

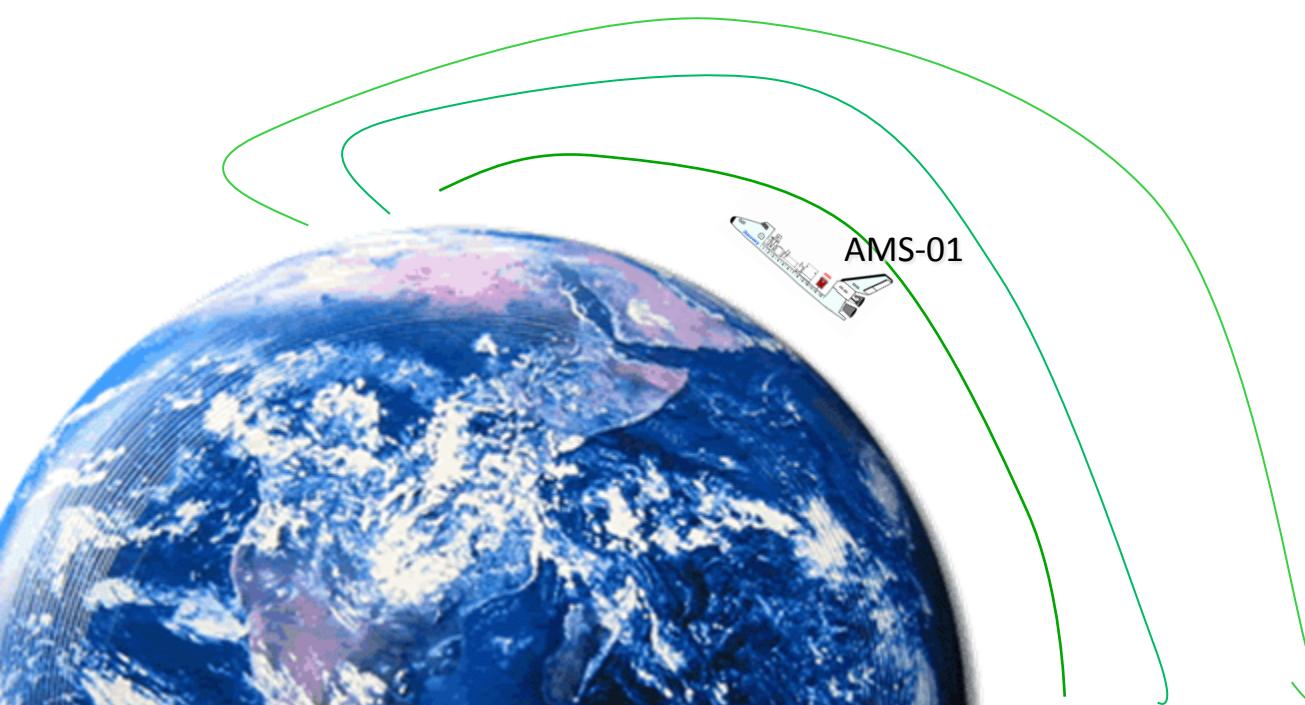


# AMS-01 on the shuttle

## analysis of light nuclei and interactions

*Nicola Tomassetti*

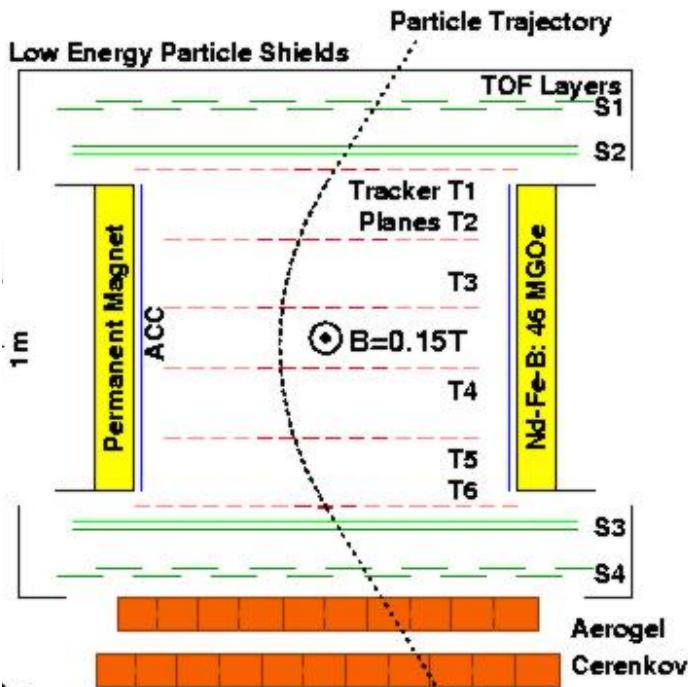
*INFN Perugia*



# Flux Measurement with AMS-01



## AMS-01



## The Instrument

- **Magnet**: dipole field;  $BL^2 = 0.14 \text{ Tm}^2$
- **TOF**: 4 layers of scintillators; fast trigger +  $\beta$  +  $dE/dx$
- **Tracker**: 6 layers of 2-sided Si;  $p/Z + dE/dx + Z/|Z|$
- **ATC**: aerogel threshold Cerenkov; e/p up 3 GeV
- **ACC**: veto against multiparticle events;

$$\phi^Z(E) \approx \frac{\Delta N^Z(E)}{\Delta E \cdot G \cdot \varepsilon^Z(E) \cdot T^Z(E)}$$

Annotations for the equation:

- detector counts* (blue oval)
- species* (red oval)
- bin width* (green oval)
- geom factor  $0.3 \text{ m}^2 \text{ sr}$*  (purple oval)
- detection efficiency* ( $\varepsilon_{\text{TOT}}^Z = \varepsilon_{\text{SEL}}^Z \cdot \varepsilon_{\text{REC}}^Z \cdot \varepsilon_{\text{TRIG}}^Z$ ) (yellow oval)
- effective exposure time* (pink oval)

## Event reconstruction

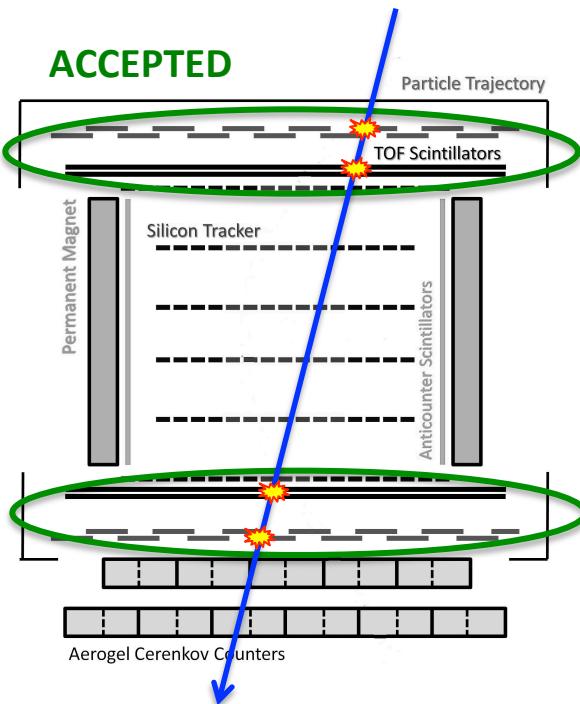
- **Velocity** and direction: by timing (TOF)
- **Rigidity**  $R=p/Z$ : by the reco trajectory (TRACKER)
- **Charge**: by multiple  $dE/dx$  (TOF & TRACKER)

Enter in the flux normalization.  
Not canceled out for ratios:

