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Data quality monitoring in the presence of aerosols and other adverse atmospheric conditions with H.E.S.S.

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Outline

- Introduction
- Hardware DQ selection
- Weather DQ selection
 - Clouds
 - Aerosols
- Cherenkov Transparency Coefficient
 - Correlation with satellite measurements
 - Systematic effects on source spectra
- Summary



- H.E.S.S. (Heidelberg analysis) uses a uniform data quality selection mechanism
 → Especially important for surveys etc.!
- Analysers use two standardised sets of cuts:
 - source detection:
 - Hardware OK, no checks on weather
 - spectral analysis:
 - Hardware + weather OK

hardware cuts

Hardware Cuts (applied for detection and spectral quality)

	Criterion	cut range (unit)
System {	run duration	600:7200 (sec)
	$\mathrm{DST}\ \mathrm{check}^b$	DST has to exist
Camera {	participation fraction	0:0.4 (rel. fraction)
	pixels with 'hardware' flag	0:120 (number)
	pixels with 'HV-off' flag	0:50 (number)
Tracking {	RMS of Az dev. distr.	0:10 (arcsec)
	RMS of Alt dev. distr.	0:10 (arcsec)
	mean deviation in RA	0:1 (arcmin)
	mean deviation in Dec	0:1 (arcmin)

 These cuts check quantities that are necessary for Hillas reconstruction
 → They are good enough for maps and detections

atmosphere cuts



Clouds

Clouds seen with H.E.S.S.: Small clouds



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Clouds

Clouds seen with H.E.S.S.: Large clouds / Cloud layers



Clouds: Cuts



Large scale aerosol aborber structure





Overall drop in system rate Not detectable within one run

Aerosols: old cut

Old method:Get characteristic rate for any time period



Cut on this rate

Aerosols: old cut

Old method:Get characteristic rate for any time period



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Aerosols: old cut

Old method:Get characteristic rate for any time period



Is there a more hardware independent quantity isolating the effect of an opaque atmosphere on the trigger rate?

• Assumption 1:

Trigger rate is integrated (local proton spectrum*Eff.area)

$$R = \int_{E_0}^{\infty} \mathrm{d}E A_{\mathrm{eff}}(E) f(E)$$
$$\simeq k \cdot E_0^{-1.7+\Delta}$$

Aerosols: new cut

 $E_0 \propto (\eta \cdot \mu \cdot g)$

air shower

- Effect of atmosphere opacity symbolised by η
- Most important hardware effects on E_O:
 - pixel gain (g) adjustments
 - Winston cone cleaning, mirror refurbishments
 → traced by muon-efficiency (µ)
- Assumption 2:

• Then for every telescope i you can calculate

$$\eta \propto \frac{R_i^{\frac{1}{1.7-\Delta}}}{\mu_i \cdot g_i} \equiv t_i$$

Average over N telescopes for a telescope-wide quantity





- Remarkably stable over 8 years
 - → largely hardware-independent
 - → isolates atmosphere conditions ↔ dips around Septembers



May-August



September-December



DQ: Impact of cuts

2



DQ: Impact of cuts

2



δ2: ~6%

DQ: Impact of cuts



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- Transparency coefficient traces aersols ?
 ↔ Proof: should correlate with independent aerosol measurements from satellites!
- MISR
 - On board TERRA
 - 275m-1.1km resolution
 - Different observation angles → disentangle clouds from aerosols etc.
 - Level 3 data product



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Select data taken within 24h of H.E.S.S. Measurements
 → overlap ~2% of H.E.S.S. Data



Select data taken within 24h of H.E.S.S. Measurements
 → overlap ~2% of H.E.S.S. Data



- Strong and positive correlation with independent satellite measurements of aerosols
- Positive and strong correlations also with measurements from LIDAR and Radiometer (see other talks)

• Effect on reconstructed spectra?



log (Energy)

• Under assumption $E_{reco} \propto T \times E_{true}$ it follows for the rec. spectrum $\frac{\mathrm{d}F}{\mathrm{d}E_{true}} \propto E_{true}^{-\Gamma} \iff \frac{\mathrm{d}F}{\mathrm{d}E_{reco}} \propto E_{reco}^{-\Gamma} \cdot T^{\Gamma-1}$





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Fit: 1.69 ± 0.13stat ↔ Lit.: 1.63 ± 0.01stat ± 0.10sys
 → perfect match!

Use in Flux correction under study → Stay tuned!

Impact of aerosols and adverse atmospheric conditions on the data quality for spectral analysis of Cherenkov telescopes

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Abstract

The Earth's atmosphere is an integral part of the detector in ground-based imaging atmospheric Cherenkov telescope (IACT) experiments and has to be taken into account in the calibration. Atmospheric and hardware-related deviations from simulated conditions can result in the mis-reconstruction of primary particle energies and therefore of source spectra. During the eight years of observations with the High Energy Stereoscopic System (H.E.S.S.) in Namibia, the overall yield in Cherenkov photons has varied strongly with

Paper submitted

For more infos:

→ astro-ph version of the paper coming soon

- H.E.S.S. uses a uniform DQ selection
- Hardware and Weather checks
- Clouds are relatively easy to detect within one run through trigger rate
- Aerosols and cloud layers more challenging
 → previous cut failed!
- New, hardware-independent quantity: Cherenkov Transparency Coefficient T

- Cherenkov Transparency Coefficient T
 - Correlates strongly with satellite data (also with LIDAR and Radiometer data)
 - Describes expected Flux reduction due to atmospheric extinction quantitatively
 → use for flux correction under study
 - Is in principle generic and 'free' to all IACT experiments
 → needs standard calibration quantities+trigger rates
 → corresponding implementation in CTA under study