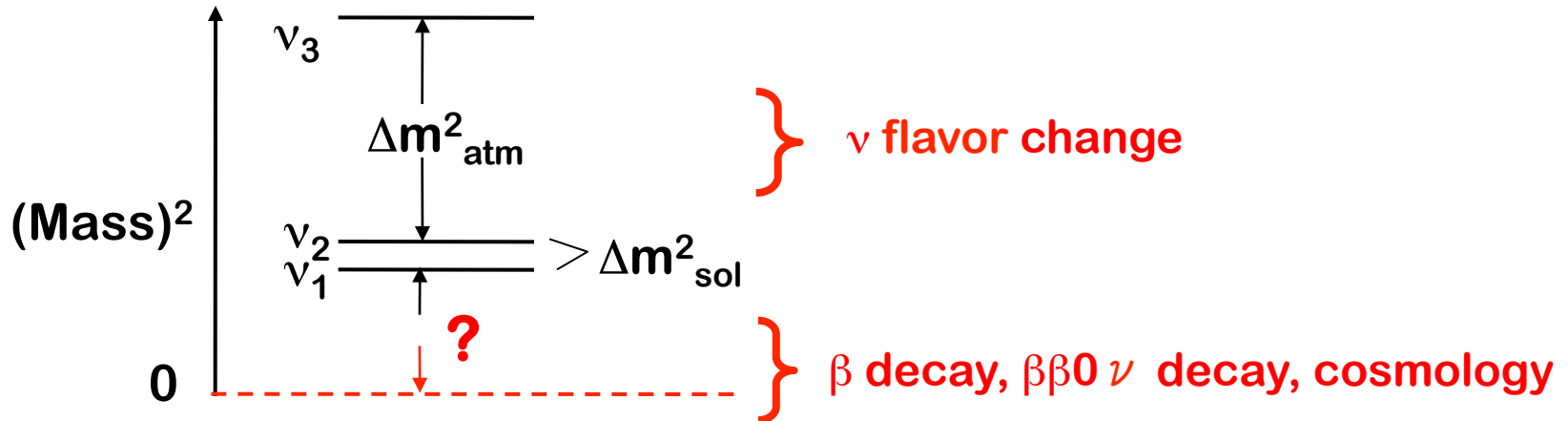


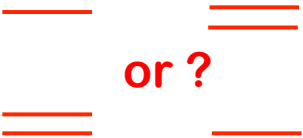
Sterile Neutrinos



Open questions

- What are the masses of the mass eigenstates ν_i ?



- Is the spectral pattern  or ? ν behavior in matter, $\beta\beta 0 \nu$, osc.

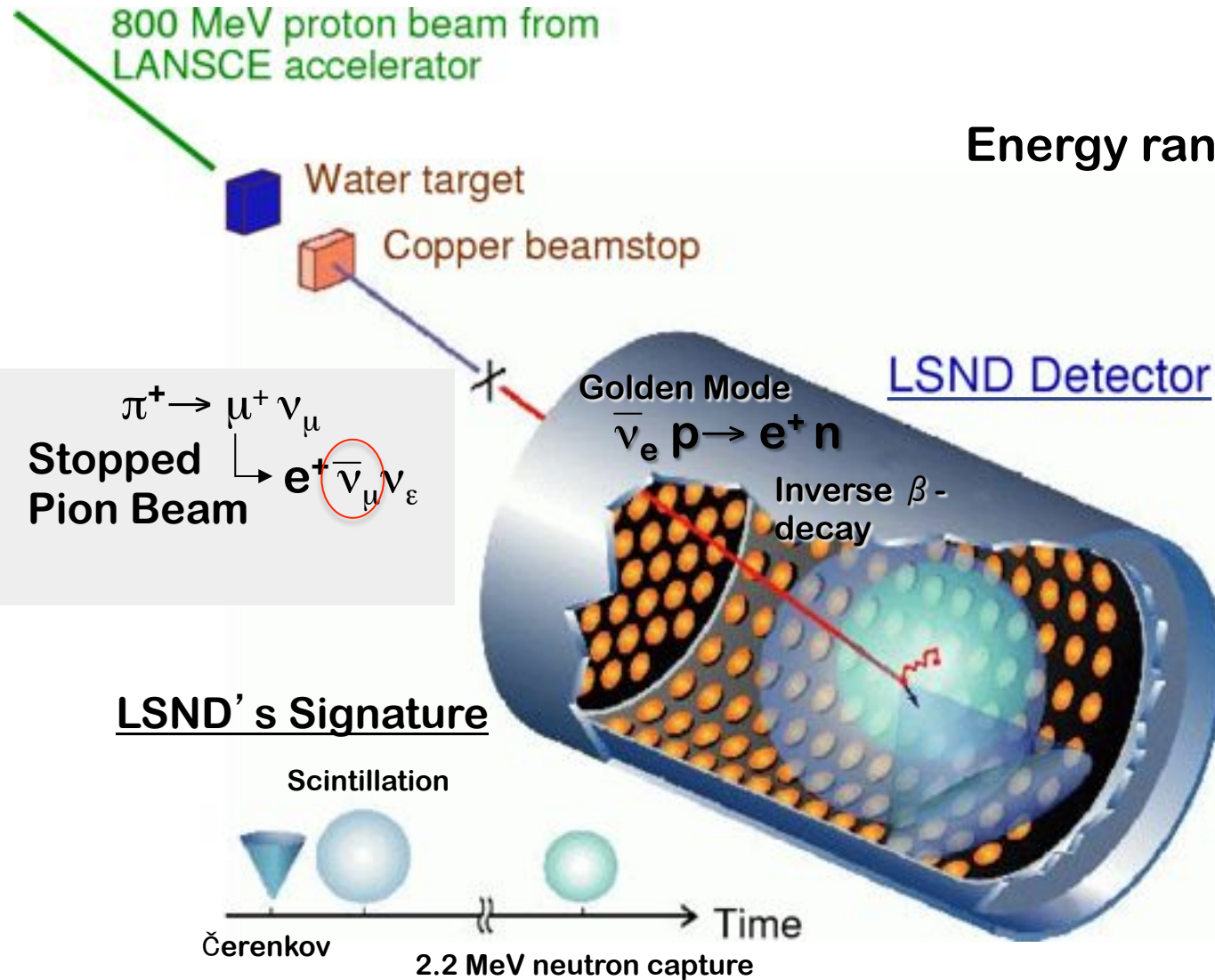
- Is there any conserved Lepton Number (Dirac or Majorana neutrino) ? $\beta\beta 0 \nu$

- Precise measurements of the leptonic mixing matrix?
 - Do the behavior of ν violate CP?
 - Is leptonic ~~CP~~ responsible for the matter-antimatter asymmetry?
- } ν flavor change

- Are there additional (sterile) neutrino states ν flavor change, Cosmology

A bunch of Neutrino Anomalies

LSND (Los Alamos, 1993-98)



Baseline of 30 m

Energy range of 20 to 55 MeV

$L/E \approx 1 \text{ m} / \text{MeV}$

$\Delta m^2 \approx 1 \text{ eV}^2$

Detection :

$$\bar{\nu}_e + {}^1\text{H} \rightarrow e^+ + n$$

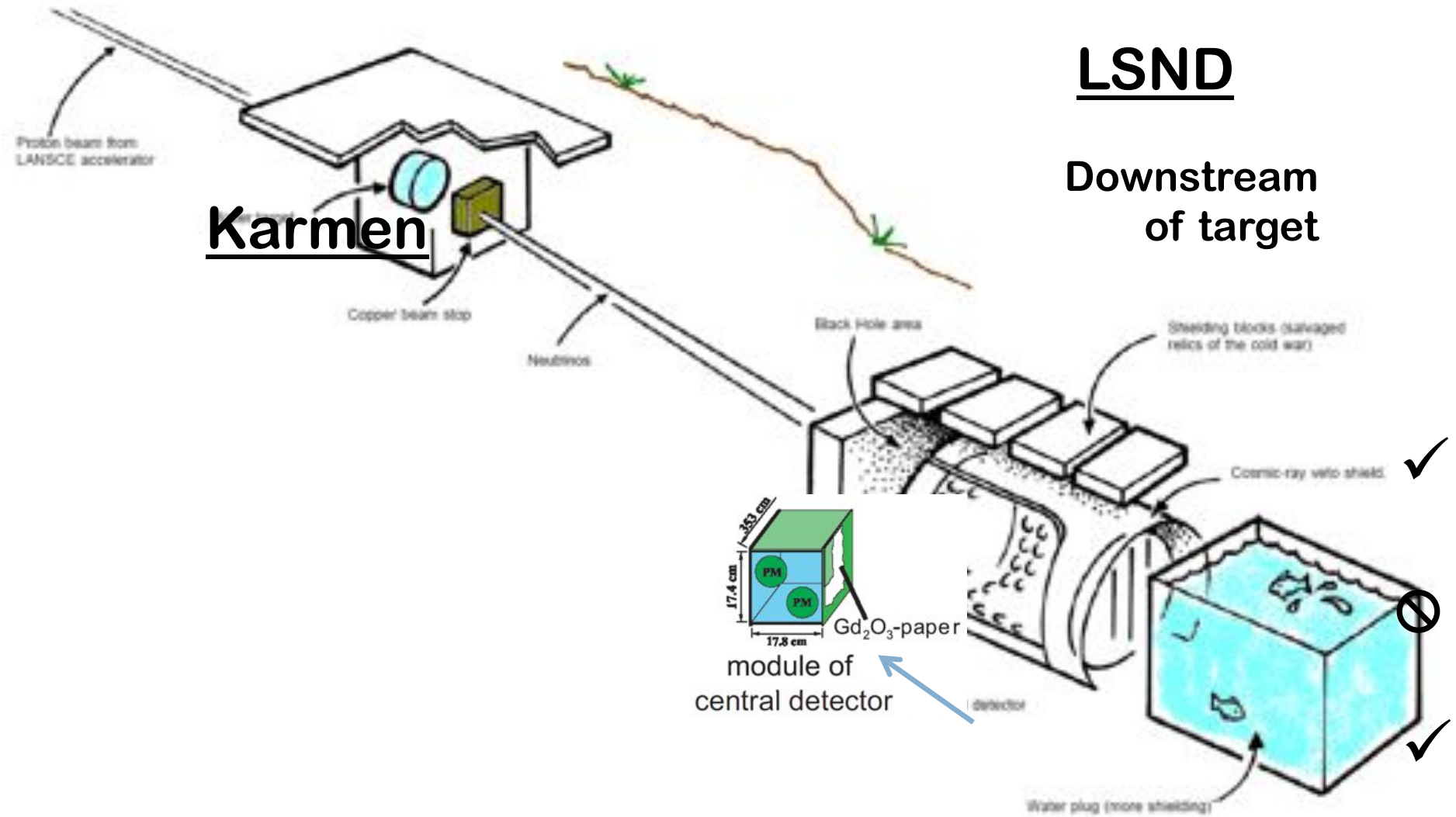
Karmen (stopped π^+ beam)

Oscillation not confirmed – exclude part of LSND

LSND

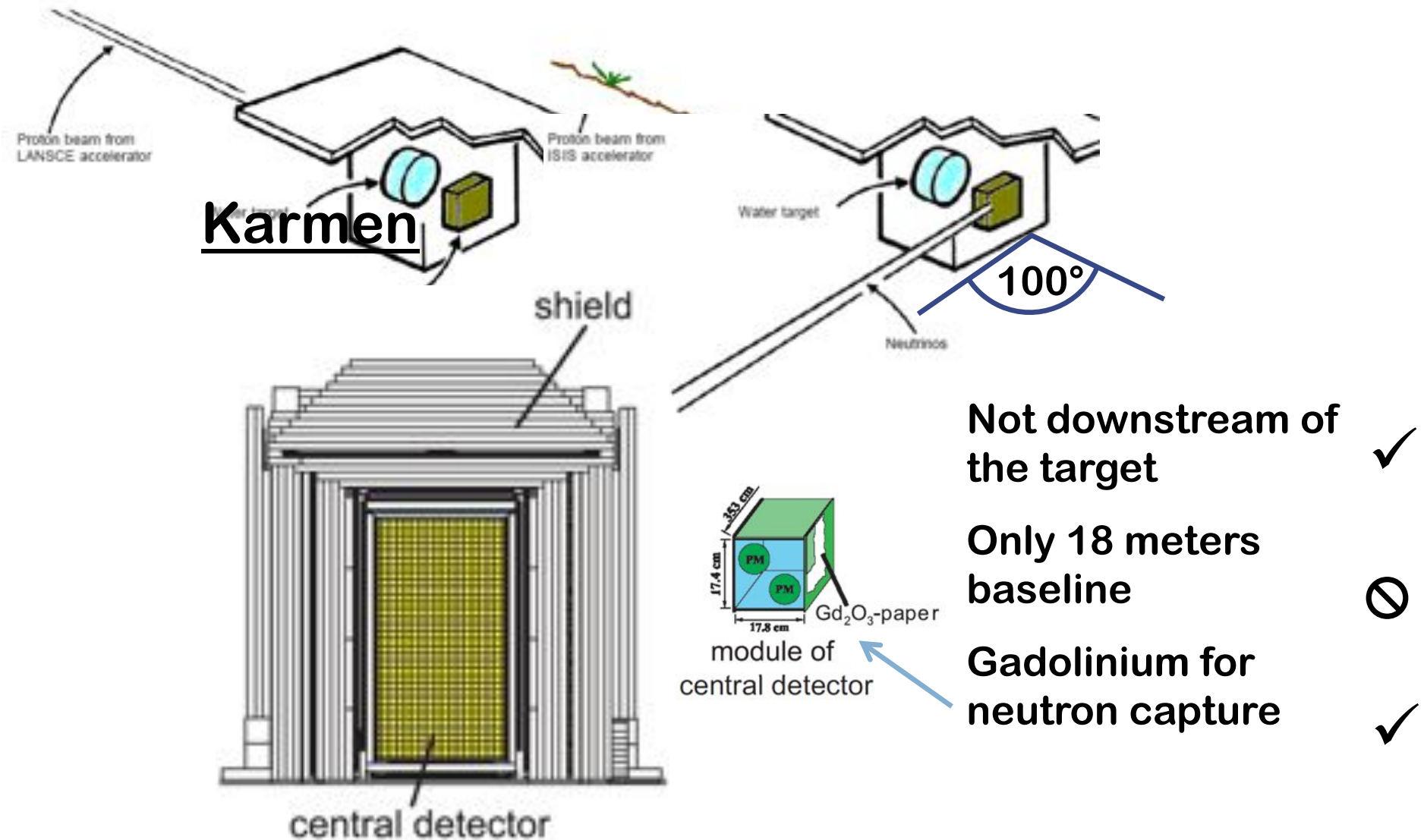
Downstream
of target

Karmen



Karmen (stopped π^+ beam)

Oscillation not confirmed – exclude part of LSND



- 1st results published in PRL 75 (1995)

- Channel:** anti- $\nu_\mu \rightarrow \text{anti-}\nu_e$

- Detection :** anti- $\nu_e + {}^1\text{H} \rightarrow e^+ + n$

- Baseline:** 30 m

- Energy:** $20 < E \text{ (MeV)} < 200$

- Status:**

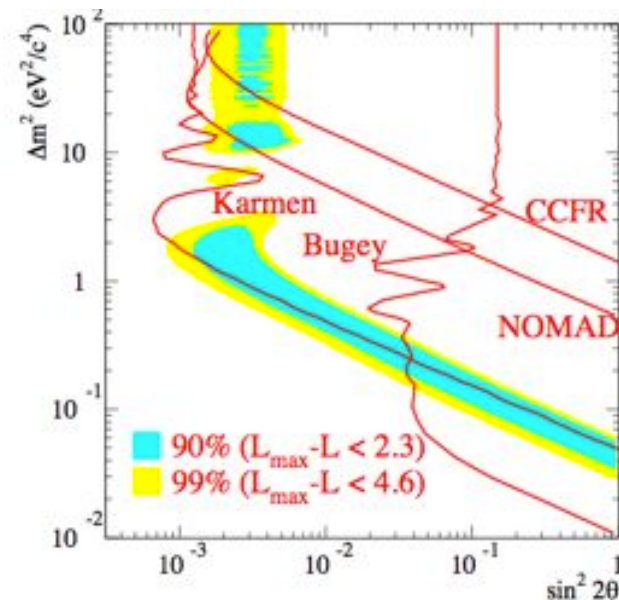
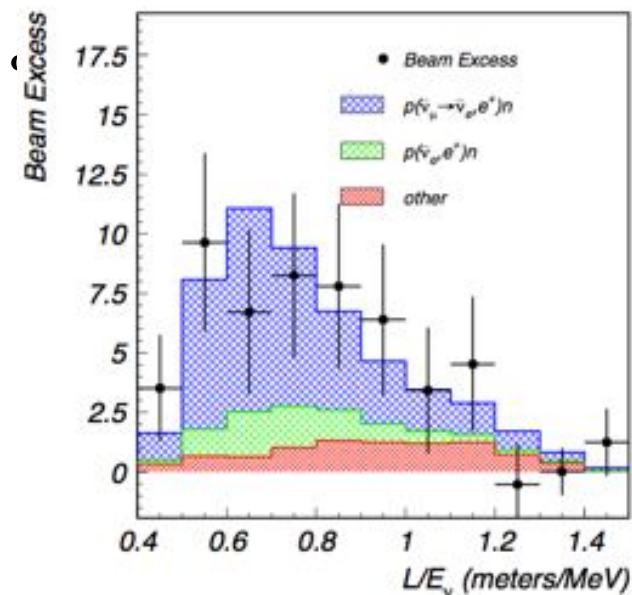
- anti- ν_e excess observed
 $\rightarrow 32.2 \pm 9.4 \pm 2.3 \text{ (3.8}\sigma\text{)}$

- not confirmed by Karmen

- Oscillation parameters:**

- $\Delta m^2 \gg 0.2 \text{ eV}^2 \gg \Delta m_{\text{atm}}^2$

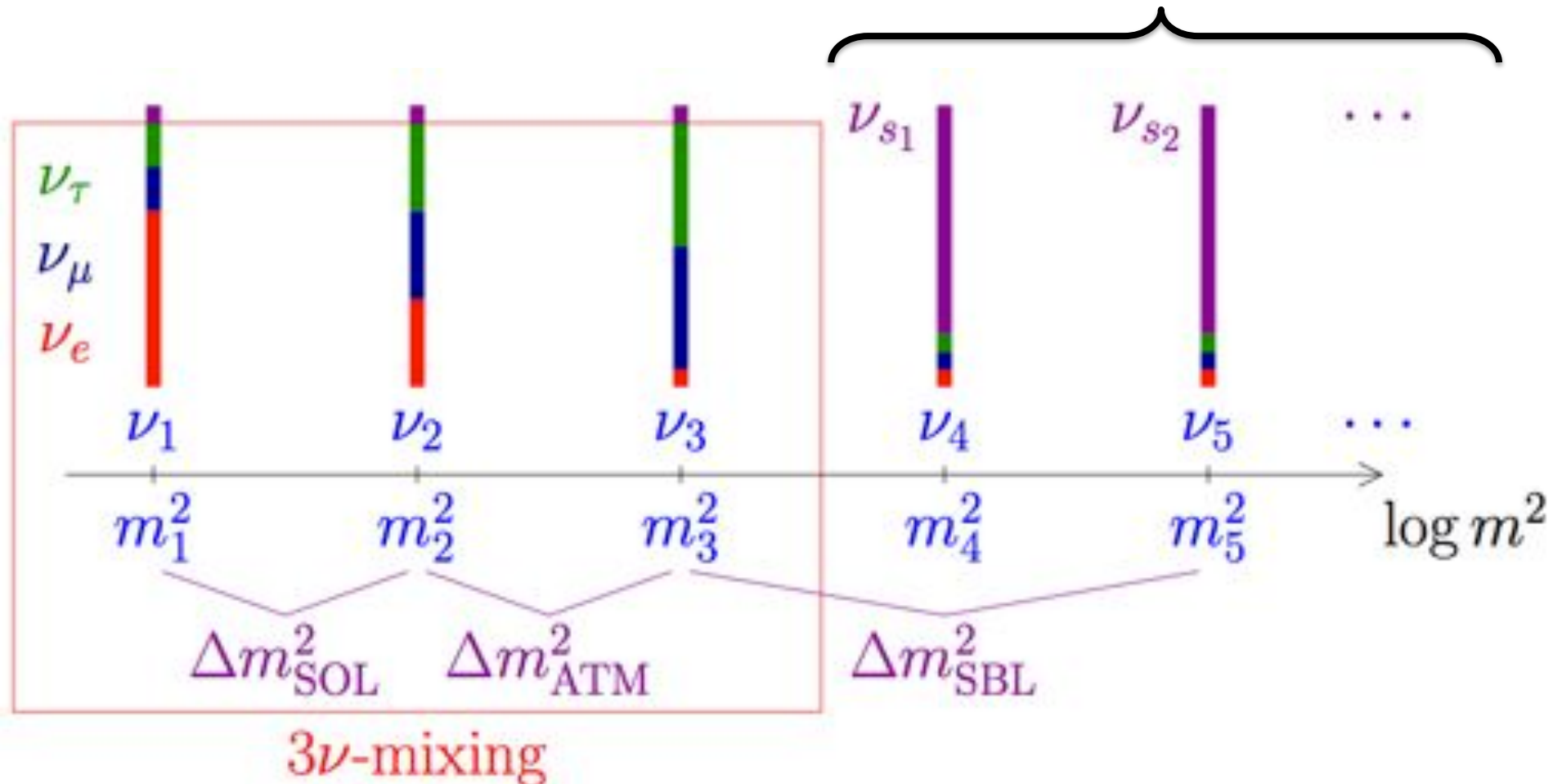
- Require a 4th neutrino state



Adding Sterile Neutrinos

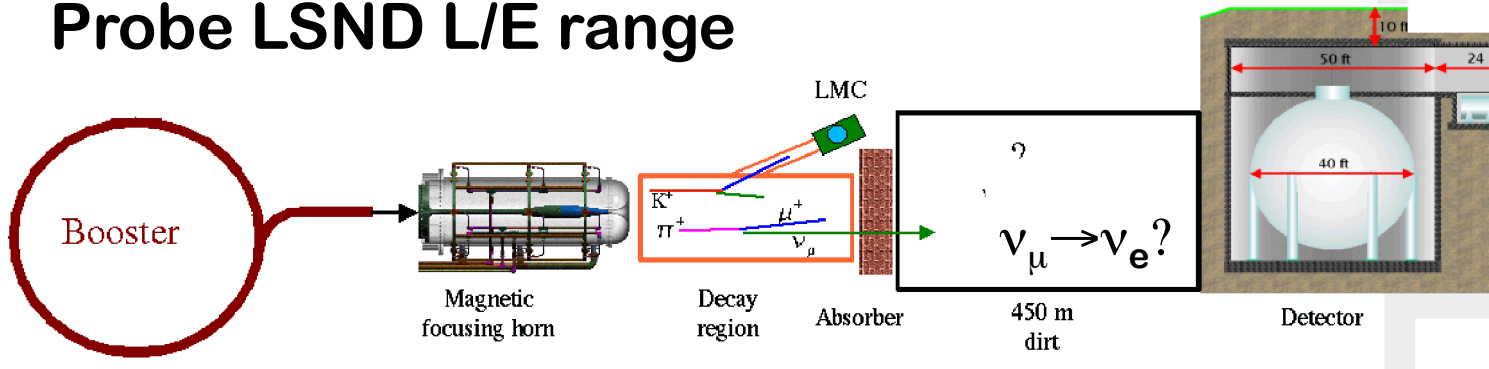
Introduce a light ν_R in SM, No SM interactions mixing with active ν 's

No coupling with Z boson (LEP)



The MiniBooNE Experiment

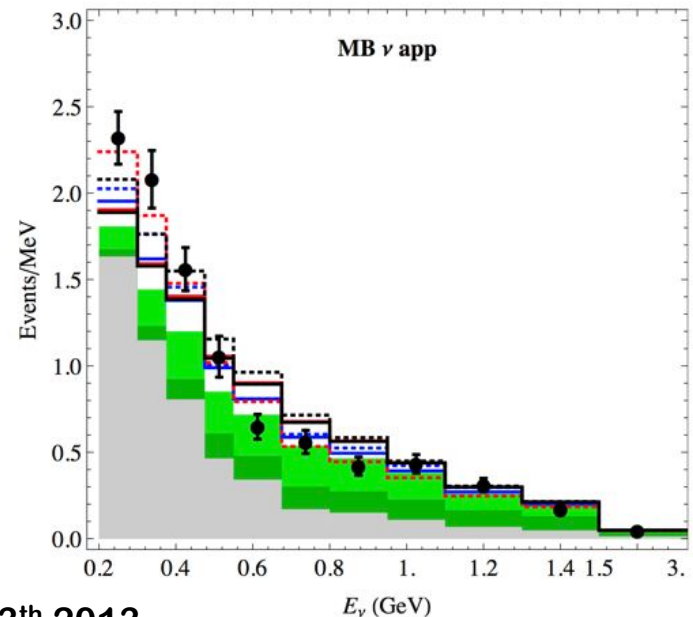
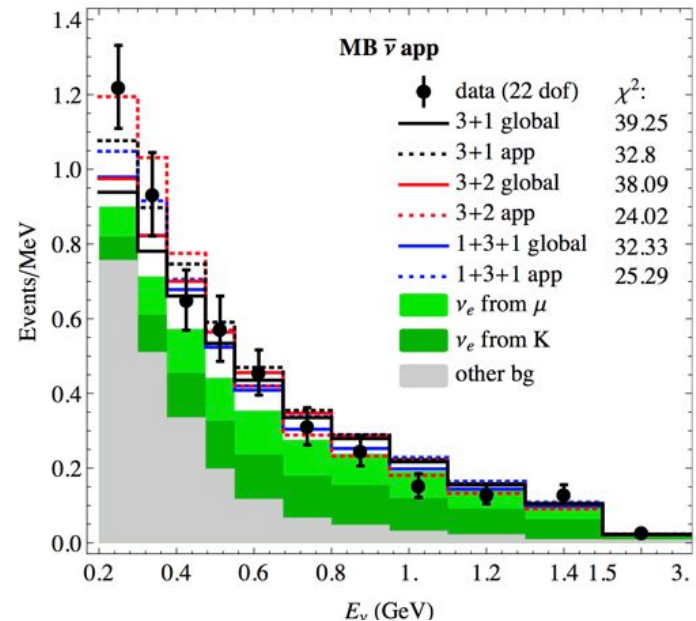
Primary goal: look for ν_e appearance in a ν_μ beam
 Probe LSND L/E range



- Cherenkov + scintillation
- π^+ (π^-) decay in flight beam
- $L/E \approx 1 \text{ m / MeV}$
 - Baseline: 500 m
 - $\langle E \rangle 500 \text{ MeV}$
- Started in 2002

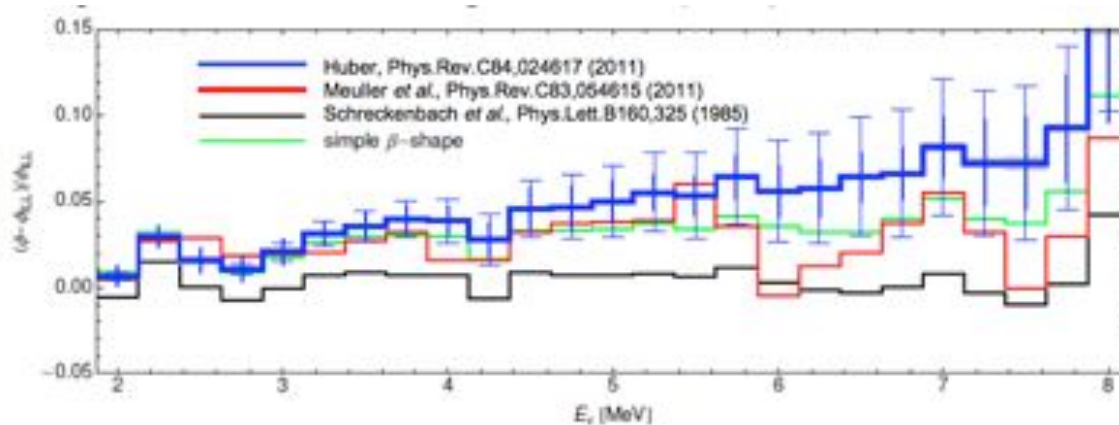
Miniboone Results (FNAL)

- Results published from 2010-7-12
- **Channel:** $(\text{anti-})\nu_\mu \rightarrow (\text{anti-})\nu_e$
- **Detection:** $\nu_e(p)n \rightarrow e\ p$ (CCQE)
- **Baseline:** 541 m
- **Energy:**
 - $200 < E \text{ (MeV)} < 3000$
- **Status:**
 - Excess of event at low energy
 - $3.8\ \sigma$
 - 4th neutrino?
 - Backgrounds issue?



