

Light Neutralinos in the NMSSM

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The NMSSM

- Minimal Supersymmetric SM + SU(2) singlet in the Higgs sector:

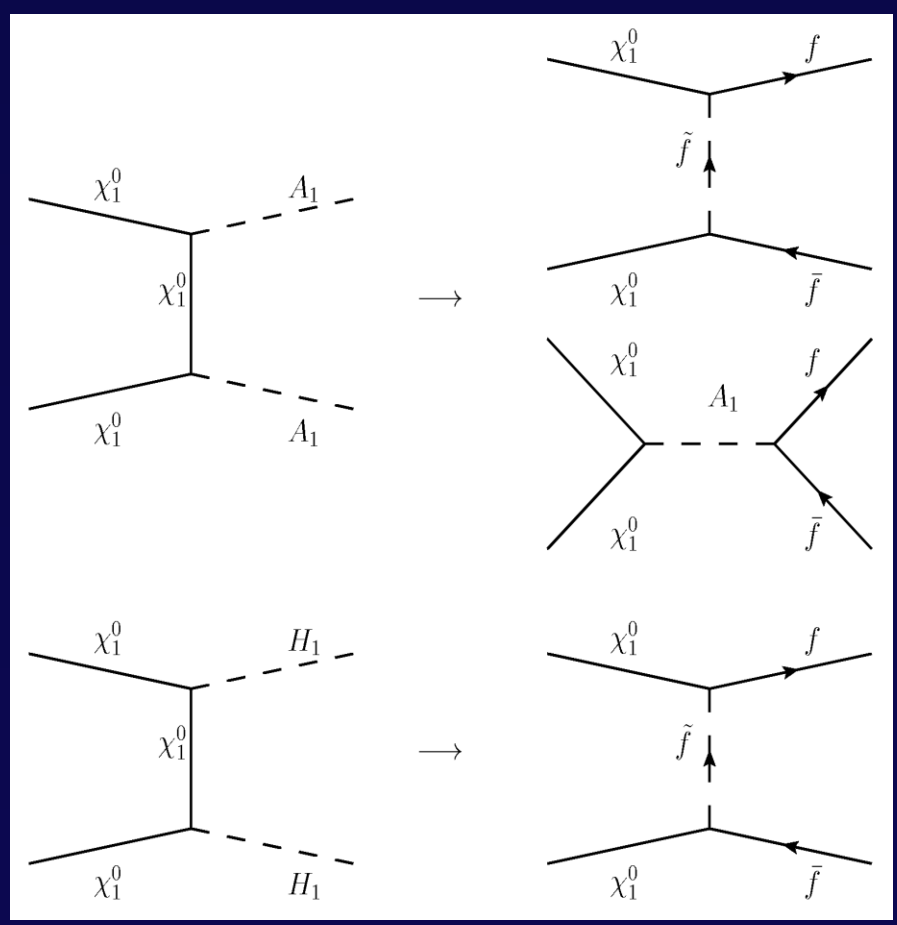
$$\mu H_1 \cdot H_2 \longrightarrow \lambda S H_1 \cdot H_2 + \frac{1}{3} \kappa S^3$$

- Enhanced Higgs spectrum: 3 CP-even, 2 CP-odd neutral bosons.
- Singlet-like lightest Higgses could be almost decoupled from the SM! It is possible to have $M_H \ll 100$ GeV and/or $M_A \ll 100$ GeV.
- 5-state mixed neutralino lightest stable particle (LSP). The mass of the LSP is a mixture of the bino (M_1), wino (M_2), higgsino and singlino masses:

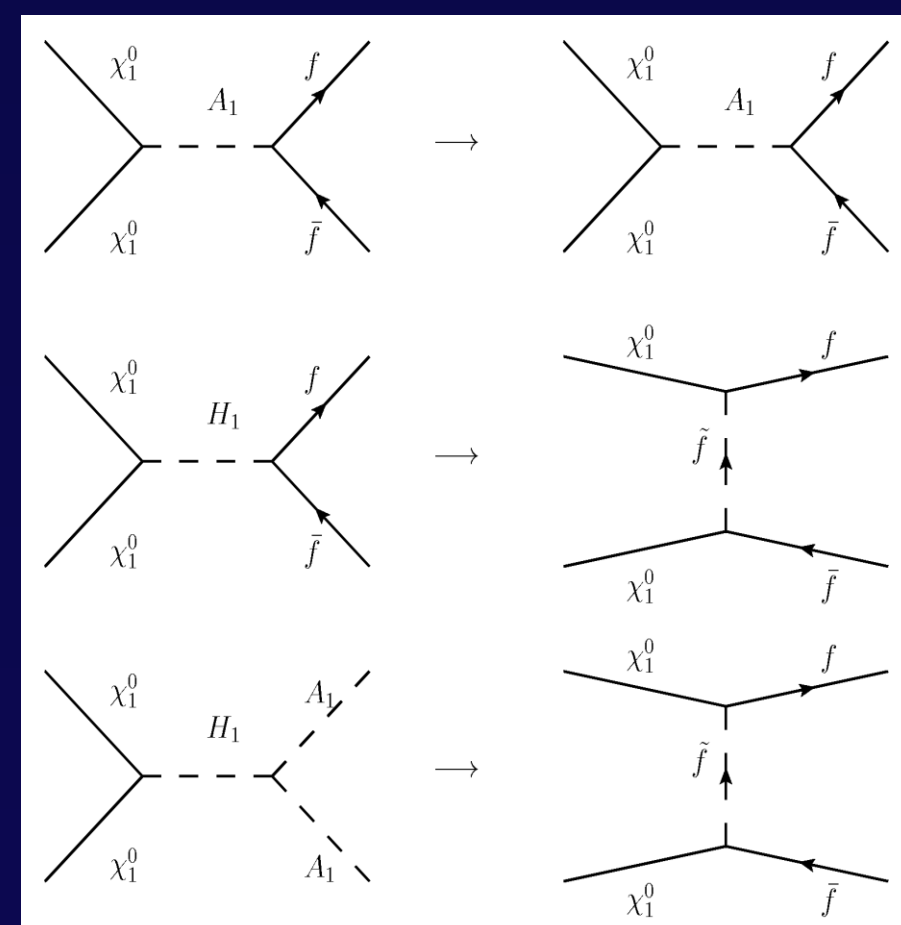
$$\chi_1^0 = N_{11} \tilde{B} + N_{12} \tilde{W}_3^0 + N_{13} \tilde{H}_d + N_{14} \tilde{H}_u + N_{15} \tilde{S}$$

Neutralino annihilations

- Different behavior at freeze-out ($v \approx 0.4 \times c$) and at galactic annihilations ($v \approx (10^{-4} - 10^{-3}) \times c$):



Non-resonant channels



Resonant channels

- Whenever the dominant channel is not the same at Early Universe and at quasi-rest there, there is a large suppression in the interaction rate: no indirect signals. The A-resonance, though, is not p-wave suppressed: a boost is obtained at low velocities, we expect large indirect signals.
- Final states: fermions of the Standard Model and light Higgs bosons (only in the Early Universe)

Parameter boundaries

(GeV units)