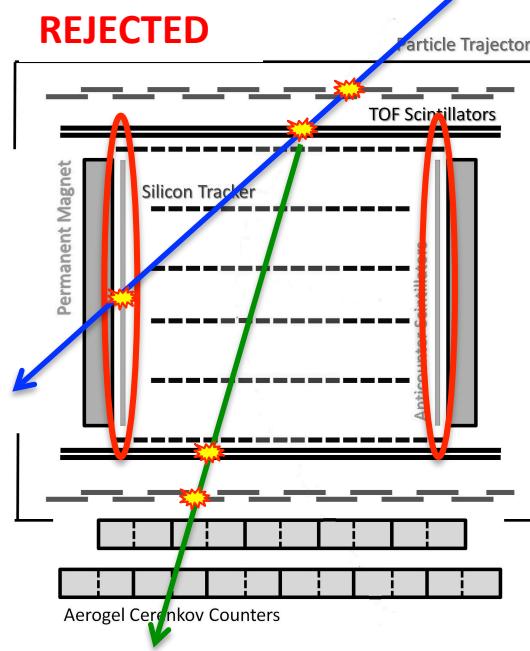
$$B / C \propto \varepsilon^C / \varepsilon^B$$

# Trigger logic

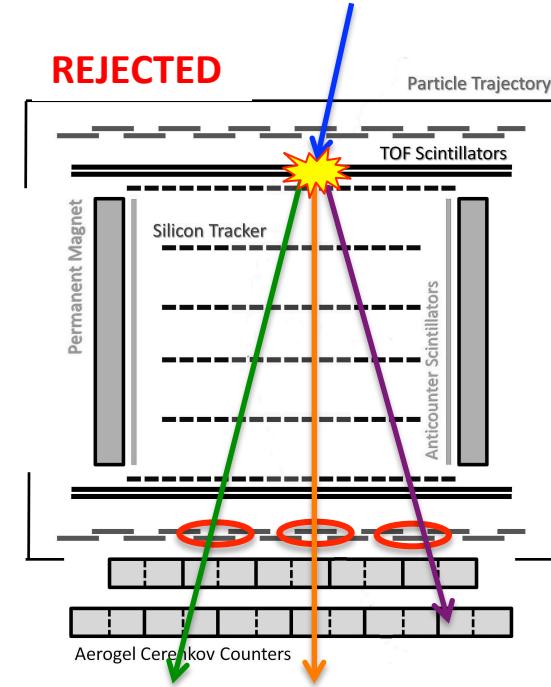
The AMS-01 trigger was designed to reject multi-particle events. 3 requirements:



**(1) TOF 4/4 planes fired**



**(2) No ACC counters fired**



**(3) No Multiple TOF hits**

## Minimum bias trigger: *for efficiency studies*

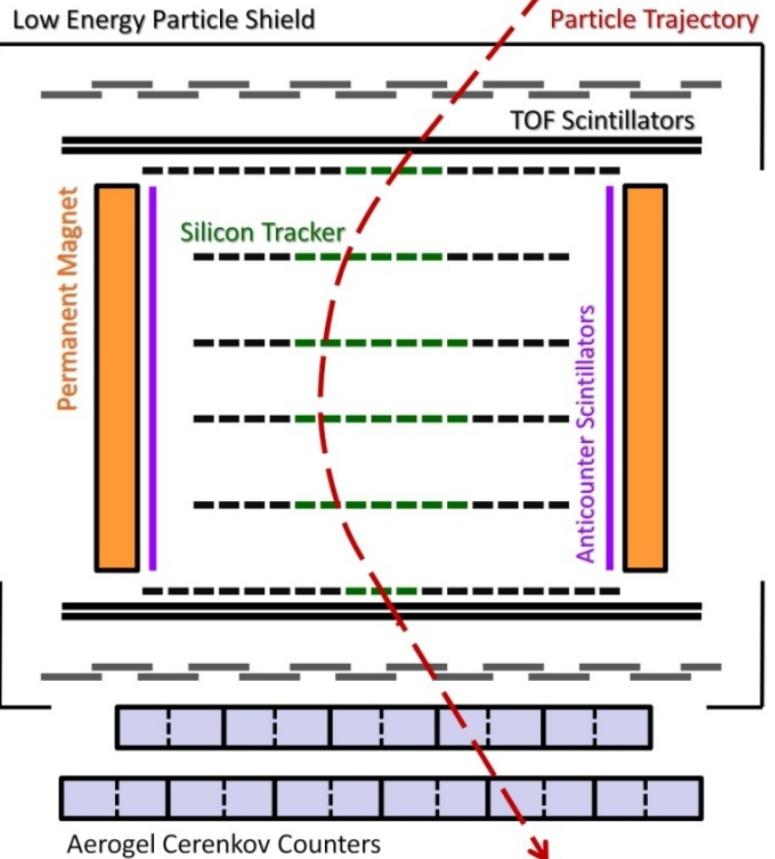
- 1 out of 1000 events recorded without the application of (2) and (3).
- Allow to study/validate trigger efficiency with data.
  - Prescaling factor 1/1000 was too small for Z>2 nuclei

# Some examples

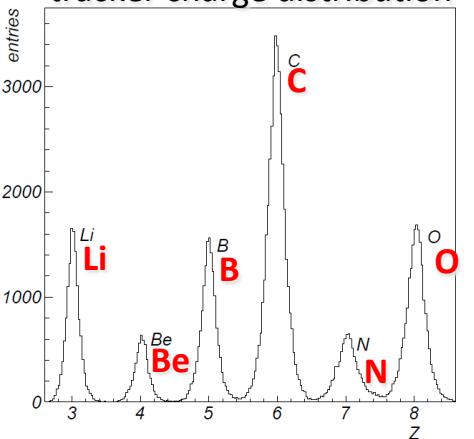


## THE GOOD EVENTS

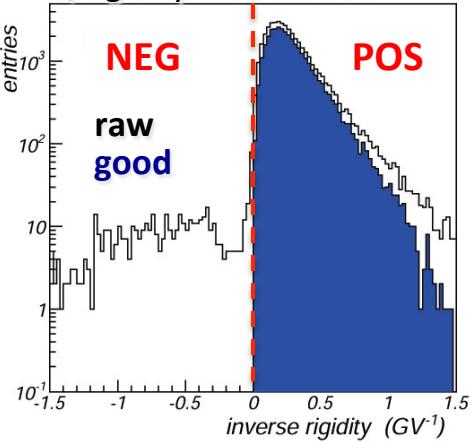
→ Used to compute fluxes



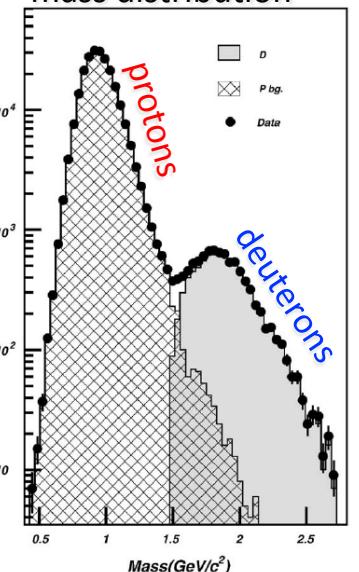
tracker charge distribution



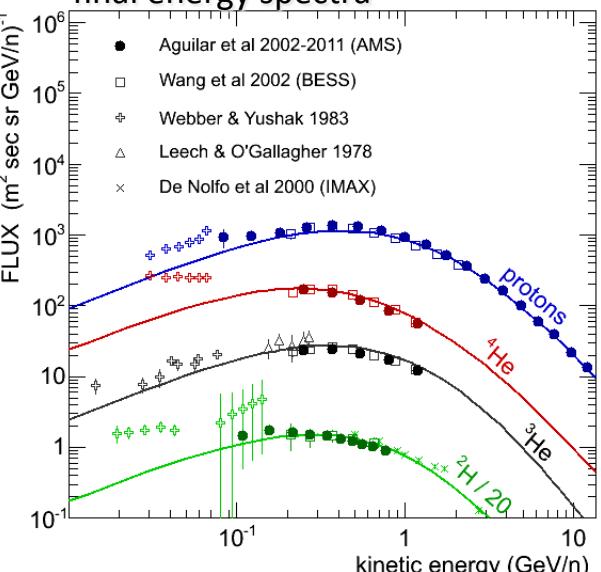
1/rigidity distribution



mass distribution



final energy spectra



- Ideal event: a CR particle traverses AMS, giving a clean track. All active layers are fired.

