# Neutrino Masses and a Leptogenesis at the TeV scale with a global U(1)

### François-Xavier Josse-Michaux CFTP, IST

in collaboration with E. Molinaro



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$$\Delta m_{21}^2 = 7.59^{+0.20}_{-0.18} \times 10^{-5} \,\mathrm{eV}^2$$
$$\Delta m_{31}^2 = 2.45 \pm 0.09 \times 10^{-3} \,\mathrm{eV}^2$$
$$\Delta m_{31}^2 = -2.34^{-0.10}_{+0.09} \times 10^{-3} \,\mathrm{eV}^2$$
$$\sin(\theta_{12})^2 = 0.312^{+0.017}_{-0.015}$$

 $\sin(\theta_{23})^2 = 0.51 \pm 0.06$  $\sin(\theta_{13})^2 = 0.010^{+0.009}_{-0.006}$ 

 $\sin(\theta_{13})^2 = 0.013^{+0.009}_{-0.007}$ 

Schwetz et al, 1103

Experimentally Challenging Absolute neutrino mass ? Hierarchy ? Nature of neutrinos ?

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Experimentally Challenging  $\Delta m_{21}^2 = 7.59^{+0.20}_{-0.18} \times 10^{-5} \,\mathrm{eV}^2$  $\Delta m_{31}^2 = 2.45 \pm 0.09 \times 10^{-3} \,\mathrm{eV}^2$ Absolute neutrino mass ?  $\Delta m_{31}^2 = -2.34^{-0.10}_{+0.09} \times 10^{-3} \,\mathrm{eV}^2$ Hierarchy ? Nature of neutrinos ?  $\sin(\theta_{12})^2 = 0.312^{+0.017}_{-0.015}$  $\sin(\theta_{23})^2 = 0.51 \pm 0.06$ Theoretically Challenging  $\sin(\theta_{13})^2 = 0.010^{+0.009}_{-0.006}$ SM-like Dirac-type mass:  $y_{\nu} \bar{\ell} H \nu_R \qquad y_{\nu} \lesssim 10^{-13}$ Minkowski; Gell-Mann, Ramond&Slansky; type I seesaw Yanagida; Mohapatra&Senjanovic; GUT-liked mass: new fermions type III seesaw Foot, Lew, He&Joshi; Ma; Magg&Wetterich; Lazarides, Shafi new scalars type II seesaw &Wetterich; Mohapatra&Senjanovic; Typically very heavy  $\Lambda \gg \text{TeV}$  and impossible to probe experimentally

Low-energy (~TeV	scale) models		
In General:	Suppressed Yukawas	hard to probe	
«Tuned» models:	Large Yukawas +Cancellations	suffer destabilizing radiative corrections	Aristizabal Sierra &Yaguna '11
«Stabilized» mode	els: Large Yukawas +Symmetry	mass is protected	

The symmetry can be explicitely broken by small parameters eg: Kersten&Smirnov '07 or conserved above the EWSB Branco, Grimus&Laboura '89

Mohapatra&Valle, '8

We study the UV-completion of the Inverse-Seesaw model imposing a conserved global U(1)



Perturbate the zeros: add scalar representations

$$M_{\nu} = \begin{pmatrix} 0 & m_{D_{1}}^{T} & M \\ m_{D_{1}} & M & M \end{pmatrix} -\mathcal{L}_{eff} \supset -\frac{y_{1}^{i}y_{2}^{j} + y_{1}^{j}y_{2}^{i}}{2M} \left(\overline{\ell}_{j}^{c}.\widetilde{H}_{2}^{*}\right) \left(\widetilde{H}_{1}^{\dagger}.\ell_{i}\right) \\ +\frac{y_{1}^{i}y_{1}^{i}\alpha^{*}}{\sqrt{2}M^{2}} \left(\overline{\ell}_{j}^{c}.\widetilde{H}_{1}^{*}\right) \left(\widetilde{H}_{1}^{\dagger}.\ell_{i}\right) H_{3}^{*} \\ +\frac{y_{2}^{i}y_{2}^{j}\alpha}{\sqrt{2}M^{2}} \left(\overline{\ell}_{\beta}^{c}.\widetilde{H}_{2}^{*}\right) \left(\widetilde{H}_{2}^{\dagger}.\ell_{\alpha}\right) H_{3} + h.c.$$

