

The SuperB Project

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Overview

Introduction

- The Super Flavour Factories
- The problem
- Λ_{NP} : the energy scale
- Other NP sensitive flavour observables
- Interplay between measurements
- The facility

Summary



Introduction

- Current flavour physics landscape is defined by BaBar, Belle and the Tevatron.
 - We learned that CKM is correct at leading order.
 - Placed indirect constraints on NP that will last well into the LHC era. (e.g. H⁺ searches).
- SuperB will start taking data in 2017.
 - LHCb will have re-defined some areas of flavour physics on that timescale [and take data through to 2017 shutdown].
 - LHC may (or may not) have found new particles.
 - Existing mass scale exclusions are model dependent.
 - In both scenarios results from SuperB can be used to constrain flavour dynamics at high energy.

J. Ellis



MSSM: > 100 parameters

Minimal Flavour Violation: 13 parameters

(+ 6 violating CP)

SU(5) unification: 7 parameters

What we call SUSY depends on how far you want to go in the battle against the curse of dimensionality.

e.g. Only 100 samples per parameter: you need 100^N samples per model.

 100^{124} would cover MSSM without $\nu_{\rm R}$

- The fewer parameters in the model the better!
- The fewer samples the quicker!
- The more constraints the better!

Is this numerical approximation realistic (i.e. good enough)?





The problem

Many NP models

- One may be a good description of nature.
- Almost nothing to go on but the SM and a few hints.



The solution...

Observable/mode	charged Higgs	MFV NP	non-MFV NP	NP in	Right-handed	LHT	SUSY					
	high $ an eta$	low $\tan \beta$	2-3 sector	Z penguins	currents		AC	RVV2	AKM	δLL	FBMSSM	GUT-CMM
$ au ightarrow \mu \gamma$							***	***	*	***	***	***
$\tau \rightarrow \ell \ell \ell$						***						?
$B \rightarrow \tau \nu, \mu \nu$	★★★ (CKM)											
$B ightarrow K^{(*)+} u \overline{ u}$			*	***			*	*	*	*	*	?
$S \text{ in } B \to K^0_S \pi^0 \gamma$			**		***							
S in other penguin modes			★★★(CKM)		***		***	**	*	***	***	?
$A_{CP}(B ightarrow X_s \gamma)$			***		**		*	*	*	***	***	?
$BR(B \rightarrow X_s \gamma)$		*	**		*							**
$BR(B \to X_s \ell \ell)$			**	*	*							?
$B \to K^{(*)}\ell\ell$ (FB Asym)							*	*	*	***	***	?
a_{sl}^s			***			***						***
Charm mixing							***	*	*	*	*	
CPV in Charm	**									***		

It is not obvious which model is correct (if any)

- Have to look at patterns of deviation from the SM to elucidate the underlying theory.
- Just like the journey from Cabibbo's 2x2 mixing matrix through to the CKM matrix of today...



- SuperB doesn't operate at high energy
 - So what's the point of having it?
 - Model dependent indirect probes for NP reach higher scales than can be attained at the LHC.
 - Model dependent direct searches for NP at the LHC have found nothing so far.

Scenario I:

LHC finds NP incompatible with flavour data \rightarrow something to fix in the theory.

Scenario 2:

LHC finds NP compatible with flavour data \rightarrow can use flavour data to start constraining couplings.

Scenario 3:

LHC finds nothing \rightarrow probe high energy indirectly and bound possible effects. e.g. B mixing and the top: everyone knew the top was light until ARGUS found B mixing to be large.

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- SuperB doesn't operate at high energy
 - So what's the point of having it?
 - Model dependent indirect probes for NP reach higher scales than can be attained at the LHC.
 - Model dependent direct searches for NP at the LHC have found nothing so far.

• Also the problem of decoding NP is not as simple as the days of the top quark searches.

We know something is missing but we don't know what we are looking for.

• Many viable models of NP: SUSY or some simple variant (mSUGRA/ CMSSM/...), extra dimensions, 4th generation, Little Higgs, etc. etc. etc. some are coupled to SM Higgs vs. no Higgs.

Each model guides a search, but only one model can be right (and it's not necessarily one of these).



Example: Consider MSSM as an illustration of SUSY

Simple, and being constrained by the LHC but general enough to illustrate the issue:



 Δ 's are related to NP mass scale.

and similarly for $M^2{}_{\widetilde{u}}$

- In many NP scenarios the energy frontier experiments will probe the diagonal elements of mixing matrices.
- Flavour experiments are required to probe off-diagonal ones.



