



ILL

ESRF

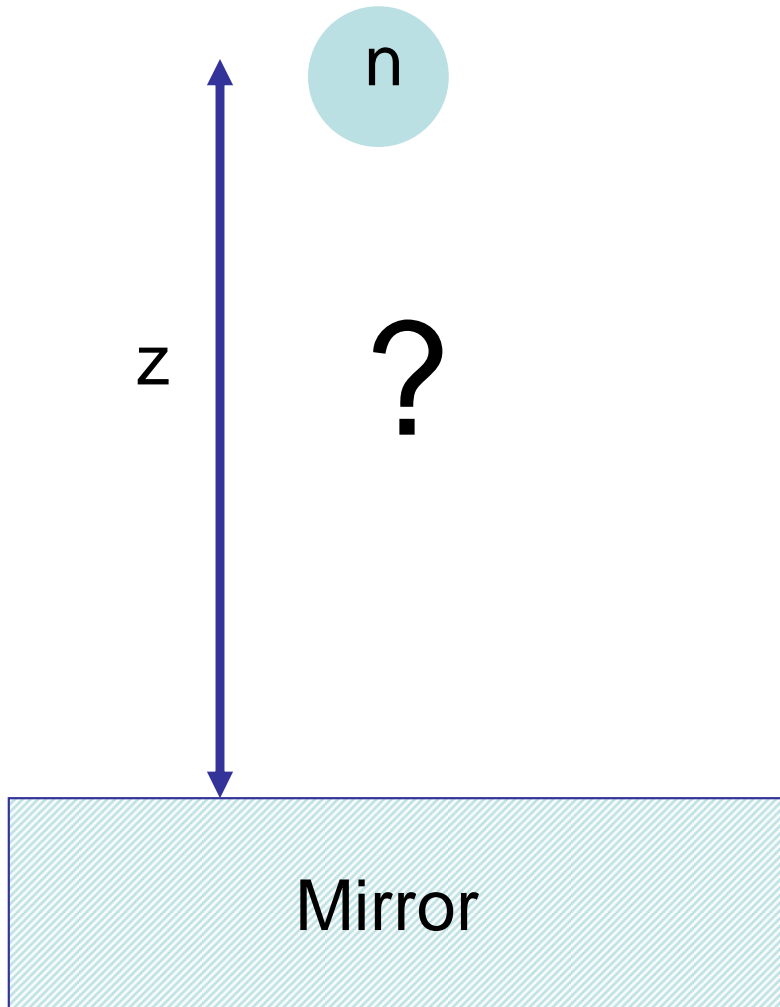
The GRANIT project: Status and perspectives

EPS-HEP 2011

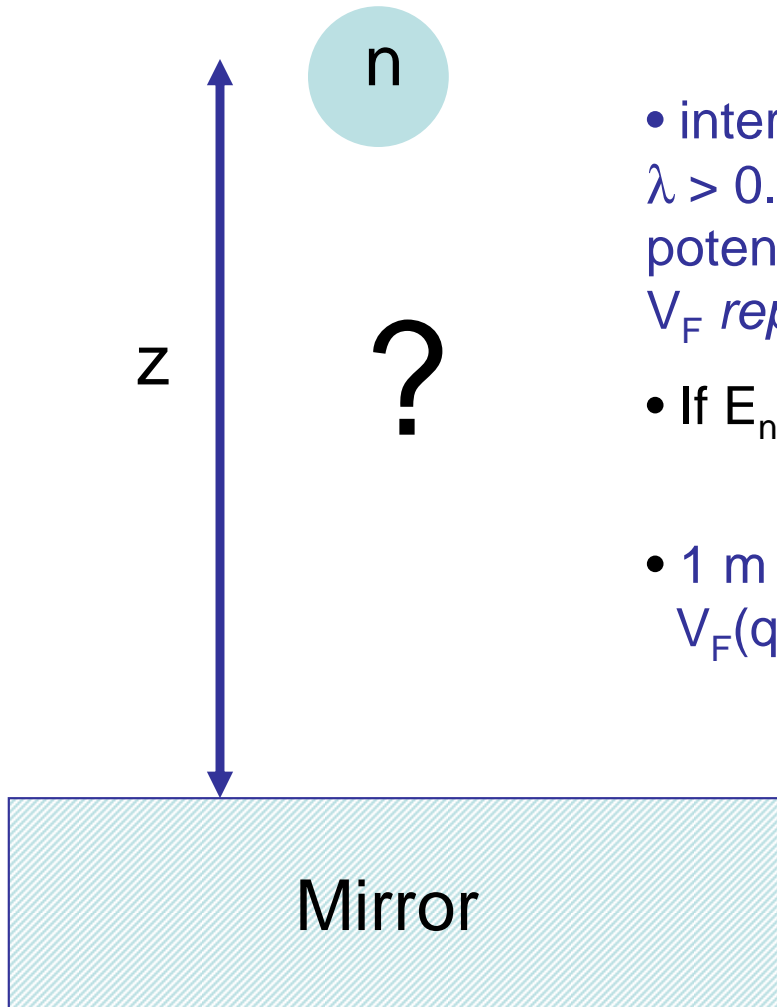
D. Rebreyend (LPSC/IN2P3-UJF)

- The quantum phenomenon
- First observations
- GRANIT
- Constraint on Chameleon models

Will neutrons bounce off the mirror ?



