

# Lectures on Neutrino Physics

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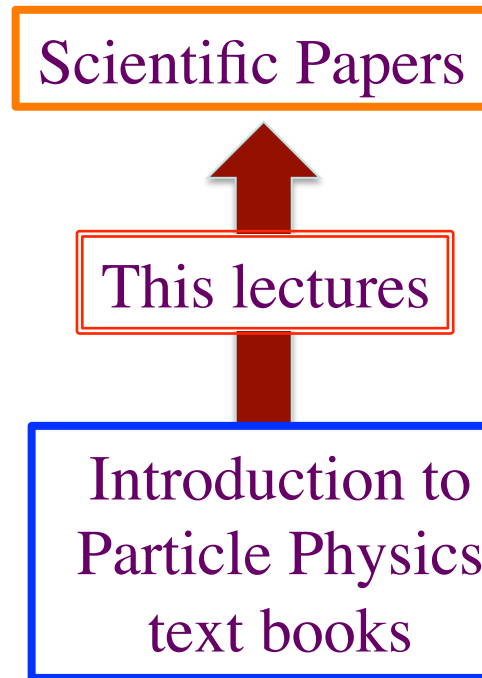
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France Asia Particle Physics School

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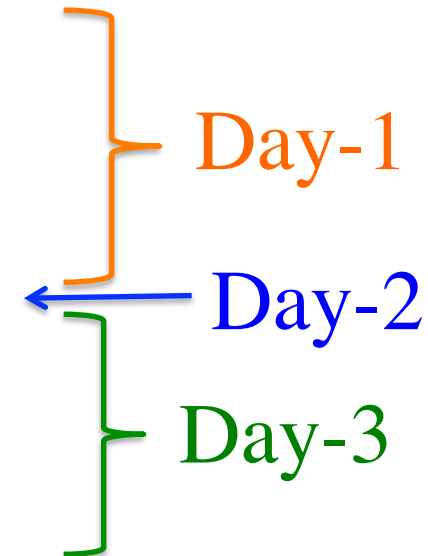
18-20/Oct./2011

- \* This lecture is intended to give intuitive understanding of neutrino physics for students and young physicists of other field.
- \* I will try to make this lecture to be a bridge between general text books and scientific papers.
- \* 3 lectures are very short to mention about all the varieties of neutrino physics and only limited but important topics are mentioned.



# Contents

- \* History
- \* Neutrinos in the Standard Model
- \* Neutrino Oscillations (Main)
- \* Double Beta decays
- \* Prospects



# What is known for neutrinos

Only a few things are known about neutrinos.

.... There is much room to study.

PDG2010

## Neutrino Properties

See the note on "Neutrino properties listings" in the Particle Listings.

Mass  $m < 2 \text{ eV}$  (tritium decay)

Mean life/mass,  $\tau/m > 300 \text{ s/eV}$ , CL = 90% (reactor)

Mean life/mass,  $\tau/m > 7 \times 10^9 \text{ s/eV}$  (solar)

Mean life/mass,  $\tau/m > 15.4 \text{ s/eV}$ , CL = 90% (accelerator)

Magnetic moment  $\mu < 0.54 \times 10^{-10} \mu_B$ , CL = 90% (solar)

## Number of Neutrino Types

Number  $N = 2.984 \pm 0.008$  (Standard Model fits to LEP data)

Number  $N = 2.92 \pm 0.05$  ( $S = 1.2$ ) (Direct measurement of invisible Z width)

## Neutrino Mixing

The following values are obtained through data analyses based on the 3-neutrino mixing scheme described in the review "Neutrino Mass, Mixing, and Oscillations" by K. Nakamura and S.T. Petcov in this *Review*.

$$\sin^2(2\theta_{12}) = 0.87 \pm 0.03$$

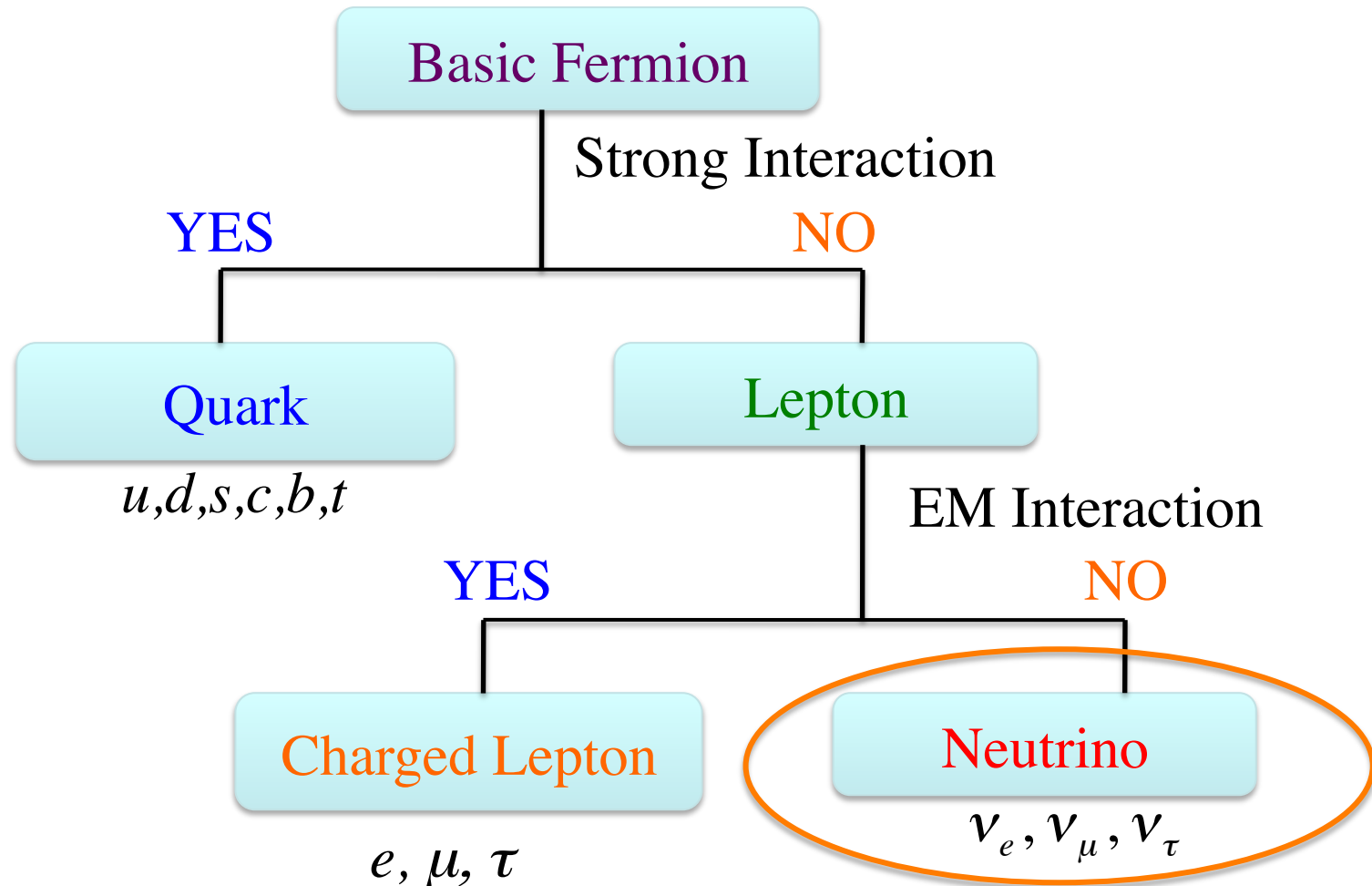
$$\Delta m_{21}^2 = (7.59 \pm 0.20) \times 10^{-5} \text{ eV}^2$$

$$\sin^2(2\theta_{23}) > 0.92 \text{ [}^l\text{]}$$

$$\Delta m_{32}^2 = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2 \text{ [}^l\text{]}$$

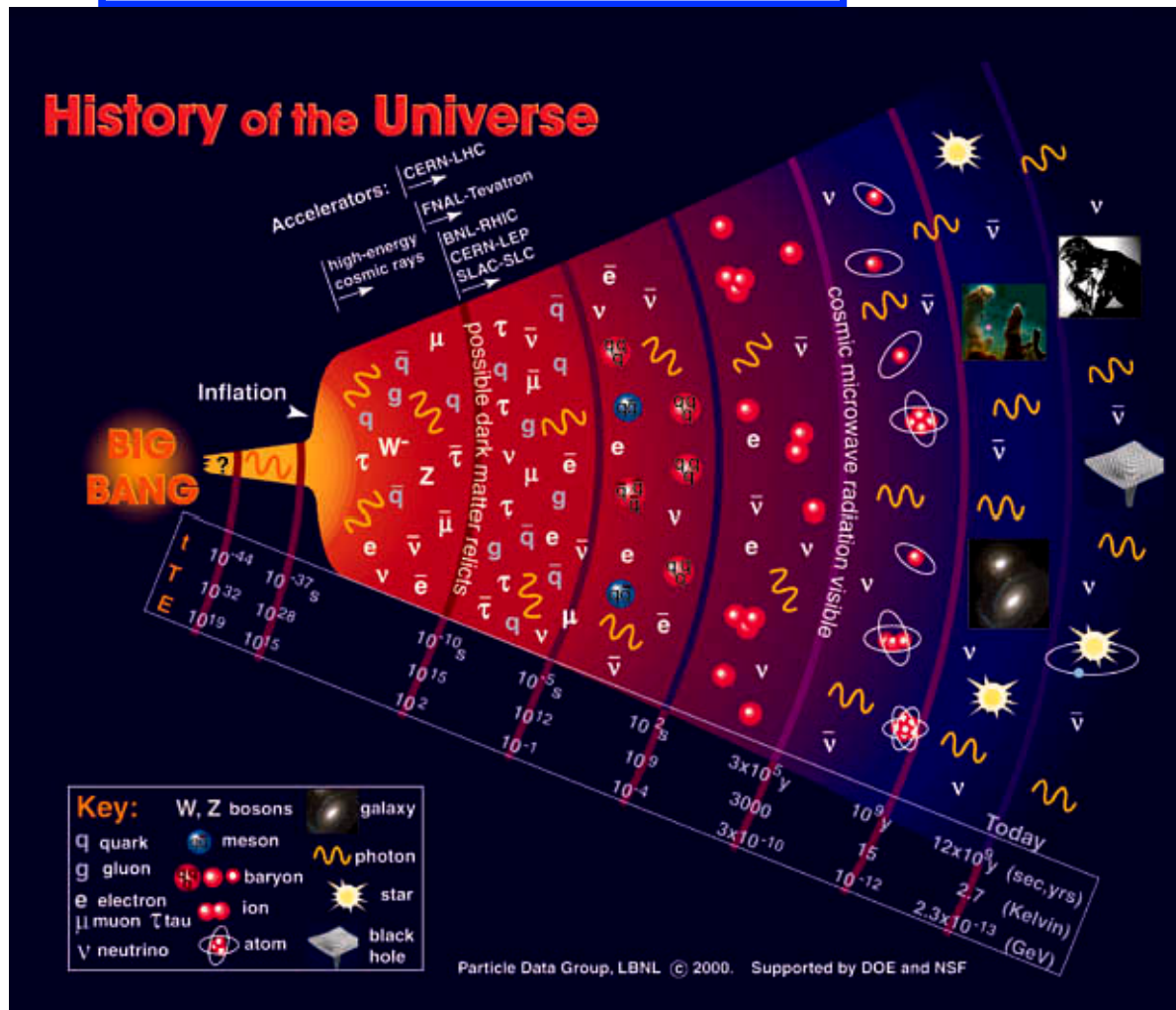
$$\sin^2(2\theta_{13}) < 0.15, \text{ CL} = 90\%$$





We call Fermions which do not perform strong nor EM interaction,  
**Neutrinos**

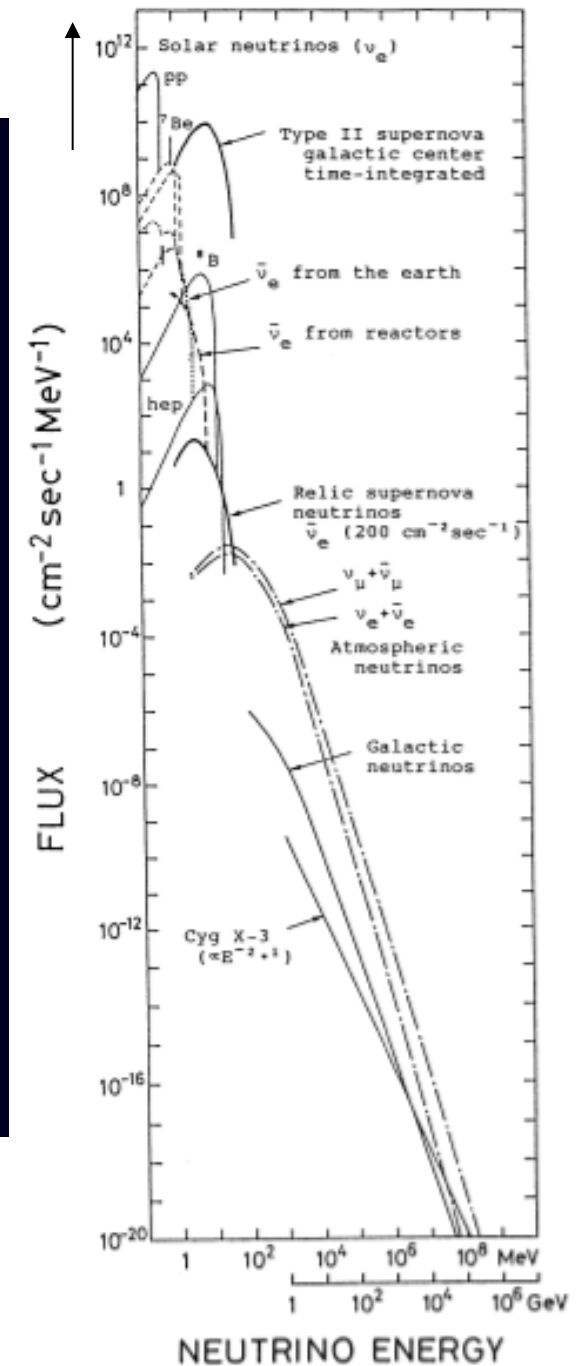
# Neutrinos in Nature



$e^+e^- \rightarrow \nu\bar{\nu}$  in early universe.

Most abundant next to photons ( $\sim 100\nu/\text{cc}$ )

big bang  $\nu$



# $\nu$ Timeline

(years are approximate)

1899 Discovery of $\beta$ -decay	[Rutherford]
1914 $\beta$ -ray has continuous energy spectrum	[Chadwick]
1930 Neutrino hypothesis	[Pauli]
1956 1 <sup>st</sup> Evidence of neutrino @ reactor	[Reines & Cowan]
1961 Discovery of $\nu_\mu$	[Shwartz, Ledermann, Steinberger]
1969~ Deficit of solar neutrino	[Davis]
1977 Discovery of $\tau$ lepton ( indirect evidence of $\nu_\tau$ )	[Perl]
1985 Proposal of MSW effect	[Mikheyev, Smirnov, Wolfenstein]
1987 Detection of neutrinos from SN1987A	[Koshiba]
1989 $N_\nu=3$ by $Z^0$ shape	[LEP]
1995 Nobel prize to Reines	
(1996, 1997) Claim of $\nu_\mu \rightarrow \nu_e$ oscillation	[LSND])
1998 1 <sup>st</sup> evidence of neutrino oscillation by atmospheric $\nu$	[SuperKamiokande]
2000 Direct evidence of $\nu_\tau$	[DONUT]
(2001 Claim of neutrinoless $\beta\beta$ decay	[Klapdor])
2002 Nobel prize to Davis & Koshiba	
2002 Flavor transition	[SNO]
Reactor Neutrino Deficit	[KamLAND]
2004 $\nu_\mu$ disappearance @ Accelerator	[K2K]
2010 $\nu_\mu \rightarrow \nu_\tau$	[OPERA]
2011 Indication of $\nu_e$ appearance @ Accelerator	[T2K]

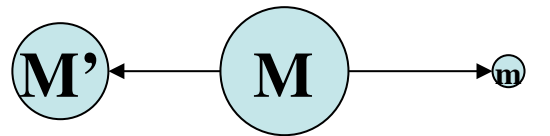
# 1<sup>st</sup> Indication of Neutrino

~1914, an anomaly found

(The 1st **anomaly** in neutrino which lead great discovery.)

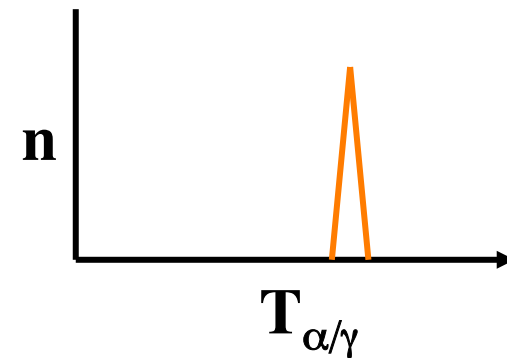


$\gamma$  &  $\alpha$  decays  $\rightarrow$  The energy of the decay particle is unique



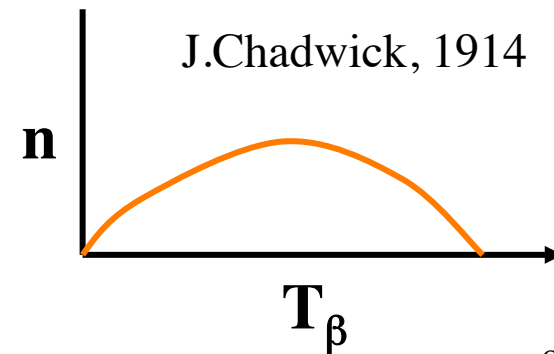
A diagram showing a central circle labeled 'M' with an arrow pointing left to a circle labeled 'M'' and an arrow pointing right to a small circle labeled 'm'.

$$T_{\alpha} = \frac{(M - m)^2 - M'^2}{2M}$$



However, for  $\beta$ -decays,  
it is continuous.

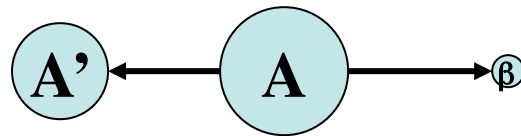
Why??



# Neutrino Hypothesis



\* Energy conservation law is broken (N.Bohr, 1932)

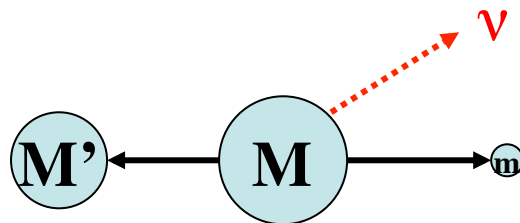


$$E_{\beta} \neq E_A - E_{A'}$$



\*  $\beta$ -decay is a 3 body reaction (W.Pauli, 1930)

$$A \rightarrow B + \beta^{-} + \nu$$



$$E_{\beta} = E_A - E_{A'} - E_{\nu} < E_A - E_{A'}$$

Neutrino Hypothesis

4/Dec./1930  
Letter from Pauli  
to participants of  
a conference.

