

Moduli stabilization and mixed gauge-gravity mediation

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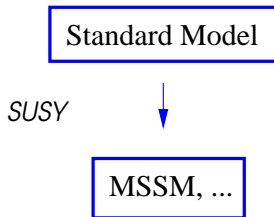
E. Dudas, Y. Mambrini, S. Pokorski, A.R.
JHEP 04 (2008) 015 [arXiv:0711.4934]

E. Dudas, Y. Mambrini, S. Pokorski, A.R. *work in progress*

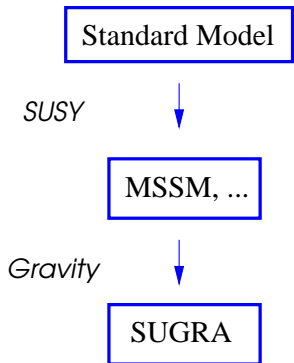
Motivations and Ingredients

Standard Model

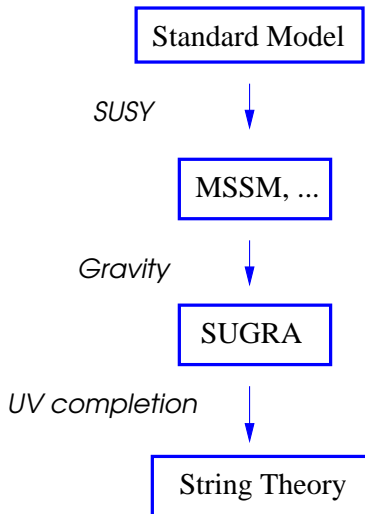
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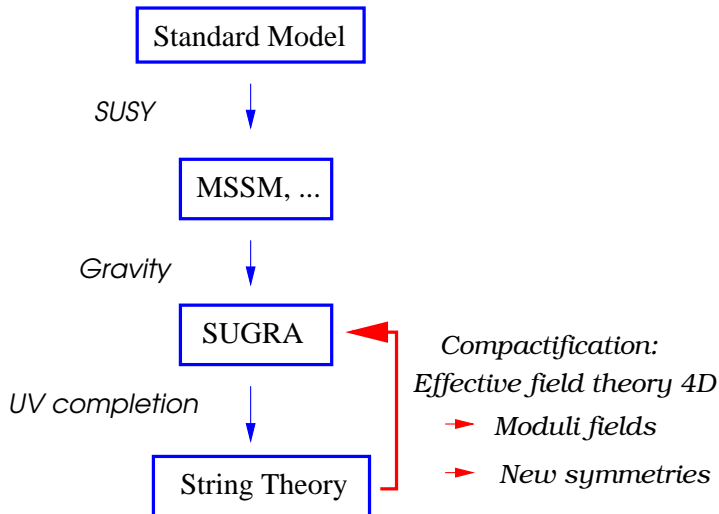
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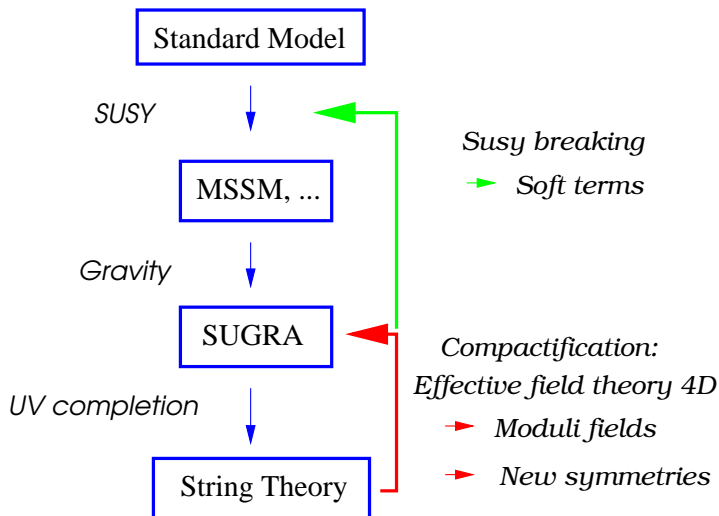
Motivations and Ingredients



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Motivations and Ingredients



Motivations and Ingredients

Moduli

- Fields parametrizing "flat directions": there's no potential for them
- Their vev's are related to physical parameters
→ they have to be taken into account and **stabilized**

