

GDR Neutrino Meeting 2013



CUORE-0

Matteo Biassoni on behalf of the CUORE Collaboration Paris, May 22, 2013

OUTLINE

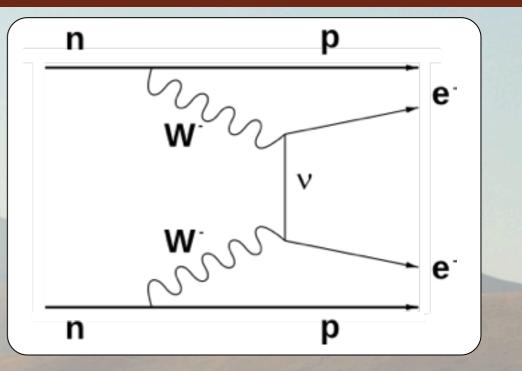
- Neutrinoless Double Beta Decay (theoretical and experimental point of view)
- the bolometric technique and the $0\nu\beta\beta$ research with TeO₂
- CUORE features and construction status
- CUORE-0: CUORE first tower
- CUORE-0: sensitivity and preliminary results
- conclusions

0νββ



 $(A,Z) \longrightarrow (A,Z+2) + 2e^{-}(+2v)$

- Channel for $\beta\beta$ decay forbidden by SM (Δ L=2)
- Extremely rare process $(T_{1/2} > 10^{22} 10^{24} y)$
- Never observed (but Ge⁷⁶ claim ^[1])
- Its observation would prove v Majorana nature



For light v_M exchange the Decay Rate is:

 $(T_{0\nu})^{-1} \propto G_{0\nu}(Q,Z) |M^{0\nu}|^2 \langle m_{ee} \rangle$

Phase Space Factor ATOMIC PHYSICS

Nuclear Matrix Element
NUCLEAR PHYSICS

Effective Majorana mass PARTICLE PHYSICS

The observation of 0vββ:

Is currently the only feasible method to establish:

- Majorana nature of the neutrino
 - Lepton number violation

0

0

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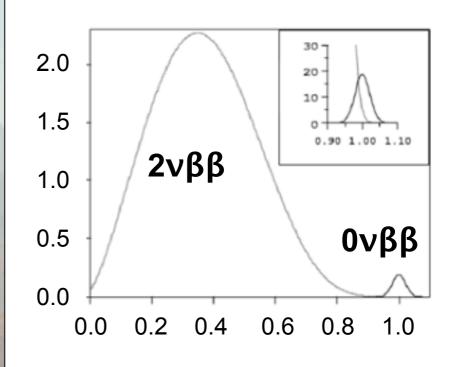
Can give important information about:

- Absolute neutrino mass scale
 - Neutrino mass hierarchy
 - **CP** Majorana phases

Experimental search



Main signature: 0vββ exhibits a peak at Q over 2vββ tail enlarged only by detector resolution



Defining the **experimental sensitivity** $S^{0\nu}$ as the lifetime corresponding to the minimum detectable number of events over background at a given C.L.

M: total active mass in kg
e: detector efficiency
a.i.: isotopic abundance
b: background in c/keV/kg/y
ΔE: detector resolution @ ROI in keV
T: total live time in y

$$S^{0\nu} \propto \frac{\epsilon \ a.i.}{A} \left(\frac{MT}{b \ \Delta E}\right)^{1/2}$$

Qualitative expression in the Gaussian approximation (not fully accurate for very low background experiments)

 $b \neq 0$

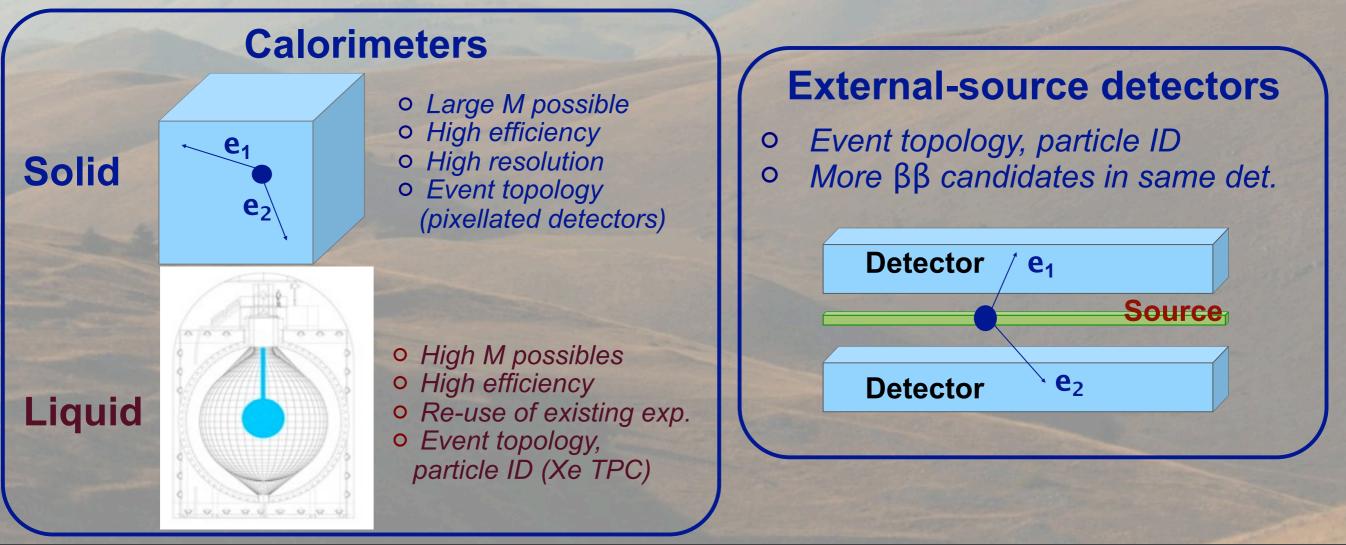
Experimental strategies



Additional signatures can be looked for:

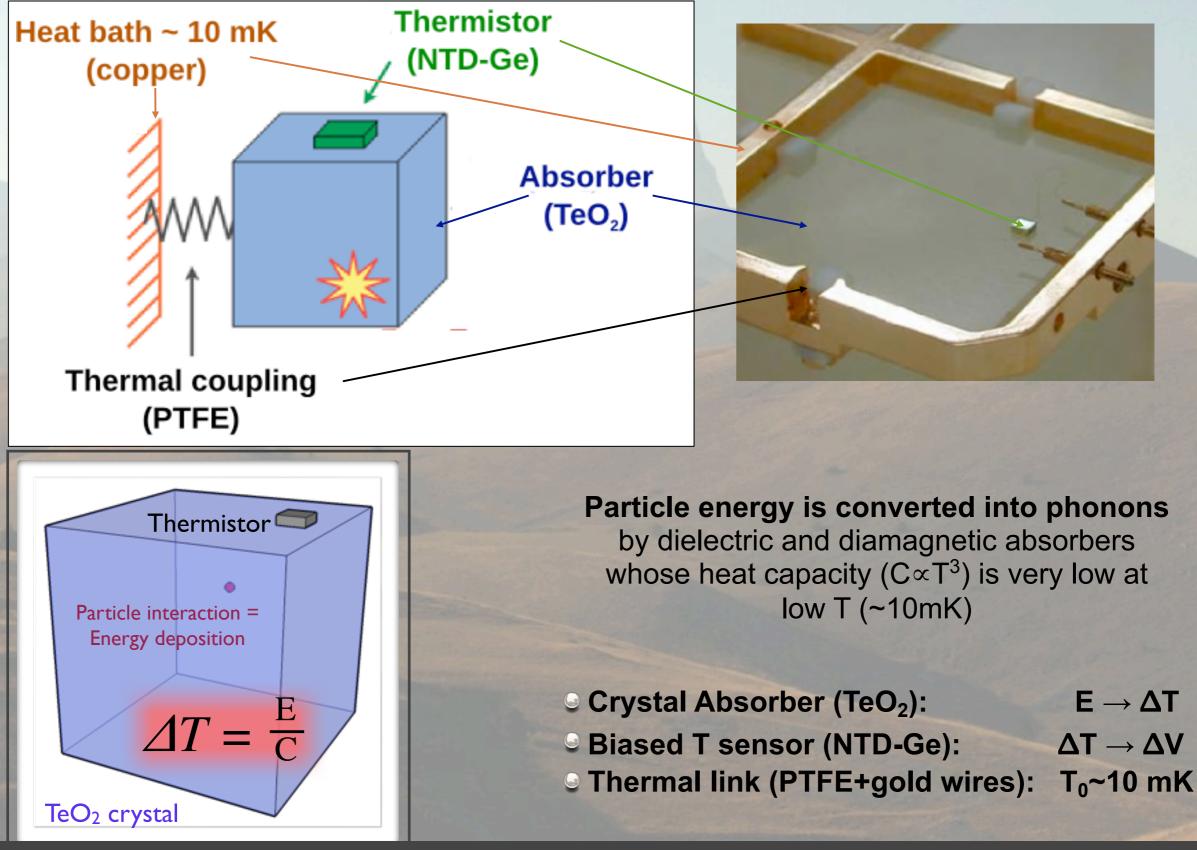
Single electron energy spectrum Angular correlation between the two electrons Track and event topology Time Of Flight Daughter nuclear specie

Two main approaches: calorimetric (source = detector) or external-source detector