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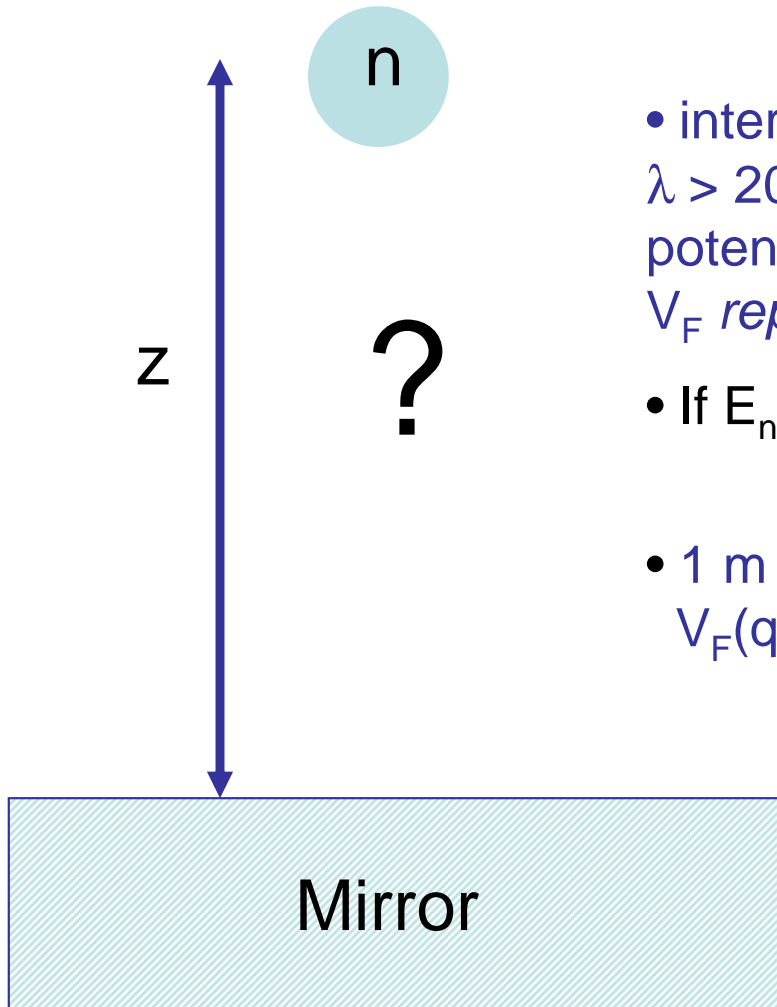
- interaction between low energy neutrons ($E_n < 25$ meV, $\lambda > 0.2$ nm) and matter described by an effective potential V_F (Fermi potential).

V_F *repulsive* for most usual materials ($V_F \sim 50-300$ neV).

- If $E_n < V_F \rightarrow$ reflexion at any incidence angle
 \equiv Ultra Cold Neutrons (UCN)

- 1 m free fall = 100 neV (5 m/s)
 $V_F(\text{quartz}) = 90$ neV

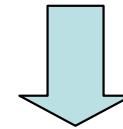
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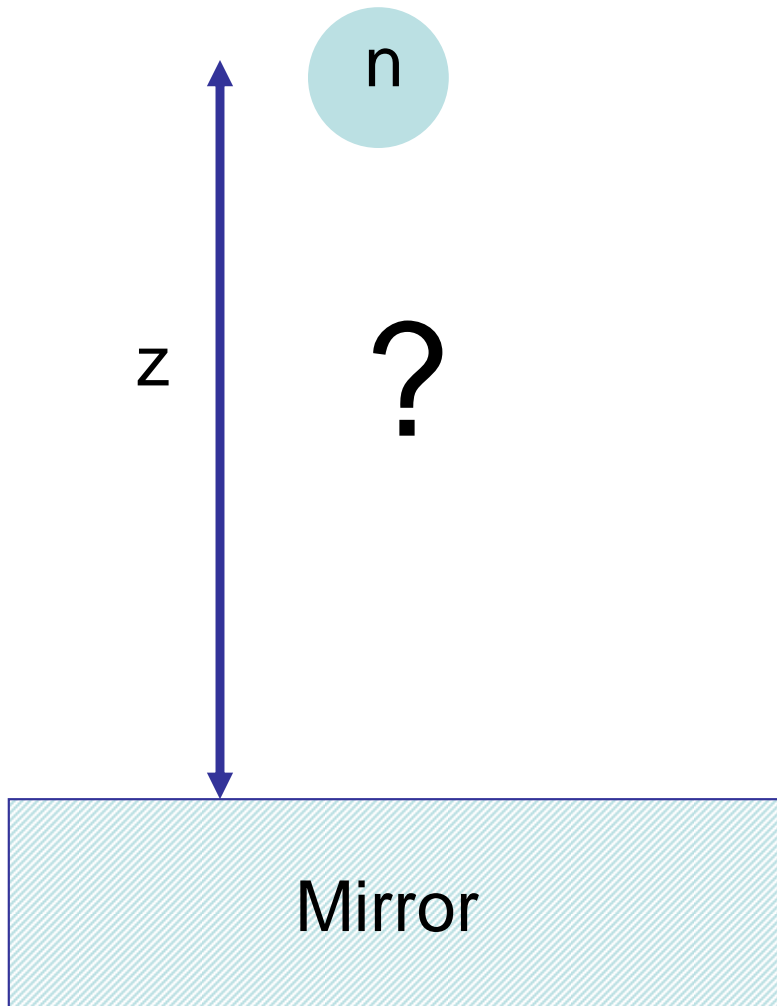
- interaction between low energy neutrons ($E_n < 25$ meV, $\lambda > 20$ nm) and matter described by an effective potential V_F (Fermi potential).

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Neutrons will rebound at the same height for $z < 90$ cm.



