

Effectives
approaches within
the MSSM and
Beyond :
Applications to
Higgs physics and
Dark Matter
observables

G.Drieu La Rochelle

Introduction

The Standard
Model and Beyond
Higgs and Dark
Matter

Effective approach
for relic density

SUSY at one-loop
Effective vertices
Results

Effective approach
for Higgs physics

The BMSSM
Computing
observables

Masses and
Couplings

Recasting SM Higgs
searches

The Higgs signal

Conclusion

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Effectives approaches within the MSSM and Beyond : Applications to Higgs physics and Dark Matter observables

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Universite de Savoie

PhD defence at Annecy

Outline

- 1 Introduction
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- 3 Effective approach for Higgs physics
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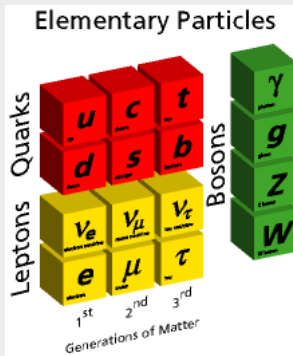
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Particle Physics : where do we stand

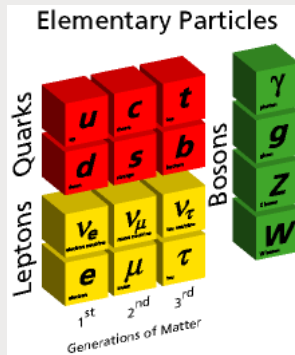
- The Standard Model for particle physics



- Matter : Three families of quark pairs and lepton pairs
- Gauge : Three interactions : γ , W/Z , G
- A symmetry breaking mechanism
 - Non-zero masses for Z , W
 - A scalar particle H : the Higgs boson

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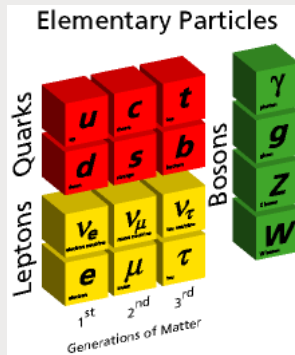
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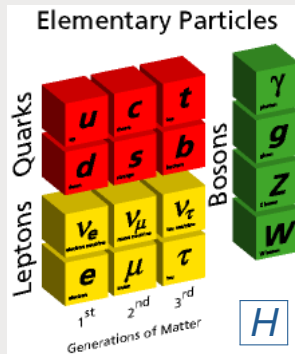
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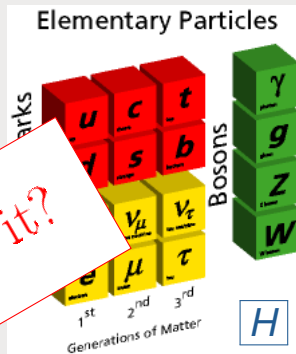
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Particle Physics : where do we stand

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So, is this it?

- Matter : Three families of quark pairs and lepton pairs
 - Gauge : Three interactions : γ , W/Z , G
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 - Non-zero masses for Z , W
 - A scalar particle H : the Higgs boson
- The SM Higgs?

Why Standard Model may not be enough

- There are hints that Standard Model is only an approximation.
- Theoretical side
 - Could be part of a unified theory (GUTs)
 - Does not explain the wide scale between all fermion masses (flavour hierarchy)
 - Naturalness : Higgs mass suspiciously affected by possible heavy particles
- Experimental side
 - Dark Matter : it can be anything but the Standard Model
 - Have we just found the Higgs, or some scalar?

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 - **Naturalness** : Higgs mass suspiciously affected by possible heavy particles → **SUSY**
- **Experimental side**
 - **Dark Matter** : it can be anything but the Standard Model → **SUSY**
 - Have we just found the Higgs, or **some scalar**?

Beyond the Standard Model : Supersymmetry

- Fields $\phi \rightarrow$ Superfields Φ
 - Each Standard Model particle has a superpartner
- Extended Higgs sector : $h/H/A^0$ and H^\pm
new parameters t_β, M_{A^0}
- **Naturalness** : Supersymmetry protects the running Higgs mass
- **Dark Matter** : superpartners of γ, Z, h , called neutralinos ($\tilde{\chi}_1^0$) are good candidates.
 - They are neutral, massive (Supersymmetry breaking) and one is stable if R-parity is assumed

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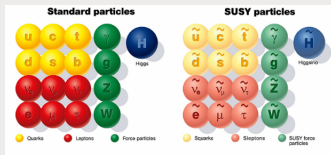
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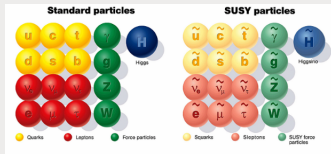
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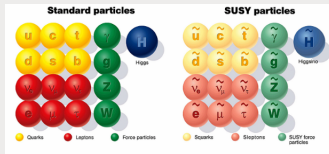
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The goal of the thesis (I)

- Constraining Supersymmetry through Dark Matter observables.

- Constraint on Relic Density

$$\Omega h^2 = 0.1126 \pm 0.0036 \quad (\text{WMAP}) \quad \rightarrow \quad \Delta_{exp} = 3\%$$

- Precise constraint on Supersymmetry

- Requires a **one-loop computation**

- However, most scans in Supersymmetry are tree-level.

- **Aim** \rightarrow gain in precision with simplicity

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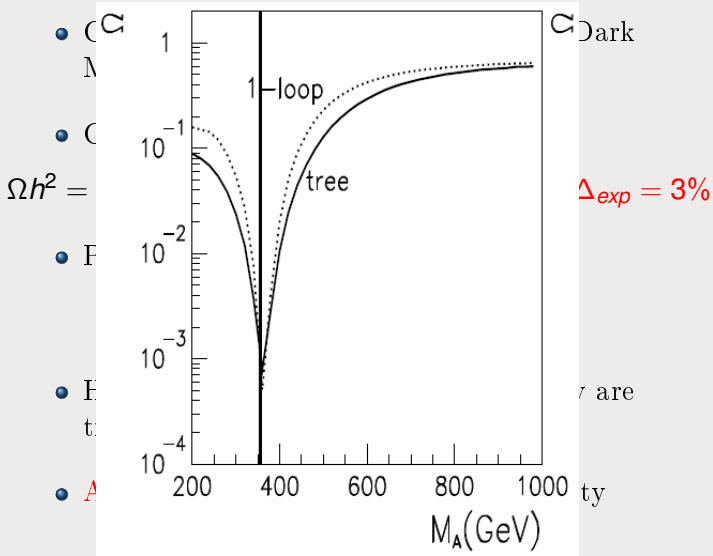
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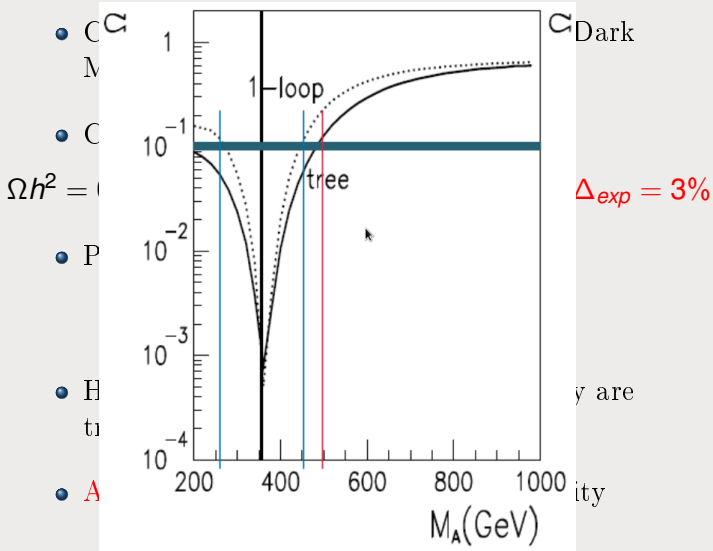
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The goal of the thesis (II)

