

IPN Orsay contributions to LHAASO

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and

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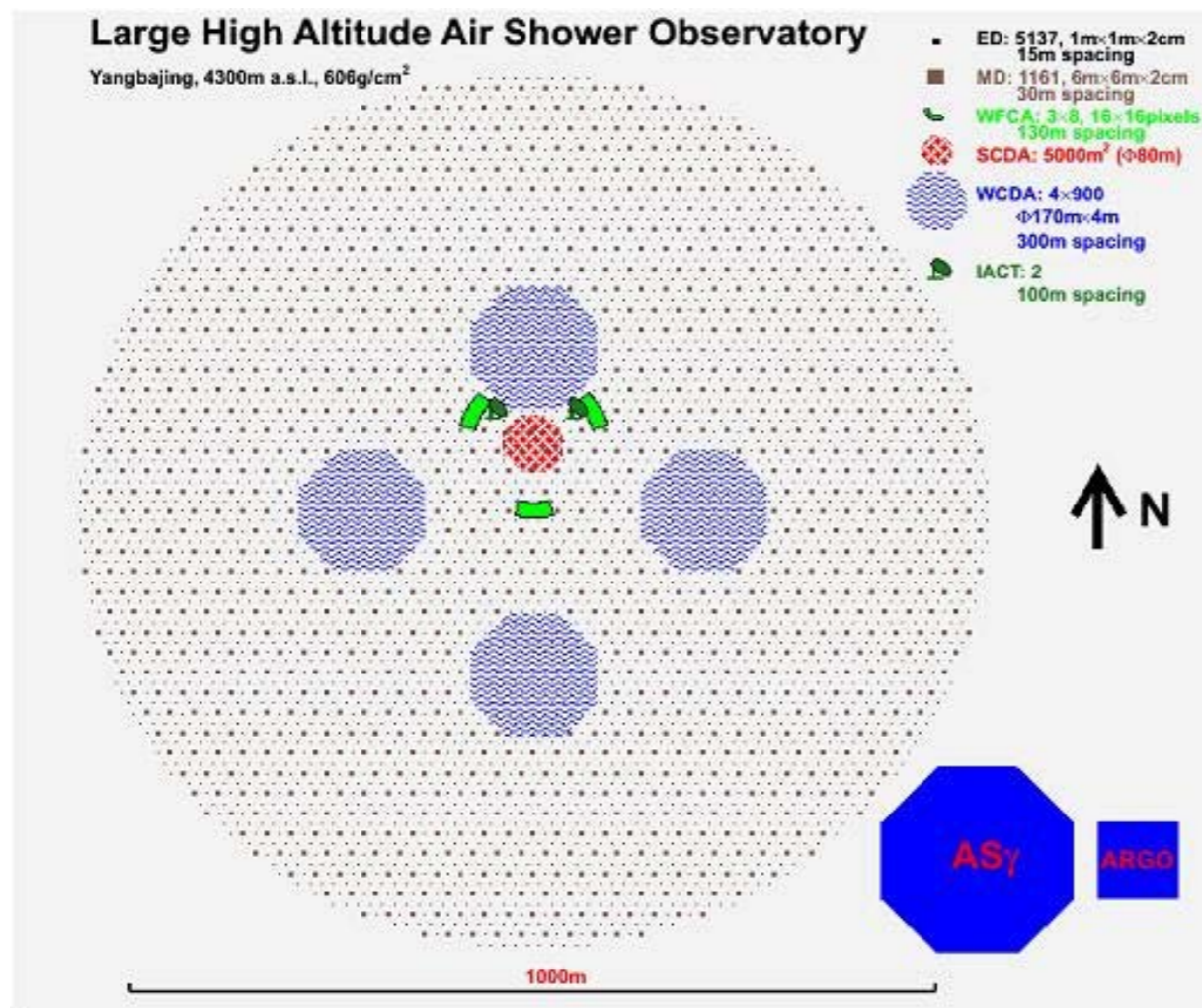
Omega 

+ Co-PhD (soon) : Yingtao Chen (Yunnan University)

ASIC : high integration, small size and low power consumption

Benefit from the R&D for next generation neutrino experiments : PARISROC

POSSIBLE USE OF PARISROCx INTEGRATED CIRCUIT FOR SOME DETECTORS FEE



KM2A : scintillator detectors

WCDA : Cerenkov in water Detectors

WFCDA : Cerenkov/fluorescence telescopes

PARISROC 2 PMT gain 10^6
Input signal: negative
Single channel hit rate: 5 kHz
Charge dynamic range: 1- 600 p.e. (100 pC)
Resolution: $<20\%$ @1pe
Disc threshold: 0.3 pe (50 fC)
Timing resolution: 220 ps
Timing precision < 1 ns