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Bolometric technique

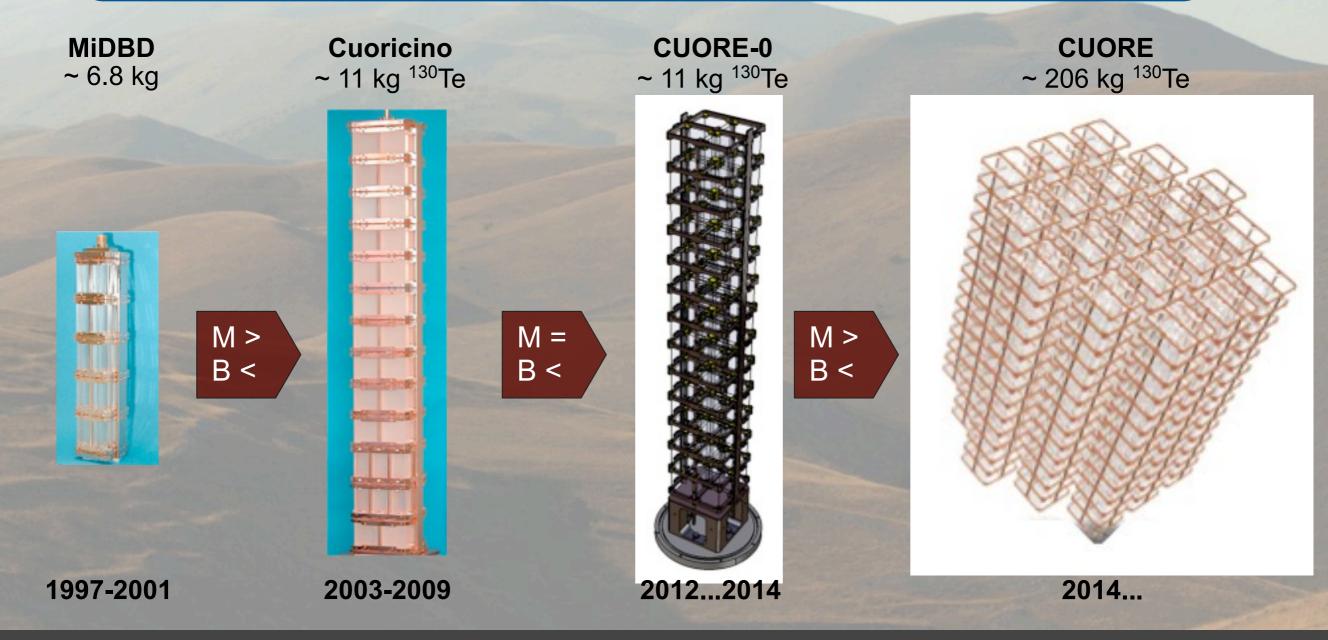


$0\nu\beta\beta$ research with TeO₂

130Te is a good DBD candidate (130 Te $\rightarrow ^{130}$ Xe + 2 e⁻) with high natural i.a. (34.2 %) and reasonably high Q-value (Q~2528 keV) leading to high G(Q,Z) and low background

TeO₂ is a compound with **good mechanical and thermal properties** containing ¹³⁰Te

 \checkmark 5x5x5 cm³ TeO₂ crystals have a high detection efficiency for 0vββ events: ~87.4%



Background reduction

Passive methods adopted for CUORE while testing different active methods (i.e. Surface sensitive bolometers, scintillating bolometers) for future improvements

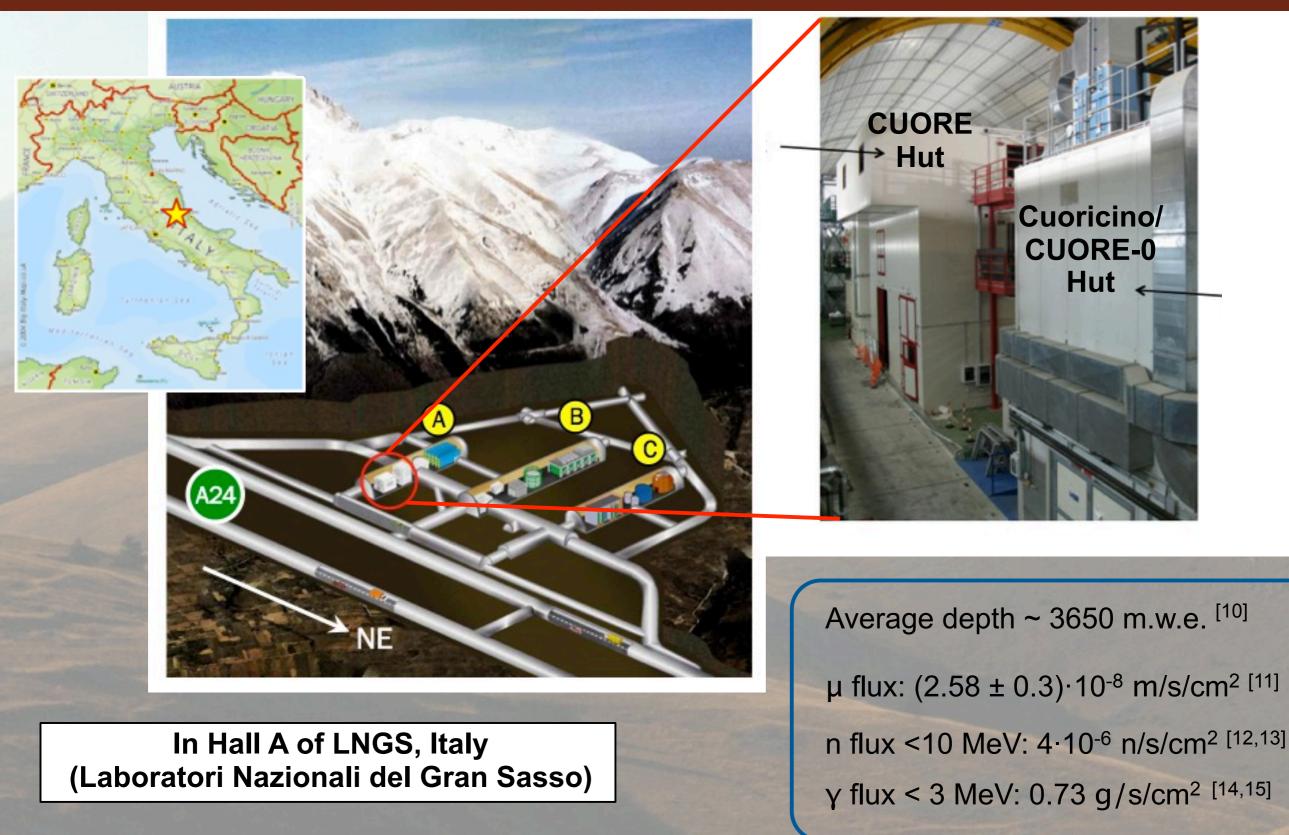
- Shields design and materials selection
- New holder design to reduce the amount of copper facing the crystals
- TeO₂ crystals bulk contamination control: strict protocol for TeO₂ production ^[20]
- Crystals surface contamination reduction: new treatment developed
 - => bolometric tests on 4 sample crystals from each batch: CCVR^[9] tests
- Surface contamination of the copper facing the crystals reduction:
 - => bolometric tests of three different surface treatments: Three Tower Test (TTT)
- Further improvement thanks to detector granularity