Action of system: $\mathcal{L} = (8)^{1/2} m g^{1/2} z^{3/2}$

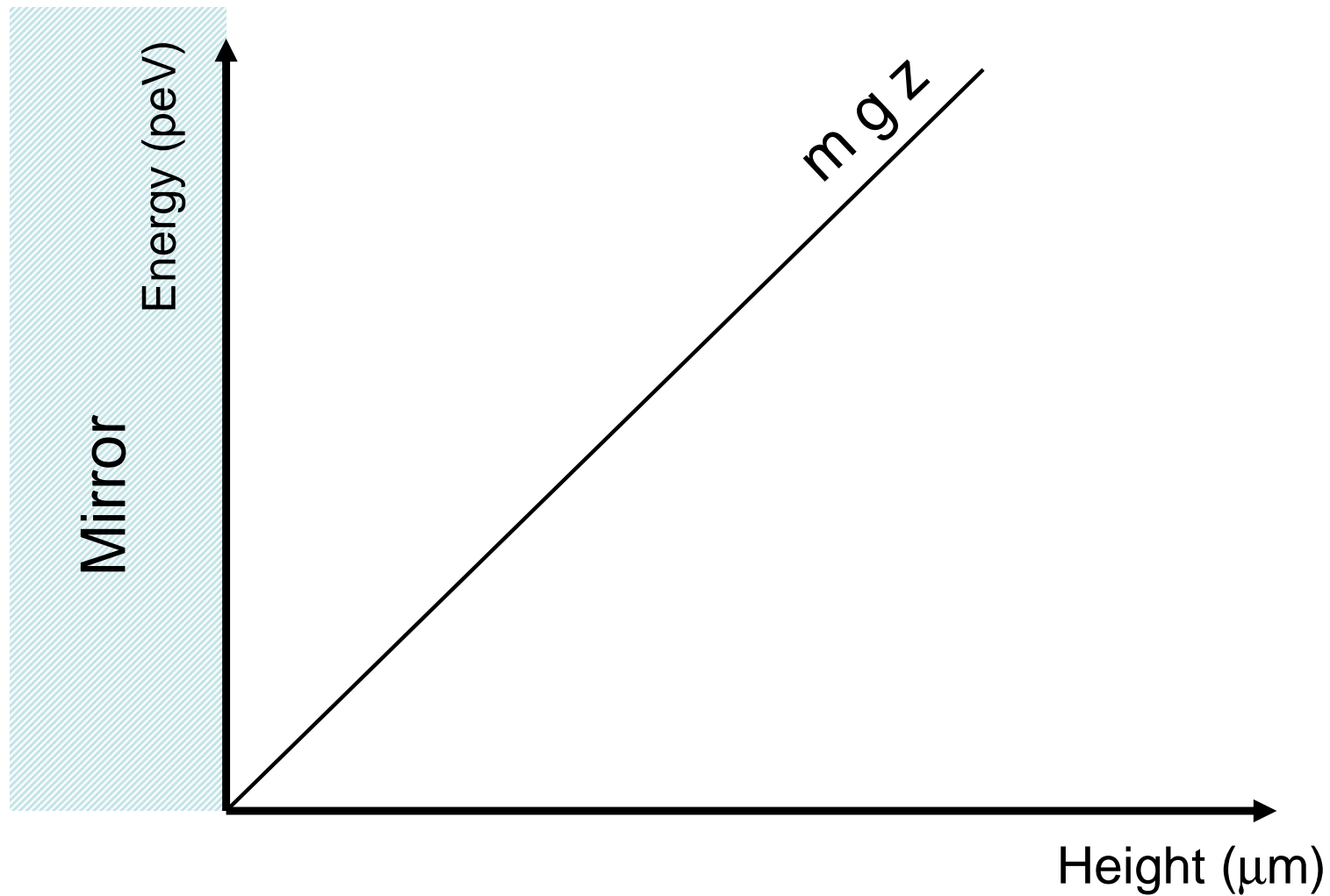
For $\mathcal{L} = \hbar$:

$\rightarrow z \approx 3 \mu\text{m}$

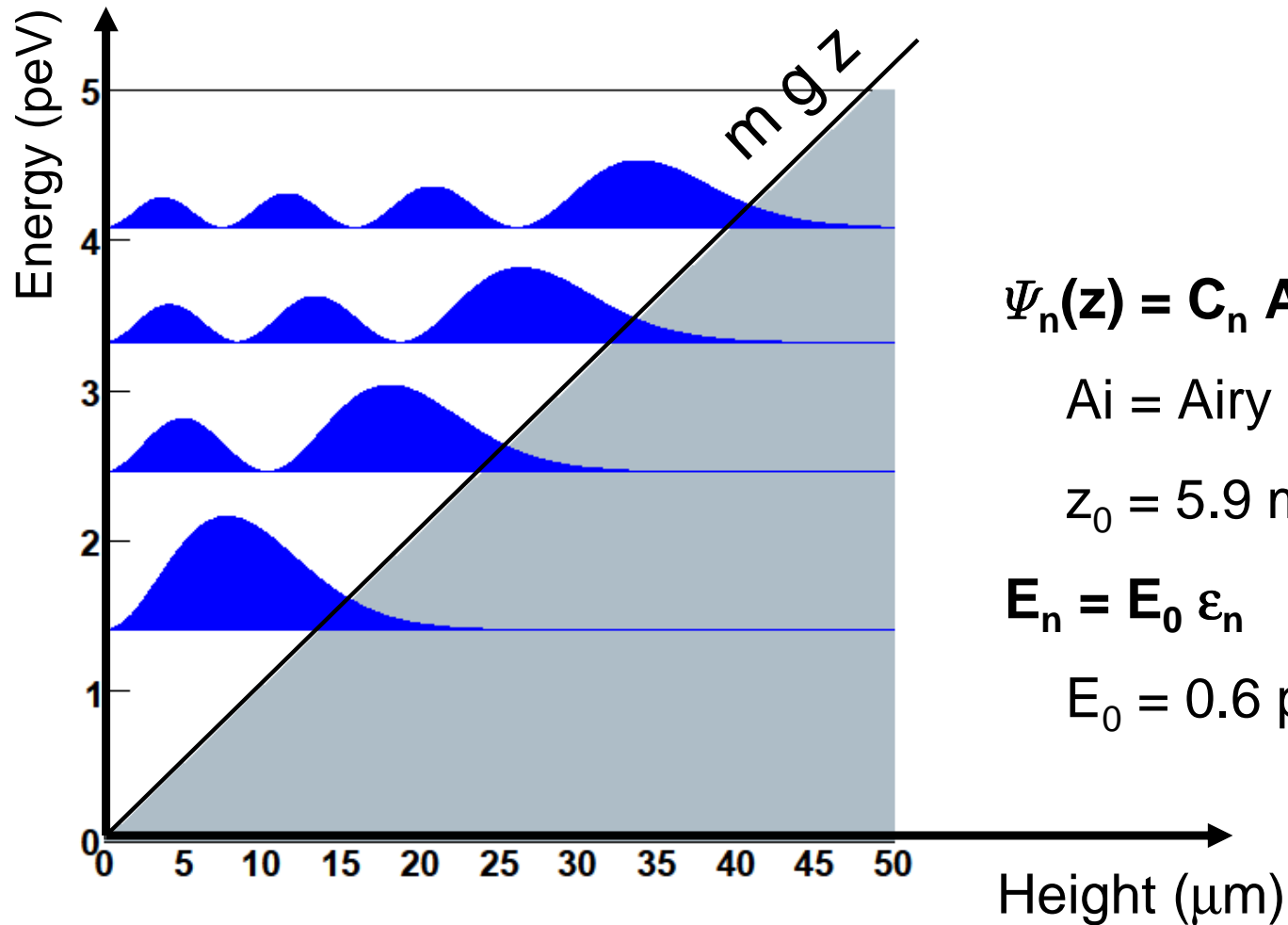
$\rightarrow E \approx 0.3 \text{ peV (7 cm/s)}$

