

asymptotically safe gravity at colliders and beyond

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classical gravity

Einstein's theory $G_N = 6.7 \times 10^{-11} \frac{m^3}{kg\ s^2}$



quantum gravity

Planck length $\ell_{\text{Pl}} = \left(\frac{\hbar G_N}{c^3} \right)^{1/2} \approx 10^{-33} \text{ cm}$

Planck mass $M_{\text{Pl}} \approx 10^{19} \text{ GeV}$

Planck time $t_{\text{Pl}} \approx 10^{-44} \text{ s}$

Planck temperature $T_{\text{Pl}} \approx 10^{32} \text{ K}$

expect **quantum modifications** at energy scales $E \approx M_{\text{Pl}}$

quantum gravity

low-scale quantum gravity

what if the fundamental Planck scale is as low as

$$M_* \approx \mathcal{O}(M_{\text{EW}}) \ll M_{\text{Pl}} ?$$

circumnavigates the SM hierarchy problem

scenario with extra dimensions (Arkani-Hamed, Dimopoulos, Dvali '98)

$D = 4 + n$ compact extra dimensions of size L ,

$$M_{\text{Pl}}^2 \sim M_*^2 (M_* L)^n$$

scale separation $1/L \ll M_* \ll M_{\text{Pl}}$

high-energetic particle colliders can **test quantisation of gravity**

quantum gravity

perturbation theory

effective expansion parameter: $g_{\text{eff}} \equiv G_N E^2 \sim \frac{E^2}{M_{\text{Pl}}^2}$

$[G] < 0$: **dangerous** interactions

perturbative non-renormalisability

gravity with matter interactions

pure gravity (Goroff-Sagnotti term)

renormalisation group

S.Weinberg ('79)

asymptotically safe gravity

running coupling $g(\mu) = G_N \cdot Z_N(\mu)^{-1} \cdot \mu^{D-2}$

anomalous dimension $\eta = -\frac{d \ln Z_N}{d \ln \mu}$

RG running $\frac{dg(\mu)}{d \ln \mu} = (D - 2 + \eta) g(\mu)$

non-Gaussian fixed point

$g_* \neq 0$ and $\eta = 2 - D$ **strong quantum effects**

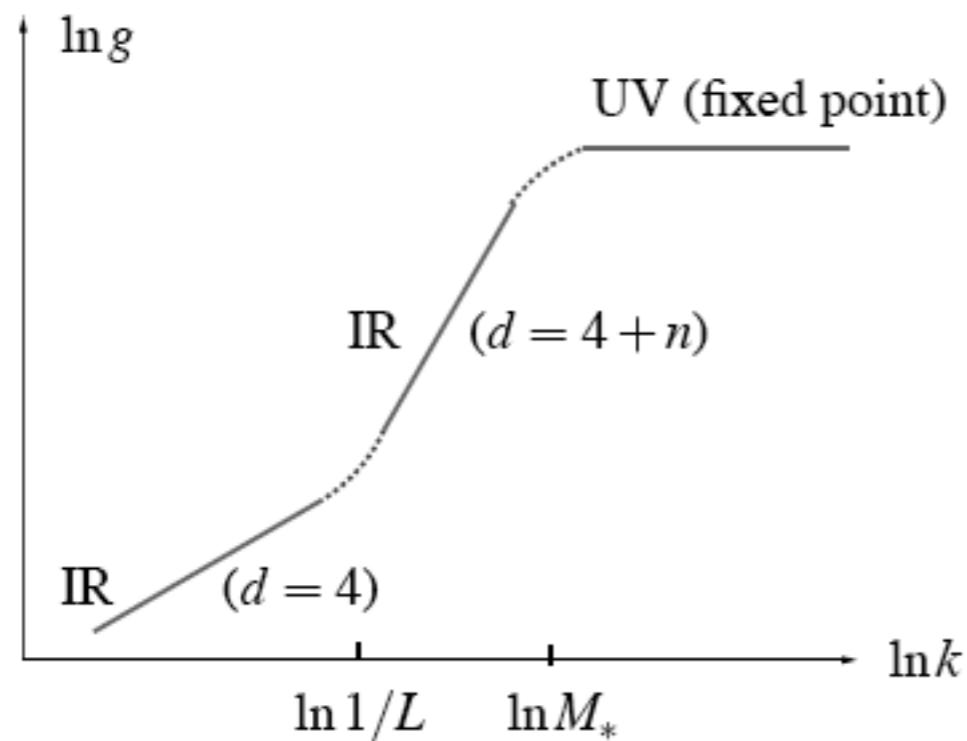
UV fixed point implies **weak coupling at high energies**