*Dear Radioactive Ladies and Gentlemen,*

*As the bearer of these lines, to whom I graciously ask you to listen, will explain to you in more detail, how because of the "wrong" statistics of the N and  $\text{Li}^6$  nuclei and the continuous beta spectrum, I have hit upon a desperate remedy to save the "exchange theorem" of statistics and the law of conservation of energy. Namely, the possibility that there could exist in the nuclei electrically neutral particles, that I wish to call neutrons, which have spin  $1/2$  and obey the exclusion principle and which further differ from light quanta in that they do not travel with the velocity of light. The mass of the neutrons should be of the same order of magnitude as the electron mass and in any event not larger than 0.01 proton masses> The continuous beta spectrum would then become understandable by the assumption that in beta decay a neutron is emitted in addition to the electron such that the sum of the energies of the neutron and the electron is constant...*

*I agree that my remedy could seem incredible because one should have seen those neutrons very earlier if they really exist. But only the one who dare can win and the difficult situation, due to the continuous structure of the beta spectrum, is lighted by a remark of my honoured predecessor, Mr Debye, who told me recently in Bruxelles: "Oh, It's well better not to think to this at all, like new taxes". From now on, every solution to the issue must be discussed. Thus, dear radioactive people, look and judge. Unfortunately, I cannot appear in Tübingen personally since I am indispensable here in Zurich because of a ball on the night of 6/7 December. With my best regards to you, and also to Mr Back.*

*Your humble servant  
. W. Pauli*

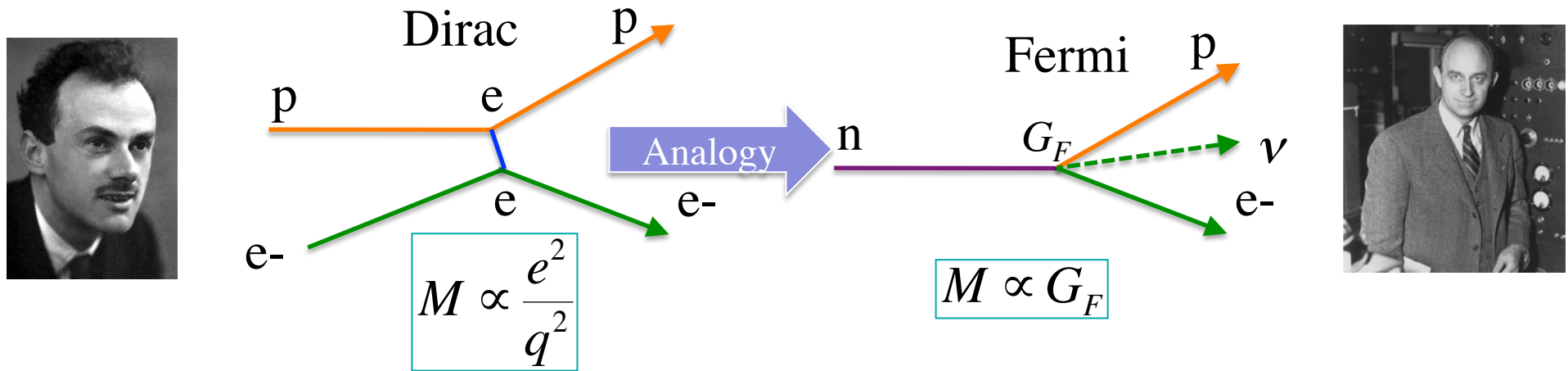


## Expected properties of $\nu$ from $\beta$ -decays

- (1)  $Q=0$   $\leftarrow$  charge conservation
- (2)  $s=1/2$   $\leftarrow$  spin conservation
- (3) mass is small if exists  $\leftarrow$  maximum energy of  $\beta$ -rays.
- (4) Interact very weakly  $\leftarrow$  lifetime of  $\beta$ -decays.



# How $\nu$ -N cross section was estimated in early days. Fermi's model



Various  $\beta^\pm$ -decays & Electron capture  $\rightarrow G_F \sim 10^{-11}/\text{MeV}^2$

$$\text{Then, } \sigma(\bar{\nu} + p \rightarrow e^+ + n) = \frac{G_F^2}{\pi} p_{CM}^2 \sim 10^{-20} [b]!!$$

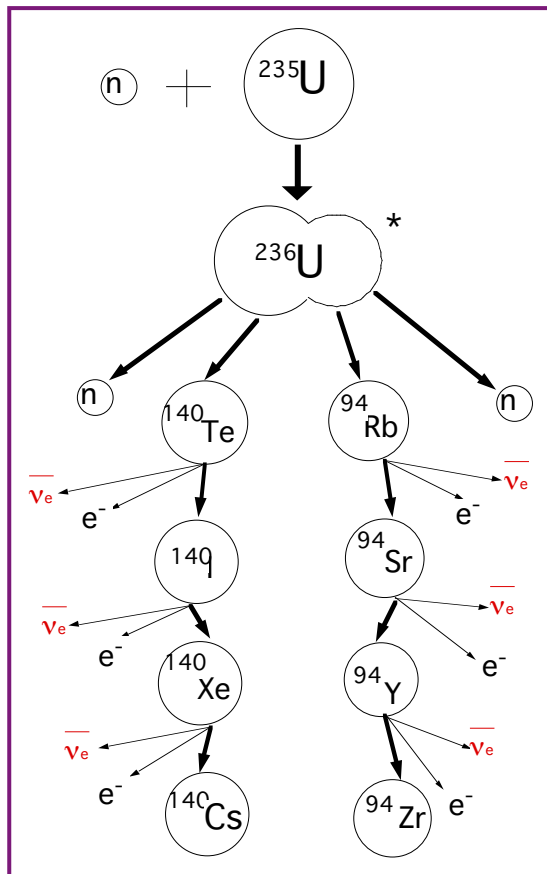
"I did something a physicist should never do. I predicted something which will never be observed experimentally..". (W.Pauli)

"There is no practically possible way of observing the neutrino"  
(Bethe & Peierls, 1934)



Then 30 years had passed ....

## Discovery of $\nu$



$E_\nu \sim \text{MeV}$

Very strong  $\nu$  sources are necessary,  
 ← Chain reactions of nuclear fissions.

Energy release:  
 200MeV/fission

$\beta$ -decays:  
 $\sim 6 \nu/\text{fission}$

$$1.9 \times 10^{11} \nu/\text{J}$$



Reactor or Nuclear Explosion

An early idea to detect  $\nu$  (not realized)



Reines &amp; Cowan

# Nuclear Explosion

## Vacuum shaft

# Neutrino Detector

# Free fall to prevent the shock wave.

by Reines &amp; Cowan

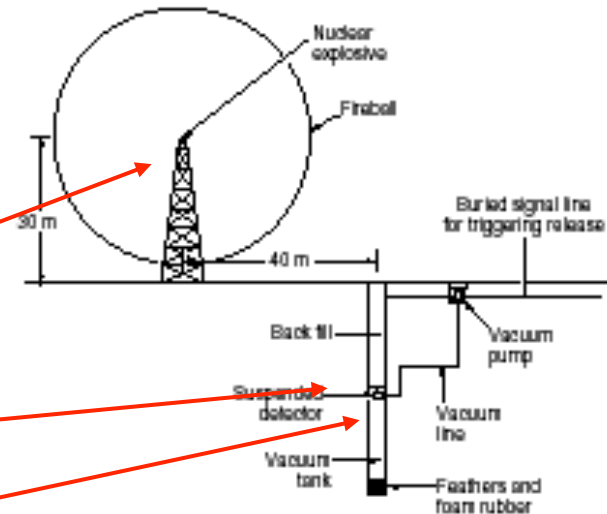


Figure 1. Detecting Neutrinos from a Nuclear Explosion  
Antineutrinos from the fireball of a nuclear device would impinge on a liquid scintillation detector suspended in the hole dug below ground at a distance of about 40 meters from the 30-meter-high tower. In the original scheme of Reines and Cowan, the antineutrinos would induce inverse beta decay, and the detector would record the positrons produced in that process. This figure was redrawn courtesy of Smithsonian Institution.

=> Physicists make use of everything available

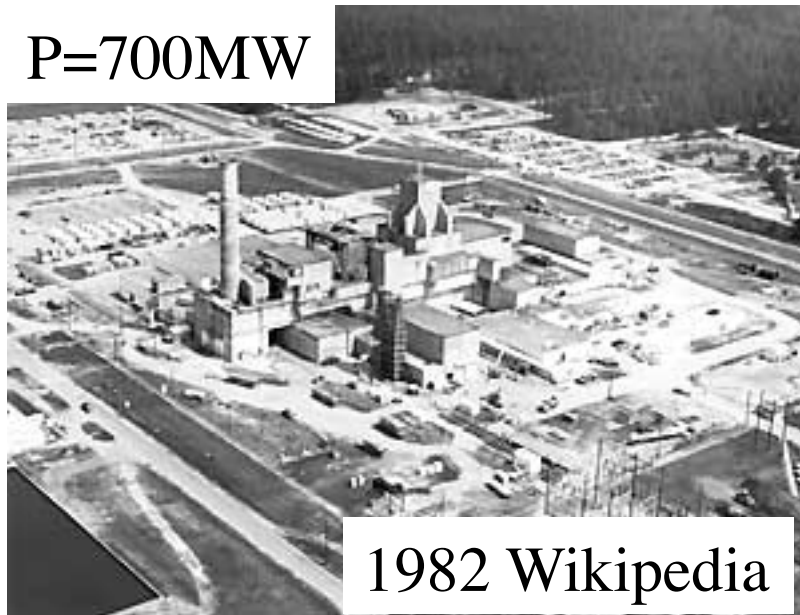
While preparing the experiment, they realized nuclear reactor is more relevant to perform experiment.

Then they moved to a Savannah River Reactor

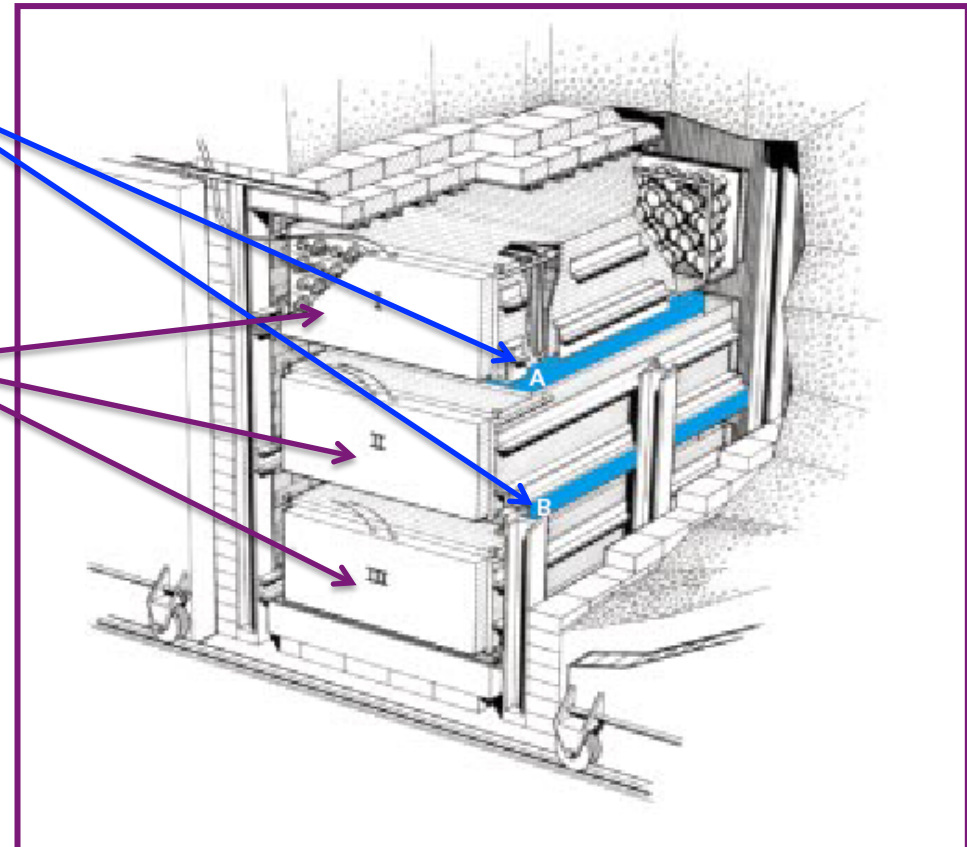
200L Cd loaded water tanks

1400L liquid scintillator tanks

P=700MW



1982 Wikipedia



**Figure 4. The Savannah River Neutrino Detector—A New Design**

The neutrino detector is illustrated here inside its lead shield. Each of two large, flat plastic tanks (pictured in light blue and labeled A and B) was filled with 200 liters of water. The protons in the water provided the target for inverse beta decay; cadmium chloride dissolved in the water provided the cadmium nuclei that would capture the neutrons. The target tanks were sandwiched between three scintillation detectors (I, II, and III). Each detector contained 1,400 liters of liquid scintillator that was viewed by 110 photomultiplier tubes. Without its shield, the assembled detector weighed about 10 tons.

# Principle of $\nu$ detection

\*  $\nu$  flux:

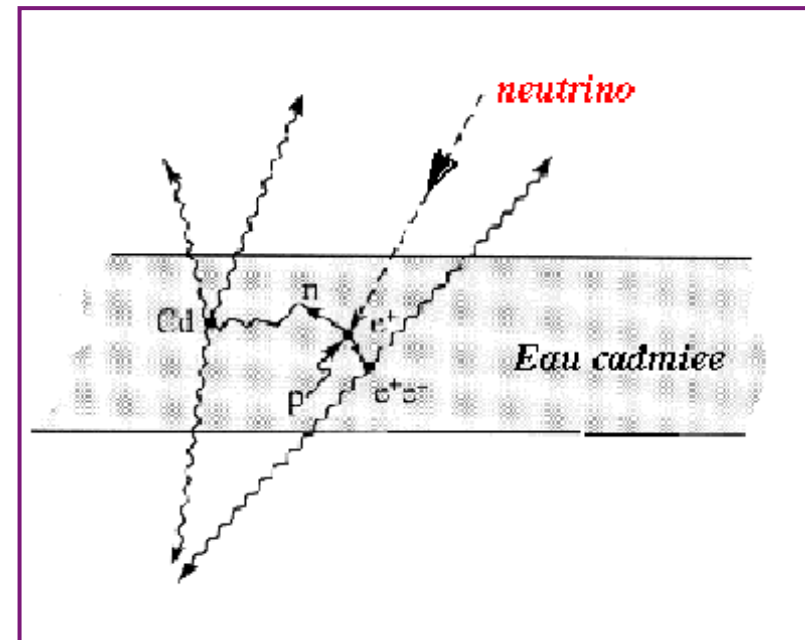
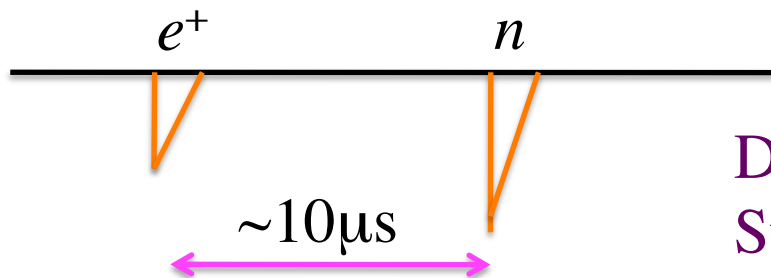
@15m from Savannah Liver P reactor core. (P=700MW)

flux  $\sim 5 \times 10^{12} \nu / \text{cm}^2 / \text{s}$

\* Detection Principle:

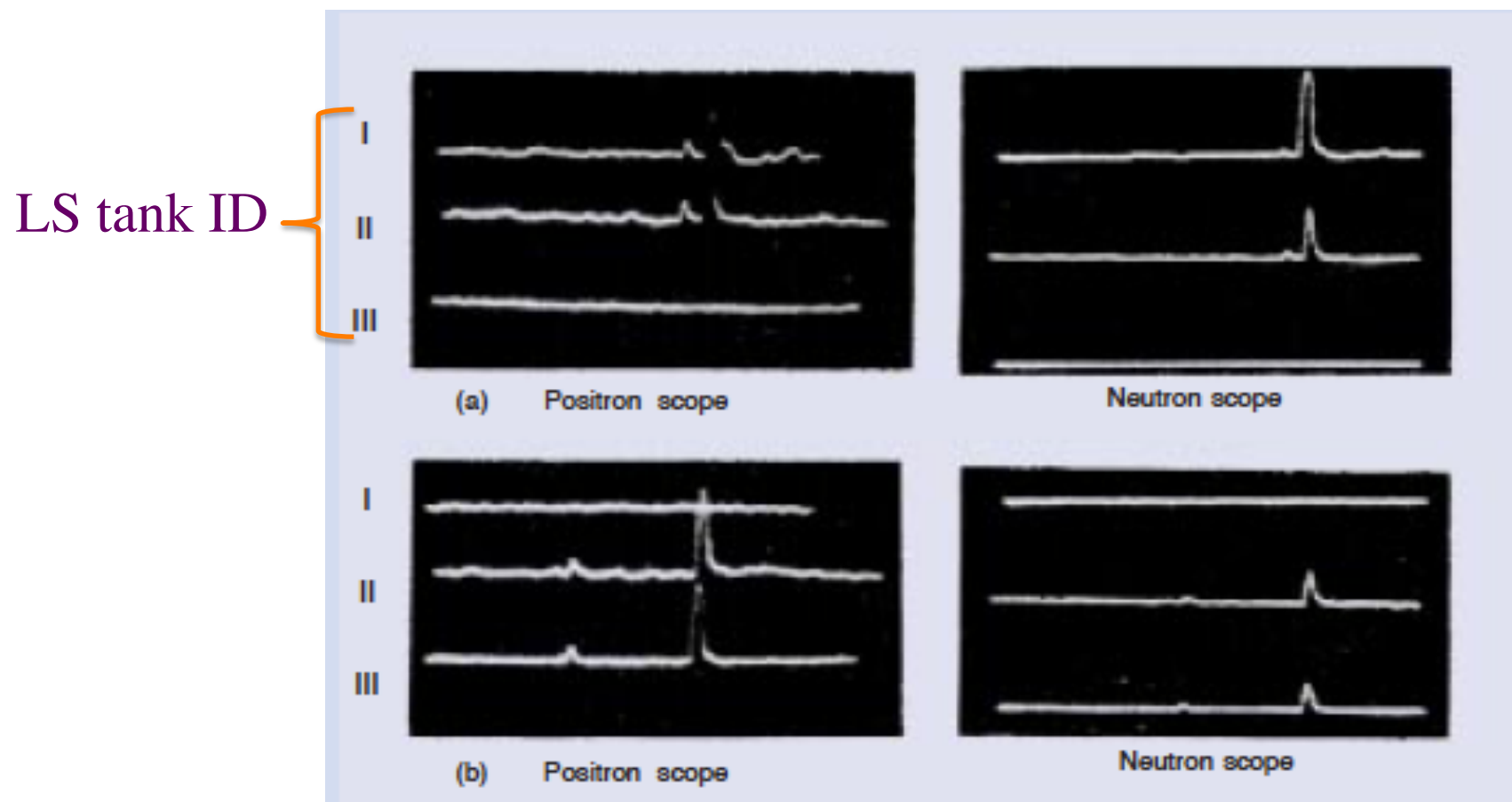
$$\bar{\nu}_e + p \rightarrow e^+ + n \Rightarrow$$

$$\left\{ \begin{array}{l} e^+ + e^- \rightarrow 2\gamma (0.5 \text{ MeV}) \\ n + \text{Cd} \rightarrow \text{Cd}^* \rightarrow \text{Cd} + n\gamma (9 \text{ MeV}) \end{array} \right.$$



Delayed Coincidence Technique  
Still used in modern experiments

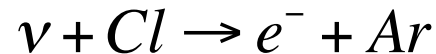
## 2 examples of delayed coincidence



<http://library.lanl.gov/cgi-bin/getfile?00326606.pdf#search='delayed%20coincidence%20cadmium%20neutrino'>

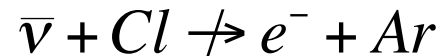
## An episode

At the same time of Reines& Cowan, R.Davis and L.Alvarez performed neutrino experiment at a Savannah river reactor, too. Their detection principle was,



However, they failed to detect positive result.

