



2011 Moriond - Electroweak Session

**New approach to antineutrino from
muon decay at rest**

Sanjib Kumar Agarwalla

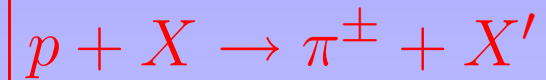
Sanjib.Agarwalla@ific.uv.es

IFIC/CSIC

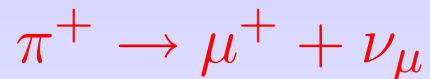
University of Valencia, Spain

Neutrinos at Spallation Source

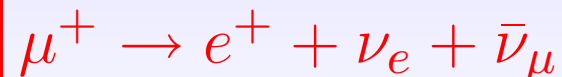
- Interaction of a few GeV protons with the target produces charged pions in addition to the spallation neutrons



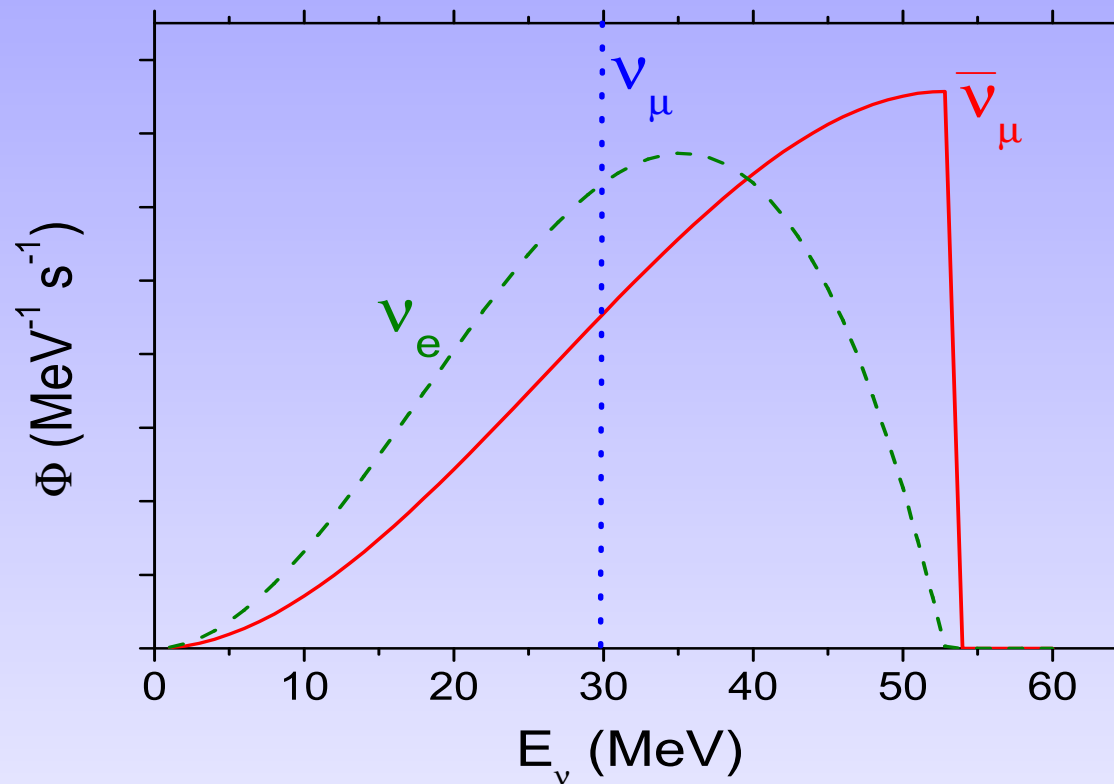
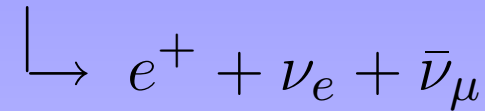
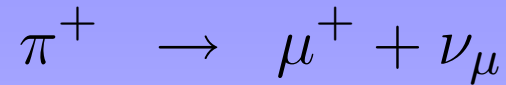
- π^{-} & daughter μ^{-} are absorbed in thick, high-Z target
- π^{+} are brought to rest in a high-Z beam stop
- π^{+} decay produces mono-energetic 29.8 MeV ν_{μ} & μ^{+}



- μ^{+} decays at rest, providing Michel spectrum



Decay At Rest (DAR) Source



Provides an equal, high-intensity, isotropic, DAR ν_μ , ν_e and $\bar{\nu}_\mu$ beam with tiny $\bar{\nu}_e$ contamination ($\sim 10^{-4}$ to 10^{-5})

DAR Source Facilities

Facility	Power	Proton energy	Time structure	Repetition rate
LAMPF (USA)	56 kW	0.8 GeV	Continuous	N/A
ISIS (UK)	160 kW	0.8 GeV	200 ns	50 Hz
DAE δ ALUS (USA)	1 - 5 MW	0.8 GeV	Continuous	N/A
SNS (USA)	> 1 MW	1 GeV	380 ns	60 Hz
JSNS (Japan)	1 MW	3 GeV	1 μ s	25 Hz
SPL (CERN)	4 MW	3.5 GeV	0.76 ms	50 Hz
ESS (Sweden)	5 MW	1.3 GeV	2 ms (1.4 μ s)	17 Hz (50 Hz)

Past, present, and future DAR Source Facilities in different regions of the world

R. Lazauskas, C. Volpe, arXiv:1004.0310v1 [hep-ph]

- LSND (167 tons) at 30 m from LAMPF in Los Alamos
- KARMEN (65 tons) at 17.6 m from ISIS at RAL

Proton Source

- Main technological challenge is the production of a sufficiently intense proton beam with few GeV energy
- Superconducting Linacs : clean & proven technology, can provide multi-MW beams, very expensive & size is large
- Conrad and Shaevitz suggested that new low-cost, high-power proton cyclotrons can deliver multi-MW beams at a cost of 5% of a conventional proton accelerator

Conrad & Shaevitz, PRL 104:141802 (2010)

- DAE δ ALUS collaboration performing an extensive R&D for a new DAR neutrino source for large Gd-doped water Cherenkov detector at Homestake mine aka DUSEL

Expression of Interest: [arXiv:1006.0260](https://arxiv.org/abs/1006.0260) [physics.ins-det]

Physics Goals

Assume that a multi-MW proton beam is affordable

Key for high-intensity DAR neutrino source

- Test anomalies: LSND, MiniBooNE, Gallium...
 - *High Δm^2 oscillation ($L \sim 20 - 50$ m)*
- Study CP violation in active neutrinos
 - *Oscillation at atm. Δm^2 ($L \sim 20$ km)*
- EW precision physics ($\sin^2 \theta_W$ measurement)
- Searches for non-standard neutrino interactions
- Coherent ν -nucleus scattering (^{12}C , ^{16}O , ^{40}Ar)
 - *Supernova ν spectrum overlaps with DAR ν*
 - *Calibration of ν detectors for supernova*

Where do we stand today?

LSND : $L = 30 \text{ m}$, $\langle E_{\nu_{\bar{\mu}}} \rangle = 40 \text{ MeV}$

- 3.8 σ excess of $\bar{\nu}_e$ events in a beam of $\bar{\nu}_{\mu}$

MiniBooNE : $L = 541 \text{ m}$, $\langle E_{\nu_{\mu}, \nu_{\bar{\mu}}} \rangle = 700 \text{ MeV}$

- No oscillation in the ν -mode for energies above 475 MeV
- An unexplained 3 σ excess of ν_e events in the ν -mode below 475 MeV
- A 2.8 σ excess of $\bar{\nu}_e$ events in the anti-neutrino mode above 475 MeV, consistent with LSND
- L/E ratio very similar between LSND and MiniBooNE
- Both points to a Δm^2 of the order $0.1 - 10 \text{ eV}^2$

Testing LSND, MiniBooNE..

