EW Moriond 2013 Theory Summary HEP-EX is on the move The triumph of the SM Naturalness challenged Where is BSM physics?

Fabio Zwirner
University and INFN, Padova
(ITN-UNILHC + ERC-DaMeSyFla)

Apologies

For theory, I counted 26 "regular" talks and 10 "short" talks by young scientists: it would be impossible (and boring) to give a fair account of all of them

I will use the unwritten privileges of the summary speaker to transmit you my (of course debatable) theorist's viewpoint on the status of the field and its perspectives The Standard Model of strong and electroweak interactions (effectively coupled to gravity) quantitatively describes most observations

v oscillations call (so far) for minor modifications Stronger exceptions with gravity/astro/cosmo: dark matter, dark energy, inflation, baryogenesis

True last year, still true today: what changed?

Let us start from the big question of particle physics to which we are finding answers now: symmetry breaking in the SM

Symmetry breaking in the SM

$$\mathcal{L}_{gf} = -\frac{1}{4} F^a_{\mu\nu} F^{a\,\mu\nu} + i \overline{\Psi} \not\!\!\!\!D \Psi$$

$$G_{loc} = SU(3) \times SU(2) \times U(1)$$
 $G_{gl} = SU(3)^5 \times U(1)^4$

Spontaneous breaking of the gauge symmetry:

$$\mathcal{L}_S = (D_\mu \phi)^\dagger (D^\mu \phi) - \mu^2 \phi^\dagger \phi - \lambda (\phi^\dagger \phi)^2$$

Explicit breaking of the flavour symmetry:

$$\mathcal{L}_{Yuk} = \overline{\Psi}_i Y_{ij} \Psi_j \phi + h.c.$$

$$H_{loc} = SU(3)_C \times U(1)_{em}$$
 $H_{gl} = U(1)^4$ $[B, L_e, L_{\mu}, L_{\tau}]$

The message from EXP to TH

The minimal, weakly-coupled
Standard Model implementation
of the Brout-Englert-Higgs mechanism
with
a single "elementary" scalar doublet
CKM description of flavour change and CPV
a generalised GIM mechanism at work

works far beyond most expectations

Recent experimental milestones (just summarised by Paris Sphicas):

- A new particle compatible with the SM scalar We have now 5 fundamental forces in Nature, mediated by spin-0, spin-1 and spin-2 bosons!
- New precise flavour and CPV tests
 passed with flying colours!
- More stringent bounds on new particles
 LHC bounds well above 1 TeV for
 sizeable couplings to quarks and gluons
 viable signatures in the LHC environment

SM Flavour

COMPARISON OF ST WITH DATA

	$ \epsilon_K $	Δm_K	$ \Delta M(B_d^0) $	$ \Delta M(B_s^0) $	$ \Delta M(D^0) $	$Br(B_s \to \mu^+\mu^-) $
EW diagr.	$6.34\ 10^{-3}$	$3.12 \ 10^{-12}$	$7.51 \ 10^{-10}$	$294 \ 10^{-10}$	$2.0\ 10^{-13} \cdot (\frac{m_s}{0.15 GeV})^2$	4.0 10 ⁻⁹
QCD correts	$2.65 \ 10^{-3}$	3.85 10 ⁻¹	4.13 10 ⁻¹⁰	119 10 ⁻¹⁰	7?	$(3.53 \pm 0.38) \ 10^{-9}$
expt	$2.228 \ 10^{-3}$	$3.483 \ 10^{-12}$	$3.34\ 10^{-10}$	$117.0\ 10^{-10}$	$(1.57 \pm 0.39) \; 10^{-11}$	$(3.2 \pm 1.4) \ 10^{-9}$

