

# Monotop production in R-Parity Violating Supersymmetry

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# Outline

- 1 Motivations
- 2 Framework
- 3 Monotop signature
- 4 Results
- 5 Conclusions & Perspectives

# Motivations & Status of SUSY searches

## ▶ SUSY most appealing features:

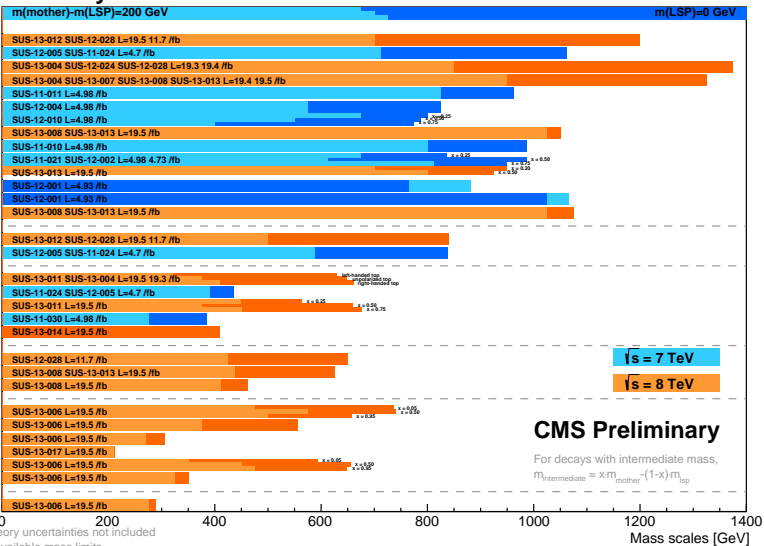
- **Stabilization** of the **hierarchy** between the **Planck** and the **Electroweak scale**
- **Gauge coupling unification**
- **Possible candidate** for **dark matter** with the **LSP**

## ▶ After 3 years of LHC running...

- **No experimental evidences** for **SUSY particles**
- **Mass exclusion limits** for **SUSY particles** have been **pushed higher and higher in energy**
- The **Higgs boson** is looking **more and more SM-like**
- **Significant portions** of the **parameter space** of **constrained models** like the **CMSSM** have been **excluded**

# Summary of CMS SUSY Results\* in SMS framework

SUSY 2013



\*Observed limits, theory uncertainties not included  
Only a selection of available mass limits  
Probe \*up to\* the quoted mass limit

# ATLAS SUSY Searches\* - 95% CL Lower Limits

Status: SUSY 2013

ATLAS Preliminary

$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$   $\sqrt{s} = 7, 8 \text{ TeV}$

