

Qu'est-ce que la physique des particules ?

Où va-t-elle ?

Pourquoi ?

Comment ?

Est-ce réellement important ?

Etienne Augé

DAS IN2P3

# des collisions:

- Mode opératoire
  - Peu raffiné
  - Toujours les mêmes ingrédients de départ
  - Inchangé depuis des décennies
- Et pourtant :
  - Accélérateurs et détecteurs toujours plus gros et plus chers
  - Ressources informatiques énormes
- **Ne manquerait-on pas d'imagination ?**

# Physique des particules **élémentaires**

- Élémentaire ?
  - Atome ? Proton ? Quark ? Electron ?...
  - Particule élémentaire ultime ?
- Lois fondamentales ?
  - La découverte des quarks n'a révolutionné ni la Chimie ni la Biologie...
- **Il existe aussi une physique de la complexité**

Et pourtant...

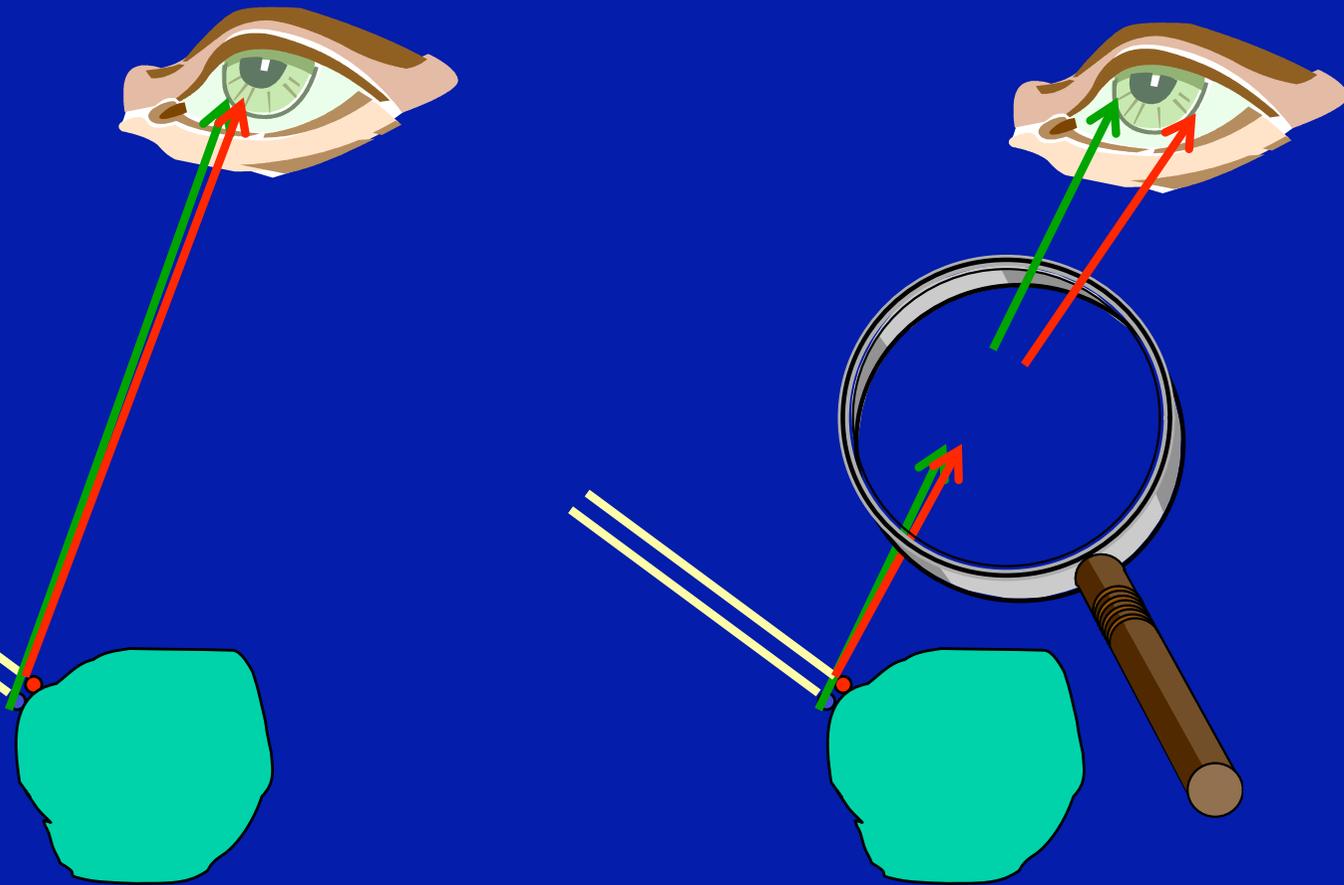
# Une physique « des deux infinis »

- Modèle cosmologique du **big bang**
  - Très grande densité d'énergie aux premiers instants de l'Univers
    - Haute température → Collisions violentes entre constituants de l'Univers
    - Comme ce qu'on provoque dans les laboratoires de physique des particules
- La connaissance des particules permet d'écrire des morceaux du scénario cosmologique
- Le scénario cosmologique permet de contraindre le comportement des particules dans des collisions beaucoup plus violentes que ce qu'on peut provoquer en laboratoire
- **Connaissance de l'univers**

# Collisions violentes : une nécessité

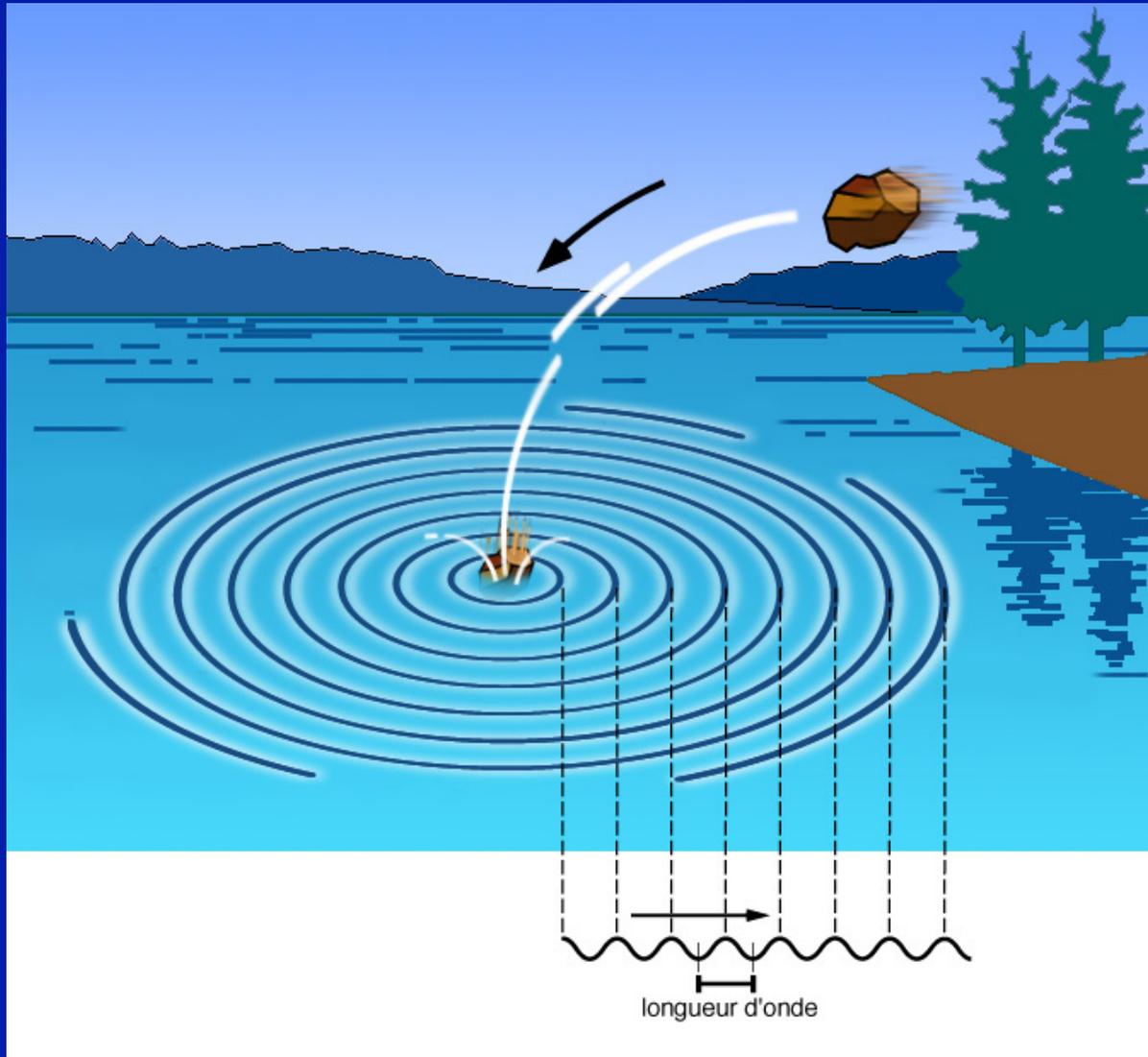
- Pour observer des détails fins (exemple H1)

lumière



Une loupe ou un microscope permettent de voir des détails plus petits que l'œil nu, mais pour des distances inférieures à la longueur d'onde...

# Nature ondulatoire → diffraction



Pouvoir séparateur proportionnel à la longueur d'onde

# Dualité onde-corpuscule

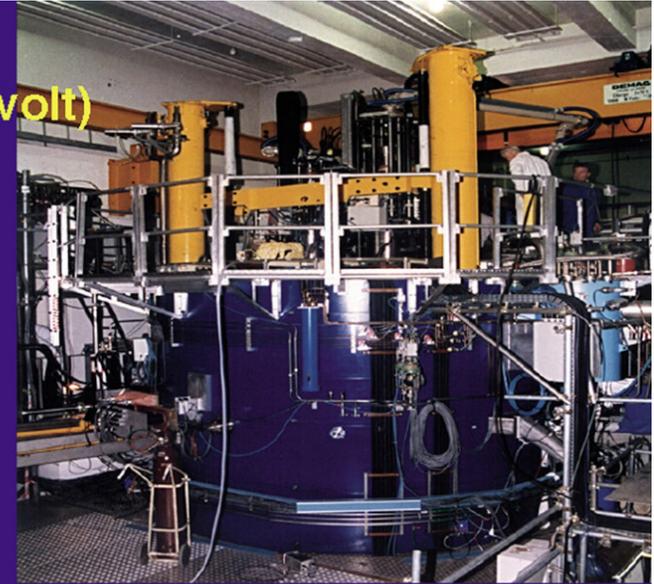
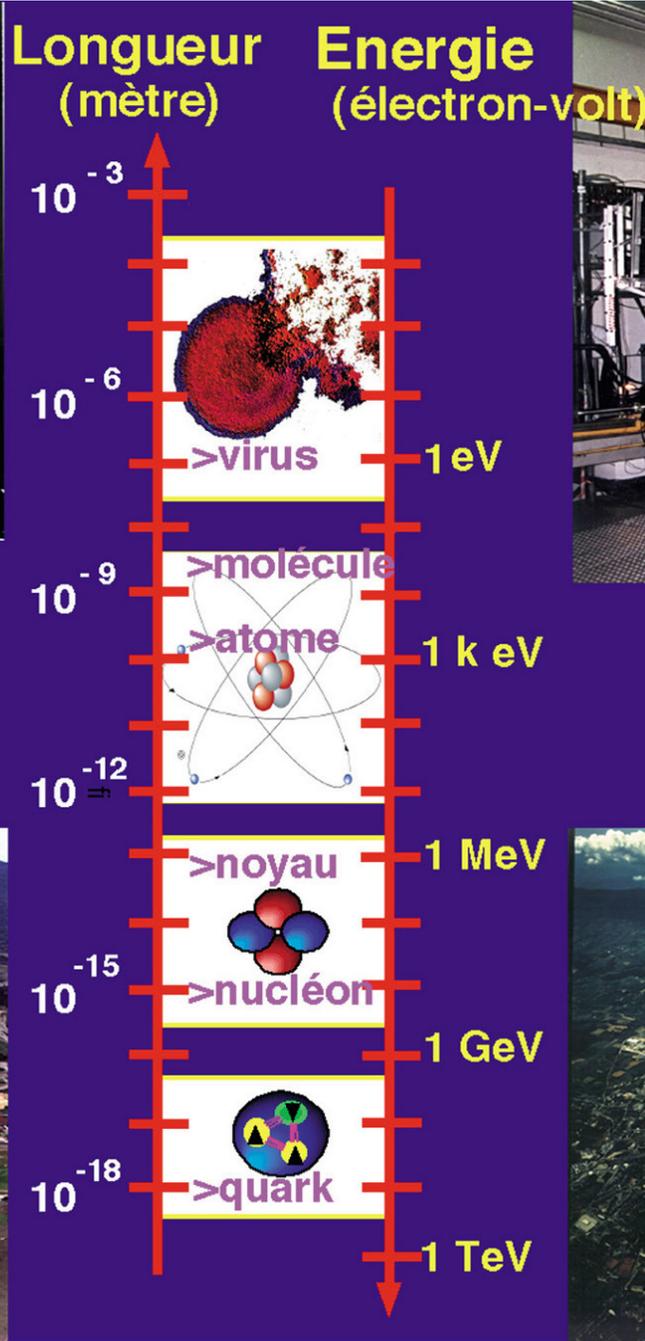
- Onde lumineuse  $\Leftrightarrow$  Faisceau de projectiles
  - Effet photo-électrique (A. Einstein)
  - photon = « grain de lumière »
- Faisceau de projectiles  $\Leftrightarrow$  onde
  - microscopie électronique
  - imagerie par faisceau de neutrons
- Energie d'un corpuscule  $\propto$  inverse de la longueur d'onde :
- finesse des détails  $\Leftrightarrow \lambda$  petit  $\Leftrightarrow E$  grand 
$$E = \frac{hc}{\lambda}$$
- $\Leftrightarrow$  Bombarder l'objet observé avec des projectiles énergiques
- $\Leftrightarrow$  L'objet est cassé par le projectile



Microscope électronique



Accélérateur linéaire



Cyclotron



Synchrotron

# Collisions violentes : une nécessité

- Pour observer des détails fins (exemple H1)
- Pour concentrer de l'énergie en un point du vide
  - 1 TeV (LHC): soulever 1,6 grammes de 0,01 mm
  - Dans un cube de  $10^{-45} \text{ m}^3$
  - Densité  $1,6 \times 10^{38} \text{ Joule/m}^3$  : monstrueux...
- Faire apparaître de nouveaux phénomènes
  - Nouvelles particules élémentaires → classification des particules élémentaires → intuition de nouvelles lois plus fondamentales (classification de Mendeleiev)
- Produit tous les états finals possibles → sélection a posteriori, on-line et off-line

# Les trois frontières

- Frontière de l'énergie  
(LEP → Tevatron → LHC → sLHC → ILC → ????)
- Frontière de l'intensité  
effet indirect des nouvelles particules → déviations/  
prédictions → mesures de précision  
LEP, LHCb → superB
- Frontière cosmique  
Matière noire, énergie noire

