

# Azimuthal correlations of Mueller-Navelet jets at the LHC

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Annual Meeting of the GDR PH-QCD

25 November 2013

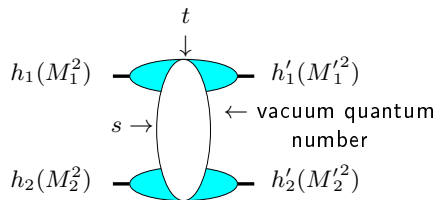
in collaboration with

L. Szymanowski (NCBJ Warsaw), S. Wallon (UPMC & LPT Orsay)

B. D, L. Szymanowski, S. Wallon, JHEP **1305** (2013) 096 [arXiv:1302.7012 [hep-ph]]

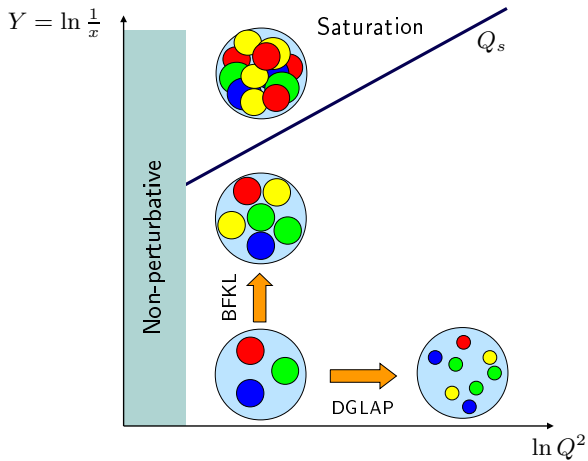
B. D, L. Szymanowski, S. Wallon, arXiv:1309.3229 [hep-ph]

- One of the important longstanding theoretical questions raised by QCD is its behaviour in the perturbative **Regge** limit  $s \gg -t$
- Based on theoretical grounds, one should identify and test suitable observables in order to test this peculiar dynamics



hard scales:  $M_1^2, M_2^2 \gg \Lambda_{QCD}^2$  or  $M_1'^2, M_2'^2 \gg \Lambda_{QCD}^2$  or  $t \gg \Lambda_{QCD}^2$   
 where the  $t$ -channel exchanged state is the so-called **hard Pomeron**

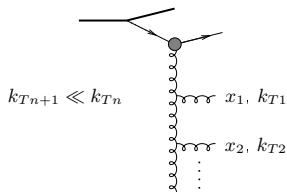
# The different regimes of QCD



Small values of  $\alpha_S$  (perturbation theory applies due to hard scales) can be compensated by large logarithmic enhancements.

⇒ resummation of  $\sum_n (\alpha_S \ln A)^n$  series

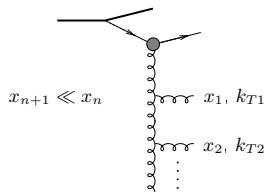
DGLAP



strong ordering in  $k_T$

$$\sum (\alpha_S \ln \frac{Q^2}{\mu^2})^n$$

BFKL



strong ordering in  $x$

$$\sum (\alpha_S \ln \frac{s}{s_0})^n$$

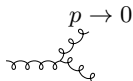
When  $\sqrt{s}$  becomes very large, it is expected that a BFKL description is needed to get accurate predictions

## What kind of observables?

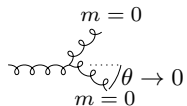
- perturbation theory should be applicable:

selecting external or internal probes with transverse sizes  $\ll 1/\Lambda_{QCD}$  or by choosing large  $t$  in order to provide the hard scale

- governed by the *soft* perturbative dynamics of QCD



and *not* by its *collinear* dynamics



$\Rightarrow$  select semi-hard processes with  $s \gg p_{T i}^2 \gg \Lambda_{QCD}^2$  where  $p_{T i}^2$  are typical transverse scale, **all of the same order**

