

MODIFICATION OF COULOMB LAW AND ENERGY LEVELS OF THE HYDROGEN ATOM IN A SUPERSTRONG MAGNETIC FIELD

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RESULTS

SCALES OF THE PROBLEM

ELECTRON IN THE “COULOMB FIELD OF A PROTON” + EXTREME B

- Resummation → polarization operator
- the adiabatic approximation
- effective potential with screening

THE GROUND LEVEL OF THE HYDROGEN ATOM

- Results
- Principle of the method (Karnakov-Popov)

OUTLOOK AND PROSPECTS

RESULTS

e^- in the “Coulomb” field of a proton (H atom) in an \simeq homogeneous, \simeq constant **ENORMOUS** external magnetic field $B \geq 10^{17} G = 10^{13} T$.

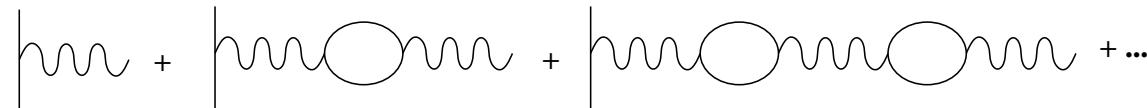


Fig. 1. *Modification of the Coulomb potential due to the dressing of the photon propagator.*

$$\rightarrow \boxed{\Phi(z) \approx \frac{e}{z} \left[1 - e^{-m_e|z|\sqrt{6}} + e^{-m_e|z|\sqrt{6+\frac{2\alpha}{\pi}\frac{B}{B_{cr}}}} \right]}$$

The Coulomb potential becomes “screened” at short distances $\leq 1.5 \frac{1}{m_e}$ when $B \geq \frac{3\pi}{\alpha} B_{cr}$.

→ energy levels of the ground state (Lowest Landau Level) in the Coulomb field of the proton

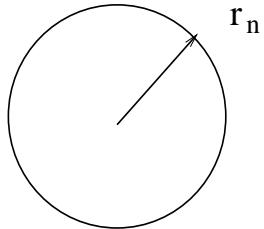
$$\boxed{\ln \left[\frac{\frac{B/B_a}{\alpha^3}}{1 + \frac{3\pi}{\alpha} \frac{B/B_a}{B_a}} \right] = \lambda + 2 \log \lambda + 2\psi \left(1 - \frac{1}{\lambda} \right) + \ln 2 + 4\gamma + \psi(1 + |m|)}$$

$$\lambda = \sqrt{-E_0/E_I}, \quad \boxed{E_I = \frac{\mu e^4}{2} = 13.6 \text{ eV}}$$

- no screening + shallow well approx: $E_0 = -E_I \ln \textcolor{red}{B}/B_a \xrightarrow{B \rightarrow \infty} -\infty$
- screening + shallow well approx: $E_0 \xrightarrow{B \rightarrow \infty} -4E_I \ln^2 \left(\sqrt{3\pi/\alpha^3} \right) \rightarrow -4 \text{ keV};$
- screening and not shallow well: $E_0 \xrightarrow{B \rightarrow \infty} -1.7 \text{ keV}$

SCALES OF THE PROBLEM

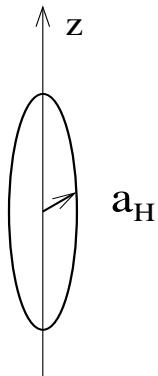
- Bohr's atom (no B)



$$E_n = -E_I/n^2, \quad r_n = n^2 a_B, \quad \boxed{a_B = \text{Bohr radius} = 1/m_e e^2 = 5.29 \cdot 10^{-9} \text{ cm}}$$

$$a_B = \lambda_c/2\pi\alpha = r_e/\alpha^2, \quad \lambda_c/2\pi = 1/m_e = 3.86 \cdot 10^{-11} \text{ cm}, \quad r_e = 2.82 \cdot 10^{-13} \text{ cm}$$

