

# Source foil design studies

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# The SuperNEMO experiment

## $0\nu\beta\beta$ with calo-track technique @ LSM:

- Tracker + calorimeter + magnetic fiels: full kinematic
- Detector  $\neq$  from  $\beta\beta$  emitter: optimise isotope choice

## Exploit a well known technique (NEMO3)

- Increase  $\beta\beta$  emitter mass up to  $\sim 100$  kg
- Improve detection technique and reduce background

## Goal: background free measurements in 5 years

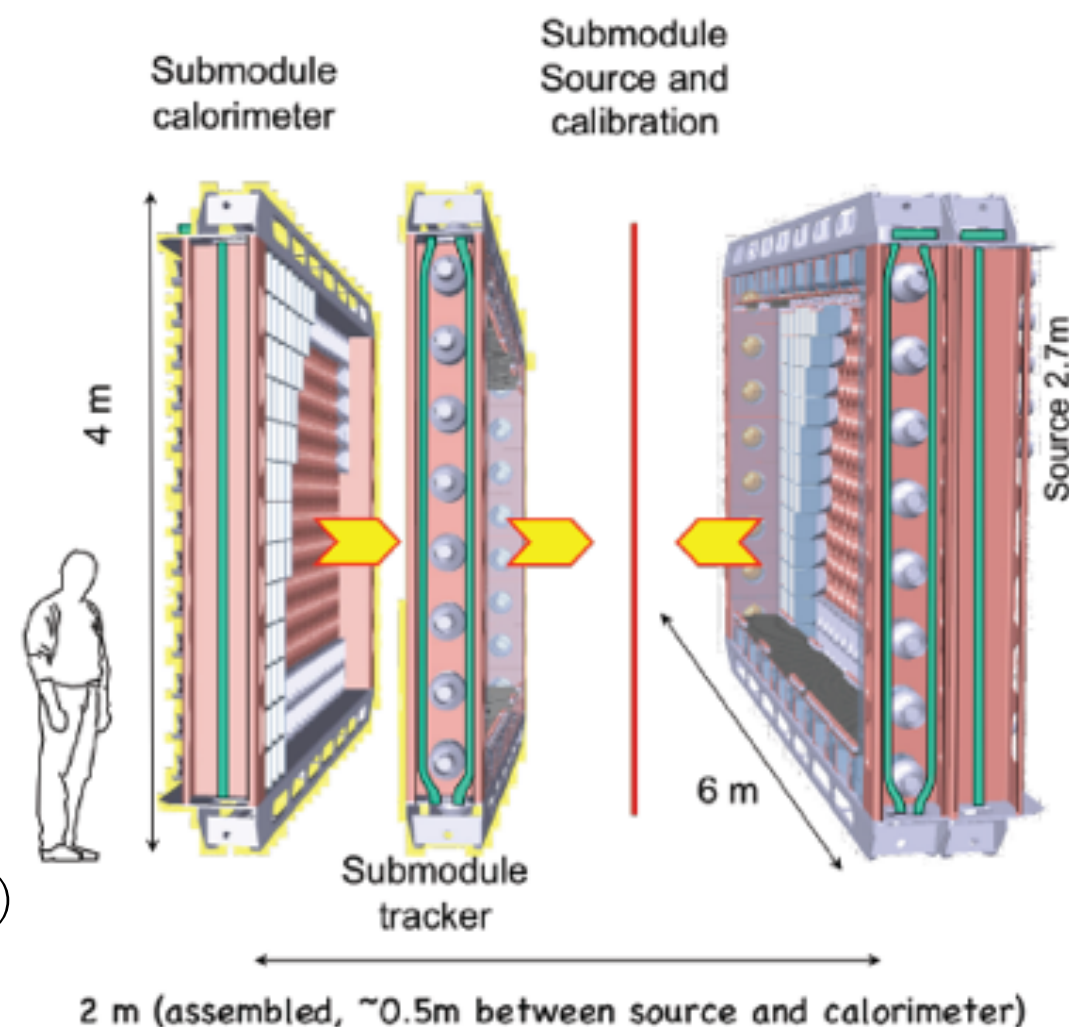
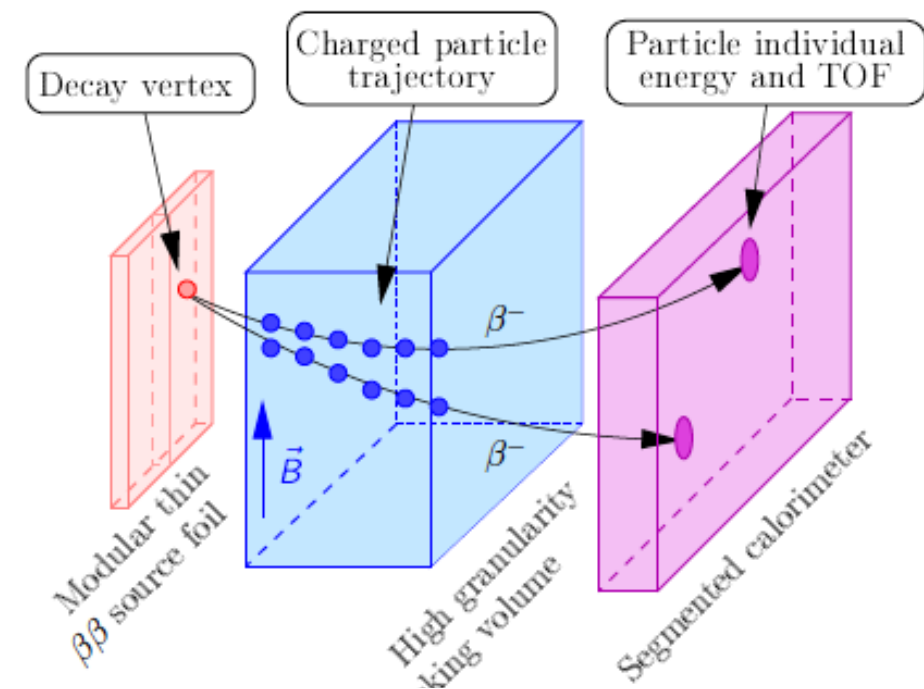
- Improve current limits by factor  $\sim 100$

## The demonstrator module

- 1 module with 7 kg  $^{82}\text{Se}$
- Match SuperNEMO's radio-purity requirements
- Less than 0.2 background events in search region with in 3yrs
- Achieve competitive sensitivity:
  - NEMO3 ( $^{100}\text{Mo}$ ) in 4.5 months
  - By 2017:  $T_{1/2}^{0\nu} > 6.5 \times 10^{24}$  yr

## LAPP responsibilities:

- Produce half of the  $\beta\beta$  foil sources for Demonstrator module (this talk)
- Slow control system development (next talk)



- Goal of the study
- Present general strategy
- Some useful figure of merits (FOM)
- Some preliminary results

# Kick off

# Goal of the study

Source foil design is **not completely frozen** yet:

- Foil geometry frozen (see ex. DocDB 2589 Marek)
- Sensitivity studies set some target values
  - Thickness,  $^{208}\text{Tl}$  &  $^{214}\text{Bi}$  max activity, others (?)
- No detailed foil composition yet ==> how do we produce it?

Let's have a **look at the details** and try to **come out with a proposal**:

- Which are those details?
- How to proceed?

# A look to the detail

Starting from NEMO3 composite foil design

- $\beta\beta$  isotope mixed with Polyvinyl-Alcohol (PVA)
- 2 Mylar backing film for mechanical support

Few simple parameters come to mind:

- (1) Foil thickness
- (2) Fraction of PVA mix
- (3) Type of mechanical support
  - Mylar backing films, central mesh, fishing wires

Keeping in mind target parameters already defined ==>

==> Let's define the design from scratch for SuperNEMO w.r.t

- **Detector performances:** dedicated MC studies (this work)
- **Production procedure:** dedicated R&D on going @ LAPP (but not only)

# How to proceed

Let's start simple:

- (1) Identify some figure of merits (FOM) related to detector performances
- (2) Let's study how such FOM behave w.r.t. source parameters
- (3) Come up with a preliminar proposal

In parallel : x-check feasibility with R&D test

In a next (future) step ==> detailed sensitivity study

From the technical point of view:

- MC simulation with **SN@ilWare** framework, legacy version:
  - Cadfael-0.2.3 ; Bayeaux-0.9.4 ; Falaise-0.2.1
  - Custom module to produce ROOT output (sncore ex05)

# Figure of merits

(1) Study electron generated in source foil

- Distribution of electron energy loss ==> **Peak** and **FWHM**
- Total energy resolution

(2) Study alpha generated in source foil

- **Exit probability** of alpha produced in the foil
- $^{214}\text{Bi}$  background measurement

Later we could consider:

(3) bb0v detection efficiency (require a bit more elaborated study)

# Events generation

- Geometry: **config\_2.0**
  - No Mylar backing film
  - Just Selenium mixed with 5 % Polyvinyl Alcohol
  - Target thickness: 50 mg/cm<sup>2</sup> ==> 0.167 mm
- Vertex generator: **source\_strips\_bulks**
- Event generator:
  - **Electron\_monokinetic:** 1 MeV, random momenta
  - **Alpha\_monokinetic:** 7.7 MeV (<sup>214</sup>Bi), random momenta
- Total number of event: 100k evt.