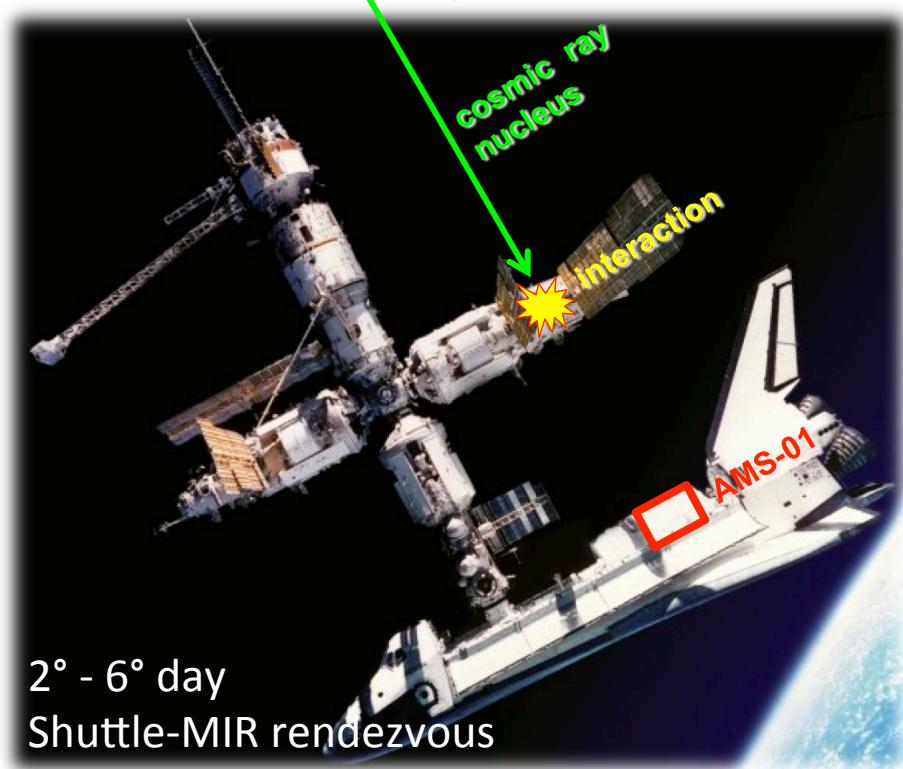
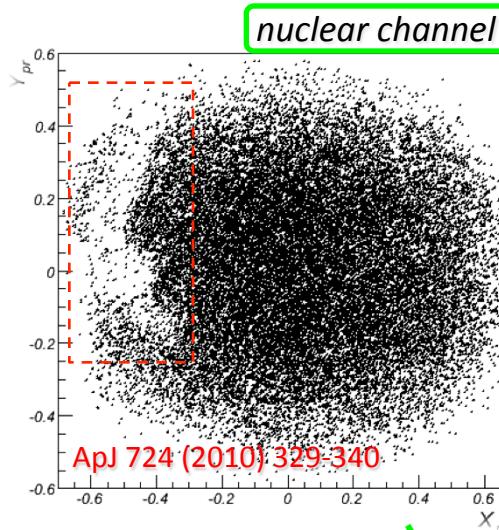
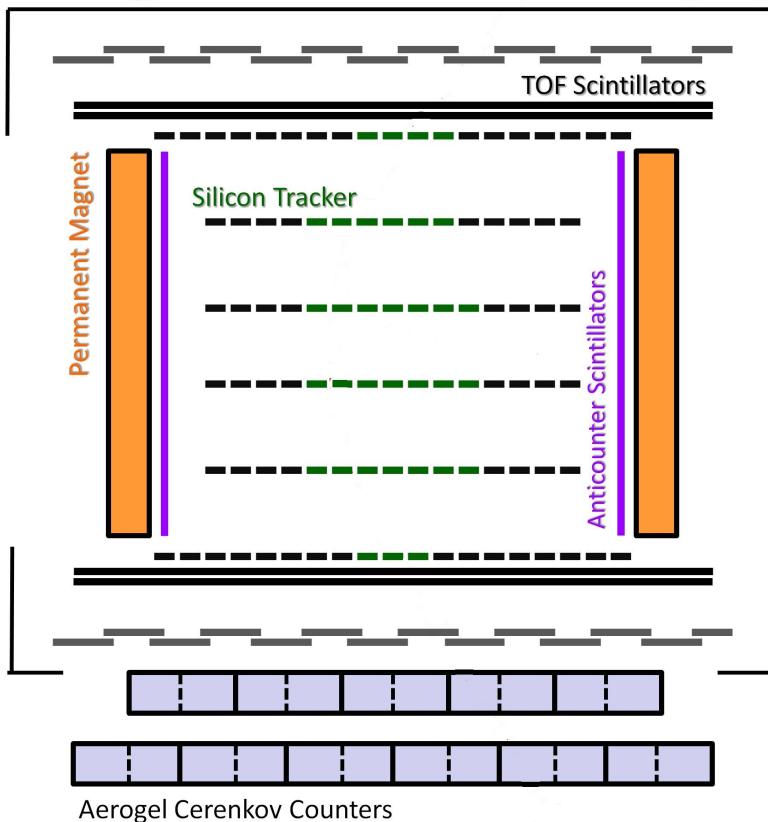
# Some examples



## INTERACTIONS ABOVE → missing flux & BG

Low Energy Particle Shield

Particle Trajectory



- Balloons: atmospheric attenuation/BG
- AMS-01: MIR modules (while docked)
- AMS-02: ISS solar arrays (when pass)