$\rightarrow f \approx 100 \text{ Hz}$

A potential well problem



A potential well problem



$$\Psi_n(z) = C_n A_i(z/z_0 - \varepsilon_n)$$

A_i = Airy functions

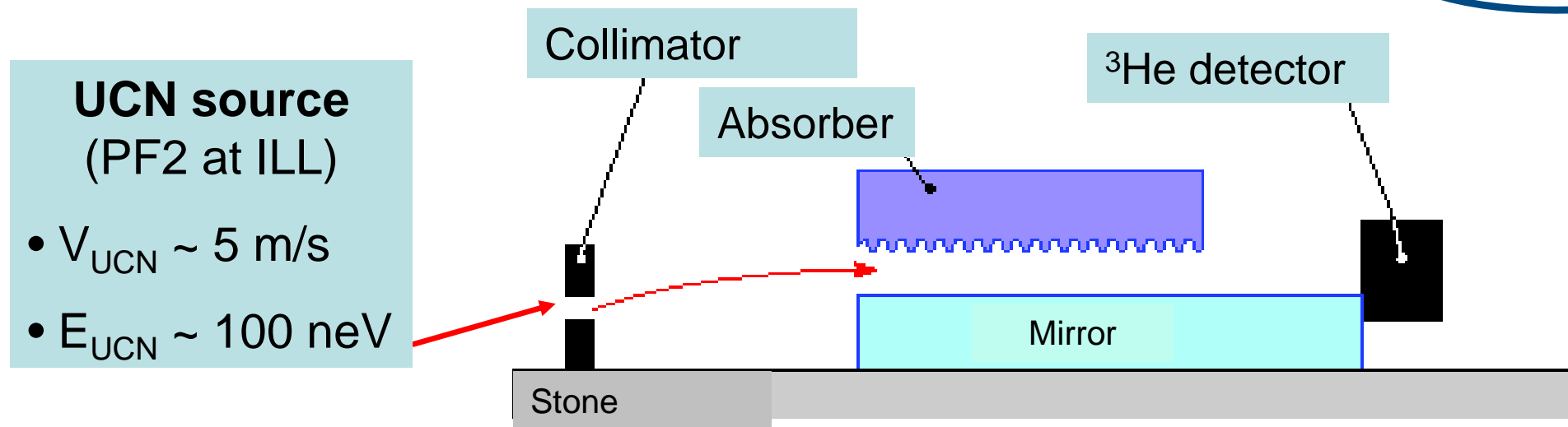
$$z_0 = 5.9 \text{ mm}$$

$$E_n = E_0 \varepsilon_n$$

$$E_0 = 0.6 \text{ peV}$$

$$-\frac{\hbar^2}{2m} \frac{d^2 \Psi}{dz^2} + mgz \Psi = E \Psi$$

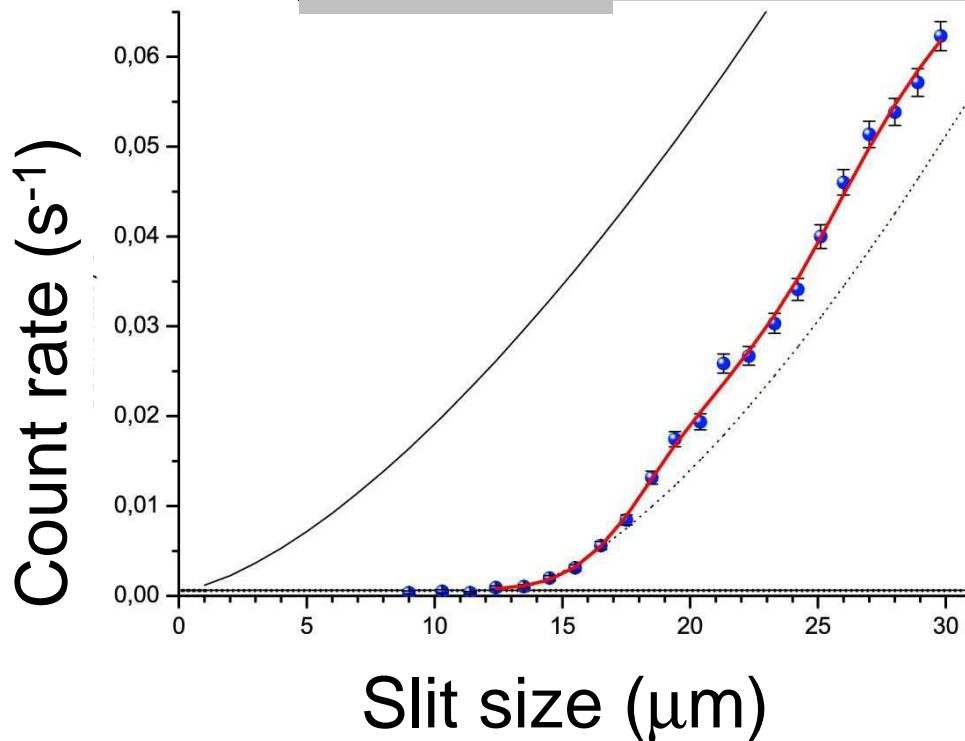
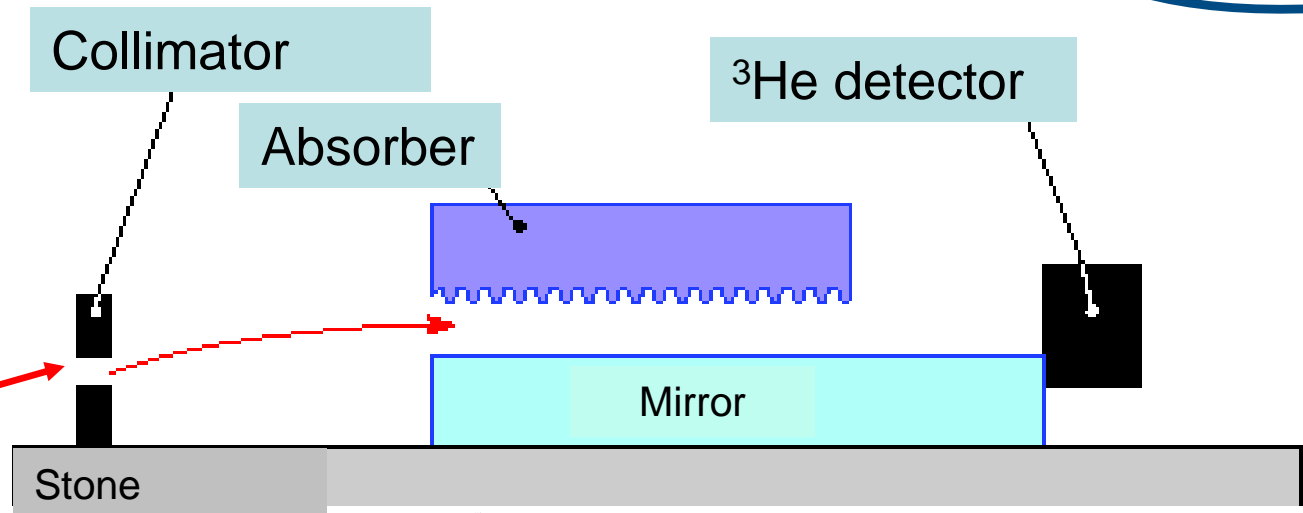
The first observations (ILL, 2000) (Integral mode)



