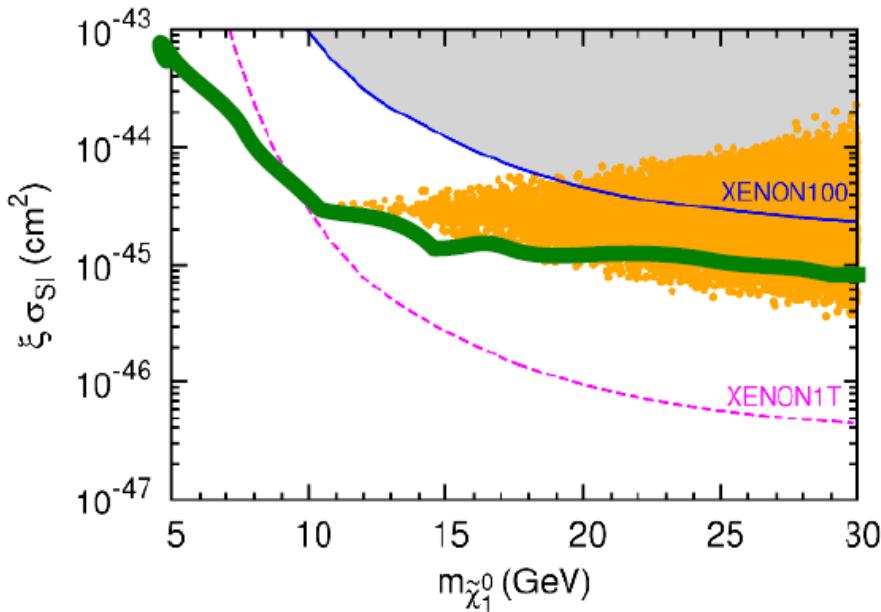
$m_{\tilde{\chi}_1^0} \gtrsim 11 \text{ GeV}$

$(\tan \beta \leq 60)$

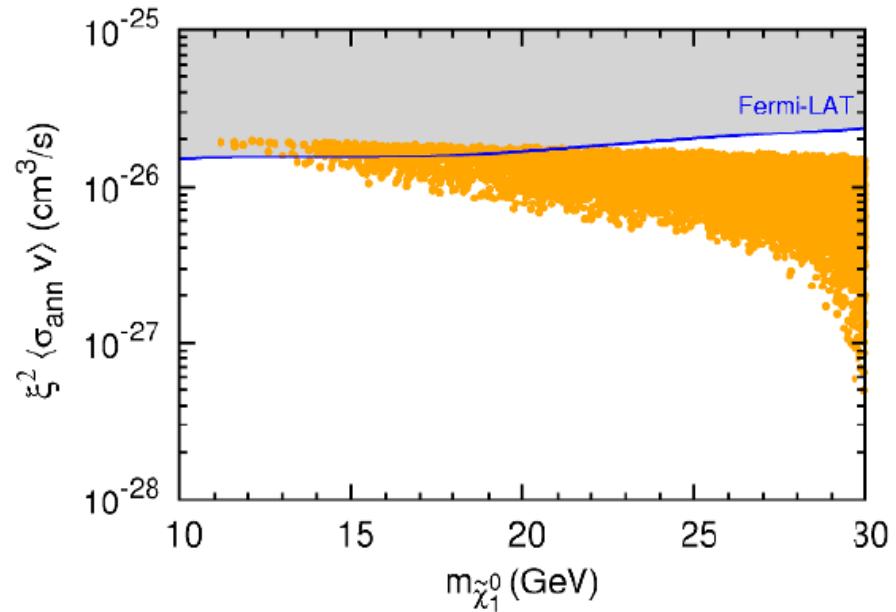


Direct and indirect Dark Matter detection

Spin-independent scattering cross section



Annihilation cross-section



Th. prediction can be lowered by 2x by
uncert. in light quark masses/form factors

Heavy Higgs slow decoupling (possible
additional $\sim 3x$ suppression)

Gamma rays from satellite galaxies

LHC phenomenology: production and decays

Relic density constr. imply that we have *at least* 4 states at O(100) GeV:

$$\tilde{\tau}_1, \tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_1^\pm$$

The rest of the spectrum *can* be decoupled. Still sizeable EW production:

$$pp \rightarrow \tilde{\tau}_1^+ \tilde{\tau}_1^- + X, \quad pp \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_j^0 + X, \quad pp \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_1^\pm + X, \quad pp \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^- + X.$$

Drell-Yan, up to O(1) pb at LHC8

Decays:

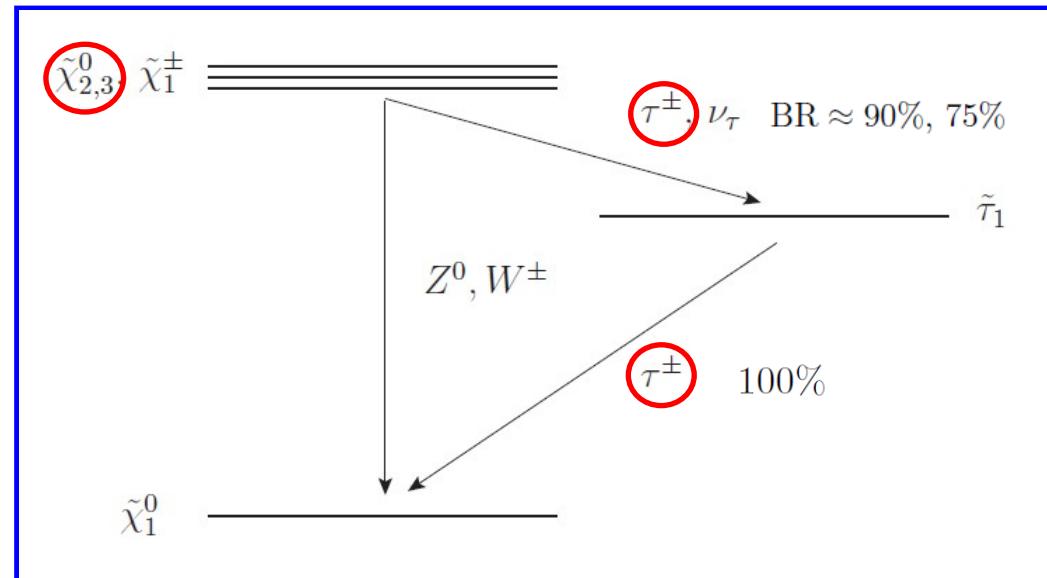
$$m_{\tilde{\chi}_1^\pm} \simeq m_{\tilde{\chi}_{2,3}^0} > m_{\tilde{\tau}_1} > m_{\tilde{\chi}_1^0}$$



$$pp \rightarrow \tilde{\chi}_{2,3}^0 \tilde{\chi}_{2,3}^0 \rightarrow 4\tau + \cancel{E}_T$$

$$pp \rightarrow \tilde{\chi}_{2,3}^0 \tilde{\chi}_1^\pm \rightarrow 3\tau + \cancel{E}_T$$

multi-tau signals!



