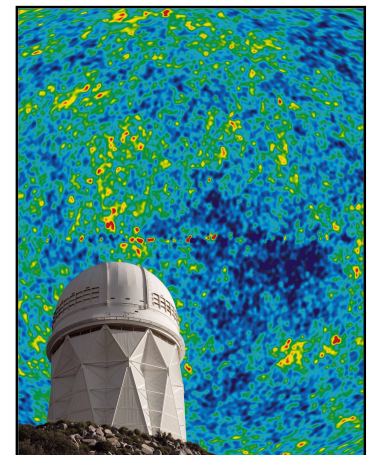




The Direction of Gravity

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Institute of the Early Universe, Korea**



The Direction of Gravity



Cosmic acceleration: Gravity is pulling *out* not down!

Is gravity (G_{Newton}) constant, or strengthening, or weakening with time?

Does gravity govern the growth of large scale structure exactly as it does for cosmic expansion, or are there more degrees of freedom?

Effect of gravity on light (strong/weak lensing).

Does gravity behave the same on all scales?

Dark energy motivates us to ask “what happens when gravity no longer points down?”.

Higher Dimensional Data



Cosmological Revolution:



From 2D to 3D – CMB anisotropies to tomographic surveys of density/velocity field.

Data, Data, Data

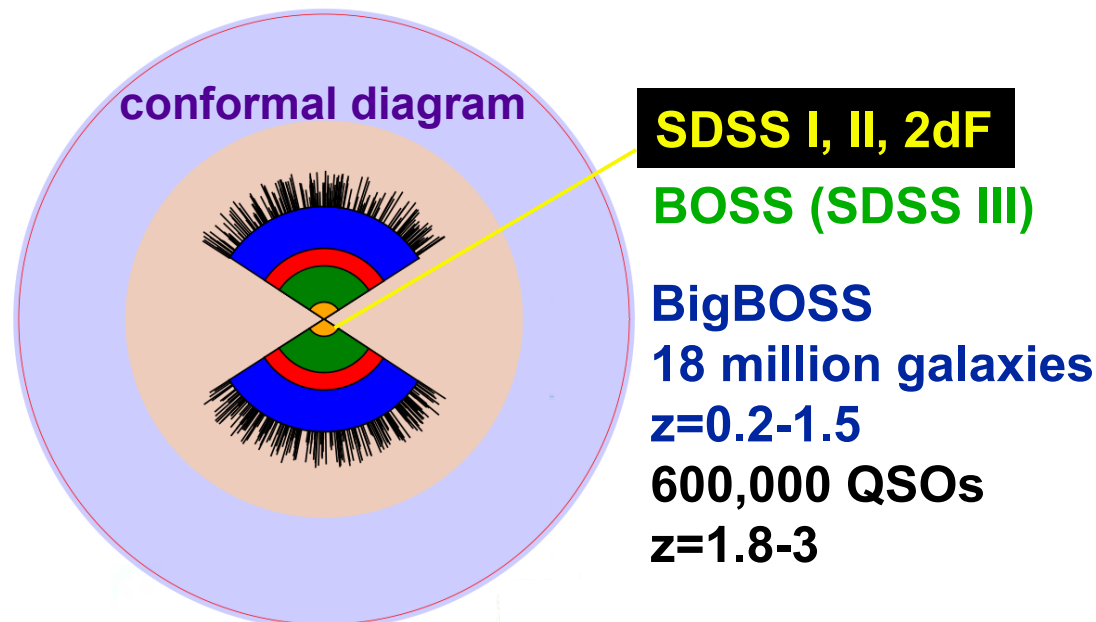


As wonderful as the CMB is, it is 2-dimensional.
The number of modes giving information is $l(l+1)$ or
~10 million.

BOSS (SDSS III) will map 400,000 linear modes.

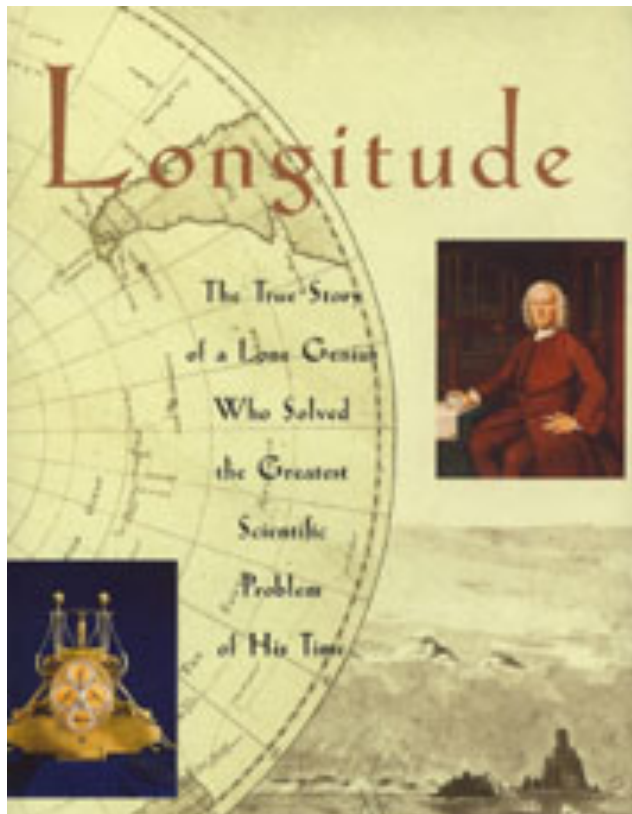
BigBOSS/Euclid will map 15 million linear modes.

