

# Non-standard interactions in future neutrino oscillation experiments

Joachim Kopp

Max-Planck-Institut für Kernphysik, Heidelberg, Germany

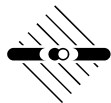
EuroGDR SUSY 2007, Brussels, 14 November 2007



MAX-PLANCK-GESELLSCHAFT

[hep-ph/0702269](#), [arXiv:0710.1867](#)

in collaboration with M. Lindner, T. Ota, and J. Sato



MAX-PLANCK-INSTITUT  
FÜR KERNPHYSIK

# Outline

- 1 Introduction to non-standard interactions
- 2 Sensitivity of reactor and superbeam experiments
- 3 Discovery reach of a neutrino factory
- 4 Summary and conclusions

# Outline

- 1 Introduction to non-standard interactions
- 2 Sensitivity of reactor and superbeam experiments
- 3 Discovery reach of a neutrino factory
- 4 Summary and conclusions

# Motivation

- “New physics” often leaves low-energy fingerprints in the form of effective, non-standard 4-fermion interactions (*NSI*).  
⇒ Modification of weak interaction Lagrangian

# Motivation

- “New physics” often leaves low-energy fingerprints in the form of effective, non-standard 4-fermion interactions (*NSI*).  
⇒ Modification of weak interaction Lagrangian
- NSI can affect neutrino production, propagation, and detection

# Motivation

- “New physics” often leaves low-energy fingerprints in the form of effective, non-standard 4-fermion interactions (*NSI*).  
 ⇒ Modification of weak interaction Lagrangian
- NSI can affect neutrino production, propagation, and detection
- Example:

$$\begin{aligned}
 \mathcal{L}_{\text{NSI}} = & \frac{G_F}{\sqrt{2}} \sum_{f,f'} \tilde{\epsilon}_{\alpha\beta}^{s,f,f'} [\bar{\nu}_\beta \gamma^\rho (1 - \gamma^5) \ell_\alpha] [\bar{f}' \gamma_\rho (1 - \gamma^5) f] \\
 & + \frac{G_F}{\sqrt{2}} \sum_f \tilde{\epsilon}_{\alpha\beta}^{m,f} [\bar{\nu}_\alpha \gamma^\rho (1 - \gamma^5) \nu_\beta] [\bar{f} \gamma_\rho (1 - \gamma^5) f] + \text{h.c.},
 \end{aligned}$$

# Motivation

- “New physics” often leaves low-energy fingerprints in the form of effective, non-standard 4-fermion interactions (NSI).  
 ⇒ Modification of weak interaction Lagrangian
- NSI can affect neutrino production, propagation, and detection
- Example:

$$\mathcal{L}_{\text{NSI}} = \frac{G_F}{\sqrt{2}} \sum_{f, f'} \tilde{\epsilon}_{\alpha\beta}^{s, f, f'} [\bar{\nu}_\beta \gamma^\rho (1 - \gamma^5) \ell_\alpha] [\bar{f}' \gamma_\rho (1 - \gamma^5) f] \\ + \frac{G_F}{\sqrt{2}} \sum_f \tilde{\epsilon}_{\alpha\beta}^{m, f} [\bar{\nu}_\alpha \gamma^\rho (1 - \gamma^5) \nu_\beta] [\bar{f} \gamma_\rho (1 - \gamma^5) f] + \text{h.c.},$$

- Lorentz structures different from  $(V - A)(V - A)$  are possible, but not considered in this talk.

see e.g. JK Lindner Ota Sato, arxiv:0708:152 for a discussion of NSI with non- $(V - A)(V - A)$  Lorentz structure.

