The Alpha Magnetic Spectrometer (AMS) Experiment

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on behalf of the AMS International Collaboration 16 Countries, 60 Institutes and 600 Physicists, 17 years



over the world and integrated in CERN



What is AMS?

Magnetic spectrometer conceived for the high precision study of cosmic radiation in the GV-TV rigidity range:

✓ Search for primordial antimatter (anti-nuclei)

✓ Indirect Search for Dark Matter (light anti-matter, gammas)

✓ Chemical composition and energy spectra of cosmic rays

✓ Exotic signals? (strangelets...)

Steadily taking data on the ISS since May 19th 2011

How it will fullfill its objectives?

-Same precision and detection capability as the large state-of-the-art particle physics Detectors (but fitting into the space shuttle...)

- Operation in space: on the ISS, at 400km, no backgrounds from atmospheric interactions (extensive space qualification tests...)

- Collection power:

large acceptance (≈ 0.5 m2sr)

x exposure time (= ISS lifetime) (extensive calibration campaigns on ground)



IMPLEMENTING ARRANGEMENT BETWEEN THE DEPARTMENT OF ENERGY AND THE NATIONAL AERONAUTICS AND SPACE ADMINISTRATION REGARDING THE ALPHA MAGNETIC SPECTROMETER IN SPACE PROGRAM

I. PROGRAM DESCRIPTION

The AMS is a state-of-the-art particle physics detector containing a large permanent magnet that will be designed constructed and tested by an international team organized under DOE sponsorship and that will use the unique environment of space to advance knowledge of the universe and lead potentially to a clearer understanding of the origin of the universe. Specifically, the science objectives of the AMS are to search for cosmic sources of antimatter (i.e., anti-helium or heavier elements) and dark matter.

... On the second

flight, NASA will launch the AMS on the Shuttle and transfer and install it onto the International Space Station (ISS). The AMS then will be operated as an externally attached payload on the ISS for a nominal three-year period, after which NASA will detach the AMS from the ISS, transfer it to a Space Shuttle, and return it to Earth. ...

Associate Administrator for Life and Microgravity Sciences and Applications National Aeronautics and Space Administration

Office of Energy Research Department of Energy

Date: _ 20 SENT Date:



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Heads of Agency International Space Station Joint Statement

TOKYO -- The heads of the International Space Station (ISS) agencies from Canada, Europe, Japan, Russia, and the United States met in Tokyo, Japan, on March 11, 2010, to review ISS cooperation.

With the assembly of the ISS nearing completion and the capability to support a full-time crew of six established, they noted the outstanding opportunities now offered by the ISS for on-orbit research and for discovery including the operation and management of the world's largest international space complex. In particular, they noted the unprecedented opportunities that enhanced use of this unique facility provides to drive advanced science and technology. This research will deliver benefits to humanity on Earth while preparing the way for future exploration activities beyond low-Earth orbit. The ISS will also allow the partnership to experiment with more integrated international operations and research, paving the way for enhanced collaboration on future international missions.

The heads of agency reaffirmed the importance of full exploitation of the station's scientific, engineering, utilization, and education potential. They noted that there are no identified technical constraints to continuing ISS operations beyond the current planning horizon of 2015 to at least 2020, and that the partnership is currently working to certify on-orbit elements through 2028. The heads of agency expressed their strong mutual interest in continuing operations and utilization for as long as the benefits of ISS exploitation are demonstrated. They acknowledged that a U.S. fiscal year 2011 budget consistent with the U.S. administration's budget request would allow the United States to support the continuation of ISS operations and utilization activities to at least 2020. They emphasized their common intent to undertake the necessary procedures within their respective governments to reach consensus later this year on the continuation of the ISS to the next decade.

In looking ahead, the heads of agency discussed the importance of increasing ISS utilization and operational efficiency by all possible means, including finding and coordinating efficiencies across the ISS Program and assuring the most effective use of essential capabilities, such as space transportation for crew and cargo, for the life of the program.

For the latest about the International Space Station, visit the Internet at:

http://www.nasa.gov/station

The AMS-01 Detector (1994-1998)

6 planes of Silicon Tracker: 3.2 % X_0 , 10 μm. BL² = 0.14 Tm², ΔP/P = 7% at 10 GeV



STS91, June 2nd - 12th 1998

- 10 days of data taking in orbit:
 - 400 Km altitude
 - latitudes <51.7°
 - all longitudes
- 10⁸ events recorded
- Physics results (Phys. Rep. 366 (2002) 331)
 - precise measurements of primary fluxes
 - detection of secondary fluxes (quasi trapped)
 - antimatter limit at 10⁻⁶
 - light nuclei (B/C,isotopes)

ApJ, 724:329–340, 2010 November 20 ApJ, 736:105 (11pp) 2011 August 1



The Detector

Timeline

2000-2008 : Sub-detector Construction and first mechanical integration **2009: Integration with the SC Magnet** 2010: Beam test EMI – TVT test **Integration with the Permanent Magnet Beam test** Shipment to KSC Integration with the Shuttle and ISS interfaces **2011: Installation on the Shuttle (March)** Launch (May 16) Installation on the ISS (May 19)



AMS: A TeV precision, multipurpose particle physics TRD spectrometer in space.





