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TheHiggs sectoroftheμνSSM

Outline

Theoretical and phenomenological motivations

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The µvSSM: Phenomenology

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Theoretical and phenomenological motivations

The µvSSM: Phenomenology



The MSSM

$$W = \epsilon_{ij} \left(Y_u H_2^j Q^i u + Y_d H_1^i Q^j d + Y_e H_1^i L^j e - \mu H_1^i H_2^j \right)$$

Natural value for $\mu \sim M_{
m Planck}$

$$-\mathcal{L}_{\text{soft}} = m_{H_1}^2 H_1^* H_1 + m_{H_2}^2 H_2^* H_2 + \dots - \epsilon_{ij} (B \mu H_1^i H_2^j + H.C.) - \frac{1}{2} (M_3 \lambda_3 \lambda_3 + M_2 \lambda_2 \lambda_2 + M_1 \lambda_1 \lambda_1 + \text{H.c.})$$

 $M_{\rm soft} \sim 1 \, {\rm TeV}$ For Correct EW and without Hierarchy problem

Correct EW symmetry breaking $V_{
m Higgs} \Rightarrow \mu \lesssim 1$ TeV $\Rightarrow \mu$ -problem

Respect to LHC

The MSSM is an R-parity conserving model

the LSP is stable

missing energy is expected at LHC.

Goods of the MSSM:

Dark matter candidate:

Neutralino (or gravitino)

Solution to the hierarchy problem

Problems:

 μ -problem

Massless Neutrinos

This motivate to postulate a New Supersymmetric Model, with the minimal natural content of fields and without μ-problem

The μ from ν Supersymmetric Standard Model ($\mu\nu$ SSM)

We use the right-handed neutrinos superfields in order the give mass to the neutrinos without µ-problem

As a consequence R-parity is explicitly broken

L-F and Carlos Muñoz, Phys. Rev. Lett. B 97 (2006) 041801 [arXiv:hep-ph/0508297]

SUSY renormalizable theory: μνSSM

$$\begin{split} W &= \epsilon_{ab} \left(Y_{u}^{ij} \,\hat{H}_{2}^{b} \,\hat{Q}_{i}^{a} \,\hat{u}_{j}^{c} + Y_{d}^{ij} \,\hat{H}_{1}^{a} \,\hat{Q}_{i}^{b} \,\hat{d}_{j}^{c} + Y_{e}^{ij} \,\hat{H}_{1}^{a} \,\hat{L}_{i}^{b} \,\hat{e}_{j}^{c} + Y_{\nu}^{ij} \,\hat{H}_{2}^{b} \,\hat{L}_{i}^{a} \,\hat{\nu}_{j}^{c} \right) \\ &- \epsilon_{ab} \lambda^{i} \,\hat{\nu}_{i}^{c} \,\hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \kappa^{ijk} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c} \,, \end{split}$$

+ SOFT TERMS



R-parity it is not symmetry of the model

SUSY renormalizable theory: µvSSM

$$\begin{split} W &= \epsilon_{ab} \left(Y_{u}^{ij} \,\hat{H}_{2}^{b} \,\hat{Q}_{i}^{a} \,\hat{u}_{j}^{c} + Y_{d}^{ij} \,\hat{H}_{1}^{a} \,\hat{Q}_{i}^{b} \,\hat{d}_{j}^{c} + Y_{e}^{ij} \,\hat{H}_{1}^{a} \,\hat{L}_{i}^{b} \,\hat{e}_{j}^{c} + Y_{\nu}^{ij} \,\hat{H}_{2}^{b} \,\hat{L}_{i}^{a} \,\hat{\nu}_{j}^{c} \right) \\ &- \epsilon_{ab} \lambda^{i} \,\hat{\nu}_{i}^{c} \,\hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \kappa^{ijk} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c} \,, \end{split}$$

+ SOFT TERMS



All known particles + SUSY partners (including neutrinos physics)

SUSY renormalizable theory: μνSSM

$$W = \epsilon_{ab} \left(Y_{u}^{ij} \hat{H}_{2}^{b} \hat{Q}_{i}^{a} \hat{u}_{j}^{c} + Y_{d}^{ij} \hat{H}_{1}^{a} \hat{Q}_{i}^{b} \hat{d}_{j}^{c} + Y_{e}^{ij} \hat{H}_{1}^{a} \hat{L}_{i}^{b} \hat{e}_{j}^{c} + Y_{\nu}^{ij} \hat{H}_{2}^{b} \hat{L}_{i}^{a} \hat{\nu}_{j}^{c} \right) - \epsilon_{ab} \frac{\lambda^{i}}{\mu^{i}} \hat{\nu}_{i}^{c} \hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \frac{\kappa^{ijk}}{\nu^{i}} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c},$$

Effective µ-term Effective Majorana Mass

+ SOFT TERMS



All known particles + SUSY partners (including neutrinos physics)