New Symmetries

- New $U(1)_X$ gauge symmetries appear very naturally
- New charged fields Φ_i in the spectrum → interplay in the game

Moduli problem: Main Ideas of KKLT model

Kachru, Kallosh, Linde, Trivedi, hep-th/0301240

General Framework

Type IIB 10D compactified on orientifold \rightarrow SUGRA in 4D

Three logical "steps"

- 1- Turn on magnetic fields **Giddings, Kachru, Polchinski, hep-th/0301240**
 \hookrightarrow All moduli stabilized, except "T" (overall volume)
- 2- Provide nonperturbative potential for T
 \hookrightarrow AdS vacuum ($\Lambda_c < 0$)
- 3- Uplift the vacuum energy adding $\overline{D3}$ -branes **[KKLT]**
 \hookrightarrow Susy explicitly broken, not standard SUGRA

From the MSSM point of view

SUSY breaking in a new hidden sector

\hookrightarrow soft terms for the MSSM masses & couplings

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Other possibilities: stay in SUGRA & change step 3

Main ingredients for SUSY theories

- real **Kahler potential**: $K \rightarrow$ kinetic terms
- holomorphic **superpotential**: $W \rightarrow$ interaction terms

Auxiliary fields

- Scalar multiplets $\rightarrow F$
- Vector multiplets $\rightarrow D$

E.O.M.: F, D functions of the scalar fields depending on K & W

Scalar potential in SUGRA

$$V \sim \Sigma (|F - \text{terms}|^2 + |D - \text{terms}|^2) - 3|W|^2$$

SUSY preserved if $F_i = D_i = 0 \rightarrow \Lambda_c \sim -3|W|^2$

Other possibilities: stay in SUGRA & change step 3

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Scalar potential in SUGRA

$$V \sim \Sigma (|F - terms|^2 + |D - terms|^2) - 3|W|^2$$

Idea: F_i and/or $D_i \neq 0 \rightarrow$ uplift and spontaneous SUSY breaking

Adding an extra anomalous $U(1)_X$

T charged under $U(1)_X$

- it transforms as a Goldstone boson
- mass to $U(1)_X$ gauge boson
- coupling depends on T \rightarrow anomalies \rightarrow G.-S. mechanism
- gauge invariance + SUSY + other fields
 - \hookrightarrow Fayet-Iliopoulos term depending on T
 - \hookrightarrow constraints for terms in the superpotential

More in details

$$K = -3 \ln(T + \bar{T}) + |\Phi_+|^2 + |\Phi_-|^2 + K_{MSSM}$$

$$W = W_0 + e^{-bT} \Phi_-^q + m \Phi_+ \Phi_-$$

$$\hookrightarrow V = e^K (|F_T|^2 + |F_+|^2 + |F_-|^2 - 3|W|^2) + g_T^2 D^2$$

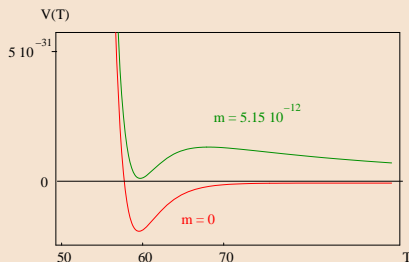
FI mechanism \rightarrow spontaneous SUSY Breaking: $D, F_i \neq 0$

First Phenomenological Results

Uplift

Parameters in the superpotential fixed by some phenomenological requirements:

- $m_{3/2} \sim \frac{W}{M_P^2} \Rightarrow \text{fix } W_0$
- Imposing $\Lambda_{\text{cosm}} = 0 \Rightarrow \text{fix } m$



Soft terms by gravitational effects:

$$m_{1/2} \sim \frac{F_T}{T} \quad (\text{if } T \text{ couples to SM vector fields})$$

$$m_0 \sim m_{3/2}$$

Usually F_T is small....here:

$$m_{1/2} \lesssim m_0 \sim m_{3/2} \sim \text{TeV} \quad D \sim (100 \text{ TeV})^2$$

Anomalies & Gauge Messengers

If T couples to the SM gauge fields, since it is charged under $U(1)_X$, its shift generates mixed $U(1)_X - G_a^2$ anomalies

$(G_a = SU(3), SU(2)_L, U(1)_Y)$

$$\int d^2\theta T W_a^\alpha W_{\alpha a} \rightarrow \text{Re}[T] F_a^{\mu\nu} F_{\mu\nu a} + i \text{Im}[T] F_a^{\mu\nu} \tilde{F}_{\mu\nu a}$$

