

Nikola Makovec

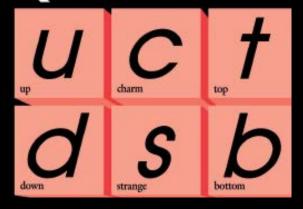


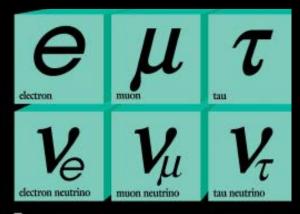
## Les acteurs de la session BSM

Speaker	Title
Ahmed BASSALAT	Overview of the Insertable B-Layer (IBL) Project of the ATLAS Experiment at the LHC
Zuzana BARNOVSKA	Search for a narrow resonance decaying to two photons with the ATLAS detector using 21 fb-1 of proton-proton collision data
Lucien HEURTIER	Extra U(1), effective operators, anomalies and dark matter
Vincent BIZOUARD	Desintegration du Higgs en Gamma Z dans le NMSSM
Marija MARJANOVIC	Analyse du canal SUSY-Olepton
Otilia Anamaria DUCU	Search for strongly-produced superpartners in final states with two same-sign leptons or three leptons with the ATLAS detector using 20 fb-1 of LHC pp collisions at 8 TeV
Alexandre Aubin	Recherche de stops dans l'experience CMS au LHC
Geoffroy Gilles	Recherche de nouveaux bosons de jauge lourds W' avec le détecteur ATLAS au LHC.
Mathieu Guigue	A la recherche de particules légères avec l'Hélium 3 polarisé
Yannick Stoll	Violation de saveur au delà du modèle minimal: Un tour d'horizon
Dorian Simon	Recherche de Nouvelle Physique dans les événements à quatre quarks top avec les détecteur ATLAS auprès du LHC



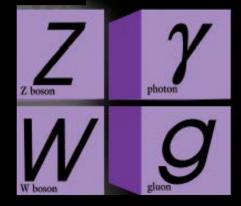
# Quarks





Leptons

# **Forces**



# Quarks **Forces** photon

Leptons

# The Standard Model of particle physics

Most general renormalizable lagrangian including all SM fields with  $SU(3)_{c} \times SU(2)_{L} \times U(1)_{y}$  gauge groups:

$$\mathcal{L}_{SM} = -\frac{1}{4g'^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{2g^2} \text{Tr}(W_{\mu\nu} W^{\mu\nu}) - \frac{1}{2g_s^2} \text{Tr}(G_{\mu\nu} G^{\mu\nu})$$

$$+ \bar{Q}_i i \not\!\!\!D Q_i + \bar{L}_i i \not\!\!\!D L_i + \bar{u}_i i \not\!\!\!D u_i + \bar{d}_i i \not\!\!\!D d_i + \bar{e}_i i \not\!\!\!D e_i$$

$$+ (Y_u^{ij} \bar{Q}_i u_j \tilde{H} + Y_d^{ij} \bar{Q}_i d_j H + Y_l^{ij} \bar{L}_i e_j H + \text{h.c.})$$

$$+ (D_{\mu} H)^{\dagger} (D^{\mu} H) - \lambda (H^{\dagger} H)^2 - \mu^2 H^{\dagger} H$$

$$+ \frac{\theta}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(G_{\mu\nu} G_{\rho\sigma})$$

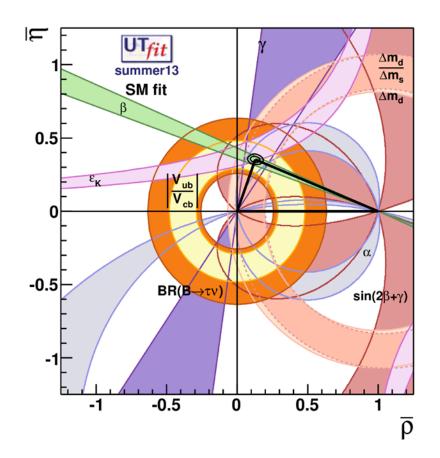
19 parameters:

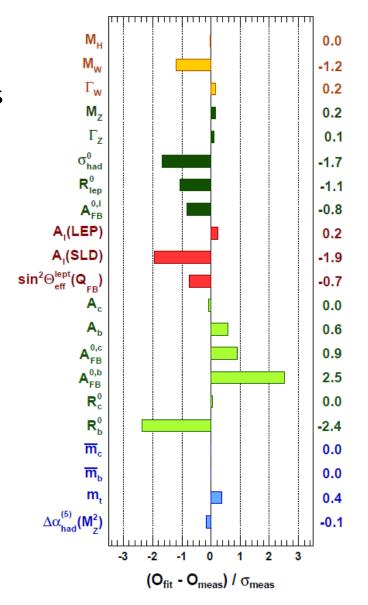
- 3 gauge coupling constants
- 9 fermion Yukawa couplings
- 3 CKM mixing angles + 1 phase
- $\mu,\lambda$  or  $m_7,m_H$
- $\theta_{strong}$



## **Success**

- Electroweak measurements at LEP, SLD, and Tevatron
- CP-violation at B-factories, K-factories

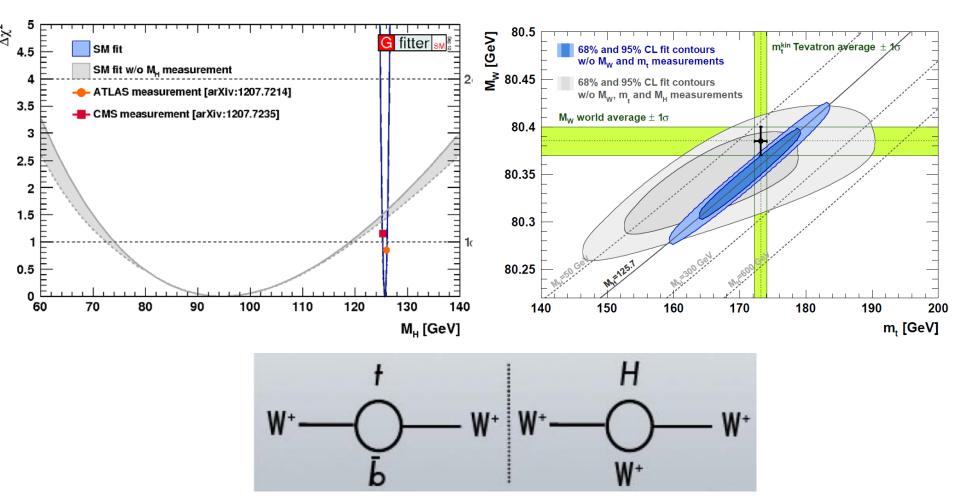






## Latest success: Higgs discovery

 LHC Higgs boson discovery at ~126 GeV is compatible with precision measurements





# Unanswered questions?



## Experimental evidence not explained by SM

- Why do neutrinos have mass yet so light?
  - Observed neutrino masses imply new physics (at least, right-handed neutrinos) somewhere between 1 keV and 10<sup>15</sup> GeV)
  - Add 7(+2) parameters to the SM
- What is the origin of matter anti-matter asymmetry in Universe?
  - Domination of matter over anti-matter requires new physics between 100 GeV and 10<sup>16</sup> GeV
- What is dark matter?
  - Existence of dark matter requires new physics (at least one new stable particle) somewhere between sub-eV and 10<sup>19</sup> GeV
  - What is the physics which reconciles gravity and quantum mechanics?
- What is the physics which reconciles gravity and quantum mechanics?
  - New physics expected (at least) at energies ~ 10<sup>19</sup>GeV!



# Why things are the way they are?

#### The Higgs mass term:

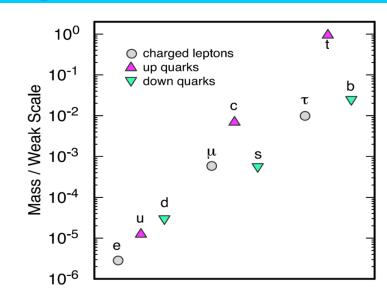
- Why the gravity is so weak?
- $M_{FW}^2 \sim 10^4 \text{GeV}^2 < < M_P^2 \sim 10^{38} \text{ GeV}^2$

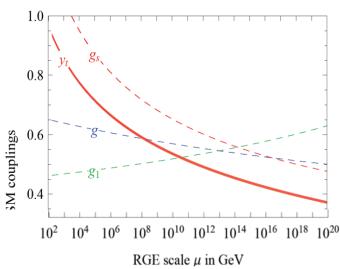
#### Charge quantization:

- Why  $|Q_p + Q_e| < 10^{-21}Q_e$ ?
- The strong CP problem:
  - Why  $\theta_{\text{strong}} < 10^{-13}$
- Number of families:
  - Why are there three families?

