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# Exploring Dirac Gaugino GUTs

Mark D. Goodsell

LPTHE, Paris

See 1311.XXXX, 1211.0552 with K. Benakli and F. Staub; 1206.6697;



## Overview

- Motivation for Dirac gauginos
- Dirac gaugino models and the Higgs
- A new constrained GUT scenario



# Dirac gauginos

- In the MSSM have Majorana gauginos described by one Weyl fermion  $\lambda$  in adjoint rep of each gauge group, mass term  $\mathcal{L} \supset -\frac{1}{2}M_{\lambda}\lambda\lambda + h.c.$
- To make give a Dirac mass, add an extra adjoint fermion  $\boldsymbol{\chi}$  to give mass term

$$\mathcal{L} \supset -\mathfrak{m}_{\mathrm{D}}\chi\lambda + \mathrm{h.c.}$$

This also requires a scalar Σ by supersymmetry, fit in an
 adjoint chiral multiplet (Σ, χ).



# Motivation: bottom up

- Dirac gauginos allow to relax LHC search bounds as production of squarks is suppressed since no chirality flip is possible. Gluino production is enhanced a little relative to MSSM, but this is greatly suppressed when  $m_{\tilde{q}_{1,2}} \gg m_{\tilde{g}}$ .
- They typically suppress processes such as B → sγ and ○ΔF=2.
- They allow for increased naturalness: supersoft masses do not lead to large corrections to stop mass.
- They allow new Higgs couplings, permitting increased Higgs mass  $\rightarrow$  compatibility with e.g. light stops.
- There would have been/could still be clear signals from accompanying adjoint scalars if light (this would have been a surprise)
- If gauginos are found at the LHC, we will have to determine whether they are Majorana or Dirac in nature, and this is very difficult to do directly: maybe only possible at ILC
- Challenge is to study the possible spectra and Higgs properties under the spectra and Higgs properties

# Motivation: top down

Some attractive theoretical motivations:

- Nelson-Seiberg Theorem: existence of R symmetry (chiral symmetry under which bosons are also charged: Φ → e<sup>iαRΦ</sup>Φ, θ → e<sup>iαθ</sup>, W → e<sup>2iα</sup>W) required for F-term SUSY breaking
- Many SUSY models preserve R symmetry (e.g. original O'Raifertaigh model)
- Dirac gaugino mass may preserve R , Majorana does not: [Fayet, 78] suggested this as the original way to obtain gaugino masses!
- Alternatively Majorana gaugino mass may be too small (e.g. from many O'Raifeartaigh models [Komargodsky and Shih, 2008], [Abel, Jaeckel, Khoze 09])
- Adjoint multiplets appear in <u>many</u> UV models would be nice to use them rather than throwing them away!



## Status

Studying non-(N)MSSM SUSY models is typically hard due to lack of tools - and sometimes theory. However, now is the time to be doing this! On the theory side,

- Dirac gauginos usually considered in context in gauge mediation; have explored many possibilities.
- Increasing numbers of people interested in this class of models (too many to mention), e.g. effect of Seiberg dualities, lepton number as R-symmetry, detailed studies of naturalness, ...
- We now understand the technical aspects well: RGEs, how the masses are generated, etc.

On the tools side:

- We now have the tools for general theories: SARAH, PYR@TE, FeynRules, CalcHEP, MadGraph, MicrOmegas, ...
- Have been several studies (e.g. Martin and Kribs; Heikinheimo, Kellerstein, Sanz '11) of collider bounds for simplified models.
- Since 2012 SARAH has incorporated the possibility of Dirac gauginos.



# MSSM with Adjoints

Names		Spin 0	Spin 1/2	Spin 1	SU(3), SU(2), U(1) <sub>Y</sub>
Quarks	Q u <sup>c</sup>		(u <sub>L</sub> , d <sub>L</sub> ) u <sup>c</sup>		( <b>3</b> , <b>2</b> , 1/6) ( <b>3</b> , <b>1</b> , -2/3)
(×3 families)	dc	ũc ãc	սլ սլ		(3, 1, 1/3)
Leptons (×3 families)	L e <sup>c</sup>	(v <sub>eL</sub> ,ẽ <sub>L</sub> ) ẽ <sub>L</sub>	(v <sub>eL</sub> ,e <sub>L</sub> ) e <sup>c</sup> <sub>L</sub>		( <b>1</b> , <b>2</b> , -1/2) ( <b>1</b> , <b>1</b> , 1)
Higgs	H <sub>u</sub> H <sub>d</sub>	$(H_{u}^{+}, H_{u}^{0})$ $(H_{d}^{0}, H_{d}^{-})$	$\frac{(\tilde{H}_{d}^{\pm}, \tilde{H}_{u}^{0})}{(\tilde{H}_{d}^{0}, \tilde{H}_{d}^{-})}$ $\frac{\lambda_{3\alpha}}{\lambda_{3\alpha}}$		( <b>1</b> , <b>2</b> , 1/2) ( <b>1</b> , <b>2</b> , -1/2)
Gluons	W <sub>3α</sub>		$\begin{bmatrix} \tilde{\lambda}_{3\alpha} \\ [\equiv \tilde{g}_{\alpha}] \end{bmatrix}$	g	( <b>8</b> , <b>1</b> , 0)
W	$W_{2\alpha}$		$[\equiv \tilde{W}^{\pm}, \tilde{W}^{0}]$	<i>W</i> <sup>±</sup> , <i>W</i> <sup>0</sup>	( <b>1</b> , <b>3</b> , 0)
В	$W_{1\alpha}$		$\stackrel{\lambda_{1\alpha}}{[\equiv \tilde{B}]}$	В	(1, 1, 0)
DG-octet	Og	$\begin{bmatrix} O_g \\ [\equiv \Sigma_g] \end{bmatrix}$	$\begin{bmatrix} \chi_g \\ \equiv \tilde{g}' \end{bmatrix}$		( <b>8</b> , <b>1</b> , 0)
DG-triplet	Т	$ \begin{array}{c} \{T^0, T^{\pm}\} \\ [\equiv \{\Sigma_0^{\mathcal{W}}, \Sigma_{\mathcal{W}}^{\pm}\}] \end{array} $	$ \{ \chi^{0}_{T}, \chi^{\pm}_{T} \} \\ [\equiv \{ \tilde{W}'^{\pm}, \tilde{W}'^{0} \} ] $		( <b>1</b> , <b>3</b> , 0 )
DG-singlet	s	$[\equiv \Sigma_B]$	$\begin{bmatrix} \chi_{S} \\ \equiv \breve{B}' \end{bmatrix}$		(1, 1, 0)



