



Particle Physics

at the

high-precision / low-energy frontier

using

UltraCold Neutrons

at the Institut Laue-Langevin in Grenoble

Peter Geltenbort

Very Hot Neutrons

10^7 eV



Ultracold Neutrons

10^{-7} eV

Setting the scene



Thanks to the organizers for the very warm and generous "registration reception"

Setting the scene



Le berceau de l'Institut Laue-Langevin

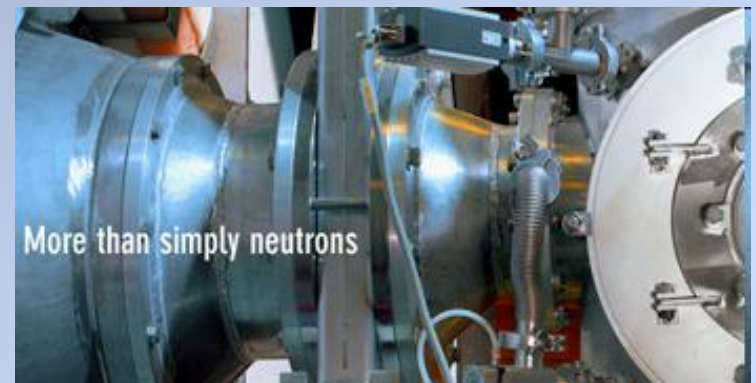
GRENOBLE

The home of the Institut Laue-Langevin

Die Wiege des Instituts Laue-Langevin



With its international funding and expertise the Institut Laue-Langevin (ILL) offers scientists and industry the world's leading facility in neutron science and technology. From its Grenoble site in the south-east of France the Institute operates the most intense neutron source on Earth



The ILL is firmly committed not only to building high-performance instruments but also to offering the best scientific environment to the user community



Max von Laue

"A neutron factory and an user facility"

founded 17 January 1967
International Convention (renewal every 10 years)
signed until end 2013

first neutrons in 1971

cold and hot neutrons
sources started operation
in 1972



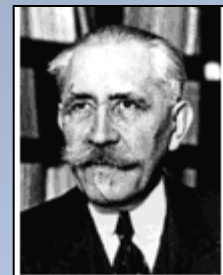
H. Maier-Leibnitz

general refit from 1991 - 94

"earthquake" refit from 2003 - 07

Millennium Programme

phase M-0 done
phase M-1 running
phase M-2 in planning (141M€ in total)



Paul Langevin



L. Neel

Associates : France, Germany, United Kingdom

Scientific Member Countries : A, B, CH, CZ, DK, E, H, I, IND, RUS, PL, S, SK

Further "Candidate" Countries: FIN, N, NL, FIN, RO, SLO, ... covers about **95%** of European neutron users

Fields of research

solid-state physics, material science,
chemistry, bio- and earth sciences,
engineering,
nuclear and particle (fundamental) physics

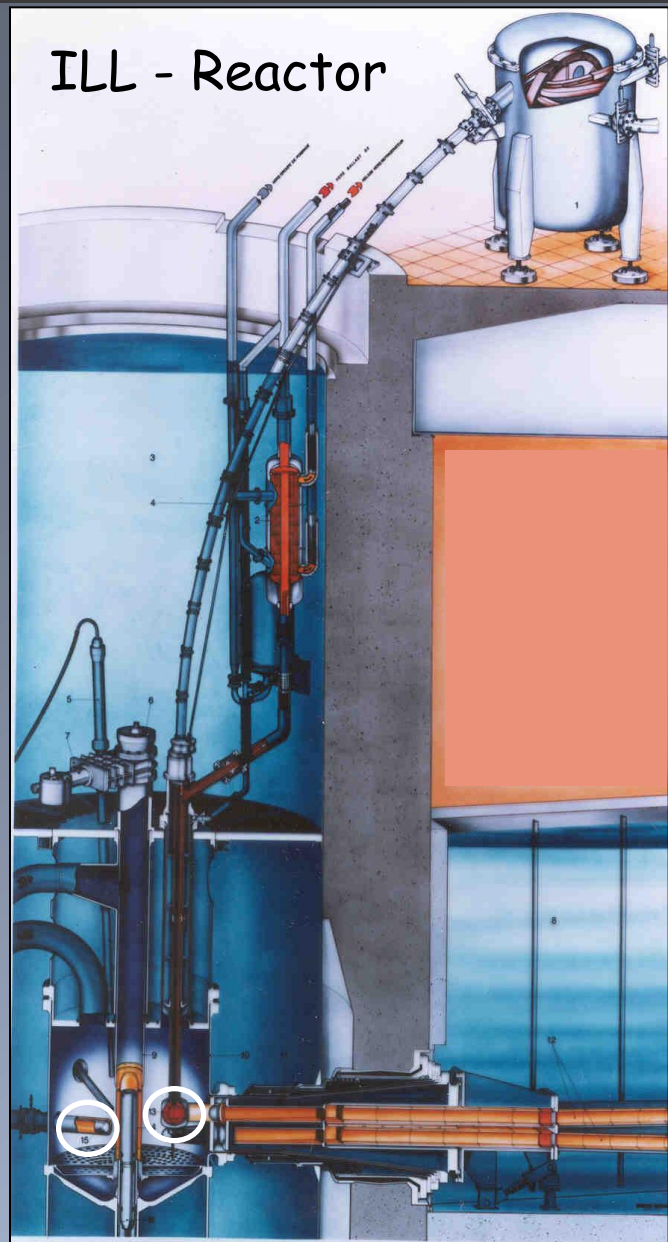
Experimental Programme in 2010 (3 cycles of 50 days)

- 740 experiments (allocated by subcommittees)
on 28 ILL-funded and 10 CRG instruments
- 1056 visitors coming from 35 countries
- 1142 proposals submitted and 670 accepted
- 652 publications by ILL staff and users

489 staff (66 sci. + 32 PhD stud.); 88.5M€ annual budget

Neutron source(s) at ILL

ILL - Reactor



Fuel (chain reaction): $^{235}\text{U}(n_{\text{th}},f) \rightarrow$ fission neutrons

Moderator: D_2O at 300K \rightarrow thermal neutrons

Hot source: 10 dm³ of graphite at 2400 K

Cold source^h (horizontal): 6 dm³ of liquid D_2 at 25 K

Cold source (vertical): 20 dm³ of liquid D_2 at 25 K

Ultracold
Neutrons

Cold
Neutrons

Reactor
Neutrons

Temperature (K)

10^{-8}

10^1

10^3

Energy (eV)

10^{-7}

10^{-3}

10^{-1}

10^7

Velocity (m/s)

5

800

2200

Nuclear physics

PN1 (LOHENGRIN)
Recoil mass spectrometer for fission fragments

PN3 (GAMS)
Ultra-high resolution gamma ray spectrometer

Particle physics

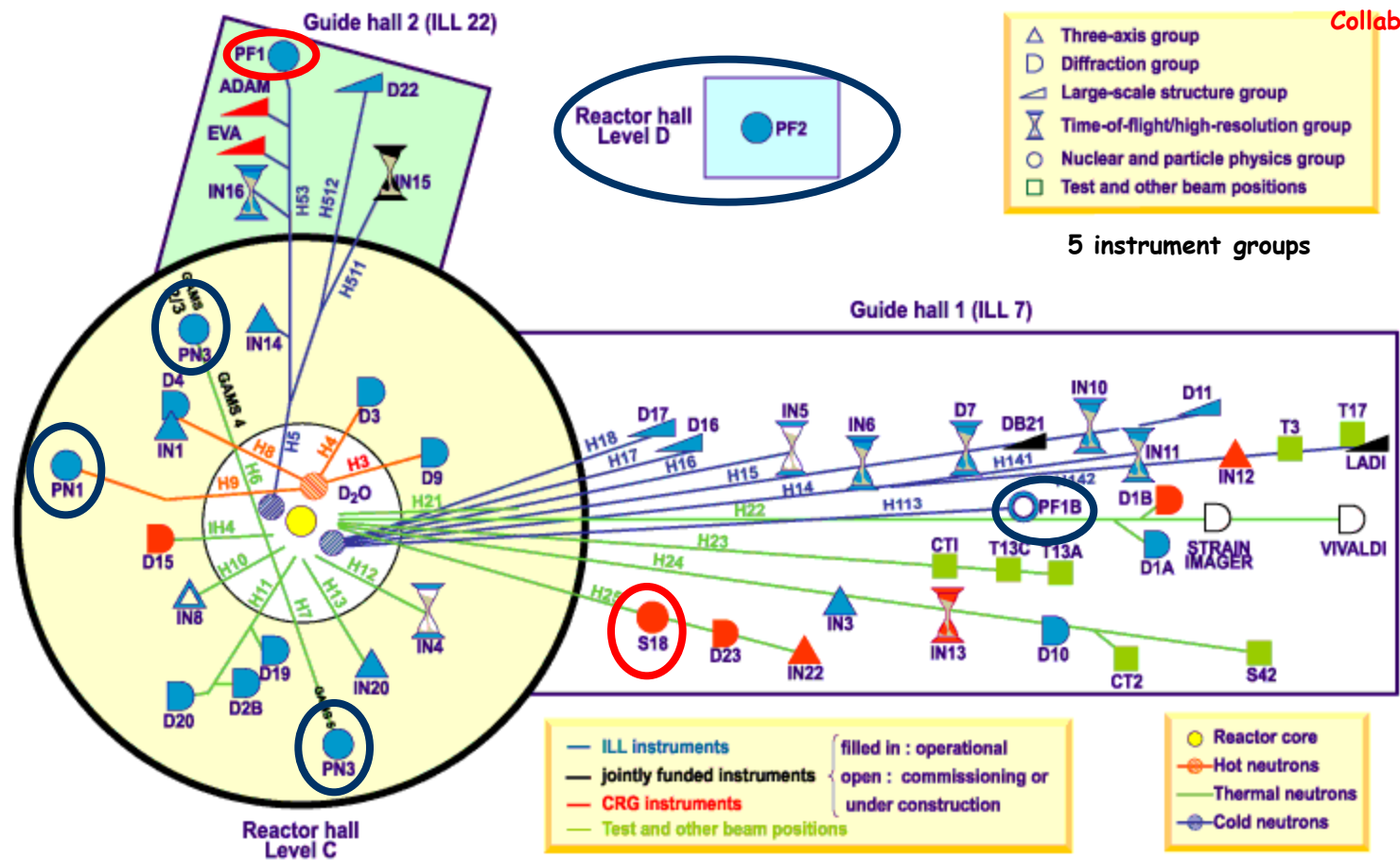
PF1B
Facility for cold neutrons

(former PF1)
cryoEDM experiment

PF2
Facility for ultracold and very cold neutrons

S18 - perfect crystal
neutron interferometer

ILL-funded



Collab. Research Group

5 instrument groups



ESRF

(6 GeV Synchrotron)

ILL

(High Flux Reactor)