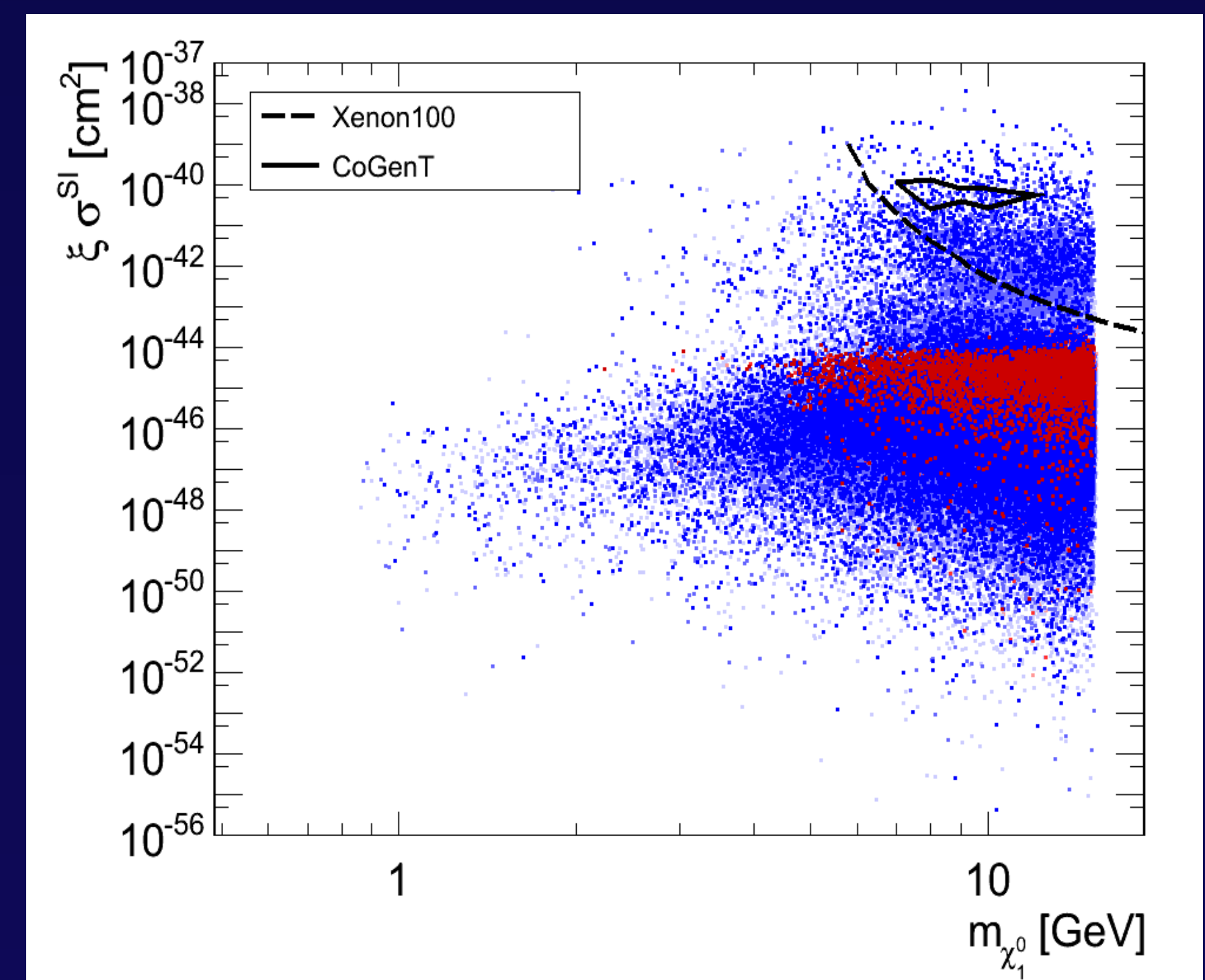
$$\begin{aligned} 1 < M_1 < 100 \\ 100 < M_2 < 2000 \\ 0 < \mu < 1000 \\ 1 < \tan \beta < 75 \\ 0 < \lambda < 0.75 \\ 0 < \kappa < 0.65 \\ -2000 < A_\lambda < 5000 \\ -5000 < A_\kappa < 2000 \\ -3000 < A_t < 3000 \\ 100 < M_{\tilde{t}} < 2000 \\ 300 < M_{\tilde{q}} < 2000 \end{aligned}$$