# Supersymmetric Couplings

Here are the most general renormalisable superpotential couplings:

- SUSY couplings contained in superpotential:  $W = W_{Yukawa} + W_{Higgs} + W_{Adjoint}$
- No new Yukawas:

$$W_{\text{Yukawa}} = Y_{\text{U}}^{ij} \mathbf{Q}_i \cdot \mathbf{H}_u \mathbf{u}_j^c + Y_{\text{D}}^{ij} \mathbf{Q}_i \cdot \mathbf{H}_d \mathbf{d}_j^c + Y_{\text{E}}^{ij} \mathbf{L}_i \cdot \mathbf{H}_d \mathbf{e}_j^c$$

• Two new Higgs couplings (c.f. NMSSM):

$$W_{\text{Higgs}} = \mu \mathbf{H}_{u} \cdot \mathbf{H}_{d} + \lambda_{S} \mathbf{S} \mathbf{H}_{d} \cdot \mathbf{H}_{u} + 2\lambda_{T} \mathbf{H}_{d} \cdot \mathbf{T} \mathbf{H}_{u}$$

• Several possible new Adjoint couplings which violate R:

$$W_{Adjoint} = L\mathbf{S} + \frac{M_S}{2}\mathbf{S}^2 + \frac{\kappa_S}{3}\mathbf{S}^3 + M_T tr(\mathbf{TT}) + \lambda_{ST}\mathbf{S}tr(\mathbf{TT}) + M_O tr(\mathbf{OO}) + \lambda_{SO}\mathbf{S}tr(\mathbf{OO}) + \frac{\kappa_O}{3}tr(\mathbf{OOO}).$$



# D-term masses

• Write the Dirac mass as a <u>Holomorphic</u> term; gives new D-term interactions:

$$d^{2}\theta 2\sqrt{2}m_{D}\theta^{\alpha} tr(W_{\alpha}\Sigma) \supset -m_{D}(\lambda_{a}\chi_{a}) + \sqrt{2}m_{D}\Sigma_{a}D_{a}$$

The New D-term couplings have two main effects:

Adjoint scalar masses and B-type masses are modified:

$$\begin{split} \mathcal{L} \supset & \frac{1}{2} D_{\alpha}^{2} + \sqrt{2} (m_{D} \Sigma_{\alpha} + \bar{m}_{D} \overline{\Sigma}_{\alpha}) D_{\alpha} \\ \xrightarrow{m_{D} \text{ real}} & - \frac{1}{2} m_{D}^{2} (\Sigma_{\alpha} + \overline{\Sigma}_{\alpha})^{2} \end{split}$$

Trilinear terms modify Higgs mass matrix

$$\frac{1}{\sqrt{2}}g_{Y}\mathfrak{m}_{D}(S+\overline{S})(H_{\mathfrak{u}}^{*}H_{\mathfrak{u}}-H_{\mathfrak{d}}^{*}H_{\mathfrak{d}})\supset -g_{Y}\mathfrak{m}_{D}c_{2\beta}\nu(s_{R}\mathfrak{h})$$

•  $\rightarrow$  Bino mass is important for the Higgs mass, cannot be decoupled!