CUORE status



Feature	Status
988 TeO ₂ crystals, 19 towers - total mass 741kg (206kg ¹³⁰ Te)	>95% stored underground @LNGS (or already assembled)
Ultra-clean copper structure/ shields	>50% stored underground @LNGS (or already assembled), delivery by end 2013
Thermistors	production complete, under testing
Dedicate hut and clean room	fully equipped
Detector assembly line	ready, being used for the assembly of the first towers
Calibration system	construction started
Ultra-clean cryostat to keep >1ton @10mK with high duty- cycle	commissioning of first 3 shields on-going at LNGS, commissioning of dilution unit completed (8muW@10mK)
Radon abatement system	installed
DAQ and Analysis tools	are being tested with CUORE-0

CUORE(-0) location



CUORE-0



1 CUORE-like tower of 13 planes - 4 crystals each 52 TeO₂ 5x5x5 cm³ crystals (750 g each) Detector Mass: 39 kg TeO₂
¹³⁰Te mass (natural i.a.): 11 kg of ¹³⁰Te

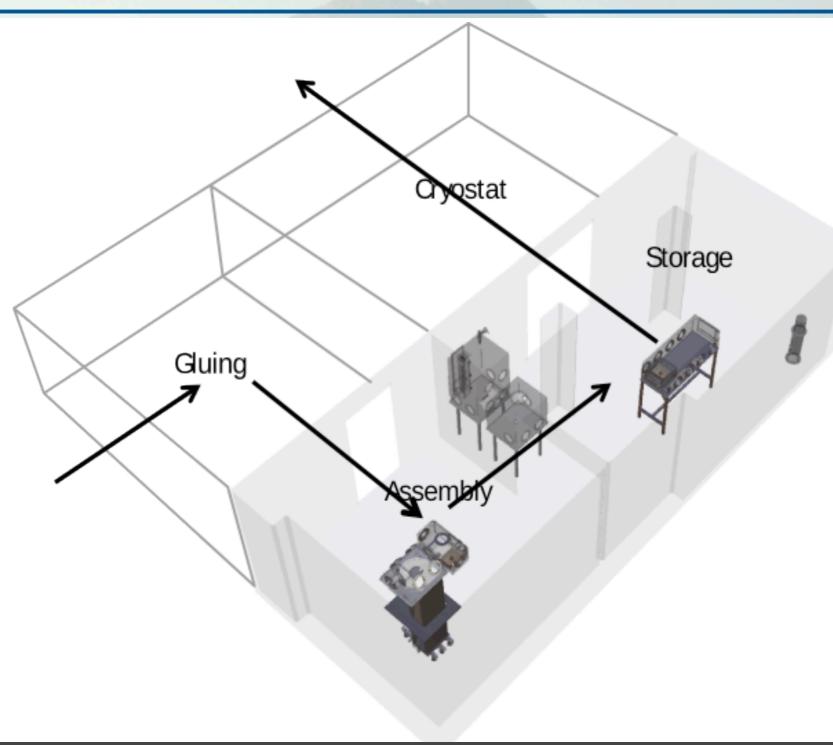
- All detector components manufactured, cleaned and stored with protocols defined for CUORE
- Assembled with the same procedures foreseen for CUORE
- In the 25 years-old CUORICINO cryostat

GOALS:

- Proof of Concept for CUORE in all stages
- Test and debug the CUORE assembly line (thermistor gluing, signal wires bonding, tower assembly)
- Test of the CUORE DAQ and analysis framework
- Extend the physics reach beyond CUORICINO while CUORE is being assembled
- Demonstrate potential for DM and Axion detection