	Model	$e, \mu, \tau, \gamma$	Jets	$E_{\text{miss}}^{\text{Tr}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit		Reference
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	$\tilde{g}$	1.7 TeV	$m(\tilde{g})=m(\tilde{g})$
	MSUGRA/CMSSM	1 e, $\mu$	3-6 jets	Yes	20.3	$\tilde{g}$	1.2 TeV	any $m(\tilde{g})$
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	$\tilde{g}$	1.1 TeV	any $m(\tilde{g})$
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}\tilde{g}$	0	2-6 jets	Yes	20.3	$\tilde{q}$	740 GeV	$m(\tilde{t}_1^0)=0 \text{ GeV}$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{g}\tilde{g}$	0	2-6 jets	Yes	20.3	$\tilde{g}$	1.3 TeV	$m(\tilde{t}_1^0)=0 \text{ GeV}$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{g}\tilde{g}(\ell\ell/\nu\nu)_{\text{I}}$	1 e, $\mu$	3-6 jets	Yes	20.3	$\tilde{g}$	1.18 TeV	$m(\tilde{t}_1^0)=200 \text{ GeV}, m(\tilde{\tau}^0)=0.5(m(\tilde{t}_1^0)+m(\tilde{g}))$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{g}\tilde{g}(\ell\ell/\nu\nu)_{\text{II}}$	2 e, $\mu$	0-3 jets	-	20.3	$\tilde{g}$	1.12 TeV	$m(\tilde{t}_1^0)=0 \text{ GeV}$
	GMSB ( $\tilde{f}$ NLSP)	2 e, $\mu$	2-4 jets	Yes	4.7	$\tilde{g}$	1.24 TeV	$\tan\beta < 15$
	GMSB ( $\tilde{f}$ NLSP)	1- $\tau$	0-2 jets	Yes	20.7	$\tilde{g}$	1.4 TeV	$\tan\beta > 18$
	GGM (bino NLSP)	2 $\gamma$	-	Yes	4.8	$\tilde{g}$	1.07 TeV	$m(\tilde{t}_1^0) > 50 \text{ GeV}$
	GGM (wino NLSP)	1 e, $\mu$ + $\gamma$	-	Yes	4.8	$\tilde{g}$	619 GeV	$m(\tilde{t}_1^0) > 50 \text{ GeV}$
	GGM (higgsino-bino NLSP)	0	1-5	Yes	4.8	$\tilde{g}$	800 GeV	$m(\tilde{t}_1^0) > 220 \text{ GeV}$
	GGM (higgsino NLSP)	2 e, $\mu$ (Z)	0-3 jets	Yes	5.8	$\tilde{g}$	890 GeV	$m(\tilde{t}_1^0) > 200 \text{ GeV}$
	Gravitino LSP	0	mono-jet	Yes	10.5	$\tilde{g}$	645 GeV	$m(\tilde{g}) > 10^{-4} \text{ eV}$
	3 <sup>rd</sup> gen. g med.	$\tilde{g} \rightarrow b\tilde{b}_{1,2}^0$	0	3 b	Yes	20.1	$\tilde{g}$	1.2 TeV
$\tilde{g} \rightarrow t\tilde{t}_1^0$		0	7-10 jets	Yes	20.3	$\tilde{g}$	1.1 TeV	$m(\tilde{t}_1^0) > 350 \text{ GeV}$
$\tilde{g} \rightarrow t\tilde{t}_2^0$		0-1 e, $\mu$	3 b	Yes	20.1	$\tilde{g}$	1.34 TeV	$m(\tilde{t}_1^0) > 400 \text{ GeV}$
$\tilde{g} \rightarrow b\tilde{b}_{1,2}^0$		0-1 e, $\mu$	3 b	Yes	20.1	$\tilde{g}$	1.3 TeV	$m(\tilde{t}_1^0) > 300 \text{ GeV}$
3 <sup>rd</sup> gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{t}_1^0$	0	2 b	Yes	20.1	$\tilde{b}_1$	100-620 GeV	$m(\tilde{t}_1^0) > 90 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{b}_1^0$	2 e, $\mu$ (SS)	0-3 b	Yes	20.7	$\tilde{b}_1$	275-430 GeV	$m(\tilde{t}_1^0) > 2 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1$ (light), $\tilde{b}_1 \rightarrow b\tilde{t}_1^0$	1-2 e, $\mu$	1-2 b	Yes	4.7	$\tilde{b}_1$	110-167 GeV	$m(\tilde{t}_1^0) > 55 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1$ (light), $\tilde{b}_1 \rightarrow W\tilde{b}_1^0$	2 e, $\mu$	0-2 jets	Yes	20.3	$\tilde{b}_1$	130-220 GeV	$m(\tilde{t}_1^0) = m(\tilde{b}_1), m(W) = 50 \text{ GeV}, m(\tilde{t}_1^0) < m(\tilde{\tau}_1^0)$
	$\tilde{b}_1\tilde{b}_1$ (medium), $\tilde{b}_1 \rightarrow t\tilde{t}_1^0$	2 e, $\mu$	2 jets	Yes	20.3	$\tilde{b}_1$	225-525 GeV	$m(\tilde{t}_1^0) > 0 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1$ (medium), $\tilde{b}_1 \rightarrow b\tilde{t}_1^0$	0	2 b	Yes	20.1	$\tilde{b}_1$	150-580 GeV	$m(\tilde{t}_1^0) > 200 \text{ GeV}, m(\tilde{\tau}_1^0) = 5 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1$ (heavy), $\tilde{b}_1 \rightarrow t\tilde{t}_1^0$	1 e, $\mu$	1 b	Yes	20.7	$\tilde{b}_1$	200-610 GeV	$m(\tilde{t}_1^0) > 0 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1$ (heavy), $\tilde{b}_1 \rightarrow t\tilde{t}_2^0$	1 e, $\mu$	2 b	Yes	20.5	$\tilde{b}_1$	320-560 GeV	$m(\tilde{t}_1^0) > 0 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow c\tilde{t}_1^0$	0	mono-jet/c-tag	Yes	20.3	$\tilde{b}_1$	90-200 GeV	$m(\tilde{t}_1^0) > 85 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1$ (natural GMSB)	2 e, $\mu$ (Z)	1 b	Yes	20.7	$\tilde{b}_1$	500 GeV	$m(\tilde{t}_1^0) > 150 \text{ GeV}$
