

Neutrino Physics – Accomplishments and Future Challenges

**Joint ECFA-EPS Session
EPS HEPP
Grenoble
July 23 2011**

**Dave Wark
Imperial/RAL
(but speaking
only for me!)**



**Science & Technology
Facilities Council**

**Imperial College
London**

What will I talk about?

This is not the plenary neutrino talk – that is by Nishikawa-san on Tuesday.

This talk is about the decisions we need to make.

This is an ECFA session, so I will only talk about experiments with an accelerator in them – no solar neutrinos, no double-beta decay, not even very much reactor neutrinos – sorry.....

And I won't talk all that much about accomplishments.

We have more than enough Future Challenges for a 20 minute talk....

Three neutrino mixing.

If neutrinos have mass:

$$|\nu_l\rangle = \sum U_{li} |\nu_i\rangle$$

$$U_{li} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \cdot \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \cdot \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

where $c_{ij} = \cos \theta_{ij}$, and $s_{ij} = \sin \theta_{ij}$

$$P_{e\mu} = \sin^2 2\theta_{12} \sin^2 2\theta_{13} \sin^2 \Delta$$

Remember degeneracies
And covariances!

$$\begin{aligned} P_{e\mu} &= \sin^2 2\theta_{12} \sin^2 \theta_{13} \cos^2 \theta_{13} \sin^2 2\theta_{13} \sin^2 2\theta_{23} \sin^2 \Delta \\ &= \sin^2 2\theta_{12} \cos^2 \theta_{13} \sin^2 2\theta_{13} \sin^2 2\theta_{23} \sin^2 \Delta \\ &= \sin^2 2\theta_{12} \cos^2 \theta_{13} \sin^2 2\theta_{13} \sin^2 2\theta_{23} \sin^2 \Delta \end{aligned}$$

with $\theta_{13} \approx 9^\circ$

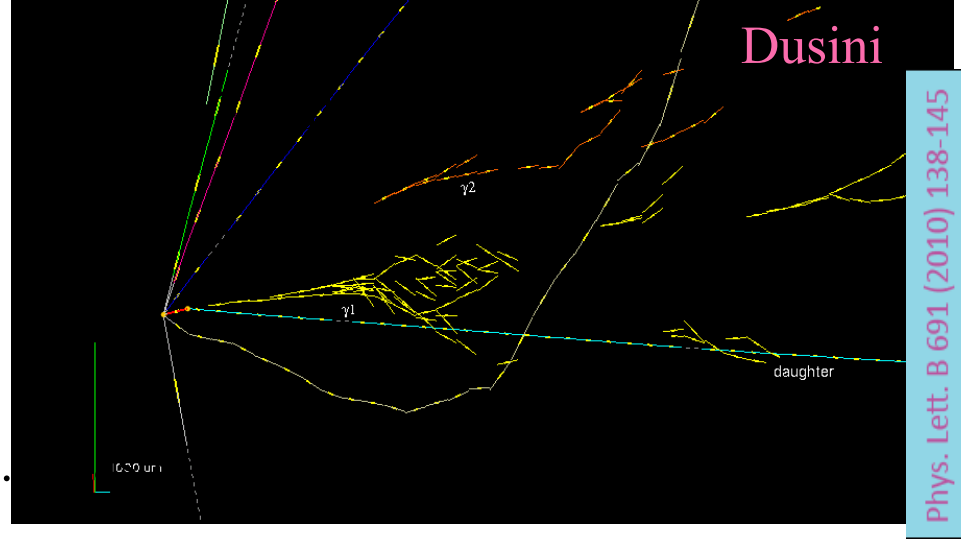
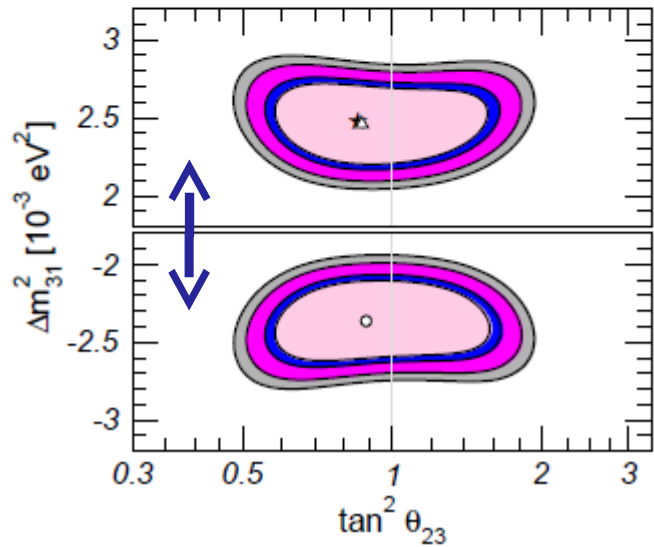
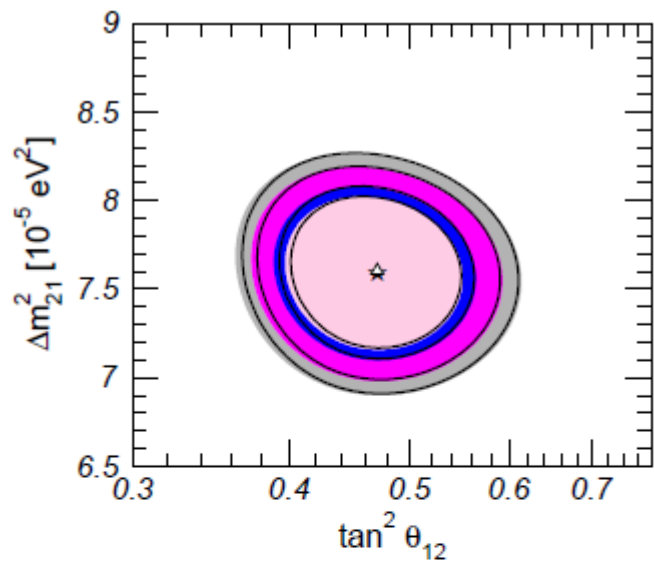
$$\sim 0.03$$

and $\Delta \approx \pi/4$

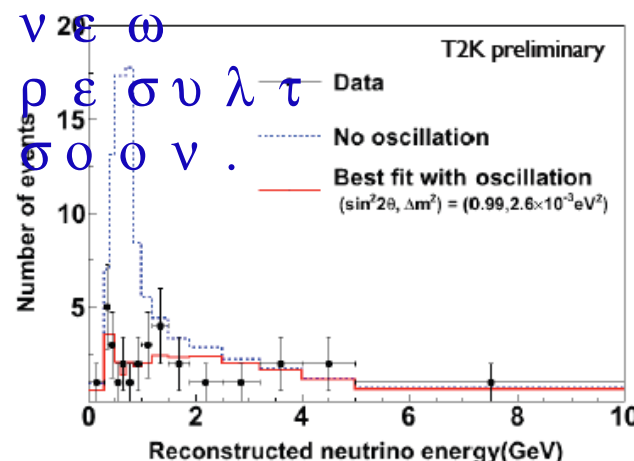
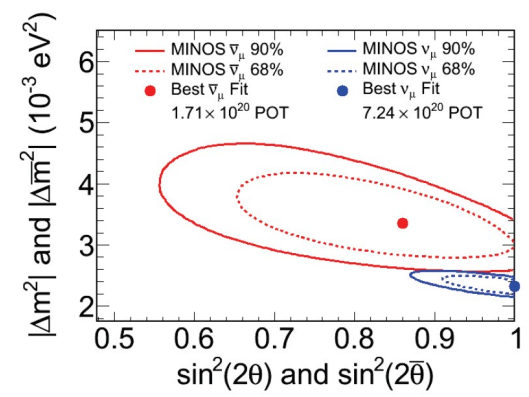
$$\sim \pi/4$$

How well do we know θ_{ij} ?

arXiv:1001.4524v4 [hep-ph] 16 Jun 2011



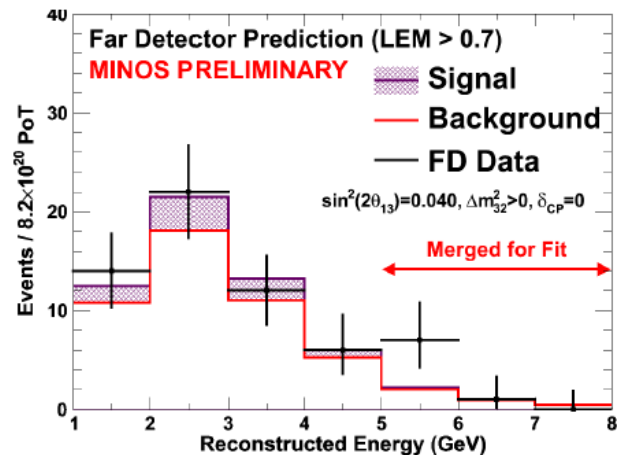
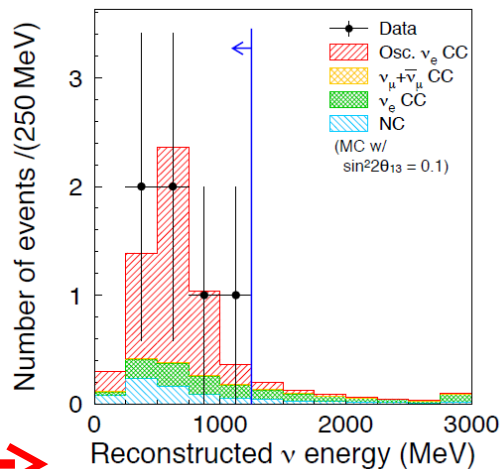
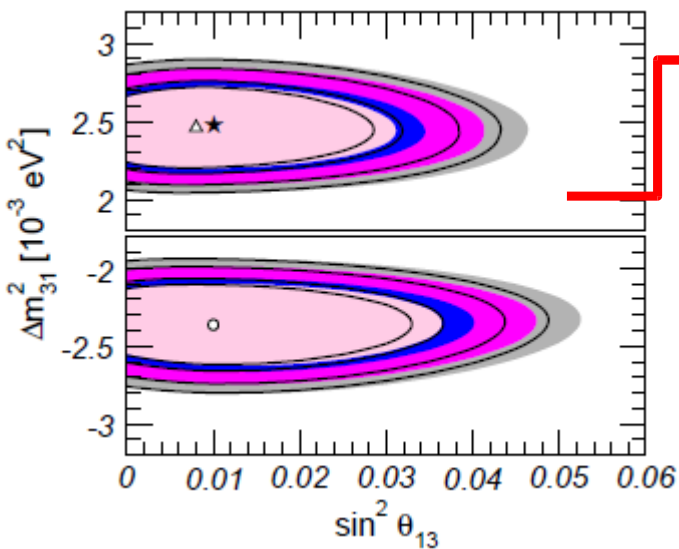
MINOS $\nu_{\mu} \rightarrow \nu_{\tau}$
 $\nu_{\mu} \rightarrow \nu_{\tau}$
 $\mu \rightarrow \tau$
 $\tau \rightarrow \mu$
Holin