## **RGEs**

Since the operator is holomorphic, the RGEs are given by

$$\beta_{\mathfrak{m}_{D}^{iA}} = \gamma_{j}^{i}\mathfrak{m}_{D}^{jA} + \frac{\beta_{g}}{g}\mathfrak{m}_{D}^{iA}$$

We can apply the normal rules for the "standard" terms and then modify by

$$\begin{split} \Delta(\mathfrak{m}^2)^i_j =& 2\mathfrak{m}_D^{iA}\mathfrak{m}_{DjA} \equiv 2(\mathfrak{m}_D^2)^i_j \\ \Delta B^{ij} =& 2\mathfrak{m}_D^{iA}\mathfrak{m}_D^{jA} \equiv 2(\mathfrak{m}_D^2)^{ij} \end{split}$$

at the desired energy scale. However, there is an exception: the tadpole RGE! The Dirac mass term enters explicitly here, so it had to be computed from scratch.

$$\beta_{t^{\alpha}}^{(i)} \equiv X_{S}^{(i)} + X_{\xi}^{(i)} + X_{D}^{(i)}$$

$$\begin{split} (4\pi)^2 X_{\xi}^{(1)} = & 2\sqrt{2}g_Y m_D^{\alpha Y} tr( \mathfrak{Y} m^2) \\ (4\pi)^4 X_{\xi}^{(2)} = & 2\sqrt{2}g_Y m_D^{\alpha Y} tr( \mathfrak{Y} m^2 (4g^2 C_2 - Y_2)) \end{split}$$

and

$$(4\pi)^{2} X_{D}^{(1)} = 2 \left[ (m_{D}^{2})_{ef} (A^{aef} + MY^{aef}) + Y_{efk} \mu^{ka} (m_{D}^{2})^{ef} \right]$$
$$(4\pi)^{4} X_{D}^{(2)} = 4 (\beta_{m_{D}}^{(1)} / m_{D})_{g}^{f} \left[ (m_{D}^{2})_{ef} (A^{aeg} + MY^{aeg}) + Y_{efk} \mu^{ka} (m_{D}^{2})^{eg} \right]$$



# Naturalness

Bottom line:

- Tadpole term naturally generated by running, but <u>not</u> dangerous in size (in fact, it is useful phenomenologically): typically  $\sim g_Y m_D m^2$  (or smaller if tr( $\Im m^2$ ) = 0 and  $\mu = 0$ ).
- Dirac gaugino masses do not enter the Higgs or stop mass RGEs → increased naturalness: finite contribution to stop mass from gluino of

$$\delta m_{\tilde{t}}^2 = \frac{(m_{D3})^2 \alpha_s}{2\pi} \log \left(\frac{m_{O_P}}{m_{D3}}\right)^2$$

Singlet and triplet scalar masses drive EWSB, enter into naturalness bounds:

$$\delta \, \mathfrak{m}^2_{H_{\mathfrak{u},d}} \supset - \, \frac{1}{16\pi^2} (2\lambda_S^2 \, \mathfrak{m}^2_S + 6\lambda_T^2 \, \mathfrak{m}^2_T) \log\left(\frac{\Lambda}{\mathsf{TeV}}\right)$$

where  $\Lambda$  is the UV cutoff of the theory. Using  $\Delta\equiv \frac{\delta m_h^2}{m_h^2}$  then we have

$$m_{S} \lesssim \text{TeV}\left(\frac{1}{\lambda_{S}}\right) \left(\frac{\log \Lambda/\text{TeV}}{3}\right)^{-1/2} \left(\frac{\Delta^{-1}}{20\%}\right)^{-1/2} \quad m_{T} \lesssim 5 \text{ TeV}\left(\frac{0.1}{\lambda_{T}}\right) \left(\frac{\log \Lambda/\text{TeV}}{3}\right)^{-1/2}$$

(in the absence of a large tree-level contribution). This is good, because

$$\Delta \rho \simeq \frac{\nu^2 (g_2 m_{D2} c_{2\beta} + \sqrt{2} \tilde{\mu} \lambda_T)^2}{(m_T^2 + |M_T|^2 + B_T + 4|m_{D2}|^2)^2} \lesssim 8 \times 10^{-4} \rightarrow m_T \gtrsim 1.4 \text{ TeV}$$

NB for light stops  $\Delta \rho^{\text{stops}} \simeq 4 \times 10^{-4} \left( \frac{500 \text{ GeV}}{m_{\tilde{t}_1}} \right)^2$ .

# Getting 126 GeV

In limit of large m<sub>S</sub>, m<sub>T</sub>, can integrate out adjoint scalars to obtain

$$\begin{split} m_h^2 \simeq & M_Z^2 c_{2\beta}^2 + \frac{\nu^2}{2} (\lambda_S^2 + \lambda_T^2) s_{2\beta}^2 + \frac{3}{2\pi^2} \frac{m_t^4}{\nu^2} \bigg[ \log \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} + \frac{\mu^2 \cot^2 \beta}{m_{\tilde{t}_1} m_{\tilde{t}_2}} \bigg( 1 - \frac{\mu^2 \cot^2 \beta}{12m_{\tilde{t}_1} m_{\tilde{t}_2}} \bigg) \bigg| \\ & + \nu^2 \bigg[ \lambda_1 c_{\beta}^4 + \lambda_2 s_{\beta}^4 + 2(\lambda_3 + \lambda_4 + \lambda_5) c_{\beta}^2 s_{\beta}^2 + 4(\lambda_6 c_{\beta}^2 + \lambda_7 s_{\beta}^2) s_{\beta} c_{\beta} \bigg] \\ & \xrightarrow{\tan \beta \to \infty} M_Z^2 + \lambda_2 \nu^2 + \frac{3}{2\pi^2} \frac{m_t^4}{\nu^2} \log \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \end{split}$$

Can enhance the Higgs mass naturally!