But this actually means the reactor neutrino (anti neutrino) dose not cause the reaction



and neutrino and anti-neutrino are different particles (concept of that time)

Later on, Davis also won Nobel prize by detecting solar neutrinos with the same technique.

Lessons : Negative result can be an important signature.  
: Hanging on is important for success.

## Discovery of $\mu$ -neutrino



Lederman, Schwartz, Steinberger

\* Neutrino Source:

@ Brookhaven AGS  $p(15\text{GeV}) + Be \rightarrow \pi + X$

$$\pi \rightarrow \mu + \nu$$

$\pi$  decays with 21m decay space.  $\frac{\Gamma(\pi \rightarrow e + \nu)}{\Gamma(\pi \rightarrow \mu + \nu)} \approx 10^{-4}$

99.99% of neutrinos are associated with muon production

## Detection of neutrino

\* Target: 90 x 2.5cmt Al slab

Looked for  $\nu + Al \rightarrow \begin{cases} \mu + X \\ e + X \end{cases}$

$\mu$  signal  $\Rightarrow$  a single track

$e$  signal  $\Rightarrow$  EM shower

They observed

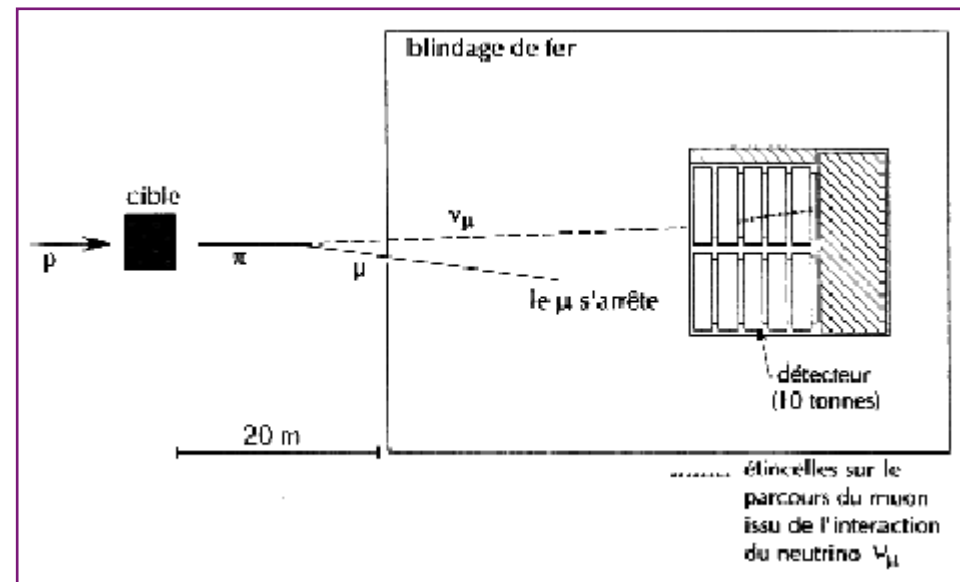
34 single track  $\mu$  events

22  $\mu + X$

6 backgrounds (not like  $e$ )

$\rightarrow$  The neutrinos from  $\beta$ -decay and  $\pi$  decay are different particle

spark chambers





# $\tau$ neutrino

(2000 DONUT group)

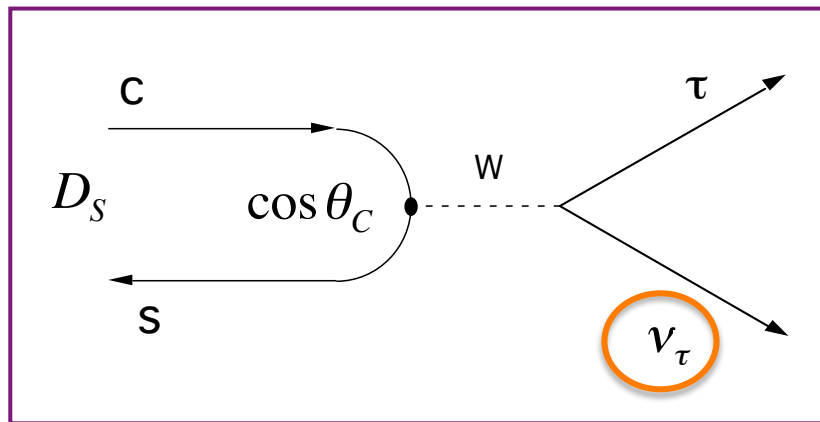
## \* Production of $\tau$ neutrino

FNAL TEVATRON

$$p(E = 800 \text{ GeV}) + W \rightarrow D_S^\pm (m = 1.97 \text{ GeV}) + X$$

Then

$$D_S^\pm \rightarrow \tau^\pm (m = 1.78 \text{ GeV}) + \nu_\tau \quad (Br \sim 4\%)$$



$$\begin{cases} m_{D_S} > m_\tau \\ \text{Cabibbo favor} \end{cases}$$

$$E_\nu \sim 70 \text{ GeV}, \quad \frac{\phi_{\nu_\tau}}{\phi_{\nu_{e,\mu}}} \sim 0.05$$

# Detection principle

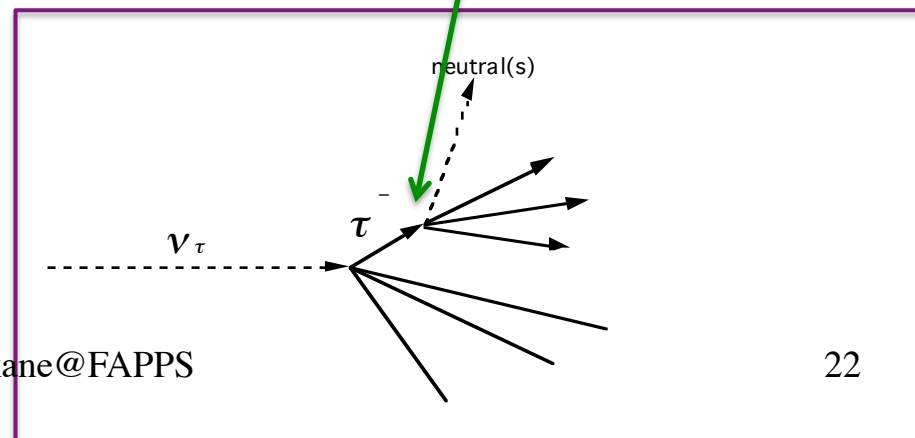
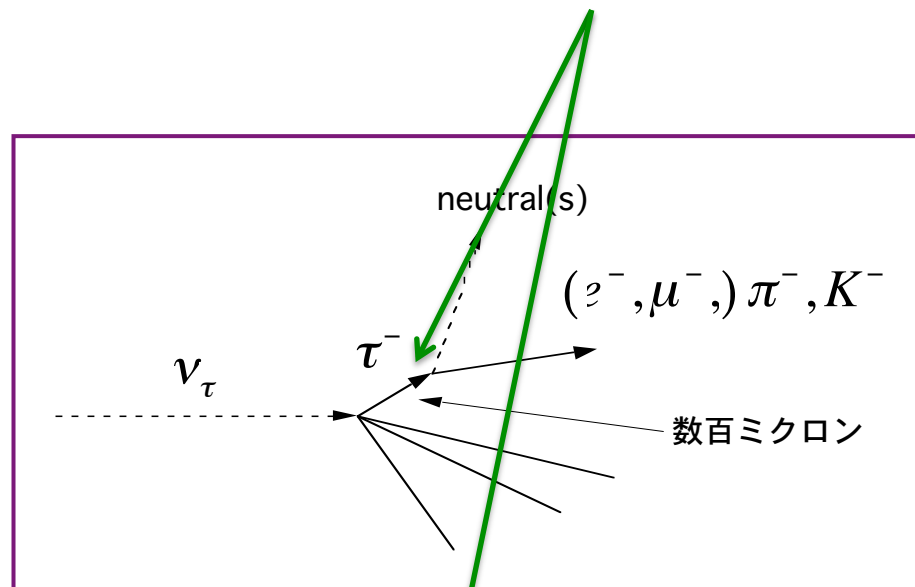
$$\nu_\tau + A \rightarrow \tau + X$$

$\tau$  decays after  $ct \sim 87 \mu\text{m}$ ,

85% for 1 prong mode

15% for 3 prongs mode.

Look for the "kink"



# the Detector

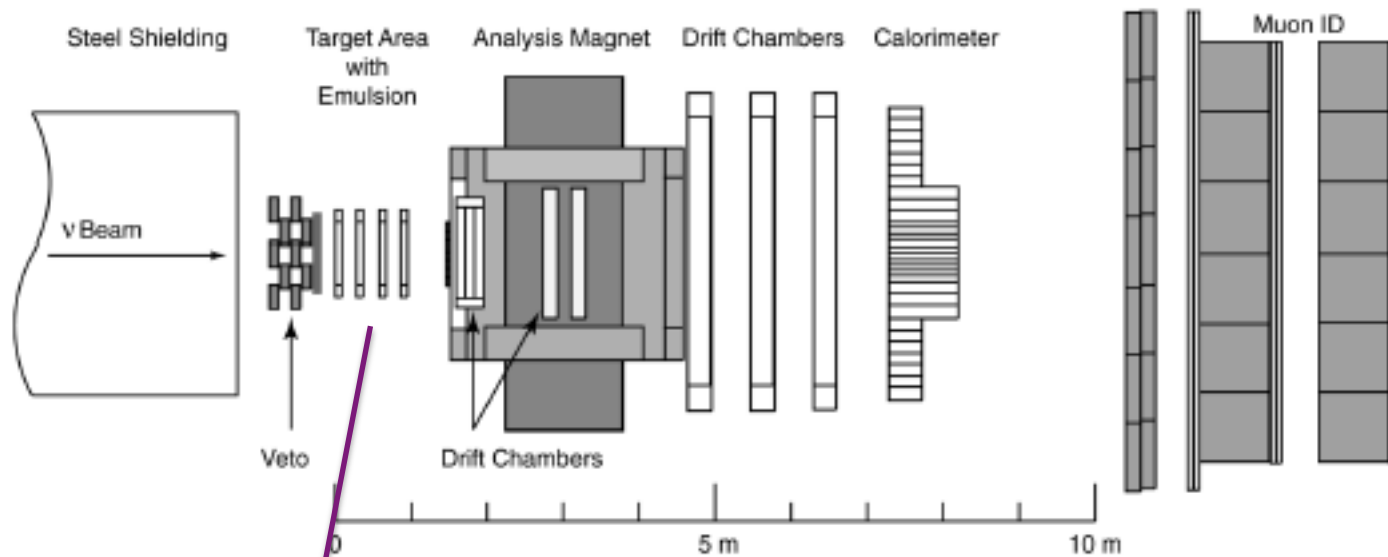


FIG. 4. Schematic plan view of the spectrometer. The neutrinos are incident from the left, emerging from the passive shield. The design is relatively compact to optimize identification of electrons and muons.

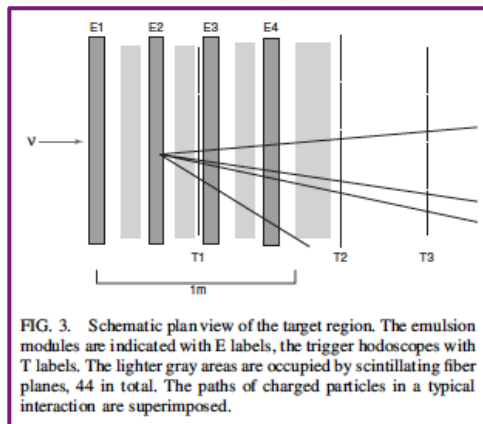
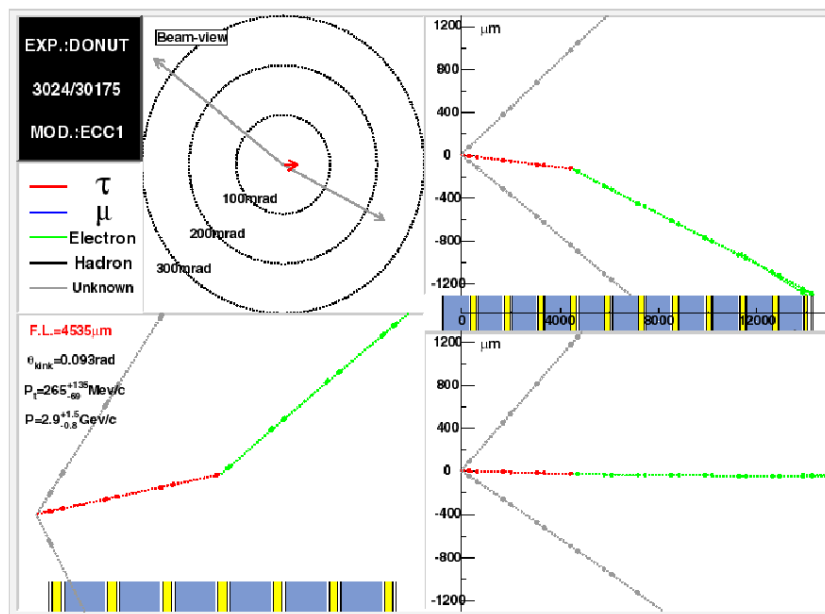


FIG. 3. Schematic plan view of the target region. The emulsion modules are indicated with E labels, the trigger hodoscopes with T labels. The lighter gray areas are occupied by scintillating fiber planes, 44 in total. The paths of charged particles in a typical interaction are superimposed.

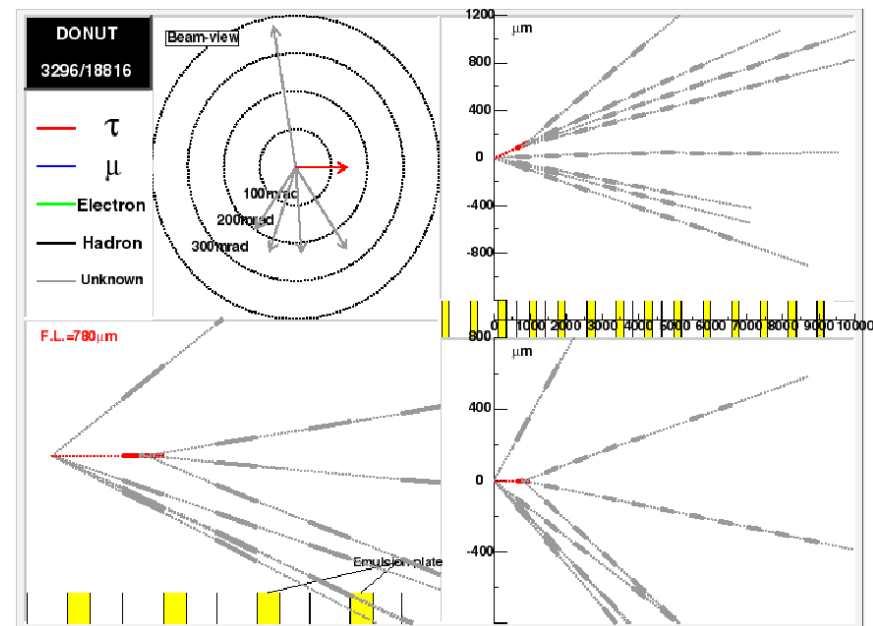
Nuclear Emulsion  
(~thick camera film)

position resolution  $< 1\mu\text{m}$

## Results (2000)



1prong



3 prongs

$4 \pm 0.44 \nu_{\tau}$  was observed in 1000 neutrino events. ( $9\sigma$  significance)

"We did R&D for  $\tau$ -neutrino detection around 1980 but once gave up because it seemed too difficult to success". K.Niwa

# Search for neutrino mass

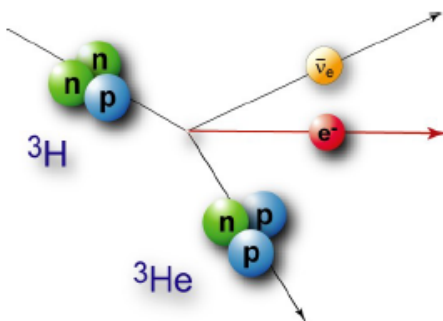
## $\beta$ -decay: absolute $\nu$ -mass

model independent, kinematics

status:  $m_\nu < 2.3$  eV

potential:  $m_\nu \approx 200$  meV

e.g.: KATRIN, MARE-II



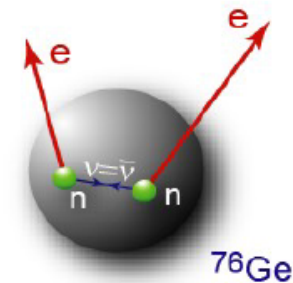
## $0\nu\beta\beta$ -decay: eff. Majorana mass

model-dependent (CP-phases)

status:  $m_{\beta\beta} < 0.35$  eV (evidence?)