1st T2K ν_{μ} disappearance

Giganti

What about θ_{13} ?



T2K (2.5σ)

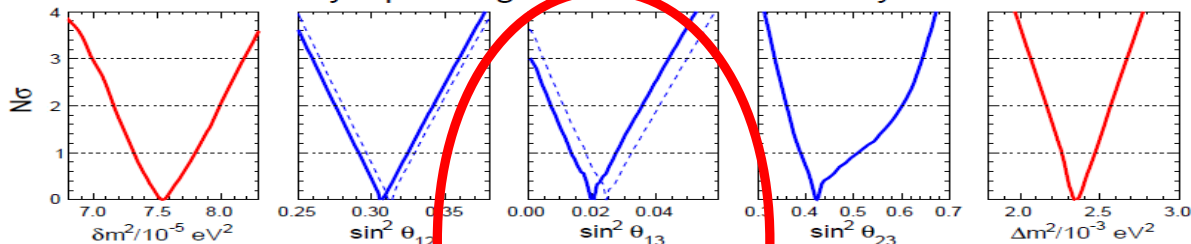
MINOS (1.7σ)

ν_e appearance, $\theta_{13} > 0$?

Evidence of $\theta_{13} > 0$ from global neutrino data analysis

G.L. Fogli,^{1,2} E. Lisi,² A. Marrone,^{1,2} A. Palazzo,³ and A.M. Rotunno¹

Synopsis of global θ_{13} oscillation analysis



arXiv:1106.6028v1 [hep-ph] 29 Jun 2011

For this talk, assume $\theta_{13} > 0$, and T2K and NOvA will prove it!

What then for ν oscillations?

Prove $\theta_{13} > 0$.

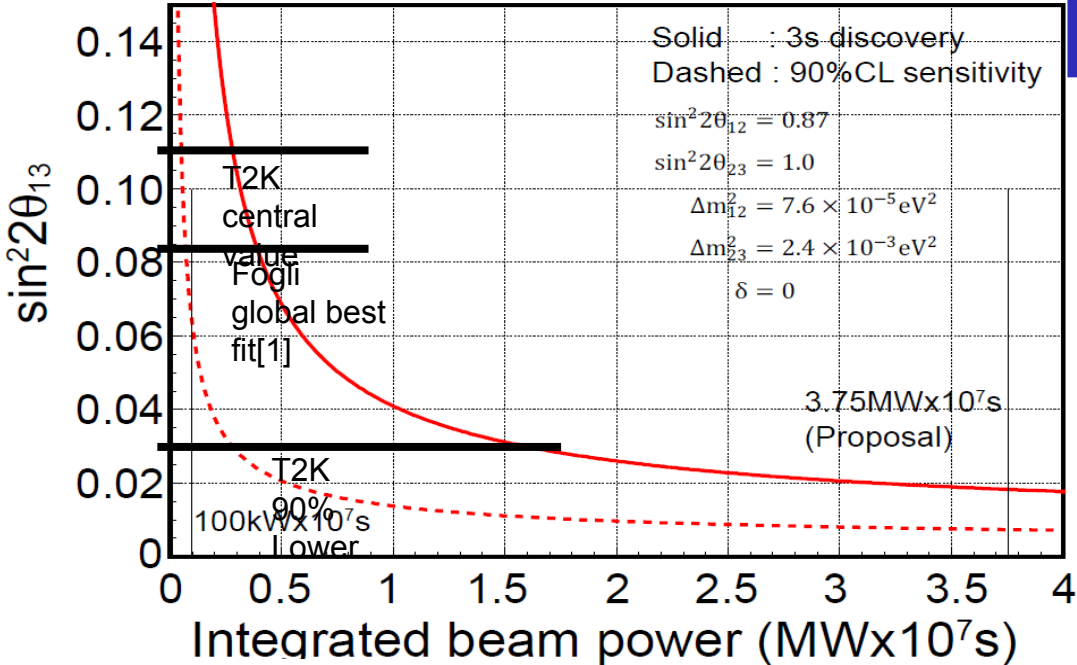
Measure all parameters with greater precision (because you need it for the rest of the list, but is there any particular necessary accuracy indicated by theory?).

Make the most sensitive possible test of the deviation of θ_{23} from 45° .

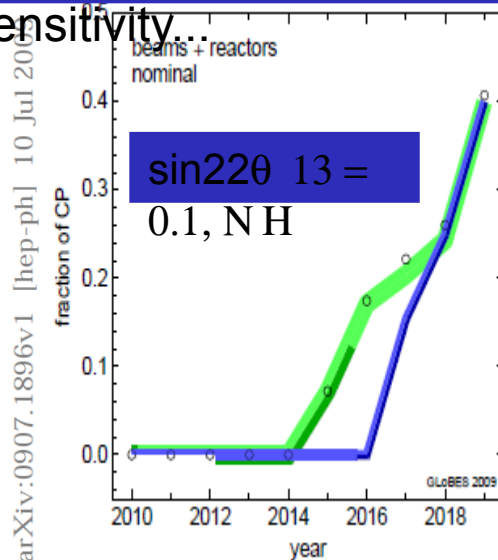
Determine the mass hierarchy by observing matter effects \rightarrow higher energy, longer baseline.

Measure the angle δ !

Are there any surprises?

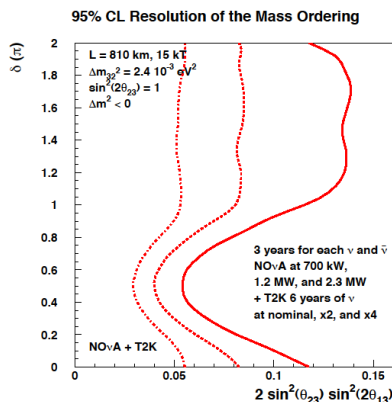
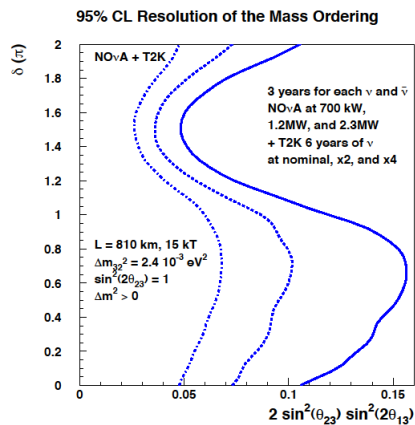


Even some 90% CP violation sensitivity...

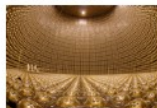


PATRICK HUBER^a, MANFRED LINDNER^b,
THOMAS SCHWETZ^c, AND WALTER WINTER^d

T2K and NOvA



Kamioka



- Ran from Jan 2010 till March 2011
 - Presently stopped; restart foreseen in 2012
 - Maximum beam power 145 kW
 - Muon disappearance → 500 kW (2yr)
 - Electron appearance → >1 MW (>5yr)
- (90%C.L.):

$0.03 < \sin^2 2\theta_{13} < 0.28$ (best fit 0.11) normal
 $0.04 < \sin^2 2\theta_{13} < 0.34$ (best fit 0.14) inverted
Aim: sensitivity $\sin^2 2\theta_{13} > 0.006$ @ 90%C.L. for $\delta_{CP}=0$

- Far detector under construction
- Start foreseen in fall 2013
- Beam power 750 kW (after upgrade)
- Baseline $L=830 \text{ km}$ gives some matter effect sensitivity
- Consider neutrino + antineutrino runs