EW direct	$\tilde{t}_1^0\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow b\tilde{t}_1^0$	2 e, $\mu$	0	Yes	20.3	$\tilde{t}_1^0$	85-315 GeV	$m(\tilde{t}_1^0) > 0 \text{ GeV}$
	$\tilde{t}_1^0\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow t\tilde{b}_1^0$	2 e, $\mu$	0	Yes	20.3	$\tilde{t}_1^0$	125-450 GeV	$m(\tilde{t}_1^0) > 0 \text{ GeV}, m(\tilde{\tau}_1^0) > 0.5(m(\tilde{t}_1^0)+m(\tilde{t}_1^0))$
	$\tilde{t}_1^0\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow \tau\tilde{\nu}_\tau$	2 $\tau$	-	Yes	20.7	$\tilde{t}_1^0$	180-330 GeV	$m(\tilde{t}_1^0) > 0 \text{ GeV}, m(\tilde{\tau}_1^0) > 0.5(m(\tilde{t}_1^0)+m(\tilde{t}_1^0))$
	$\tilde{t}_1^0\tilde{t}_1^0 \rightarrow \tilde{t}_1^0\tilde{\nu}_\tau, \tilde{t}_1^0 \rightarrow \tau\tilde{\nu}_\tau$	3 e, $\mu$	0	Yes	20.7	$\tilde{t}_1^0, \tilde{\nu}_\tau$	600 GeV	$m(\tilde{t}_1^0) = m(\tilde{\nu}_\tau), m(\tilde{t}_1^0) > 0, m(\tilde{\tau}_1^0) > 0.5(m(\tilde{t}_1^0)+m(\tilde{t}_1^0))$
	$\tilde{t}_1^0\tilde{t}_1^0 \rightarrow W\tilde{t}_1^0, Z\tilde{t}_1^0$	3 e, $\mu$	0	Yes	20.7	$\tilde{t}_1^0, \tilde{W}, \tilde{Z}$	315 GeV	$m(\tilde{t}_1^0) = m(\tilde{W}), m(\tilde{t}_1^0) = 0, \text{ sleptons decoupled}$
	$\tilde{t}_1^0\tilde{t}_1^0 \rightarrow W\tilde{t}_1^0, h\tilde{t}_1^0$	1 e, $\mu$	2 b	Yes	20.3	$\tilde{t}_1^0$	285 GeV	$m(\tilde{t}_1^0) = m(\tilde{t}_1^0), m(\tilde{t}_1^0) = 0, \text{ sleptons decoupled}$
Long-lived particles	Direct $\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow \text{prod. long-lived } \tilde{t}_1^0$	Disapp. trk	1 jet	Yes	20.3	$\tilde{t}_1^0$	270 GeV	$m(\tilde{t}_1^0) = m(\tilde{t}_1^0) = 160 \text{ MeV}, r(\tilde{t}_1^0) = 0.2 \text{ ns}$
	Stable, stopped $\tilde{g}$ R-hadron	0	1-5 jets	Yes	22.9	$\tilde{g}$	832 GeV	$m(\tilde{t}_1^0) > 100 \text{ GeV}, 10 \mu\text{s} < \tau < 1000 \text{ s}$
	GMSB, stable $\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow \tau(\tilde{e}, \mu) + \tau(e, \mu)$	1-2 $\mu$	-	-	15.9	$\tilde{t}_1^0$	475 GeV	$10^{-4} < \tau < 50 \text{ ns}$
RPV	GMSB, $\tilde{t}_1^0 \rightarrow \nu\tilde{t}_1^0$ , long-lived $\tilde{t}_1^0$	2 $\gamma$	-	Yes	4.7	$\tilde{t}_1^0$	230 GeV	$0.4 < \tau < 2 \text{ ns}$
	$\tilde{q}\tilde{q}, \tilde{t}_1^0 \rightarrow \tilde{q}\tilde{q}$ (RPV)	1 $\mu$ , displ. vtx	-	-	20.3	$\tilde{t}_1^0$	1.0 TeV	$1.5 < c\tau < 156 \text{ nm}, \text{BR}(\mu\mu) = 1, m(\tilde{t}_1^0) = 108 \text{ GeV}$
	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu + \tau$	2 e, $\mu$	-	-	4.6	$\tilde{\nu}_\tau$	1.81 TeV	$\tilde{a}_{\mu\tau} = 0.10, \tilde{a}_{\tau\tau} = 0.05$
Other	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu + \tau$	1 e, $\mu$ + $\tau$	-	-	4.6	$\tilde{\nu}_\tau$	1.1 TeV	$\tilde{a}_{\mu\tau} = 0.10, \tilde{a}_{\tau\tau} = 0.05$
	Binlinear RPV CMSSM	1 e, $\mu$	7 jets	Yes	4.7	$\tilde{g}$	1.2 TeV	$m(\tilde{g}) = m(\tilde{g}), c_{\text{RPV}} < 1 \text{ mm}$
	$\tilde{t}_1^0\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow W\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow e\tilde{\nu}_\tau, \mu\tilde{\nu}_\tau$	4 e, $\mu$	-	Yes	20.7	$\tilde{t}_1^0$	760 GeV	$m(\tilde{t}_1^0) > 300 \text{ GeV}, \tilde{a}_{\mu\tau} < 10$
	$\tilde{t}_1^0\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow W\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow \tau\tilde{\nu}_\tau, e\tilde{\nu}_\tau$	3 e, $\mu$ + $\tau$	-	Yes	20.7	$\tilde{t}_1^0$	350 GeV	$m(\tilde{t}_1^0) > 80 \text{ GeV}, \tilde{a}_{\mu\tau} < 10$
	$\tilde{g} \rightarrow \tilde{q}\tilde{q}$	0	6-7 jets	-	20.3	$\tilde{g}$	916 GeV	$\text{BR}(\tau) = \text{BR}(h) = \text{BR}(c) = 0\%$
	$\tilde{g} \rightarrow \tilde{t}_1^0\tilde{t}_1^0, \tilde{t}_1^0 \rightarrow b\tilde{s}$	2 e, $\mu$ (SS)	0-3 b	Yes	20.7	$\tilde{g}$	880 GeV	
Other	Scalar gluon pair, sgluon $\rightarrow \tilde{q}\tilde{q}$	0	4 jets	-	4.6	sgluon	100-287 GeV	ind. limit from 1110.2693
	Scalar gluon pair, sgluon $\rightarrow \tilde{t}_1^0\tilde{t}_1^0$	2 e, $\mu$ (SS)	1 b	Yes	14.3	sgluon	600 GeV	
	WIMP interaction (D5, Dirac $\chi$ )	0	mono-jet	Yes	10.5	$\tilde{g}$	704 GeV	$m(\chi) = 80 \text{ GeV}, \text{limit of } \sim 687 \text{ GeV for D8}$