potential:  $m_{\beta\beta} \approx 20$ -50 meV

e.g.: GERDA, CUORE, EXO, SNO+, Majorana, Nemo 3, COBRA, KamLAND-Zen



$m_\nu$

$m_{\beta\beta}$

neutrino mass  
measurements

$\Sigma m_i$

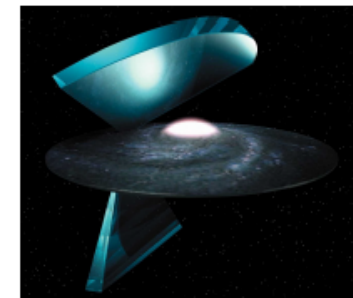
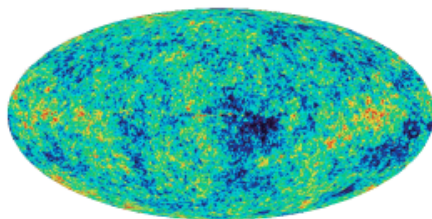
## cosmology: $\nu$ hot dark matter $\Omega_\nu$

model dependent, analysis of LSS data

status:  $\Sigma m_\nu < 0.44$  eV (Hannestad et al., JCAP08(2010)001)

potential:  $\Sigma m_\nu \approx 20$ -50 meV

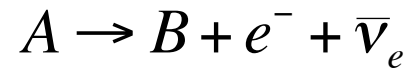
e.g.: WMAP, SDSS, LSST, Planck



# Direct neutrino mass detection

electron neutrino

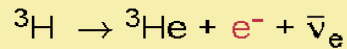
Principle



$$N(p_e)dp_e \propto p_e^2 (E_0 - E_e)^2 \sqrt{1 - \left( \frac{m_\nu}{E_0 - E_e} \right)^2} dp_e$$

$$E_e^{MAX} = E_0 \rightarrow E_0 - m_\nu$$

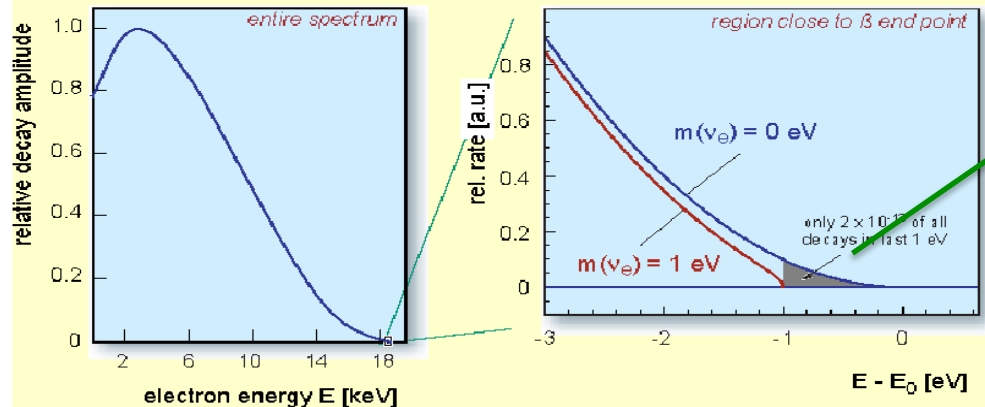
tritium  $\beta$ -decay and the neutrino rest mass



superallowed

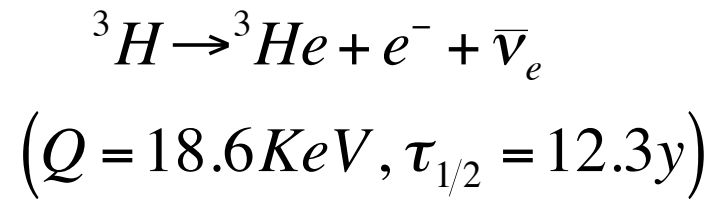
half life :  $t_{1/2} = 12.32 \text{ a}$

$\beta$  end point energy :  $E_0 = 18.57 \text{ keV}$



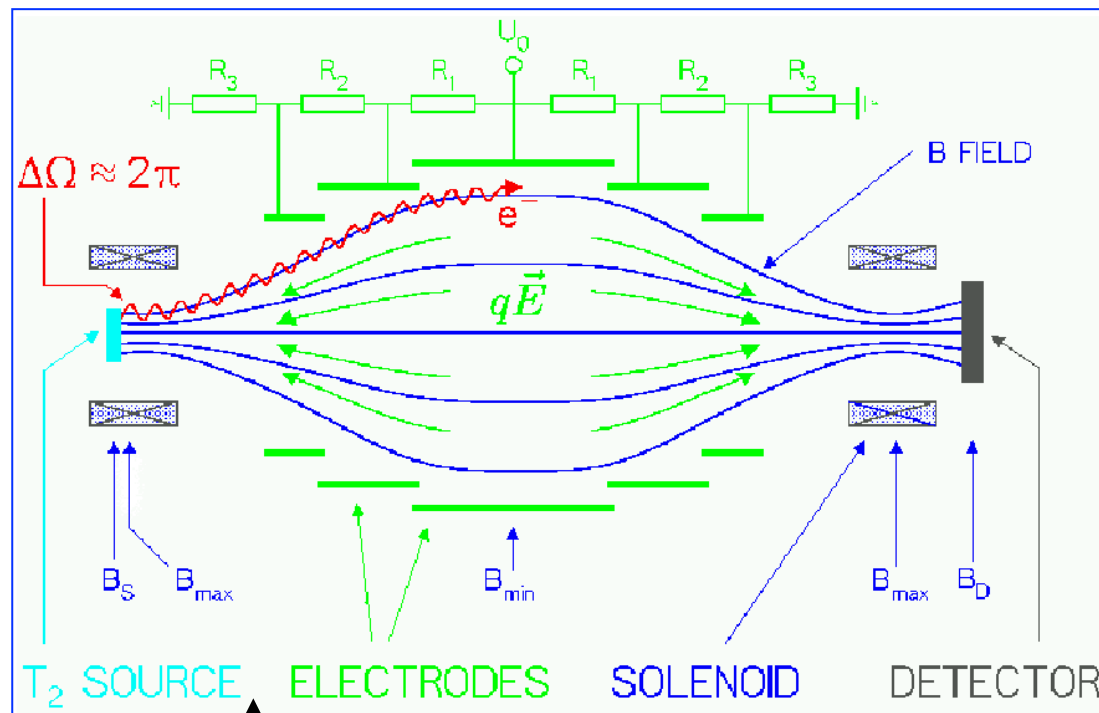
Distortion of energy spectrum at the end point

Source: Tritium

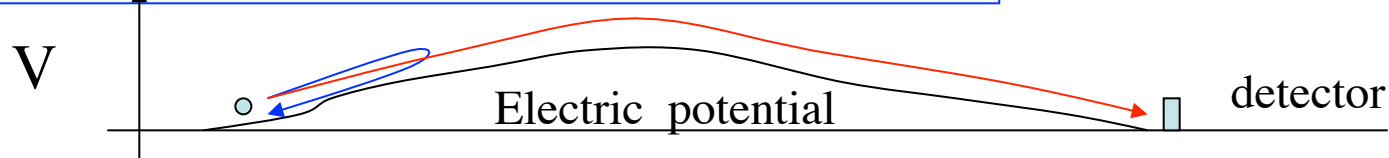


$E_0 = \text{small} \rightarrow \text{good } m_\nu \text{ sensitivity}$   
 $\text{Lifetime} \rightarrow \text{reasonably short \& long}$   
 $Z = \text{small} \rightarrow \text{small correction}$

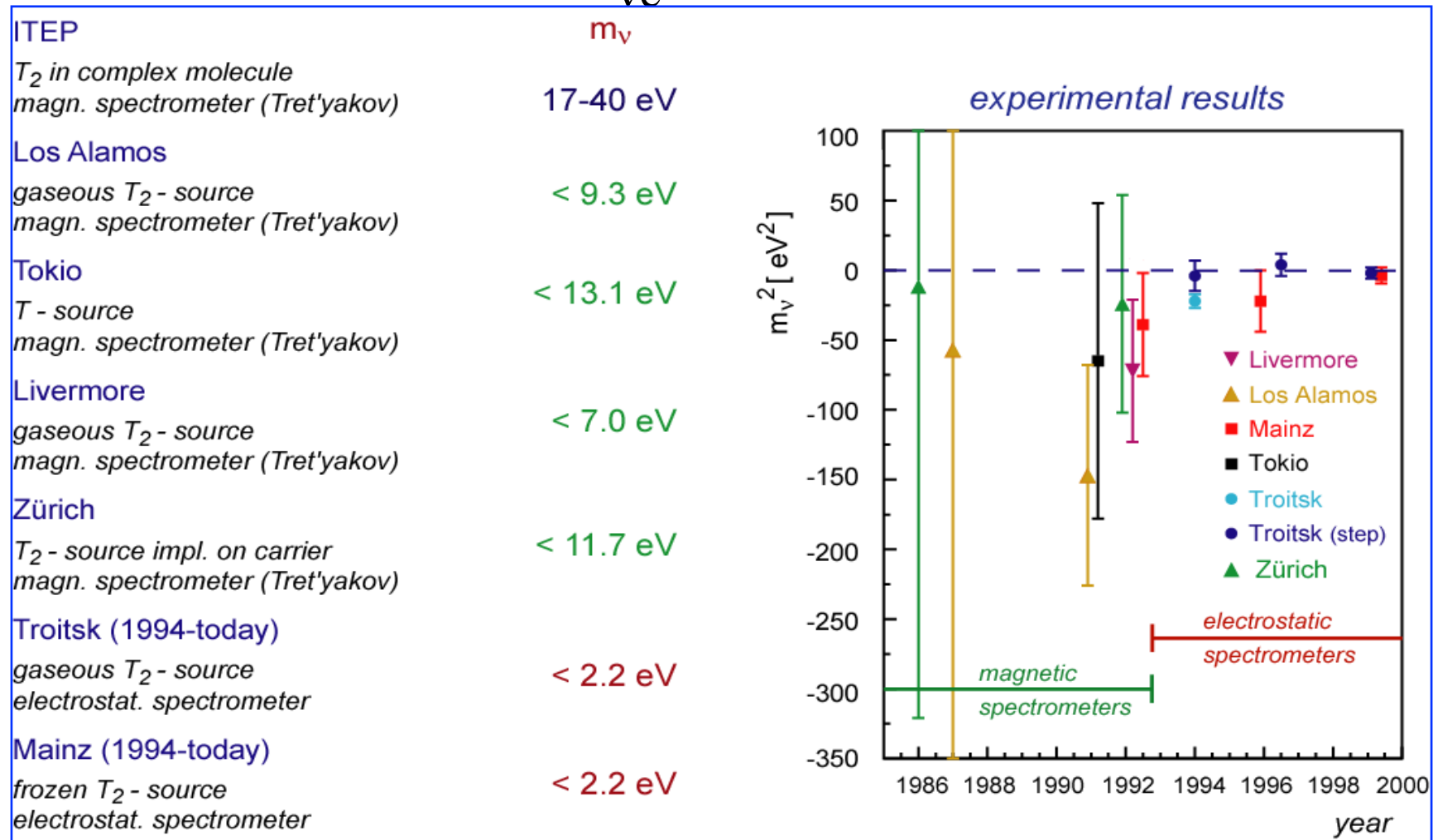
} ideal isotope to seek for  $m_\nu$



- \* Strong Source
- \* Large Acceptance
- \* Energy measurement by Electric potential



# $m_{\nu e}$ Results





# From current to future experiments

## Mainz:

$$m_\nu^2 = -1.2(-0.7) \pm 2.2 \pm 2.1 \text{ eV}^2 \quad m_\nu^2 = -2.3 \pm 2.5 \pm 2.0 \text{ eV}^2$$

$$m_\nu < 2.2(2.3) \text{ eV (95\%CL)}$$

C. Weinheimer, Nucl. Phys. B (Proc. Suppl.) 118 (2003) 279

C. Kraus, EPS HEP2003 (neighbour excitations self-consistent)

## Troitsk:

$$m_\nu^2 = -2.3 \pm 2.5 \pm 2.0 \text{ eV}^2$$

$$m_\nu < 2.1 \text{ eV (95\%CL)}$$

V. Lobashev, private communication

(allowing for a step function near endpoint)

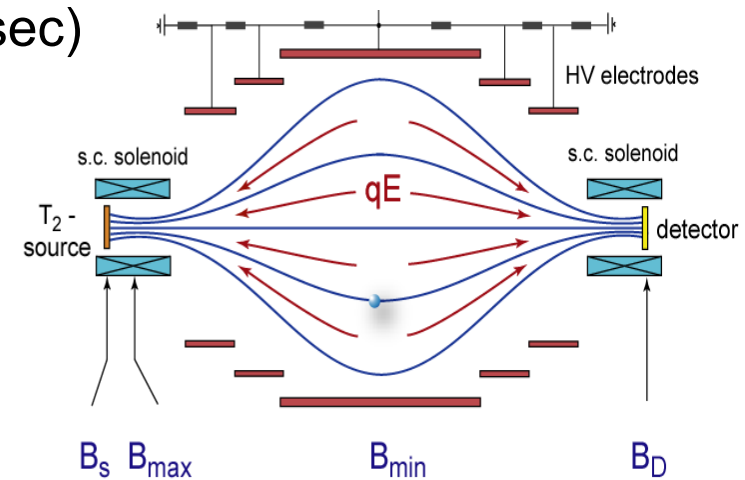
**aim:** improvement of  $m_\nu$  by one order of magnitude ( $2\text{eV} \rightarrow 0.2\text{eV}$ )  
 $\rightarrow$  improvement of uncertainty on  $m_\nu^2$  by 100 ( $4\text{eV}^2 \rightarrow 0.04\text{eV}^2$ )

## statistics:

- stronger Tritium source ( $>> 10^{10} \beta'/\text{sec}$ )
- longer measurement  
 ( $\sim 100 \text{ days} \rightarrow \sim 1000 \text{ days}$ )

## energy resolution:

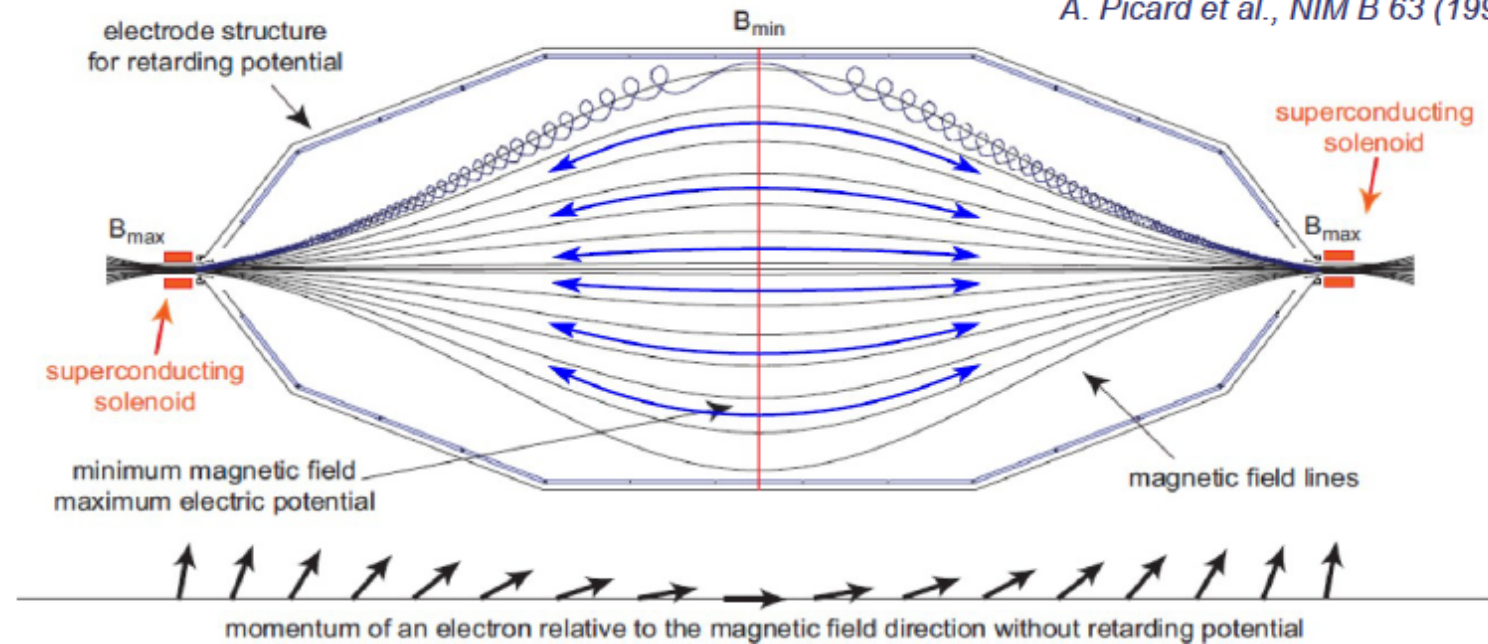
- $\Delta E/E = B_{\min}/B_{\max}$   
 $\rightarrow$  spectrometer with  $\Delta E = 1\text{eV}$   
 $\rightarrow \varnothing 10\text{m UHV vessel}$