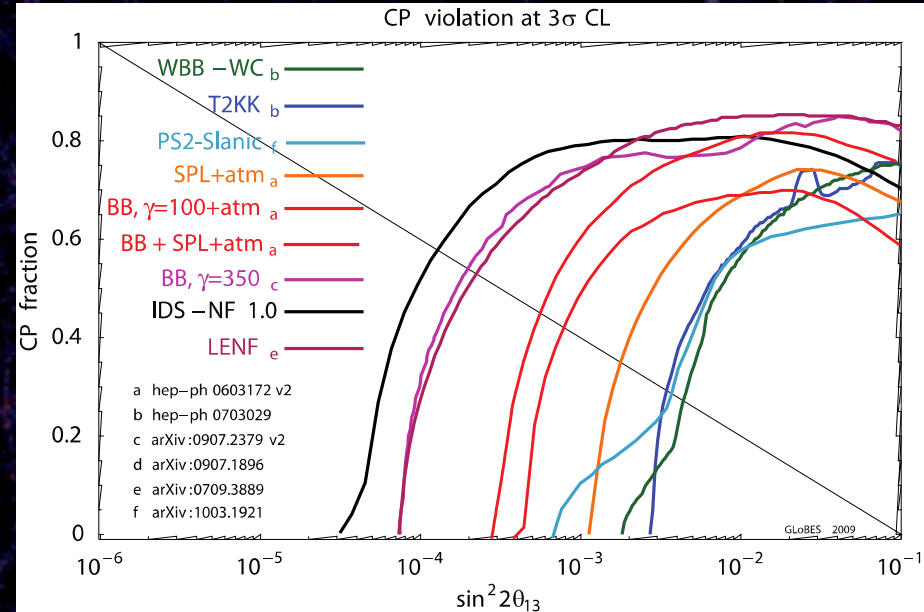
Aim: sensitivity $\sin^2 2\theta_{13} > 0.007$ @ 90%C.L. for $\delta_{CP}=0$

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m^2_{23} L}{4E_\nu} \right) + \text{higher order } f(\delta_{CP}, \theta_{12})$$

OK, then what?



Simon van der Meer, 1925 -



Three “conventional” beam proposals:

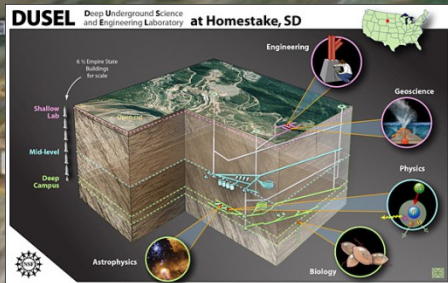
An upgrade of T2K based on reaching 1.6 MW beam power and a new far detector.

LBNE – a plan to build a new neutrino beam at Fermilab aimed at Homestake, where either a large water Cerenkov detector or a LAr tracking calorimeter would be built.

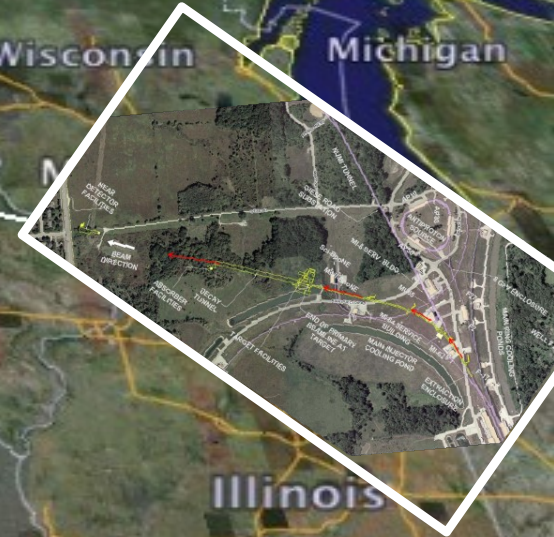
LAGUNA-LBNO – three different options for new long baseline in Europe.

US: Long Baseline Neutrino Experiment

CD 0: January 2010

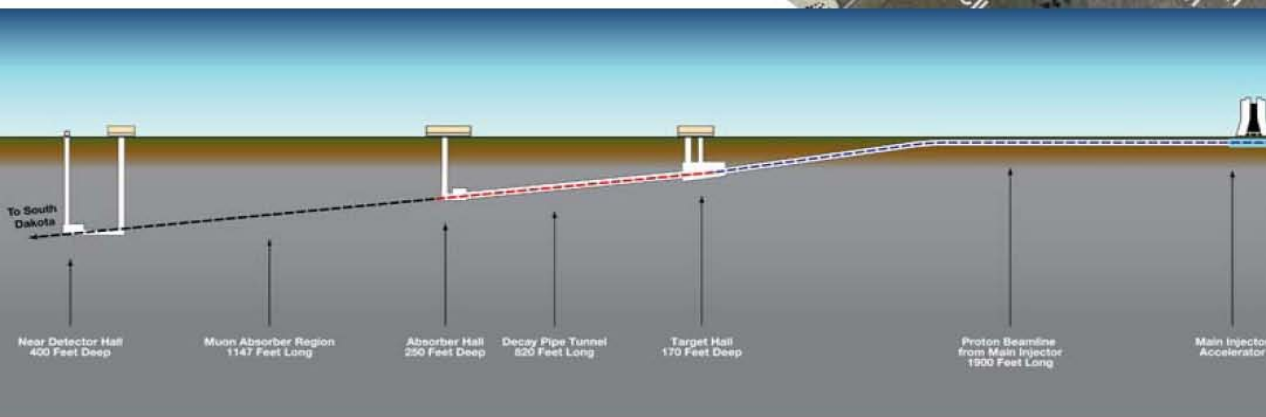
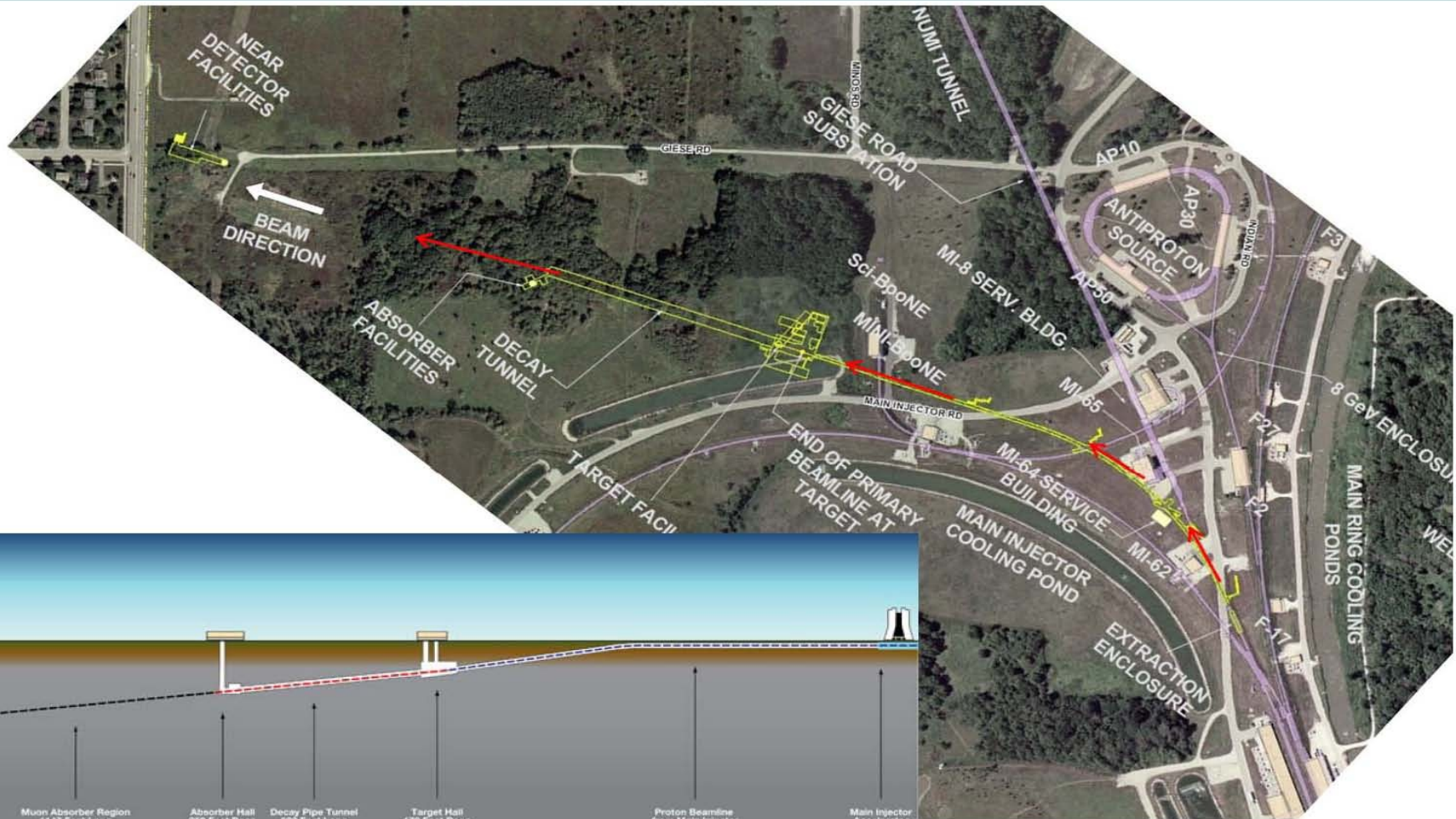