Requirements	PARISROC 2
KM2A (KM^2 Array) (ED: electron detector; MD: muon detector) PMT: ET9903B or Hamamatsu R11102 (for ED) (1.5-in) ($\text{GAIN } 10^6$) Hamamatsu R5912 (for MD) (8-in) ($\text{GAIN } 1 \times 10^7$)	
PMT high voltage: negative	OK
Single channel hit rate: 1kHz (for ED); 10kHz (for MD)	OK
Charge dynamic range: 1-3000 electron (for ED); 1-3000muon (for MD)	OK if.....
Resolution: $<25\%$ @1e, $<5\%$ @3000e (for ED); $<25\%$ @1muon, $<5\%$ @3000muon (for MD)	OK max 0.2 p.e. (32 fC) for 10-bit ADC
Disc threshold: 0.25e (for ED) (1 e 160 fC) 0.25muon (for MD) (1 e 160 fC) (1 e 1.6 pC)	OK if... OK
Timing resolution (RMS): $<1\text{ns}$ (for ED), $<10\text{ns}$ (for MD)	OK
Multi-hit resolution: 30ns	To be measured

Requirements	PARISROC 2
WCDA (Water Cerenkov detectors) PMT: XP3062 (8-in) ($\text{GAIN } 1 \times 10^7$)	
PMT high voltage: positive HV	NO (New version of PARISROC)
Single channel hit rate: 50kHz	NO (New version of PARISROC) 80 MHz read out frequency and ADC 8 bits to achieve the 50kHz hit rate
Charge dynamic range: 1pe-4000pe (1 e 1.6 pC)	OK
Shaping width: $< 350\text{ns}$	OK
Disc threshold: 0.25pe (1 e 1.6 pC)	OK
Timing Precision (RMS): $<0.5\text{ns}$	OK
Multi-hit resolution: 25ns	To be measured

Modifications of PARISROC2 needed to achieve the LHAASO sub-detectors requirements

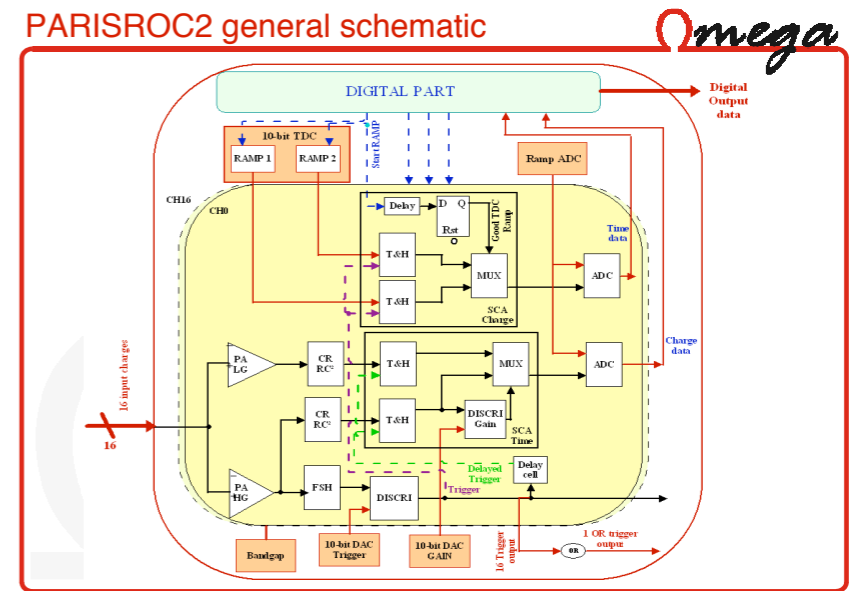
Requirements	PARISROC 2
WFCA (Wide Field Cerenkov Telescopes Array) PMT: Hamamatsu R5912 (1.5-in) ($\text{GAIN } 2.6 \times 10^5$)	
PMT high voltage: negative	OK
Single channel hit rate: $<5\text{kHz}$ (10kHz)	OK (New version)
Charge dynamic range: 1pe-3000pe (1 e 40 fC)	OK if... (New version)
Resolution: $<50\%$ @1pe, $<1\%$ @3000pe	OK if... (New version)
Disc threshold can be adjusted: 0.25pe -- 100pe	OK if... (New version)
Timing resolution (RMS): $<100\text{ns}$	OK

→ More details in Gisèle Martin-Chassard's talk tomorrow

Electronic readout tests

→ Training tests at IPNO next month

(new board assembled)