# KATRIN: MAC-E filter concept

## Magnetic Adiabatic Collimation with Electrostatic Filter

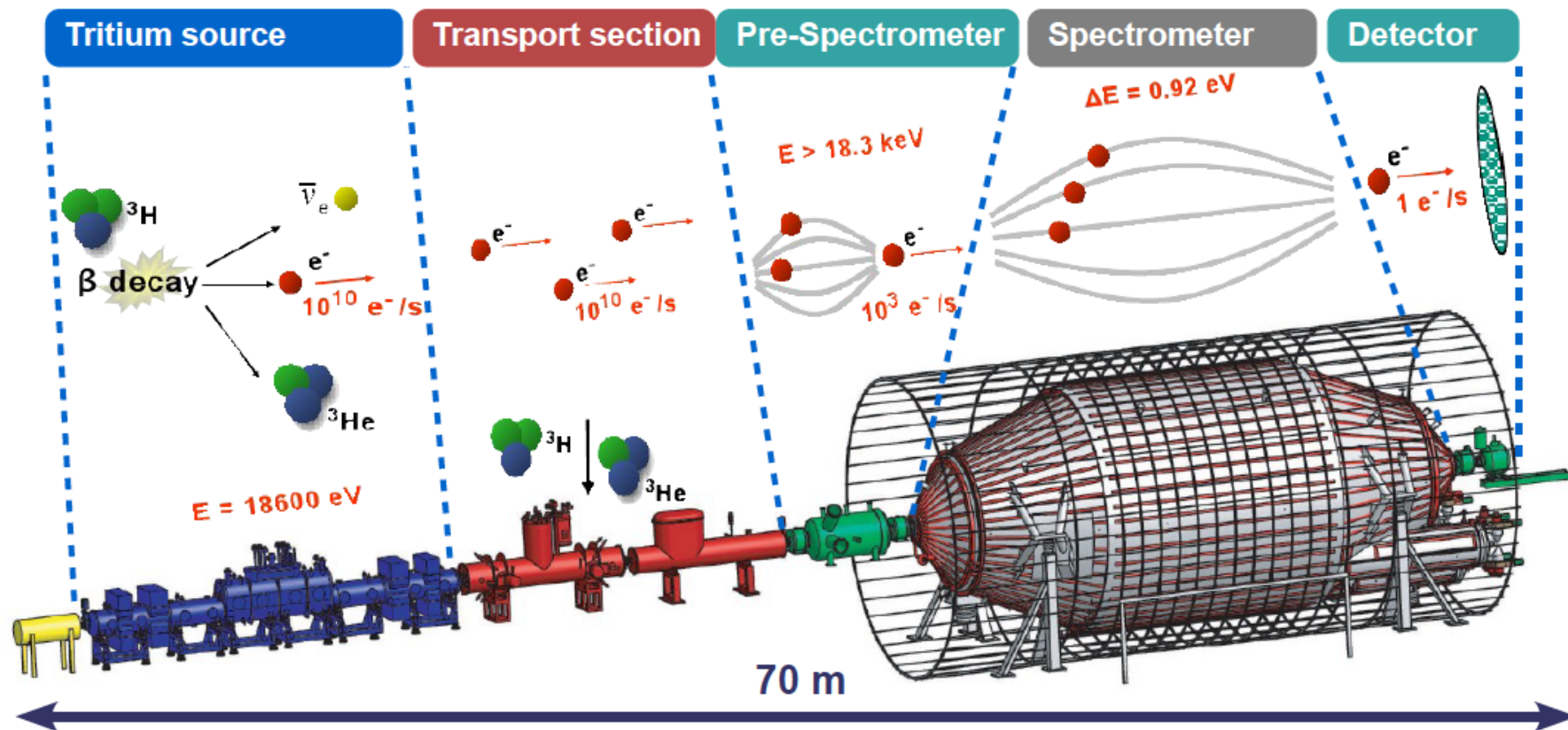
A. Picard et al., NIM B 63 (1992)



- adiabatic transport  $\rightarrow \mu = E_{\perp} / B = \text{const.}$
- $B$  drops by  $2 \cdot 10^4$  from solenoid to analyzing plane  $\rightarrow E_{\perp} \rightarrow E_{\parallel}$
- only electrons with  $E_{\parallel} > eU_0$  can pass the retardation potential
- Energy resolution  $\Delta E = E_{\perp, \text{max, start}} \cdot B_{min} / B_{max} < 1 \text{ eV}$

Volker Hannen, TAUP 2011 conference, 7.9.2011

# KATRIN experiment at KIT





## A famous picture



111018

suekane@FAPPS

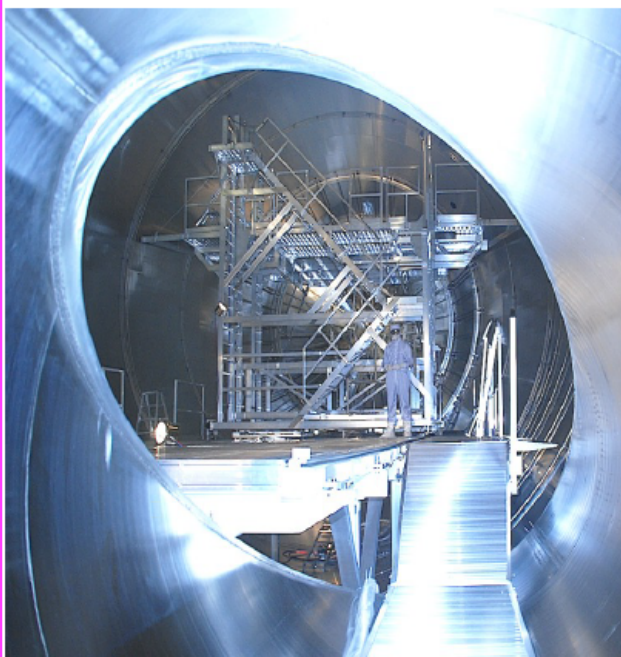
32



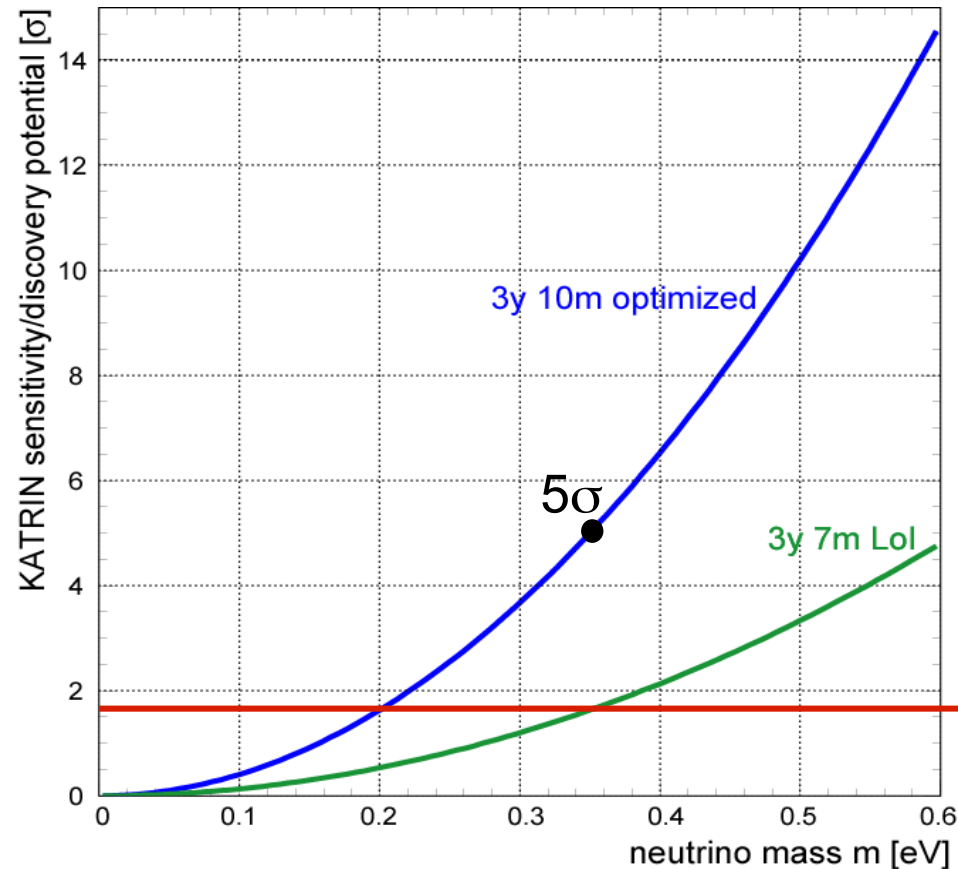
# Some KATRIN highlights

## Main spectr. inner wire electrode

- Purpose: - electrostatic shielding of background electrons  
- shaping of electric fields  
- removal of trapped particles
- 224 of 248 double layer wire modules installed
- last 24 modules to be installed after preparation of the pump port region in fall 2011
- main spectrometer commissioning spring 2012



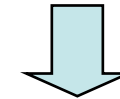
# KATRIN sensitivity & discovery potential



expectation:

after 3 full beam years

$$\sigma_{\text{syst}} \sim \sigma_{\text{stat}}$$



$$m_\nu = 0.35\text{eV} (5\sigma)$$

$$m_\nu = 0.3\text{eV} (3\sigma)$$

**discovery potential**

$$m_\nu < 0.2\text{eV} (90\%\text{CL})$$

**sensitivity**

# $\nu_\mu$ mass limit

PHYSICAL REVIEW D

VOLUME 53, NUMBER 11

1 JUNE 1996

## Upper limit of the muon-neutrino mass and charged-pion mass from momentum analysis of a surface muon beam

K. Assamagan,<sup>4,\*</sup> Ch. Brönnimann,<sup>1,2</sup> M. Daum,<sup>1</sup> H. Forrer,<sup>1,3,†</sup> R. Frosch,<sup>1</sup> P. Gheno,<sup>1</sup> R. Horisberger,<sup>1</sup>  
M. Janousch,<sup>1,2,‡</sup> P.-R. Kettle,<sup>1</sup> Th. Spirig,<sup>1,2,§</sup> and C. Wigger<sup>1,2</sup>

$$\pi(stop) \rightarrow \mu + \nu_\mu$$



$$m_\pi = \sqrt{m_\mu^2 + p_\mu^2} + \sqrt{m_\nu^2 + p_\mu^2}$$

→

$$m_\nu = 0.34 m_\pi \sqrt{1 - \frac{p_\mu}{p_\mu^0}}$$

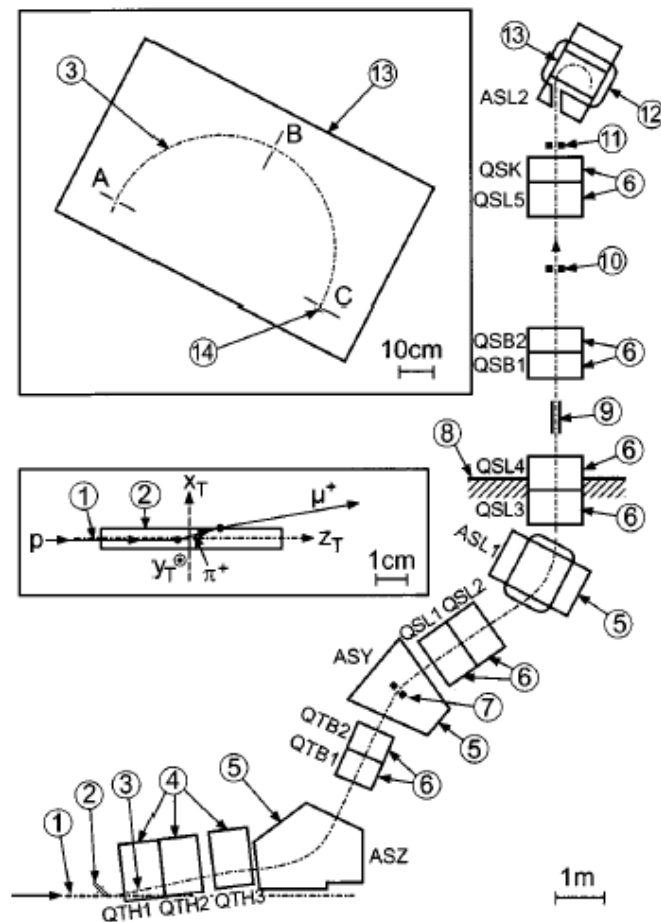


FIG. 1. Experimental setup. (1) Central trajectory of 590 MeV proton beam; (2) graphite target; (3) central trajectory of muon beam; (4) half-quadrupole magnets; (5) dipole magnets; (6) quadrupole magnets; (7) collimator defining the beam momentum acceptance; (8) concrete shielding of proton channel; (9) crossed-field particle separator; (10) lead collimator; (11) remotely movable collimator system (normally open); (12) magnetic spectrometer; (13) pole of spectrometer; (14) muon detectors (silicon microstrip and single silicon barrier detectors); *A, B, C*: copper collimators.

## A precise spectrometer

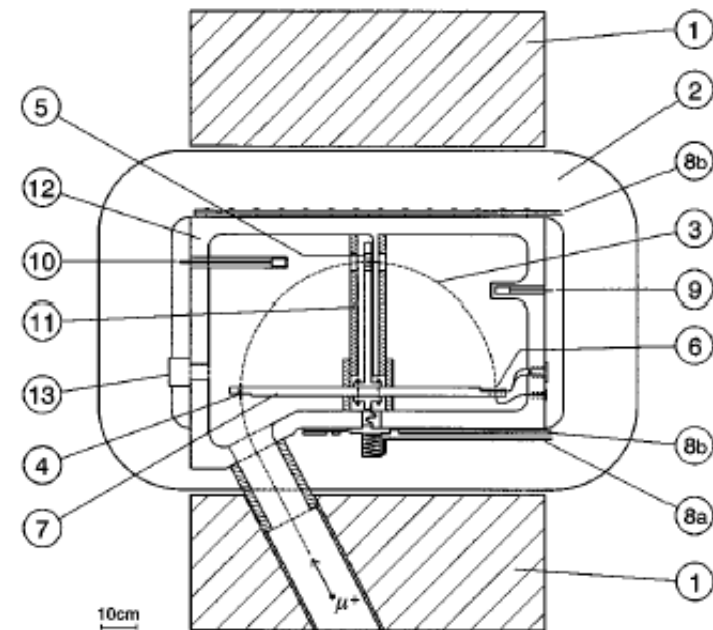
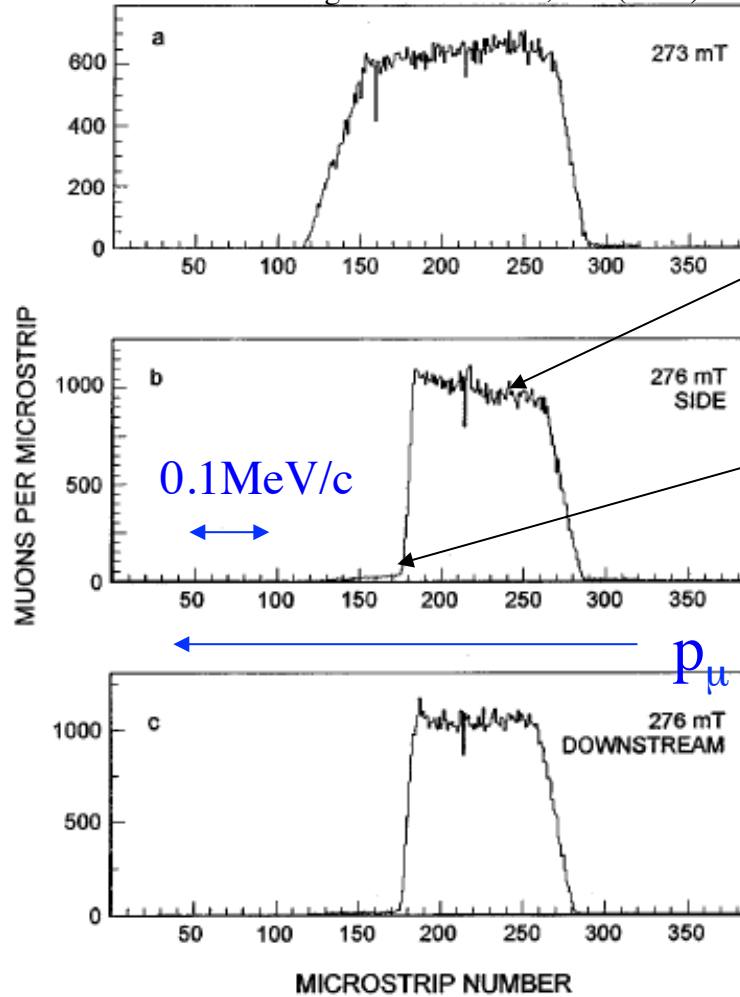


FIG. 3. The muon spectrometer: (1) magnet yoke; (2) magnet coils; (3) central muon trajectory; (4)–(6) copper collimators *A, B, C*; (7) titanium support; (8a) and (8b) cooling water pipes; (9) and (10) NMR probes; (11) lead shielding; (12) vacuum chamber; (13) port for vacuum pump.

@PSI





energy loss in the target

$$p_\mu^{\max} = (29.79200 \pm 0.00011) \text{ MeV} / c$$

$$m_{\nu_\mu} < 0.19 \text{ MeV} (90\%CL)$$

(PDG average)

FIG. 6. Distribution of muons in the microstrip detector for three typical runs. (a) Central muon-beam momentum 29.45 MeV/c, spectrometer field 273.0 mT; (b) and (c) central muon-beam momentum 29.75 MeV/c, spectrometer field 276.0 mT. One microstrip width (0.05 mm) corresponds to  $\Delta p_\mu \approx 0.0021 \text{ MeV}/c$ . The muon momentum increases to the left. For details, see Sec.

