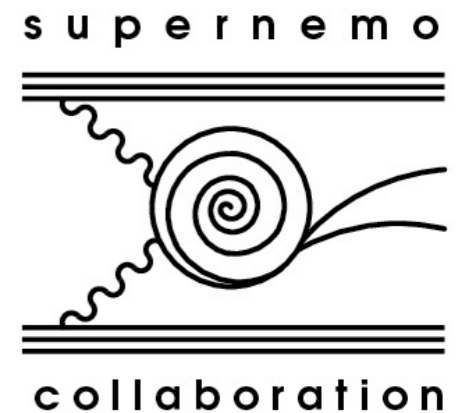


# SuperNEMO

within the labex ENIGMASS  
(and few words on  $0\nu\beta\beta$  physics)



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# Is there any Majorana particles?

In 1936 E. Majorana:

- A *real* wave equation to describe massive & electrically neutral fermion:
  - **particle** == **antiparticle**

In 1957 B. Pontecorvo:

- Set the basis for the neutrino flavour oscillation

In the last 15 years:

- Huge experimental effort → confirm **neutrino oscillation** and measure parameters

The neutrino is a **fermion**, is **electrically neutral** and is **massive**

- **Is it a Majorana particle?**

Interesting implication for particle physics:

- Lepton number violation must occur ( $0\nu\beta\beta$  decay)
- GUT, Leptogenesis model, See-Saw mechanism

**$0\nu\beta\beta$  decay is the only practical way to test Majorana nature of neutrinos**

# The neutrino-less double beta decay

$$2\nu\beta\beta \text{ decay: } (A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e$$

- 2nd order process **allowed** in the SM
- Single  $\beta$  decay forbidden (energy & angular momentum)

$$0\nu\beta\beta \text{ decay: } (A, Z) \rightarrow (A, Z + 2) + 2e^-$$

- process **forbidden** in the SM

$$(T_{1/2}^{0\nu})^{-1} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \eta^2$$

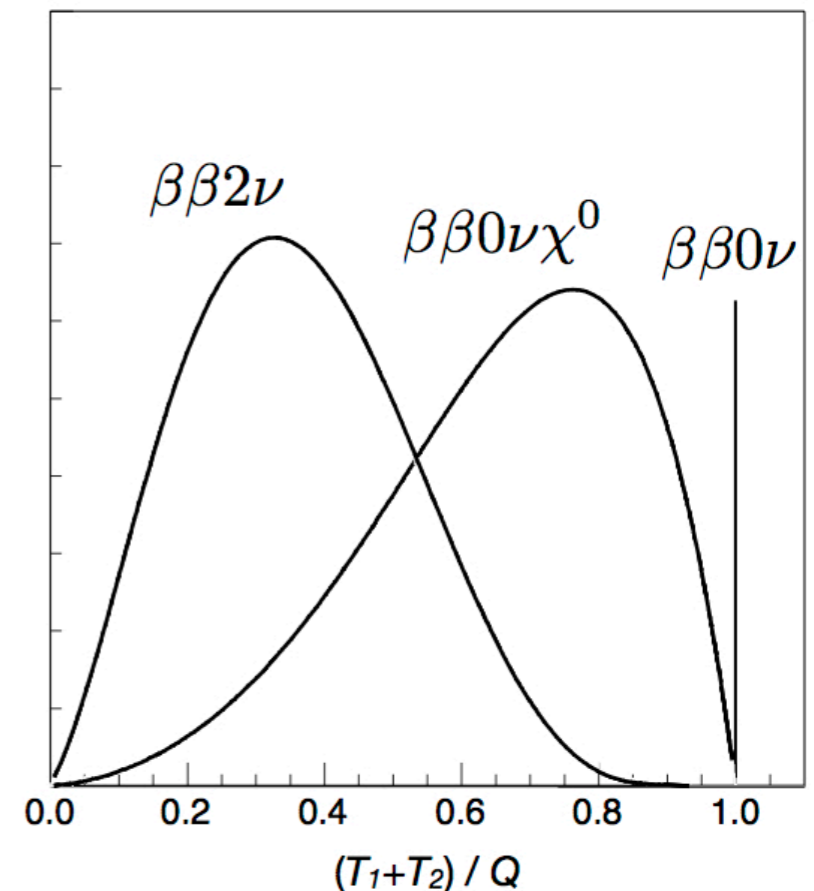
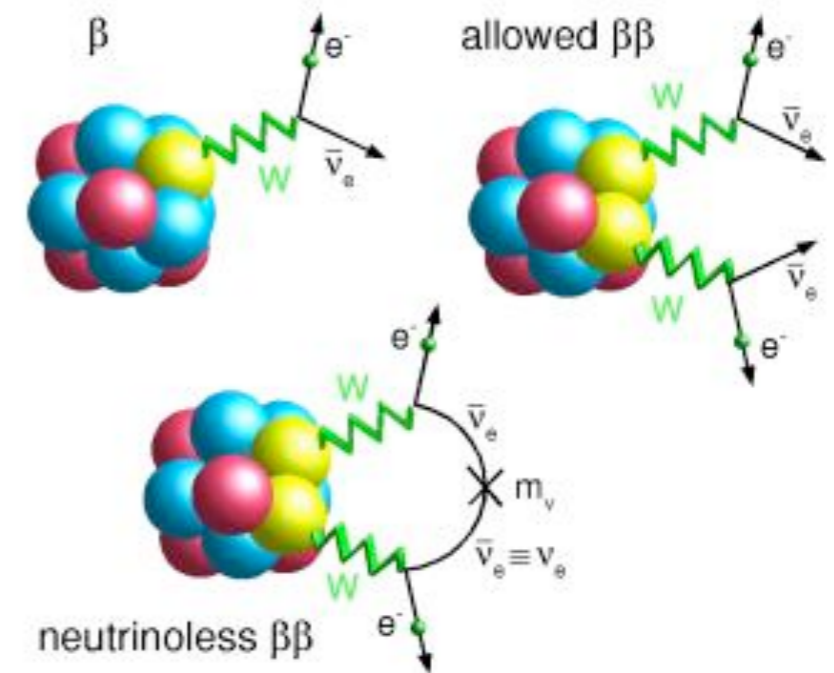
$G_{0\nu}$ : Phase space term (atomic physics)

$M_{0\nu}$ : Nuclear matrix element (nuclear physics)

$\eta$ : decay mechanism (particle physics)

- **Light Majorana neutrino exchange**
- Right-handed current (V+A), SUSY, I Majoron, etc.

Different event topology in the final state

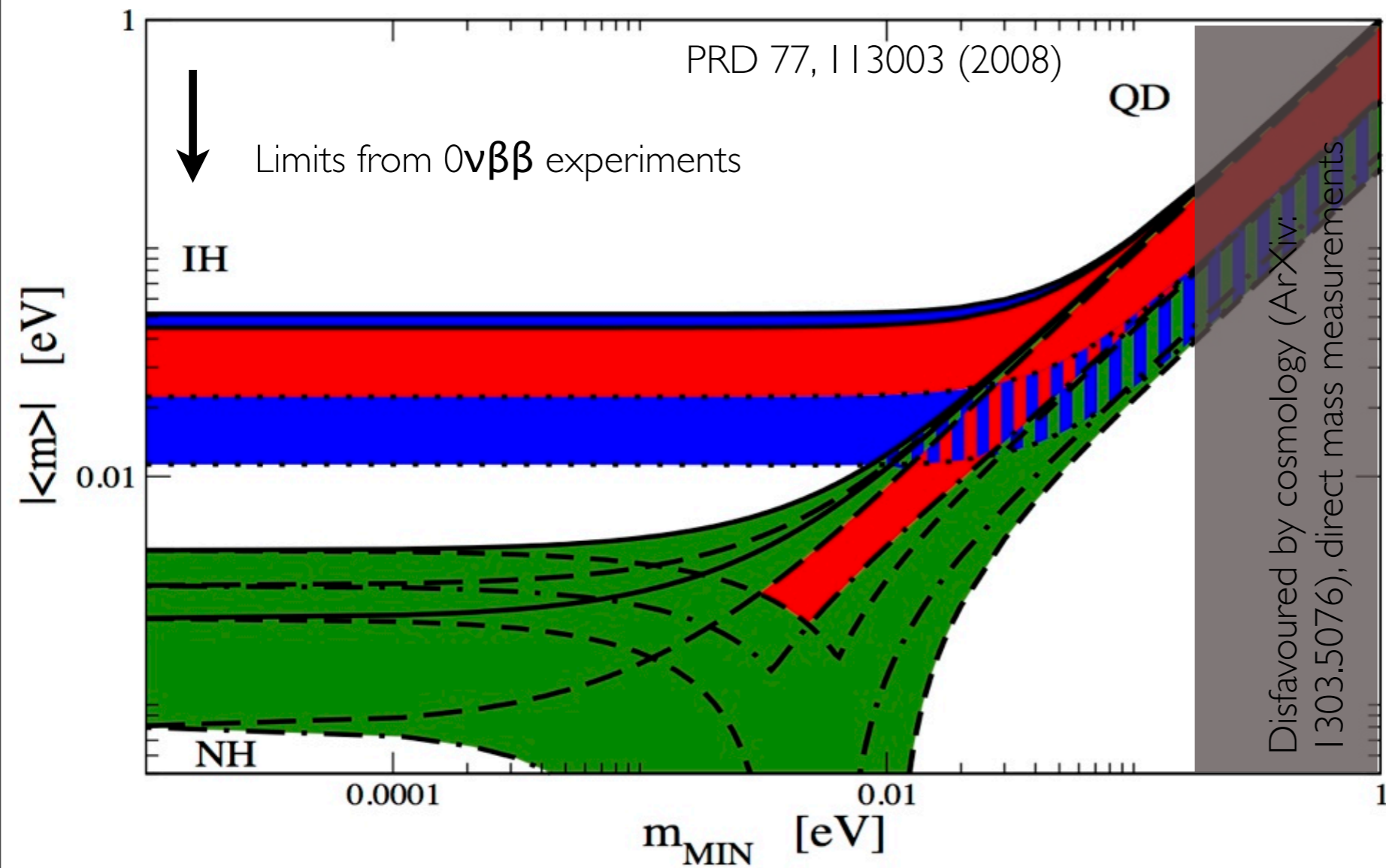


# Sensitivity on neutrino mass scale

$$\left(T_{1/2}^{0\nu}\right)^{-1} = G_{0\nu}(Q_{\beta\beta}, Z) |M_{0\nu}|^2 \frac{\langle m_{ee} \rangle^2}{m_e^2}$$