The Reactor Anomaly

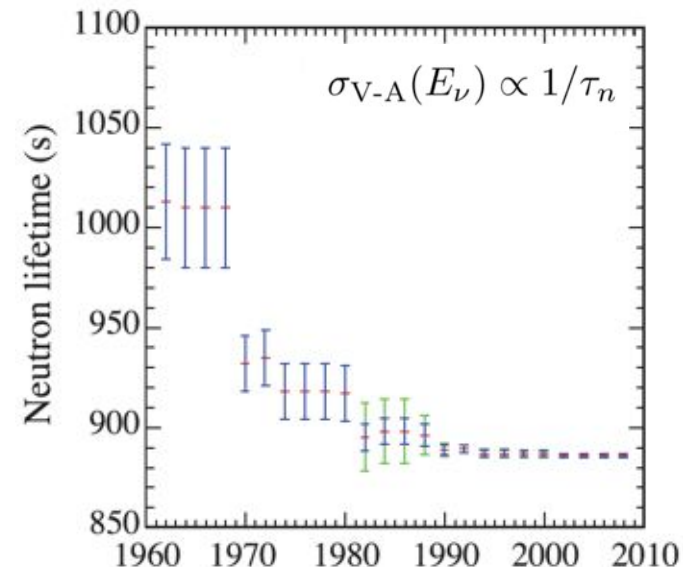
- i) **V_{emission}**: Improved reactor neutrino spectra $\rightarrow$ +3.5%



PRC83, 054615 (2011)

PRC84, 024617 (2011)

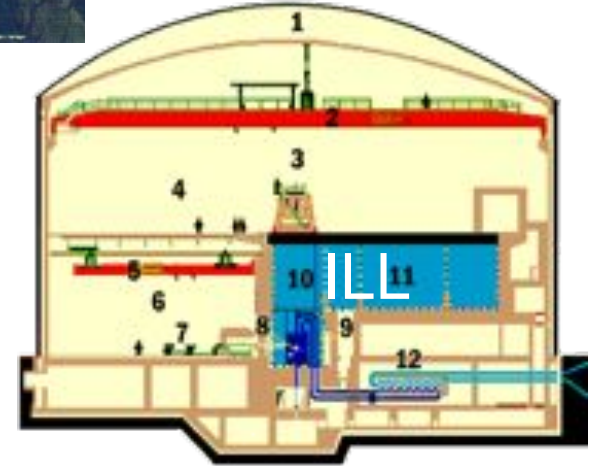
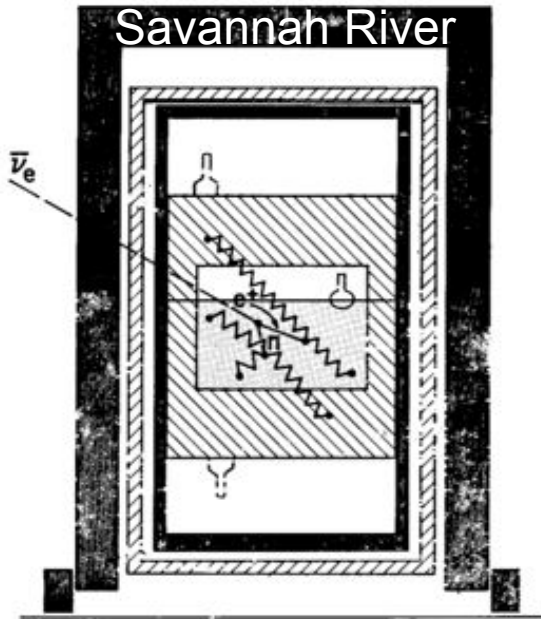
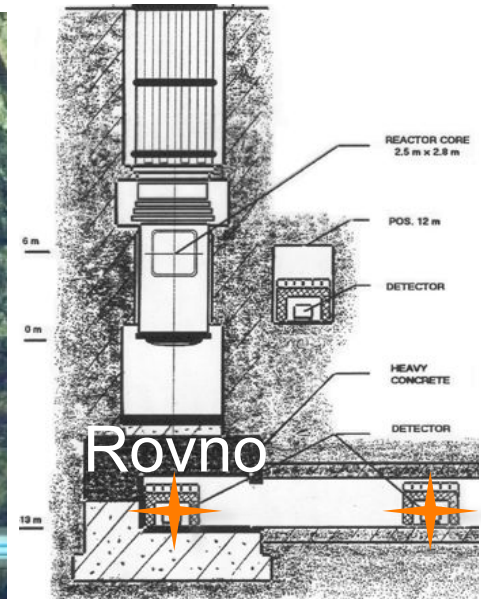
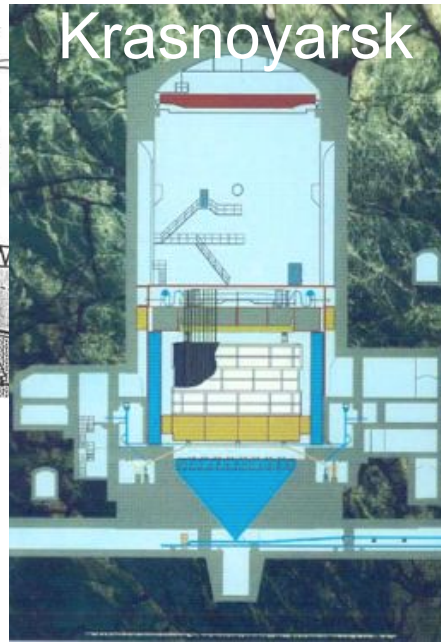
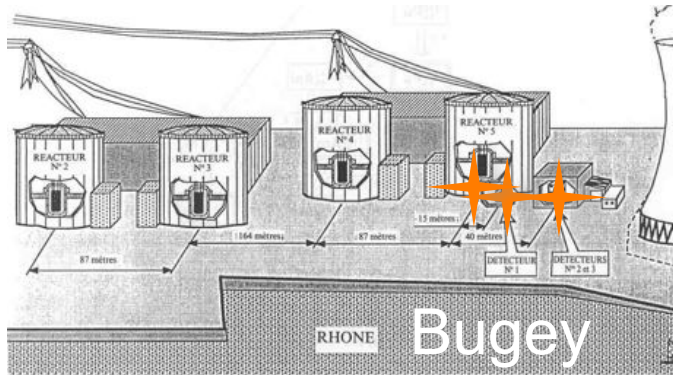
- ii) **V_{detection}**: Reevaluation of σ_{IBD} $\rightarrow$ +1%
Evolution of the neutron life time



- iii) **V_{detection}**: Accounting for long-lived isotopes accumulating in reactors $\rightarrow$ +1%

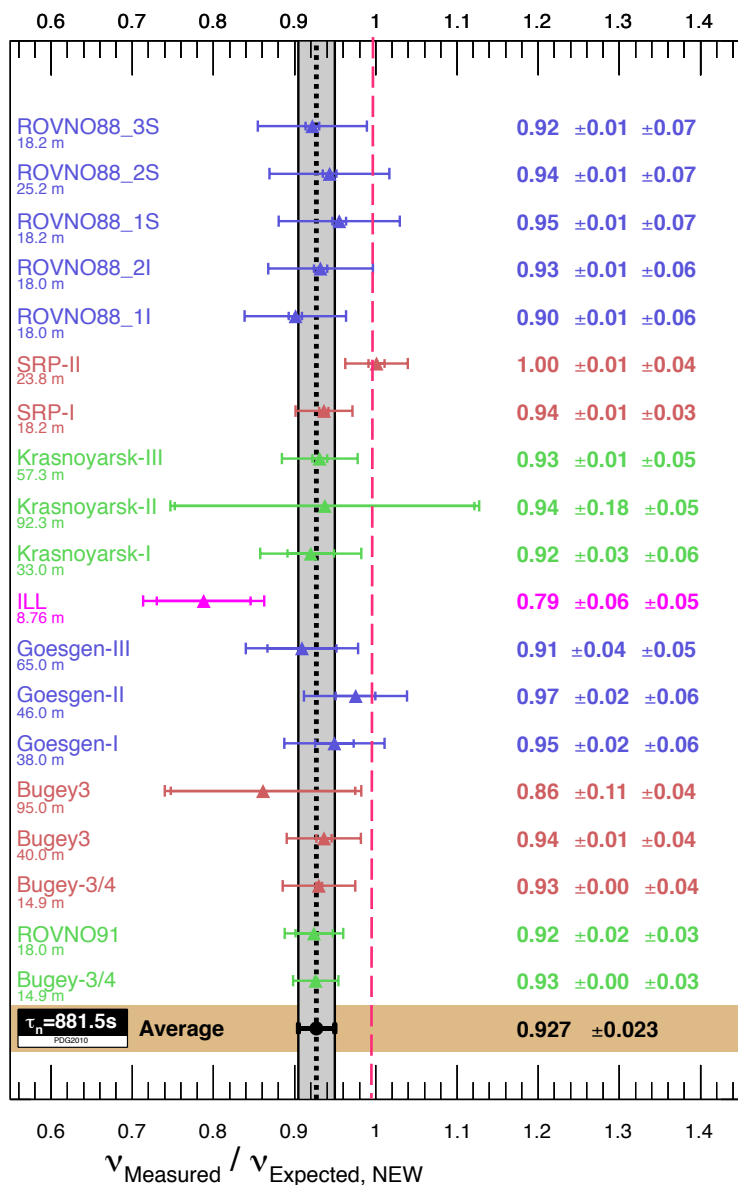
PRD 83, 073006 (2011)

19 Experimental Results below 100m



Measured cross sections / fission taken at their face values

The reactor anomaly

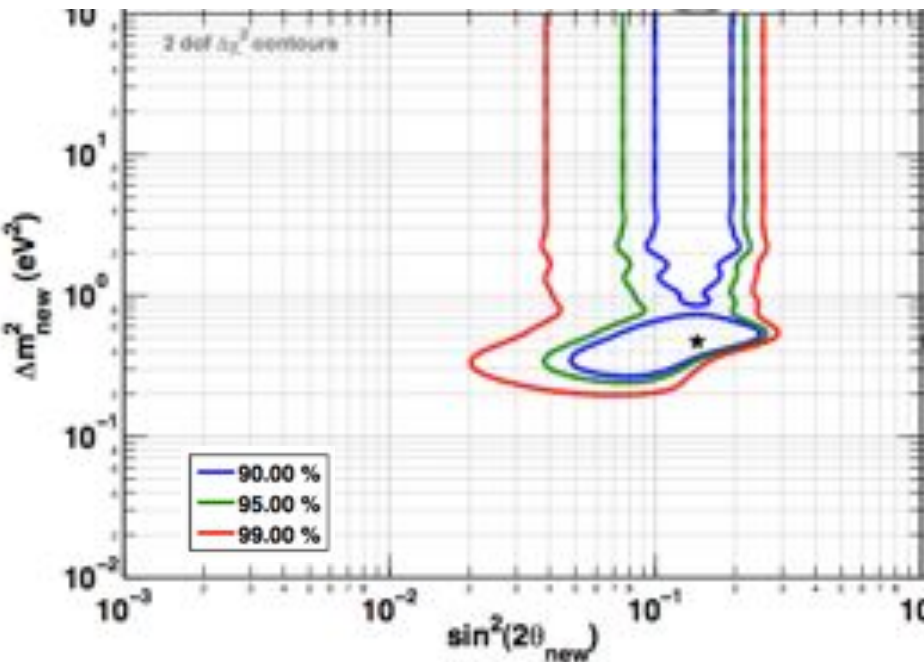


- 19 experiments reanalyzed
PRD 83, 073006 (2011)
- **7% deficit wrt the new prediction**
 - $\approx 3\%$: reevaluation of emitted flux
 - $\approx 3\%$: reevaluation of
 - IDB cross section revision
 - Accounting for off eq. effect
- **99.7 % C.L. deviation from unity**
- **Artifact or new physics?**

The 4th neutrino hypothesis

Rate Only Analysis

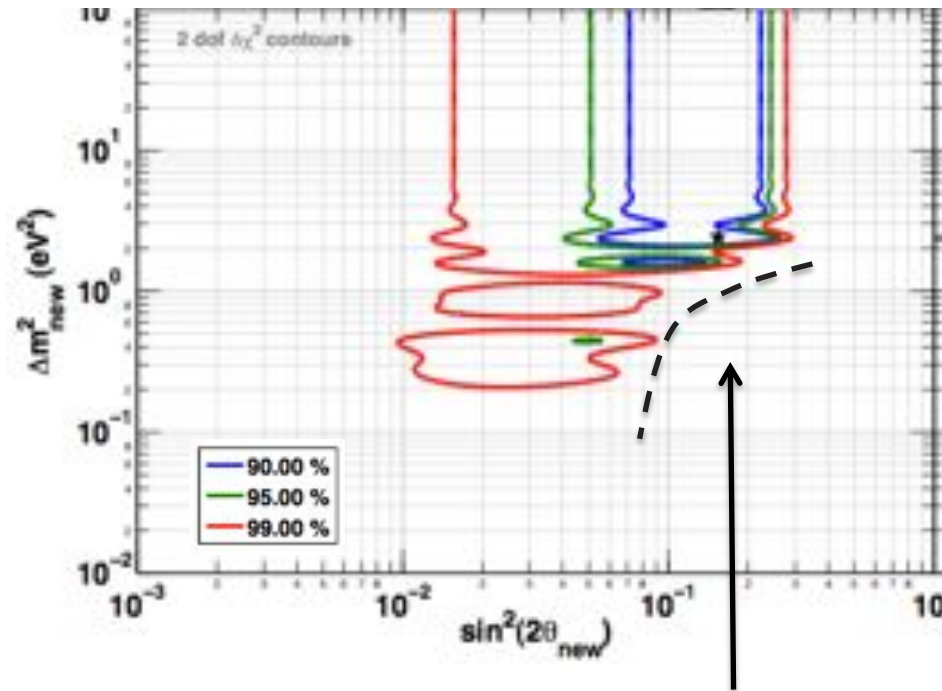
Phys. Rev. D 83, 073006 (2011)



- Best Fit at $\Delta m^2_{\text{new}} \approx \text{few } 0.1 \text{ eV}^2$

Rate + Shape Analysis

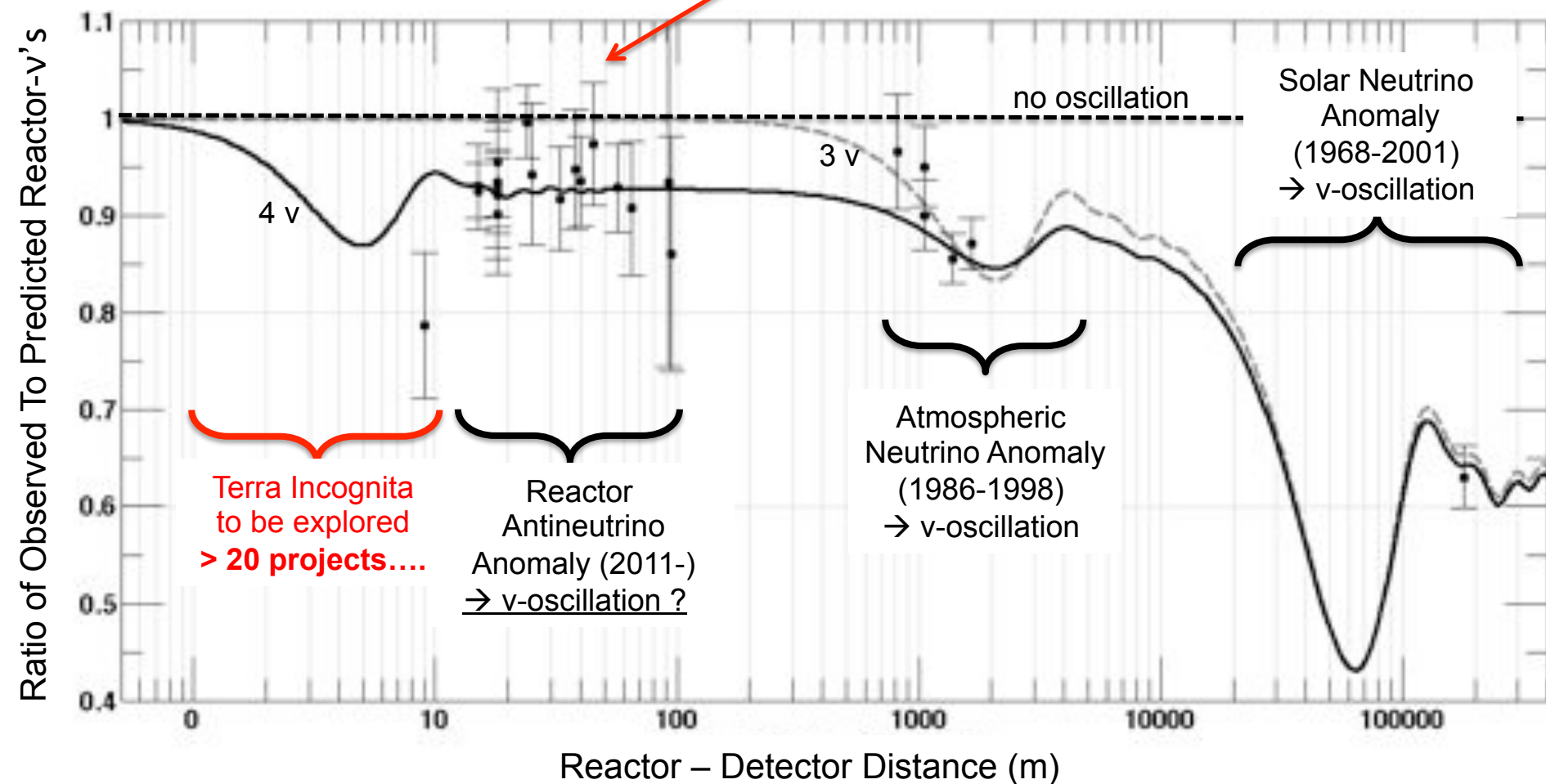
Phys. Rev. D 83, 073006 (2011)



- Bugey-3 40m/15m E_{spectrum} ratio
→ No energy spectrum distortion
→ large PWR core extension
- Best Fit at $\Delta m^2_{\text{new}} > 1 \text{ eV}^2$

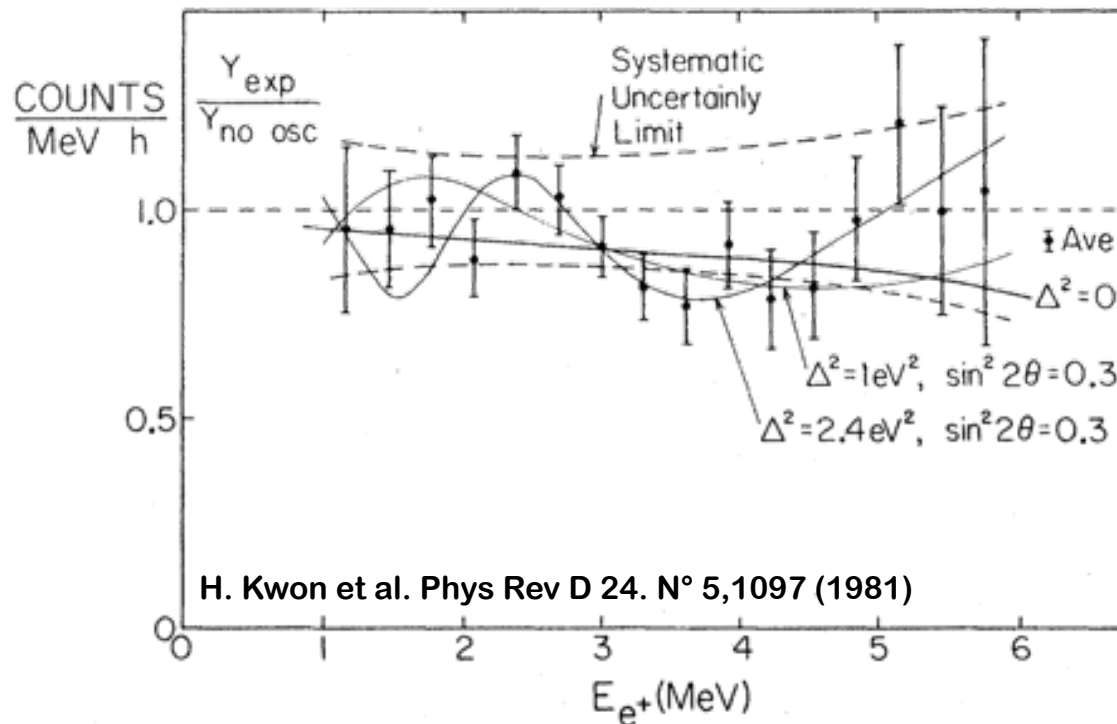
The Reactor Anomaly

- Observed/predicted averaged event ratio: $R=0.927\pm0.023$ (3.0σ)



Puzzling 1981 ILL v-experiment

- Reactor at ILL with almost pure ^{235}U , with compact core
- **Detector 8.8 m from a COMPACT core**
- Reanalysis in 1995 to account for overestimation of flux at ILL reactor by 10%... Affects the rate only but **20% deficit!**



- Large errors, but a striking pattern is seen by eye ?

The Gallium Neutrino Anomaly

- Test of solar neutrino detectors
GALLEX and SAGE (ν_e 's)

- $E \approx \text{MeV}$, Baseline range $\approx \text{few m}$

- 4 calibration runs
 $\approx 1\text{-}2 \text{ MCi EC } \nu_e$ emitters

- Gallex

- ^{51}Cr source (750 keV)

- Sage

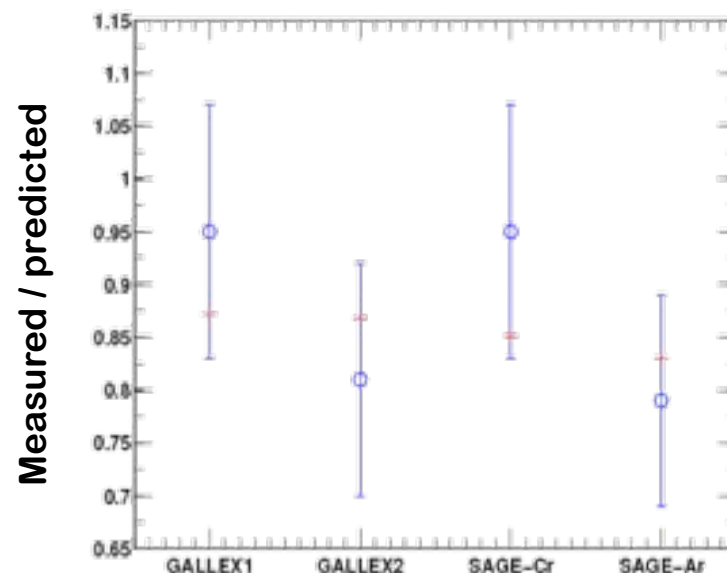
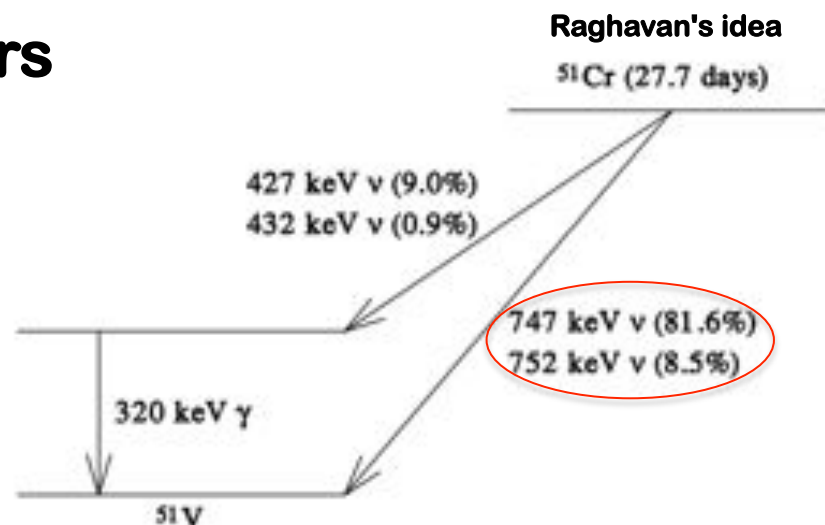
- ^{51}Cr & ^{37}Ar (810 keV)



- Deficit observed

- $R_{\text{obs/pred}} = 0.86 \pm 0.05$ (σ_{Bahcall})

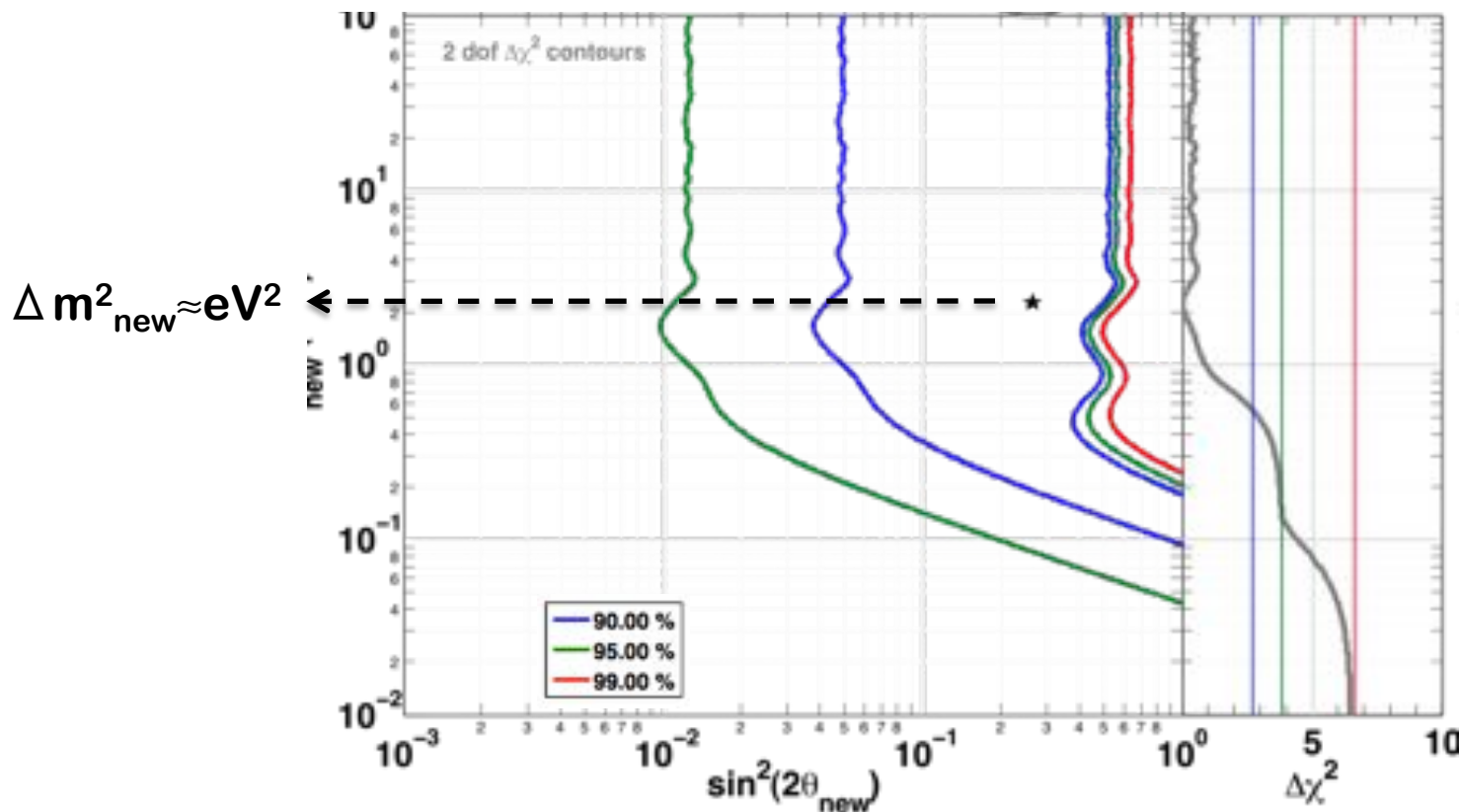
- $R_{\text{obs/pred}} = 0.76 \pm 0.085$ (σ_{Haxton})