SUSY renormalizable theory: μνSSM

$$\begin{split} W &= \epsilon_{ab} \left(Y_{u}^{ij} \,\hat{H}_{2}^{b} \,\hat{Q}_{i}^{a} \,\hat{u}_{j}^{c} + Y_{d}^{ij} \,\hat{H}_{1}^{a} \,\hat{Q}_{i}^{b} \,\hat{d}_{j}^{c} + Y_{e}^{ij} \,\hat{H}_{1}^{a} \,\hat{L}_{i}^{b} \,\hat{e}_{j}^{c} + Y_{\nu}^{ij} \,\hat{H}_{2}^{b} \,\hat{L}_{i}^{a} \,\hat{\nu}_{j}^{c} \right) \\ &- \epsilon_{ab} \lambda^{i} \,\hat{\nu}_{i}^{c} \,\hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \kappa^{ijk} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c} \,, \end{split}$$



Only one scale: The SUSY breaking scale, only source of gauge breaking

SUSY renormalizable theory: µvSSM

$$\begin{split} W &= \epsilon_{ab} \left(Y_{u}^{ij} \,\hat{H}_{2}^{b} \,\hat{Q}_{i}^{a} \,\hat{u}_{j}^{c} + Y_{d}^{ij} \,\hat{H}_{1}^{a} \,\hat{Q}_{i}^{b} \,\hat{d}_{j}^{c} + Y_{e}^{ij} \,\hat{H}_{1}^{a} \,\hat{L}_{i}^{b} \,\hat{e}_{j}^{c} + Y_{\nu}^{ij} \,\hat{H}_{2}^{b} \,\hat{L}_{i}^{a} \,\hat{\nu}_{j}^{c} \right) \\ &- \epsilon_{ab} \lambda^{i} \,\hat{\nu}_{i}^{c} \,\hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \kappa^{ijk} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c} \,, \end{split}$$

+ SOFT TERMS



The model "behaves properly" up to Planck scale.

Soft Terms

$$-\mathcal{L}_{\text{soft}} = m_{H_{1}}^{2} H_{1}^{a*} H_{1}^{a} + m_{H_{2}}^{2} H_{2}^{a*} H_{2}^{a} + (m_{\tilde{\nu}^{c}}^{2})^{ij} \tilde{\nu}_{i}^{c*} \tilde{\nu}_{j}^{c} + \dots + \left[\epsilon_{ab} (A_{\nu} Y_{\nu})^{ij} H_{2}^{b} \tilde{L}_{i}^{a} \tilde{\nu}_{j}^{c} + \dots + \frac{1}{3} (A_{\kappa} \kappa)^{ijk} \tilde{\nu}_{i}^{c} \tilde{\nu}_{j}^{c} \tilde{\nu}_{k}^{c} + \text{H.c.} \right] \\ - \frac{1}{2} \left(M_{3} \tilde{\lambda}_{3} \tilde{\lambda}_{3} + M_{2} \tilde{\lambda}_{2} \tilde{\lambda}_{2} + M_{1} \tilde{\lambda}_{1} \tilde{\lambda}_{1} + \text{H.c.} \right).$$

Are given by SUGRA



Phenomenology of the *µvSSM*

Since R-parity it is not symmetry of the model the phenomenology is very different to the one of the MSSM

Neutrinos mix with neutralino (10 X 10) 3 light neutrinos (mainly neutrinos left-handed) + 7 mainly neutralinos (including neutrinos right-handed)

Neutral Higgs mix with sneutrinos (8 X 8)
 8 CP even Higgses (5 + 3 mainly sneutrinos left-handed)

7 CP odd Higgses (4 + 3 mainly sneutrinos left-handed)

The Neutralino-Neutrino mass matrix is:

$$\mathcal{M}_n = \begin{pmatrix} M & m \\ m^T & 0_{3\times 3} \end{pmatrix},$$

$$M = \begin{pmatrix} M_1 & 0 & -Av_d & Av_u & 0 & 0 & 0 \\ 0 & M_2 & Bv_d & -Bv_u & 0 & 0 & 0 \\ -Av_d & Bv_d & 0 & -\lambda_i \nu_i^c & -\lambda_1 v_u & -\lambda_2 v_u & -\lambda_3 v_u \\ Av_u & -Bv_u & -\lambda_i \nu_i^c & 0 & -\lambda_1 v_d + Y_{\nu_{i1}} \nu_i & -\lambda_2 v_d + Y_{\nu_{i2}} \nu_i & -\lambda_3 v_d + Y_{\nu_{i3}} \nu_i \\ 0 & 0 & -\lambda_1 v_u & -\lambda_1 v_d + Y_{\nu_{i1}} \nu_i & 2\kappa_{11j} \nu_j^c & 2\kappa_{12j} \nu_j^c & 2\kappa_{13j} \nu_j^c \\ 0 & 0 & -\lambda_2 v_u & -\lambda_2 v_d + Y_{\nu_{i2}} \nu_i & 2\kappa_{21j} \nu_j^c & 2\kappa_{22j} \nu_j^c & 2\kappa_{23j} \nu_j^c \\ 0 & 0 & -\lambda_3 v_u & -\lambda_3 v_d + Y_{\nu_{i3}} \nu_i & 2\kappa_{31j} \nu_j^c & 2\kappa_{32j} \nu_j^c & 2\kappa_{33j} \nu_j^c \end{pmatrix},$$

where
$$A = \frac{G}{\sqrt{2}} \sin \theta_W$$
, $B = \frac{G}{\sqrt{2}} \cos \theta_W$, and