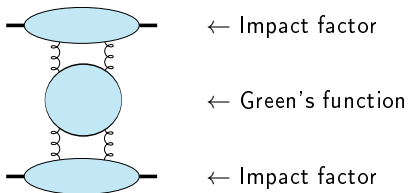
## QCD in the perturbative Regge limit

The amplitude can be written as:

$$\mathcal{A} = \underbrace{\text{Diagram 1}}_{\sim s} + \left( \text{Diagram 2} + \text{Diagram 3} + \dots \right) + \left( \text{Diagram 4} + \dots \right) + \dots$$

$\sim s$ 
 $\sim s (\alpha_s \ln s)$ 
 $\sim s (\alpha_s \ln s)^2$

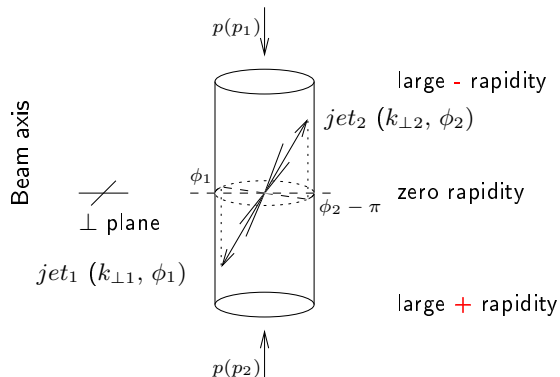
this can be put in the following form :



- Higher order corrections to BFKL kernel are known at NLL order (Lipatov Fadin; Camici, Ciafaloni), now for arbitrary impact parameter  $\alpha_S \sum_n (\alpha_S \ln s)^n$  resummation
- impact factors are known in some cases at NLL
  - $\gamma^* \rightarrow \gamma^*$  at  $t = 0$  (Bartels, Colferai, Gieseke, Kyrielleis, Qiao; Balitski, Chirilli)
  - forward jet production (Bartels, Colferai, Vacca; Caporale, Ivanov, Murdaca, Papa, Perri; Chachamis, Hentschinski, Madrigal, Sabio Vera)
  - inclusive production of a pair of hadrons separated by a large interval of rapidity (Ivanov, Papa)
  - $\gamma_L^* \rightarrow \rho_L$  in the forward limit (Ivanov, Kotsky, Papa)

## Mueller-Navelet jets

- Consider two jets (hadrons flying within a narrow cone) **separated by a large rapidity**, i.e. each of them almost fly in the direction of the hadron “close” to it, and with very similar transverse momenta
- in a pure LO collinear treatment, these two jets should be emitted **back to back** at leading order:  $\Delta\phi - \pi = 0$  ( $\Delta\phi = \phi_1 - \phi_2 =$  relative azimuthal angle) and  $k_{\perp 1} = k_{\perp 2}$ . There is no phase space for (untagged) emission between them





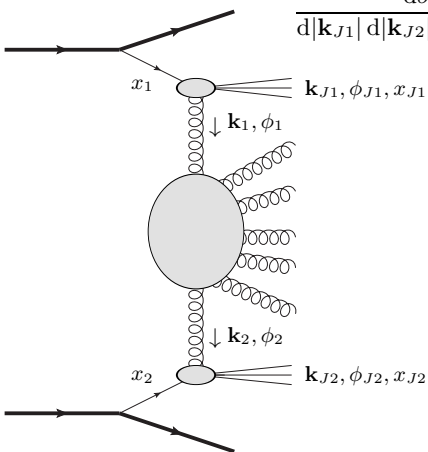
$k_T$ -factorized differential cross section

$$\frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}} = \int d\phi_{J1} d\phi_{J2} \int d^2\mathbf{k}_1 d^2\mathbf{k}_2$$

$$\times \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1)$$

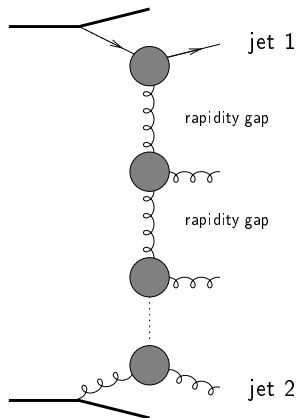
$$\times G(\mathbf{k}_1, \mathbf{k}_2, \hat{s})$$