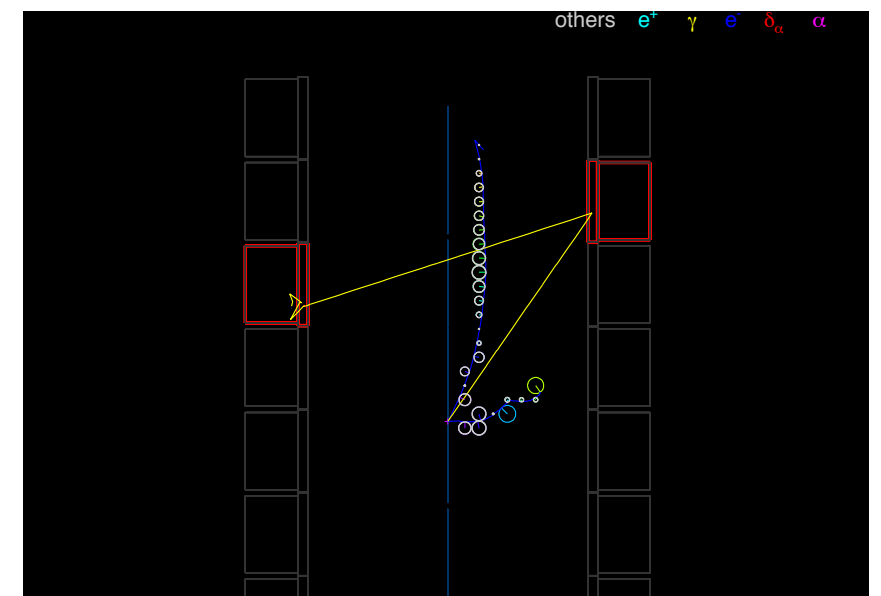
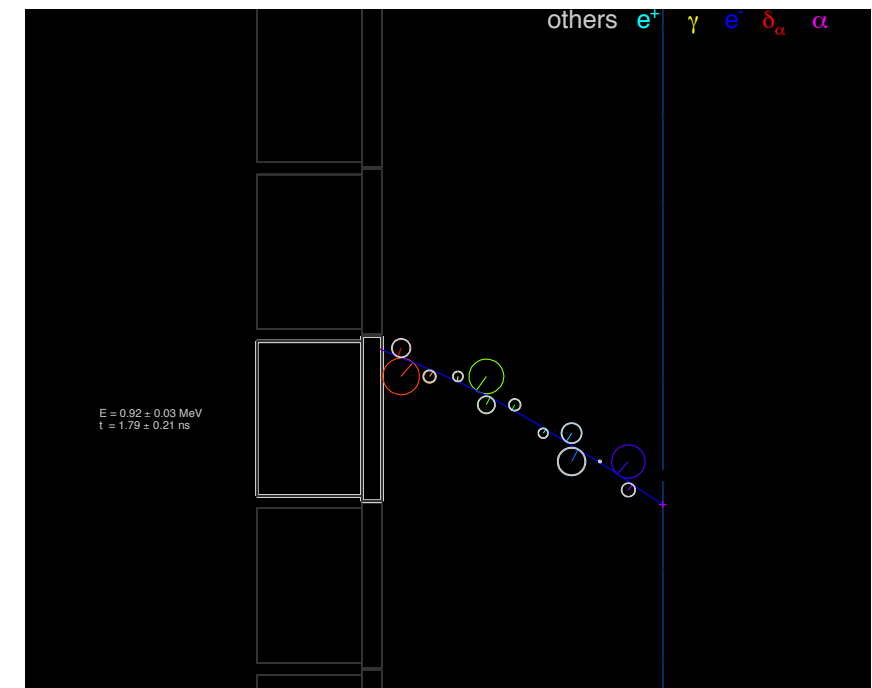
# Electron energy loss

# Selection criteria

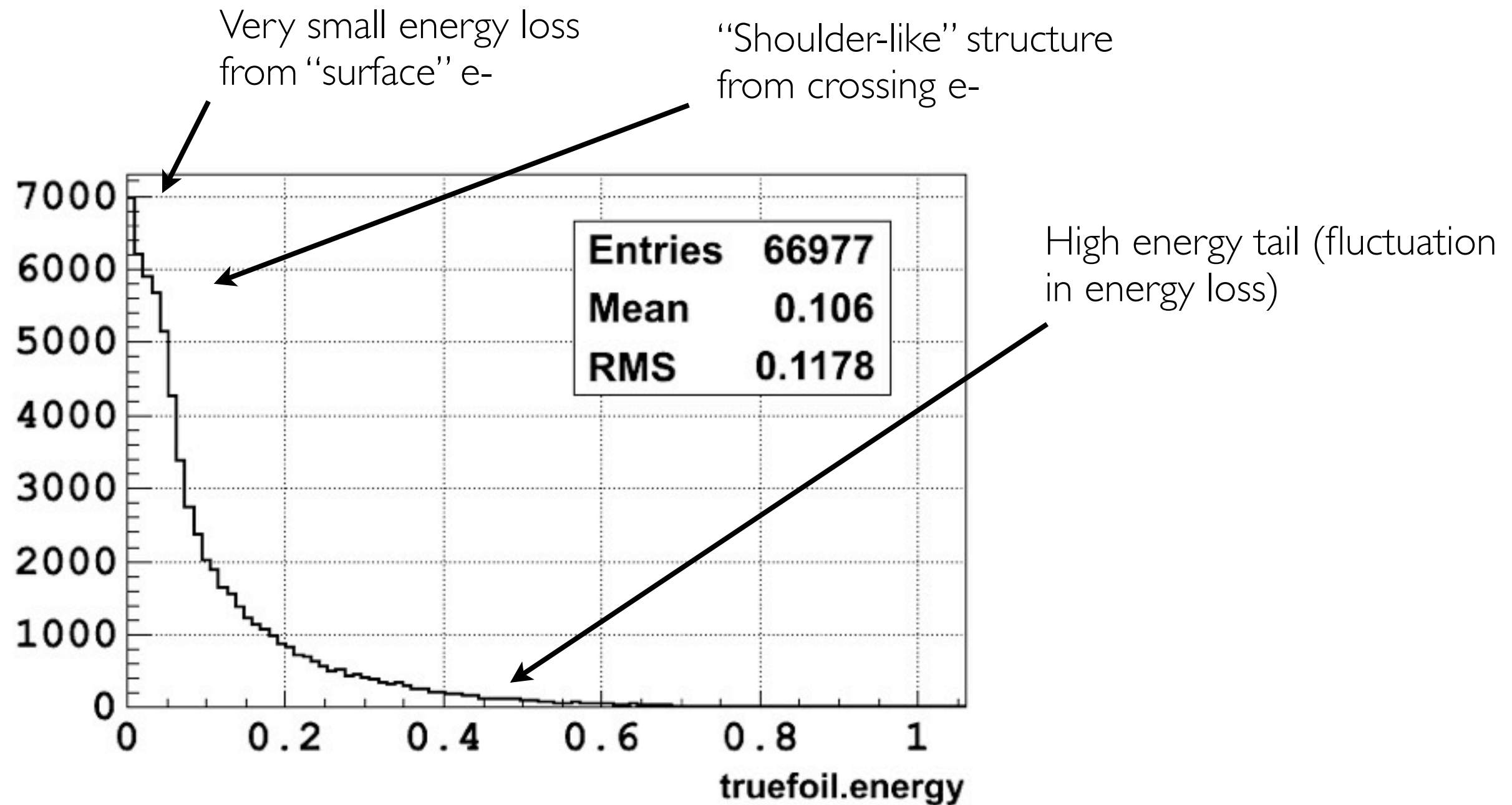
- Selection criteria:
- One particle from the source ==> one hit in the main calo
- Get rid of funny events
- Total number of evt. after selection: 66977 evt. (66%)

Check:

- True energy loss in foil source
- True energy loss in main calo
- Calibrated energy loss in main calo (resolution effects)

