

Jean SANCHEZ

intern at LPNHE



IN2P3

Institut national de **physique nucléaire**
et de **physique des particules**

**Measurement of the spurious signal
7 TeV**

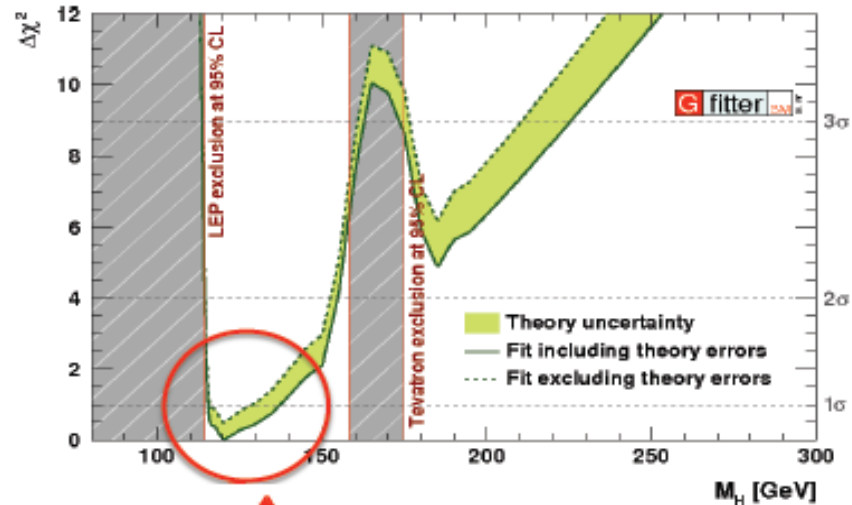
Outline

- **General presentation about Higgs research**
- **Background modeling**
- **Fit the background**
- **Influence of the fitting range**
- **Conclusion**