# HEPAP (P5) Report

## Major Findings

- **Progress in achieving the goals of particle physics requires advancements at the:**
  - **Energy Frontier**
  - **Intensity (or precision) Frontier**
  - **Cosmic (or particle astrophysics) Frontier**(each provides a unique window for insight about the fundamental forces/particles of nature)
- **LHC offers an outstanding opportunity for discoveries at the Energy Frontier**
  - Resources will be needed to support the extraction of the science by U.S. scientists
  - Resources will be needed for planned accelerator and detector upgrades
- **An opportunity exists for the U.S. to become a world leader at the Intensity Frontier**
  - Central is an intense neutrino beam and large underground long-based line detector
  - Building on infrastructure at Fermilab and partnering with NSF
  - Develops infrastructure that positions the U.S. to regain Energy Frontier (Muon Collider)
- **Promising opportunities for advancing particle physics identified at Cosmic Frontier**
  - Requires partnering with NASA, NSF, etc.
- **HEP at its core is an accelerator based experimental science**
  - Accelerator R&D develops technologies needed by the field and that benefit the nation

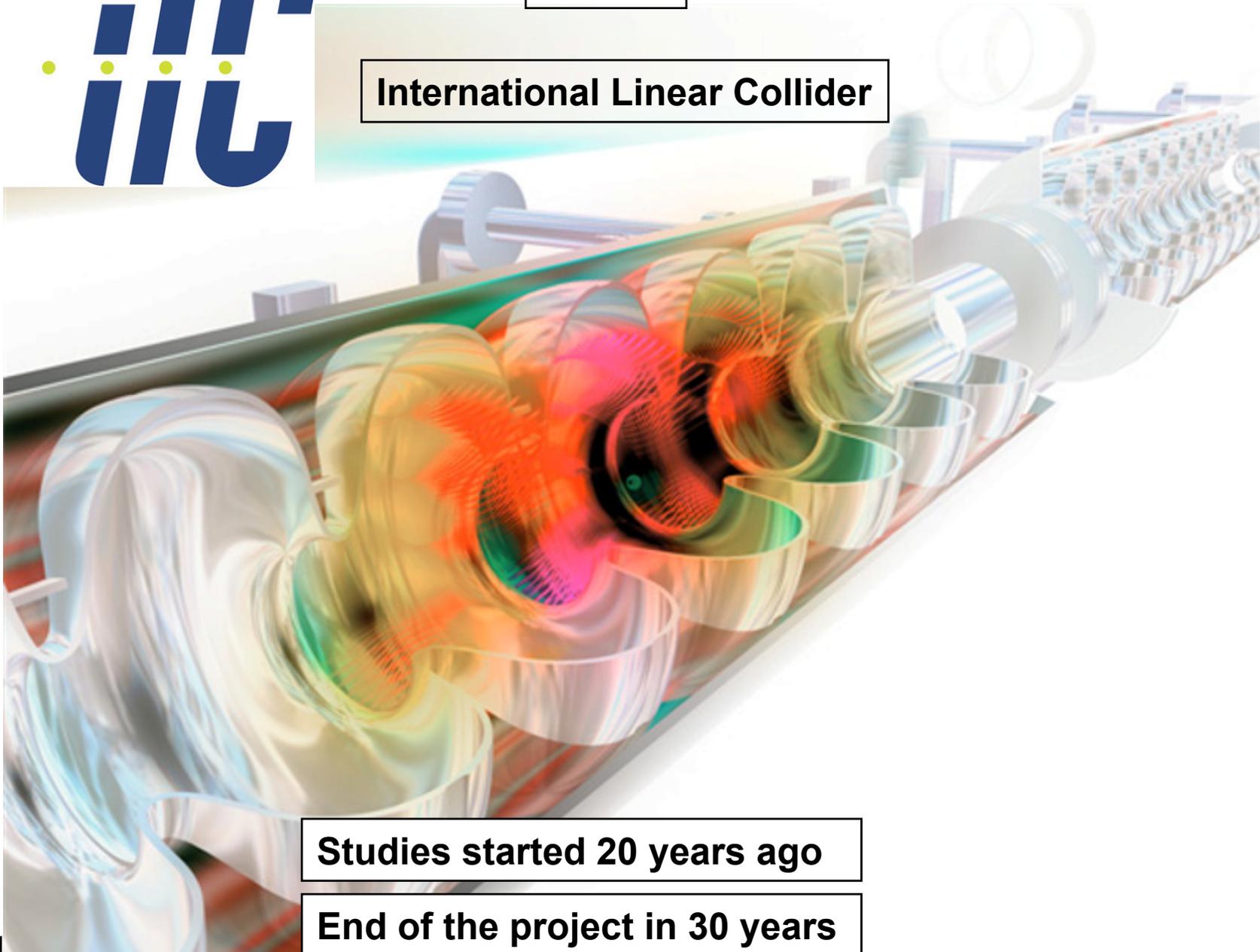
# Un peu de sociologie

- Du résultat d'un doctorant à une recommandation de l'ICFA
- Filtre de la collaboration (working group, publications committee) → publication
- Présentations en conférences, discussions dans la communauté : **très important**
- Emergence de nouveaux projets (processus collectif qui implique aussi les théoriciens)
- Eventuellement validés et soutenus par les tutelles
- Financement et construction planifiés à l'échelle mondiale : exemple ILC



2019 ?

International Linear Collider



Studies started 20 years ago

End of the project in 30 years

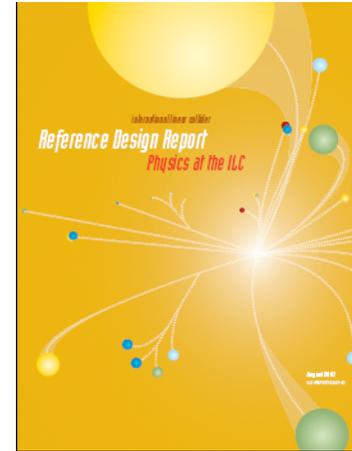
# Organisation ILC

- ILC Steering Committee (E. Iarocci)
- Global Design Effort
  - B. Barish, chef du comité exécutif
- World-Wide Study (physique et détecteurs)
  - S. Yamada, directeur de la recherche
- International Detector Advisory Group
  - Nommé par l'ILCSC, M. Davier, président

- **Reference Design Report (4 volumes)**



Executive  
Summary



Physics  
at the  
ILC



Accelerator

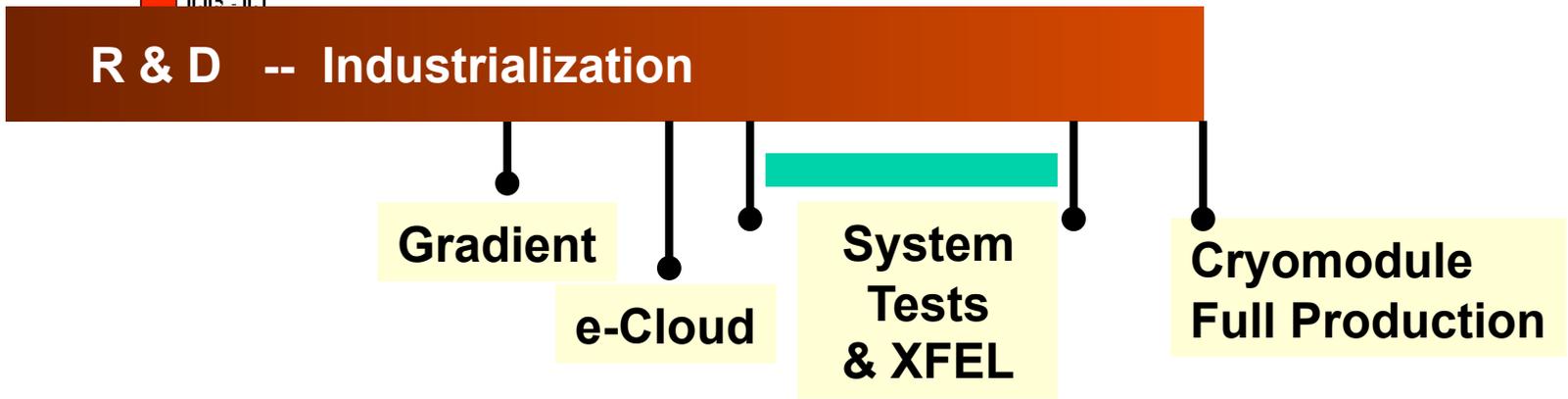
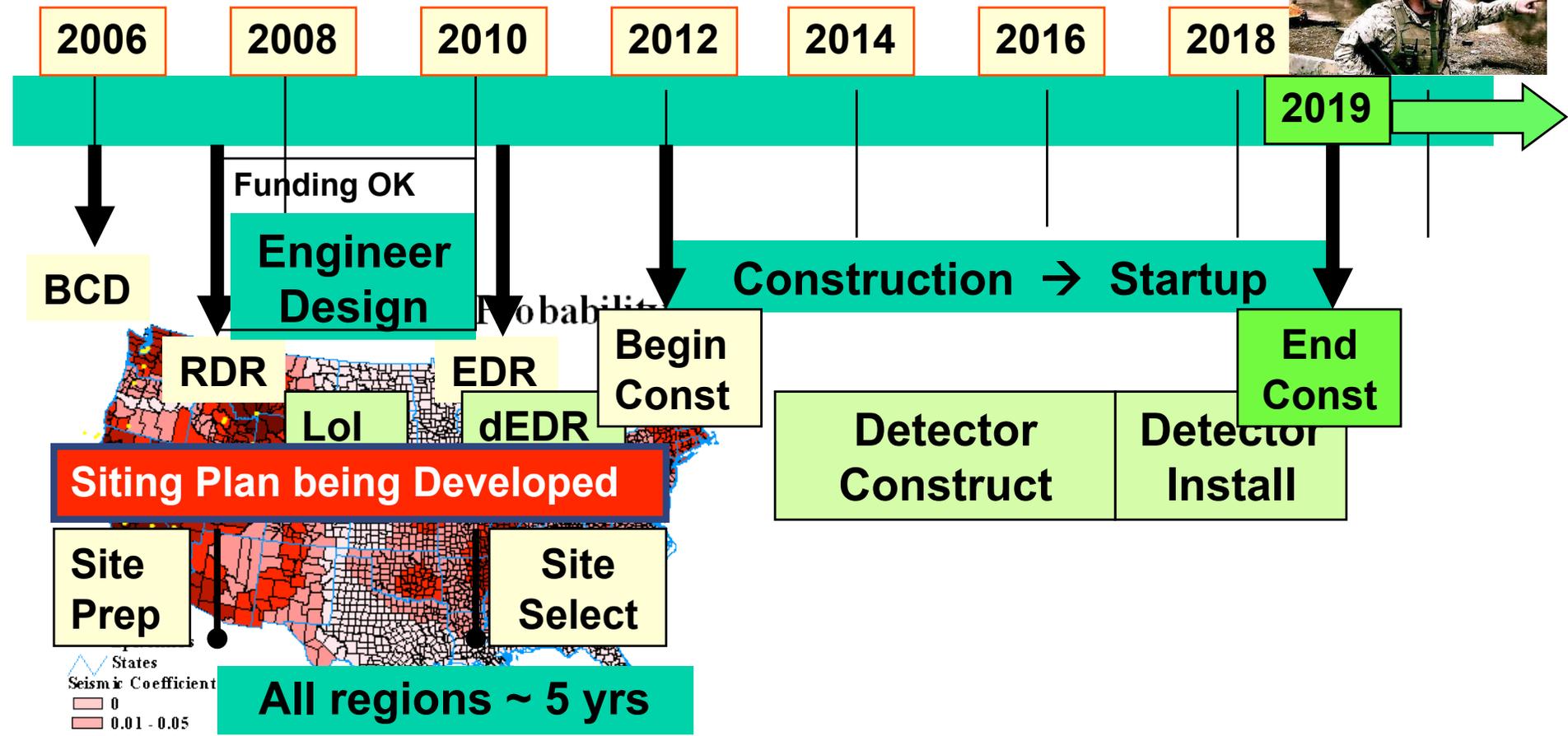


Detectors

LOI attendues en 2009

International Detector Advisory Group : présidé par M. Davier

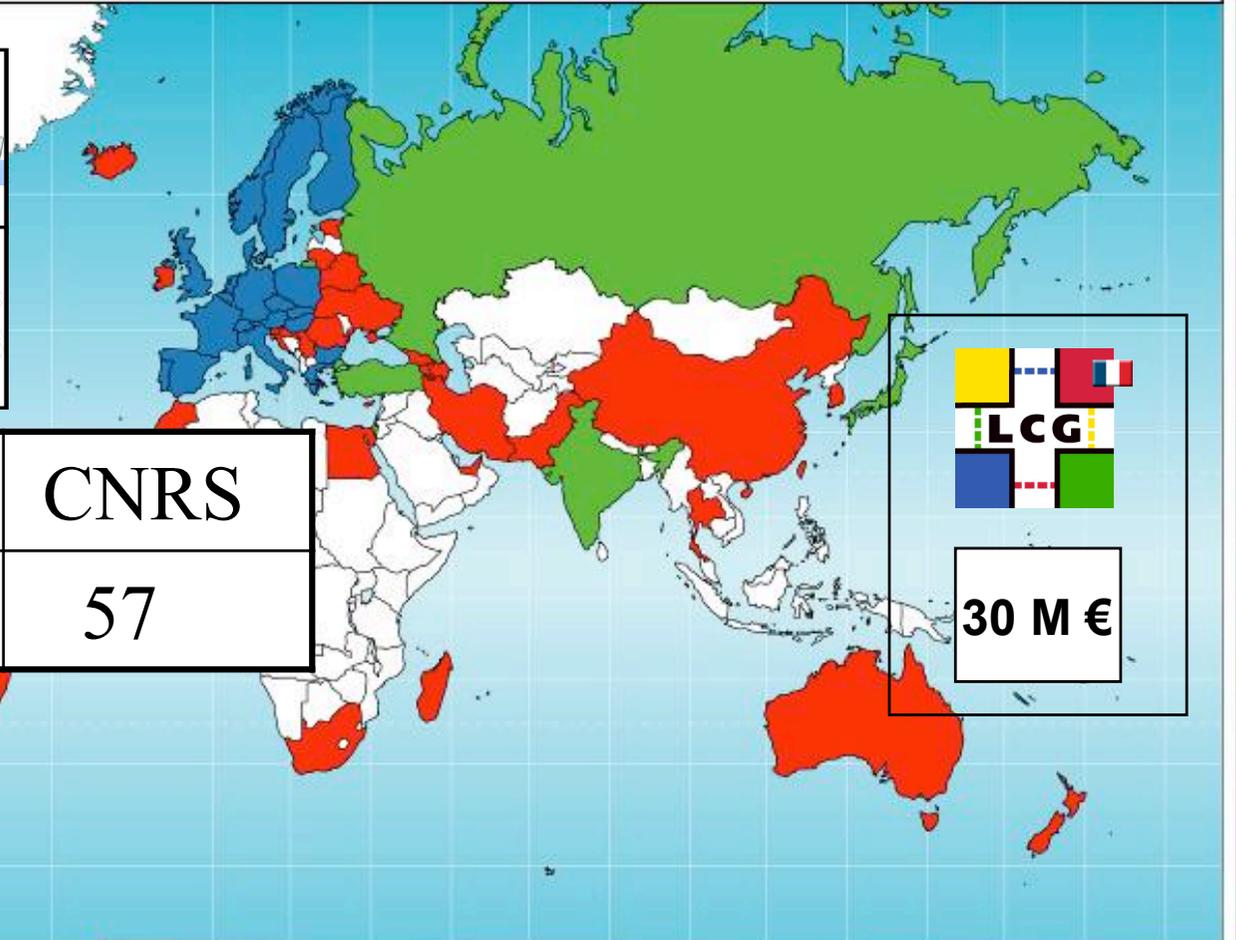
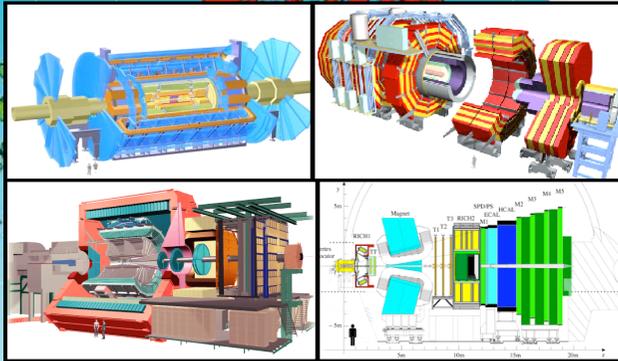
# Technically Driven (non aggressive) Timeline