Possibilities :

- MSSM fields charged under $U(1)_X \Rightarrow$
problems: large masses for squarks and sleptons $\sim \sqrt{D}$,
tachyonic directions,...
- MSSM fields uncharged but additional fields M, \tilde{M}

M, \tilde{M} "Messengers" fields introduced in Gauge Mediated Supersymmetry Breaking

GMSB: Messengers M, \tilde{M} couple to Susy Breaking hidden sector (superfield S)

$$\lambda \int d^4\theta S M \tilde{M}$$



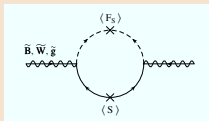
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↪ mediation of Susy Breaking at perturbative level

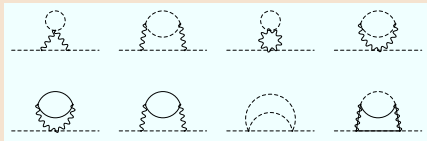
Gaugino masses

$$m_{1/2}^{GMSB} \sim \frac{g_a^2}{8\pi^2} \frac{F_S}{S}$$



Scalar masses

$$(m_0^{GMSB})^2 \sim \frac{g_a^4}{128\pi^4} \left(\frac{F_S}{S}\right)^2$$



Gauge Mediation & Uplift + D-term

Scalar soft masses

Poppitz, Trivedi, hep-ph/9703246

$$\text{STr } \mathcal{M}_{\text{Mess}}^2 = \text{Tr } m_{\text{scal}}^{\text{mess} 2} - \text{Tr } m_{\text{ferm}}^{\text{mess} 2} \neq 0$$

\Rightarrow different GMSB result for scalars

In this model...

- Anomalies argument: M, \tilde{M} charged "+" under $U(1)_X$

\hookrightarrow $\text{STr } \mathcal{M}_{\text{Mess}}^2 \sim D$ (D-term)

\hookrightarrow Coupling with SUSY breaking sector: $\lambda \int d^4\theta \Phi_- M \tilde{M}$

True GMSB contribution for scalar soft masses

$$(m_0^{\text{GMSB}})^2 = \frac{g_a^4}{128\pi^4} \left\{ \left(\frac{F_-}{\Phi_-} \right)^2 + D \left[2 - \ln \left(\frac{\Lambda_{\text{UV}}^2}{m_{\text{Mess}}^2} \right) \right] \right\}$$

Mixed Mediation

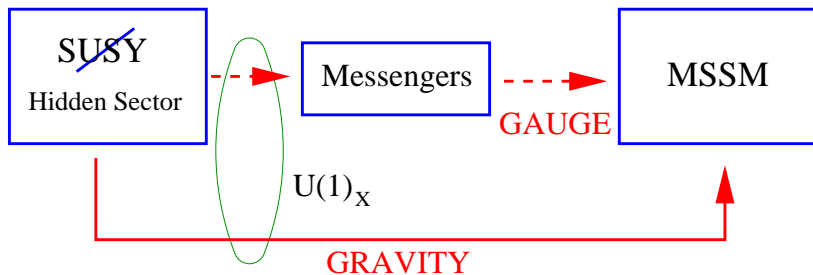
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Nontrivial interplay between $(m_0^2)_{\text{grav.}}$ and $(m_0^{GMSB})^2$:

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 \hookrightarrow Here: $(m_0^{GMSB})^2 < 0$ and $|(m_0^{GMSB})^2| \sim m_{3/2}^2 \sim (\text{TeV})^2$

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- $(m_0^{GMSB})^2$ depends on couplings between scalars and gauge vectors

Mixed Mediation

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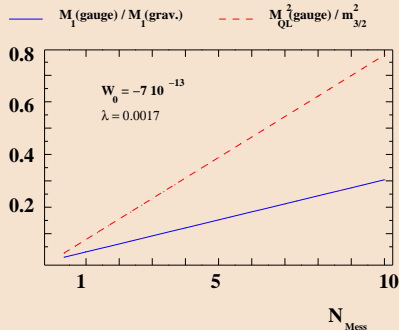
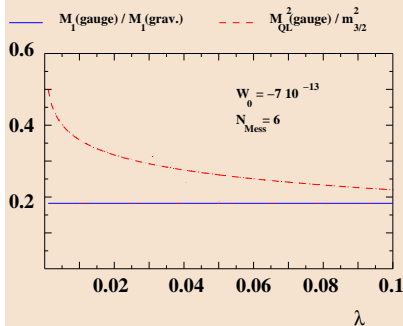
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- $(m_0^{\text{GMSB}})^2$ depends on couplings between scalars and gauge vectors