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UCN source
(PF2 at ILL)

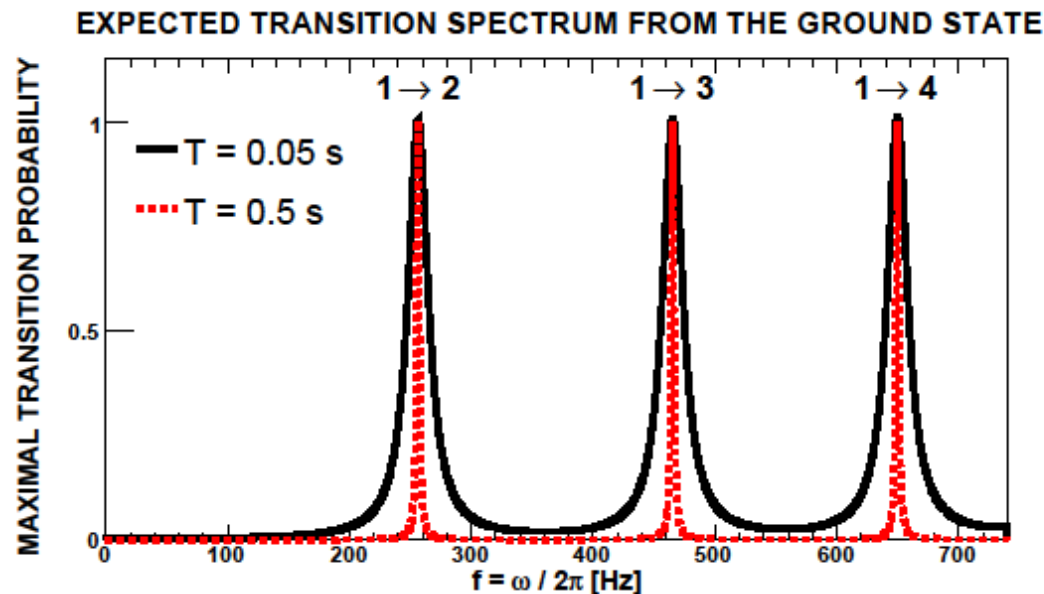
- $V_{\text{UCN}} \sim 5 \text{ m/s}$
- $E_{\text{UCN}} \sim 100 \text{ neV}$



V. Nesvizhevsky et al.,
Nature 2002

Goal: induce resonant transitions to measure energy of quantum levels via frequency measurements.

Method: use of an oscillating perturbation (Rabi resonance).



Perturbation:

- Vibration of the mirror (Jenke et al, Nature Phys. 2011)
- Magnetic field gradients → GRANIT

Transitions induced by magnetic field gradients (Flow through mode)