N. Padmanabhan



Maps of
density
velocity
gravity

“Greatest Scientific Problem”



“When I’m playful I use the meridians of longitude and parallels of latitude for a seine, drag the Atlantic Ocean for whales.”

– Mark Twain, *Life on the Mississippi*

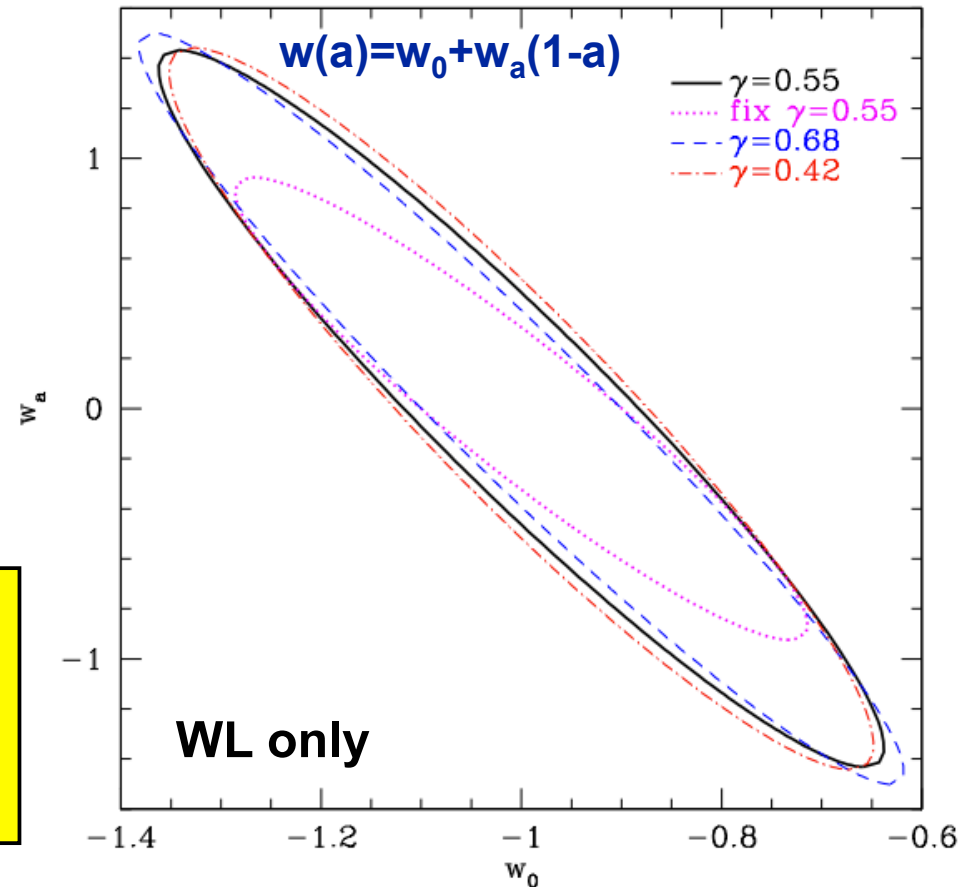
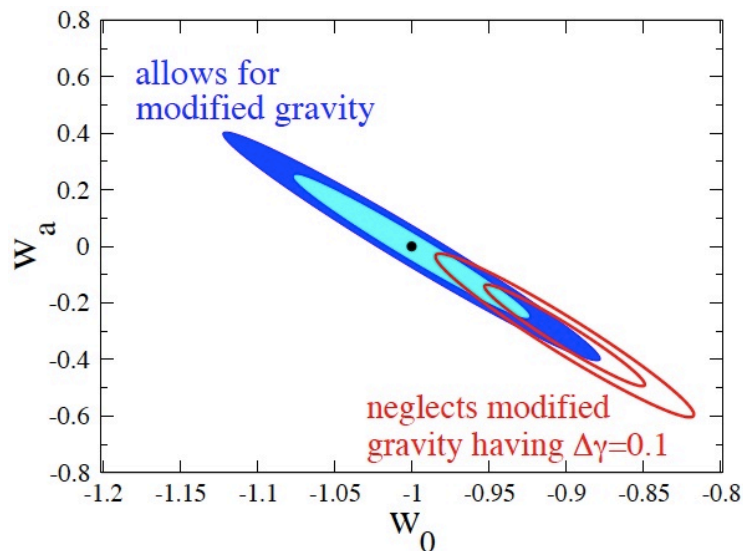
Cosmological Framework



Allow parameters to describe growth separate from expansion, e.g. gravitational growth index γ .