# $\nu_\mu$ mass limit

$$m_\nu = 0.34 m_\pi \sqrt{1 - \frac{p_\mu}{p_\mu^0}}$$

$$\delta m_{\nu_\mu} \sim \sqrt{m_\pi^2 - m_\mu^2} \left( \sqrt{\frac{\delta m_\pi}{m_\pi}} + \sqrt{\frac{1 - (m_\mu/m_\pi)^2}{1 + (m_\mu/m_\pi)^2}} \sqrt{\frac{\delta p_\mu}{p_\mu}} \right)$$

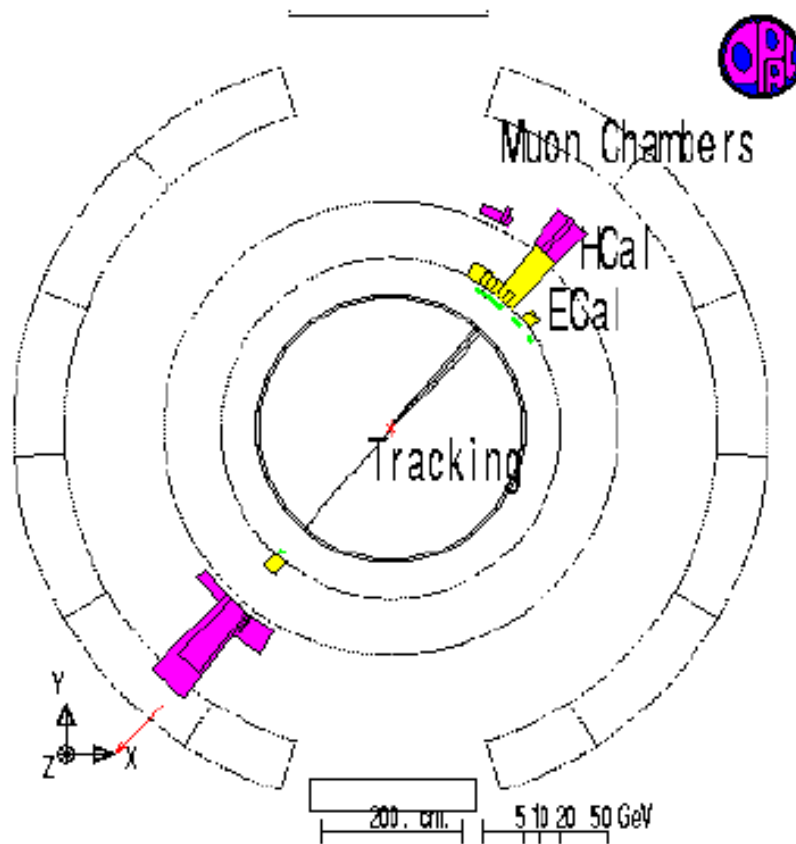
$$\underbrace{\sim 0.15 \text{ MeV}}_{\boxed{\delta m_\pi}} \oplus \underbrace{49 \text{ MeV} \sqrt{\frac{\delta p_\mu}{p_\mu}}}_{\boxed{\sim 0.1 \text{ MeV}}}$$

$\delta m_\pi$  limits the precision

# $\tau$ neutrino mass limit

$$e^+e^- \rightarrow \tau^+\tau^-$$

$$\tau \rightarrow \nu_\tau + X \quad (LEP, CLEO)$$

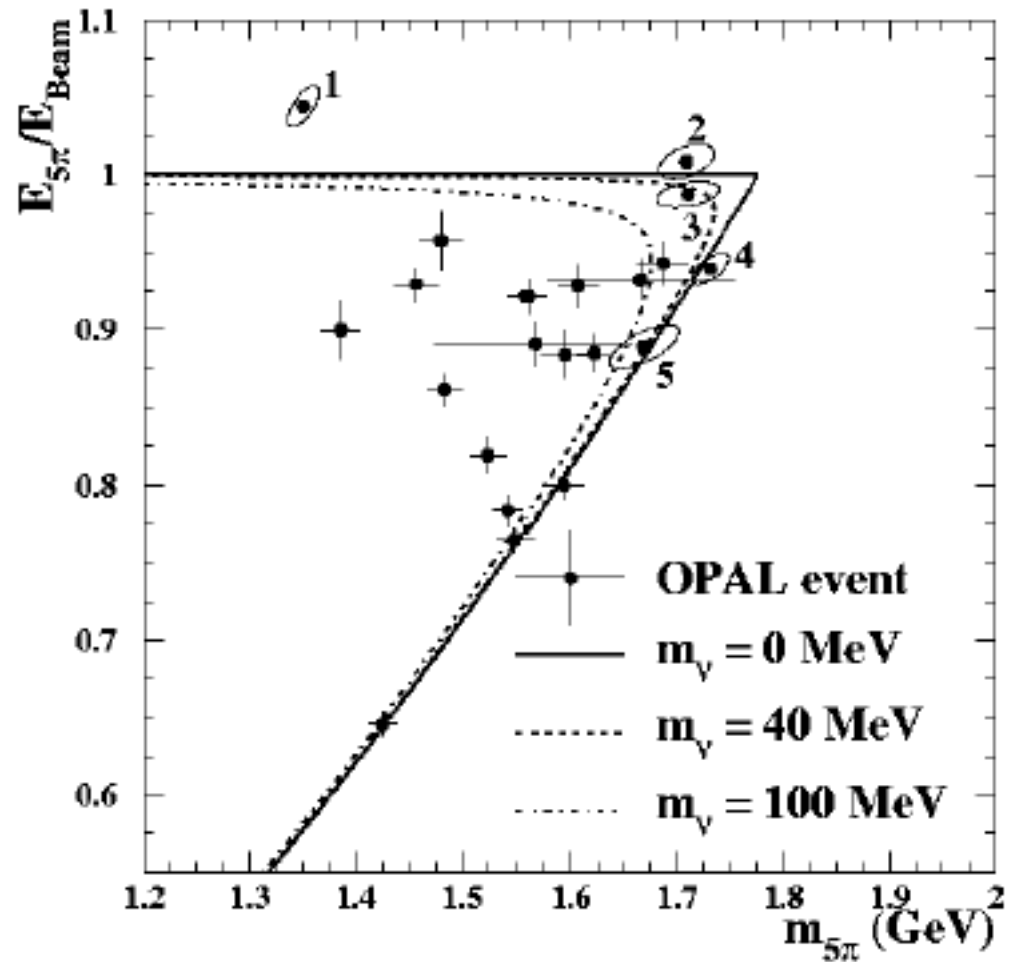


$$\tau \rightarrow \nu_\tau + X \quad (LEP, CLEO)$$

distribution of  $m_X$  &  $E_X$ ,

$$\frac{d\Gamma}{dm_X dE_X} = f(m_X, E_X; m_\nu)$$

→ obtain most likely  $m_\nu$



Better precision for smaller Q-value, but low statistics,

$\tau \rightarrow 5\pi + \nu_\tau$  are used.

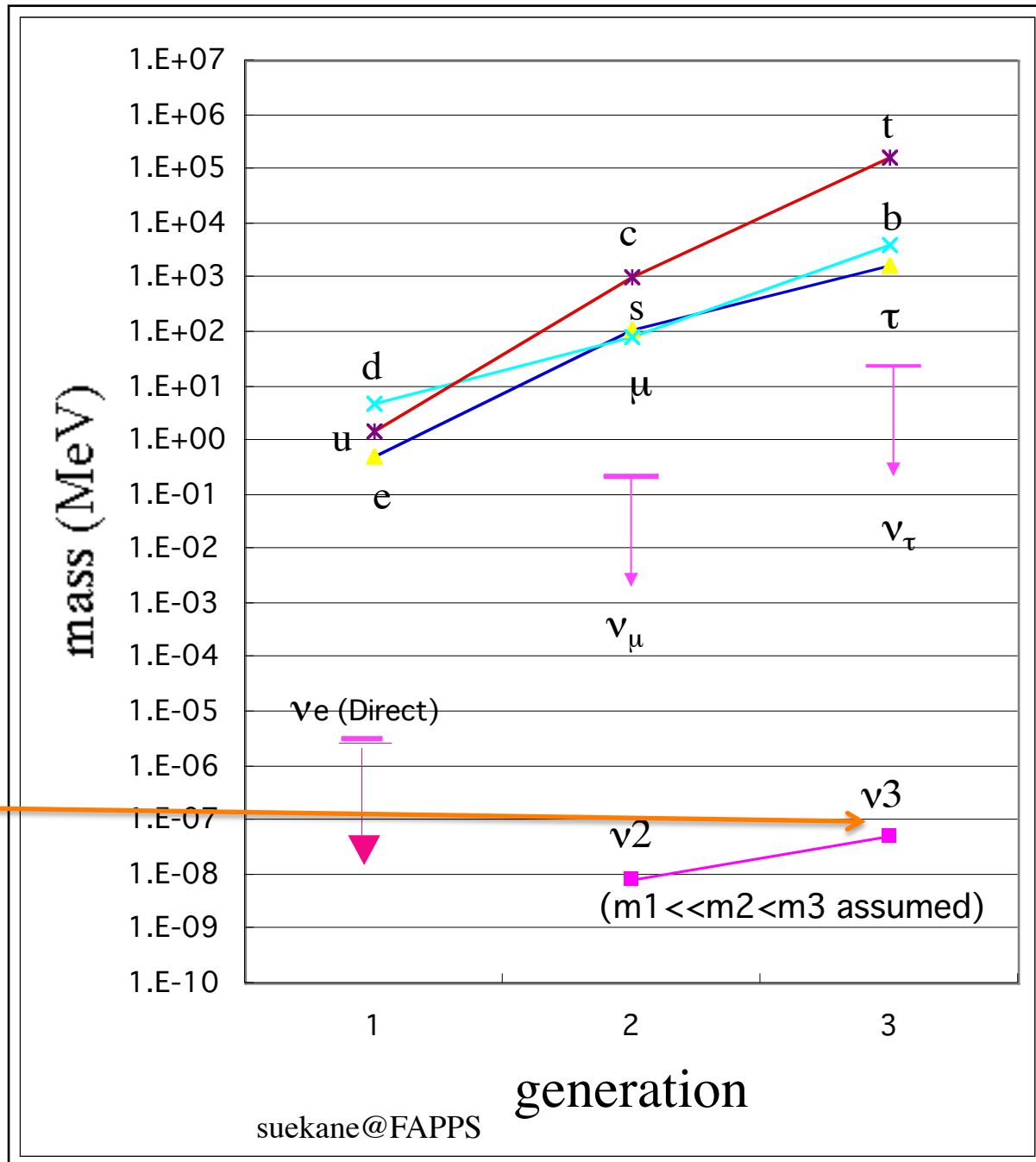
PDG average ;  $m_{\nu_\tau} < 18.2 \text{ MeV} (95\%CL)$

## Quark-Lepton masses

$\nu$  mass is very small

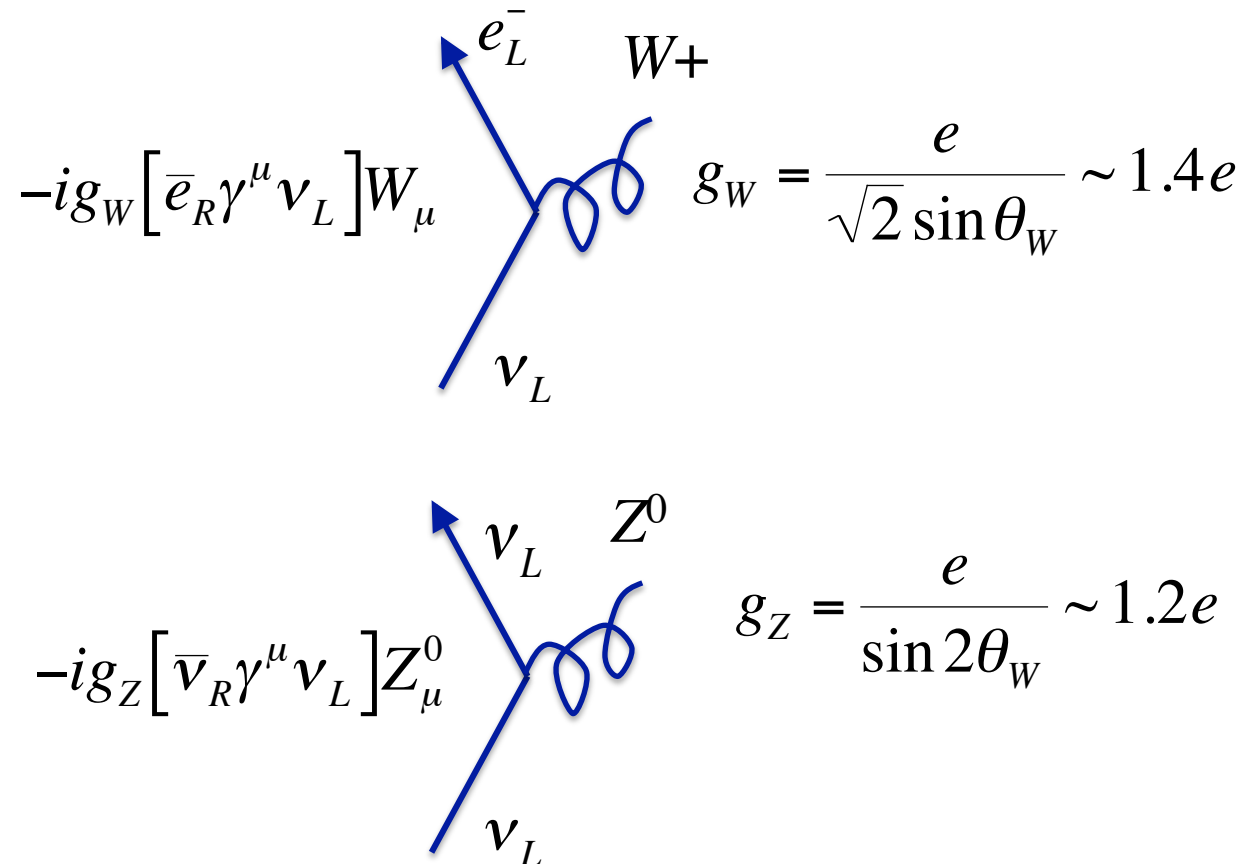
Lower limit of heaviest neutrino mass is  $\sim 50\text{meV}$

←  $\nu$  oscillation



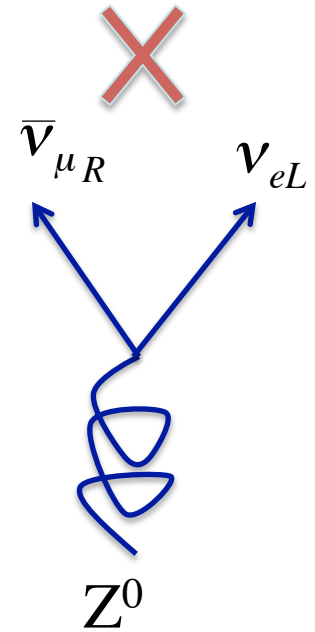
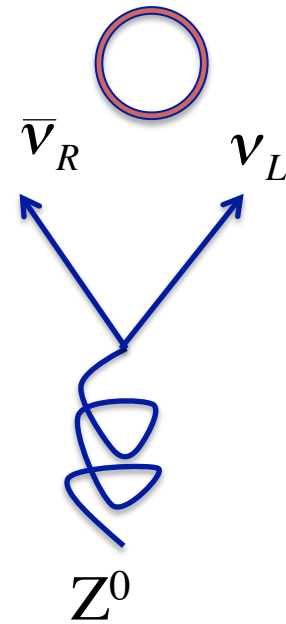
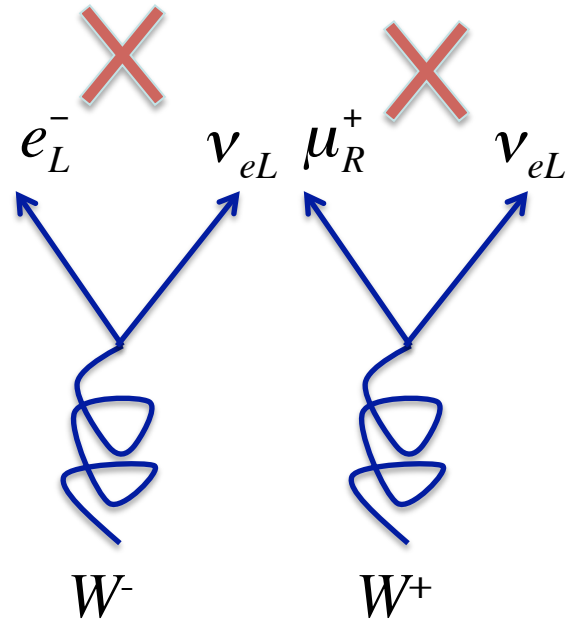
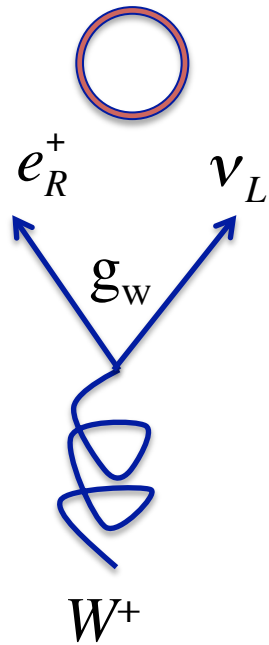
# Neutrinos in the Standard Model

- \*  $Q=0$ ,
- \* No color
- \*  $m=0$ ,
- \*  $s=1/2$
- \*  $\bar{\nu} \neq \nu$
- \* only  $\nu_L$  exists  
(or  $\nu_L$  may exist  
but it does not  
interact at all)



$\sin^2 \theta_W \sim 0.23$  (Weinberg angle)

suekane@FAPPS



Lepton  
number  
violation

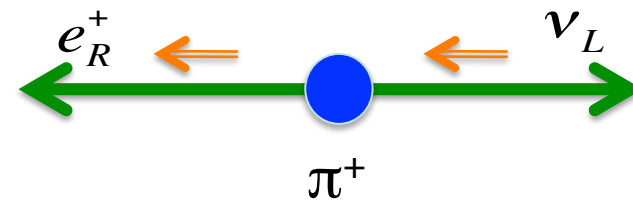
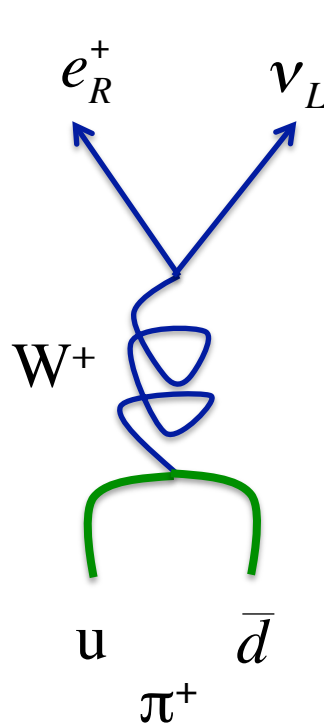
Flavor  
violation

Flavor  
violation

# 'Helicity' Suppression of $\pi$ -decay

Experimental fact: 
$$\frac{\Gamma_{\pi \rightarrow e \nu}}{\Gamma_{\pi \rightarrow \mu \nu}} = 1.2 \times 10^{-4}$$

How it is explained?



You may say  $W$  couple only LH  $\nu$  and RH  $e$ .  
So that  $J(e\nu)=1$ , while pion spin=0  
 $\Rightarrow$  violates spin conservation

However, this decay exists if very small.  
And for  $\pi \rightarrow \mu + \nu$  decay, the spin  
conservation seems to strongly violated.



# Helicity and Chirality

Sometimes Helicity and Chirality are used in confuse.  
Here they are defined and their relations are discussed.

Dirac equation in free space is,

$$(i\gamma^\mu \partial_\mu - m)\psi(x) = 0$$

General solution is,

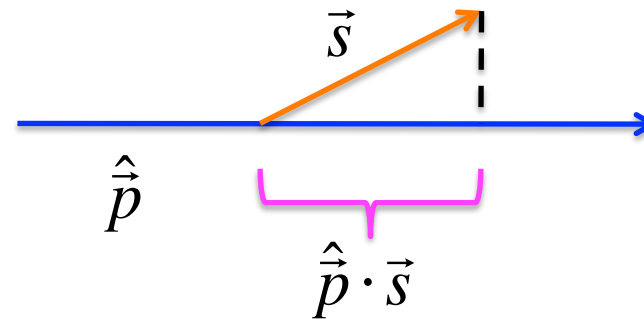
$$\psi(x) = \begin{pmatrix} u \\ \vec{\eta} \cdot \vec{\sigma} u \end{pmatrix} e^{i(\vec{p}\vec{x} - Et)} + \begin{pmatrix} -\vec{\eta} \cdot \vec{\sigma} v \\ v \end{pmatrix} e^{i(\vec{p}\vec{x} + Et)}$$

$$\left[ \vec{\eta} \equiv \frac{\vec{p}}{E + m}, \quad E \equiv +\sqrt{\vec{p}^2 + m^2}, \quad u = \begin{pmatrix} u_1 \\ u_2 \end{pmatrix}, \quad v = \begin{pmatrix} v_1 \\ v_2 \end{pmatrix} \right]$$

Now we take initial condition as positive energy and  $\vec{p} = (0, 0, p)$

$$\psi(0) = \begin{pmatrix} u \\ \eta \sigma_z u \end{pmatrix}$$

**Helicity** is the spin component to the direction of the movement.



If the movement is along the z-direction, **helicity** components are,

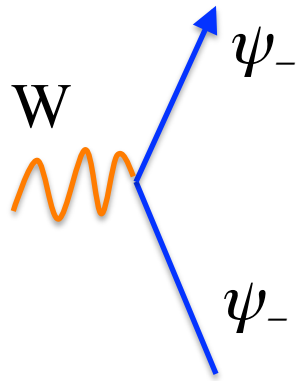
$$\begin{cases} \psi_R \equiv \frac{1}{2}(1 + \hat{p} \cdot \vec{\sigma})\psi = \frac{1}{2}(1 + \sigma_z)\begin{pmatrix} u \\ \eta u \end{pmatrix} = u_1 \begin{pmatrix} \chi_1 \\ \eta \chi_1 \end{pmatrix} \\ \psi_L \equiv \frac{1}{2}(1 - \hat{p} \cdot \vec{\sigma})\psi = \frac{1}{2}(1 - \sigma_z)\begin{pmatrix} u \\ -\eta u \end{pmatrix} = u_2 \begin{pmatrix} \chi_2 \\ -\eta \chi_2 \end{pmatrix} \end{cases}$$

These **helicity** states show actual spin direction.