$\mu \rightarrow \infty :$ $G(\mu) \rightarrow g_* \mu^{2-D} \ll G_N$

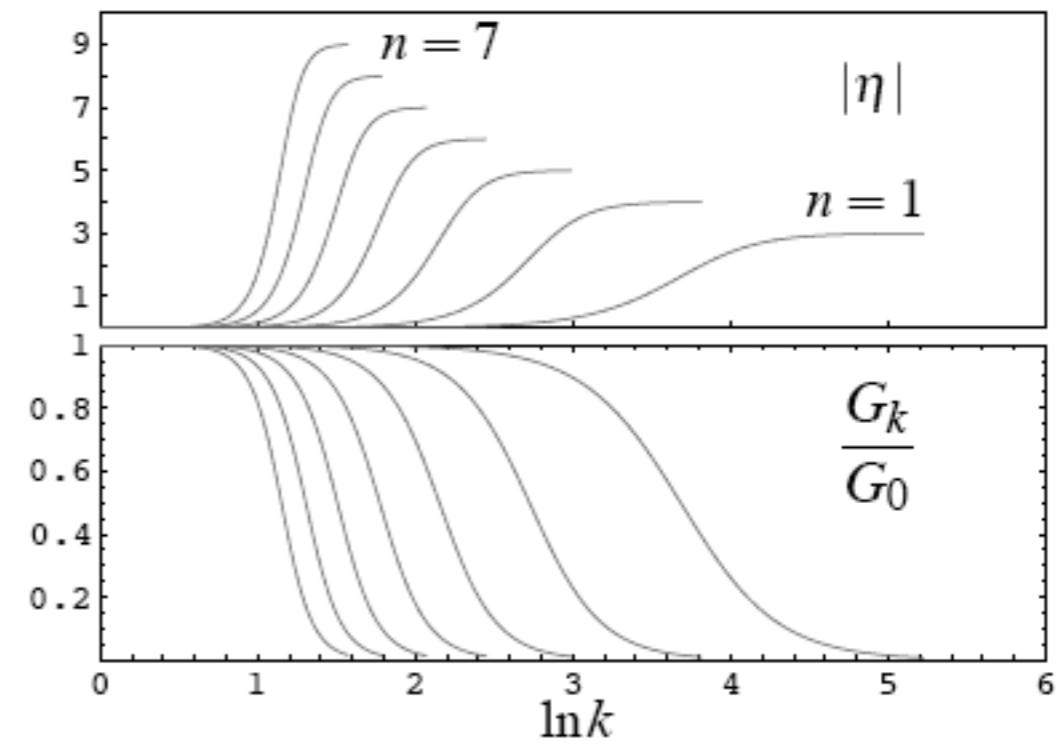
running gravitational coupling

DL ('03), Fischer, DL ('05)

a) schematically



b) numerically



$$g(\mu) \equiv G(\mu) \mu^2$$

$$G(\mu) = G_N \cdot Z_N(\mu)^{-1}$$

collider signatures of quantum gravity

- **real gravitons**

graviton production via $p\ p \rightarrow \text{jet} + G$

signature: missing energy

- **virtual gravitons**

lepton production $q\bar{q} \rightarrow \ell^+ \ell^-$ via graviton exchange

signature: deviations in SM reference processes

- **mini-black holes**

black hole production and decay

signature: many body final states

Drell Yan production

effective theory

Giudice, Rattazzi, Wells ('98)

scattering amplitude for Drell-Yan lepton production

$$A = \mathcal{S}(s) \times T, \quad T = T^{\mu\nu}T_{\mu\nu} - \frac{1}{n+2}T_\mu^\mu T_\nu^\nu$$

$$\mathcal{S}(s) = \frac{1}{M_*^{n+2}} \int_0^\infty dm \frac{m^{n-1}}{s - m^2}$$

UV divergent for $n \geq 2$.

Drell Yan production

renormalisation group

DL, Plehn ('07), Gerwick, DL, Plehn ('11)

RG improved scattering amplitude for Drell-Yan production

$$A = \mathcal{S}(s) \times T, \quad T = T^{\mu\nu}T_{\mu\nu} - \frac{1}{n+2}T_\mu^\mu T_\nu^\nu$$

