

The right-handed sneutrino Cold Dark Matter

Jonathan Da Silva

Laboratoire d'Annecy-le-Vieux de Physique Théorique
Beginning of 2nd year of PhD, Annecy-le-Vieux
G. Bélanger, J. Da Silva and A. Pukhov, [arXiv:1110.2414v1 \[hep-ph\]](#)



Outline

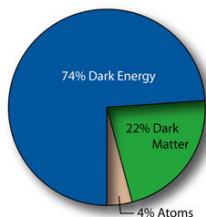
- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

Motivations

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

Dark Matter and supersymmetry

- Solving Dark Matter (DM) issue :

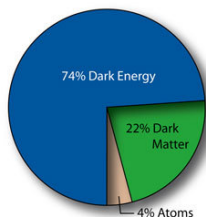


CMB, rotation curves, Bullet cluster, ...

⇒ need to introduce weakly interacting, stable particles

Dark Matter and supersymmetry

- Solving Dark Matter (DM) issue :



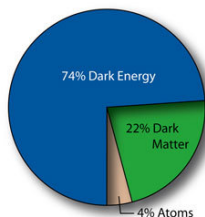
CMB, rotation curves, Bullet cluster, ...

⇒ need to introduce weakly interacting, stable particles

⇒ I considered massive particles (around GeV), Cold Dark Matter

Dark Matter and supersymmetry

- Solving Dark Matter (DM) issue :

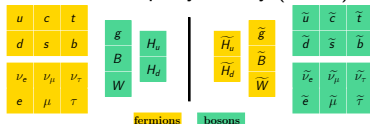


CMB, rotation curves, Bullet cluster, ...

⇒ need to introduce weakly interacting, stable particles

⇒ I considered massive particles (around GeV), Cold Dark Matter

- Interests on Supersymmetry (SUSY) :

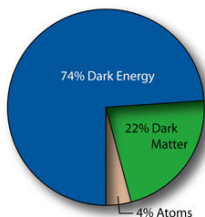


Hierarchy problem, unification of the couplings, ...

⇒ also addition of new particles interacting weakly with standard particles

Dark Matter and supersymmetry

- Solving Dark Matter (DM) issue :

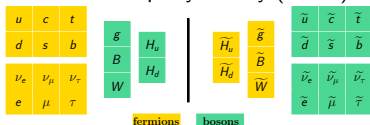


CMB, rotation curves, Bullet cluster, ...

⇒ need to introduce weakly interacting, stable particles

⇒ I considered massive particles (around GeV), Cold Dark Matter

- Interests on Supersymmetry (SUSY) :



Hierarchy problem, unification of the couplings, ...

⇒ also addition of new particles interacting weakly with standard particles

⇒ Dark Matter candidates in supersymmetric models

Some candidates

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

Some candidates

Assuming R-parity :

- 2 WIMPs candidates in the MSSM :

- ▶ Lightest neutralino : a lot of studies \Rightarrow good DM candidate
- ▶ LH sneutrino : too high coupling with $Z^0 \Rightarrow$ don't satisfy experimental constraints on spin independent direct detection \Rightarrow bad DM candidate

Some candidates

Assuming R-parity :

- 2 WIMPs candidates in the MSSM :
 - ▶ Lightest neutralino : a lot of studies \Rightarrow good DM candidate
 - ▶ LH sneutrino : too high coupling with $Z^0 \Rightarrow$ don't satisfy experimental constraints on spin independent direct detection \Rightarrow bad DM candidate
- Others SUSY candidates to DM :
 - ▶ Gravitino, axino, ...
 - ▶ SUSY partner of the RH neutrino : coupled with other particles, mixed with LH sneutrino, sterile, ...

Some candidates

Assuming R-parity :

- 2 WIMPs candidates in the MSSM :
 - ▶ Lightest neutralino : a lot of studies \Rightarrow good DM candidate
 - ▶ LH sneutrino : too high coupling with $Z^0 \Rightarrow$ don't satisfy experimental constraints on spin independent direct detection \Rightarrow bad DM candidate
- Others SUSY candidates to DM :
 - ▶ Gravitino, axino, ...
 - ▶ SUSY partner of the RH neutrino : coupled with other particles, mixed with LH sneutrino, sterile, ...