MiniBooNE is not conclusive w.r.t. LSND

Simple increase in statistics would not resolve the issue

How to address this problem?

We suggest to perform a modern version of LSND

LSND reloaded

SKA, Patrick Huber, arXiv:1007.3228

DAR + Super-Kamiokande (Gd)

- Place DAR source 20m away from the surface of SK, should provide sufficient shielding against neutrons
- $\bar{\nu}_\mu$ from DAR source will oscillate into $\bar{\nu}_e$ & will be detected via inverse beta decay in Gd doped SK

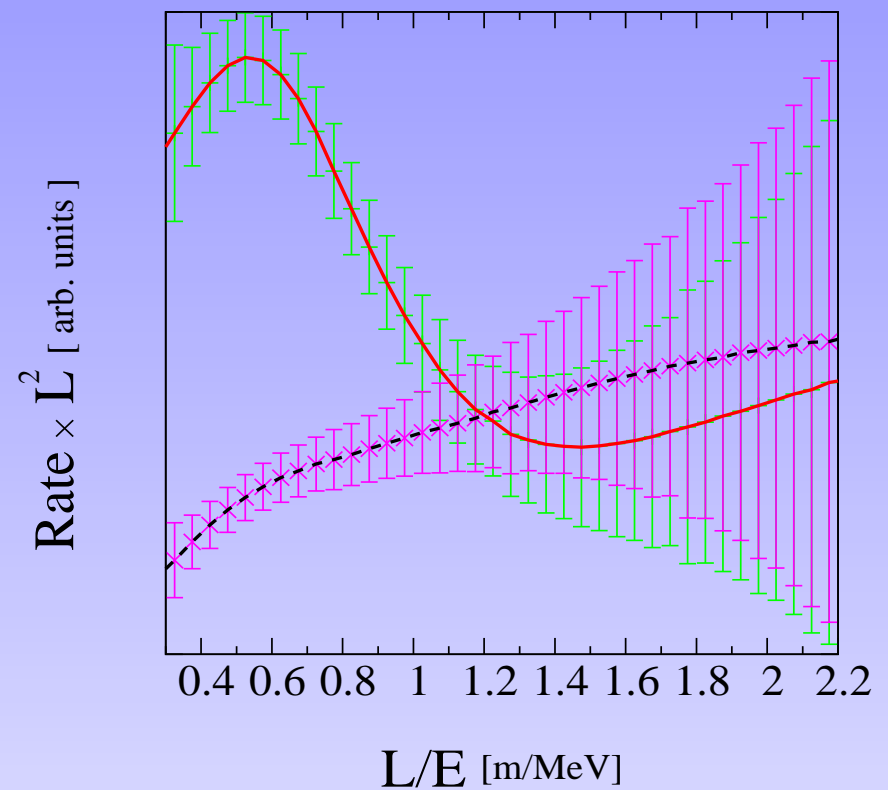
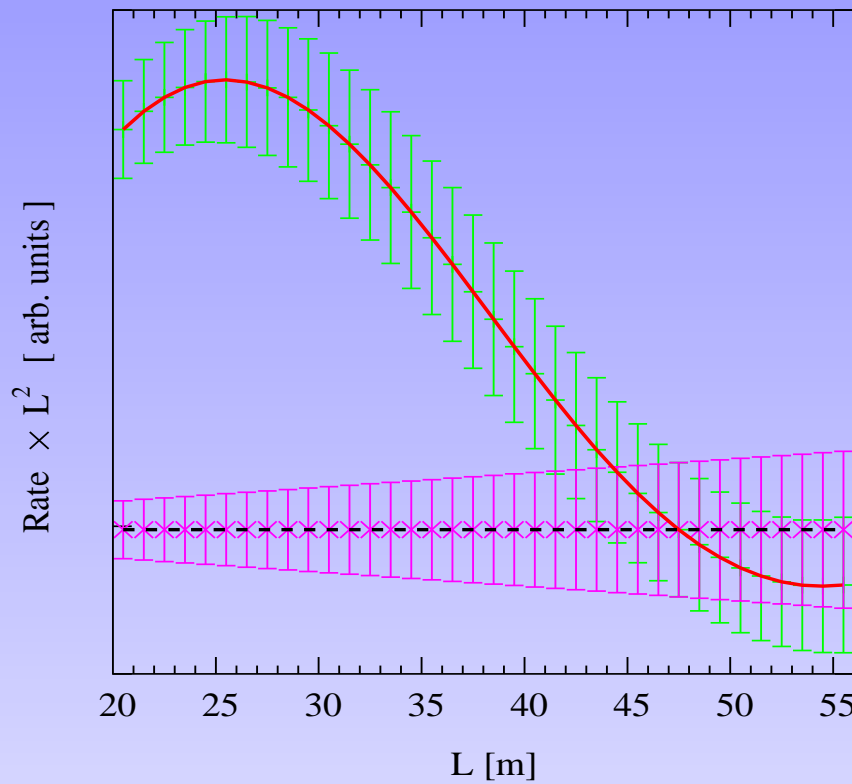


- Useful coincidence tag between the prompt positron and the delayed neutron capture in a Gd doped WC detector
- Event rates will be very high in SK due to large fiducial mass of 22.5 kt .vs. 167 t in LSND
- 300 kW proton beam power is sufficient to have 4×10^{21} $\bar{\nu}_\mu$ per year from each cyclotron
- Intrinsic $\bar{\nu}_e$ contamination small, 4×10^{-4} (LSND value)

Gd doped Super-Kamiokande

- Gd doping allows efficient detection of neutron capture ($\Delta T \sim 30 \mu s$, 8 MeV γ). 67% IBD detection efficiency
- SK is deep (2,700 mwe) and thus has only very small cosmic background rate compared to LSND (120 mwe)
- In SK, the large dimensions of the fiducial volume a cylinder of 14 m radius with a height of 36 m allow to observe the characteristic L dependence of oscillation
- Beam stop size 0.5 m & vertex resolution in SK 0.75 m added in quadrature yields $\Delta x = 0.9$ m. We use 1 m bins
- Energy range: 20 - 52 MeV with $\sigma(E) = 50\% \sqrt{E/MeV}$
- Baseline range: 20 - 56 m
- 38 equally size L/E bins ranging 0.4 - 2.8 m MeV⁻¹