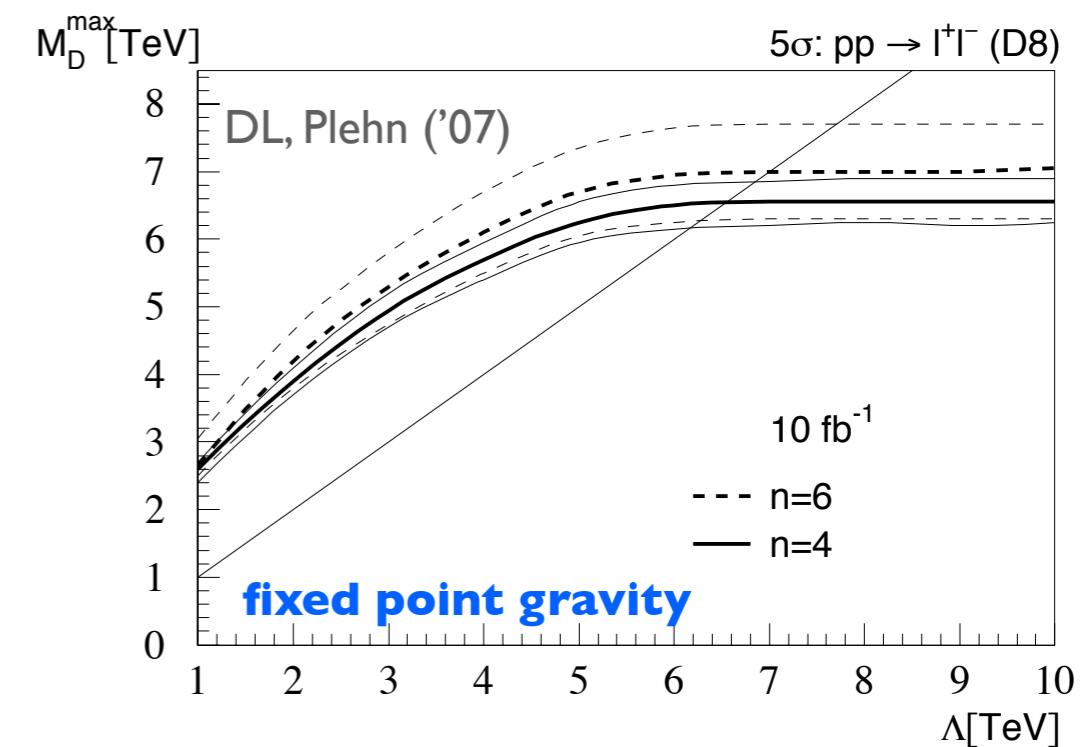
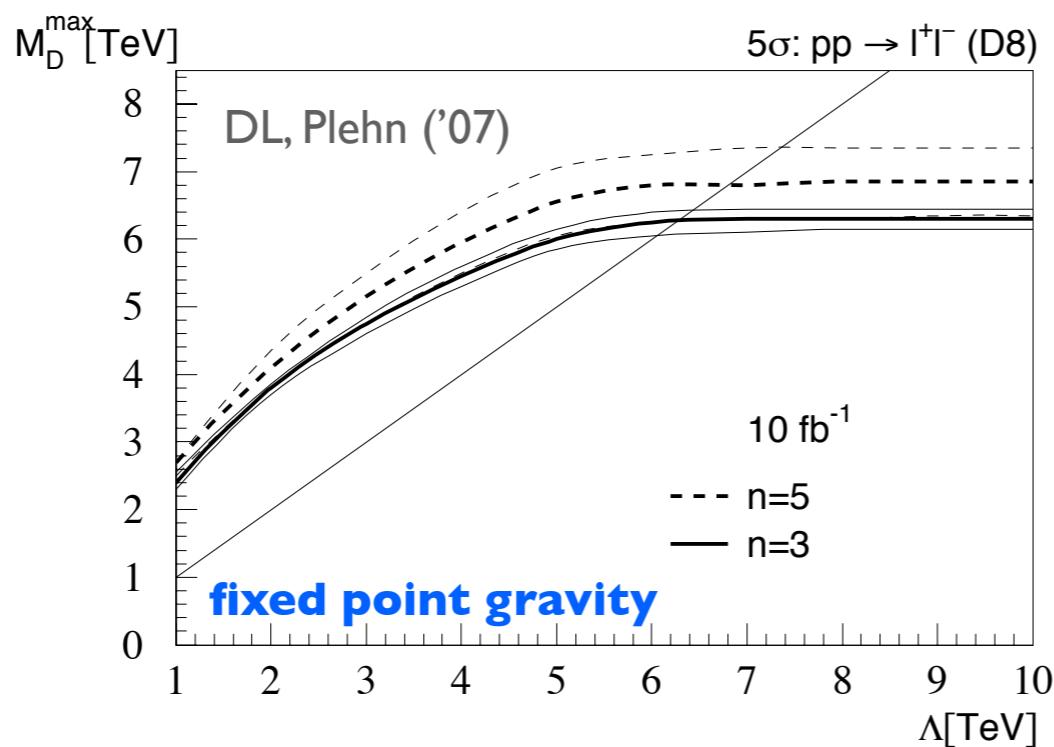
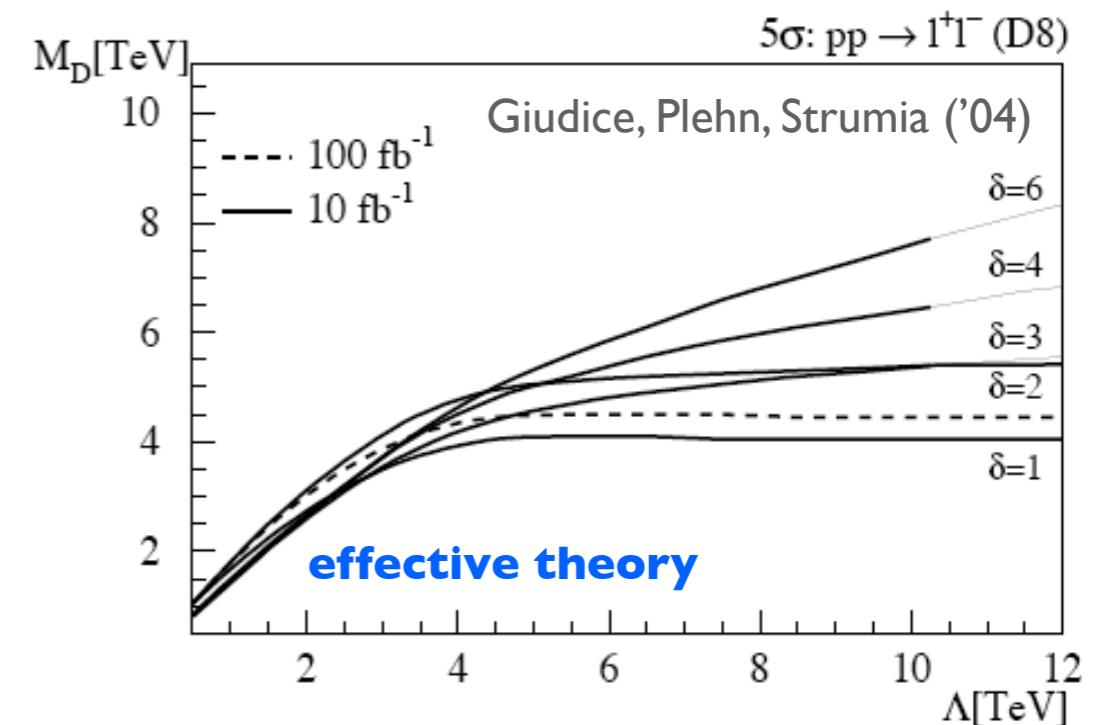
$$\mathcal{S}(s) = \frac{1}{M_*^{n+2}} \int_0^\infty dm \frac{m^{n-1}}{s-m^2} Z^{-1}(\mu(s, m^2, \Lambda_T))$$

UV finite for all n .