Table 1: Masses in MeV

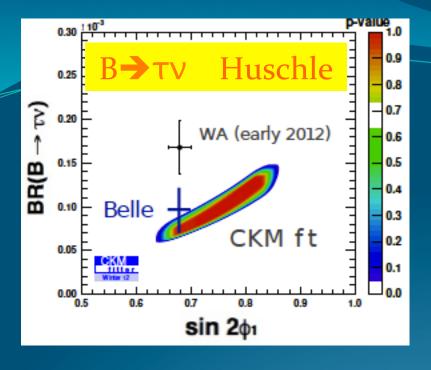
dominated by the t-quark

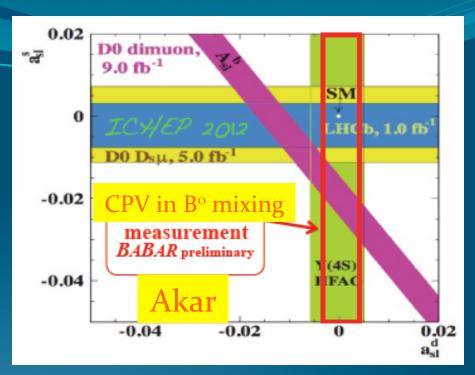
dominated by the c-quark

Input data:

CKM coefficients (weak decays of s, c and b) m_c=1.5, m_t=173, m_s=0.150, m_b=5.0







Many small tensions going...



Some hopes that have not materialised (so far)

- $\mu \rightarrow e\gamma$: new MEG bound BR < 5.7 x 10⁻¹³ (90%cl)
- LHCb 3.5 σ evidence for $B_s \rightarrow \mu\mu$, agrees with SM Sarti

Other small tensions fading away?

- CDF top FB distribution, 2.1 σ from NLO SM starts being tested at the LHC Wilson; Battilana
- ΔA_{CP} in D decays (3.5 σ in LHCb), followed by theory reassessment, exp update is imminent

Hampson

Precision flavour physics calls for state-of-the-art phenomenology

A hot example discussed at this meeting:

b \rightarrow s l⁺ l⁻ transitions, e.g. B \rightarrow K^{*} μ ⁺ μ ⁻

Tevatron and LHCb have entered the precision era and started measuring angular distributions

Predictions for ang distr at low hadronic recoil
 Extraction of hadronic form factors from data

Hiller

...as well as for hard SM theory (e.g. lattice)

 $\Delta I=1/2$ rule from the lattice $\Gamma(K_S \rightarrow \pi^+\pi^-)/\Gamma(K^+ \rightarrow \pi^+\pi^0)=670$ [ReA_o/ReA₂ \approx 22.5] with implications for ϵ'/ϵ

Domain wall quarks full QCD (no xPT) physical pion mass

Soni

New strong cancellation found (factor 3-4) in ReA₂. A_o needs more work but it is under way

Towards solving a long-standing puzzle?

The boson

Impressive progress in the study of its properties as just described in the experimental summary And there is more to come from Moriond QCD!

Three comments:

Once more, admiration for ATLAS/CMS colleagues

I was impressed by new direct indications for SM-like couplings to τ leptons and b quarks

To take deviations from the SM seriously we should apply the same stringent standards as for discovery!

Difficult to imagine a SM scalar crisis in < 3 years

Is it a spin-0 CP-even particle?

ATLAS/CMS are now testing J^{CP} of the new particle Important as a consistency check and must be done

However, we should keep in mind what the σ 's mean

With M_H known, no free parameter left in the SM to describe all production mechanisms and decay modes: renormalizable theory, passes all precision tests, can be safely extrapolated to (much) higher energies

More complicated to do the same with $J^{CP} \neq O^+$! Technically possible to write an effective Lagrangian, but this adds many parameters and "theory sigmas"

Is it the SM scalar?

Several ways to go non-standard:

Rzehak, Carena, Ellwanger, Yamawaki,...

- H mixes with other spin-o states
 e.g. additional doublets and/or singlets
- H meson of a new strong force, kept light by its pseudo-Goldstone boson nature
- H decays into invisible particles
- Loops for H production (ggH) and decay (Hgg,Hγγ,HZγ) modified by new particles

How to parametrise a non-SM scalar?