$$\times \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2)$$



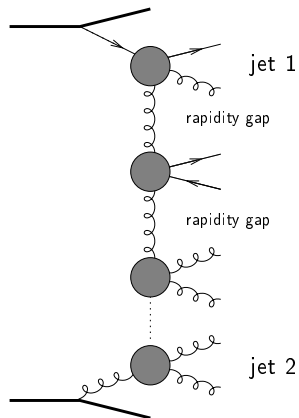
$$\text{with } \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2) = \int dx_2 f(x_2) V(\mathbf{k}_2, x_2) \quad f \equiv \text{PDF} \quad x_J = \frac{|\mathbf{k}_J|}{\sqrt{s}} e^{y_J}$$

LL BFKL



$$\sum (\alpha_s \ln s)^n$$

NLL BFKL



$$\sum (\alpha_s \ln s)^n + \alpha_s \sum (\alpha_s \ln s)^n$$

## Results for a symmetric configuration

In the following we show results for

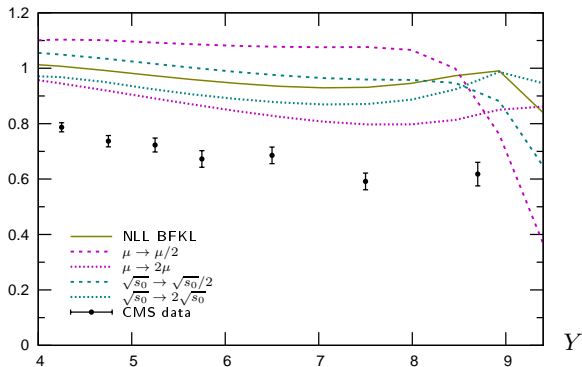
- $\sqrt{s} = 7 \text{ TeV}$
- $35 \text{ GeV} < |\mathbf{k}_{J1}|, |\mathbf{k}_{J2}| < 60 \text{ GeV}$
- $0 < y_1, y_2 < 4.7$

These cuts allow us to compare our predictions with the first experimental data on azimuthal correlations of Mueller-Navelet jets from the LHC presented by the **CMS** collaboration (CMS-PAS-FSQ-12-002)

note: unlike experiments we have to set an upper cut on  $|\mathbf{k}_{J1}|$  and  $|\mathbf{k}_{J2}|$ . We have checked that our results don't depend on this cut significantly.

Azimuthal correlation  $\langle \cos \varphi \rangle$ 

$$\langle \cos \varphi \rangle \equiv \langle \cos(\phi_{J1} - \phi_{J2} - \pi) \rangle$$



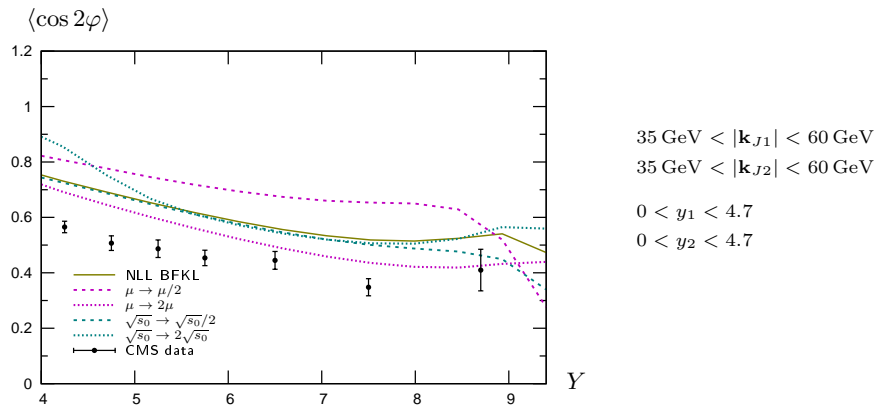
$$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$$