- in the presence of B



$$\boxed{\text{Landau radius } a_H = \sqrt{1/eB}}.$$

Remember:

Larmor frequency $\omega = eB/m_e$,

Larmor radius $= m_e v / eB = (L \equiv m \cdot \hbar) / eB r \Rightarrow r = m \cdot \hbar / eB r \Rightarrow r^2 = m / eB = m a_H^2$

$$\frac{a_B^2}{a_H^2} = \frac{B}{B_a}; \quad B_a = \text{atomic B} = m_e^2 e^3 = 1/e a_B^2 = 2.35 \cdot 10^9 G.$$

We deal with $B \approx 10^{17} G \Rightarrow \boxed{a_H \ll a_B}$. Remember: $B_{\text{earth}} \approx 1 G, \quad B_{\text{LHC}} \approx \text{a few } 10^5 G$

$$B_{cr} \text{ or } B_{Schwinger} = m_e^2 / e = B_a / \alpha^2 = 4.35 \cdot 10^{13} G.$$

For $B \gg B_{cr}$, electromagnetic fields start to interact with each other: non-linearity; e^- becomes relativistic.

For us $B \gg B_{cr} \gg B_a$

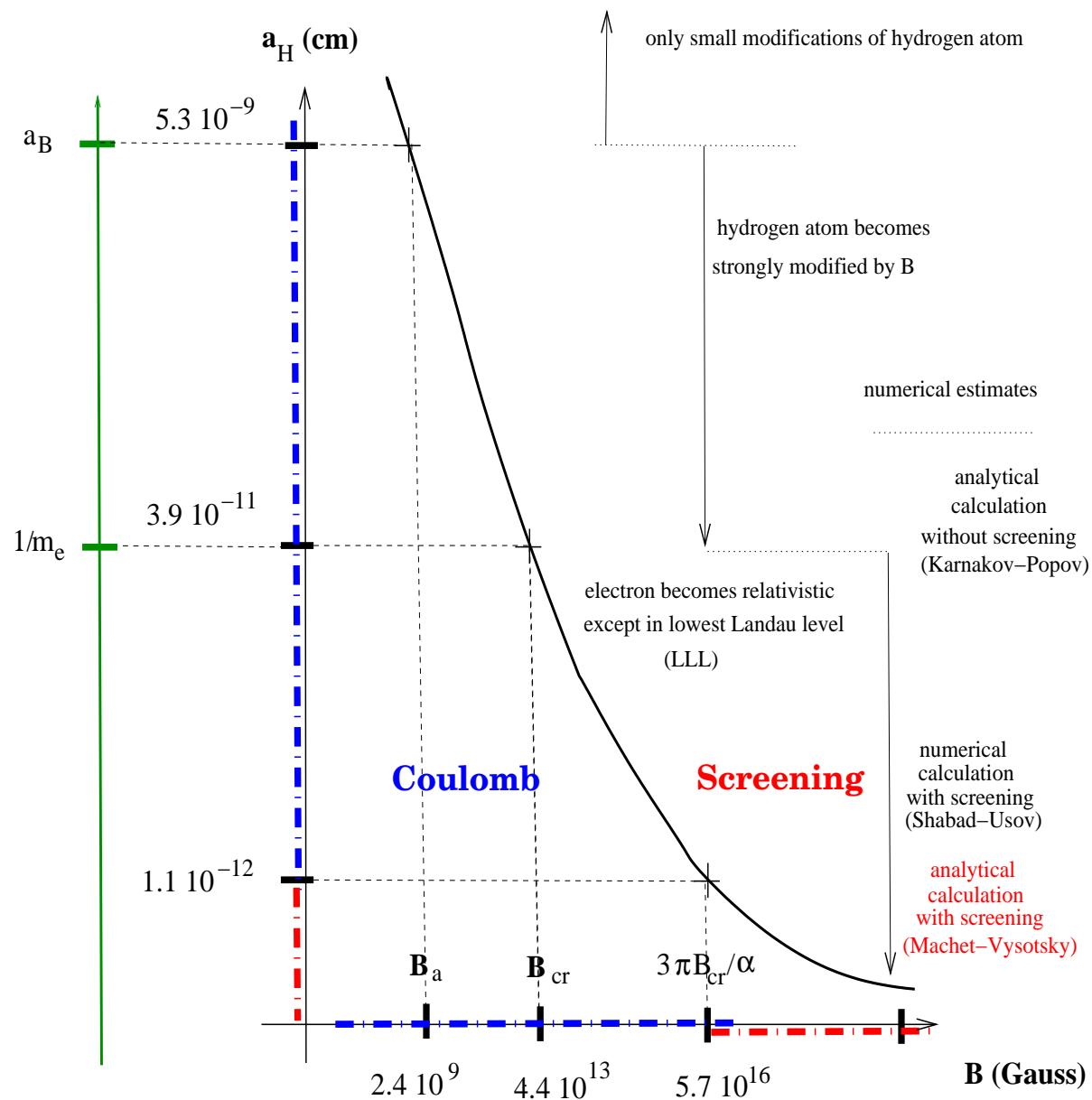


Fig. 2. Landau radius a_H versus magnetic field B .

ELECTRON IN THE “COULOMB FIELD OF A PROTON” + EXTREME B

Very strong $B \rightarrow$ fast oscillations $\perp B$, slow oscillations $\parallel B \Rightarrow$ adiabatic approximation $\psi_{n,n_\rho=0,m,\sigma_3=-1}^{LLL} = R_{0m}(\vec{\rho})\chi_n(z)$

e^- relativistic except in the LLL \rightarrow solve non-relativistic Schrödinger equation for $\chi(z)$ (1-dim) with effective potential U_{eff} incorporating the resummation of Fig. 1.