$$m^{T} = \begin{pmatrix} -\frac{g_{1}}{\sqrt{2}}\nu_{1} & \frac{g_{2}}{\sqrt{2}}\nu_{1} & 0 & Y_{\nu_{1i}}\nu_{i}^{c} & Y_{\nu_{11}}v_{u} & Y_{\nu_{12}}v_{u} & Y_{\nu_{13}}v_{u} \\ -\frac{g_{1}}{\sqrt{2}}\nu_{2} & \frac{g_{2}}{\sqrt{2}}\nu_{2} & 0 & Y_{\nu_{2i}}\nu_{i}^{c} & Y_{\nu_{21}}v_{u} & Y_{\nu_{22}}v_{u} & Y_{\nu_{23}}v_{u} \\ -\frac{g_{1}}{\sqrt{2}}\nu_{3} & \frac{g_{2}}{\sqrt{2}}\nu_{3} & 0 & Y_{\nu_{3i}}\nu_{i}^{c} & Y_{\nu_{31}}v_{u} & Y_{\nu_{32}}v_{u} & Y_{\nu_{33}}v_{u} \end{pmatrix}$$

EW see-saw mechanism

In first approximation the light neutrinos mass matrix is:

$$M_{\nu} = m^T M^{-1} m$$

With neutrino masses of order 10^{-2} e

$$10^{-2} \text{ eV} = 10^{-11} \text{ GeV} \Rightarrow$$

$$10^{-11} \text{GeV} = \frac{Y_{\nu}^2 (10^2 \text{GeV})^2}{10^3 \text{GeV}} \rightarrow Y_{\nu} \sim 10^{-6}$$

The Neutralino-Neutrino mass matrix is:

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$$M = \begin{pmatrix} M_1 & 0 & -Av_d & Av_u \\ 0 & M_2 & Bv_d & -Bv_u \\ -Av_d & Bv_d & 0 & -\lambda_i \nu_i^c \\ Av_u & -Bv_u & -\lambda_i \nu_i^c & 0 \\ 0 & 0 & -\lambda_1 v_u & -\lambda_1 v_d + Y_{\nu_{i1}} \nu_i \\ 0 & 0 & -\lambda_2 v_u & -\lambda_2 v_d + Y_{\nu_{i2}} \nu_i \\ 0 & 0 & -\lambda_2 v_u & -\lambda_2 v_d + Y_{\nu_{i2}} \nu_i \\ 0 & 0 & -\lambda_3 v_u & -\lambda_3 v_d + Y_{\nu_{i3}} \nu_i \\ 0 & 0 & -\lambda_3 v_u & -\lambda_3 v_d + Y_{\nu_{i3}} \nu_i \\ 2\kappa_{31j} \nu_j^c & 2\kappa_{32j} \nu_j^c & 2\kappa_{33j} \nu_j^c \end{pmatrix},$$

where
$$A = \frac{G}{\sqrt{2}} \sin \theta_W$$
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$$m^{T} = \begin{pmatrix} -\frac{g_{1}}{\sqrt{2}}\nu_{1} & \frac{g_{2}}{\sqrt{2}}\nu_{1} & 0 & Y_{\nu_{1i}}\nu_{i}^{c} & Y_{\nu_{11}}v_{u} & Y_{\nu_{12}}v_{u} & Y_{\nu_{13}}v_{u} \\ -\frac{g_{1}}{\sqrt{2}}\nu_{2} & \frac{g_{2}}{\sqrt{2}}\nu_{2} & 0 & Y_{\nu_{2i}}\nu_{i}^{c} & Y_{\nu_{21}}v_{u} & Y_{\nu_{22}}v_{u} & Y_{\nu_{23}}v_{u} \\ -\frac{g_{1}}{\sqrt{2}}\nu_{3} & \frac{g_{2}}{\sqrt{2}}\nu_{3} & 0 & Y_{\nu_{3i}}\nu_{i}^{c} & Y_{\nu_{31}}v_{u} & Y_{\nu_{32}}v_{u} & Y_{\nu_{33}}v_{u} \end{pmatrix}$$

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where
$$A = \frac{G}{\sqrt{2}} \sin \theta_W$$
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$$M_{\nu} = m^T M^{-1} m$$

Using Diagonal Yukawas for Neutrinos

$$(m_{eff|real})_{ij} \simeq \frac{v_u^2}{6\kappa\nu^c} Y_{\nu_i} Y_{\nu_j} \left(1 - 3\,\delta_{ij}\right) - \frac{1}{2M_{eff}} \left[\nu_i \nu_j + \frac{v_d \left(Y_{\nu_i} \nu_j + Y_{\nu_j} \nu_i\right)}{3\lambda} + \frac{Y_{\nu_i} Y_{\nu_j} v_d^2}{9\lambda^2}\right]$$

$$M_{eff} \equiv M \left[1 - \frac{v^2}{2M \left(\kappa \nu^{c^2} + \lambda v_u v_d \right) \, 3\lambda \nu^c} \left(2\kappa \nu^{c^2} \frac{v_u v_d}{v^2} + \frac{\lambda v^2}{2} \right) \right] \qquad \qquad \frac{1}{M} = \frac{g_1^2}{M_1} + \frac{g_2^2}{M_2}$$