# Modified neutrino oscillation probabilities

## Standard oscillations

$$P_{\nu_\alpha \rightarrow \nu_\beta} = |\langle \nu_\beta | e^{-iHL} | \nu_\alpha \rangle|^2$$



# Modified neutrino oscillation probabilities

## Standard oscillations

$$P_{\nu_\alpha \rightarrow \nu_\beta} = |\langle \nu_\beta | e^{-iHL} | \nu_\alpha \rangle|^2$$

## Oscillations with NSI

$$P_{\nu_\alpha^s \rightarrow \nu_\beta^d} = |\langle \nu_\beta^d | e^{-i(H + V_{\text{NSI}})L} | \nu_\alpha^s \rangle|^2$$

# Modified neutrino oscillation probabilities

## Standard oscillations

$$P_{\nu_\alpha \rightarrow \nu_\beta} = |\langle \nu_\beta | e^{-iHL} | \nu_\alpha \rangle|^2$$

## Oscillations with NSI

$$P_{\nu_\alpha^s \rightarrow \nu_\beta^d} = |\langle \nu_\beta^d | e^{-i(H+V_{\text{NSI}})L} | \nu_\alpha^s \rangle|^2$$

- **CC type NSI:** Flavour mixture at source and detector (Grossman PL B359 (1995) 141)

$$|\nu_\alpha^s\rangle = |\nu_\alpha\rangle + \sum_{\beta=e,\mu,\tau} \varepsilon_{\alpha\beta}^s |\nu_\beta\rangle,$$

$$\text{e.g. } \pi^+ \xrightarrow{\varepsilon_{\mu e}^s} \mu^+ \nu_e$$

$$\langle \nu_\beta^d | = \langle \nu_\beta | + \sum_{\alpha=e,\mu,\tau} \varepsilon_{\alpha\beta}^d \langle \nu_\alpha |$$

$$\text{e.g. } \nu_\tau N \xrightarrow{\varepsilon_{\tau e}^d} e^- X$$

# Modified neutrino oscillation probabilities

## Standard oscillations

$$P_{\nu_\alpha \rightarrow \nu_\beta} = |\langle \nu_\beta | e^{-iHL} | \nu_\alpha \rangle|^2$$

## Oscillations with NSI

$$P_{\nu_\alpha^s \rightarrow \nu_\beta^d} = |\langle \nu_\beta^d | e^{-i(H+V_{\text{NSI}})L} | \nu_\alpha^s \rangle|^2$$

- **CC type NSI:** Flavour mixture at source and detector (Grossman PL **B359** (1995) 141)

$$|\nu_\alpha^s\rangle = |\nu_\alpha\rangle + \sum_{\beta=e,\mu,\tau} \epsilon_{\alpha\beta}^s |\nu_\beta\rangle,$$

$$\text{e.g. } \pi^+ \xrightarrow{\epsilon_{\mu e}^s} \mu^+ \nu_e$$

$$\langle \nu_\beta^d | = \langle \nu_\beta | + \sum_{\alpha=e,\mu,\tau} \epsilon_{\alpha\beta}^d \langle \nu_\alpha |$$

$$\text{e.g. } \nu_\tau N \xrightarrow{\epsilon_{\tau e}^d} e^- X$$

- **NC type NSI:** Extra matter effects in propagation

Wolfenstein PR **D17** (1978) 2369, Valle PL **B199** (1987) 432, Guzzo Masiero Petcov PL **B260** (1991) 154, Roulet PR **D44** (1991) R935, etc.

$$(V_{\text{NSI}})_{\alpha\beta} = \sqrt{2} G_F N_e \epsilon_{\alpha\beta}^m$$

# NSI in oscillation experiments

- Compared to charged lepton flavour violation experiments: Interference between standard and non-standard amplitudes is possible  
⇒ NSI effects suppressed only by  $|\epsilon|$  instead of  $|\epsilon|^2$ .

Grossman 1995, Wolfenstein 1977, Valle 1987, Guzzo 1991, Roulet 1991, Bergmann 1999, Gago 2001.

# NSI in oscillation experiments

- Compared to charged lepton flavour violation experiments: Interference between standard and non-standard amplitudes is possible  
⇒ NSI effects suppressed only by  $|\epsilon|$  instead of  $|\epsilon|^2$ .  
Grossman 1995, Wolfenstein 1977, Valle 1987, Guzzo 1991, Roulet 1991, Bergmann 1999, Gago 2001.
- Possible consequences in oscillation experiments:

# NSI in oscillation experiments

- Compared to charged lepton flavour violation experiments: Interference between standard and non-standard amplitudes is possible  
⇒ NSI effects suppressed only by  $|\epsilon|$  instead of  $|\epsilon|^2$ .  
Grossman 1995, Wolfenstein 1977, Valle 1987, Guzzo 1991, Roulet 1991, Bergmann 1999, Gago 2001.
- Possible consequences in oscillation experiments:
  - **Poor quality** of standard oscillation fit (⇒ Detection of NSI possible)

# NSI in oscillation experiments

- Compared to charged lepton flavour violation experiments: Interference between standard and non-standard amplitudes is possible

⇒ NSI effects suppressed only by  $|\epsilon|$  instead of  $|\epsilon|^2$ .