$$\left[-\frac{1}{2m_e} \frac{d^2}{dz^2} + U_{eff} \right] \chi_n(z) = E_n \chi_n(z)$$

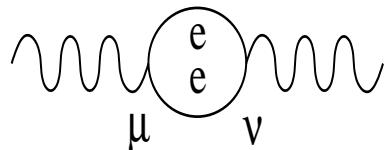
We have to determine what is $U_{eff}(z)$.

Remember; Coulomb: $\Phi_c(z) = e/z$, $U_c(z) = -e^2/z$.

- If no screening, U_{eff} is given by Landau: $U_{eff} = -e^2 \int \frac{|R_{0m}(z)|^2}{\sqrt{z^2 + \rho^2}} d^2\rho; \quad z \gg a_H \rightarrow -e^2/z; \quad z \rightarrow 0 \rightarrow -e^2/a_H$ finite

• Including screening

$$\Phi(k) \stackrel{no\ resum}{=} -4\pi g/k^2 \stackrel{resum}{\rightarrow} \frac{-4\pi g}{k_\parallel^2 + k_\perp^2 + \Pi(k^2)} = \frac{-4\pi g}{k_\parallel^2 + k_\perp^2 + \Pi(k^2)}. \quad \text{Remember: in 4D, } \Pi_{\mu\nu} = -(g_{\mu\nu} - k_\mu k_\nu/k^2)\Pi(k^2).$$



$= \Pi_{\mu\nu}$. Static charge \rightarrow we need only Π_{00} .

$$\Phi(k) = \frac{4\pi g}{k_\parallel^2 + k_\perp^2 + \frac{2\alpha}{\pi} \frac{B}{B_{cr}} m_e^2 e^{\left(\frac{-k_\parallel^2}{2m_e^2} \frac{B_{cr}}{B}\right)} P\left(\frac{k_\parallel^2}{4m_e^2}\right)}$$

$$\Pi(k^2) = -4g^2 P(t), \quad t = \frac{-k^2}{4m_e^2}. \quad P \text{ is the same as in 2 D: } P(t) = 1 - \frac{\ln(\sqrt{1+t} + \sqrt{t})}{\sqrt{t(1+t)}}.$$