- Higgs sector of the MSSM is constrained
 - Tension between Naturalness, and a possible heavy Higgs.
 - What if we had found a heavier Higgs ($m_h > 150$ GeV)?
 - What couplings can be allowed with $m_h = 125$ GeV?
- Supersymmetry could well be non-minimal.
 - Account for a new sector coupled to Higgses
 - Allows for a non minimal Higgs sector
→ BMSSM
- Aim → Gain in flexibility in the Higgs sector

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Effective approach

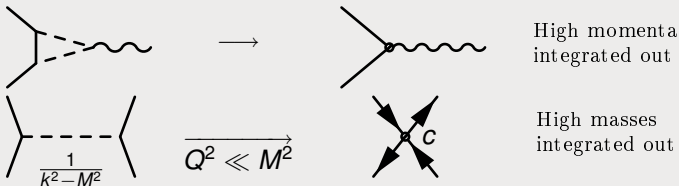
- Equivalence Low-energy theory \Leftrightarrow Full theory

- Possibility for a generic set-up with $\delta\mathcal{L} = \sum_i c_i \mathcal{O}_i$
 - The \mathcal{O}_i respect all symmetries of the low energy theory
- Analogy with LEP

Effective terms $s_w(Q), M_Z(Q), \dots$

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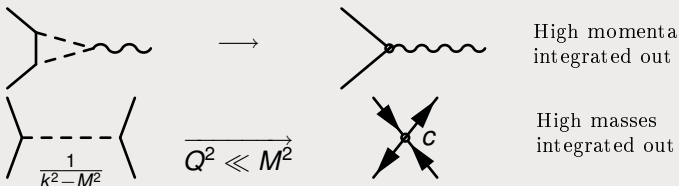


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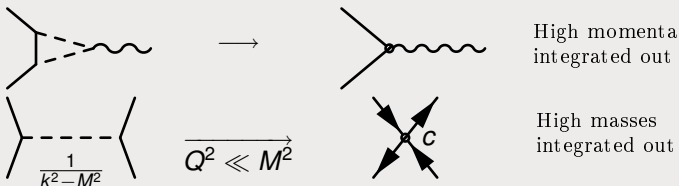


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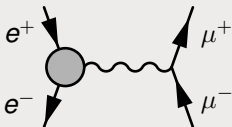
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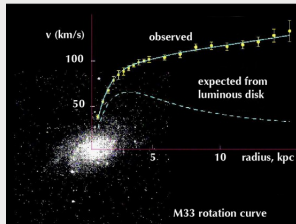
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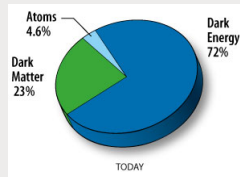
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The Dark Matter problem

- What do we need Dark Matter for?



CMB →



- How to compute the Relic Density

- $\Omega h^2 \Leftrightarrow \sigma_{\tilde{\chi}\tilde{\chi} \rightarrow \text{SM}}$

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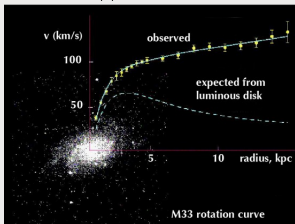
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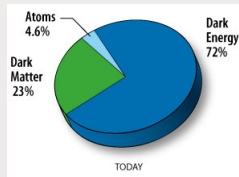
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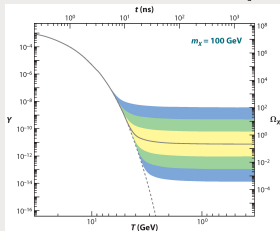
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Phenomenology of neutralinos

- Properties of $\tilde{\chi}_1^0$ depend on its nature

$$\begin{pmatrix} \tilde{\chi}_1^0 \\ \tilde{\chi}_2^0 \\ \tilde{\chi}_3^0 \\ \tilde{\chi}_4^0 \end{pmatrix} = Z_n \begin{pmatrix} -i\tilde{B} \\ -i\tilde{W}^3 \\ \tilde{h}_1^0 \\ \tilde{h}_2^0 \end{pmatrix}$$

- Depending on $|Z_{n11}|^2$, $|Z_{n12}|^2$ and $|Z_{n13}|^2 + |Z_{n14}|^2$
 - Bino case $M_1 \ll M_2, |\mu|$
 - Wino case $M_2 \ll M_1, |\mu|$
 - Higgsino case $|\mu| \ll M_1, M_2$

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- Bino case $M_1 \ll M_2, |\mu|$ $\frac{\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \bar{f}f}{\Omega h^2 > 1}$
- Wino case $M_2 \ll M_1, |\mu|$ $\frac{\tilde{\chi}_1^0 \tilde{\chi}_1^+ \rightarrow ff'}{\Omega h^2 \ll 1}$
- Higgsino case $|\mu| \ll M_1, M_2$ $\frac{\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow VV}{\Omega h^2 \ll 1}$

One-loop computations

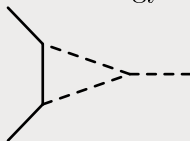
- One-loop diagrams



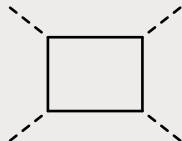
Self energy



Field renormalisation



Vertices



Boxes

- Regularisation : $\mathcal{A} = \mathcal{A}_0 + \frac{1}{\epsilon} \mathcal{A}_1 + \dots$
 - Dimensional Regularisation : Divergence is $1/\epsilon$ in the limit $\epsilon \rightarrow 0$
- Renormalisation : $g = g_0 + \frac{1}{\epsilon} g_1 + \dots$

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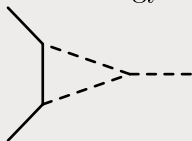
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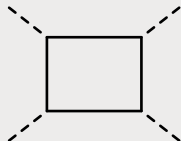
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The renormalisation scheme

The OS scheme : $\delta\mathcal{A} = 0$

e.g. for a mass, the sum of self-energies and counterterms vanishes.