I – General presentation about Higgs research

Missing Higgs

- The problem
- The method
- Previous experiments :
Exclusion domains



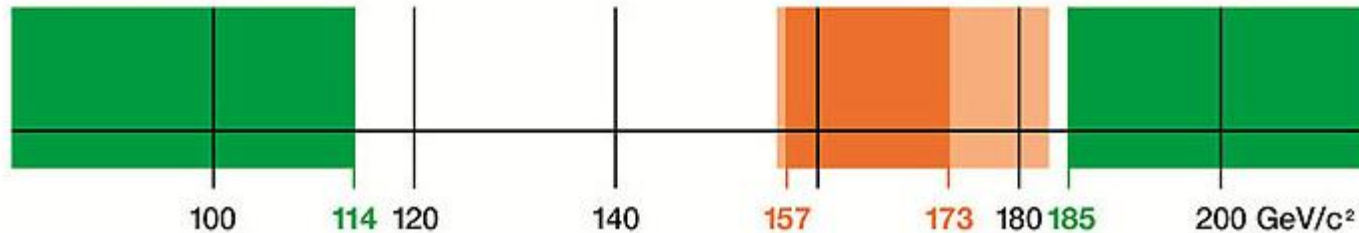
Status as of March 2011

95% confidence level

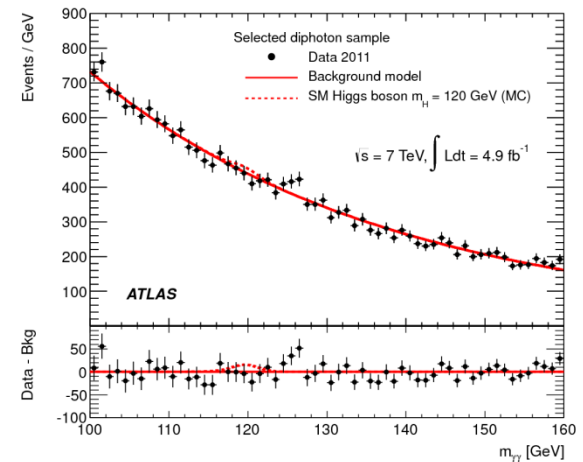
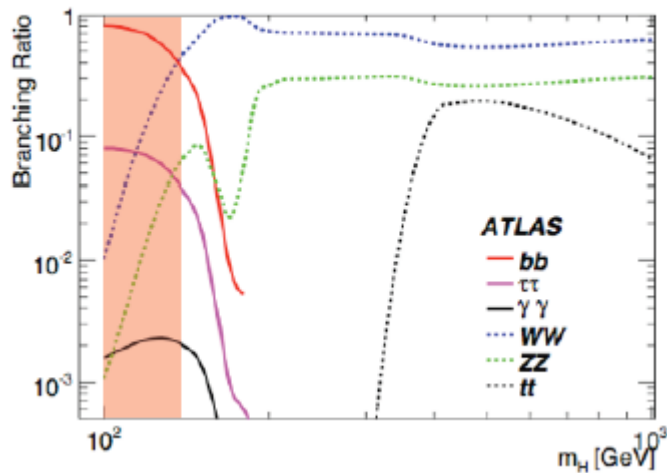
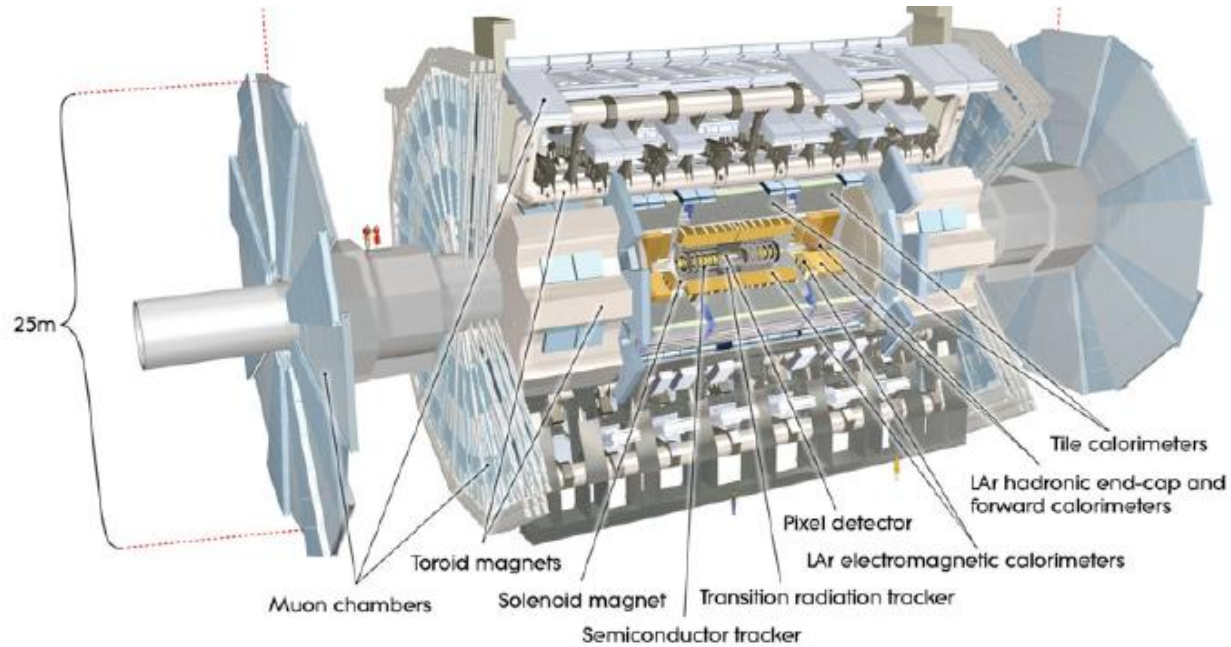
Excluded by
LEP Experiments
95% confidence level

Excluded by
Tevatron
Experiments

Excluded by
Indirect Measurements
95% confidence level



The ATLAS experiment



II – Background modeling

About fitting error

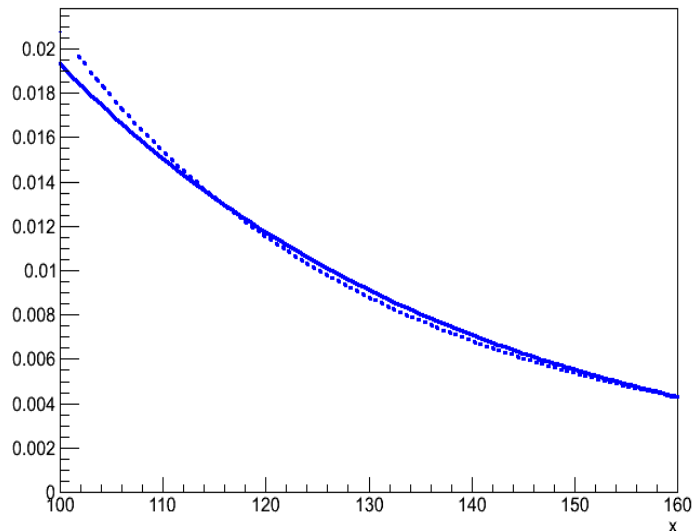
Background function is too complicated to be known

-> we do not know the exact shape of the fit function

Expected number of Higgs events : QCD -> aprx 200 events / 20000 background events

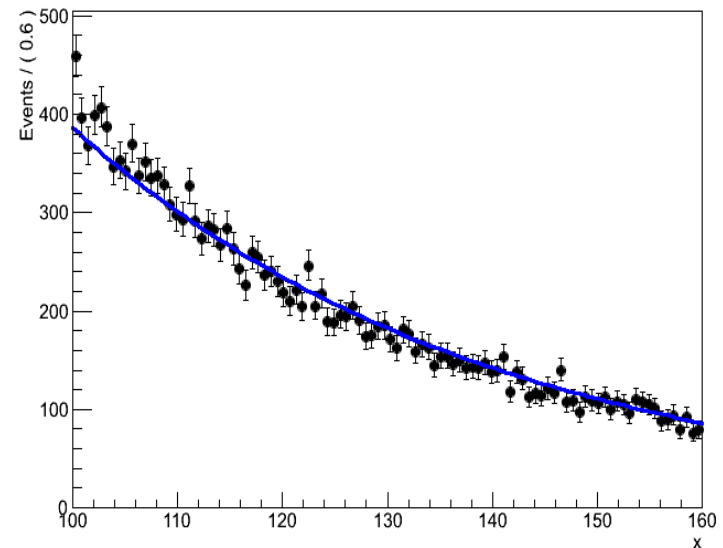
-> fitting error becomes significant

Un exemple de fit d'une somme d'exponentielle par une exponentielle



Obvious !