Fourier transform $\longrightarrow \Phi(z)$. But exact P too cumbersome to integrate \longrightarrow use approximate $\boxed{\overline{P}(t) = \frac{2t}{3+2t}}$.

$$\boxed{\Phi(z) \approx \frac{e}{z} \left[1 - e^{-m_e|z|\sqrt{6}} + e^{-m_e|z|\sqrt{6+\frac{2\alpha}{\pi}\frac{B}{B_{cr}}}} \right]}$$

* for $B \ll 3\pi B_{cr}/\alpha$, $\Phi(z) = \frac{e}{|z|} (1 + \mathcal{O}(\alpha B/B_{cr}))$

* for $B \gg 3\pi B_{cr}/\alpha$

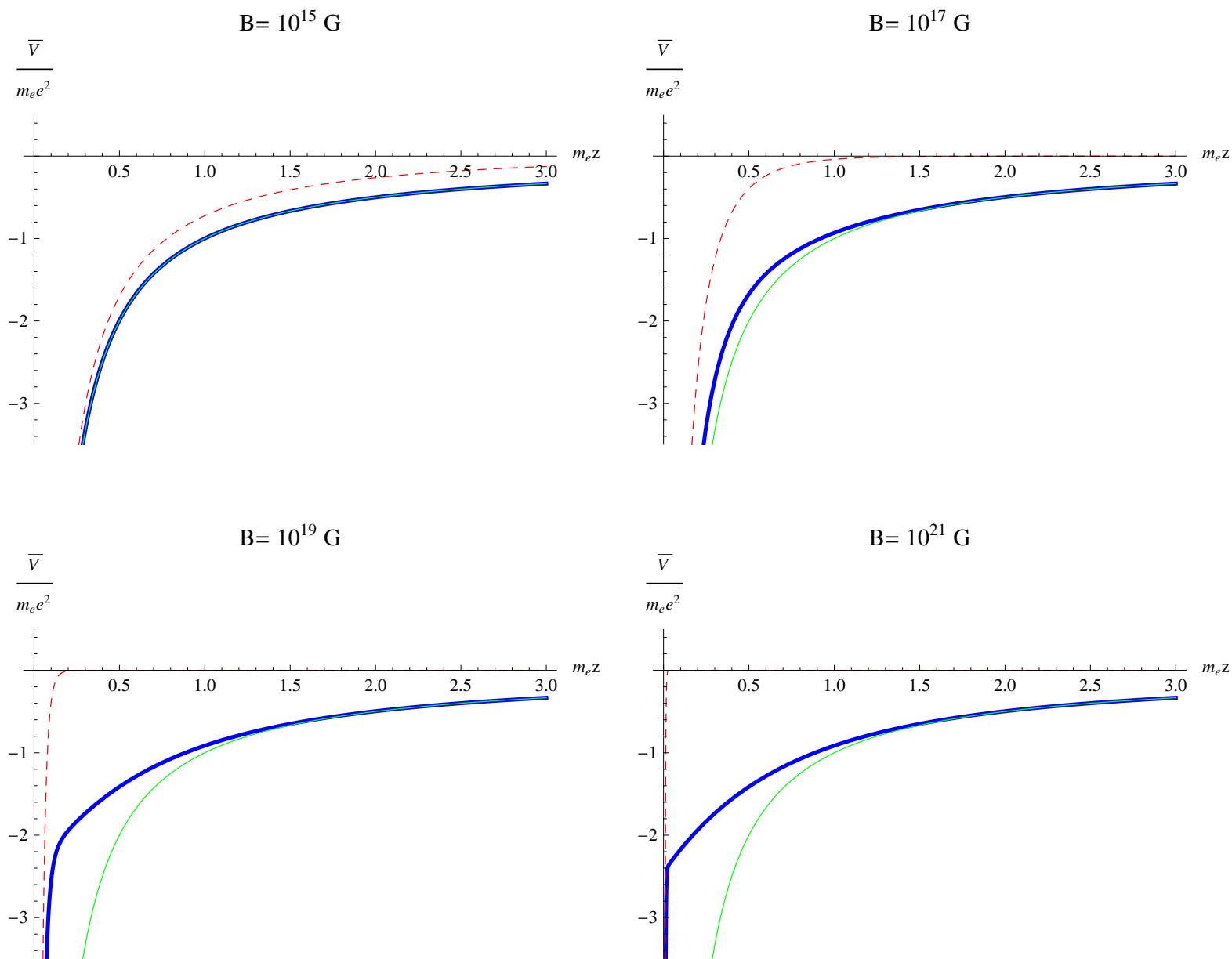
$$\Phi(z) = \begin{cases} \frac{e}{|z|} e^{-m_e|z|\sqrt{\frac{2\alpha}{\pi}\frac{B}{B_{cr}}}} & \text{for } |z| < \frac{1}{\sqrt{\frac{2\alpha}{\pi}\frac{B}{B_{cr}}}} \frac{1}{m_e} \ln(\sqrt{\frac{\alpha}{3\pi}\frac{B}{B_{cr}}}), \quad \text{screening} \\ \frac{e}{|z|} \left(1 - e^{-m_e|z|\sqrt{6}}\right) & \text{for } \frac{1}{m_e} > |z| > \frac{1}{\sqrt{\frac{2\alpha}{\pi}\frac{B}{B_{cr}}}} \frac{1}{m_e} \ln(\sqrt{\frac{\alpha}{3\pi}\frac{B}{B_{cr}}}) \\ \frac{e}{|z|} & \text{for } |z| > \frac{1}{m_e}, \quad \text{Coulomb} \end{cases}$$