UCN
source

Wire system

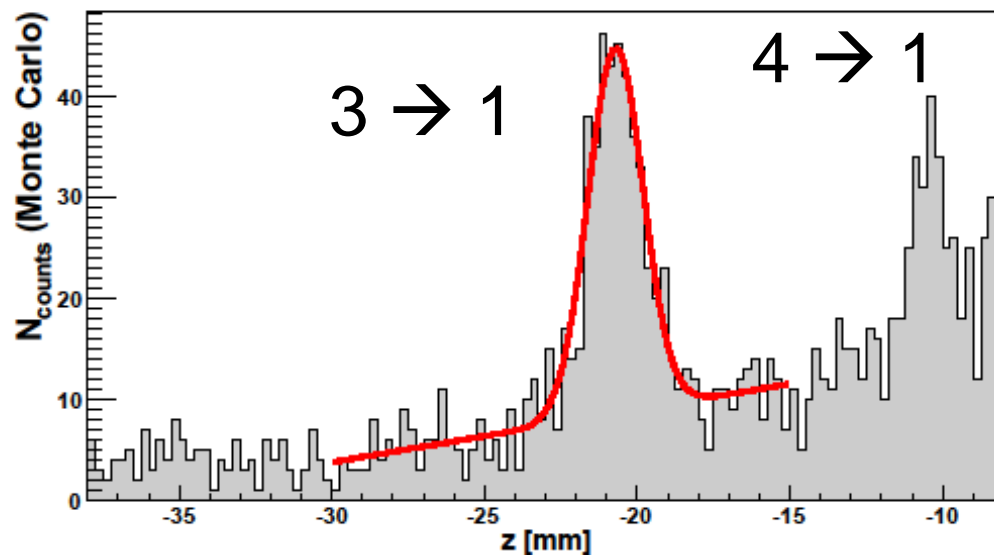
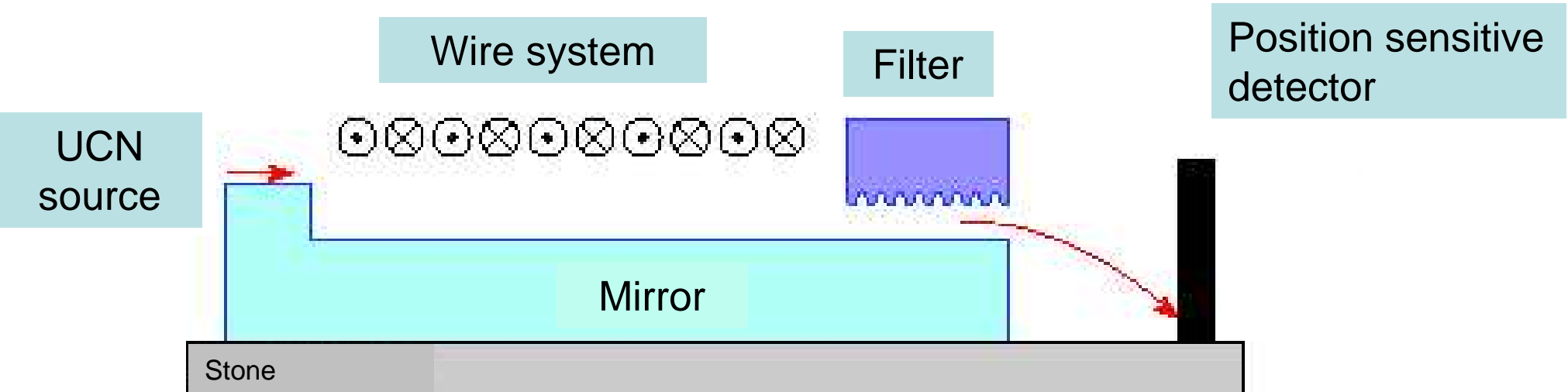
Filter

Position sensitive
detector



Velocity meas.
via free fall height

Transitions induced by magnetic field gradients (Flow through mode)

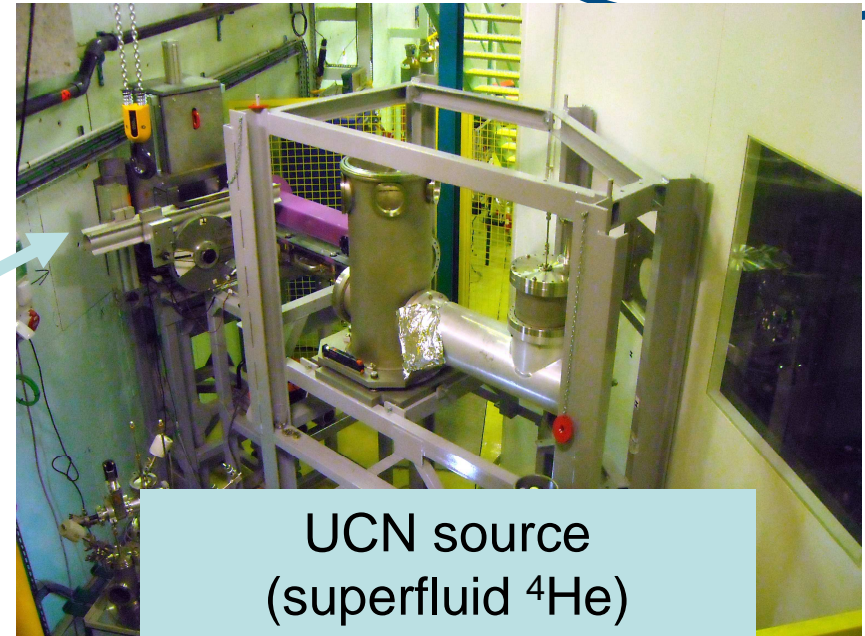


Height ↔ UCN velocity ↔ Frequency

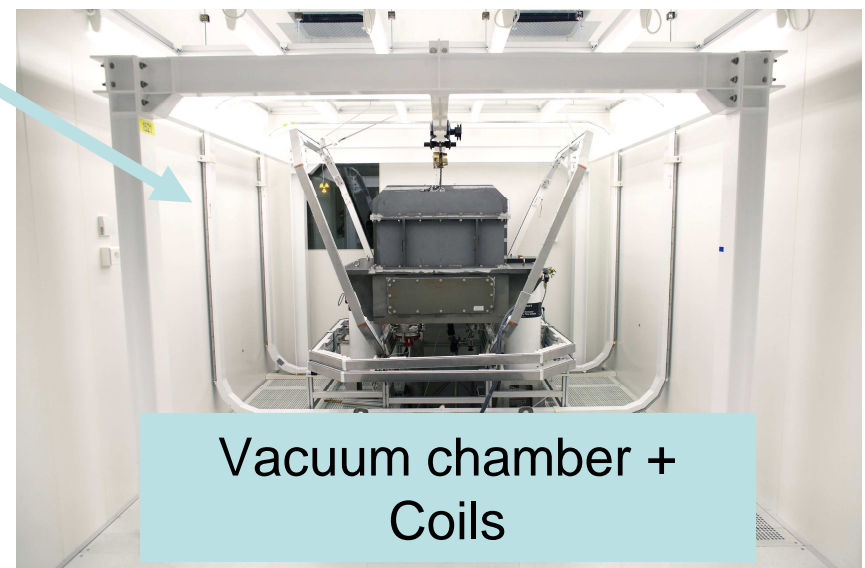
The GRANIT spectrometer



Clean room



UCN source
(superfluid ^4He)