Nuclear physics

PN1 (LOHENGRIN)
Recoil mass spectrometer for fission fragments

PN3 (GAMS)
Ultra-high resolution gamma ray spectrometer

Particle physics

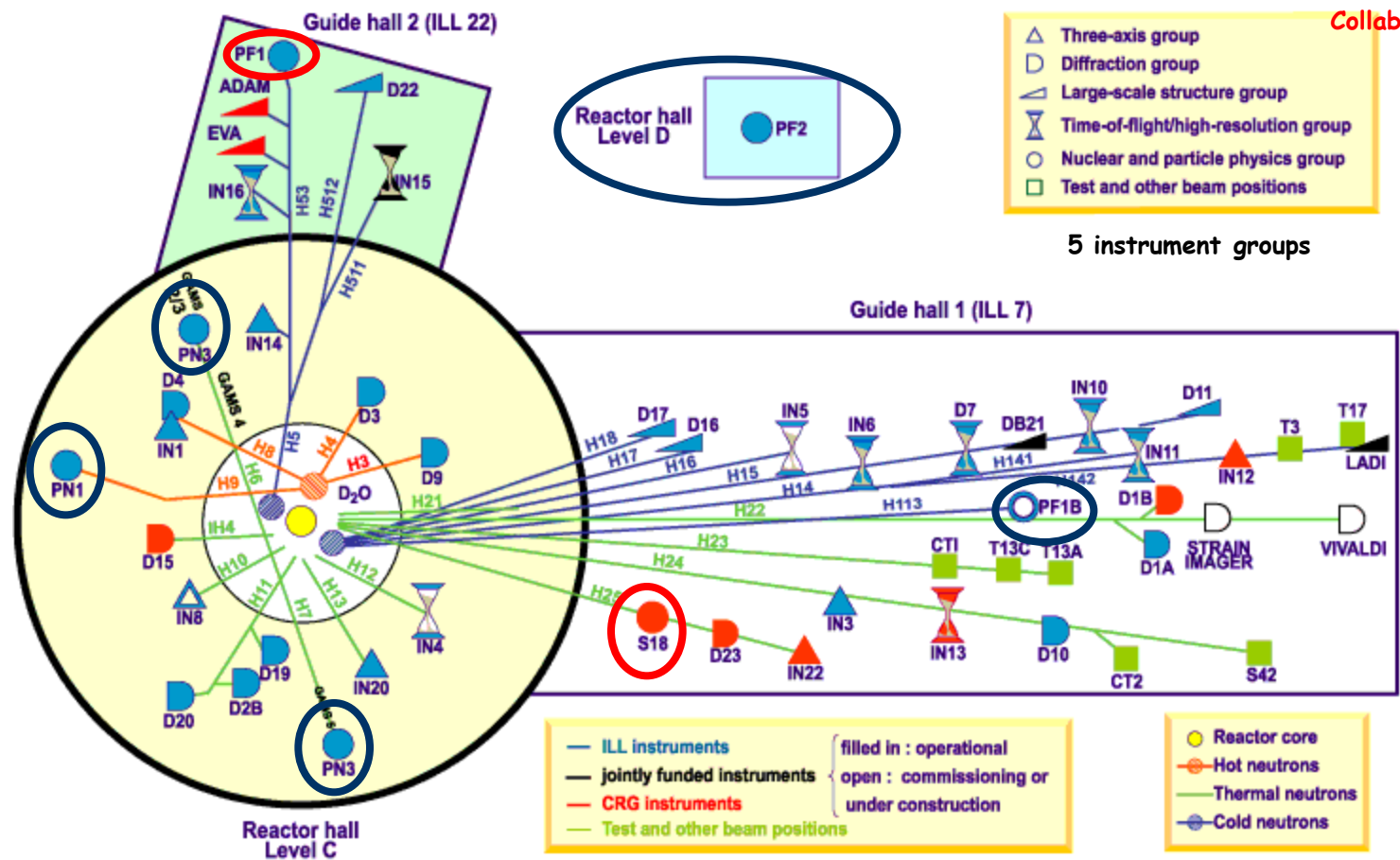
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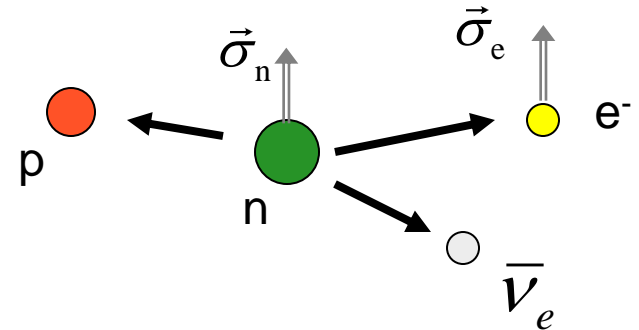
Collab. Research Group

5 instrument groups

Observables in Neutron Beta Decay

Jackson et al., PR 106, 517 (1957):

Observables in Neutron beta decay, as a function of generally possible coupling constants (assuming only Lorentz-Invariance)



$$dw \propto \rho(E_e) \cdot (1 + 3|\lambda|^2) \cdot \left\{ 1 + a \frac{\vec{p}_e \cdot \vec{p}_v}{E_e E_v} + b \frac{m_e}{E_e} + \vec{\sigma}_n \cdot \left(A \frac{\vec{p}_e}{E_e} + B \frac{\vec{p}_v}{E_v} + D \frac{\vec{p}_e \times \vec{p}_v}{E_e E_v} \right) \right\}$$

Fierz interference term $b = 0$

Beta-Asymmetry $A = -2 \frac{|\lambda|^2 + \text{Re } \lambda}{1 + 3|\lambda|^2}$

Neutrino-Electron-Correlation $a = \frac{1 - |\lambda|^2}{1 + 3|\lambda|^2}$

Neutron lifetime $\tau_n^{-1} = \frac{2\pi}{\hbar} G_F^2 V_{ud}^2 (1 + 3|\lambda|^2) \int \rho(E_e)$

Hot topic questions beyond the SM

- What do we learn from V_{ud} and quark mixing?
- What is the origin of P-violation?
- T-violation?
- Additional forces
- Number of quark generations
- Neutrino helicity
- Search for RHC: W-mass and mixing ζ
- CP-violation

Properties of UCN

Ultracold neutrons, that is, neutrons whose energy is so low that they can be contained for long periods of time in material and magnetic bottles

$$E_{\text{kin}} (\sim 5 \text{ ms}^{-1}) = 100 \text{ neV} (10^{-7} \text{ eV})$$

$$\lambda_{\text{UCN}} \sim 1000 \text{ \AA}$$

$$T_{\text{UCN}} \sim 2 \text{ mK}$$

UCN are totally reflected from suitable materials at *any* angle of incidence, hence **storable!**

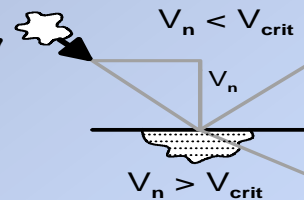
Long storage and observation times possible (up to several minutes)!

High precision measurements of the properties of the free neutron (lifetime, electric dipole moment, gravitational levels, ...)

Interaction with matter:
UCN see a *Fermi-Potential* E_F

$E_F \sim 10^{-7} \text{ eV}$ for many materials, e.g.

- beryllium 252 neV
- stainless steel 200 neV



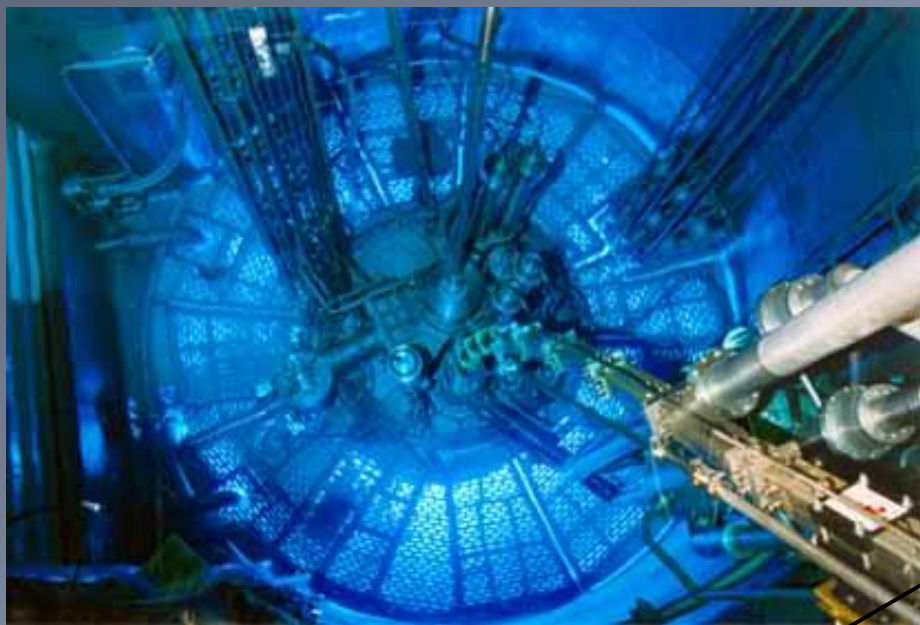
UCN are furthermore storable by gravity and/or magnetic fields

Fermi potential	$\sim 10^{-7} \text{ eV}$
Gravity $\Delta E = m_n g \Delta h$	$\sim 10^{-7} \text{ eV} / \text{Meter}$
Magnetic field $\Delta E = \mu_n B$	$\sim 10^{-7} \text{ eV} / \text{Tesla}$

The UCN/VCN facility PF2



NEUTRONS
FOR SCIENCE

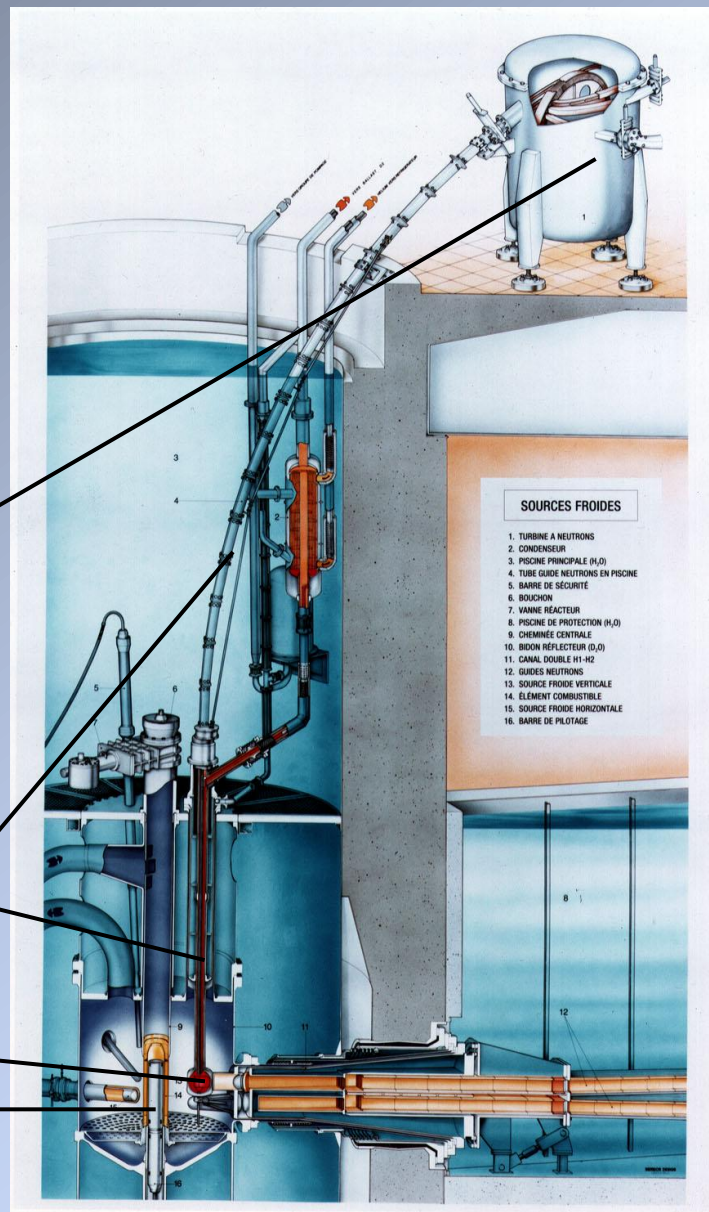


Neutron turbine
A. Steyerl (TUM - 1985)