\Rightarrow Squarks can be lighter than sleptons at high energy
 \hookrightarrow measurable effects at low energy after RG flow

Gravity vs. Gauge

Playing with λ and N_{Mess}

$$(m_0^2) = (m_0^2)_{\text{grav.}} + N_{\text{Mess}}(m_0^{\text{GMSB}})^2$$
$$m_{1/2} = (m_{1/2})_{\text{grav.}} + N_{\text{Mess}}(m_{1/2}^{\text{GMSB}})$$



Some numbers

	A	B
W_0	$-7 \cdot 10^{-13}$	$-4.3 \cdot 10^{-13}$
m	$7.3 \cdot 10^{-12}$	$4.5 \cdot 10^{-12}$
a	1	1
b	0.3	0.5
q	1	1
$\tan \beta$	30	15
t	98.3	59.4

μ (GeV)	810	1070
$B\mu$ (GeV) ²	(400) ²	(870) ²
$m_{\chi_1^0}$	110	140
$m_{\chi_1^\pm}$	220	290
$m_{\tilde{g}}$	760	950
m_h	120	120
m_A	2220	3290
$m_{\tilde{t}_1}$	1380	1770
$m_{\tilde{t}_2}$	1920	2610
$m_{\tilde{c}_1}, m_{\tilde{u}_1}$	2580	3300
$m_{\tilde{b}_1}$	1910	2610
$m_{\tilde{b}_2}$	2310	3230
$m_{\tilde{s}_1}, m_{\tilde{d}_1}$	2580	3300
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	A+GMSB	B+GMSB
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t	97.3	59.4
λ	$1.7 \cdot 10^{-3}$	$1.1 \cdot 10^{-3}$
N_{Mess}	6	6

μ (GeV)	186	216
$B\mu$ (GeV) ²	(330) ²	(730) ²
$m_{\chi_1^0}$	120	150
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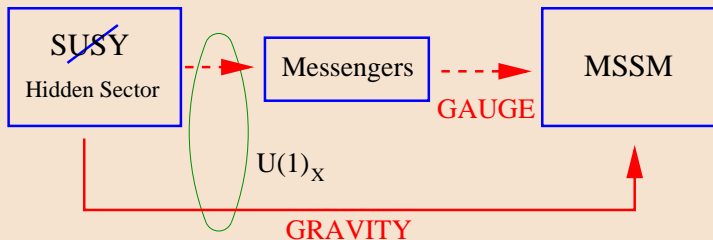
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Hybrid Model

Gauge - Gravity mixed mediation

Idea: Generalize the previous stringy motivated example



Generic scenario with **both** gauge and gravity contributions to the soft terms at the GUT scale

↪ Suitable **parametrization** and **scan** on the parameters space

Parametrization

Soft terms for the masses

$$(m_0^2)_i = M_S^2 [1 + N_{\text{mess}} \tilde{\alpha}_i (1 - \tilde{\beta})]$$

$$(m_{1/2})_a = f M_S$$

Parametrization

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- M_S → scale of the **gravity** mediation

Parametrization

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- M_S → scale of the gravity mediation
- N_{mess} → number of **messengers**

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- f → both gravity & gauge contributions **universal**

Parametrization... $\tilde{\alpha}_i$

$$m_{Q_L}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{4}{3}n_q\alpha_3^2 + \frac{3}{4}n_l\alpha_2^2 + \frac{3}{5}\left(\frac{1}{6}\right)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$m_{U_R}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{4}{3}n_q\alpha_3^2 + \frac{3}{5}\left(\frac{-2}{3}\right)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$m_{D_R}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{4}{3}n_q\alpha_3^2 + \frac{3}{5}\left(\frac{1}{3}\right)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$m_{E_L}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{3}{4}n_l\alpha_2^2 + \frac{3}{5}\left(\frac{-1}{2}\right)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$m_{E_R}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{3}{5}(1)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$m_{H_u}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{3}{4}n_l\alpha_2^2 + \frac{3}{5}\left(\frac{1}{2}\right)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$m_{H_d}^2 = M_S^2 + (1 - \tilde{\beta})\tilde{\alpha} \left[\frac{3}{4}n_l\alpha_2^2 + \frac{3}{5}\left(\frac{-1}{2}\right)^2 \frac{2n_q + 3n_l}{5}\alpha_1^2 \right] M_S^2,$$