1300 km



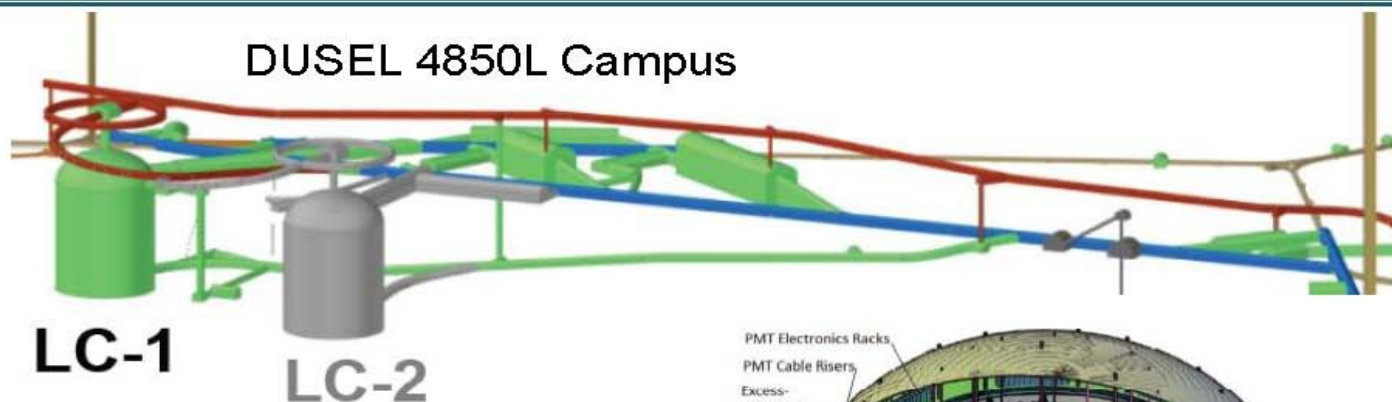
Collaboration:
288 members from 54 institutions (India, Italy, Japan, UK, US)
Continue to grow!

Conceptual Design Overview – Neutrino Beam

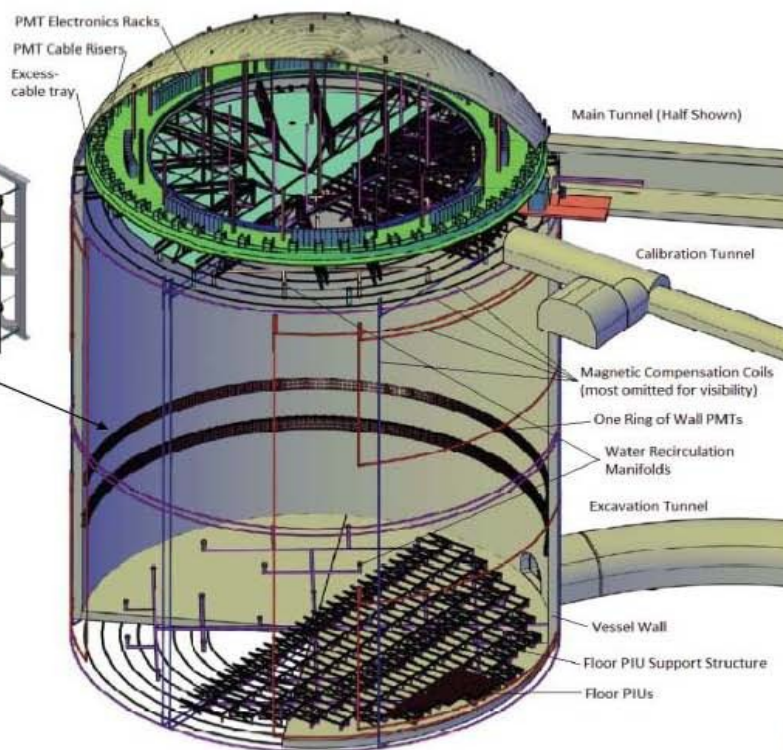
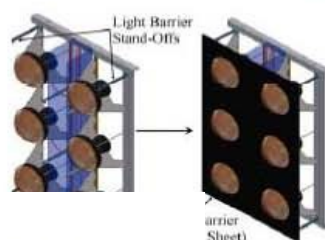
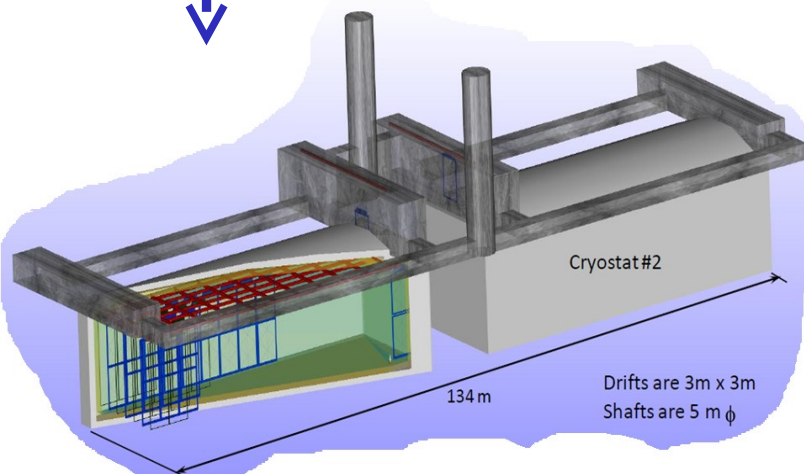


Long-Baseline Neutrino Experiment

Conceptual Design Overview – Water Cherenkov



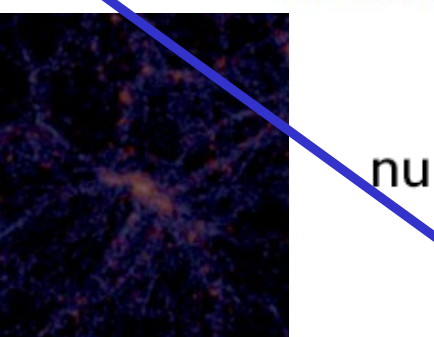
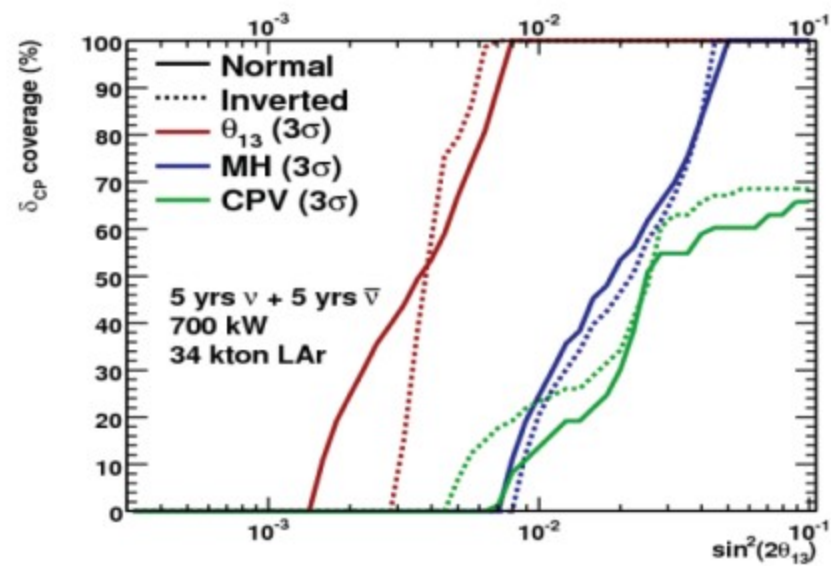
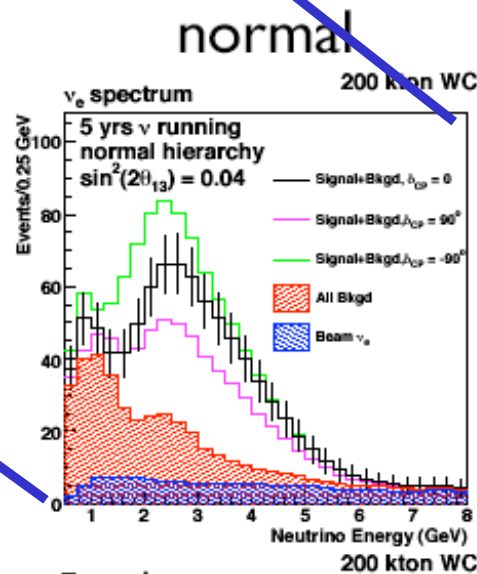
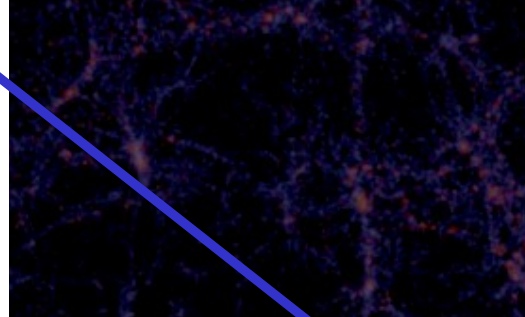
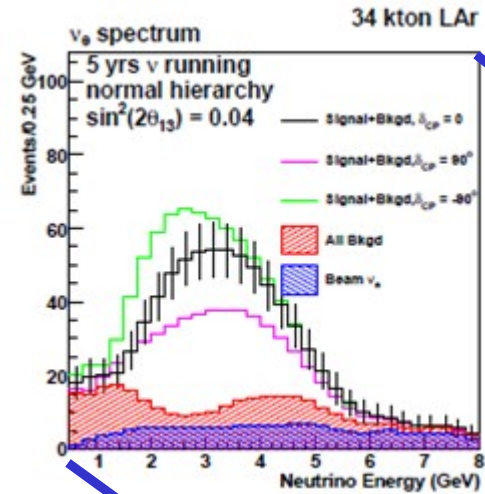
Alternative is
34 kT of LAr