Otherwise bias $\Delta w_a \sim 8\Delta\gamma$

$$f = \frac{d \ln D}{d \ln a} \sim \Omega_m(a)^\gamma$$



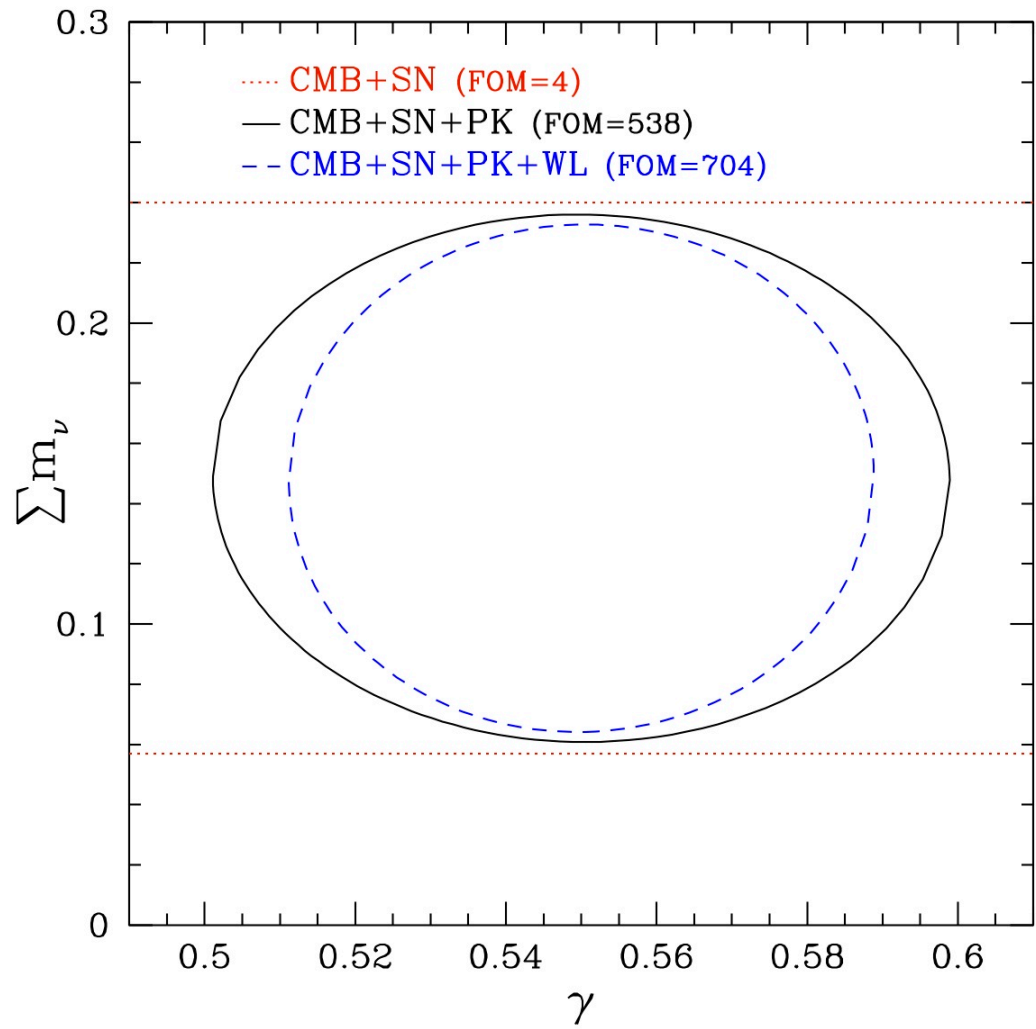
Fit simultaneously;
good distinction from
equation of state.

5 Year Realization (Cosmology 2017)



CMB lensing also probes gravity.

CMBlens+BOSS+DES can get $\sigma(\gamma)=0.026$ by ~2017!



Fit for vanilla +

w_0, w_a

m_ν

γ

Das & Linder 2012

Beyond GR Functions



Test gravity in model independent way.

Gravity and growth: $\nabla^2 \phi = 4\pi G a^2 \delta\rho$

Gravity and acceleration: $-\vec{\nabla}\psi = \ddot{x}$

Are ϕ and ψ the same? (yes, in GR)

Tie to observations via modified Poisson equations:

$$\nabla^2(\phi + \psi) = 8\pi G_N a^2 \delta\rho \times G_{\text{light}}$$

$$\nabla^2\psi = 4\pi G_N a^2 \delta\rho \times G_{\text{matter}}$$

G_{light} tests how light responds to gravity: central to lensing and integrated Sachs-Wolfe.