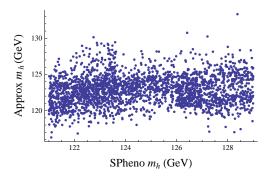
- At small tan  $\beta$ , do not need heavy stops or large stop mixing etc: for large  $\lambda_S$  or  $\lambda_T$  we can take just the tree-level part:  $m_h^2 \simeq M_Z^2 c_{2\beta}^2 + \frac{\nu^2}{2} (\lambda_S^2 + \lambda_T^2) s_{2\beta}^2 \rightarrow \lambda_S \sim 0.7$  to obtain correct Higgs mass as at small tan  $\beta$  as in NMSSM/ $\lambda$ SUSY.
- Also the origin of the potential may be a <u>maximum</u> rather than saddlepoint as in MSSM
- For large tan  $\beta$ , scalar and triplet scalars can do the same job if they are heavy (e.g. for  $\lambda_S = 1.8$  or  $\lambda_T = 1.2$  with no stop contribution)

$$\begin{split} 32\pi^2\lambda_2 \supset & 2\lambda_S^4\log\frac{m_S^2}{\nu^2} + (g_2^4 - 4g_2^2\lambda_T^2 + 10\lambda_T^4)\log\frac{m_T^2}{\nu^2} \\ & + \frac{4\lambda_S^2\lambda_T^2}{m_S^2 - m_T^2} \bigg[m_S^2\log\frac{m_S^2}{\nu^2} - m_T^2\log\frac{m_T^2}{\nu^2} - (m_S^2 - m_T^2)\bigg] \end{split}$$



### Numerical comparison

Comparison of effective potential with a SARAH/SPheno scan over models with heavy singlet and triplet scalars:



Can have light squarks and correct Higgs mass.

$$\label{eq:barrent} \begin{split} &\tan\beta=50,\,m_{D\,2}=600~\text{GeV}, \text{first two generation sfermion mass squareds of } 4\times10^7~(\text{GeV})^2, \text{third generation} \\ &\text{sfermion mass squareds } 4\times10^6~(\text{GeV})^2 \text{ and scanning over} \end{split}$$

$$\begin{split} \mathfrak{m}_{D\,1} &\in [-600, 600] \text{ GeV}, \mu \in [-750, 750] \text{ GeV}, B\,\mu \in [5000, 10^6] \; (\text{GeV})^2, \lambda_T \in [0,1] \text{ while adjusting } \lambda_S \text{ to keep} \\ \mathfrak{m}_h &= 125 \pm 4 \; \text{GeV}. \text{ The expectation values } \nu_S, \nu_T \text{ were set by the tree level minimisation equation with} \\ \mathfrak{m}_S^2, \mathfrak{m}_T^2 &= 2.5 \times 10^7 \; (\text{GeV})^2 \end{split}$$



UPMC

## Unification

• MSSM one-loop beta-function coefficients are  $(b_3, b_2, b_1 = (5/3)b_Y) = (3, -1, -11)$ , lead to unification of couplings at  $10^{16}$  GeV with perturbative couplings  $\alpha_{GUT} \sim 1/24$ .

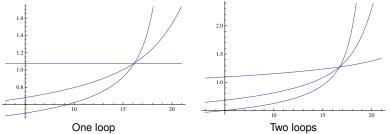
$$\frac{1}{g_{\mathfrak{i}}^2(\mu)} = \frac{1}{g_{\mathfrak{i}}^2(M_{SUSY})} + \frac{b_{\mathfrak{i}}}{8\pi^2} \log \mu / M_{SUSY}$$

- Triumph of the MSSM (modulo two-loop discrepancy...) that we might like to preserve!!
- Adding adjoint fields does (except for S, a singlet): T decreases  $b_2$  by 2,  $\mathbf{0}_g$  decreases  $b_3$  by 3

Our choice: add

$$(\mathbf{1}, \mathbf{2})_{1/2} + (\mathbf{1}, \mathbf{2})_{-1/2} + 2 \times (\mathbf{1}, \mathbf{1})_{\pm 1}$$

This could come from  $(SU(3))^3$  (would need also four SM singlets).



# Toward a GUT scenario

• This configuration is particularly interesting. Let us add to the minimal Dirac gaugino MSSM a pair of doublets  $R_u$ ,  $R_d$  and two non-chiral selectrons  $\hat{E}_i$ ,  $\hat{\tilde{E}}_i$ , i = 1, 2. The most general higgs potential is then

$$\begin{split} & \mathcal{W} \supset (\mu + \lambda_S S) H_d H_u + 2\lambda_T H_d T H_u \\ & + (\mu_R + \lambda_{SR} S) R_u R_d + 2\lambda_{TR} R_u T R_d + \mu_{\hat{E} \, ij} \hat{E}_i \hat{\bar{E}}_j \\ & + (\mu_u + \lambda_{Su} S) R_u H_u + 2\lambda_{Tu} R_u T H_u + (\mu_d + \lambda_{Sd} S) R_d H_d + 2\lambda_{Td} R_d T H_d \\ & + Y_{\hat{E} i} R_u H_d \hat{E}_i + Y_{\hat{E} i} R_d H_u \hat{\bar{E}}_i \end{split}$$

We can now take one of two directions:

- An extended MRSSM  $\rightarrow$  removing  $\mu$ ,  $\mu_R$ ,  $\lambda_S$ ,  $\lambda_T$  and related couplings, where an R-symmetry is preserved by the Higgs sector.
- Charge the new fields under lepton number, so that we have new heavy vector-like leptons and sleptons. The superpotential becomes

$$\begin{split} W \supset & (\mu + \lambda_{S}S)H_{d}H_{u} + 2\lambda_{T}H_{d}TH_{u} \\ & + (\mu_{R} + \lambda_{SR}S)R_{u}R_{d} + 2\lambda_{TR}R_{u}TR_{d} + (\mu_{\hat{E}\,ij} + \lambda_{SEij}S)\hat{E}_{i}\hat{\tilde{E}}_{j} \\ & + Y_{\hat{E}i}R_{u}H_{d}\hat{E}_{i} + Y_{\hat{E}i}R_{d}H_{u}\hat{\tilde{E}}_{i} \\ & + Y_{LFV}^{ij}L_{i} \cdot H_{d}\hat{E}_{j} + Y_{EFV}^{j}R_{u}H_{d}E_{j} \end{split}$$



# Introducing the CMDGSSM

We can now specify a minimal set of boundary conditions at the GUT scale:

- As in the CMSSM/mSUGRA, we have  $m_0$ , tan  $\beta$  but instead of  $m_{1/2}$  we have  $m_D$ . We set  $A_0 = 0$  due to SUSY preserving R-symmetry.
- We also choose to take non-universal Higgs masses, and so specify μ, B<sub>μ</sub>.
- Since we have two new tadpole conditions from v<sub>S</sub>, v<sub>T</sub> we specify m<sub>S0</sub> (singlet scalar mass) and m<sub>T0</sub> (triplet scalar mass) at the GUT scale. We set the octet scalar mass equal to the triplets, and take B<sub>T</sub> = B<sub>S</sub> = B<sub>O</sub> = 0 for minimality.
- We have the Yukawa couplings Y<sub>Êi</sub>, Y<sub>Êi</sub>, Y<sup>ij</sup><sub>Ei</sub>, Y<sup>ij</sup><sub>EFV</sub>, Y<sup>j</sup><sub>EFV</sub> which are equivalent to lepton Yukawas; they are constrained to be ≤ 0.01 and so irrelevant for spectrum-generator purposes.
- We have a choice of μ<sub>R</sub>, μ<sub>E</sub> → can either adjust for precision gauge unification; set to be equal to the Higgs mu; set at convenient values. The Higgs mass and coloured sparticle spectrum is largely independent of this choice.
- We have a choice of couplings  $\lambda_S$ ,  $\lambda_T$ ,  $\lambda_{SR}$ ,  $\lambda_{TR}$ ,  $\lambda_{SEij}$ : can take N = 2 values, or  $(SU(3))^3$  values, or choose freely.



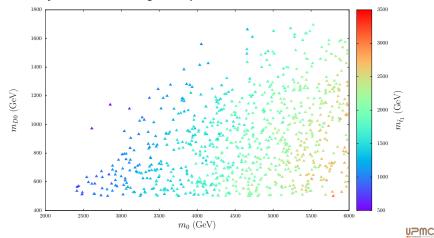
# First forays

- We can now start to explore the parameter space using SPheno code produced by SARAH for this model, which calculates two-loop RGEs and one-loop pole masses.
- One important technical limitation is due to the Higgs mass: if we enhance it using heavy stops, then the accuracy of the spectrum generator is no longer trustworthy.
- We instead choose to explore the corner of parameters space with  $\lambda_S \sim 0.7$ , small tan  $\beta$  so that no sparticle contributing significantly to the Higgs mass is heavier than about 2 TeV.
- For convenience we take  $\lambda_T \sim 0$  and set  $\mu_R \sim \mu_E \sim \text{TeV}$ , scan over  $\lambda_S$ , tan  $\beta$  within a narrow range and otherwise scan randomly over  $\mu$ ,  $B\mu$ ,  $m_0$ ,  $m_D$ ,  $m_{S0}$ ,  $m_{T0}$ .
- We keep only points with the correct Higgs mass satisfying the constraints from HiggsBounds.



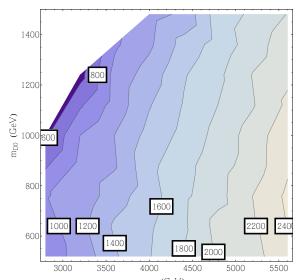
### $m_0 - m_D$ plane and stop masses

We find no models in the with small  $m_0$  and large  $m_D$  since the bino mass is important in the Higgs mass calculation. We do find many models with light stops:



### $m_0 - m_D$ plane II

Same plot as on previous slide, clearly showing contours of stop mass:

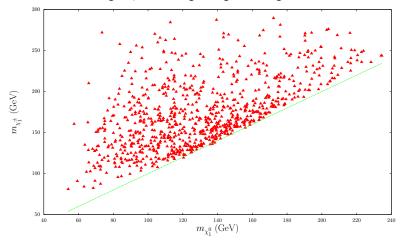




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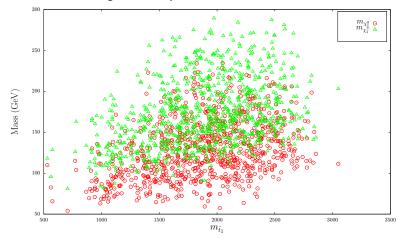
## Charginos and neutralinos

The neutralinos are typically light due to the restriction on the bino mass from the Higgs mass. Also the scans preferentially find models with light  $\mu$ , leading to light charginos.