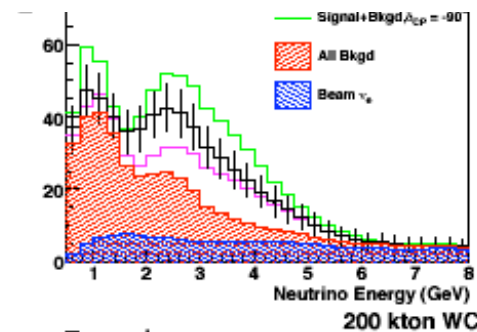
J. Strait — DOE Briefing

16

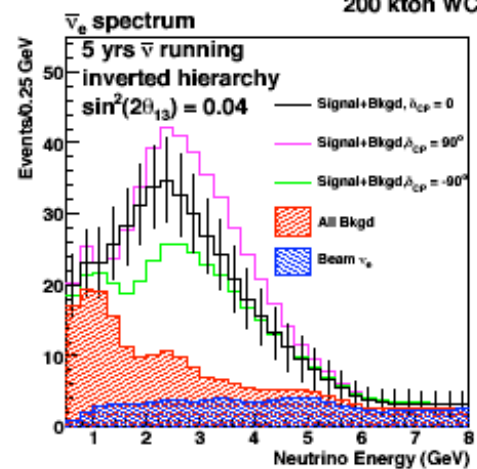
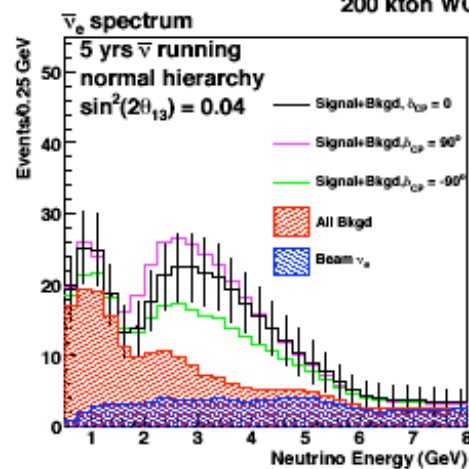
LAr Slight cheaper but riskier – Marx Committee



nu



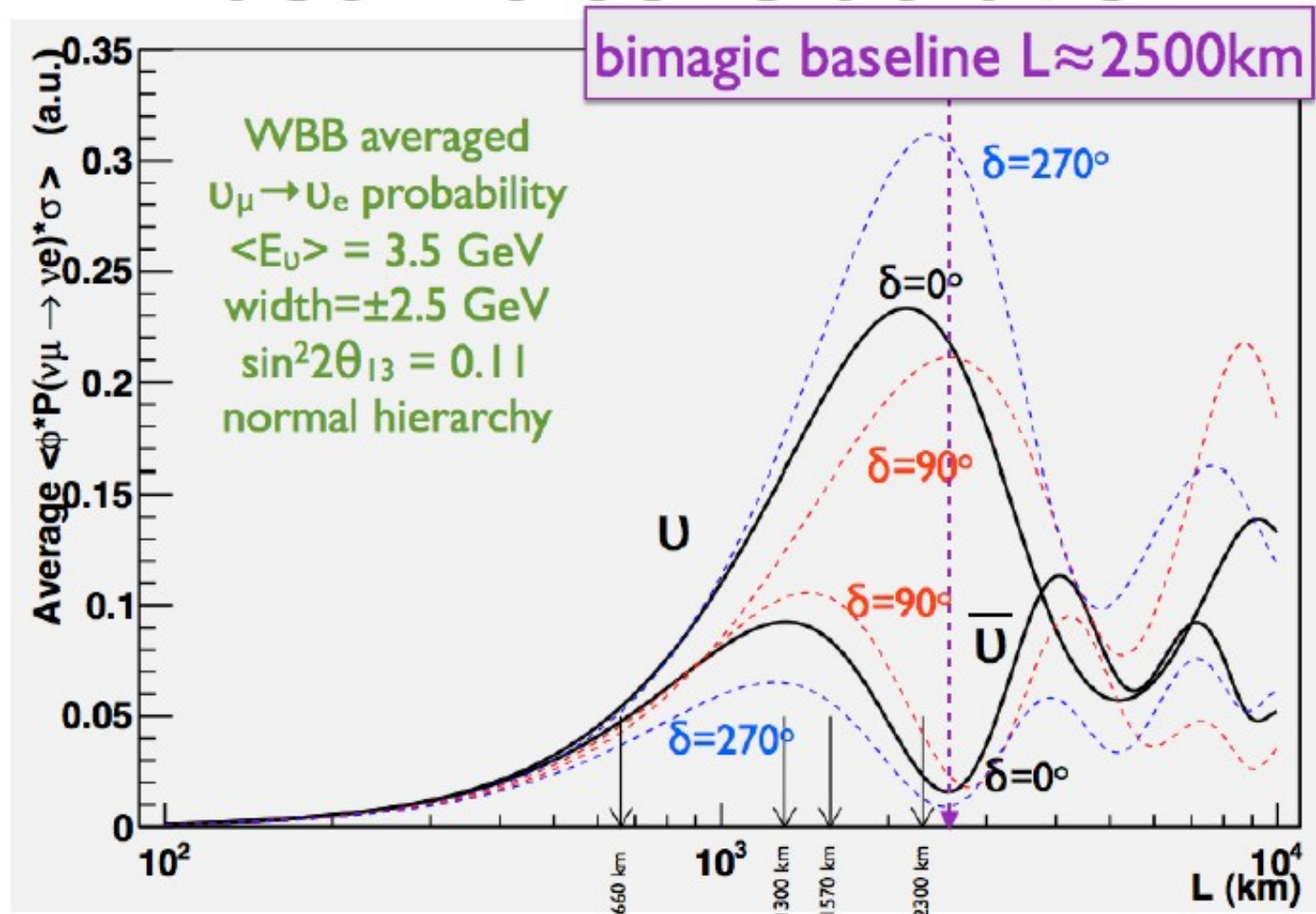
anti-nu



Lisa Whitehead

ark
ge/RAL

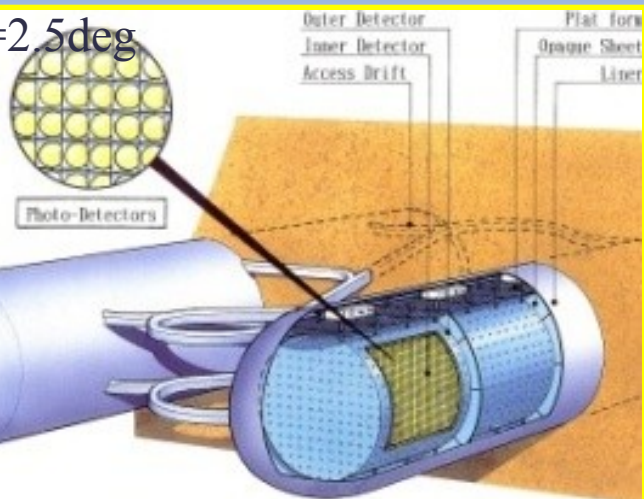
Baseline consideration



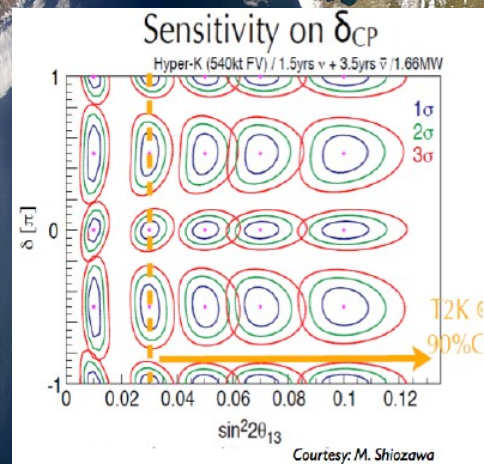
The optimal baselines are in the range 1300-2500 km

Kamioka L=295km

OA=2.5deg

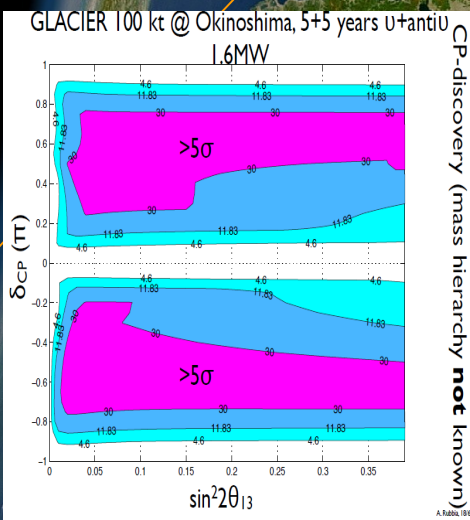
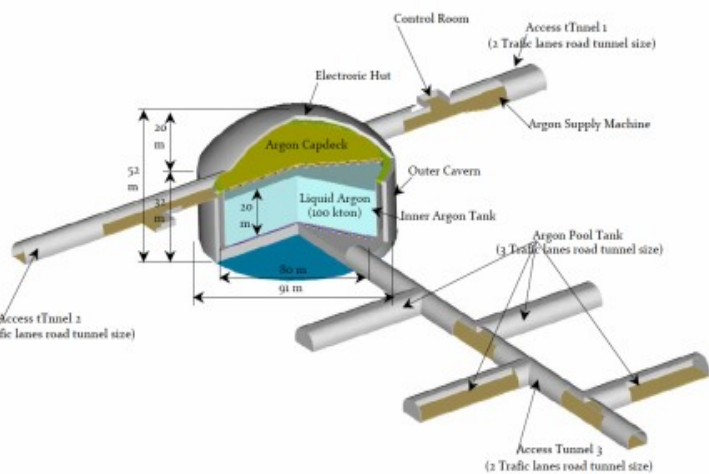