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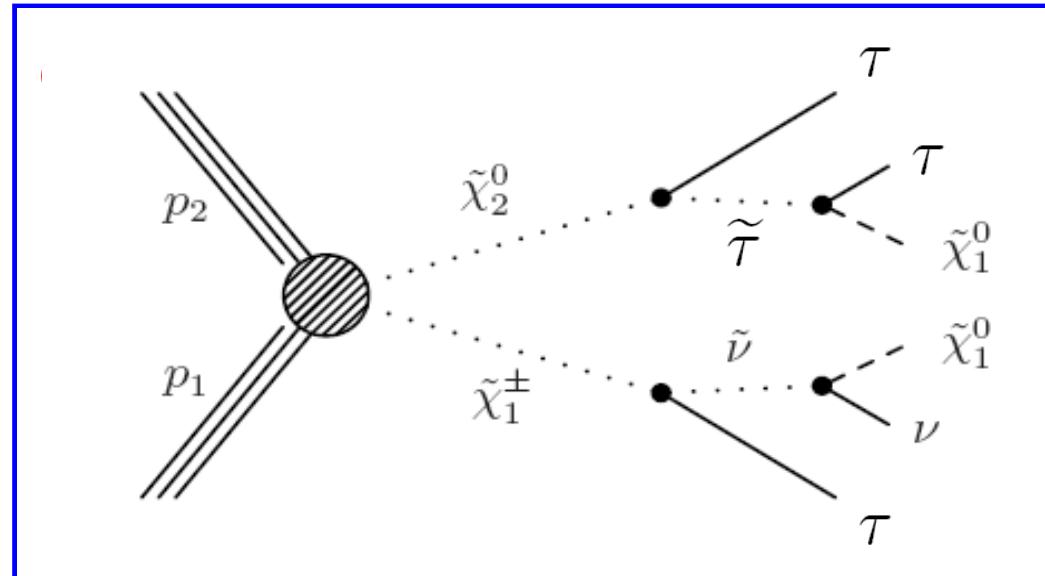
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multi-tau signals!



LHC phenomenology: production and decays

Relic density constr. imply that we have *at least* 4 states at $O(100)$ GeV:

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The rest of the spectrum *can* be decoupled. Still sizeable EW production:

$$pp \rightarrow \tilde{\tau}_1^+ \tilde{\tau}_1^- + X, \quad pp \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_j^0 + X, \quad pp \rightarrow \tilde{\chi}_i^0 \tilde{\chi}_1^\pm + X, \quad pp \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^- + X.$$

Drell-Yan, up to $O(1)$ pb at LHC8

Decays:

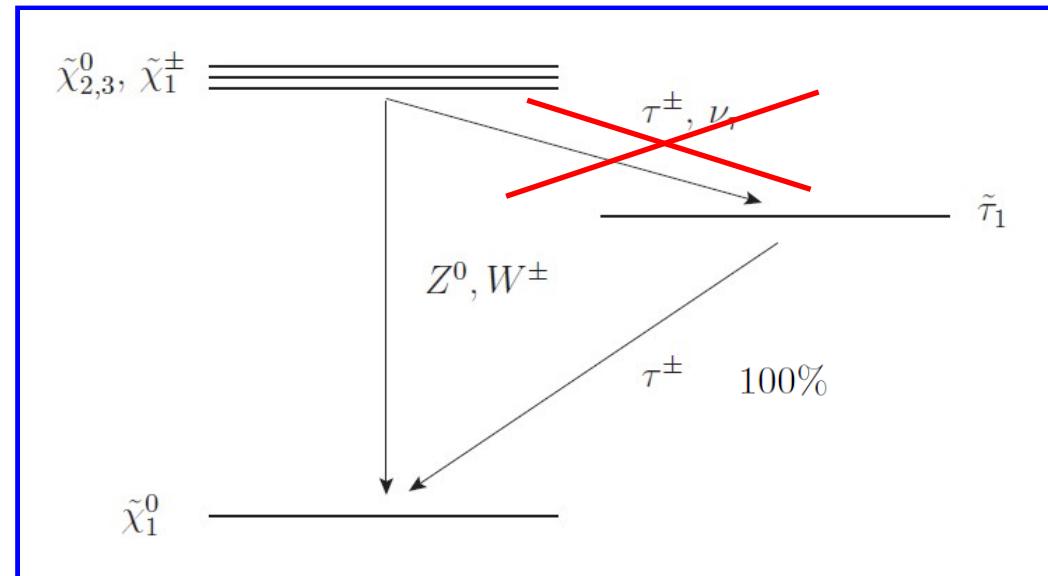
$$m_{\tilde{\tau}_1} > m_{\tilde{\chi}_{2,3}^0} > m_{\tilde{\chi}_1^0}$$



3-body decays

it might be more difficult

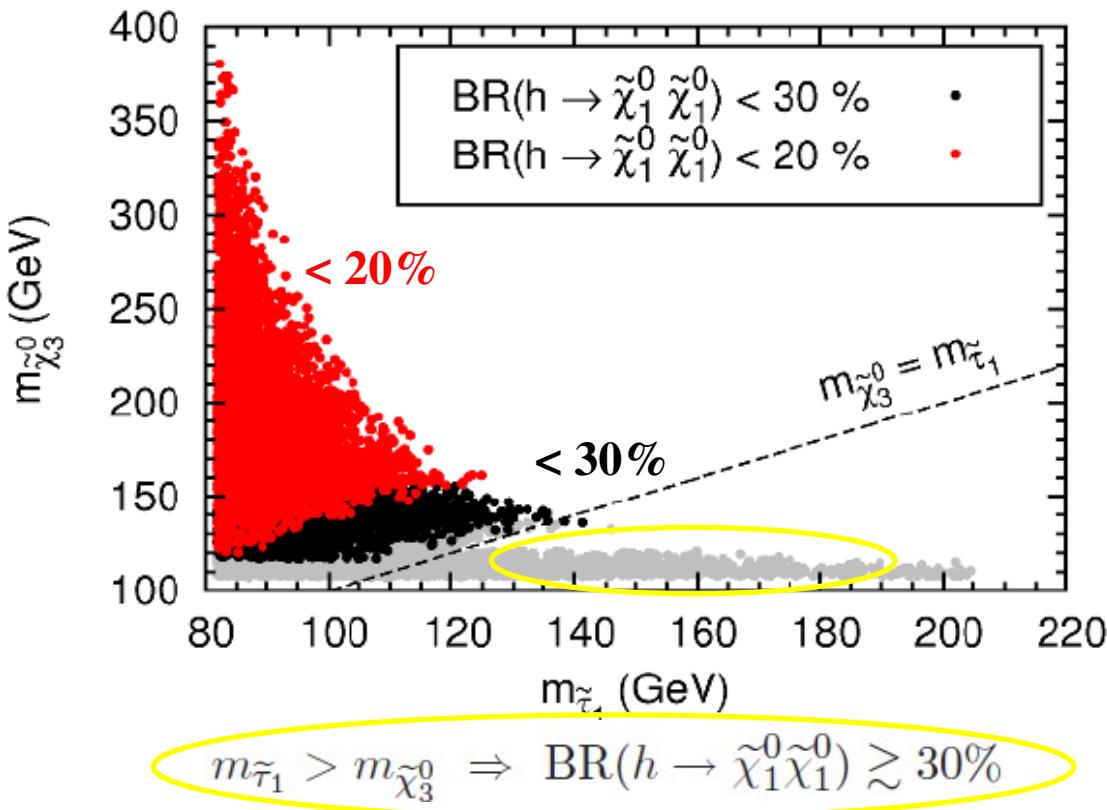
However...