Effective mass term:

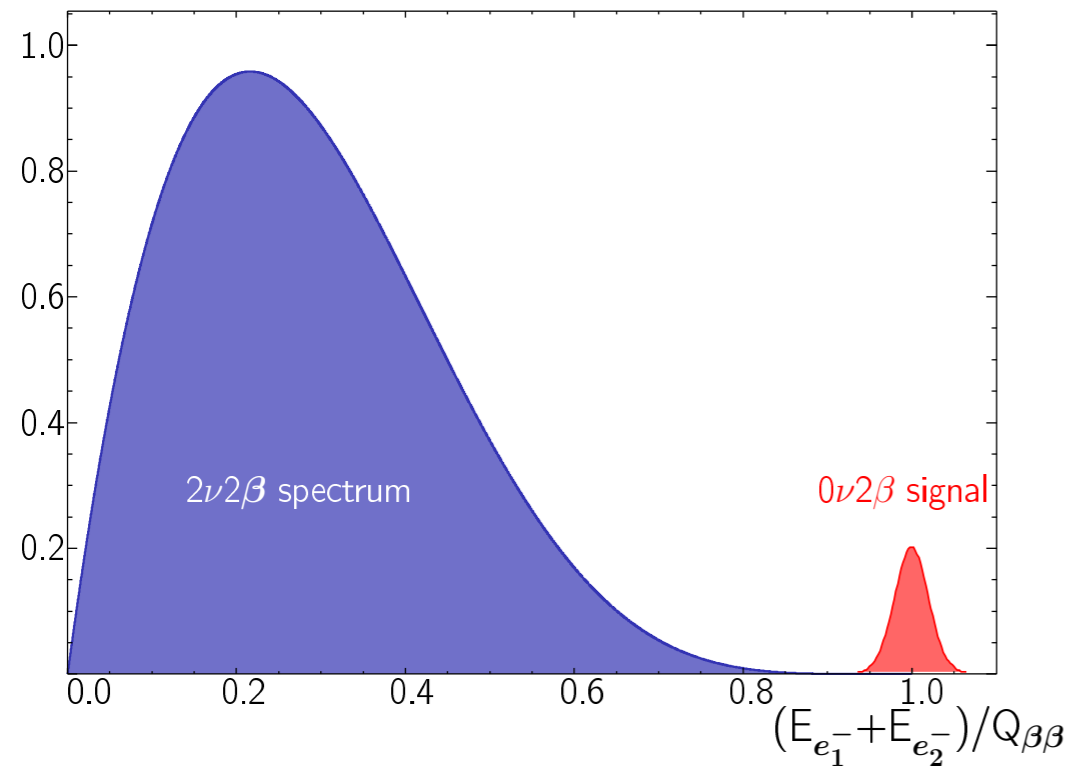
$$\langle m_{ee} \rangle = \left| \sum_i U_{ei}^2 m_i \right|$$



Related to neutrino oscillation  
 Sensitive to the neutrino mass scale ( $m_{\text{min}}$ )

**CP violation**  
**Normal Hierarchy**  
**Inverted Hierarchy**

# Searching for $0\nu\beta\beta$ process



Measure the 2  $e^-$  energy spectrum

- $2\nu\beta\beta$  signature  $\rightarrow$  Broad spectrum
- $0\nu\beta\beta$  signal signature  $\rightarrow$  **Peak** @  $Q_{\beta\beta}$

If no signal  $\rightarrow$  set a **limit** on half life

$$T_{1/2}^{0\nu} > \frac{N_A \ln 2}{n_\sigma} \times \frac{\epsilon}{A} \times \sqrt{\frac{M \times t}{B \times \Delta E}}$$

Labels and arrows in the diagram:

- $n_\sigma$ : Excluded events at a given C.L.
- $N_A$ : Atomic mass
- $\epsilon$ : Detection efficiency
- $A$ :  $\beta\beta$  emitter mass
- $B$ : Bkg. index
- $\Delta E$ : E res. @  $Q_{\beta\beta}$
- $M \times t$ :  $\beta\beta$  emitter mass  $\times$  Exposure time

# Some general remarks

Rare process ( $T_{1/2} \approx 10^{25} \text{ y}$ )  $\rightarrow$  Long **exposure**, high isotopic **mass**

Low energy process ( $Q_{\beta\beta} \approx 5 \text{ MeV}$ ):

- **Natural radioactivity** is an issue ( $^{232}\text{Th}$ ,  $^{238}\text{U}$ , ...)
  - $\gamma$  (up to 2.6 MeV,  $^{208}\text{Tl}$ ) and  $\beta$  (up to 3.2 MeV,  $^{214}\text{Bi}$  Q-value)
  - Ultra **high radio-purity material**, dedicated shielding
- Cosmic muons are an issue  $\rightarrow$  **Deep underground** lab

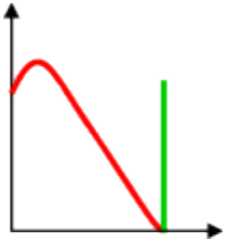
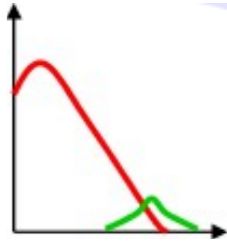
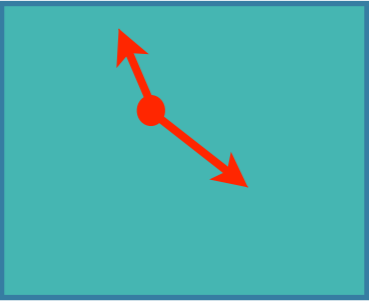

Distinguish  $0\nu$  from  $2\nu$  mode  $\rightarrow$  the ultimate background

- Good detector **energy resolution**
- Long  $2\nu$  mode half life (less  $2\nu$  background @ end point)

Many available isotopes  $\rightarrow$  **multiply** experimental efforts

- Adopt different detection techniques
- **Independent** measurements

# Experimental technique

<b>En. resolution</b> <b>Technique</b>	<b>&lt; 1 %</b> 	<b>&gt; 1 %</b> 
<p><b>Calorimeter</b></p>  <p><b>Detector = Source</b></p>	<p><b>Ge diodes</b></p> <ul style="list-style-type: none"> <li>• Efficiency, PID</li> </ul> $^{76}\text{Ge}$	<p><b>Liquid scintillators</b></p> <ul style="list-style-type: none"> <li>• Mass</li> </ul> $^{136}\text{Xe}$ , $^{150}\text{Nd}$ , $^{48}\text{Ca}$ , $^{100}\text{Mo}$
	<p><b>Bolometers</b></p> <ul style="list-style-type: none"> <li>• Efficiency, (PID)</li> </ul> $^{130}\text{Te}$ , $^{82}\text{Se}$ , $^{100}\text{Mo}$	<p><b>Liquid Xe TPC</b></p> <ul style="list-style-type: none"> <li>• Mass, PID</li> </ul> $^{136}\text{Xe}$
<p><b>Electron Tracking</b></p>  <p><b>Source</b></p> <p><b>Detector</b></p>	<p><b>Gas Xe TPC</b></p> <ul style="list-style-type: none"> <li>• (PID)</li> </ul> $^{136}\text{Xe}$	<p><b>Tracko-Calo</b></p> <ul style="list-style-type: none"> <li>• PID, Full kinematic</li> </ul> $^{82}\text{Se}$ , $^{150}\text{Nd}$ , $^{48}\text{Ca}$ , ...
		<p><b>Pixelized Scintillator</b></p> <ul style="list-style-type: none"> <li>• PID</li> </ul> $^{116}\text{Cd}$



# What's the status?

1993 - 2000:

- **HdM** (~11 kg) & **IGEX** (~2 kg),  $^{76}\text{Ge}$ 
  - $T^{0\nu}_{1/2} > 1.9 \cdot 10^{25} \text{ y @ 90\% C.L.}$
- **HdM claim:**  $\langle m_{ee} \rangle = 0.32 \pm 0.03 \text{ eV}$