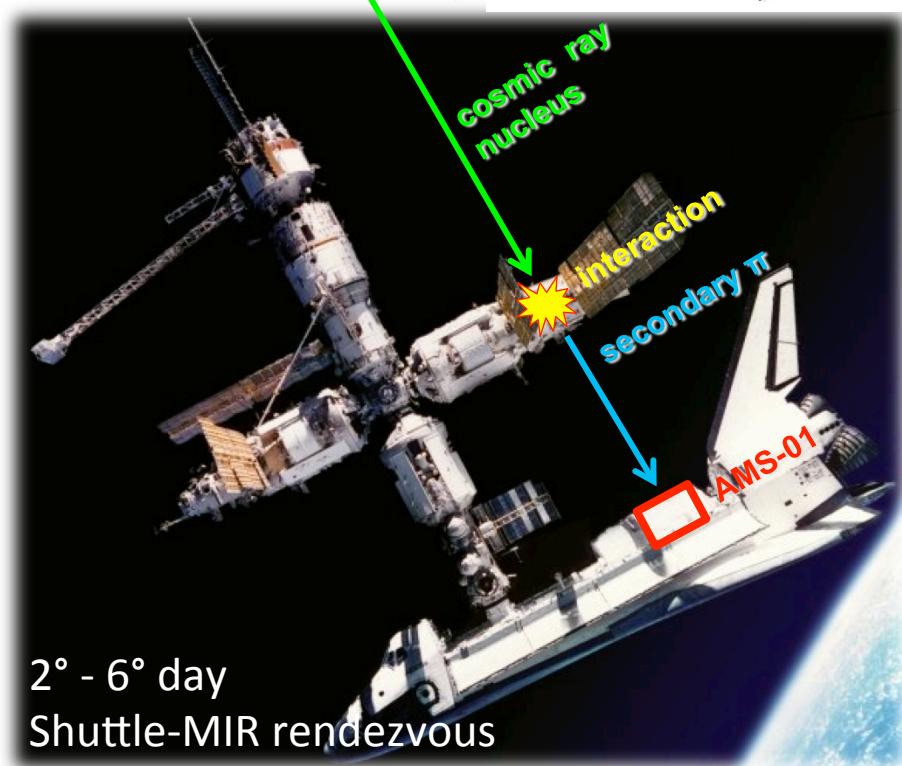
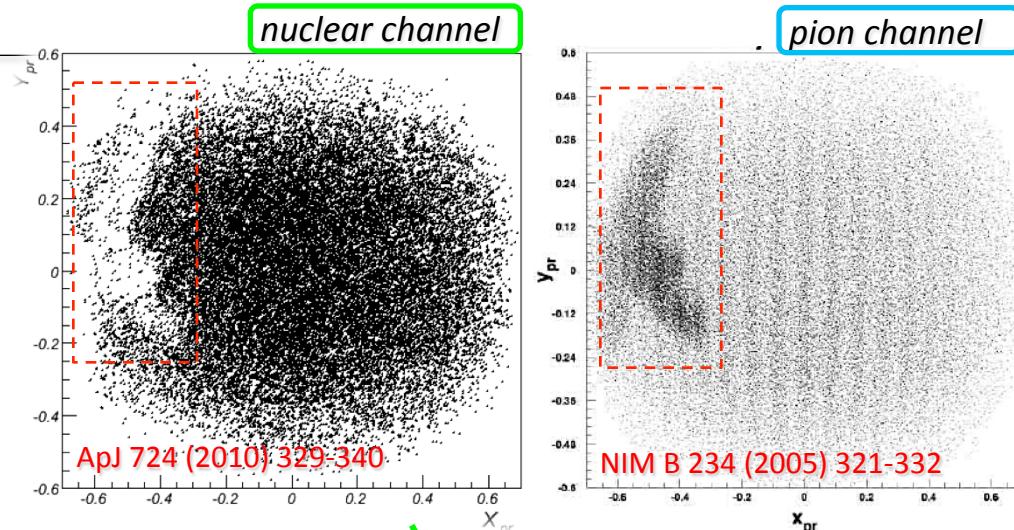
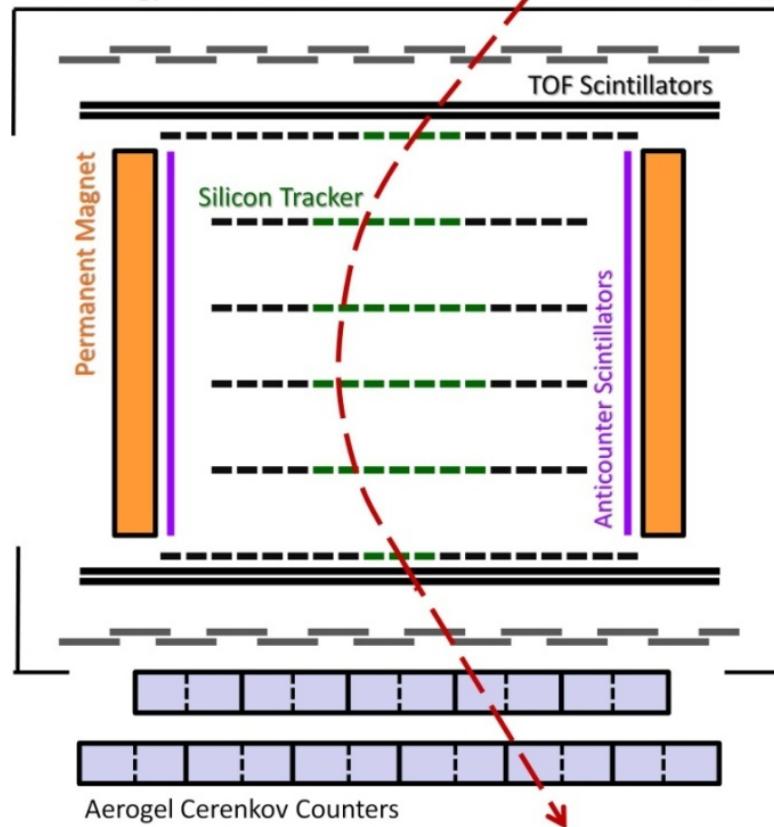
# Some examples



## INTERACTIONS ABOVE

→ missing flux & BG

Low Energy Particle Shield

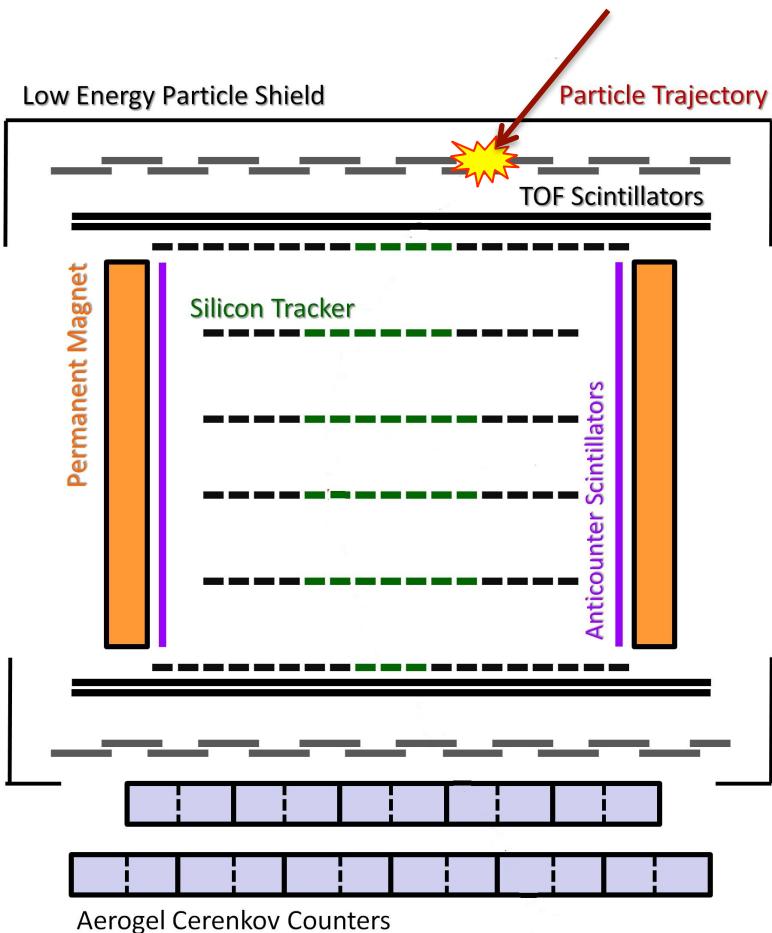


- Balloons: atmospheric attenuation/BG
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# Some examples



## TOI INTERACTIONS



TOI interactions (above tracking volume)  
 → Absorption -> missed event  
 → Secondary production -> rejected evt

### Survival fraction

$$P = e^{-x/\lambda} \approx 90\%$$

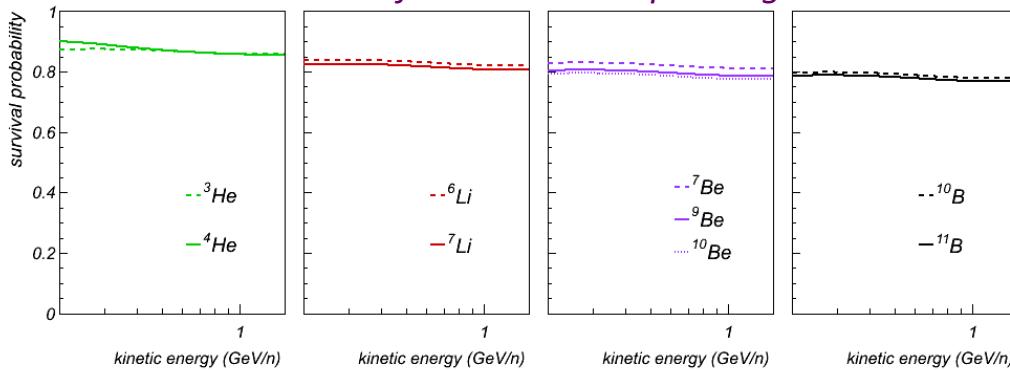
$$\left. \begin{aligned} \lambda_{He} &\sim \frac{M_{He}}{\sigma_{He}} \sim 50 \text{ g/cm}^2 \\ x &\sim 5 \text{ g/cm}^2 \end{aligned} \right\}$$

### Uncertainty for He

$$\frac{\delta P}{P} \approx \frac{x/\lambda}{1-x/\lambda} \left( \frac{\delta \sigma}{\sigma} \right) \approx 1\% \quad \sim 10\%$$