We have neglected all the terms of order $Y_{\nu}^{2}\nu^{2}$, $Y_{\nu}^{3}\nu$ and $Y_{\nu}\nu^{3}$

(Diagonal Yukawas for Neutrinos)

$$(m_{eff|real})_{ij} \simeq \frac{v_u^2}{6\kappa\nu^c} Y_{\nu_i} Y_{\nu_j} \left(1 - 3\,\delta_{ij}\right) - \frac{1}{2M_{eff}} \left[\nu_i \nu_j + \frac{v_d \left(Y_{\nu_i} \nu_j + Y_{\nu_j} \nu_i\right)}{3\lambda} + \frac{Y_{\nu_i} Y_{\nu_j} v_d^2}{9\lambda^2}\right]$$

$$M_{eff} \equiv M \left[1 - \frac{v^2}{2M \left(\kappa \nu^{c^2} + \lambda v_u v_d \right) \ 3\lambda \nu^c} \left(2\kappa \nu^{c^2} \frac{v_u v_d}{v^2} + \frac{\lambda v^2}{2} \right) \right] \qquad \frac{1}{M} = \frac{g_1^2}{M_1} + \frac{g_2^2}{M_2}$$

(Diagonal Yukawas for Neutrinos)

$$(m_{eff|real})_{ij} \simeq \frac{v_u^2}{6\kappa\nu^c} Y_{\nu_i} Y_{\nu_j} \left(1 - 3\,\delta_{ij}\right) - \frac{1}{2M_{eff}} \left[\nu_i \nu_j + \frac{v_d \left(Y_{\nu_i} \nu_j + Y_{\nu_j} \nu_i\right)}{3\lambda} + \frac{Y_{\nu_i} Y_{\nu_j} v_d^2}{9\lambda^2}\right]$$

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$$\begin{array}{l} v_d \to 0 \\ M_{\text{eff}} \approx M \end{array} \qquad \left(m_{eff|real} \right)_{ij} \simeq \frac{v_u^2}{6\kappa\nu^c} Y_{\nu_i} Y_{\nu_j} (1 - 3\delta_{ij}) - \frac{1}{2M} \nu_i \nu_j \end{array}$$

(Diagonal Yukawas for Neutrinos)

$$(m_{eff|real})_{ij} \simeq \frac{v_u^2}{6\kappa\nu^c} Y_{\nu_i} Y_{\nu_j} \left(1 - 3\,\delta_{ij}\right) - \frac{1}{2M_{eff}} \left[\nu_i \nu_j + \frac{v_d \left(Y_{\nu_i} \nu_j + Y_{\nu_j} \nu_i\right)}{3\lambda} + \frac{Y_{\nu_i} Y_{\nu_j} v_d^2}{9\lambda^2}\right]$$

$$M_{eff} \equiv M \left[1 - \frac{v^2}{2M \left(\kappa \nu^{c^2} + \lambda v_u v_d \right) \ 3\lambda \nu^c} \left(2\kappa \nu^{c^2} \frac{v_u v_d}{v^2} + \frac{\lambda v^2}{2} \right) \right] \qquad \frac{1}{M} = \frac{g_1^2}{M_1} + \frac{g_2^2}{M_2}$$

$$\begin{array}{l} v_{d} \rightarrow 0 \\ M_{\text{eff}} \approx M \end{array} \qquad (m_{eff|real})_{ij} \simeq \frac{v_{u}^{2}}{6\kappa\nu^{c}}Y_{\nu_{i}}Y_{\nu_{j}}(1-3\delta_{ij}) - \frac{1}{2M}\nu_{i}\nu_{j} \end{array}$$

$$\begin{array}{l} \textbf{Gaugino see saw} \end{array}$$

(Diagonal Yukawas for Neutrinos)

$$(m_{eff|real})_{ij} \simeq \frac{v_u^2}{6\kappa\nu^c} Y_{\nu_i} Y_{\nu_j} (1 - 3\,\delta_{ij}) - \frac{1}{2M_{eff}} \left[\nu_i \nu_j + \frac{v_d \left(Y_{\nu_i} \nu_j + Y_{\nu_j} \nu_i\right)}{3\lambda} + \frac{Y_{\nu_i} Y_{\nu_j} v_d^2}{9\lambda^2} \right]$$

$$M_{eff} \equiv M \left[1 - \frac{v^2}{2M \left(\kappa \nu^{c^2} + \lambda v_u v_d \right) \ 3\lambda \nu^c} \left(2\kappa \nu^{c^2} \frac{v_u v_d}{v^2} + \frac{\lambda v^2}{2} \right) \right] \qquad \frac{1}{M} = \frac{g_1^2}{M_1} + \frac{g_2^2}{M_2}$$

 $\begin{array}{ll} v_{d} \rightarrow 0 & (m_{eff|real})_{ij} \simeq \frac{v_{u}^{2}}{6\kappa\nu^{c}}Y_{\nu_{i}}Y_{\nu_{j}}(1-3\delta_{ij}) - \frac{1}{2M}\nu_{i}\nu_{j} \\ \mathbf{v}_{\mathsf{R}}\text{-Higgsino} & \text{see saw} \end{array}$

With diagonal Yukawas for neutrinos we could reproduce the experimental mixing angles.