Vertical and curved guide tubes

Cold source

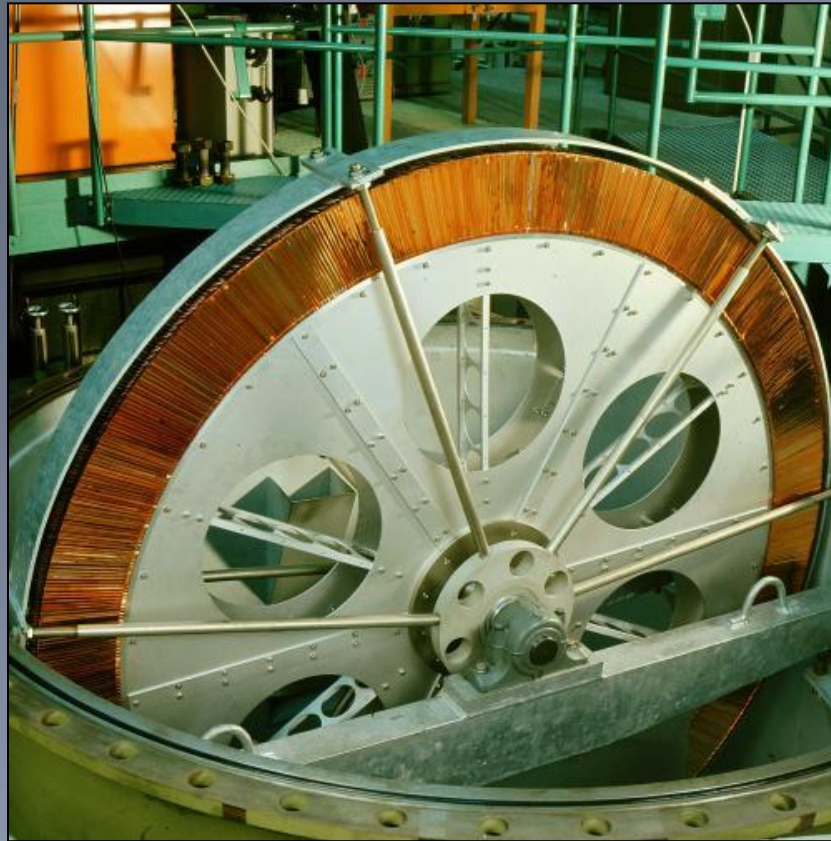
Reactor core



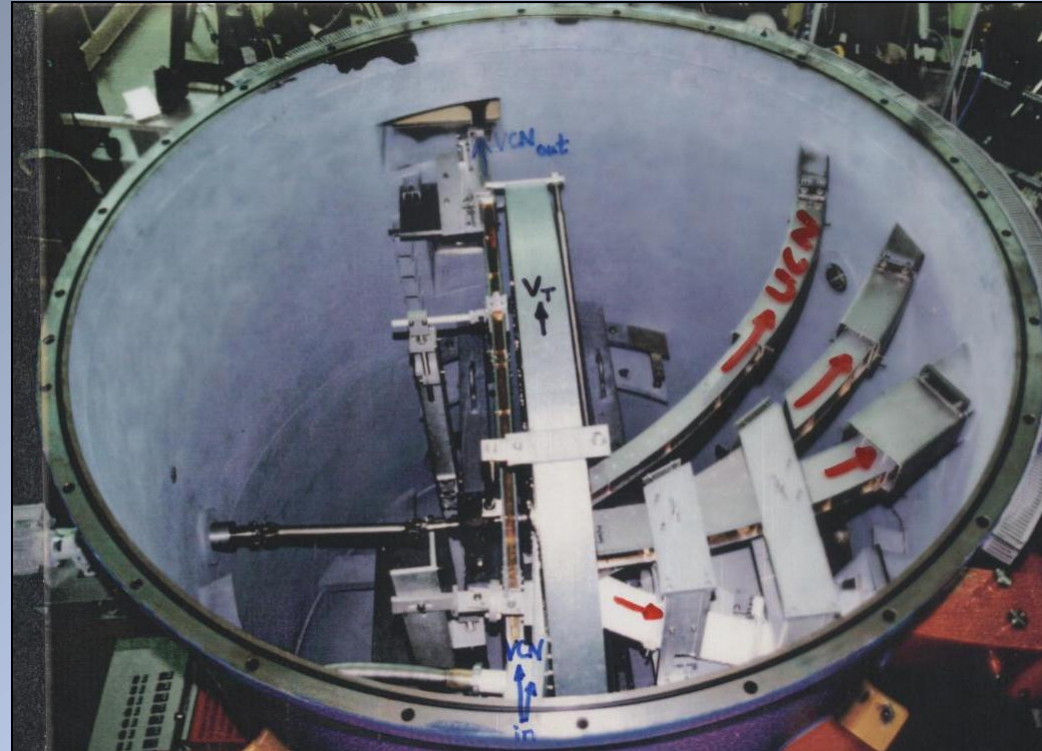
The Vertical Cold Source (VCS)



NEUTRONS
FOR SCIENCE

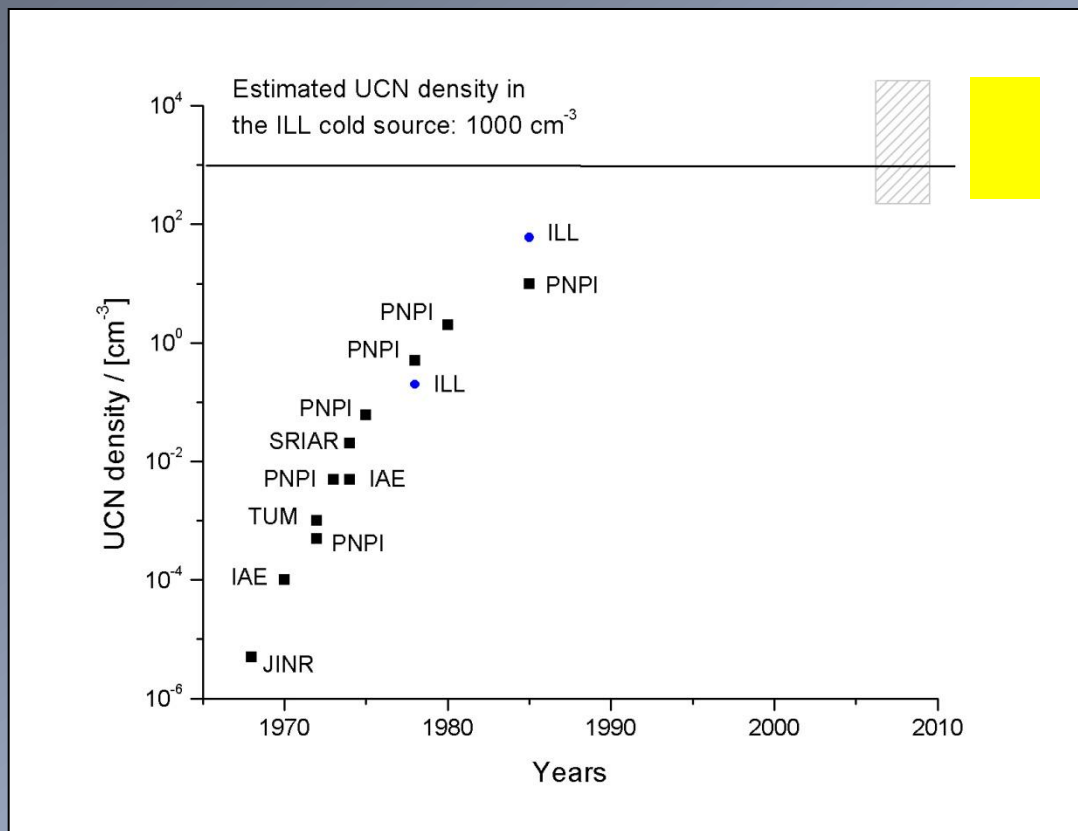


Steyerl turbine
at FRM-I (Munich)



Steyerl turbine (2nd generation)
at PF2 / ILL
10 years later

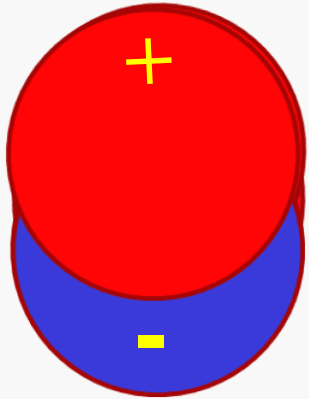
UCN facilities - Status and Future



More UCN facilities in the future worldwide

- PSI (CH)
- Mainz / Munich (D)
- ILL (F)
- PNPI (RUS)
- LANL / NCSU [SNS/ NIST] (USA)
- RCNP / JPARC (J)
later at TRIUMF (CAN)

EDM CP violation



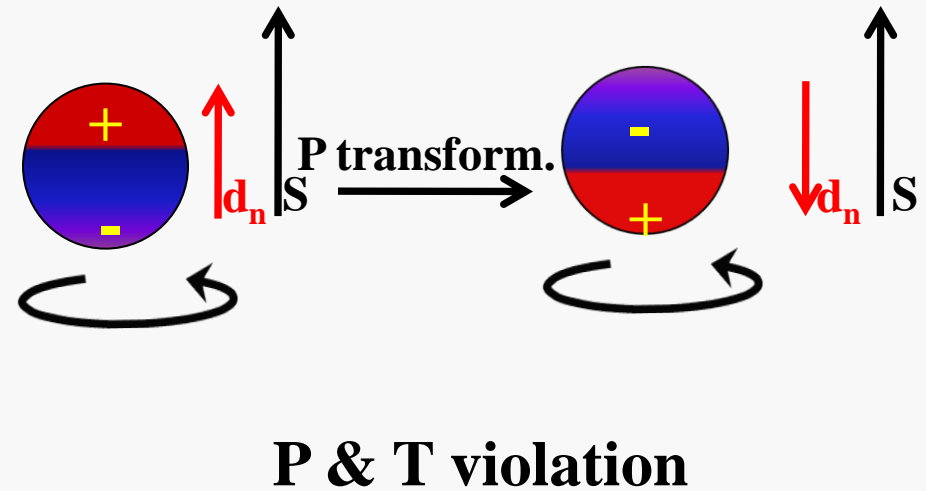
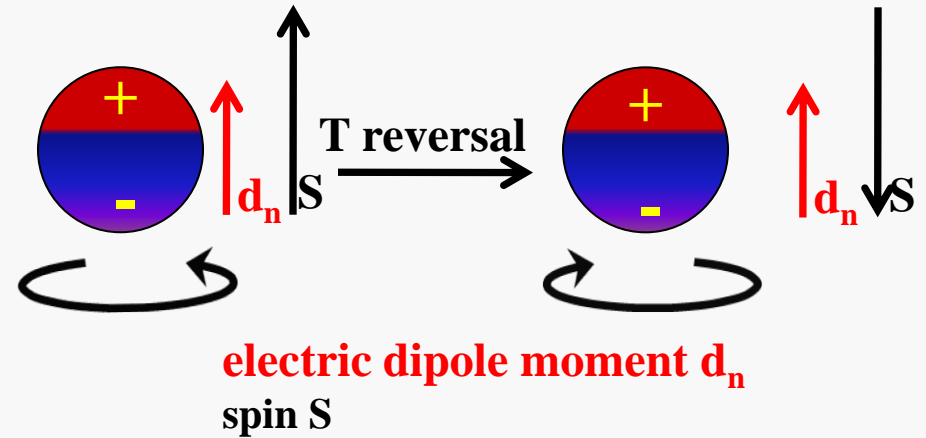
Electric Dipole Moment:

electrically neutral or charged particles

If there is a charge distribution:



EDM



CPT conservation \rightarrow CP violation

EDM CP violation

The Electric Dipole Moment: d_n

$d_n \neq 0 \Rightarrow P$ and T violation

CP violation observed in K decay,
B mesons $\rightarrow d_n \neq 0$

CP violation, interest

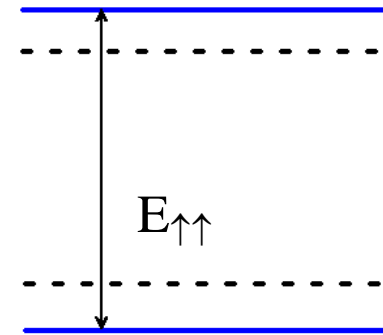
- The study of CP violation is important to:
 - Understanding the fundamental laws of physics
 - Understanding the baryon asymmetry of the cosmos
- EDM is a particularly promising laboratory for CP violation
 - The Standard Model contribution is very small
 - Contributions from new physics tend not to be

nEDM: spin frequency

Experiments:

Measurement of Larmor precession frequency of polarised neutrons in a magnetic & electric field

Compare the precession frequency for parallel fields:

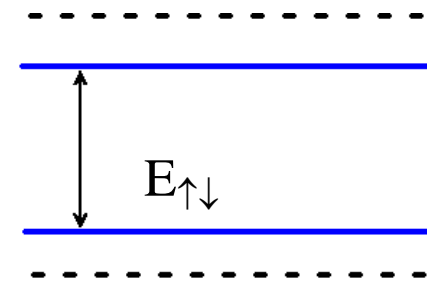


$$\nu_{\uparrow\uparrow} = E_{\uparrow\uparrow}/h = [-2B_0\mu_n - 2Ed_n]/h$$

The difference is proportional to d_n and E :

$$\mathbf{h}(\nu_{\uparrow\uparrow} - \nu_{\uparrow\downarrow}) = 4E d_n$$

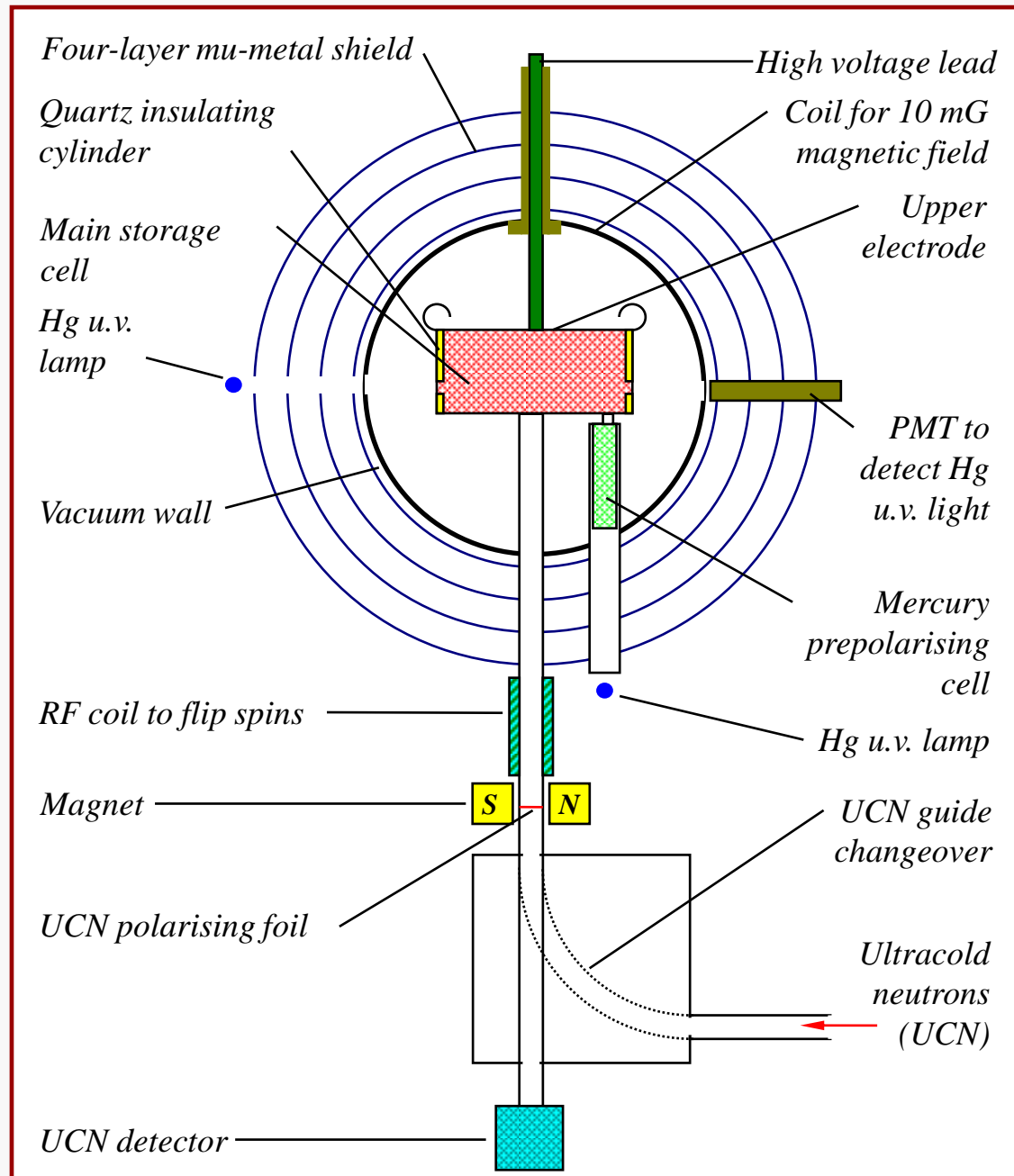
to the precession frequency for anti-parallel fields



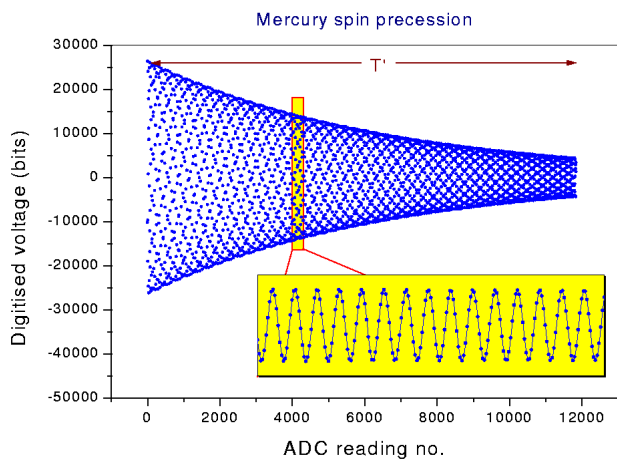
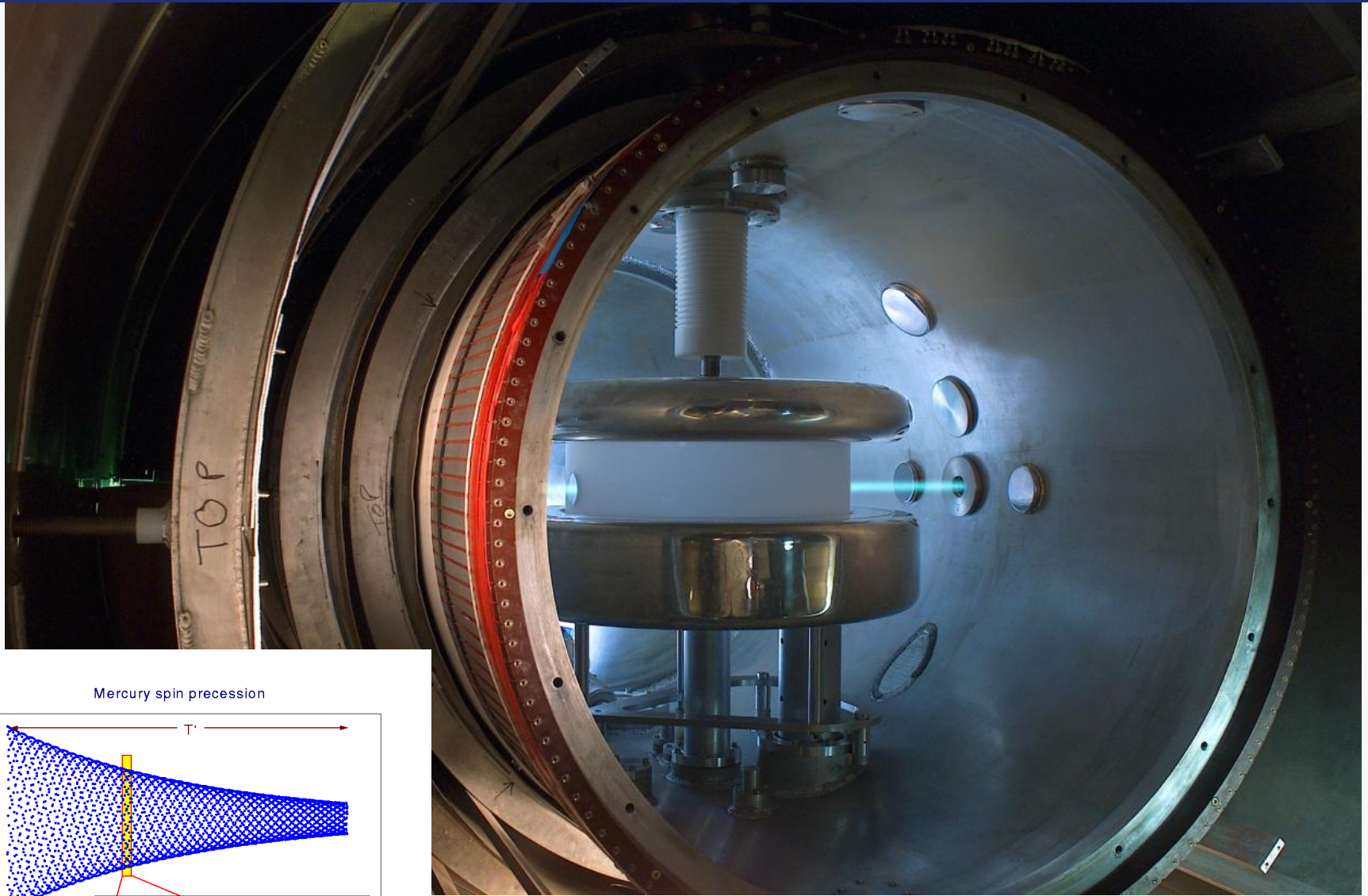
$$\nu_{\uparrow\downarrow} = E_{\uparrow\downarrow}/h = [-2B_0\mu_n + 2Ed_n]/h$$

Need to measure change in Larmor precession frequency to a very high degree : $< 1 \mu\text{Hz}$
 < 1 turn per month!

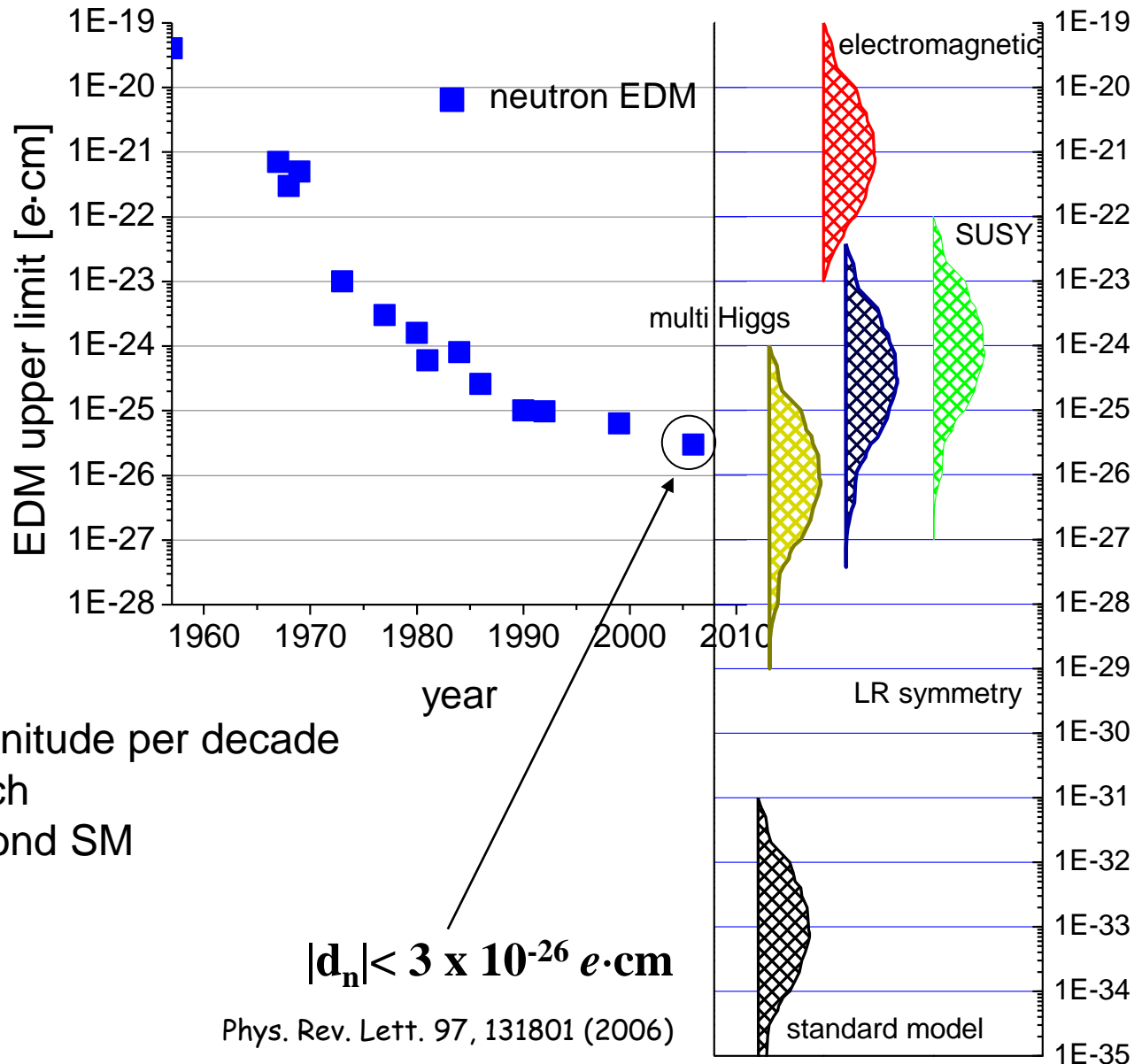
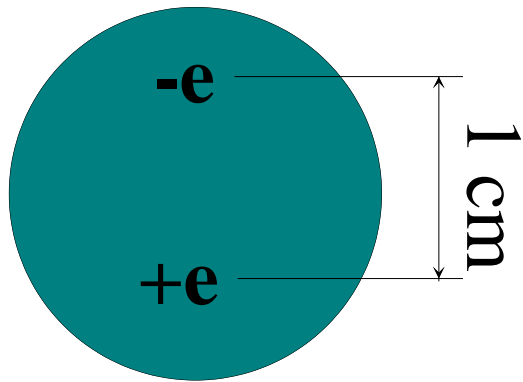
nEDM: ILL