Grossman 1995, Wolfenstein 1977, Valle 1987, Guzzo 1991, Roulet 1991, Bergmann 1999, Gago 2001.

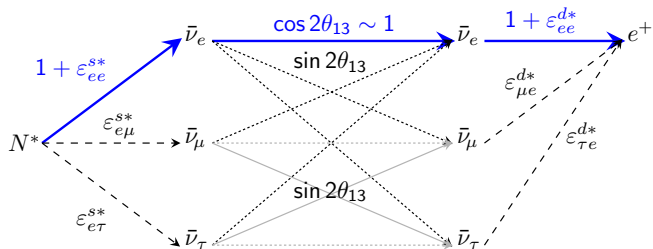
- Possible consequences in oscillation experiments:
  - **Poor quality** of standard oscillation fit (⇒ Detection of NSI possible)
  - **Mismatch** between standard fits to different experiments

# NSI in oscillation experiments

- Compared to charged lepton flavour violation experiments: Interference between standard and non-standard amplitudes is possible  
⇒ NSI effects suppressed only by  $|\epsilon|$  instead of  $|\epsilon|^2$ .  
Grossman 1995, Wolfenstein 1977, Valle 1987, Guzzo 1991, Roulet 1991, Bergmann 1999, Gago 2001.
- Possible consequences in oscillation experiments:
  - **Poor quality** of standard oscillation fit (⇒ Detection of NSI possible)
  - **Mismatch** between standard fits to different experiments
  - **Offset**: Consistent, but wrong reconstruction of neutrino mixing parameters

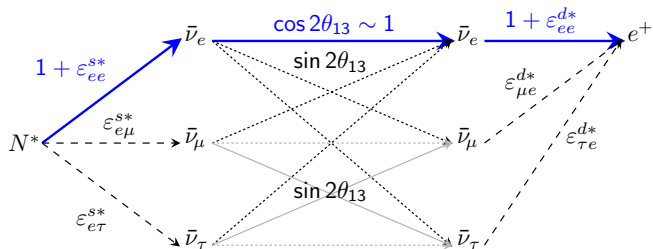


# Qualitative arguments



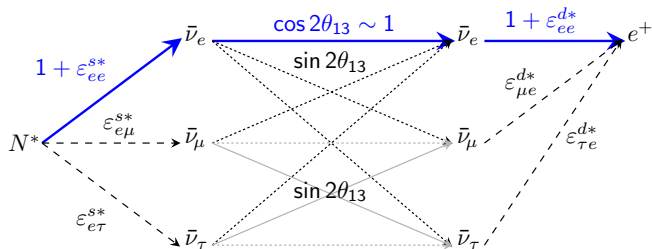
- **Standard path:**  $N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+$

# Qualitative arguments



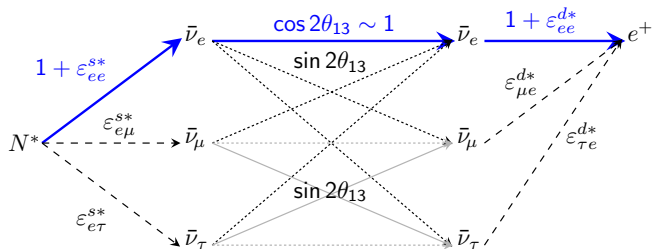
- **Standard path:**  $N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+$
- **NSI production processes:**  $N^* \xrightarrow{\epsilon_{ee}^{s*}} \bar{\nu}_e$      $N^* \xrightarrow{\epsilon_{e\mu}^{s*}} \bar{\nu}_\mu$      $N^* \xrightarrow{\epsilon_{e\tau}^{s*}} \bar{\nu}_\tau$