In a sense we have an explanation of :

Why mixtures in Neutrino and quark sectors are so different?

LHC: Higgs signals

$$\begin{split} W &= \epsilon_{ab} \left(Y_{u}^{ij} \,\hat{H}_{2}^{b} \,\hat{Q}_{i}^{a} \,\hat{u}_{j}^{c} + Y_{d}^{ij} \,\hat{H}_{1}^{a} \,\hat{Q}_{i}^{b} \,\hat{d}_{j}^{c} + Y_{e}^{ij} \,\hat{H}_{1}^{a} \,\hat{L}_{i}^{b} \,\hat{e}_{j}^{c} + Y_{\nu}^{ij} \,\hat{H}_{2}^{b} \,\hat{L}_{i}^{a} \,\hat{\nu}_{j}^{c} \right) \\ &- \epsilon_{ab} \lambda^{i} \,\hat{\nu}_{i}^{c} \,\hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \kappa^{ijk} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c} \,, \end{split}$$

$$\begin{aligned} -\mathcal{L}_{\text{soft}} &= m_{H_1}^2 H_1^* H_1 + m_{H_2}^2 H_2^* H_2 + \dots - \epsilon_{ij} (B \mu H_1^i H_2^j + H.C.) \\ &- \frac{1}{2} (M_3 \lambda_3 \lambda_3 + M_2 \lambda_2 \lambda_2 + M_1 \lambda_1 \lambda_1 + H.c.) \\ \lambda_i &= \lambda, \ \tan \beta, \ \kappa_{iii}, \ \nu_i^c &= \nu^c, \ \nu_1, \ \nu_2 &= \nu_3, \ Y_{\nu_1}, \ Y_{\nu_2} &= Y_{\nu_2}, \ A_{\lambda}, \ A_{\kappa}, \ M_2, \end{aligned}$$

LHC: Higgs signals

$$W = \epsilon_{ab} \left(Y_{u}^{ij} \hat{H}_{2}^{b} \hat{Q}_{i}^{a} \hat{u}_{j}^{c} + Y_{d}^{ij} \hat{H}_{1}^{a} \hat{Q}_{i}^{b} \hat{d}_{j}^{c} + Y_{e}^{ij} \hat{H}_{1}^{a} \hat{L}_{i}^{b} \hat{e}_{j}^{c} + Y_{\nu}^{ij} \hat{H}_{2}^{b} \hat{L}_{i}^{a} \hat{\nu}_{j}^{c} \right)$$
$$- \epsilon_{ab} \lambda^{i} \hat{\nu}_{i}^{c} \hat{H}_{1}^{a} \hat{H}_{2}^{b} + \frac{1}{3} \kappa^{ijk} \hat{\nu}_{i}^{c} \hat{\nu}_{j}^{c} \hat{\nu}_{k}^{c},$$
$$\mathcal{L}_{acft} = m_{u}^{2} H_{1}^{*} H_{1} + m_{u}^{2} H_{2}^{*} H_{2} + \dots - \epsilon_{ii} (B \mu H_{1}^{i} H_{2}^{j} + H_{i}C_{i})$$

$$\mathcal{L}_{\text{soft}} = m_{H_1}^2 H_1^* H_1 + m_{H_2}^2 H_2^* H_2 + \dots - \epsilon_{ij} (B \mu H_1^* H_2^j + H.C.) - \frac{1}{2} (M_3 \lambda_3 \lambda_3 + M_2 \lambda_2 \lambda_2 + M_1 \lambda_1 \lambda_1 + H.c.)$$

$$\lambda_i = \lambda$$
, $\tan \beta$, κ_{iii} , $\nu_i^c = \nu^c$, ν_1 , $\nu_2 = \nu_3$, Y_{ν_1} , $Y_{\nu_2} = Y_{\nu_3}$, A_{λ} , A_{κ} , M_2 ,

Gauge unification $M_1 = \frac{\alpha_1^2}{\alpha_2^2} M_2, M_3 = \frac{\alpha_3^2}{\alpha_2^2} M_2, \qquad M_1 \approx 0.5 M_2, M_3 = 2.7 M_2.$

Lightest Higgs (tree level bound)

$$m_h^2 \leq M_Z^2 \left(\cos^2 2\beta + \frac{2\lambda^2 \cos^2 \theta_W}{g_2^2} \sin^2 2\beta \right) \approx M_Z^2 \left(\cos^2 2\beta + 3.62 \lambda^2 \sin^2 2\beta \right)$$
$$\lambda^2 = \lambda_i \lambda_i \cdot$$

Landau pole condition (GUT) $\implies \lambda^2 \leq (0.7)^2 \quad \lambda \equiv \lambda_i \leq 0.7/\sqrt{3} \approx 0.4$ for $\tan \beta = -2(4) \implies m_h \leq -111(98)$ GeV
Pure doublet $\longrightarrow \lambda_i \to 0$ or $A_{\lambda_i} = \frac{2\mu}{\sin 2\beta} - \frac{2}{\lambda_i} \sum_{j,k} \kappa_{ijk} \lambda_j \nu_k^c$

 $1 \log \longrightarrow m_h \lesssim 140 \text{ GeV}$

$h_1 \rightarrow 2$ Standard Model fermions



For example we have

 $h_3 \rightarrow 2h_2 \rightarrow 4h_1 \rightarrow 8$ Standard Model fermions



:
$$h_4 \to \tilde{\chi}^0 \tilde{\chi}^0 \to 2P2\nu \to 2\tau^+ 2\tau^- 2\nu$$
. $h_4 \to \tilde{\chi}^0 \tilde{\chi}^0 \to 2P2\nu \to 2b2\bar{b}2\nu$.