Assembly procedure

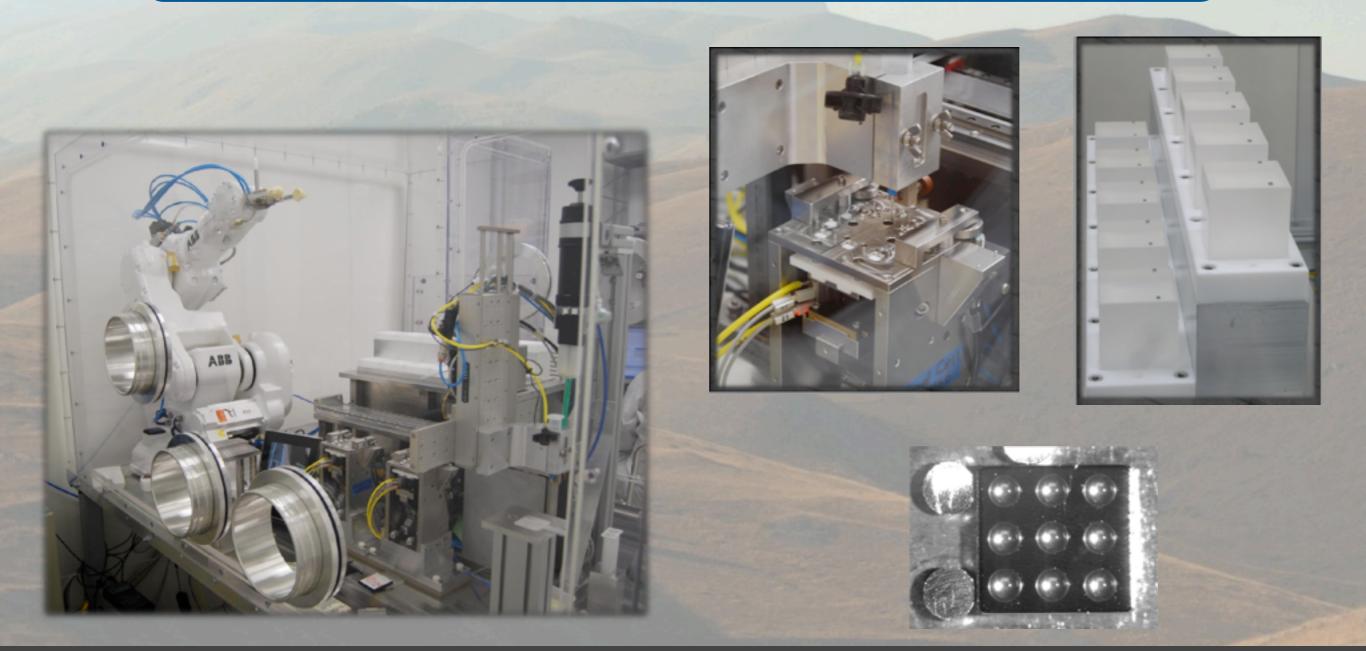
CUORE-0 assembly was performed in the new CUORE clean room following all the stages and using **all the equipment developed for CUORE**



Thermistor gluing

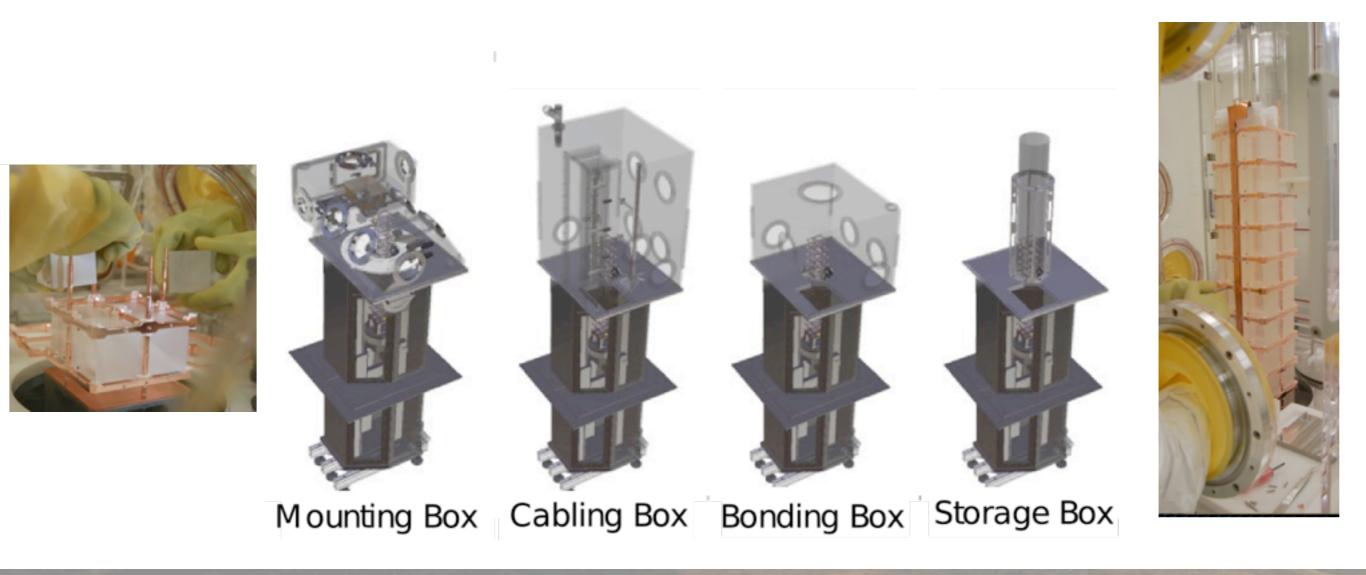
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The gluing of CUORE-0 thermistor to crystals was performed with the **new CUORE gluing semi-automatic machine** (in a N_2 flushed glove box): fast, operator independent, minimizes radioactive contaminations, makes this stage more reproducible thus improving detector uniformity.