nEDM: ILL



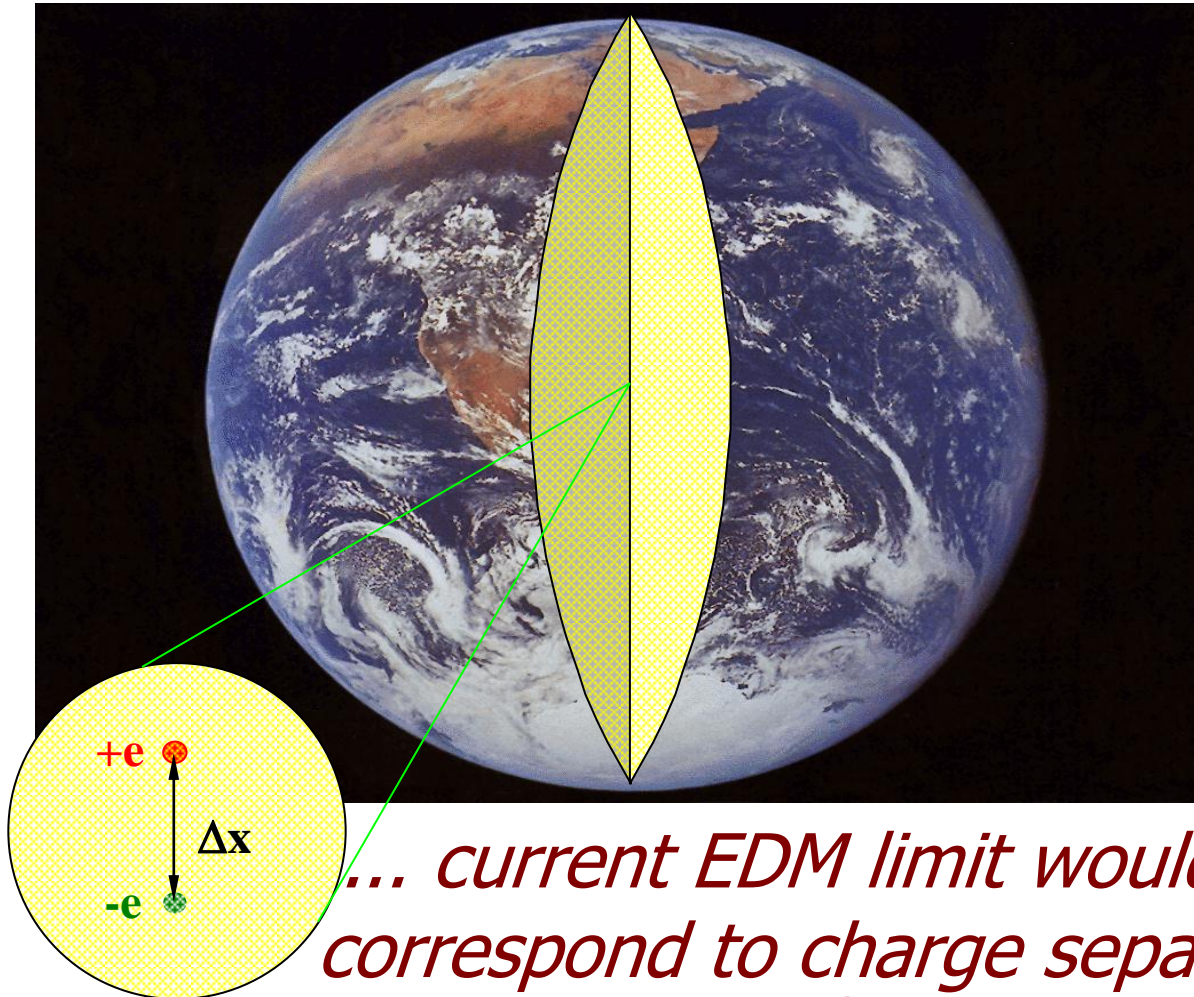
EDM models - experiment



Progress at ~ order of magnitude per decade
 Standard Model out of reach
 constraints on models beyond SM

Reality check

If neutron were the size of the Earth...



... current EDM limit would correspond to charge separation of
 $\Delta x \approx 3\mu$

nEDM: cryoEDM

$$\sigma(d_n) = \frac{\hbar}{2ET\alpha\sqrt{N}}$$

new superthermal UCN source

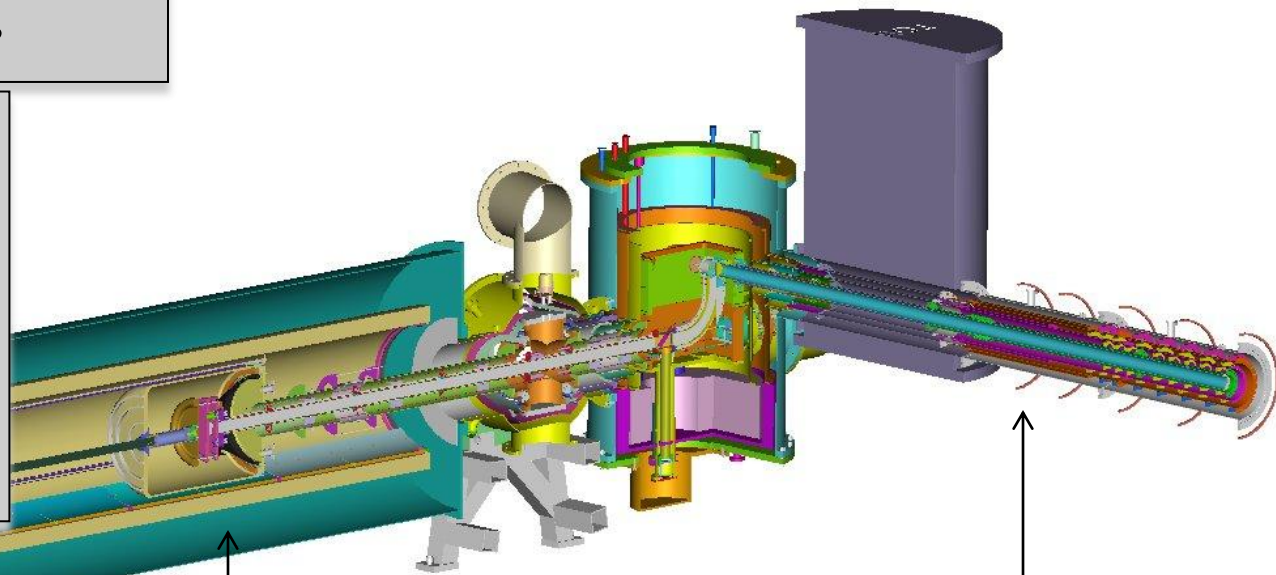
higher electric fields in l He

Ultra-cold neutron source:

Strong 9 Å neutron beam coupled to superthermal UCN source

magnetic screening:

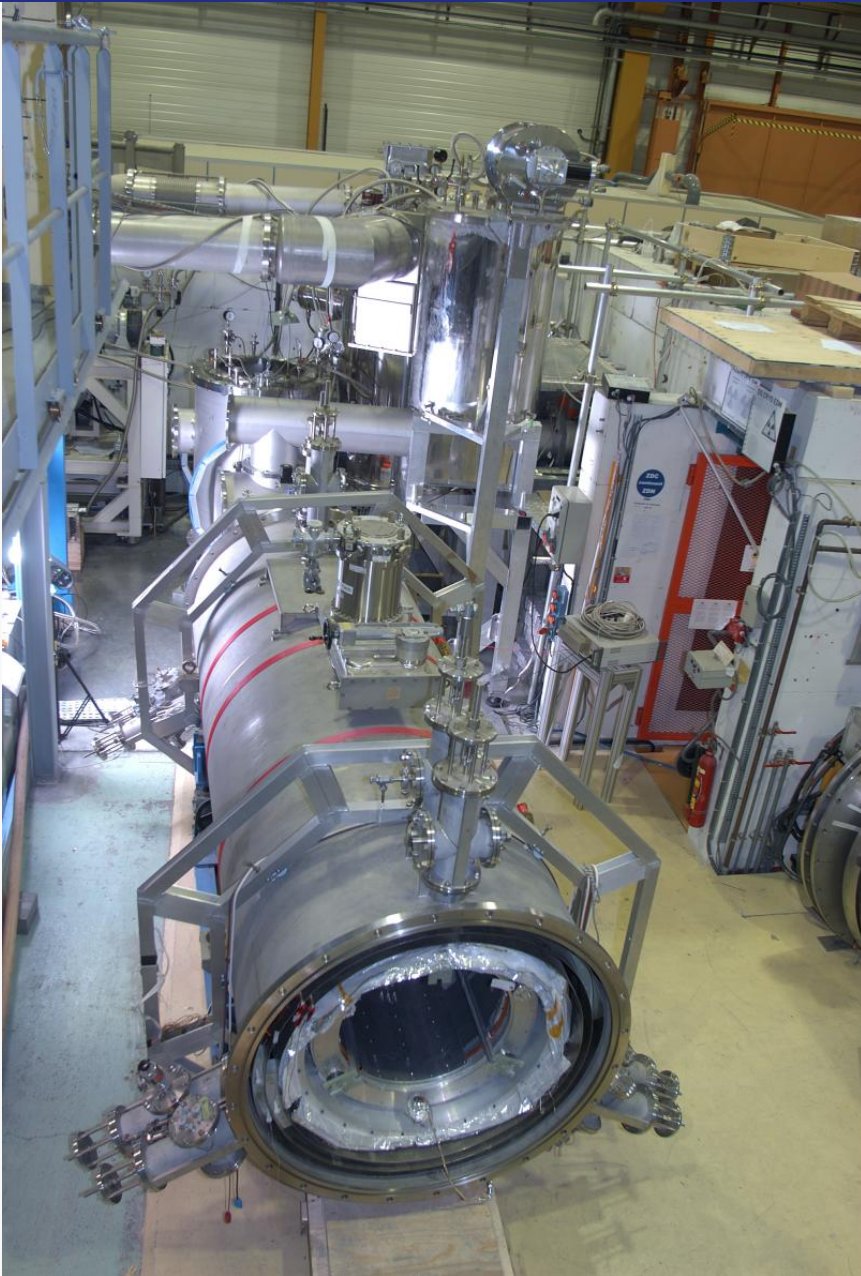
- *mu*-metal & superconducting lead
- high precision SQUID magnetometry



whole experiment in superfluid He at 0.5 K

- production of UCN
- storage & Larmor precession of UCN
- magnetometry
- detection of UCN

nEDM: cryoEDM outlook



*Cryogenic apparatus working:
base temperature $\sim 0.5\text{K}$*

*UCN production, transfer &
storage system working => needs
optimising*

*HT applied on system => needs
enhancing*

*Magnetic field system operational
=> will require further improving*

***Experiment runs (commissioning)
but most parameters need
improvement for order of $\sim 10^{-27}$
e·cm measurement***

Search for Neutron - Mirror Neutron Oscillations using storage of Ultracold Neutrons

PNPI/IPTI/ILL collaboration: A. Serebrov et al., E. Alexandrov et al., P. Geltenbort, O. Zimmer