The Gallium Neutrino Anomaly

Fit to ν_e disappearance hypothesis (3+1)

$$\begin{pmatrix} \nu_e \\ \nu_s \end{pmatrix} = \begin{pmatrix} \cos \theta_{\text{new}} & \sin \theta_{\text{new}} \\ -\sin \theta_{\text{new}} & \cos \theta_{\text{new}} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_{\text{new}} \end{pmatrix}, \quad P_{ee} = 1 - \sin^2(2\theta_{\text{new}}) \sin^2\left(\frac{\Delta m_{\text{new}}^2 L}{E}\right)$$

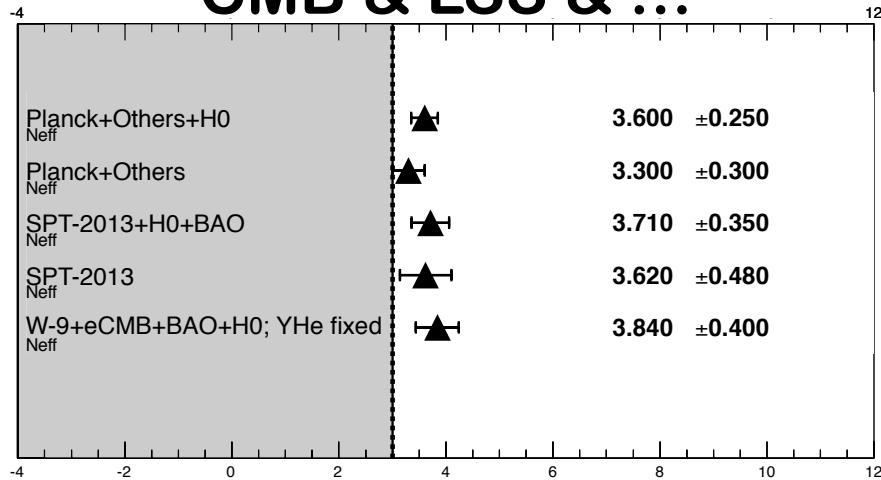


No-oscillation hypothesis disfavored at 2.7σ (PRC 83 065504, 2011)

Number of ν 's From Cosmology

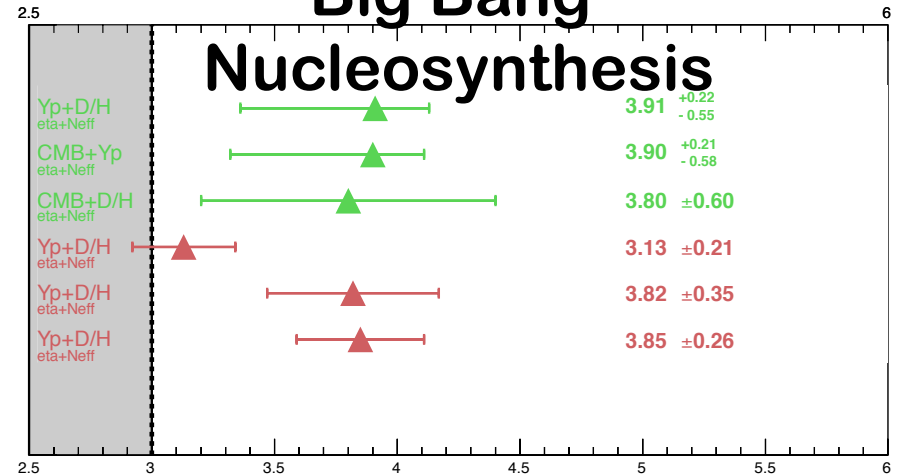
Universe Expansion Rate $H^2 \approx (\rho_Y + \rho_\nu) - \rho_Y$ given by CMB data

CMB & LSS & ...



Big Bang

Nucleosynthesis



- WMAP + other observables consistent with 3 and 4
- Planck favors 3.3, but consistent with 4
- Strong cosmological bounds, but model dependent

Anomalies & 4th Neutrino

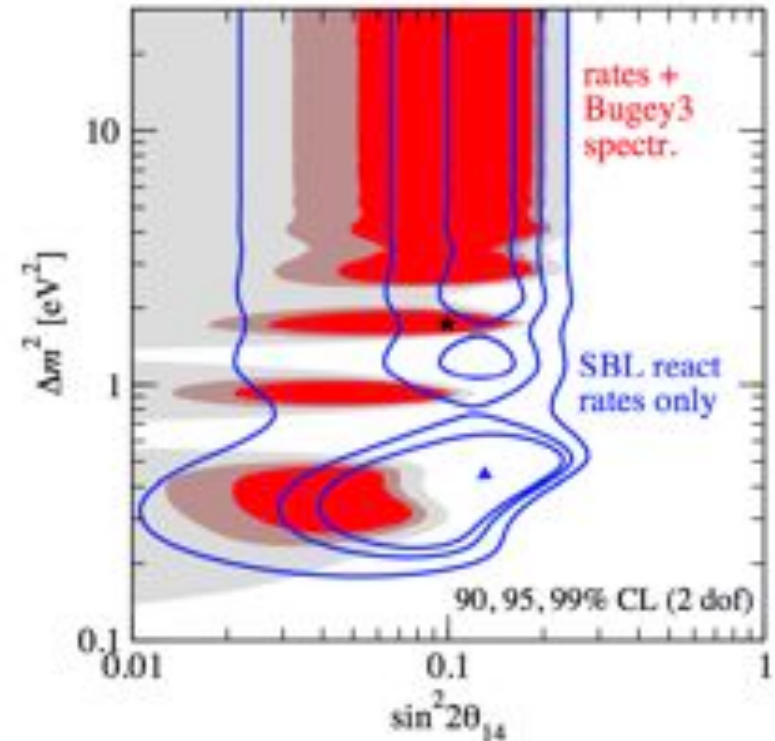
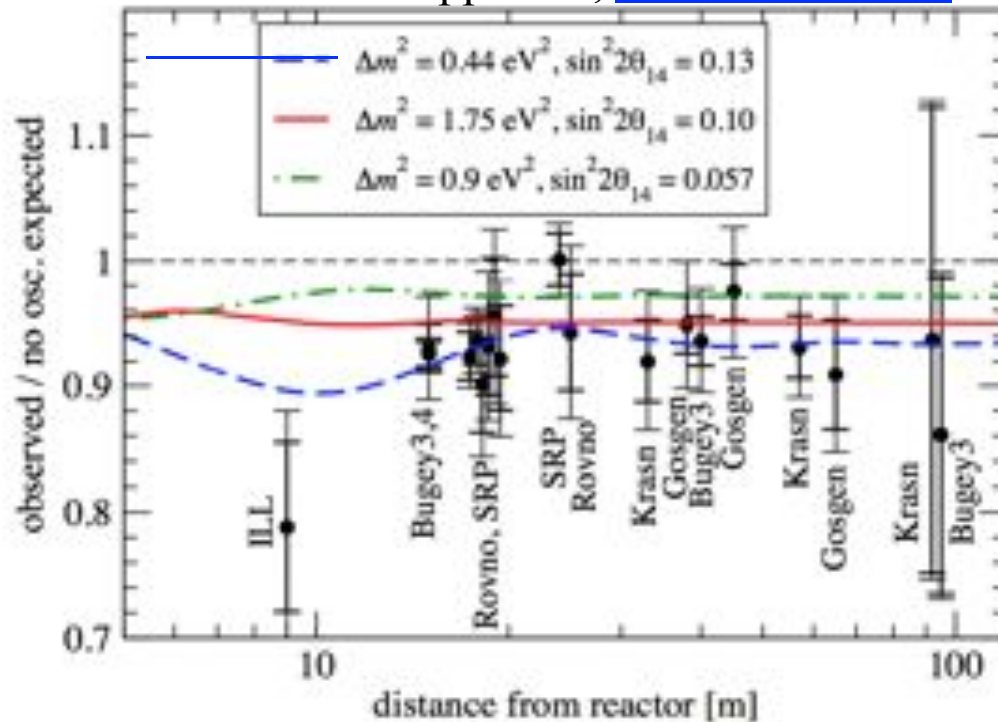
Anomaly	Source	Type	Sensitivity to Oscillation	Channel	Significance
LSND	Decay-at-Rest	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	<u>Total Rate, Energy</u>	CC	3.8 σ
MiniBoone	Short baseline	$\nu_\mu \rightarrow \nu_e$	<u>Total Rate, Energy</u>	CC	3.8 σ
Gallium	Electron Capture	ν_e dis.	<u>Total Rate</u>	CC	2.7 σ
Reactor	Beta-decay	$\bar{\nu}_e$ dis.	<u>Total Rate, Energy</u>	CC	3.0 σ
Cosmology	Big-Bang	All	Number of ν , N_{eff}	$N_{\text{eff}} = 3$ or 4 allowed	

could be interpreted by an existing eV² 4th neutrino state?

Global Picture

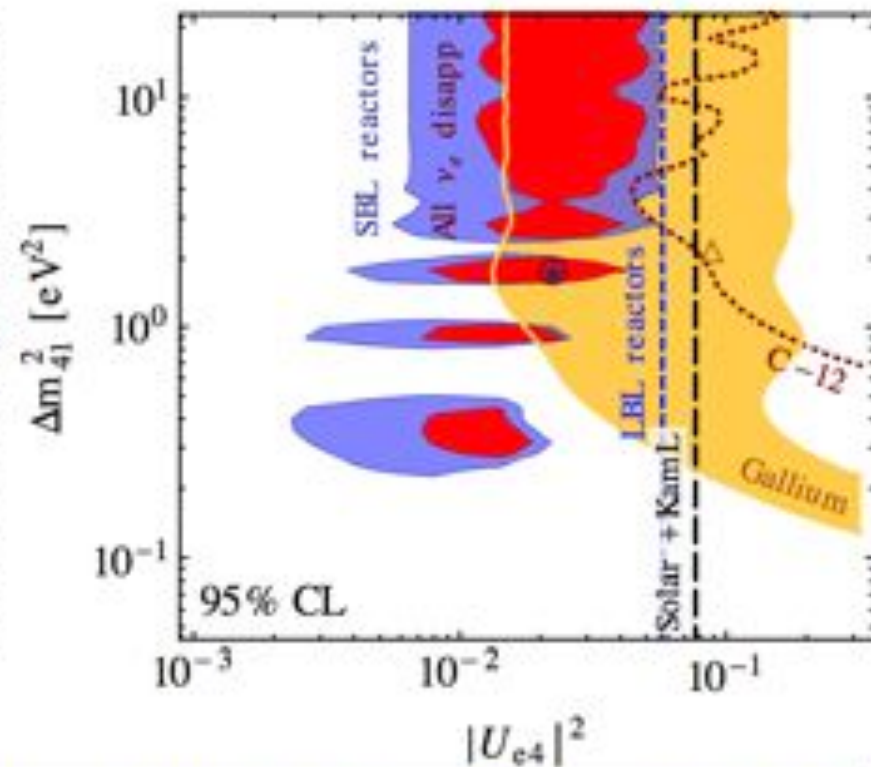
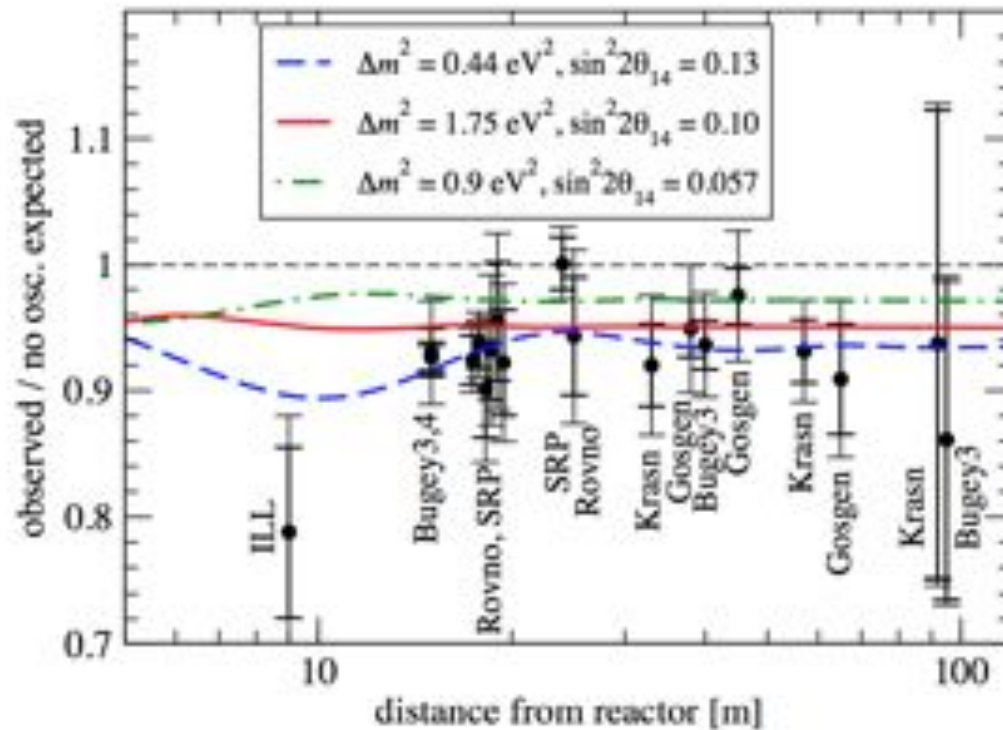
$(\bar{\nu}_e)$ disappearance (3+1 scenario)

J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)



	$\sin^2 2\theta_{14}$	Δm_{41}^2 [eV ²]	$\chi_{\min}^2/\text{dof}$ (GOF)	$\Delta\chi_{\text{no osc}}^2/\text{dof}$ (CL)
SBL rates only	0.13	0.44	11.5/17 (83%)	11.4/2 (99.7%)
SBL incl. Bugey3 spectr.	0.10	1.75	58.3/74 (91%)	9.0/2 (98.9%)

$(\bar{\nu}_e)$ disappearance (3+1 scenario)

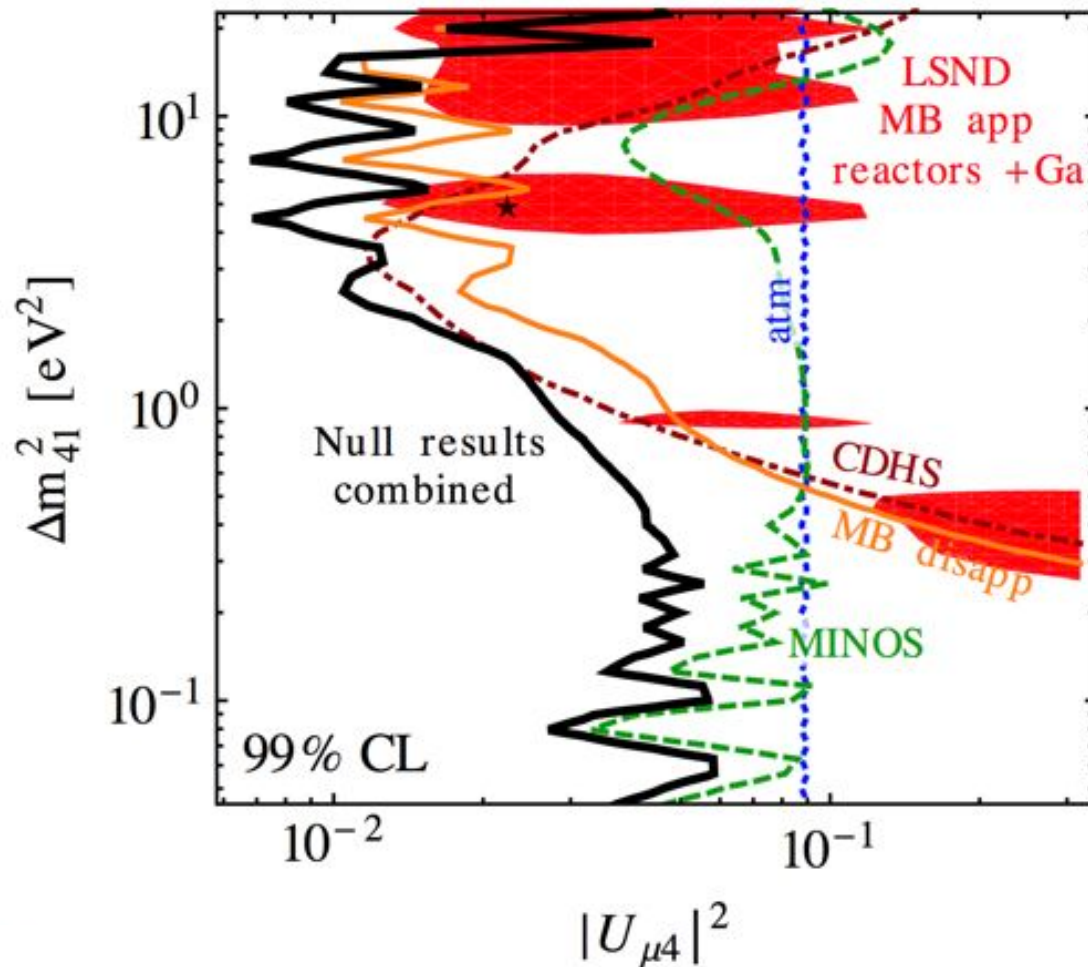


J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

	$\sin^2 2\theta_{14}$	Δm_{41}^2 [eV ²]	$\chi^2_{\min}/\text{dof}$ (GOF)	$\Delta\chi^2_{\text{no osc}}/\text{dof}$ (CL)
SBL rates only	0.13	0.44	11.5/17 (83%)	11.4/2 (99.7%)
SBL incl. Bugey3 spect.	0.10	1.75	58.3/74 (91%)	9.0/2 (98.9%)
SBL + Gallium	0.11	1.80	64.0/78 (87%)	14.0/2 (99.9%)
global ν_e disapp.	0.09	1.78	403.3/427 (79%)	12.6/2 (99.8%)

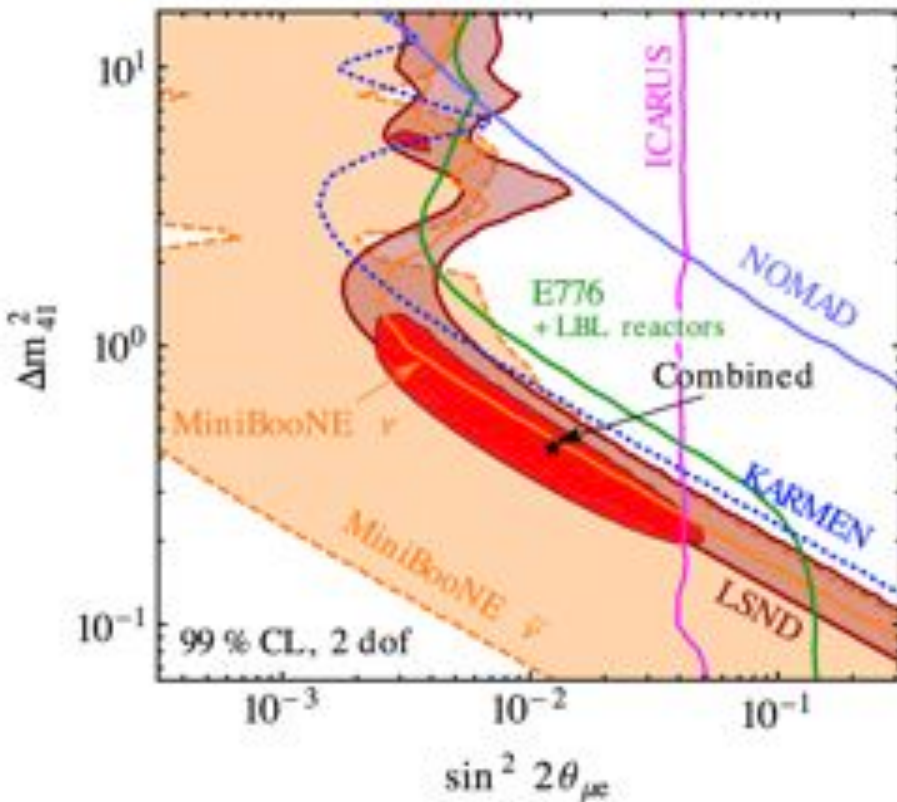
$(\bar{\nu}_\mu)$ disappearance (3+1 scenario)

No evidence for disappearance



J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

$(\bar{\nu}_e)$ appearance (3+1, 3+2 scenarios)



J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

	χ^2_{3+1}/dof	χ^2_{3+2}/dof	$\chi^2_{1+3+1}/\text{dof}$
LSND	11.0/11	8.6/11	7.5/11
MiniB ν	19.3/11	10.6/11	9.1/11
MiniB $\bar{\nu}$	10.7/11	9.6/11	12.7/11
E776	32.4/24	29.2/24	31.3/24
KARMEN	9.8/9	8.6/9	9.0/9
NOMAD	0.0/1	0.0/1	0.0/1
ICARUS	2.0/1	2.3/1	1.5/1
Combined	87.9/66	72.7/63	74.6/63

- Global fit to all appearance data is consistent
- Improvement of the fit by adding 2 sterile neutrinos

Appearance / Disappearance Coupling

- Electron disappearance channel

- $P_{ee} = 1 - \sin^2 2\theta_{ee} \sin^2 \frac{\Delta m_{41}^2}{4E}$ & $\sin^2 2\theta_{ee} = |U_{e4}|^2 (1 - |U_{e4}|^2)$

- Muon disappearance channel

- $P_{\mu\mu} = 1 - \sin^2 2\theta_{\mu\mu} \sin^2 \frac{\Delta m_{41}^2}{4E}$ & $\sin^2 2\theta_{\mu\mu} = |U_{\mu 4}|^2 (1 - |U_{\mu 4}|^2)$

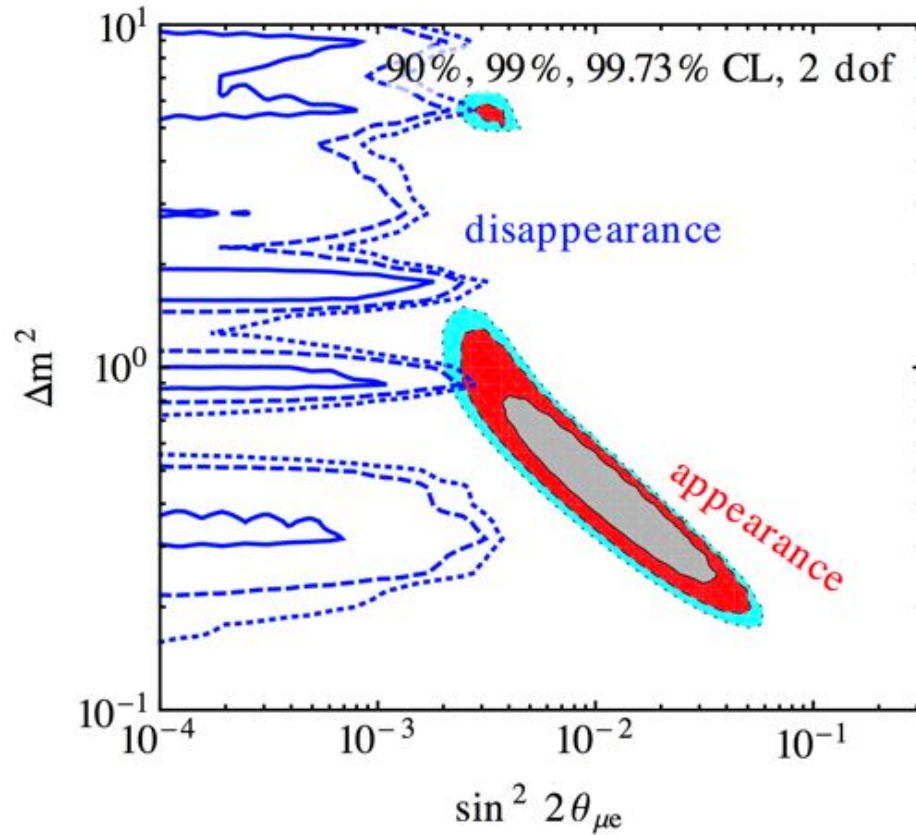
- Electron appearance channel

- $P_{\mu e} = 4 \sin^2 2\theta_{\mu e} \sin^2 \frac{\Delta m_{41}^2}{4E}$ & $\sin^2 2\theta_{\mu e} \approx \frac{1}{4} \sin^2 2\theta_{ee} \sin^2 2\theta_{\mu\mu}$

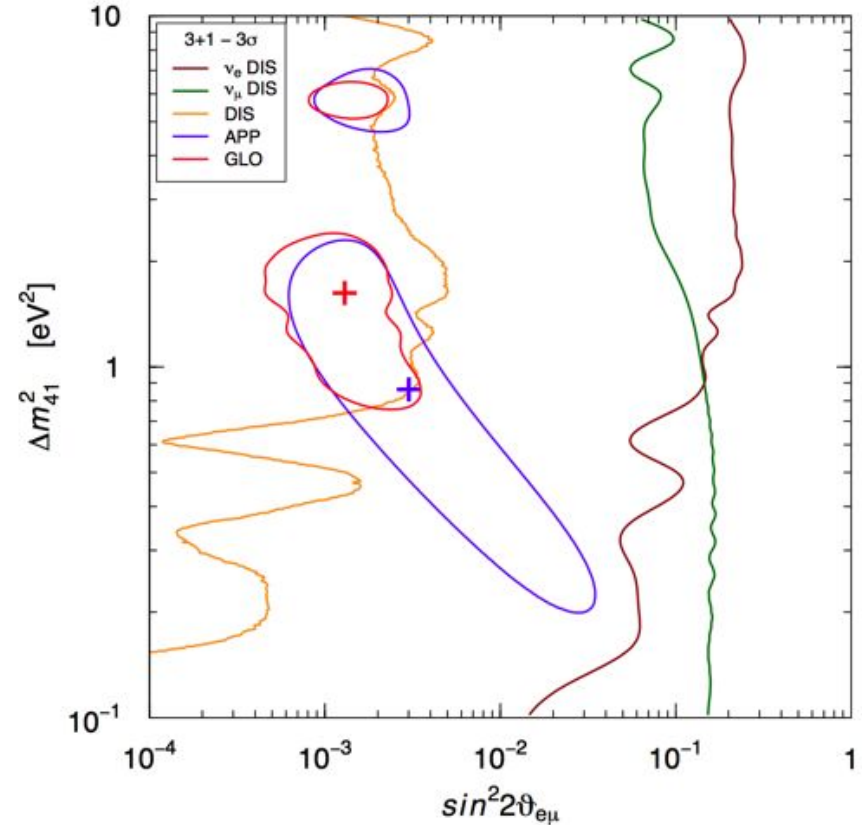
$\nu_\mu \rightarrow \nu_e$ appearance required both ν_μ & ν_e disappearance