\Rightarrow I focused on a RH sneutrino supersymmetric particle
of the a Dirac RH neutrino

Some candidates

Assuming R-parity :

- 2 WIMPs candidates in the MSSM :
 - ▶ Lightest neutralino : a lot of studies \Rightarrow good DM candidate
 - ▶ LH sneutrino : too high coupling with $Z^0 \Rightarrow$ don't satisfy experimental constraints on spin independent direct detection \Rightarrow bad DM candidate
- Others SUSY candidates to DM :
 - ▶ Gravitino, axino, ...
 - ▶ SUSY partner of the RH neutrino : coupled with other particles, mixed with LH sneutrino, sterile, ...

\Rightarrow I focused on a RH sneutrino supersymmetric particle
of the a Dirac RH neutrino

\Rightarrow This candidate couples to new vector, scalar field,
adding a new abelian gauge group

The model

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

Contents

- Symmetry group : $SU(3)_c \times SU(2)_L \times U(1)_Y \times U'(1)$

Coupling constants associated : g_3, g_2, g_Y and $g'_1 = g_1 = \sqrt{\frac{5}{3}}g_Y$

- $U'(1)$ stem from E_6 model

⇒ Model implemented at tree-level in Unitary and Feynman gauge in LanHEP

Contents

- Symmetry group : $SU(3)_c \times SU(2)_L \times U(1)_Y \times U'(1)$

Coupling constants associated : g_3, g_2, g_Y and $g'_1 = g_1 = \sqrt{\frac{5}{3}}g_Y$

- $U'(1)$ stem from E_6 model

⇒ Model implemented at tree-level in Unitary and Feynman gauge in LanHEP

Some differences with the MSSM :

- Gauge sector : Physical abelian gauge bosons : Z_1 and Z_2 , mixing of the Z^0 of the SM and Z'

Contents

- Symmetry group : $SU(3)_c \times SU(2)_L \times U(1)_Y \times U'(1)$

Coupling constants associated : g_3, g_2, g_Y and $g'_1 = g_1 = \sqrt{\frac{5}{3}}g_Y$

- $U'(1)$ stem from E_6 model

⇒ Model implemented at tree-level in Unitary and Feynman gauge in LanHEP

Some differences with the MSSM :

- Gauge sector : Physical abelian gauge bosons : Z_1 and Z_2 , mixing of the Z^0 of the SM and Z'
- Higgs sector : 1 CP odd Higgs A^0 , 5 CP even Higgs : h^\pm, h_1, h_2 and h_3
singlet-like Higgs (h_2 or h_3) mass near Z_2 mass
including pure UMSSM terms + radiative corrections
⇒ m_{h_1} above LEP limits

Contents

- Symmetry group : $SU(3)_c \times SU(2)_L \times U(1)_Y \times U'(1)$

Coupling constants associated : g_3, g_2, g_Y and $g'_1 = g_1 = \sqrt{\frac{5}{3}}g_Y$

- $U'(1)$ stem from E_6 model

⇒ Model implemented at tree-level in Unitary and Feynman gauge in LanHEP

Some differences with the MSSM :

- Gauge sector : Physical abelian gauge bosons : Z_1 and Z_2 , mixing of the Z^0 of the SM and Z'
- Higgs sector : 1 CP odd Higgs A^0 , 5 CP even Higgs : h^\pm, h_1, h_2 and h_3
singlet-like Higgs (h_2 or h_3) mass near Z_2 mass
including pure UMSSM terms + radiative corrections
⇒ m_{h_1} above LEP limits
- Gauginos sector : 6 neutralinos in the basis $(\tilde{B}, \tilde{W}^3, \tilde{H}_d^0, \tilde{H}_u^0, \tilde{S}, \tilde{B}')$

Relevant free parameters : $M_{\tilde{\nu}_R}, \mu, A_\lambda, M_{Z_2}, \theta_{E_6}, \alpha_Z, M_1, M'_1$. Soft terms at 2 TeV

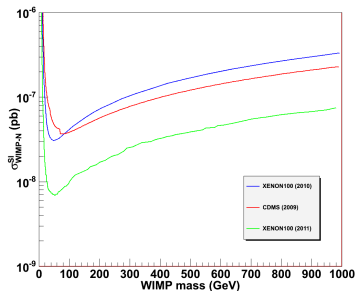
Constraints

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - **Constraints**
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

Constraints

On our CDM candidate :

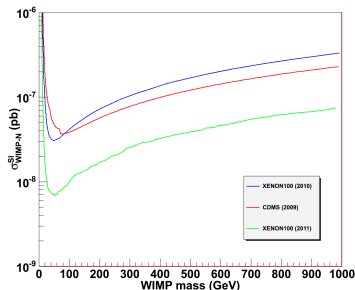
- Relic density at 3σ with $\Omega_{WIMP} h^2 = 0.1123 \pm 0.0035$
- Spin independent direct detection cross section