G_{matter} tests how matter responds to gravity: central to growth and velocities (γ is closely related).

Scale and Time Dependence

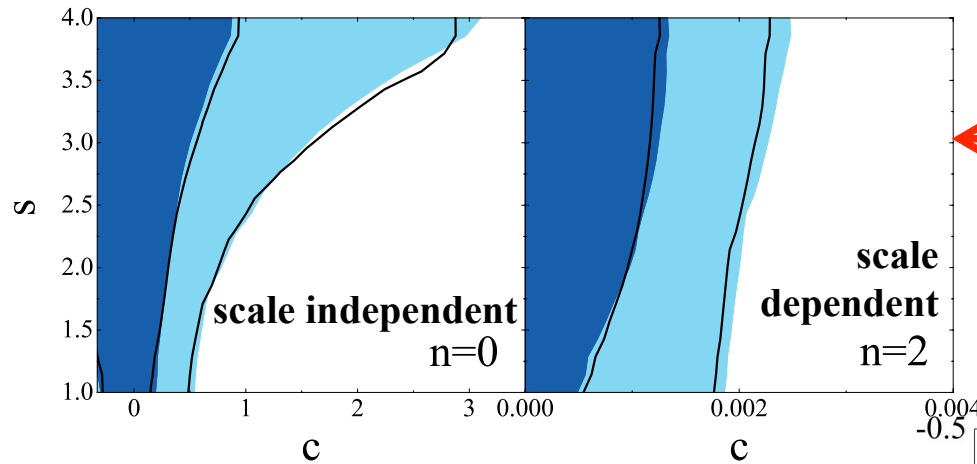


Padé approximant weights high/low z fairly.

$$G_{\text{matter}} = 1 + \frac{ca^s(k/H_0)^n}{1 + 3|c|a^s(k/H_0)^n}$$

Zhao+ 1109.1846

Accurate to ~1% for f(R) and DGP gravity.

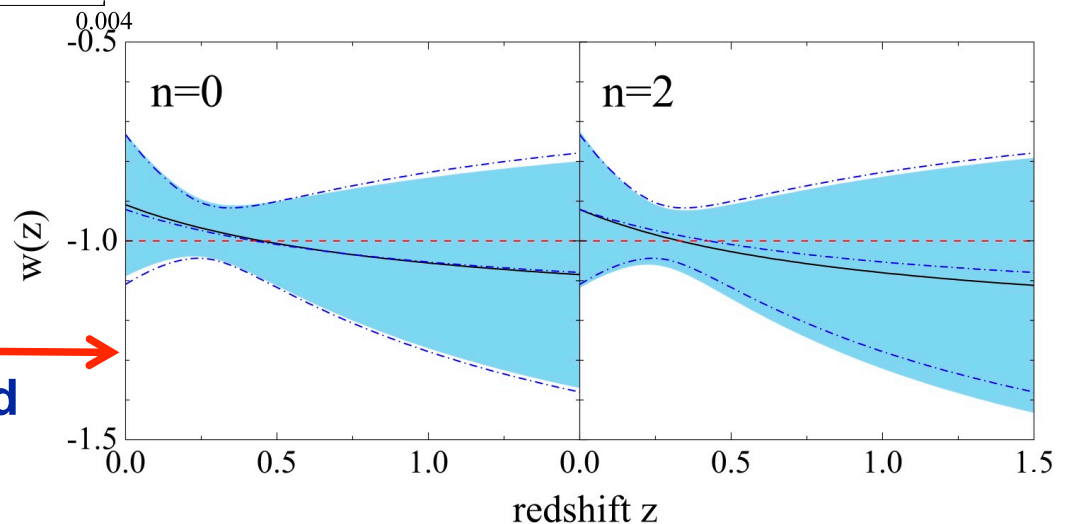


Shaded – fix to Λ ; Outline – fit w_0, w_a

Gravity fit unaffected by expansion fit.

Outline – fix to GR ;
Shaded – fit gravity c,s

Expansion fit unaffected
by gravity fit.



2 x 2 x 2 Gravity



Bin in **k** and **z**:

Model independent “2 x 2 x 2 gravity”