$\sqrt{s} = 7 \text{ TeV}$  full data  $\sqrt{s} = 8 \text{ TeV}$  partial data  $\sqrt{s} = 8 \text{ TeV}$  full data

Mass scale [TeV]

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 $\sigma$  theoretical signal cross section uncertainty.

# Motivations for beyond Minimal SUSY searches

## ▶ **Non-observations of SUSY particles @ the LHC:**

- **Rekindled the interest** for **non-minimal SUSY** models

## ▶ **Idea behind non-minimal SUSY models**

- **It could be that we are missing an additional ingredient**
  - **R-Parity Violating (RPV) MSSM**
  - NMSSM, Left-Right MSSM, MRSSM, Vector-like MSSM

## ▶ **Price to pay for non-minimal SUSY**

- **More interactions** ⇒ **New free parameters**
- **Phenomenological analyses** ⇒ **More complicated**

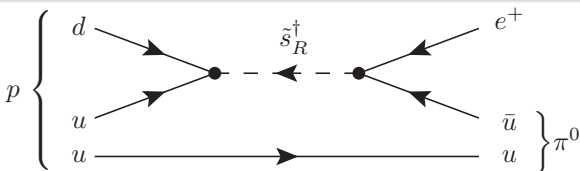
## ▶ **Attractive feature of non-minimal SUSY**

- **May solve problems that minimal SUSY does not**
  - **Baryon Asymmetry in the Universe (BAU)** [Barbier et al.]
  - **Neutrino mass generation** [Barbier et al.]
  - **$\mu$  problem**

# Motivations for R-Parity Violating SUSY

## R-Parity

- Forbids **both BNV** and **LVN interactions** in the MSSM
- Imposed to avoid **fast proton decay**:  $\tau_p \gtrsim 10^{33}$  years