LHC phenomenology: invisible scalar decay

$$\Gamma(h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = \frac{G_F M_W^2 m_h}{2\sqrt{2}\pi} \left(1 - 4m_{\tilde{\chi}_1^0}^2/m_h^2\right)^{3/2} |C_{h\tilde{\chi}_1^0\tilde{\chi}_1^0}|^2$$

$$C_{h\tilde{\chi}_1^0\tilde{\chi}_1^0} = (N_{12} - \tan \theta_W N_{11}) (\sin \beta N_{14} - \cos \beta N_{13})$$



Last fits to Higgs data:

$\text{BR}_h^{\text{inv}} < 16\%$ (95% CL)

Falkowski et al. '13

$\text{BR}_h^{\text{inv}} < 19\%$ (95% CL)

Giardino et al. '13

but with 20% theo unc.:

$\text{BR}_h^{\text{inv}} < 52\%$ (68% CL)

Djouadi Moreau '13

$m_{\tilde{\tau}_1} > m_{\tilde{\chi}_{2,3}^0} > m_{\tilde{\chi}_1^0}$

strongly disfavoured!



ATLAS NOTE

ATLAS-CONF-2013-028

March 10, 2013

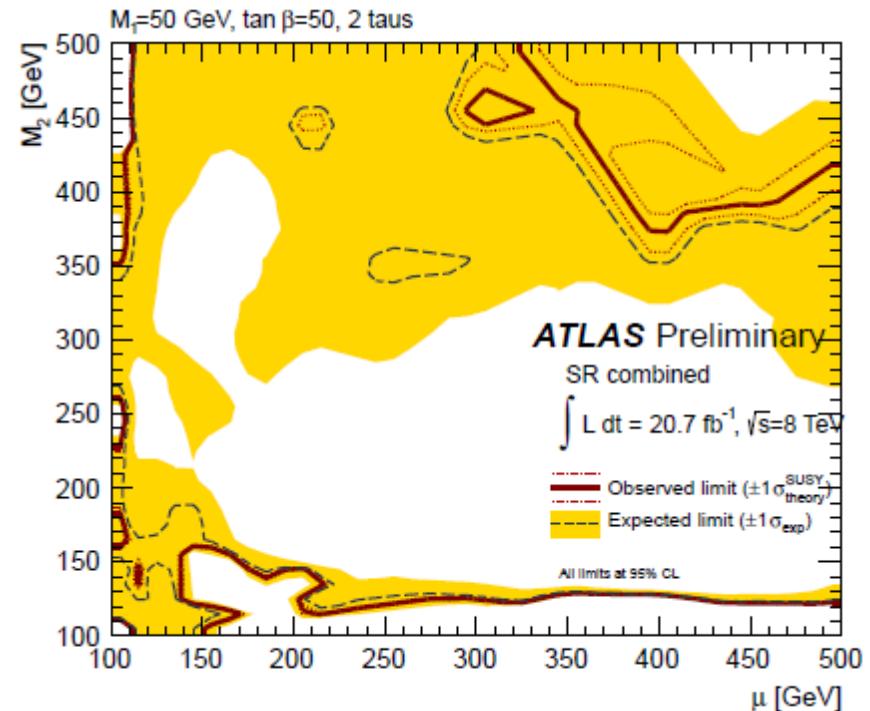


Search for electroweak production of supersymmetric particles in final states with at least two hadronically decaying taus and missing transverse momentum with the ATLAS detector in proton-proton collisions at $\sqrt{s} = 8 \text{ TeV}$

The ATLAS Collaboration

Signal region	requirements
OS m_{T2}	at least 1 OS tau pair jet veto Z-veto $E_T^{\text{miss}} > 40 \text{ GeV}$ $m_{T2} > 90 \text{ GeV}$
OS m_{T2} -nobjet	at least 1 OS tau pair b-jet veto Z-veto $E_T^{\text{miss}} > 40 \text{ GeV}$ $m_{T2} > 100 \text{ GeV}$

Table 1: Definition of the signal regions.



SM process	SR OS m_{T2}	SR OS m_{T2} -nobjet
top	$0.2 \pm 0.5 \pm 0.1$	$1.6 \pm 0.8 \pm 1.2$
Z+jets	$0.28 \pm 0.26 \pm 0.23$	$0.4 \pm 0.3 \pm 0.3$
diboson	$2.2 \pm 0.5 \pm 0.5$	$2.5 \pm 0.5 \pm 0.9$
multi-jet & W+jets	$8.4 \pm 2.6 \pm 1.4$	$12 \pm 3 \pm 3$
SM total	$11.0 \pm 2.7 \pm 1.5$	$17 \pm 4 \pm 3$
data	6	14

$$m_{\tilde{\tau}_1} = 95 \text{ GeV}, \quad m_{\tilde{\chi}_1^0} = 50 \text{ GeV}$$



$$S_{\text{SR1}}^{95} < 5.6 \quad S_{\text{SR2}}^{95} < 10.4$$

Simulation

Herwig++ (event samples)

Prospino 2 (NLO K-factors)

Delphes 3 (fast detector simul.)