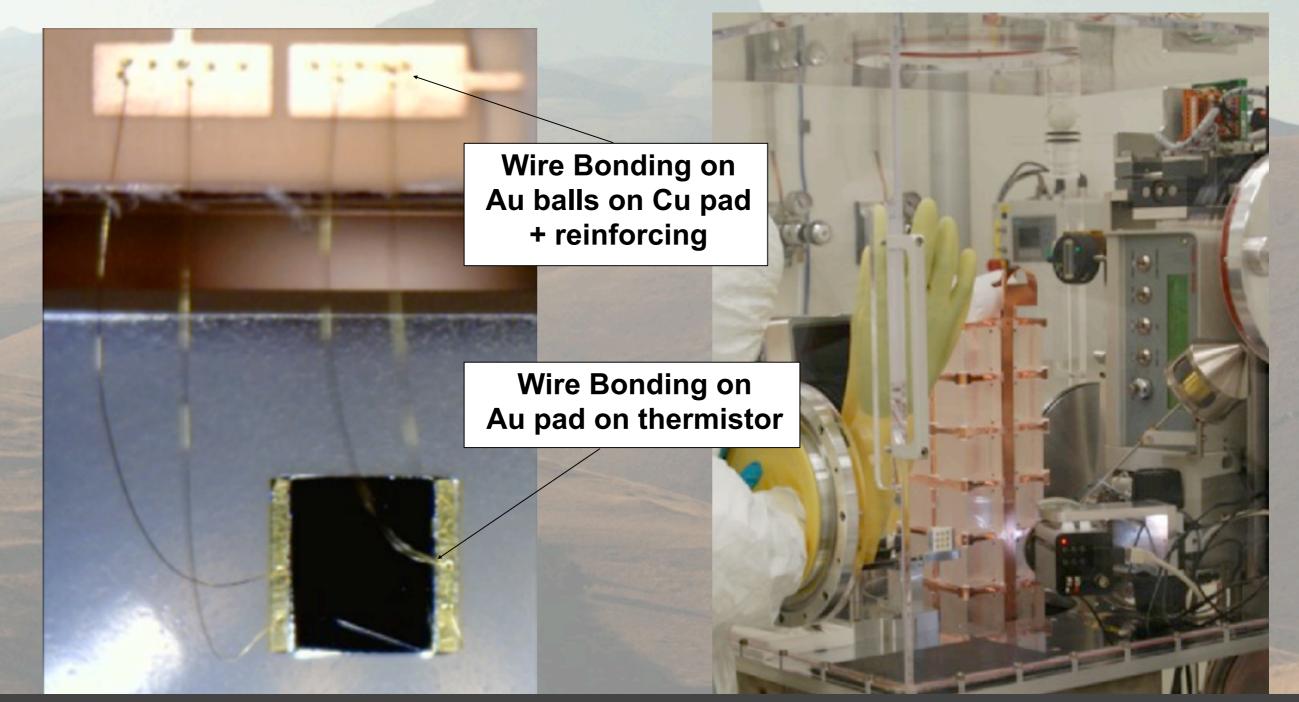
Tower assembly

The assembly of the tower was done with the **CTAL (CUORE Tower Assembly Line)** provided of a sealed and flushed stainless steel chamber (Garage) supporting a working plane where **4 different glove boxes switch** allowing 4 operations to be performed (mounting, cabling, bonding and tower storage) with radioactivity control and reproducible protocols.



Signal wires connection

The signal readout is provided by **gold wires directly bonded** on the assembled tower in a N_2 fluxed glove box. The bonding proceeds in **3 main steps**: 1. ball bonding on a Au pad on the thermistor; 2. bonding of the wire on Au balls on a Cu pad; 3. reinforcing.



Installation



After shield installation the tower has been closed in a flushed Plexiglas box, then moved to the special opening door of the **CUORE clean room** with a trans-pallet. It has then been lifted with a fork-lift to the **CUORICINO clean room** where it has been joined to the cryostat dilution unit (operations not necessary for CUORE).