L & L/E pattern in SK

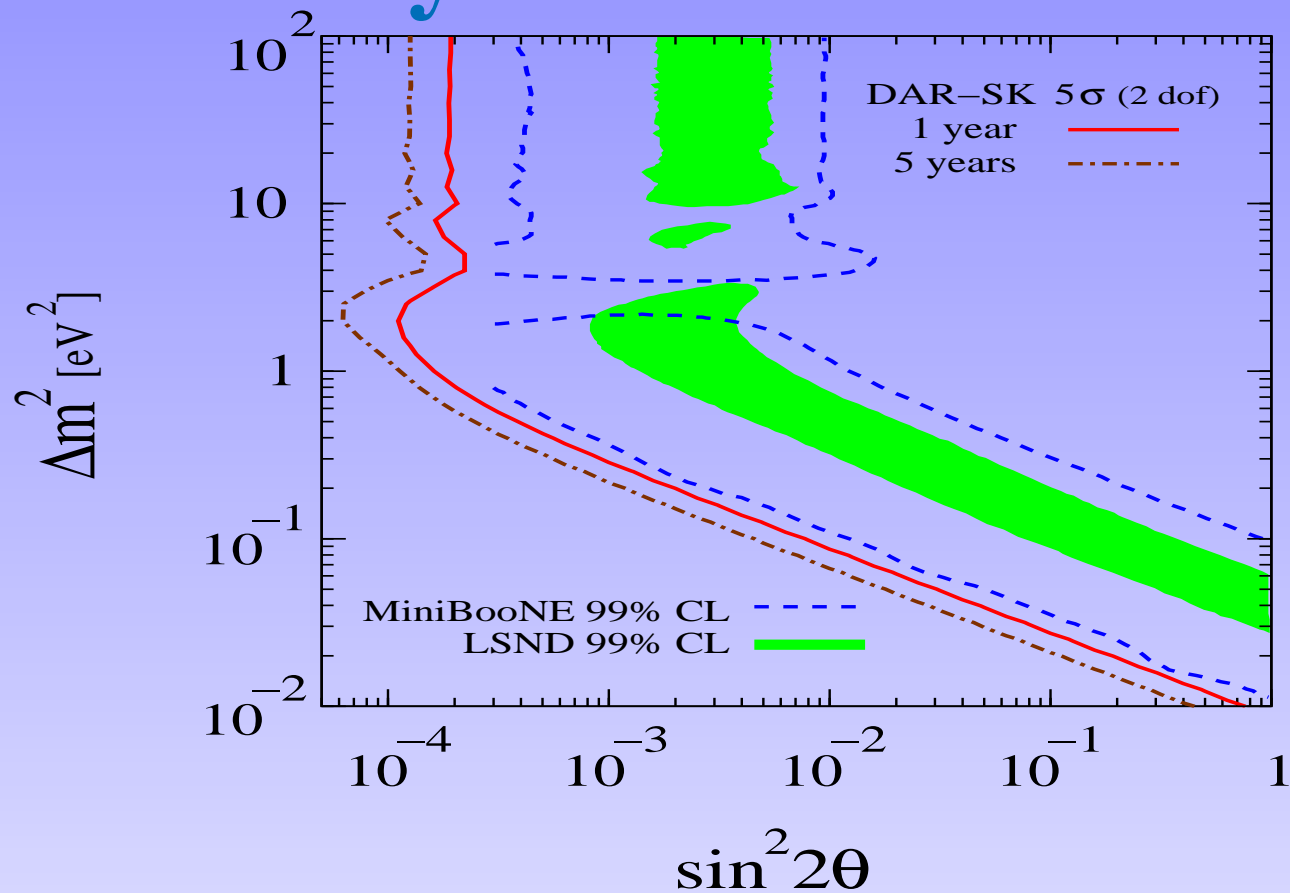


2 flav: $\sin^2 2\theta = 10^{-3}$, $\Delta m^2 = 2 \text{ eV}^2$, $L^{osc} = 49.6 \text{ m}$ (40 MeV)

● Black: Background (765), Red: Signal (1856 events)

● L/E dependence: powerful handle to reject background and to cancel systematic errors among various bins

Sensitivity to Sterile neutrinos



- We cover the whole 99% CL region from both MiniBooNE and LSND at more than 5 σ CL in one year
- If a signal will be seen, we will know whether it is oscillation or something else from L-dependence

It will return the final verdict on LSND

Leptonic CP violation

The discovery of CP violation in the lepton sector would constitute a smoking gun for leptogenesis to explain the baryon asymmetry of the Universe

Can we study CPV using DAR neutrino sources?

The effort of DAE δ ALUS collaboration

arXiv:0912.4079, 1006.0260, 1008.4967, 1012.4853

LBNE and DAR neutrino source

SKA, Huber, Link, Mohapatra, arXiv:1005.4055

$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ in vacuum

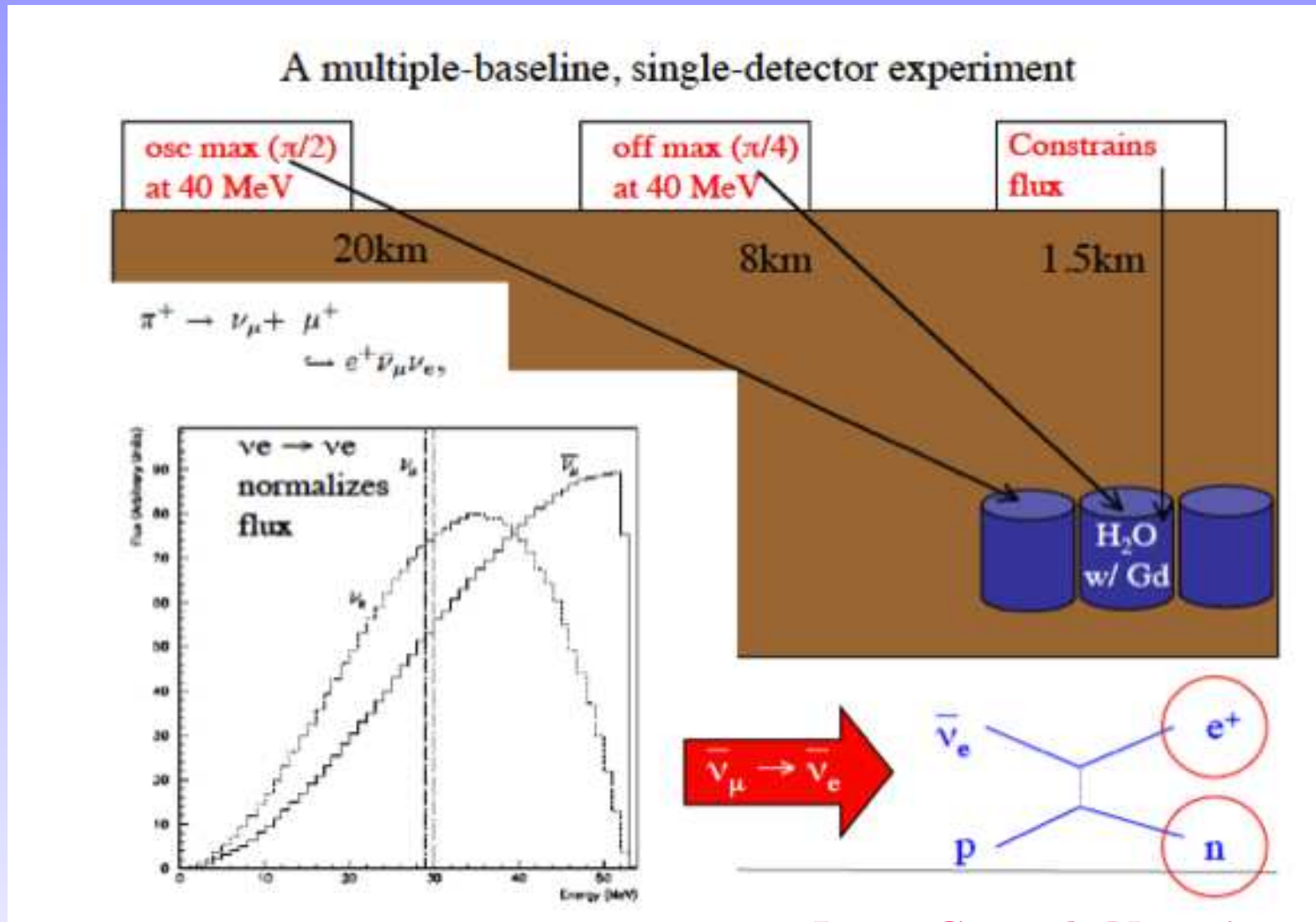
$$P = \begin{aligned} & \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \Delta_{13} \\ & + \sin \delta_{CP} \sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12} \sin^2 \Delta_{13} \sin \Delta_{12} \\ & + \cos \delta_{CP} \sin 2\theta_{13} \sin 2\theta_{23} \sin 2\theta_{12} \sin \Delta_{13} \cos \Delta_{13} \sin \Delta_{12} \\ & + \cos^2 \theta_{23} \sin^2 2\theta_{12} \sin^2 \Delta_{12} \end{aligned}$$