- e.g. MSSM with generic squark mass matrices.
- Use Mass insertion approximation with $m_{\tilde{q}} \sim m_{\tilde{g}}$ to constrain couplings:

$$(\delta_{ij}^q)_{AB} = \frac{(\Delta_{ij})_{AB}^q}{m_{\widetilde{q}}^2}$$

• Can constrain the δ^{d}_{ij} 's using $\mathcal{B}(B \to X_s \gamma)$ $\mathcal{B}(B \to X_s \ell^+ \ell^-)$ $\mathcal{A}_{CP}(B \to X_s \gamma)$ Existing LHC constraints on the gluino mass, mean couplings are non-zero, so we can provide an upper bound on Λ_{NP} .





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• Can constrain the δ^{d}_{ij} 's using • $\mathcal{B}(B \to X_s \gamma)$ • $\mathcal{B}(B \to X_s \ell^+ \ell^-)$ • $\mathcal{A}_{CP}(B \to X_s \gamma)$



e.g. see Hall et al., Nucl. Phys. B **267** 415-432 (1986) Ciuchini et al., hep-ph/0212397



Other NP sensitive flavour observables

- LFV: τ decays
- B Physics
- D Physics

See the following preprints for a more comprehensive overview:

arXiv:1109.5028 [Impact of SuperB]

arXiv:1008.1541 [SuperB Physics progress report]

arXiv:1002.5012 [Belle II Physics Programme]

• Precision $\sin^2\theta_W$



Lepton Flavour Violation (LFV)

- v mixing leads to a low level of charged LFV ($B \sim 10^{-54}$).
 - Enhancements to observable levels are possible with new physics scenarios.
 - Searching for transitions from 3rd generation to 2nd and 1st, i.e.



 \geq Hadron machines are not competitive with e⁺e⁻ machines for this.

▶ N.B. e⁻ beam polarisation helps suppress background.

 $au
ightarrow \mu$ and au
ightarrow e



The golden LFV modes: $au ightarrow \mu\gamma, 3\mu$

Symmetry breaking scale assumed: 500GeV.



NP scale assumed: 500GeV.

- Current experimental limits are at the edges of the model parameter space
- SuperB will be able to significantly constrain these models, and either find both channels, or constrain a large part of parameter space.



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M. Blanke et al. arXiv:0906.5454



Specific example: $\tau \rightarrow \mu \gamma$

• Only cleanly accessible in e^+e^- (golden modes: $\mu\gamma$, 3 lepton).

Model dependent NP constraint.

Correlated with other flavour observables: MEG, LHCb etc.

TABLE III: Expected 90% CL upper limits and 3σ evidence reach on LFV decays with 75 ab⁻¹ with a polarized electron beam.

Drocoss	Expected	3σ evidence
FIGUESS	$90\%{\rm CL}$ upper limit	reach
$\mathcal{B}(\tau \to \mu \gamma)$	$2.4 imes10^{-9}$	$5.4 imes10^{-9}$
$\mathcal{B}(\tau \rightarrow e \gamma)$	$3.0 imes10^{-9}$	$6.8 imes10^{-9}$
$\mathcal{B}(\tau \to \ell \ell \ell)$	$2.3{-}8.2\times10^{-10}$	$1.2{-}4.0\times10^{-9}$

 $m_{\tilde{q}} = 300 \, GeV$ BLUE $m_{\tilde{q}} = 500 \, GeV$ RED

Not updated to latest results from LHCb









B_{u,d} physics: Rare Decays

- Example: $B \to K^{(*)} \nu \overline{\nu}$
 - ▶ Need 75ab⁻¹ to observe pseudoscalar and vector modes.
 - ▶ With more than 75ab⁻¹ we could measure polarisation.

$$\epsilon = \frac{\sqrt{|C_L^{\nu}|^2 + |C_R^{\nu}|^2}}{|(C_L^{\nu})^{\text{SM}}|} , \qquad \eta = \frac{-\text{Re}\left(C_L^{\nu}C_R^{\nu*}\right)}{|C_L^{\nu}|^2 + |C_R^{\nu}|^2}$$

Sensitive to models with Z', RH currents and light scalar particles.





$b \rightarrow sl^+l^-$

SFFs can measure inclusive and exclusive modes:

- Crosscheck results to understand source of NP.
- Important as theory uncertainties differ.
- e.g. expect: 10-15,000 K*μμ and 10-15,000 K*ee events at SuperB.