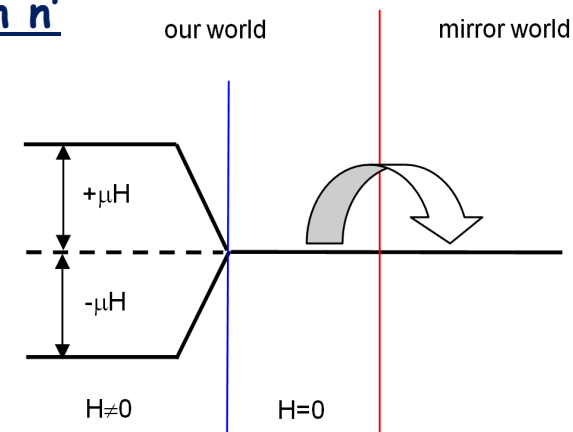
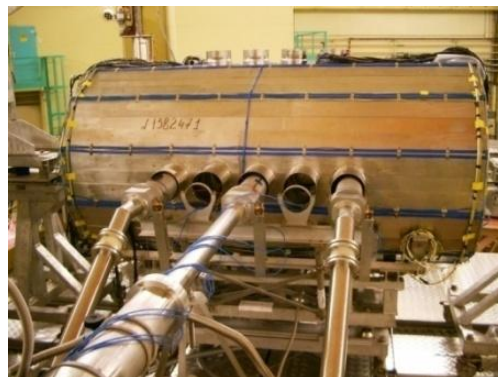
Hypothesis: There is a "mirror world" of partners of the known particles with

- same fundamental interactions with opposite handedness
→ natural explanation of parity violation
- no interactions with our world, apart gravity and mixing of neutral particles
→ mirror matter is a viable dark-matter candidate

Z. Berezhiani, A.D. Dolgov and R.N. Mohapatra, Phys. Lett. B **375**, 26 (1996)

Test: Search transition of neutron n to mirror neutron n'

- Situation 2006: $\tau_{osc} \geq 1 \text{ s}$
- A magnetic field suppresses nn' mixing
→ Look for difference of UCN storage time without ($< 20 \text{ nT}$) and with field ($2 \mu\text{T}$)



Result with PNPI EDM-setup at PF2:

$$\tau_{osc} (90\% \text{ C.L.}) \geq 414 \text{ s}$$

A. Serebrov et al., arXiv:0706.3600; submitted to PLB

Search for Macroscopic CP Violating Forces Using a Neutron EDM Spectrometer[†]

A. P. Serebrov^a, O. Zimmer^b, P. Geltenbort^b, A. K. Fomin^a, S. N. Ivanov^a, E. A. Kolomensky^a,
I. A. Krasnoshekhova^a, M. S. Lasakov^a, V. M. Lobashev^a, A. N. Pirozhkov^a,
V. E. Varlamov^a, A. V. Vasiliev^a, O. M. Zherebtsov^a, E. B. Aleksandrov^c,
S. P. Dmitriev^c, and N. A. Dovator^c

^a Petersburg Nuclear Physics Institute, Russian Academy of Sciences, Gatchina, Leningrad region, 188300 Russia

e-mail: serebrov@pnpi.spb.ru

^b Institut Laue Langevin, BP 156, 38042 Grenoble, France

^c Ioffe Physico-Technical Institute, Russian Academy of Sciences, St. Petersburg, 194021 Russia

Received December 8, 2009

The search for CP violating forces between nucleons in the so-called axion window of force ranges λ between 2×10^{-5} m and 0.02 m is interesting because only little experimental information is available there. Axion-like particles would induce a pseudo-magnetic field for neutrons close to bulk matter. A laboratory search investigates neutron spin precession close to a heavy mirror using ultracold neutrons in a magnetic resonance spectrometer. From the absence of a shift of the magnetic resonance we established new constraints on the coupling strength of axion-like particles in terms of the product $g_s g_p$ of scalar and pseudo-scalar dimensionless constants, as a function of the force range λ , $g_s g_p \lambda^2 \leq 2 \times 10^{-21}$ [cm²] (C.L. 95%) for 10^{-4} cm $< \lambda < 1$ cm. For 0.1 cm $< \lambda < 1$ cm previous limits are improved by 4 to 5 orders of magnitude.

DOI: 10.1134/S0021364010010029

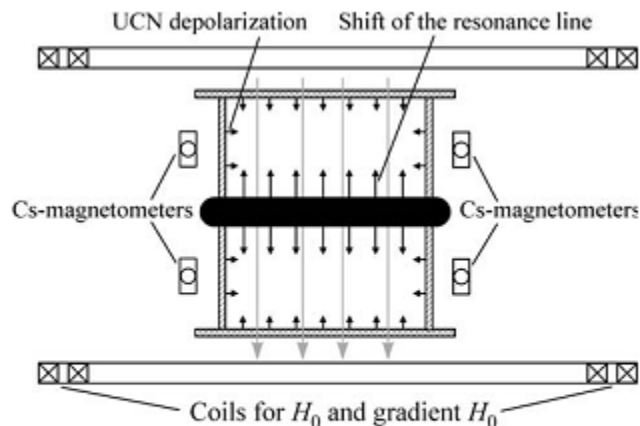


Fig. 1. Experimental setup with double storage chamber for UCN, Cs-magnetometers and coils for field setting and correction.

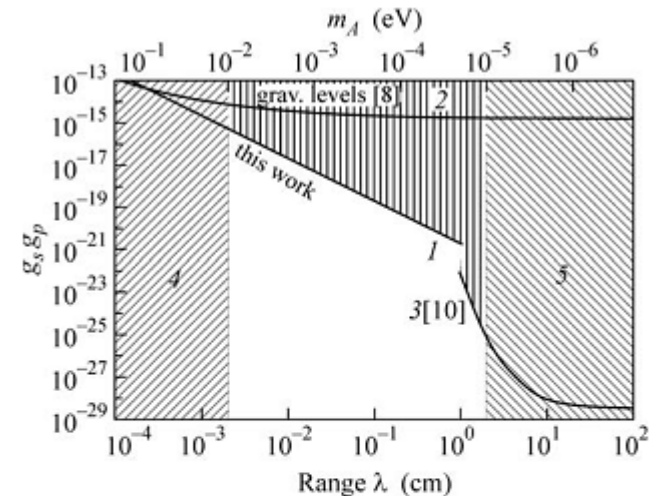


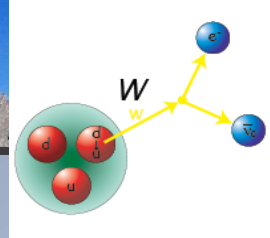
Fig. 3. Constraints to the coupling constant product $g_s g_p$ of axion-like particles to nucleons as a function of the range λ of the macroscopic force. On the upper horizontal axis the mass range of the axion-like particle is shown using the relation $\lambda = \hbar/m_A c$. Line 1: shift of resonance (this work); line 2: gravitational levels [8]; line 3 [10]; line 4: astrophysically excluded region of axion mass [3]; line 5: cosmologically excluded region in model of cold dark matter [3].

Is our world really isotropic?

Abstract: *Physics at the Planck scale could be revealed by looking for tiny violations of fundamental symmetries in low energy experiments. In 2008, we have performed a sensitive test of the isotropy of the Universe using stored Ultra Cold Neutrons, obtaining the first limits on the coupling of a free neutron with a hypothetical cosmic axial field.*



Figure 1 : The hypothetical cosmic axial field, defining a privileged direction in the universe, is searched for in daily modulation of precision observables.



The free neutron lifetime: $n \rightarrow p + e^- + \bar{\nu}_e$ (+782 keV)

$$\frac{1}{\tau_n} \propto G_F^2, V_{ud}^2, \lambda^2 \quad \lambda = \frac{g_A}{g_V}$$

$$n \rightarrow p + e^- + \bar{\nu}_e + \gamma \quad BR(15keV) \approx 3 \times 10^{-3}$$

$$n \rightarrow H^0 + \bar{\nu}_e \quad BR \approx 4 \times 10^{-6}$$

Together with measurements of asymmetry coefficients in neutron decay

Weak interaction theory

Neutrino physics

Cosmology

Extraction of g_V, g_A and V_{ud}

Test of Conserved Vector Current (CVC: ' $g_V = 1$ ')

Solar pp-process:

$$p + p \rightarrow d + e^+ + \nu_e \quad \sigma \propto g_A^2$$

Test of Unitary of CKM matrix ($V_{ud}^2 + V_{us}^2 + V_{ub}^2 = 1$)

Big bang:

Primordial elements' abundances

Neutrino induced reactions:

$$\bar{\nu}_\mu + p \rightarrow \mu^+ + n$$

$$\nu_\mu + n \rightarrow \mu^- + p$$

Neutrino detectors:

$$p + \bar{\nu}_e \rightarrow n + e^+$$

$$\sigma \propto \frac{1}{\tau_n}$$

Important input parameter for tests of the Standard Model of the weak interaction

Necessary to understand matter abundance in the Universe

Necessary to calibrate Neutrino Detectors and to predict event rates

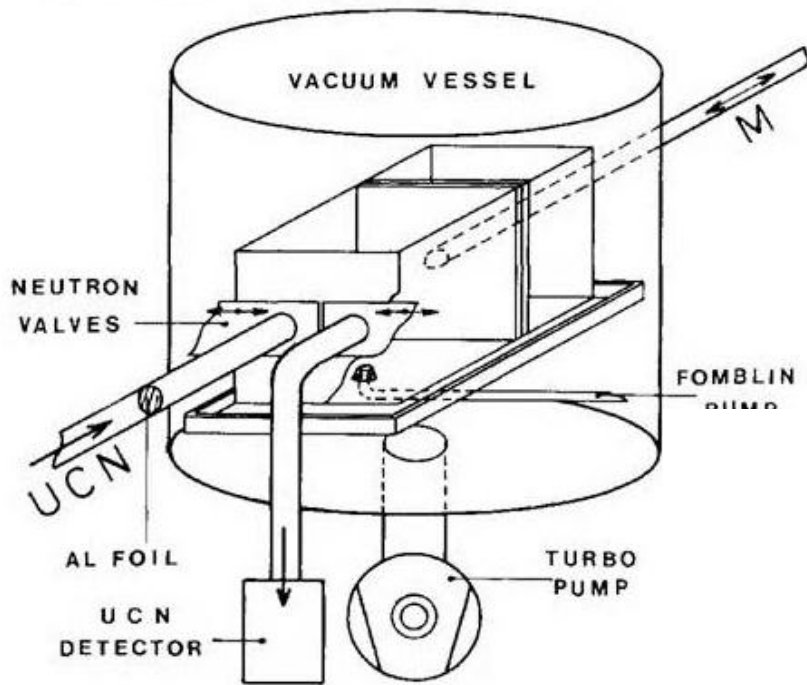


FIG. 1. Sketch of the apparatus.

Glass walls:

$H=0.3$ m, $W=0.4$ m

$L=0.5$ m ... 0.01 m

(surface A and volume V sizeable)

$$\frac{1}{\tau_m} = \frac{1}{\tau_\beta} + \frac{1}{\tau_{\text{wall}}} + \dots$$

$\tau_{\text{wall}} \rightarrow$ number of wall collisions,
i.e. mean free path λ

Measure storage lifetime τ_m

- for different

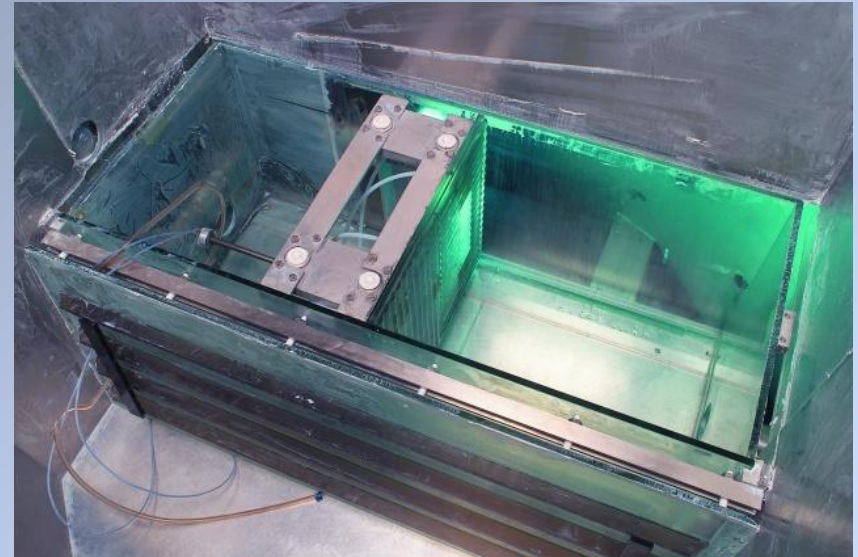
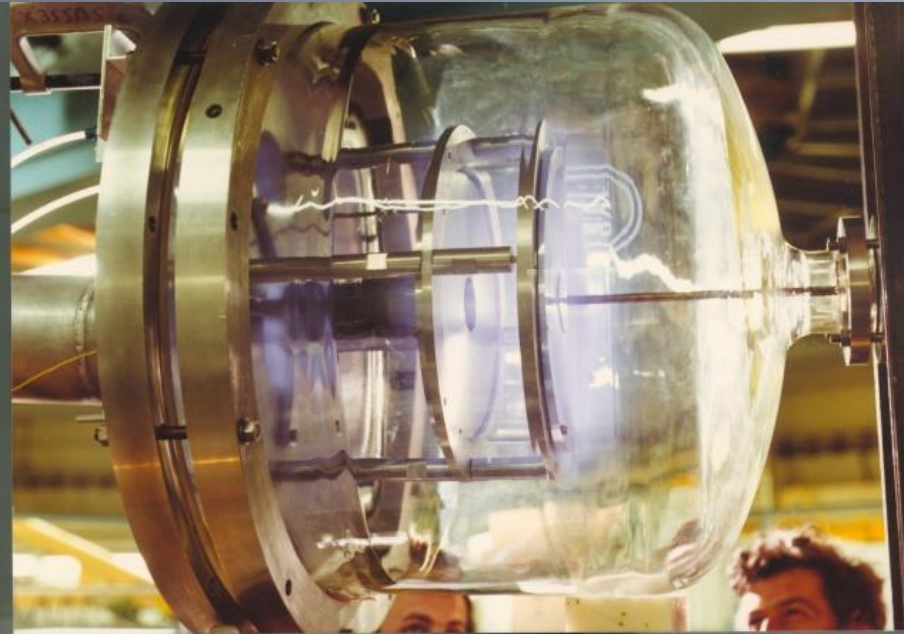
volume to surface ratios V/A