Un exemple de fit d'une somme d'exponentielle par une exponentielle



Less intuitive...

Objectives

- **Find a method to estimate these error**
- **Find the best fit for a chosen shape of the background**
- **Estimate the influence of the fitting range**

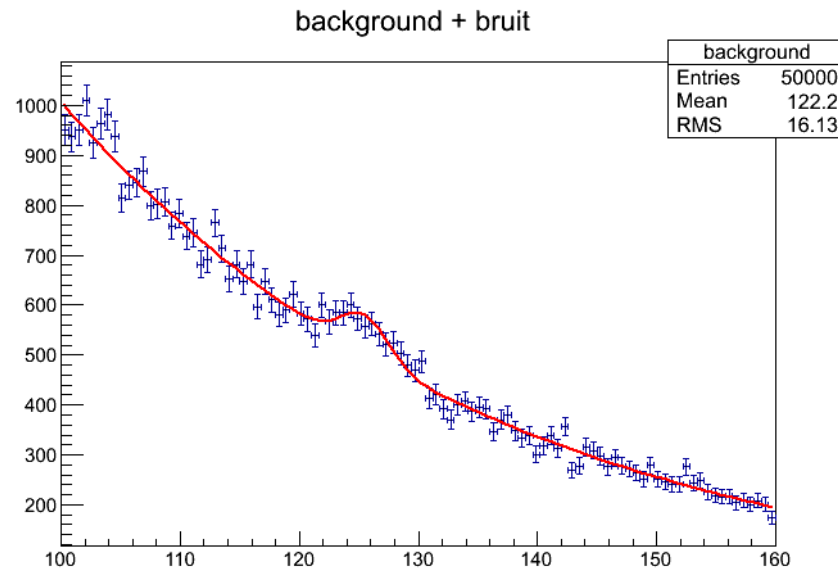
III – Fit the background

The method

- Produce toys with a given background model,
- fit with a given function (possible model of the background) + a Gaussian signal (fixed origin and sigma)
- measure the number of event of the Gaussian signal

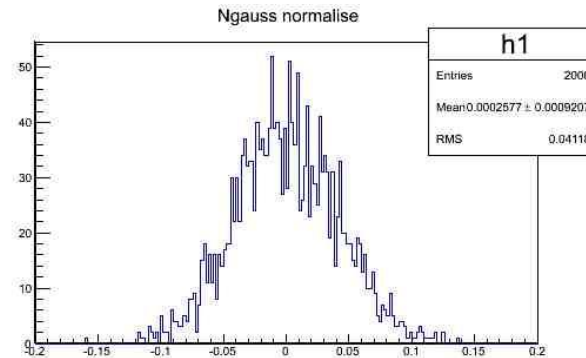
Perfect fit : Gaussian norm = 0

Imperfect fit : bias on the Gaussian norm



The method

- In order to measure the uncertainty caused by noise, we did a statistical study on 2000 toys :



Ex : toys generated with a double exponential fitted with a single exponential + Gaussian

- **Distribution of the number of signal events obtained per toy**
 - **Mean = bias = systematical uncertainty**
 - **RMS = statistical uncertainty**

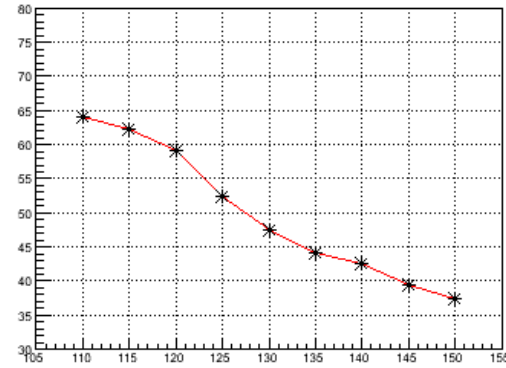
Study of an example

- **Distribution generated from a double exponential**
- **Fit with :**
 - Exponential
 - Exponential of 2 order polynomial
- **Study**

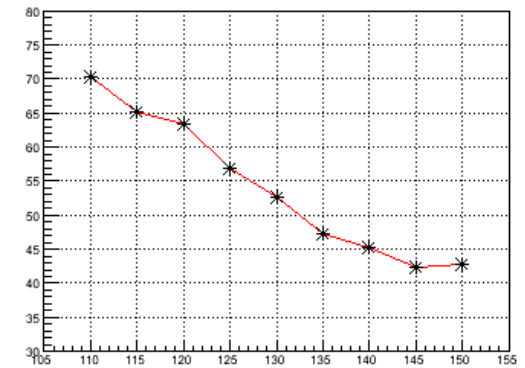
Fit of a double exponential by :

- A single exponential (right)
- A exponential of a 2nd order polynomial (left)

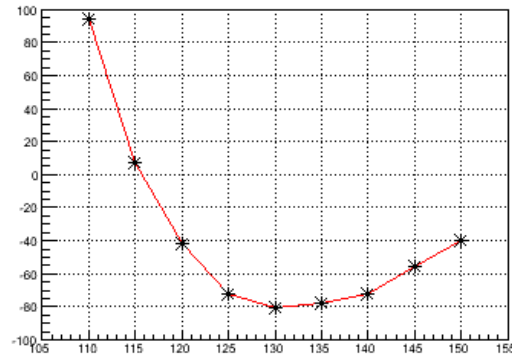
RMS_Ngauss vs mass (2expo fit by 1expo+gauss)