$$(m_{1/2})_a = M_S f = M_{1/2}$$

Sample spectra

SUSPECT + micrOMEGAs2.0

	A	B	C
M_S	1.3 TeV	1.3 TeV	1.3 TeV
$M_{1/2}$	500	500	500
$\tilde{\alpha}$	0	1	1
$\tilde{\beta}$	0	0	1.85
$\tan \beta$	5	5	5
μ	750	1070	258
M_1	212	213	209
$m_{\chi_1^0}$	211	211	186
$m_{\chi_1^+}$	400	408	430
$m_{\tilde{g}}$	1222	1251	1175
m_h	113.2	114.1	112.4
m_A	1530	1920	1095
$m_{\tilde{t}_1}$	1066	1260	844
$m_{\tilde{t}_2}$	1403	1780	1008
$m_{\tilde{c}_1}, m_{\tilde{u}_1}$	1612	2052	1119
$m_{\tilde{b}_1}$	1388	1771	954
$m_{\tilde{b}_2}$	1588	1891	1282
$m_{\tilde{s}_1}, m_{\tilde{d}_1}$	1613	2053	1122
$m_{\tilde{\tau}_1}$	1275	1448	1046
$m_{\tilde{\tau}_2}$	1302	1550	1106
$m_{\tilde{\mu}_1}, m_{\tilde{e}_1}$	1303	1551	1047
Ωh^2	15.1	29.4	0.105

	A	B	C
M_S	1.3 TeV	1.3 TeV	1.3 TeV
$M_{1/2}$	500	500	500
$\tilde{\alpha}$	0	1	1
$\tilde{\beta}$	0	0	1.85
$\tan \beta$	5	5	5
μ	750	1070	258
M_1	212	213	209
$m_{\chi_1^0}$	211	211	186
$m_{\chi_1^+}$	400	408	430
$m_{\tilde{g}}$	1222	1251	1175
m_h	113.2	114.1	112.4
m_A	1530	1920	1095
$m_{\tilde{t}_1}$	1066	1260	844
$m_{\tilde{t}_2}$	1403	1780	1008
$m_{\tilde{c}_1}, m_{\tilde{u}_1}$	1612	2052	1119
$m_{\tilde{b}_1}$	1388	1771	954
$m_{\tilde{b}_2}$	1588	1891	1282
$m_{\tilde{s}_1}, m_{\tilde{d}_1}$	1613	2053	1122
$m_{\tilde{\tau}_1}$	1275	1448	1046
$m_{\tilde{\tau}_2}$	1302	1550	1106
$m_{\tilde{\mu}_1}, m_{\tilde{e}_1}$	1303	1551	1047
Ωh^2	15.1	29.4	0.105

- Squarks more sensitive than sleptons to GMSB negative contribution

Sample spectra

SUSPECT + micrOMEGAs2.0

	A	B	C
M_S	1.3 TeV	1.3 TeV	1.3 TeV
$M_{1/2}$	500	500	500
$\tilde{\alpha}$	0	1	1
$\tilde{\beta}$	0	0	1.85
$\tan \beta$	5	5	5
μ	750	1070	258
M_1	212	213	209
$m_{\tilde{\chi}_1^0}$	211	211	186
$m_{\tilde{\chi}_1^\pm}$	400	408	430
$m_{\tilde{g}}$	1222	1251	1175
m_h	113.2	114.1	112.4
m_A	1530	1920	1095
$m_{\tilde{t}_1}$	1066	1260	844
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$m_{\tilde{\mu}_1}, m_{\tilde{e}_1}$	1303	1551	1047
Ωh^2	15.1	29.4	0.105