Scenarios in Japan



Okinoshima L=658km OA=0.78deg

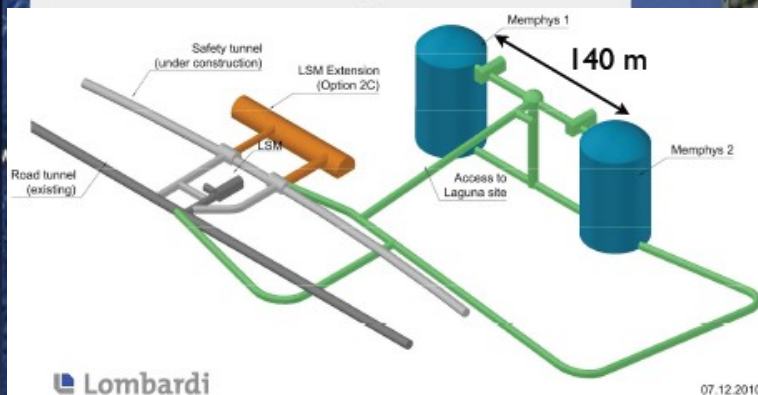
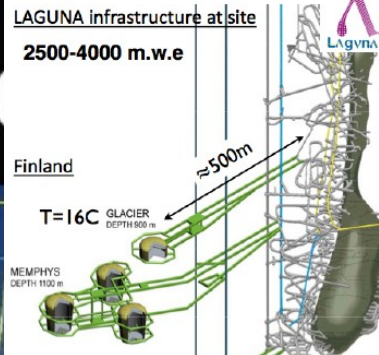
Almost On-Axis



P32 proposal (Lar TPC R&D)
Recommended by J-PARC PAC
(Jan 2010) arXiv:0804.2111

Three options

3 main options selected for LAGUNA-LBNO



CN2FR
 L=130 km,
 HP-SPL 5 GeV 4 MW LINAC +
 accumulator ring
 + MMW target + horn
 + near detector infrastructure

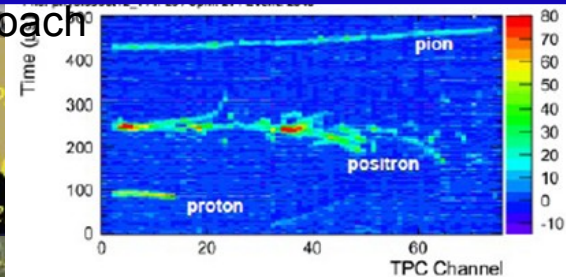
Possible synergy with a β beam

CN2PY

L=2288 km, CERN SPS 400 GeV
 + new beam line 0.75 MW
 + near detector infrastructure
 Longer term: 2MW with
 LP-SPL+HPPS accelerator

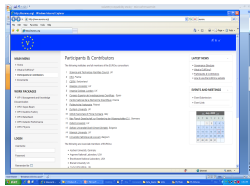
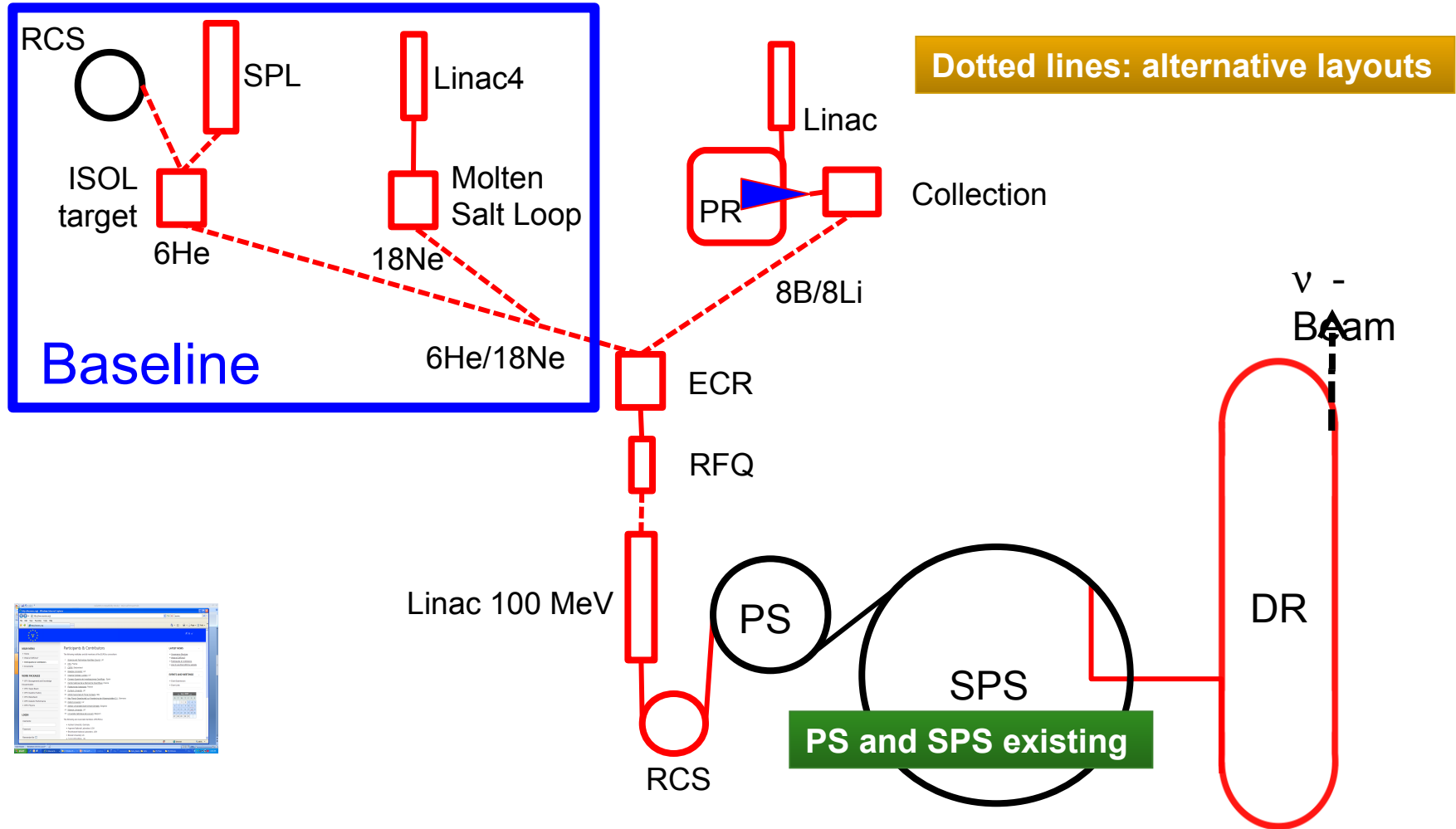
Joint Japanese/European approach

up to $h=2$
 Max drift length



CNGS-Umbria
 L=658 km, 1 deg OA
 CERN SPS 400 GeV
 presently operating 0.3 MW
 (0.5 MW max)
 no near detector infrastructure