We took three light singlet-like Higgses, the four is doublet-like

Gluon Fusion dominate the Higgs production at LHC



$$\sigma(gg \to h_4) = \sigma(gg \to H_{\rm SM}) \frac{\Gamma(h_4 \to gg)}{\Gamma(H_{\rm SM} \to gg)} \simeq \sigma(gg \to H_{\rm SM})$$

 $0.75 \ \sigma(gg \to H_{\rm SM}) \lesssim \sigma(gg \to h_4) \lesssim \sigma(gg \to H_{\rm SM})$

 $\sigma(gg \rightarrow H_{\rm SM})$ is about 17 – 19.5 pb For center of mass energy of 7 TeV

Gluon Fusion dominate the Higgs production at LHC



$$\sigma(gg \to h_4) = \sigma(gg \to H_{\rm SM}) \frac{\Gamma(h_4 \to gg)}{\Gamma(H_{\rm SM} \to gg)} \simeq \sigma(gg \to H_{\rm SM})$$

 $0.75 \ \sigma(gg \to H_{\rm SM}) \lesssim \sigma(gg \to h_4) \lesssim \sigma(gg \to H_{\rm SM})$

 $\sigma(gg \rightarrow H_{\rm SM})$ is about 17 – 19.5 pb For center of mass energy of 7 TeV

Many Colliders constraints

- h → invisible [36] 37]. Here we are assuming as invisible the light neutrinos. A more elaborated analysis requires a re-analysis of LEP data, taking into account for instance that neutralinos could partially contribute to the missing energy when the decay distance is comparable to the size of the detector. We have checked that in the points where the decay length of the lightest neutralino is considerably greater than O(1 m), considering also the LSP as invisible, the constraint is satisfied.
- h → γγ, from LEP Higgs working group results [19].
- h → bb, from the LEP Higgs working group [21].
- h to two jets, from OPAL and the LEP Higgs working group, both at LEP2 [38, 39].
- $h \to \tau^+ \tau^-$, from the LEP Higgs working group [21].
- h → PP with PP decaying to 4 jets, 2 jets + cc, 2 jets + τ⁺τ⁻, 4 τ's, cccc, ττ + cc, from OPAL results [26].

For e⁺e⁻ → hP with hP decaying into 4 b, 4 τ, and PPP → 6b studied by DELPHI
 [23].

3) For $e^+e^- \rightarrow hZ \rightarrow PPZ \rightarrow 4b + 2jets$ the DELPHI constraints [23].

4) For e⁺e⁻ → hZ independent of h decay mode, combining the results of ALEPH and OPAL collaborations [36, 21].

One of the most important LEP constraints



G. Abbiendi et al., LEP Working Group for Higgs Boson Searches, Phys. Lett. B565 (2003) 61 [arXiv:hep-ex/0306033].



CMS PAS HIG-11-022



Figure 21: The combined 95% C.L. upper limits on the signal strength modifier $\mu = \sigma / \sigma_{sM}$, as a function of the SM Higgs boson mass in the range 110-600 GeV/ c^2 . The observed limits are shown by the solid symbols and the black line. The dashed line indicates the median expected limit on μ for the background-only hypothesis, while the green/yellow bands indicate the ranges that are expected to contain 68%/95% of all observed limit excursions from the median.

CMS PAS HIG-11-032

LHC: CMS



Figure 2: The combined 95% C.L. upper limits on the signal strength modifier $\mu = \sigma/\sigma_{SM}$ as a function of the SM Higgs boson mass in the range 110–600 GeV/ c^2 . The observed limits are shown by the solid symbols and the black line. The dashed line indicates the median expected limit on μ for the background-only hypothesis, while the green (yellow) bands indicate the ranges that are expected to contain 68% (95%) of all observed limit excursions from the median.

From Sandra Kortner talk at Recontres Moriond 2012



Fidalgo, L-F, Munoz, Ruiz de Austri, 2011

Benchmark point	Cascade	$\sigma(gg \to h_4) \times BR_{\text{cascade}} \text{ (fb)}$	
1	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2P 2\nu \rightarrow 2b 2\overline{b} 2\nu$	270	
	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2h2\nu \rightarrow 4P2\nu \rightarrow 4b4\overline{b}2\nu$	44	
2	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2P 2\nu \rightarrow 2\tau^+ 2\tau^- 2\nu$	1620	
3	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2P 2\nu \rightarrow 2b 2\overline{b} 2\nu$	70	
4	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2P 2\nu \rightarrow 2b 2\overline{b} 2\nu$	5860	
5	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P 2\nu \rightarrow 2b 2\overline{b} 2\nu$	4870	
6	$h_1 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2l2q2\bar{q}$	150	
	$h_1 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2\nu 2l2\overline{l}$	80	
	$h_1 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2\nu 2q 2\bar{q}$	40	
	$h_1 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 6\nu$	15	
7	$h_4 \rightarrow 2P \rightarrow 2b2\overline{b}$	5450	
	$h_4 \rightarrow 2h_1 \rightarrow 4P \rightarrow 4b4\overline{b}$	460	
8	$h_4 \rightarrow 2P_3 \rightarrow 2b2\overline{b}$	1660	
	$h_4 \rightarrow h_1 h_1 \rightarrow 4 P_{1,2} \rightarrow 4 \tau^+ 4 \tau^-$	460	
	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P_{1,2} 2\nu \rightarrow 2\tau^+ 2\tau^- 2\nu$	80	
	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2h2\nu \rightarrow 4P_{1,2}2\nu \rightarrow 4\tau^+ 4\tau^- 2\nu$	150	
	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P_3 2\nu \rightarrow 2b 2\overline{b} 2\nu$	20	