Constraints

On our CDM candidate :

- Relic density at 3σ with $\Omega_{WIMP} h^2 = 0.1123 \pm 0.0035$
- Spin independent direct detection cross section



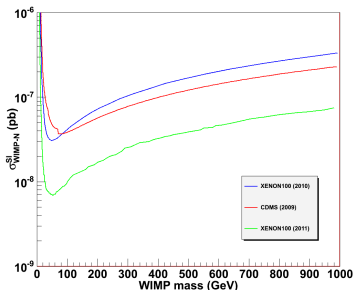
On the model in general :

- Higgs mass constraints from LEP and LHC : $114.4 \text{ GeV} < m_{h_1} < 144 \text{ GeV}$

Constraints

On our CDM candidate :

- Relic density at 3σ with $\Omega_{WIMP} h^2 = 0.1123 \pm 0.0035$
- Spin independent direct detection cross section



On the model in general :

- Higgs mass constraints from LEP and LHC : $114.4 \text{ GeV} < m_{h_1} < 144 \text{ GeV}$
- New Z boson mass constraints from ATLAS :

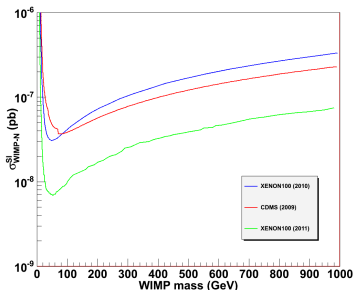
Q' choice	Q_ψ	Q_N	Q_η	Q_I	Q_S	Q_χ
M_{Z_2} (TeV)	1.49	1.52	1.54	1.56	1.60	1.64

- Z^0 properties $\Rightarrow \alpha_Z \lesssim 10^{-3}$

Constraints

On our CDM candidate :

- Relic density at 3σ with $\Omega_{WIMP} h^2 = 0.1123 \pm 0.0035$
- Spin independent direct detection cross section



On the model in general :

- Higgs mass constraints from LEP and LHC : $114.4 \text{ GeV} < m_{h_1} < 144 \text{ GeV}$
- New Z boson mass constraints from ATLAS :

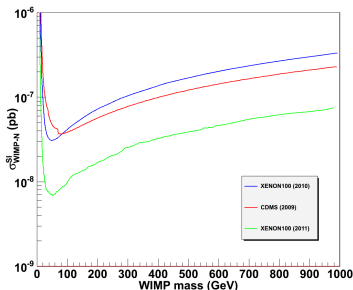
Q' choice	Q_ψ	Q_N	Q_η	Q_I	Q_S	Q_χ
M_{Z_2} (TeV)	1.49	1.52	1.54	1.56	1.60	1.64

- Z^0 properties $\Rightarrow \alpha_Z \lesssim 10^{-3}$
- LEP constraints on particles masses implemented in the micrOMEGAs code

Constraints

On our CDM candidate :

- Relic density at 3σ with $\Omega_{WIMP} h^2 = 0.1123 \pm 0.0035$
- Spin independent direct detection cross section



On the model in general :

- Higgs mass constraints from LEP and LHC : $114.4 \text{ GeV} < m_{h_1} < 144 \text{ GeV}$
- New Z boson mass constraints from ATLAS :

Q' choice	Q_ψ	Q_N	Q_η	Q_I	Q_S	Q_χ
M_{Z_2} (TeV)	1.49	1.52	1.54	1.56	1.60	1.64

- Z^0 properties $\Rightarrow \alpha_Z \lesssim 10^{-3}$
- LEP constraints on sparticles masses implemented in the micrOMEGAs code
- B mesons physics constraints : $\Delta M_{d,s}$ mass differences (code adapted from a NMSSMTools routine)

CDM interactions

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

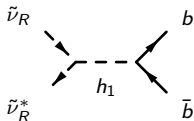
WIMP annihilation

Parameter space regions with $\Omega_{WIMP} h^2 \approx 0.1 \Rightarrow$ need to increase the annihilation cross section :

WIMP annihilation

Parameter space regions with $\Omega_{\text{WIMP}} h^2 \approx 0.1 \Rightarrow$ need to increase the annihilation cross section :

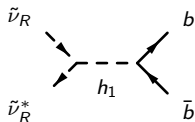
- WIMP mass near $m_{h_1}/2$:



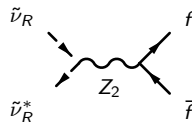
WIMP annihilation

Parameter space regions with $\Omega_{\text{WIMP}} h^2 \approx 0.1 \Rightarrow$ need to increase the annihilation cross section :

- WIMP mass near $m_{h_1}/2$:



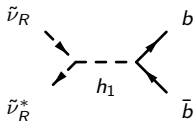
- WIMP mass near $M_{Z_2}/2$ (also $m_{h_i}/2$) :



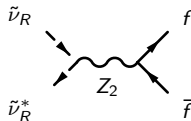
WIMP annihilation

Parameter space regions with $\Omega_{\text{WIMP}} h^2 \approx 0.1 \Rightarrow$ need to increase the annihilation cross section :

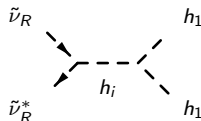
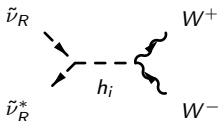
- WIMP mass near $m_{h_1}/2$:



- WIMP mass near $M_{Z_2}/2$ (also $m_{h_i}/2$) :



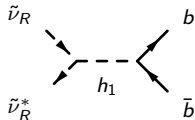
- WIMP mass near $m_{h_i}/2$ or above W pair threshold :



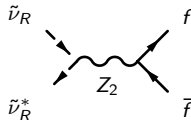
WIMP annihilation

Parameter space regions with $\Omega_{\text{WIMP}} h^2 \approx 0.1 \Rightarrow$ need to increase the annihilation cross section :

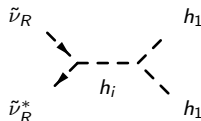
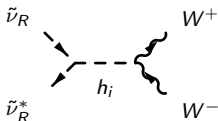
- WIMP mass near $m_{h_1}/2$:



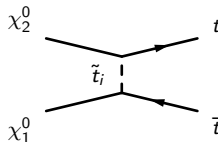
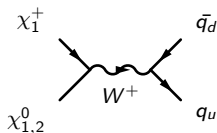
- WIMP mass near $M_{Z_2}/2$ (also $m_{h_i}/2$) :



- WIMP mass near $m_{h_i}/2$ or above W pair threshold :



- Coannihilation processes (mainly higgsino-like) :



Scattering on nucleons

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - Output
- 5 Conclusion and perspectives

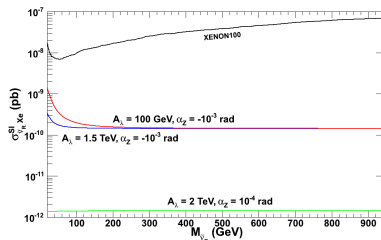
Scattering on nucleons

- Mainly abelian gauge bosons contribution, h_1 for LSP mass $\lesssim 200$ GeV

Scattering on nucleons

- Mainly abelian gauge bosons contribution, h_1 for LSP mass $\lesssim 200$ GeV

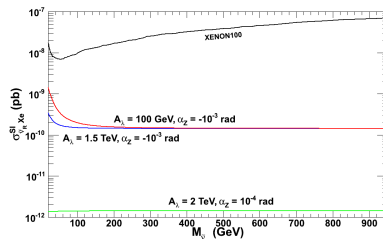
\Rightarrow for some $U'(1)$ models we can have
a good suppression of the gauge boson
or/and Higgs part :



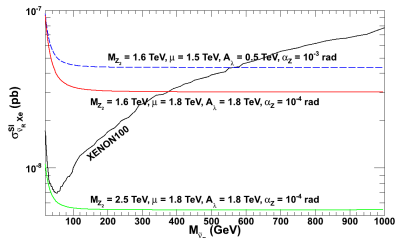
Scattering on nucleons

- Mainly abelian gauge bosons contribution, h_1 for LSP mass $\lesssim 200$ GeV

\Rightarrow for some $U'(1)$ models we can have a good suppression of the gauge boson or/and Higgs part :



\Rightarrow for other models, huge constraints on the parameter space appear :



Some results

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 **Some results**
 - **Characteristics of the global scan**
 - Output
- 5 Conclusion and perspectives