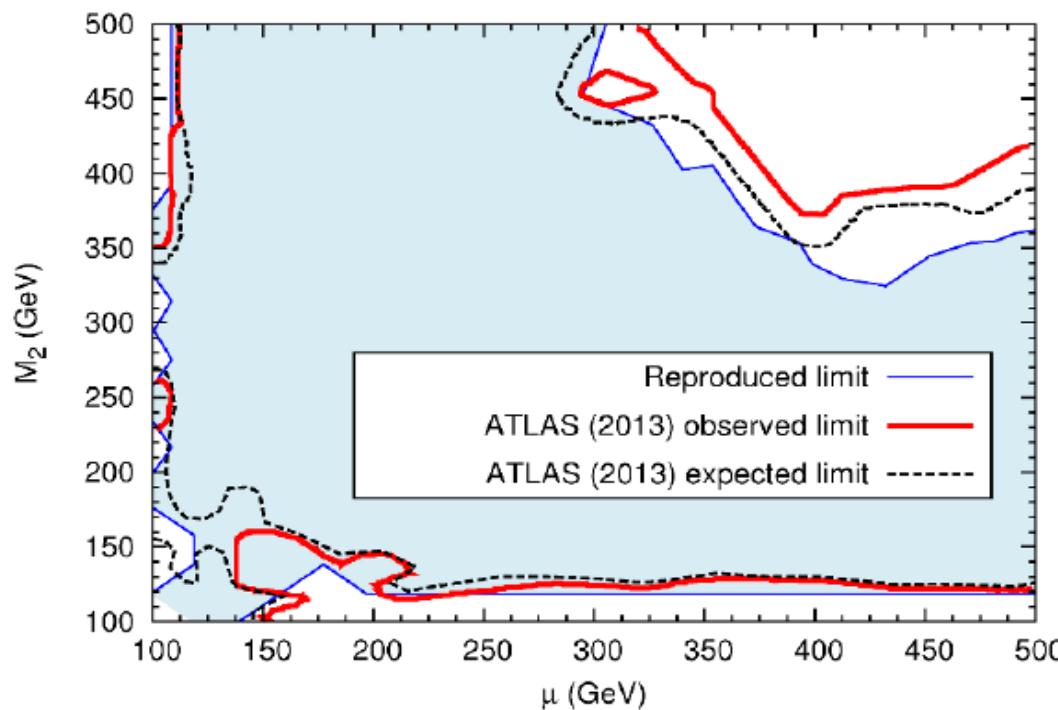
Cuts of two Atlas SR applied

95% CL limits on the number of events:

$$S_{\text{SR1}}^{95} < 5.6 \quad S_{\text{SR2}}^{95} < 10.4$$

Reproduced limit:

LC Lindert Ota Takanishi '13



Results

Herwig++ (event samples)

Prospino 2 (NLO K-factors)

Delphes 3 (fast detector simul.)

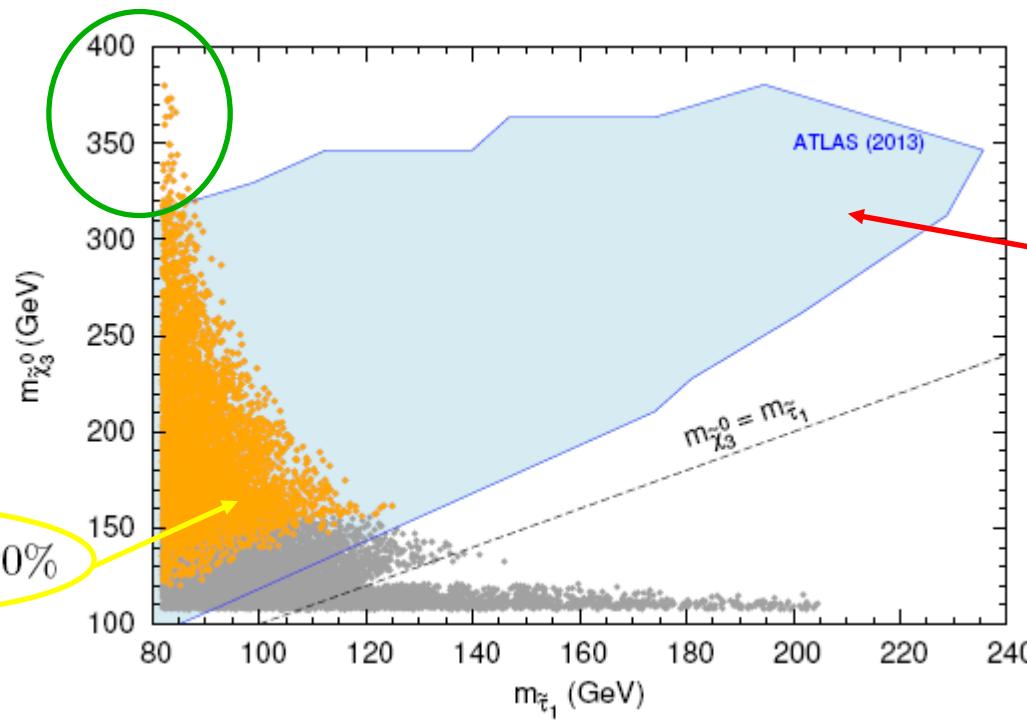
Cuts of two Atlas SR applied

95% CL limits on the number of events:

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Interpretation on our parameter space:

LC Lindert Ota Takanishi '13



Results

Herwig++ (event samples)

Prospino 2 (NLO K-factors)