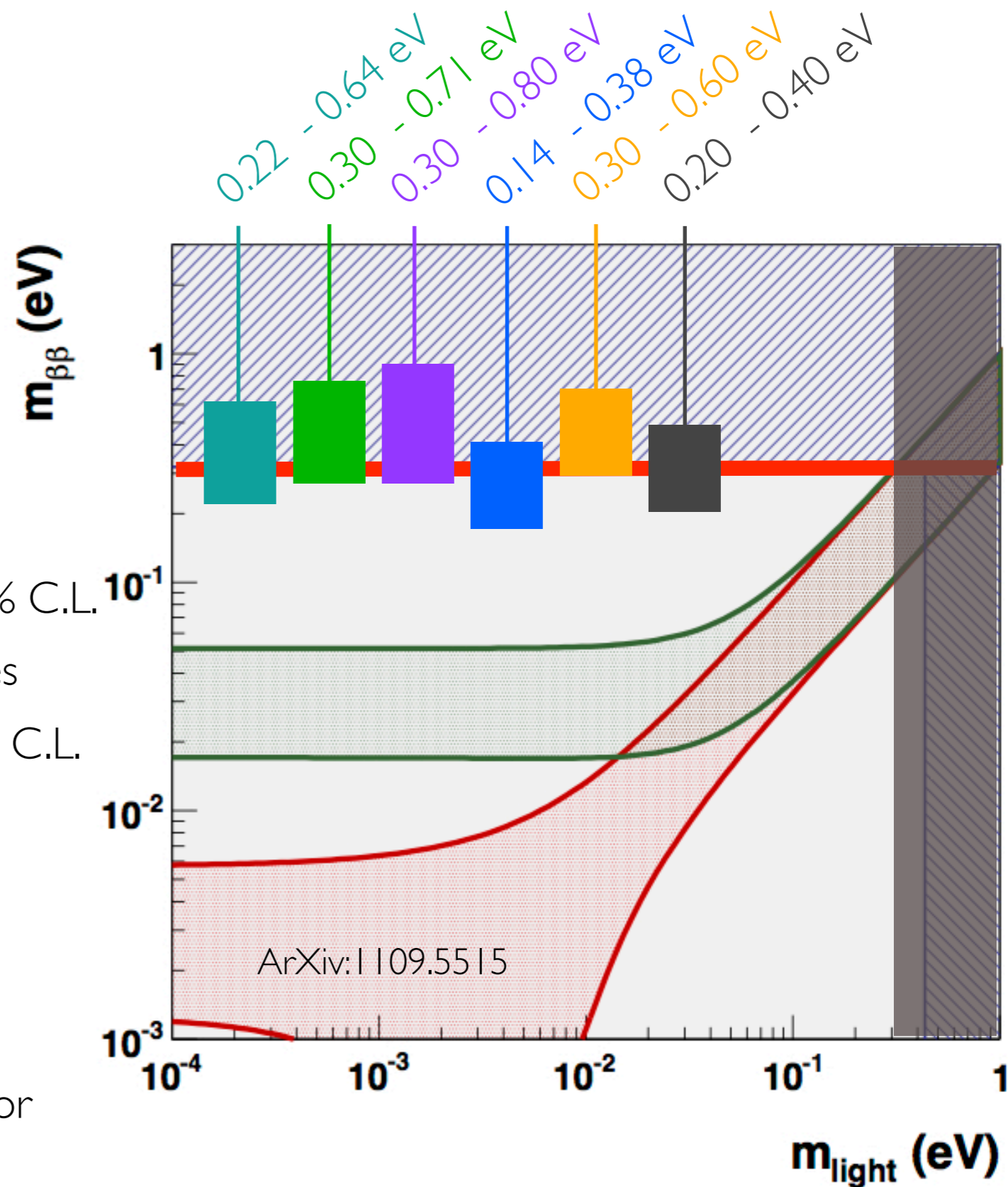
2000 - 2010:

- **Cuoricino:**  $\text{TeO}_2$  bolometric detector
  - ~11 kg  $^{130}\text{Te}$ :  $T^{0\nu}_{1/2} > 2.8 \cdot 10^{24} \text{ y @ 90\% C.L.}$
- **NEMO3:** Tracko-Calor, 7 different isotopes
  - ~7 kg  $^{100}\text{Mo}$ :  $T^{0\nu}_{1/2} > 1.1 \cdot 10^{24} \text{ y @ 90\% C.L.}$

Since 2011: new generation

10 - 100 kg, R&D for future scaling

- **EXO200** ( $^{136}\text{Xe}$ ): Liquid TPC
- **Kamland-ZEN** ( $^{136}\text{Xe}$ ): Liquid Scintillator
- **GERDA Phase I** ( $^{76}\text{Ge}$ ): Ge diodes





# Future projects

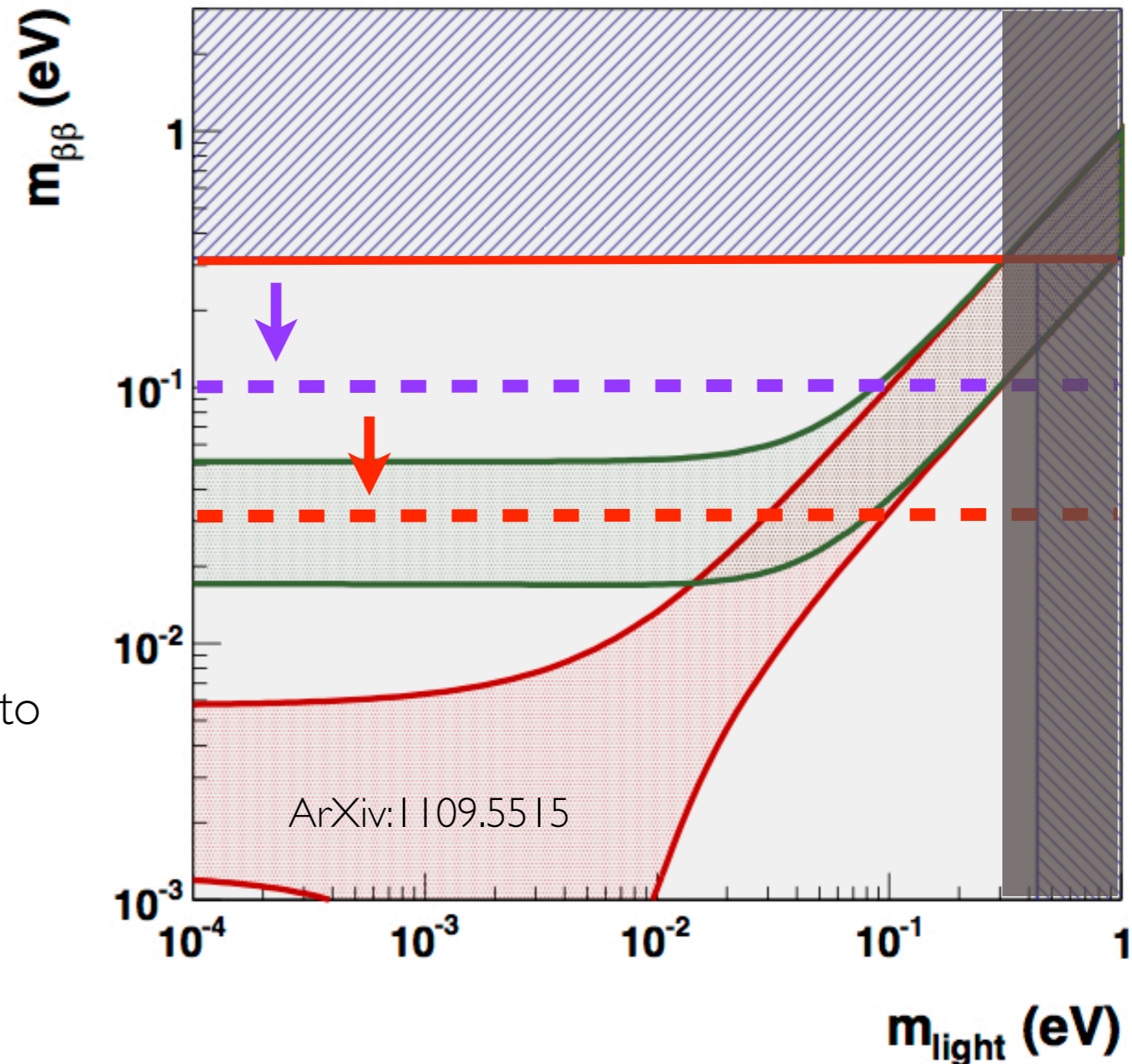
## 5 years time scale:

- $M \sim 10 - 50$  kg of  $\beta\beta$  isotope
- Background level  $10^{-3}$  cts./ (keV kg y)
- Explore quasi-degenerate region

## 10 years time scale:

- $M \sim 100$  kg - 1t of  $\beta\beta$  isotope
- Background level  $10^{-4}$  cts./ (keV kg y)
- Approach Inverse Hierarchy region
- Extended R&D: Energy resolution, particle ID, radio-purity
- Multi-phase approach: demonstrate scalability to higher mass and background levels

CUORE, Gerda, Majorana, Lucifer, AMORE, NEXT, COBRA, EXO, SNO+, KamLAND-Zen, Candels, **SuperNEMO**, MOON, DCBA, ...