Eboli, Jenkins, Azatov, Gavela Merlo, Yamawaki

Compromise between simplicity and completeness, depending on kind of experimental data and purpose

For present book-keeping, under reasonable assumptions (spin-o, CP even, custodial, no FCNC)

$$\mathcal{L}_{\leq m_{h}}^{eff} \approx \frac{\mathbf{c}_{\mathbf{V}}(\frac{2m_{W}^{2}}{v}W_{\mu}^{+}W_{\mu}^{-} + \frac{m_{Z}^{2}}{v}Z_{\mu}^{2})h + \mathbf{c}_{b}\frac{m_{b}}{v}\bar{b}bh + \mathbf{c}_{\tau}\frac{m_{\tau}}{v}\bar{\tau}\tau h$$

$$+ \frac{\mathbf{c}^{\gamma}}{9\pi v}F_{\mu\nu}^{2}h + \mathbf{c}^{g}\frac{\alpha_{S}}{12\pi v}G_{\mu\nu}^{2}h$$

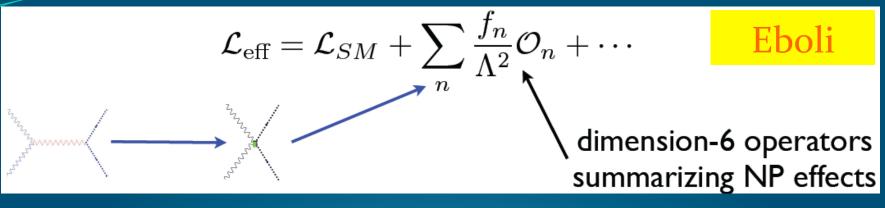
$$+ \mathcal{L}(h \to inv)$$

$$c^{\gamma} = \mathbf{c}_{t} + \frac{9}{2}\delta c^{\gamma}$$

$$c^{g} = c_{t} + \delta c^{g}$$

SM recovered for all five c=1 and $L(h \rightarrow inv)=0$

For testing models, use effective Lagrangians



Linear realization for models with elementary H

Non-linear realization for composite H (WTC, PC)

$$\mathbf{U}(x) \equiv \mathbf{M}(x)/v = e^{i\sigma_a \pi^a(x)/v} + F_{i}(H)$$

In both cases, important to identify a suitable basis of independent operators

Theoretically-motivated fitting strategies being suggested by theorists to experimentalists

Azatov

Correlations between anomalous scalar couplings, EWPT and anomalous vector boson couplings

Jenkins, Eboli

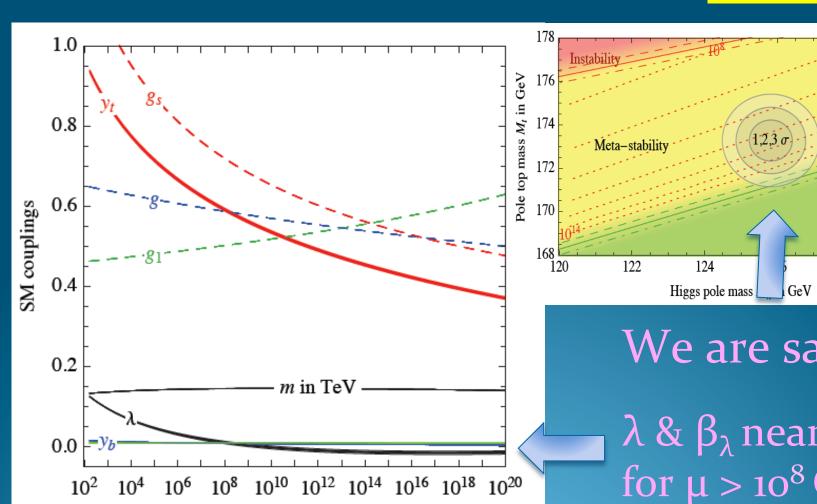
RG evolution from Λ to m_H can give sizable effects

Jenkins

$$\mu_{\gamma\gamma} \simeq 1 - 0.02\,\mathcal{S}\,\lograc{\Lambda}{M_h} + 0.02\left(rac{1\, ext{TeV}}{\Lambda}
ight)^2\left(16\pi^2c_{\gamma\gamma}(\Lambda)
ight)$$

What is m≈125.8 GeV telling us?