Table 9: Production cross section multiplied by branching ratios of the cascades, for the benchmark points discussed in the text.

Fidalgo, L-F, Munoz, Ruiz de Austri, 2011

Benchmark point	Cascade	$\sigma(gg \to h_4) \times BR_{\text{cascade}} \text{ (fb)}$	
1	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P 2\nu \rightarrow 2b 2\overline{b} 2\nu$	270	
	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2h2\nu \rightarrow 4P2\nu \rightarrow 4b4\overline{b}2\nu$	44	
2	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P 2\nu \rightarrow 2\tau^+ 2\tau^- 2\nu$	1620	
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5	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P 2\nu \rightarrow 2b 2\overline{b} 2\nu$	4870	
6	$h_1 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2l2q2\bar{q}$	150	
	$h_1 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2\nu 2l2\overline{l}$	80	
	$h_1 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2\nu 2q 2\bar{q}$	40	
	$h_1 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 6\nu$	15	
7	$h_4 \rightarrow 2P \rightarrow 2b2\overline{b}$	5450	
	$h_4 \rightarrow 2h_1 \rightarrow 4P \rightarrow 4b4\overline{b}$	460	
8	$h_4 \rightarrow 2P_3 \rightarrow 2b2\overline{b}$	1660	-
	$h_4 \rightarrow h_1 h_1 \rightarrow 4 P_{1,2} \rightarrow 4 \tau^+ 4 \tau^-$	460	
	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P_{1,2} 2\nu \rightarrow 2\tau^+ 2\tau^- 2\nu$	80	
	$h_4 \rightarrow \tilde{\chi}^0_4 \tilde{\chi}^0_4 \rightarrow 2h2\nu \rightarrow 4P_{1,2}2\nu \rightarrow 4\tau^+ 4\tau^- 2\nu$	150	
	$h_4 \rightarrow \tilde{\chi}_4^0 \tilde{\chi}_4^0 \rightarrow 2P_3 2\nu \rightarrow 2b2\overline{b}2\nu$	20	

Table 9: Production cross section multiplied by branching ratios of the cascades, for the benchmark points discussed in the text.

λ	κ ₁₁₁	К ₂₂₂	κ_{333}	A_{κ} (GeV)	M_2 (GeV)
1.0×10^{-1}	7.7×10^{-3}	7.5×10^{-3}	7.3×10^{-3}	-1.0	-1.7×10^3
an eta	A_{λ} (GeV)	ν_1 (GeV)	$\nu_{2,3}(\text{GeV})$	Y_{ν_1}	$Y_{\nu_{2,3}}$
3.7	1.0×10^3	2.92×10^{-5}	1.46×10^{-4}	2.70×10^{-8}	1.51×10^{-7}
ν^{c} (GeV)	m_{h_1} (GeV)	m_{h_2} (GeV)	m_{h_3} (GeV)	m_{h_4} (GeV)	m_{P_1} (GeV)
$8.0 imes 10^2$	13.6	13.9	17.0	116.2	8.4
m_{P_2} (GeV)	m_{P_3} (GeV)	$m_{\tilde{\chi}_{\pm}^{0}}$ (GeV)	$m_{\tilde{\chi}_{5}^{0}}$ (GeV)	$m_{\tilde{\chi}^0_{B}}$ (GeV)	-
9.5	9.6	11.8	12.2	14.0	-
$BR(h_4 \rightarrow \sum_{i,j=4}^{6} \tilde{\chi}_i^0 \tilde{\chi}_j^0)$	$BR(\tilde{\chi}_4^0 \rightarrow \sum_{i=1}^{3} P_i \nu)$	$BR(P_1 \rightarrow \tau^+ \tau^-)$	$BR(P_2 \rightarrow \tau^+ \tau^-)$	$BR(P_3 \rightarrow \tau^+ \tau^-)$	$l_{\tilde{\chi}_{4}^{0}\rightarrow}$ (cm)
0.12	1.0	0.89	0.83	0.82	189