$$(m_{\nu})_{ij} = -\left(y_1^i y_2^j + y_2^i y_1^j - y_1^i y_1^j \alpha^* \frac{v_1 v_3}{v_2 M} - y_2^i y_2^j \alpha \frac{v_2 v_3}{v_1 M}\right) \frac{v_1 v_2}{2M} \neq 0$$

UV-completion of the Inverse-Seesaw:

Add 3 RH neutrinos + 1 Scalar doublet H2 + 1 singlet H3

Gain: 2 massive neutrinos

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+ 1 singlet S: L(S)=1

Gain: 2 massive neutrinos + Leptogenesis + Dark Matter

### **2-Steps Leptogenesis**







WMAP7 EPS-HEP 2011, Grenoble

### **2-Steps Leptogenesis**

#### Relevant processes: $\mathcal{O}(g^2), \mathcal{O}(g^2 \alpha^2), \mathcal{O}(g^4), \mathcal{O}(y^2), \mathcal{O}(g^2 y^2)$

 $\Delta N_D = 2$  scatterings

Decays / Inverse decays  $[N_3 \leftrightarrow N_D S^*]$ 

1st Step

 $\begin{bmatrix} N_D S^* \leftrightarrow \overline{N_D} S \end{bmatrix} \quad \begin{bmatrix} N_D N_D \leftrightarrow S S \end{bmatrix}$ 

Scatterings on N3: CP violation is included

 $[N_3 N_D \leftrightarrow S^* H_3] \quad [N_3 S \leftrightarrow N_D H_3^*] \quad [N_D S \leftrightarrow N_3 H_3]$ 

2nd Step

Decays / Inverse decays of ND

 $\begin{bmatrix} N_D \leftrightarrow \ell H_1 \end{bmatrix} \begin{bmatrix} N_D \leftrightarrow \overline{\ell} H_2^* \end{bmatrix}$ 

Scatterings on top quarks  $\left[N_D \overline{\ell} \leftrightarrow \overline{t} Q_3\right] \left[N_D q \leftrightarrow \ell q\right]$ 

Scatterings on N3  $\begin{bmatrix} N_3 \ S \leftrightarrow \ell \ H_1 \end{bmatrix} \quad \begin{bmatrix} N_3 \ S \leftrightarrow \overline{\ell} \ H_2^* \end{bmatrix}$ 

#### **Boltzmann Equations**

$$Y_{N_3}' = (1 - y_{N_3}) \left( \gamma_D + 2 \sum_{k=a,...,e} \gamma_{N_3}^k \right)$$

 $y_{\Delta N_D}{}' = -\epsilon_{CP} \gamma_D (1 - y_{N_3}) - 2 \epsilon_{CP}^a \gamma_{N_3}^a (1 - y_{N_3}) + 2 \epsilon_{CP}^b \gamma_{N_3}^b (1 - y_{N_3}) + 2 \epsilon_{CP}^c \gamma_{N_3}^c (1 - y_{N_3})$  $-2 \left(\gamma_{\Delta 2}^a + 2\gamma_{\Delta 2}^b\right) \left(y_{\Delta N_D} - y_{\Delta S}\right) - \gamma_{D\ell} \left(-y_{\Delta H_1} - y_{\Delta \ell} + y_{\Delta N_D}\right) - \gamma_{D\overline{\ell}} \left(y_{\Delta H_2} + y_{\Delta \ell} + y_{\Delta N_D}\right)$  $- \left(\gamma_{N_D}^s + 2\gamma_{N_3}^t\right) \left(y_{\Delta N_D} - y_{\Delta \ell}\right) + \gamma_{N_3}^a \left(y_{\Delta H_3} - y_{\Delta S} - y_{N_3} y_{\Delta N_D}\right) + \gamma_{N_3}^b \left(y_{\Delta H_3} - y_{\Delta N_D} - y_{N_3} y_{\Delta S}\right)$  $+ \gamma_{N_3}^c \left(y_{N_3} y_{\Delta H_3} - y_{\Delta N_D} - y_{\Delta S}\right)$ 