Scanning Parameter Space

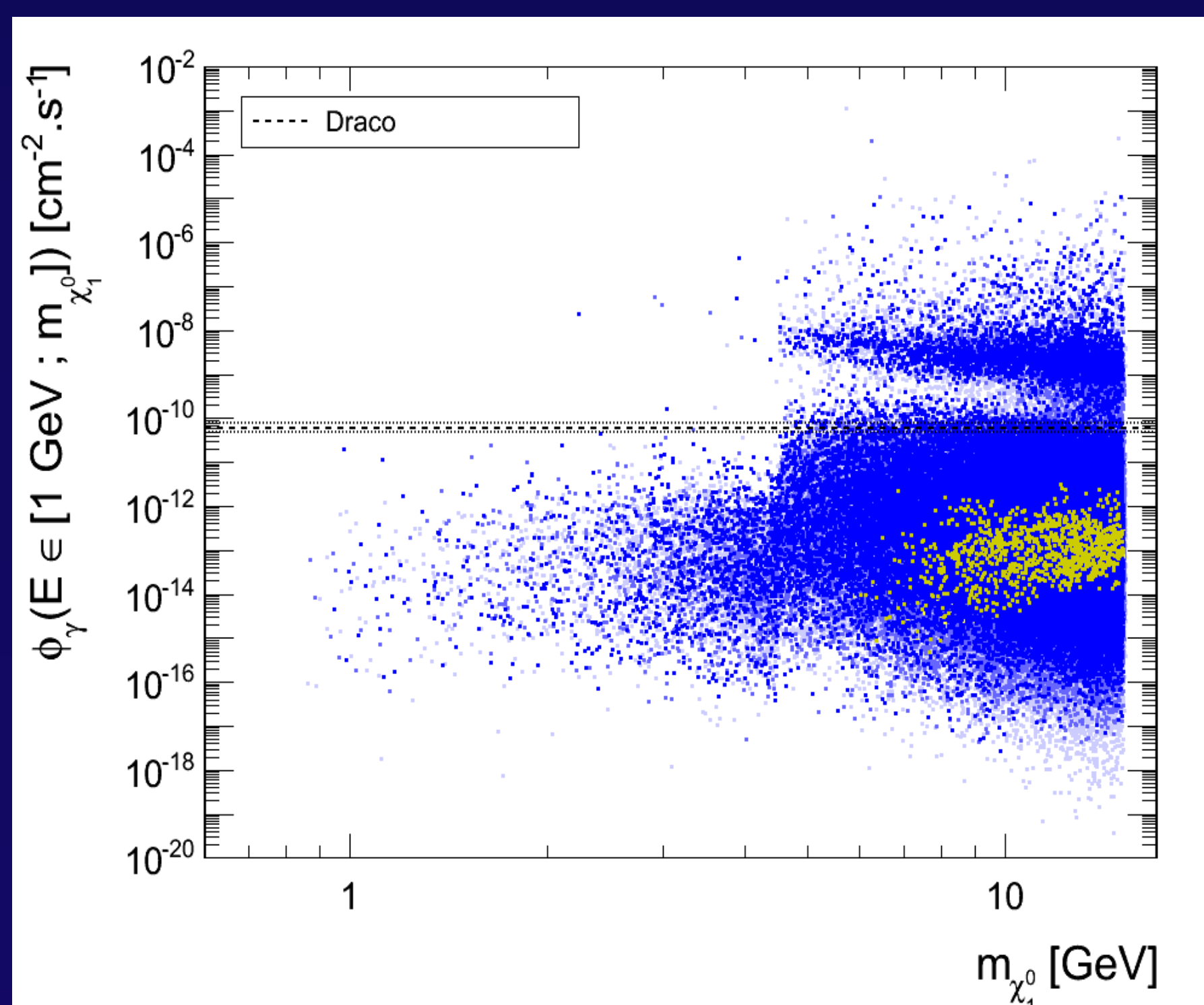
- Tools: micrOMEGAs 2.4 + NMSSMTools
- Large parameter space issue. Markov Chain Monte-Carlo scans: a random walk approach
- Points discriminated by priors (phenomenological framework) and likelihoods (fitting available data)
- Priors:
 - Parameter space boundaries
 - Neutralino mass below 15 GeV
 - Physical supersymmetric configuration
- Likelihoods:
 - Relic Density: at least 10% of the cosmological Dark Matter
 - Mass limits (chargino, sleptons...)
 - Higgs phenomenology
 - Muon anomalous magnetic moment
 - Rare B-meson decays and oscillations
 - Z invisible width

Direct Detection

- Neutralinos in CoGenT (and DAMA) regions
- ...or neutralinos excluded by Xenon100
- Spin independent interactions: t-channel Higgs boson or s and u-channel squark exchanges
- A exchanges are chirally suppressed at $v \rightarrow 0$, while squarks are too heavy
- Large interactions: only possible via a light CP-even (singlet-like) Higgs: M_H^{-4} enhancement
- In red: excluded by Fermi-LAT limits on γ -rays (see Indirect Detection)