Strumia Elias-Miro



RGE scale μ in GeV

We are safe!

 $\lambda \& \beta_{\lambda}$ nearly o for $\mu > 10^8 \,\text{GeV}$

Stability

130

128

- There is nothing forcing us to extend the SM before 10^{10} GeV or so if we ignore naturalness (scales for v_R & invisible axion can be beyond)
- Is there some meaning in the near vanishing of λ , β_{λ} , m_H/Λ at very high cutoff scales for the SM?
- Precise RG calculations and top mass measurements will become important when/if threshold effects at Λ will be calculable
- Potential implications for cosmology & susy
- A scalar singlet is enough to cure the instability if needed for the consistency of model building

A tribute to SM theory

A theoretical construction essentially completed 40 years ago found recently its coronation and stands as solid as a rock admiration for the vision and insight of the founding theorists

However, we should also pay tribute to the long-term efforts to characterize direct and indirect signals of the SM scalar boson and compute the relevant backgrounds

Are we done now?

No, it js just the start of a major programme that may take several decades for completion

The program for the years to come

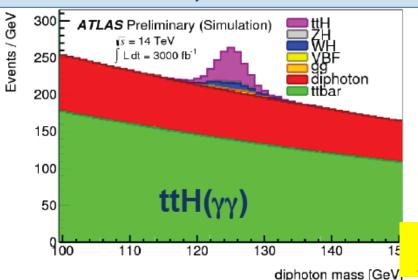
- 1. Study the properties of the new scalar with the highest possible precision, to reveal possible inconsistencies of the SM that would point indirectly to new physics
- 2. Find out whether it is accompanied by other new physics near the TeV scale

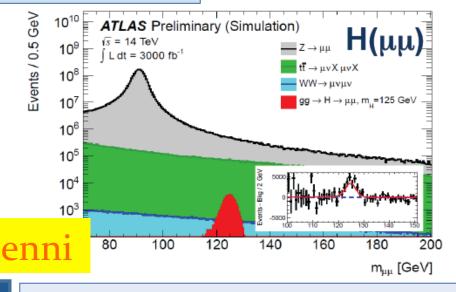
Both missions may require an electronpositron collider to complement the unique information that LHC will collect until ~2030

Physics @ LHC: high luminosity

From 'High Energy Frontier' presented by Marcella Diemoz

HL-LHC: $\sqrt{s} \sim 14$ TeV, L=5x10³⁴ cm⁻²s⁻¹ and 3000 fb⁻¹ ~ 2030





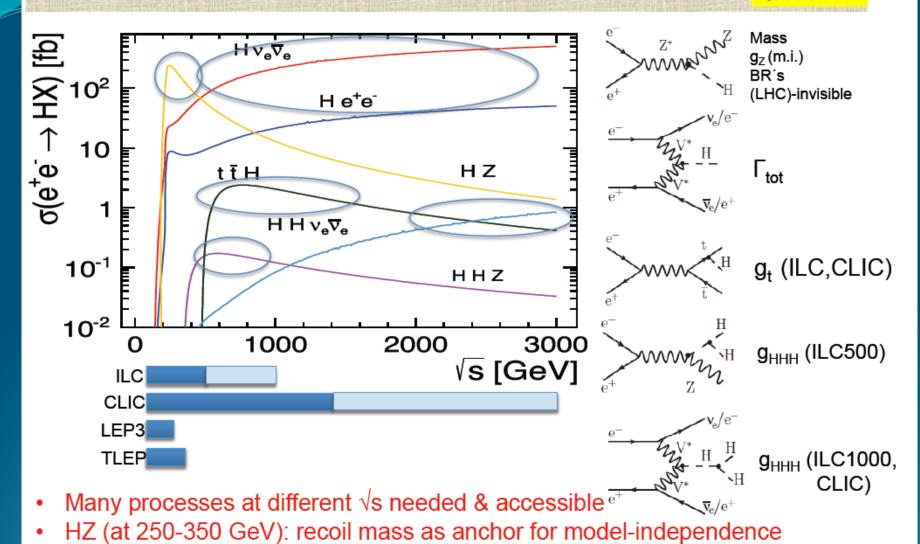
Coupling	300	3000 fb^{-1}		
CMS ESTIM	ATE syst.	(%)	syst. (%)	
CWIS	actual	$_{\rm scaled}$	actual	$_{\rm scaled}$
κ_{γ}	6.5	(5.1)	5.4	(1.5)
κ_V	5.7	2.7	4.5	1.0
κ_g	11	5.7	7.5	2.7
κ_b	15	6.9	11	2.7
κ_t	14	8.7	8.0	3.9
$\kappa_{ au}$	8.5	5.1	5.4	2.0