$$\boxed{U_{eff}(z) = -e^2 \int d^2\rho \frac{|R_{0m}(\rho)|^2}{\sqrt{z^2 + \rho^2}} \left[1 - e^{-m_e|z|\sqrt{6}} + e^{-m_e|z|\sqrt{6+\frac{2\alpha}{\pi}\frac{B}{B_{cr}}}} \right]}$$

$$R_{0m}(\rho) = \frac{1}{\sqrt{\pi(2a_H^2)^{(1+|m|)}|m|!}} \rho^{|m|} e^{im\rho} e^{-\rho^2/4a_H^2} = \text{wave function of } e^- \text{ in magnetic field (Landau Q.M.)}$$

At short distance, the photon gets an **effective mass** $m_\gamma^2 = \alpha \frac{B}{B_{cr}} m_e^2 = e^3 B$ (2-dim massive QED, Schwinger model).

$$\Phi(z) \longrightarrow \text{potential energy } V(z) = -e\Phi(z)$$



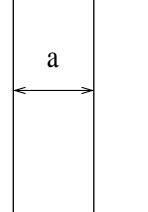
Modified Coulomb potential energy at $B = 10^{17} \text{ G}$ (blue) and its long distance (green-pale) and short distance (red-dashed) asymptotics.

THE GROUND LEVEL OF THE HYDROGEN ATOM

- no screening + shallow well approximation

$$E_0 = -E_I \ln \frac{B}{B_a} \xrightarrow{B \rightarrow \infty} -\infty \text{ ("fall to the center")}$$

- shallow well approximation (Landau) + screening



$$E_0 = -2m_e \left[\int_{a_H}^{a_B} U(z) dz \right]^2 \xrightarrow{B \rightarrow \infty} -2m_e \alpha^2 \ln^2 \left(\sqrt{\frac{3\pi}{\alpha^3}} \right) = -4E_I \times 72.3 = -4 \text{ keV}$$

shallow well approximation valid for $m_e|U|a^2 \ll 1$; for 1-D Coulomb potential: $a \ll a_B$, not realized because Coulomb extends to ∞ .