	IMAX	BESS	AMS-01
MASS RESOLUTION	<12%	<12%	12-14%
MASS PDF MODEL	semi empirical	double gaussian	full MC
TOI MATERIAL	12 g/cm <sup>2</sup>	9 g/cm <sup>2</sup>	4 g/cm <sup>2</sup>
ATMOSPHERE	5 g/cm <sup>2</sup>	5 g/cm <sup>2</sup>	0 g/cm <sup>2</sup>
EVENTS/BINS	47,650/10	29,937/7	18,035/5

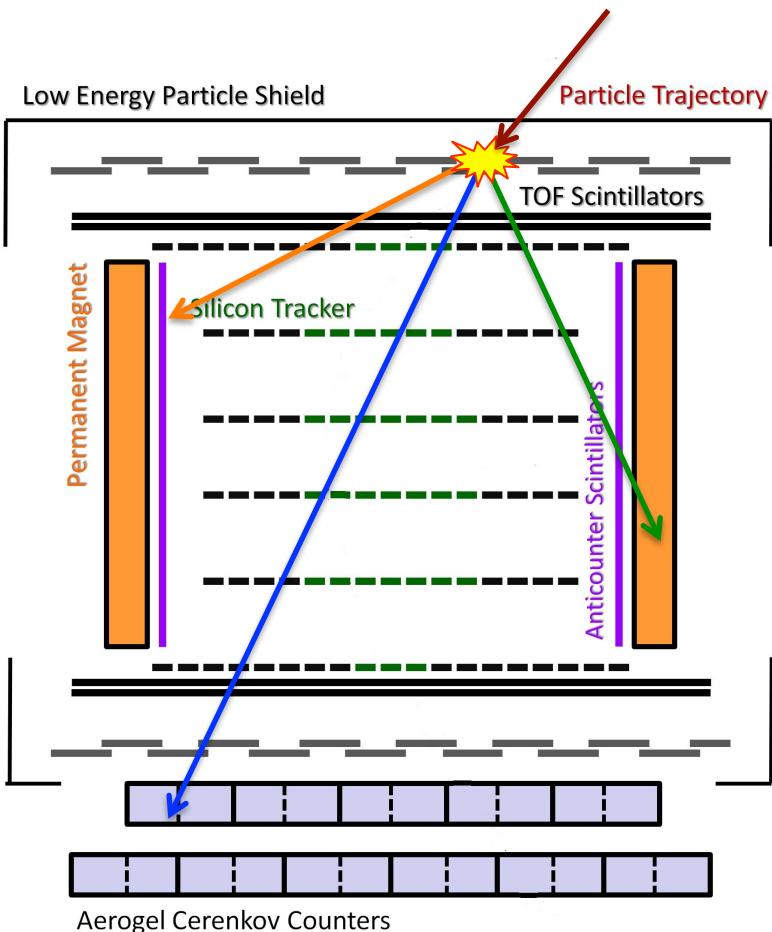
### AMS-01 TOI Survival fraction – G4TripathiLightCrossSection



# Some examples



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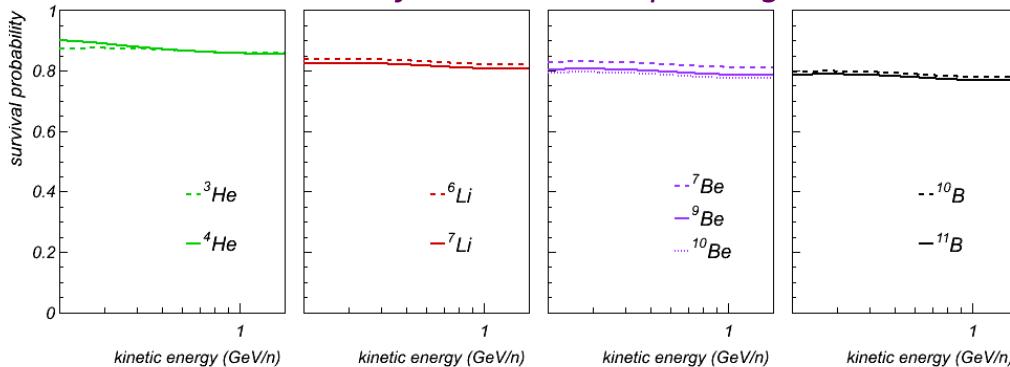
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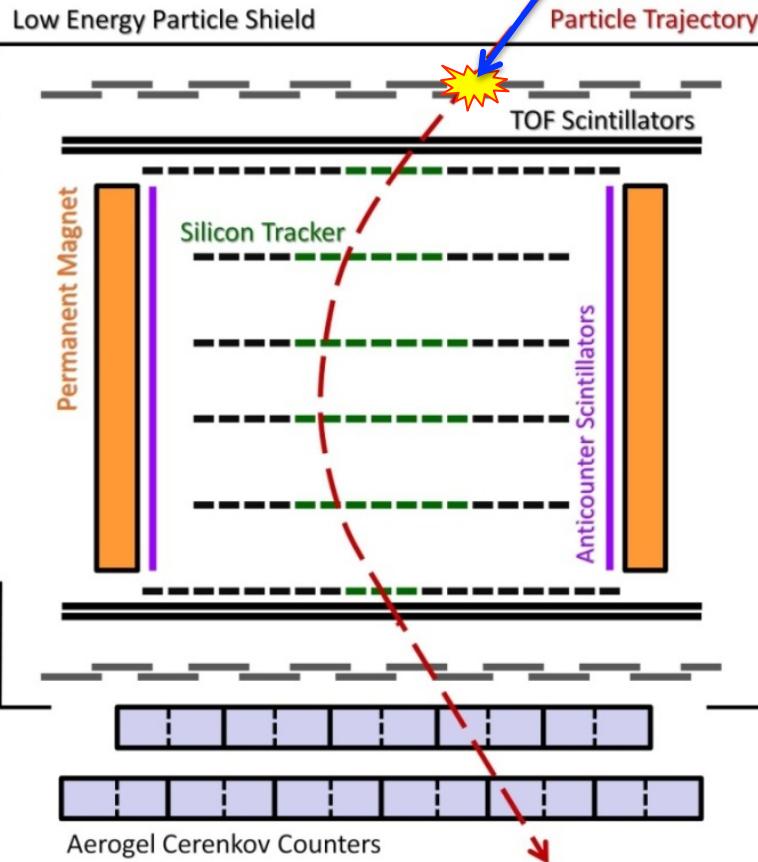
### AMS-01 TOI Survival fraction – G4TripathiLightCrossSection



# Some examples



## TOI CHARGE-CHANGING

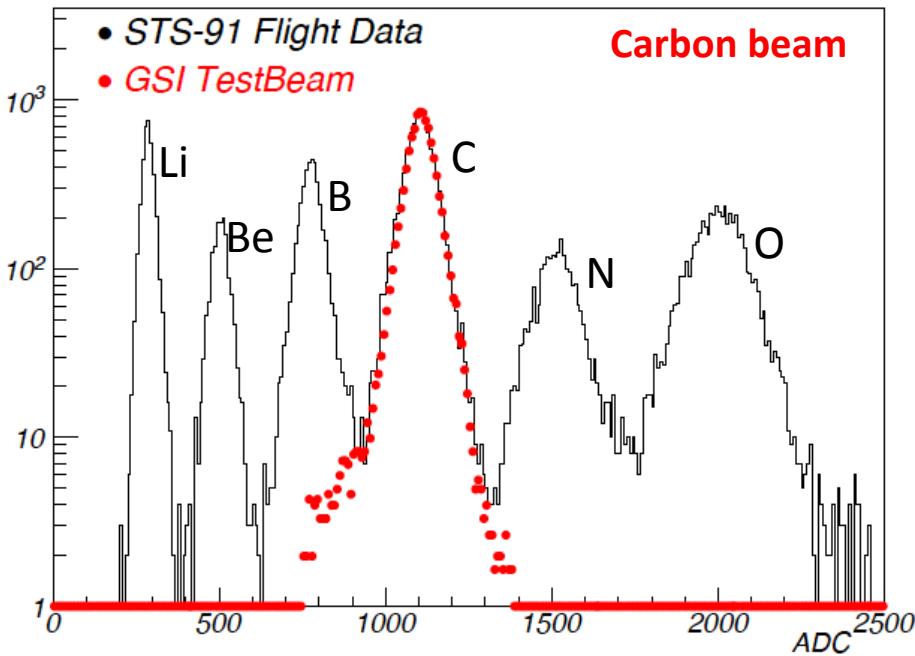


TOI charge-changing processes  
Clean track reconstructed. Wrong Z.