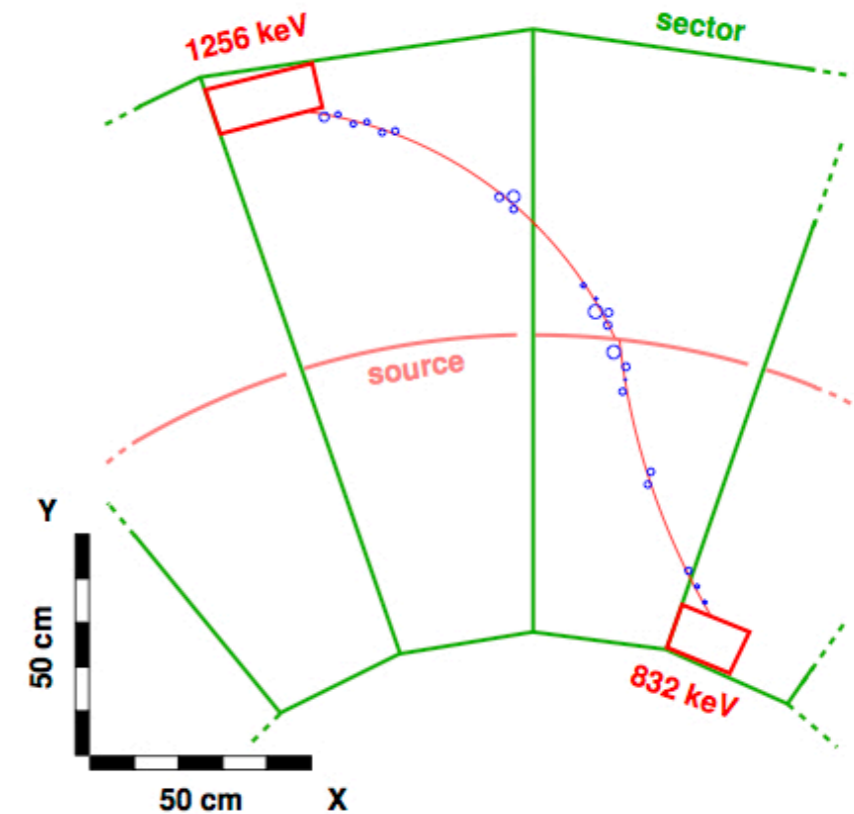


# NEMO3 and the Tracko-Calo technique

Running @ LSM 2003 - 2011

**Full** reconstruction of  $2e^-$  kinematics (unique!)

- Individual  $e^-$  energy, arrival time, track curvature in magnetic field, emission vertex and tracks angle
- Low energy resolution:  $[14 - 17] \% / \text{Sqrt}(E)$
- **Excellent** background rejection
- **Equivalent** to best calorimetric experiment



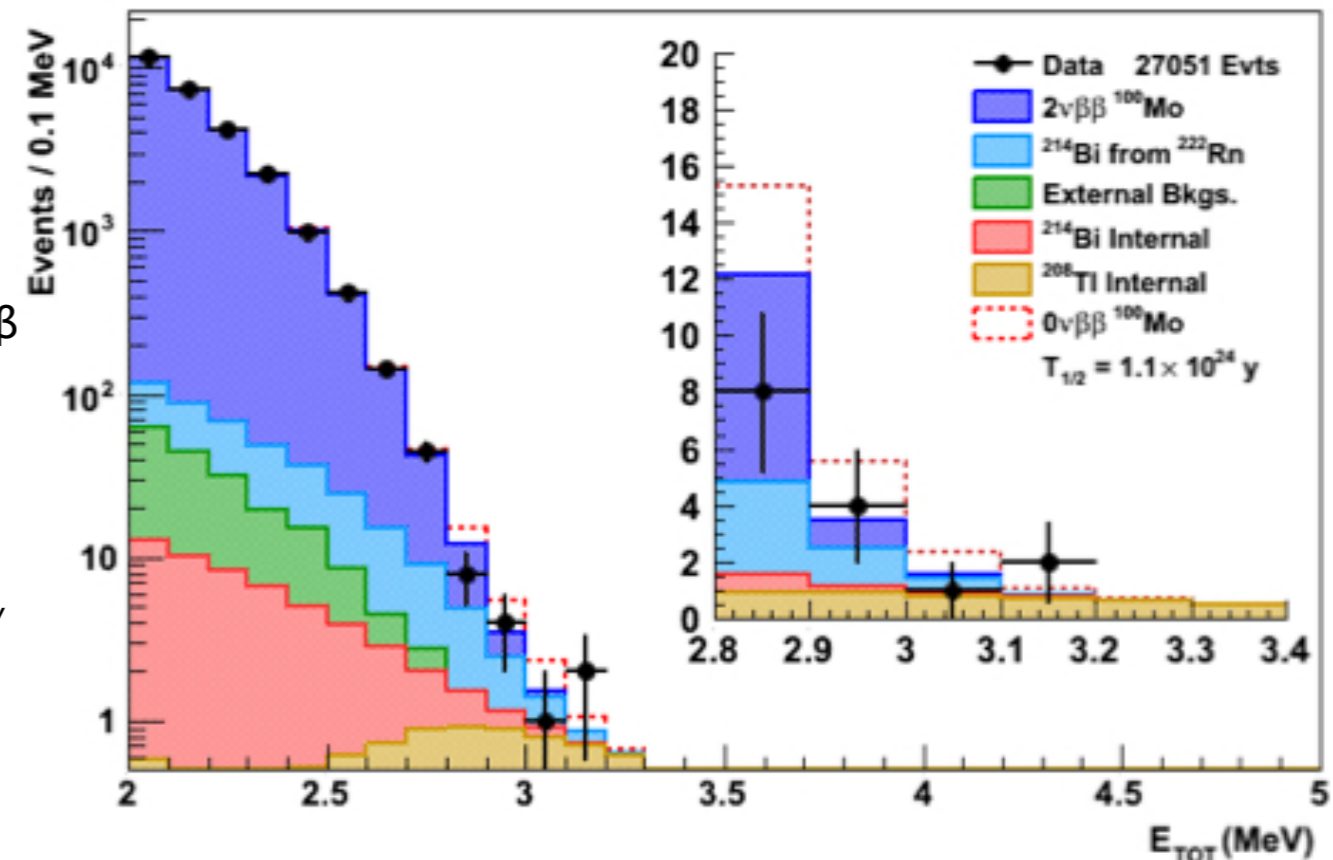
Recent results with full  $^{100}\text{Mo}$  exposure 34.7 kg y

- $T_{1/2}^{0\nu} > 1.1 \cdot 10^{24} \text{ y}$  @ 90% C.L.
- Background level  $\sim 0.02 \text{ cts.} / (\text{keV kg y})$  @  $Q_{\beta\beta}$
- No background event  $> 3.2 \text{ MeV}$

Potential background **free** technique for high energy  $Q_{\beta\beta}$  isotopes ( $^{48}\text{Ca}$ ,  $^{150}\text{Nd}$ ,  $^{96}\text{Zr}$ )

NEMO-3 -  $^{100}\text{Mo}$  - 7 kg, 4.96 y

NEMO3 (TAUP 2013)





# The next step: SuperNEMO

Exploit a well known technique:

- 20 detection module 5 kg  $\beta\beta$  emitter each

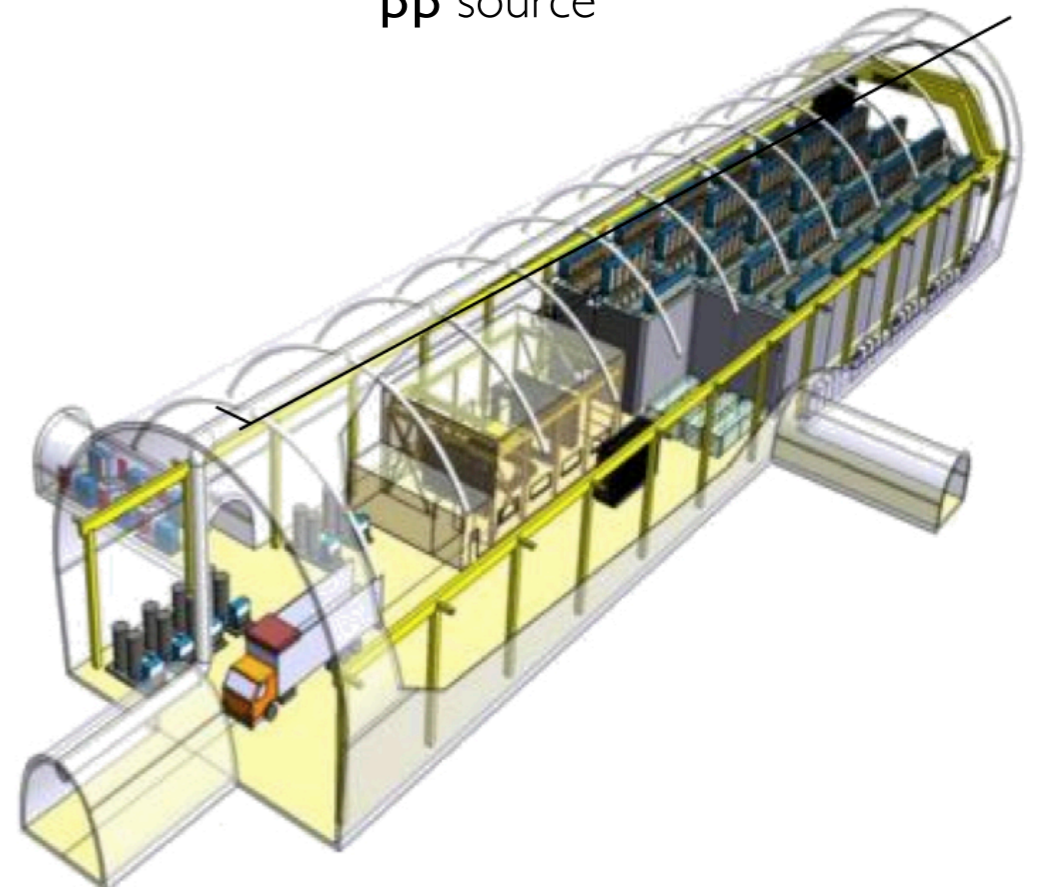
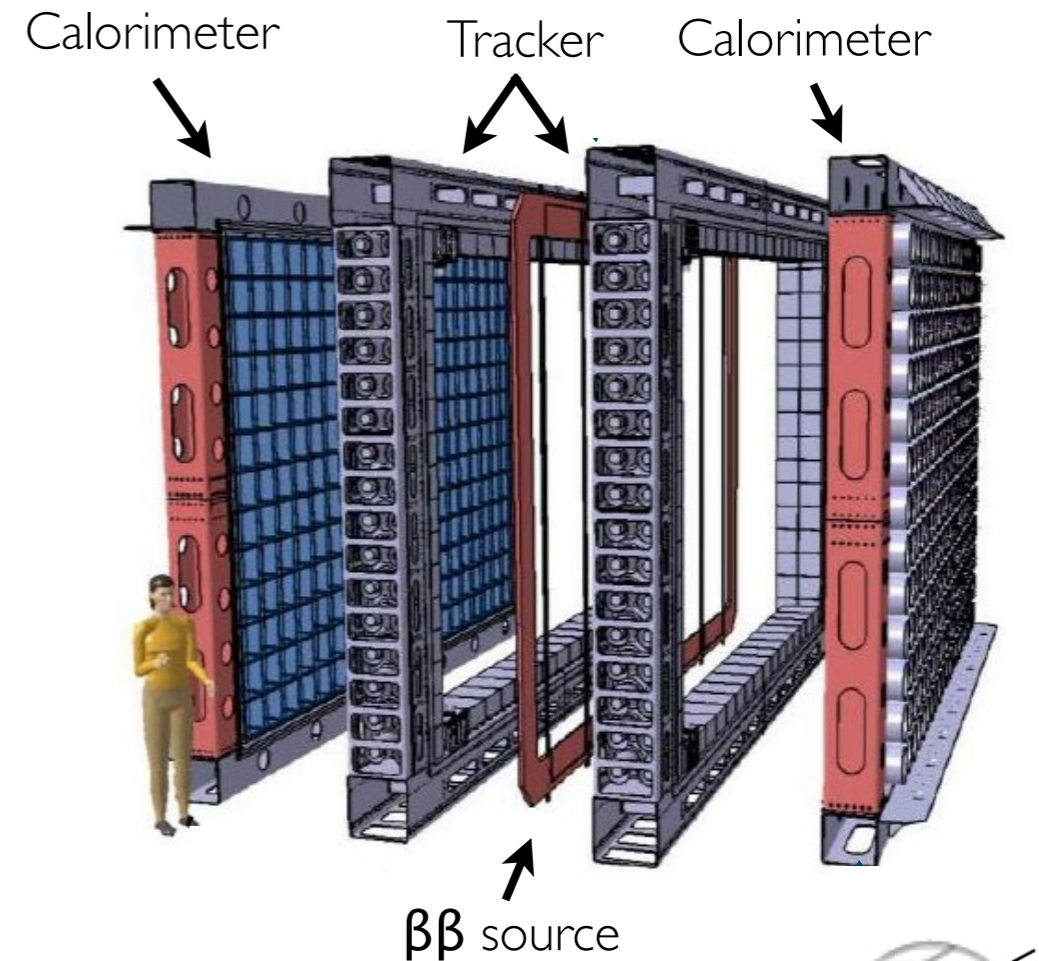
Goal:

- Bkg. level:  $10^{-4}$  cts./(keV kg y)
- Approach IH region

A **challenge** under many aspects:

	NEMO3	SuperNEMO
Efficiency	18%	~30%
Isotope	7 kg $^{100}\text{Mo}$	~100 kg $^{82}\text{Se}$
Exposure	35 kg y	~500 kg y
Energy res.	8% @ 3 MeV	4% @ 3 MeV
$^{208}\text{Tl}$	~100 $\mu\text{Bq/kg}$	< 2 $\mu\text{Bq/kg}$
$^{214}\text{Bi}$	> 300 $\mu\text{Bq/kg}$	< 10 $\mu\text{Bq/kg}$
$T_{1/2}$	$1 \cdot 10^{24}$ y	$1 \cdot 10^{26}$ y
$\langle m_\nu \rangle$	0.31 - 0.79 eV	0.04 - 0.1 eV

- Require LSM extension
- First step  $\rightarrow$  demonstrator module



# The SuperNEMO demonstrator

One SuperNEMO module → **7 kg  $^{82}\text{Se}$  running ~2.5 y**

- To be installed @ LSM (replacing NEMO3)

Match SuperNEMO requirements

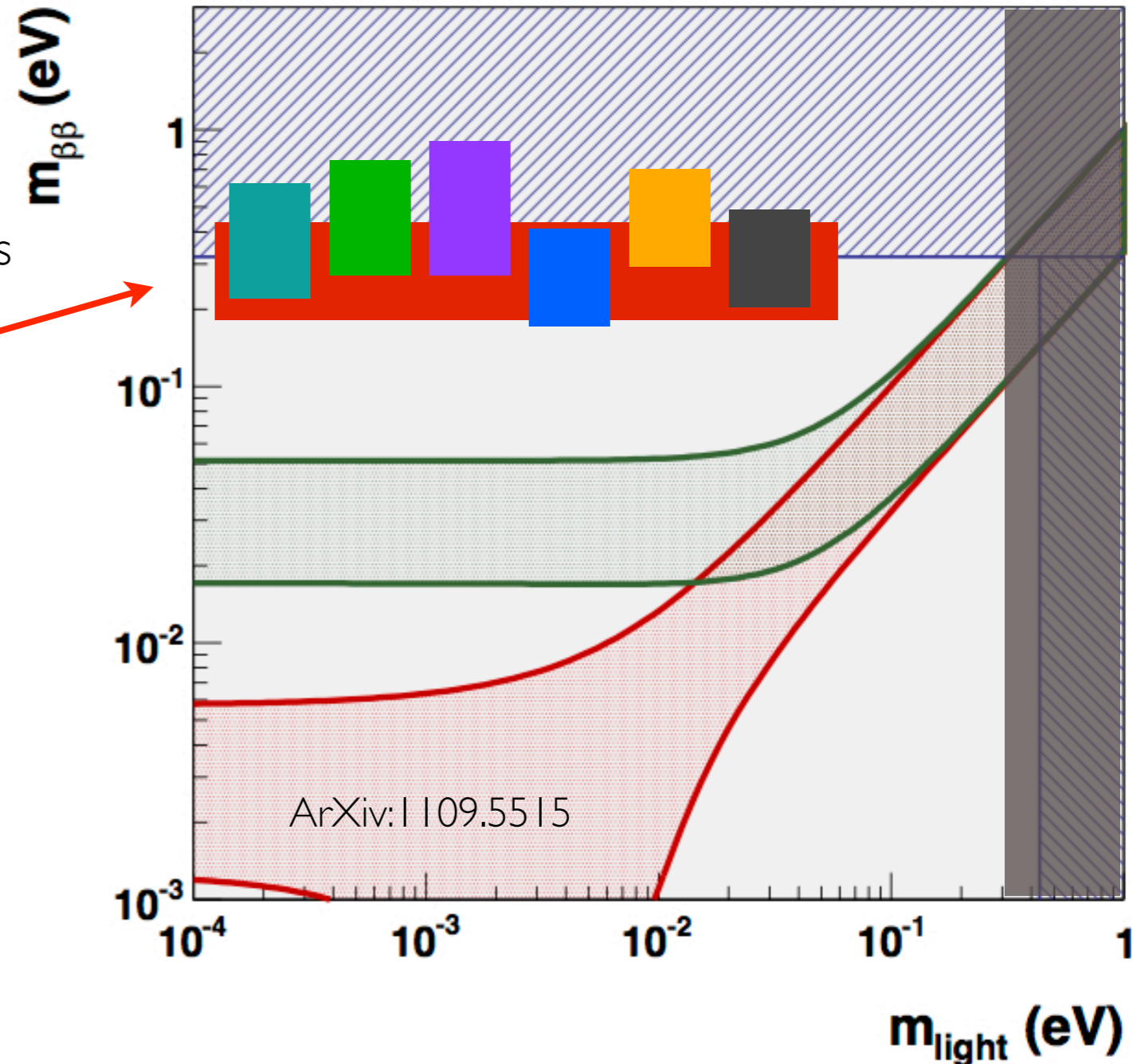
- Less than **0.2 bkg. events** in ROI

Reach NEMO3 ( $^{100}\text{Mo}$ ) sensitivity in 4.5 months

- Sensitivity:  $\langle m_{ee} \rangle \sim$  **0.20 - 0.40 eV**
- Test HdM claim with  $^{82}\text{Se}$

Schedule:

- Installation & commissioning 2014
- First **physics data in 2015**



# Contributions from the labex



- Laboratories: **LSM** & **LAPP**
- Period: 2012 - 2016
- Activities: Development & installation of the SuperNEMO demonstrator module