- ▶ **Proton decay requires both BNV and LVN interactions**
  - R-parity conservation is **too restrictive**
  - **Either BNV or LVN are allowed**  $\Rightarrow$  **No proton decay**
- ▶ **RPV-MSSM with BNV model features**
  - BNV + lepton number conservation **compatible** with a **GUT**
  - Provides the **third Sakharov condition**  $\Rightarrow$  **BAU**
- ▶ **Price to pay**: Extremely difficult to accommodate DM

# The model

## ▶ BNV superpotential:

$$W_{BNV} = \frac{1}{2} \lambda''_{ijk} U^i D^j D^k + W_{MSSM}$$

- $\lambda''_{ijk}$ : BNV couplings  $\Rightarrow$  9 new free independent parameters
- $U, D$ : Superfields
- $i, j$  and  $k$ : Flavor indices

## ▶ BNV Lagrangian:

$$\begin{aligned} \mathcal{L}_{U_i U_j D_k} &= -\frac{1}{2} \lambda''_{ijk} \varepsilon^{c_1 c_2 c_3} \left( \tilde{u}_{lc_3}^{0\dagger} R_{l(k+3)}^u \bar{\Psi}_{Dic_1}^d P_L \Psi_{Djc_2}^d \right. \\ &\quad \left. + \tilde{d}_{lc_2}^{0\dagger} R_{l(k+3)}^d \bar{\Psi}_{Dic_1}^u P_L \Psi_{Dkc_2}^d + \bar{\Psi}_{Dic_1}^u P_L \Psi_{Djc_2}^d R_{l(k+3)}^d \tilde{d}_{lc_3}^{0\dagger} \right) + \text{h.c.} \end{aligned}$$



# Constraints on BNV couplings

## Present experimental constraints on $\lambda''_{ijk}$ couplings:

- Neutron dipole moment [Slavich arXiv:0008270]
- Antinucleon oscillations, double nucleon decays
- Rare hadronic decays of  $B$ -mesons,  $K$ - $K$  systems
- **Observed flux of cosmic rays antiprotons :**  
[Gondolo arXiv:9704411]

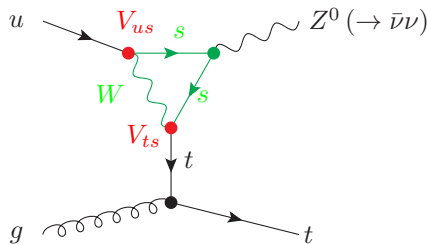
$$\lambda''_{ijk} < 10^{-19} - 10^{-24}$$

- **Yet... Not applicable** to  $\lambda''_{3jk}$  if the **top quark** is **heavier than the Lightest Supersymmetric Particle (LSP)**
  - ▶  $\lambda''_{3jk}$  is left **almost unconstrained** [Barbier et al.]
- **Enforce MFV**  $\Rightarrow$  **Only**  $\lambda''_{312}$  is **sizable**  $\sim \mathcal{O}(0.1)$

# Monotop production in the Standard model

[Fuks, Andrea, Maltoni arXiv:11066199]

- Final state signature :  $t + \cancel{E}_T \longrightarrow bjj + \cancel{E}_T$
- Production mode  $\longrightarrow$  **subdominant contribution**
  - **CKM suppressed:**  $V_{us} \simeq 0.23$ ,  $V_{ts} \simeq 0.04$
  - **Loop-suppressed**
  - **Branching ratio:**  $\text{BR}(Z \rightarrow \nu\bar{\nu}) \simeq 0.2$

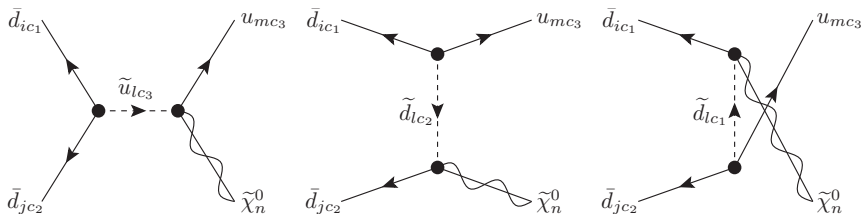


# Monotop production within the RPV-MSSM

## ► Production @ tree-level through squark exchange

$$pp \rightarrow \tilde{q} \rightarrow \tilde{\chi}_1^0 + t$$

- **6 diagrams** in the **flavor conserving case**
- $\cancel{E}_T$  associated to the lightest neutralino
  - ⇒ Kinematic condition:  $m_t > m_{\tilde{\chi}_1^0} \Rightarrow$  **Long-lived neutralino**
  - ⇒ Decay far outside of the detector due to its **long lifetime**