Profit from statistics:

- rare H production channels
- rare H decay channels
- H couplings
- Higgs self coupling (HH detection)
- VBS: dynamics of EWSB (is it SM?)
- New physics with suppressed couplings

Precision Higgs Studies

Jenni



From 'High Energy Frontier' presented by Marcella Diemoz

What lies Beyond the SM (with a chance of being at reach)?

LHC-8 relied on a powerful no-lose theorem either the SMS, or new physics at the TeV scale We won't be again in such a condition for a long time

Diversify efforts to maximise chances

We are now sailing in uncharted waters must be persistent as Columbus in his trip to Indies...

The SM as an effective theory

= effective UV cutoff (not necessarily universal) = the scale of some (unspecified) new physics

$$L_{eff}^{SM} = \Lambda^4 + \Lambda^2 \Phi^2$$
 $(\Lambda^{n>0} \Rightarrow hierarchy problems!)$

 $+(D\Phi)^2 + \overline{\Psi} \cancel{D}\Psi + F \cdot F + F \cdot \widetilde{F} + \overline{\Psi}\Psi\Phi + \Phi^4$ (controllable log Λ dependence via quantum corrections)

$$+\frac{\overline{\Psi}\Psi\Phi^{2}}{\Lambda} + \frac{\overline{\Psi}\sigma^{\mu\nu}\Psi F_{\mu\nu}}{\Lambda} + \frac{\overline{\Psi}\Psi\overline{\Psi}\Psi}{\Lambda^{2}} + \frac{\Phi^{2}F^{\mu\nu}F_{\mu\nu}}{\Lambda^{2}} + \dots$$

$$(\Lambda^{n<0} \Rightarrow EW \ tests, \ flavour \ tests, \ \not\!\!{E}, \ \not\!\!{L}, \ \dots)$$

Beyond the SM with neutrinos

$$\mathcal{L}_{eff} = \mathcal{L}_{SM} + \delta \mathcal{L}(m_{\nu}) + \dots$$

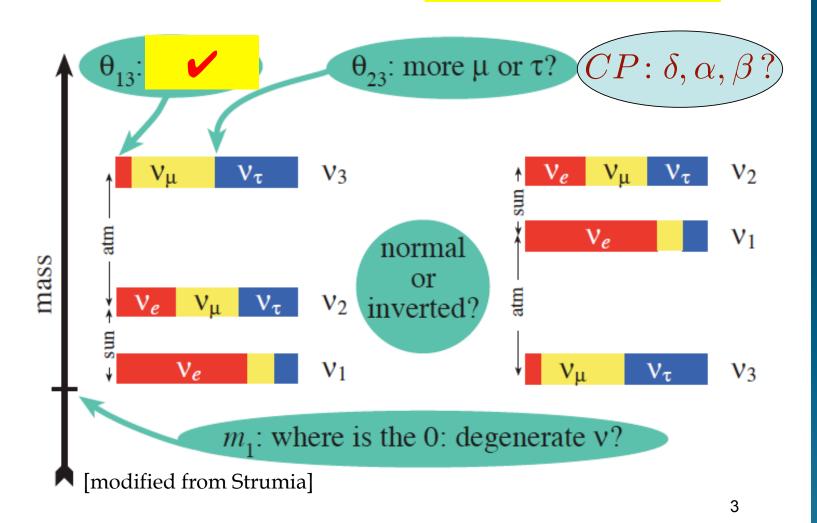
 $\delta \mathcal{L}(m_{
u})$: experimentally needed, still undetermined