$$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$$

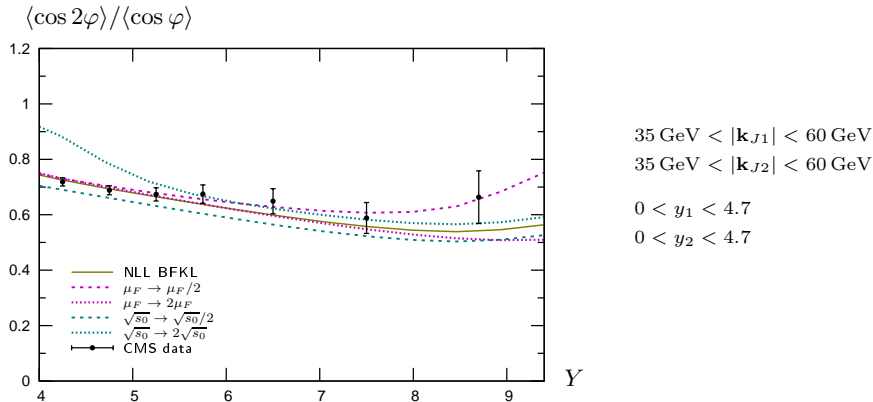
$$0 < y_1 < 4.7$$

$$0 < y_2 < 4.7$$

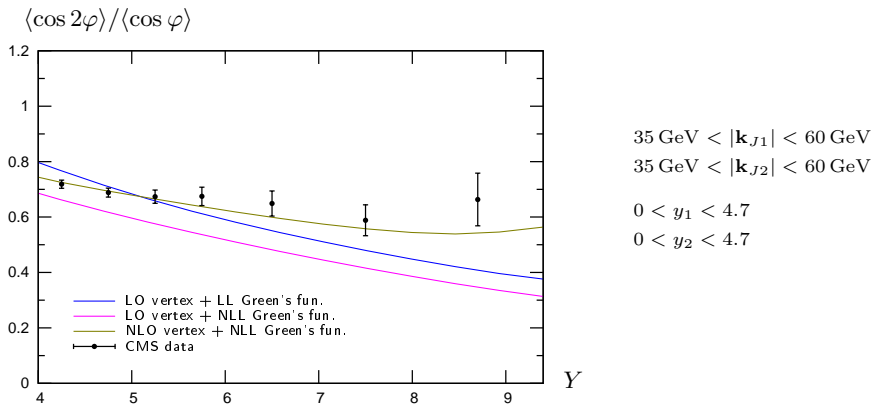
- NLL BFKL predicts a too small decorrelation
- The NLL BFKL calculation is still rather dependent on the scales, especially the renormalization / factorization scale

Azimuthal correlation  $\langle \cos 2\varphi \rangle$ 

- The agreement with data is a little better for  $\langle \cos 2\varphi \rangle$  but still not very good
- This observable is also very sensitive to the scales

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

- This observable is more stable with respect to the scales than the previous ones
- The agreement with data is good across the full  $Y$  range

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

It is necessary to include the NLO corrections to the jet vertex to reproduce the behavior of the data at large  $Y$

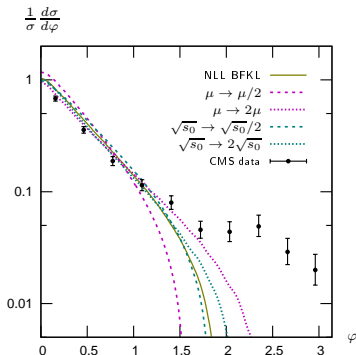
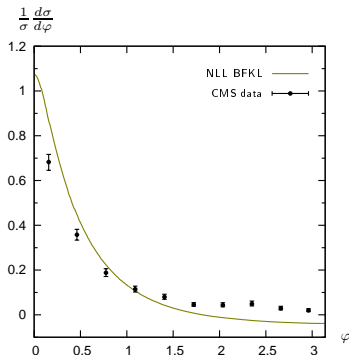
## Azimuthal distribution

The azimuthal distribution  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  has also been measured by the CMS collaboration. It can be written as