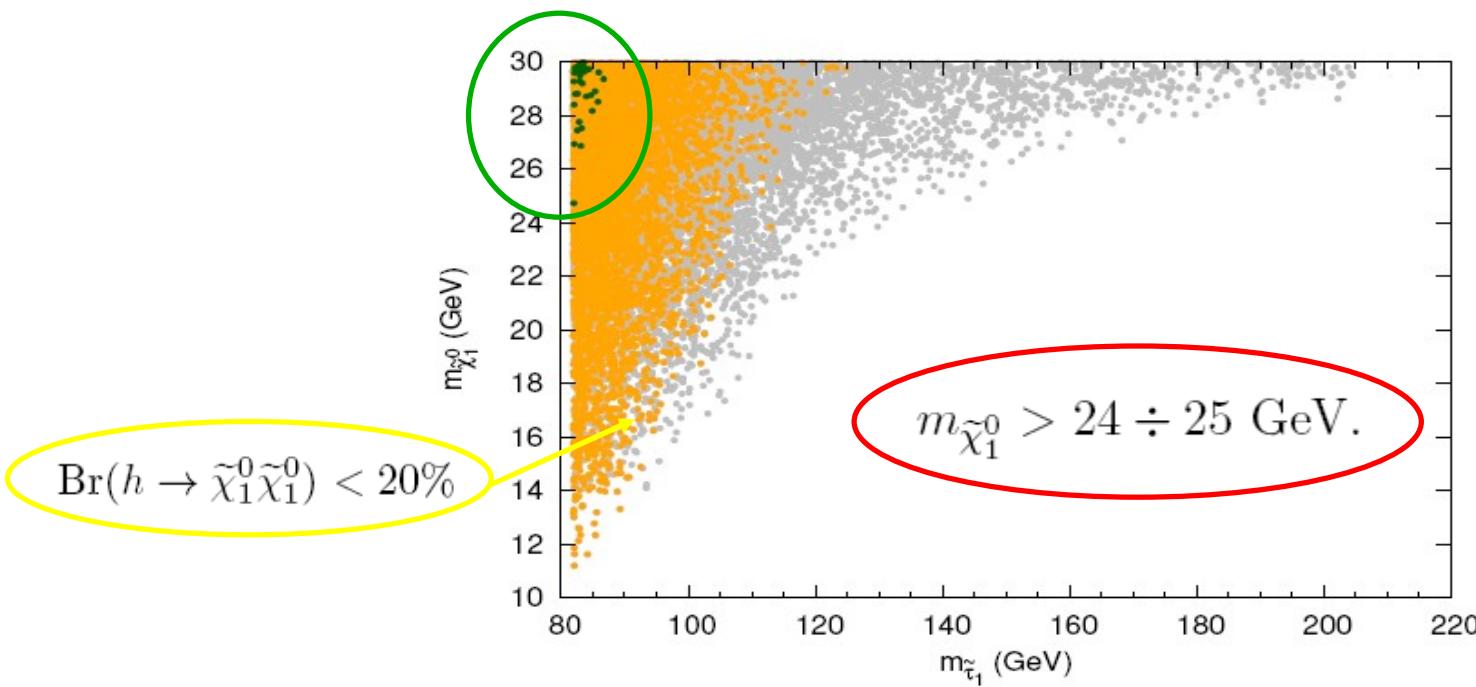
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Cuts of two Atlas SR applied

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Interpretation on the stau-neutralino mass plane:

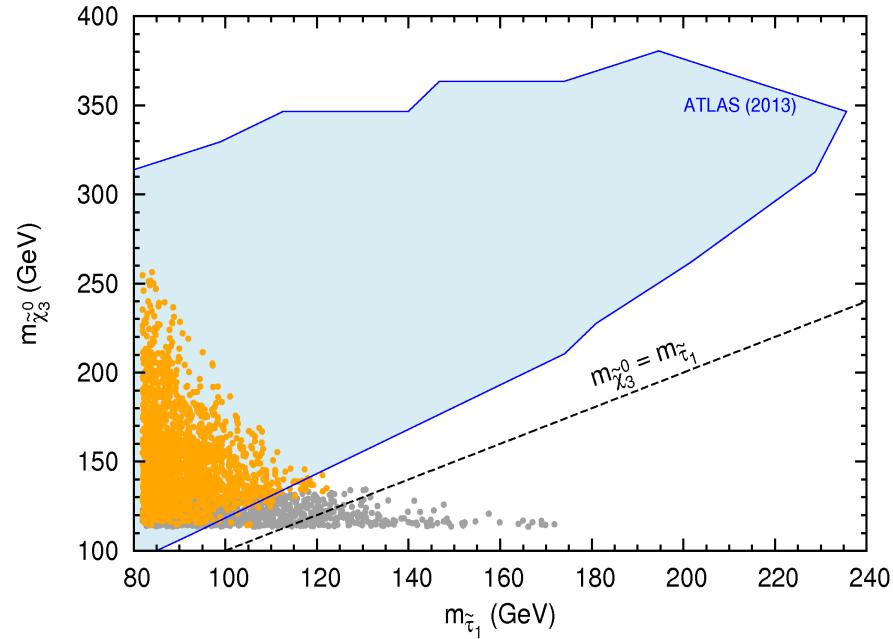
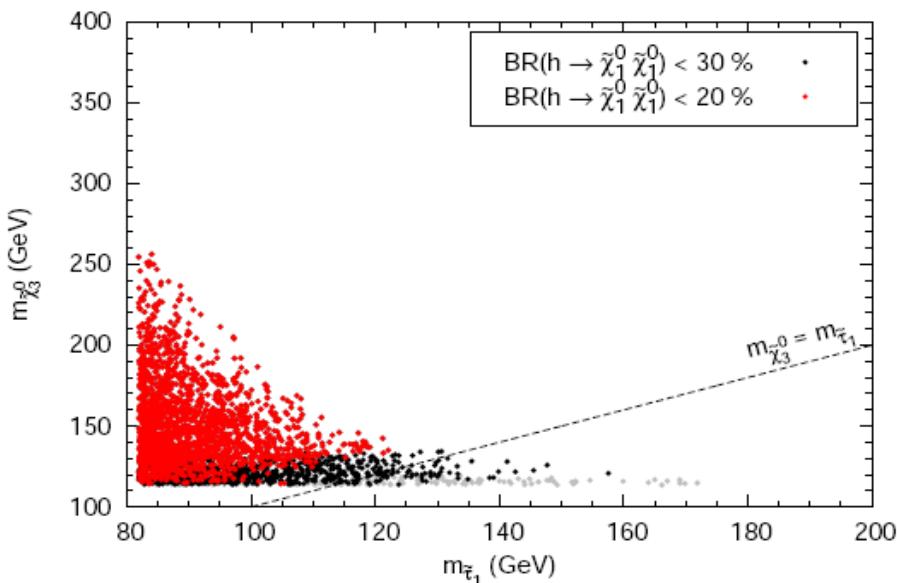


Negative μ

Flipping the sign of μ :

Destructive interference between two annihilation modes: $\tilde{B}\tilde{B}$ and $\tilde{B}\tilde{H}_d$

Partial cancellation in $\tilde{\chi}_1^0 \tilde{\chi}_1^0 h \implies$ lower $\text{BR}(h \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)$



Bound from the invisible decay milder, but DM heavier because annihilation less efficient

\Rightarrow limit on the neutralino mass similar to the case $\mu > 0$

Conclusions

Neutralino Dark Matter lighter than ~ 30 GeV requires light staus and higgsinos (relic density constraints)

Few parameters involved: manageable simplified model

Generic prediction: multi-tau + missing energy signal at the LHC

Reinterpretation of a ATLAS search sets strong constraint on light Neutralino Dark Matter (in combination with scalar \rightarrow invisible)