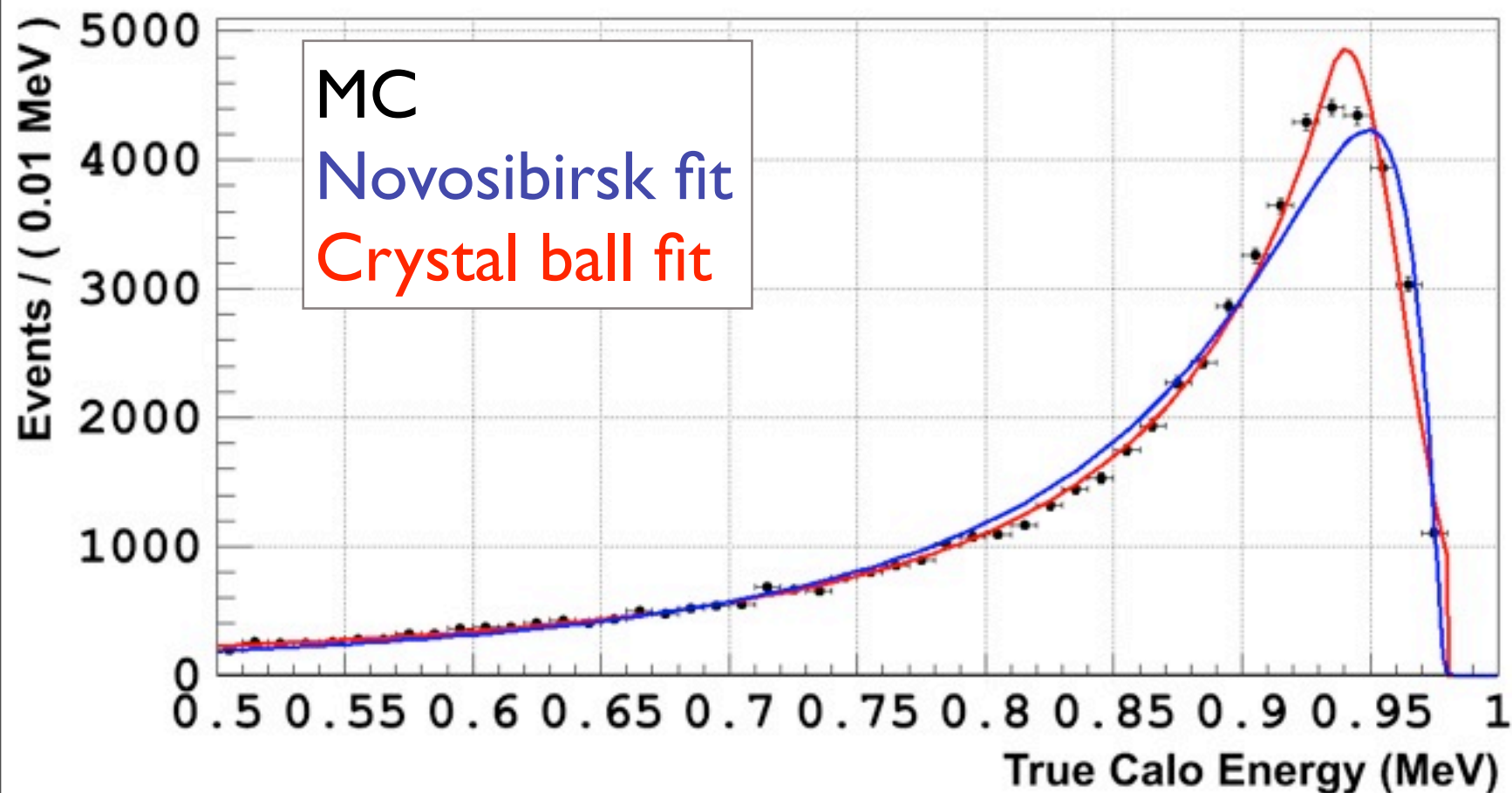


# Electron energy loss in source



- Distribution not easy to characterise through Peak and FWHM

# True energy in the calorimeter



Floating Parameter	FinalValue +/-	Error
-----	-----	-----
CBS_alpha	3.3012e-01 +/-	7.18e-03
CBS_n	2.1502e+00 +/-	6.05e-02
CBS_peak	9.4048e-01 +/-	3.08e-04
CBS_width	2.1740e-02 +/-	3.05e-04
Reduced chi2: 6.13506673248		

Floating Parameter	FinalValue +/-	Error
-----	-----	-----
NOV_peak	9.4911e-01 +/-	3.55e-04
NOV_tail	-1.0939e+00 +/-	7.35e-03
NOV_width	4.4778e-02 +/-	3.23e-04
Reduced chi2: 9.8176867036		

- Convolution among E loss in source, tracker & scint. block
  - No calorimetric effect here, only fluctuation in energy loss
- Fit with two models:
  - **Novosibirsk**: gauss with exponential tail
  - **Crystal Ball**: gauss with power law tail
- CB generally fits better (for this study)

commonly used to  
models energy loss

# Calibrated energy (resolution effect)

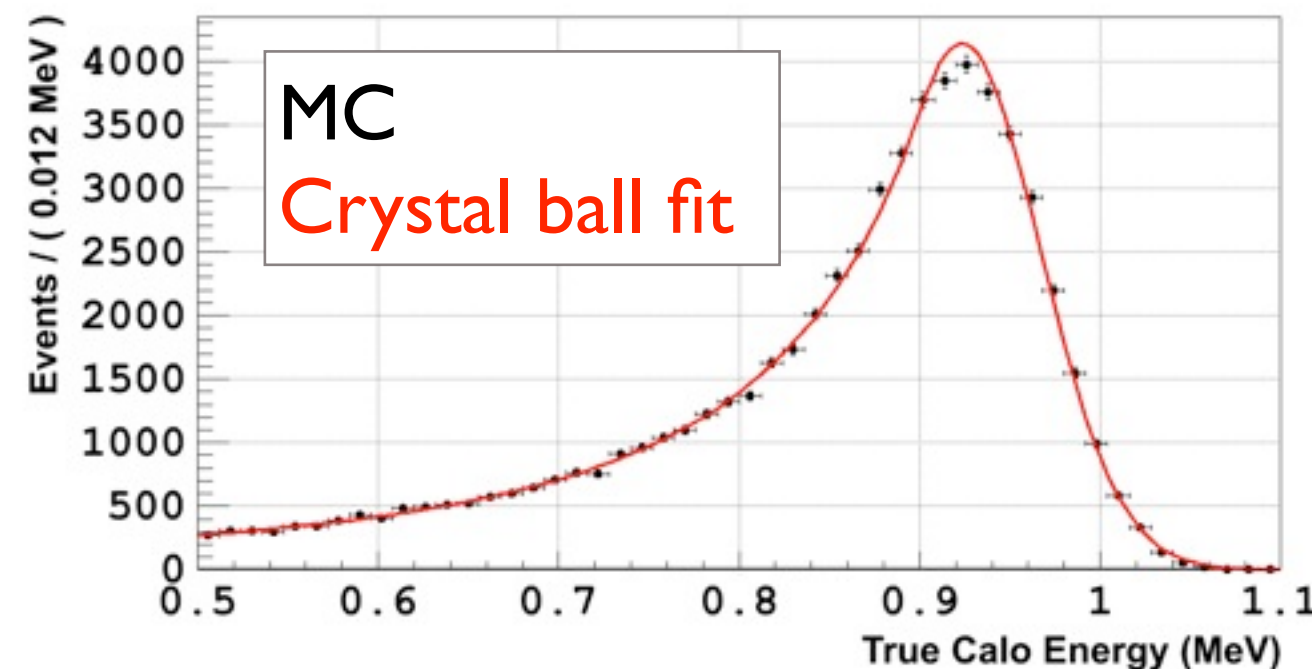
	Peak [MeV]	FWHM [MeV]	Resolution @ 1 MeV [%]
TRUE	0.940 +/- 0.001	0.051 +/- 0.002	<b>5.2 +/- 0.2</b>
Calib.	0.923 +/- 0.001	0.103 +/- 0.002	<b>10.7 +/- 0.2</b>

Add 8 % gaussian spread (?)

Floating Parameter	FinalValue +/-	Error
-----	-----	-----
CBS_alpha	5.0949e-01 +/-	1.01e-02
CBS_n	2.3899e+00 +/-	1.03e-01
CBS_peak	9.2326e-01 +/-	4.23e-04
CBS_width	4.3732e-02 +/-	2.94e-04
Reduced chi2: 2.21700146943		

chi2 looks better

Calo Energy [MeV]



Peak position decrease of ~17 keV: was it expected? (Light collection on PMT? QE?)

Subtracting 8% resolution in quadrature ==> 7.1%  
so about 2% higher than expected.

Something I'm not keeping into account?  
Something else?