Vacuum chamber +
Coils

Limits on a fifth force

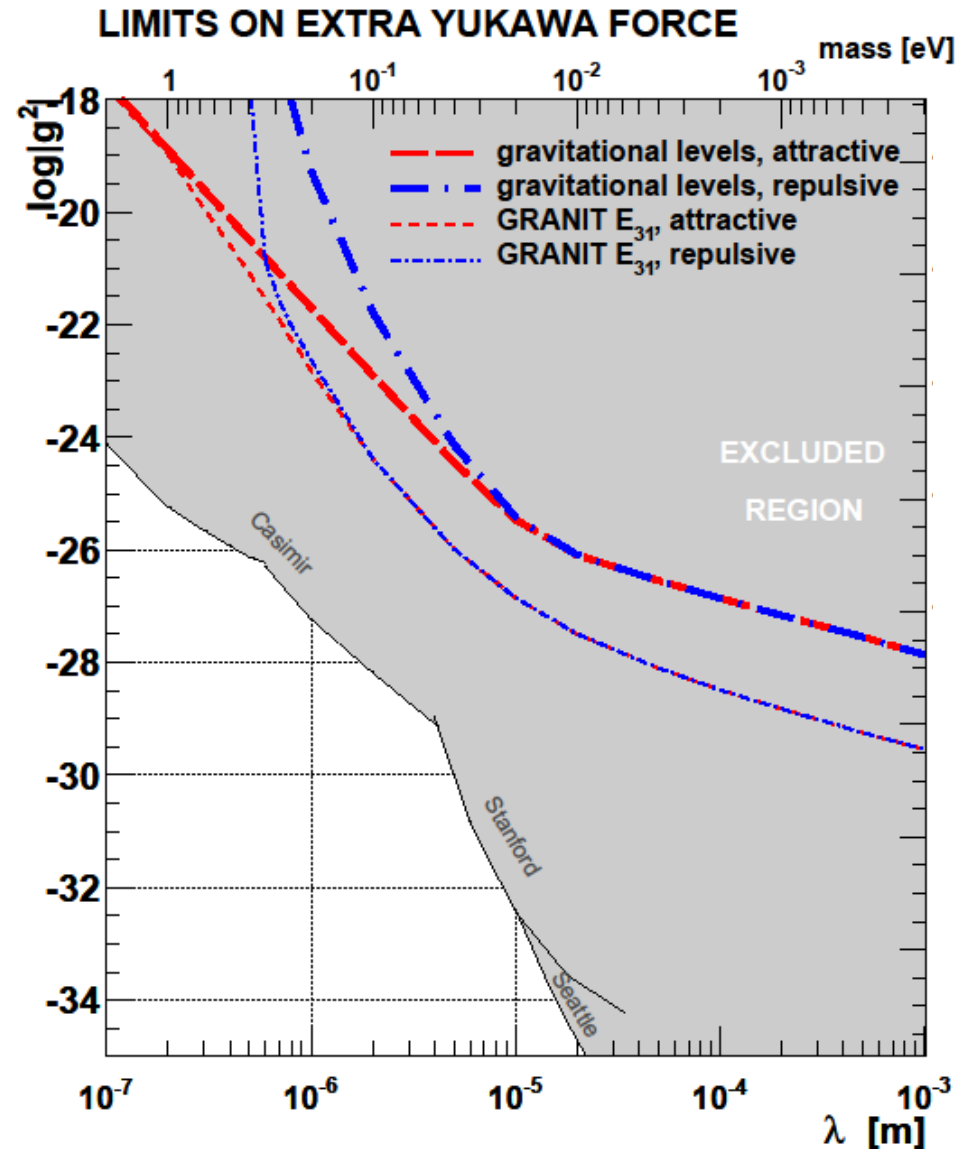
Yukawa interaction between neutron and mirror:

$$V(z) = \frac{g^2 \rho}{2m_n} \hbar c \lambda^2 e^{-z/\lambda}$$

g = coupling constant

λ = range

$\rho = 2.5 \text{ g.cm}^{-3}$ (quartz density)



G. Pignol, UJF PhD thesis, 2009

- Cosmological model to explain the dark energy problem.
- Scalar field *coupled* to matter.
- Chameleon mechanism: only thin shell at the surface of a body contributes to the field
 - Could evade lab limits with macroscopic bodies even for strong coupling
 - Modification of the gravity law close to the surface of a body

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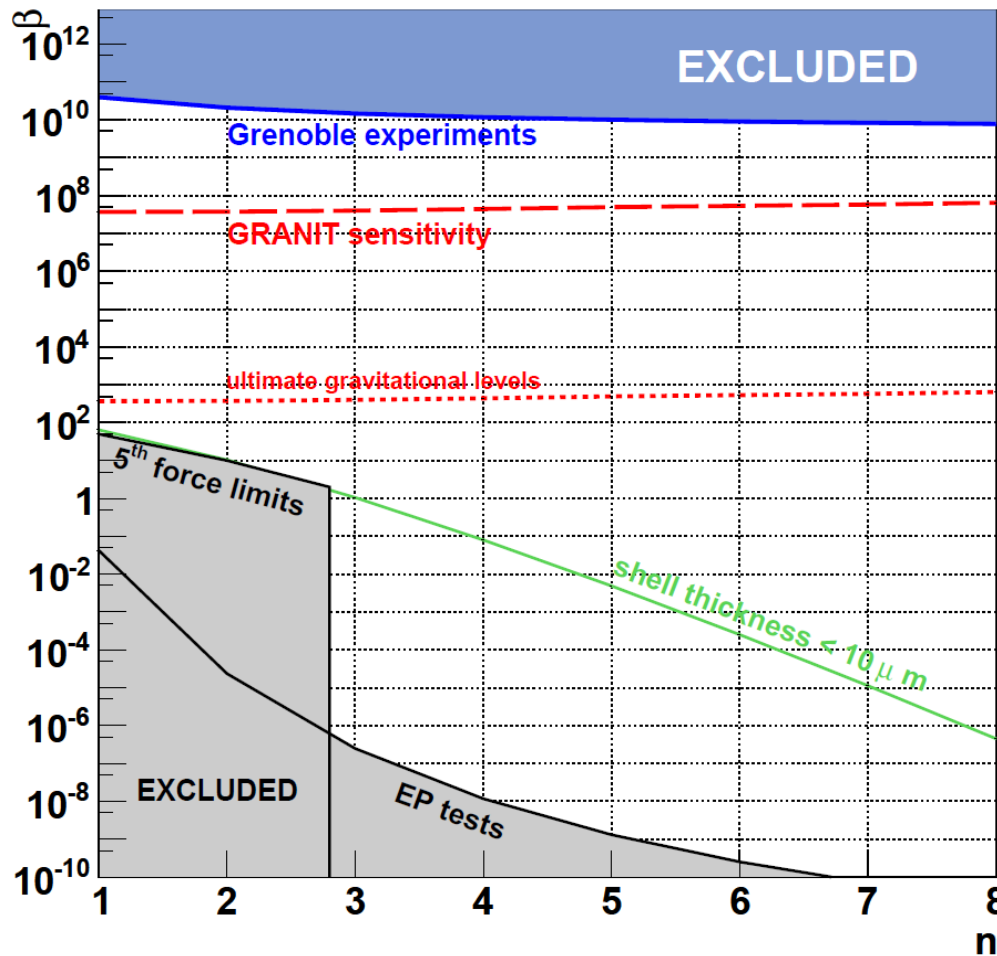
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Calculation for GRANIT (Brax & Pignol, arXiv:1105.3420v1):