SFFs can study all lepton flavours:

- Equal amounts of μ and e final states can be measured.
 - Need both of these to measure all NP sensitive observables.
 - LHCb will accumulate slight more events in the μμ mode.
 - ▶ Expect ~20 times the statistics than LHCb for ee mode.
 - ► S/B~ 0.3, c.f. S/B~1.0 for LHCb.
- Can also search for $K^{(*)}\tau^+\tau^-$ decay.
- ... and constrain Majorana v's using like sign final states.
 - Also of interest for D_s decays to $K^{(*)}$ ll final states near charm threshold.



$b \rightarrow sl^+l^-$

SFFs can measure inclusive and exclusive modes:

- Crosscheck results to understand source of NP.
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- e.g. expect: 10-15,000 K*μμ and 10-1

vents at SuperB.

SFFs can study all lept

- -y all lept data from SFFs and raai amounts the data from test all Need by Together he used to test all Need by Together h udy all lept data from SFFs and decays. unts the data from test all NP unts the data from test all decays. ints the used to these decays. iogether be used to these permeasured iogether be used to these permeasured ints can be used to the permeasured ints can be permeasured ints can be permeasured ints can be used to the Equal amounts De measured.
 - sensitive observables.

 - Expect ~2
 - S/B~ 0.3, c. P²/B~1.0 for LHCb.
- Can also search for $K^{(*)}\tau^+\tau^-$ decay.
- ... and constrain Majorana v's using like sign final states.
 - Also of interest for D_s decays to $K^{(*)}$ ll final states near charm threshold.



TDCPV in B decays (i.e. CKM angles $\beta \& \alpha$)

• There are many redundant measurements of the CKM angles that are potential probes of NP.

Mode	C	urrent	Precision	Predi	cted P	recision $(75 \mathrm{ab}^{-1})$	Disco	very Potential
	Stat.	Syst.	$\Delta S^{f}(\text{Th.})$	Stat.	Syst.	$\Delta S^f(\text{Th.})$	3σ	5σ
$J/\psi K_S^0$	0.022	0.010	0 ± 0.01	0.002	0.005	0 ± 0.001	0.02	0.03
$\eta' K_S^0$	0.08	0.02	0.015 ± 0.015	0.006	0.005	0.015 ± 0.015	0.05	0.08
$\phi K^0_S \pi^0$	0.28	0.01	_	0.020	0.010	_	_	_
$f_0 K_S^0$	0.18	0.04	0 ± 0.02	0.012	0.003	0 ± 0.02	0.07	0.12
$K^{0}_{S}K^{0}_{S}K^{0}_{S}$	0.19	0.03	0.02 ± 0.01	0.015	0.020	0.02 ± 0.01	0.08	0.14
ϕK_S^0	0.26	0.03	0.03 ± 0.02	0.020	0.005	0.03 ± 0.02	0.09	0.14
$\pi^0 K_S^0$	0.20	0.03	0.09 ± 0.07	0.015	0.015	0.09 ± 0.07	0.21	0.34
ωK_S^0	0.28	0.02	0.1 ± 0.1	0.020	0.005	0.1 ± 0.1	0.31	0.51
$K^+K^-K^0_S$	0.08	0.03	0.05 ± 0.05	0.006	0.005	0.05 ± 0.05	0.15	0.26
$\pi^0\pi^0K^0_S$	0.71	0.08	_	0.038	0.045	_	_	_
$ ho K_S^0$	0.28	0.07	-0.13 ± 0.16	0.020	0.017	-0.13 ± 0.16	0.41	0.69
$J/\psi\pi^0$	0.21	0.04	_	0.016	0.005	_	-	_
$D^{*+}D^{*-}$	0.16	0.03	_	0.012	0.017	_	-	_
D^+D^-	0.36	0.05	_	0.027	0.008	-	_	-

• Can also measure α using all modes: $\pi\pi$, $\rho\pi$, $\rho\rho$, $a_1\pi$

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$$B_{s} physics$$
• Can cleanly measure A^{s}_{SL} using 5S data
$$A^{s}_{SL} = \frac{\mathcal{B}(B_{s} \to \overline{B}_{s} \to X^{-}\ell^{+}\nu_{\ell}) - \mathcal{B}(\overline{B}_{s} \to B_{s} \to X^{-}\ell^{+}\nu_{\ell})}{\mathcal{B}(B_{s} \to \overline{B}_{s} \to X^{-}\ell^{+}\nu_{\ell}) + \mathcal{B}(\overline{B}_{s} \to B_{s} \to X^{-}\ell^{+}\nu_{\ell})} = \frac{1 - |q/p|^{4}}{1 - |q/p|^{4}}$$

 $\sigma(A_{SL}^s) \sim 0.004$ with a few ab^{-1}



SuperB can also study rare decays with many neutral particles, such as $B_s \rightarrow \gamma \gamma$, which can be enhanced by SUSY.