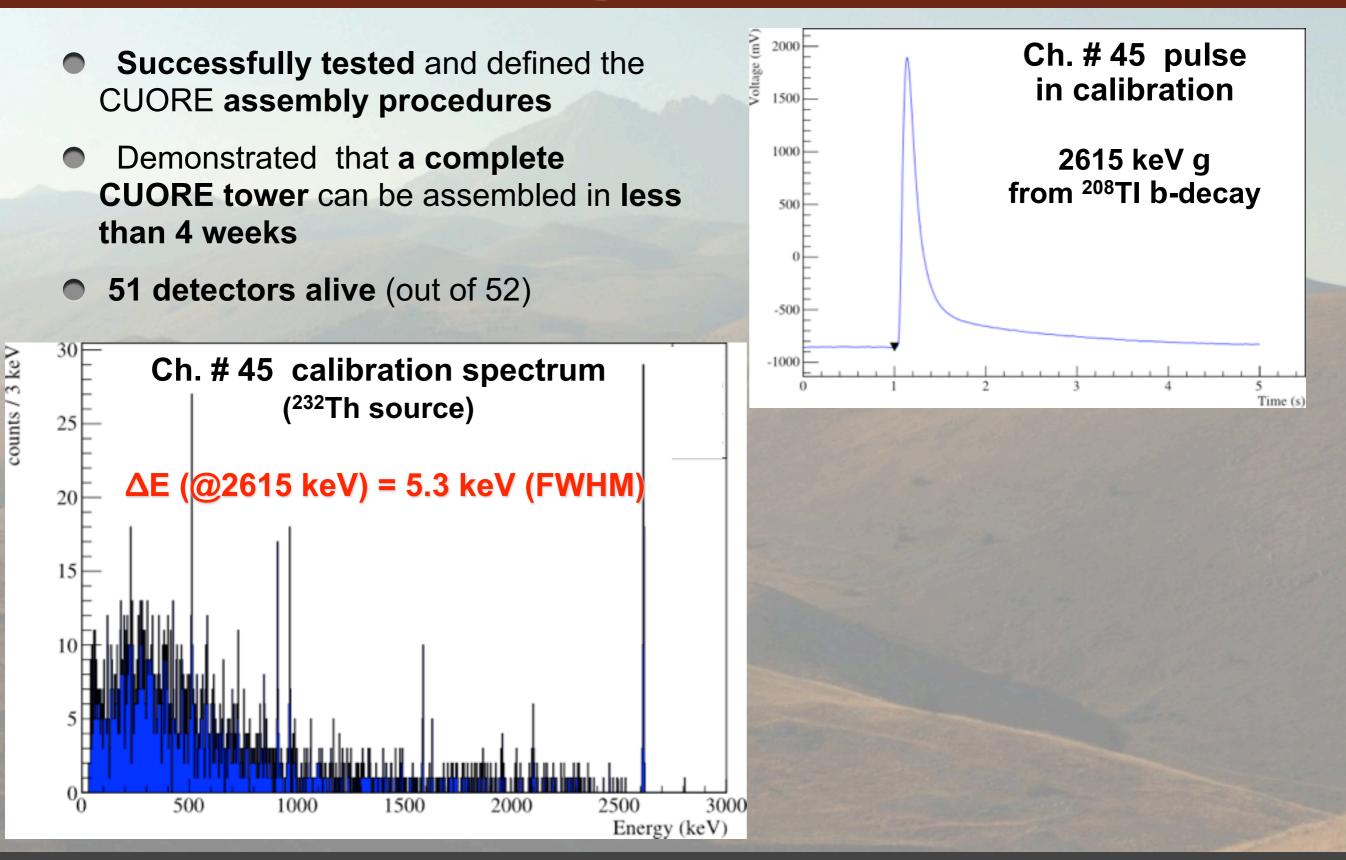
CUORE-0 status



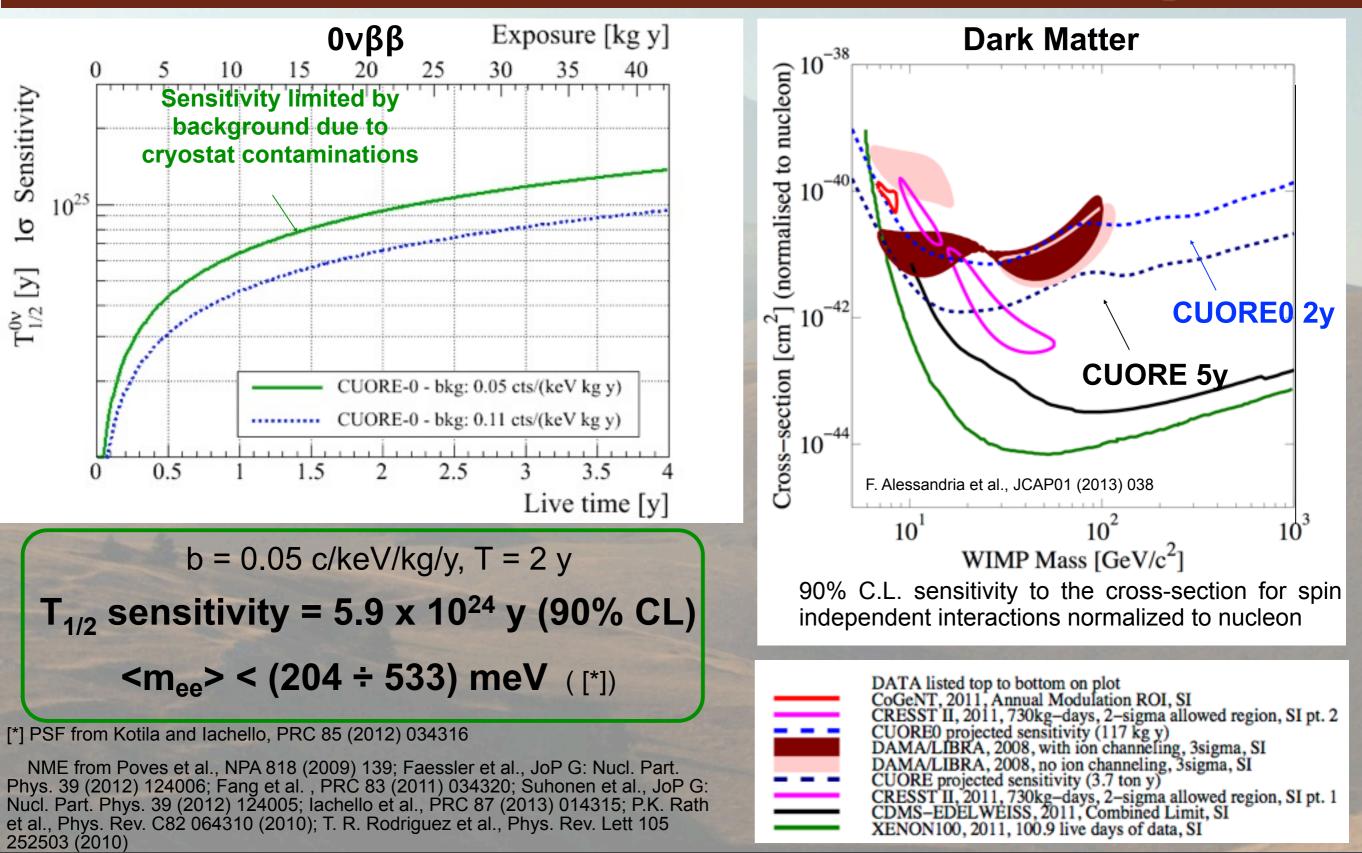
- CUORE-0 was cooled down to a base T ~ 8 mK to start the pre-operation and optimization phase in August 2012
- Pre-operation was disturbed by 2 vacuum leaks (the cryostat is ~25 years old) which deteriorated detector performances
- We were able to perform calibrations despite the leaks, showing reasonable detector performance
- The second leak has been fixed at the beginning of February
- Pre-operation and optimization phase restarted in mid-March



Detector performance



0vββ and DM sensitivity



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Paris - May 22, 2013

Conclusions



- TeO₂ bolometers represent since many years a competitive detector for $0\nu\beta\beta$ research
- Going from CUORICINO to CUORE a strong R&D has been performed in order to reduce the background in the ROI (the main challenge being surface contaminations of detector and facing parts)
- Bolometric tests after improving surface treatments demonstrate that the CUORE goal of 0.01 c/keV/kg/y is just behind the corner. Copper/PTFE surface is the most crucial issue
- CUORE is under construction: LNGS hut and clean room ready, cryostat commissioning ongoing, crystal production and storage almost complete, assembly of the first two towers almost complete
- CUORE cool down foreseen for end 2014
- First CUORE tower, CUORE-0, is cold in the CUORICINO cryostat at LNGS and in preliminary data-taking