where $\Delta_{ij} = \Delta m_{ij}^2 L / 4E_\nu$

- Want to measure δ_{CP} but θ_{13} and $\text{sgn}(\Delta m_{31}^2)$ not known
- $1.27 \Delta m_{31}^2 \frac{L}{E} \sim \frac{\pi}{2}$ at $E = 40$ MeV, $L = 20$ km (Osc. max)
- $1.27 \Delta m_{31}^2 \frac{L}{E} \sim \frac{\pi}{4}$ at $E = 40$ MeV, $L = 10$ km (Off max)
- $1.27 \Delta m_{31}^2 \frac{L}{E} \sim 0$ at $E = 40$ MeV, $L = 1.5$ km (flux)

Use the L/E dependence of the interference terms

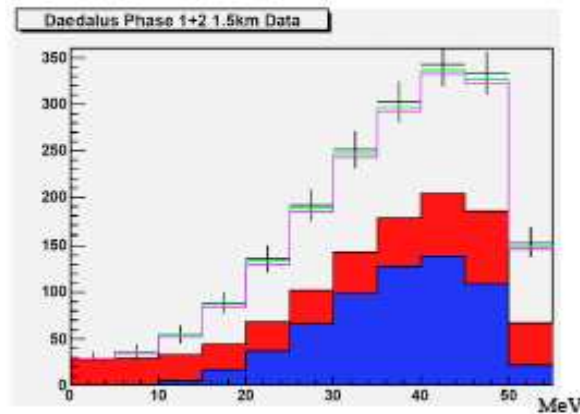
DAE δ ALUS setup



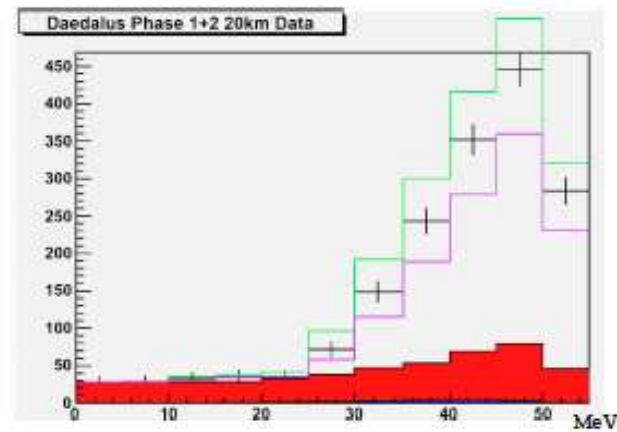
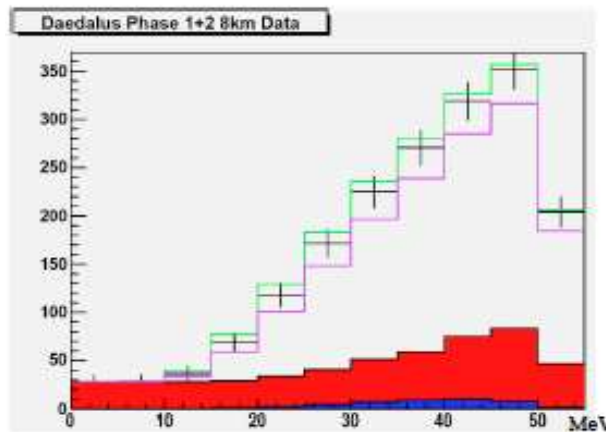
Janet Conrad, Neutrino 2010

DAR $\bar{\nu}_\mu$ beam + 300 kt Gd-doped water detector @ DUSEL
 No. of Cyclotrons: 10 (20 km), 4 (8 km), 2 (1.5 km)

Event .vs. Background

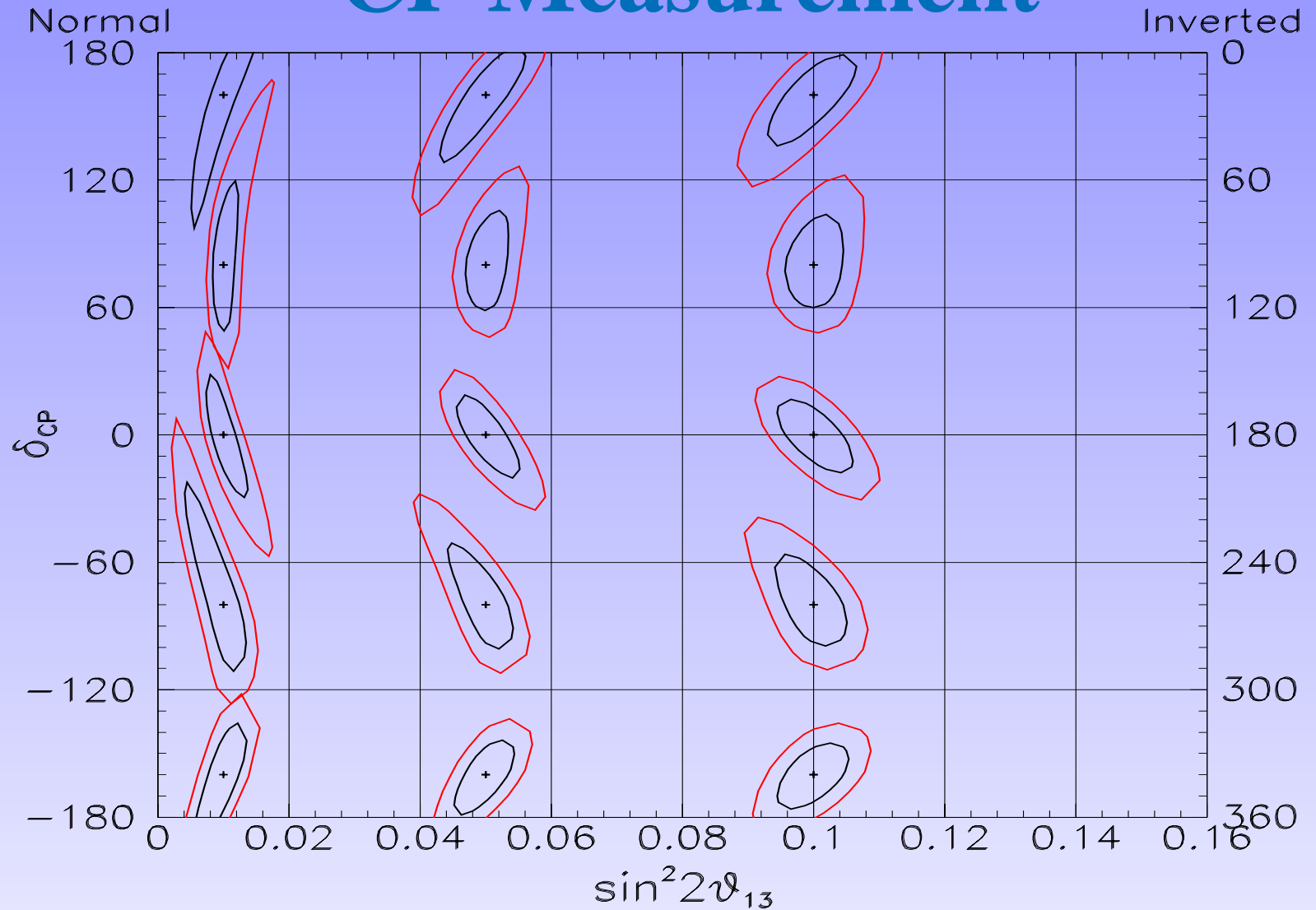


Blue: Intrinsic ν_e bkgnd
Red: Beam off bkgnd
Black: $\delta_{CP} = 0^\circ$
Violet: $\delta_{CP} = 45^\circ$
Green: $\delta_{CP} = -45^\circ$