#### Fermion masses:

- Why  $m_{top}/m_e \sim 3 \times 10^5$ ?
- Gauge coupling unification:
  - Is there a unified description of all forces at higher energy?

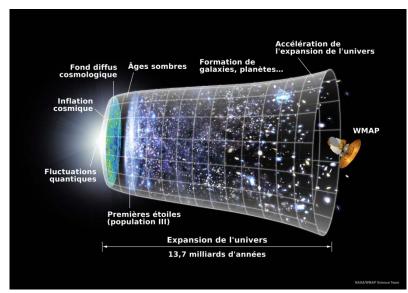


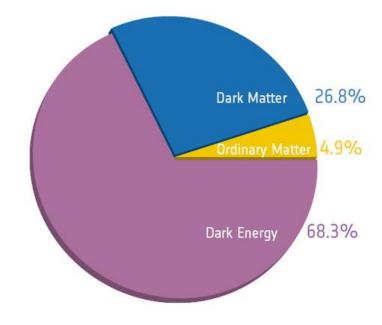




## Cosmology and particle physics

- What is the dark energy?
- Why is the cosmological constant small but non zero?
- Why now? (Cosmic coincidence problem)
- Why is the Universe so big? (flatness problem, horizon problem)
- What is the physics which underlies inflation?
  - Link with Higgs?

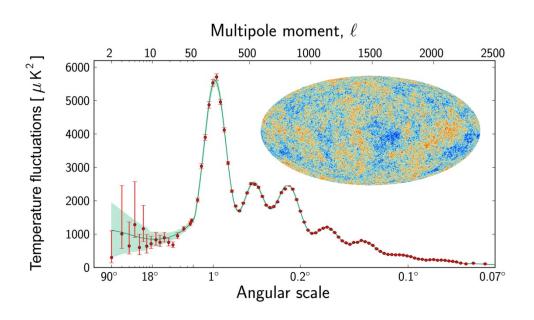


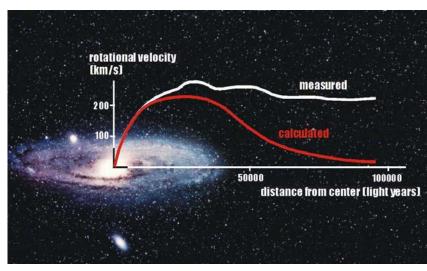


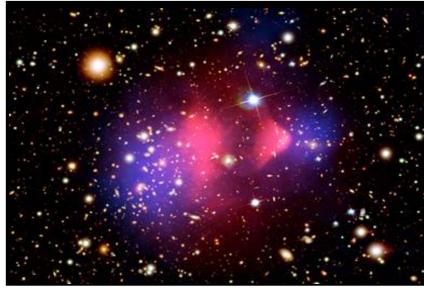


## **Dark matter**

- Dark matter does not emit or reflect sufficient electromagnetic radiation to be detected.
- ~ 80% of matter is Dark Matter
- The most compelling candidate is a WIMP for Weakly Interacting Massive Particle



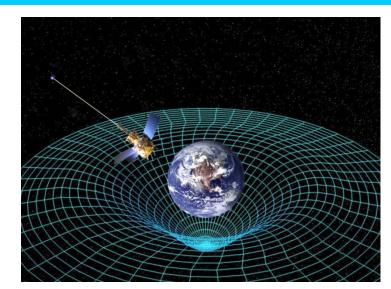


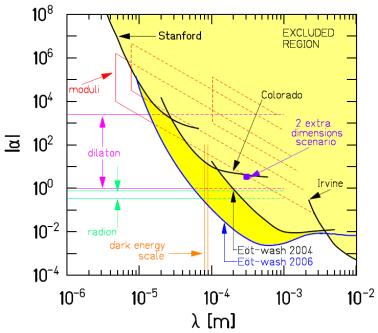




# **Gravity**

- Described successfully by the **General** relativity but:
  - No acceptable quantum theory of gravity
  - Inverse-square law of gravity only down to distances just shorter than 0.1 mm, corresponding to energies of 0.01eV
- Associated with a dynamical symmetry group (symmetry of the spacetime), other gauge theories are built on internal symmetries.
- Very weak compared to the three other forces
  - Suppressed E/M<sub>planck</sub> with M<sub>planck</sub>=1.22 10<sup>19</sup>GeV
- New physics expected at least at energies ~ 10<sup>19</sup>GeV!