### Charginos, neutralinos and stops

Here we show the correlation of the neutralino and chargino masses with the lightest stop mass:





# Predictions

- Unification takes place at  $(1.8\pm0.4)\times10^{17}~\text{GeV}$
- We have a compressed pattern of soft masses (with deviations of a few percent):

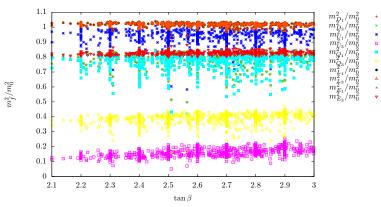
$$\begin{split} & m^2_{U33}:m^2_{Q33}:m^2_{Q11}:m^2_{D\,i\,i}:m^2_{E\,i\,i}:m^2_{U11}:m^2_{L\,i\,i}\\ = & 0.16:0.39:0.77:0.79:0.83:0.93:1.02 \end{split}$$

- Hence sleptons are heavy and quasi-degenerate with the first two generations of squarks. This is because the Dirac gaugino masses do not enter into the squark RGEs.
- While the lightest stop masses are  $1.9\pm0.5, 2.9\pm0.6$  TeV.
- The gaugino masses are in the ratio 0.22 : 0.9 : 3.5, i.e. the Wino barely runs from  $m_D$  (as can be seen from the one-loop RGE, which is zero for small  $\lambda_T$ ).



### Squark masses

Over the range of  $\tan \beta$  scanned, the squark masses vary little:





### Conclusions

- Dirac gauginos have many attractive phenomenological and theoretical advantages over their Majorana counterparts, and can arise naturally in many different contexts (strong dynamics, higher dimensions, string theory, ...)
- There now exists a tool (SARAH with SPheno) to seriously study many aspects of their phenomenology which can interface with other tools.
- We have embedded DG models into a unified scenario and conducted a first study.
- Long program of research into these and other beyond-MSSM theories



# Future Possibilities

Many possible avenues for future work:

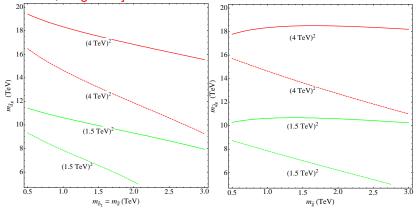
- Modifications of Higgs sector
- Calculation of two-loop effects and implementation in codes
- Connection with collider limits
- Models to realise messenger mass patterns
- Explicit D-term SUSY sectors (e.g. 4 1 model)
- Warped models
- Gauge messengers
- Gravity mediation, embedding in string models, Dirac gravitinos,....



### Flavour constraints

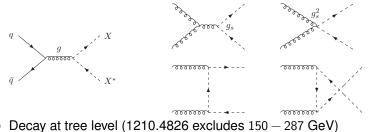
*Work with E. Dudas, L. Heurtier, P. Tziveloglou* Dirac gauginos typically suppress flavour-chaging neutral currents. As an example, consider the model of Dudas, von Gersdorff,

Pokorski, Ziegler 13]:



# Collider Signatures

- May be difficult to distinguish directly Dirac and Majorana gauginos except at e<sup>-</sup>e<sup>-</sup> collider
- Indirectly we would obtain spectacular signals from the adjoint scalars



$$\begin{array}{l} X \rightarrow \hspace{-0.5mm} \tilde{g} \hspace{-0.5mm} \tilde{g} \rightarrow q q \tilde{q} \hspace{-0.5mm} \tilde{q} \rightarrow q q q q + \hspace{-0.5mm} \tilde{\chi} \hspace{-0.5mm} \tilde{\chi} \\ X \rightarrow \hspace{-0.5mm} \tilde{q} \hspace{-0.5mm} \tilde{q} \rightarrow q q + \hspace{-0.5mm} \tilde{\chi} \hspace{-0.5mm} \tilde{\chi} \end{array}$$

and (one loop):

$$X \to t\bar{t}$$



### Higgs mass matrix

- Recall <u>holomorphic</u> coupling  $\int d^2\theta 2\sqrt{2}m_D\theta^\alpha tr(W_\alpha\Sigma)\supset -m_D(\lambda_\alpha\chi_\alpha)+\sqrt{2}m_D\Sigma_\alpha D_\alpha \text{ gives new } D\text{-term interactions}$
- This translates into a shift in D-term Higgs potential, masses and also the sfermion masses! Higgs mass matrix in the basis { $h, H, S_R, T_R^0$ } is

$$\begin{pmatrix} M_{Z}^{2} + \Delta_{h} s_{2\beta}^{2} & \Delta_{h} s_{2\beta} c_{2\beta} & \Delta_{hs} & \Delta_{ht} \\ \Delta_{h} s_{2\beta} c_{2\beta} & M_{A}^{2} - \Delta_{h} s_{2\beta}^{2} & \Delta_{Hs} & \Delta_{Ht} \\ \Delta_{hs} & \Delta_{Hs} & \tilde{m}_{S}^{2} & \lambda_{S} \lambda_{T} \frac{\nu^{2}}{2} \\ \Delta_{ht} & \Delta_{Ht} & \lambda_{S} \lambda_{T} \frac{\nu^{2}}{2} & \tilde{m}_{T}^{2} \end{pmatrix}$$

where

$$\Delta_{h} = \frac{v^{2}}{2} (\lambda_{S}^{2} + \lambda_{T}^{2}) - M_{Z}^{2}$$
$$\Delta_{hs} = v [v_{S} \lambda_{S}^{2} - g' m_{D1} c_{2\beta}]$$
$$\Delta_{Hs} = g' m_{1D} v_{S2\beta}$$

Higgs mass is not independent of bino mass! Although triplet scalar is heavy  $\rightarrow$  roughly independent of Wino mass.