D Physics

- The programme includes
 - Mixing
 - CP Violation

[This talk]

[This talk]

- Quantum Correlation based measurements
- Rare decays

There were a number of talks on charm physics at the threshold workshop in October. Only a few highlights are shown here: see:

http://indico.ihep.ac.cn/conferenceTimeTable.py?confld=2171#all.detailed

for more detailed discussion on this interesting area of the physics programme [AB, Giorgi, Meadows, Neri, Rama, Sokoloff].



Charm Mixing

Collect data at threshold and at the 4S.

Benefit charm mixing and CPV measurements.



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Charm Mixing





The quest for the final angle of the CKM matrix: β_c

• The charm cu triangle has one unique element: β_c



$$\begin{aligned} \alpha_c &= \arg \left[-V_{ub}^* V_{cb} / V_{us}^* V_{cs} \right]. \\ \beta_c &= \arg \left[-V_{ud}^* V_{cd} / V_{us}^* V_{cs} \right], \\ \gamma_c &= \arg \left[-V_{ub}^* V_{cb} / V_{ud}^* V_{cd} \right], \\ \alpha_c &= (111.5 \pm 4.2)^{\circ} \\ \beta_c &= (0.0350 \pm 0.0001)^{\circ} \\ \gamma_c &= (68.4 \pm 0.1)^{\circ} \end{aligned}$$

 $V_{ud}^* V_{cd} + V_{us}^* V_{cs} + V_{ub}^* V_{cb} = 0$

- Precision measurement of mixing phase in many channels (<2°)</p>
- Constrain $\beta_{c,eff}$ using a $D \rightarrow \pi \pi$ Isospin analysis
 - Search for NP and constrain $\beta_{c,eff} \sim 1^{\circ}$.
 - Can only fully explore in an e^+e^- environment.
 - Data from the charm threshold region completes the set of 5 |V_{ij}| to measure: needs SuperB to perform an indirect test of the triangle. AB, Inguglia, Meadows, PRD 84_114009 (2011)



Precision EW Physics



Precision Electroweak

sin²θ_W can be measured with polarised e⁻ beam:
 √s=Y(4S) is theoretically clean, c.f. b-fragmentation at Z pole. Measure LR asymmetry in



Measure LR asymmetry in $e^+e^- \rightarrow b\overline{b}$ $e^+e^- \rightarrow c\overline{c}$ $e^+e^- \rightarrow \tau^+\tau^$ $e^+e^- \rightarrow \mu^+\mu^-$

at the $\Upsilon(4S)$ to same precision as LEP/SLC at the Z-pole.

Complements

measurements planned/ underway at lower energies (QWeak/MESA). This table concentrates on observables that SFFs can measure, with a few of the prime examples from hadron experiments to highlight that there are many things that need to be measured well.



Golden Measurements: General



This table concentrates on observables that SFFs can measure, with a few of the prime examples from hadron experiments to highlight that there are many things that need to be measured well.



Golden Measurements: CKM

Comparison of relative benefits of SuperB (75ab⁻¹) vs. existing measurements and LHCb (5fb⁻¹) and the LHCb upgrade (50fb⁻¹).