Drell Yan production

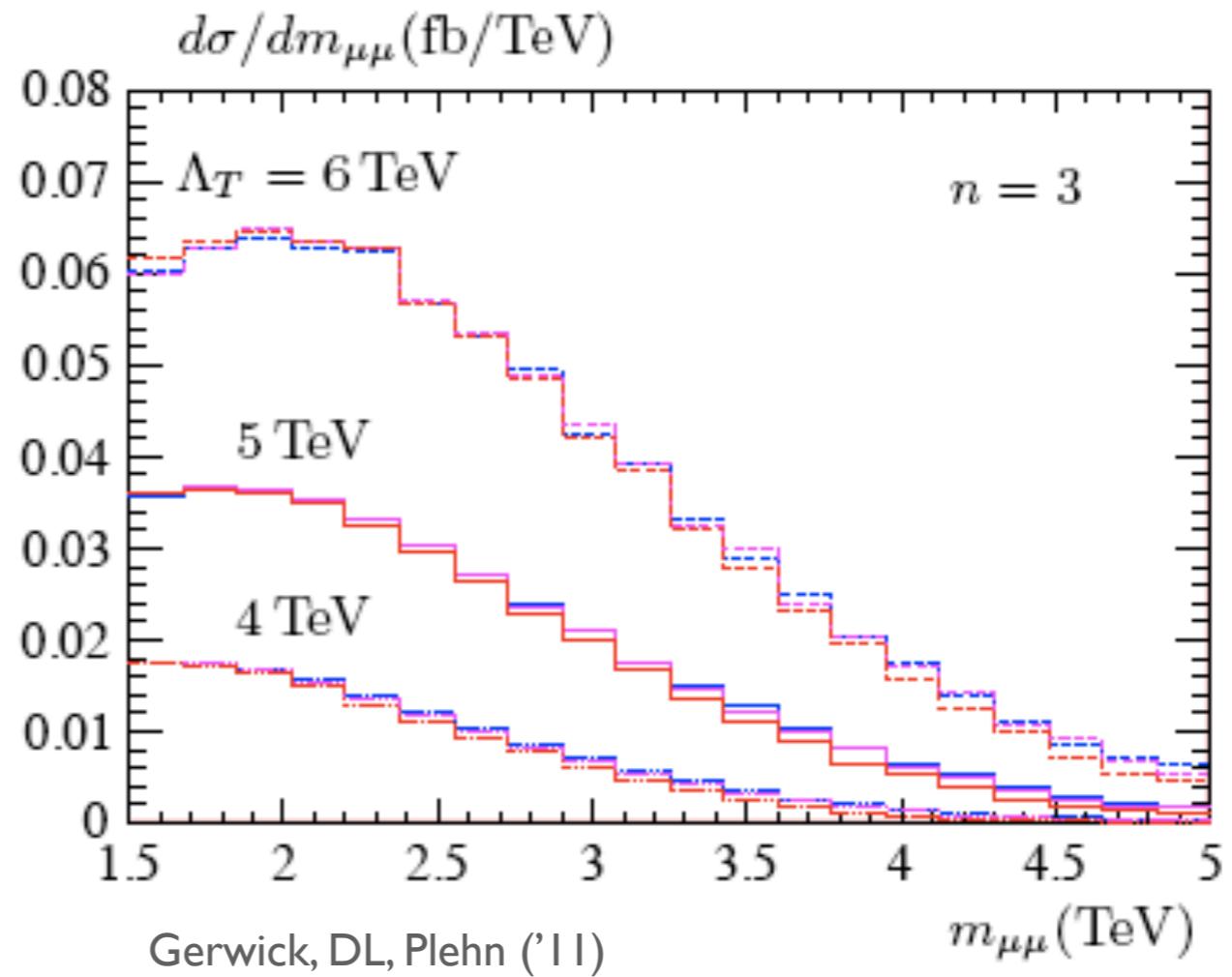
discovery reach

effective theory vs fixed point gravity



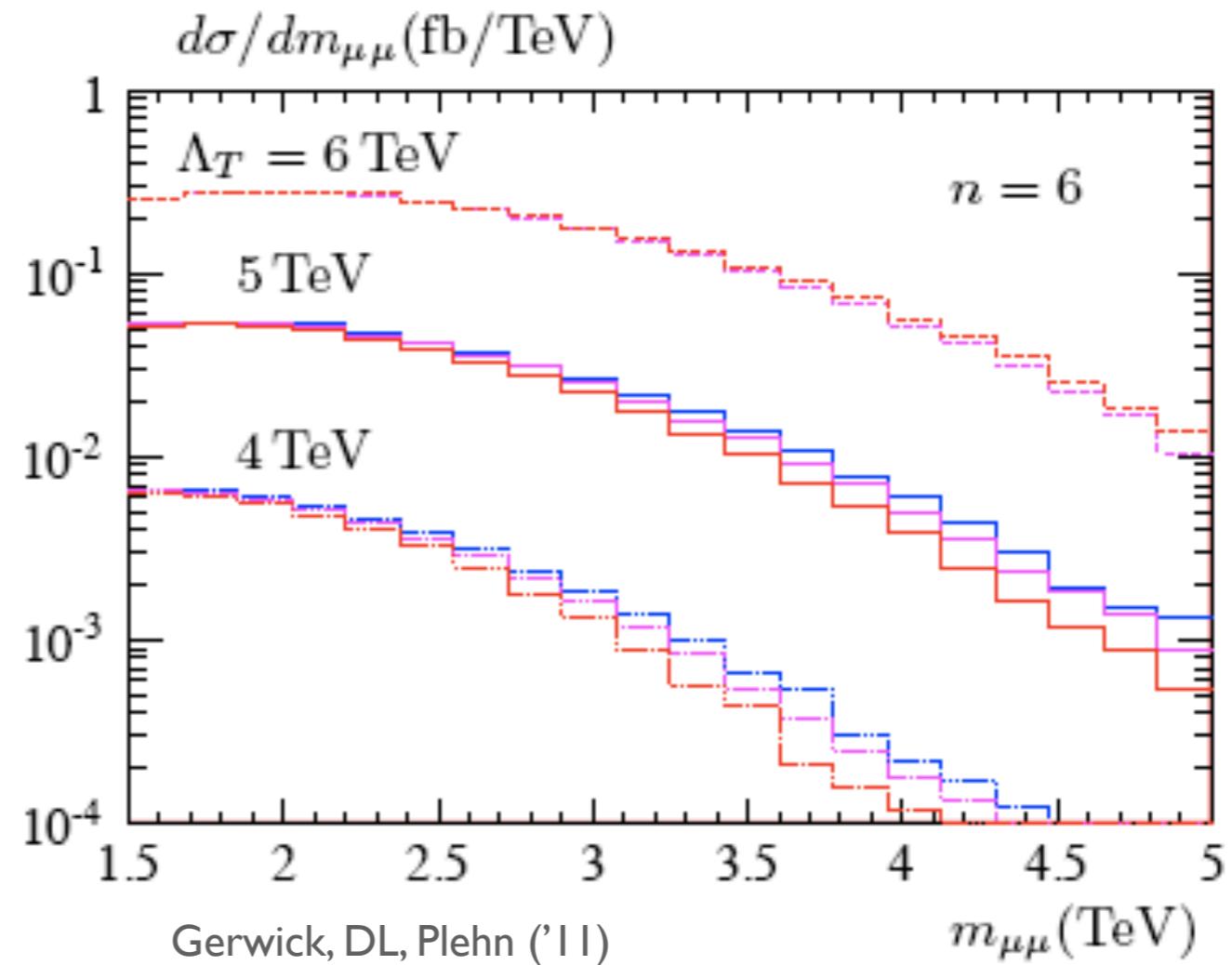
sensitivity

- **vary transition scale**
strong sensitivity
- **vary RG scheme**
weak sensitivity