# Analytical results

## ► Cross section

### Full 2 → 2 process

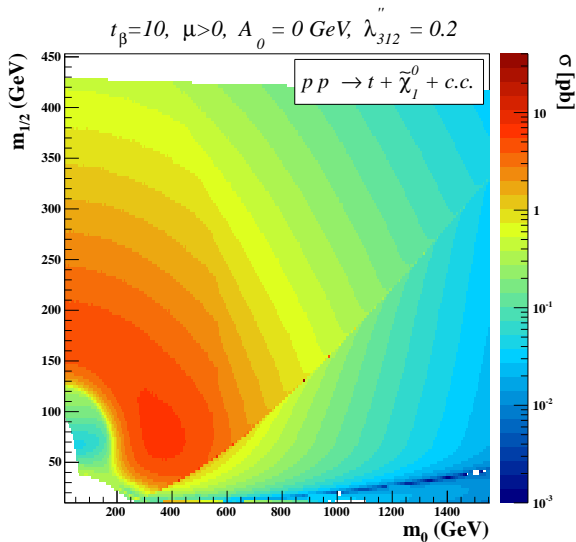
$$\frac{d\hat{\sigma}_{h_a h_b}}{dt} = \frac{\alpha}{12s_W^2 c_W^2 s^2} \sum_{l,o=1,2}^6 \left[ (1-h_a)(1-h_b) \left( \frac{Q_{lo}^{ss}}{s_l s_o} + \frac{Q_{lo}^{tt-}}{t_l t_o} + \frac{Q_{lo}^{uu-}}{u_l u_o} + \frac{Q_{lo}^{st}}{s_l t_o} \right. \right. \\ \left. \left. + \frac{Q_{lo}^{su}}{s_l u_o} + \frac{Q_{lo}^{tu}}{t_l u_o} \right) + (1-h_a)(1+h_b) \frac{Q_{lo}^{tt+}}{t_l t_o} + (1+h_a)(1-h_b) \frac{Q_{lo}^{uu+}}{u_l u_o} \right]$$

- **Compactified expression** of the **cross section**
- $h_a$  and  $h_b \implies$  **helicities** of **incoming particles**
- $Q^{ss}$ : interferences between diagrams in the s-channel

# Tool chain

- ▶ **SLHA2 input file:** [Allanach et al.]
  - Model and parameters specification
- ⇓
- ▶ **SUSY-spectrum:** SPheno-3.2.1: [Porod, Staub]
  - Compute **RGEs** at the **two loop level**
- ⇓
- ▶ **Decay Width an Branching Ratios:** SUSY-HIT-1.3  
[Djouadi, Mullheitner, Spira]
  - **Width** and **BRs** of the **MSSM Higgs bosons:** HDECAY
  - **Width** and **BRs** of the **SUSY particles:** SDECAY
- ⇓
- ▶ **Cross section & scans:** XSUSY-1.9.23 (**Private code**)
  - Original C++ **code** developped by **B. Fuks** & **B. Herrmann**
  - **Numerical integration** performed using **VEGAS**
  - **Extended** with **new PDF sets**
    - MRST2002, DSSV.2008, NNPDFpo11.0, BB10, NNPDF2.1
  - **Extended** with **new features**
    - Scans for polarized cross section automated

# Numerical results



# The Narrow Width Approximation

- ▶ **The Narrow Width Approximation (NWA):**
  - **Reduction** of the **complexity of scattering amplitudes**
  - **Assumes peaked resonance** with a **Breit-Wigner lineshape**
  - For small  $\Gamma$ , **off-shell effects are suppressed**
    - ➔ Intermediate resonance can be **approximated** to be **on-shell**
  - **Production and decay of unstable particles**  $\Rightarrow$  **factorized**
  - **Non-resonant contributions** are **neglected**
  - Introduction of an **error** of  $\mathcal{O}(\Gamma/M)$  for each **Breit-Wigner**

## Requirements for the NWA

- 1 **Total width** of the particle way **smaller** than its **mass**:  $\Gamma \ll M$
- 2 Propagator **separable** from the **matrix element**
- 3 No significant interferences from **non-resonant processes**
- 4 **Scattering energy larger** than mass of the **resonance**:  $\sqrt{\hat{s}} \gg M$
- 5 Mass of **resonance larger** than masses of the **daughter** particles