benchmark point 2

λ	ĸ111	K222	K333	A_{κ} (GeV)	M_2 (GeV)
1.0×10^{-1}	1.66×10^{-2}	1.65×10^{-2}	1.64×10^{-2}	-5.0	-1.7×10^{3}
$\tan \beta$	A_{λ} (GeV)	ν_1 (GeV)	$\nu_{2,3}(\text{GeV})$	Y_{ν_1}	$Y_{\nu_{2,3}}$
4.9	1.0×10^{3}	5.84×10^{-5}	2.25×10^{-4}	1.25×10^{-7}	2.26×10^{-7}
ν^{c} (GeV)	m_{h_1} (GeV)	m_{h_2} (GeV)	m_{h_3} (GeV)	m_{h_4} (GeV)	m_{P_1} (GeV)
8.0×10^{2}	19.8	21.6	21.8	120.2	8.8
m_{P_2} (GeV)	m_{P_3} (GeV)	$m_{\tilde{\chi}_{4}^{0}}$ (GeV)	$m_{\tilde{\chi}_{5}^{0}}$ (GeV)	$m_{\tilde{\chi}_8^0}$ (GeV)	—
8.9	16.9	26.3	26.5	27.8	—
$BR(h_4 \rightarrow h_1h_1)$	$BR(h_4 \rightarrow P_3P_3)$	$BR(h_4 \rightarrow \sum_{i,j=4}^{6} \bar{\chi}_i^0 \bar{\chi}_j^0)$	$BR(h_4 \rightarrow b\bar{b})$	$BR(h_1 \rightarrow \sum_{i,j=1}^2 P_i P_j)$	$BR(h_{2,3} \rightarrow \sum_{i,j=1}^{2} P_i P_j)$
0.05	0.12	0.06	0.55	0.98	1.0
$BR(P_{1,2} \rightarrow \tau^+\tau^-)$	$BR(P_3 \rightarrow b\bar{b})$	$BR(\tilde{\chi}_4^0 \rightarrow \sum_{i=1}^3 h_i \nu)$	$BR(\bar{\chi}_4^0 \rightarrow P_{1,2}\nu)$	$BR(\tilde{\chi}_4^0 \rightarrow P_3\nu)$	$l_{\tilde{\chi}_4^0}$ (cm)
0.88	0.93	0.51	0.33	0.16	15

Possible signals at LHC

Displaced vertices

multi-jets plus missing energy

multi-leptons plus missing energy

SUMMARY

The µvSSM includes right-handed neutrino superfields
 to solve the µ-problem of the MSSM generating also the masses for the neutrinos. The only source for EW breaking are the soft terms

After EW symmetry breaking → see saw mechanism at
 W scale (is possible to reproduce experiments with diagonal Yukawas for the leptonic sector)

- Gravitino can be the Dark Matter giving visible signals Fermi satellite
- LHC the Higgs phenomenology is very rich.
 Multileptons and/or multijets decays are possible

Higgs Prodution by: quark quark fusion

$$\sigma(b\bar{b} \to h_4) = \sigma(b\bar{b} \to H_{\rm SM}) \left(\frac{Y_{bbh_4}}{Y_{bbH_{\rm SM}}}\right)^2 = \sigma(b\bar{b} \to H_{\rm SM}) \frac{S^2(d,4)}{\cos^2\beta}$$

Neutrino Physics

 $7.14 < \Delta m_{sol}^2 / 10^{-5} \text{ eV}^2 < 8.19 \text{ , } 2.06 < \Delta m_{atm}^2 / 10^{-3} \text{ eV}^2 < 2.81$ $0.263 < \sin^2 \theta_{12} < 0.375 \text{ , } \sin^2 \theta_{13} < 0.046 \text{ , } 0.331 < \sin^2 \theta_{23} < 0.644$

• Charginos mix with charged leptons

 $\Psi^{+T} = (-i\tilde{\lambda}^{+}, \tilde{H}_{u}^{+}, e_{R}^{+}, \mu_{R}^{+}, \tau_{R}^{+}) \qquad \qquad \Psi^{-T} = (-i\tilde{\lambda}^{-}, \tilde{H}_{d}^{-}, e_{L}^{-}, \mu_{L}^{-}, \tau_{L}^{-})$

$$-\frac{1}{2}(\psi^{+T},\psi^{-T})\begin{pmatrix}0&M_C^T\\M_C&0\end{pmatrix}\begin{pmatrix}\psi^{+T}\\\psi^{-T}\end{pmatrix}$$

$$M_{C} = \begin{pmatrix} M_{2} & g_{2}v_{u} & 0 & 0 & 0 \\ g_{2}v_{d} & \lambda_{i}\nu_{i}^{c} & -Y_{e_{i1}}\nu_{i} & -Y_{e_{i2}}\nu_{i} & -Y_{e_{i3}}\nu_{i} \\ g_{2}\nu_{1} & -Y_{\nu_{1i}}\nu_{i}^{c} & Y_{e_{11}}v_{d} & Y_{e_{12}}v_{d} & Y_{e_{13}}v_{d} \\ g_{2}\nu_{2} & -Y_{\nu_{2i}}\nu_{i}^{c} & Y_{e_{21}}v_{d} & Y_{e_{22}}v_{d} & Y_{e_{23}}v_{d} \\ g_{2}\nu_{3} & -Y_{\nu_{3i}}\nu_{i}^{c} & Y_{e_{31}}v_{d} & Y_{e_{32}}v_{d} & Y_{e_{33}}v_{d} \end{pmatrix}$$

Neutral Higgs mix with sneutrinos

• Charged Higgs mix with charged sleptons