$$Y_{\Delta S}' = \epsilon_{CP} \gamma_D (1 - y_{N_3}) - 2 \epsilon_{CP}^a \gamma_{N_3}^a (1 - y_{N_3}) + 2 \epsilon_{CP}^b \gamma_{N_3}^b (1 - y_{N_3}) + 2 \epsilon_{CP}^c \gamma_{N_3}^c (1 - y_{N_3}) - 2 \epsilon_{CP} \gamma_{N_3}^d (1 - y_{N_3}) - 2 \epsilon_{CP} \gamma_{N_3}^e (1 - y_{N_3}) + 2 \left(\gamma_{\Delta 2}^a + 2\gamma_{\Delta 2}^b\right) \left(y_{\Delta N_D} - y_{\Delta S}\right) + 2 \gamma_{SS} \left(y_{\Delta H_1} - y_{\Delta H_2} - 2y_{\Delta S}\right) + \gamma_{N_3}^a \left(y_{\Delta H_3} - y_{\Delta S} - y_{N_3}y_{\Delta N_D}\right) + \gamma_{N_3}^b \left(y_{\Delta H_3} - y_{\Delta N_D} - y_{N_3}y_{\Delta S}\right) + \gamma_{N_3}^c \left(y_{N_3}y_{\Delta H_3} - y_{\Delta N_D} - y_{\Delta S}\right) + \gamma_{N_3}^d \left(y_{\Delta H_1} + y_{\Delta \ell} - y_{N_3}y_{\Delta S}\right) - \gamma_{N_3}^e \left(y_{\Delta H_2} + y_{\Delta \ell} + y_{N_3}y_{\Delta S}\right)$$

$$Y_{\Delta\ell}' = 2 \epsilon_{CP} \gamma_{N_3}^d (1 - y_{N_3}) - 2 \epsilon_{CP} \gamma_{N_3}^e (1 - y_{N_3}) + \left(\gamma_{N_D}^s + 2\gamma_{N_3}^t\right) \left(y_{\Delta N_D} - y_{\Delta\ell}\right)$$
$$-\gamma_{D\ell} \left(y_{\Delta H_1} + y_{\Delta\ell} - y_{\Delta N_D}\right) - \gamma_{D\bar{\ell}} \left(y_{\Delta H_2} + y_{\Delta\ell} + y_{\Delta N_D}\right) - \gamma_{\Delta L=2} \left(y_{\Delta H_1} + y_{\Delta H_2} + 2y_{\Delta\ell}\right)$$
$$-\gamma_{N_3}^d \left(y_{\Delta H_1} + y_{\Delta\ell} - y_{N_3}y_{\Delta S}\right) - \gamma_{N_3}^e \left(y_{\Delta H_2} + y_{\Delta\ell} + y_{N_3}y_{\Delta S}\right)$$

$$Y_{\Delta H_1}' = \dots \quad Y_{\Delta H_2}' = \dots \quad Y_{\Delta H_3}' \quad \dots$$

#### **2-Steps Leptogenesis**

$$-\mathcal{L}_{\text{int}} \supset \frac{1}{2} M_3 \overline{N}_3 N_3^c + \left( g \overline{N}_3 N_D S^* + \frac{\alpha}{\sqrt{2}} H_3 \overline{N}_D N_D^c - \frac{\mu''}{\sqrt{2}} S^2 H_3^* + h.c. \right) + M \overline{N}_D N_D + \left( y_1^i \overline{N}_D \tilde{H}_1^\dagger \ell_i + y_2^j \overline{N}_D^c \tilde{H}_2^\dagger \ell_j + h.c. \right)$$