- Why the gravity is so weak?
  - $M_{FW}^2 \sim 10^4 \text{GeV}^2 < < M_P^2 \sim 10^{38} \text{ GeV}^2$
- At the quantum level, scalar masses are extremely sensitive to heavy states

$$m_H^2 = \frac{h}{m_H^2} + \frac{h}{m_{bare}^2} + c.M_p^2$$

- Strong dependence of physics at the electroweak scale on the physics at the Planck scale
  - It's like saying that to describe the rotation of the moon around the earth one needs to know about the quarks inside the proton



$$m_H^2 = m_{bare}^2 + c.M_p^2$$

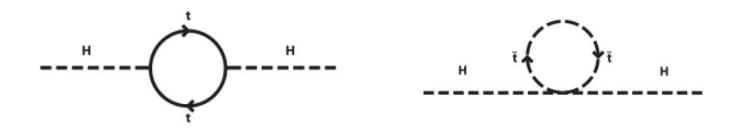


$$m_H^2 = 125^2 = 36,127,890,984,789,307,394,520,932,878,928,933,023$$
  
- 36,127,890,984,789,307,394,520,932,878,928,917,398



$$m_H^2 = 125^2 = 36,127,890,984,789,307,394,520,932,878,928,933,023$$
  
 $-36,127,890,984,789,307,394,520,932,878,928,917,398$ 

- Extreme fine tuning of parameters!!!!
- Different situation for fermions or gauge bosons
   ⇒ gauge symmetries protect them
- Solution to the hierarchy problem lead to new physics at the weak scale

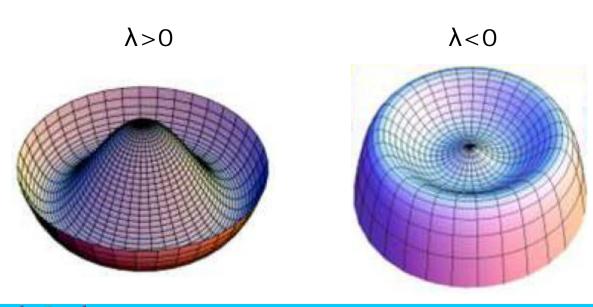


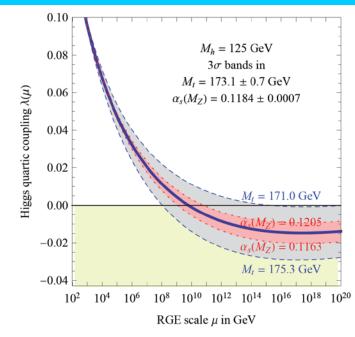


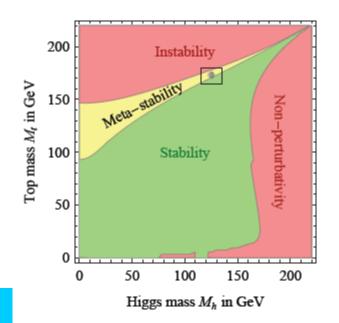
## Fate of the EW vacuum

- Higgs self-coupling (λ) energy variation deduces from renormalization group evolution
- EW vacuum unstable if λ < 0</li>

$$V(\phi) = \mu^2 \left| \phi \right|^2 + \lambda \left| \phi \right|^4$$



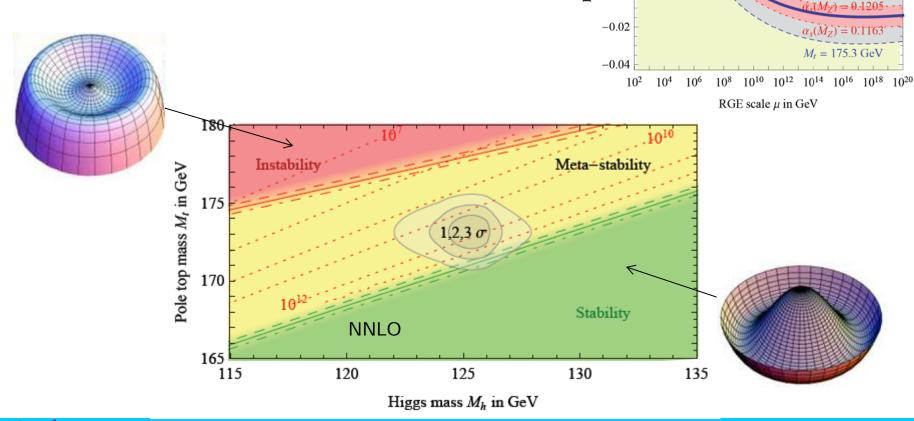






#### Fate of the EW vacuum

- Assuming SM valid up to M<sub>p</sub>
   ⇒ meta stable @ 2σ
- Or stable vacuum if new physics appears  $\Lambda = 10^{10} \text{GeV...}$