10 years run, NH, $\sin^2 2\theta_{13} = 0.04$, no matter effects
Signal to Background ratio is good at L = 8 and 20 km

CP Measurement



Inner (outer) region is 1σ (2σ) assuming hierarchy is known

$[\delta_{cp} = x^\circ, \text{NH}]$ is equivalent to $[\delta_{cp} = (180 - x)^\circ, \text{IH}]$

DAE δ ALUS + LBNE

What would happen if we combine
DAE δ ALUS + LBNE experiment?

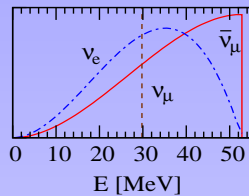
Why should we combine them?

Long Baseline Neutrino Experiment

- New, upgraded, intense wide band beam from FNAL towards 300 kt water Cerenkov detector at Homestake mine aka DUSEL at a distance of 1300 km
- LBNE has to run conventional neutrino (HFN) and anti-neutrino (HFA) beams based on the decay of horn focused pions to disentangle CP and matter effects
- Anti-neutrino run in a superbeam is difficult due to lower anti-neutrino production, lower detection cross-sections and large neutrino contamination in the anti-neutrino beam

New Approach

New Approach : DAR + HFN
LBNE : HFA + HFN



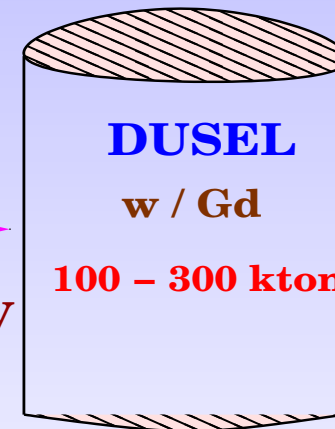
4 Proton Accelerators
1 MW, 1 GeV
Low-cost, High-power

DAR

$4 \cdot 10^{22}$ neutrinos/flux/year/accelerator

DAR : $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
20 - 55 MeV

20 km IBD



FNAL

1300 km CC QE

HFN : $\nu_\mu \rightarrow \nu_e$

0.5 - 6 GeV

HFA : $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$

Wide Band Beam
700 kW, 120 GeV

$6.2 \cdot 10^{20}$ protons/year

SKA, Huber, Link, Mohapatra, arXiv:1005.4055

Signal .vs. Background

	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	Bkg	$\nu_\mu \rightarrow \nu_e$	Bkg
DAR+HFN	1194	217	1532	428
HFA+HFN	231	158	766	214

$\sin^2 2\theta_{13} = 0.1$, $\delta_{\text{CP}} = 0^\circ$, NH and total 6 years run

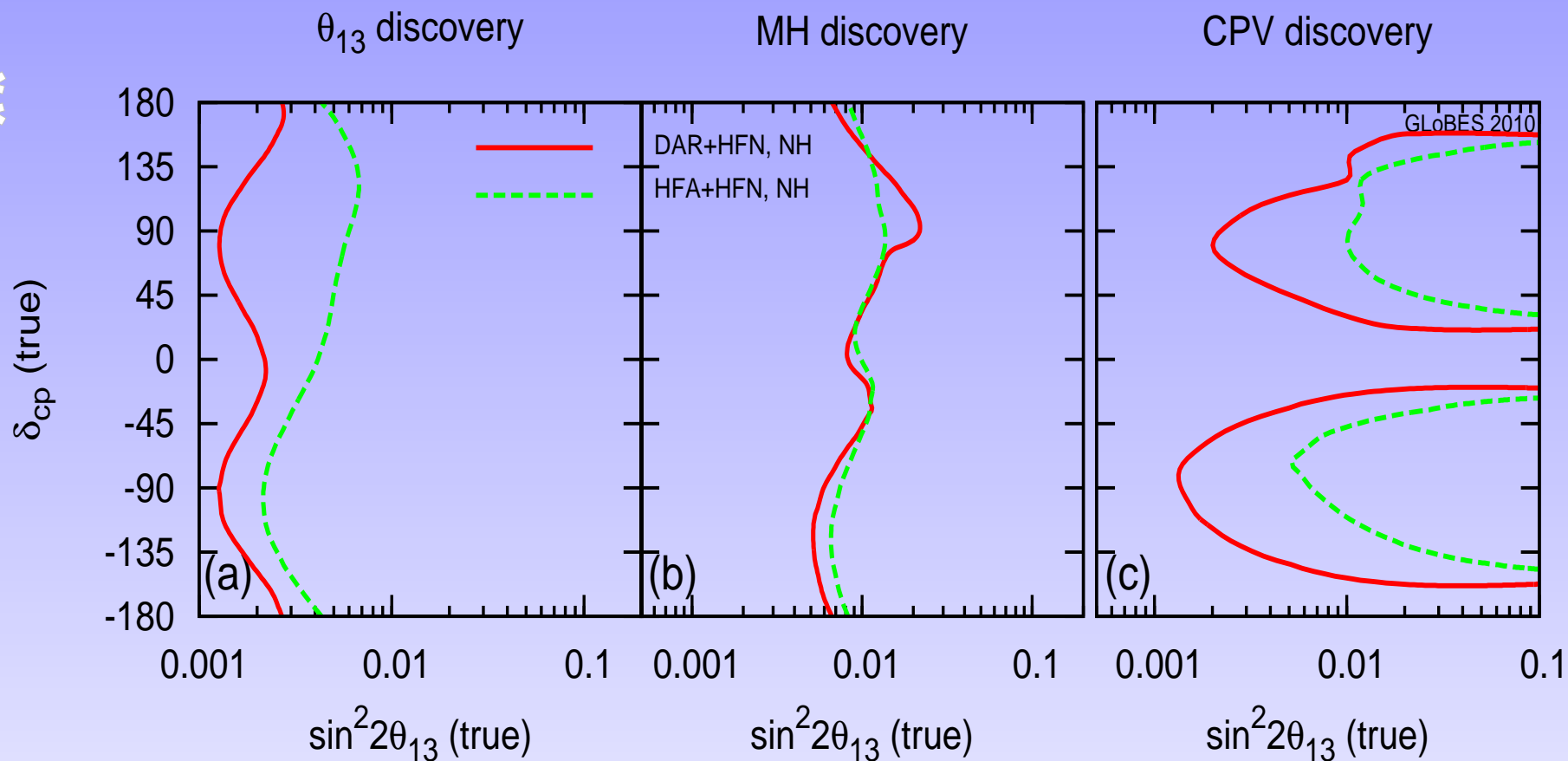
DAR+HFN: 6 years of *simultaneous* running of neutrinos and anti-neutrinos

HFA+HFN: 3 years of *consecutive* running of neutrinos and anti-neutrinos

DAR+HFN: *twice* the statistics in the neutrino mode and *five times* as much statistics with a five times better sig/bkg ratio in the anti-neutrino mode