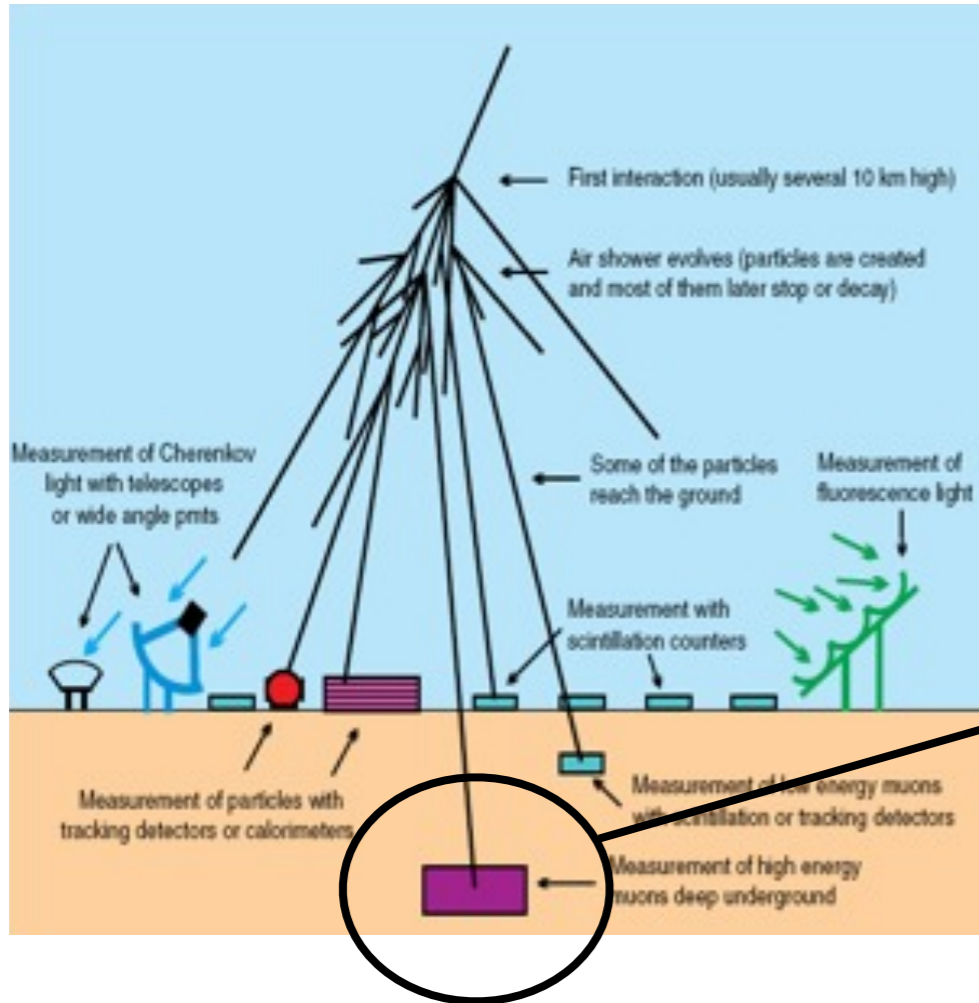


Signal acquisition
on computers

→ Tests on prototype detectors in Tibet (4300m a.s.l.)