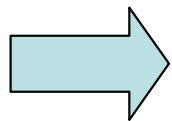
$$\phi_{Ch}(z) = \frac{\hbar c}{d_{DE}} \left(\frac{z}{d_{DE}} \right)^{2/2+n} \quad n = 1, 2, 3...$$

where $d_{DE} = 82 \mu\text{m}$ is the characteristic distance associated to the dark energy ($\hbar c / d_{DE}^4 = \text{dark energy density}$).

Chameleon exclusion plot



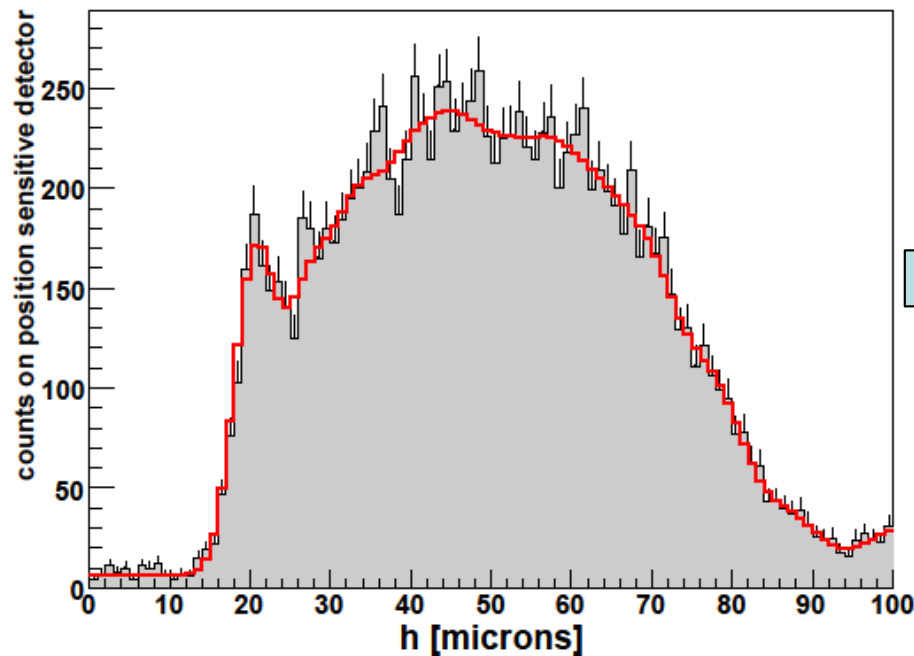
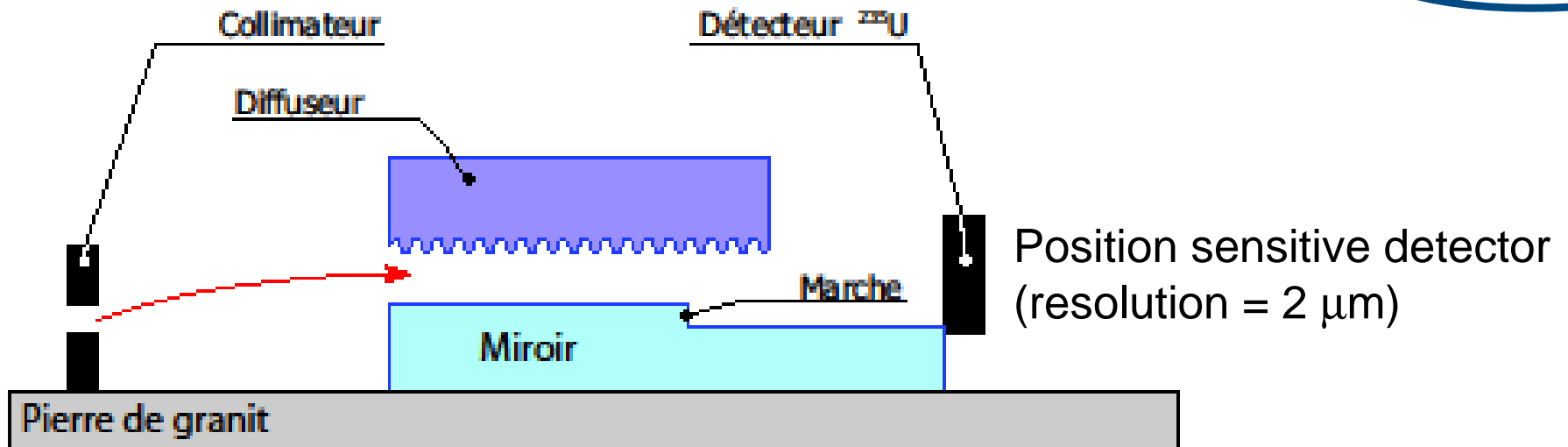
Additional bound state



Competitive limit for strongly coupled chameleons

Thank you for your attention.

Observations in differential mode



Measurement of z_0
with a precision $\approx 3\%$