- δM_1 usually renormalised on $m_{\tilde{\chi}_1^0}$
 - But if $|Z_{n11}|$ is small, it becomes ill-suited and δM_1 is large.
 - “Bino-like” scheme : δM_1 is renormalised on the most bino-like of the four neutralinos.
- t_β can be renormalised in different ways
 - m_H
 - $A^0 Z$ transition (DCPR scheme)
 - \overline{DR}
 - $A^0 \rightarrow \bar{\tau}\tau$
- Accuracy check : we will change the renormalisation scheme to assess the precision of the full one-loop result.

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Using an effective approach versus full one-loop computation

- Full one-loop computation nearly not used at all!
 - Complications (need for a renormalisation scheme, no fully automated code)
 - Computing time issue :
 $\tau_{\text{one-loop}}/\tau_{\text{tree-level}} \sim 10^2 - 10^4$
 - Supersymmetric studies based on ~ 100 processes
 - For each process, 1000 one-loop diagrams
- Effective approach : tree-level computation with effective operators
 - $\tilde{\chi}_1^0 \tilde{f} f$
 - $\tilde{\chi}_1^0 \tilde{\chi}_1^0 Z$
 - $Z \tilde{f} f$

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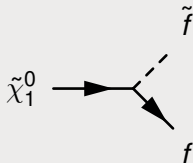
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Determination of effective coefficients

- The $\tilde{\chi}_1^0 \tilde{f} f$ vertex



- Considering **only fermions and sfermions**, there are no loops involved
- Hence the one-loop coupling $\tilde{\chi}_1^0 \tilde{f} f$ is physical (hence finite)

$$(g_2 Y_f, g_1 \tau_f^3, y_{1f}, y_{2f}) Z_n$$

↓

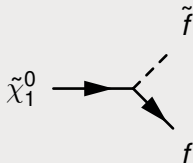
$$((g_2 + \delta g_2 Y_f, (g_1 + \delta g_1) \tau_f^3, y_{1f} + \delta y_{1f}, y_{2f} + \delta y_{2f}) Z_n (1 + \delta Z_n))$$

- Requires that all δX quantities are computed with only fermions/sfermions running in the loop.

Hollik et al. (JHEP (2002))

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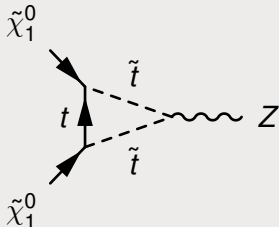
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Determination of effective coefficients : II

- The $\tilde{\chi}_1^0 \tilde{\chi}_1^0 Z$ vertex

There is now a genuine
 triangle loop with
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- The effective vertex reads then

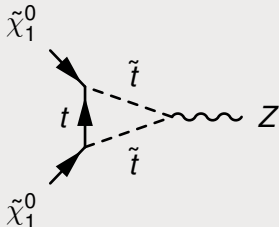
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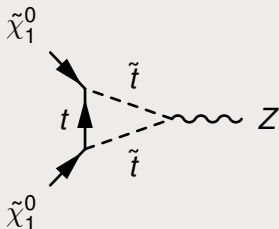
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- Parameters :

$M_{A^0} = 1 \text{ TeV}$	$t_\beta = 4$
$M_{\tilde{\ell}} = 500 \text{ GeV}$	$M_{\tilde{q}} = 800 \text{ GeV}$
$A_f = 0$	

- Process

$$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \mu^+ \mu^- \quad \text{at } v = 0.2$$

- We will compare different cross sections

$\Delta_{\text{one-loop}}$: full one-loop

Δ_{eff} : universal effective part

$\Delta_{\alpha(Q)}$: $\alpha(Q)$ running correction

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Analysis

Effectives
approaches within
the MSSM and
Beyond :
Applications to
Higgs physics and
Dark Matter
observables

G.Driu La Rochelle

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Results : Bino Case

$$M_1 = 90 \text{ GeV}, M_2 = 200 \text{ GeV}, \mu = -600 \text{ GeV}$$

$$\Delta_{\text{eff}} = 17.5\% \quad \Delta_{\alpha} = 14.6\% \quad \Delta_{\text{one-loop}} = 19.6\%$$

- Non-decoupling of heavy squarks.
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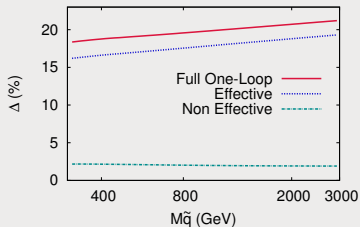
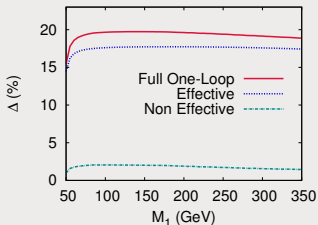
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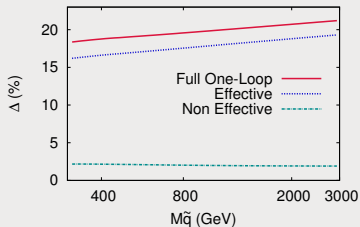
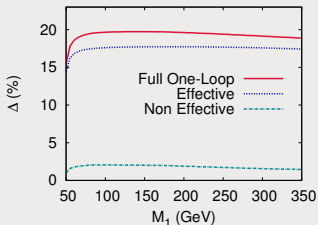
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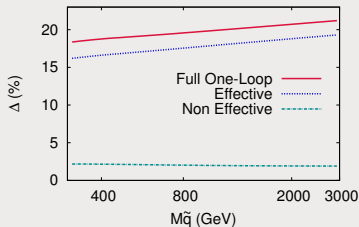
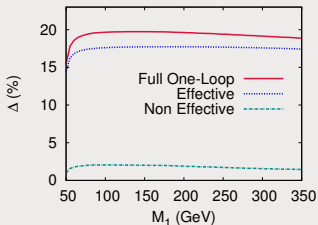
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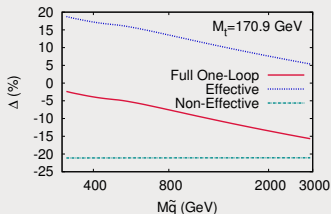
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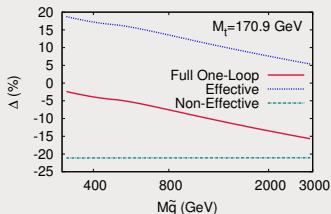
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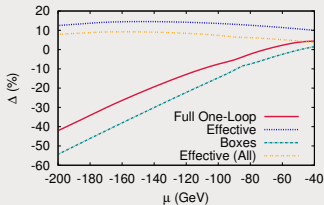
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- Effective computation always faster than full one-loop.
- Better than the naive running of $\alpha(Q)$ (in bino case, 2% agreement)
- Exhibits interesting feature : the squarks non-decoupling contribution (in all cases)
- Is not always precise
 - It lacks the process specific correction (box contribution)
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 - Higgs and Dark Matter
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Naturalness in the Standard Model