Limit stronger than direct/indirect DM searches

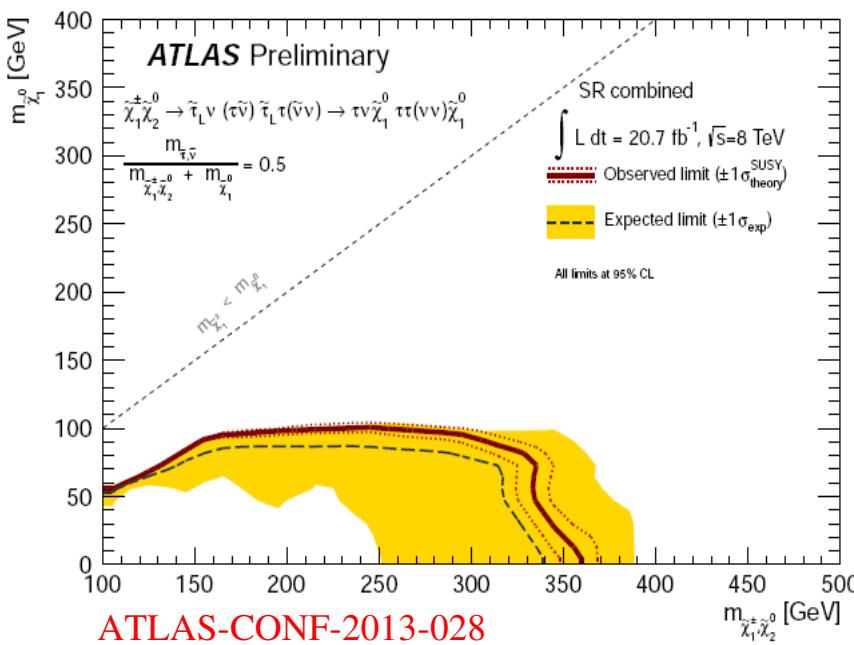
Nice interplay among different experimental info (collider and not)

Backup slides

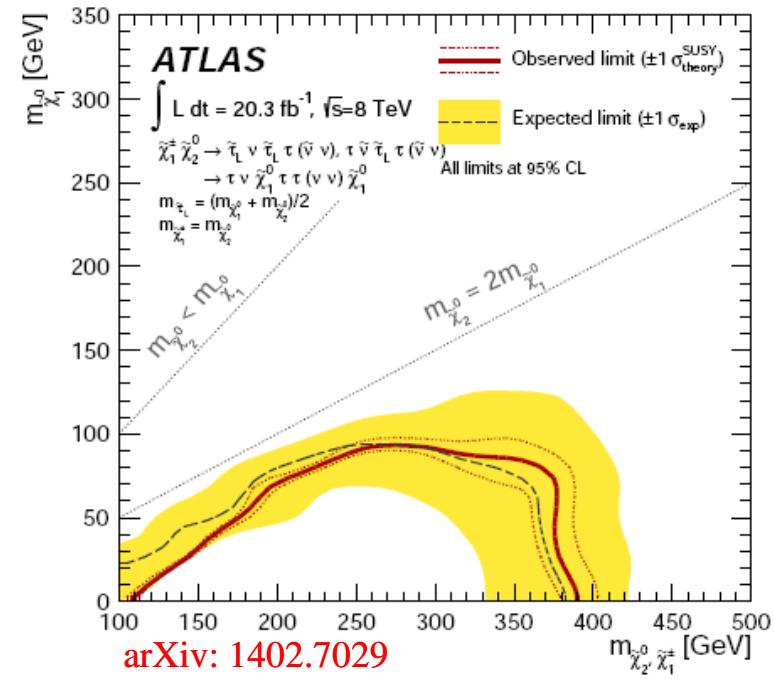
Other related searches: ATLAS arXiv:1402.7029

Table 5. Summary of the selection requirements for the validation regions. Energies, momenta and masses are given in units of GeV.

Region name	N(ℓ)	N(τ)	Flavour/sign	Z boson	E_T^{miss}	N(b -tagged jets)	Target process
VR0 τ ncZa	3	0	$e^+ e^- e, e^+ e^- \ell'$	$m_{\text{SFOS}} & m_{3\ell}$ veto	35–50	–	$WZ^*, Z^*Z^*, Z^*+\text{jets}$
VR0 τ Za	3	0	$e^+ e^- e, e^+ e^- \ell'$	request	35–50	–	$WZ, Z+\text{jets}$
VR0 τ ncZb	3	0	$e^+ e^- e, e^+ e^- \ell'$	$m_{\text{SFOS}} & m_{3\ell}$ veto	> 50	1	$t\bar{t}$
VR0 τ Zb	3	0	$e^+ e^- e, e^+ e^- \ell'$	request	> 50	1	WZ
VR0 τ b	3	0	$e^+ e^- e, e^+ e^- \ell'$	binned	binned	1	$WZ, t\bar{t}$
VR1 τ a	2	1	$\tau^\pm e^\mp \ell^\mp, \tau^\pm e^\mp \ell'^\mp$	–	35–50	–	$WZ, Z+\text{jets}$
VR1 τ b	2	1	$\tau^\pm e^\mp \ell^\mp, \tau^\pm e^\mp \ell'^\mp$	–	> 50	1	$t\bar{t}$
VR2 τ a	1	2	$\tau\tau\ell$	–	35–50	–	$W+\text{jets}, Z+\text{jets}$
VR2 τ b	1	2	$\tau\tau\ell$	–	> 50	1	$t\bar{t}$