# Qualitative arguments



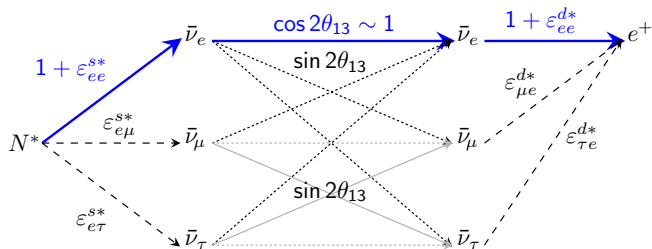
- **Standard path:**  $N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+$
- **NSI production processes:**  $N^* \xrightarrow{\epsilon_{ee}^{s*}} \bar{\nu}_e$      $N^* \xrightarrow{\epsilon_{e\mu}^{s*}} \bar{\nu}_\mu$      $N^* \xrightarrow{\epsilon_{e\tau}^{s*}} \bar{\nu}_\tau$
- **NSI detection processes:**  $\bar{\nu}_e \xrightarrow{\epsilon_{ee}^{d*}} e^+$      $\bar{\nu}_\mu \xrightarrow{\epsilon_{\mu e}^{d*}} e^+$      $\bar{\nu}_\tau \xrightarrow{\epsilon_{\tau e}^{d*}} e^+$

# Qualitative arguments



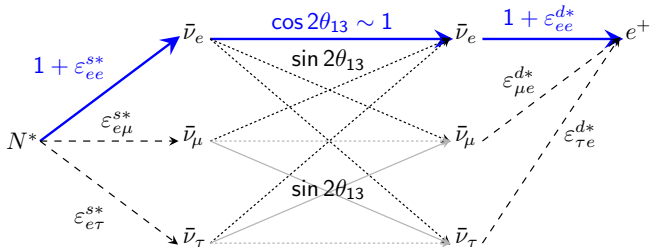
- **Standard path:**  $N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+$
- **NSI production processes:**  $N^* \xrightarrow{\epsilon_{ee}^{s*}} \bar{\nu}_e$      $N^* \xrightarrow{\epsilon_{e\mu}^{s*}} \bar{\nu}_\mu$      $N^* \xrightarrow{\epsilon_{e\tau}^{s*}} \bar{\nu}_\tau$
- **NSI detection processes:**  $\bar{\nu}_e \xrightarrow{\epsilon_{ee}^{d*}} e^+$      $\bar{\nu}_\mu \xrightarrow{\epsilon_{\mu e}^{d*}} e^+$      $\bar{\nu}_\tau \xrightarrow{\epsilon_{\tau e}^{d*}} e^+$
- $\bar{\nu}_e \leftrightarrow \bar{\nu}_\mu$  and  $\bar{\nu}_e \leftrightarrow \bar{\nu}_\tau$  oscillations are suppressed by  $\theta_{13}$ .

# Qualitative arguments



- **Standard path:**  $N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+$
- **NSI production processes:**  $N^* \xrightarrow{\epsilon_{ee}^{s*}} \bar{\nu}_e$      $N^* \xrightarrow{\epsilon_{e\mu}^{s*}} \bar{\nu}_\mu$      $N^* \xrightarrow{\epsilon_{e\tau}^{s*}} \bar{\nu}_\tau$
- **NSI detection processes:**  $\bar{\nu}_e \xrightarrow{\epsilon_{ee}^{d*}} e^+$      $\bar{\nu}_\mu \xrightarrow{\epsilon_{\mu e}^{d*}} e^+$      $\bar{\nu}_\tau \xrightarrow{\epsilon_{\tau e}^{d*}} e^+$
- $\bar{\nu}_e \leftrightarrow \bar{\nu}_\mu$  and  $\bar{\nu}_e \leftrightarrow \bar{\nu}_\tau$  oscillations are suppressed by  $\theta_{13}$ .
- Assume only one  $\epsilon$  parameter is sizeable