Characteristics of the global scan

Fixed parameters				Free parameters	
Soft terms				Name	Domain of variation
m_{Q_i}	2 TeV	m_{L_i}	2 TeV	$M_{\tilde{\nu}_R}$	[0, 1.5] TeV
$m_{\tilde{u}_i}$	2 TeV	$m_{\tilde{d}_i}$	2 TeV	M_{Z_2}	[1.3, 3] TeV
$m_{\tilde{e}_i}$	2 TeV	$m_{\tilde{\nu}_j}$	2 TeV	μ	[0.1, 2] TeV
$i \in \{1, 2, 3\}, j \in \{1, 2\}$				A_λ	[0, 2] TeV
Trilinear couplings + M_K				θ_{E_6}	$[-\pi/2, \pi/2]$ rad
A_t	1 TeV	A_b	0 TeV	α_Z	$[-3.10^{-3}, 3.10^{-3}]$ rad
A_c	0 TeV	A_s	0 TeV	M_1	[0.1, 2] TeV
A_u	0 TeV	A_d	0 TeV	M'_1	[0.1, 2] TeV
A_l	0 TeV	M_K	1 eV	$M_2 = 2M_1$ et $M_3 = 6M_1$	

Output

- 1 Motivations
 - Dark Matter and supersymmetry
 - Some candidates
- 2 The model
 - Contents
 - Constraints
- 3 CDM interactions
 - WIMP annihilation
 - Scattering on nucleons
- 4 Some results
 - Characteristics of the global scan
 - **Output**
- 5 Conclusion and perspectives

Output

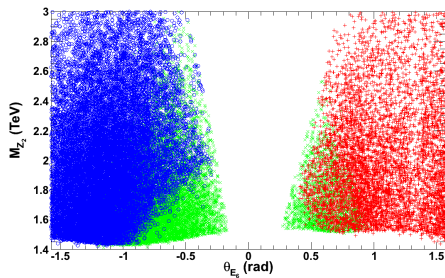
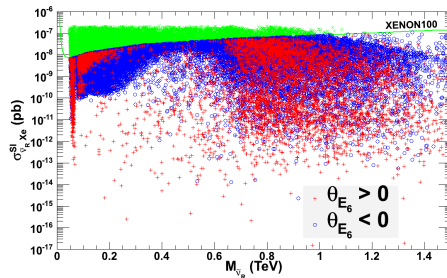
Interesting WIMP mass from 50 GeV to TeV-scale, for following processes :

- h_1 resonance
- Z_2 /singlet-like Higgs resonance
- Coannihilation with neutralinos, charginos (a few with sfermions)
- Annihilation into W pairs through Higgs exchange

Output

Interesting WIMP mass from 50 GeV to TeV-scale, for following processes :

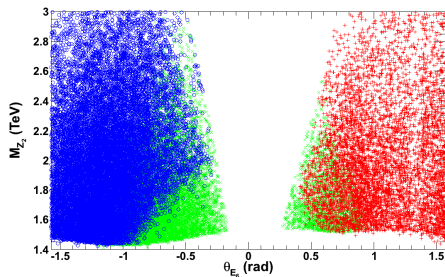
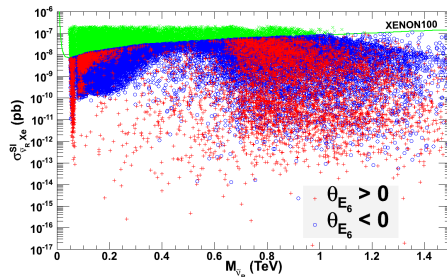
- h_1 resonance
- Z_2 /singlet-like Higgs resonance
- Coannihilation with neutralinos, charginos (a few with sfermions)
- Annihilation into W pairs through Higgs exchange



Output

Interesting WIMP mass from 50 GeV to TeV-scale, for following processes :

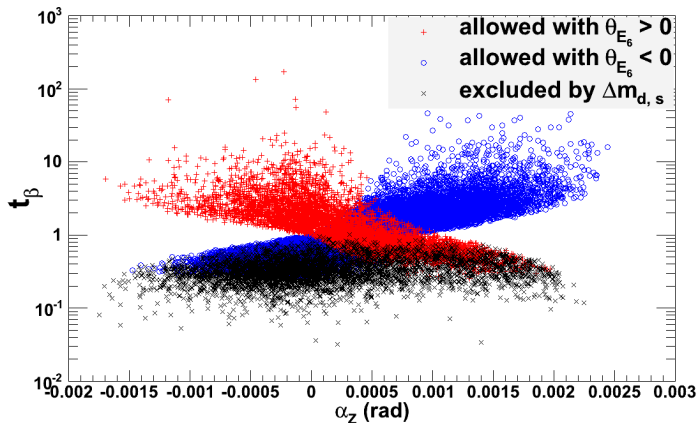
- h_1 resonance
- Z_2 /singlet-like Higgs resonance
- Coannihilation with neutralinos, charginos (a few with sfermions)
- Annihilation into W pairs through Higgs exchange