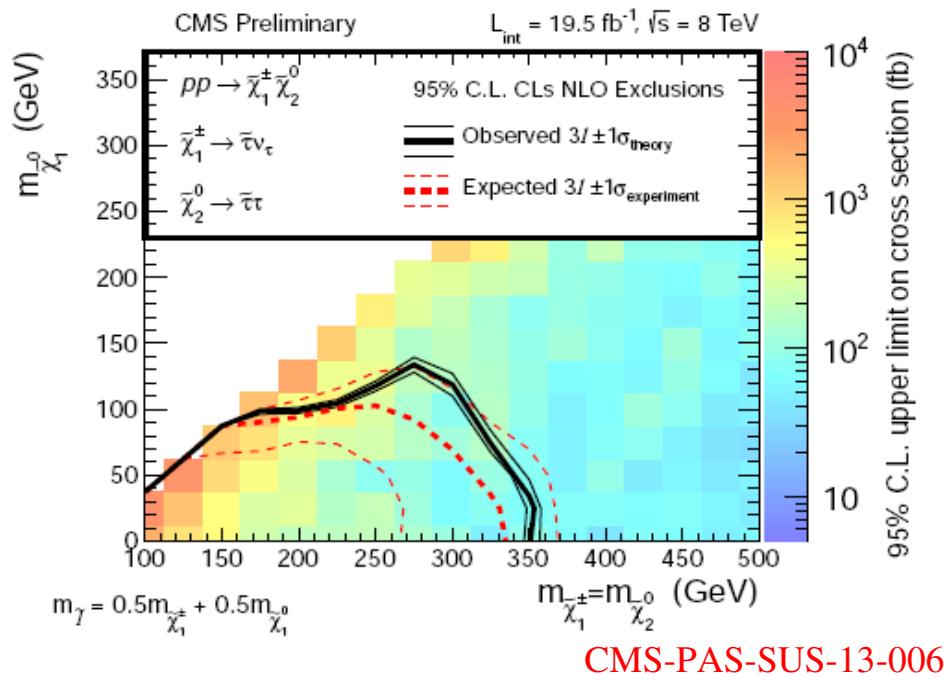
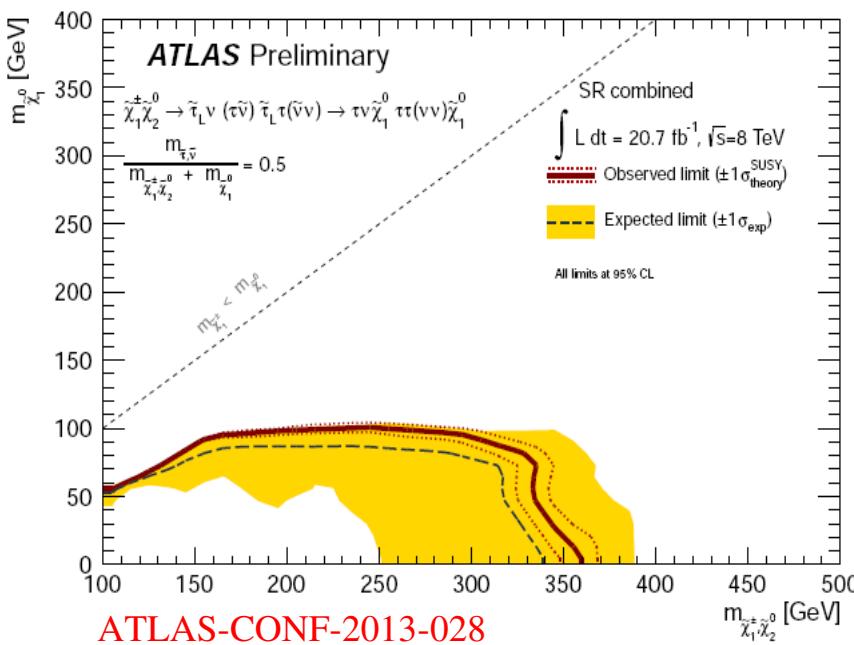


ATLAS-CONF-2013-028



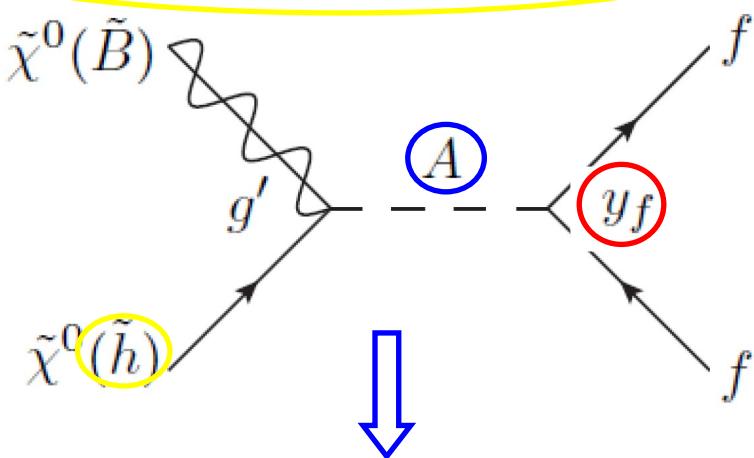
arXiv: 1402.7029

Search for electroweak production of charginos, neutralinos, and sleptons using leptonic final states in pp collisions at $\sqrt{s} = 8$ TeV



Annihilation mechanisms: CP-odd scalar exchange

- Higgs mediated annihilation:



$$\propto \frac{m_{\tilde{\chi}_1^0}^2}{m_A^2} \frac{m_f}{v} (N_{11} N_{13,14}) \tan \beta$$

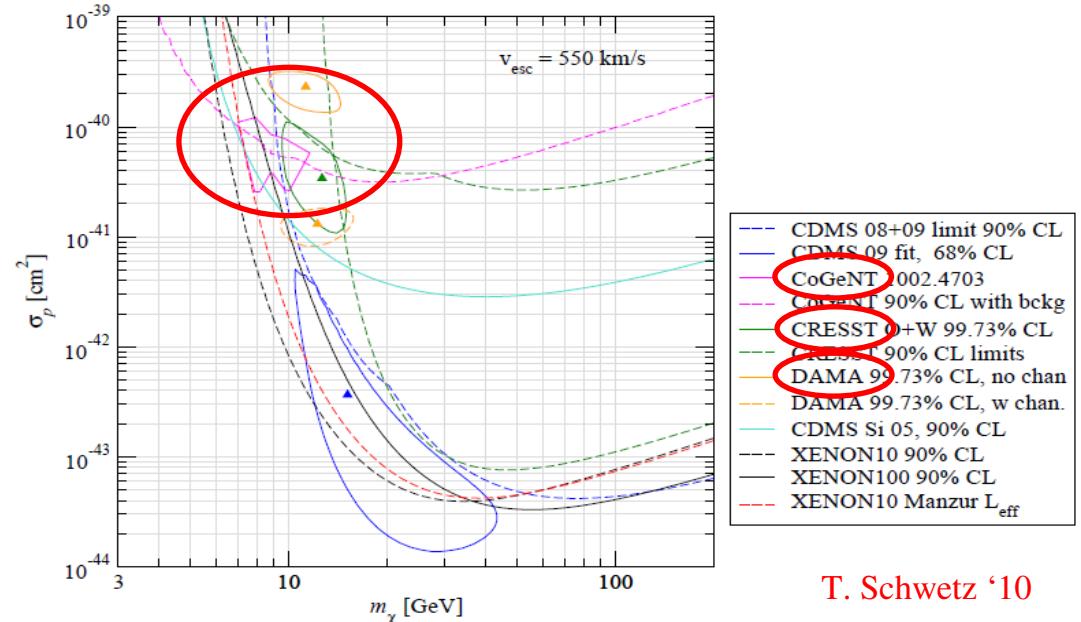
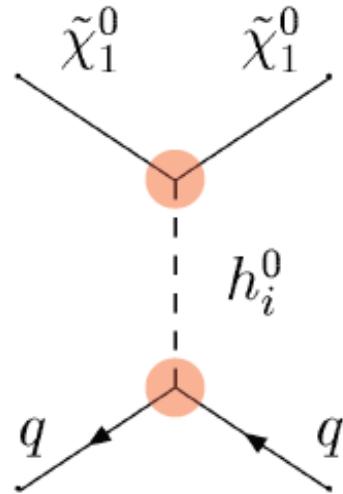
Drees Nojiri '92