# Let's change the foil thickness

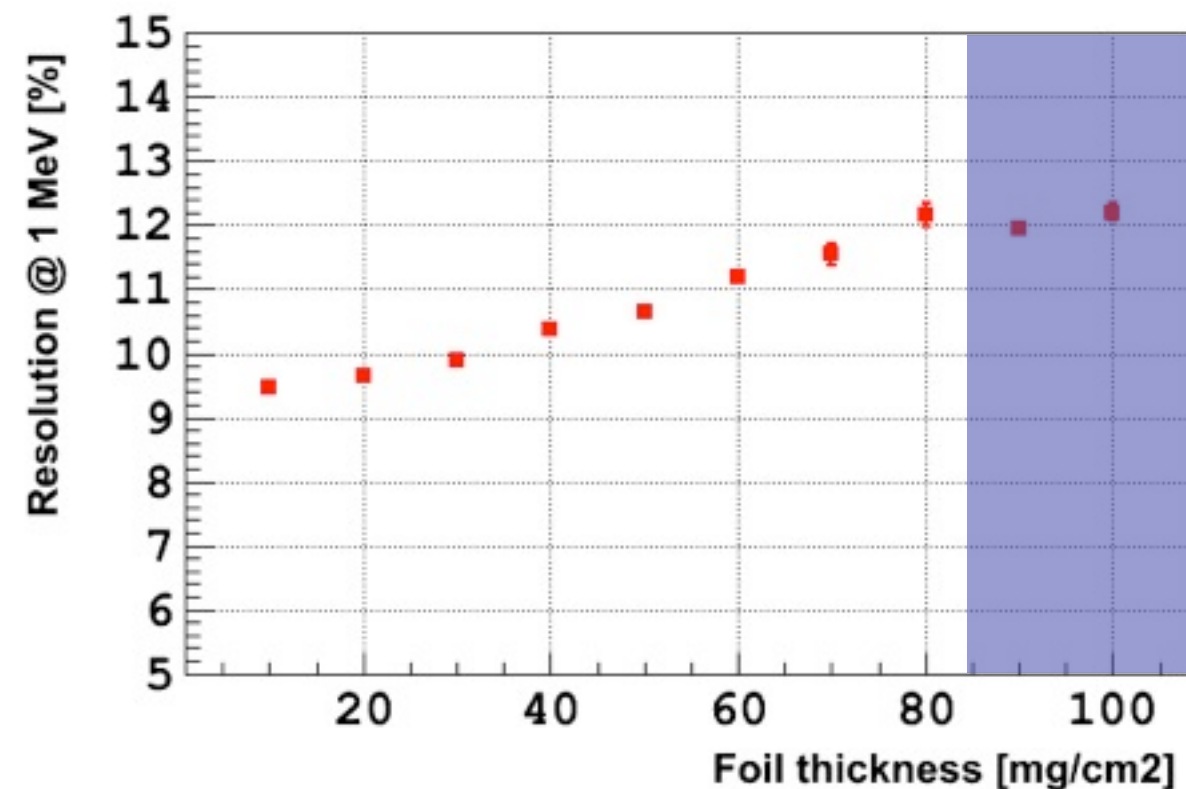
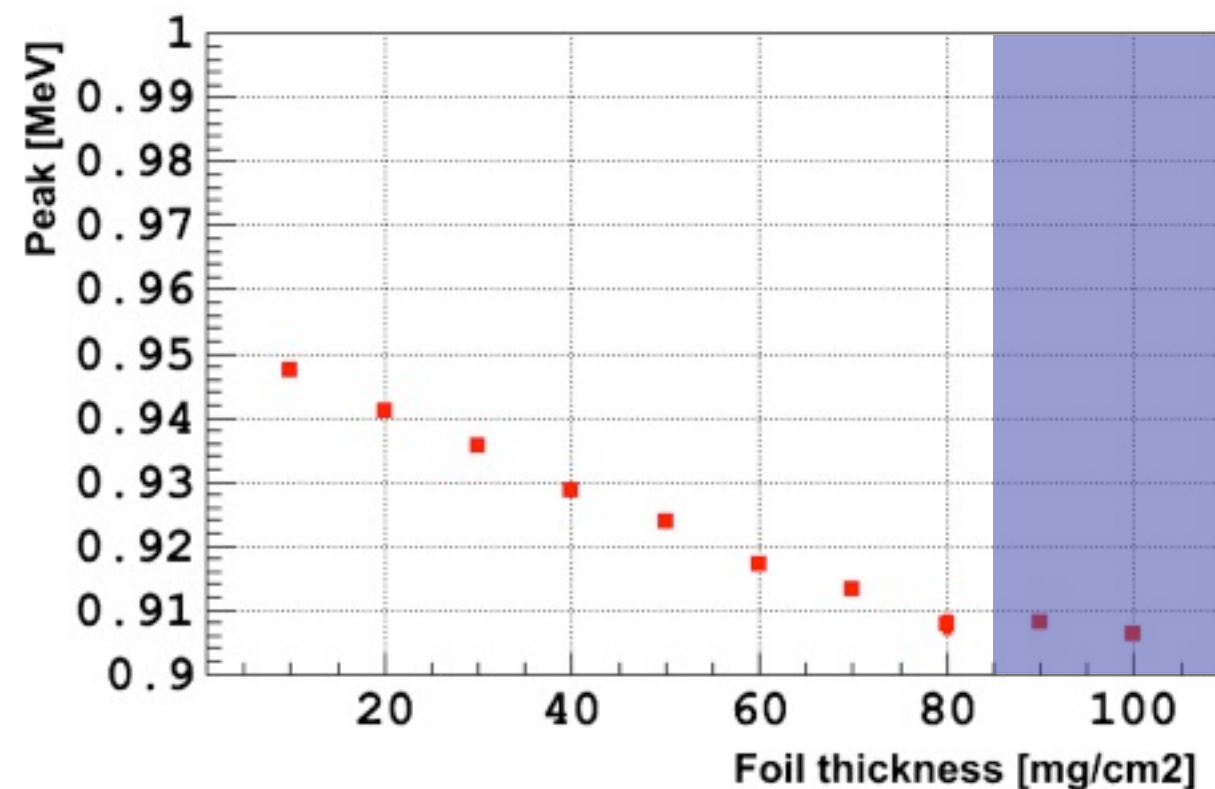
[mg/cm<sup>2</sup>]

[MeV]

[%]

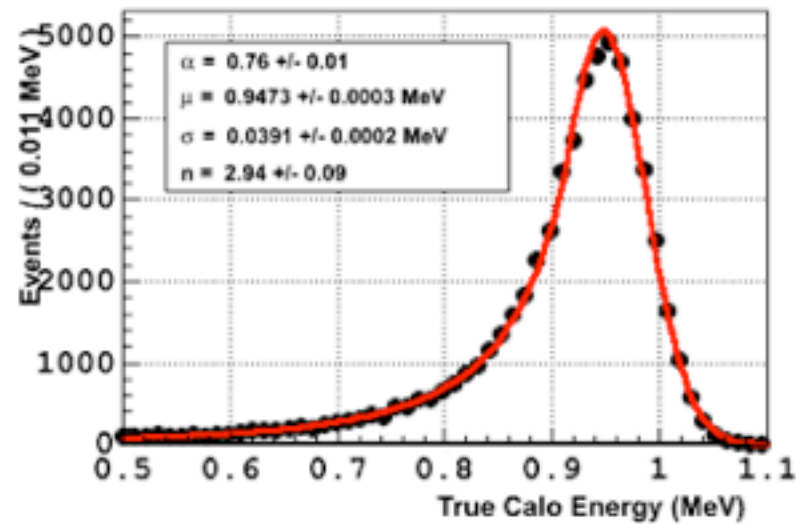
* thickness *	peak *	100*fwhm/ *	chi2 *
10 *	0.947 *	9.46 *	2.34 *
20 *	0.941 *	9.64 *	1.61 *
30 *	0.936 *	9.88 *	2.25 *
40 *	0.929 *	10.4 *	1.71 *
50 *	0.924 *	10.6 *	2.43 *
60 *	0.917 *	11.2 *	2.13 *
70 *	0.913 *	11.5 *	1.88 *
80 *	0.908 *	12.2 *	2.45 *
90 *	0.908 *	11.9 *	4.3 *
100 *	0.906 *	12.2 *	3.72 *

- Good fit except last two point
- Peak shift down ~40 keV
- Resolution decrease ~2.5 - 3 %

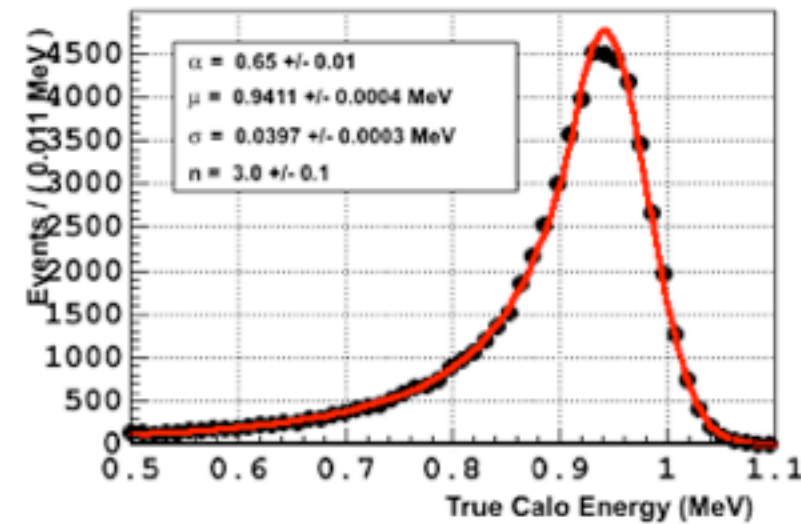


# Let's change the foil thickness (fit results)

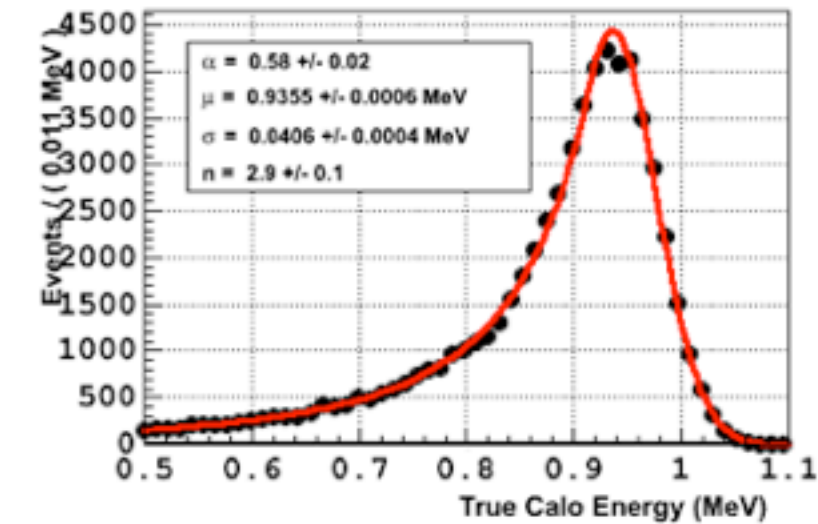
Electron energy lost [MeV], Thickness 10 % mg/cm2