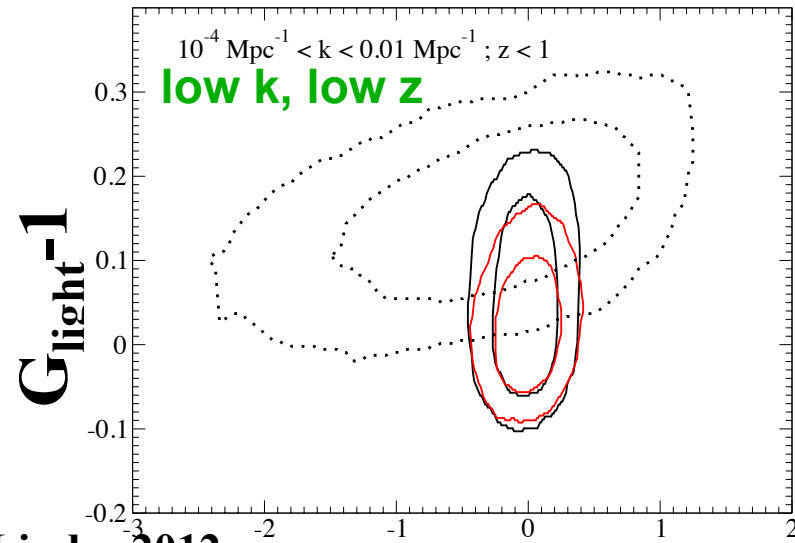
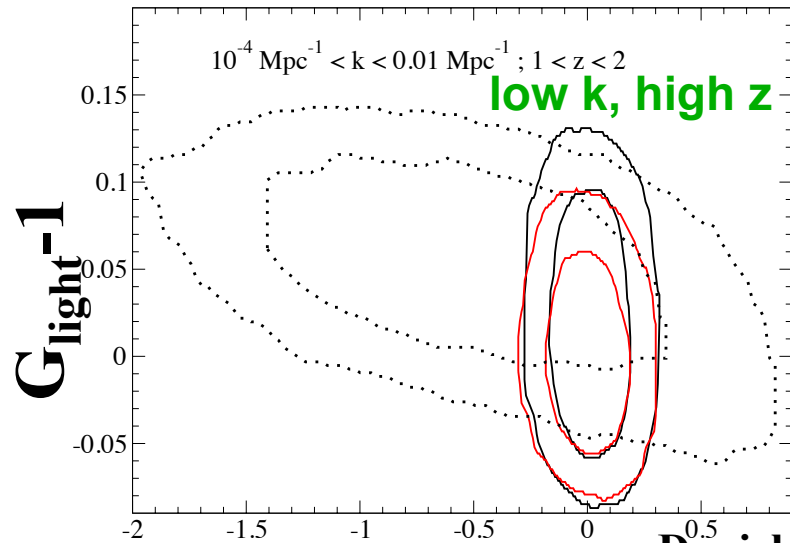
Why bin?

1) Model independent.

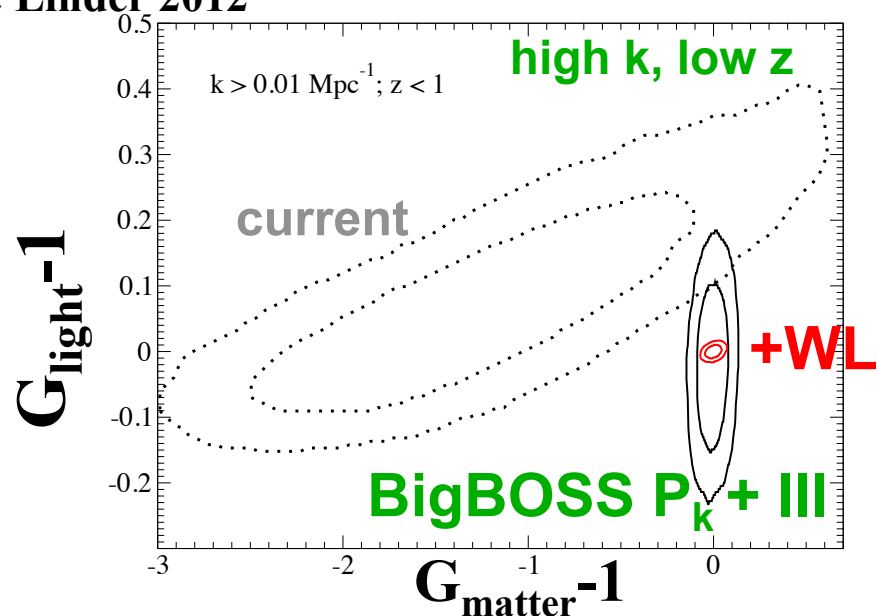
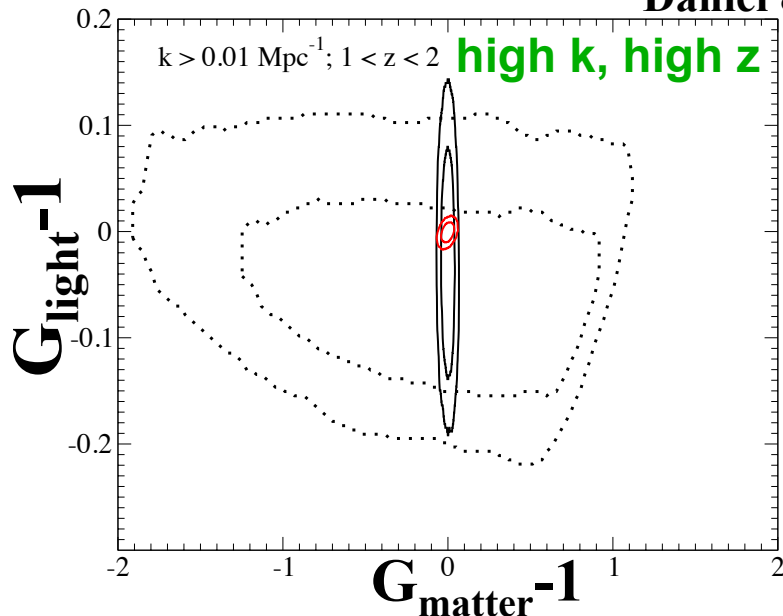
2) Cannot constrain >2 PCA with strong S/N (N bins gives $2N^2$ parameters, $N^2(2N^2+1)$ correlations).