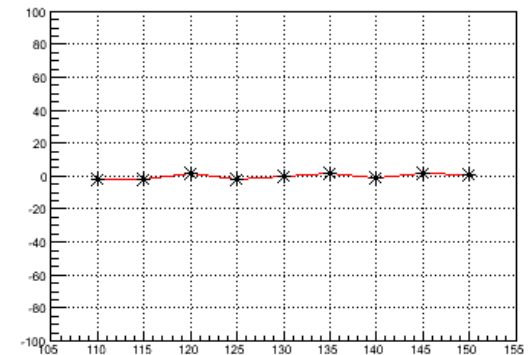
RMS_Ngauss vs mass (2expo fit by exppoly2+gauss)



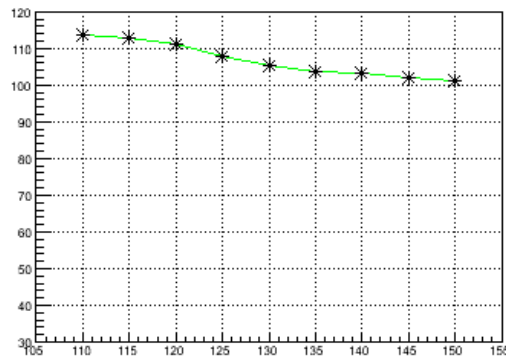
Ngauss vs mass (2expo fit by 1expo+gauss)



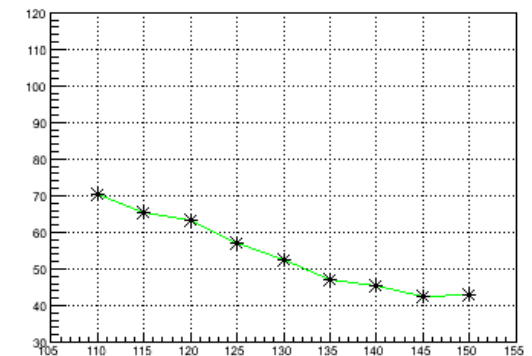
Ngauss vs mass (2expo fit by exppoly2+gauss)



Uncertainties = sqrt(RMSA²+max(error_sys)A²) vs mass (2expo fit by 1expo+gauss)



Uncertainties = sqrt(RMSA²+max(error_sys)A²) vs mass (2expo fit by exppoly2+gauss)



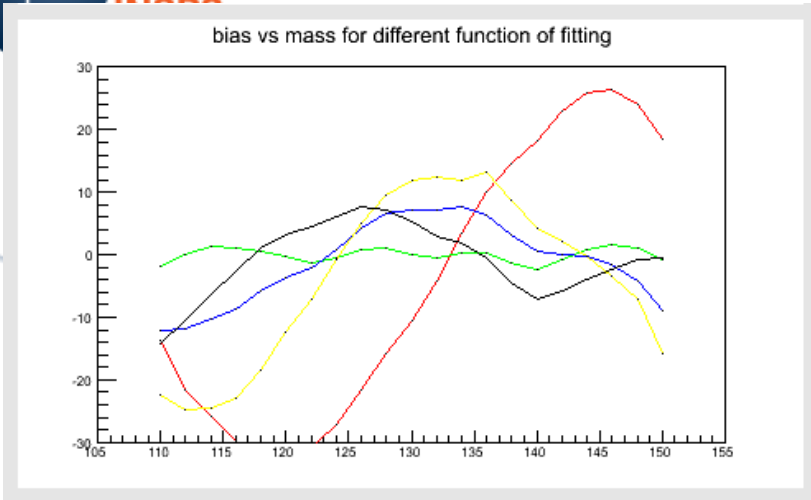
Conclusion

- **The exponential of 2nd order polynomial provides a better fit than the single exponential**

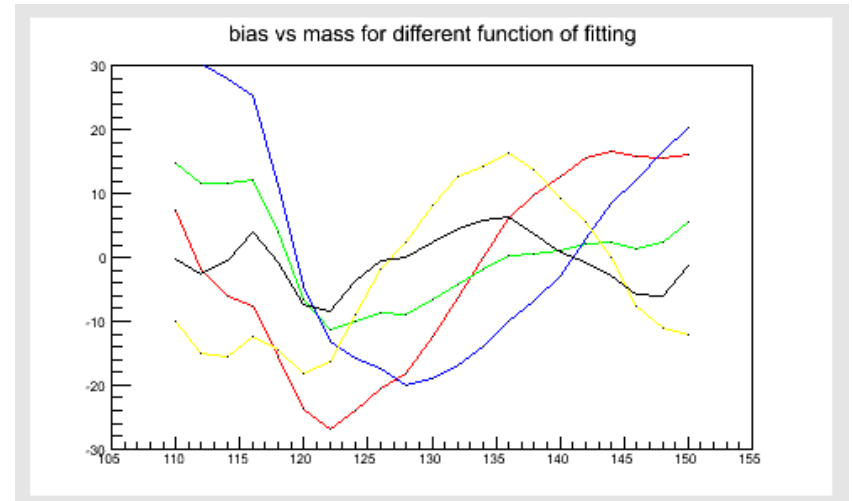
Test on DIPHOX, RESBOS and SHERPA

- **We test different background models (DIPHOX, RESBOS, SHERPA)**
- **We try to estimate :**
 - The RMS associated with the model choice.
 - The possible bias
 - The global uncertainty
- **We also look at possible ways for reducing the systematics uncertainties associated to these biases.**
- **We tested :**
 - Polynomials (2, 3, 4, 5, 6 order)
 - Bernstein polynomials (3,4, 5, 6 order)
 - Single – double exponentials + exponential of 2nd order polynomial