# Un peu de géopolitique

- ATLAS: construit avec l'argent de 30 pays
  - Coordination des physiciens, des ingénieurs, des administrations → liens forts.
- Opportunité pour que des pays développent leurs relations internationales à travers la Science
- Rôle du CERN – Roadmap ESFRI
- Rôle de l'OCDE (Open Science Forum)
- Réunions FALC

# CERN 6500 users from institutes (+2500 from CERN)



M €	Total	CNRS
MoU	828	57

**30 M €**

**MEMBER STATES**

**4414**

AUSTRIA	NETHERLANDS
BELGIUM	NORWAY
BULGARIA	POLAND
CZECH REPUBLIC	PORTUGAL
DENMARK	SLOVAKIA
FINLAND	SPAIN
FRANCE	SWEDEN
GERMANY	SWITZERLAND
GREECE	UNITED KINGDOM
HUNGARY	
ITALY	

**OBSERVER STATES**

INDIA	54
ISRAEL	39
JAPAN	88
RUSSIA	770
TURKEY	25
USA	737

**1713**

**OTHER STATES**

ARGENTINA	1	CROATIA	18	LITHUANIA	1	SOUTH AFRICA	3
ARMENIA	11	CUBA	2	MADAGASCAR	1	TAIWAN	16
AUSTRALIA	13	CYPRUS	7	MEXICO	13	THAILAND	1
AZERBAIJAN	2	EGYPT	1	MOROCCO	7	UKRAINE	12
BELARUS	16	ESTONIA	9	NEW ZEALAND	3	UNITED ARAB	
BRAZIL	25	GEORGIA	7	PAKISTAN	15	EMITATES	4
CANADA	71	ICELAND	2	ROMANIA	25		
CHINA	46	IRAN	5	SERBIA AND			
		IRELAND	7	MONTENEGRO	12		
		KOREA	10	SLOVENIA	8		

**374**

CERN: where the WeB was born

~2000 physicists/experiment

The physics of the ultra small

2004

LHC  
Large Hadron Collider

needs the ultra large

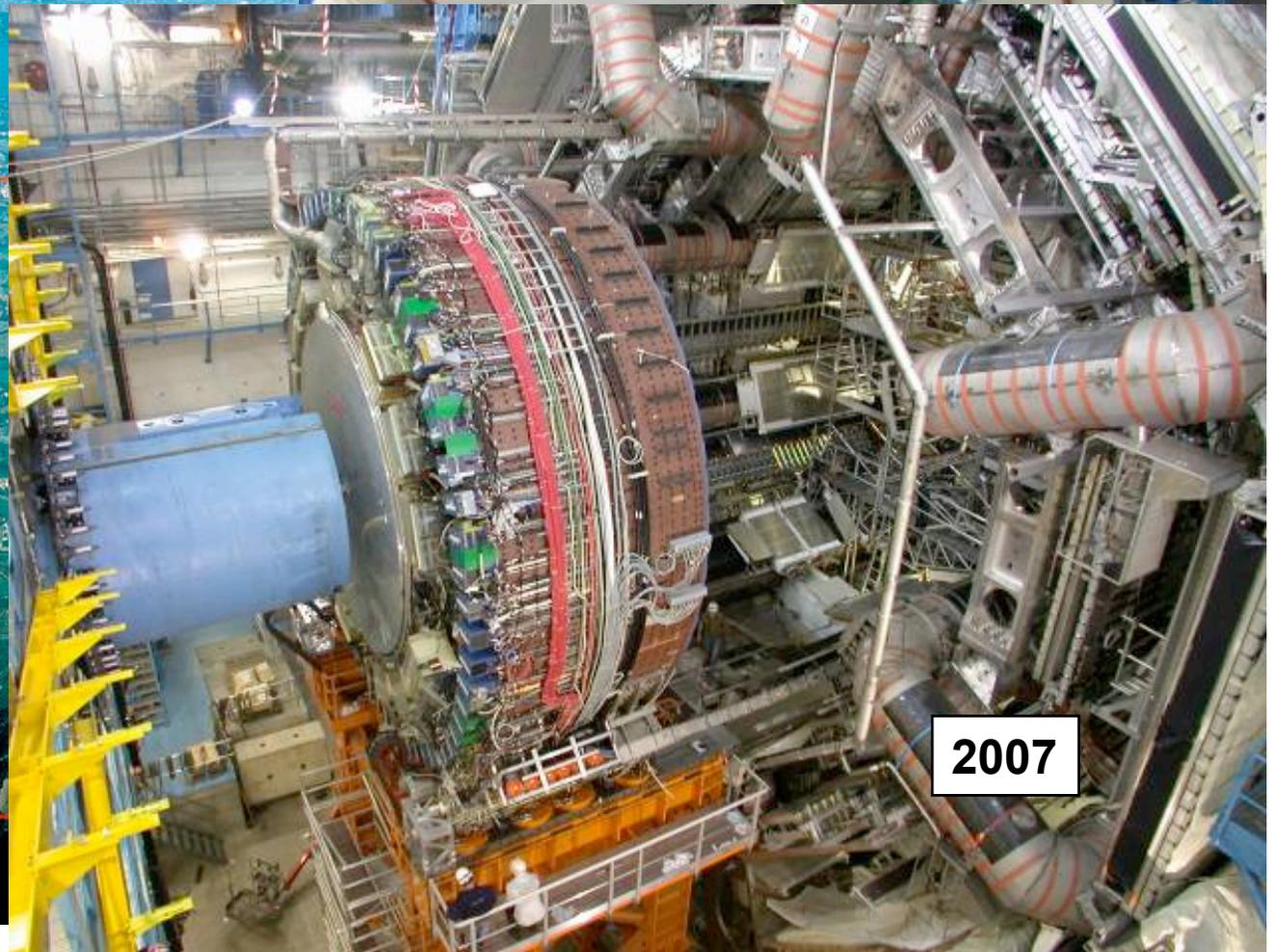
*L. Leman*

27 km Ø

~100 m underground

2° K

*CMS*



2007

# CMS

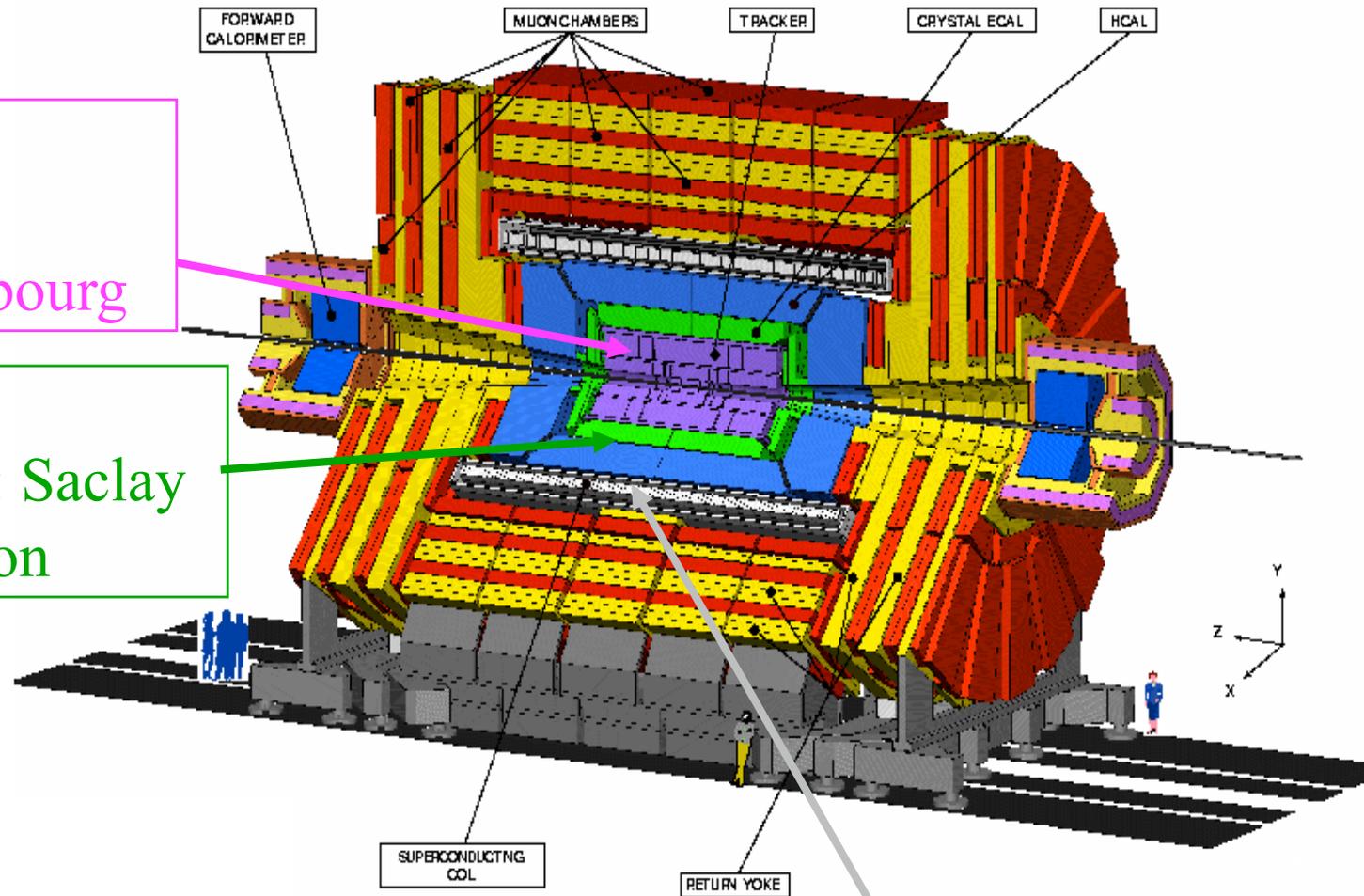
Si Tracker  
End cap  
Lyon-Strasbourg

EM Calo  
Crystals test : Saclay  
Readout : Lyon

Length : 20 m  
Radius : 7 m  
Weight : 14000 tonnes  
Electronic channels :  $10^8$