Simplest solutions

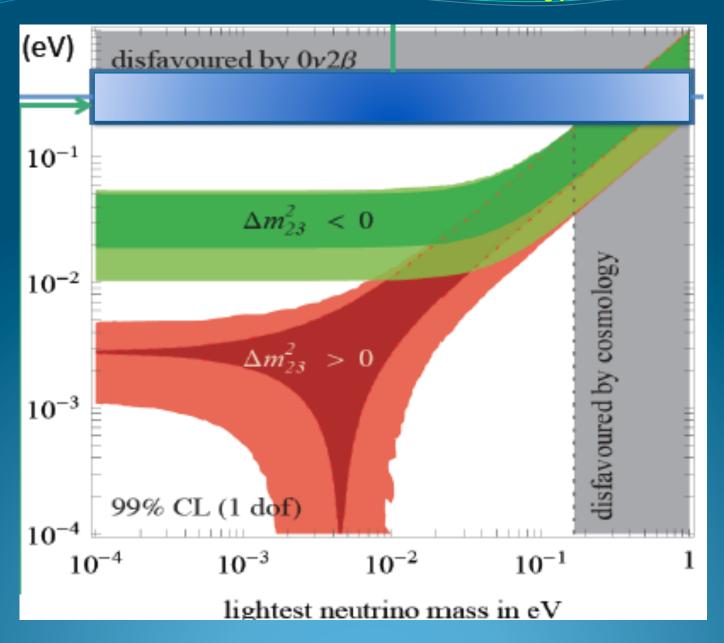
- 1. Dirac [add 3 right-handed v_{R_i} assume (B-L)]
- 2. Majorana [Broken (B-L)], favoured because of
- Unique d=5 op in L_{eff} , Λ as large as almost M_{GUT}
- Simplest see-saw mechanism with heavy ν_R
- Makes possible baryogenesis via leptogenesis Clear th bias for 2., but exp open question: $(\beta\beta)_{0\nu}$ Neither 1. nor 2. affect the success of L_{SM} until Λ

Known unknowns

Is L violated?



My n.1 question: is L violated? $(\beta\beta)_{0v}$ experiments



Main issues of v-related theory talks

Gavela, Hernandez, Tamborra, Palomares-Ruiz

- Cosmological constraints/hints for sterile v's (difficult to draw firm conclusions? Planck?)
- Are we in a better position to study flavour models now that we have 2 flavour sectors? Non-abelian discrete symmetries? Yukawa couplings as dynamical variables? (beware MFV spoiled by higher-dim ops when integrating out heavy fields carrying flavour)
- Can the scale of lepton flavour be low enough to give detectable signals in charged LFV? Y
- Mass hierarchy from atmospheric neutrinos?

Beyond the SM with Dark Matter

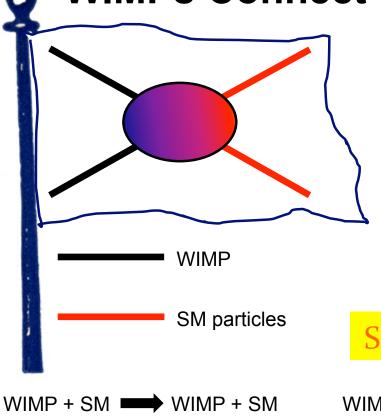
WIMP = Weakly Interacting Massive Particle A good argument for new physics at the TeV scale but not fully compelling: DM could well be axions

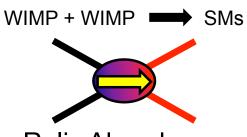
For WIMPs in thermal equilibrium after inflation

$$\langle \sigma_{\rm ann} v \rangle \simeq 3 \times 10^{-26} \rm cm^3 s^{-1}$$

EW x-section for particle with M~10²⁻³ GeV
The WIMP miracle...