Observable/mode	Current	LHCb	SuperB	Belle II	LHCb upgrade	theory					
	now	(2017)	(2021)	(2021)	(10 years of	now					
		$5{\rm fb}^{-1}$	$75 {\rm ab}^{-1}$	$50 {\rm ab}^{-1}$	running) 50fb^{-1}						
au Decays											
$\tau \rightarrow \mu \gamma ~(\times 10^{-9})$	< 44		< 2.4	< 5.0							
 $\tau \rightarrow e \gamma \; (\times 10^{-9})$	< 33		< 3.0	< 3.7 (est.)							
$\tau \rightarrow \ell \ell \ell (\times 10^{-10})$	< 150 - 270	$< 244^{\ a}$	< 2.3 - 8.2	< 10	$< 24^{-b}$						
		В	u,d Decays								
$BR(B \to \tau \nu) (\times 10^{-4})$	1.64 ± 0.34		0.05	0.04		1.1 ± 0.2					
$BR(B \rightarrow \mu \nu) (\times 10^{-6})$	< 1.0		0.02	0.03		0.47 ± 0.08					
$BR(B \to K^{*+} \nu \overline{\nu}) \ (\times 10^{-6})$	< 80		1.1	2.0		6.8 ± 1.1					
$BR(B \to K^+ \nu \overline{\nu}) \ (\times 10^{-6})$	< 160		0.7	1.6		3.6 ± 0.5					
$BR(B \rightarrow X_s \gamma) (\times 10^{-4})$	3.55 ± 0.26		0.11	0.13	0.23	3.15 ± 0.23					
$A_{CP}(B \rightarrow X_{(s+d)}\gamma)$	0.060 ± 0.060		0.02	0.02		$\sim 10^{-6}$					
$B \to K^* \mu^+ \mu^-$ (events)	250 ^c	8000	$10-15k^d$	7-10k	100,000	-					
$BR(B \to K^* \mu^+ \mu^-) \ (\times 10^{-6})$	1.15 ± 0.16		0.06	0.07		1.19 ± 0.39					
$B \rightarrow K^* e^+ e^-$ (events)	165	400	10-15k	7-10k	5,000	-					
$BR(B \to K^* e^+ e^-) \ (\times 10^{-6})$	1.09 ± 0.17		0.05	0.07		1.19 ± 0.39					
$A_{FB}(B \to K^* \ell^+ \ell^-)$	0.27 ± 0.14^e	ſ	0.040	0.03		-0.089 ± 0.020					
$B \to X_s \ell^+ \ell^-$ (events)	280		8,600	7,000		-					
$BR(B \rightarrow X_s \ell^+ \ell^-) (\times 10^{-6})^g$	3.66 ± 0.77^{h}		0.08	0.10		1.59 ± 0.11					
$S \text{ in } B \rightarrow K^0_S \pi^0 \gamma$	-0.15 ± 0.20		0.03	0.03		-0.1 to 0.1					
$S \text{ in } B \rightarrow \eta' K^0$	0.59 ± 0.07		0.01	0.02		± 0.015					
$S \text{ in } B \rightarrow \phi K^0$	0.56 ± 0.17	0.15	0.02	0.03	0.03	± 0.02					
		I	B_s^0 Decays								
${ m BR}(B^0_s o \gamma \gamma) \; (imes 10^{-6})$	< 8.7		0.3	0.2 - 0.3		0.4 - 1.0					
A_{SL}^{s} (×10 ⁻³)	-7.87 ± 1.96 i	j	4.	5. (est.)		0.02 ± 0.01					
		j	D Decays								
x	$(0.63 \pm 0.20\%)$	0.06%	0.02%	0.04%	0.02%	$\sim 10^{-2 \ k}$					
y	$(0.75 \pm 0.12)\%$	0.03%	0.01%	0.03%	0.01%	$\sim 10^{-2}$ (see above).					
y _{CP}	$(1.11 \pm 0.22)\%$	0.02%	0.03%	0.05%	0.01%	$\sim 10^{-2}$ (see above).					
$\left q/p\right $	$(0.91 \pm 0.17)\%$	8.5%	2.7%	3.0%	3%	~ 10^{-3} (see above).					
$\arg\{q/p\}$ (°)	-10.2 ± 9.2	4.4	1.4	1.4	2.0	$\sim 10^{-3}$ (see above).					
		Other p	rocesses De	cays							
 $\sin^2 \theta_W$ at $\sqrt{s} = 10.58 \text{GeV}/c^2$			0.0002	1		clean					