fragmentation matrix:  $P(Z_{\text{rec}} | Z_{\text{true}})$

$Z_{\text{true}}$

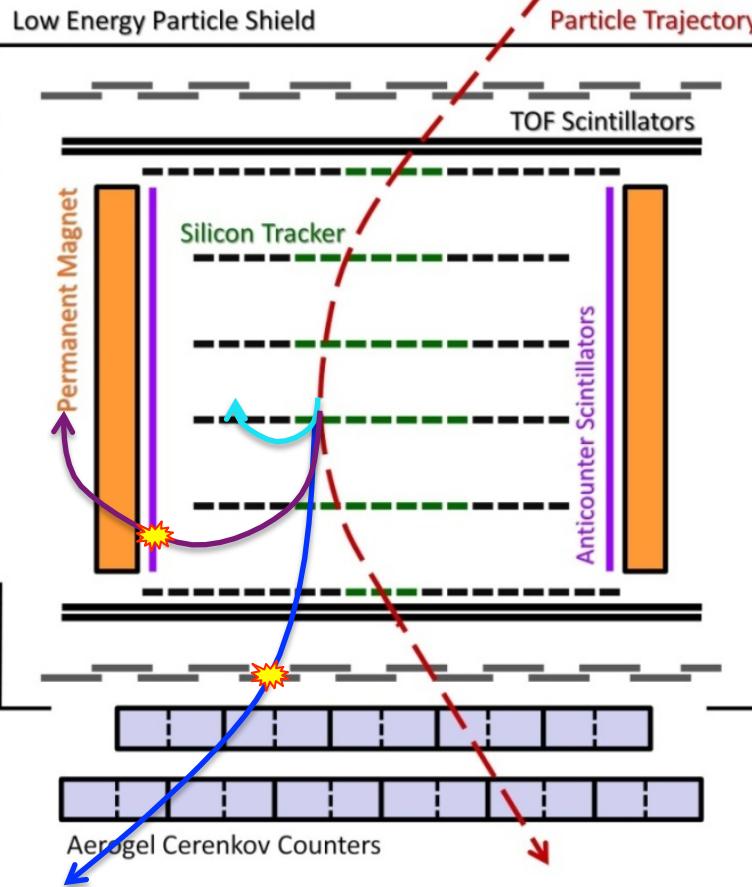
%	Li	Be	B	C	N	O
Li	99.99	0.12	0.06	0.02	0.01	0.01
Be		99.88	0.36	0.10	0.08	0.07
B			99.58	0.53	0.15	0.08
C				99.24	5.31	4.07
N					94.44	0.01
O						95.77



# Some examples



## DELTA-RAY EMISSION



High-energy delta-ray emission.

Event may (or may not) be rejected by trigger.  
Strongly dependent on primary Z and Energy.

*Number of delta-rays emitted at energy T*

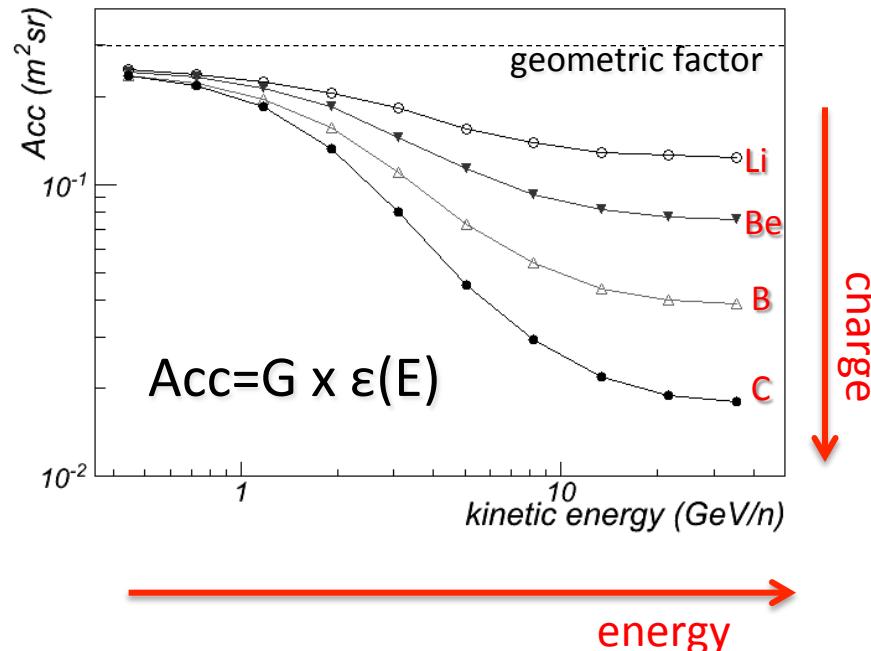
$$\frac{dN}{dT} \propto \frac{Z^2}{T^2}$$