SKA, Huber, Link, Mohapatra, arXiv:1005.4055

300 kton & NH



SKA, Huber, Link, Mohapatra, arXiv:1005.4055

Discovery reaches at 3σ confidence level

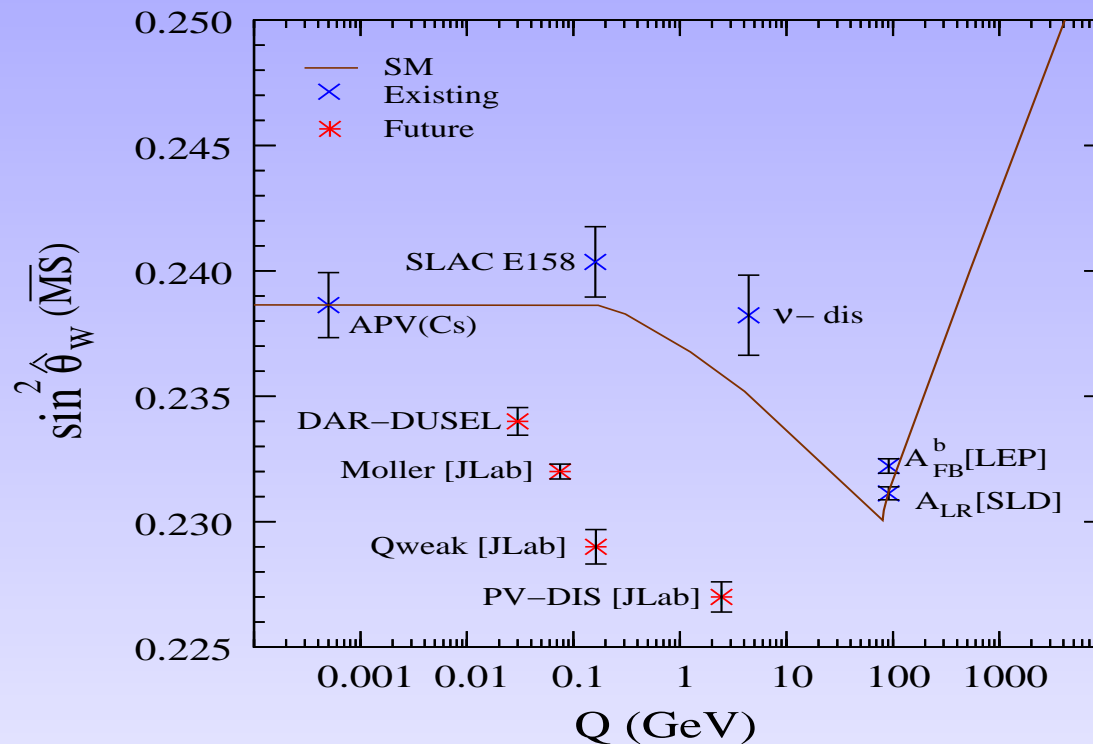
Concluding Remarks

- A stopped pion source provides neutrino beams with energy of a few tens of MeV from pion and muon decay-at-rest
- A rich physics program can be accomplished with such a neutrino source
- Low-cost, multi-MW proton cyclotrons needed
- 300kW proton beam + Gd-doped SK can test LSND, MiniBooNE anomalies in one year at more than 5σ CL
- DAR neutrino sources together with LBNE can accelerate the discovery reach of CP violation

For applications in EW precision physics, see
SKA, Patrick Huber, arXiv:1005.1254

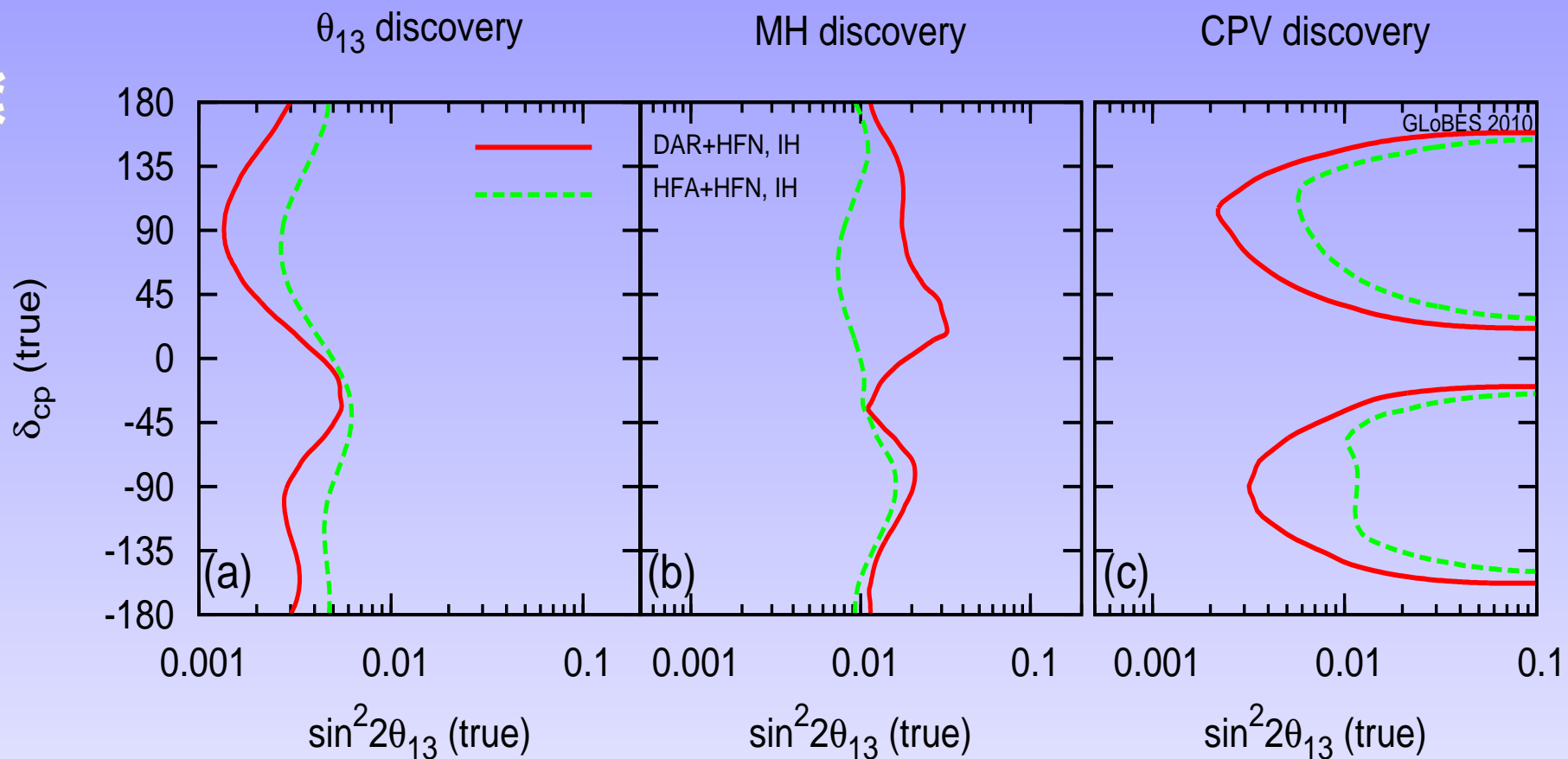
DAR-DUSEL

Our proposed experiment will provide a $\simeq 0.24\%$ measurement of $\sin^2 \theta_W$



This configuration can be a natural part of the proposed physics program for DUSEL