Appearance / disappearance tension

J. Kopp et al., arXiv:1303.3011



C. Giunti et al., arXiv:1302.6720



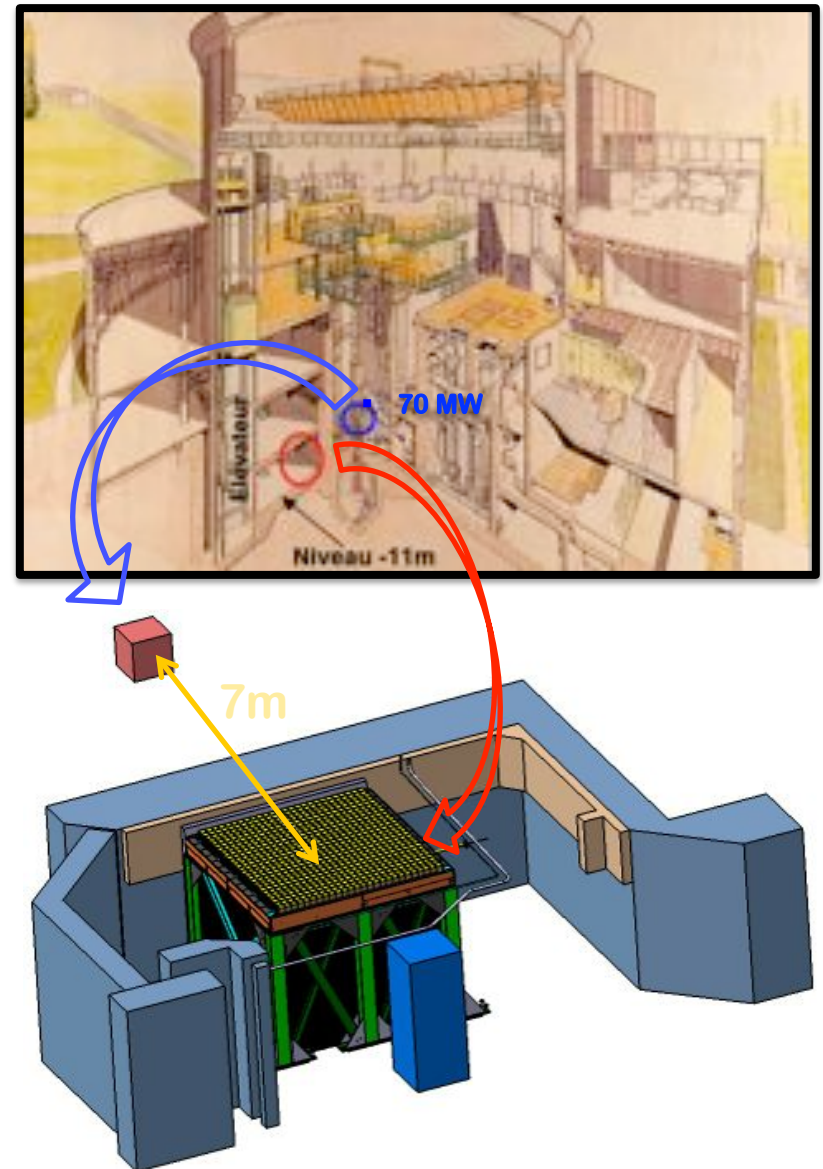
Testing the eV-scale Sterile Neutrino Hypothesis

Testing $(\bar{\nu}_e)$ disappearance anomalies

- GA & RAA arise from comparisons between data and event prediction → **Need a conclusive technique**
- Input from Sterile Neutrino Fits
 - $\Delta m^2 \approx 0.1-10 \text{ eV}^2 \rightarrow L_{\text{osc}}(\text{m}) = 2.5 \frac{E(\text{MeV})}{\Delta m^2(\text{eV}^2)} \approx 2-10 \text{ m}$
 - $\sin^2(2\theta_{\text{new}}) \approx 0.1$
- **Experimental Specifications**
 - Search for L, E, L/E pattern (shape only)
 - Complement with a rate analysis (direct test of RAA+GA)
 - $\Delta m^2 \approx \text{eV}^2$: compact source <1m & good vertex resolution (<1m)
 - $\sin^2(2\theta_{\text{new}})$: experiment with few % stat. syst. uncertainties

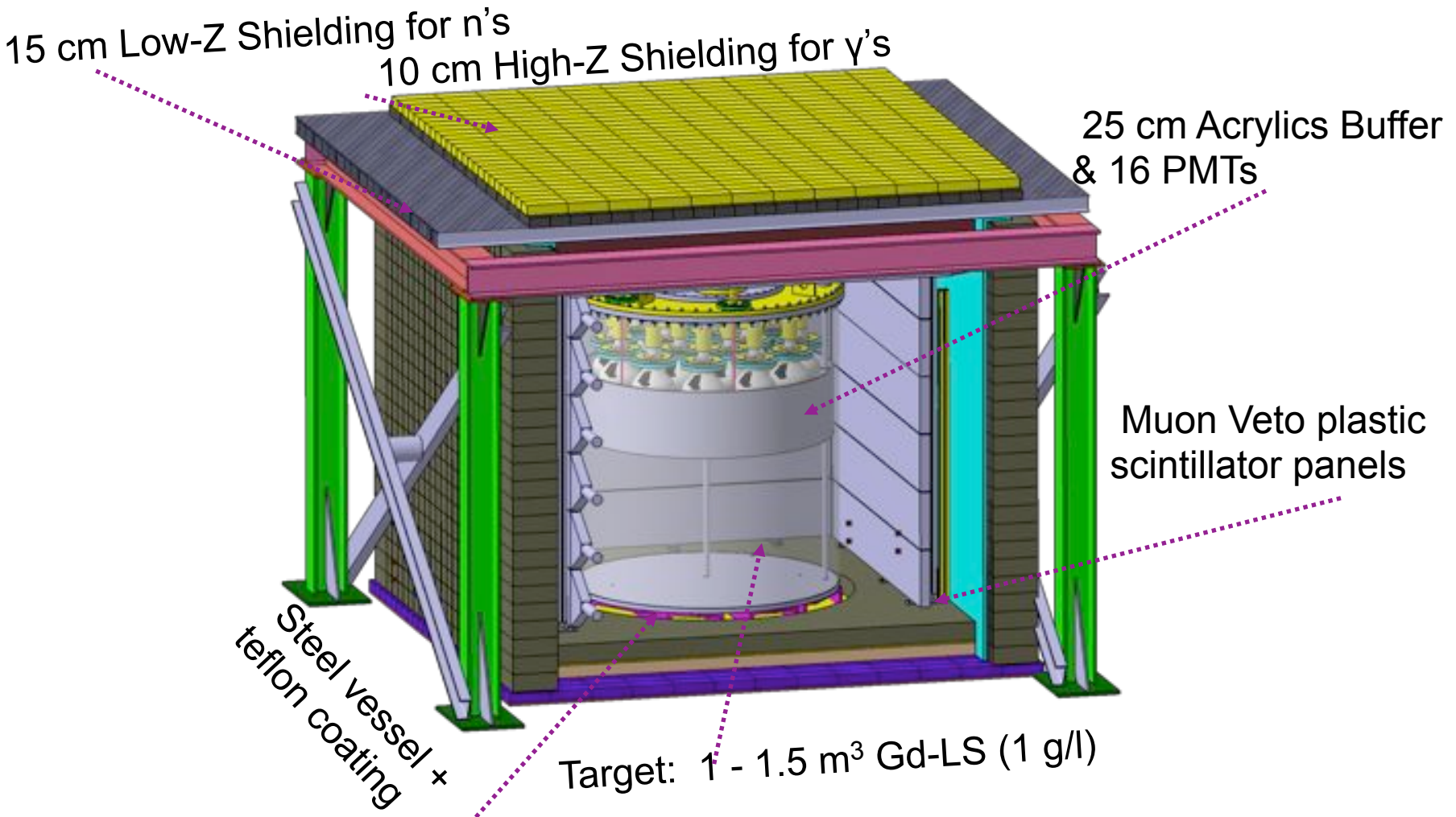
Nucifer

- **Osiris research reactor**
 - At Saclay, France
 - 70 MW, 20% ^{235}U
 - **Compact**: 60x60x60 cm³
- Detector designed for reactor monitoring studies
 - 850 kg Gd-loaded liquid scintillator
 - 350 int. expected / day
- But modest oscillation detection capabilities:
 - Short baseline: only **7 m**
 - Target: $h = 70 \text{ cm}$, $\Phi = 1.2 \text{ m}$



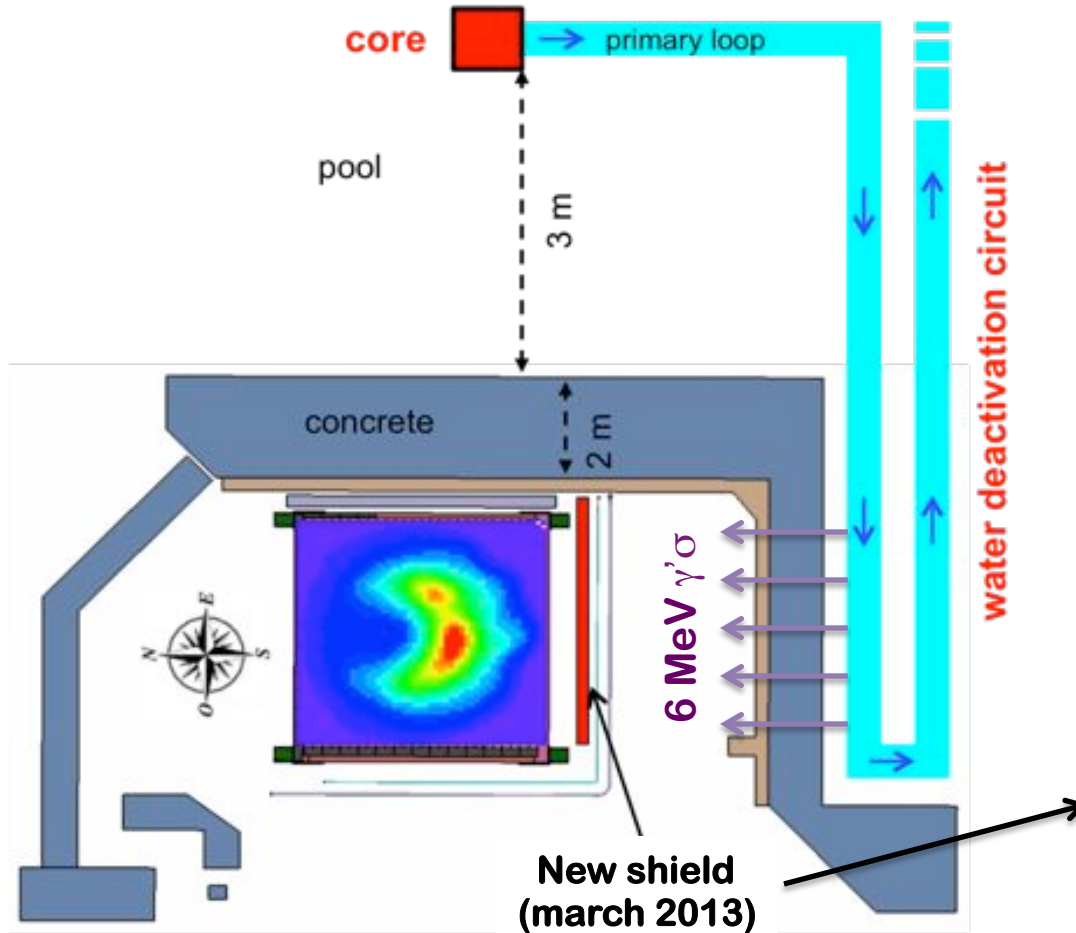
Nucifer Detector

- Design accounting for IAEA constraints : A simple monolithic Gd-LS detector with PMTs reading from the top (Nitrogen sealed, small, safe, robust, non intrusive)



Nucifer Upgrade

First run in 2012 indicated a unexpected background from primary loop circuit

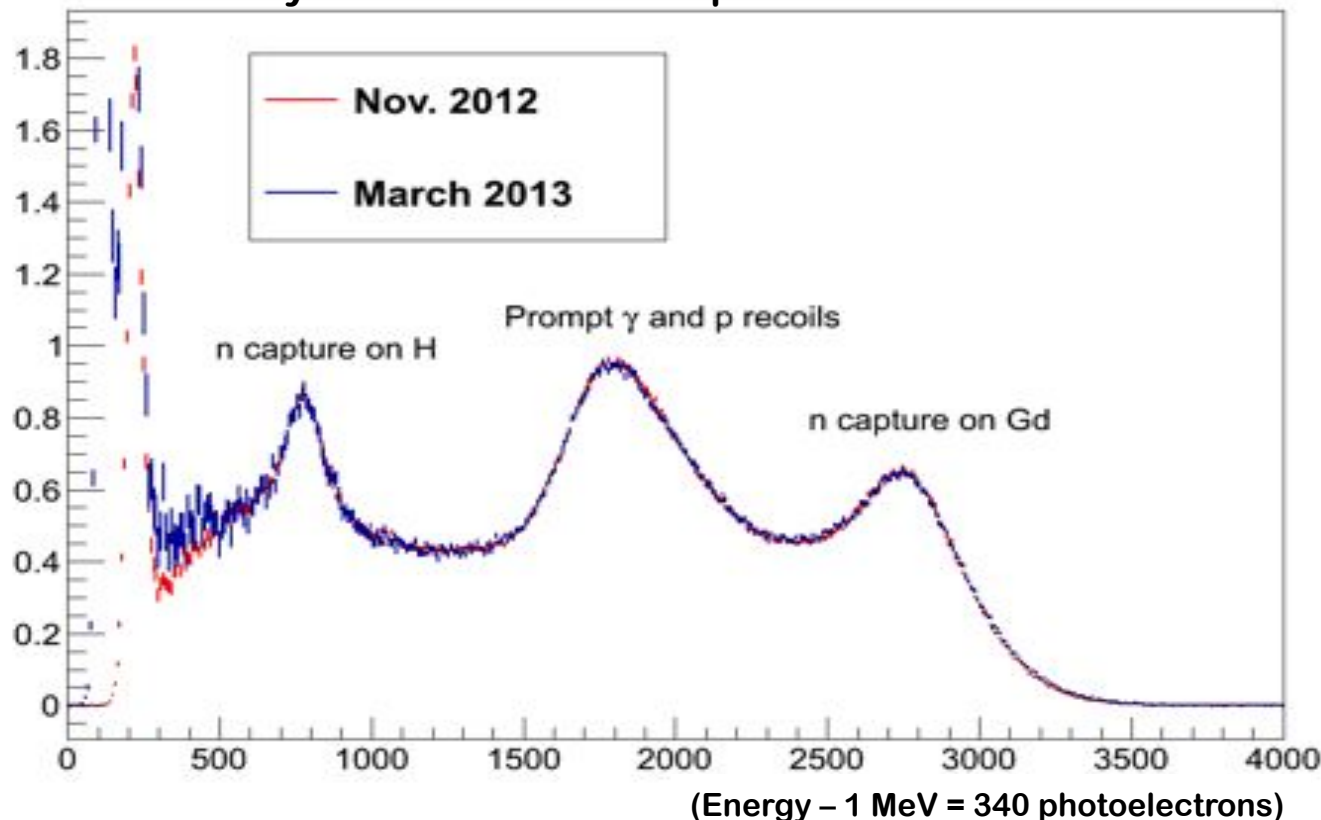


New lead wall erected early 2013
Accidental background has been reduced by a factor 500

Detector Operation

- Detector is working as designed
- Detector is running smoothly. Detector response is stable
- Nucifer has been operated remotely for several months.
- No safety incident has been reported

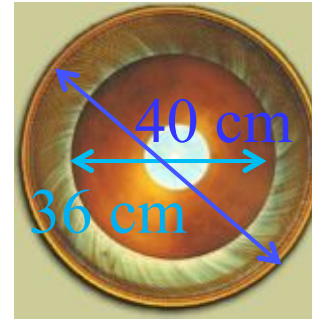
Very stable detector response over 5 months



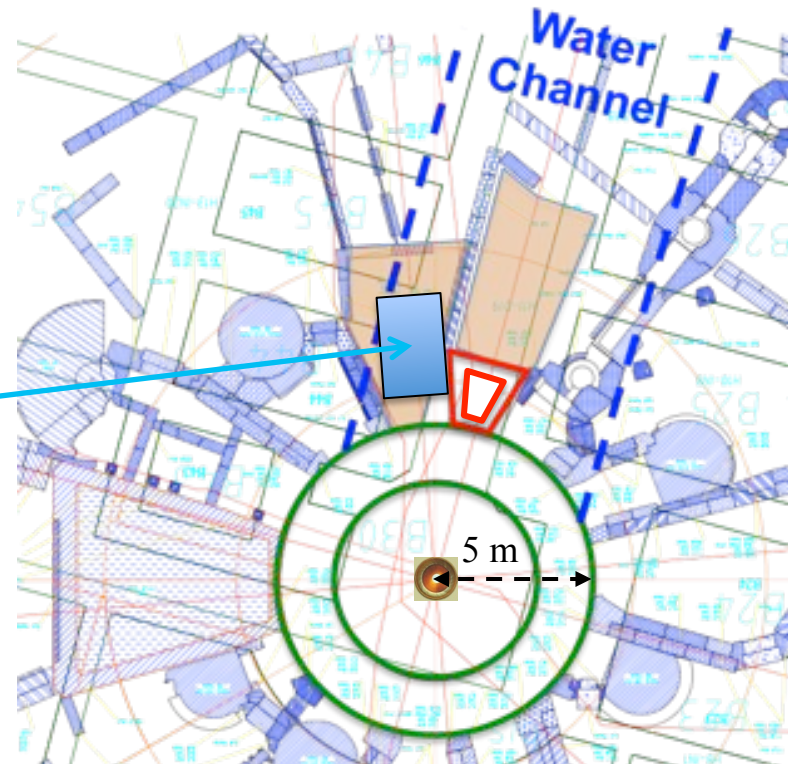
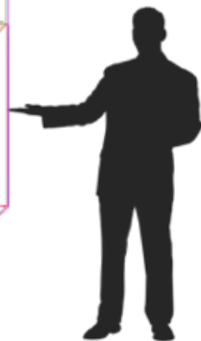
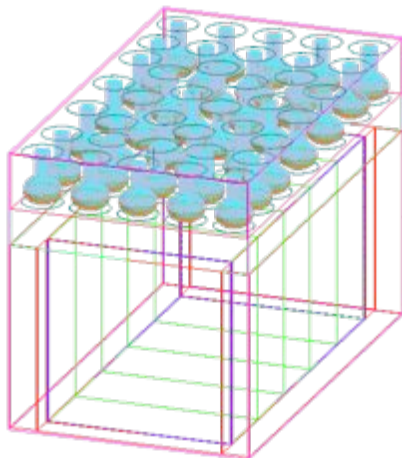
STEREO project

Contact: D. Lhuillier

- ILL research reactor (Grenoble):
 - 57 MW, highly enriched U
 - **Compact**: $h = 80$ cm, $\Phi = 40$ cm
- Dedicated detector:
 - 5 segments: **L and E oscillation**
 - Active outer layer: high eff.+veto
 - Muon flux divided by 4, thick CH_2 and Pb walls (70 t)



Target:
Gd-doped
liquid
scintillator



Oscillometry inside a v-detector

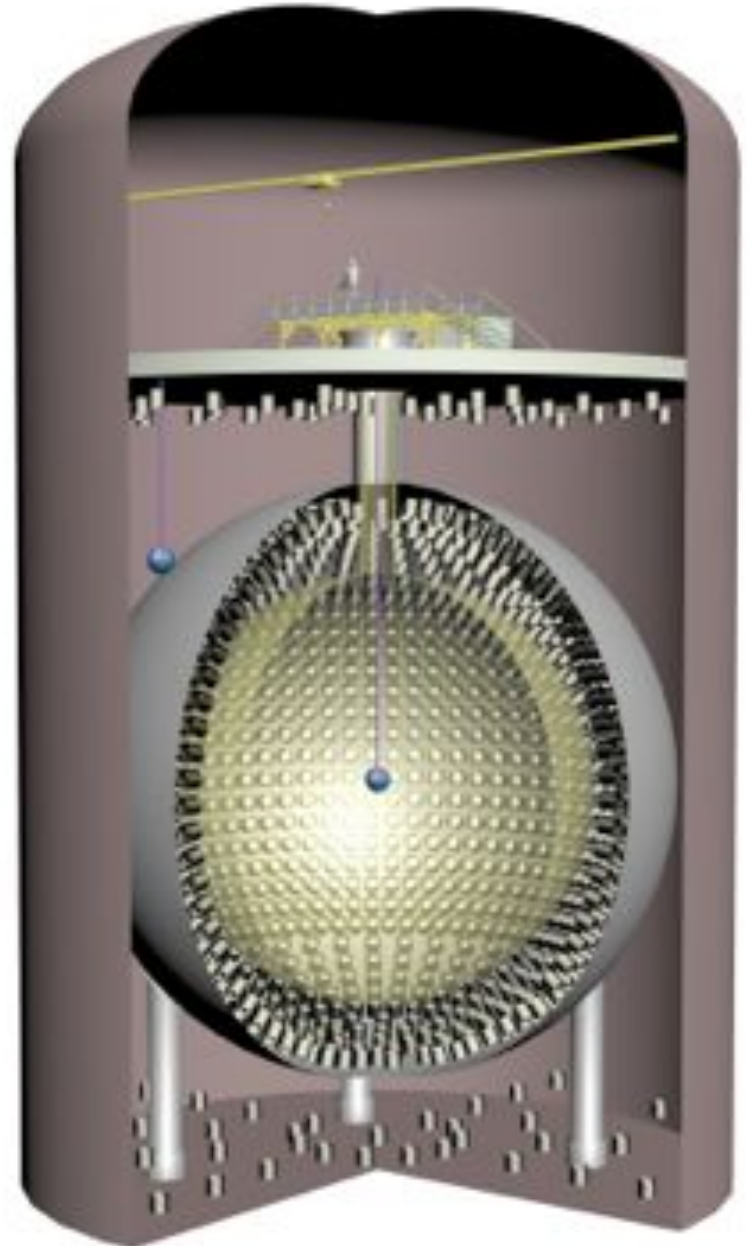
- Place the v-emitter inside or close to existing detectors
 - Very short Baseline (few m)
 - Low Background

i) v-source at center

- $$\frac{dN_\nu}{dR} \propto \left[1 - \sin^2(2\theta) \sin^2 \left(1.27 \frac{\Delta m^2 R}{\langle E \rangle} \right) \right]$$

ii) v-source Outside LS

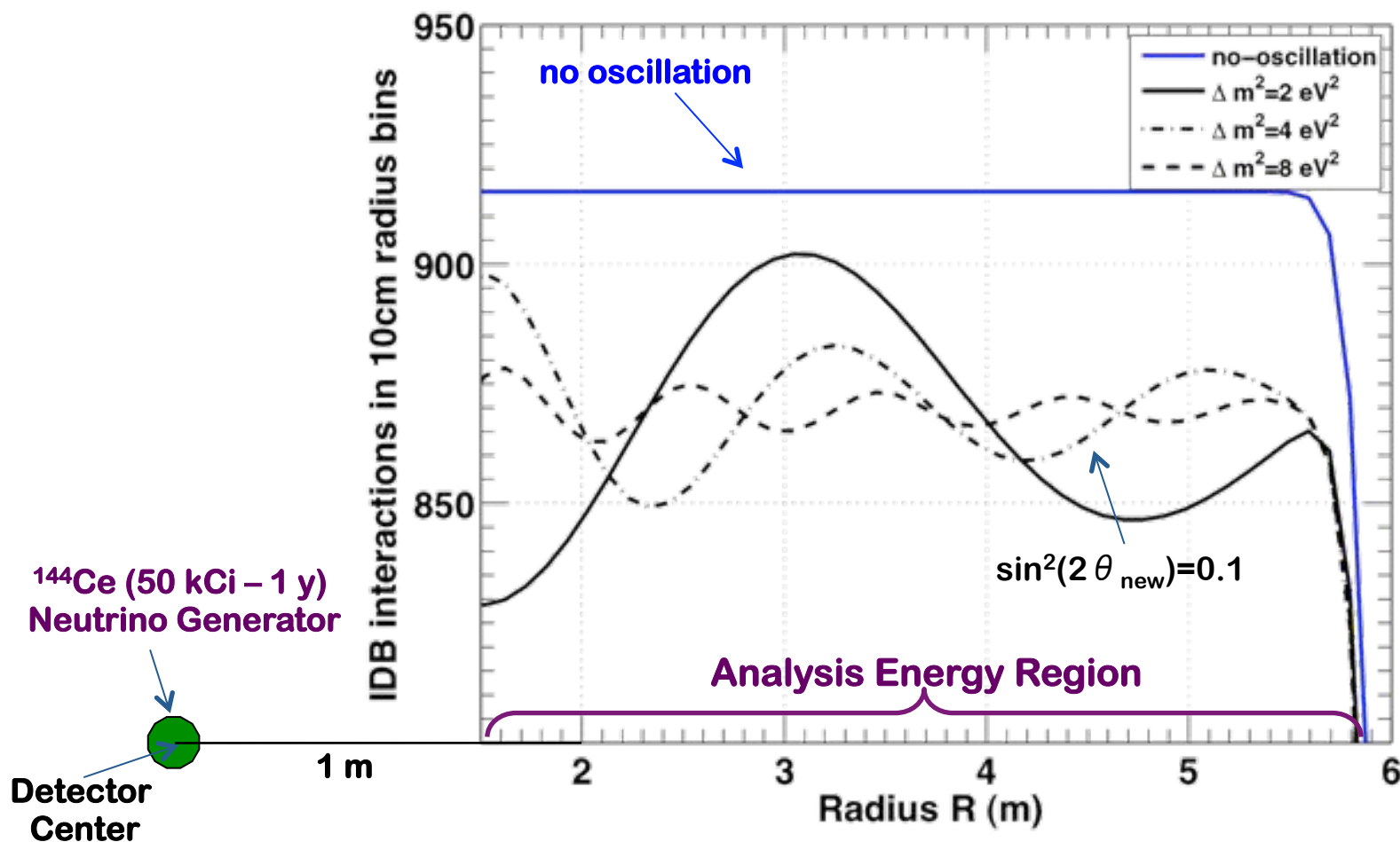
- Specific oscillation pattern analytically computable



Unambiguous Proof of $\nu_e \rightarrow \nu_s$ Oscillation

erc

$$\frac{dN}{dR}(R,t) \propto \frac{A(t)}{4\pi R^2} \times \langle \sigma \rangle \times N_p \times \cancel{4\pi R^2} \times P_{ee} \left(\frac{\Delta m^2 R}{\langle E \rangle} \right)$$