## **1) R&D and production of $^{82}\text{Se}$ foil source**

## **2) Monte Carlo studies for the foil design optimisation**

## **3) Development of the Slow Control system**

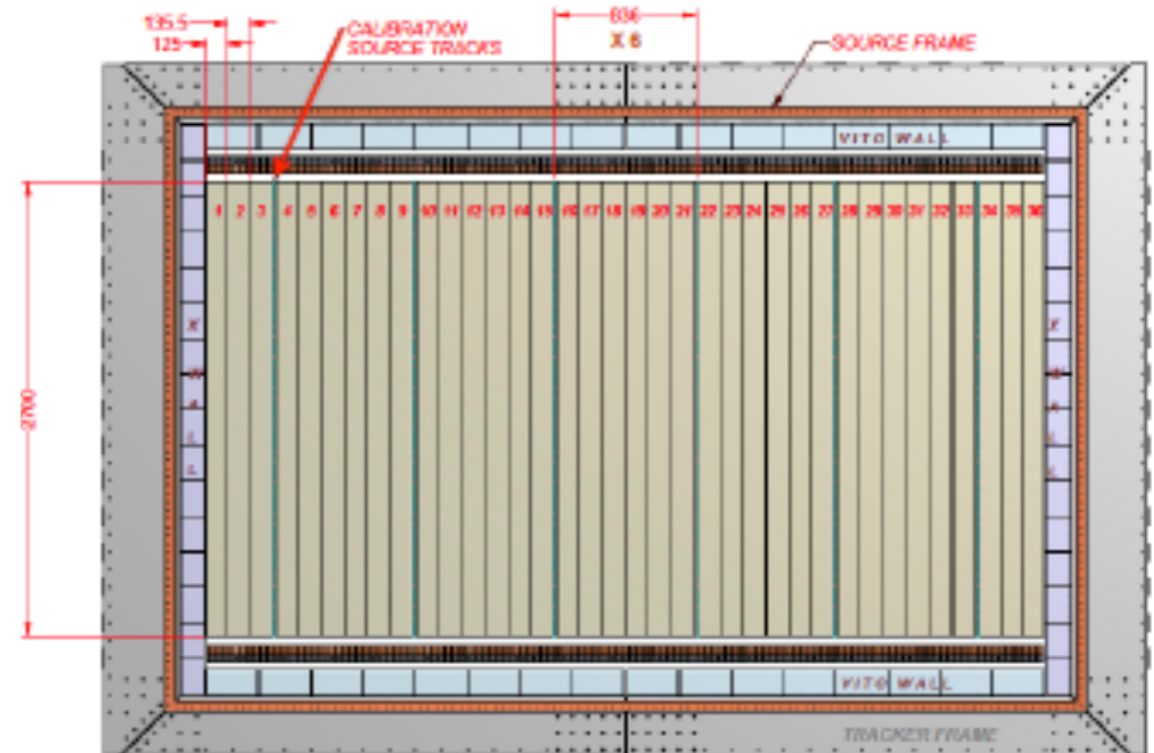
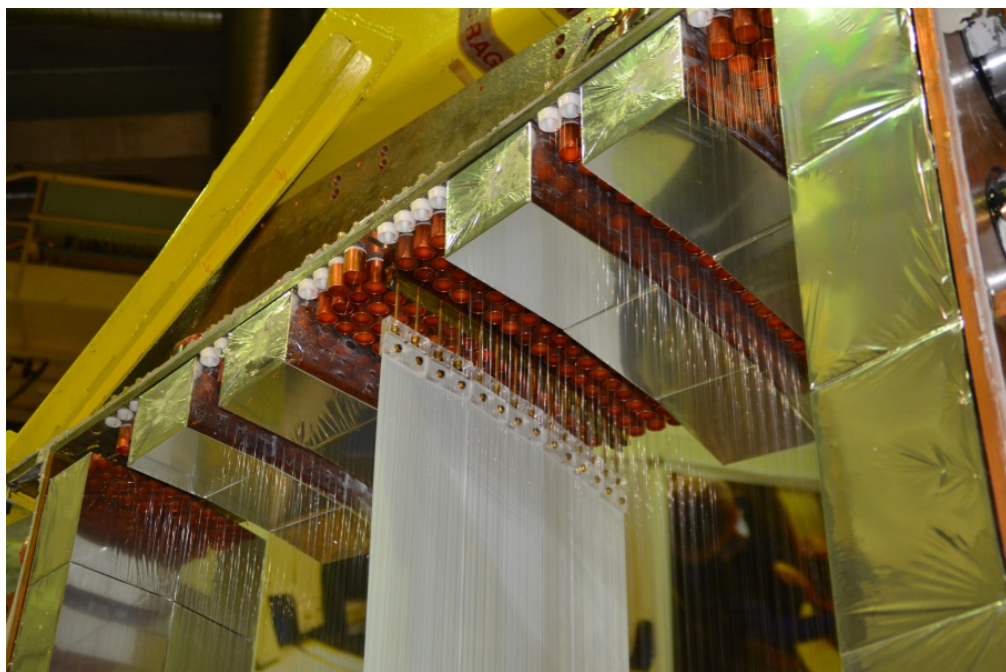
4) Detector installation + commissioning + running (2014 - 2016)

5) Physics data analysis (2015 - 2016)



SuperNEMO  $\beta\beta$  source:

- 36 strips 3 long,  $\sim 200$   $\mu\text{m}$  thick ( $50$   $\text{mg}/\text{cm}^2$ )
- Strong material radio-purity constrain



NEMO3 foil production experience:

- Se + PVA glue + Mylar film
- Mylar is **not radiopure enough** for SuperNEMO
- $A(^{208}\text{Tl}) \sim 9 \pm 3$   $\mu\text{Bq}/\text{kg}$  (limits is  $2$   $\mu\text{Bq}/\text{kg}$ )

An **alternative strategy** is necessary:

- Mechanical department @ LAPP in collaboration with LSM
- $3.5$   $\text{kg}$  of  $^{82}\text{Se}$  available in France  $\rightarrow$  **produce 1/2 demonstrator foils**



## Set up a test bench:

- ISO 5-6 clean room

## R&D for PVA glue

- Samples for radio-purity measurement
  - Ge (LSM): ~1.5 kg PVA powder
  - BiPo (LSC): 20 thin foil → 30x30 cm 200 um thick

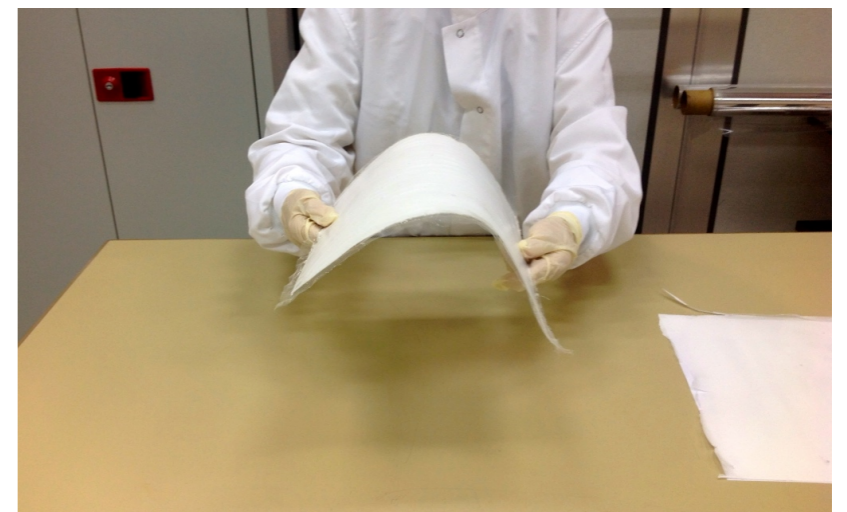
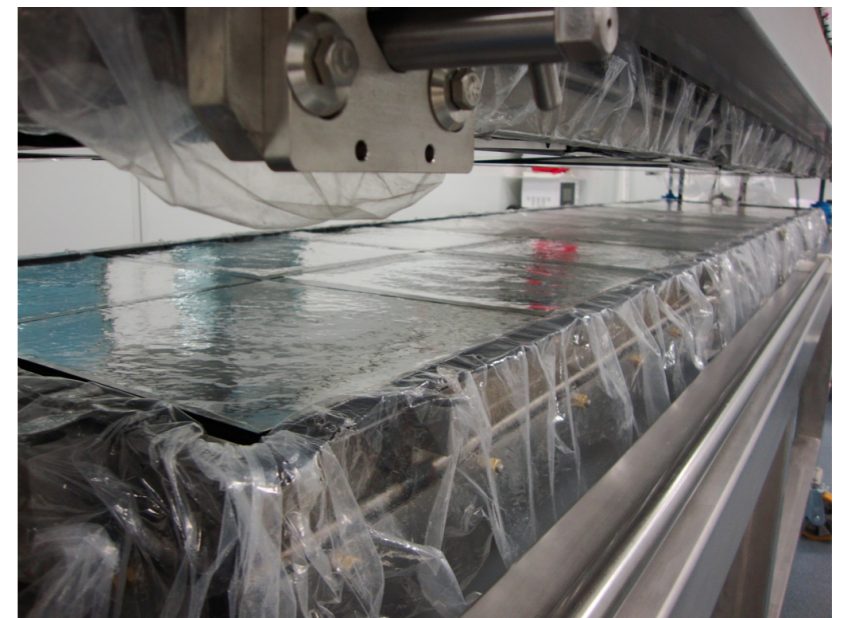
## R&D for new mechanical support

- Fine mesh fabric (Tulle) as central backbone
- Flexible foil with smaller support mass (~ 1 %)
- Setting up 30 m<sup>2</sup> for radio-purity measure