0.10

0.08

0.06

0.04

0.02

0.00

Higgs quartic coupling  $\lambda(\mu)$ 

 $M_h = 125 \; \text{GeV}$ 

 $3\sigma$  bands in  $M_t = 173.1 \pm 0.7 \text{ GeV}$ 

 $\alpha_s(M_Z) = 0.1184 \pm 0.0007$ 

 $M_t = 171.0 \text{ GeV}$ 



## **BSM** models

Many BSM models developed to answer Standard Model limitations. For instance:

## Supersymmetry:

 Add a new broken symmetry to SM to protect Higgs mass

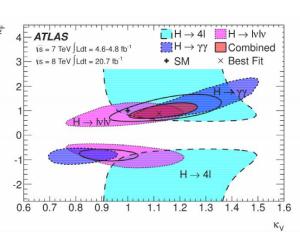
## Composite Higgs:

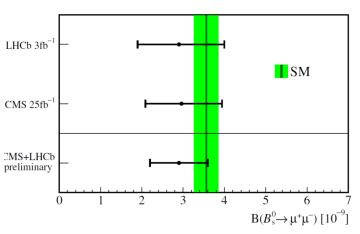
- The Higgs is not elementary, first manifestation of a new strong force
- Large extra dimensions:
  - addresses Hierarchy Problem by bringing the Plank scale down to TeV

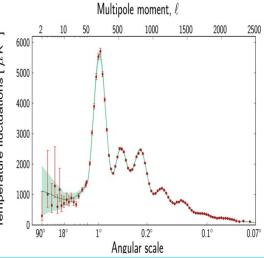


# Various strategies to track new physics

- Multi-pronged approach to search for new physics
  - Direct Searches at high energy colliders
  - Precision Measurements (need good theoretical control)
  - Rare decays (K & B) and Forbidden Processes
  - Neutrino physics
  - Cosmological observations
- Beyond Standard Model physics not only in this session!!!