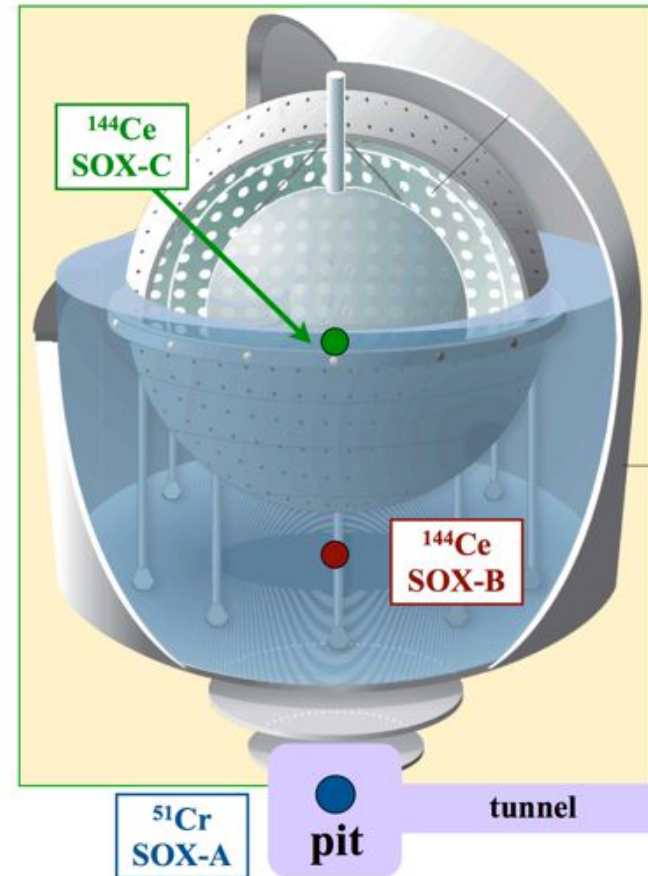
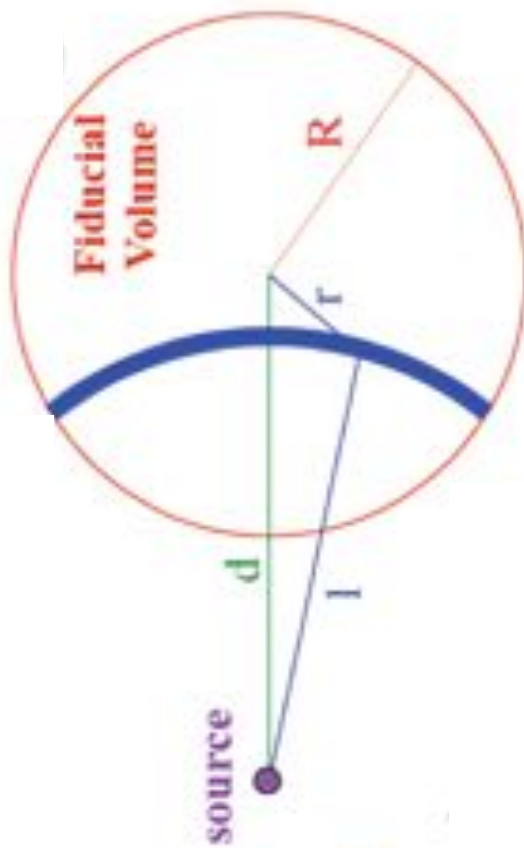
ν generator proposals

Type	channel	Background	Source	Production	Activity (Mci)		Proposal
ν_e	$\nu_e e \rightarrow \nu_e e$ Compton edge 5% E_{res} 15cm R_{res}	radioactivity (managable) Solar ν (irreducible) ν -Source (out ok but in ?)	^{51}Cr 0.75 MeV $t_{1/2}=26\text{d}$	n_{th} irradiation in Reactor	in	>3	Sage LENS
					out	>10	SOX* SNO+
			^{37}Ar 0.8 MeV $t_{1/2}=35\text{d}$	n_{fast} irradiation in Reactor (breeder)	in	>1	-
					out	5	Ricochet (NC)
$\bar{\nu}_e$	$\bar{\nu}_e p \rightarrow e^+ n$ $E_{th}=1.8\text{ MeV}$ (e^+, n) coincidence 5% E_{res} 15cm R_{res}	reactor ν & ν -Source $\rightarrow$ Background free!	^{144}Ce $E < 3\text{ MeV}$ $t_{1/2}=285\text{d}$	spent nuclear fuel reprocessing	in	0.075	CeLAND* SOX
					out	0.5	Daya-Bay
			^{90}Sr ^{106}Rh		-	-	-
			^{42}Ar	?	-	-	-

$^{51}\text{Cr}/^{144}\text{Ce}$ Source in Borexino (SOX)

- Existing Tunnel $\rightarrow$ source at 8.25 m from the LS target

Eur. Phys. J C8, 1999



- Observable: ν_e events as a function of distance (l)

^{51}Cr Source underneath Borexino

■ 10 MCi ^{51}Cr

- Re-use Gallex 36 kg of enriched chromium (38%)
- But need add. enriched ^{50}Cr

■ Reactors (Petten, Ludmila, US)

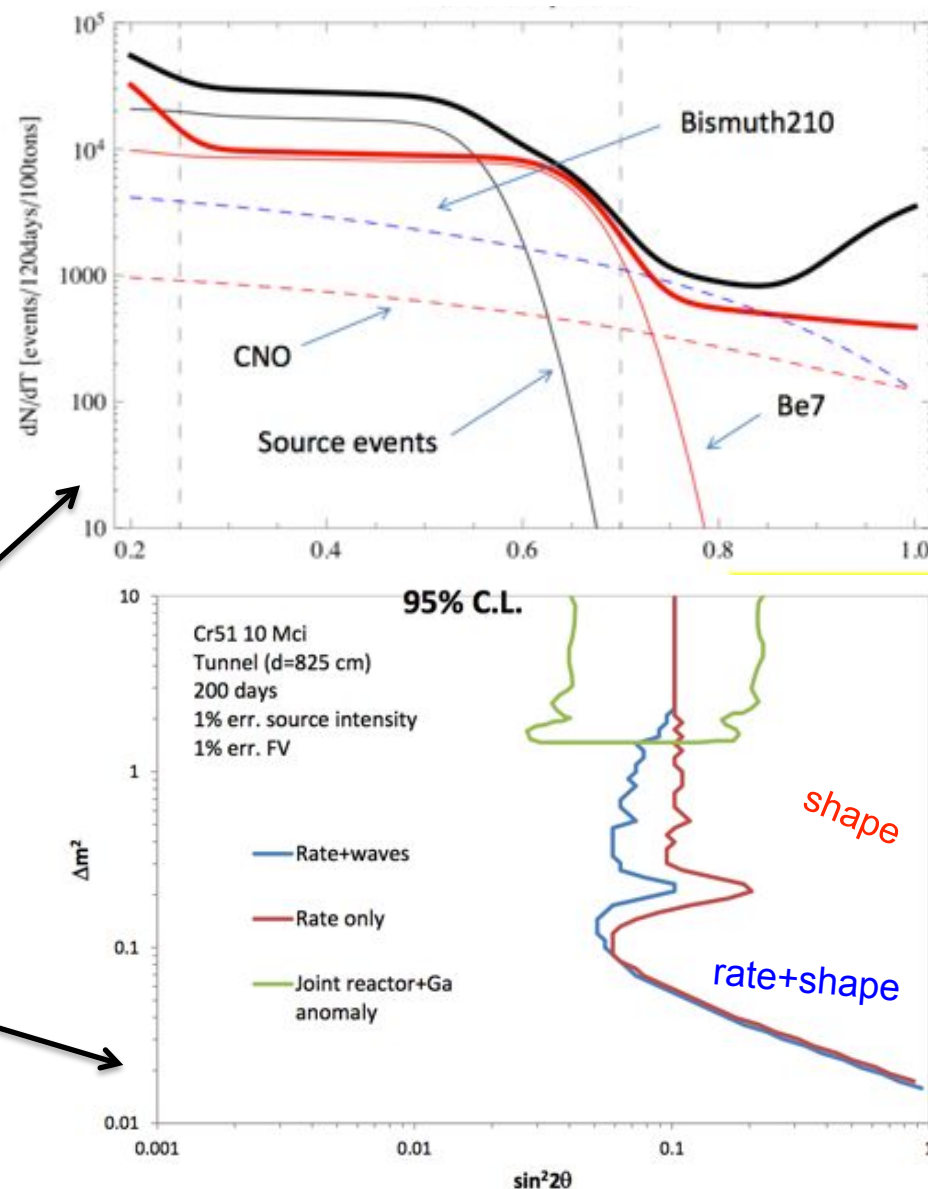
- $n_{\text{th}} \approx 10^{15} \text{ n/cm}^2/\text{sec}$
- Space to accommodate ^{50}Cr

■ Detection as ^7Be solar ν

- Well known background in 0.25-0.7 MeV: solar ν 's & ^{210}Bi
- 1% fiducial volume error

■ R+S Oscillometry analysis

■ ERC Funding



CeLAND: 75 kCi ^{144}Ce - ^{144}Pr in KamLAND

Phys. Rev. Lett. 107, 201801 (2011)

Antineutrino Source: ^{144}Ce - ^{144}Pr

(ITEP N°90 1994, PRL 107, 201801, 2011)

erc

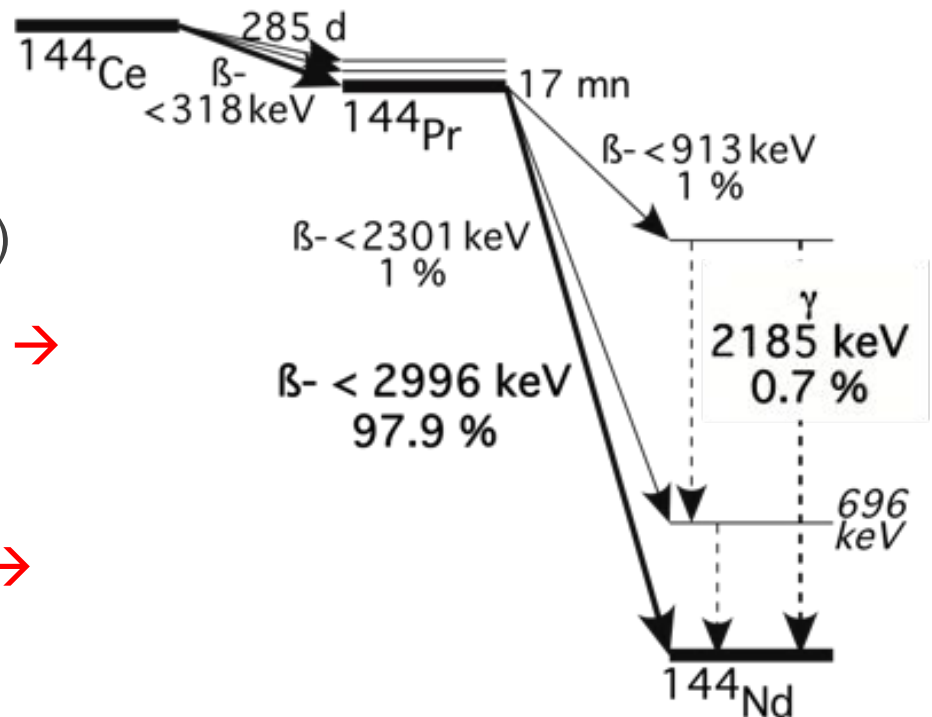
- 1st Trick: $\bar{\nu}_e$ source detected via $\bar{\nu}_e + p \rightarrow e^+ + n$ (Thr=1.8 MeV)
 - High IBD cross section $\rightarrow$ 10-100 **kCi activity**
 - (e^+ ,n) detected in coincidence $\rightarrow$ **Background free**

- 2nd Trick: **^{144}Ce - ^{144}Pr**

- Abundant fission product (5%)

- ^{144}Ce : long-lived & low- Q_β $\rightarrow$
Enough time to produce, transport, use

- ^{144}Pr : short-lived & high- Q_β $\rightarrow$
 $\bar{\nu}_e$ -emitter above threshold



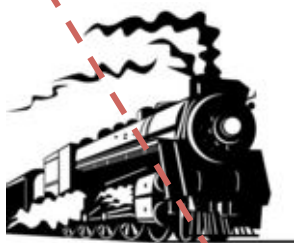
^{144}Ce Production at PA Mayak: 2014

erc

75 kCi (2.77 PBq), 4 kg of CeO_2 ($\rho = 4 \text{ g/cm}^3$), 600 W

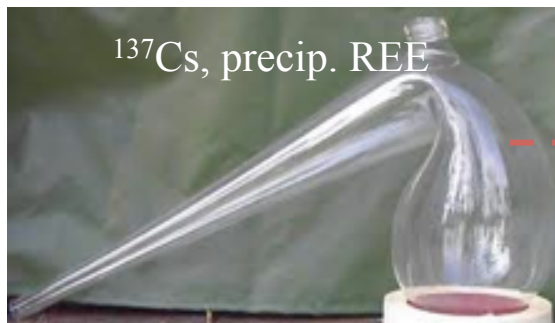


VVR-440, storage



TUK-6

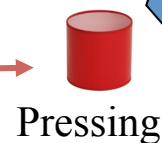
Cutting, digestion
Purex



Displacement
Chromatography



CeO_2 calcination

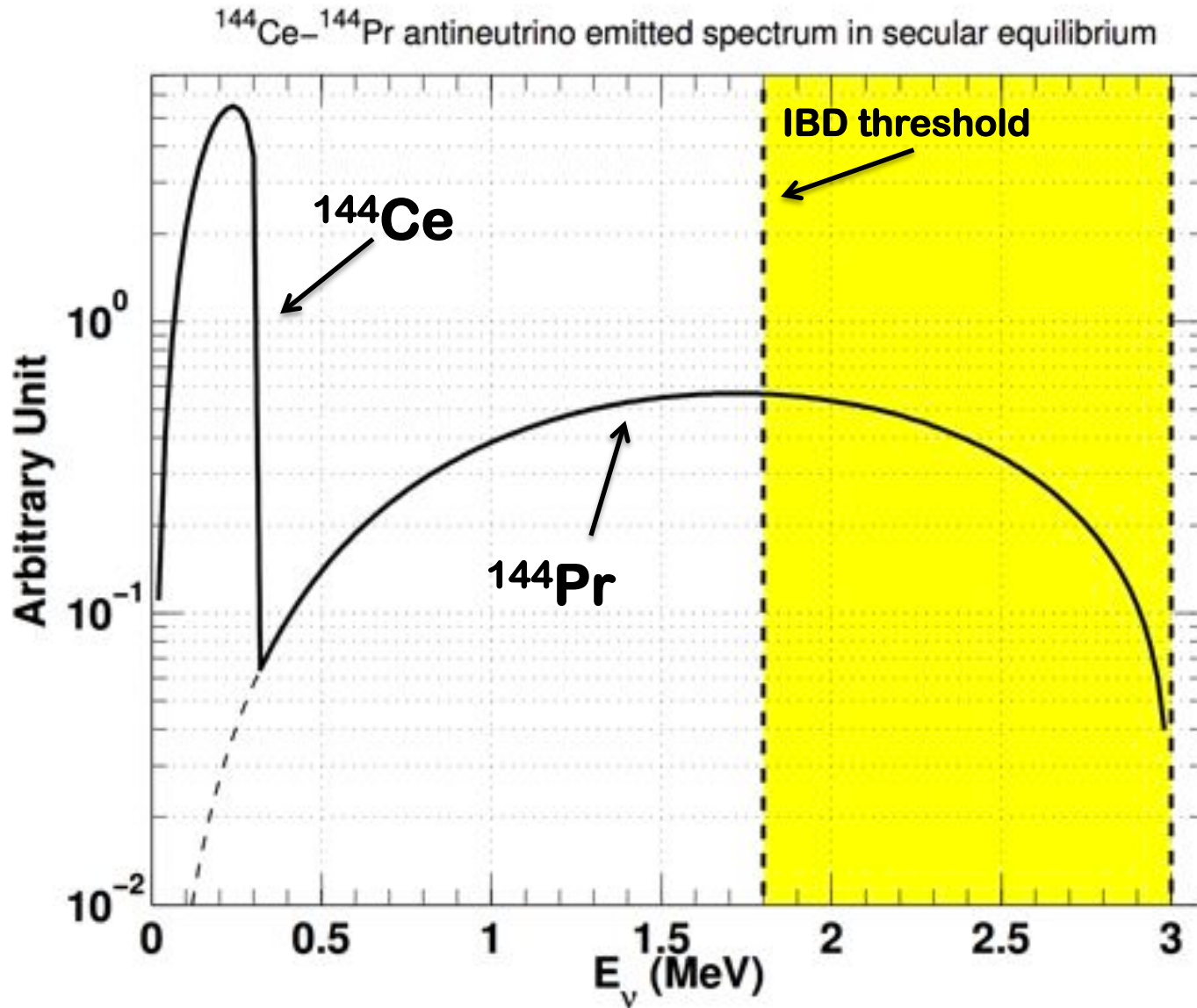


Pressing



2 tons
W-Shield

^{144}Ce - ^{144}Pr Antineutrino Spectra



75 kCi (2.78 PBq) ^{144}Ce - ^{144}Pr Source

erc

 $\phi 162$ exter. $\phi 152$ inter.

30

12.5

 $\phi 5$

ep. 2

ep. 2

153 inter.

Source capsule sketch

- $\rho(\text{CeO}_2) = 4 \text{ g/cm}^3$
- $H_{\text{cyl}} = D_{\text{cyl}} = 15 \text{ cm}$
- 23 g ^{144}Ce
- 10 kg of CeO_2
- 600 Watt

P	A	10-01-13	Origine	Scola	
St.	Ind.	Date	Réf. approb. ou modif.	Dess.	Vérif.
Ss-Ens. / Sub Assy: .					
Séparateur Separator		Ss-Traitant Supplier		PROVISoire	
0,00		0,00			
Quantité/Quantity: 1		Echelle/Scale : 1:1		Masse / Mass :	
Plan dessiné avec le système C.A.O. CATIA V5. Ne peut être modifié que par le même système. CATIA V5 C.A.D. drawing. Do not make manual revisions or alterations.					
Réf. C.A.O. : 21000001PA_Capsule.CATDrawing					
CeLAND ^{144}Ce Source Mechanics Capsule					
Ref. ext. : .			Contrat : .		
Il n'est permis d'utiliser ce dessin qu'avec licence spéciale ou autorisation expresse (loi du 11 mars 1957) This drawing may not be used without special license					
Planche/Sheet : 1/1					
71 3139 DM- 2100 000 PA					

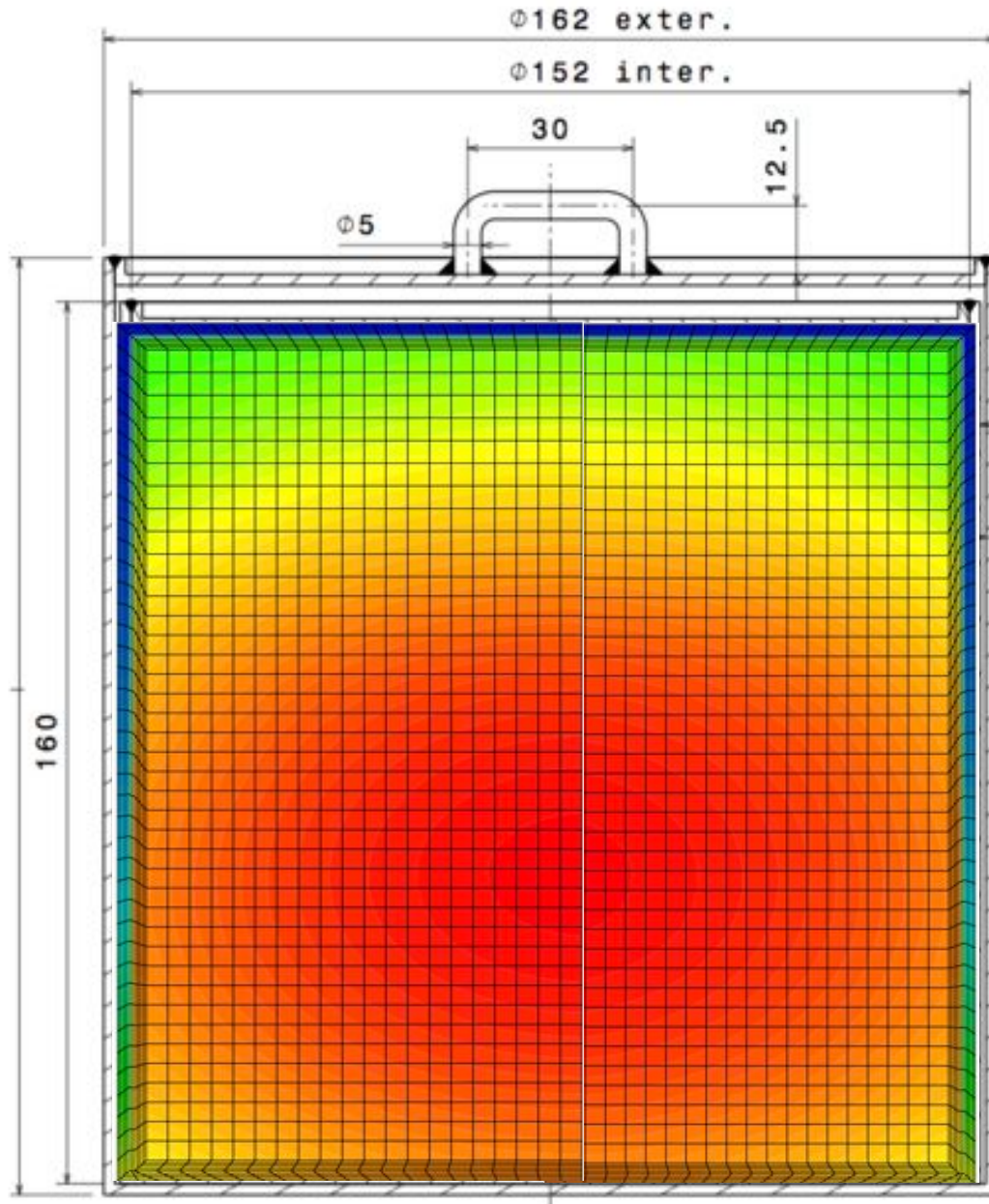
I r f u

cea

sacalay

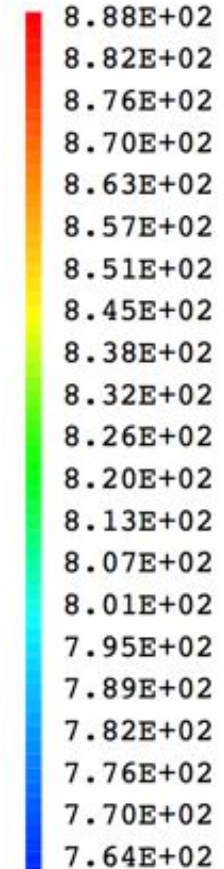
^{144}Ce - ^{144}Pr Capsule Temperature

erc



$T (^{\circ}\text{K})$

Conduction + Convection
No radiation loss

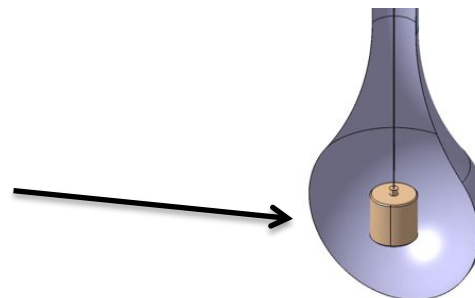


Steel OK

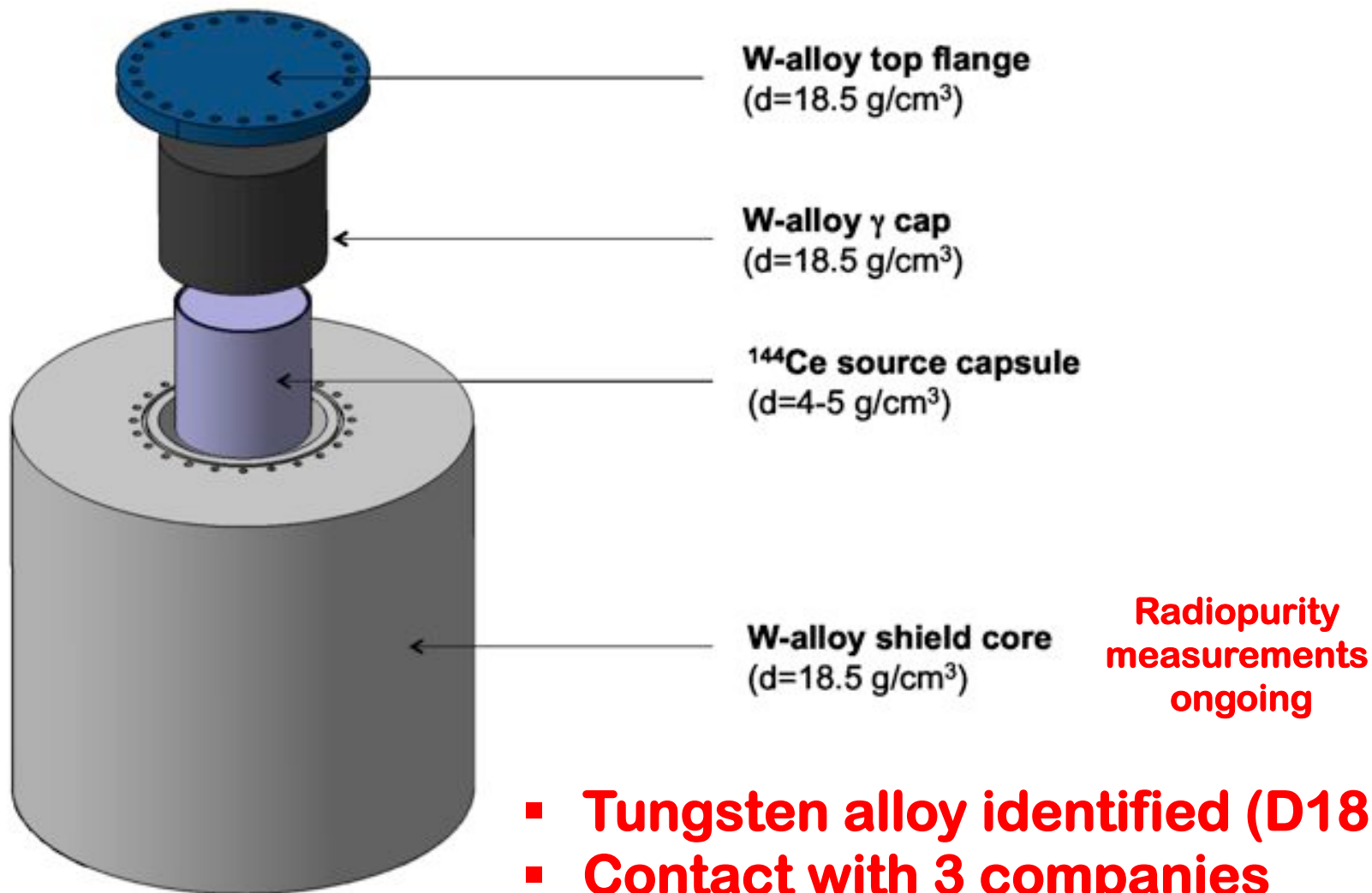
High-Z Shielding

erc

- **Two separate goals**
 - Usual biological protection
 - Can be achieved with ≈ 16 cm of W-alloy
 - Absorption of the $2.18 \text{ MeV } \gamma$'s
- **Deployment scenario in KamLAND**
 - 2015: deployment in Water Veto (OD) - Feasible
 - 6 months data taking
 - 20 cm W-alloy shield
 - 2016: deployment in Liquid Scintillator (ID) - Study
 - 1 year data taking
 - Need additional shielding
 - 35 cm W-alloy
 - Heavy liquid balloon (NaWO_3 , Hg)