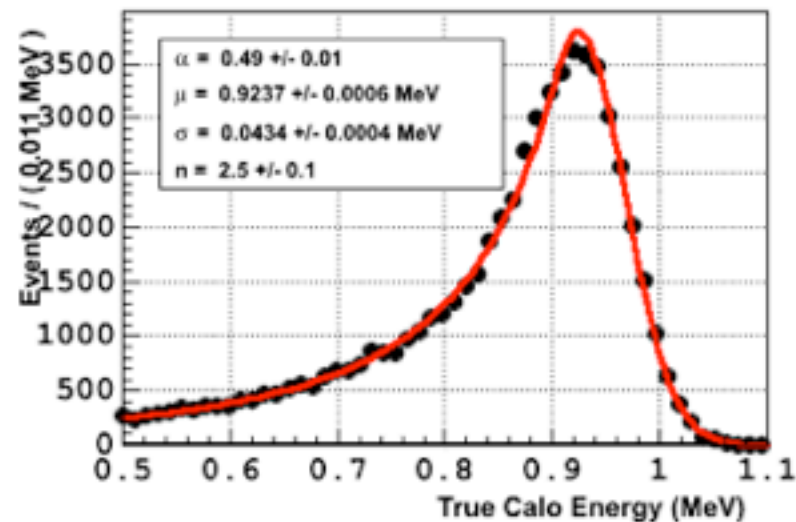
Electron energy lost [MeV], Thickness 20 % mg/cm2



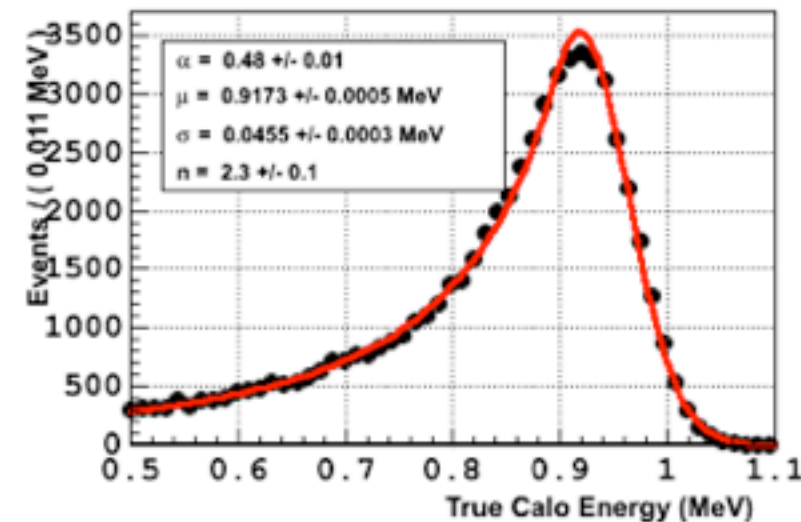
Electron energy lost [MeV], Thickness 30 % mg/cm2



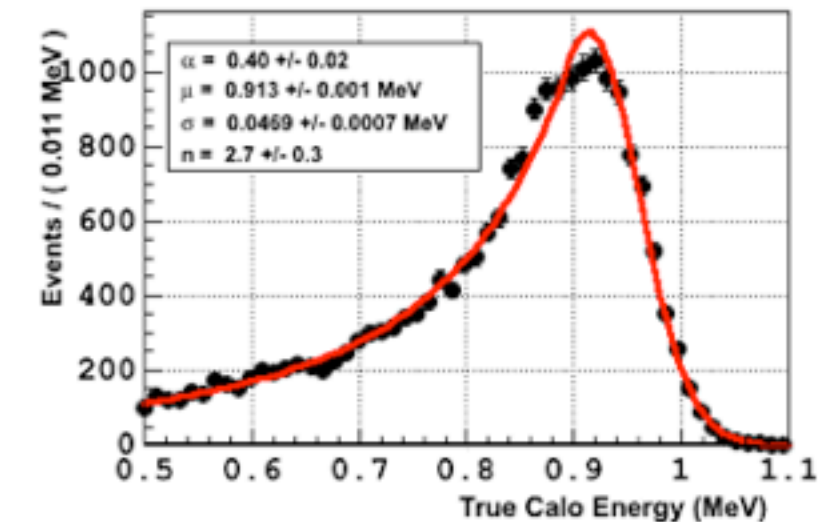
Electron energy lost [MeV], Thickness 50 % mg/cm2



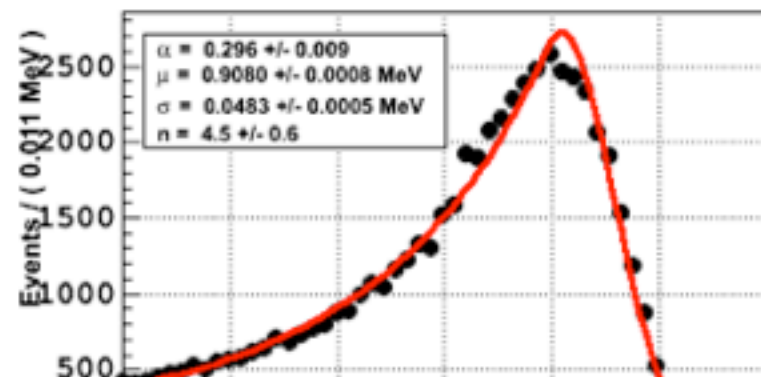
Electron energy lost [MeV], Thickness 60 % mg/cm2



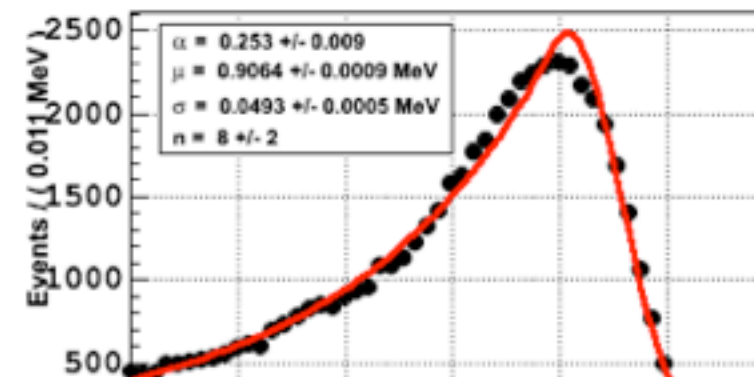
Electron energy lost [MeV], Thickness 70 % mg/cm2