- Quantum corrections \rightarrow effective masses and couplings change with energy
 - e.g. the screening of the electric charge
 - Measuring parameters at two different scales yields different values

- Running Higgs mass $m_h(Q)$ with a fermion of mass m_f

$$\frac{dm_h^2(Q)}{d \ln Q} = -\frac{3y^2}{4\pi^2} m_f^2$$

- Assuming complete Standard Model except for new particles appearing at Planck scale Λ_{Planck}
 - $\rightarrow m_h(\Lambda_{\text{Planck}}) \gg m_h$
 - Around 28 orders of magnitude of difference
 - \rightarrow a fine-tuning issue
- Solutions
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- Exact Supersymmetry \rightarrow exact cancellation of fermionic and bosonic loops to m_h
- But Supersymmetry appears broken
 - Fine-tuning now between M_Z and M_{SUSY}
 - Mainly concerns third generation (stops, sbottoms)
- The little hierarchy issue
 - At tree-level $m_h < M_Z$
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 - Having the lightest Higgs heavy ($\gtrsim 125$ GeV) brings back fine-tuning.

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Effective approach again : the bottom-up

- A natural MSSM (i.e. with light stops) bound to cope with a light h
- Such a Higgs boson has very standard-like properties.
- These issues can be relaxed by not restricting to minimal Supersymmetry.

- Non-minimal extensions : NMSSM, U(1)'MSSM

- Effective approach :

$$K = K_{\text{MSSM}} + \frac{1}{M} K^{(1)} + \frac{1}{M^2} K^{(2)} + \dots$$

$$W = K_{\text{MSSM}} + \frac{1}{M} K^{(1)} + \frac{1}{M^2} W^{(2)} + \dots$$

- New set of operators \mathcal{O}_i and coefficients c_i
 - \mathcal{O}_i respect **all symmetries** of the low-energy theory
 - If high-energy completion is weakly coupled,
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Reducing the number of operators

- In order to account for a generic UV completion, one must include **ALL** effective operators.
- Restrictions : truncate at order 2 ($1/M^2$), and include only Higgs superfields

$$O_i = \frac{1}{M^k} O_i(H_1, H_2), \quad k = 1 \text{ or } 2$$

- Fields redefinition and equations of motion

$$\begin{aligned} W_{\text{eff}} &= \zeta_1 \frac{1}{M} (H_1 \cdot H_2)^2 \\ K_{\text{eff}} &= a_1 \frac{1}{M^2} \left(H_1^\dagger e^{V_1} H_1 \right)^2 + a_2 \frac{1}{M^2} \left(H_2^\dagger e^{V_2} H_2 \right)^2 \\ &+ a_3 \frac{1}{M^2} \left(H_1^\dagger e^{V_1} H_1 \right) \left(H_2^\dagger e^{V_2} H_2 \right) + a_4 \frac{1}{M^2} (H_1 \cdot H_2) \left(H_1^\dagger \cdot H_2^\dagger \right) \\ &+ \frac{1}{M^2} \left(H_1 \cdot H_2 + H_1^\dagger \cdot H_2^\dagger \right) \left(a_5 H_1^\dagger e^{V_1} H_1 + a_6 H_2^\dagger e^{V_2} H_2 \right) \end{aligned}$$

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$$\begin{aligned} W_{\text{eff}} &= \zeta_1 \frac{1}{M} (H_1 \cdot H_2)^2 \\ K_{\text{eff}} &= a_1 \frac{1}{M^2} \left(H_1^\dagger e^{V_1} H_1 \right)^2 + a_2 \frac{1}{M^2} \left(H_2^\dagger e^{V_2} H_2 \right)^2 \\ &+ a_3 \frac{1}{M^2} \left(H_1^\dagger e^{V_1} H_1 \right) \left(H_2^\dagger e^{V_2} H_2 \right) + a_4 \frac{1}{M^2} (H_1 \cdot H_2) \left(H_1^\dagger \cdot H_2^\dagger \right) \\ &+ \frac{1}{M^2} \left(H_1 \cdot H_2 + H_1^\dagger \cdot H_2^\dagger \right) \left(a_5 H_1^\dagger e^{V_1} H_1 + a_6 H_2^\dagger e^{V_2} H_2 \right) \end{aligned}$$

Brignole et al. (Nucl.Phys.B (2003)), Antoniadis et al. (Nucl.Phys.B (2009))
Carena et al. (Phys.Rev.D (2009))

Supersymmetry breaking of effective operators

- Allows a Supersymmetry breaking part
 - Replace effective coefficients ζ_1, \mathbf{a}_i by spurions

$$\zeta_1 \longrightarrow \zeta_{10} + \zeta_{11} m_s \theta^2$$

$$\mathbf{a}_i \longrightarrow \mathbf{a}_{i0} + \mathbf{a}_{i1} m_s \theta^2 + \mathbf{a}_{i1}^* m_s \bar{\theta}^2 + \mathbf{a}_{i2} m_s^2 \bar{\theta}^2 \theta^2$$

- We want to keep the effective field theory approximately supersymmetric

$$\frac{m_s}{M} = 0.2 < 1$$

- We have chosen $M = 1.5$ TeV, $m_s = 300$ GeV.
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Deriving operators for fields

- Integration over the Grassmann variables \Rightarrow automated with `lanHEP`

- $\mathcal{L} = \mathcal{L} \left(W, K, \partial_i W, \partial_{ij} W, \partial_{\bar{i}\bar{j}} K, \dots \right)$

- Not all field operators are non-renormalisable!

- e.g. $\left(H_1^\dagger e_1^V H_1 \right)^2 \rightarrow c_{\text{eff}} \frac{v^2}{M^2} Z_\mu Z^\mu$
 \rightarrow Correction to M_Z

- First observable consequence : the EWPT

$$\delta\epsilon_1 = 4e^2 \frac{M_W^2(M_W^2 - M_Z^2)}{M_Z^2 M^2} \left(a_{10} t_\beta^{-4} - a_{30} t_\beta^{-2} + a_{20} \right) = \frac{1}{2} \delta\epsilon_2 = \delta\epsilon_3$$

- $\epsilon_1, \epsilon_2, \epsilon_3 = S, T, U$ reduce the effective parameter space.