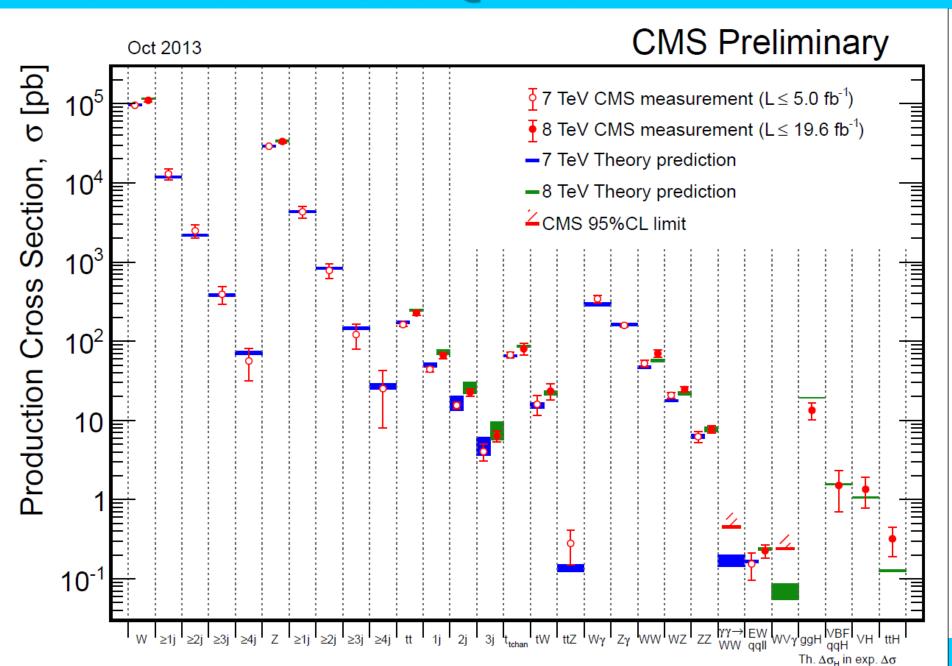








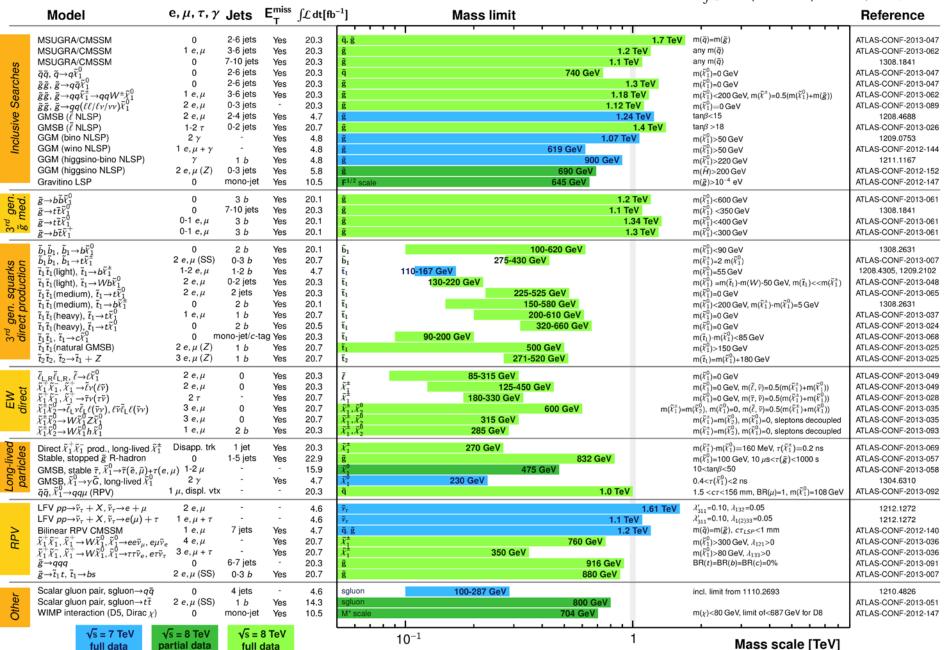
## The Standard Model @ LHC

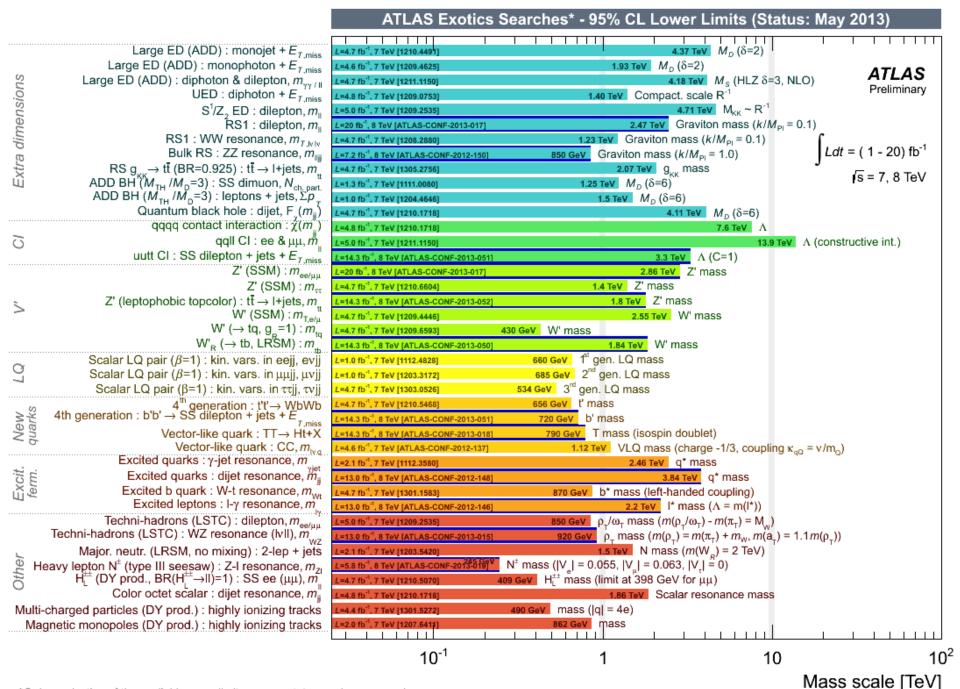


Limits, limits, limits Observed limit (±1  $\sigma_{theory}^{SUSY}$ ) o acc ATLAS Preliminary 'Stable' ~  $\overbrace{\tilde{g}}^{\tilde{g}} \widetilde{g} \text{ production, } \widetilde{g} \rightarrow \widetilde{t}\widetilde{t}, \widetilde{1} \text{ (RPV)} \rightarrow \text{bs}^{\frac{\widetilde{g}}{g}}$ **ATLAS** dt = 20.3 fb 1, 15 5 1.1 fb<sup>-1</sup>, \s=8 TeV  $\tilde{\chi}^0_s \rightarrow W^{\pm} \tilde{\chi}^0_s h^0 \tilde{\chi}^0_s$  $\sqrt{s} = 8 \text{ TeV}, \int L dt = 20.3 \text{ fb}^{-1}$ ATLAS |L dt = 20.3 fb-1 Multijet Combined Continued 7 Test Studenters ted limit (±1  $\sigma_{exp}$ ) 8 1600 ATLAS Preliminary Observed limit ± 1 of theory ved limit (±1 σ<sub>theory</sub>) - 0-1 + 3 b-jets, 4.7 fb<sup>-1</sup>, 7 TeV L<sup>et</sup> = 20.1 fb<sup>-1</sup>, (s=8 TeV ATLAS Preliminary ---- Expected limit ±1 σ<sub>exp</sub> - 0-1 + 3 b-jets, 12.8 fb<sup>-1</sup> Ldt = 20.3 fb<sup>-1</sup>, (s=8 TeV 140 160 180 200 220 240 260 28 Ldt = 20.7 fb 1, 16=8 TeV -ATLAS Prefimina ATLAS 1400 ATLAS Preliminary ATLAS Preliminary 15.9 