$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}$$



## Azimuthal distribution: comparison to CMS data

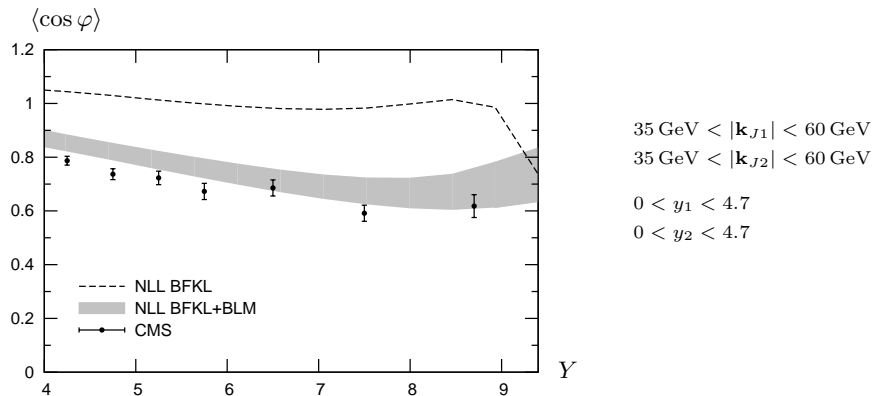


- Our calculation predicts a too large value of  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  for  $\varphi \lesssim \frac{\pi}{2}$  and a too small value for  $\varphi \gtrsim \frac{\pi}{2}$
- For large values of  $\varphi$ , the distribution even becomes negative

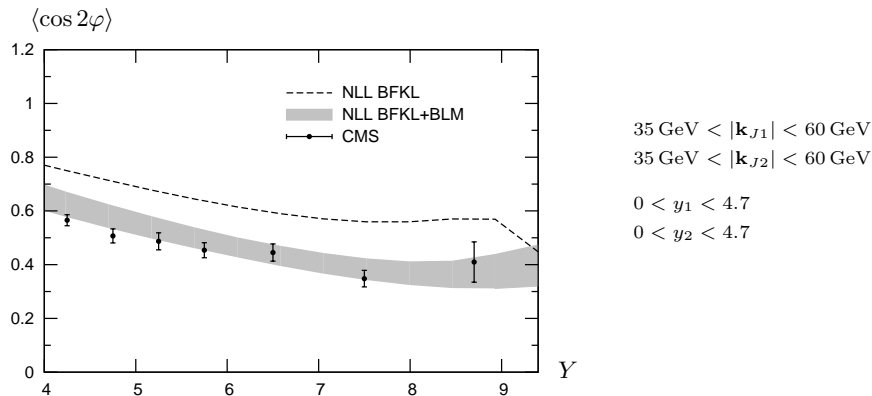
- The agreement of our calculation with the data for  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is good and very stable with respect to the scales
- The agreement for  $\langle \cos n\varphi \rangle$  and  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  is not very good and very sensitive to the choice of the renormalization scale  $\mu_R$
- An all-order calculation would be independent of the choice of  $\mu_R$ . This feature is lost if we truncate the perturbative series  
 $\Rightarrow$  How to choose the renormalization scale?  
 'Natural scale': sometimes the typical momenta in a loop diagram are different from the natural scale of the process

The **Brodsky-Lepage-Mackenzie** (BLM) procedure resums the self-energy corrections to the gluon propagator at one loop into the running coupling. These contributions are formally of higher-order but they are proportional to

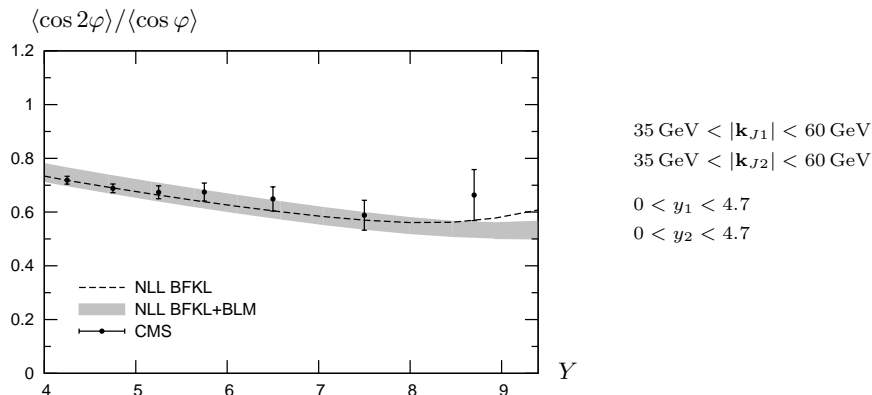
$$\beta_0 = \frac{11N_c - 2N_f}{3} \simeq 7.67$$