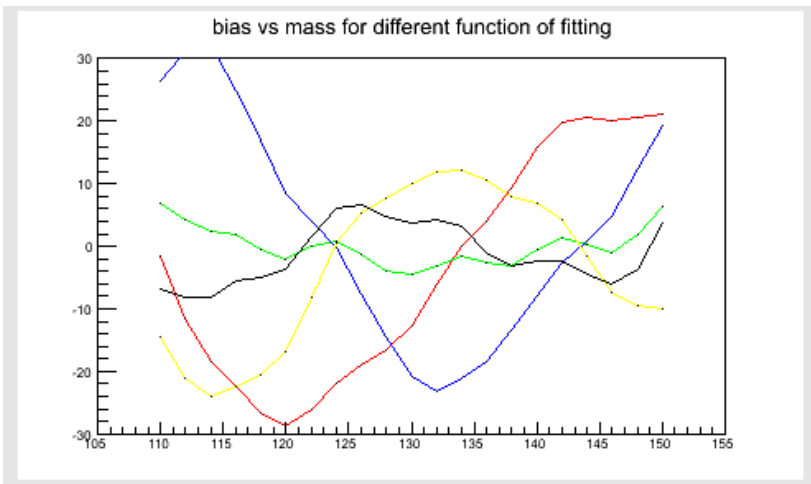
Bias : Polynomials



Diphox



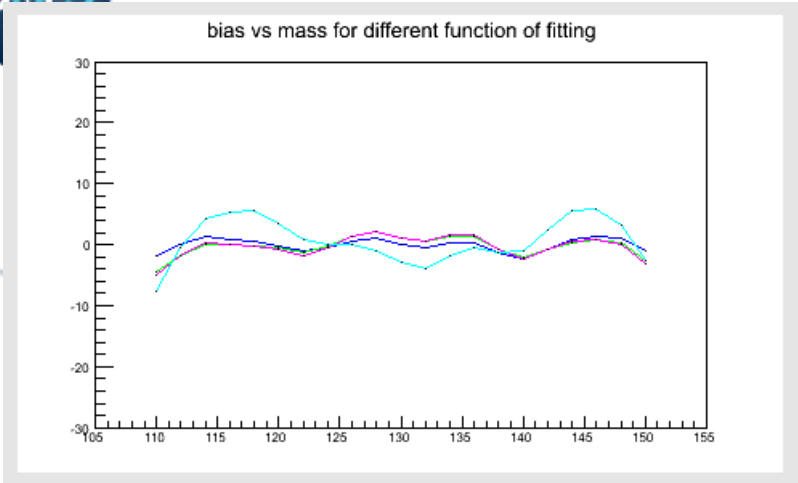
Resbos



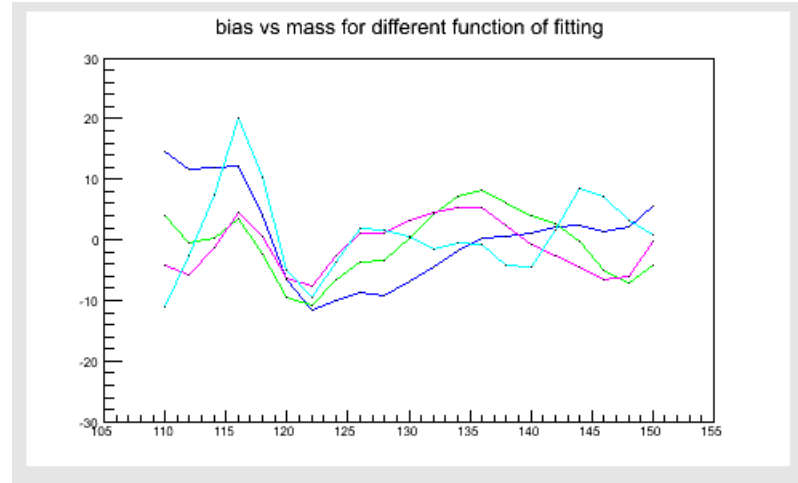
Sherpa

	2 nd degree polynomial
	3 nd degree polynomial
	4 nd degree polynomial
	5 nd degree polynomial
	6 nd degree polynomial