fb<sup>-1</sup> Ld = 20.3 8" (6-8 Tel tion,  $\tilde{g} \rightarrow t \tilde{\chi}_{1}^{0}$ ,  $m(\tilde{t}) >> m(\tilde{g})$ ATLAS Preliminary 15 = 8 TeV | 1dt - 15 0 fh 5 500 ATLAS Preliminary Lat = 20.3 fb 1, vs=8 TeV  $\widetilde{\chi}, \widetilde{\chi}, \rightarrow 2 \times \widetilde{h}(\widetilde{v}) \rightarrow 2 \times h$ χ̄<sup>0</sup> mass [GeV] 0 60 80 Exp Limit

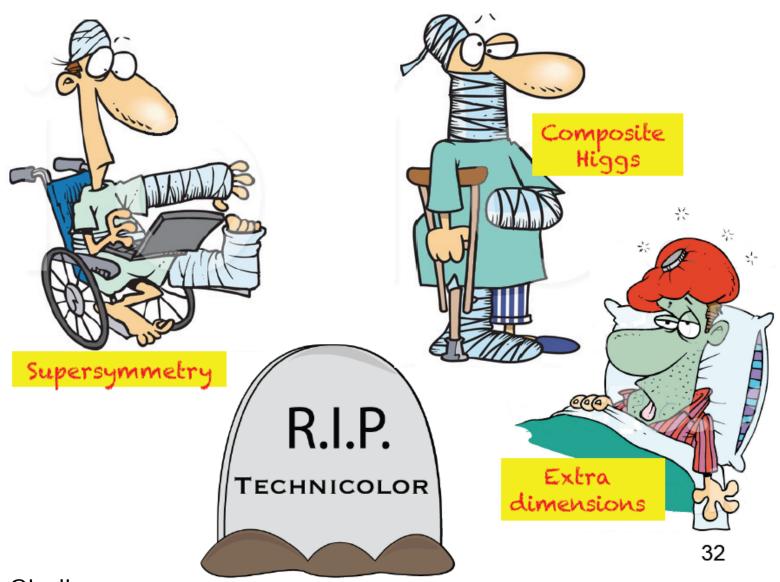
±1 σ Exp Limit 200 ATLAS Prefi gg Cross-Section (NLO+NLL) € 10<sup>3</sup> ATLASPINI TLAS Preliminary Chrystope Smit ± 1 of the BR(t)=0%, BR(b)=0%, BR(c)=50% roduction,  $\tilde{t}_1 \rightarrow c + \tilde{\chi}$  $\int L dt \sim 20.3 \text{ fb}^{-1}, 1 \text{ s} = 8 \text{ TeV}$ 17LAS Preliminary | L dt = 20.3 fb.1, is=8 TeV Ldt = 20.3 fb 1 Ns = 8 TeV 100 ATLASP ATLAS Preliminary ted limit ±1 a.... is = 8 TeV, Ldt = 15.9 fb<sup>-1</sup> biets 4.7 fb<sup>-1</sup>.7 Te ATLAS Preliminary 0.1 fb<sup>-1</sup>, \s=8 at 95% CL F 2.65 fb 300 ATLAS Prelimina 5.2 fb<sup>-1</sup> Simplified model,  $\tilde{g}\tilde{g} \rightarrow q\bar{q}q\bar{q} \tilde{\chi}_{1}^{1}\tilde{\chi}_{1}^{1} \rightarrow q\bar{q}q\bar{q} W^{1}W^{1}\tilde{\chi}_{1}^{0}\tilde{\chi}_{2}^{1}$ LAS 2.05 f Status: SUSY 2013