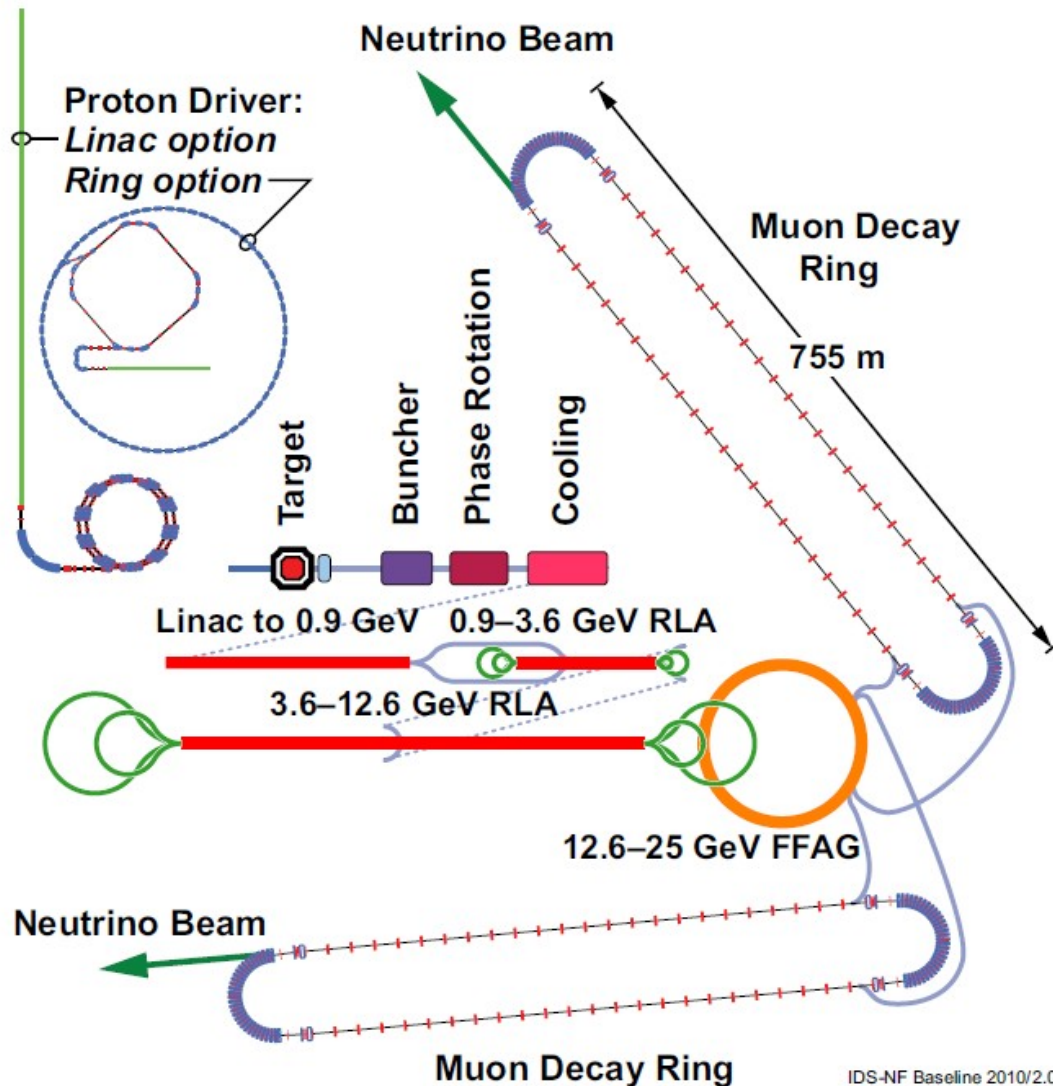
CERN Beta Beams, Synoptic



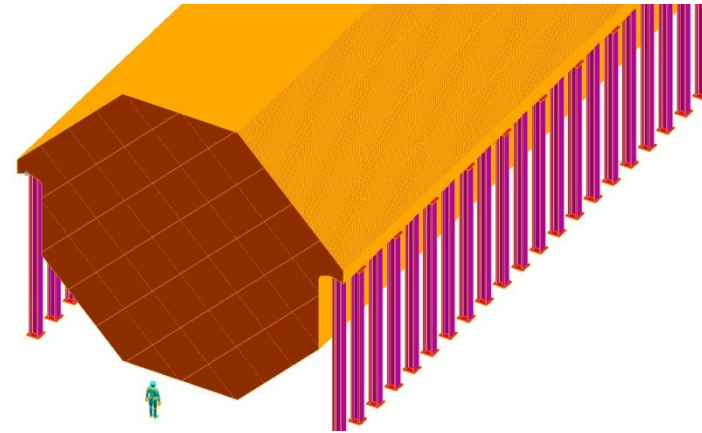
Decay Ring: $B\rho \sim 500 \text{ Tm}$, $B = \sim 6 \text{ T}$, $C = \sim 6900 \text{ m}$, $L_{ss} = \sim 2500 \text{ m}$, $\gamma = 100$, all ions

2011-07-23

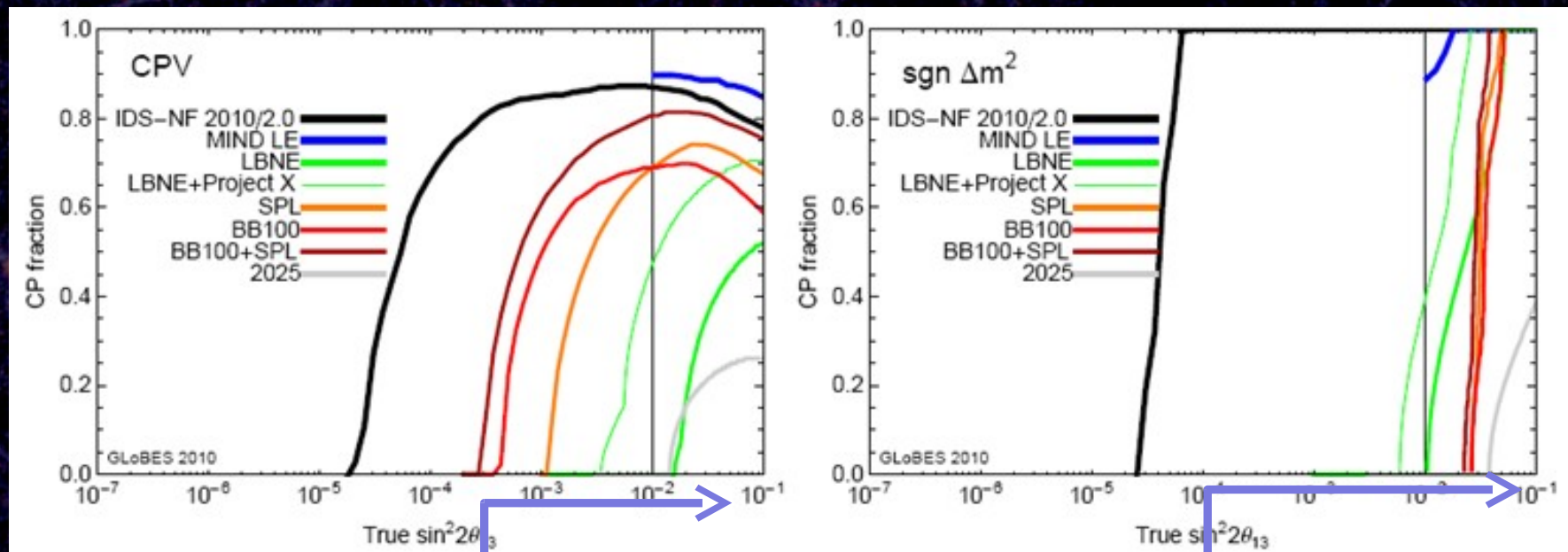
Neutrino Factory Baseline



- Two Magnetised Iron Neutrino Detectors (MIND):
 - 100 kton at 2500-5000 km
 - 50 kton at 7000-8000 km



Baseline constantly under review in light of new physics results



All thinking has been focused on $\sin^2 2\theta_{13} \approx 0.13$

But large θ_{13} phase will quickly change

the CP fraction considered

In the systematics dominated era support measurements are even more

CERN NA61 measurements

Evaluation of Particle Yields in 30 GeV p+C Inelastic Interactions and in the T2K replica target

Large acceptance spectrometer:

- 5 TPCs
- 2 dipole magnets
- $\sigma(p)/p^2 \approx 10^{-4} \text{ (GeV/e)}^{-1}$
- $\sigma(dE/dx)/<dE/dx> \approx 0.04$
- 3 ToFs
- $\sigma(\text{ToF-F}) = 120 \text{ ps}$
- $\sigma(\text{ToF-L/R}) = 60 \text{ ps}$

Full Coverage of T2K phase space

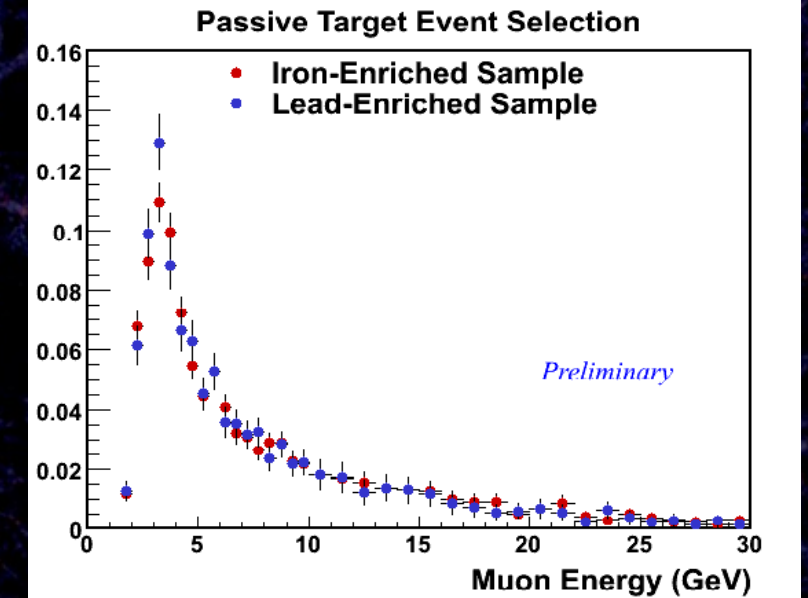
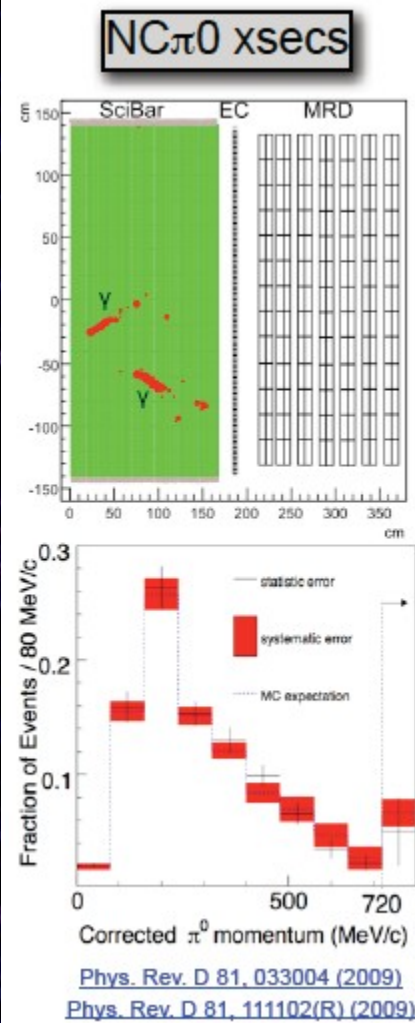
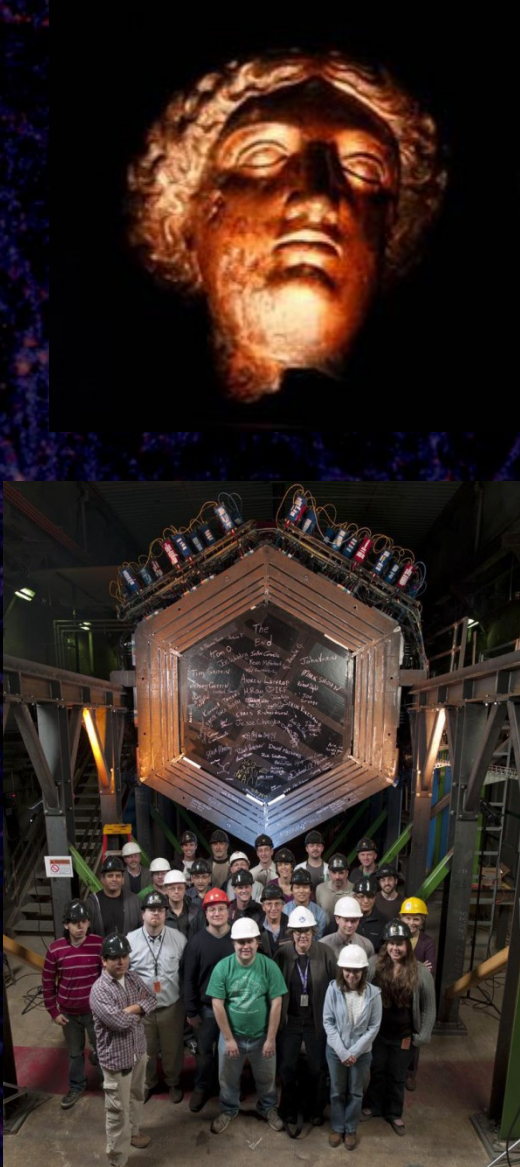
thin target: $2.5 \times 2.5 \times 2 \text{ cm}^3$ int. length ~ 0.04 $\sim 600k$ triggers in 2007