Lower is $|\theta_{E_6}|$, higher are Z_2 processes in direct detection cross section \Rightarrow huge constraint

Output

Large SUSY corrections proportional to $\frac{1}{t_\beta^4} \Rightarrow$ small values of t_β very constrained by ΔM_s :



Conclusion and perspectives

1 Motivations

- Dark Matter and supersymmetry
- Some candidates

2 The model

- Contents
- Constraints

3 CDM interactions

- WIMP annihilation
- Scattering on nucleons

4 Some results

- Characteristics of the global scan
- Output

5 Conclusion and perspectives

Conclusion and perspectives

- **RH sneutrino is a viable dark matter candidate**

it respects experimental limits in the case of some processes :

- ▶ Resonance (h_1 , Z_2 and singlet-like Higgs)
- ▶ Coannihilation (neutralinos, charginos, others sfermions)
- ▶ Annihilation into W pairs generally with exchange of h_1

- Direct detection experiments strongly constrain the model as well as ΔM_s

Conclusion and perspectives

- **RH sneutrino is a viable dark matter candidate**

it respects experimental limits in the case of some processes :

- ▶ Resonance (h_1 , Z_2 and singlet-like Higgs)
- ▶ Coannihilation (neutralinos, charginos, others sfermions)
- ▶ Annihilation into W pairs generally with exchange of h_1

- Direct detection experiments strongly constrain the model as well as ΔM_s

- This model could be also tested with indirect detection, other flavour physics observables, ...

- More careful study of the UMSSM Higgs sector in preparation

- Others projects in my Explora DOC exchange program in Durham like LHC limits in the NMSSM, ...

Conclusion and perspectives

- **RH sneutrino is a viable dark matter candidate**

it respects experimental limits in the case of some processes :

- ▶ Resonance (h_1 , Z_2 and singlet-like Higgs)
- ▶ Coannihilation (neutralinos, charginos, others sfermions)
- ▶ Annihilation into W pairs generally with exchange of h_1

- Direct detection experiments strongly constrain the model as well as ΔM_s

- This model could be also tested with indirect detection, other flavour physics observables, ...

- More careful study of the UMSSM Higgs sector in preparation

- Others projects in my Explora DOC exchange program in Durham like LHC limits in the NMSSM, ...

Thanks for your attention !

BACKUP

UMSSM fields

Chiral supermultiplets				
Supermultiplets		spin 0	spin 1/2	$SU(3)_c, SU(2)_L, U(1)_Y, U'(1)$
squarks, quarks (3 families)	Q	$(\tilde{u}_L \tilde{d}_L)$	$(u_L d_L)$	$(\mathbf{3}, 2, \frac{1}{6}, Q'_Q)$
	\bar{u}	\tilde{u}_R^*	\bar{u}_R	$(\bar{\mathbf{3}}, 1, -\frac{2}{3}, Q'_u)$
	\bar{d}	\tilde{d}_R^*	\bar{d}_R	$(\bar{\mathbf{3}}, 1, \frac{1}{3}, Q'_d)$
sleptons, leptons (3 families)	L	$(\tilde{\nu}_L \tilde{e}_L)$	$(\nu_L e_L)$	$(\mathbf{1}, 2, -\frac{1}{2}, Q'_L)$
	$\bar{\nu}$	$\tilde{\nu}_R^*$	$\bar{\nu}_R$	$(\mathbf{1}, 1, 0, Q'_\nu)$
	\bar{e}	\tilde{e}_R^*	\bar{e}_R	$(\mathbf{1}, 1, \frac{1}{6}, Q'_e)$
Higgs, higgsinos	H_u	$(H_u^+ H_u^0)$	$(\tilde{H}_u^+ \tilde{H}_u^0)$	$(\mathbf{1}, 2, \frac{1}{2}, Q'_{H_u})$
	H_d	$(H_d^0 H_d^-)$	$(\tilde{H}_d^0 \tilde{H}_d^-)$	$(\mathbf{1}, 2, -\frac{1}{2}, Q'_{H_d})$
	S	S	\tilde{S}	$(\mathbf{1}, 1, 0, Q'_S)$
Vector supermultiplets				
Supermultiplets		spin 1/2	spin 1	$SU(3)_c, SU(2)_L, U(1)_Y, U'(1)$
gluino, gluon		\tilde{g}	g	$(\mathbf{8}, \mathbf{1}, 0, 0)$
winos, W bosons		$\tilde{W}^\pm \tilde{W}^3$	$W^\pm W^3$	$(\mathbf{1}, \mathbf{3}, 0, 0)$
bino, B boson		\tilde{B}	B	$(\mathbf{1}, \mathbf{1}, 0, 0)$
bino', B' boson		\tilde{B}'	B'	$(\mathbf{1}, \mathbf{1}, 0, 0)$