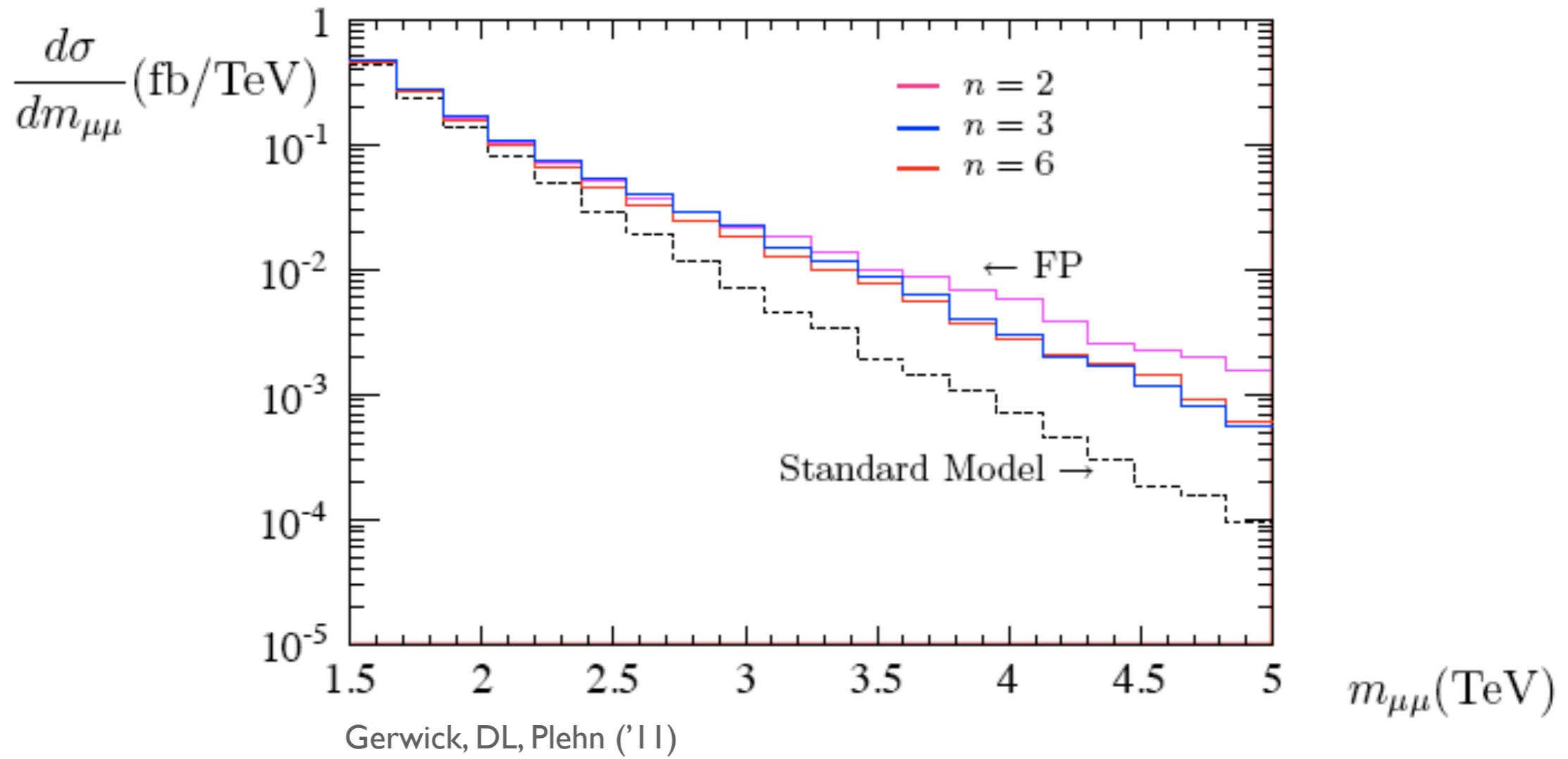


sensitivity

- **vary transition scale**
strong sensitivity
- **vary RG scheme**
weak sensitivity



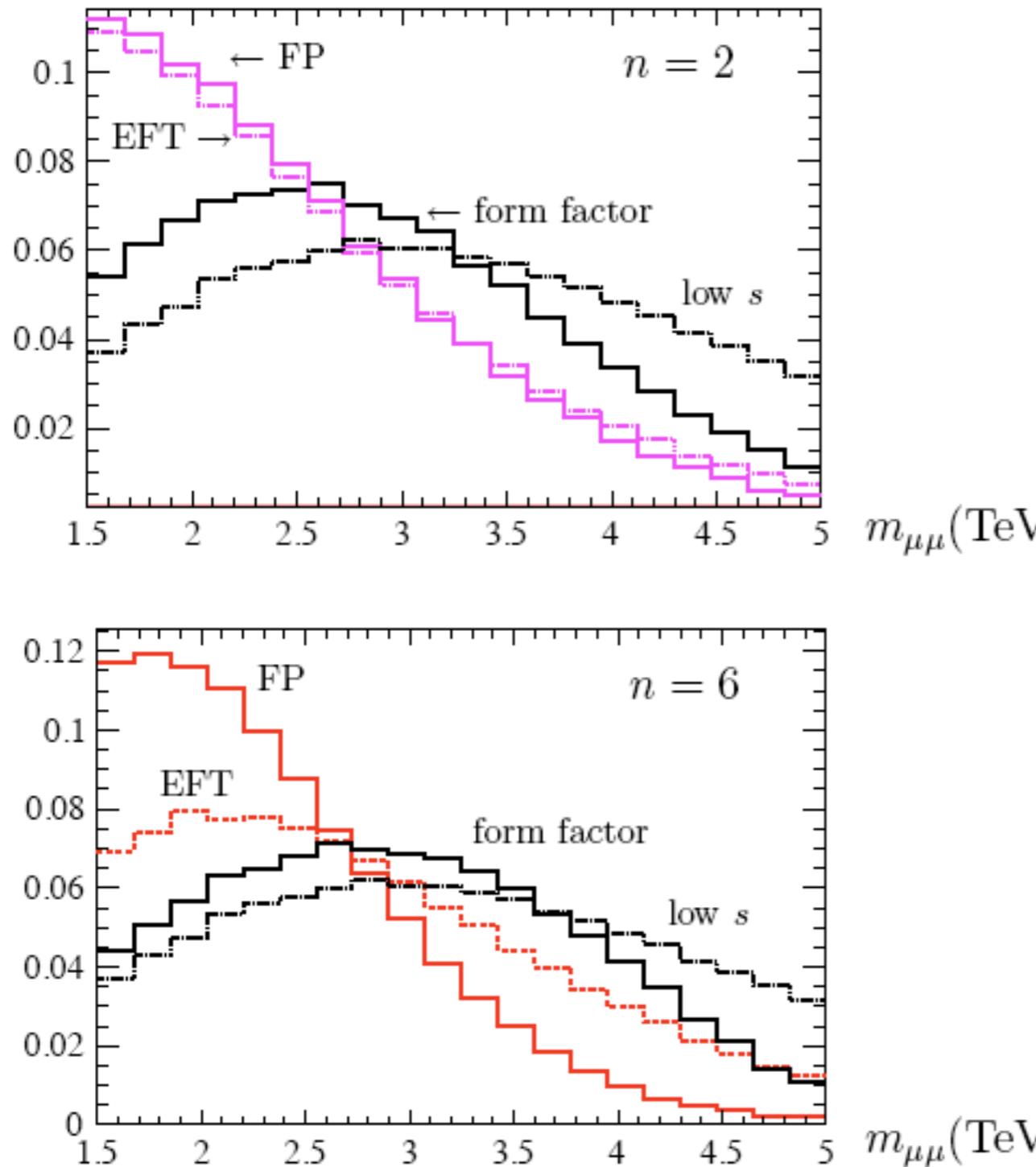
signal vs background



(Gerwick, DL, Plehn '11, to appear in PRD)

comparison

$(1/\sigma) d\sigma/dm_{\mu\mu} (\text{TeV}^{-1})$



Gerwick, DL, Plehn ('11)

mini-black hole production

semi-classical picture

Dimopoulos, Landsberg ('01) Giddings, Thomas ('01)

semi-classical production cross section

$$\hat{\sigma} = \pi r_{\text{cl}}^2 (M = \sqrt{s}) \times \theta(\sqrt{s} - M_{\min})$$

production cross section at the LHC $pp \rightarrow \text{final state}$

$$\sigma = \sum_{i,j} \int_0^1 dx_1 \int_0^1 dx_2 f_i(x_1) f_j(x_2) \hat{\sigma}(q_i q_j \rightarrow \text{final state})$$

parton distribution functions from **CTEQ61**

evaluated at $Q^2 = M_{\text{BH}}^2$.