$$\mu^2 \approx -m_{H_2}^2 - \frac{1}{2}M_Z^2$$

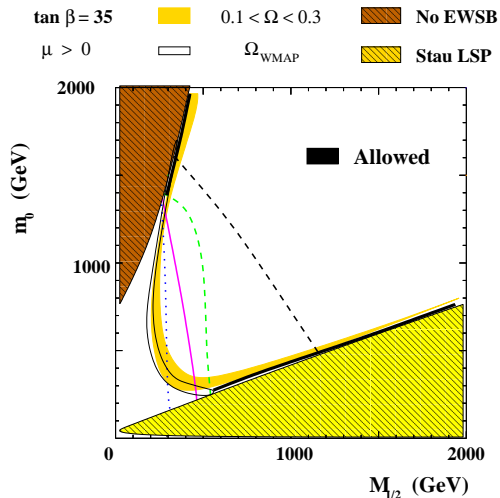
$$\frac{\partial m_{H_2}^2}{\partial \log \mu} \approx 6y_t^2 (m_{H_2}^2 + m_{U_3}^2 + m_{Q_3}^2 + A_t^2)$$

Initial conditions + RG flow effects:

- Lighter squarks
- ↪ lower $|m_{H_2}^2|$ value
- ↪ lower μ value
- ↪ Lightest neutralino is generally a mixed bino-higgsino state
- ↪ Suitable **dark matter relic density** value allowed

Examples of scan

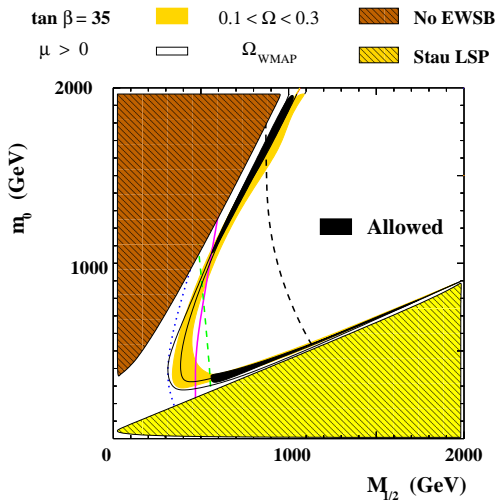
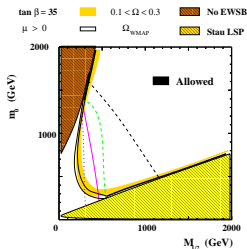
$$N_{mess} = 5 \quad \tilde{\beta} = 5$$



$$\tilde{\alpha} = 0 \quad (\text{mSUGRA})$$

Examples of scan

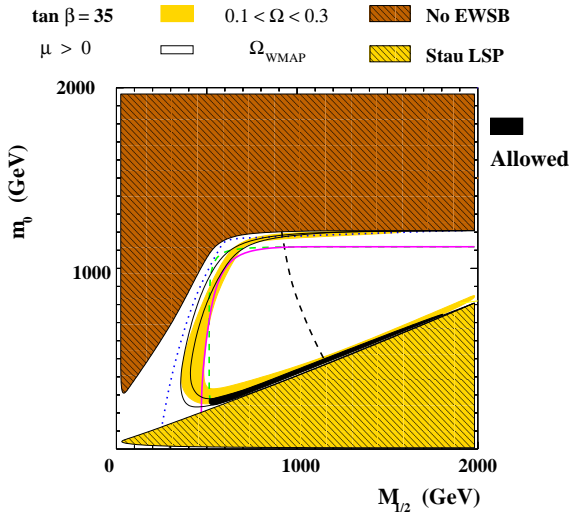
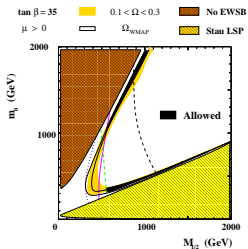
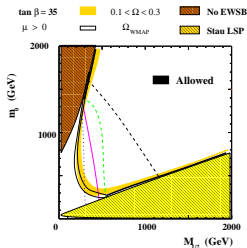
$$N_{mess} = 5 \quad \tilde{\beta} = 5$$



$$\tilde{\alpha} = 1$$

Examples of scan

$$N_{mess} = 5 \quad \tilde{\beta} = 5$$



$$\tilde{\alpha} = 1.1$$

Conclusions & Outlook

Results

- High energy:
 - Example of a **non-decoupled** uplift: modulus contribution F_T to SUSY Breaking is not negligible
 - Natural mixing between **gauge & gravity** mediation mechanisms
- Low energy:
 - Gaugino masses **comparable** to the scalars and gravitino masses
 - **Spectrum "compressed"**: squarks more sensitive to GMSB

Work in progress

- Top-down: Can F_T be increased?
- Bottom-up: Systematic study of the phenomenology of the **hybrid** model