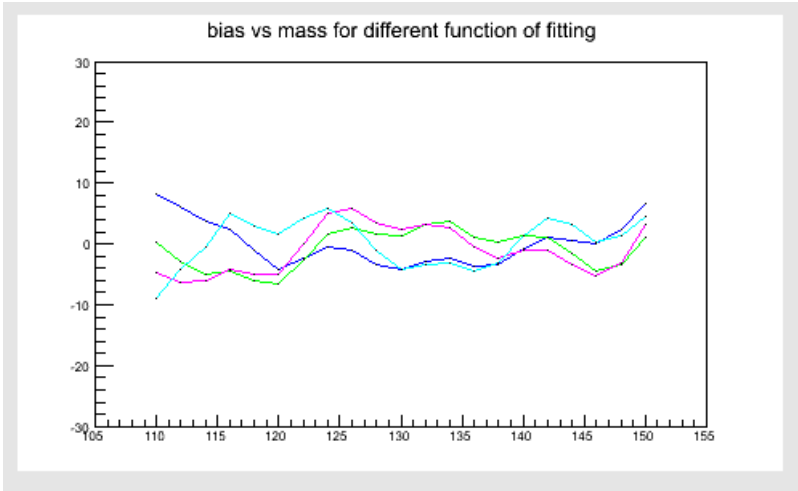
Bias : Bernstein polynomials







Diphox



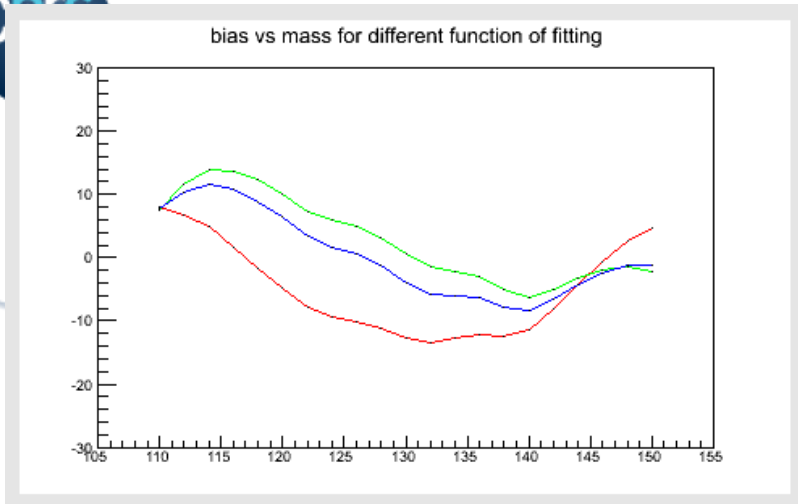
Resbos



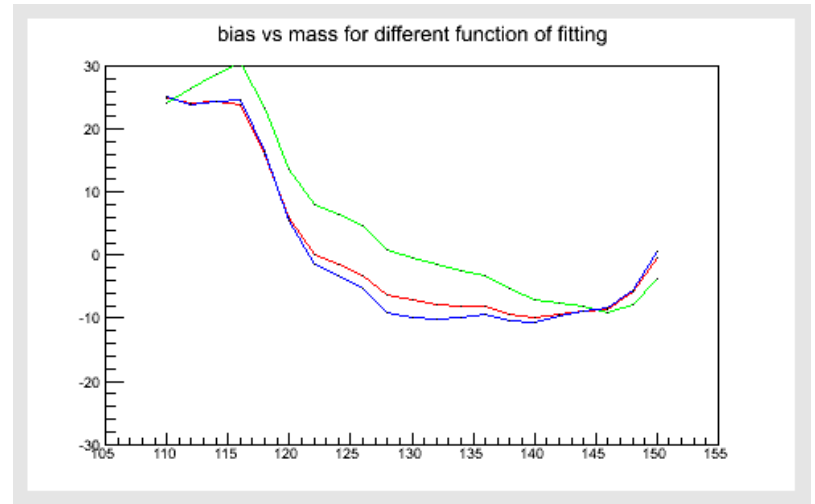
Sherpa

	3 bernstein's polynomials
	4 bernstein's polynomials
	5 bernstein's polynomials
	6 bernstein's polynomials