The Magnet







6400 high-grade Nd–Fe–B blocks



Transition Radiation Detector (TRD): identifies Positron and Electron







Time of Flight (TOF) TOF HHAPpappapt **Provides trigger for** charged particles ALL CONCERNMENT **Trigger time is** synchronized to UTC time to 1µs **Measures the time** of relativistic particles TOF to 160 picoseconds /<mark>/∖. ∆t/t=160ps</mark> 3000 2500 TOF Time 2000 1500 1000

500

0

-0.5

Ò (T1-T3)-(T2-T4) (ns)

0.5

Veto System rejects random cosmic rays



Measured veto efficiency better than 0.99999

The coordinate resolution is 10 micron

Silicon Tracker

- 9 Layers of silicon microstrip detectors for an active area of 6.4m², 3 μm mechanical alignment.
- 200k channels for 129 W of power
- high dynamic range front end for charge measurement
- wide temperature range (-20/+40 s)

50,000 fibers, $\phi = 1$ mm, distributed uniformly inside 1,200 lb of lead which provides a precision, 3-dimensional, $17X_0$ measurement of the directions and energies of light rays and electrons up to 1 TeV

The completed flight electronics (650 microprocessors, 300,000 channels)

High Energy Physics Electronics and Detectors to Low Earth Orbit

Redundancy in Design

2009: AFTER 9000 hrs of TVT...THE END OF SUB-SYSTEM TESTS

THE END

r 02 60:

AMS in the ESA Anecoic chamber, Noordwijk, the Netherlands

AMS in the ESA Thermal Vacuum Chamber, Noordwijk, the Netherlands

Consumables lifetime – TRD Leak rate

Time [month]

Caused by CO_2 Diffusion CO_2 Storage at Launch: 5 kg Leakrate of 5 µg / s corresponds to a

TRD Lifetime of 30 Years

Test at CERN AMS in accelerator test beam Feb 4-8 and Aug 8-20, 2010

CERN Accelerator Complex

Beam test calibration

The (long) way to space

Loading of AMS into a US Air Force C5-M at Geneva Airport – 25 Aug 2010

AMS C5 landing on the Shuttle runway at Kennedy Space Center – 26 Aug 2010

Calibration with muons

Tracker : Rigidity resolution

AMS Monitoring Interface - Mozilla Firefox

A 2.5 h after launch..

Everything OK

Google

INFN

http://pcposj0.lecal.8081/welcome/show/GTSN/0x28/3 hours

SPD2 @ TSPD4

SPD3@TSPD6

S2

53

SHV2

13.625 °C

14.0625 °C

13.3125 °C

13.875 °C

14.3175 'C

Until now or 16:35 16/05/2011 CHECK MPD @ TMPD2 13.875 'C М 11.9375 °C GPS 12.5 'C Π 14.0625 °C TTCBP 16.0625 'C TTCBS 16.0625 'C UGPD 13.75 °C UG 12.5 °C CCEB Signal Side 13.625 °C CCEB Power Side 13.5625 'C UPDO 13.6875 'C UO 12.1875 °C UPD1 13.8125 °C U1 11.625 'C SPD0@TSPD1 13.6875 'C SO 11.9375 °C SHVO 13.8125 °C SPD1@TSPD3 13.5625 °C S1 12.0 °C 13.0625 °C

GTSN

JPD-A

JPD-B

May 19: AMS installed on ISS 5:15 CDT, start taking data 9:35 CDT

One of the Firsts AMS-02 Event in Space as seen in Houston

Experience in flight

ISS orbit ≈ 390 km, 51.7°, 90 minutes

Science Data Flow

AMS POCC at CERN for the lifetime of the ISS

Nov. 2010

Collected and Reconstructed events 2 months ≈ 2.5 10⁹ events

In flight experience: Tracker cooling and currents

TRACKER THERMAL STABILITY : OK

TRD Temperature Monitoring on ISS, to be within ±2C°

AMS data on ISS Electron 240 GeV, 22 May

Calorimeter (ECAL) 3D Sampling of Showers

18 longitudinal samplings (9 groups of 2 layers) Readout cell dimensions: Z (9 mm), X (9 mm), Y (9 mm)

AMS data on ISS

Photon 40 GeV, 23 May

The Cosmos is the Ultimate Laboratory.

The issues of antimatter in the universe and the origin of Dark Matter probe the foundations of modern physics.

AMS is the only large scientific experiment to study these issues in space.

Backup slides