3) a^s form gives bias: value of **s** runs with redshift so fixing **s** puts CMB, WL in tension. Data insufficient to constrain **s**.

Next Generation Leverage



Daniel & Linder 2012



5-10% test of 8 parameters of model-independent gravity.

The Direction of Gravity



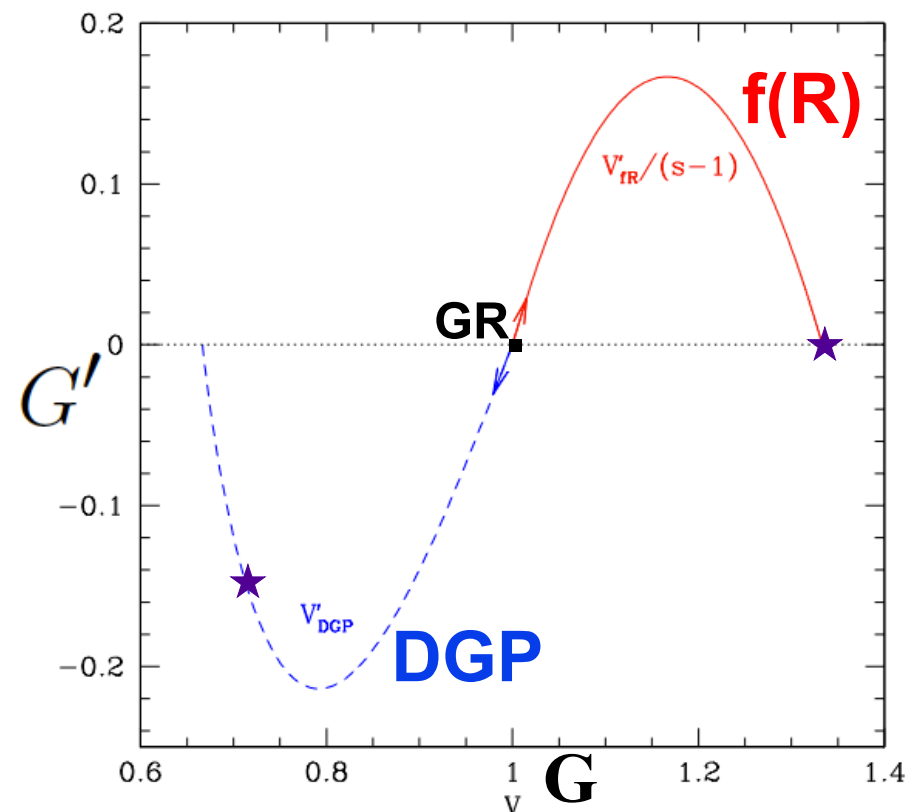
Scalar field dark energy (and Λ) have problems with naturalness of potential and high energy corrections.

Can avoid *both* problems by having a purely geometric object with no potential.

Galileon fields arise as geometric objects from higher dimensions and have shift symmetry protection (like DGP).

They also have screening (Vainshtein), satisfying GR on small scales.

Nicolis+ 2009, Deffayet+ 2009



Galileon Gravity



Scalar field π with shift symmetry $\pi \rightarrow \pi + c$, derivative self coupling, guaranteeing 2nd order field equations.

$$S = \int d^4x \sqrt{-g} \left[\left(1 - 2c_0 \frac{\pi}{M_{\text{pl}}} \right) \frac{M_{\text{pl}}^2 R}{2} - \frac{c_2}{2} (\partial\pi)^2 - \frac{c_3}{M^3} (\partial\pi)^2 \square\pi - \frac{c_4 \mathcal{L}_4}{2} - \frac{c_5 \mathcal{L}_5}{2} - \frac{M_{\text{pl}}}{M^3} c_G G^{\mu\nu} \partial_\mu \pi \partial_\nu \pi - \mathcal{L}_m \right]$$

GR

Linear coupling

Standard Galileon

Derivative coupling

Coupled Galileons ruled
~out by Appleby & Linder 1112.1981
due to instabilities.