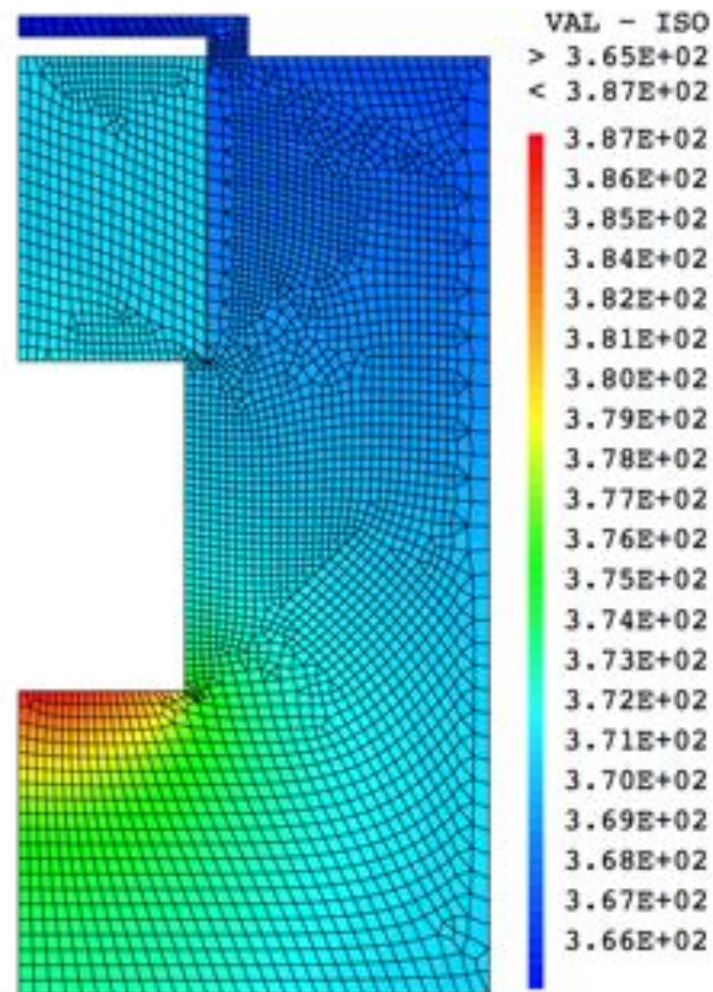
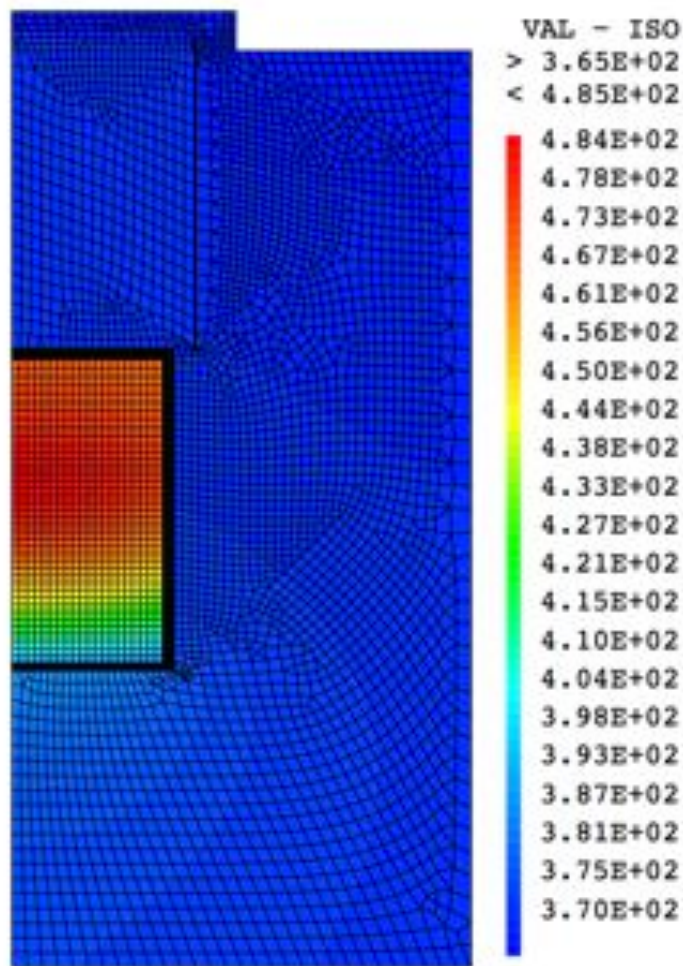
Biological & Phase 1 Shield



- Tungsten alloy identified (D185)
- Contact with 3 companies in Europe, Russia, China

Shielding Temperature

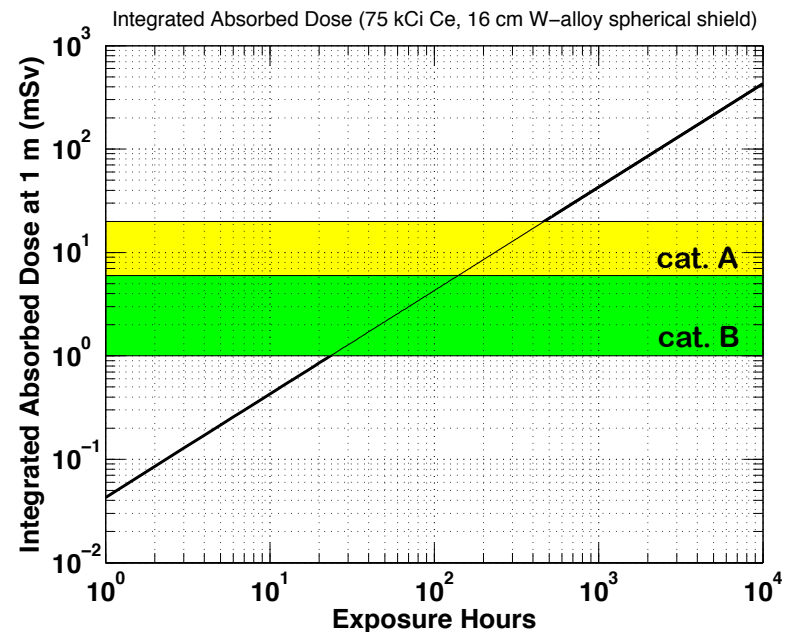
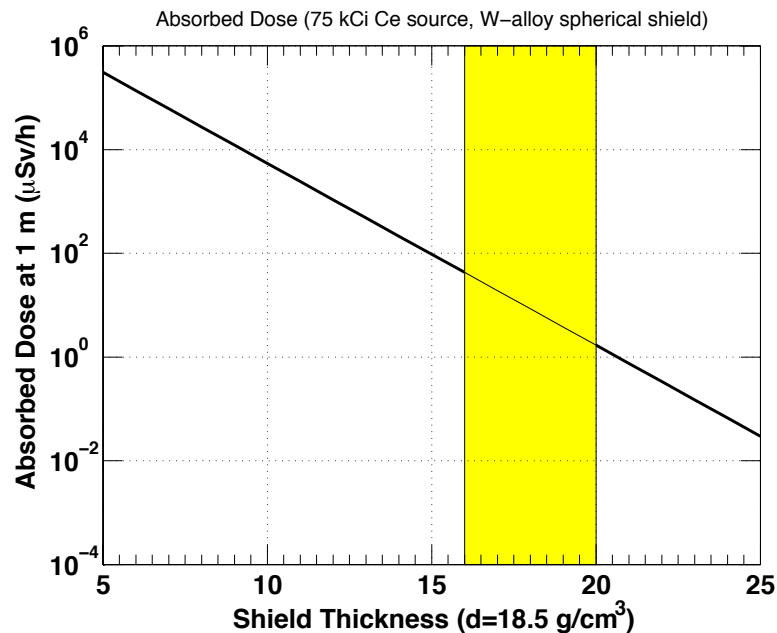
Shielding external temperature: $T_{\text{ext}} (38^\circ\text{K}) + 60^\circ\text{K}$



Biological Protection 16 cm D185

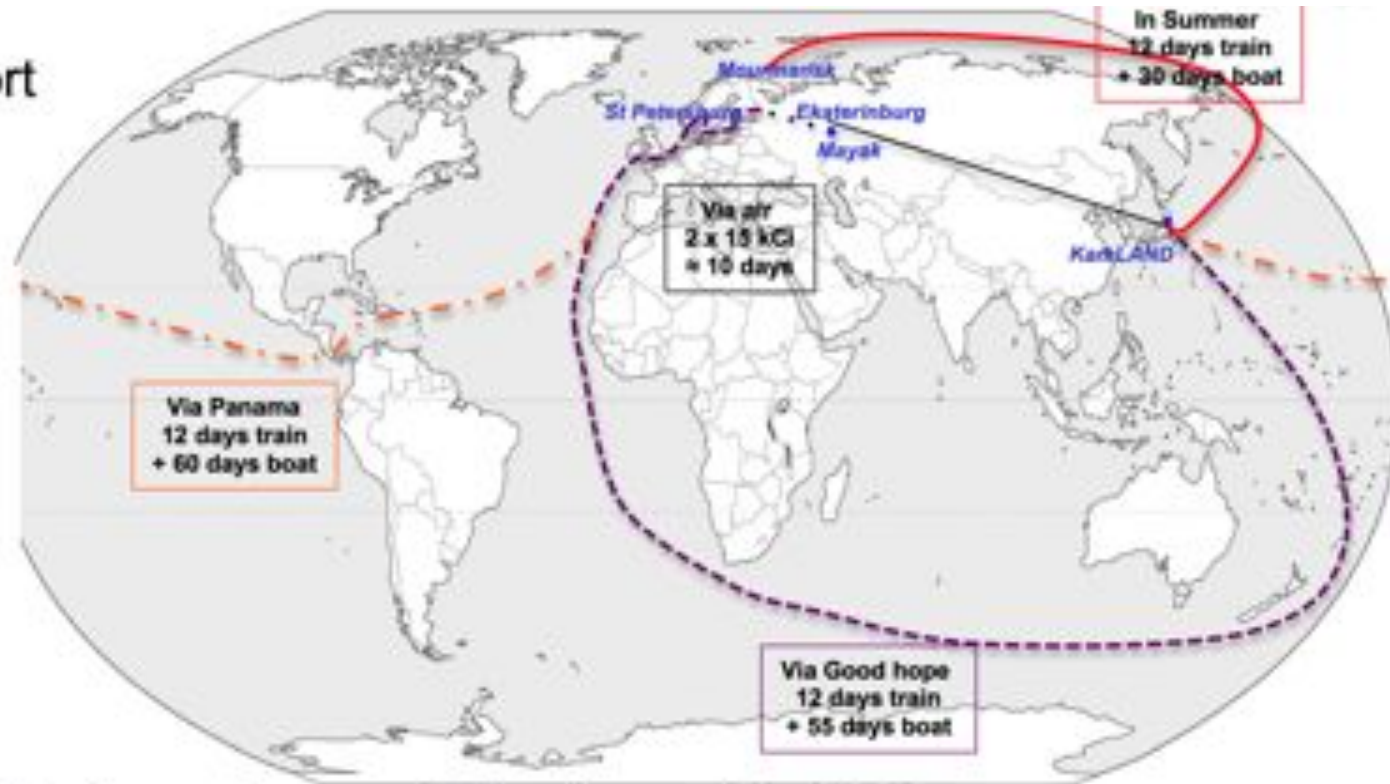
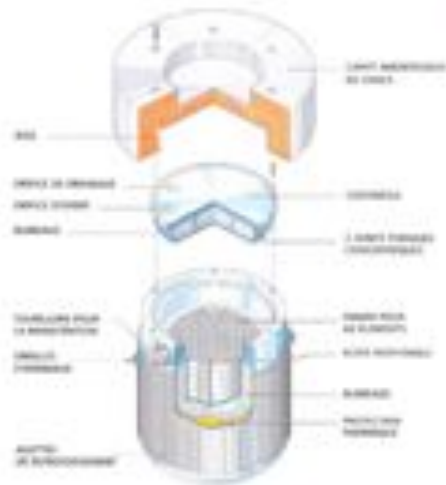
erc

- Back of the envelope dose @1m: $14 \mu\text{Sv/h}$
- GEANT4 dose @1m: $28 \mu\text{Sv/h}$
- **CEA-SPR computation dose with MCNP @1m: $42 \mu\text{Sv/h}$**
 - Regulation limit 2 mSv/h @1m? Applicable in Kamioka?



❖ Certified transport container

- 23 tons !



❖ Severe constraints based on regulation issued by IAEA

- nothing impossible, but long, bureaucratic and costly
- by air limit for each radioisotope : 16.2 kCi for ^{144}Ce , 2.4 MCi for ^{51}Cr
- by boat : only limited number of harbours agreed for radioactive materials

Transport Route to KamLAND

erc

- Transportation from Tokyo/Yokoyama harbor by truck (ok)
- Several difficulties to be solved for the final lab access



^{144}Ce - ^{144}Pr Source Calorimetry

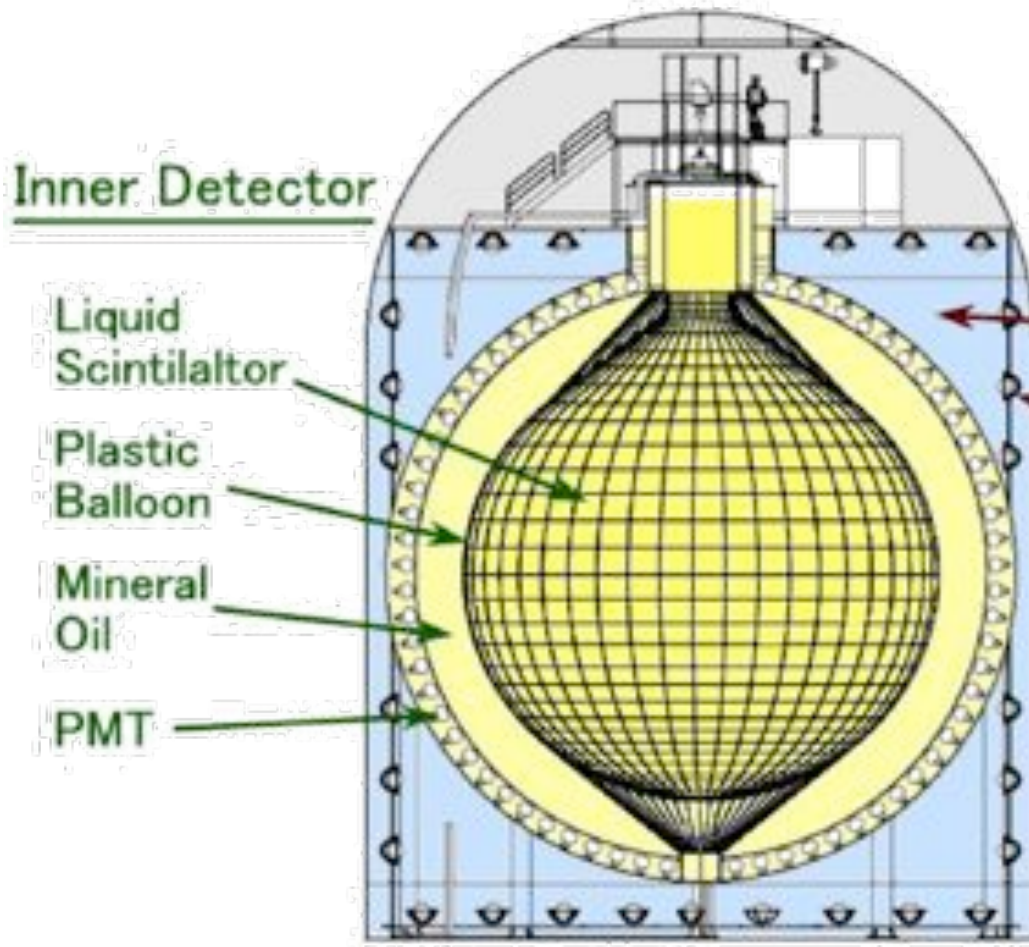
erc ■ Source Heat Release (75 kCi, 600 W, 96% β)

Time (months)	0	6	12	18	24	36	48	60	120
Activity (kCi)	75.0	48.1	30.8	19.8	12.7	5.2	2.1	0.9	0.01
Activity (PBq)	2.78	1.78	1.14	0.73	0.47	0.19	0.08	0.03	0.0004
Heat (W)	592.5	379.9	243.5	156.1	100.1	41.1	16.9	6.9	0.08

■ Calorimetry

- Precision requested for the activity : <1.5%
 - Achieved for Gallex / Sage
- Realization of a calorimeter (2013/14)
- Measurement at the KamLAND site just before deployment, for few days

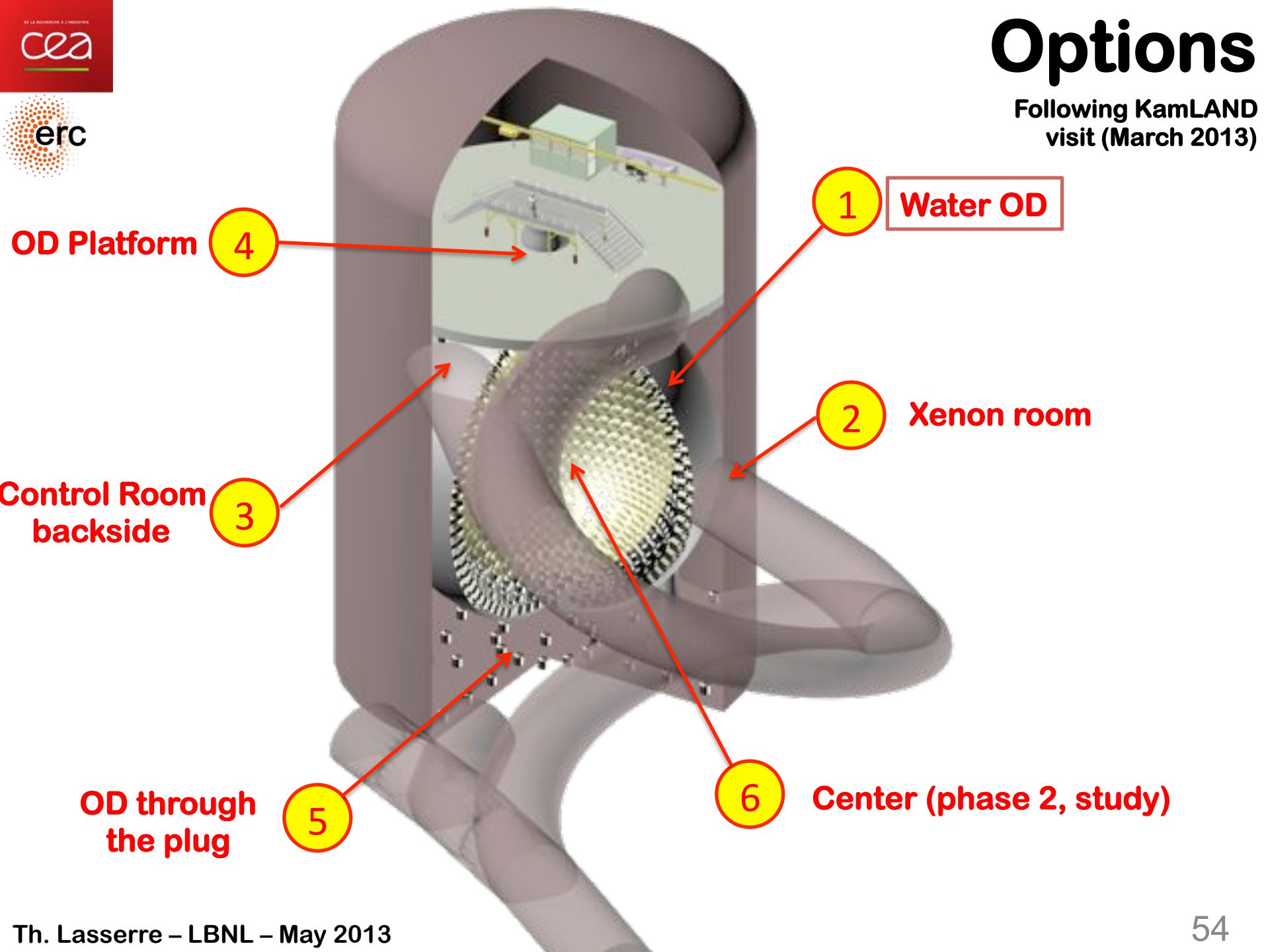
CeLAND: ^{144}Ce in KamLAND



- **A great existing underground detector**
 - 1000 tons of PC
 - +mineral oil
- **But several constraints**
 - Full of extra pure mineral oil
 - Avoid contaminations
- **The entrance hole**
 - 55 cm in diameter
 - Complex operations to insert the source
- **Hanging suspension**
 - Phase 1: OD Water
 - Phase 2: ID Oil

Options

Following KamLAND
visit (March 2013)



Deployment in Water OD

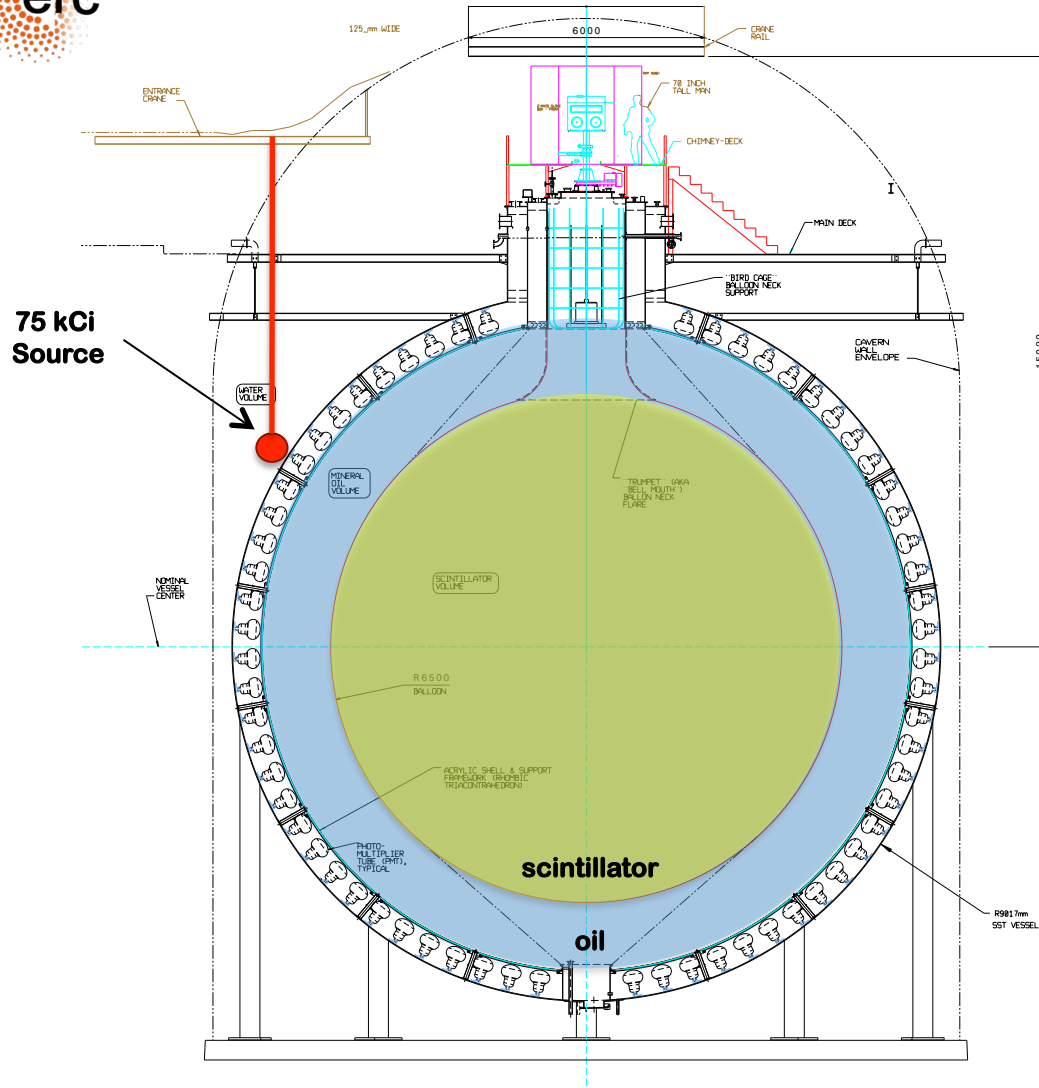
erc

- Deployment in Water OD through the platform man-hole
- Hanged from a new supporting crane using existing anchors
- Run in parallel with KamLAND-Zen



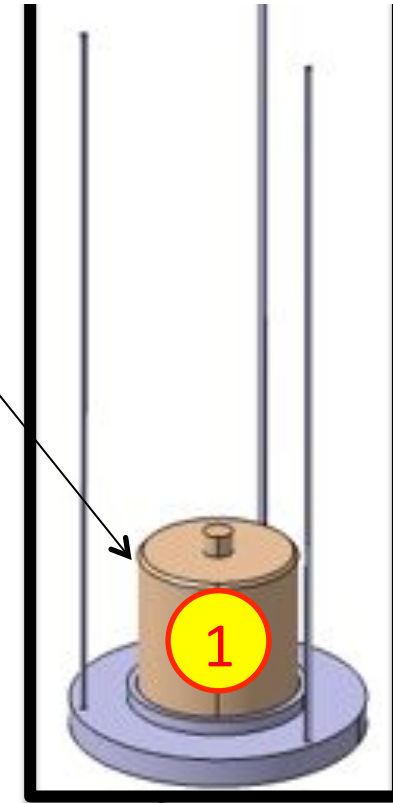
CELAND Phase 1: 2015 in KamLAND

erc



Source @2.5 m away from LS
75 kCi & 6 months of data taking

tungsten alloy,
54 cm
 $d=18.5 \text{ g/cm}^3$

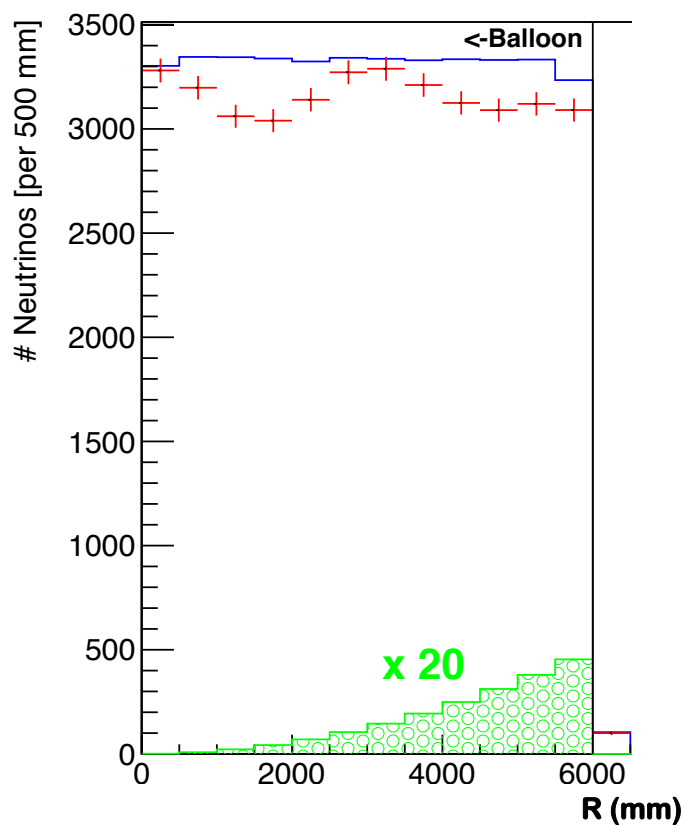


SS structure to isolate the source from
the OD Water (no contamination risk)