# **Scenarios**

- Recall that in the MSSM we have two non-trivial tadpole conditions (from  $\partial_{H^0_u} V = \partial_{H^0_d} V = 0$ )
- In the NMSSM we have three (including the singlet vev  $\partial_S V = 0$ )
- Here we have four, from including the triplet scalar vev ( $\partial_T V = 0$ )
- $\rightarrow$  these should be traded with four soft parameters.

Can have several different scenarios, e.g. :

- 1. <u>MSSM in disguise</u>: here we shall allow a  $\mu$ -term, and assume that the only source of R-symmetry violation arises in the supersymmetry-breaking sector, but permit only a B $_{\mu}$  term.
- 2. <u>MSSM without  $\mu$  term</u>: this is the scenario of [Nelson, Rius, Sanz, Unsal], taking  $\mu = 0$ , essentially now ruled out.
- 3. Dynamical  $\mu$  models: take  $\mu = 0$  but allow substantial non-zero expectation value for the singlet, possibly via non-zero  $t_s$ .
- 4. <u>Dynamical  $\mu$  and  $B\mu$  models</u>: we allow a non-zero  $W \supset \frac{1}{3}\kappa S^3$ , breaking R-symmetry in the visible sector, but allowing  $\mu$  and  $B_{\mu}$  to be generated via a non-zero singlet vev as in the NMSSM.

Will mostly consider (1) and (3) scenarios in the following, leaving plenty of scope for future work.



# Unification

• MSSM one-loop beta-function coefficients are  $(b_3, b_2, b_1 = (5/3)b_Y) = (3, -1, -11)$ , lead to unification of couplings at  $10^{16}$  GeV with perturbative couplings  $\alpha_{GUT} \sim 1/24$ .

$$\frac{1}{g_{\mathfrak{i}}^2(\mu)} = \frac{1}{g_{\mathfrak{i}}^2(M_{SUSY})} + \frac{b_{\mathfrak{i}}}{8\pi^2}\log\mu/M_{SUSY}$$

- Triumph of the MSSM (modulo two-loop discrepancy...) that we might like to preserve!!
- Adding complete GUT multiplets (as in gauge mediation) does not alter this (beta-function coefficients decreased by (1,1,1) per pair of SU(5) messengers).
- Adding adjoint fields does (except for S, a singlet): T decreases  $b_2$  by 2,  $\mathbf{0}_g$  decreases  $b_3$  by 3

Four alternatives

- 1. Abandon matter and gauge unification
- 2. Modify our definition of "unification" ...
- 3. Add extra "bachelor" states to make up complete GUT adjoint multiplets [Fox, Nelson and Weiner, 02], allows matter and gauge unification
- 4. Add minimal extra states to restore gauge unification



## Messengers to the Rescue

- Gauge mediation requires messenger fields these could also restore gauge unification!
- Require at least 2 pairs of messengers in (anti) fundamental of SU(2) and SU(3) for adjoint scalar masses (see later)
- Easy to find sets of messengers that satisfy this, e.g.

$$\begin{array}{lll} 4\times [(1,1)_1+(1,1)_{-1}] & \mbox{at} & \mbox{m}_1=3\,10^{12}\mbox{GeV} \\ 4\times [(1,2)_{1/2}+(1,\overline{2})_{-1/2}] & \mbox{at} & \mbox{m}_2=1.3\,10^{13}\mbox{GeV} \\ 2\times [(3,1)_{1/3}+(\overline{3},1)_{-1/3}] & \mbox{at} & \mbox{m}_3=10^{13}\mbox{GeV} \\ M_U\sim 9.9\cdot 10^{17}\mbox{GeV} & \mbox{$\alpha_{11}^{-1}\sim 4.77$} \end{array}$$

High messenger scale required to allow perturbativity up to GUT scale

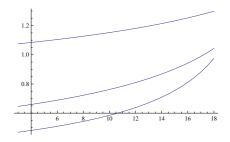


# **F-theory Unification**

- If we modify our definition of unification, in the heterotic string, we could "unify" the hypercharge at a different Kac-Moody level.
- Alternatively, we can modify our definition of unification to the F-theory criterion

$$5\alpha_1^{-1} - 3\alpha_2^{-1} - 2\alpha_3^{-1} = 0$$

at a GUT scale of  $\sim 1.7 \times 10^{17}$  GeV. This entails simply adding a vector-like pair of electron fields  $(1,1)_{\pm 1}$  to the model, following [Davies, 2012]