Here after we call Right(Left) handed Helicity = **RH (LH)**

# What W couples to: Chirality

W couples to **negative Chirality** (NC) particle and **positive Chirality** (PC) state anti-particle.



Chirality components of  $\psi$  is defined by,

$$\psi_{\pm} \equiv \frac{1}{2}(1 \pm \gamma_5)\psi = \frac{1}{2} \begin{pmatrix} 1 & \pm 1 \\ \pm 1 & 1 \end{pmatrix} \begin{pmatrix} u \\ \eta \sigma_z u \end{pmatrix} = \frac{1 \pm \eta \sigma_z}{2} \begin{pmatrix} u \\ \pm u \end{pmatrix}$$

For  $m \rightarrow 0$ ,  $\eta \rightarrow 1$  and

$$\psi_{\pm} = \frac{1 \pm \eta \sigma_z}{2} \begin{pmatrix} u \\ \pm u \end{pmatrix} \xrightarrow{m=0} \frac{1 \pm \sigma_z}{2} \begin{pmatrix} u \\ \pm u \end{pmatrix} = \psi_{R/L}$$

For high energy, the helicity and chirality are same and sometimes they are confused.

**For low energy, NC has RH component.**

In  $\pi^- \rightarrow e^- \nu$  decay,  $e^-$  is **NC** state and  $\nu$  is **PC** state.

Then the **RH** component of electron in the  $\pi$  decay is,

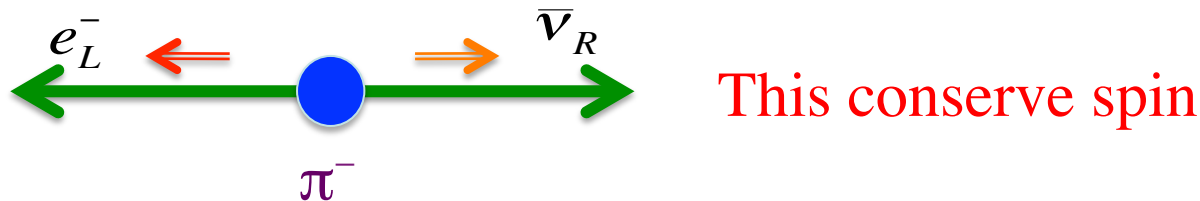
$$\psi_{R-} = \frac{1}{2} (1 + \hat{p} \cdot \vec{\sigma}) \psi_- = \frac{(1-\eta)}{4} (1 + \sigma_z) \begin{pmatrix} u \\ -u \end{pmatrix}$$

So that the probability which is **RH** is,

$$|\psi_{R-}|^2 = \frac{(1-\eta)^2}{8} [u^\dagger (1 + \sigma_z) u] = \frac{\gamma}{\gamma+1} \frac{(1-\beta)}{2} |u_1|^2 \xrightarrow{m/E \ll 1} \left( \frac{m^2}{4E^2} \right) |u_1|^2$$

This means the electron has right handed component with probability

$$P \propto (m_e^2 / 4E_e^2) \sim 1.3 \times 10^{-5}$$



For muon case,  $m_\mu/E_\mu \sim 1$  and the suppression is not strong.

Taking into account the phase space the theoretical prediction is

$$\frac{\Gamma_{\pi \rightarrow e \nu}}{\Gamma_{\pi \rightarrow \mu \nu}} = \left( \frac{m_e}{m_\mu} \right)^2 \left( \frac{m_\pi^2 - m_e^2}{m_\pi^2 - m_\mu^2} \right) = 1.28 \times 10^{-4}$$

while observation is

$\pi^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence
$\mu^+ \nu_\mu$	[b] (99.98770 ± 0.00004) %	
$\mu^+ \nu_\mu \gamma$	[c] ( 2.00 ± 0.25 ) × 10 <sup>-4</sup>	
$e^+ \nu_e$	[b] ( 1.230 ± 0.004 ) × 10 <sup>-4</sup>	

$$\rightarrow 1.23 \times 10^{-4}$$

Likewise for K decay,

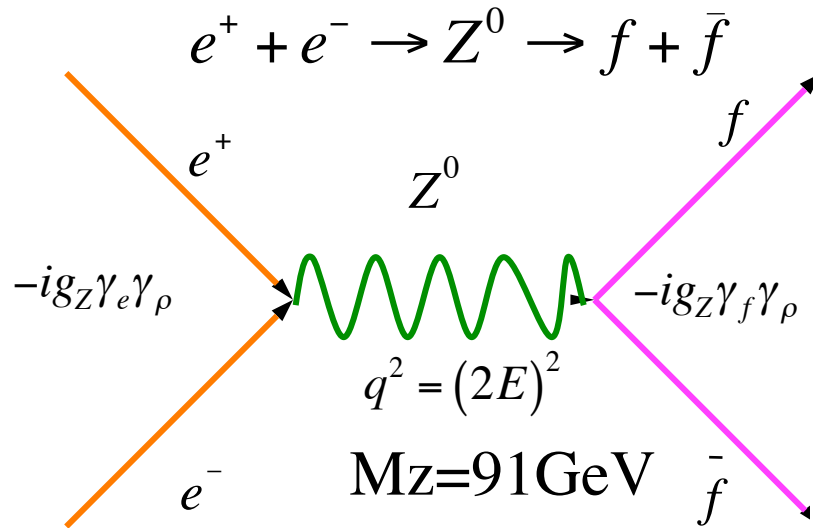
$$\frac{\Gamma_{K \rightarrow e \nu}}{\Gamma_{K \rightarrow \mu \nu}} = \left( \frac{m_e}{m_\mu} \right)^2 \left( \frac{m_K^2 - m_e^2}{m_K^2 - m_\mu^2} \right) = 2.4 \times 10^{-5}$$

while observation is

$K^+$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confidence
<b>Leptonic and semileptonic modes</b>		
$K^+ \rightarrow e^+ \nu_e$	( 1.584 ± 0.020 ) × 10 <sup>-5</sup>	
$K^+ \rightarrow \mu^+ \nu_\mu$	(63.55 ± 0.11 ) %	

$$\rightarrow 2.5 \times 10^{-5}$$

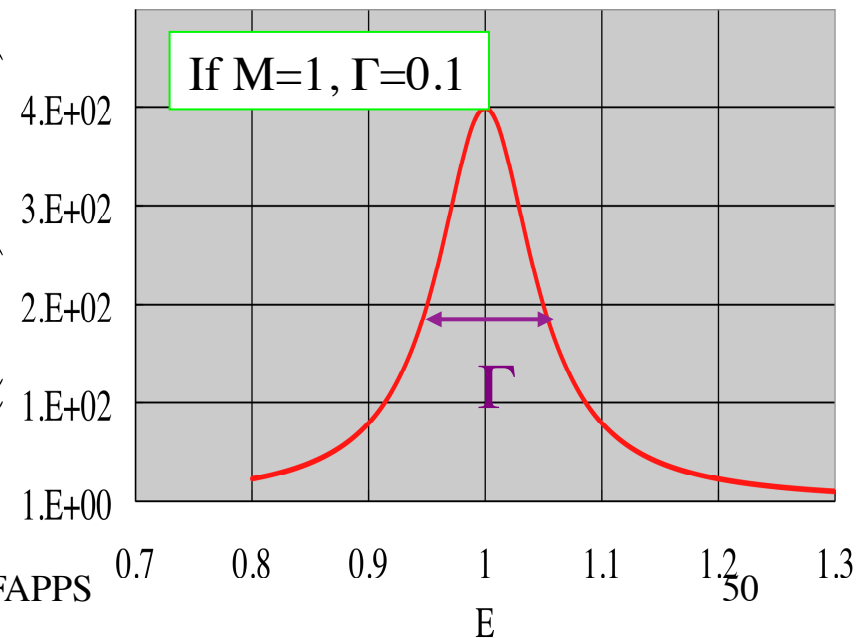
# neutrino flavor counting using $Z^0$



$$\sigma \propto \frac{g_e^2 g_f^2}{4M_Z^2 \left( \left( \sqrt{q^2} - M_Z \right)^2 + \Gamma_Z^2/4 \right)}$$

$$1/((E-M)^2 - \Gamma^2/4)$$

Resonance



The width of the  $Z^0$ :  $\Gamma$  is inverse of the lifetime of  $Z^0$ :  $\tau$

The lifetime is proportional to inverse sum of decay width.

$$\frac{1}{\Gamma_Z} = \tau = \frac{1}{\Gamma_{Z \rightarrow u\bar{u}} + \Gamma_{Z \rightarrow d\bar{d}} + \cdots + \Gamma_{Z \rightarrow \nu_e \bar{\nu}_e}}$$

If the number of neutrino flavors is  $n_\nu$ ,

$$\Gamma_Z = 6\Gamma_U + 9\Gamma_D + 3\Gamma_L + n_\nu \Gamma_\nu$$

$$\Gamma_\nu \propto 1, \quad \Gamma_L \propto 1 - 4x_w + 8x_w^2, \quad \Gamma_U \propto 1 - \frac{8}{3}x_w + \frac{32}{9}x_w^2, \quad \Gamma_D \propto 1 - \frac{4}{3}x_w + \frac{8}{9}x_w^2$$

$$x_w = \sin \theta_w \sim 0.23$$

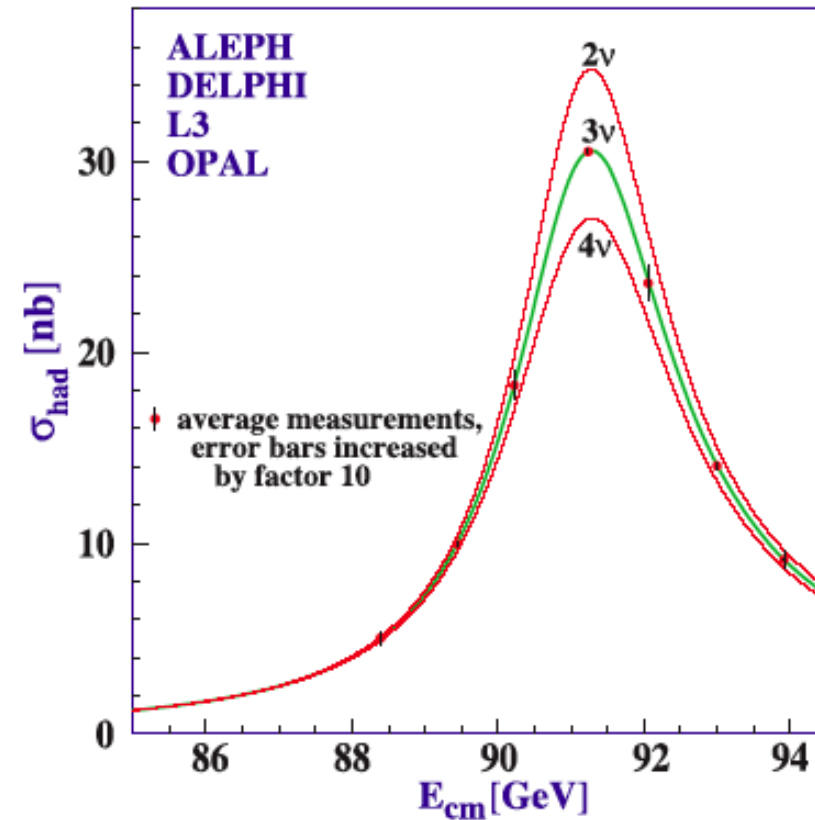
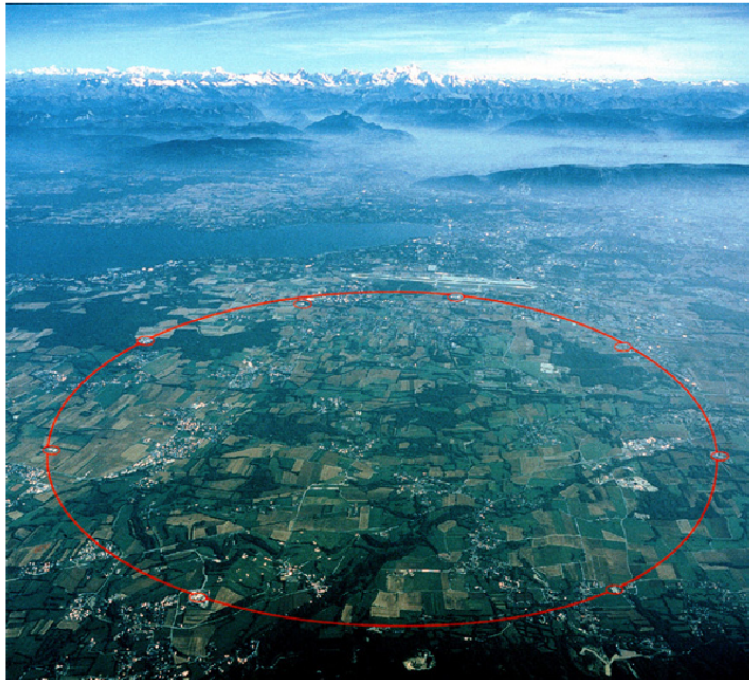
measurement

$$n_\nu = \frac{\Gamma_Z - 6\Gamma_U - 9\Gamma_D - 3\Gamma_L}{\Gamma_\nu} \leftarrow \text{known}$$

# $Z^0$ data

CERN's aerial view in 1994. (Copyright CERN)

Vue aeriennne du CERN avec le trac du tunnel LEP et du futur LHC (1994). (Copyright CERN)



4 experiments at LEP (ALEPH, DELPHI, OPAL, L3) showed

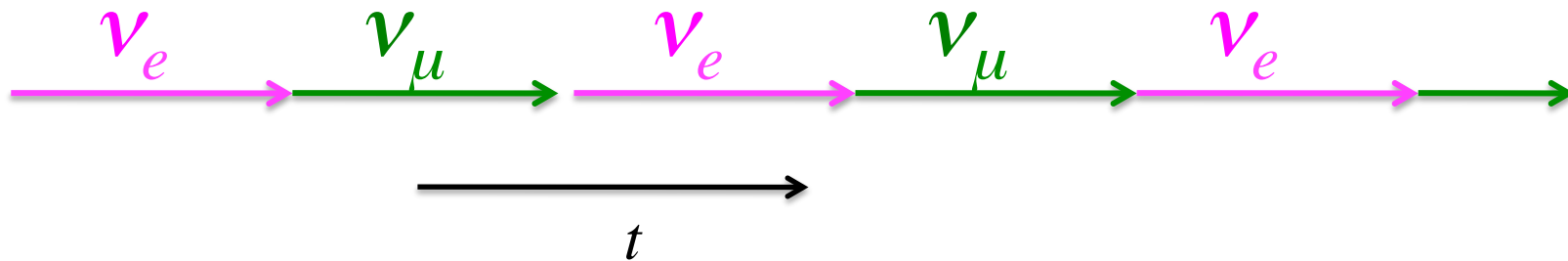
$$n_v = 2.984 \pm 0.008$$

If 4<sup>th</sup> neutrino exists  $m_4 > 45\text{GeV}$  or  
it does not couple to  $Z^0$ , called **sterile neutrino**.



# $\nu$ Oscillation: An Introduction

Neutrino oscillation is phenomena in which flavor of neutrino oscillatory changes as time passed by.



If we start from  $\nu_e$ , the probability to find  $\nu_\mu$  at time  $t$  is expressed as:

$$P_{\nu_e \rightarrow \nu_\mu}(t) = \sin^2 2\theta \sin^2 \frac{m_2^2 - m_1^2}{4E} t$$

Where  $m_2$  and  $m_1$  are masses of energy-eigenstate neutrinos,  $\theta$  is called mixing angle.

$\nu$  oscillation is the 1<sup>st</sup> firm evidence beyond the standard model and its studies are important to understand the nature.

# Phenomena of spin-1/2 for $\nu$ Oscillation

The formalism of the  $\nu$  oscillation is very similar to that of spin-1/2 under magnetic field.  
So let's review the spin motion as introduction.

The spin motion under magnetic field is described by the Pauli equation:

$$i\dot{\psi} = \mu \vec{B} \cdot \vec{\sigma} \psi$$

Where  $\vec{B} = (B_x, B_y, B_z)$  is the magnetic field and  $\mu$  is magnetic dipole moment of the particle.

This equation was 1stly introduced empirically by Pauli, and later on obtained by taking non relativistic limit of the Dirac equation with electro-magnetic interaction.

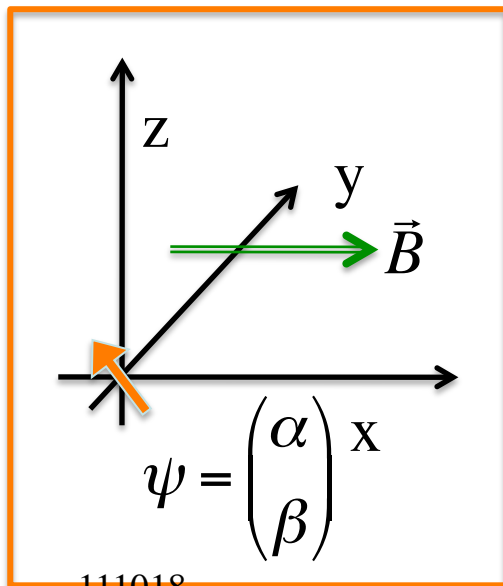
# Phenomena of spin-1/2 for $\nu$ Oscillation

$$i\dot{\psi} = \mu \vec{B} \vec{\sigma} \psi$$

The wave function is a mixture of spin up and down states

$$\psi(t) = \alpha(t)|\uparrow\rangle + \beta(t)|\downarrow\rangle \equiv \begin{pmatrix} \alpha(t) \\ \beta(t) \end{pmatrix}$$

Here, we think the case that magnetic field is along the  $x$  axis.



$$\vec{B} = (B, 0, 0)$$

Then the Pauli equation becomes

$$\begin{pmatrix} \dot{\alpha} \\ \dot{\beta} \end{pmatrix} = -i\mu B \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix}$$

## Phenomena of spin-1/2 for $\nu$ Oscillation

$$\begin{cases} \dot{\alpha} = -i\mu B\beta \\ \dot{\beta} = -i\mu B\alpha \end{cases}$$

by taking the derivative of the 1<sup>st</sup> equation,  
and replacing  $\dot{\beta}$  by the 2<sup>nd</sup> equation,

$$\ddot{\alpha} = -(\mu B)^2 \alpha$$

This is the harmonic oscillator and we know the general solution;

$$\begin{cases} \alpha(t) = pe^{-i\mu Bt} + qe^{i\mu Bt} \\ \beta(t) = pe^{-i\mu Bt} - qe^{i\mu Bt} \end{cases}$$

Where  $p$  &  $q$  are integral constants to be determined by initial condition.