L1 Trigger  
Electronics  
LLR

Solenoid : Saclay



CMS-PARA-001-11/07/97

JLB.PP

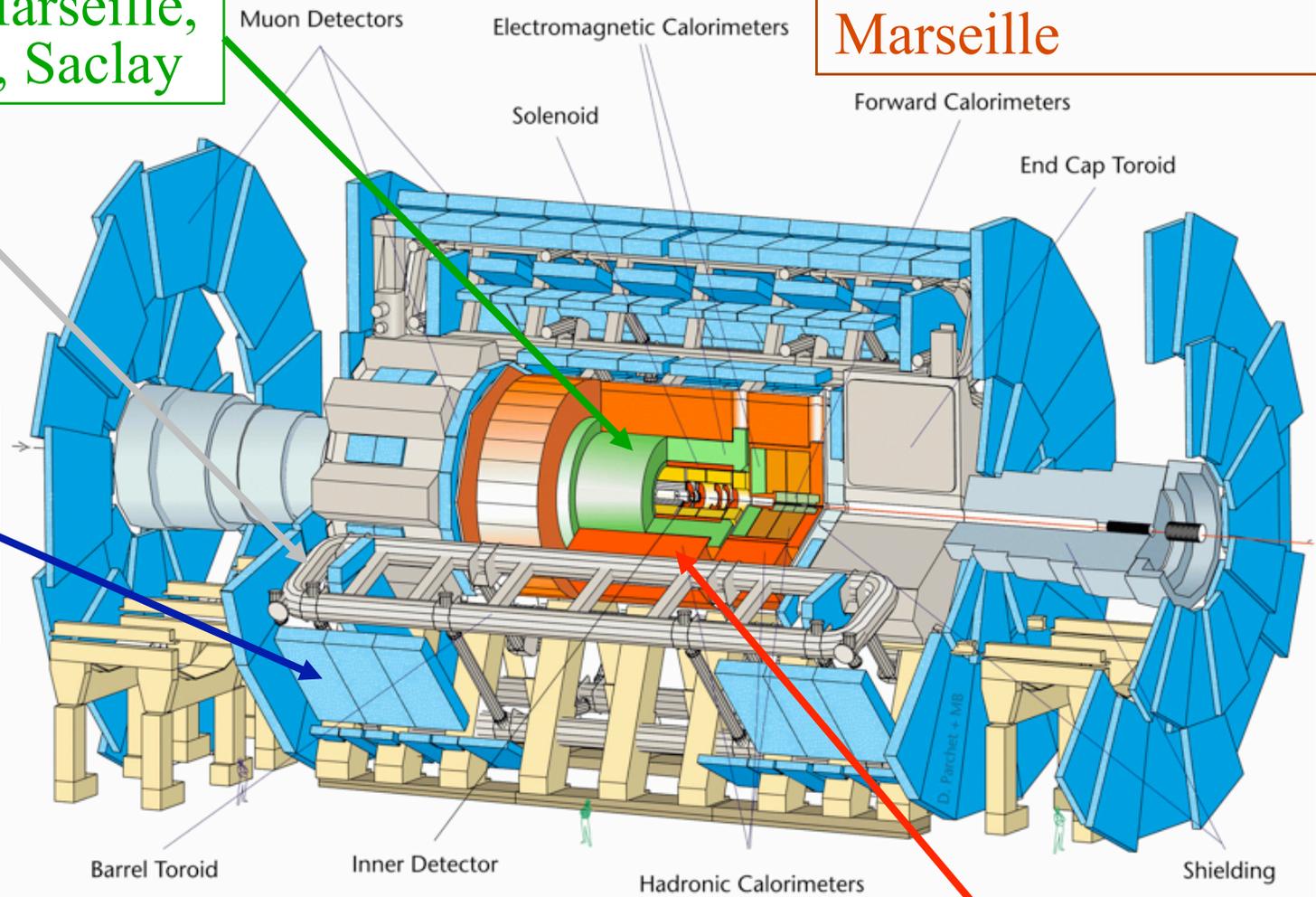
# A T L A S

EmCal : Annecy,  
Grenoble, Marseille,  
Orsay, Paris, Saclay

Pixels and HLT :  
Marseille

Toroids :  
Saclay

Muon  
chambers:  
Saclay



Length : 40 m

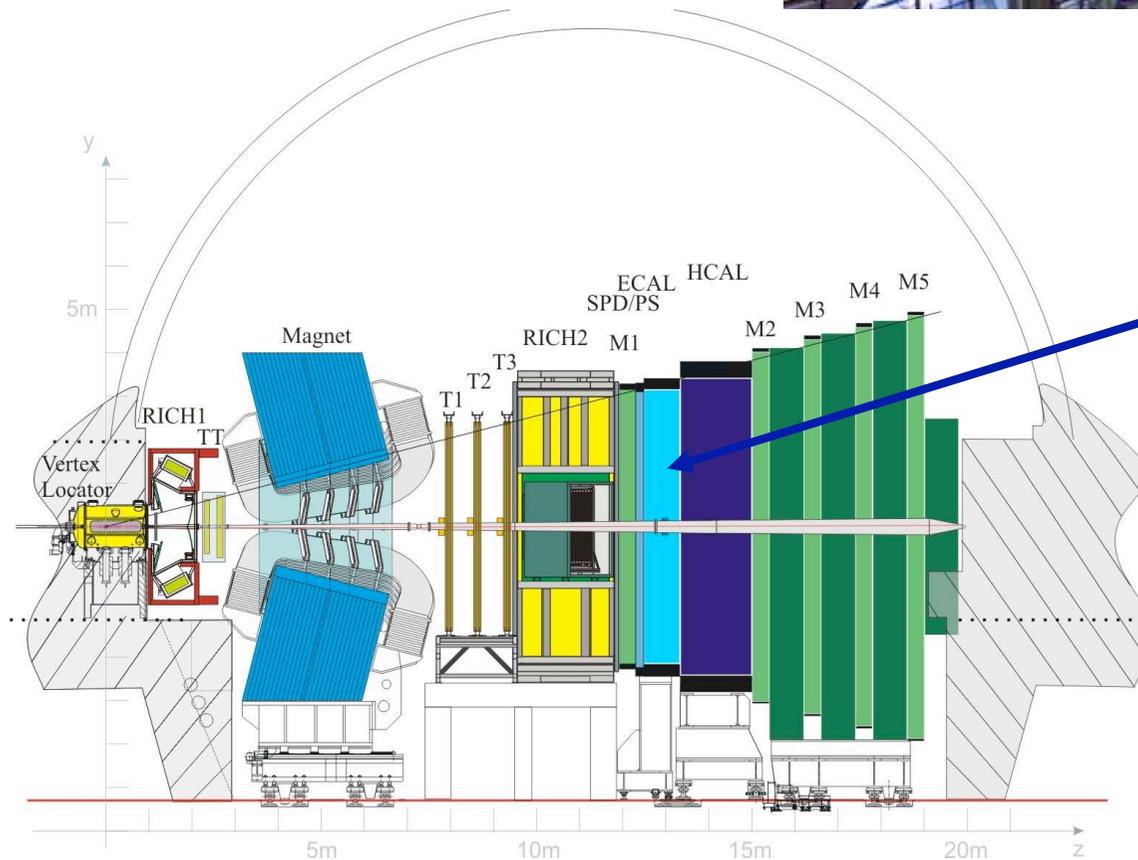
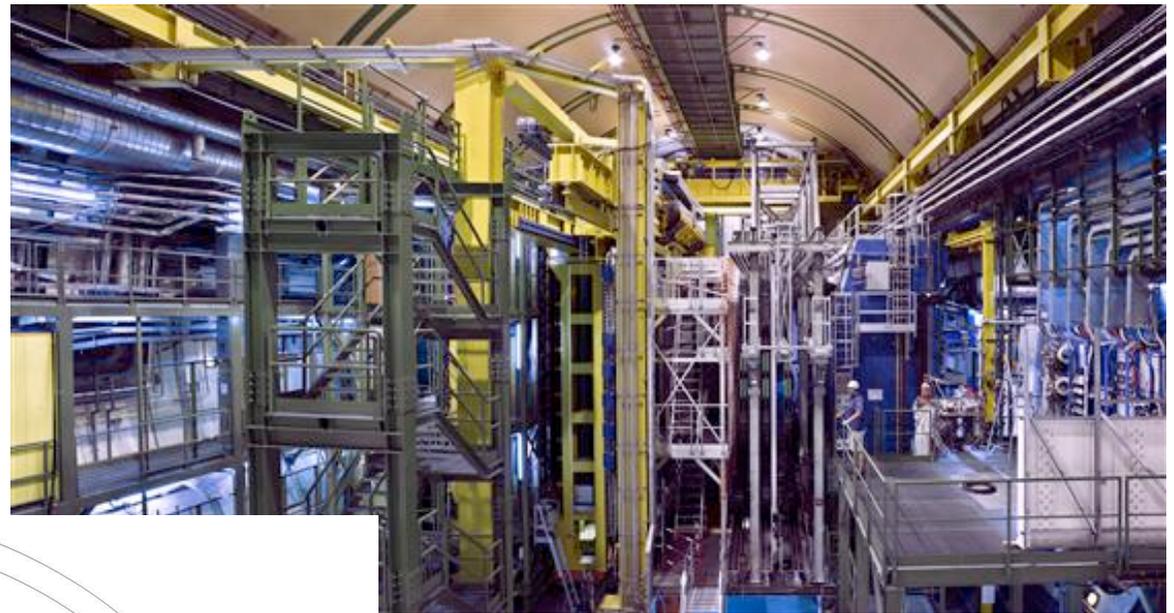
Radius : 10 m

Weight : 7000 tonnes

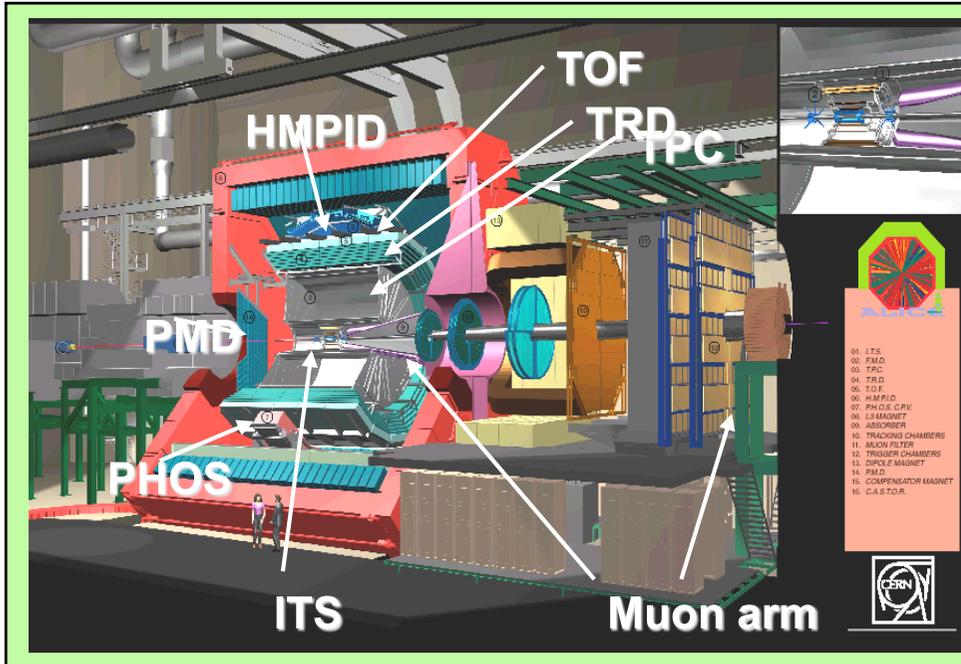
Electronic channels :  $10^8$

HCal :  
Clermont-Ferrand

# LHCb



Calorimètres :  
Structure/électronique  
Trigger L0  
Annecy, Clermont-Ferrand, Marseille, Paris, Orsay



## Nature of the confinement of quarks and gluons

### ALICE (CERN)

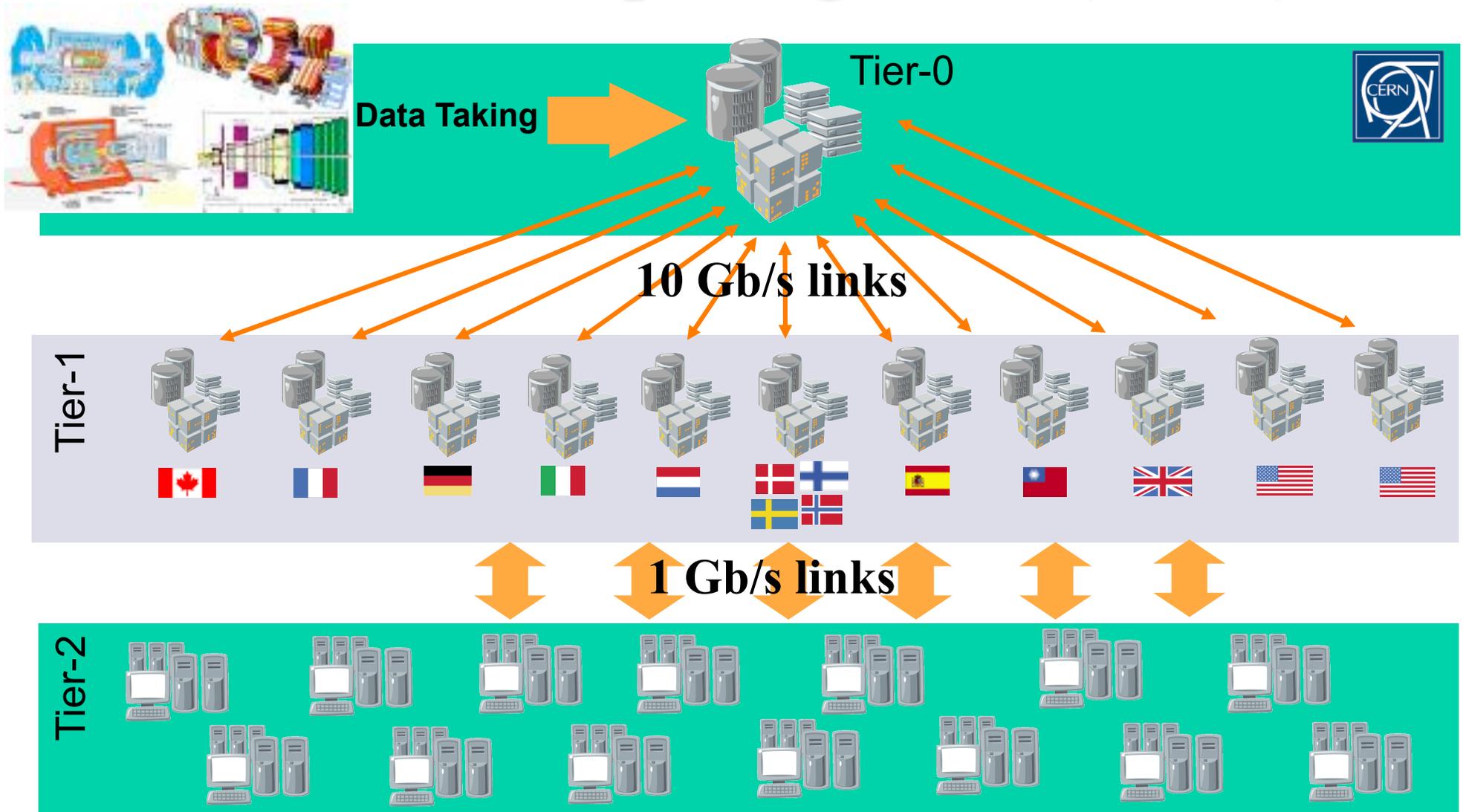
IN2P3 6 laboratories

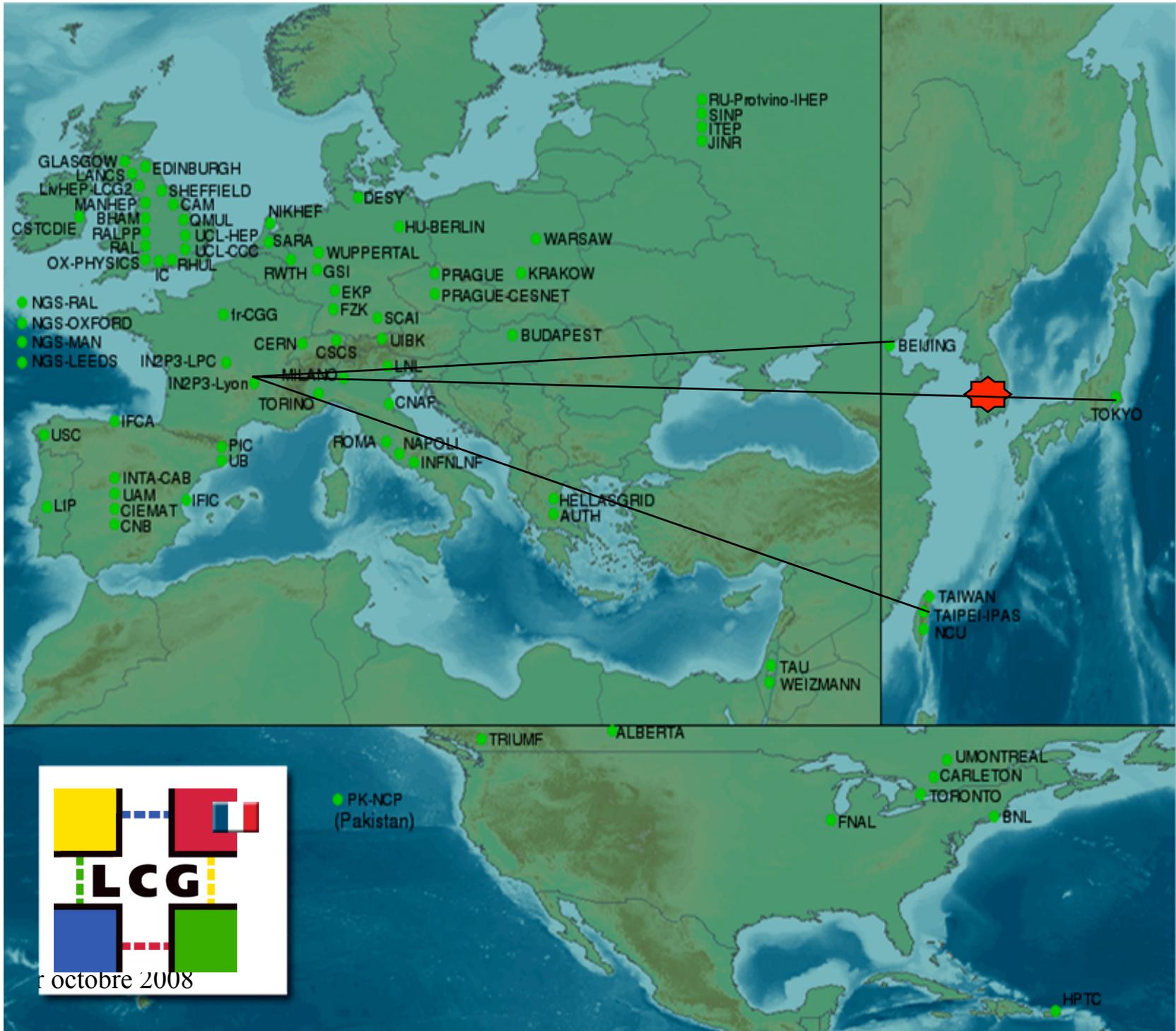
- Dimuon arm
- Internal tracking system (ITS)