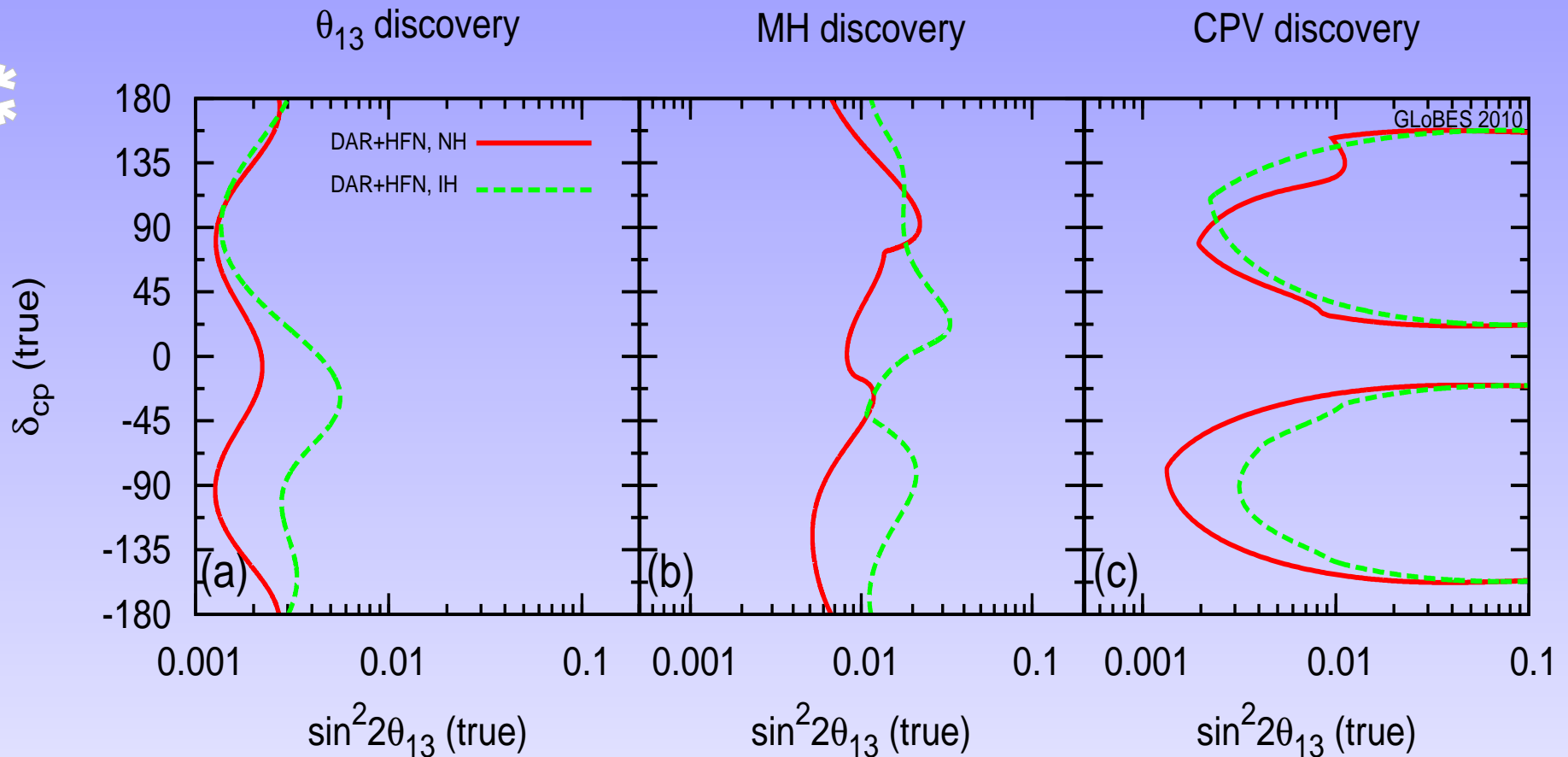
300 kton & IH



Agarwalla, Huber, Link, Mohapatra, arXiv:1005.4055 [hep-ph]

Discovery reaches at 3σ confidence level

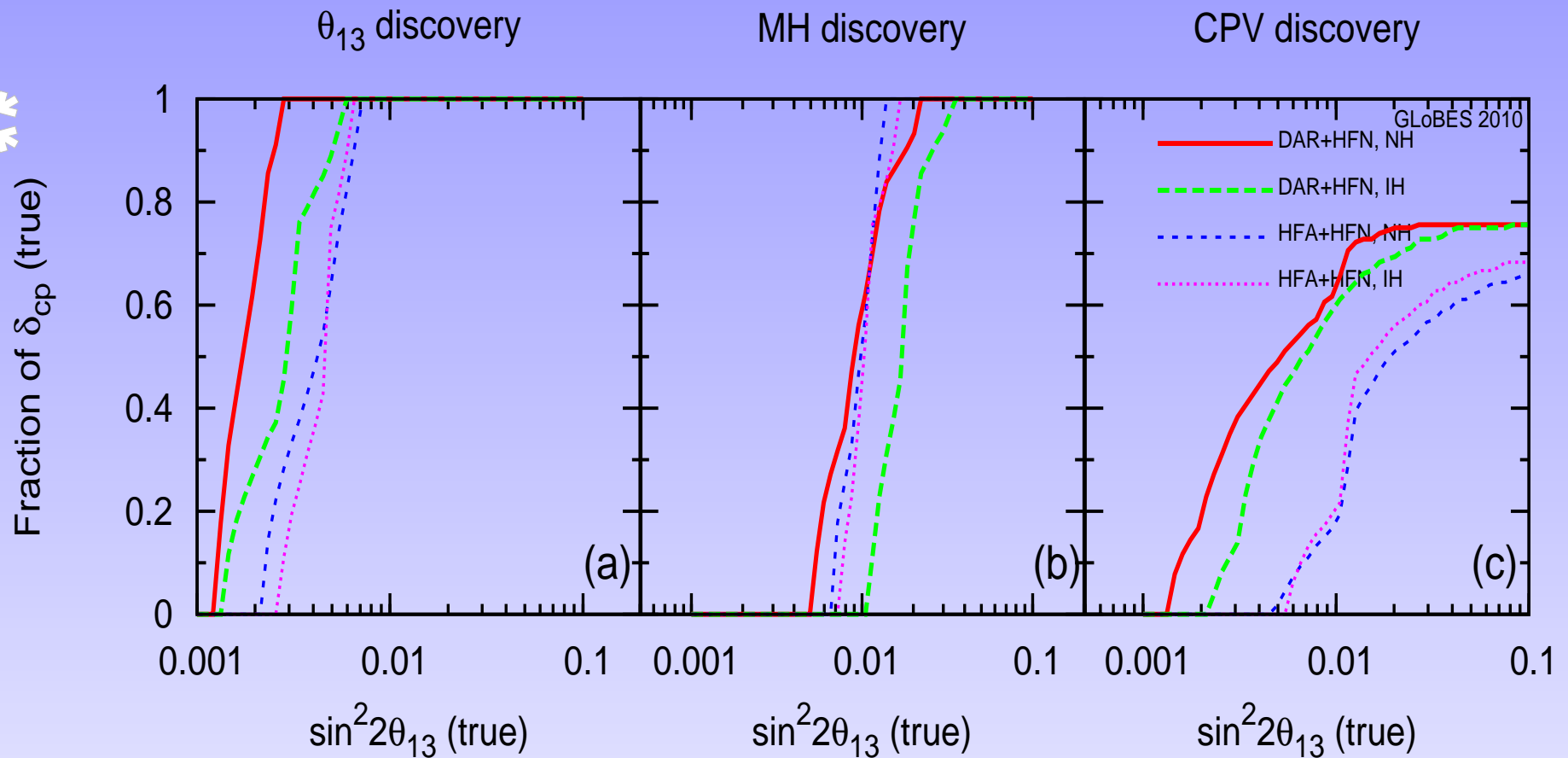
300 kton, NH & IH



Agarwalla, Huber, Link, Mohapatra, arXiv:1005.4055 [hep-ph]