Small $m_{\tilde{\chi}_1^0} \iff$ large $\tan \beta$, small m_A (and μ)

$m_A \approx 100 \div 120 \text{ GeV}$

Bottino et al. '02- '10

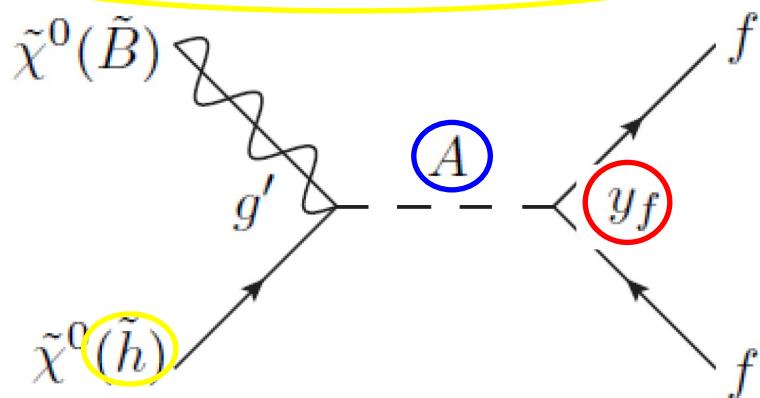
Large scattering cross-section with nuclei:



T. Schwetz '10

Annihilation mechanisms: CP-odd scalar exchange

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$$\propto \frac{m_{\tilde{\chi}_1^0}^2}{m_A^2} \frac{m_f}{v} (N_{11} N_{13,14}) \tan \beta$$

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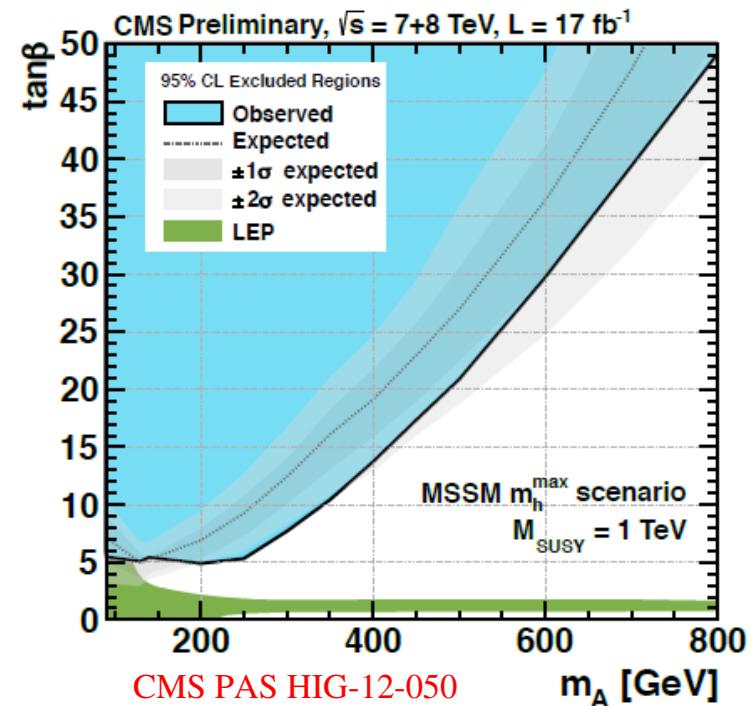
Bottino et al. '02- '10

This parameter space for light neutralinos is now excluded by:

- $B_s \rightarrow \mu^+ \mu^-$
- Searches for extended scalar sector at the LHC,
 $pp \rightarrow X \Phi \rightarrow \tau \tau$
- Mass and couplings of the 126 GeV scalar

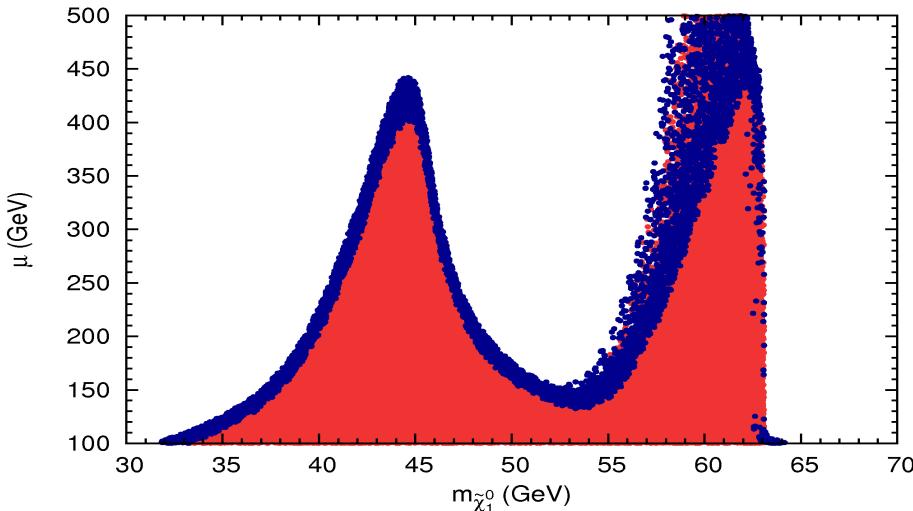
LC Ota Takanishi '11

➡ MSSM not compatible with 10 GeV DM

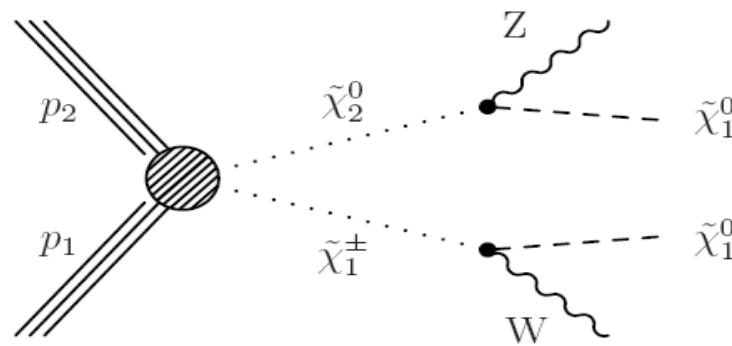


What if all sfermions are heavy?

Dark Matter can hide close to the Z or h resonances



Again, LHC searches can test it!



Problematic for
direct/indirect searches