# Phenomena of spin-1/2 for $\nu$ Oscillation

Then the general spin state is,

$$\psi(t) = (pe^{-i\mu Bt} + qe^{i\mu Bt})|\uparrow\rangle + (pe^{-i\mu Bt} - qe^{i\mu Bt})|\downarrow\rangle$$

Now we assume that at  $t=0$ , the spin pointed upward.

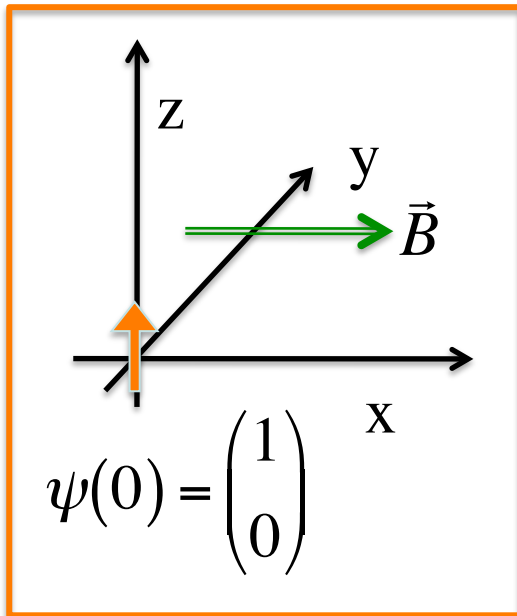
$$\psi(0) = |\uparrow\rangle$$

Then we can determine  $p$  and  $q$  for this case,

$$\psi(0) = (p+q)|\uparrow\rangle + (p-q)|\downarrow\rangle \equiv |\uparrow\rangle$$



$$p = q = \frac{1}{2}$$



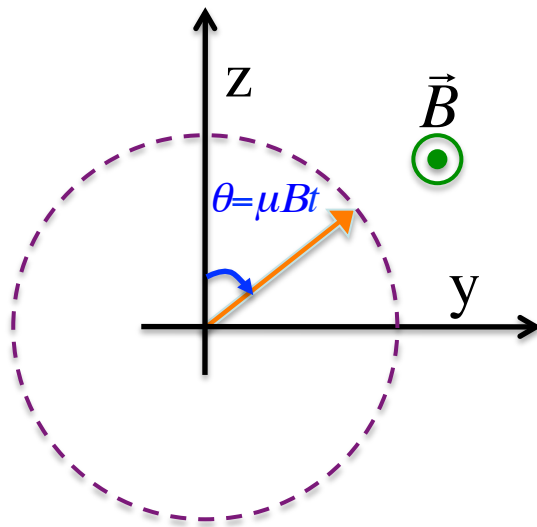
# Phenomena of spin-1/2 for $\nu$ Oscillation

Then we get specific wave function;

$$\psi(t) = \cos(\mu B t) |\uparrow\rangle - i \sin(\mu B t) |\downarrow\rangle$$

This means at later time,  $|\downarrow\rangle$  state is generated with **oscillating probability**

$$P_{\uparrow \rightarrow \downarrow}(t) = \sin^2(\mu B t)$$



$$\psi(\theta = \mu B t, \varphi = 0)$$

If we recall that the wave function of the spin, which is in z-y plane and polar angle is  $\theta$  is

$$\psi(\theta) = \cos \theta |\uparrow\rangle - i \sin \theta |\downarrow\rangle$$

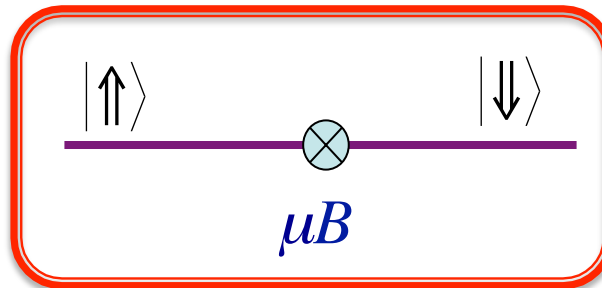
Physically it corresponds to the **precession** of the spin, caused by the torque by  $B$  and  $\mu$ .

## Phenomena of spin-1/2 for $\nu$ Oscillation

$$\begin{pmatrix} \dot{\alpha} \\ \dot{\beta} \end{pmatrix} = -i\mu B \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix}$$

Quantum Mechanically, this is understood as the effect that magnetic field causes transition between  $|\uparrow\rangle \leftrightarrow |\downarrow\rangle$  with amplitude  $\mu B$ .

We will draw this kind of effect schematically as follows.



The neutrino oscillation can be understood in exactly the same manner.

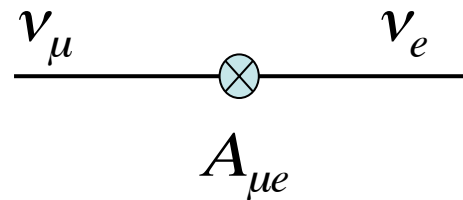
# A simple case of $\nu$ Oscillation

We assume 2 neutrino system;  $\nu_e$  &  $\nu_\mu$

The general state is;

$$\psi_\nu(t) = \alpha(t)|\nu_e\rangle + \beta(t)|\nu_\mu\rangle$$

We assume something makes transition:  $\nu_e \Leftrightarrow \nu_\mu$



Then there are correspondences to the spin case

$$\left\{ \begin{array}{l} |\nu_e\rangle \Leftrightarrow |\uparrow\uparrow\rangle \\ |\nu_\mu\rangle \Leftrightarrow |\downarrow\downarrow\rangle \\ A_{\mu e} \Leftrightarrow \mu B \end{array} \right.$$

If the initial state is pure  $\nu_e$  state, like beta decay, then

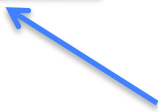
$$P_{\nu_e \rightarrow \nu_\mu}(t) = \sin^2 A_{\mu e} t$$

This is the very basic of neutrino oscillation.



## A simple case of $\nu$ Oscillation

However, we often see the neutrino oscillation probability as

$$P_{\nu_\mu}(t) = \sin^2 2\phi \sin^2 \frac{m_2^2 - m_1^2}{4E} L$$


Where does this come from?

What is the analogy of spin motion?

## Still spin-1/2 for $\nu$ Oscillation

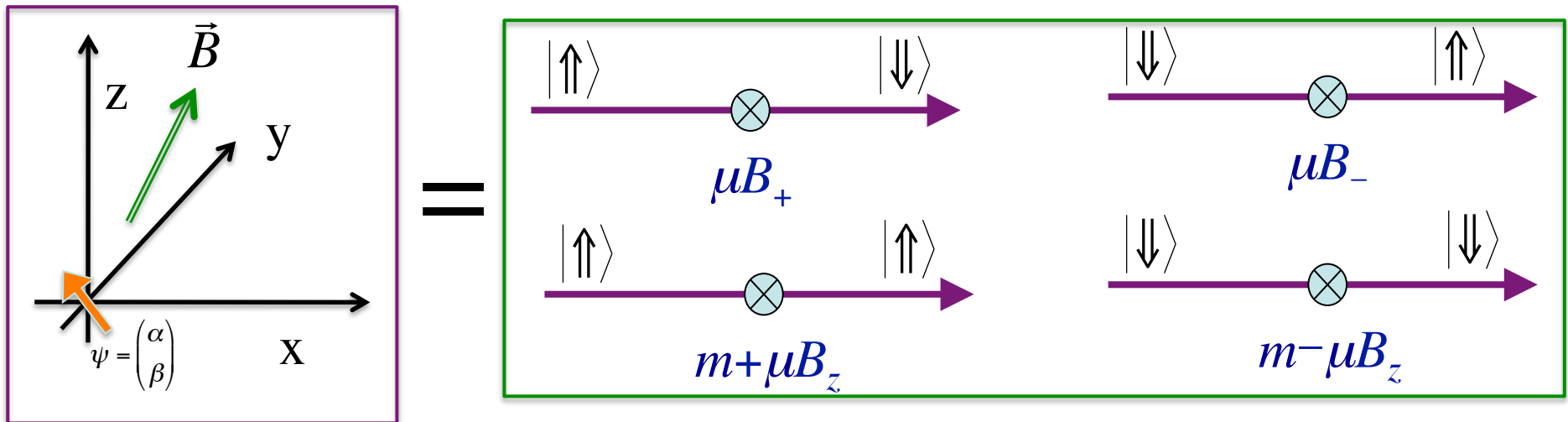
In actual case, mass term has to be included in the Pauli equation

$$i\dot{\psi} = (m + \mu \vec{B} \vec{\sigma}) \psi$$

and the most general equation with arbitrary magnetic field is,

$$\begin{pmatrix} \dot{\alpha} \\ \dot{\beta} \end{pmatrix} = -i \begin{pmatrix} m + \mu B_z & \mu B_- \\ \mu B_+ & m - \mu B_z \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix}; \quad B_{\pm} \equiv B_x \pm iB_y$$

Spin transition amplitudes are



## Still spin-1/2 for $\nu$ Oscillation

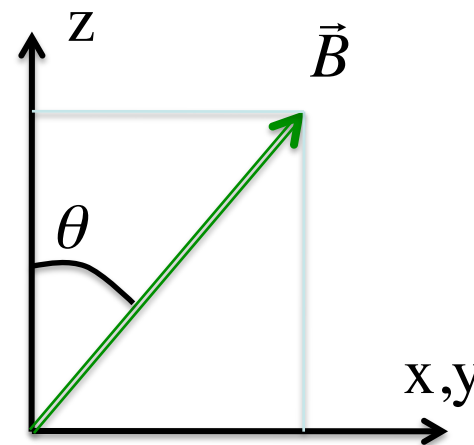
$$\begin{pmatrix} \dot{\alpha} \\ \dot{\beta} \end{pmatrix} = -i \begin{pmatrix} m + \mu B_z & \mu B_- \\ \mu B_+ & m - \mu B_z \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix}$$

This equation has the general form

$$\begin{pmatrix} \dot{\alpha} \\ \dot{\beta} \end{pmatrix} = -i \begin{pmatrix} P & Q^* \\ Q & R \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix}, \quad P, R \in \text{Real}$$

Relation between Polar angle  $\theta$  of the magnetic field and transition amplitudes is,

$$\tan \theta \equiv \frac{B_{\perp}}{B_z} = \frac{2|Q|}{R - P}$$



## Still spin-1/2 for $\nu$ Oscillation

$$\begin{pmatrix} \dot{\alpha} \\ \dot{\beta} \end{pmatrix} = -i \begin{pmatrix} P & Q^* \\ Q & R \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix}, \quad P, R \in \text{Real}$$

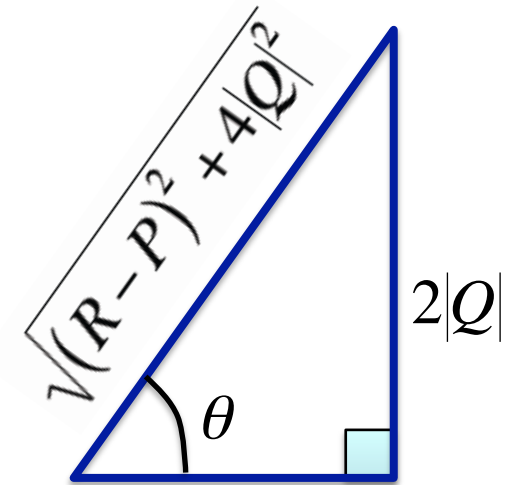
Then the general solution can be expressed using  $\theta$  as,

$$\begin{cases} \alpha(t) = C_1 \cos(\theta/2) e^{-iE_+ t} - C_2 \cos(\theta/2) e^{-iE_- t} \\ \beta(t) = C_1 \sin(\theta/2) e^{-iE_+ t} + C_2 \cos(\theta/2) e^{-iE_- t} \end{cases}$$

where,

$$\begin{cases} E_+ = \frac{1}{2} \left( (P + R) + \sqrt{(P - R)^2 + 4|Q|^2} \right) \\ E_- = \frac{1}{2} \left( (P + R) - \sqrt{(P - R)^2 + 4|Q|^2} \right) \end{cases}$$

(Note: by definition,  $E_+ > E_-$ )



$$\tan \theta \equiv \frac{2|Q|}{R - P}$$

## Still spin-1/2 for $\nu$ Oscillation

The general wave function is,

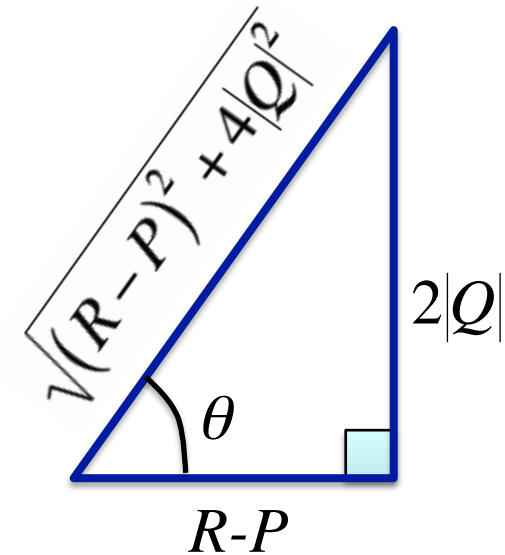
$$\psi(t) = \left( C_1 \cos(\theta/2) e^{-iE_+ t} - C_2 \sin(\theta/2) e^{-iE_- t} \right) |\uparrow\rangle + \left( C_1 \sin(\theta/2) e^{-iE_+ t} + C_2 \cos(\theta/2) e^{-iE_- t} \right) |\downarrow\rangle$$

Again we start with

$$\psi(0) = |\uparrow\rangle$$

Then, the integral constants are determined.

$$\begin{cases} C_1 \cos(\theta/2) - C_2 \sin(\theta/2) = 1 \\ C_1 \sin(\theta/2) + C_2 \cos(\theta/2) = 0 \end{cases} \Rightarrow \begin{cases} C_1 = \cos(\theta/2) \\ C_2 = -\sin(\theta/2) \end{cases}$$



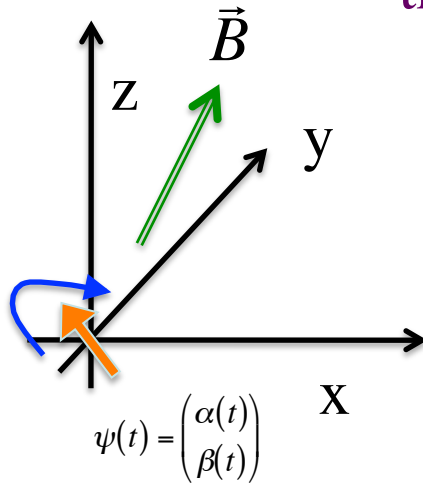
$$\tan \theta \equiv \frac{2|Q|}{R-P}$$

## Still spin-1/2 for $\nu$ Oscillation

In this case, the specific wave function is

$$\psi(t) = \left( \cos^2(\theta/2) e^{-iE_+ t} + \sin^2(\theta/2) e^{-iE_- t} \right) |\uparrow\uparrow\rangle + \frac{1}{2} \sin \theta (e^{-iE_+ t} - e^{-iE_- t}) |\downarrow\downarrow\rangle$$

This state corresponds to spin precession within the plane perpendicular to the magnetic field.



The time dependent probability of spin-down state is

$$P_{\uparrow \rightarrow \downarrow}(t) = \left| \frac{\sin \theta}{2} (e^{-iE_+ t} - e^{-iE_- t}) \right|^2 = \sin^2 \theta \sin^2 \mu B t$$

corresponds to the angle between the precession plane and z axis.

Difference of the energies in the energy eigenstate

## Still spin-1/2 for $\nu$ Oscillation

Look for Energy eigenstate,

Remember the  
general state

$$\psi_\nu(t) = \left( C_1 \cos(\theta/2) e^{-iE_+ t} - C_2 \sin(\theta/2) e^{-iE_- t} \right) |\uparrow\uparrow\rangle \\ + \left( C_1 \sin(\theta/2) e^{-iE_+ t} + C_2 \cos(\theta/2) e^{-iE_- t} \right) |\downarrow\downarrow\rangle$$

If we choose,  $C_1=1, C_2=0$ ,

$$\psi_+(t) = \left( \cos(\theta/2) |\uparrow\uparrow\rangle + \sin(\theta/2) |\downarrow\downarrow\rangle \right) e^{-iE_+ t}$$

This means

$$|+\rangle \equiv \cos(\theta/2) |\uparrow\uparrow\rangle + \sin(\theta/2) |\downarrow\downarrow\rangle$$

is energy eigenstate with energy  $E_+$ .

Similarly, if we choose,  $C_1=0, C_2=1$ ,

$$\psi_-(t) = \left( -\sin(\theta/2) |\uparrow\uparrow\rangle + \cos(\theta/2) |\downarrow\downarrow\rangle \right) e^{-iE_- t} \equiv |-\rangle e^{-iE_- t}$$



## Still spin-1/2 for $\nu$ Oscillation

The spin states  $|\uparrow\rangle$ ,  $|\downarrow\rangle$  itselfs are NOT energy eigenstate and do not have definite energy.

(If you try to measure the energy of  $|\uparrow\rangle$  state, you will see 2 energies.)

But the mixed state,

$$\begin{pmatrix} |+\rangle \\ |-\rangle \end{pmatrix} = \begin{pmatrix} \cos(\theta/2) & \sin(\theta/2) \\ -\sin(\theta/2) & \cos(\theta/2) \end{pmatrix} \begin{pmatrix} |\uparrow\rangle \\ |\downarrow\rangle \end{pmatrix}$$

ARE energy eigenstate and have definite energy.

$\theta/2$  is called mixing angle between energy eigenstate and spin state.

=The mixing angle corresponds to 1/2 of the polar angle of  $\vec{B}$ .