Discovery reaches at 3σ confidence level

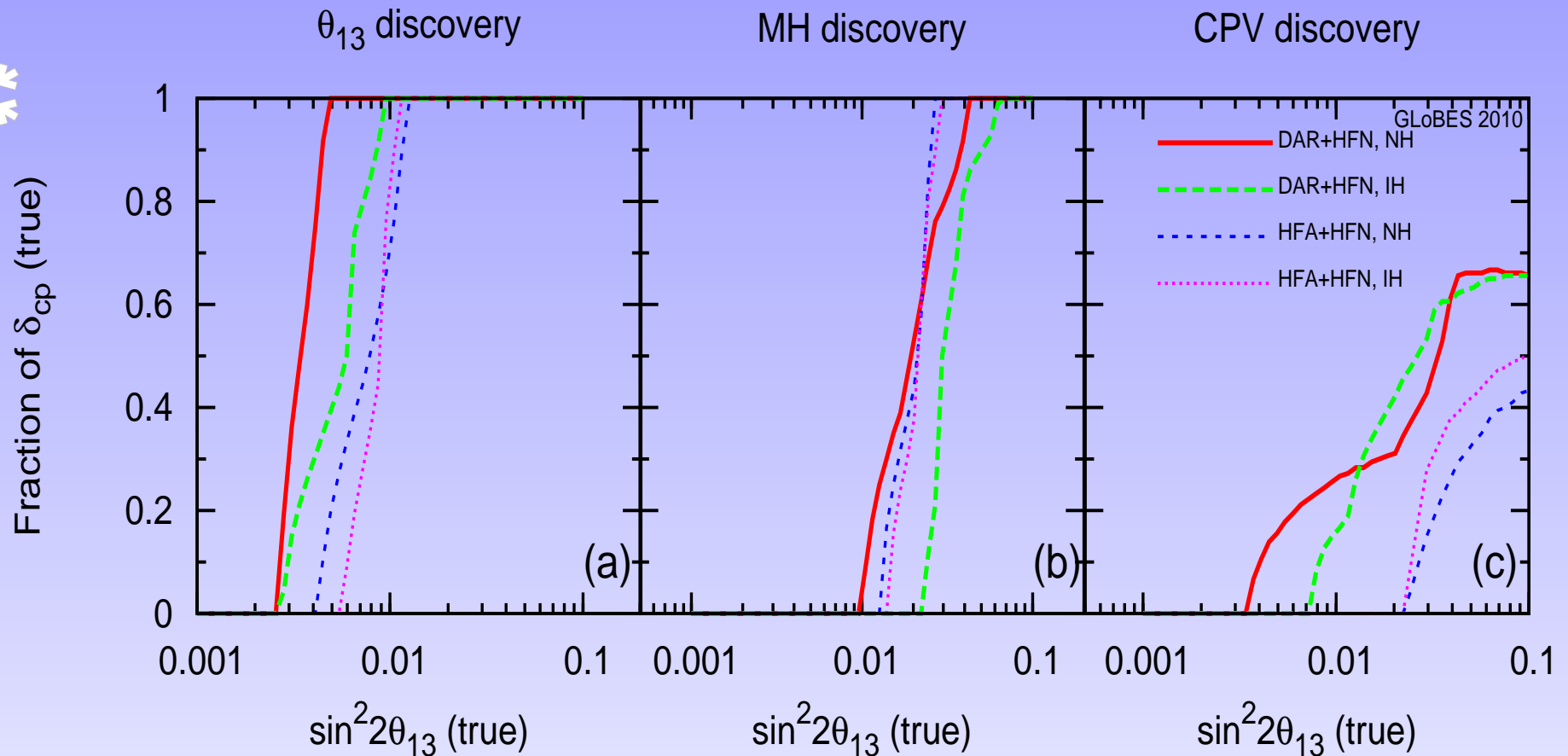
300 kton, NH & IH



Agarwalla, Huber, Link, Mohapatra, arXiv:1005.4055 [hep-ph]

CP fraction for which a discovery at 3σ confidence level is possible as a function of $\sin^2 2\theta_{13}$ (true)

100 kton, NH & IH



Agarwalla, Huber, Link, Mohapatra, arXiv:1005.4055 [hep-ph]

CP fraction for which a discovery at 3σ confidence level is possible as a function of $\sin^2 2\theta_{13}(\text{true})$

Backup Slides

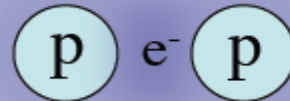
Technologies explored

Linacs

- Cleanest of technologies,
but there are issues of size and cost

Cyclotrons

- Compact Superconducting (proton) Cyclotron
- Stacked (proton) Cyclotron
- H_2^+ Cyclotron -- reduces many problems related to beam loss and extraction compared to other designs



DAEδALUS

Backup Slides

Cyclotron Options

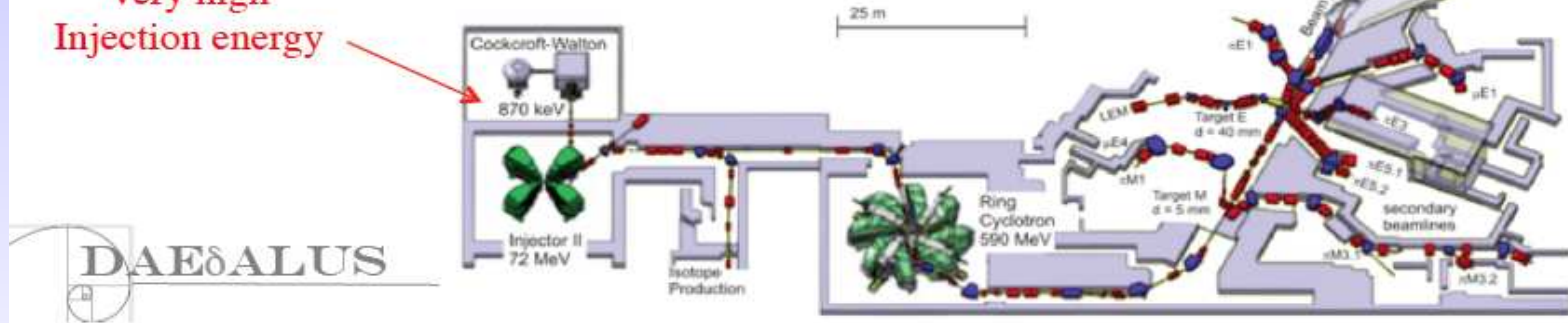
- PSI is best in the world
 - 590 MeV protons
 - 2.2 mA
 - 1.3 MW



Proton Accelerator Complex
Paul-Scherrer-Institute
Switzerland

Neutron Spallation
Source SINQ

Very high
Injection energy



Backup Slides

H_2^+ Ring Cyclotron Promising Design from 1990's

- Concept proposed by Carlo Rubbia ~1994
- Initial designs done by Luciano Calabretta, Catania
 - Reports in European Particle Accelerator Conference
 - Calabretta et al: PAC 99 & EPAC 2000
- 1 GeV, ~6 mA
- High rigidity for H_2^+
 - Superconducting magnets reduce consequences
- Clean extraction (via stripping)