Expansion & Gravity



Solve for background expansion and for linear perturbations – field evolution and gravity evolution.



$$\bar{\nabla}^2 \phi = \frac{4\pi a^2 G_{\text{eff}}^{(\phi)} \rho_m}{H_0^2} \delta_m$$

$$\bar{\nabla}^2 \psi = \frac{4\pi a^2 G_{\text{eff}}^{(\psi)} \rho_m}{H_0^2} \delta_m$$

$$\bar{\nabla}^2 (\psi + \phi) = \frac{8\pi a^2 G_{\text{eff}}^{(\psi+\phi)} \rho_m}{H_0^2} \delta_m$$

G_{matter}

G_{light}

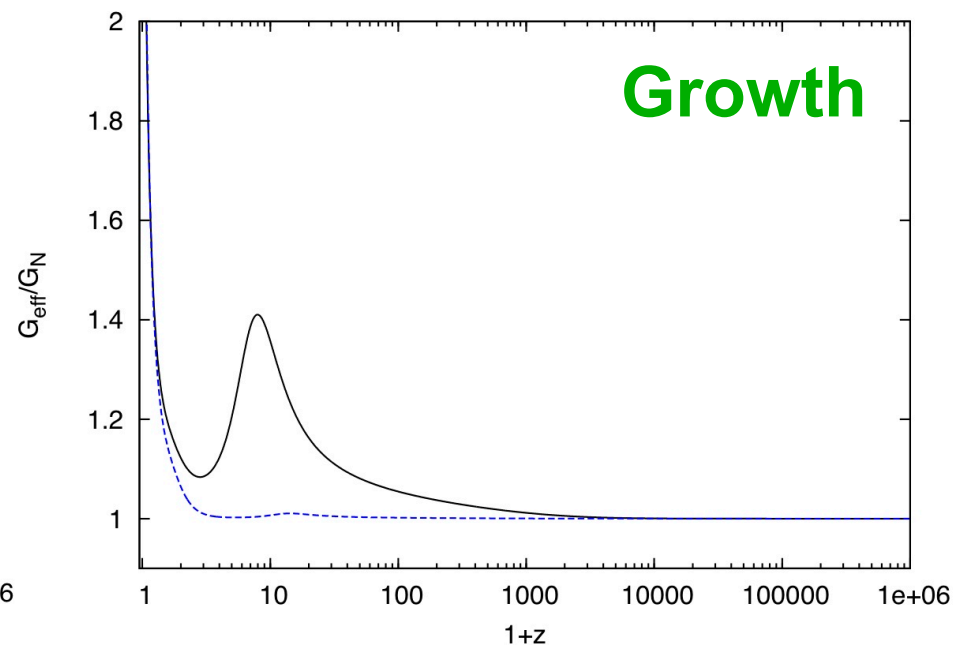
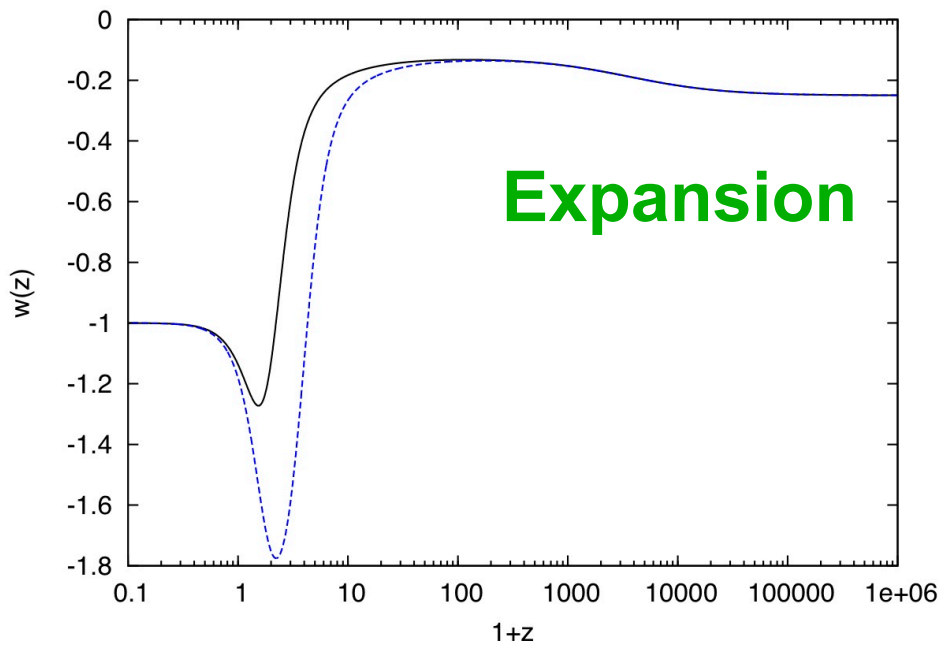
Modified Poisson equations. Can study “paths of gravity” evolution of $G(a)$.

