### Effective Approach to Dark Matter Decay into γ-Ray Line

and constraints from cosmic rays



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19th of March 2014 Rencontres de Moriond - YSF

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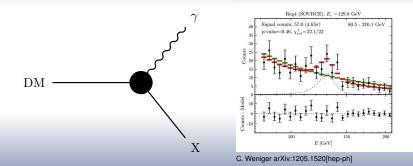


based on M. Gustafsson, T. Hambye, T.S. Phys.Lett. B724 (2013) 288-295 C. El Asaiti, T.Hambye, T.S. arXiv:1403.1280 [hep-ph]



## Interest of $\gamma$ -Ray Lines

 A γ-ray line (monochromatic photon) is a "smoking-gun" signature of DM





# DM Decay into $\gamma$ -Ray Lines

- Dark Matter is (almost) neutral: coupling to the photon?
  - neutral DM: coupling to the photon through a loop
  - millicharged DM: very small tree-level DM-γ coupling
- DM is really long lived: why study a decay?
  - An *accidental* symmetry might provide a long but finite lifetime to the DM
  - Allows for indirect detection signal with distinct features with respect to annihilation



- Model-independent approach
- Justified in the case of a decay

$$au_{DM} \ge 10^{26} {
m sec} ~~ {O_6 \over \Lambda^2} \ m_{DM} \simeq 100 {
m GeV} ~~ \Lambda \simeq 10^{15} {
m GeV}$$

 $\begin{array}{l} \mbox{Full list of operators of dimension five and six} \\ SU(2)_L \times U(1)_Y \times G_{\mbox{hidden}} \mbox{ invariant, giving} \\ \psi^0_{DM} / \psi^{\mbox{milliQ}}_{DM} \rightarrow X + \gamma \end{array}$ 



### **Lists of Operators**

### **Neutral DM**

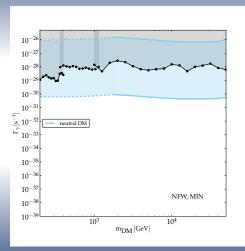
Fermion DMScalar DMVector DM $\bar{\psi}\sigma_{\mu\nu}\psi_{DM}F^{\mu\nu}(\phi)$  $\phi_{DM}F_{\mu\nu}F^{\mu\nu}(\phi)$  $F^{DM}_{\mu\nu}F^{\nu\rho}F^{\prime\prime\rho}_{\mu}$  $D_{\mu}\bar{\psi}\gamma_{\nu}\psi_{DM}F^{\mu\nu}$  $D_{\mu}\phi_{DM}D_{\nu}\phi F^{\mu\nu}$  $F^{DM}_{\mu\nu}F^{\mu\nu}\phi(\phi')$  $\bar{\psi}\gamma_{\mu}D_{\nu}\psi_{DM}F^{\mu\nu}$  $D^{DM}_{\mu}\phi D^{DM}_{\nu}\phi' F^{\mu\nu}$ 

#### Millicharged DM

Fermion DMScalar DMVector DM $D_{\mu}D_{\nu}\bar{\psi}\sigma_{\mu\nu}\psi_{DM}(\phi)$  $\phi_{DM}F^{A}_{\mu\nu}F^{A\mu\nu}(\phi)$  $\phi F^{A}_{\mu\nu}F^{A\mu\nu}(\phi')$  $\bar{\psi}\sigma_{\mu\nu}D_{\mu}D_{\nu}\psi_{DM}(\phi)$  $F^{A\mu\nu}D_{\mu}\phi_{DM}D_{\nu}\phi'$  $F^{A\mu\nu}D_{\mu}\phi_{D\nu}\phi'$  $D_{\mu}\bar{\psi}\sigma_{\mu\nu}D_{\nu}\psi_{DM}(\phi)$  $F^{A\mu\nu}F^{A}_{\mu\nu}F^{A}_{\mu\nu}$ 

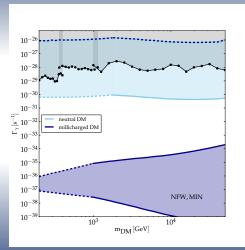


### **Cosmic Rays Constraints**





### **Cosmic Rays Constraints**



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### Conclusions

- The complete list of operators up to dimension six that could account for neutral or millicharged DM decay into γ-ray lines has been written down.
- Due to gauge invariance, to each operator producing a γ-ray signal, there is an associated operator producing a continuum of cosmic rays, which puts constraints on the maximum intensity of the γ-ray line.
- In the case of neutral DM, various operators could account for a γ-ray line if it were detected at the present experimental sensitivities, and there is a potential to discriminate among them.
- Instead a millicharged DM decay is bounded to produce very weak γ-ray lines except if it is a singlet of the SM gauge group.