- general case with screening (Karnakov-Popov)

$$E_0 \xrightarrow{B \rightarrow \infty} -1.7 \text{ keV}$$

- * even/odd degeneracy is broken
- * corrections to odd levels very small
- * corrections to even levels substantial for $B \geq 3\pi \frac{B_{cr}}{\alpha} = 6 \cdot 10^{16} G$

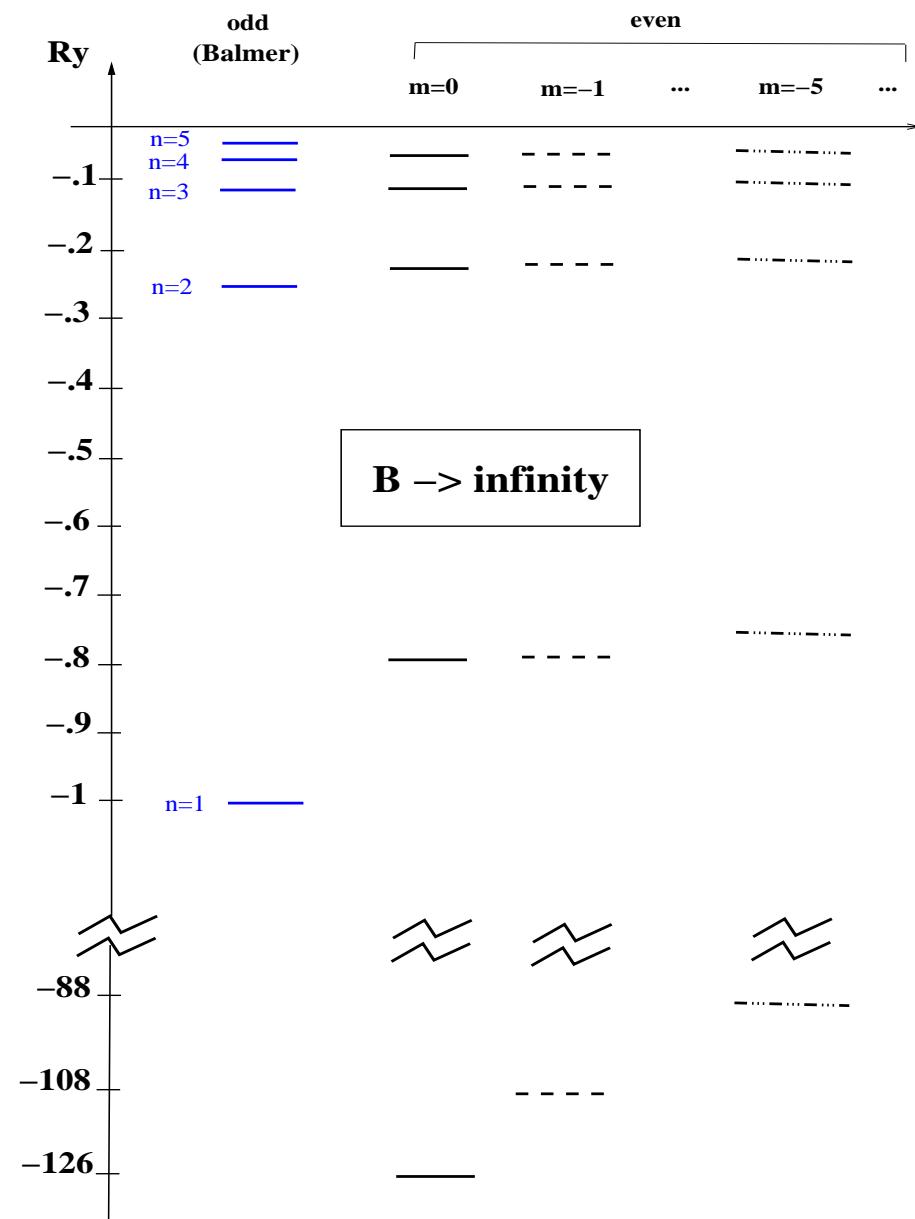


Fig. 5. Spectrum of hydrogen levels in the limit of infinite magnetic field. Energies are given in rydberg units, $Ry \equiv 13.6 \text{ eV}$.