increases with Z  
high energy tails

*Maximum energy of delta-rays*

$$T_{Max} \approx 2m_e\beta^2c^2\gamma^2$$

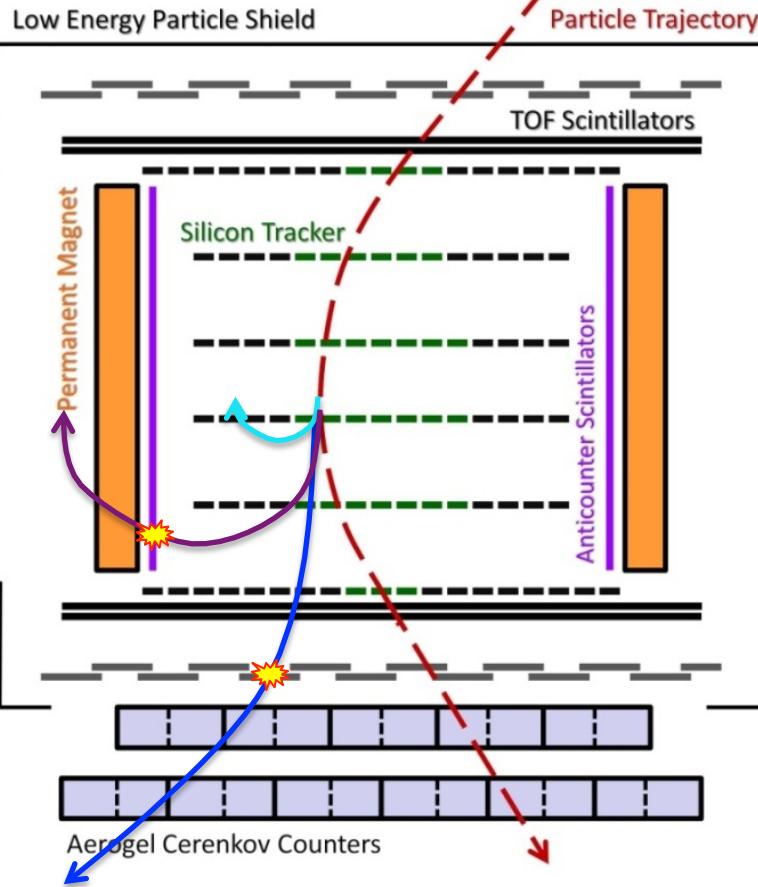
increases with the primary CR energy



# Some examples



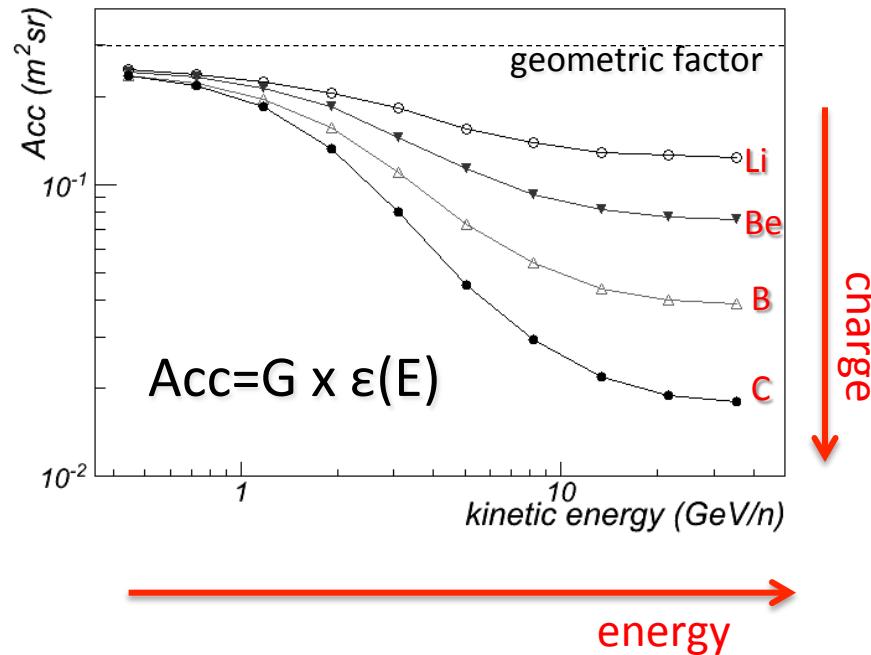
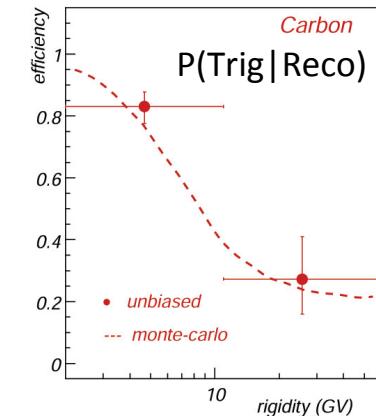
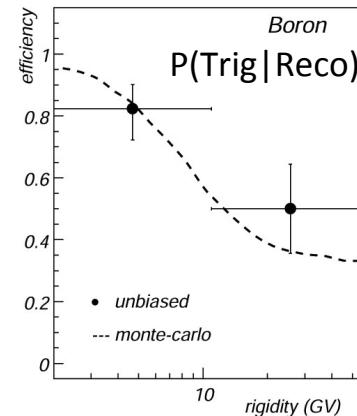
## DELTA-RAY EMISSION



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Trigger eff. with minimum-bias data



# Trigger Efficiency



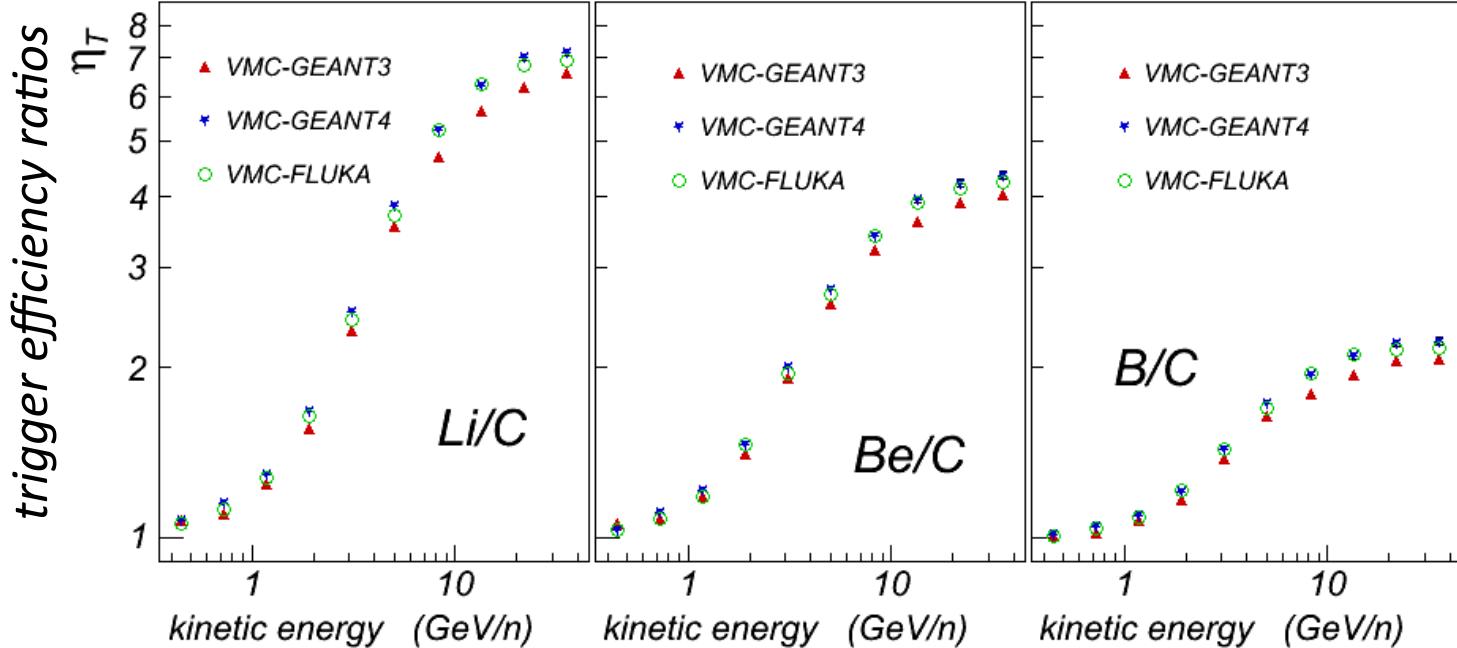
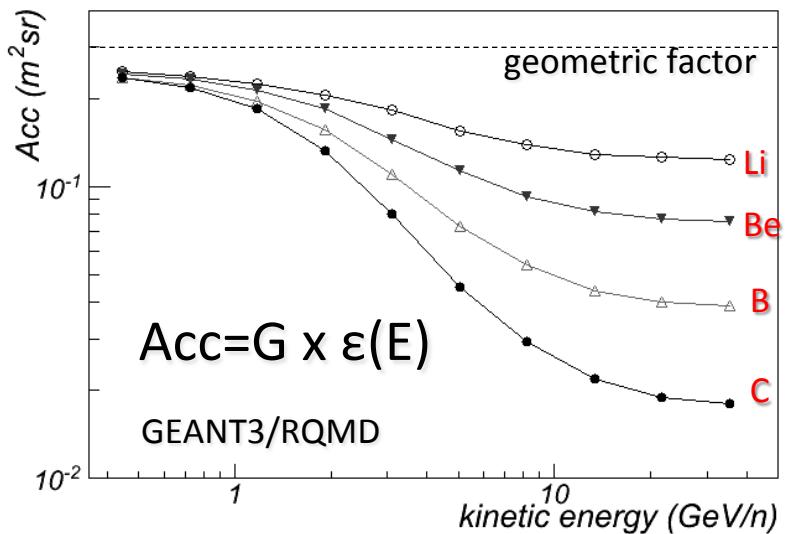
MC estimations were performed with three different simulations codes of particle interactions.

**GEANT3**

**GEANT4**

**FLUKA**

$$\eta_{\text{TRIG}}^Z = \epsilon_{\text{TRIG}}^Z / \epsilon_{\text{TRIG}}^6$$



# Summary

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CR interactions in material affect the flux measurements.

1. Inefficiency due to missing flux.
  - Increases with CR mass number A ( $\sim Z$ ) and TOI material thickness
  - Requires reliable estimate of destruction cross-sections
  
2. Background due to charge- (mass-) changing processes
  - Complex hadronic MC simulation needed.
  - Can be minimized with analysis cuts.
  - Can be studied with DATA, using redundant Z-evaluations.