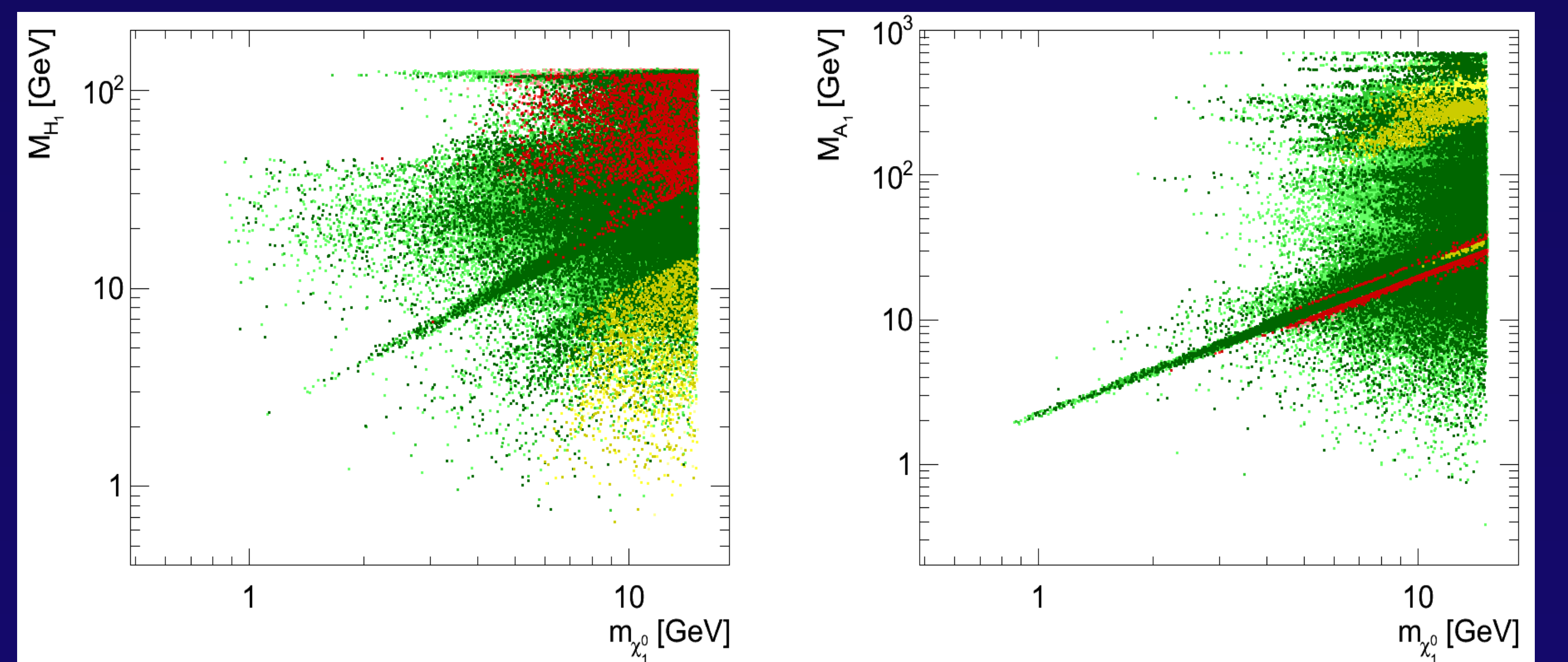
Indirect Detection



- Fermi-LAT null results \rightarrow stringent limits on the γ -ray flux from Dark Matter annihilations from Draco dwarf spheroidal galaxy
- γ -rays from Dark Matter annihilations: decays of fermionic final states
- Large annihilations for light CP-odd (singlet-like) Higgs
- Excluded points have annihilation rates $\sigma v \gg 10^{-26} \text{ cm}^3 \text{ s}^{-1}$ (the canonical freeze out rate) at $v \rightarrow 0$: A-pseudoscalar boson resonance
- Other checks (not further constraining):
 - antiprotons and positrons
 - radio light from the galactic center and clusters
- In yellow: points excluded by Xenon100 (see Direct Detection)

Light Higgs configurations

- Relic Density achieved by annihilation through or into light Higgs bosons (see the concentration of configurations towards the $y=2x$ line in both panels suggesting resonances)
- In red: excluded by Fermi-LAT limits on γ -rays (see Indirect Detection), in yellow: points excluded by Xenon100 (see Direct Detection) and in green: safe points.
- Yellow points in left panel suggest that the configurations yielding large SI interactions annihilate into H_1 pairs in Early Universe and thus have suppressed galactic annihilations
- Red points in the right panel are very correlated to the resonance: Indirect signals depend upon the fine-tuning of the resonance and neutralino-neutralino-pseudoscalar couplings



References: [micrOMEGAs: arXiv:hep-ph/055142, arXiv:0803.2360, arXiv:1004.1092] – [NMSSMTools: arXiv:hep-ph/0505022] – [CoGenT: arXiv:1002.4703] – [Xenon100: arXiv:1104.2549] – [Fermi-LAT: arXiv:1001.4531]