## Se powder processing (grinding/purification)

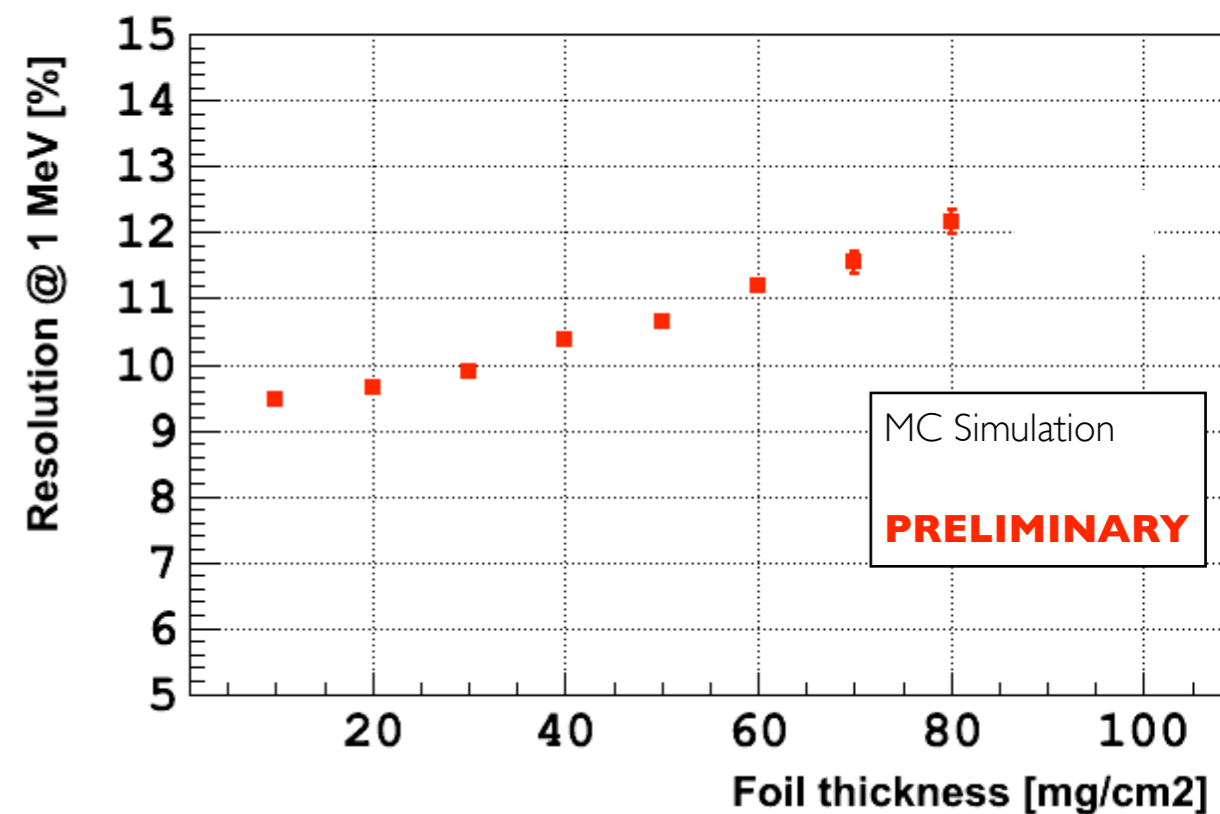
- Choose suitable materials w.r.t. radio-purity requirements
- Purification strategy under study @ LSM

## Full foils production expected by the end of 2014



**Optimise** detailed **foil design** & composition:

- Test different foil design
- Mylar film / Tulle, PVA fraction, foil thickness, ...



- Tune foil parameter w.r.t.  $T^{0v}_{1/2}$  sensitivity → guideline for best foil production

Setting up MC production + Analysis code for full sensitivity study

- **Submit a Master 2 project on the subject (January/June 2014)**

Computing + electronic departments @ LAPP

Develop dedicated software system to:

- **Control/monitor** environmental parameters, detector subsystems (local & remote)
- Operate **heterogeneous** devices

Proposed solution:

- Common choice with CTA project → Take advantage of **existing experience**
  - OPC UA **specification** → A standard issued by HW & SW industrial vendors
  - Generic solution **independent** from context (experiments, technical strategy, devices)
- Definition of a **Interface Control Document** (ICD) to collect devices infos/specifications
- First prototype of integrated hardware to be tested next spring
  - From ICD → to web interface management

**Implement device & user Interfaces by the end of 2014 / beginning 2015**

# Conclusions



- World wide interest on  $0\nu\beta\beta$  searches  $\rightarrow$  answer a fundamental question
- Rich experimental program for the next 5 & 10 years
- SuperNEMO aims to explore the IH region with a Tracko-Calo detector
  - The Demonstrator module is the first step towards the full detector
  - In 2 years time scale will reach sensitivity to test HdM claim with  $^{82}\text{Se}$
- Key contributions from ENIGMASS groups
  - R&D and production of a  $\beta\beta$  emitter source with a new design
  - Slow control system development
  - We're on the front line for the detector installation/commissioning/running

# Backup slides

# Choosing the $\beta\beta$ isotopes

Isotopes enrichment and  $T_{1/2}^{2\nu}$  from respective experiment

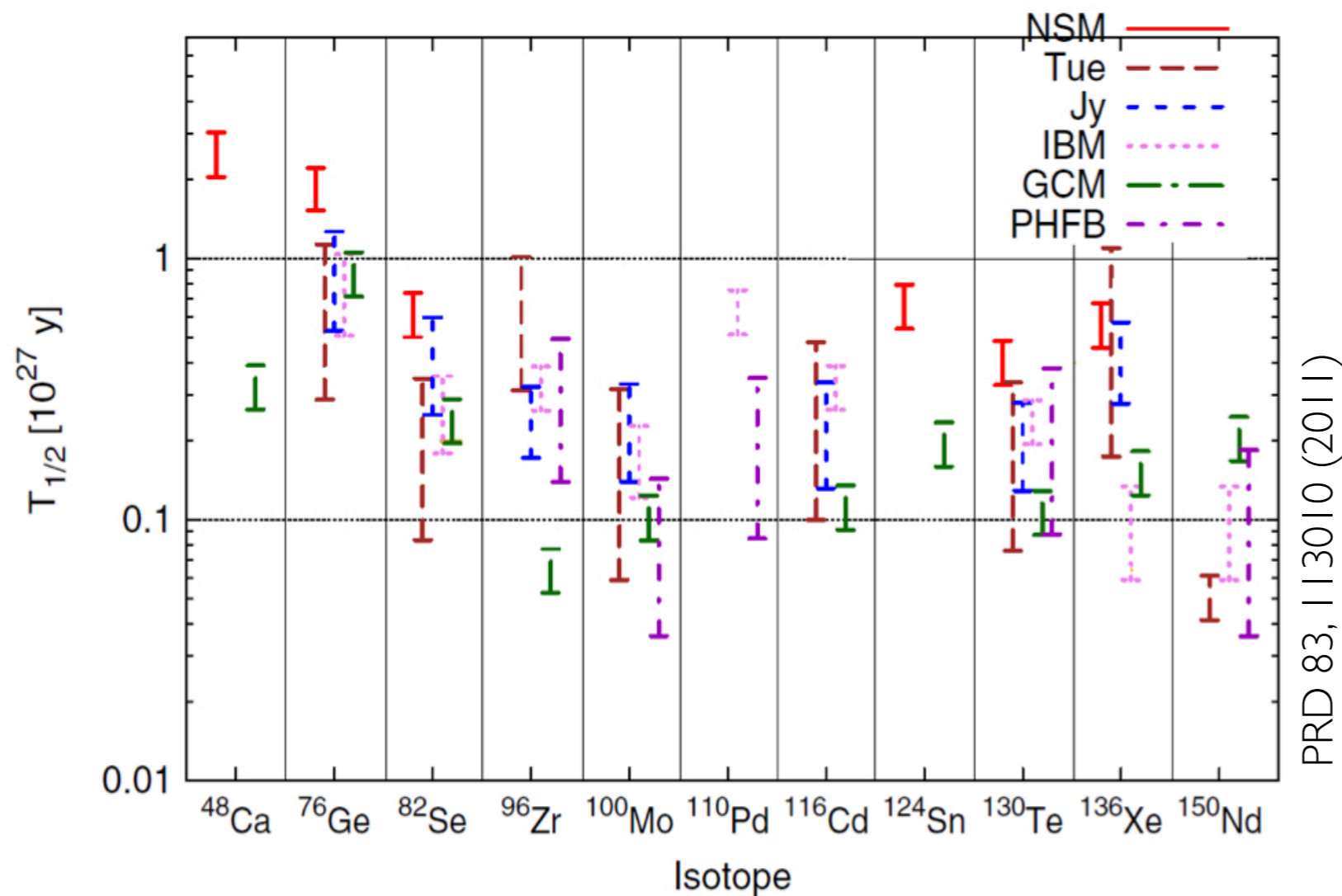
Isotope	$Q_{\beta\beta}$ [keV]	Nat. abund. (enrich.) [%]	$G_{0\nu}$ [ $10^{-14} \text{ y}^{-1}$ ](*)	$T_{1/2}^{2\nu}$ [ $10^{19} \text{ y}$ ]	Experiment
$^{48}\text{Ca}$	4270	0.187 (73)	6.35	$4.2^{+2.1}_{-1.0}$	NEMO3
$^{76}\text{Ge}$	2039	7.8 (86)	0.623	$150 \pm 10$	HM
$^{82}\text{Se}$	2995	8.7 (97)	2.70	$9.0 \pm 0.7$	NEMO3
$^{96}\text{Zr}$	3350	2.8 (57)	5.63	$2.0 \pm 0.3$	NEMO3
$^{100}\text{Mo}$	3034	9.6 (99)	4.36	$0.71 \pm 0.04$	NEMO3
$^{116}\text{Cd}$	2802	7.5 (93)	4.62	$3.0 \pm 0.2$	NEMO3
$^{130}\text{Te}$	2527	34.5 (90)	4.09	$70 \pm 10$	NEMO3
$^{136}\text{Xe}$	2480	8.9 (80)	4.31	$238 \pm 14$	KamlandZEN
$^{150}\text{Nd}$	3367	5.6 (91)	19.2	$0.78 \pm 0.7$	NEMO3



# Nuclear Matrix Element

Contain nuclear structure effects → only **approximative** theoretical calculation