# Narrow Width Approximation & New Physics

[Kauer, Rainwater, Berdine arXiv:0703058]

## ▶ **Narrow Width Approximation features:**

- Drastically **simplifies calculations**  $\Rightarrow \sigma_p \times \text{BR}$
- **Reduce CPU time** required for computations
- **Works pretty well** in the case of the **SM**
- **Extensively used** for **BSM searches**

## ▶ **Is the NWA reliable in the context of BSM searches?**

- **NWA assumes Breit-Wigner resonance**
  - In the vicinity of kinematical bounds like  $\sqrt{\hat{s}} \sim M$
  - For near-degenerate parent-daughter masses  $m \sim M$
- **Breit-Wigner lineshape is distorted** by threshold factors
  - phase space factors  $\beta = \sqrt{1 - (m/M)^2}$

## ▶ **Preliminary results for monotop RPV-production**

- Indicate **substantial discrepancies** between **NWA** and  $2 \rightarrow 2$
- Stay tuned...



# Conclusions & Perspectives

## Conclusions

- Monotop production in the SM **subdominant**
- **Observation** of monotop signature **means New physics**
- **RPV-MSSM** with **BNV** allows for monotop production @ LO
- **Strictly speaking ruled-out** by LHC constraints
  - ➔ **Meaning** in the case of **RPV-CMSSM**
- **Not excluded** for **more general models** like the pMSSM
  - **NWA** is **not always** reliable

## Perspectives:

- Comparison **naive NWA/off-shell effects**
- Compute **spin asymmetries**
- Compare with **ratios of parton-luminosities**

# Thanks for your attention

# Back-up slides

## Sakharov's conditions :

- 1 **Baryon number is violated**
- 2  $CP$  and  $C$  symmetries are **violated**
- 3 **Departure from thermal equilibrium**

### ▶ **Standard Model:**

- Baryon number is violated  $\Rightarrow$  **Sphalerons**
  - ▶ Non-perturbative processes
    - **Violates  $B + L$  but preserve  $B - L$**
- Departure from thermal equilibrium
  - **ELECTROWEAK PHASE TRANSITION**

### ▶ Standard electroweak baryogenesis $\Rightarrow$ **RULED OUT**

# Full expressions of the charges

$$\begin{aligned}
 Q_{lo}^{ss} &= s C_{d_i d_j u_l} \tilde{\sim} C_{d_i d_j u_o}^* \tilde{\sim} \left[ \left( s - m_{q_m}^2 - m_{\tilde{\chi}_n^0}^2 \right) \left( L_{u_m u_l \tilde{\chi}_n^0} \tilde{\sim} L_{u_m u_o \tilde{\chi}_n^0}^* \tilde{\sim} \right) \right. \\
 &\quad \left. + R_{u_m u_l \tilde{\chi}_n^0} \tilde{\sim} R_{u_m u_o \tilde{\chi}_n^0}^* \tilde{\sim} \right) - 2m_{u_m} m_{\tilde{\chi}_n^0} \left( L_{u_m u_l \tilde{\chi}_n^0} \tilde{\sim} R_{u_m u_o \tilde{\chi}_n^0}^* \tilde{\sim} + R_{u_m u_l \tilde{\chi}_n^0} \tilde{\sim} L_{u_m u_o \tilde{\chi}_n^0}^* \tilde{\sim} \right) \Big] \\
 Q_{lo}^{tt+} &= C_{u_m d_i d_l} \tilde{\sim} C_{u_m d_i d_o}^* \tilde{\sim} \left( t - m_{u_m}^2 \right) \left( t - m_{\tilde{\chi}_n^0}^2 \right) \left( L_{d_j d_l \tilde{\chi}_n^0} \tilde{\sim} L_{d_j d_o \tilde{\chi}_n^0}^* \tilde{\sim} \right) \\
 Q_{lo}^{tt-} &= C_{u_m d_i d_l} \tilde{\sim} C_{u_m d_i d_o}^* \tilde{\sim} \left( t - m_{u_m}^2 \right) \left( t - m_{\tilde{\chi}_n^0}^2 \right) \left( R_{d_j d_l \tilde{\chi}_n^0} \tilde{\sim} R_{d_j d_o \tilde{\chi}_n^0}^* \tilde{\sim} \right) \\
 Q_{lo}^{uu+} &= C_{u_m d_j d_l} \tilde{\sim} C_{u_m d_j d_o}^* \tilde{\sim} \left( u - m_{u_m}^2 \right) \left( u - m_{\tilde{\chi}_n^0}^2 \right) \left( L_{d_i d_l \tilde{\chi}_n^0} \tilde{\sim} L_{d_i d_o \tilde{\chi}_n^0}^* \tilde{\sim} \right)
 \end{aligned}$$