 $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \qquad \sqrt{s} = 7, 8 \text{ TeV}$ 





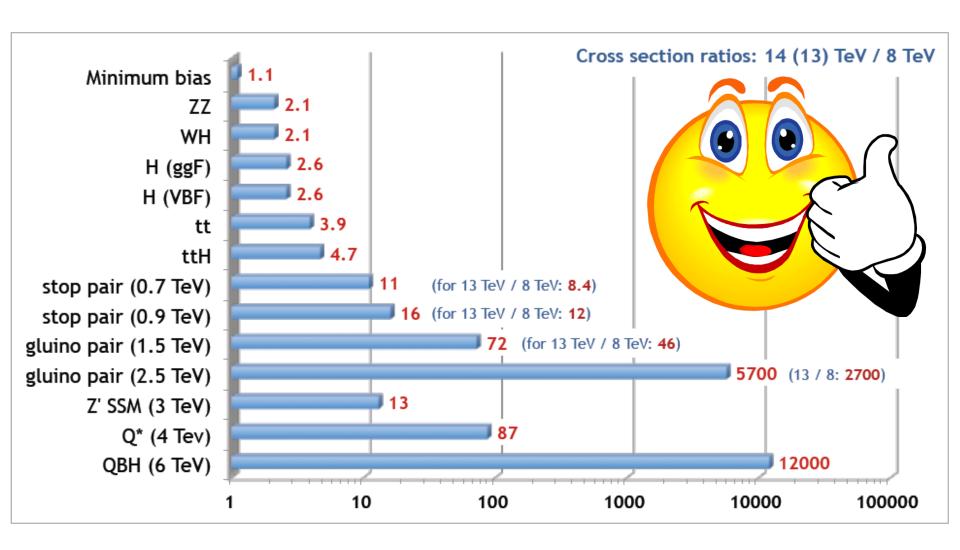
# LHC run I summary



G.F. Giudice



# LHC run II perspective





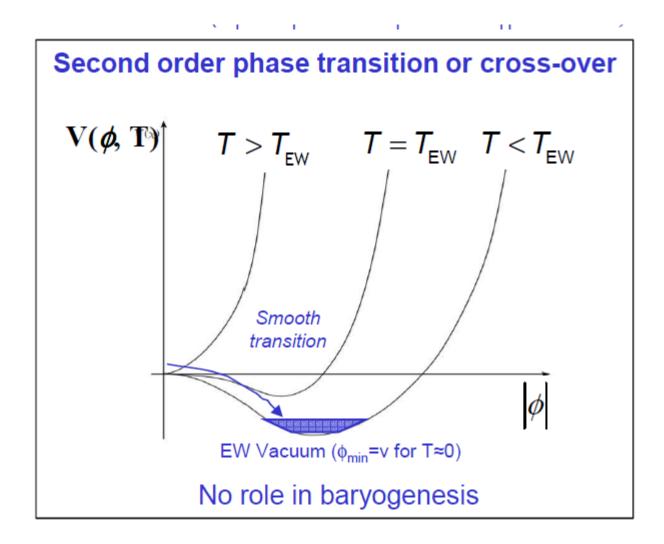
# Sources of inspiration

- Unanswered Questions in the Electroweak Theory, C. Quigg, hep-ph/0905.3187
- Beyond the Standard Model, A. Pomarol, hep-ph/1202.1391
- Physics Beyond the Standard Model and Dark Matter, H. Murayama, hep-ph/0704.2276
- Nouvelles Physiques by Henri Bachacou and Adam Falkowski, Lecture @ Ecole de Gif 2012 (slides only)



# **Backup**







### Fate of the EW vacuum

- Higgs self-coupling (λ) energy variation deduces from renormalization group evolution
- EW vacuum unstable if λ < 0</li>