CP asymmetry in N3 decays: 
$$\epsilon_{CP} \simeq \frac{\text{Im}(\alpha)}{16 \pi} \frac{\mu''}{M_3}$$

ND asymmetry@ 1st step: 
$$Y^{1st}_{\Delta N_D} \propto \epsilon_{CP} \times \eta_1(g)$$

L asymmetry@ 2nd step:  $Y_{\Delta L}^{1st} \propto Y_{\Delta N_D}^{1st} \times \eta_2(y_1, y_2) \sim Y_{\Delta N_D}^{1st}$ 

Final Baryon asymmetry:

$$Y_{\Delta B} \propto \epsilon_{CP} \eta_1(g) \eta_2(y_1, y_2)$$

### Successful Leptogenesis



or the depletion processes are too big

#### **Dark Matter**



### Conclusion

We construct a UV-completion of the Inverse-Seesaw imposing a conserved global Lepton number

1 scalar doublet & 1 scalar singlet added (+RHNs)

We add a scalar singlet

a new-scenario of Leptogenesis emerges

a viable Dark Matter candidate (S) exists

As long as N is charged under a conserved (lepton) number, any asymmetry in N transfers to the lepton sector through the neutrino Yukawa couplings

#### Addendum: Model Lagrangian

Lagrangian: 
$$\mathcal{L}_{\text{int}} = \mathcal{L}_{\text{int}}^{\text{SM}} - \mathcal{V}_{\text{Sca}} - M \overline{N}_D N_D - \frac{1}{2} M_3 \overline{N}_3 N_3$$
$$- \left( y_1^i \overline{N}_D \widetilde{H}_1^{\dagger} \ell_i + y_2^j \overline{N}_D^c \widetilde{H}_2^{\dagger} \ell_j + \frac{\alpha}{\sqrt{2}} H_3 \overline{N}_D N_D^c + g S \overline{N}_D N_3^c + \text{h.c.} \right),$$

#### Scalar Potential:

$$\begin{split} \mathcal{V}_{\text{Sca}} &= -\mu_1^2 \, H_1^{\dagger} \, H_1 + \lambda_1 \, (H_1^{\dagger} \, H_1)^2 - \mu_2^2 \, H_2^{\dagger} \, H_2 + \lambda_2 \, (H_2^{\dagger} \, H_2)^2 - \mu_3^2 \, H_3^* H_3 + \lambda_3 \, (H_3^* H_3)^2 \\ &+ \kappa_{12} \, H_1^{\dagger} \, H_1 H_2^{\dagger} \, H_2 + \kappa_{12}' \, H_1^{\dagger} \, H_2 H_2^{\dagger} \, H_1 + \kappa_{13} \, H_1^{\dagger} \, H_1 H_3^* H_3 + \kappa_{23} \, H_2^{\dagger} \, H_2 H_3^* H_3 \\ &+ \mu_S^2 \, S^* S + \lambda_S \, (S^* S)^2 + \mathcal{F}_1 \, H_1^{\dagger} \, H_1 S^* S + \mathcal{F}_2 \, H_2^{\dagger} \, H_2 S^* S + \mathcal{F}_3 \, H_3^* H_3 S^* S \\ &+ h \, S^2 H_1^{\dagger} \, H_2 + h^* \, S^{*2} H_2^{\dagger} \, H_1 - \frac{\mu'}{\sqrt{2}} \, \left( H_1^{\dagger} \, H_2 H_3 + H_2^{\dagger} \, H_1 H_3^* \right) \, - \frac{\mu''}{\sqrt{2}} (S^2 H_3^* + S^{*2} H_3) \, . \end{split}$$

### **Addendum: Higgs Sector**



Lightest CPE h0 has suppressed couplings to SM

SM-like H0 has additionnal decay channels:

HO-> J J is important before HO->WW channel opens