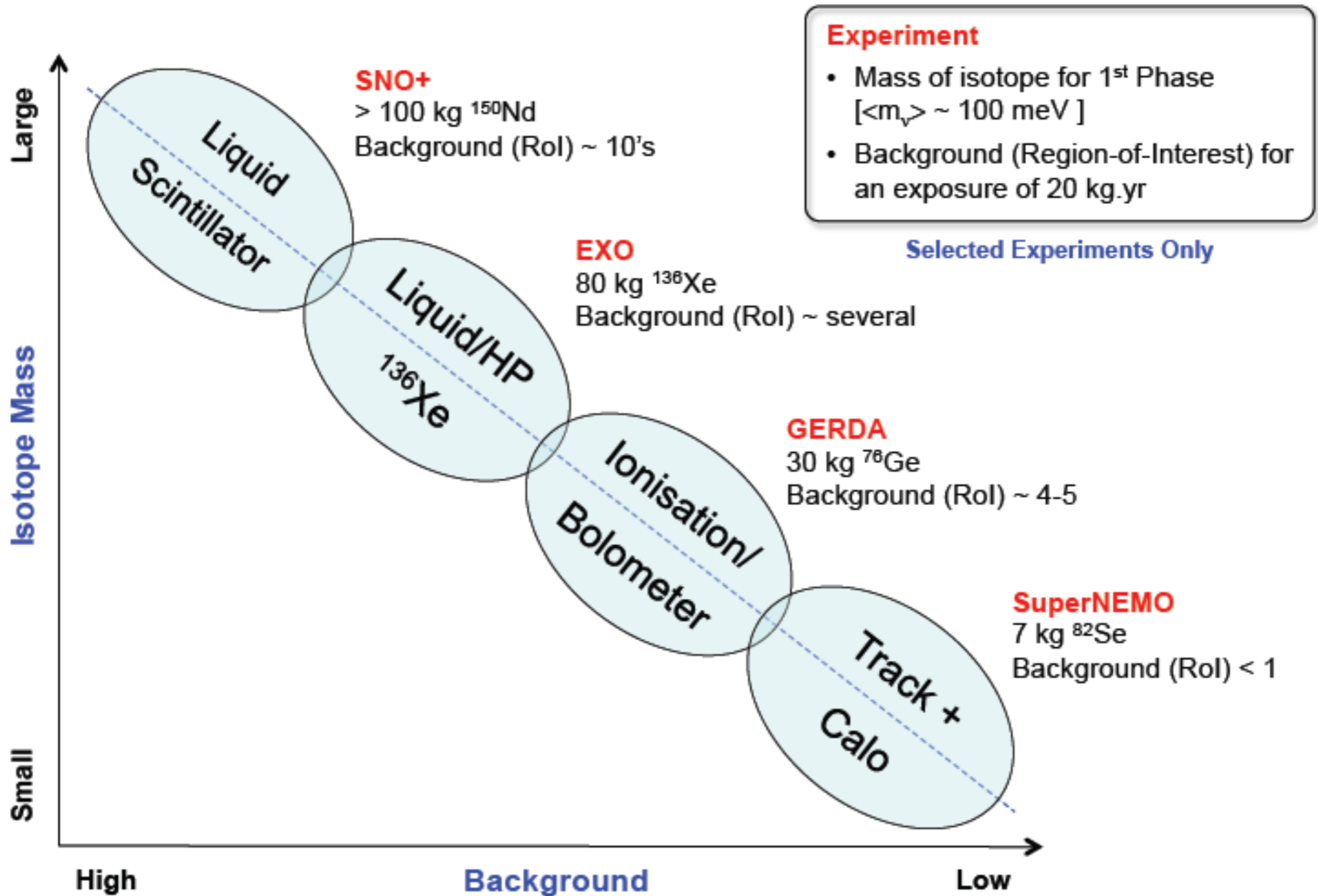
Required  $T_{1/2}^{0\nu}$  sensibility for  $\langle m_{ee} \rangle \sim 0.05$  eV (IH)



- Many approximation methods
- Variation up to factor  $\sim 10$
- Different among isotopes
- Up to factor 10 on required mass! (<sup>150</sup>Nd, <sup>100</sup>Mo w.r.t. <sup>76</sup>Ge)

Main **limitation** in interpreting result & comparing among different isotopes

# Comparison of Current Techniques



# SuperNEMO demonstrator: status

Most R&D completed

Radio purity measurements of materials ongoing

- HPGe, Radon emanation chamber, BiPo

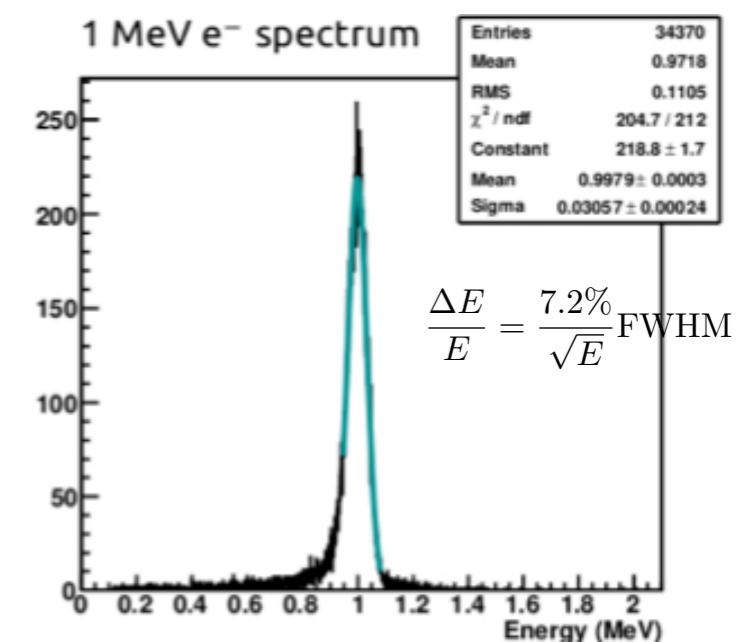
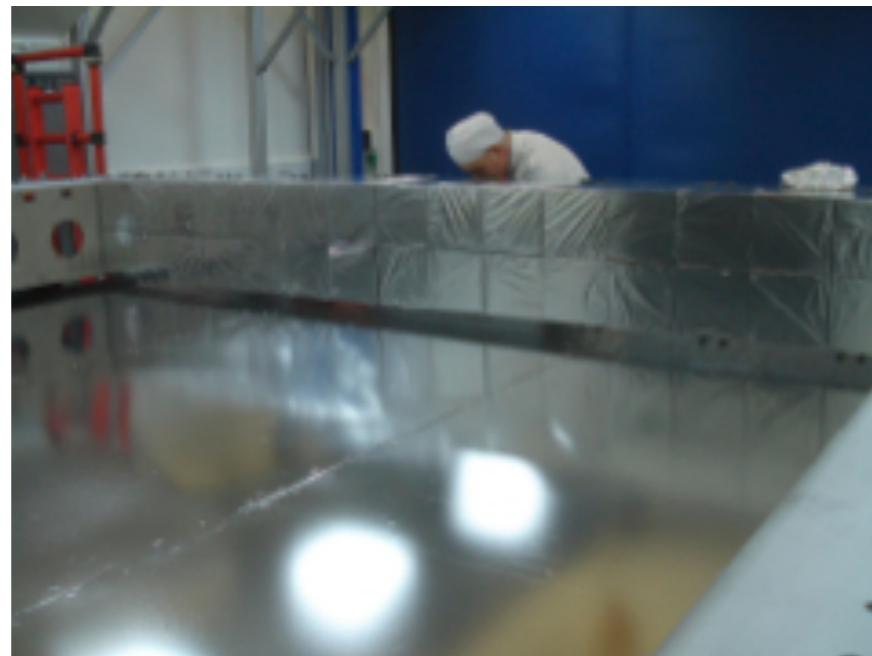
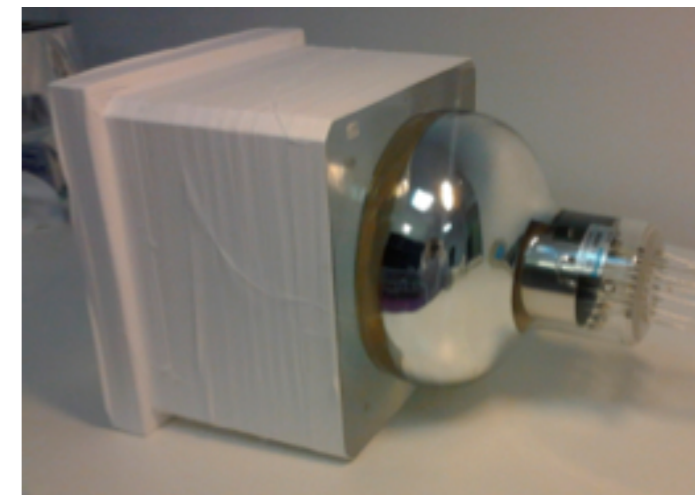
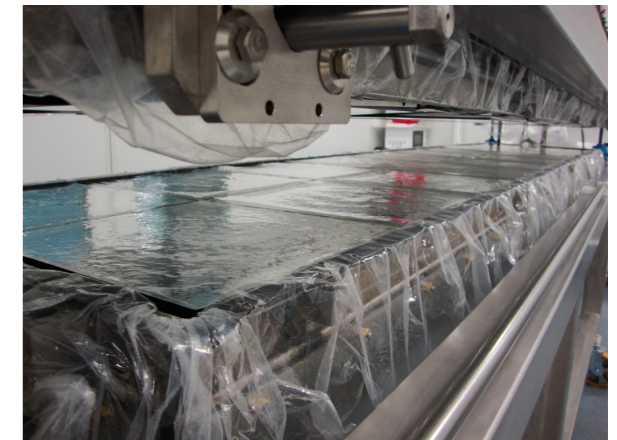
Calorimeter: main wall under construction

- Scintillator block under production
- 8" Hamamatsu PMT by February 2014
- Electronic (FE digitiser & trigger board) under production

Tracker: C0 under construction.

- Commissioning end 2013 (surface) 2014 (underground)

Source: more detail later...





# LSM Extension: DOMUS

From F. Piquemal (TAUP2013)

- Project accepted by Ministry of Research and programmed by CNRS
- Cost estimated by project supervisor of safety gallery :
- 7 M€ including 20% hazards
- Funding secured from CNRS, Rhône-Alpes region, Savoie department, FEDER
- Technical studies completed
- Negotiations with the civil work company in October
- Digging Spring 2014 or end 2015 depending of the company schedule
- 6 months for excavation, 10 months for outfittings
- Extension in operation 2016 -2017