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Deriving operators for fields (II)

- Many effects are **simple shifts** of tree-level couplings
 - Suppression factors $\frac{\mu}{M}, \frac{v}{M}, \frac{m_s}{M} \sim 0.2$
- Naively, new Lorentz structures arise with Higgs-only vertices
 - Three scalars : $\partial_\mu h \partial^\mu H^+ H^- \in \mathcal{L}$
 - Higgsino-Higgsino-Higgs : $h \not{\partial} \tilde{\chi}^+ \tilde{\chi}^- \in \mathcal{L}$
- Applying **equations of motion**
 - MSSM Lorentz structure
- This is crucial on loop-induced processes :

$$h/H/A^0 \rightarrow \gamma\gamma, B_s \rightarrow \mu^+ \mu^-, B \rightarrow X_s \gamma^*$$

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- We take $\frac{m_s}{M} = 0.2 \rightarrow$ new physics approximatively supersymmetric
 - $M = 1.5$ TeV, $m_s = 300$ GeV
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MSSM Parameter Space

- $M_2 = 300 \text{ GeV}$, $M_1 = \frac{5}{3} \tan^2 \theta_W M_2 \simeq M_2/2$,
 $M_3 = 800 \text{ GeV}$.
- $M_{\tilde{f}} = 1 \text{ TeV}$, $A_f = 0$, **except** for third generation

- m_h max case :

$$M_{U3} = M_{Q3} = M_{d3} = 1 \text{ TeV}, A_t = A_b = 2 \times 1 \text{ TeV} + \frac{\mu}{t_\beta}$$

- A): Light degenerate stops $M_{U3} = M_{Q3} = M_{d3} = 400 \text{ GeV}$, $A_t = A_b = 0$

- B): Light stops mass separated and maximal mixing
 $m_{\tilde{t}_1} = 200 \text{ GeV}$, $m_{\tilde{t}_2} \in [300, 800] \text{ (GeV)}$ and
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- Higgs sector

$$t_\beta \in [2, 40], \quad M_{A^0} \in [50, 450] \text{ GeV}$$

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Increase of the lightest Higgs mass

- MSSM : $m_h < 135$ GeV
→ BMSSM : $m_h < 250$ GeV.

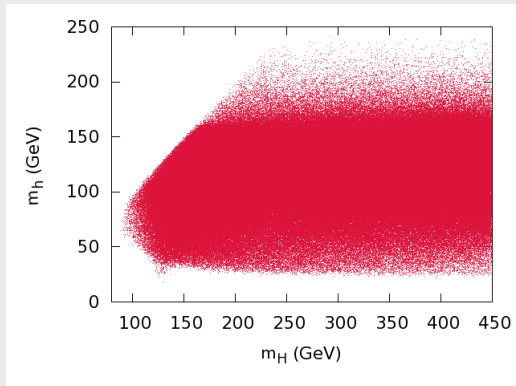


Figure: BMSSM reach in m_H, m_h plane without experimental constraints

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Tree-level Higgs couplings

Z, W	similar features as MSSM	$g_{hVV}^2 + g_{HVV}^2 \simeq 1$
u, c, t	similar features as MSSM	
d, s, b, τ	t_β effect + Slow-decoupling	$\left \frac{g_{h\bar{b}b}}{g_{h\bar{b}b}^{\text{SM}}} \right \in [0, 10]$

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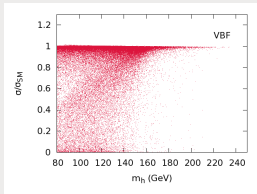
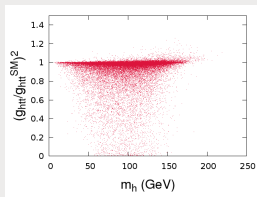
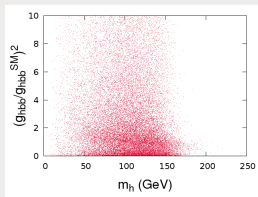
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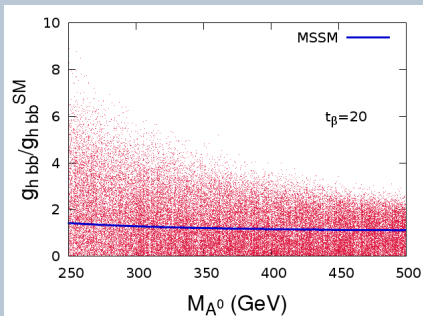
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Slow decoupling of h at high M_{A^0}



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Loop induced Higgs couplings

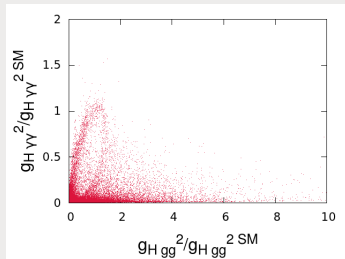
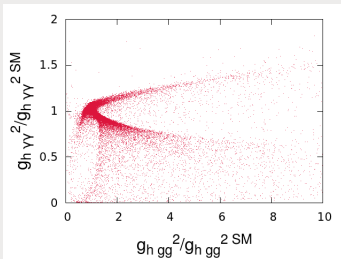
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Relating couplings to observables

- The couplings are not directly accessible
- For instance $gg \rightarrow h \rightarrow \gamma\gamma$ is sensitive to
 - ① Coupling $g_{h\bar{b}b} \Rightarrow \Gamma_h \Rightarrow BR(h \rightarrow \gamma\gamma)$
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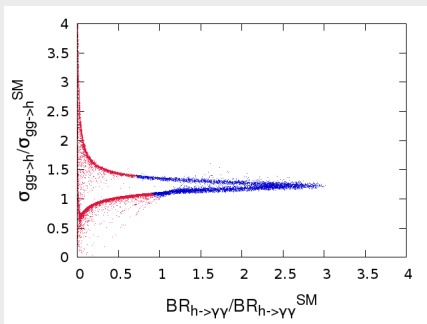
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Technique : using tools for phenomenology

- The field lagrangian is complicated, especially the scalar part

- F-terms from $V_F = -\bar{W}_i W_i$ in MSSM to

$$V_F = - \left(\bar{W}^{\bar{j}} + \frac{1}{2} K_{\bar{i}kl} \psi^k \psi^l \right) K_{ij} \left(W^j + \frac{1}{2} K_{jkl} \bar{\psi}^{\bar{k}} \bar{\psi}^{\bar{l}} \right)$$