# Qualitative arguments



$$N^* \xrightarrow{\epsilon_{ee}^{s*}} \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+$$

$$\mathcal{O}(\epsilon)$$

but: absorbed in flux uncertainty

$$N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \xrightarrow{\epsilon_{ee}^{d*}} e^+$$

$$\mathcal{O}(\epsilon)$$

but: absorbed in flux uncertainty

$$N^* \xrightarrow{\epsilon_{e\mu}^{s*}} \bar{\nu}_\mu \xrightarrow{\sin \theta_{13}} \bar{\nu}_e \rightarrow e^+$$

$$\mathcal{O}(\epsilon \sin \theta_{13})$$

$$N^* \xrightarrow{\epsilon_{e\tau}^{s*}} \bar{\nu}_\tau \xrightarrow{\sin \theta_{13}} \bar{\nu}_e \rightarrow e^+$$

$$\mathcal{O}(\epsilon \sin \theta_{13})$$

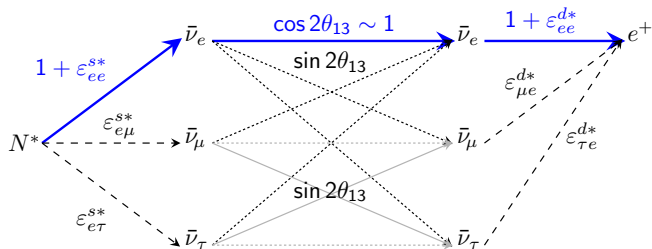
$$N^* \rightarrow \bar{\nu}_e \xrightarrow{\sin \theta_{13}} \bar{\nu}_\mu \xrightarrow{\epsilon_{\mu e}^{d*}} e^+$$

$$\mathcal{O}(\epsilon \sin \theta_{13})$$

$$N^* \rightarrow \bar{\nu}_e \xrightarrow{\sin \theta_{13}} \bar{\nu}_\tau \xrightarrow{\epsilon_{\tau e}^{d*}} e^+$$

$$\mathcal{O}(\epsilon \sin \theta_{13})$$

# Qualitative arguments



$$N^* \xrightarrow{\epsilon_{ee}^{s*}} \bar{\nu}_e \rightarrow \bar{\nu}_e \rightarrow e^+ \quad \mathcal{O}(\epsilon)$$

but: absorbed in flux uncertainty

$$N^* \rightarrow \bar{\nu}_e \rightarrow \bar{\nu}_e \xrightarrow{\epsilon_{ee}^{d*}} e^+ \quad \mathcal{O}(\epsilon)$$

but: absorbed in flux uncertainty

$$N^* \xrightarrow{\epsilon_{e\mu}^{s*}} \bar{\nu}_\mu \xrightarrow{\sin \theta_{13}} \bar{\nu}_e \rightarrow e^+ \quad \mathcal{O}(\epsilon \sin \theta_{13})$$

$$N^* \xrightarrow{\epsilon_{e\tau}^{s*}} \bar{\nu}_\tau \xrightarrow{\sin \theta_{13}} \bar{\nu}_e \rightarrow e^+ \quad \mathcal{O}(\epsilon \sin \theta_{13})$$

$$N^* \rightarrow \bar{\nu}_e \xrightarrow{\sin \theta_{13}} \bar{\nu}_\mu \xrightarrow{\epsilon_{\mu e}^{d*}} e^+ \quad \mathcal{O}(\epsilon \sin \theta_{13})$$

$$N^* \rightarrow \bar{\nu}_e \xrightarrow{\sin \theta_{13}} \bar{\nu}_\tau \xrightarrow{\epsilon_{\tau e}^{d*}} e^+ \quad \mathcal{O}(\epsilon \sin \theta_{13})$$