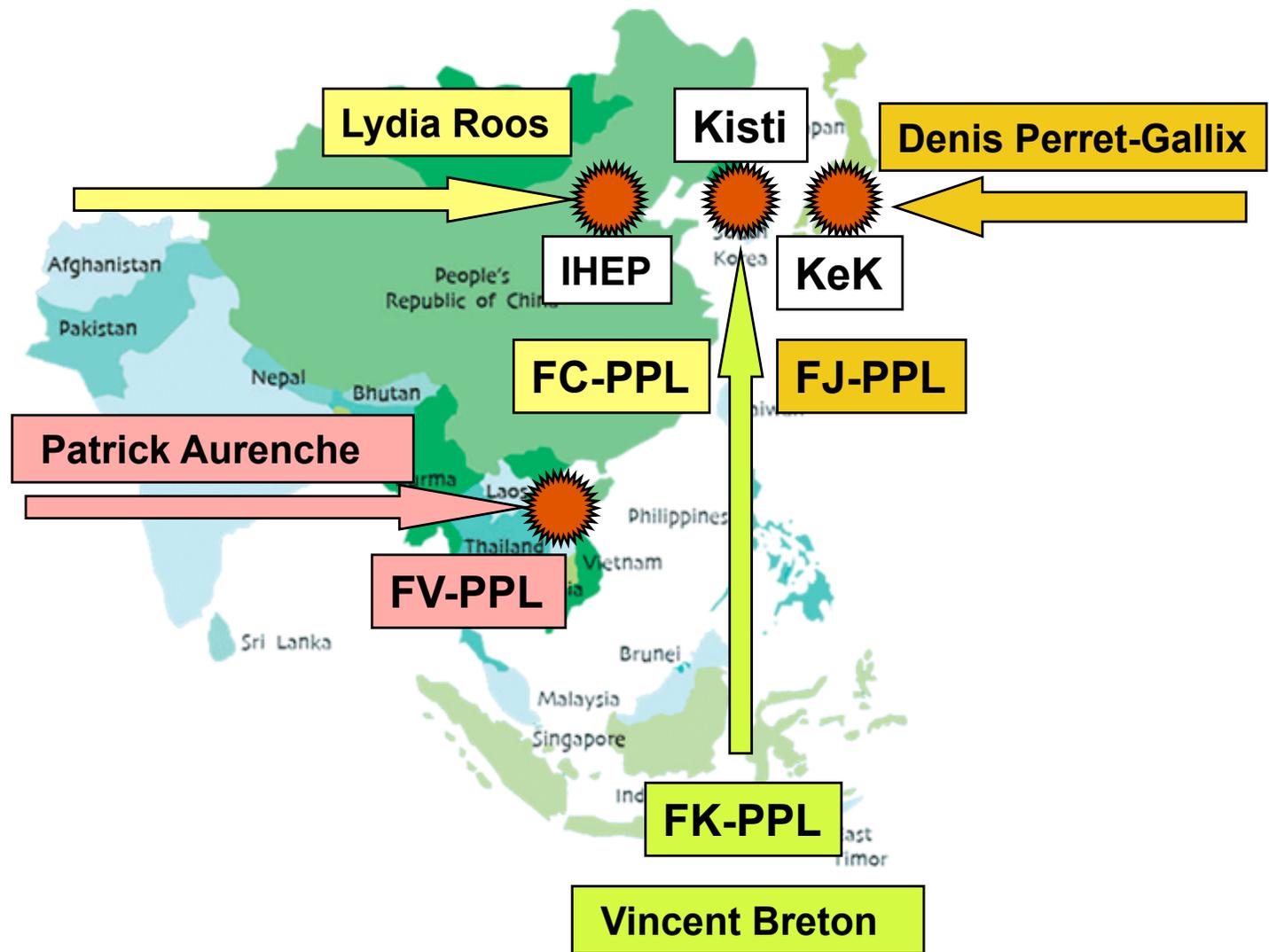
Investment: 6.6 M€

+ electromagnetic calorimeter

# LHC Computing Grid (LCG)







## Laboratoire Annecy de Physique des Particules

C. Adloff, Y. Karyotakis, A. Espargilière, R.Hermel, J.Prast, N.Geffroy, P.Delbecque, C.Girard

## Laboratoire de l'Accélérateur Linéaire

B.Bouquet ,R.Poeschl , F.Richard , H. Li ,Ph. Doublet , Ch. De la Taille, J.Fleury , S.Callier,  
N.Seguin-Moreau ,F.Dulucq , L.Raux ,G.Martin-Chassard ,F.Wicek ,A.Faloud

## Laboratoire Leprince Ringuet

V.Boudry, J-C.Brient, A.Rougé, M.Ruan, H.Videau, M.Reinhard ,  
M.Anduze , C.Clerc, G.Musat, P.Mora de Freitas, R.Cornat ,J-Ch. Vanel , F.Gastaldi

## Institut de physique Nucléaire de Lyon

I.Laktineh, M.Bedjidian , E.Latour, M.Vander Donckt , R.Kieffer, G.Grenier, P.Lebrun  
J.Fay, N.Lumb, H.Mathez , R.Gaglione, Ch. Combaret, E.Schibler, J.-C Ianigro

## Laboratoire de physique subatomique et corpusculaire

J-Y Hostachy, L.Morin , D.Grondin , D.Zahini ,O. Rosetto, C. Vescovi , F.-E Rarbi

## Laboratoire de physique Corpusculaire de Clermont

F.Badaud, M.Benyamna, C.Carloganu, P.Gay , Ph.Gris , G.Blanchar, N .Brun , M.Crouau , S.Manen ,  
L.Royer

# Conclusions

- La physique des particules a de beaux jours devant elle
  - LHC, sLHC, ILC
  - OPERA, T2K, ... Mégatonnes
  - SuperNEMO
- Des projets mondiaux très lourds
  - R&D Instrumentation (et informatique)
  - Gros détecteurs, Grosses collaborations

# Backup slides

# Particle Physics at IN2P3

**400 permanent scientists (FTE) / 10 Laboratories / 7 Experiments**

## Ongoing Experiments

H1 (10/3)

BaBar (20/5)

D0 (30/8)

## Next Experiments

ATLAS (60/7)

CMS (30/4)

LHCb (20/6)

ALICE (38/6)

*LCG (LHC computing grid) all*

## Future Experiments

ILC (30/9)

## *Data taking*

ended in **2007**

ended in **2008**

ends in **2009** (2010?)

starts in **2008**

starts in **2008**

starts in **2008**

starts in **2008**

ready

starts > 2020

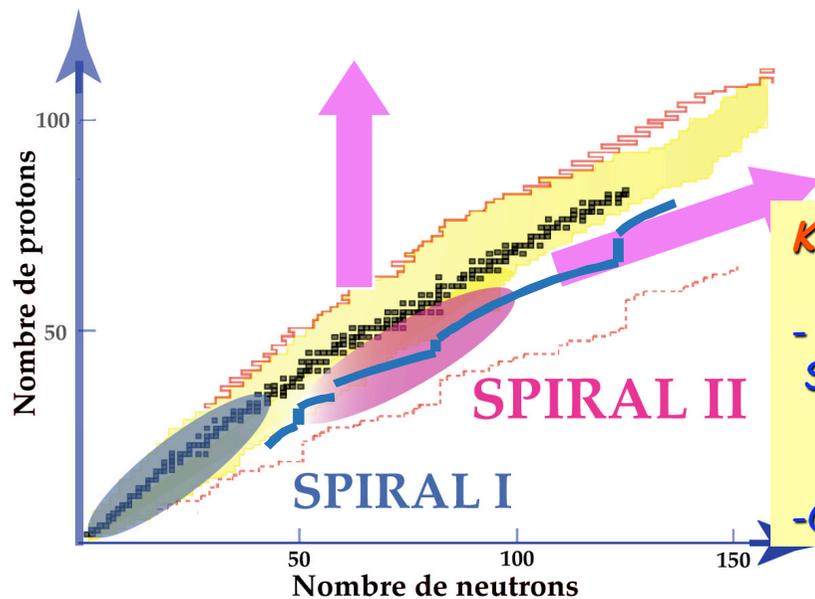
# Structure and the dynamics of nuclei. Nature of the nucleonic matter

- Origin of nuclear binding
- Limits of nuclear stability
- Heavy and super-heavy elements
- Formation of elements in the universe (nucléosynthesis)



## Some key facts:

- New limits of stability (37Ne)*
- Discovery of 48Ni*
- 2p radioactivity*
- New deformations, New magicity*
- Nuclear molecules*
- New detectors (EXOGAM, VAMOS)*



R&D for SPIRAL2: ALTO (agreement DUBNA-IPNO)  
R&D future detectors: AGATA

## KEEP THE INTERNATIONAL COMPETITIVITY OF GANIL SPIRAL2 project

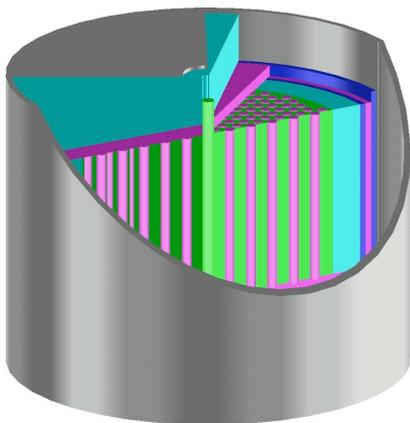
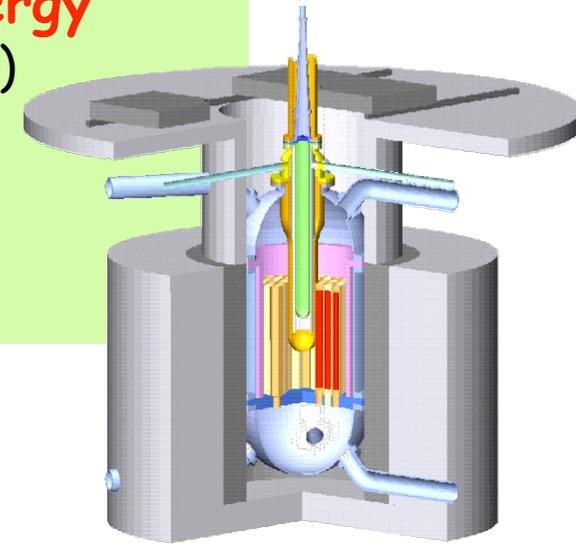
- 2002-2008 experiments with SPIRAL I
- SPIRAL II, a priority stated by NuPECC
- 118M€ (full cost) - cofinanced by Region (33%)
- Into operation: 2012-2015
- Complementarity with GSI/FAIR - Future with EURISOL...

# Pluridisciplinary Program PACE

*Back-end of electronuclear cycle*

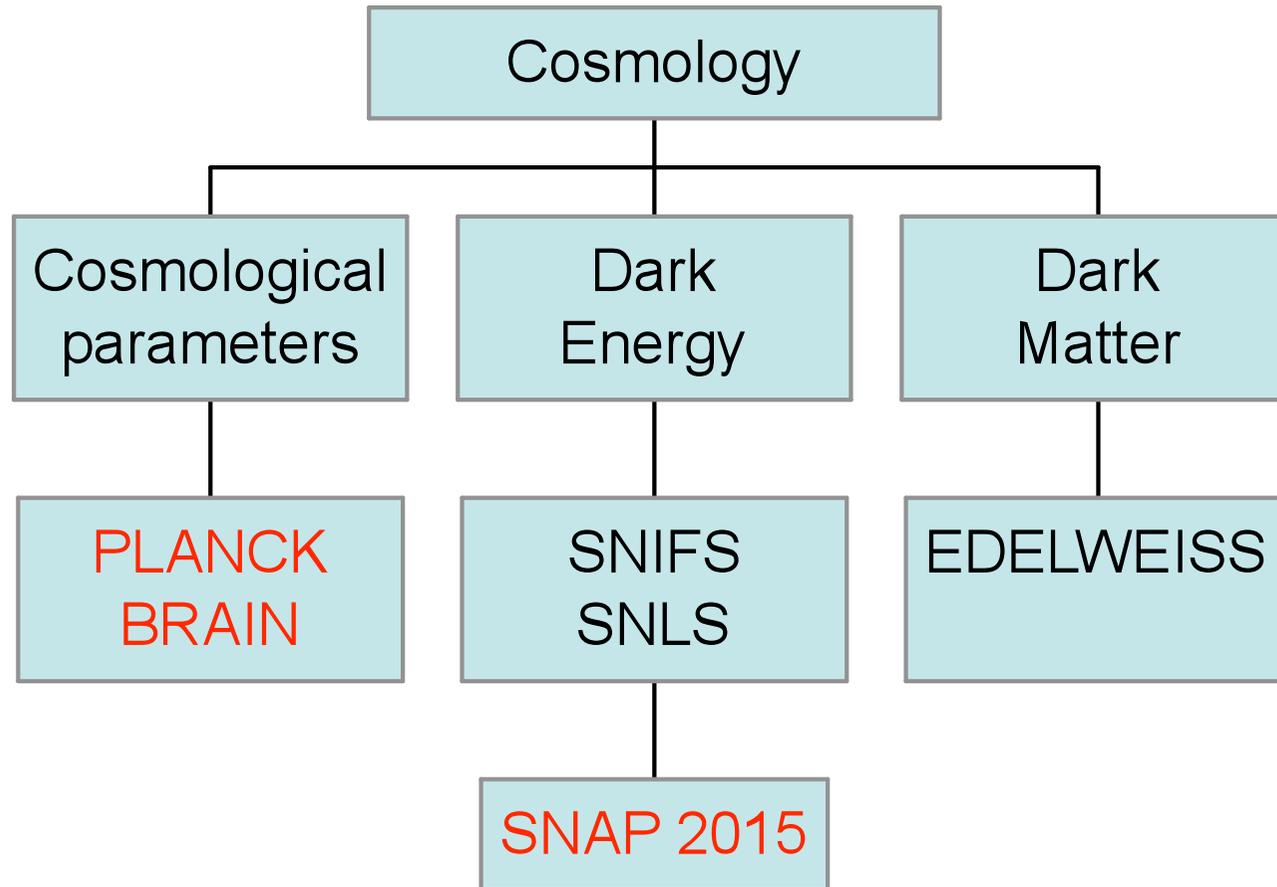
**IN2P3 focused on transmutation of nuclear wastes and innovative systems for the future of nuclear Energy**