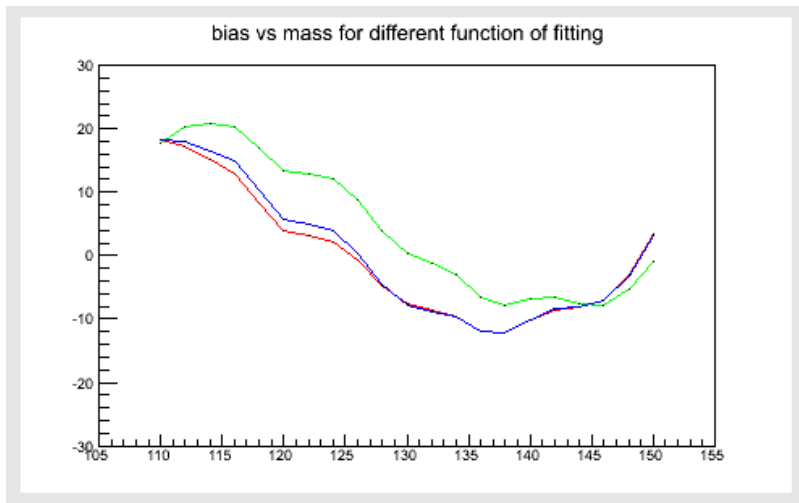
Bias : Exponentials






Diphox



Resbos

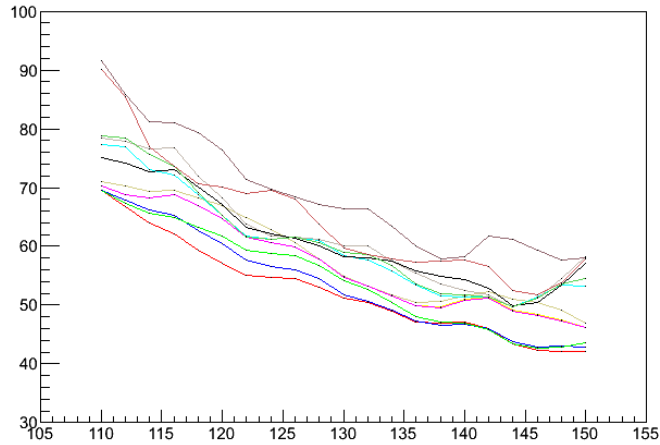


Sherpa

	Single exponential
	Double exponential
	exponential(2 nd degree polynomial)

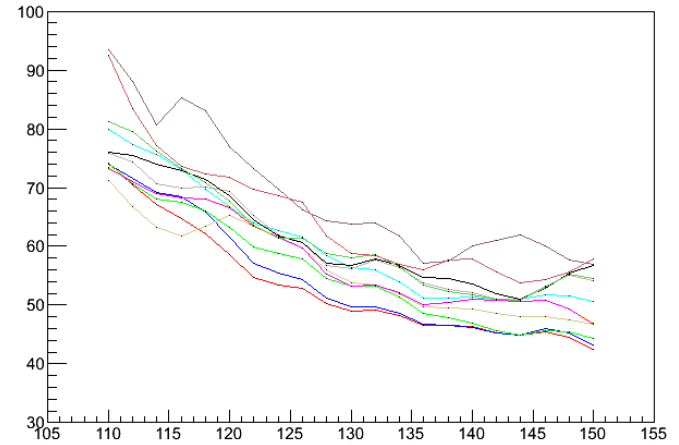
Global uncertainty

bias vs mass for different function of fitting



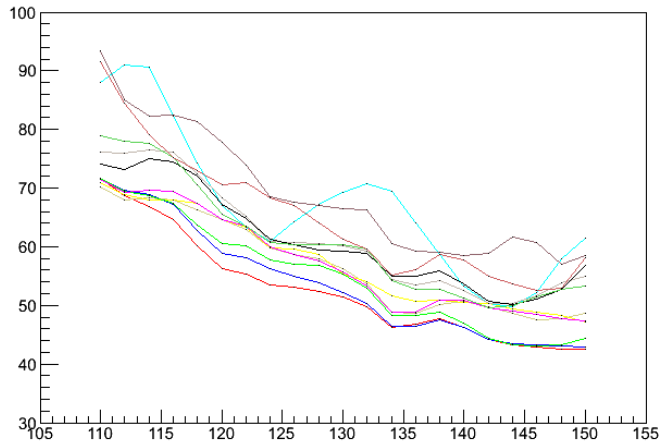
Diphox

bias vs mass for different function of fitting



Resbos

bias vs mass for different function of fitting



Sherpa

	2 nd degree polynomial
	3 rd degree polynomial
	4 nd degree polynomial
	5 nd degree polynomial
	6 nd degree polynomial
	exponential(2 nd degree polynomial)
	Double exponential
	Single exponential
	3 bernstein's polynomials
	4 bernstein's polynomials
	5 bernstein's polynomials

Analysis

- The uncertainty is essentially due to the RMS influence
- Important global uncertainty (about 70 events) for 200 expected events for the Higgs
- Comparison with Heberth's result form the C1 decay

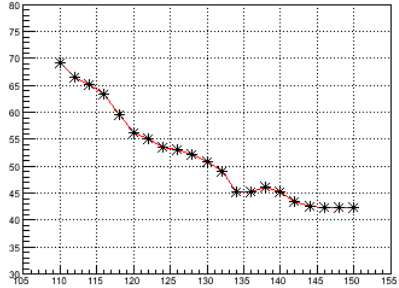
mass_Higgs = 125 GeV	Heberth		Jean	
	1 Expo	2 Expo	1 Expo	2 Expo
bias	- 5,20	0,13	- 3,7	0,12
RMS	13,57	14	13,90	13,9

- Validation of the model ?

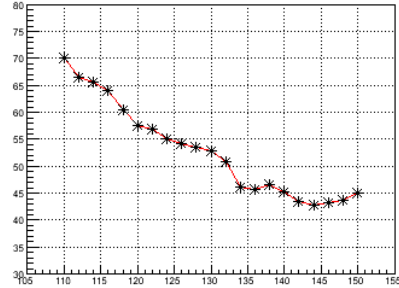
III – Influence of the fitting range

Test on sherpa for single exponential fit

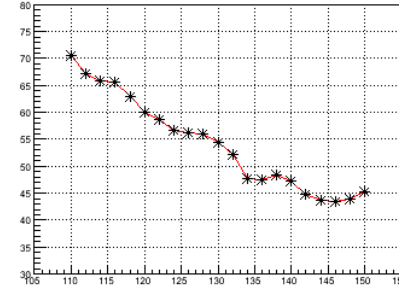
RMS_Ngauss vs mass (2expo fit by 1expo+gauss)