- and extrapolate for $V \rightarrow \infty$

$$\frac{1}{\tau_{\text{wall}}} \rightarrow 0$$

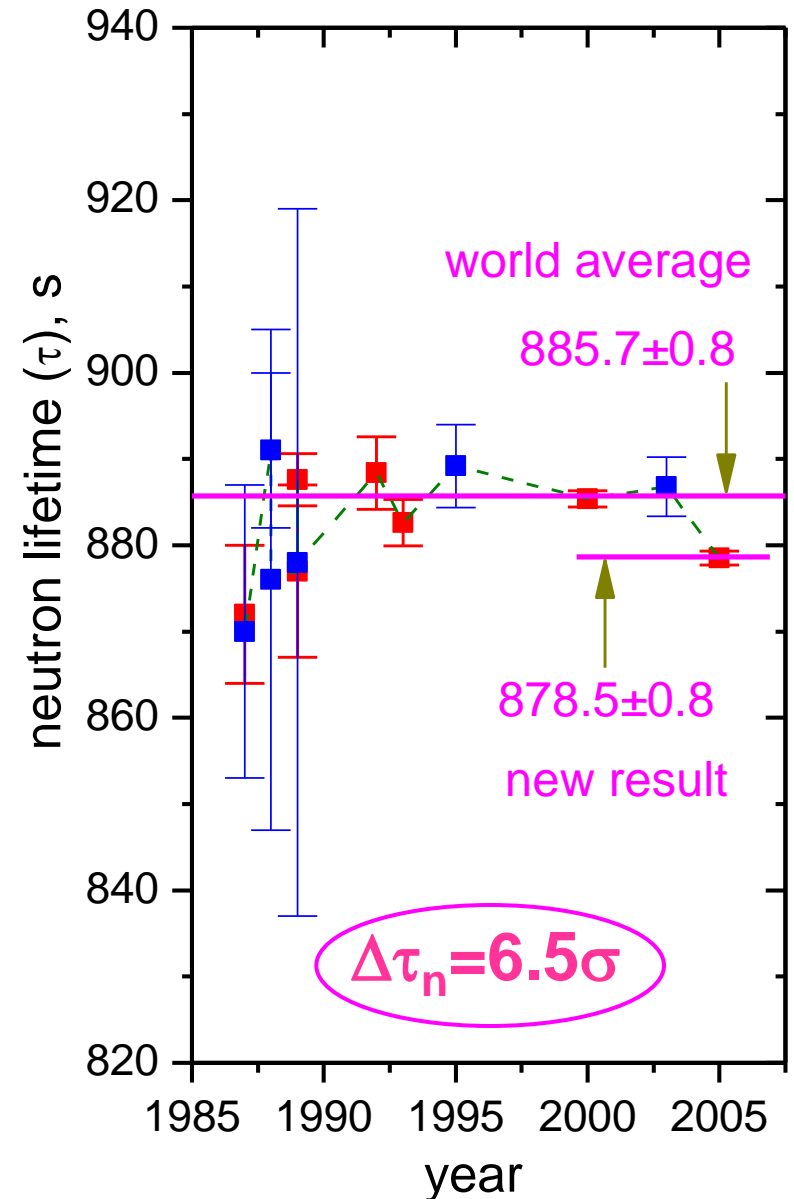
MamBo I

MamBo II



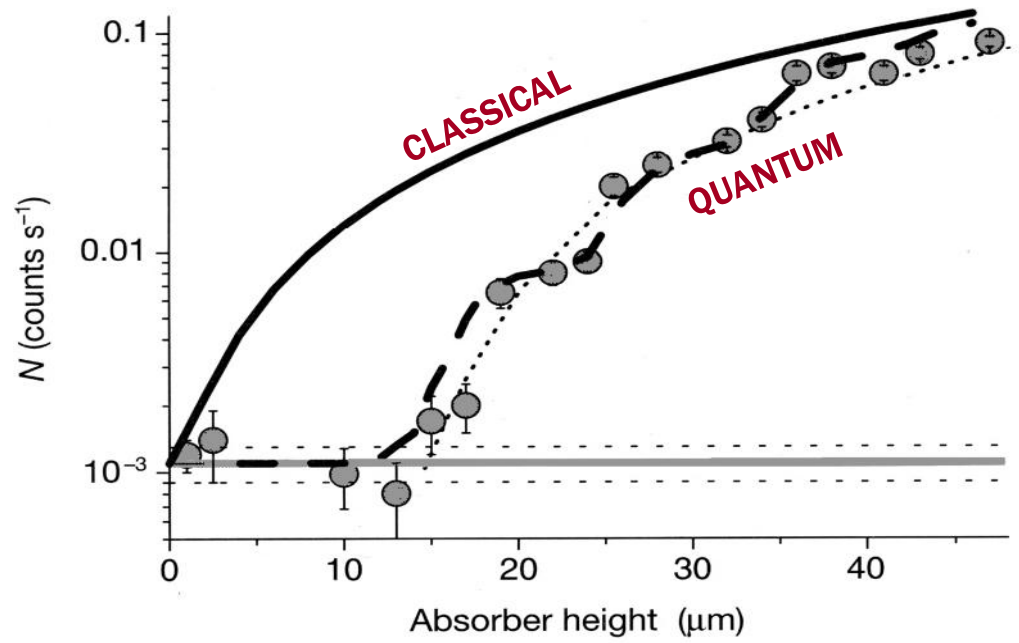
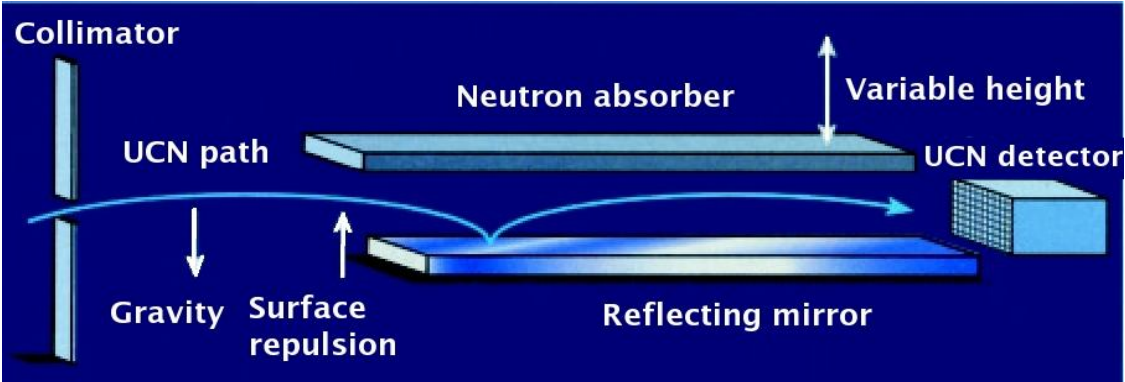
Neutron lifetime: world average and new result

Lifetime τ [s]	Method	Ref./Year
878.5 ± 0.8 $878.5 \pm 0.7_{\text{stat}} \pm 0.3_{\text{sys}}$	Storage of ultra-cold neutrons	Serebrov et al. 2005 PLB 605(2005)72
886.8 ± 3.42	Neutron beam experiment	M.S. Dewey et al. 2003
885.4 ± 0.95 $885.4 \pm 0.9_{\text{stat}} \pm 0.4_{\text{sys}}$	Storage of ultra-cold neutrons	S. Arzumanov et al. 2000 PLB 483(2000)15
889.2 ± 4.8	Neutron beam experiment	J. Byrne et al. 1995
882.6 ± 2.7	Storage of ultra-cold neutrons	W. Mampe et al. 1993
$888.4 \pm 3.1 \pm 1.1$	Storage of ultra-cold neutrons	V. Nesvizhevski et al. 1992
$878 \pm 27 \pm 14$	Neutron beam experiment	R. Kosakowski 1989
887.6 ± 3.0	Storage of ultra-cold neutrons	W. Mampe et al. 1989
877 ± 10	Storage of ultra-cold neutrons	W. Paul et al. 1989
$876 \pm 10 \pm 19$	Neutron beam experiment	J. Last et al. 1988
891 ± 9	Neutron beam experiment	P. Spivac et al. 1988
872 ± 8	Storage of ultra-cold neutrons	A. Serebrov et al. 1987
870 ± 17	Neutron beam experiment	M. Arnold et al. 1987
903 ± 13	Storage of ultra-cold neutrons	Y.Y. Kosvintsev et al. 1986
875 ± 95	Storage of ultra-cold neutrons	Y.Y. Kosvintsev et al. 1980
937 ± 18	Neutron beam experiment	J. Byrne et al. 1980
881 ± 8	Neutron beam experiment	L. Bondarenko et al. 1978
918 ± 14	Neutron beam experiment	C.J. Christensen et al. 1972
885.7 ± 0.8	world average 2004	PDG 2004



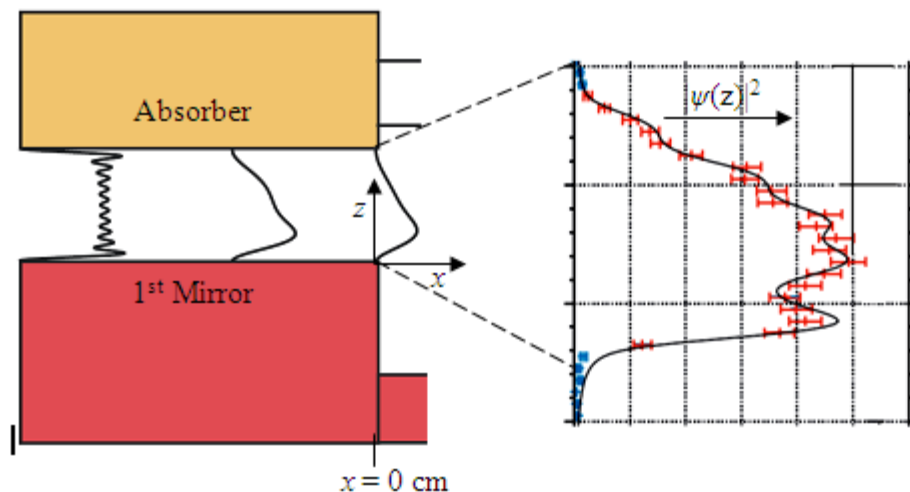
Discovery of neutron quantum states in 1999

Nesvizhevsky *et al*, Nature 415 (2002)