# Full expressions of the charges

$$\begin{aligned}
 Q_{lo}^{uu-} &= C_{u_m d_j \tilde{d}_l} C_{u_m d_j \tilde{d}_o}^* \left( u - m_{u_m}^2 \right) \left( u - m_{\tilde{\chi}_n^0}^2 \right) \left( R_{d_i \tilde{d}_l \tilde{\chi}_n^0} R_{d_i \tilde{d}_o \tilde{\chi}_n^0}^* \right) \\
 Q_{lo}^{st} &= 2m_{\tilde{\chi}_n^0} m_{u_m} s \Re \left( C_{d_i d_j u_l} L_{u_m u_l \tilde{\chi}_n^0} C_{u_m d_i \tilde{d}_o}^* R_{d_j \tilde{d}_o \tilde{\chi}_n^0}^* \right) \\
 &+ 2st \Re \left( C_{d_i d_j u_l} R_{u_m u_l \tilde{\chi}_n^0} C_{u_m d_i \tilde{d}_o}^* R_{d_j \tilde{d}_o \tilde{\chi}_n^0}^* \right) \\
 Q_{lo}^{su} &= -2m_{\tilde{\chi}_n^0} m_{u_m} s \Re \left( C_{d_i d_j u_l} L_{u_m u_l \tilde{\chi}_n^0} C_{u_m d_j \tilde{d}_o}^* R_{d_i \tilde{d}_o \tilde{\chi}_n^0}^* \right) \\
 &- 2us \Re \left( C_{d_i d_j u_l} R_{u_m u_l \tilde{\chi}_n^0} C_{u_m d_j \tilde{d}_o}^* R_{d_i \tilde{d}_o \tilde{\chi}_n^0}^* \right) \\
 Q_{lo}^{ut} &= 2 \left( m_{\tilde{\chi}_n^0}^2 m_{u_m}^2 - ut \right) \Re \left( C_{u_m d_i \tilde{d}_l} R_{d_j \tilde{d}_l \tilde{\chi}_n^0} C_{u_m d_j \tilde{d}_o}^* R_{d_i \tilde{d}_o \tilde{\chi}_n^0}^* \right)
 \end{aligned}$$

# Full expressions for the couplings

$$L_{\tilde{u}_j u_k \tilde{\chi}_i^0} = \left[ (e_q - T_q^3) s_W N_{i1} + T_q^3 c_W N_{i2} \right] R_{jk}^{u*} + \frac{m_{u_k} c_W N_{i4} R_{j(k+3)}^{u*}}{2 m_W \sin \beta}$$

$$R_{\tilde{u}_j u_k \tilde{\chi}_i^0} = -e_q s_W N_{i1}^* R_{j(k+3)}^{u*} + \frac{m_{u_k} c_W N_{i4}^* R_{jk}^{u*}}{2 m_W \sin \beta}$$

$$L_{\tilde{d}_j d_k \tilde{\chi}_i^0} = \left[ (e_q - T_q^3) s_W N_{i1} + T_q^3 c_W N_{i2} \right] R_{jk}^{d*} + \frac{m_{d_k} c_W N_{i4} R_{j(k+3)}^{d*}}{2 m_W \sin \beta}$$

$$R_{\tilde{d}_j d_k \tilde{\chi}_i^0} = -e_q s_W N_{i1}^* R_{j(k+3)}^{d*} + \frac{m_{d_k} c_W N_{i4}^* R_{jk}^{d*}}{2 m_W \sin \beta}$$

$$C_{d_i d_j \tilde{u}_l} = \lambda_{ijk}'' R_{l(k+3)}^u \quad \text{and} \quad C_{u_i d_j \tilde{d}_l} = \lambda_{ijk}'' R_{l(k+3)}^d$$

With

$$\tilde{\chi}_i^0 = N_{ij} \psi_j^0 \quad \text{with} \quad i = 1, 2, 3, 4 \quad (1)$$