Electron energy lost [MeV], Thickness 90 % mg/cm2



Electron energy lost [MeV], Thickness 100 % mg/cm2



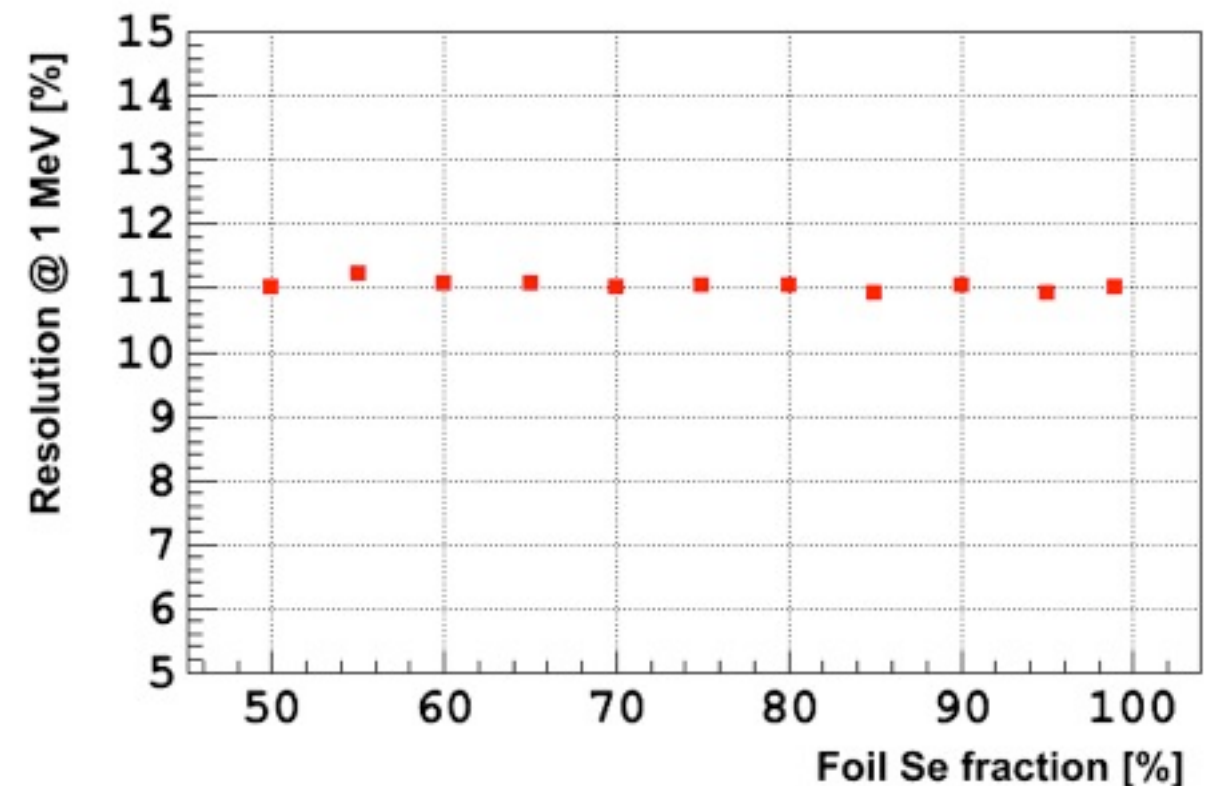
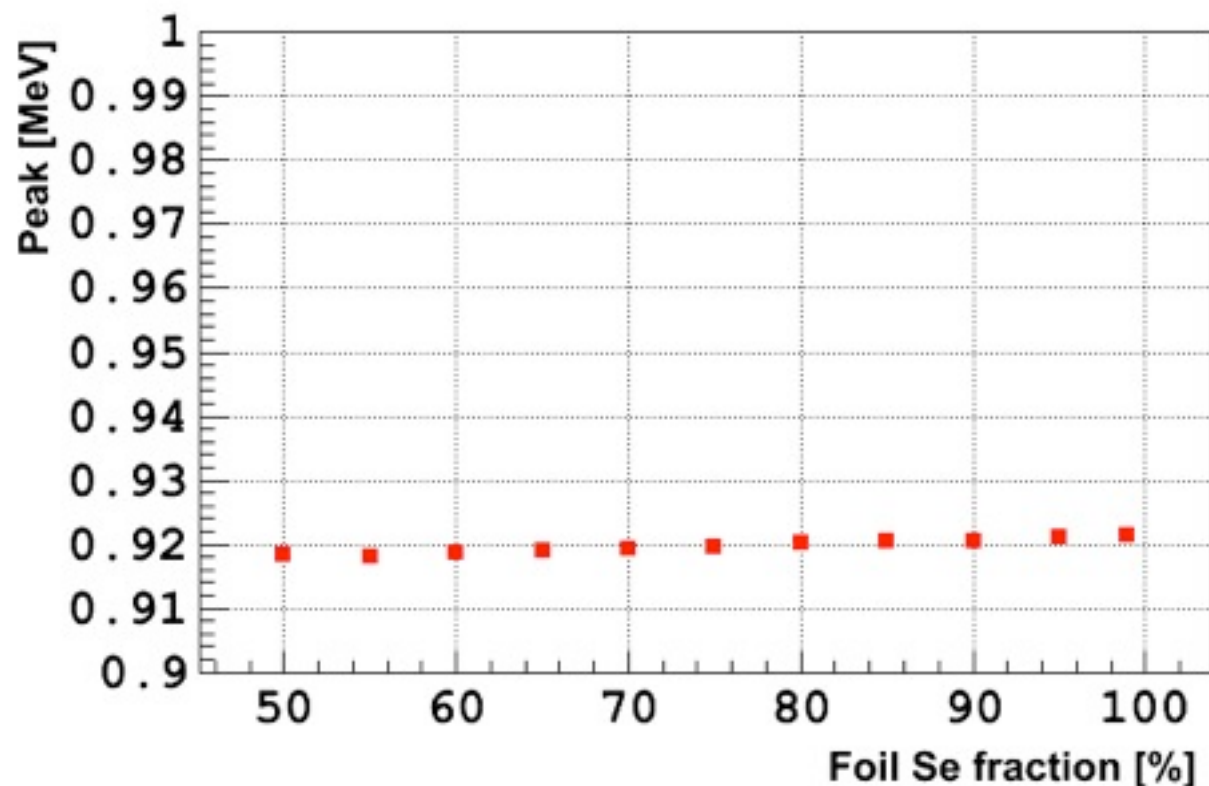


# Let's change the Se:PVA mix

[Se %]      [MeV]      [%]

frac *	peak *	100*(2.35 *	chi2 *
50 *	0.9185 *	11.02 *	3.049 *
55 *	0.918 *	11.23 *	2.877 *
60 *	0.9188 *	11.09 *	2.497 *
65 *	0.9191 *	11.06 *	2.061 *
70 *	0.9192 *	11 *	1.909 *
75 *	0.9196 *	11.06 *	3.268 *
80 *	0.9201 *	11.05 *	2.946 *
85 *	0.9205 *	10.93 *	2.58 *
90 *	0.9206 *	11.03 *	2.601 *
95 *	0.9211 *	10.93 *	3.597 *
99 *	0.9213 *	11.02 *	2.091 *

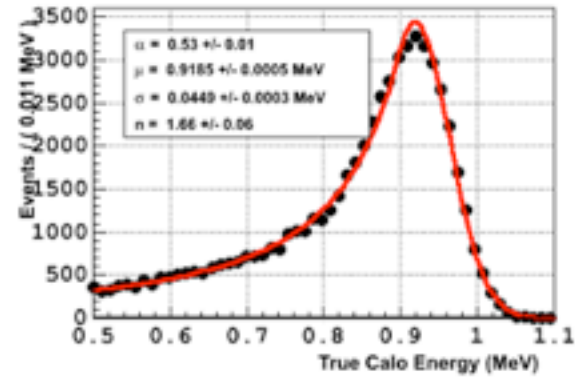
- Change total Se mass fraction & ecompute other components:
  - Pure  $^{82}\text{Se}$  & natural Se, assuming 3% contamination (Barabash)
  - PVA components (C, O, H)
- Very week dependence < 1 % variation