Azimuthal correlation  $\langle \cos \varphi \rangle$ 

Using the BLM scale setting, the scale uncertainty is reduced and the agreement with data becomes much better

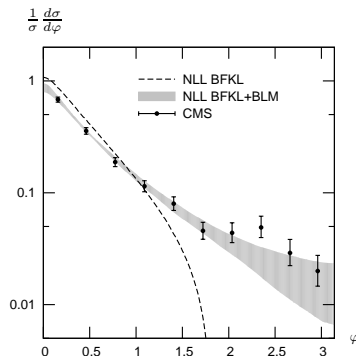
Azimuthal correlation  $\langle \cos 2\varphi \rangle$ 

Using the BLM scale setting, the scale uncertainty is reduced and the agreement with data becomes much better

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

Because it is much less dependent on the scales, the observable  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is almost not affected by the BLM procedure and is still in very good agreement with the data

## Azimuthal distribution: comparison to CMS data



With the BLM scale setting the azimuthal distribution no longer reaches negative values and is in good agreement with the data across the full  $\varphi$  range.

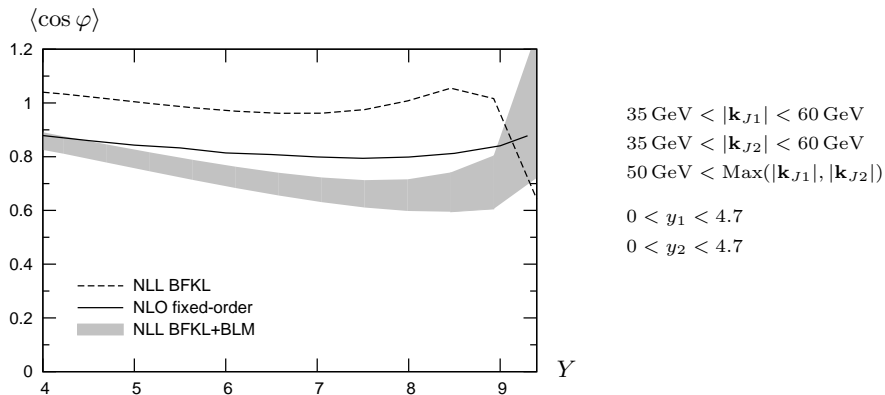
Using the BLM scale setting:

- The agreement  $\langle \cos n\varphi \rangle$  with the data becomes much better
- The agreement for  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  is still very good and unchanged as this observable is weakly dependent on  $\mu_R$
- The azimuthal distribution no longer reaches negative values and is in much better agreement with the data

But the configuration chosen by CMS with  $\mathbf{k}_{J_{\min 1}} = \mathbf{k}_{J_{\min 2}}$  does not allow to compare with a *fixed-order* treatment (i.e. without resummation)

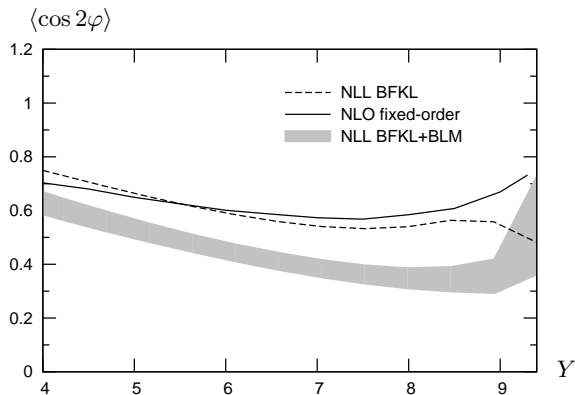
We compare our results with the NLO fixed-order code Dijet (Aurenche, Basu, Fontannaz) in an asymmetric configuration