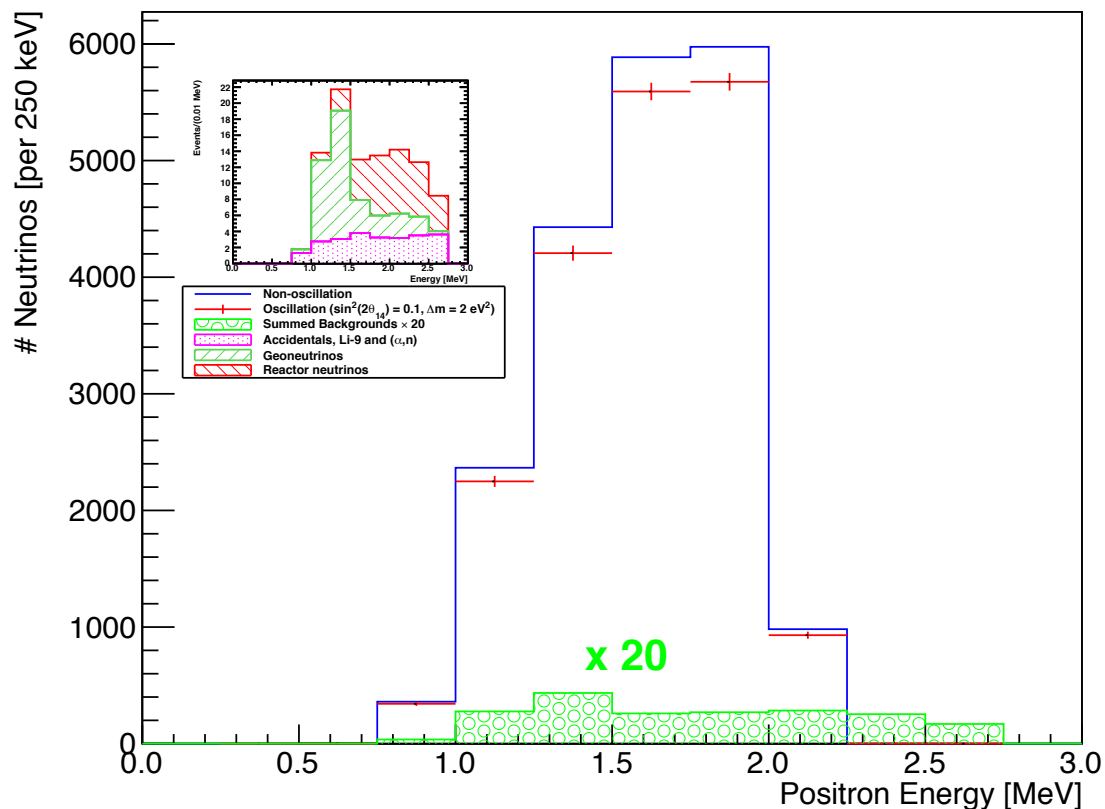
Phase 1: signal/backgrounds

75 kCi Source @2.5 m away from LS
9.3 m from the detector center

Spatial event distribution



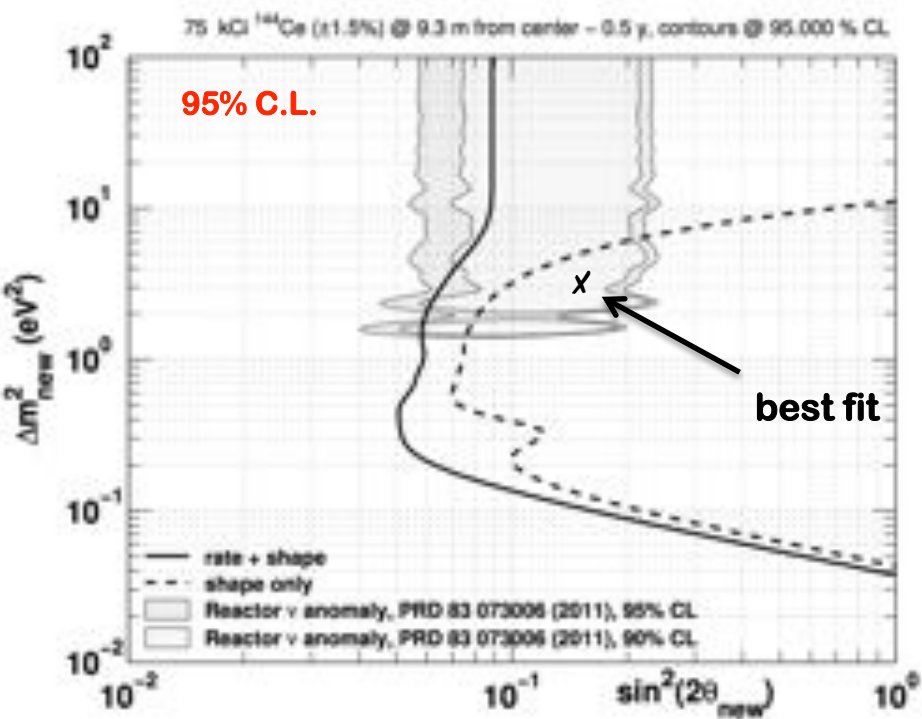
Energy Spectrum



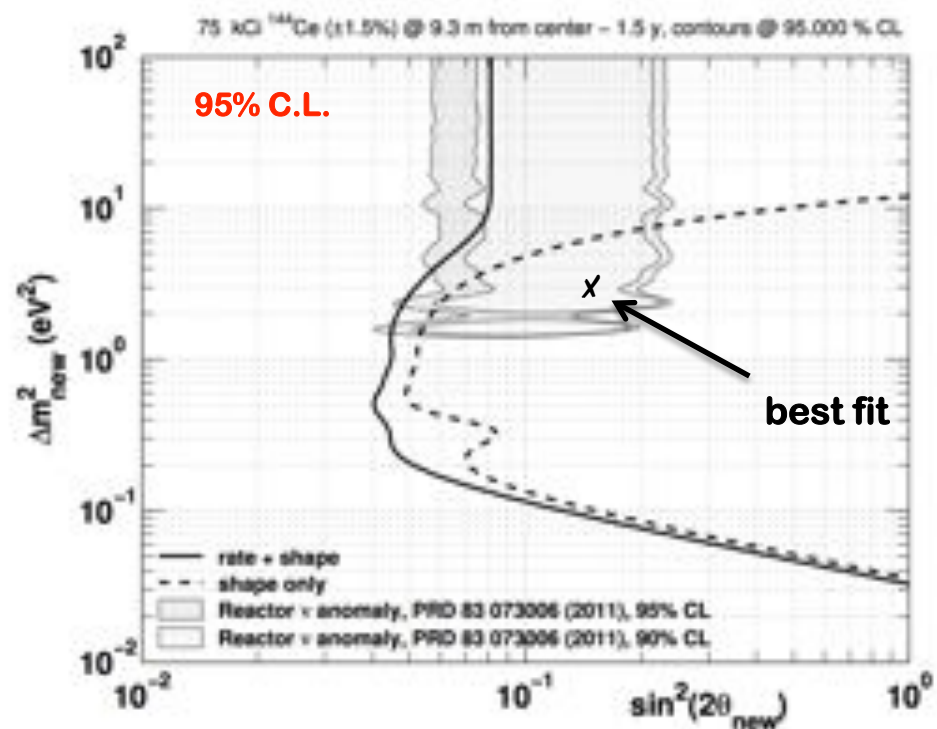
CeLAND phase 1 sensitivity

75 kCi ^{144}Ce - ^{144}Pr – 9.3 m from detector center

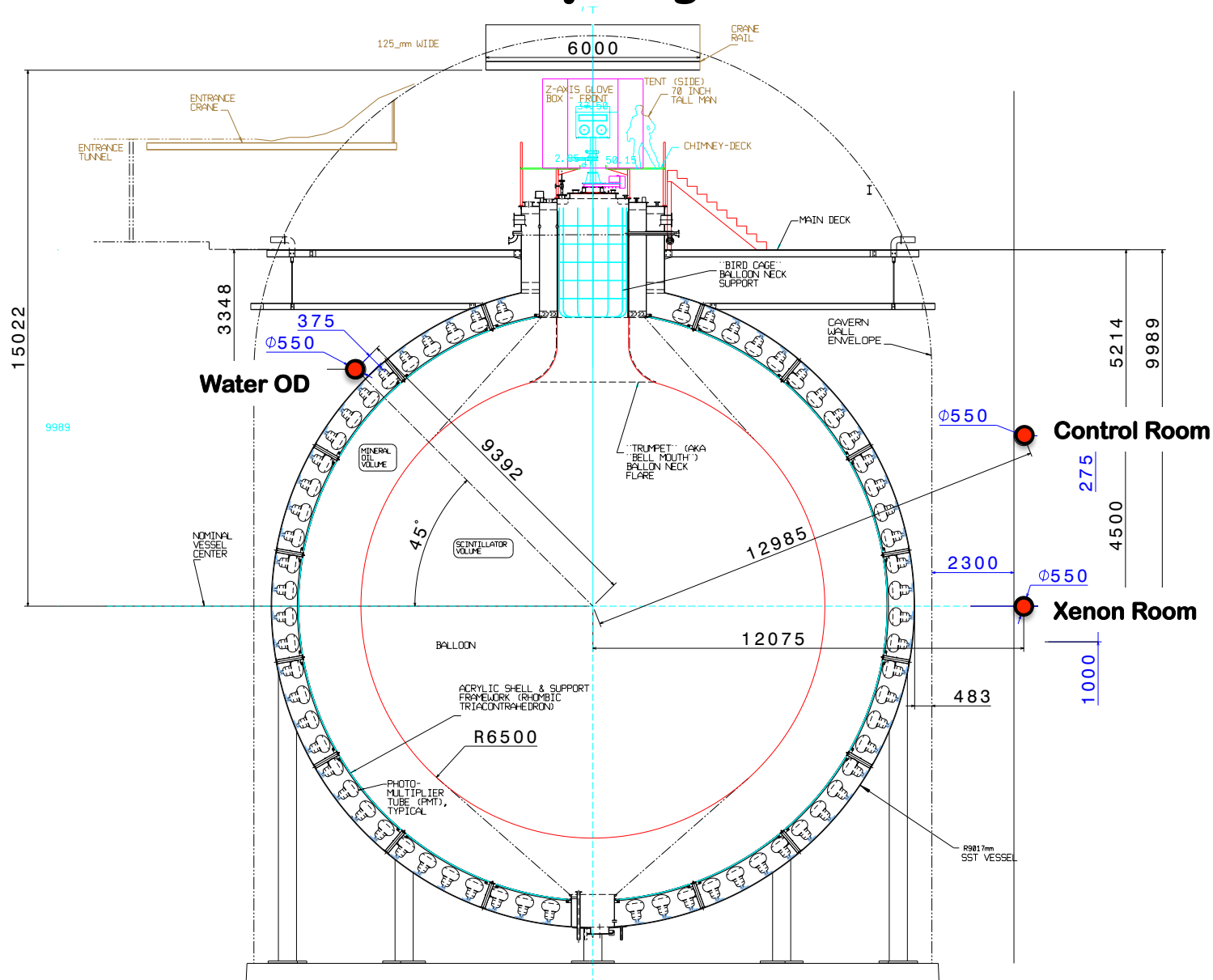
0.5 year of data (10 keVts)



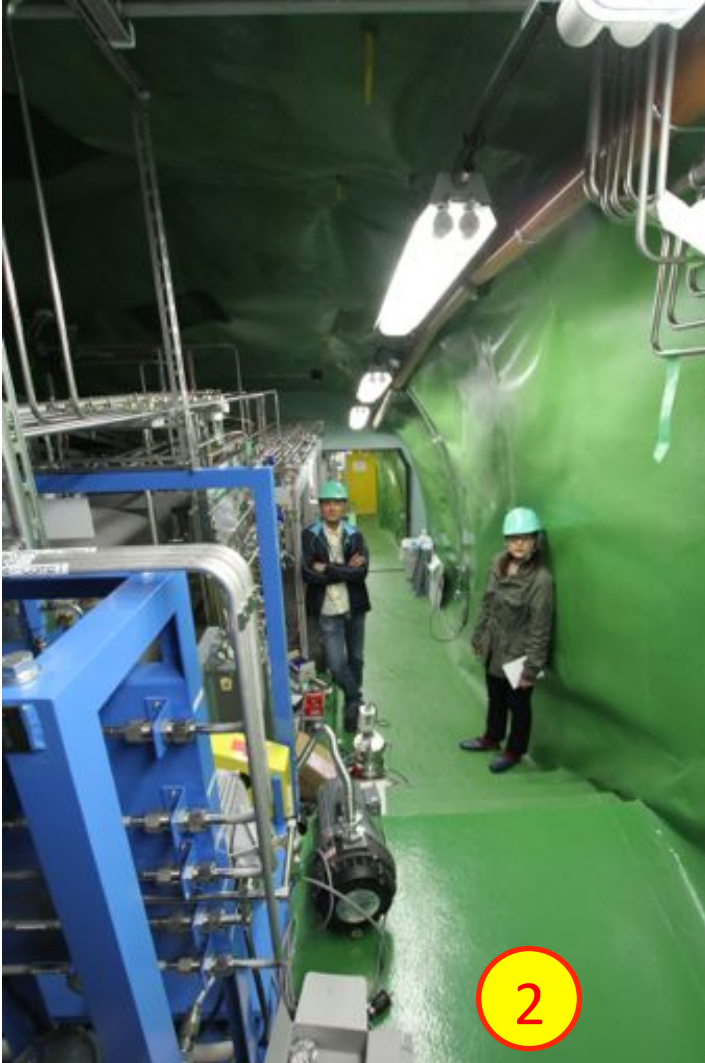
1.5 year of data (20 keVts)



Alternative Deployment Locations



CeLAND phase 1 : Xenon Room

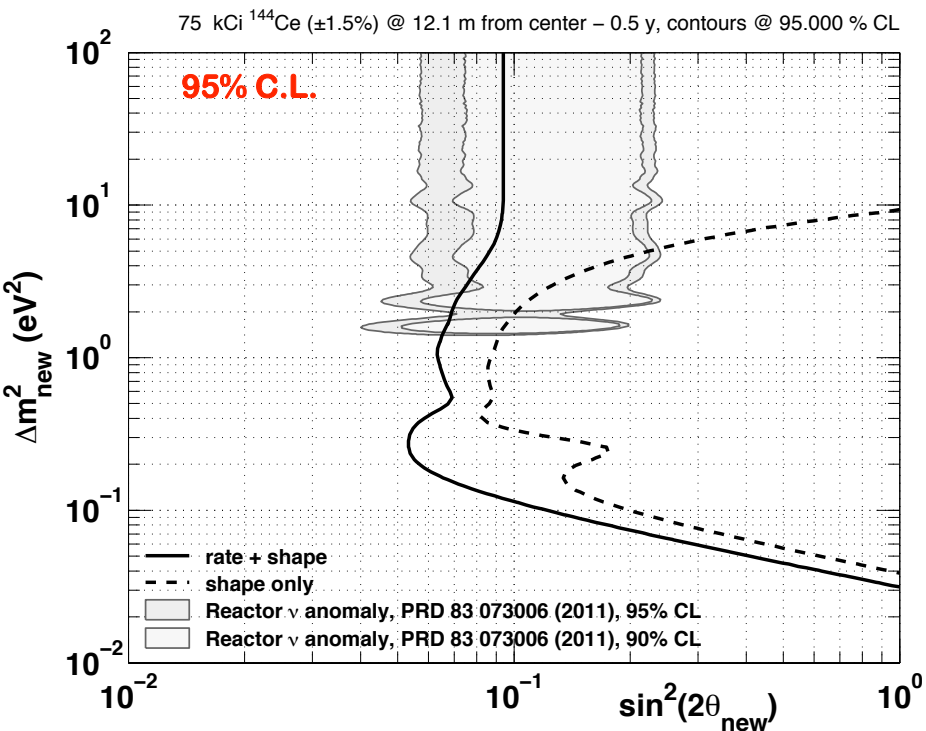


- **Configuration:**
 - 75 kCi ^{144}Ce - ^{144}Pr
 - 12 m from detector center
 - 14 kevt/1.5 y
- **Pros:**
 - Space available
 - Easier integration
 - Simultaneous calorimetry
- **Cons:**
 - Farther from LS
 - Worst shape-only sensitivity

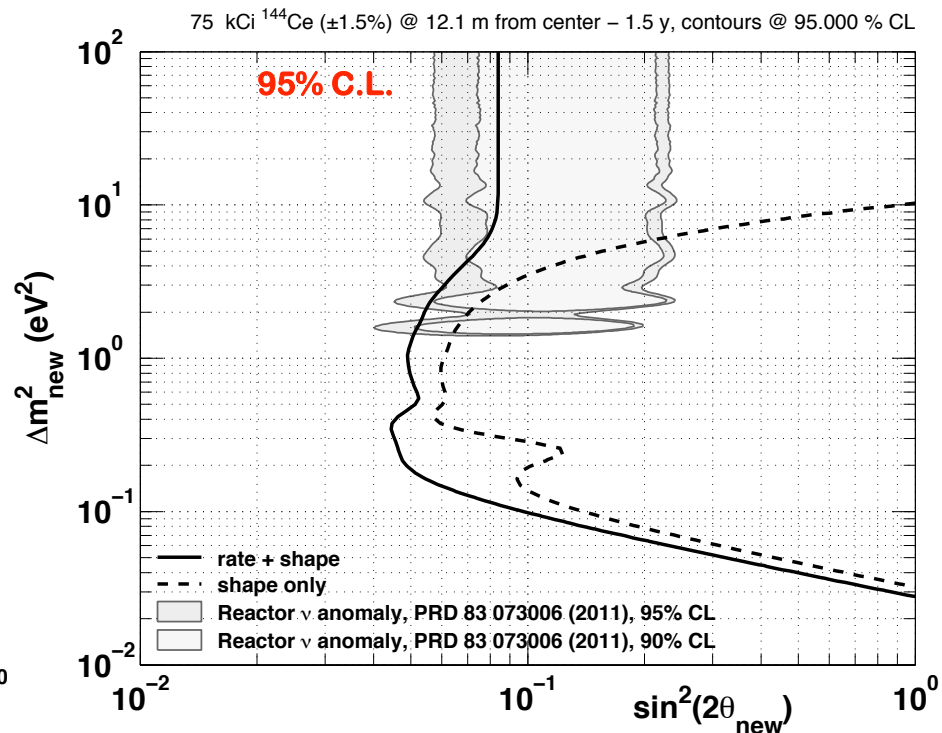
^{144}Ce - ^{144}Pr in Xenon room

75 kCi ^{144}Ce - ^{144}Pr – 12.1 m from detector center
Fiducial volume considered from 0 to 6.5 m

0.5 year of data (7 kevtS)



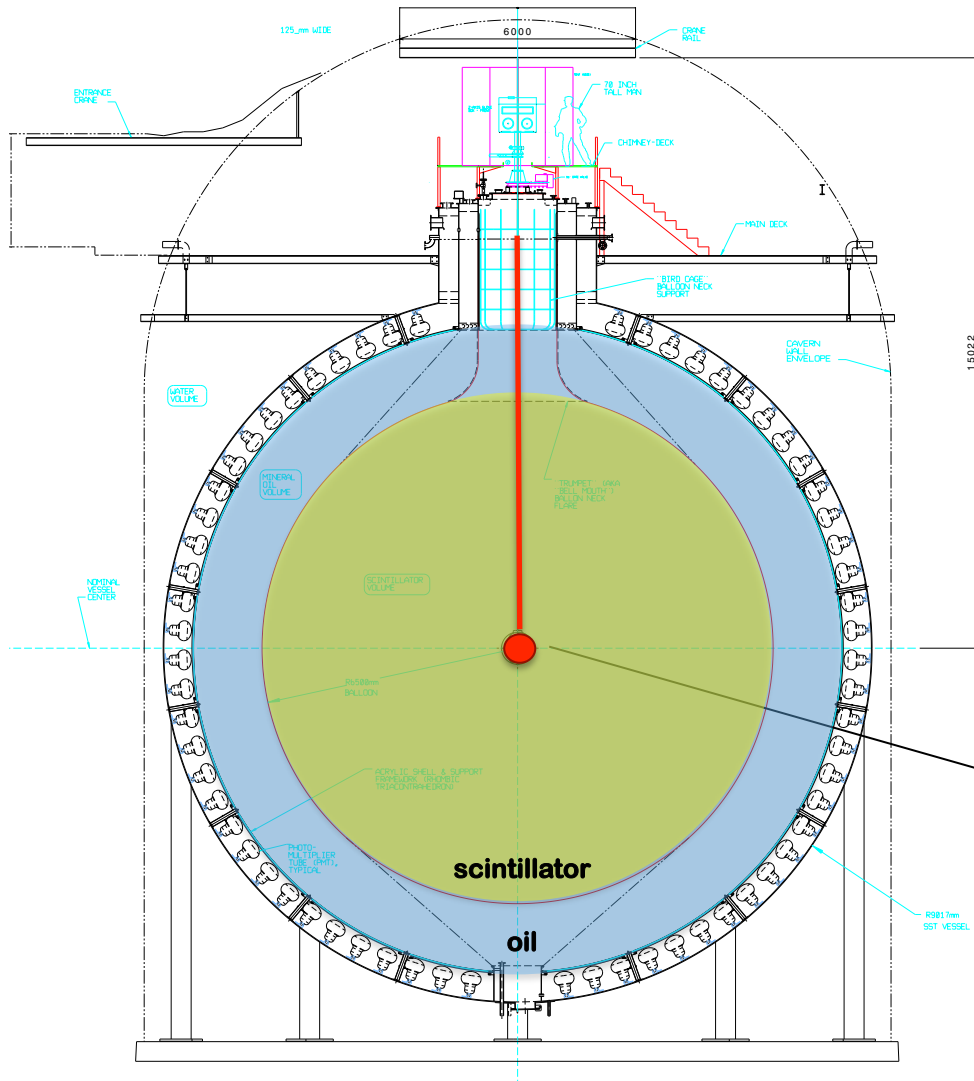
1.5 year of data (14.3 kevtS)



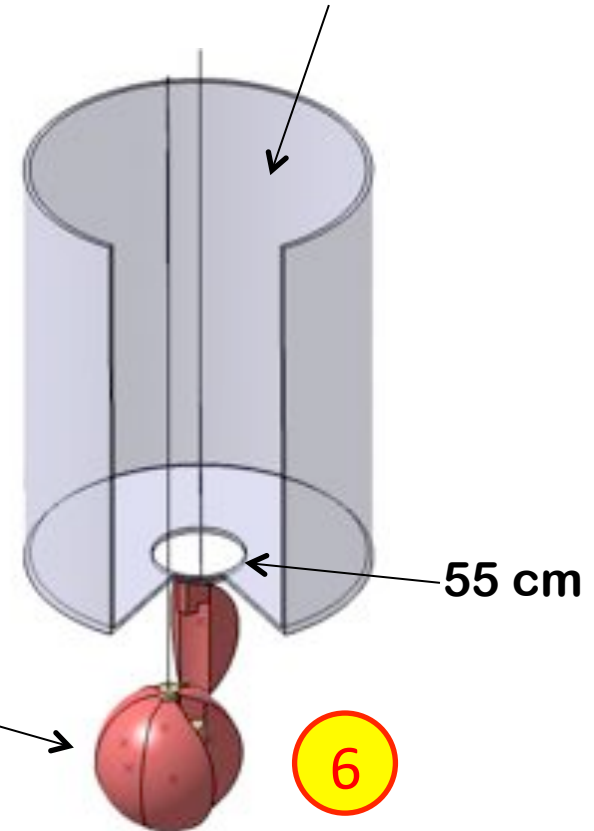
CELAND Phase 2: 2016

erc

Relocate the ^{144}Ce source: **75 kCi leads to 50 kCi after 6 months**



KamLAND neck

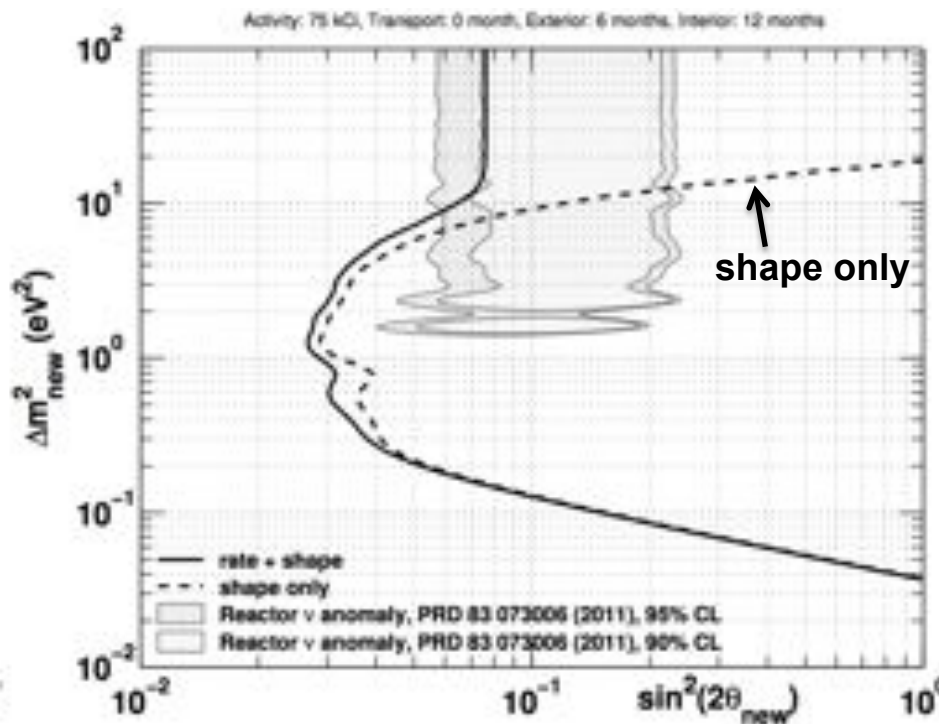


40 cm W-alloy, $d=18.5 \text{ g/cm}^3$
 γ -attenuation (2 MeV) : 10^{-13}

CELAND Phase 2: Sensitivity

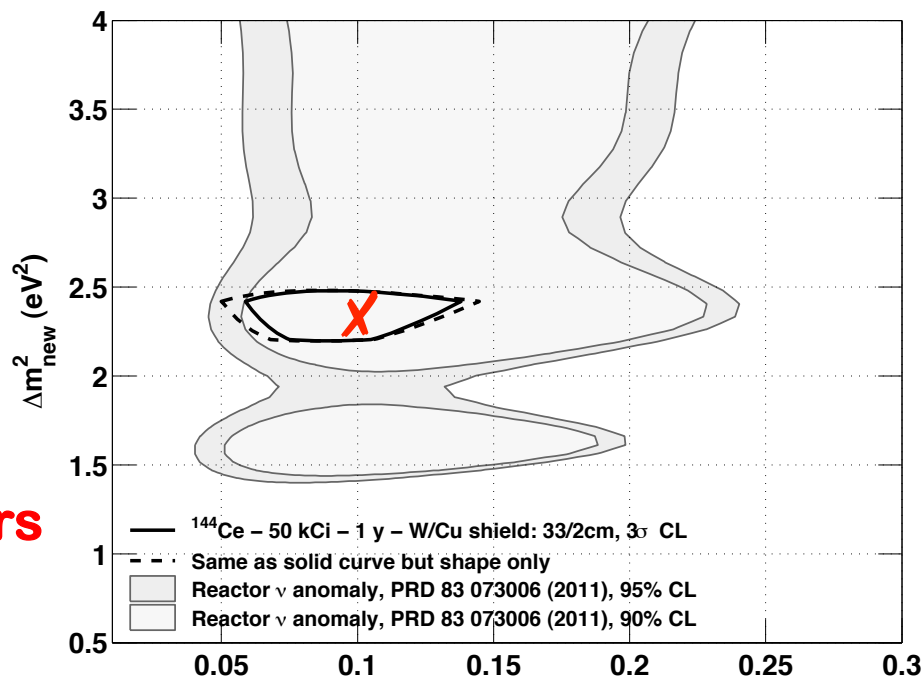
75 kCi 6 months + 50 kCi 1 year (10+50 kevt)