- Experimental validation of Accelerator Driven Systems (ADS)  
R&D European programs  
(MEGAPIE, n-TOF, MUSE, PDS-XADS, EUROTRANS)
- High intensity accelerator developments (IPHI)



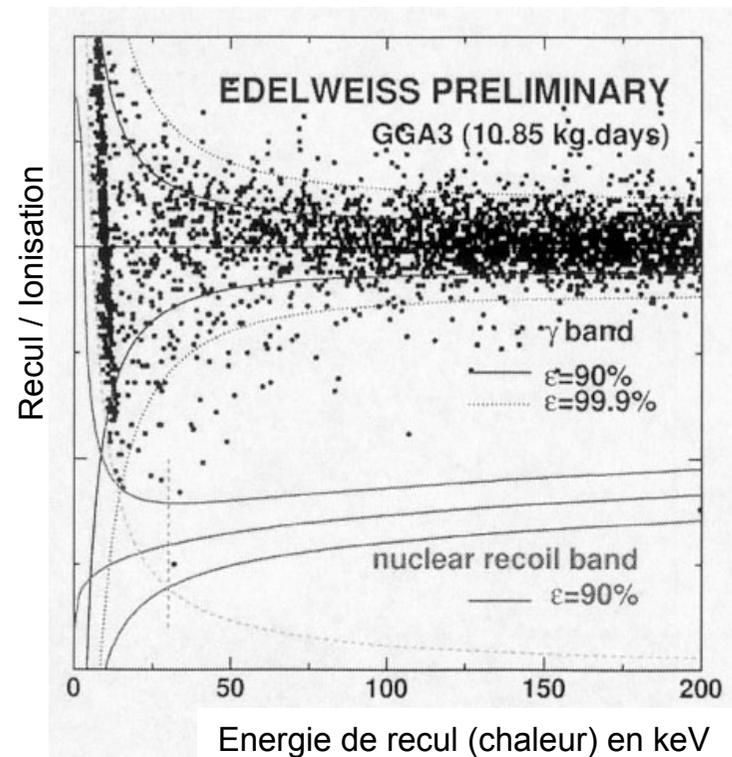
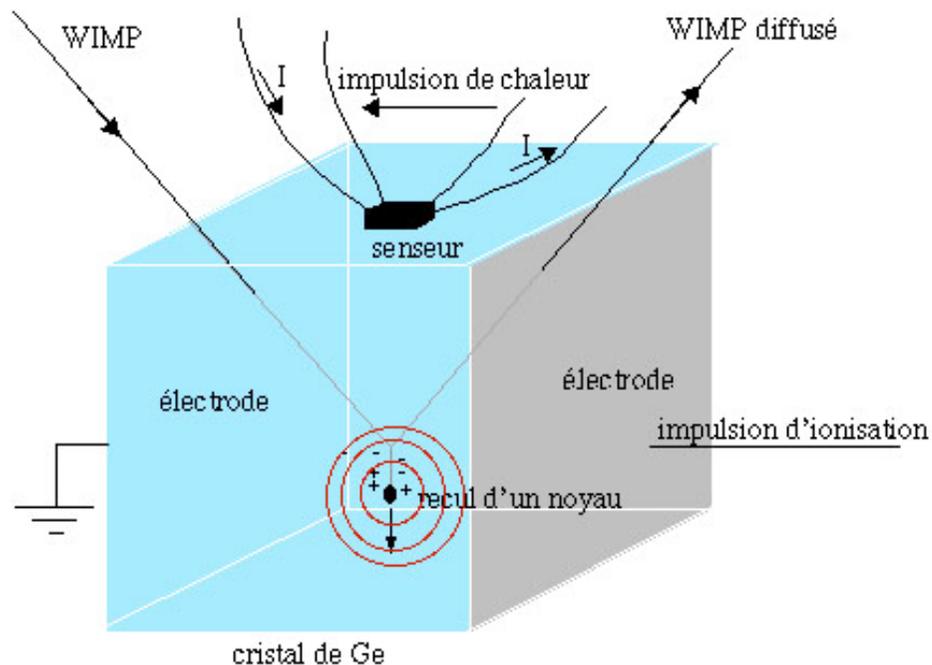
Innovative systems for Nuclear Energy based on the Th cycle Scenarios, focused on the concept of a molten salt reactor

# Current IN2P3/IRFU program in Cosmology

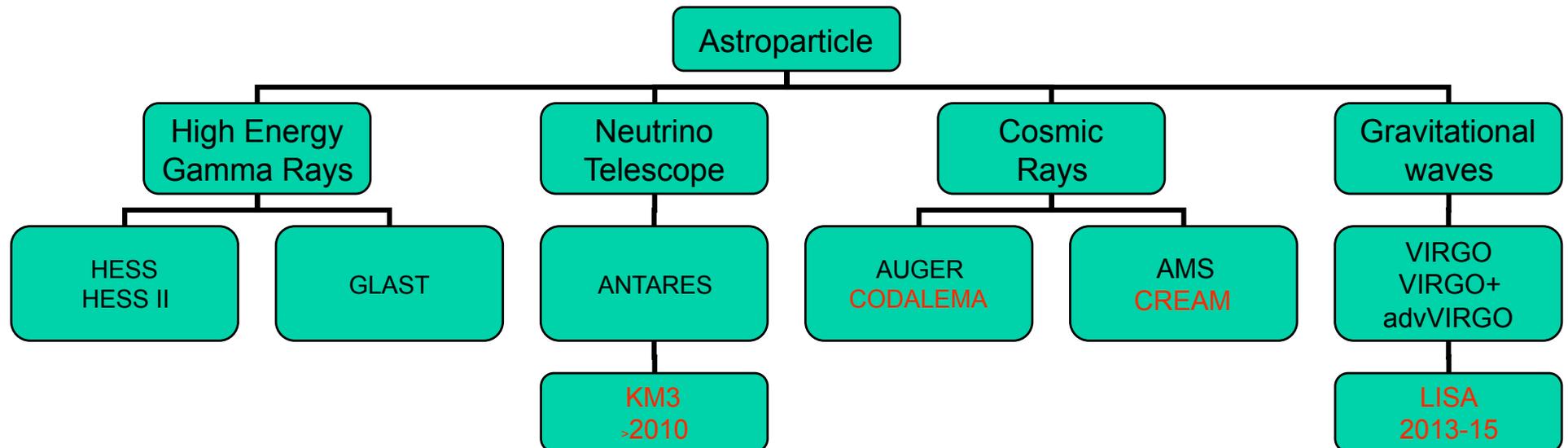


# EDELWEISS searches for WIMPS

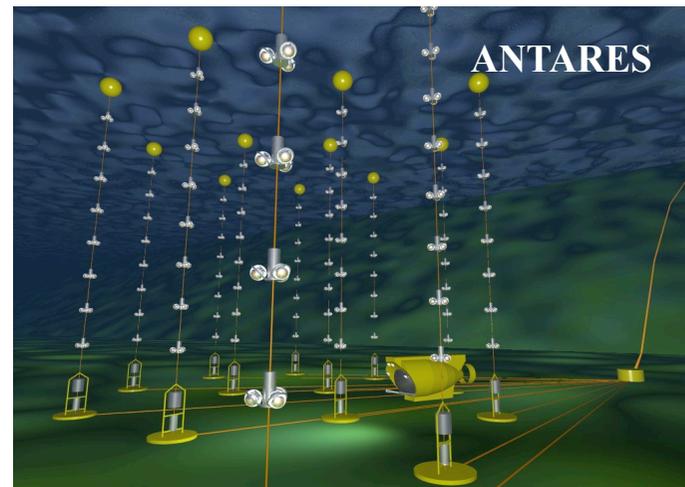
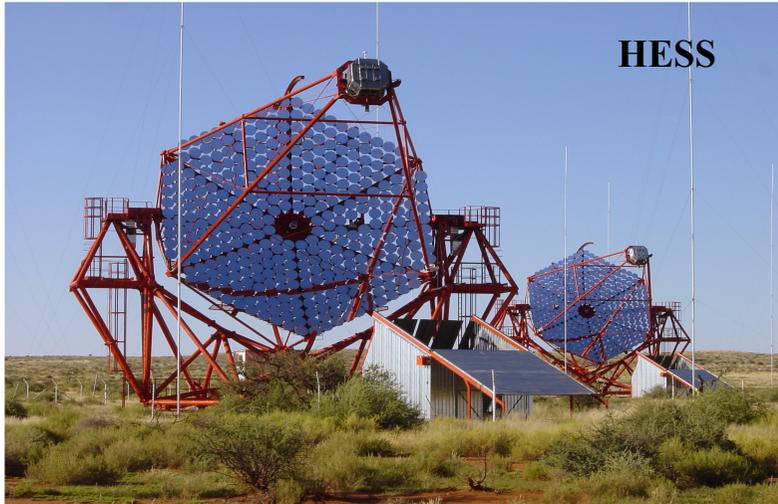
- **Technological challenge :**
  - **Big monocrystals (Germanium)**
    - Few grams  $\rightarrow$  kilograms  $\rightarrow$  100 kilograms **now 28 crystals = 10 kg**
    - **Température close to absolute zéro : 10 mK**
    - **Protected from ordinary cosmic rays  $\rightarrow$  Underground laboratory in Modane archeologic lead protection**
- CRTBT Grenoble, IPN Lyon, LSM Modane, CSNSM Orsay, IAP Paris, CEA/DAPNIA et CEA/DRECAM Saclay, Karlsruhe



# Current IN2P3/IRFU program in Cosmic Rays



# A cosmic ray program



# In the near future

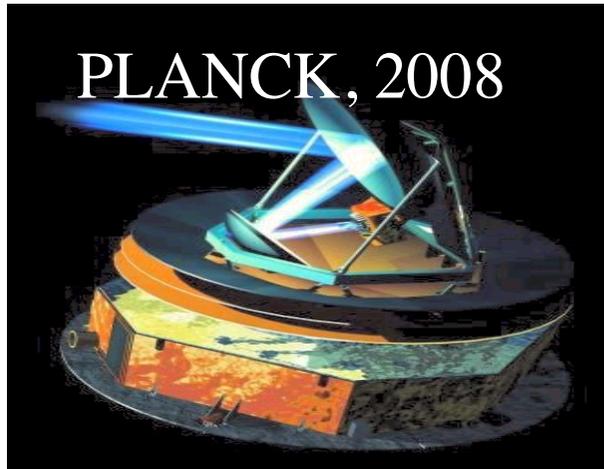
## Cosmic rays and cosmology in space

High energy gamma rays, antimatter, CMB, Dark Energy

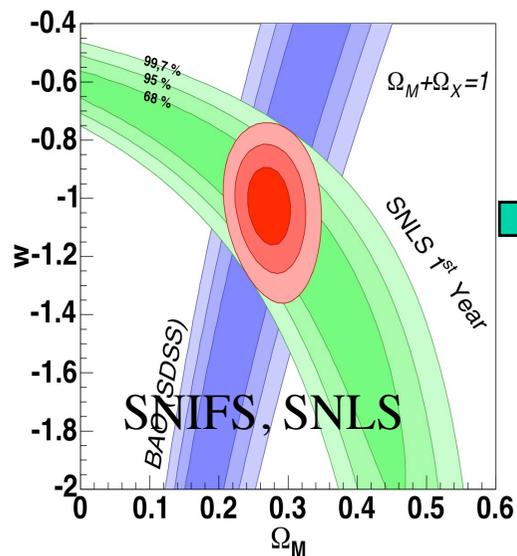
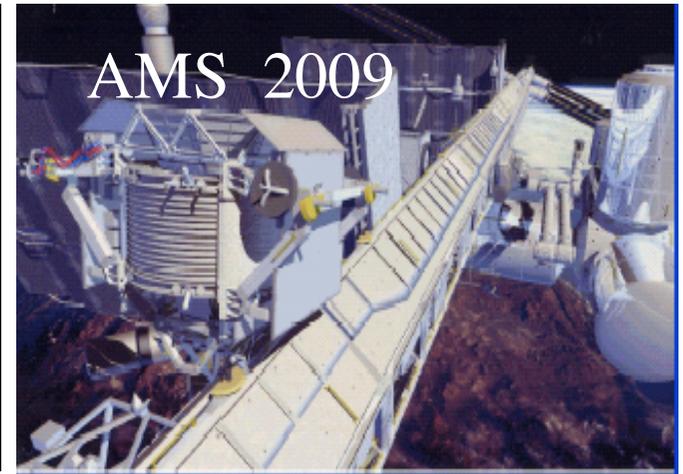
GLAST, 2007