mini-black hole production

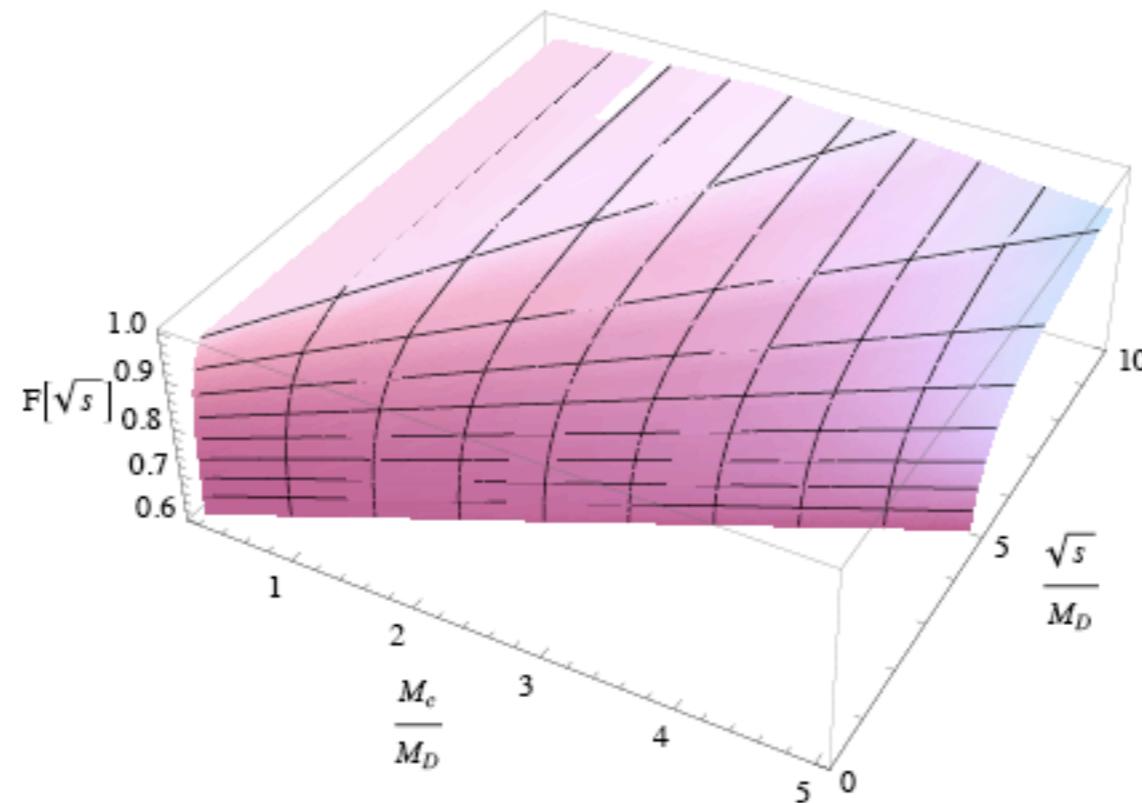
RG quantum corrections

Falls, DL, Raghuraman ('10)

quantum corrected production cross section

$$\hat{\sigma} \rightarrow \hat{\sigma} = F(\sqrt{s}) \times \pi r_{\text{cl}}^2(M = \sqrt{s}) \times \theta(\sqrt{s} - M_c)$$