WIMPs Connect to Standard Model

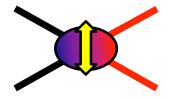




Relic Abundance

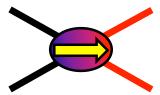
$$\Omega_{\rm DM} h^2 = 0.112 \longrightarrow$$

Stolen from R.Kolb



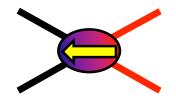
Direct Detection

WIMPs SMs



Indirect Detection

SMs WIMPs



Collider Production

Theory issues with DM

Ibarra, Rydbeck, Tytgat, Lopez-Honorez

Variety of models explored to produce diverse interesting signals in indirect or collider searches

My comments:

LHC results may eventually shift focus from "social" (MSSM, Xdim,...) to "simplified" DM models

Check the consistency of the approximation when using pointlike 4-fermion operators to put bounds on DM at the LHC with monojets or monophotons

Naturalness

coefficients small only because of symmetries It works in many cases!

- Electron mass in NR QED \rightarrow positron $\delta m_e \sim \alpha \Lambda \rightarrow \delta m_e \sim \alpha m_e \log \dots$
- 4-f FCNC box diagram with 3 light $q \rightarrow c$ $G_F^2 \Lambda^2 \sim G_F^2 m_W^2$ too large! $\rightarrow G_F^2 m_c^2 OK$
- π^+ π^0 mass difference in QED \Rightarrow ρ $\Delta m_\pi^2 = (3\alpha)/(4\pi)\Lambda^2 \rightarrow \Lambda \sim m_\rho \ OK$

It does not seem to work for the dark energy $\Lambda_{CC} \sim \text{2.4 x 10}^{\text{-3}} \text{ eV}$

The naturalness puzzle in the SM

No quantum SM symmetry recovered for $m_H \rightarrow 0$ (scale invariance broken by quantum corrections and UV physics)

SM unnatural unless New Physics at the LHC scale

$$\delta m_H^2 \sim -\frac{3h_t^2}{8\pi^2} \Lambda^2 < O(m_H^2) \rightarrow \Lambda < O(500) \text{ GeV}$$

However, precision tests of the SM suggest:

 $\Lambda > O(few) TeV [flavour-ind operators]$

 $\Lambda > O(1000)$ TeV [flavour-dep operators]

What is the way out?

Ways out of the naturalness puzzle

Insist on the few viable (almost) natural models

- 1) Natural supersymmetry
- 2) Natural composite scalar

They can all be ruled out by the LHC

Have we missed some more subtle possibilities (perhaps in connection with gravity and DE)?

Puzzle might be solved only in the full theory (mysterious IR-UV connection missed by EFT)

Naturalness vs. flavour tests

Sala

Flavour: excellent agreement between data and CKM picture

In other words:
$$\Delta \mathcal{L} = \sum_i \frac{1}{\Lambda_i^2} \mathcal{O}_i \quad \Rightarrow \quad \Lambda_i \gtrsim 10^3 \div 10^4 \, \mathrm{TeV}$$

Hierarchy problem:

$$m_h \approx \Lambda$$

 $m_h \approx \Lambda$ [Λ = highest scale h couples to]

Possible way out:
$$\Delta \mathcal{L} = \sum_{i} \xi_{i} \frac{c_{i}}{\Lambda_{i}^{2}} \mathcal{O}_{i}$$

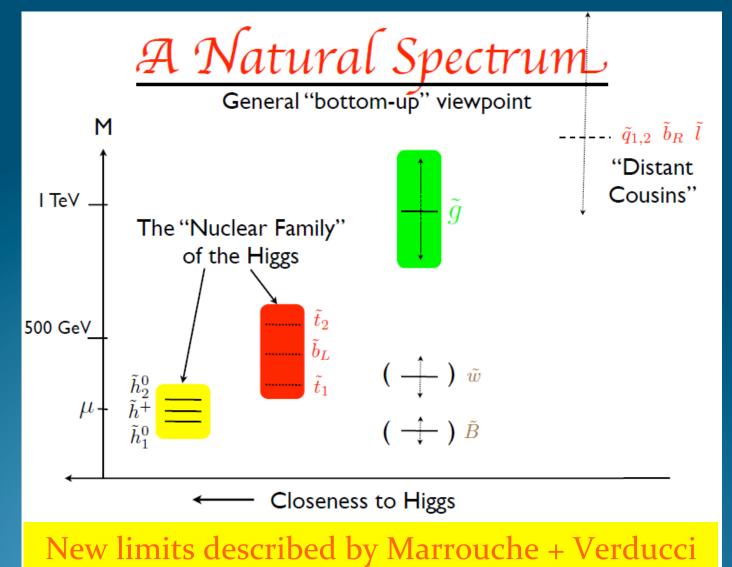
with $c_i \sim O(1)$ and ξ_i small due to some **flavour symmetry**