Observeble/mede	Current	LHCP	SuperB	Belle II	I HCh upgrada	theory					
Observable/ mode	Current	(2017)	Super D	(2021)	(10 rears of	theory					
	now	(2017)	(2021)	(2021)	(10 years of)	now	<u> </u>				
		5 ID	75 ab	50 ab	running) 501b						
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$BB(B \rightarrow \pi u) (\times 10^{-4})$	1.64 ± 0.34		0.05	0.04		11 ± 0.2					
$BR(B \rightarrow \mu\nu) (\times 10^{-6})$	1.04 ± 0.04		0.00	0.04		1.1 ± 0.2 0.47 ± 0.08					
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$BR(B \to K^+ \nu \bar{\nu}) (\times 10^{-6})$	< 160		0.7	2.0		0.0 ± 1.1 2.6 ± 0.5					
$BR(B \to K^{-}\nu\nu) (\times 10^{-4})$	< 100		0.1	1.0	0.93	3.0 ± 0.0 2.15 ± 0.02					
$BR(B \to A_s \gamma) (\times 10^{-1})$	3.55 ± 0.20		0.02	Jeca	0.25	3.10 ± 0.23					
$A_{CP}(D \to A_{(s+d)}\gamma)$ $B \to K^* u^+ u^- \text{ (ovents)}$	0.000 ± 0.000	8000	10.02	m at		~ 10					
$BR(B \rightarrow K^* \mu^+ \mu^-) (\times 10^{-6})$	1.15 ± 0.16	0000	f cha	stiga	100,000	110 ± 0.30					
$B \rightarrow K^* e^+ e^- \text{ (events)}$	1.15 ± 0.10	1	x 0' zi	NEIN	5.000	1.13 ± 0.03					
$BB(B \rightarrow K^* e^+ e^-) (\times 10^{-6})$	1 09 +	umb.	inder	0.07	0,000	119 ± 0.39					
$A_{\rm RD}(B \to K^* \ell^+ \ell^-)$	0.27	(10e	5 1040	0.03		-0.089 ± 0.020					
$R_{FB}(D \rightarrow R \ e \ e)$ $R \rightarrow X \ \ell^+ \ell^- \ (events)$	28	108	8 600	7.000		-0.005 ± 0.020					
$BB(B \to X, \ell^+ \ell^-) \times 10^{-6})^g$	3 66 ± 0 200		0.08	0.10		1.59 ± 0.11					
$S \text{ in } B \to K^0_{\sigma} \pi^0 \gamma$	-0.15 ± 0.20		0.03	0.03		-0.1 to 0.1					
$S \text{ in } B \to n' K^0$	0.59 ± 0.07		0.01	0.02		+0.015					
$S \text{ in } B \to \phi K^0$	0.56 ± 0.17	0.15	0.02	0.03	0.03	+0.02					
	0.00 ± 0.11	1	B ⁰ Decays	0.00	0.00	20102					
$BR(B^0_s \to \gamma\gamma) (\times 10^{-6})$	< 8.7		0.3	0.2 - 0.3		0.4 - 1.0					
A_{sr}^{s} (×10 ⁻³)	-7.87 ± 1.96	j	4.	5. (est.)		0.02 ± 0.01					
			D Decays								
x	$(0.63 \pm 0.20\%)$	0.06%	0.02%	0.04%	0.02%	$\sim 10^{-2 k}$					
v	$(0.75 \pm 0.12)\%$	0.03%	0.01%	0.03%	0.01%	~ 10^{-2} (see above).					
y _{CP}	$(1.11 \pm 0.22)\%$	0.02%	0.03%	0.05%	0.01%	~ 10^{-2} (see above).					
q/p	$(0.91 \pm 0.17)\%$	8.5%	2.7%	3.0%	3%	~ 10^{-3} (see above).					
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Observable/mode	Current	LHCb	$\mathrm{Super}B$	Belle II	LHCb upgrade	theory
	now	(2017)	(2021)	(2021)	(10 years of running)	now
		$5{ m fb}^{-1}$	$75\mathrm{ab}^{-1}$	$50\mathrm{ab}^{-1}$	$50{ m fb}^{-1}$	
$lpha$ from $u\overline{u}d$	6.1°	$5^{\circ a}$	1°	1°	b	$1-2^{\circ}$
β from $c\bar{c}s$ (S)	0.8° (0.020)	0.5° (0.008)	0.1° (0.002)	0.3° (0.007)	0.2° (0.003)	clean
$S \text{ from } B_d o J/\psi \pi^0$	0.21		0.014	0.021 (est.)		clean
$S ext{ from } B_s o J/\psi K^0_S$?			?	clean
$\gamma \text{ from } B \to DK$	11°	$\sim 4^{\circ}$	1°	1.5°	0.9°	clean
$ V_{cb} $ (inclusive) %	1.7		0.5%	0.6 (est.)		dominant
$ V_{cb} $ (exclusive) $\%$	2.2		1.0%	1.2 (est.)		dominant
$ V_{ub} $ (inclusive) $\%$	4.4		2.0%	3.0		dominant
$ V_{ub} $ (exclusive) %	7.0		3.0%	5.0		dominant