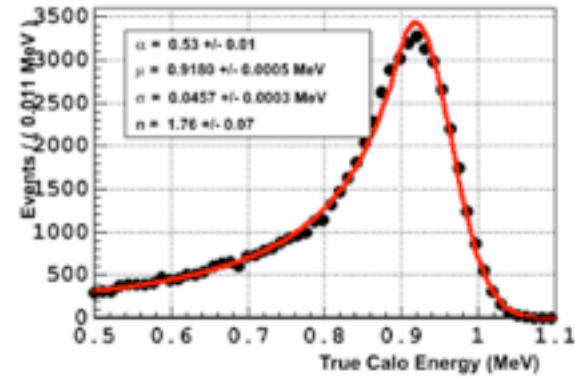


# Let's change the Se:PVA mix (fit results)

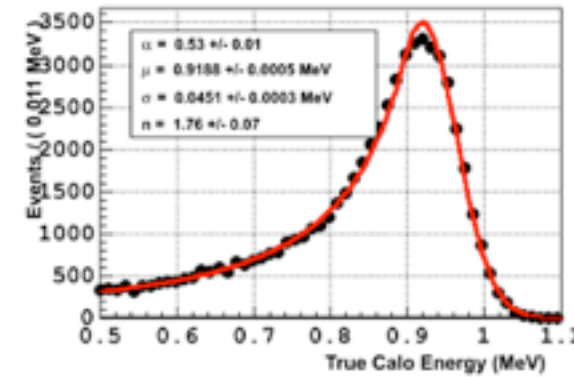
Electron energy lost [MeV], 50 % Se



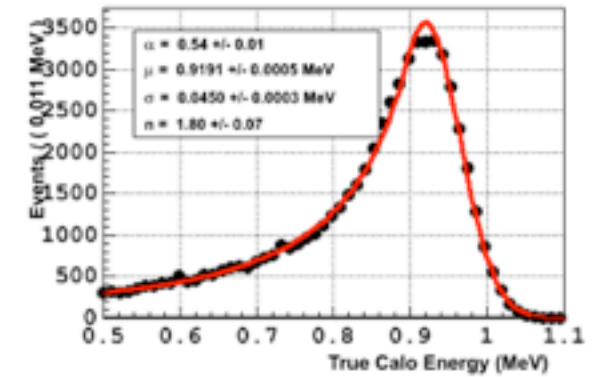
Electron energy lost [MeV], 55 % Se



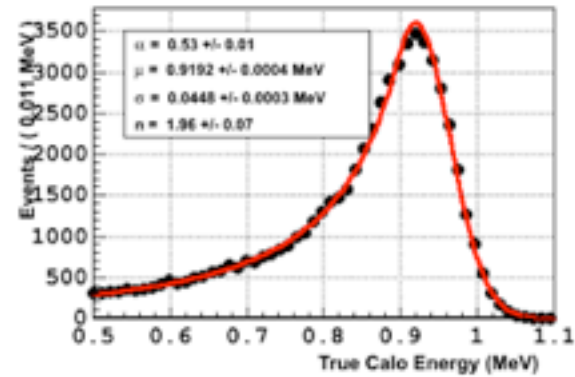
Electron energy lost [MeV], 60 % Se



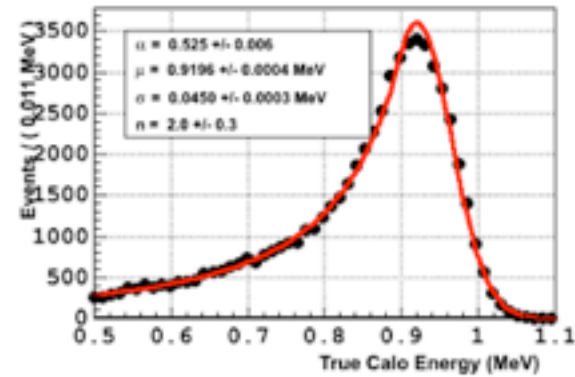
Electron energy lost [MeV], 65 % Se



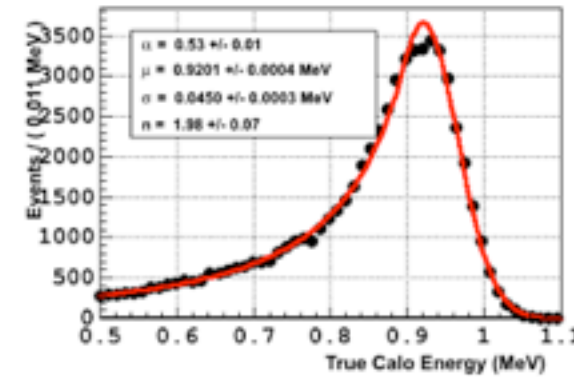
Electron energy lost [MeV], 70 % Se



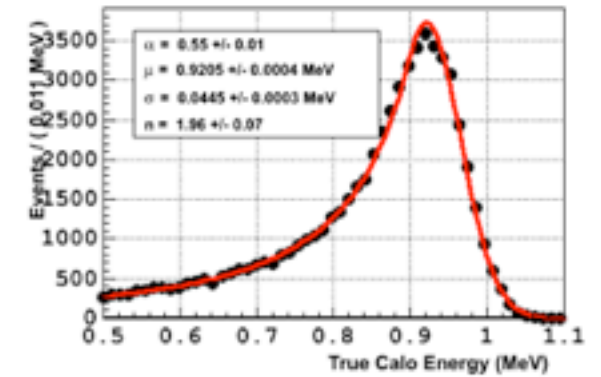
Electron energy lost [MeV], 75 % Se



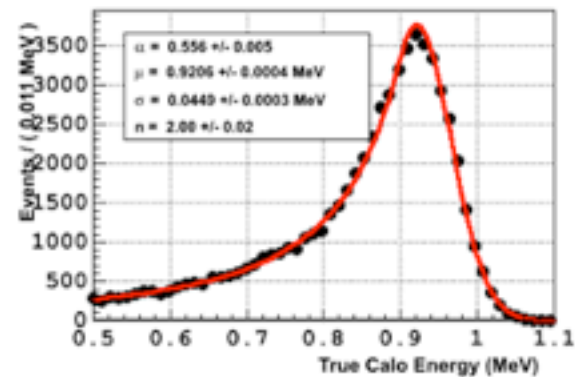
Electron energy lost [MeV], 80 % Se



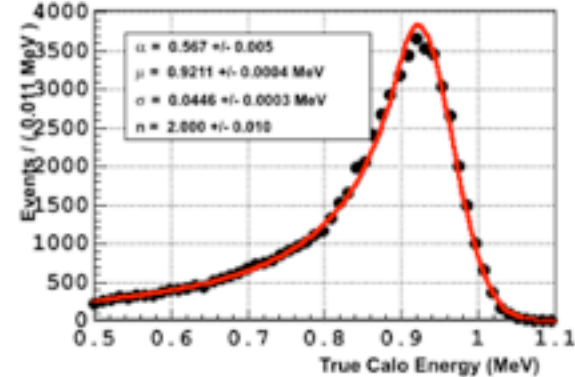
Electron energy lost [MeV], 85 % Se



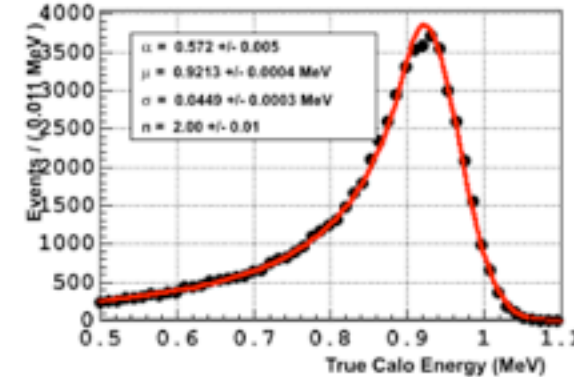
Electron energy lost [MeV], 90 % Se



Electron energy lost [MeV], 95 % Se



Electron energy lost [MeV], 99 % Se



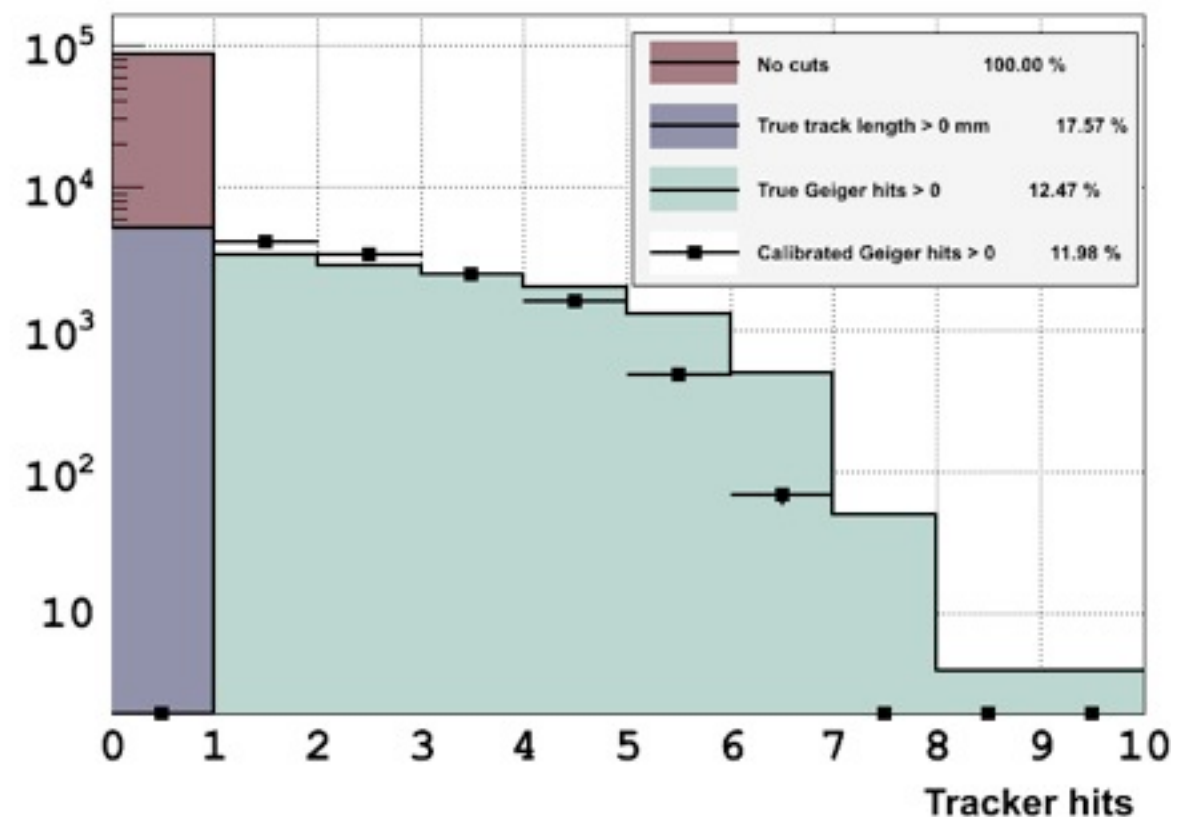
Alpha exit  
probability

# Alpha exit probability from source foil

Alpha are identified through their track length in the tracker gas

- Higher the exit probability
  - Higher the alpha detection efficiency
  - Better the measurement of internal  $^{214}\text{Bi}$  contamination

Cut	Prob. [%]
No cuts	100
Track length in gas > 0 cm	17.57
True geiger hit > 0	12.47
Calibrated geiger hit > 0	11.98



For comparison:

- NEMO3 got ~16% alpha detection efficiency
- Order of magnitude is correct
- Expecting higher value (maybe just too simple study?)