- An automated treatment is more suited

→ **lanHEP**

- Creation of a **new version** with higher-order derivatives of K and W .
- It is enough to specify K , W and the truncation order in $\frac{1}{M}$ to obtain all Feynman rules
- Tools for observables
 - CalcHEP for processes without loop corrections
 - **Modified version** of HDecay for the remaining part

$$h \rightarrow \gamma\gamma, h \rightarrow gg, h \rightarrow \bar{b}b \dots$$

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- Creation of a **new version** with higher-order derivatives of K and W .
- It is enough to specify K , W and the truncation order in $\frac{1}{M}$ to obtain **all Feynman rules**
- Tools for observables
 - CalcHEP for processes without loop corrections
 - **Modified version** of HDecay for the remaining part

$$h \rightarrow \gamma\gamma, h \rightarrow gg, h \rightarrow \bar{b}b \dots$$

Technique : using tools for phenomenology

- The field lagrangian is complicated, especially the scalar part

- F-terms from $V_F = -\bar{W}_i W_i$ in MSSM to

$$V_F = - \left(\bar{W}^{\bar{j}} + \frac{1}{2} K_{\bar{i}kl} \psi^k \psi^l \right) K_{ij} \left(W^j + \frac{1}{2} K_{jkl} \bar{\psi}^{\bar{k}} \bar{\psi}^{\bar{l}} \right)$$

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Experimental constraints

What experimental constraints will be used

- Electroweak Precision Test (EWPT)
 - Flavour Physics : $B_s \rightarrow \mu^+ \mu^-$ and $B \rightarrow X_s \gamma^*$
 - Muon anomalous magnetic moment $g_\mu - 2$
 - Dark Matter constraints (Relic density and direct detection)
 - Higgs exclusion limits on $h/H/A^0$ and H^\pm :
LEP/Tevatron/LHC
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- The direct searches for superpartners have been avoided by using heavy quarks (~ 1 TeV), except stops
 - For neutral Higgses, we combine all available channels and test each Higgs separately.

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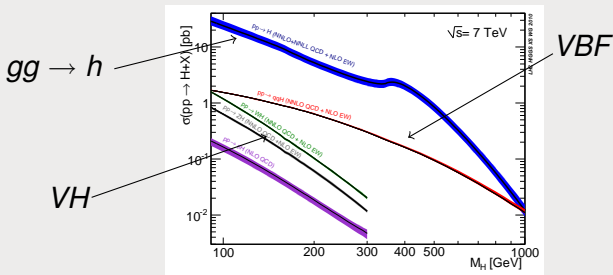
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The Standard Model Higgs search : production

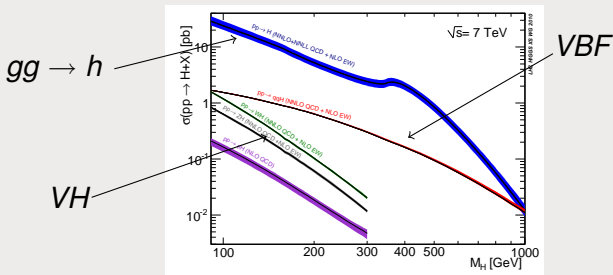


- Dominated by gluon fusion
- But VBF or VH can be distinguished
 - g_{hVV} is less flexible than g_{hgg}
- The ratio of production modes can probe the SM

$$\frac{\sigma_{gg \rightarrow h \rightarrow XX}}{\sigma_{VBF \rightarrow h \rightarrow XX}}$$

- Hints of non-standard behaviour

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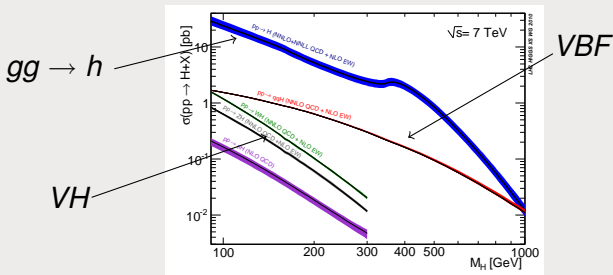


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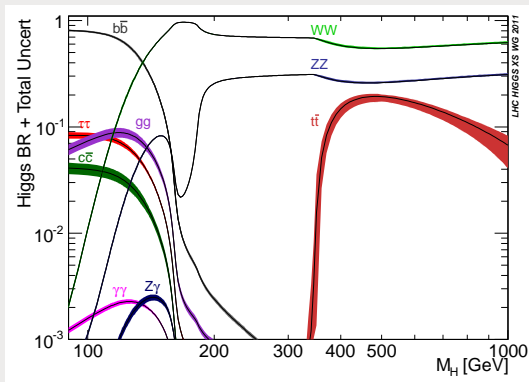
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The Standard Model Higgs search : channels



The Standard Model Higgs search : channels

- $\gamma\gamma$: low masses
 - can probe VBF with $\gamma\gamma + 2j$
- ZZ : different subchannels ($4\ell, 2\ell 2\nu \dots$)
Good sensitivity even below the threshold
- WW : similar features
 - can probe VH, VBF with $WW + 1/2j$
- $\bar{\tau}\tau$: whole mass range
 - MSSM-like analysis since this channel is t_β enhanced
- $VH \rightarrow V\bar{b}b$: low masses
 - Independent of $gg \rightarrow H$!

Recasting to non-SM models

F.Boudjema,GDLR (PhysRevD 2011)

- Use the exclusion ratio in the no signal case

$$R_{XX}^{\text{excl } 95\%} = \frac{\sigma_{pp \rightarrow \phi \rightarrow XX}}{\sigma_{pp \rightarrow \phi \rightarrow XX}^{\text{excl } 95\%}}$$

- Use signal strength in the case of a signal

$$R_{XX} = \frac{\sigma_{pp \rightarrow \phi \rightarrow XX}}{\sigma_{pp \rightarrow H \rightarrow XX}^{\text{SM}}}$$

- Use the MSSM production modes

$$\sigma_{pp \rightarrow \phi \rightarrow XX} = (\sigma_{ggh} + \sigma_{VBF} + \sigma_{Vh} + \sigma_{\bar{b}bh}) \times BR(\phi \rightarrow XX)$$

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Recasting Issue 1 : the efficiencies

- On a single count experiment we have a priori

$$\mu = \frac{n_S}{n_S^{\text{SM}}} \neq \frac{\sigma_{\text{inclusive}}}{\sigma_{\text{inclusive}}^{\text{SM}}}$$

- Analyses impose cuts on the final phase space
 - If the differential distributions $(p_T, \eta, m_{XX}, \dots)$ are not proportional to SM, the events yields will vary
 - Typically those distributions depend on the production mode gg, VBF, \dots

- μ is then obtained by the exclusive cross-section :

$$\mu = \frac{\sigma^{\text{exclusive}}}{\sigma_{\text{SM}}^{\text{exclusive}}}, \quad \sigma^{\text{exclusive}} = \sum_i \epsilon_i \sigma_i \times BR$$