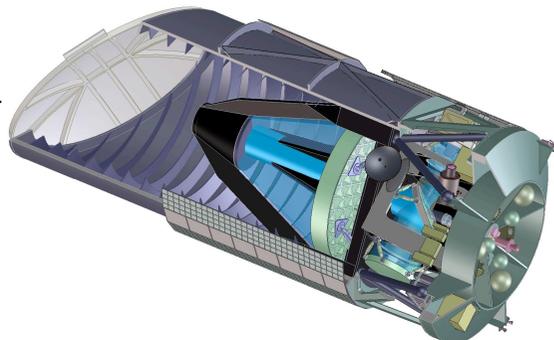
PLANCK, 2008



AMS 2009



SNAP/JDEM 2015

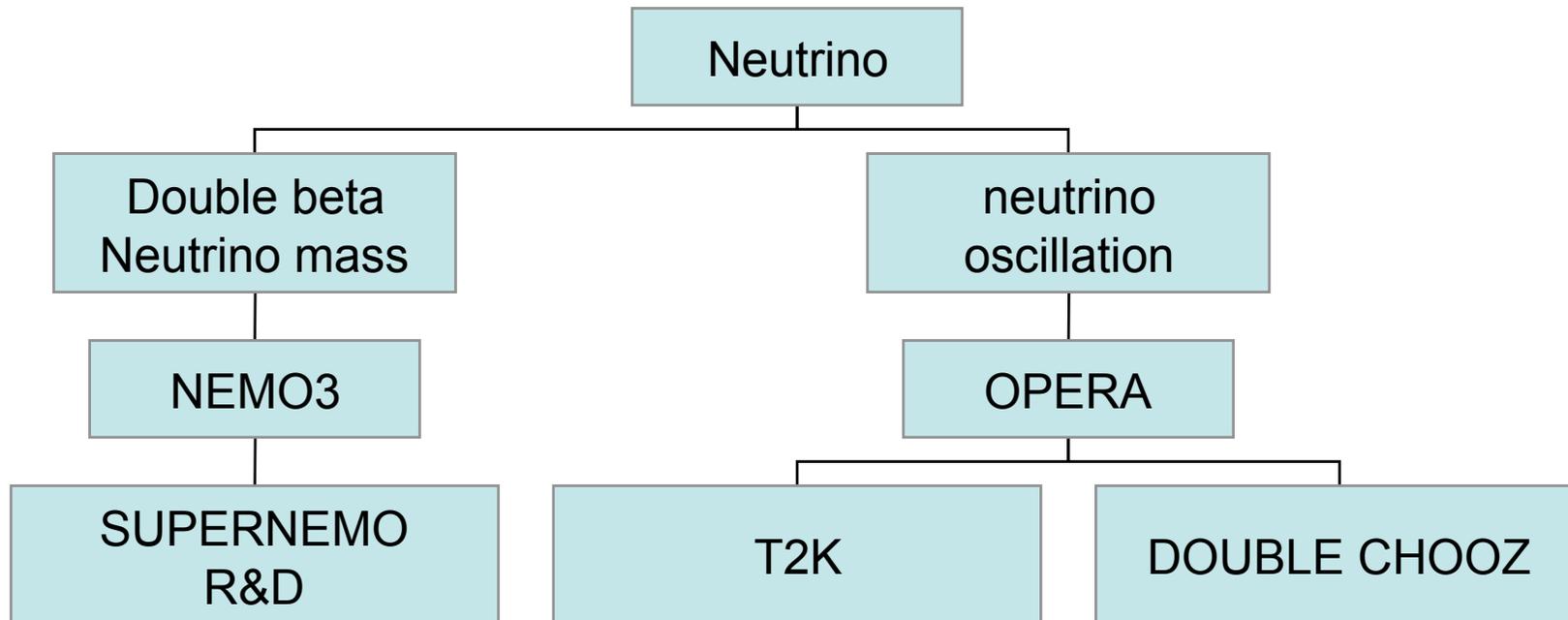


Journées Informatiques IN2P3-IRFU

LISA 2015

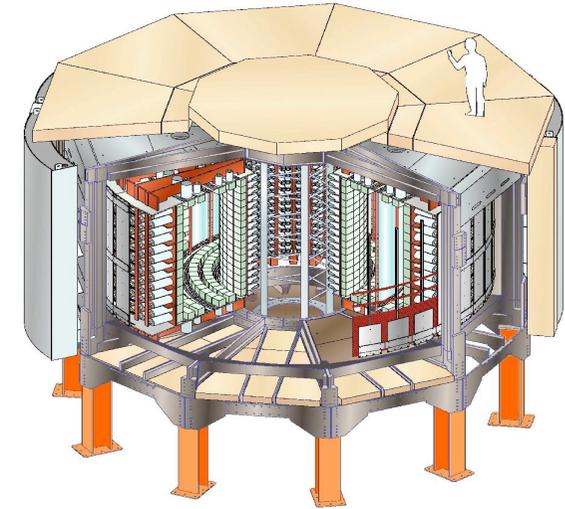
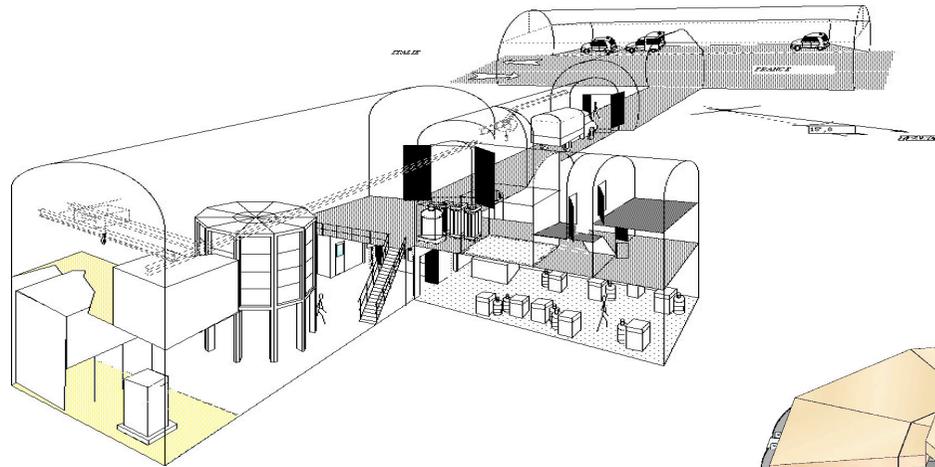


# Current IN2P3/IRFU program in Neutrino physics

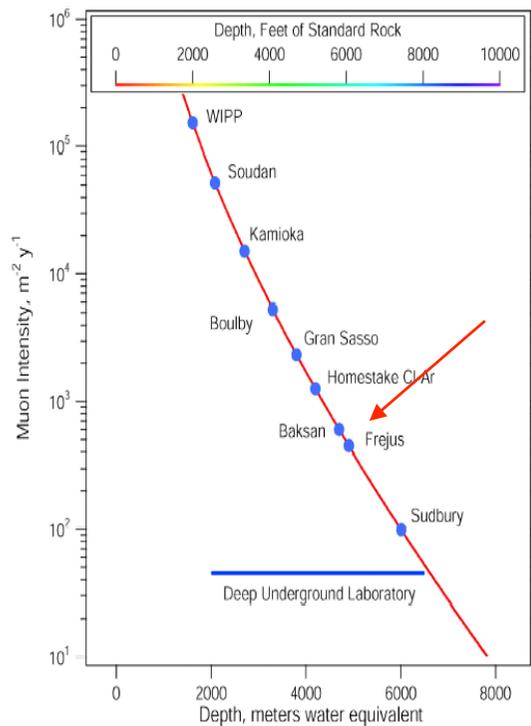




## Laboratoire Souterrain de Modane, France



## LSM in Fréjus



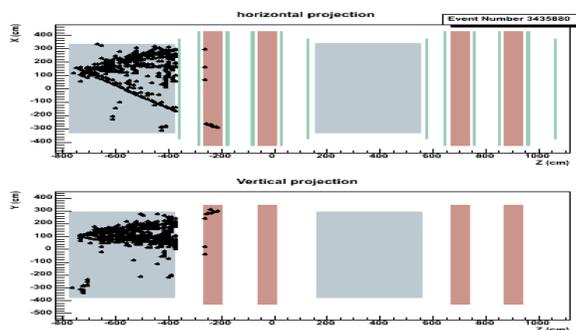
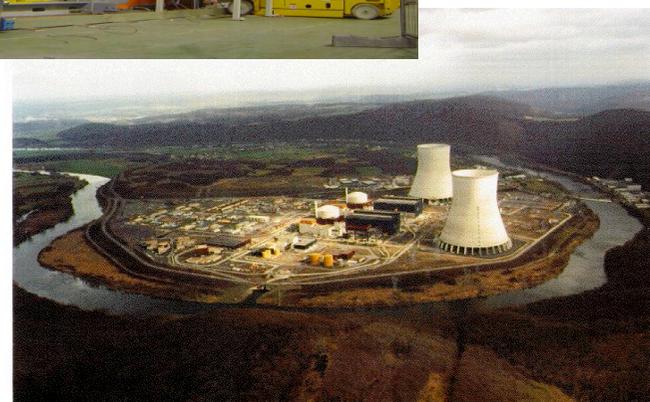
- NEMO3, double beta decay (2003-)
- EDELWEISS II, dark matter (2006)
- Low radioactivity studies (ILIAS)
- Search for SHE in nature

Journées Informatiques IN2P3-IRFU1

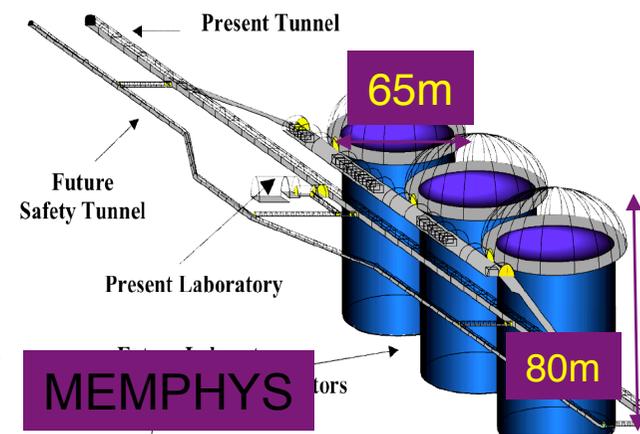


# Neutrino roadmap

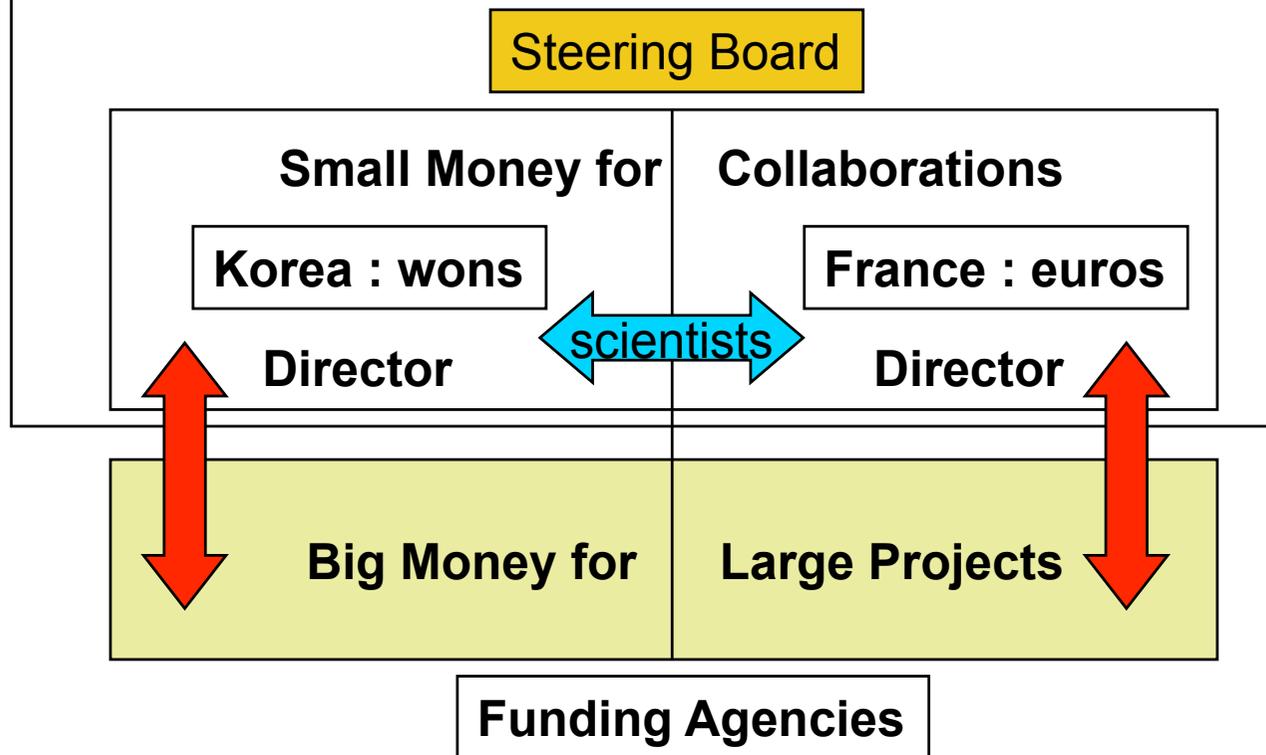
- OPERA(2006)
- In the next few years:
  - DOUBLECHOOZ (2008-2009)
  - T2K (2009-2013)
- R&D for SuperNEMO, double beta decay (>2010)
- At the horizon of 2015-18
  - Megatonne detector in Fréjus in relation to super and beta beam from CERN?



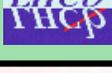
mées Informatiques IN2P3-IRFU1

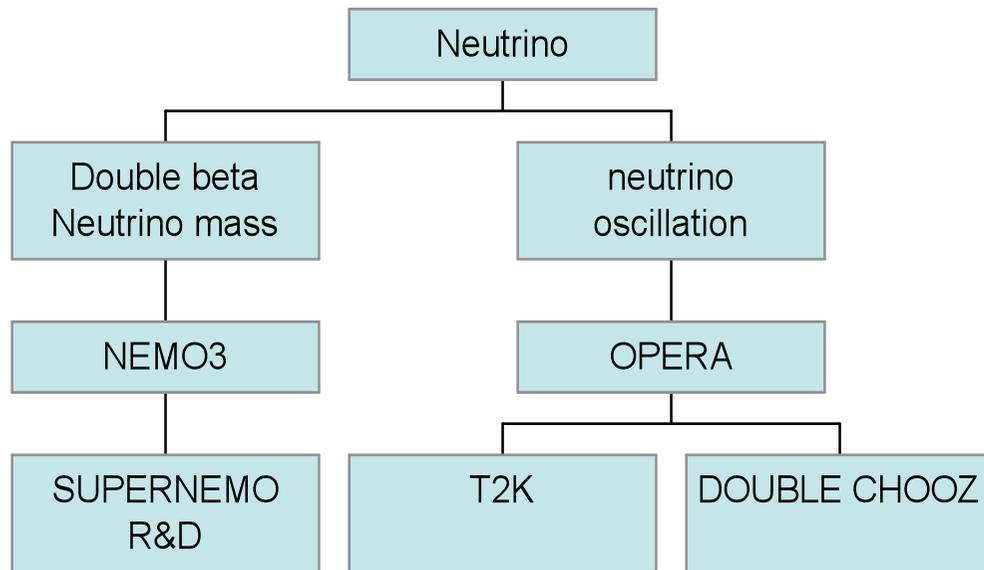
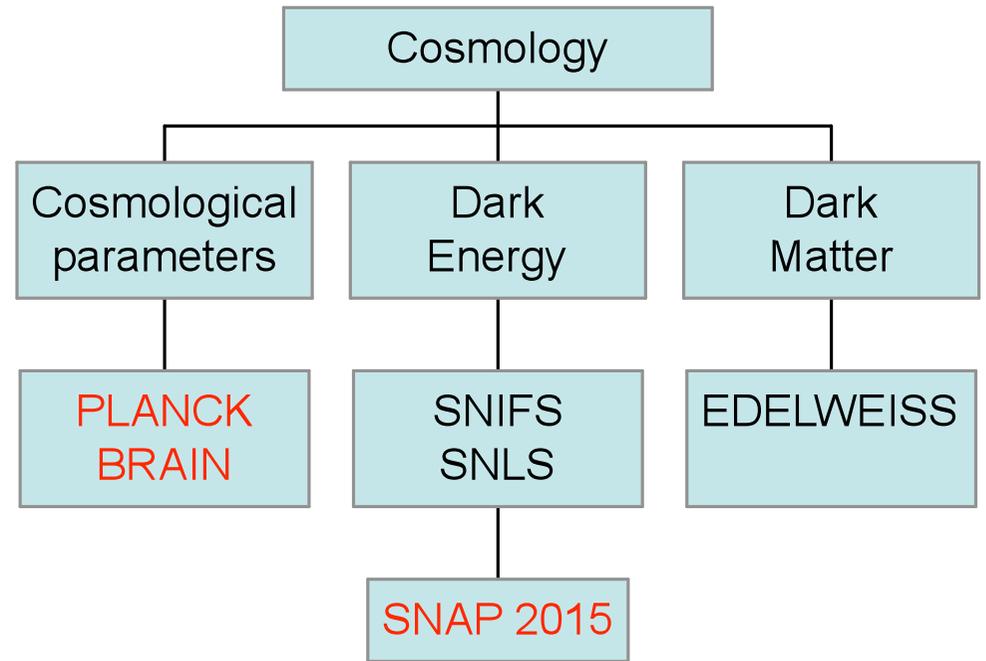
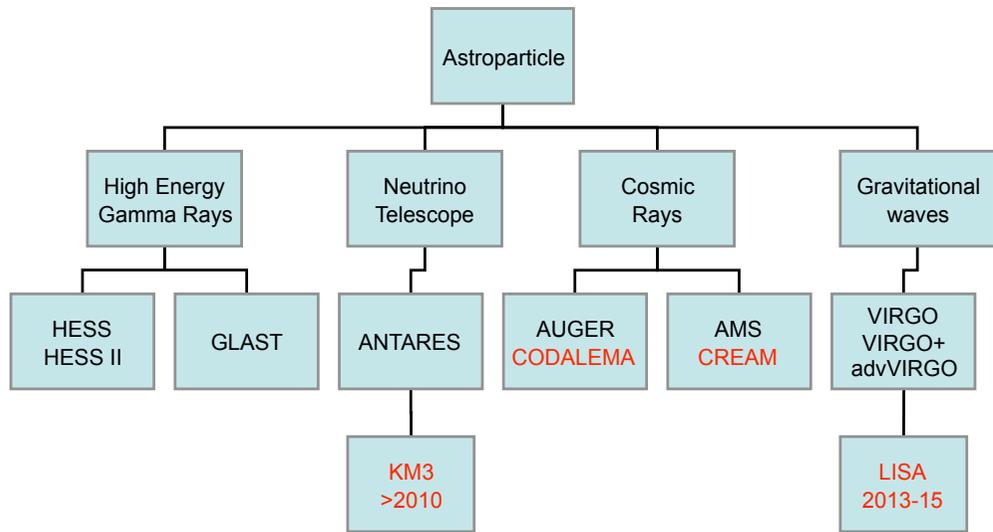