$U(2)^3 = U(2)_{Q_I} \times U(2)_{U_R} \times U(2)_{D_R}$

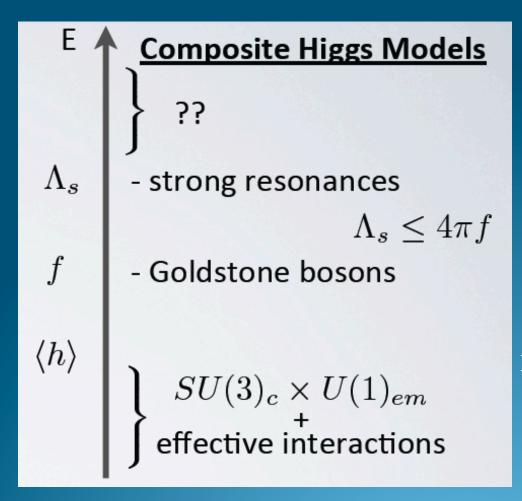
- $\checkmark \ \xi \sim V_{CKM}^{2 \div 4} \ \Rightarrow \ \Lambda \sim \text{a few TeV}$ is OK with flavour bounds
- potentially rich phenomenology behind the corner
- ✓ other virtues...

Natural SUSY

Example:



Natural composite scalar



Partial compositeness: Light scalar correlates with light top partners

Merlo, Azatov

Walking technicolor with approximate scale invariance: technidilaton with non-SM couplings

Yamawachi

Supersymmetry and ascendants

Kazakov, Carena, Ellwanger, Dudas, Sagnotti

My opinion (after following the experimental searches and contributing to the theory for roughly 30 years):

- Too good an idea to be wasted by Nature (general symmetry of RQFT, role in superstrings, etc)
- Might need to be combined with some additional ingredient to solve the SM naturalness problem
- Conventional susy models (CMSSM, NMSSM,...) do not work as such and should finally rest in peace

SUSY phenomenology

A quote from another summary talk, H.Georgi at a conference in Santa Barbara I attended in the early 90's (giving a talk on susy pheno):

"stop wandering in susy parameter space"

At the moment I was not very happy, but I think we theorists can now be more useful:

- Pointing out to experimentalists possible signals they may have overlooked so far in the searches
- Trying to understand what we are missing within simple controllable (even non-realistic) contexts

Seemingly unnatural SUSY

If we do not insist on naturalness, SUSY with heavy scalars can evade direct searches and flavour constraints while mantaining gauge coupling unification + DM candidate

E.g. mini-split supersymmetry: Dudas scalars at ~ 1000 TeV, gauginos 1-2 loop factors lighter (R-symmetry), higgsinos model-dependent, $m_H \sim 125$ GeV easy

Conclusions (of Winter 2013)

HEP-EX is on the move The triumph of the SM Naturalness challenged Where is BSM physics? We must know, we will know but we must be patient and try hard We are lucky to live in such exciting times!

Many thanks to:

- The Secretariat: Isabelle Cossin, Vera de Sa-Varanda, Sarodia Vydelingum
- The Computer Support: Gérard Dreneau, Damien Fligiel, Victor Mendoza, Olivier Drevon
- The Organizers: Etienne Augé, Jacques Dumarchez, Jean Tran Thanh Van
- The Scientific Organizers present in La Thuile: Lydia Iconomidou-Fayard, Jean-Marie Frère, Eric Armengaud, Patrick Janot, Jean-Pierre Lees, Sotiris Loucatos, François Montanet, Jean Orloff as well as those who contributed remotely.