## Limitations with AMS-01

1. Tot optimized for  $Z>2$  nuclei. Strong E-Z-dependent trigger efficiency.
2. Minimum-bias DATA not sufficient to validate the MC efficiencies
3. Lack of statistics for high-energy and high-charge nuclei

*The main goal of the AMS-01 mission was to gain experience and knowledge for the final design of the AMS-02 project.*

# AMS-02 : Evidence of fragmentation Carbon → Boron



## Carbon → Boron in Upper TOF

Optimized for high-Z measurements

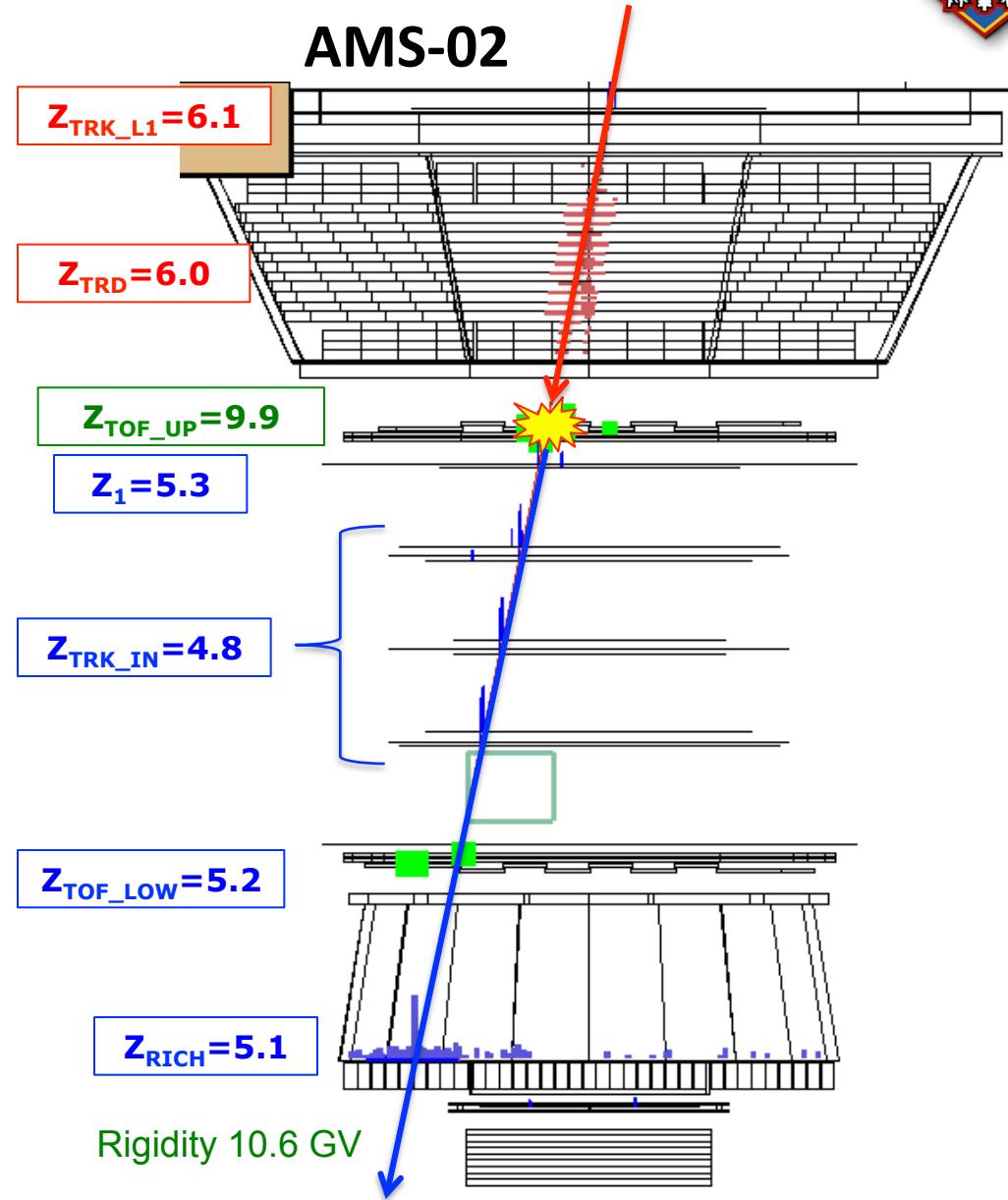
- Large dynamical range:  $Z \sim 1 - 30$
- Many layers of active material.
- Many independent evaluations of  $Z$ .

Dedicated Trigger for  $Z > 1$ :

- 4/4 TOF planes fired
- Multiple TOF hits allowed
- NACC < 5

Minimum bias trigger:

- 1/100 prescaling!!
- 3/4 TOF fired
- No conditions on NACC



## AMS Hadronic Tomography

with the cosmic-ray p/He ratio

Exposure Time: May 20 2011 – May 20 2012

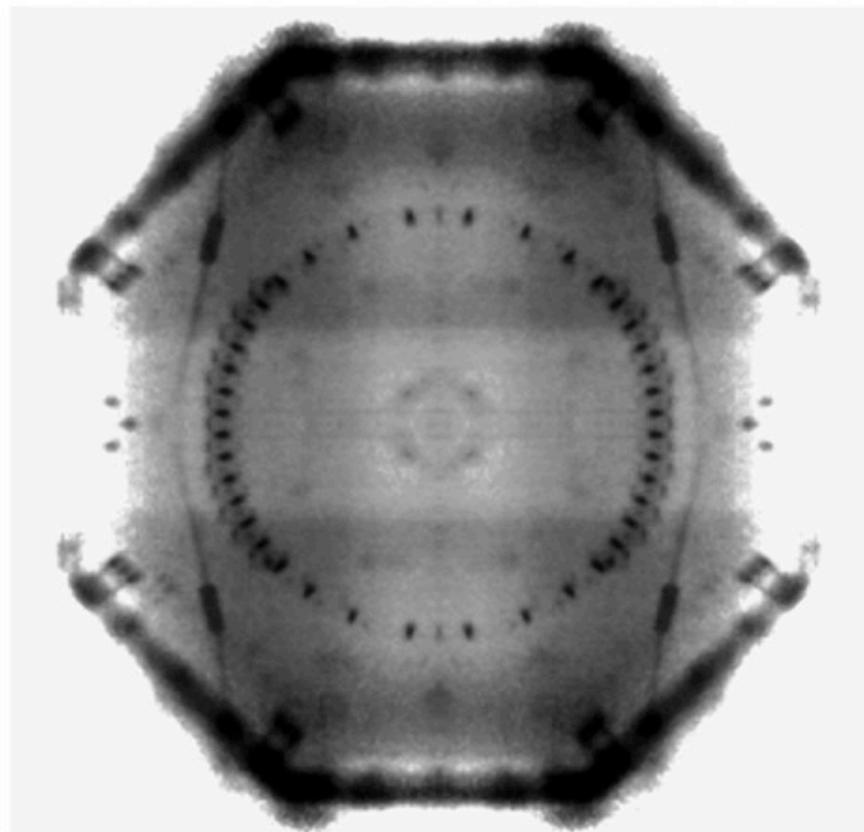
Number of Protons: 3,676,863,217

Number of Helium nuclei: 620,303,906

Rigidity range: 2 GV – 2000 GV

Tomographic plane:  $Z = +165 \text{ cm}$

XY pixel area:  $1 \text{ cm}^2$



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Operating in the International Space Station since May 2011, AMS is performing very accurate measurements of cosmic ray (e.g. Proton and Helium nuclei) with unprecedent sensitivity. This picture represents a “tomographic” reconstruction of the AMS top-of-instrument material obtained using the Proton-to-Helium flux ratio. Tiny changes of the interaction probabilities of these nuclei with different materials are used to trace the material inhomogeneities. Detector elements such as screws, electronics boards, and mechanical interfaces are clearly recognizable.