Decoding New Physics

Observable/mode	charged Higgs	MFV NP	non-MFV NP	NP in	Right-handed	LHT		SUSY				
	high $\tan \beta$	low $\tan \beta$	2-3 sector	Z penguins	currents		AC	RVV2	AKM	δLL	FBMSSM	GUT-CMM
$\tau ightarrow \mu \gamma$							***	***	*	***	***	***
$\tau \rightarrow \ell \ell \ell$						***						?
$B \rightarrow \tau \nu, \mu \nu$	* * *(CKM)											
$B ightarrow K^{(*)+} u \overline{ u}$			*	***			*	*	*	*	*	?
$S \text{ in } B \to K^0_S \pi^0 \gamma$			**		***							
S in other penguin modes			* * *(CKM)		***		***	**	*	***	***	?
$A_{CP}(B ightarrow X_s \gamma)$			***		**		*	*	*	***	***	?
$BR(B \rightarrow X_s \gamma)$		*	**		*							**
$BR(B \to X_s \ell \ell)$			**	*	*							?
$B \to K^{(*)}\ell\ell$ (FB Asym)							*	*	*	***	***	?
a_{sl}^s			***			***						***
Charm mixing							***	*	*	*	*	
CPV in Charm	**									***		

- This matrix is partially complete, but illustrates the point:
 - Decoding NP is not as simple as setting the mass scale for the top was...



The facility

- Physics dictates:
 - ▶ 75ab⁻¹ at the Y(4S)
 - 1ab⁻¹ at the ψ(3770)
 - + running at nearby resonances (e.g. ~1ab⁻¹ at the Y(5S) and some running above this)



Data samples available at SuperB

- Y(4S) region:
 - \blacktriangleright 75ab⁻¹ at the 4S.
 - Also run above / below the 4S.
 - ~75 x10⁹ B, D and τ pairs.
- ψ(3770) region:
 - ▶ 500fb⁻¹ at threshold.
 - Also run at nearby resonances.
 - ~2 x 10⁹ D pairs at threshold in a few months of running.





Accelerator



Accelerator Scheme



- General strategy:
 - Very small emittance (ILC-DR)
 - Small β^* at IP
 - Large Piwinski angle
 - Crab waist technique
 - Currents similar to present accelerators

Advantages

- Small collision area
- No parasitic crossing
- No synchro-betatron resonances
- Moderate backgrounds
- Re-use, Re-cycle

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 Possible to reuse many components from PEP-II



Raimondi, Shatilov, Zobov http:// arxiv.org/abs/physics/0702033

Crab-waist scheme

Crab sextupoles OFF: Waist line is orthogonal to the axis of other beam



		Base Line		Low En	nittance	High C	urrent	τ/charm		
Parameter	Units	HER (e+)	LER (e-)	HER (e+)	LER (e-)	HER (e+)	LER (e-)	HER (e+)	LER (e-)	
LUMINOSITY (10 ³⁶)	cm ⁻² s⁻¹		1		1					
Energy	GeV	6.7	4.18	6.7	4.18	6.7	4.18	2.58	1.61	
Circumference	m	125	8.4	125	68.4	125	8.4	1258.4		
X-Angle (full)	mrad	6	0	6	0	6	0	60		
Piwinski angle	rad	20.80	16.91	29.42	23.91	13.12	10.67	8.00	6.50	
β _x @ IP	cm	2.6	3.2	2.6	3.2	5.06	6.22	6.76	8.32	
β _v @ IP	cm	0.0253	0.0205	0.0179	0.0145	0.0292	0.0237	0.0658	0.0533	
Coupling (full current)	%	0.25	0.25	0.25	0.25	0.5	0.5	0.25	0.25	
ϵ_x (without IBS)	nm	1.97	1.82	1.00	0.91	1.97	1.82	1.97	1.82	
ϵ_x (with IBS)	nm	2.00	2.46	1.00	1.23	2.00	2.46	5.20	6.4	
ε _y	pm	5	6.15	2.5	3.075	10	12.3	13	16	
σ _x @ IP	μ m	7.211	8.872	5.099	6.274	10.060	12.370	18.749	23.076	
σ _y @ IP	μm	0.036	0.036	0.021	0.021	0.054	0.054	0.092	0.092	
Σ _x	μ m	11.4	433	8.0	85	15.	944	29.732		
Σ _y	μ m	0.0	50	0.0)30	0.0	76	0.131		
σ_L (0 current)	mm	4.69	4.29	4.73	4.34	4.03	3.65	4.75	4.36	
σ_L (full current)	mm	5	5	5	5	4.4	4.4	5	5	
Beam current	mA	1892	2447	1460	1888	3094	4000	1365	1766	
Buckets distance	#	i	2	2		1		1		
Buckets distance	ns	4.	20	4.20		2.	10	2.10		
lon gap	%	i	2	i	2	i	2	i	2	
RF frequency	MHz	47	76	47	76	47	76	47	' 6	
Harmonic number		19	98	19	98	19	98	19	98	
Number of bunches		46	55	46	55	93	31	93	31	
N. Particle/bunch (10 ¹⁰)		5.08	6.56	3.92	5.06	4.15	5.36	1.83	2.37	
Tune shift x		0.0026	0.0040	0.0020	0.0031	0.0053	0.0081	0.0063	0.0096	
Tune shift y		0.1067	0.1069	0.0980	0.0981	0.0752	0.0755	0.1000	0.1001	
Long. damping time	msec	13.4	20.3	13.4	20.3	13.4	20.3	26.8	40.6	
Energy Loss/turn	MeV	2.11	0.865	2.11	0.865	2.11	0.865	0.4	0.166	
σ _E (full current)	δΕ/Ε	6.43E-04	7.34E-04	6.43E-04	7.34E-04	6.43E-04	7.34E-04	6.94E-04	7.34E-04	
CM σ _E	δE/E	5.00	E-04	5.00E-04		5.00E-04		5.26	E-04	
Total lifetime	min	4.23	4.48	3.05	3.00	7.08	7.73	11.41	6.79	
Total RF Power	MW	16.	16.38		.37	28	.83	2.81		