# Example: Approximate expression for $P_{\nu_\mu^s \rightarrow \nu_e^d}$



# Example: Approximate expression for $P_{\nu_{\mu}^s \rightarrow \nu_e^d}$

$$\begin{aligned}
P_{\nu_{\mu}^s \rightarrow \nu_e^d}^{\text{mat}} &= 4 \bar{s}_{13}^2 s_{23}^2 \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{4E} \\
&+ \left( \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \right)^2 c_{23}^2 s_{23}^2 \times 12 \left( \frac{\Delta m_{31}^2}{a_{\text{CC}}} \right)^2 \sin^2 \frac{a_{\text{CC}}L}{4E} \\
&- \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \bar{s}_{13} s_{23} \times 12 s_{23}^2 \times 23 \cos \delta_{\text{CP}} \frac{\Delta m_{31}^2}{a_{\text{CC}}} \left[ \sin^2 \frac{a_{\text{CC}}L}{4E} - \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{4E} \right] \\
&- \frac{1}{2} \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \bar{s}_{13} s_{23} \times 12 s_{23}^2 \times 23 \sin \delta_{\text{CP}} \frac{\Delta m_{31}^2}{a_{\text{CC}}} \left[ \sin \frac{a_{\text{CC}}L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{2E} \right] \\
&- 4 |\epsilon_{\mu e}^s| \bar{s}_{13} s_{23} \cos(\phi_{\mu e}^s + \delta_{\text{CP}}) \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{4E} \\
&- 2 |\epsilon_{\mu e}^s| \bar{s}_{13} s_{23} \sin(\phi_{\mu e}^s + \delta_{\text{CP}}) \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{2E} \\
&+ 4 |\epsilon_{\mu e}^d| \bar{s}_{13} s_{23} \cos(\phi_{\mu e}^d + \delta_{\text{CP}}) \left[ c_{23}^2 \sin^2 \frac{a_{\text{CC}}L}{4E} - c_{23}^2 \sin^2 \frac{\Delta m_{31}^2 L}{4E} + s_{23}^2 \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{4E} \right] \\
&+ 2 |\epsilon_{\mu e}^d| \bar{s}_{13} s_{23} \sin(\phi_{\mu e}^d + \delta_{\text{CP}}) \left[ c_{23}^2 \sin \frac{a_{\text{CC}}L}{2E} - c_{23}^2 \sin \frac{\Delta m_{31}^2 L}{2E} - s_{23}^2 \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{2E} \right] \\
&- 4 |\epsilon_{\tau e}^d| \bar{s}_{13} s_{23}^2 c_{23} \cos(\phi_{\tau e}^d + \delta_{\text{CP}}) \left[ \sin^2 \frac{a_{\text{CC}}L}{4E} - \sin^2 \frac{\Delta m_{31}^2 L}{4E} - \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{4E} \right] \\
&- 2 |\epsilon_{\tau e}^d| \bar{s}_{13} s_{23}^2 c_{23} \sin(\phi_{\tau e}^d + \delta_{\text{CP}}) \left[ \sin \frac{a_{\text{CC}}L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}})L}{2E} \right]
\end{aligned}$$

# Example: Approximate expression for $P_{\nu_\mu^s \rightarrow \nu_e^d}$

$$\begin{aligned}
& - 4 |\epsilon_{e\mu}^m| \bar{s}_{13} s_{23} c_{23}^2 \cos(\phi_{e\mu}^m + \delta_{\text{CP}}) \left[ \sin^2 \frac{a_{\text{CC}} L}{4E} - \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{4E} \right] \\
& - 2 |\epsilon_{e\mu}^m| \bar{s}_{13} s_{23} c_{23}^2 \sin(\phi_{e\mu}^m + \delta_{\text{CP}}) \left[ \sin \frac{a_{\text{CC}} L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{2E} \right] \\
& + 8 |\epsilon_{e\mu}^m| \bar{s}_{13} s_{23}^3 \cos(\phi_{e\mu}^m + \delta_{\text{CP}}) \frac{a_{\text{CC}}}{\Delta m_{31}^2 - a_{\text{CC}}} \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{4E} \\
& + 4 |\epsilon_{e\tau}^m| \bar{s}_{13} s_{23}^2 c_{23} \cos(\phi_{e\tau}^m + \delta_{\text{CP}}) \left[ \sin^2 \frac{a_{\text{CC}} L}{4E} - \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{4E} \right] \\
& + 2 |\epsilon_{e\tau}^m| \bar{s}_{13} s_{23}^2 c_{23} \sin(\phi_{e\tau}^m + \delta_{\text{CP}}) \left[ \sin \frac{a_{\text{CC}} L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{2E} \right] \\
& + 8 |\epsilon_{e\tau}^m| \bar{s}_{13} s_{23}^2 c_{23} \cos(\phi_{e\tau}^m + \delta_{\text{CP}}) \frac{a_{\text{CC}}}{\Delta m_{31}^2 - a_{\text{CC}}} \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{4E} \\
& + 2 |\epsilon_{\mu e}^s| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 c_{23} \cos \phi_{\mu e}^s \frac{\Delta m_{31}^2}{a_{\text{CC}}} \sin^2 \frac{a_{\text{CC}} L}{4E} \\
& - |\epsilon_{\mu e}^s| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 c_{23} \sin \phi_{\mu e}^s \frac{\Delta m_{31}^2}{a_{\text{CC}}} \sin \frac{a_{\text{CC}} L}{2E} \\
& - 2 |\epsilon_{\mu e}^d| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 c_{23} \cos \phi_{\mu e}^d \frac{\Delta m_{31}^2}{a_{\text{CC}}} \left[ c_{23}^2 \sin^2 \frac{a_{\text{CC}} L}{4E} - s_{23}^2 \sin^2 \frac{\Delta m_{31}^2 L}{4E} + s_{23}^2 \sin^2 \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{4E} \right] \\
& - |\epsilon_{\mu e}^d| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 c_{23} \sin \phi_{\mu e}^d \frac{\Delta m_{31}^2}{a_{\text{CC}}} \left[ c_{23}^2 \sin \frac{a_{\text{CC}} L}{2E} + s_{23}^2 \sin \frac{\Delta m_{31}^2 L}{2E} - s_{23}^2 \sin \frac{(\Delta m_{31}^2 - a_{\text{CC}}) L}{2E} \right]
\end{aligned}$$