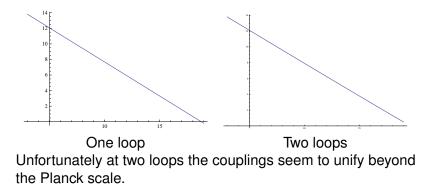




Conclusions

Extras

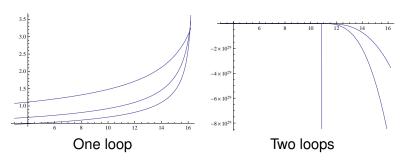
## F-theory Unification II





# Other unification schemes

Can consider other schemes for unification with gravity mediation. E.g. Adding "bachelor" states at low scale



$$\mathbf{24} 
ightarrow \mathbf{8_0} + \mathbf{3_0} + \mathbf{1_0} + (\mathbf{3,2})_{-5/6} + (\mathbf{ar{3},2})_{5/6}$$

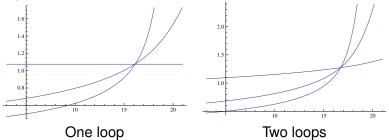


## Other unification schemes II

Or just add

$$(\mathbf{1},\mathbf{2})_{1/2} + (\mathbf{1},\mathbf{2})_{-1/2} + 2 \times (\mathbf{1},\mathbf{1})_{\pm 1}$$

This could come from  $(SU(3))^3$  (would need also four SM singlets).





# RGEs cont'd

Bottom line:

- Tadpole term naturally generated by running, but not dangerous in size (in fact, it is useful phenomenologically): typically  $\sim g_Y m_D m^2$  (or smaller if tr( $\Im m^2$ ) = 0 and  $\mu = 0$ ).
- Dirac gaugino masses do not enter the Higgs or stop mass RGEs  $\to$  increased naturalness: finite contribution to stop mass from gluino of

$$\delta m_{\tilde{t}}^2 = \frac{(m_{D3})^2 \alpha_s}{2\pi} \log \left(\frac{m_{O_P}}{m_{D3}}\right)^2$$

However, singlet and triplet scalar masses enter and help drive EWSB:

$$\begin{split} 16\pi^2 \frac{d}{dt} m_{H_u}^2 = & 6|y_t|^2 [m_{Q_3}^2 + m_{U_3}^2 + m_{H_u}^2] \\ & + 2\lambda_S^2 [m_{H_u}^2 + m_S^2 + m_{H_d}^2] + 6\lambda_T^2 [m_{H_u}^2 + m_T^2 + m_{H_d}^2] \\ & + g_Y^2 \text{Tr}(\text{Ym}^2) \\ 16\pi^2 \frac{d}{dt} m_{H_d}^2 = & 6|y_b|^2 [m_{Q_3}^2 + m_{D_3}^2 + m_{H_d}^2] \\ & + 2|y_\tau|^2 [m_{L_3}^2 + m_{E_3}^2 + m_{H_d}^2] \\ & + 2\lambda_S^2 [m_{H_u}^2 + m_S^2 + m_{H_d}^2] + 6\lambda_T^2 [m_{H_u}^2 + m_T^2 + m_{H_d}^2] \\ & - g_Y^2 \text{Tr}(\text{Ym}^2) \\ 16\pi^2 \frac{d}{dt} B_\mu = B_\mu [3|y_t|^2 + 3|y_b|^2 + |y_\tau|^2 - 3g_2^2 - y_Y^2] + 2B_\mu \lambda_S^2 + 6B_\mu \lambda_T^2 \end{split}$$



# **Naturalness**

Scalar masses thus enter into naturalness bounds:

$$\delta \, \mathfrak{m}^2_{H_{u,d}} \supset - \, \frac{1}{16 \pi^2} (2 \lambda_S^2 \, \mathfrak{m}^2_S + 6 \lambda_T^2 \, \mathfrak{m}^2_T) \log \left( \frac{\Lambda}{\text{TeV}} \right)$$

where  $\Lambda$  is the UV cutoff of the theory. Using  $\Delta\equiv \frac{\delta m_h^2}{m_h^2}$  then we have

$$\begin{split} m_S \lesssim & \mathsf{TeV}\left(\frac{1}{\lambda_S}\right) \left(\frac{\log \Lambda/\mathsf{TeV}}{3}\right)^{-1/2} \left(\frac{\Delta^{-1}}{20\%}\right)^{-1/2} \\ m_T \lesssim & \mathsf{TeV}\left(\frac{0.1}{\lambda_T}\right) \left(\frac{\log \Lambda/\mathsf{TeV}}{3}\right)^{-1/2} \left(\frac{\Delta^{-1}}{20\%}\right)^{-1/2} \end{split}$$

(in the absence of a large tree-level contribution). This is good, because

$$\Delta \rho \simeq \frac{\nu^2 (g_2 m_{D2} c_{2\beta} + \sqrt{2} \tilde{\mu} \lambda_T)^2}{(m_T^2 + |M_T|^2 + B_T + 4|m_{D2}|^2)^2} \lesssim 8 \times 10^{-4} \rightarrow m_T \gtrsim 1.4 \text{ TeV}$$

NB for light stops

$$\Delta \rho^{stops} \simeq \! 4 \times 10^{-4} \left(\frac{500 \; \text{GeV}}{\mathfrak{m}_{\tilde{t}_1}}\right)^2$$