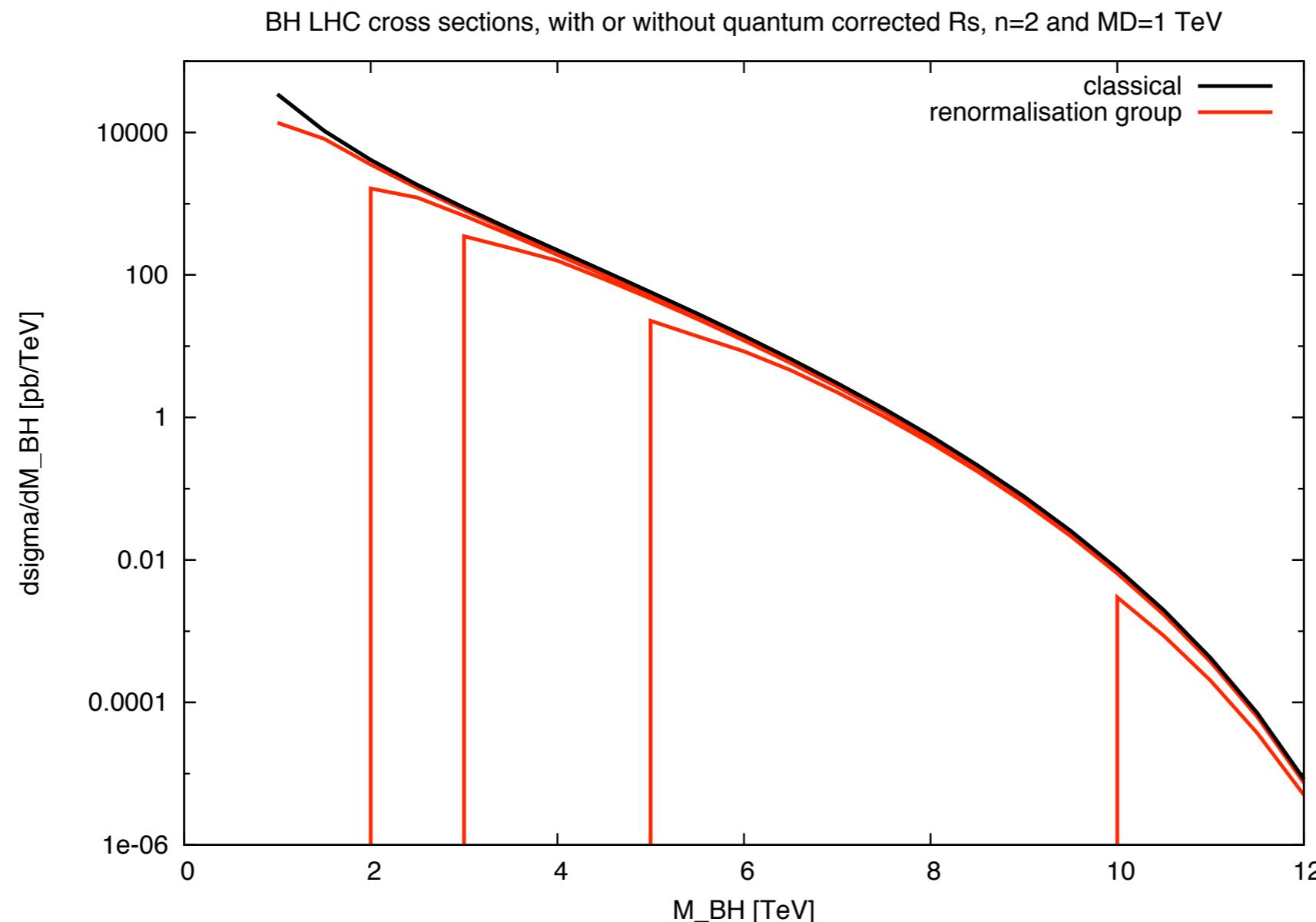
new form factor F



mini-black hole production

Falls, Hiller, DL (Pascos '09 and in prep.)

RG quantum corrections



higher-dimensional black holes

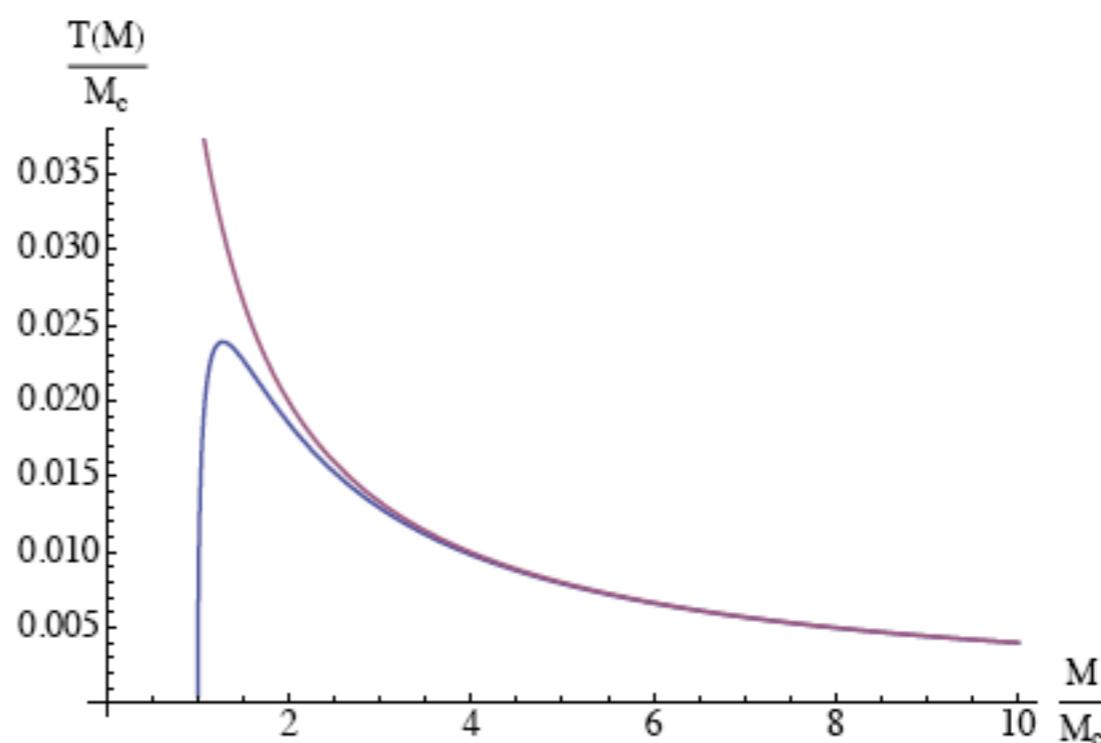
thermodynamics

Falls, DL, Raghuraman ('10)

temperature from surface gravity

$$T = \frac{\kappa}{2\pi} = \frac{d-3}{4\pi r_s} \left(1 + \frac{\eta(r_s)}{d-3} \right)$$

vanishes at $M = M_c$.



conclusion

asymptotically safe gravity

strong indications for gravitational fixed point

low-scale quantum gravity

strong sensitivity on the cross-over scale at the LHC

moderate RG scheme dependence

access to several observables

thank you!