Some new lagrangian terms

- Superpotential :

$$W_{MSSM} = \bar{u}y_u QH_u - \bar{d}y_d QH_d - \bar{e}y_e LH_d + \mu H_u H_d$$

$$W_{UMSSM} = W_{MSSM}(\mu = 0) + \lambda S H_u H_d + \tilde{\nu} y_\nu L H_u$$

- Soft supersymmetry breaking :

$$\begin{aligned} \mathcal{L}_{soft}^{MSSM} = & -\frac{1}{2}(M_3 \tilde{g} \tilde{g} + M_2 \tilde{W} \tilde{W} + M_1 \tilde{B} \tilde{B} + \text{c.c.}) \\ & -(\tilde{u}_R^* a_u \tilde{Q} H_u - \tilde{d}_R^* a_d \tilde{Q} H_d - \tilde{e}_R^* a_e \tilde{L} H_d + \text{c.c.}) \\ & -\tilde{Q}^\dagger m_Q^2 \tilde{Q} - \tilde{L}^\dagger m_L^2 \tilde{L} - \tilde{u}_R^* m_{\tilde{e}}^2 \tilde{u}_R - \tilde{d}_R^* m_{\tilde{d}}^2 \tilde{d}_R - \tilde{e}_R^* m_{\tilde{e}}^2 \tilde{e}_R \\ & -m_{H_u}^2 H_u^\dagger H_u - m_{H_d}^2 H_d^\dagger H_d - (b H_u H_d + \text{c.c.}) \\ \mathcal{L}_{soft}^{UMSSM} = & \mathcal{L}_{soft}^{MSSM}(b = 0) - \left(\frac{1}{2} M_1' \tilde{B}' \tilde{B}' + M_K \tilde{B} \tilde{B}' + \tilde{\nu}_R^* a_\nu \tilde{L} H_u + \text{c.c.} \right) \\ & -\tilde{\nu}_R^* m_{\tilde{\nu}}^2 \tilde{\nu}_R - (\lambda A_\lambda S H_u H_d + \text{c.c.}) - m_S^2 S^* S \end{aligned}$$

LanHEP, A. Semenov, arXiv :0805.0555v1 [hep-ph]

Reason of constrained t_β

$$M_Z^2 = M_{Z_1}^2 \cos^2 \alpha_{ZZ'} + M_{Z_2}^2 \sin^2 \alpha_{ZZ'}$$

$$M_{Z'}^2 = M_{Z_1}^2 \sin^2 \alpha_{ZZ'} + M_{Z_2}^2 \cos^2 \alpha_{ZZ'}.$$

$$\Downarrow$$

$$\tan 2\alpha_{ZZ'} = \frac{2\Delta^2}{M_{Z'}^2 - M_Z^2} \implies \sin 2\alpha_{ZZ'} = \frac{2\Delta^2}{M_{Z_2}^2 - M_{Z_1}^2}.$$

Knowing that

$$\Delta^2 = \frac{g'_1 \sqrt{g'^2 + g_2^2}}{2} v^2 (Q'_2 s_\beta^2 - Q'_1 c_\beta^2),$$

$$\Downarrow$$

$$c_\beta^2 = \frac{1}{Q'_1 + Q'_2} \left(\frac{\sin 2\alpha_{ZZ'} (M_{Z_1}^2 - M_{Z_2}^2)}{v^2 g'_1 \sqrt{g'^2 + g_2^2}} + Q'_2 \right).$$