- $35 \text{ GeV} < |\mathbf{k}_{J_1}|, |\mathbf{k}_{J_2}| < 60 \text{ GeV}$
- $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J_1}|, |\mathbf{k}_{J_2}|)$
- $0 < y_1, y_2 < 4.7$

Azimuthal correlation  $\langle \cos \varphi \rangle$ 

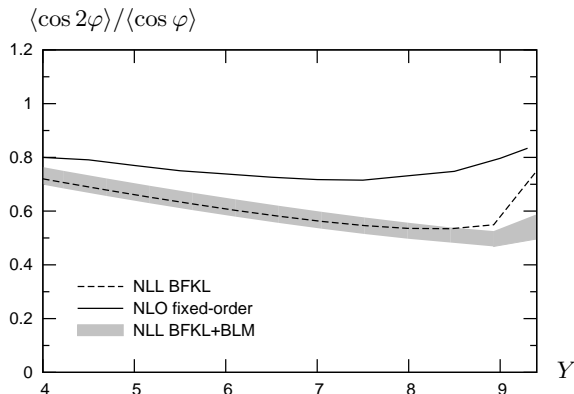
- As in the symmetric case, the BLM procedure strongly modifies the result of our BFKL calculation
- The NLO fixed-order and NLL BFKL+BLM calculations are very close



Azimuthal correlation  $\langle \cos 2\varphi \rangle$ 

$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$   
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$   
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$   
 $0 < y_1 < 4.7$   
 $0 < y_2 < 4.7$

- As in the symmetric case, the BLM procedure strongly modifies the result of our BFKL calculation
- The BLM procedure leads to a larger difference between NLO fixed-order and NLL BFKL+BLM

Azimuthal correlation  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

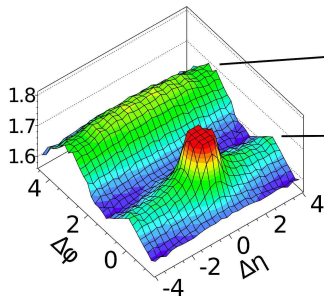
$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$   
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$   
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$   
 $0 < y_1 < 4.7$   
 $0 < y_2 < 4.7$

Using BLM or not, we see a **sizable difference** between BFKL and fixed-order  
 $\Rightarrow$  An experimental analysis with enough statistics should provide clear discrimination between these two treatments

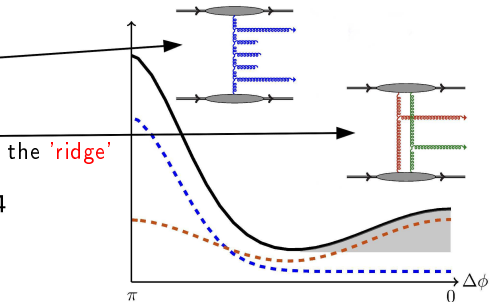
The spectra of two-particle correlations in high multiplicity proton-proton, proton-nucleus and nucleus-nucleus collisions at RHIC and the LHC show two structures:

- one corresponding to almost back-to-back jets
- one corresponding to jets emitted in the same direction (the 'ridge').

The first one can be understood as a manifestation of BFKL dynamics and the second one can be explained by 'Glasma graphs' (Dusling, Venugopalan).



CMS p-Pb,  $\sqrt{s_{NN}} = 5.02$  TeV  
 $N \geq 110$ ,  $1 < p_T < 3$  GeV



the 'ridge'

Dusling, Venugopalan

- We studied Mueller-Navelet jets at full (vertex + Green's function) **NLL** accuracy and compared our results with the first data from the **LHC**
- The observables  $\langle \cos n\varphi \rangle$  and  $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$  are very dependent on the choice of the scales and don't agree very well with data
- The agreement with **CMS** data is greatly improved by using the **BLM** scale fixing procedure
- For the observable  $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ :
  - NLL BFKL predictions are much more stable with respect to the scales
  - the data is well described by BFKL in a **symmetric** configuration
  - there is a clear difference between **NLO fixed-order** and our **NLL BFKL** calculation in an **asymmetric** configuration

⇒ In our opinion this is a strong motivation for an experimental analysis in an asymmetric configuration