p+C @ 31 GeV/c

Particle ID methods used:

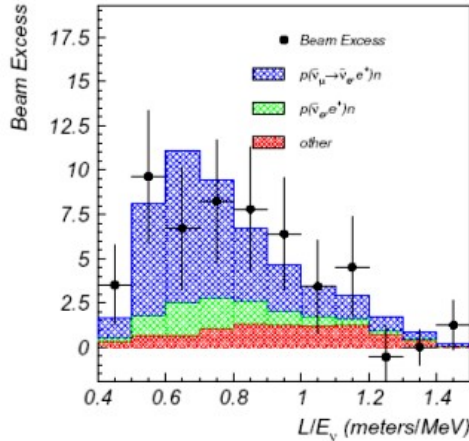
- 1) dE/dx ($p < 1 \text{ GeV/c}$, $p > 4 \text{ GeV/c}$)
- 2) Combined $dE/dx + \text{ToF}$ ($1 < p \text{ [GeV/c]} < 4$)
- 3) Negatively charged hadron h- analysis (π^- only)

A. Rubbia
XIV International Workshop on Neutrino Telescopes (2011)
Wednesday, March 16, 2011



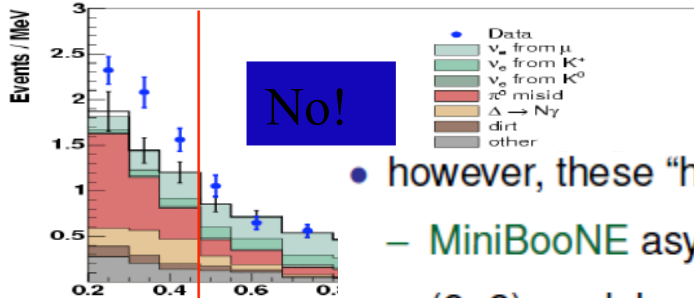
LSND Starts it all...

Short baselines
($L/E \sim 1$)
and sterile ν .

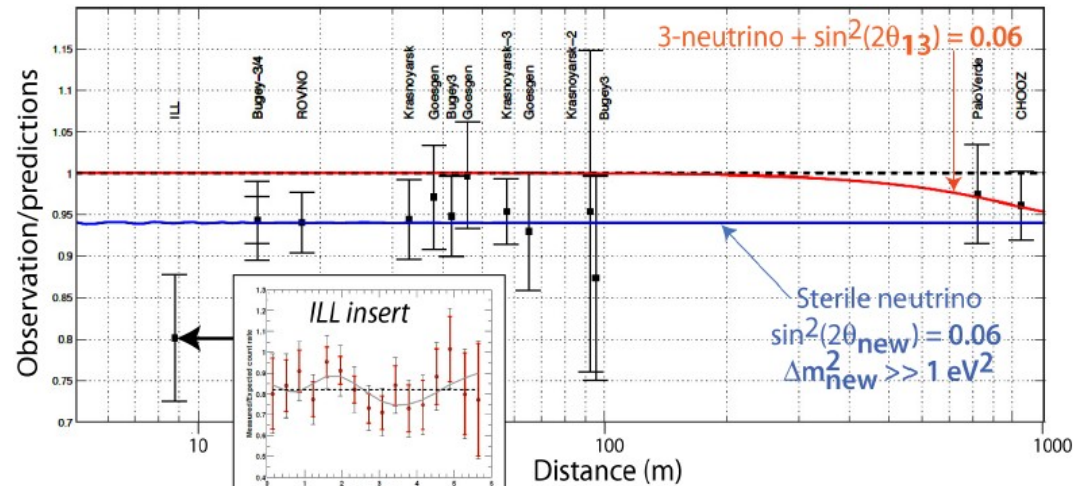
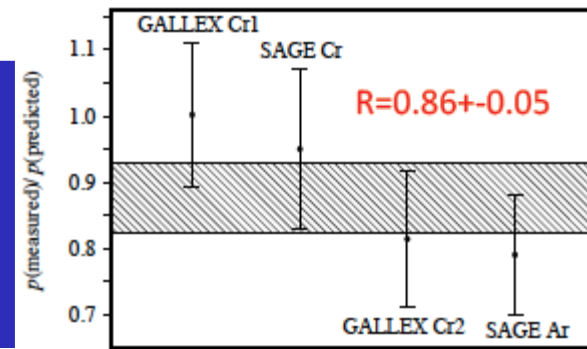
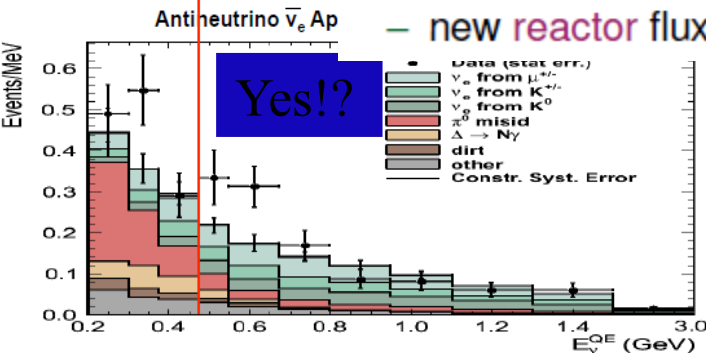


MiniBooNE says....

Neutrino ν_e Appearance Results (6.5E20 POT)



Yes!?



- however, these “hints” for sterile neutrinos are **not** in agreement among them:
 - MiniBooNE asymmetry in $\nu/\bar{\nu}$ requires CP violation, hence at least **two** sterile ν 's;
 - (3+2) models reconcile APP data, but DIS ones still show tension; **Malton**
 - new reactor fluxes reduce tension with DIS data, but not for MB low-E excess;



Conclusions (I)

Neutrino oscillations are the first confirmed physics beyond the SM, and their continued study is essential to extend our knowledge of fundamental interactions.

Current indications are that $\sin^2 2\theta_{13} \geq \sim 0.01$, which would make further long baseline experiments the most attractive option for first searches for CP violation in the neutrino sector.

Do not assume we know everything that is going on – redundancy is essential!

Continued operation of “existing” experiments (T2K, MINOS, NOvA, Double Chooz) is the highest priority.

There are three proposed next-generation projects. We can probably justify two, but not three, so we need to some international coordination.

The mine at Pyhäsalmi is potentially an extremely valuable resource for European neutrino physics due to its distance from CERN, but we should move fast if we are going to retain the option of using it in the future. Can we build a 5 kT LAr prototype?

Conclusions (II)

There will be many other opportunities for smaller-scale involvement in cross-section, hadron production, and other critical technological development projects such as the MICE experiment.

Oscillations depend on L/E , not E or L , so other possibilities should be considered – DAEdALUS?

Beyond those facilities we will almost certainly wish to have an even more capable facility, either a Beta Beam or a Neutrino Factory – more work is needed to optimize the sensitivity if θ_{13} is large.

A PS neutrino beam project to look for sterile neutrinos has been suggested for CERN. Here are my very personal views:

We don't need another 2-3 sigma effect. Any proposed experiment should have clear 5 sigma sensitivity over the entire indicated range.

It could be an interesting part of a broad suite of European neutrino experiments, but I wouldn't want to see it be CERN's only neutrino experiment.

Neutrino oscillations offers a very wide range of experiments, which must happen, no matter what we do at the energy frontier!