Higgs masses

$$m_{A^0}^2 = \frac{\lambda A_\lambda \sqrt{2}}{\sin 2\phi} v + \Delta_{EA} \quad \tan \phi = \frac{v \sin 2\beta}{2v_s}$$

$$m_{H^\pm}^2 = \frac{\lambda A_\lambda \sqrt{2}}{\sin 2\beta} v_s - \frac{\lambda^2}{2} v^2 + \frac{g_2^2}{2} v^2 + \Delta_\pm \quad \tan \beta = \frac{v_u}{v_d}$$

$$M_{CP\text{even}}^2 :$$

$$(\mathcal{M}_+^0)_{11} = \left[\frac{(g'^2 + g_2^2)^2}{4} + Q_1'^2 g_1'^2 \right] (v c_\beta)^2 + \frac{\lambda A_\lambda t_\beta v_s}{\sqrt{2}} + \Delta_{11}$$

$$(\mathcal{M}_+^0)_{12} = - \left[\frac{(g'^2 + g_2^2)^2}{4} - \lambda^2 - Q_1' Q_2' g_1'^2 \right] v^2 s_\beta c_\beta - \frac{\lambda A_\lambda v_s}{\sqrt{2}} + \Delta_{12}$$

$$(\mathcal{M}_+^0)_{13} = \left[\lambda^2 + Q_1' Q_S' g_1'^2 \right] v c_\beta v_s - \frac{\lambda A_\lambda v s_\beta}{\sqrt{2}} + \Delta_{13}$$

$$(\mathcal{M}_+^0)_{22} = \left[\frac{(g'^2 + g_2^2)^2}{4} + Q_2'^2 g_1'^2 \right] (v s_\beta)^2 + \frac{\lambda A_\lambda v_s}{t_\beta \sqrt{2}} + \Delta_{22}$$

$$(\mathcal{M}_+^0)_{23} = \left[\lambda^2 + Q_2' Q_S' g_1'^2 \right] v s_\beta v_s - \frac{\lambda A_\lambda v c_\beta}{\sqrt{2}} + \Delta_{23}$$

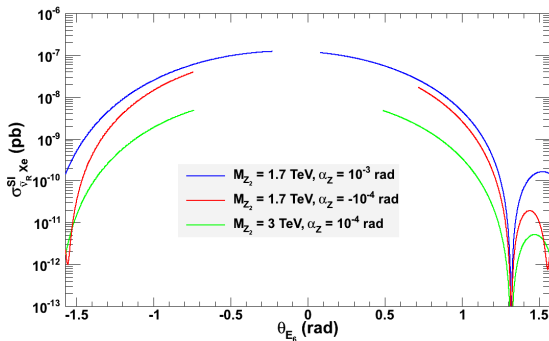
$$(\mathcal{M}_+^0)_{33} = Q_S'^2 g_1'^2 v_s^2 + \frac{\lambda A_\lambda v^2 s_\beta c_\beta}{v_s \sqrt{2}} + \Delta_{33}$$

Direct detection constraint

Abelian gauge boson contribution to direct detection :

$$\sigma_{\tilde{\nu}_R N}^{Z_1, Z_2} = \frac{\mu_{\tilde{\nu}_R N}^2}{\pi} (g'_1 Q'_{\tilde{\nu}})^2 [(y(1 - 4s_W^2) + y')Z + (-y + 2y')(A - Z)]^2$$

$$\text{with } y = \frac{g' \sin \alpha_Z \cos \alpha_Z}{4 \sin \theta_W} \left(\frac{1}{M_{Z_2}^2} - \frac{1}{M_{Z_1}^2} \right), \quad y' = -\frac{g'_1}{2} Q'_{\tilde{\nu}}{}^d \left(\frac{\sin^2 \alpha_Z}{M_{Z_1}^2} + \frac{\cos^2 \alpha_Z}{M_{Z_2}^2} \right)$$



\Rightarrow stringent constraints for small $|\theta_{E_6}|$ because of $Q'_{\tilde{\nu}}{}^d$ term

Coannihilation with sfermions

Sparticles sector :

$$M_{\tilde{f}}^2 = \begin{pmatrix} m_{\text{soft}}^2 + m_{\tilde{f}}^2 + M_{Z^0}^2 \cos 2\beta (l_f^3 - e_f \sin^2 \theta_W) + \Delta_f & m_f (A_f - \mu (t_\beta)^{-2} l_f^3) \\ m_f (A_f - \mu (t_\beta)^{-2} l_f^3) & m_{\text{soft}}^2 + M_{Z^0}^2 \cos 2\beta (l_{\tilde{f}}^3 - e_{\tilde{f}} \sin^2 \theta_W) + m_{\tilde{f}}^2 + \Delta_{\tilde{f}} \end{pmatrix}$$

where $\Delta_f = \frac{1}{2} g_1'^2 Q_f' (Q_{H_d}' v_d^2 + Q_{H_u}' v_u^2 + Q_S' v_s^2) \Rightarrow$ Coannihilations :

$\theta_{E_6} > 0$: generally \tilde{t}_1

$\theta_{E_6} < 0$: generally RH down squarks