erc



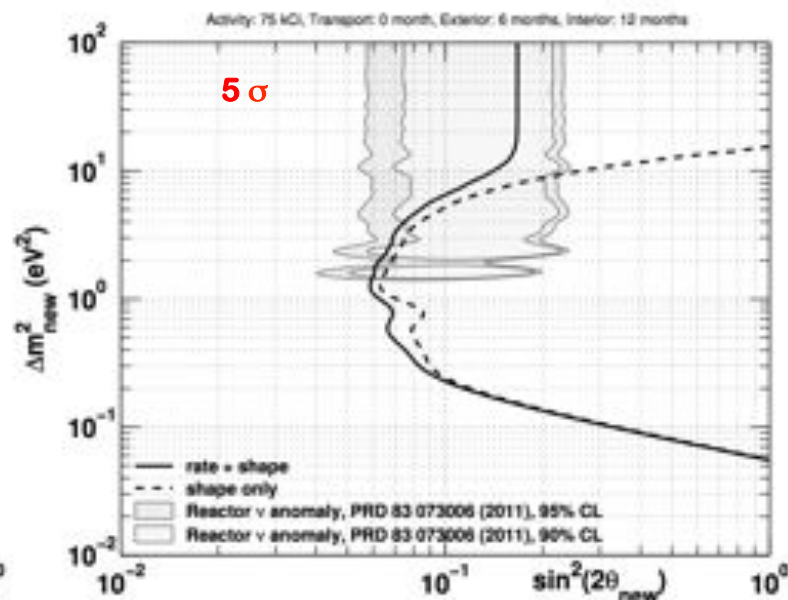
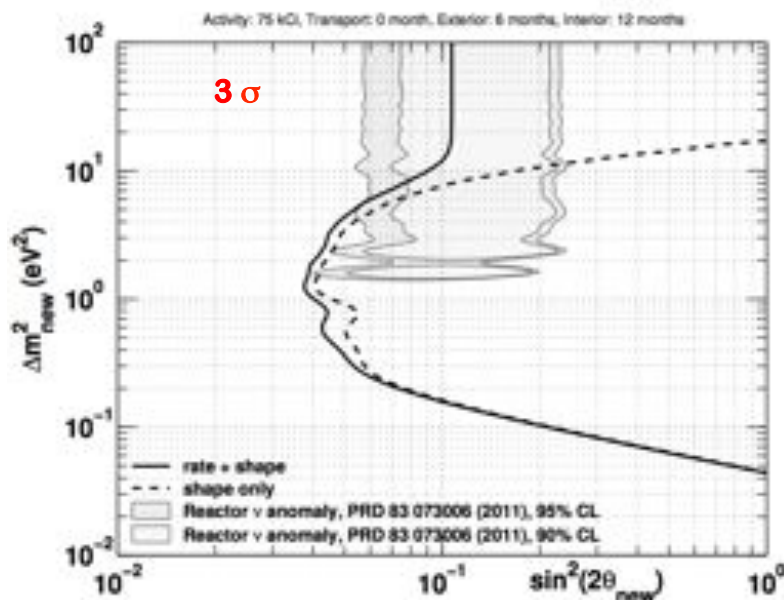
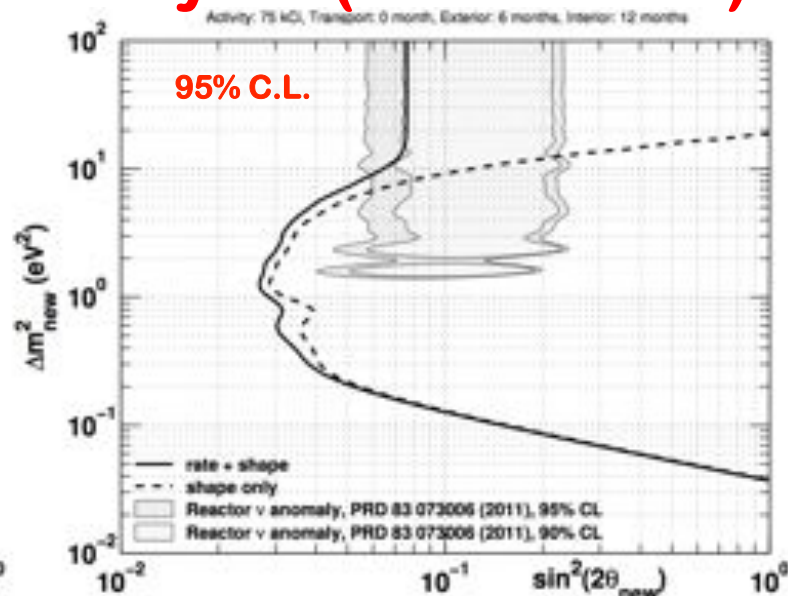
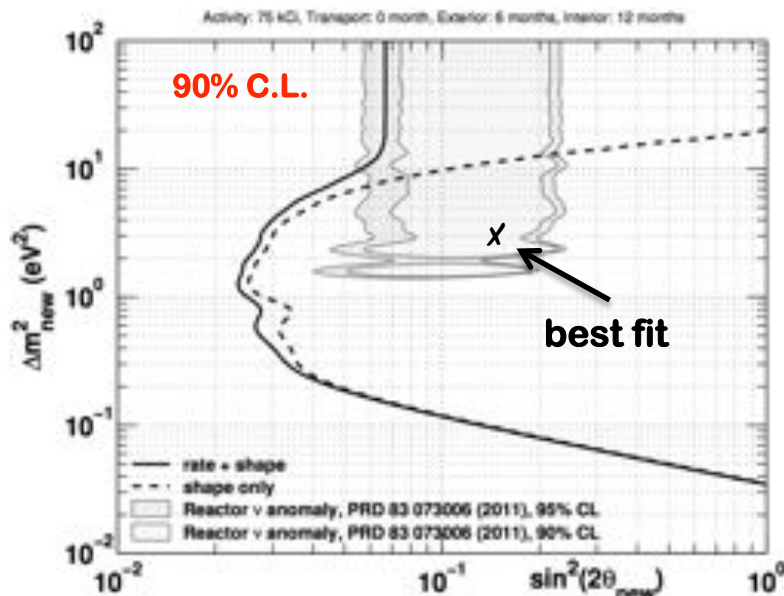
95% C.L. exclusion

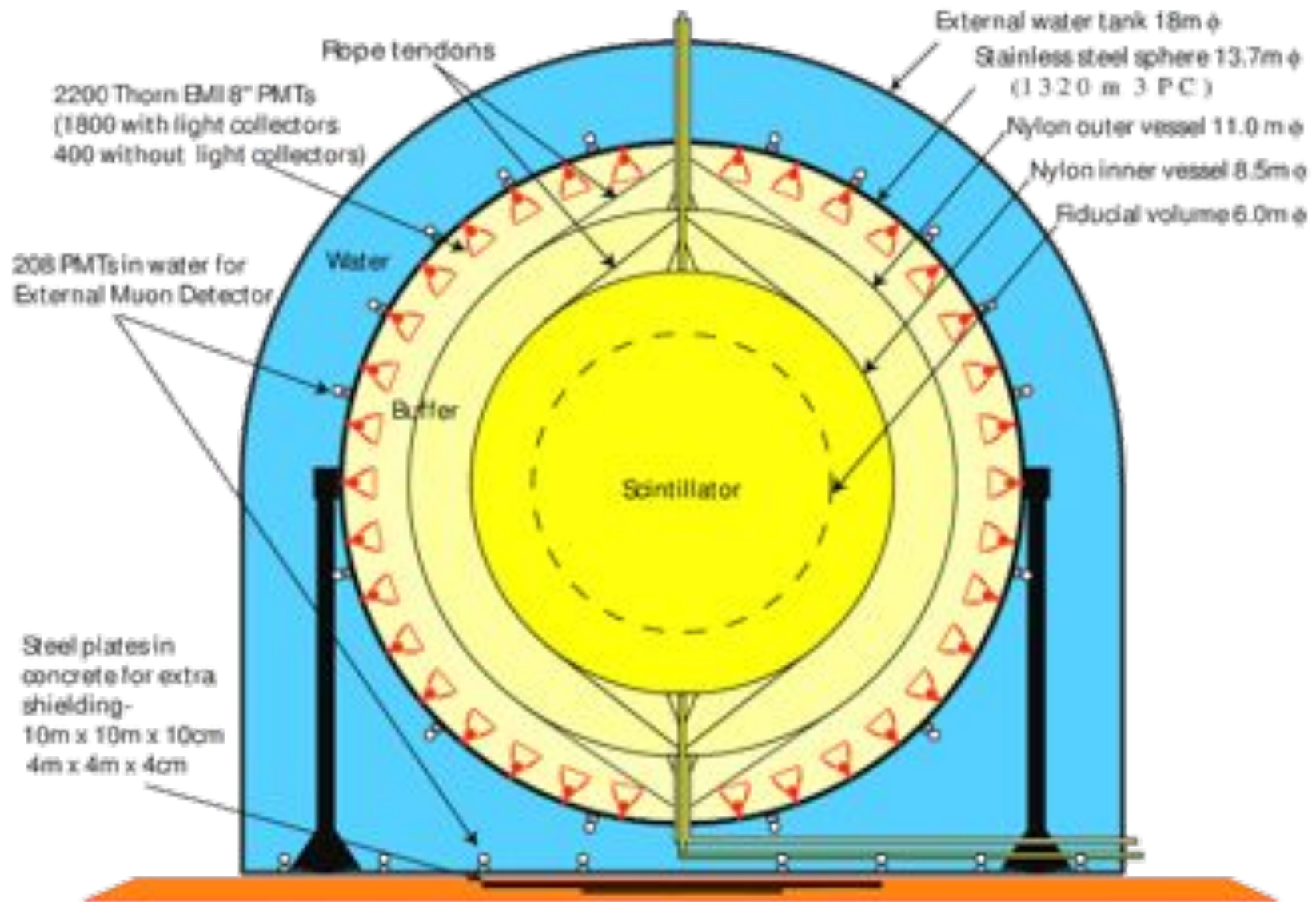
3 σ measurement contours



CeLAND Phase 1 + Phase 2

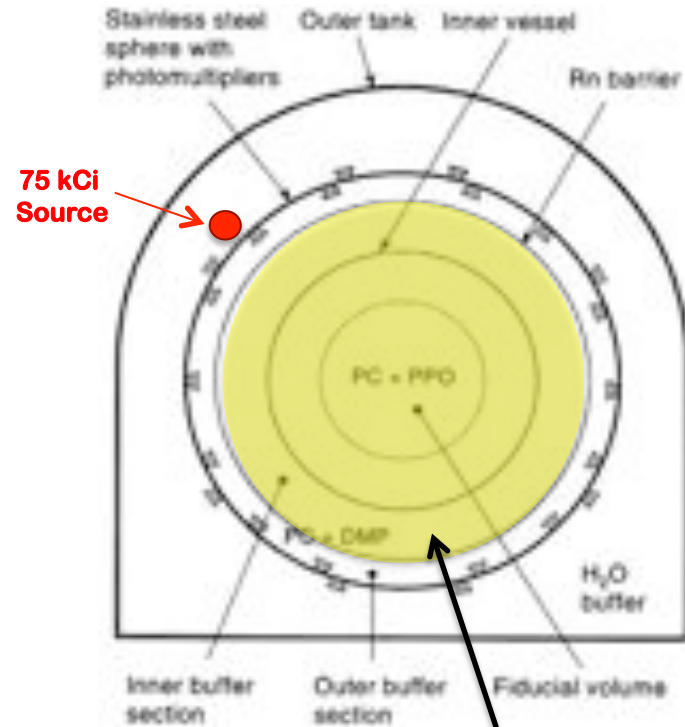
75 kCi 6 months + 50 kCi 1 year (10+50 kevt)





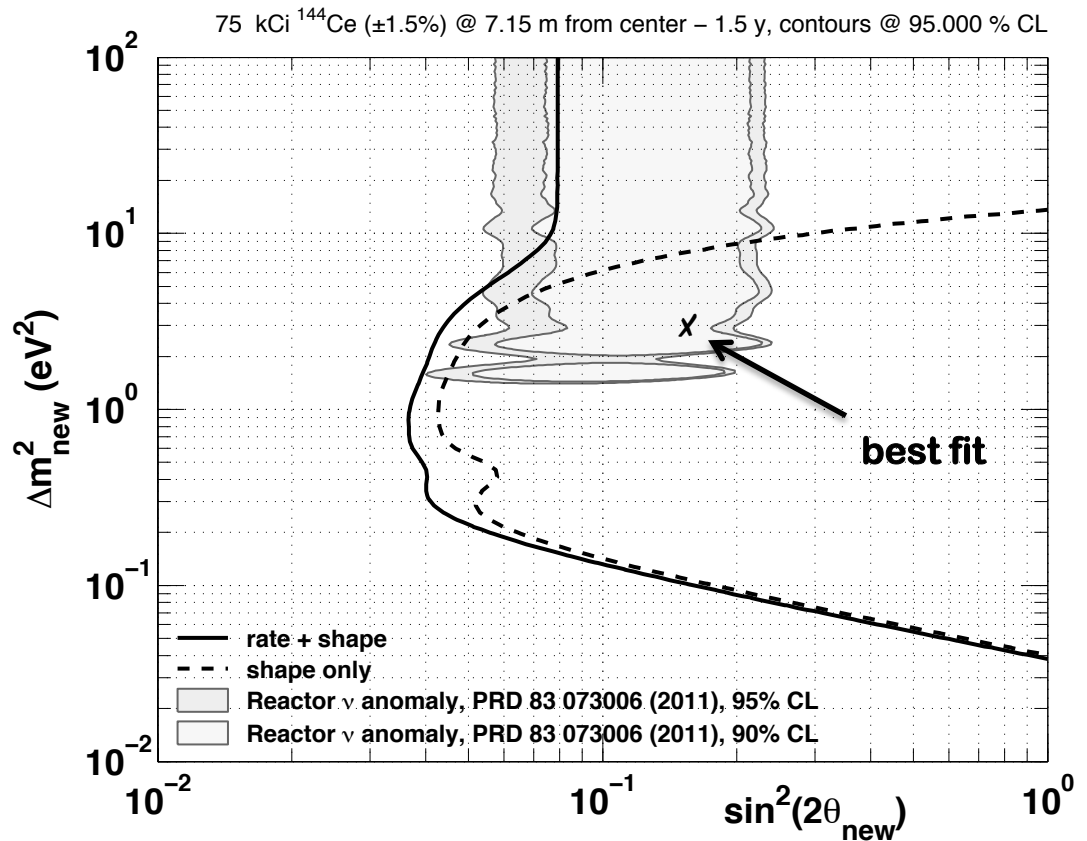
CeSOX 613 tons - 75 kCi - 1.5y

Distance from center: 7.15 m



$R_{\max} = 5.5 \text{ m}$
 $N_{\text{evt}} = 21800$

Need liquid scintillator modification
as well as detector draining

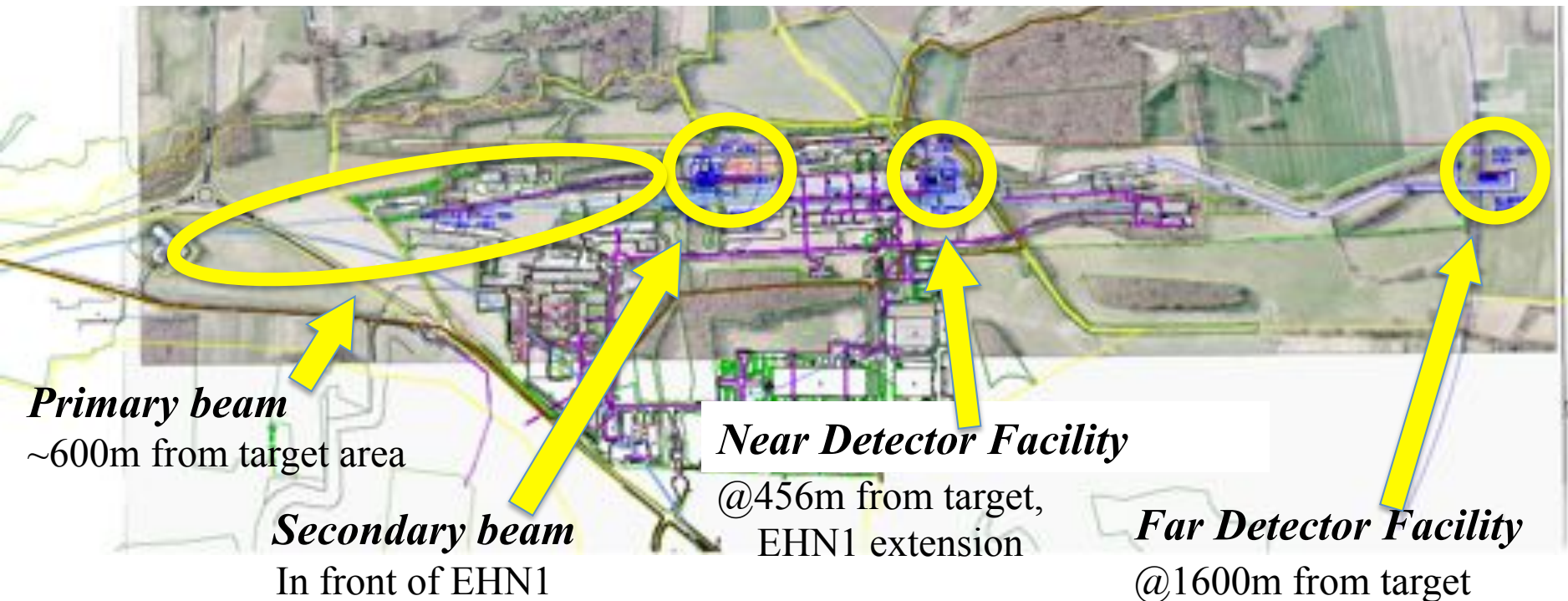


Many Projects: Overview

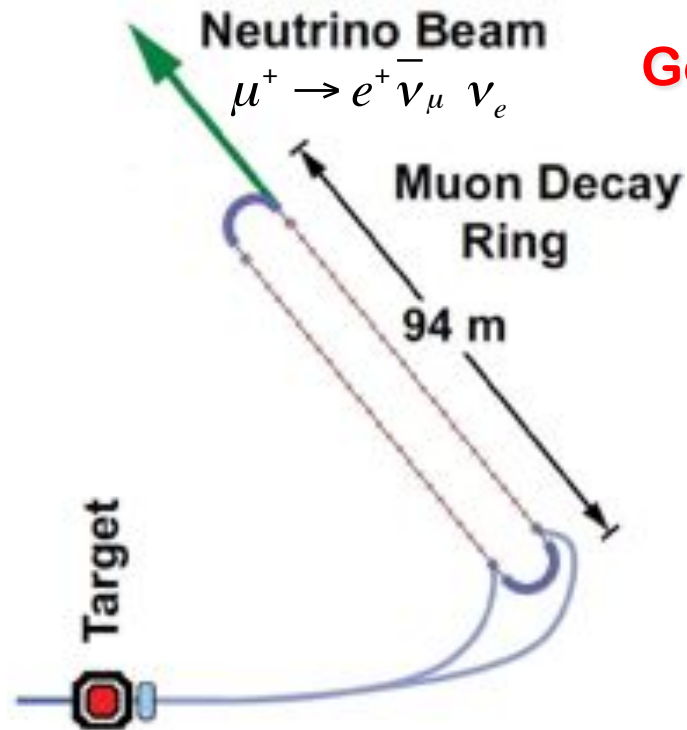
Experiment Type	Appearance / Disappearance	Oscillation Channel	Projects
Reactor	Disappearance	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	Nucifer, Stéréo, Scraam, Neutrino-4, DANSS, Poséidon, MARS, ...
Radioactive Source	Disappearance	$\bar{\nu}_e \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	CeLAND, SOX (Cr & Ce), Sage2, SNO+, LENS-s
Cyclotron	Disappearance	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	IsoDAR
Pion / Kaon Decay-at-Rest	Apparition & Disappearance	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	OscSNS, CLEAR, DAEδALUS, KDAR
Pion Decay-in-Flight (Beam)	Appearance & Disappearance	$\nu_\mu \rightarrow \nu_e$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_\mu \rightarrow \nu_\mu$ $\nu_e \rightarrow \nu_e$	MINOS+, MicroBooNE, LAr1kton+MicroBooNE, Icarus/Nessie@CERN
Low-E Neutrino Factory	Appearance & Disappearance	$\nu_e \rightarrow \nu_\mu$ $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$ $\nu_\mu \rightarrow \nu_\mu$ $\bar{\nu}_e \rightarrow \bar{\nu}_e$	vSTORM@Fermilab

Icarus-Nessie at CERN

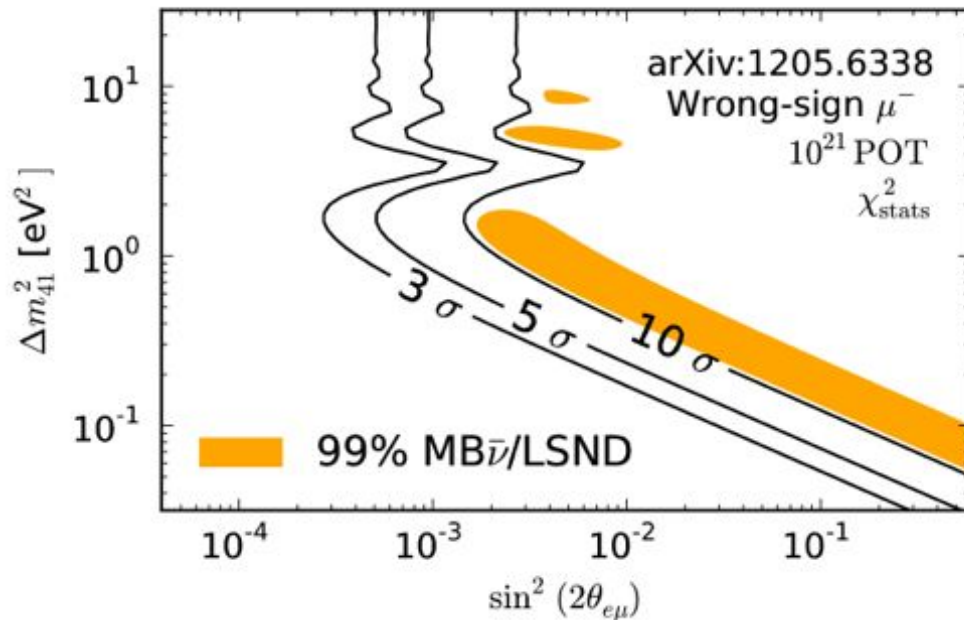
- **Concept** ($L/E \approx \text{km/GeV} \rightarrow \text{eV-scale } 4^{\text{th}} \nu$)
 - Move T600 (2 x 300 tons) from LNGS to CERN, 2013/14
 - New T150 (150 tons)
 - New magnetized muon spectrometers (Opera's techno)
 - New neutrino beam on the SPS (on-axis, few GeV, fast extraction, $4.5 \cdot 10^{19}$ pot/y, neutrinos & antineutrino runs, start 2016/7?, under review)



Decay in Flight Beams: ν -STORM



Golden Mode: $\bar{\nu}_\mu^{(-)}$ appearance in a $\bar{\nu}_e^{(-)}$ beam



- Beam+Detector: manageable extrapolation from known technology
- Capable of definitive sterile ν search in appearance and disappearance channels with : ν_μ , ν_e , $\bar{\nu}_\mu$ and $\bar{\nu}_e$.
- Source of GeV $\bar{\nu}_e^{(-)}$ for studies of cross sections (HK, LBNO/E).
- First step to Neutrino Factory / Muon Collider (LOI @FNAL/CERN)

Conclusions

- **A bunch of anomalies calling for clarification:**

- LSND (ν_s , $\Delta m^2 \approx \text{eV}^2$) & Miniboone ?
- Gallium Anomaly (ν_s , $\Delta m^2 \approx \text{eV}^2$)
- Reactor Anomaly (ν_s , $\Delta m^2 \approx \text{eV}^2$)
- Cosmological bounds do not exclude a 4th ν
- But:
 - No deficit in muon channel
 - Significant tensions in global fits

→ Bunch of 2 to 4 σ effects → **NEED** for new **CONCLUSIVE** short baseline experiments, >20 projects under consideration

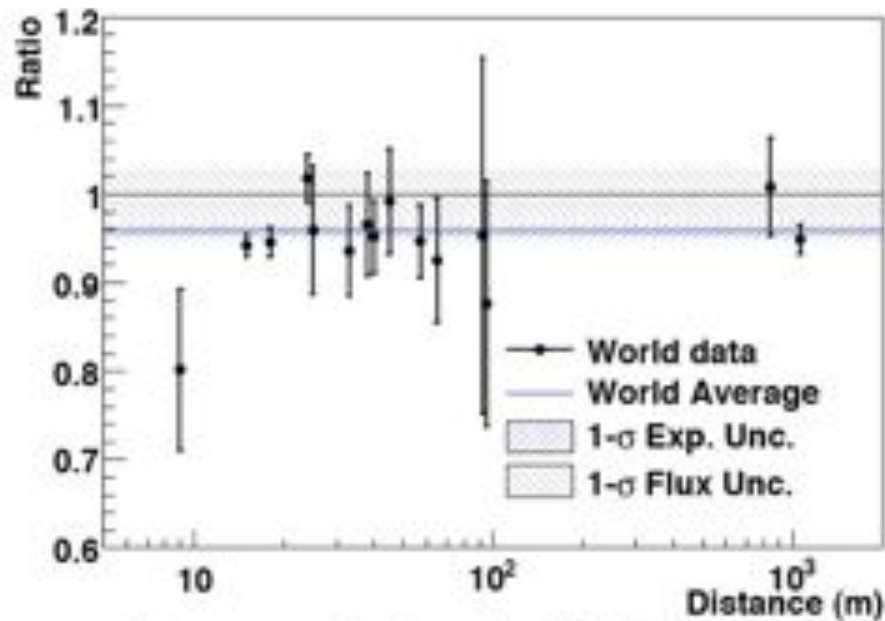
- Need to Test the Beam related anomalies (LSND, MiniBoone)
- Need to Test the Gallium and Reactor Anomalies

Thanks for your attention !



C. Zhang, X. Qian & P. Vogel

RAA Reanalysis (arXiv:1303.0900v1)



❖ New reanalysis of old 19 reactors antineutrinos experiments adding :

- $\sin^2(2\theta_{13}) = 0.089 \pm 0.011$
- allow to use Chooz, Palo Verde, Double Chooz (Gd & Hyd. datas)
- $\Rightarrow R = 0.959 \pm 0.009(\text{stat.}) \pm 0.027(\text{react. flux}) : 1.4 \sigma \text{ effect}$

❖ First quick comments from Saclay's group

- which IBD cross-section used ?
- wrong error used by authors in RAA
 - 0.027 instead of 0.023
 - taken wrongly as fully correlated
 - only ILL data is fully correlated
- Saclay preliminary similar reanalysis :
 $\approx 0.94 \pm 0.023$ (confirmed by Arxiv:1303.3011)
- short written answer in preparation