- σ_i : inclusive production cross-sections
- ϵ_i : **efficiencies**
- Concerned channels

$$h \rightarrow \gamma\gamma + 0/2j \quad h \rightarrow WW + 0/1/2j \quad h \rightarrow \bar{\tau}\tau$$

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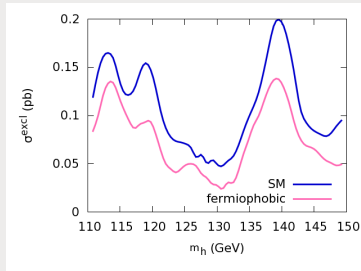
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Recasting Issue 1 : Example

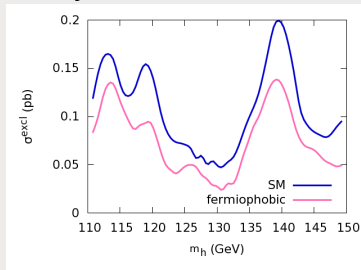
- Limit on $R_{\text{Fermiophobic}}^{\text{excl } 95\%}$ versus $R_{\text{SM}}^{\text{excl } 95\%}$



- Efficiencies rarely quoted for $m_h = 120$ GeV, and final state $\gamma\gamma + 2j$
 $\epsilon_{ggh} = 0.005$, $\epsilon_{vbf} = 0.15$
 - Not quoted on the public note nor the paper!
- Ratios can be estimated on a Monte-Carlo basis
 $WW + 0/1j$, $m_h < 160$ GeV, $\sim 10 - 20\%$ accuracy
- Though they show a limited impact on searches, they become relevant in the case of a signal

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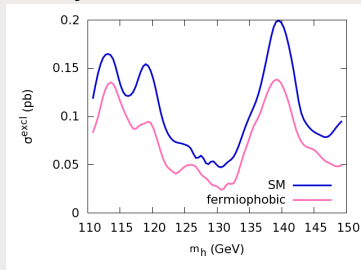
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- Statistical combinations of n channels assume identical $R_{X_1 X_1} = R_{X_2 X_2} = \dots = R_{X_n X_n} = \mu$
 - i.e. the signal is assumed to be exactly SM-like
 - It induces a bias on $\hat{\mu}$ and $R^{\text{excl } 95\%}$

- Naive combination

$$R^{\text{excl } 95\%} = \left(\sum_i R_i^{\text{excl } 95\%} - 2 \right)^{-\frac{1}{2}}$$

- Reconstruct each likelihood L_i from $R_{\text{expected}}^{\text{excl } 95\%}$, $\hat{\mu}$ in the limit $s \ll b$ Azatov et al (arXiv:1202.3415)
- Improvements on the experimental side
 - Statistical combination with different R_i
 - ATLAS analysis for $h \rightarrow \gamma\gamma$: $L = L(R_{ggh}, R_{VWh})$

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Allowed region for neutral Higgses

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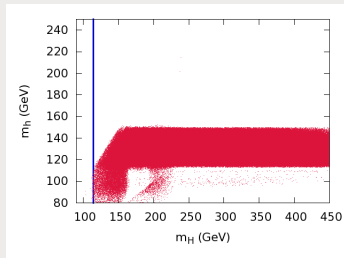
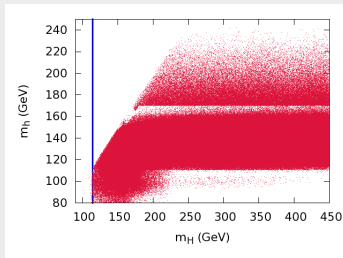
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Conclusion

F.Boudjema,GDLR (PhysRevD 2011)



- Lightest Higgs h restricted to be light $m_h < 150$ GeV
- Cases with $m_h < 114$ GeV
 - h is then decoupled to Z, W bosons.
 - H can either be light $m_H < 150$ GeV, or decay as $H \rightarrow hh$
- In particular there are two possibilities $m_h = 125$ GeV and $m_H = 125$ GeV.

- $m_h = 125 - 127$ GeV
- We take $\hat{\mu}$ from the most indicative channels
 - The uncertainties quoted are for 1σ bands.

	ATLAS		CMS	
	2012	2011	2012	2011
$\gamma\gamma$	1.9 ± 0.5	2 ± 0.8	1.4 ± 0.5	1.7 ± 0.8
$\gamma\gamma + 2j$	2.8 ± 1.2		1.6 ± 1.1	3.8 ± 2.1
$ZZ \rightarrow 4\ell$	1.35 ± 0.9	1.2 ± 1.0	0.8	0.5 ± 0.8

- Other channels are dominated by uncertainties

	ATLAS	CMS
$WW \rightarrow l\nu l\nu$		0.5 ± 0.4
$\bar{\tau}\tau$		-0.3
$VH \rightarrow V\bar{b}b$		0.6

- X : Reconstructed values

Higgs signal in $h \rightarrow \gamma\gamma$

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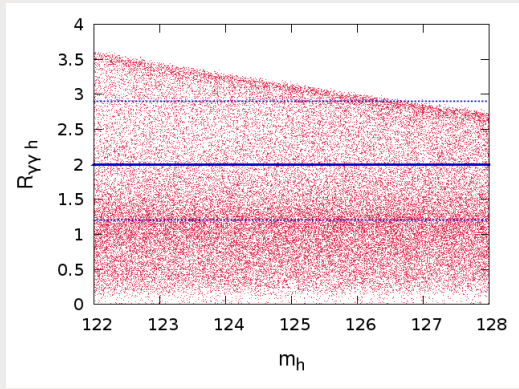
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- $h \rightarrow \gamma\gamma$ can be enhanced :



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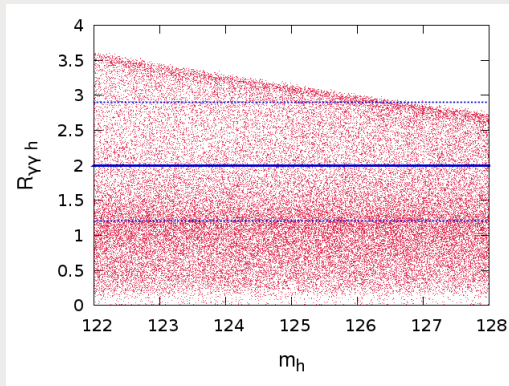
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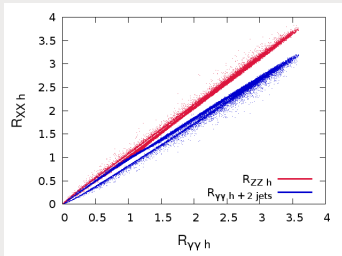
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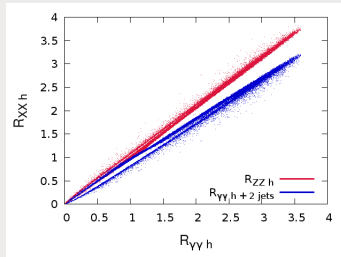
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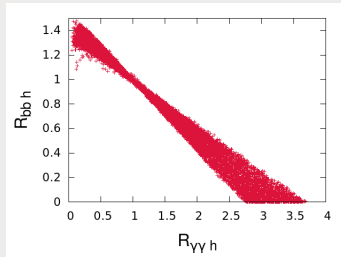
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Effect of stops loops