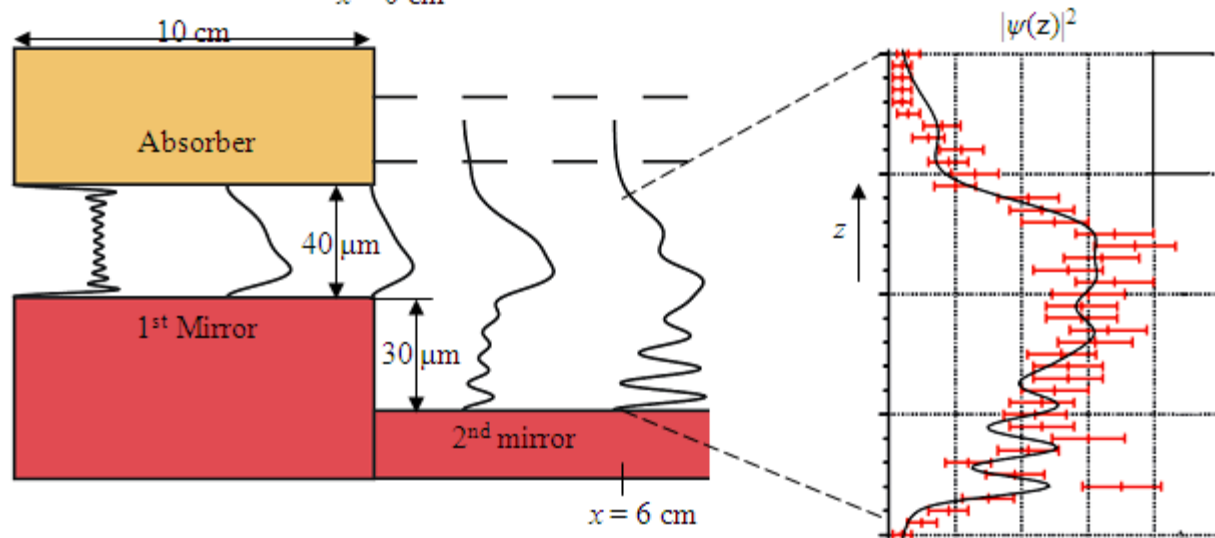


$$z_0 = \left(\frac{\hbar^2}{2m^2g} \right)^{1/3} = 5.87 \mu\text{m}$$

The Quantum Bouncer, Results



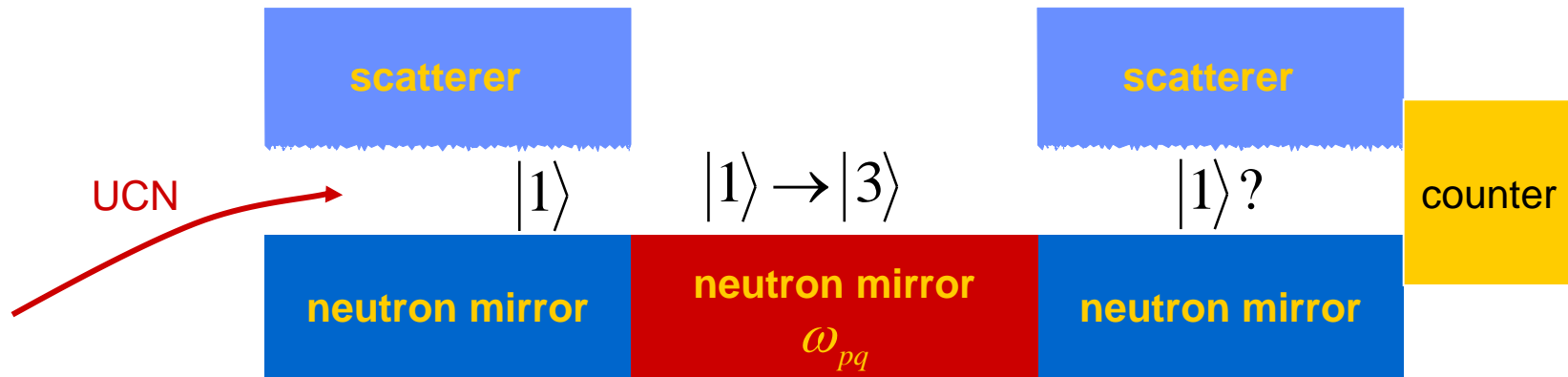
Time evolution of
Coherent Superposition



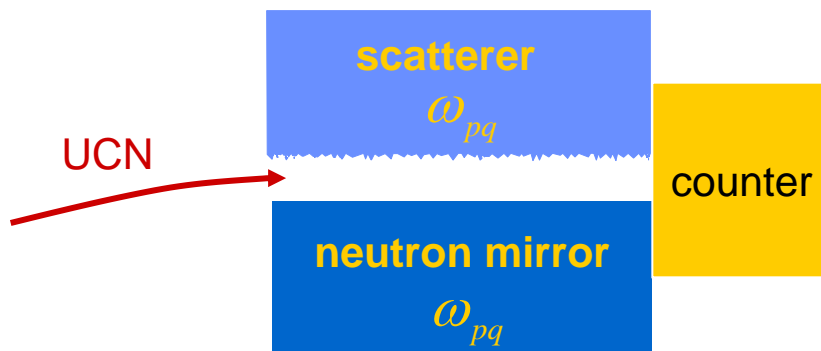
Abele, H., Jenke, T., Stadler, D., & Geltenbort, P. *Nucl. Phys.* **A827**, 593c (2009), see also Dubbers 2010

On the way towards a Resonance Spectroscopy Technique

- First Idea: „Standard“ Rabi Experiment



- Better Idea: Simplified Setup, „Rabi Flopping with Damping“

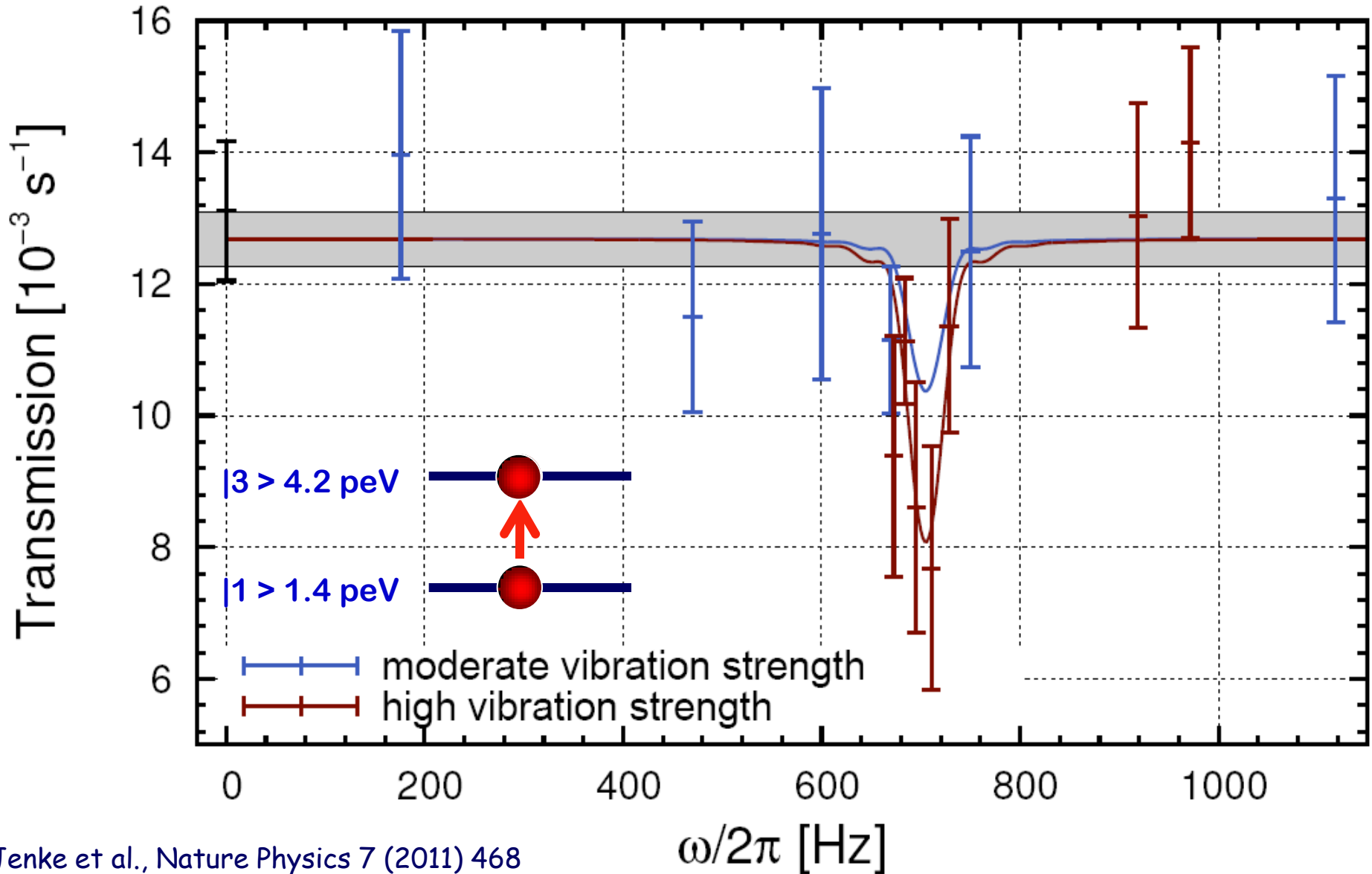


- proof of technique possible
- simple (well-known) setup
- avoids „steps“
- better transmission (short)
- perfectly PF2-compatible

- in principle (probably) limited by knowledge of gap height I (can be removed by measuring > 1 resonance)

T. Jenke, PhD Thesis, 2011

Gravity Resonance Spectroscopy and Excitation



T. Jenke et al., Nature Physics 7 (2011) 468

On the feasibility of using ultracold neutrons to measure the electric charge of the neutron

Yu. V. Borisov, N. V. Borovikova, A. V. Vasil'ev, L. A. Grigor'eva, S. N. Ivanov, N. T. Kashukeev,¹⁾ V. V. Nesvizhevskii, A. P. Serebrov, and P. S. Yaidzhiev²⁾

(Submitted April 22, 1987)

Zh. Tekh. Fiz. 58, 951-958 (May 1988)

A study is made of the feasibility of measuring the electric charge of the neutron with the aid of ultracold neutrons. An experimental apparatus based on the focusing of a beam of ultracold neutrons by means of a cylindrical mirror is described. A trial series of measurements for three days' worth of statistical data gave the result $q = (-4.3 \pm 7.1) \cdot 10^{-20} e$.

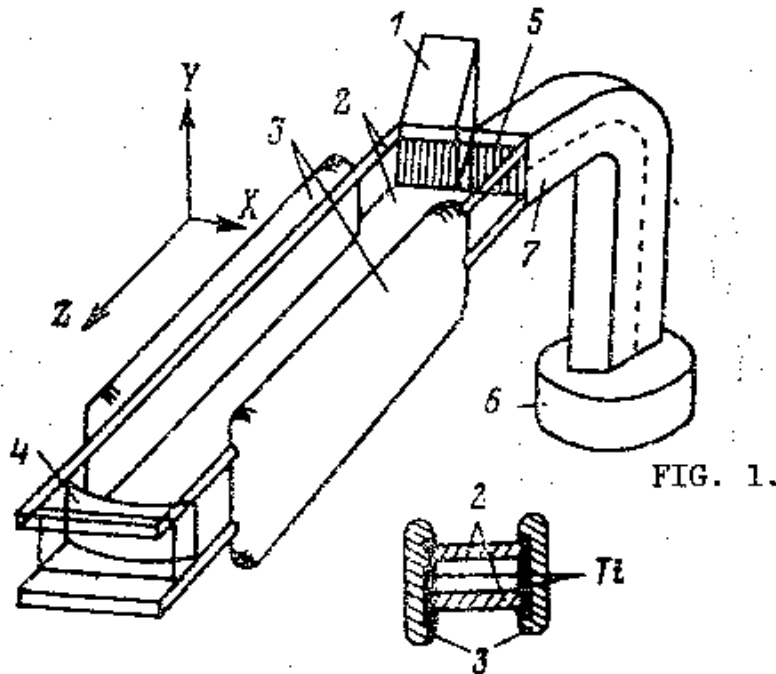


FIG. 1.

I hope I could convince you
that not only ILL and Grenoble are worth a visit
but also that **ultracold neutrons** are
– due to the fact that they are storable –
a fancy and **powerful tool in fundamental physics**



For more information: www.ill.eu
or just call me +33 (0)47620 7242 or geltenbort@ill.fr

Recent review "The neutron and its role in cosmology and particle physics"
by D. Dubbers & M. Schmidt, [arXiv:1105.3694](https://arxiv.org/abs/1105.3694), Rev. Mod. Phys. (in print)



**Thank you, merci beaucoup and besten Dank
for your attention!**