# Let's change the foil thickness

[mg/cm<sup>2</sup>]      [%]      [%]      [%]

thickness	$100 \cdot n_1 / n_0$	$100 \cdot n_2 / n_0$	$100 \cdot n_3 / n_0$
10	71.2	55.4	53.8
20	43.2	30.5	29.3
30	28.4	20.1	19.3
40	21.6	15.3	14.7
50	17.6	12.5	12
60	14.2	10	9.62
70	12.4	8.77	8.4
80	10.6	7.63	7.32
90	9.54	6.82	6.55
100	8.49	6	5.75

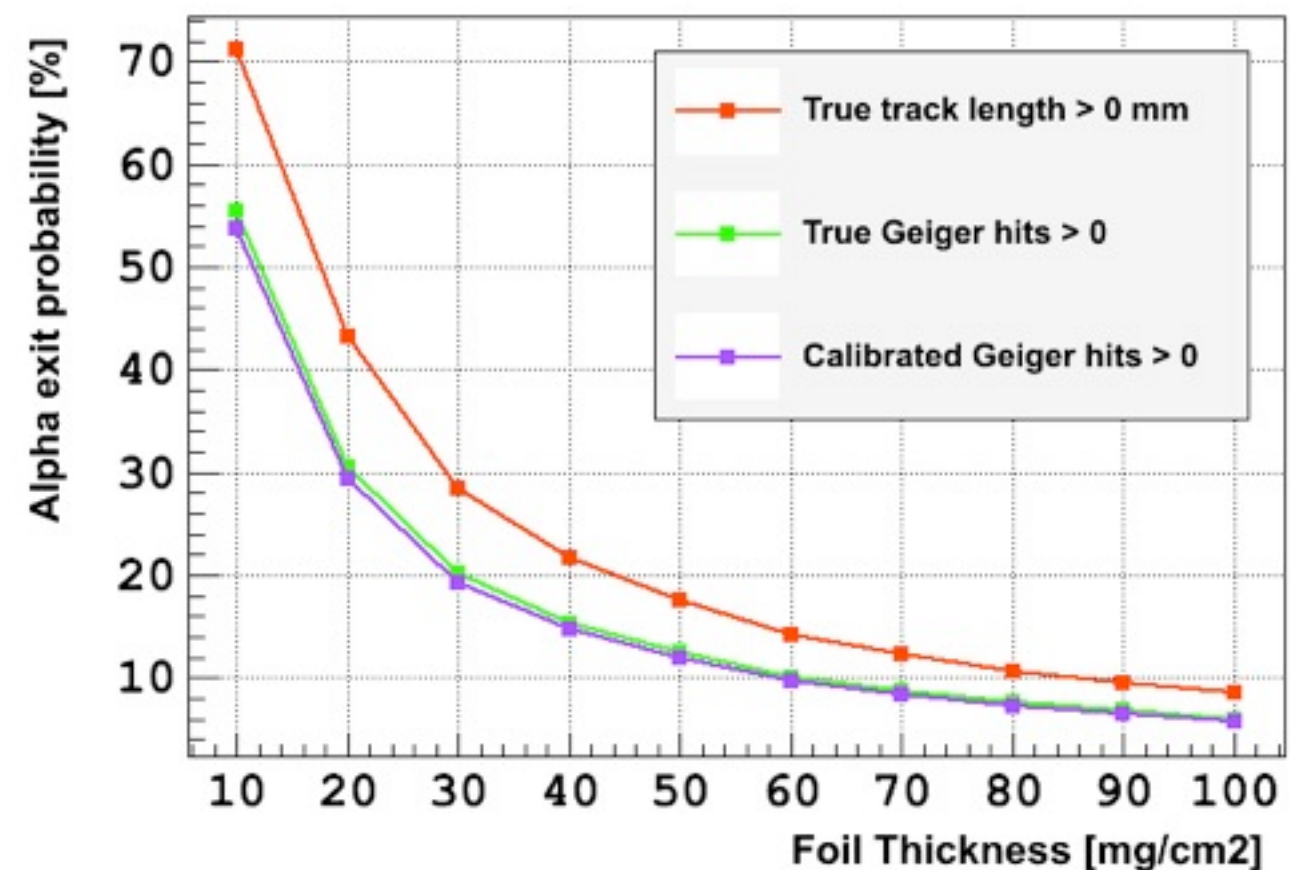
In the explored region:

- Alpha exit probability from ~55 % to ~5%
- 12 % for the target thickness

Keep in mind the current foil design:

- No mylar film
- 5 % PVA (while 20% is more likely)

Values are “upper limit”, ==> expected to lower down a bit



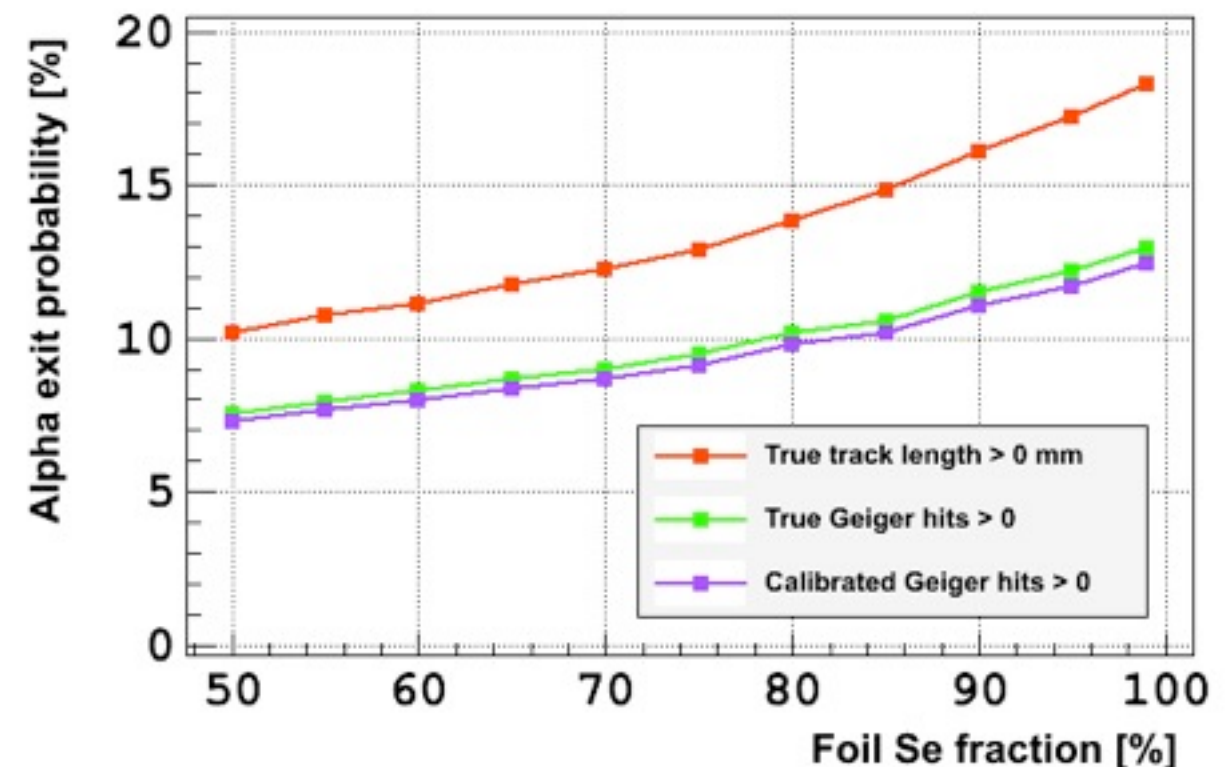


# Let's change the Se:PVA mix

[Se %]                      [%]                      [%]                      [%]

frac	* 100*n1/n0	* 100*n2/n0	* 100*n3/n0
50	10.2	7.51	7.25
55	10.7	7.92	7.62
60	11.1	8.28	7.98
65	11.7	8.67	8.33
70	12.3	8.99	8.64
75	12.9	9.48	9.1
80	13.8	10.2	9.81
85	14.8	10.6	10.2
90	16.1	11.5	11
95	17.2	12.2	11.7
99	18.3	12.9	12.4

- Change total Se mass fraction & recompute other components:
  - Pure  $^{82}\text{Se}$  & natural Se, assuming 3% contamination (Barabash)
  - PVA components (C, O, H)
- Slight increase of exit probability  $\sim 5.5\%$



- The foil source design is not frozen yet
  - Some details need to be investigated: foil composition, mechanical supports
- First very simple attempt in such direction
  - Use simple mono energetic particle gun
    - Concentrate on the detail of the detector response
  - Define few FOM related to the detector performance
    - Peak & FWHM of e- energy loss in calo (resolution)
    - Alpha exit probability from foil (Background detection efficiency)
  - Study such FOM w.r.t. source parameters: foil thickness, Se:PVA ratio

# Conclusions

- Need to modify source geometry to include mechanical support:
  - Mylar backing film
  - Mesh backbone
  - Others (?)
- Investigate the effect of the multiple scattering on the particle exit probability (alpha and electron)
- Consider to keep into account the total mass of the bb isotope

# Next steps