- Lightest stop loop

$$g_{h\tilde{t}_1\tilde{t}_1} \simeq \frac{g}{M_W} \left(\sin^2(2\theta_{\tilde{t}}) \frac{m_{\tilde{t}_1}^2 - m_{\tilde{t}_2}^2}{4} + m_{\tilde{t}}^2 + O(M_Z^2) \right)$$

- Opposite contributions to $gg \rightarrow h$ and $h \rightarrow \gamma\gamma$.
- Model B) : $\Delta m = 400$ GeV, $s_{2\theta_{\tilde{t}}} = 1$

- $R_{\gamma\gamma}$ is reduced as compared as scenario A) but now there is a hierarchy

$$R_{\gamma\gamma+2j} > R_{\gamma\gamma} > R_{ZZ}$$

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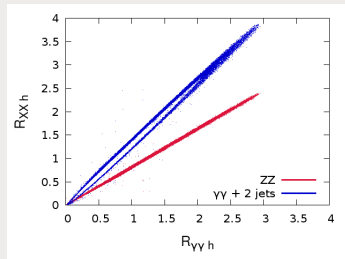
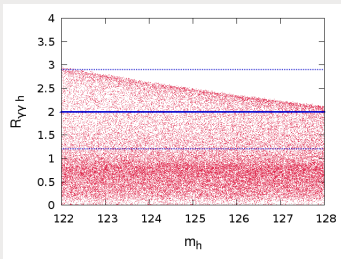
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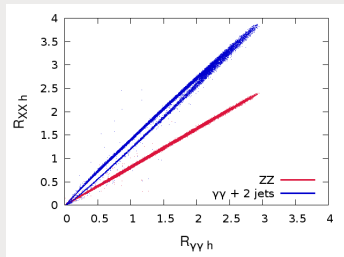
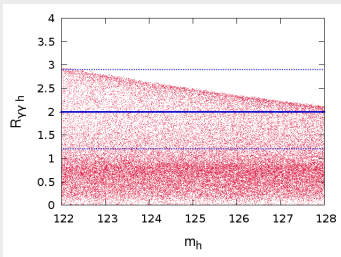
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Effect of stops loops

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- $R_{\gamma\gamma}$ is reduced as compared as scenario **A**) but now there is a hierarchy

$$R_{\gamma\gamma+2j} > R_{\gamma\gamma} > R_{ZZ}$$

Flavour Constraints

F.Boudjema,GDLR (to be published)

- Supersymmetry typically allows for extra contributions
 $B_s \rightarrow \mu^+ \mu^-$ and $B \rightarrow X_S \gamma^*$
- New contributions from the BMSSM (hH^+H^- ,
 $A^0 \tilde{\chi}_i^+ \tilde{\chi}_j^-$)
- First consequence : $M_{A^0} < 200$ GeV excluded for all
 $t_\beta \in [2, 40]$
- Specific for Model B)
 - Because of the stop-chargino loop, t_β is restricted to be small (< 5)

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G.Drieu La Rochelle

Introduction

The Standard
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 Higgs and Dark
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Effective approach
 for relic density
 SUSY at one-loop
 Effective vertices
 Results

Effective approach
 for Higgs physics

The BMSSM
 Computing
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Masses and
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Recasting SM Higgs
 searches

The Higgs signal

Conclusion

Flavour Constraints

F.Boudjema,GDLR (to be published)

- Supersymmetry typically allows for extra contributions
 $B_s \rightarrow \mu^+ \mu^-$ and $B \rightarrow X_S \gamma^*$
- New contributions from the BMSSM (hH^+H^- ,
 $A^0 \tilde{\chi}_i^+ \tilde{\chi}_j^-$)
- First consequence : $M_{A^0} < 200$ GeV excluded for all
 $t_\beta \in [2, 40]$
- Specific for Model B)
 - Because of the stop-chargino loop, t_β is restricted to be small (< 5)

Effectives
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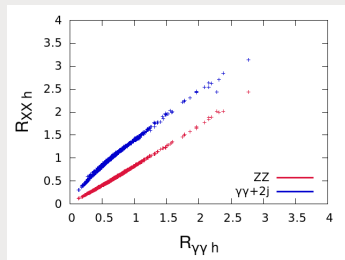
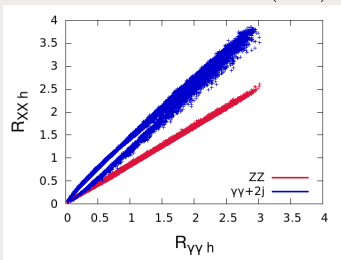
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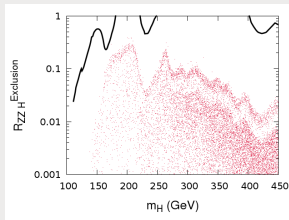
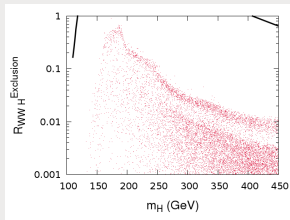
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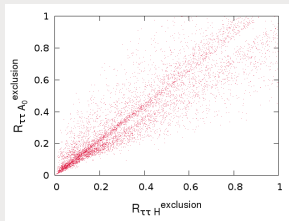
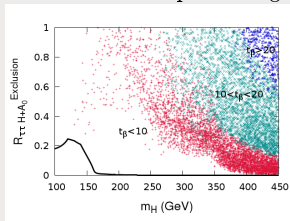
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- Selecting $m_h \in [123, 129] \text{ GeV}$, with $R_{\gamma\gamma} = X$, $R_{ZZ} = X$



- $\bar{\tau}\tau$ channel more promising



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Conclusion

- Power of the effective approach
 - A precious tool for separating an observable scale, i.e. at low energy from a hypothesized scale
 - For Dark Matter the effective scale is the neutralino mass and the heavy one the squark masses.
 - For the Higgs search the effective scale is m_h and the heavy scale the extra physics.
- Conclusions from Higgs and Dark Matter constraints
 - Relic Density precise enough \rightarrow towards an era of precision?
 - The supersymmetric Higgs can be quite non-standard (Correlations between channels)
 - With the data coming in, importance of a model-independent result