Table

Tau/charm threshold running at 10³⁵

Baseline + other 2 options: •Lower y-emittance •Higher currents (twice bunches)

Baseline: •Higher emittance due to IBS •Asymmetric beam currents

RF power includes SR and HOM



Synchroton Light properties and uses

- Synchrotron light properties from dipoles are competitive
- Assumed undulators characteristics as NSLS-II
- Light properties from undulators still better than most LS, slightly worst than PEP-X (last generation project)



Italian Institute for Technology interests:

- lithography for 3D scaffolding for bioengineering
- laser ablation on biomaterials
- femtochemistry studies
- photon induced growth for material science
- innovative interface diffraction techniques
- imaging in biomedicine
- X ray microscopy





- Several locations have been examined
 - Seismic properties, water, infrastructures, etc.
- Decision in May 2011 for the Tor Vergata Campus in Rome.
- Full support of Tor Vergata University for the project







Possible layout @ Tor Vergata





Detector



Electronics

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Some parts from BaBar will be reused to make the SuperB detector.



DETECTOR TIMELINE CARTOON Data

Bkgnd/ Commissioning Detector

MDI

IR Hall

Services

2015 Detector Start Assembly

2014 Detector Start Procurement 2013

Design

2012

Detector TDR

IR Hall Available Civil Start Construction

Circulating

Beams

CabibboLab

Taking

Detector

Commissi

oning

Transp.

of Babar

pieces

2017

2016



Recent public documents



Conceptual Design Report: arXiv:0709.0451

Valencia Physics Workshop Report: arXiv:0810.1312

Detector White Paper: arXiv:1007.4241

Accelerator White Paper: arXiv:1009.6178

Physics White Paper: arXiv:1008.1541

Impact Document: arXiv:1109.5028

Recent Review: arXiv:1110.3901



Backup



 SVT provides precise tracking and vertex reconstruction, crucial for time dependent measurements, and perform stand-alone tracking for low p_t particles.

Based on BaBar SVT: 5 layers silicon strip modules + Layer0 at small radius to improve vertex resolution and compensate the reduced SuperB boost w.r.t PEPII





- Physycs performance and back. levels set stringent requirements on Layer0:
- R~I.5 cm, material budget < I% X_{0,}
- hit resolution 10-15 um in both coordinates
- Track rate > 5MHz/cm² (with large cluster too!), TID > 3MRad/yr

FLOOR

Several options under study for Layer0











Precision CKM constraints



SuperB Measures the sides and angles of the Unitarity Triangle.



Physics programme in a nutshell

- Versatile flavour physics experiment:
 - Probe new physics observables in wide range of decays.
 - Pattern of deviation from Standard Model can be used to identify structure of new physics.
 - > Clean experimental environment means clean signals in many modes.
 - Polarised e^- beam benefit for τ LFV searches (unique feature).
 - Charm threshold running adds many more observables, and improves potential of SuperB (unique feature).
 - Measure angles and sides of the Unitarity triangle.
 - Measure other CKM matrix elements at threshold and using τ data.
- SuperB is working on a TDR for 2012.
- Will be followed by a physics book some time later.
 - Plenty of open areas for newcomers to work on!



LHC Results on SUSY (slide from A. Cakir, Lomonosov XV)

Interpretation of the Physics Results for Summer 2011



So far no evidence for SUSY.

The SUSY mass scale is now looking likely to be above ITeV.

This has interesting implications for some of our measurements.

We need to make sure our benchmark processes and assumed scales are still valid as these contours are updated. J. Ellis