Bibliography

[1] H.V. Klapdor-Kleingrothaus et al., Mod. Phys. Lett. A 21 (2006) 1547 [2] S.M. Bilenky and C.Giunti, Mod.Phys.Lett. A27 (2012) 1230015 [3] M. Auger et al., Phys. Rev. Lett. 109 (2012) 032505, arXiv:1205.5608 [hep-ex] [4] A. Gando et al., Phys. Rev. C 85 (2012) 045504, arXiv:1205.6130v1 [hep-ex] [5] E. Andreotti at al., Astrop. Phys. 34 (2011) 822-831 [6] Simkovic et al., PRC 77 (2008) 045503 Citarese et al., JoP:Conference series 173 (2009) 012012 Menendez et al., NPA 818 (2009) 139 Barea and Iachello, PRC 79 (2009) 044301 [7] C. Arnaboldi et al., Phys. Rev. Lett. 95, 142501 (2005) [8] C. Arnaboldi et al., Phys. Rev. C 78, 035502 (2008) [9] F. Alessandria et al., Astrop. Phys. 35, 839-849 (2010) [10] Mei and Hime, Phys. Rev. D 73, 053004 (2006) [11] MACRO Coll., M.Ambrosio et al. Phys. Rev. D52, 3793 (1995) MACRO Coll., M.Ambrosio et al., Astrop. Phys. 19, 313 (2003) [12] F.Arneodo et al. Il Nuovo Cim. 112A, 819 (1999) [13] P.Belli et al., Il Nuovo Cim. 101A, 959 (1989) [14] C. Bucci et al. Eur. Phys. J. A 41:155-168 (2009) [15] F.Bellini et al. Astrop. Phys 33, 169-174 (2010) [16] L.Foggetta et al. Appl. Phys. Lett. 86, 134106, 2005 [17] J.W.Beeman et al., Astrop. Phys. 35, 558-562 (2012) [18] T.Tabarelli de Fatis, Eur. Phys. J. C65, 359-361 (2010) [19] F.Vissani, A.Strumia, http://arxiv.org/pdf/hep-ph/0606054v3.pdf [20] C.Arnaboldi et al., J. Cryst. Growth 312 (2010) 2999–3008