in a few months

Response of muon detectors

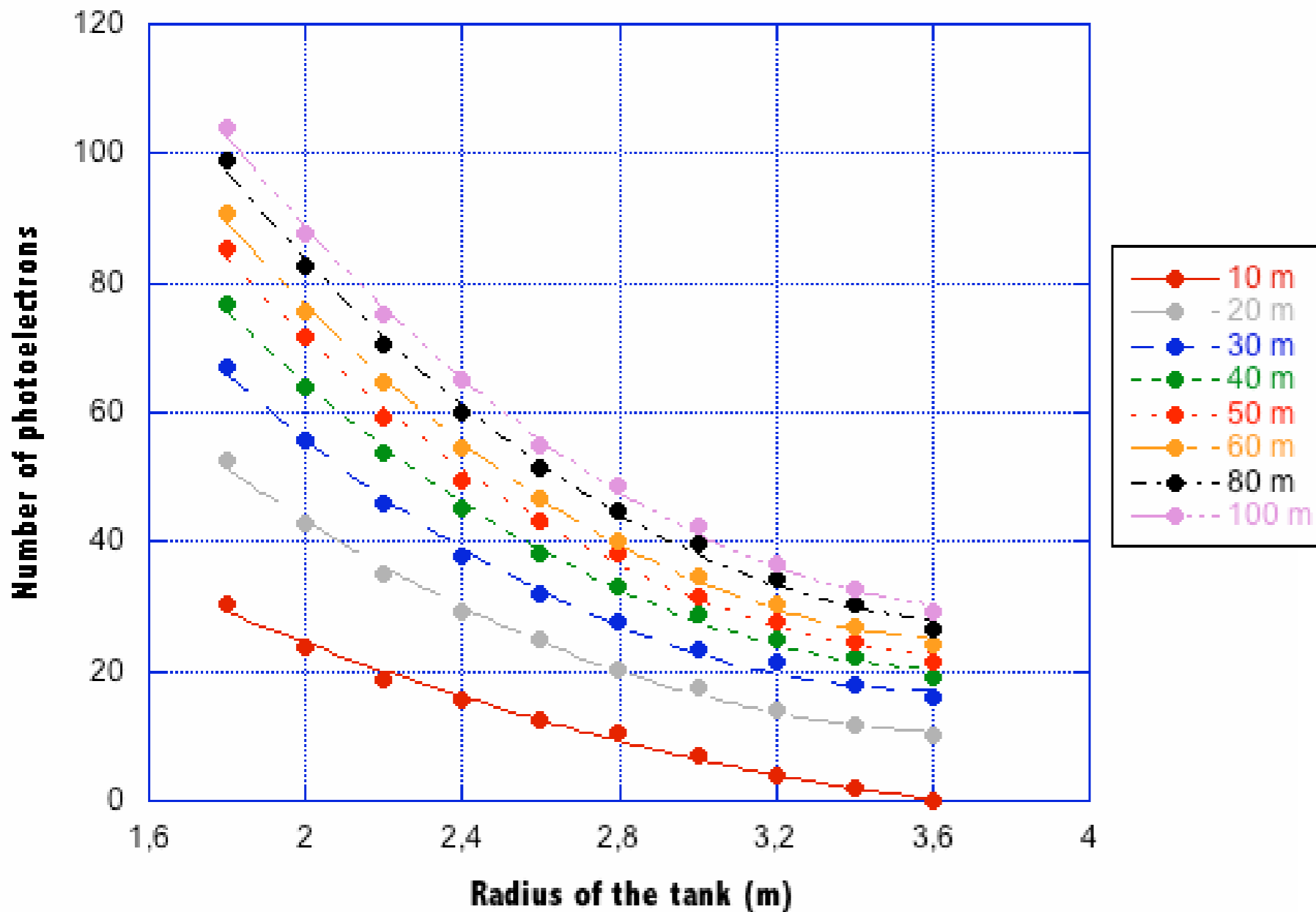


Response to \sim GeV-muons depends on :

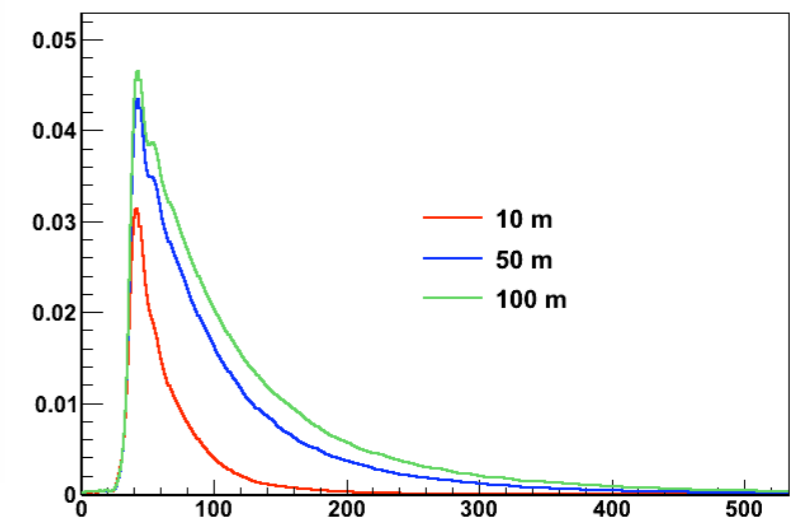
- Water transparency
- Tyvek diffusive reflectivity

Response of muon detectors

Number of photoelectrons according to the tank sizes and the water absorption length
 - Muons all directions -



Profiles of PMT for different water absorption lengths (radius 3.6 m)



Signals measured with good accuracy

Workshop organization

~50 participants (from both HE gamma ray and cosmic ray physics)



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THIRD WORKSHOP FOR AIR SHOWER DETECTION AT HIGH ALTITUDES

Institut de Physique Nucléaire d'Orsay

October 6-7, 2011

In the recent years, the interest on astroparticle physics experiments using air shower detection techniques at high altitudes has strongly increased. The ARGO-YBJ experiment has been publishing exciting observational results, the HAWC water tank array and muon detectors in the ASy array are currently under construction and the R&D for the LHAASO project is already well advanced.

To discuss the current results and the design for new experiments, we will hold a 3rd Workshop for Air Shower Detection at High Altitudes



Outlook

- *Electronic readout developments*

- *Simulation of the whole KM2A*
- Angular/Energy resolution studies
- Gamma/hadrons discrimination
- Mass discrimination (for CRs)

- *Simulation of the WCDA*
- Angular/Energy resolution studies
- Gamma/hadrons discrimination