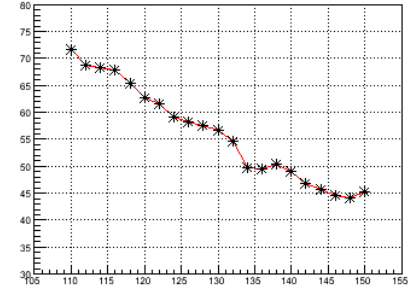
RMS_Ngauss vs mass (2expo fit by 1expo+gauss)



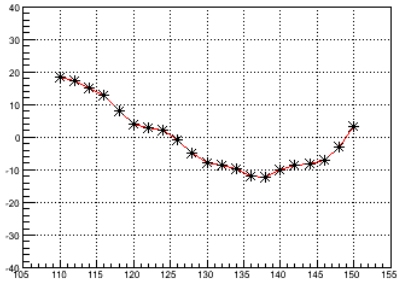
RMS_Ngauss vs mass (2expo fit by 1expo+gauss)



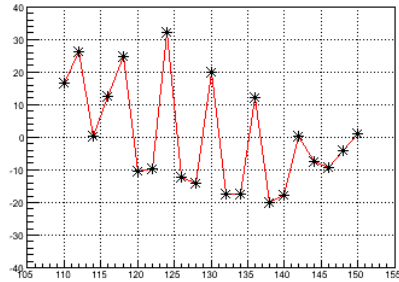
RMS_Ngauss vs mass (2expo fit by 1expo+gauss)



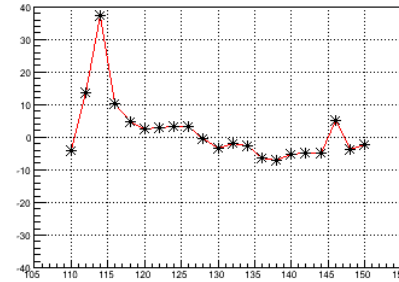
Ngauss vs mass (2expo fit by 1expo+gauss)



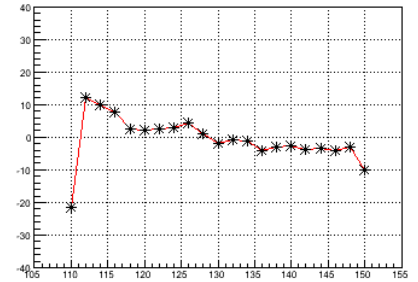
Ngauss vs mass (2expo fit by 1expo+gauss)



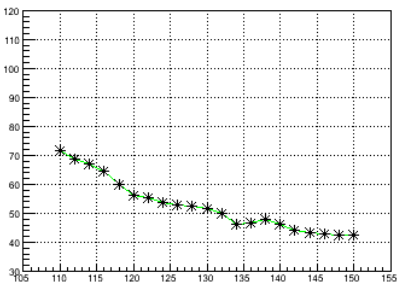
Ngauss vs mass (2expo fit by 1expo+gauss)



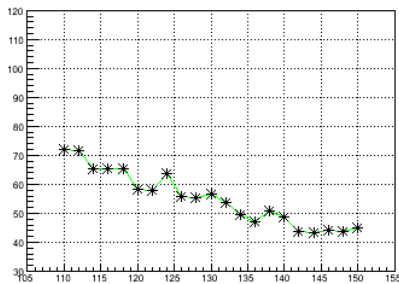
Ngauss vs mass (2expo fit by 1expo+gauss)



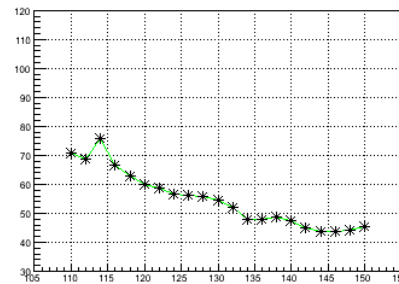
Uncertainties = $\sqrt{(\text{RMSA} + \text{max}(\text{error_sys})^2)}$ vs mass (2expo fit by 1expo+gauss)



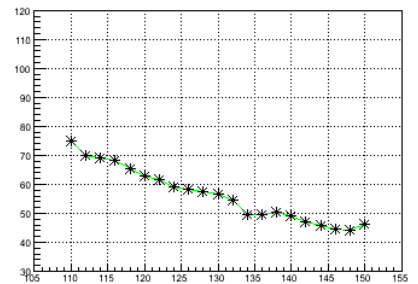
Uncertainties = $\sqrt{(\text{RMSA} + \text{max}(\text{error_sys})^2)}$ vs mass (2expo fit by 1expo+gauss)



Uncertainties = $\sqrt{(\text{RMSA} + \text{max}(\text{error_sys})^2)}$ vs mass (2expo fit by 1expo+gauss)



Uncertainties = $\sqrt{(\text{RMSA} + \text{max}(\text{error_sys})^2)}$ vs mass (2expo fit by 1expo+gauss)



No range

-/+ 20 GeV

-/+ 15 GeV

-/+ 12 GeV

Analysis

- **No real influence of the fitting range : the RMS compensate the bias modification**
- **Instability of the fit for little fit range ($< \pm 10$ GeV)**

Conclusion

- **Validation of the method to study the error of the fit functions**
- **Huge uncertainty mainly due to the RMS**
- **Differences of fit between Resbos, Diphox and Sherpa depending on the functions**
- **No real influence of the fit range**



Merci de votre attention