Theory constrained by no-ghost condition and stability $c_s^2 > 0$.

Data vs Gravity



Galileon cosmology has early time tracker solutions (no fine tuning) and late time de Sitter attractor ($s_{\text{lip}}=0$). Beautiful class of theories!



But Appleby & Linder 1204.4314 rule out Standard Galileon with $\Delta\chi^2_{\text{LCDM}} > 30$ from current data. Data kill entire class of gravity!

Weirder Gravity (and Λ)



From Horndeski general scalar-tensor theory,
Charmousis+ 2011 found “Fab 4” unique self tuning terms.
Appleby, De Felice, Linder 2012 promote to nonlinear, mixed function.

$$f(c_2 g^{\mu\nu} \phi_\mu \phi_\nu + c_G G^{\mu\nu} \phi_\mu \phi_\nu)$$

“Fab 5 Freddy”

Noncanonical, nonlinear kinetic “hip-hop” gravity



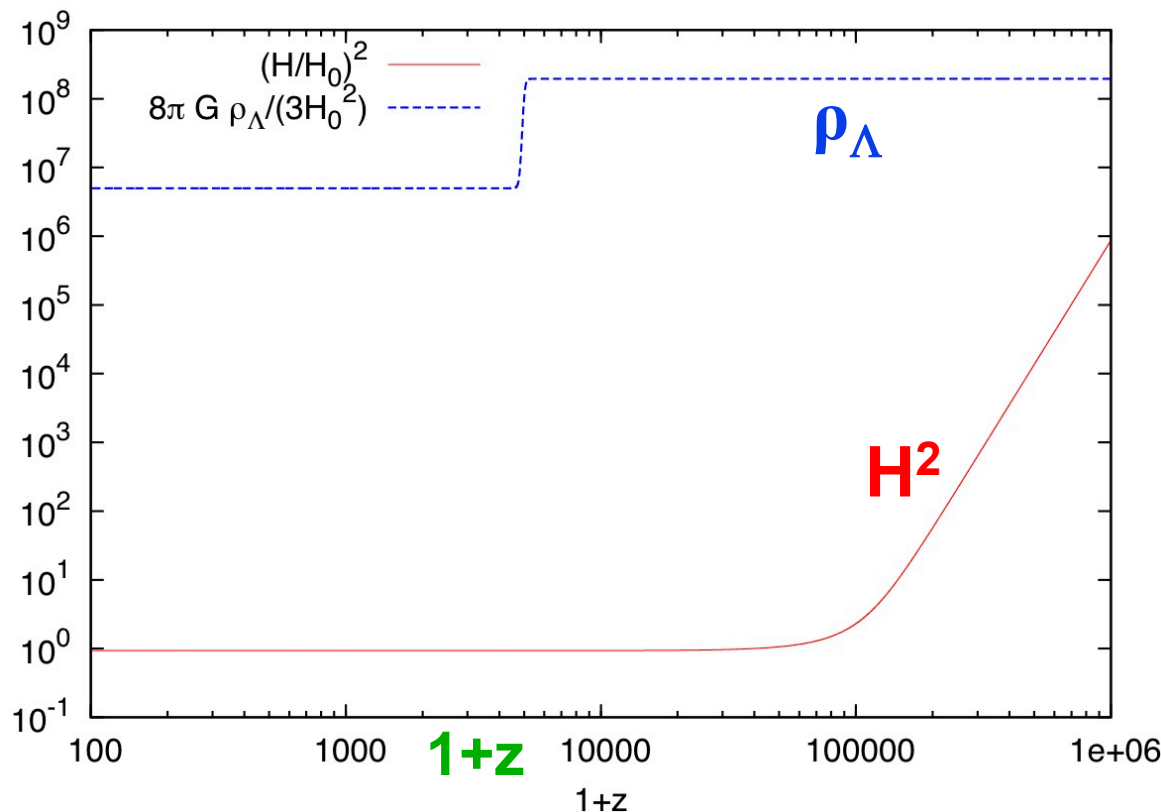
“Fab 5 Freddy told me everybody’s fly” – Blondie, *Rapture*

Fab 5 Freddy



Fab 5 Freddy indeed accelerates, and has tracker, dS attractor, no extra dof! – and **self tuning**.

ϕ dynamically adjusts to cancel Λ , even thru phase transition.



Λ is invisible!



Summary



2D to 3D mapping of cosmic structure is major advance. Galaxy redshift + CMB lensing: $\sigma(\gamma)=0.026$

Measure growth history. Comparison with expansion history opens window on gravity physics. $w(a)$ alone not enough (especially if $w \sim -1$): G_{matter} , G_{light} .
Don't fix w_0 , w_a , GR, m_ν etc!

Data already powerful enough to put Galileon gravity on trial.

Model independent approach: 2 x 2 x 2 gravity. 5-10% measures possible with next generation.

Lots of interesting theory still to explore!