# LIA : International Associated Laboratory



# IN2P3 is a NETWORK of 20 large laboratories

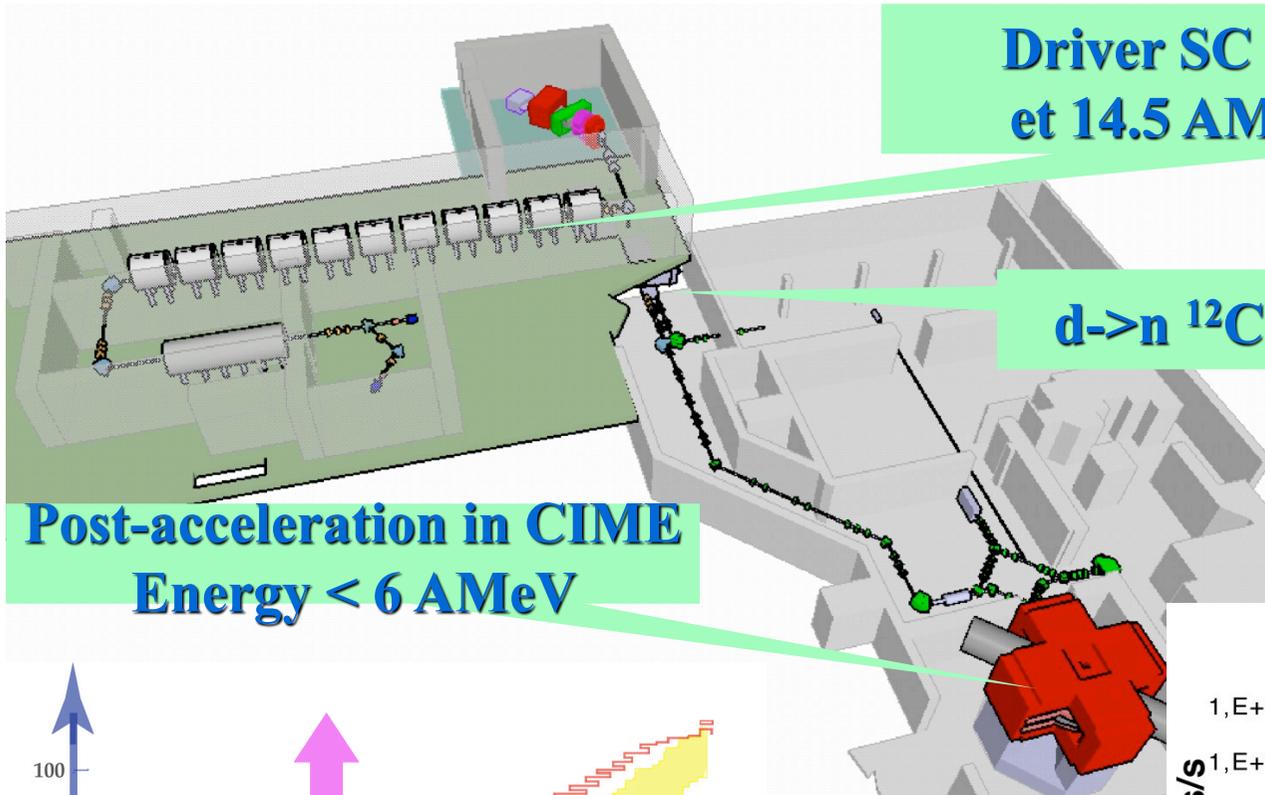
	BaBar	D0	H1	ATLAS	CMS	LHCb	LCG	ILC
CPPM								
LPSC								
IPNL								
LPC CI								
LAPP								
IReS								
LPNHE								
LLR								
LAL								
CC-IN2P3								



Current IN2P3 program

On Astroparticle, Cosmology  
and Neutrino

# SPIRAL II project



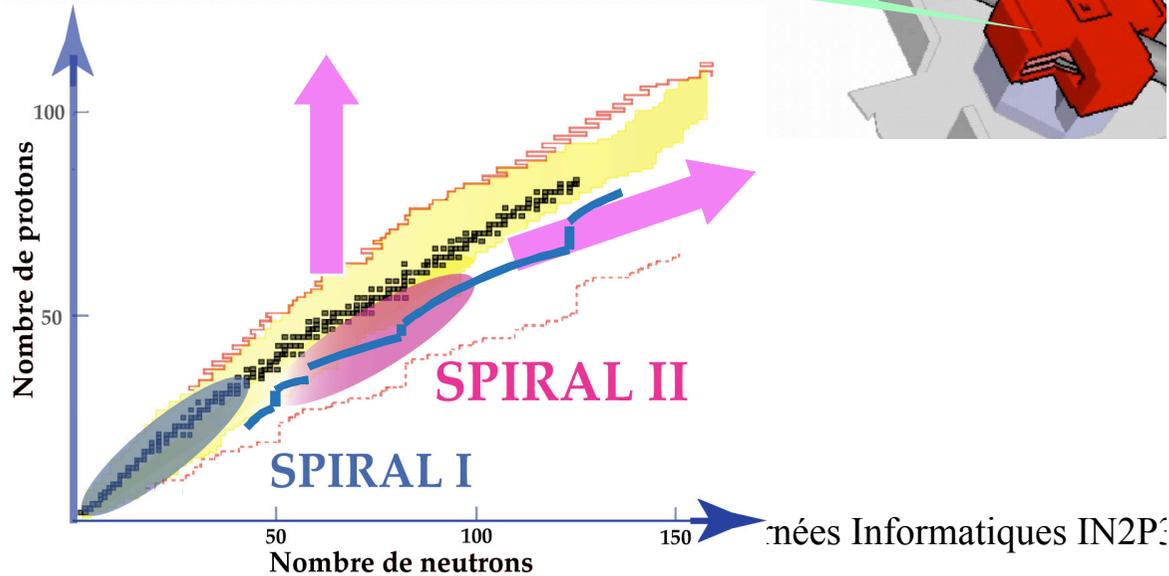
Driver SC Linac: 40MeV, 5mA d  
et 14.5 AMeV, 1pMA heavy-ions

$>10^{13}$  fissions/sec.

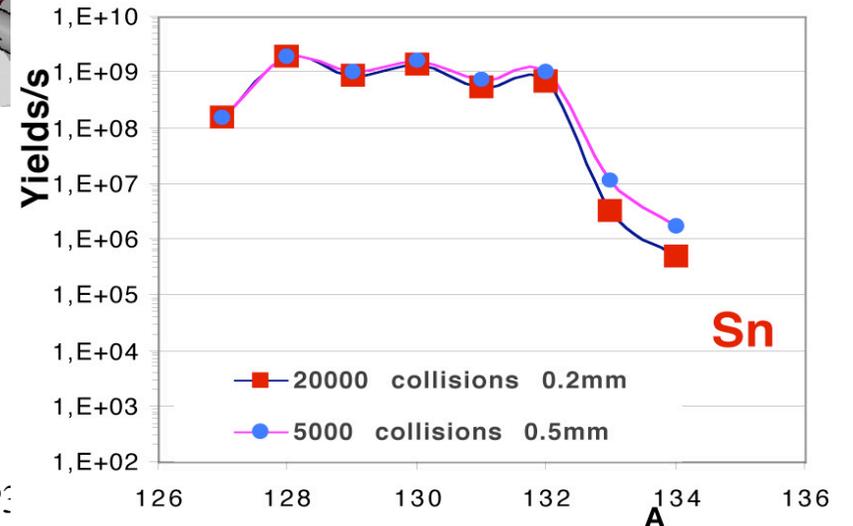
d  $\rightarrow$  n  $^{12}\text{C}$  converter +  $^{238}\text{UC}_x$  target

Post-acceleration in CIME  
Energy  $< 6$  AMeV

- Multi-beams (up to 5)
- Into operation: 2009



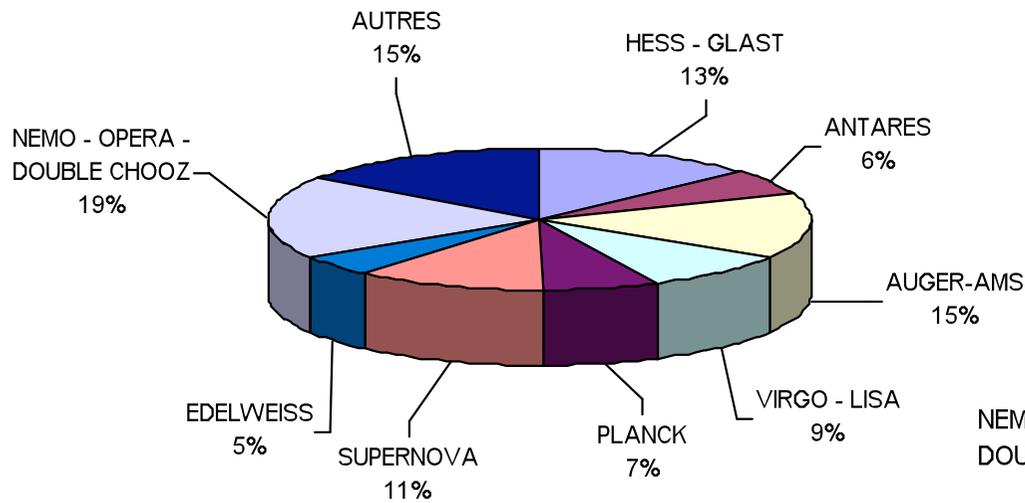
d (40MeV, 5mA)+C+UC



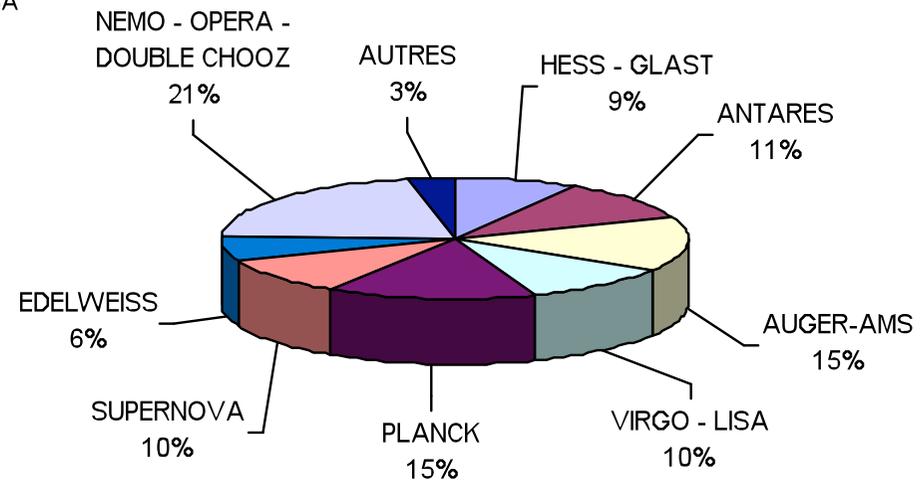
# ASTROPARTICLE PHYSICS & NEUTRINO

200 researchers, 154 engineers, 70 non permanent staff

**Permanent researchers**



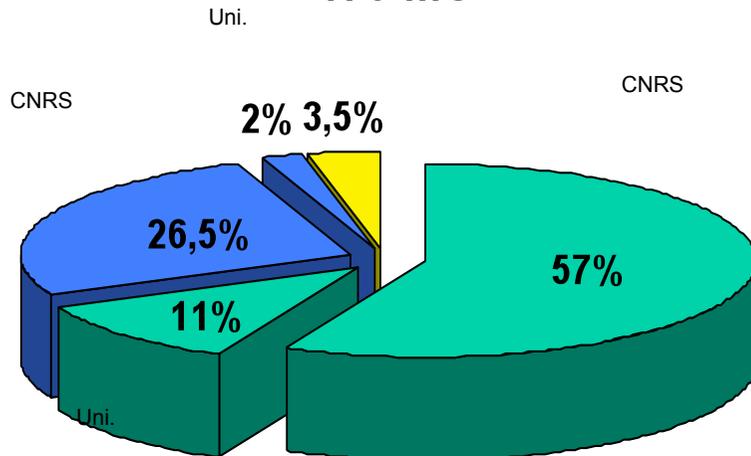
**Engineers & techniciens**



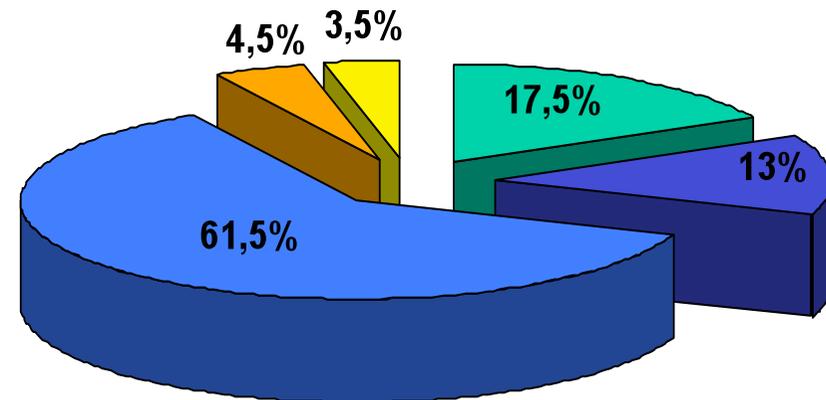
**Space programs 25 %**

# IN2P3/CNRS: 20 labs

**BUDGET**  
170 M€



**STAFF**  
2500



■ CNRS Staff + University  
■ Investment + running CNRS + University  
■ Others ressources

■ CNRS Researchers  
■ CNRS Technical Staff  
■ University Technical  
■ University Researchers  
■ None permanents