# Example: Approximate expression for $P_{\nu_\mu^s \rightarrow \nu_e^d}$

$$\begin{aligned}
& + 2|\epsilon_{\tau e}^d| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23} c_{23}^2 \cos \phi_{\tau e}^d \frac{\Delta m_{31}^2}{a_{CC}} \left[ \sin^2 \frac{a_{CC} L}{4E} + \sin^2 \frac{\Delta m_{31}^2 L}{4E} - \sin^2 \frac{(\Delta m_{31}^2 - a_{CC})L}{4E} \right] \\
& + |\epsilon_{\tau e}^d| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23} c_{23}^2 \sin \phi_{\tau e}^d \frac{\Delta m_{31}^2}{a_{CC}} \left[ \sin \frac{a_{CC} L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{CC})L}{2E} \right] \\
& + 4|\epsilon_{e\mu}^m| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 c_{23}^3 \cos \phi_{e\mu}^m \frac{\Delta m_{31}^2}{a_{CC}} \sin^2 \frac{a_{CC} L}{4E} \\
& - 2|\epsilon_{e\mu}^m| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23}^2 c_{23} \cos \phi_{e\mu}^m \frac{\Delta m_{31}^2}{\Delta m_{31}^2 - a_{CC}} \left[ \sin^2 \frac{a_{CC} L}{4E} - \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \frac{(\Delta m_{31}^2 - a_{CC})L}{4E} \right] \\
& + |\epsilon_{e\mu}^m| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23}^2 c_{23} \sin \phi_{e\mu}^m \frac{\Delta m_{31}^2}{\Delta m_{31}^2 - a_{CC}} \left[ \sin \frac{a_{CC} L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{CC})L}{2E} \right] \\
& - 4|\epsilon_{e\tau}^m| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23} c_{23}^2 \cos \phi_{e\tau}^m \frac{\Delta m_{31}^2}{a_{CC}} \sin^2 \frac{a_{CC} L}{4E} \\
& - 2|\epsilon_{e\tau}^m| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23} c_{23}^2 \cos \phi_{e\tau}^m \frac{\Delta m_{31}^2}{\Delta m_{31}^2 - a_{CC}} \left[ \sin^2 \frac{a_{CC} L}{4E} - \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \frac{(\Delta m_{31}^2 - a_{CC})L}{4E} \right] \\
& + |\epsilon_{e\tau}^m| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_2 \times 12 s_{23} c_{23}^2 \sin \phi_{e\tau}^m \frac{\Delta m_{31}^2}{\Delta m_{31}^2 - a_{CC}} \left[ \sin \frac{a_{CC} L}{2E} - \sin \frac{\Delta m_{31}^2 L}{2E} + \sin \frac{(\Delta m_{31}^2 - a_{CC})L}{2E} \right] \\
& + \mathcal{O}\left(\left[\frac{\Delta m_{21}^2}{\Delta m_{31}^2}\right]^3\right) + \mathcal{O}\left(\left[\frac{\Delta m_{21}^2}{\Delta m_{31}^2}\right]^2 s_{13}\right) + \mathcal{O}\left(\frac{\Delta m_{21}^2}{\Delta m_{31}^2} s_{13}^2\right) + \mathcal{O}(s_{13}^3) \\
& + \mathcal{O}\left(\epsilon \left[\frac{\Delta m_{21}^2}{\Delta m_{31}^2}\right]^2\right) + \mathcal{O}\left(\epsilon s_{13} \frac{\Delta m_{21}^2}{\Delta m_{31}^2}\right) + \mathcal{O}(\epsilon s_{13}^2) + \mathcal{O}(\epsilon^2).
\end{aligned}$$