$$\ln \left[\frac{\frac{B/B_a}{\alpha^3}}{1 + \frac{3\pi}{\alpha^3} \frac{B/B_a}{B_a}} \right] = \lambda + 2 \log \lambda + 2\psi \left(1 - \frac{1}{\lambda} \right) + \ln 2 + 4\gamma + \psi(1 + |m|)$$

* **if no screening:** $\Rightarrow \lambda \equiv \sqrt{-E_0/E_I} \xrightarrow{B \rightarrow \infty} \infty$.

When $B \rightarrow \infty$, l.h.s $\rightarrow \infty \Rightarrow \psi(1 - 1/\lambda)$ has poles at $1 - 1/\lambda = 0, -1, -2, \dots - n + 1 \Rightarrow \lambda = 1/n$. One recovers the Balmer series

* **with screening:** when $B \rightarrow \infty$, $\ln \frac{3\pi}{\alpha^3} = r.h.s.$.. No pole \Rightarrow departure from Balmer series.

For $m = 0, B \rightarrow \infty$,

$$\lambda = \frac{1}{1 - \frac{2}{\ln \left[1 + \frac{\frac{B/B_a}{\alpha^3} \frac{B}{B_a}}{1 + \frac{3\pi}{\alpha^3} \frac{B}{B_a}} \right] - \ln 2 - \gamma + 1}}$$

For arbitrary n and $m = 0$,

$$\lambda = \frac{1}{n} - \frac{2/n^2}{\ln \left[1 + \frac{\frac{B/B_a}{\alpha^3} \frac{B}{B_a}}{1 + \frac{3\pi}{\alpha^3} \frac{B}{B_a}} \right] - \ln 2 - \gamma + 1/n + 2 \ln n - 2H_{n-1}}$$

$$H_{n-1} = \text{Harmonic number} = \sum_{k=1}^{n-1} \frac{1}{k}$$

- **Principle of the method (Karnakov-Popov)**

If $\chi(z) \approx cst$, integrate Shrœdinger →

$$\chi'(z) \approx 2m_e \int_0^z dt U(t)\chi(t)$$

(valid for $z \ll a_B$; (self energy negligible)).

At the limit of the screening domain, χ and Whittaker (Coulomb) should match $\Rightarrow \frac{\chi'}{\chi} \stackrel{z > few}{\approx} \frac{W'}{W}$

- **Validity** of K.P. method : $\chi(z) \approx cst$. It is achieved, in particular thanks to the screening.

OUTLOOK AND PROSPECTS

- accuracy of adiabatic approximation
- 3D potential $V(z, \vec{\rho} \neq 0)$
- beyond LLL; relativistic e^- (not very physical)
- effective mass of $e^- \longrightarrow E_0 = m_e \left[1 - \frac{e^4}{2} \lambda^2 + \frac{e^2}{4\pi} \ln^2 \left(\frac{eB}{m_e^2} \right) \right]$; corrections can be large but are universal, they cancel in transitions
- trace levels when B goes from 0 to ∞ (Zeldovich effect)
- string-like behavior when $B \rightarrow \infty$ (Kondo, Shabad-Usov)
- exciton physics in semi-conductors. $B_a \rightarrow 2000 G$ for InSb $\Rightarrow B_{cr} \sim \frac{B_a}{\alpha^2} \approx 10^7 G$
- graphene : 2D QED for massless fermions. $c \rightarrow v_f \ll c \Rightarrow \alpha \rightarrow \frac{e^2}{v_f} \gg 1/137$.
Same physics at much smaller B ?