# Outline

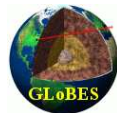
- 1 Introduction to non-standard interactions
- 2 Sensitivity of reactor and superbeam experiments**
- 3 Discovery reach of a neutrino factory
- 4 Summary and conclusions

# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0

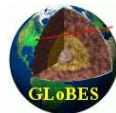
Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,

Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432



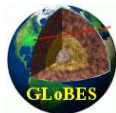
# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0  
Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,  
Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432
- Event rate based simulations (no Monte Carlo)



# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0  
Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,  
Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432
- Event rate based simulations (no Monte Carlo)
- Near detector simulated explicitly



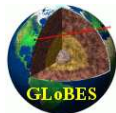
# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0

Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,

Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432

- Event rate based simulations (no Monte Carlo)
- Near detector simulated explicitly
- $\chi^2$  analysis includes





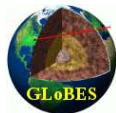
# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0

Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,

Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432

- Event rate based simulations (no Monte Carlo)
- Near detector simulated explicitly
- $\chi^2$  analysis includes
  - Systematical errors



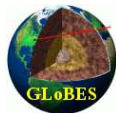
# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0

Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,

Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432

- Event rate based simulations (no Monte Carlo)
- Near detector simulated explicitly
- $\chi^2$  analysis includes
  - Systematical errors
  - Parameter correlations



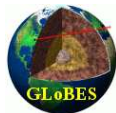
# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0

Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,

Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432

- Event rate based simulations (no Monte Carlo)
- Near detector simulated explicitly
- $\chi^2$  analysis includes
  - Systematical errors
  - Parameter correlations
  - Degeneracies

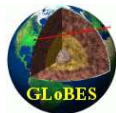


# Numerical simulation techniques

- All simulations have been performed with GLoBES 3.0

Huber Lindner Winter Comput. Phys. Commun. **167** (2005) 195,

Huber JK Lindner Rolinec Winter Comput. Phys. Commun. **177** (2007) 432



- Event rate based simulations (no Monte Carlo)
- Near detector simulated explicitly
- $\chi^2$  analysis includes
  - Systematical errors
  - Parameter correlations
  - Degeneracies
  - External input on those parameters which cannot be determined by the experiment under consideration

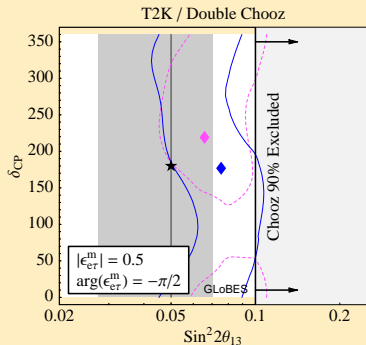
# NSI-induced mismatches and offsets

$P_{\bar{\nu}_e^s \rightarrow \bar{\nu}_e^d}$  (reactor) and  $P_{\nu_\mu^s \rightarrow \nu_e^d}$  (superbeam) are affected differently by NSI  
⇒ Naive standard oscillation fits respond differently

# NSI-induced mismatches and offsets

$P_{\bar{\nu}_e^s \rightarrow \bar{\nu}_e^d}$  (reactor) and  $P_{\nu_\mu^s \rightarrow \nu_e^d}$  (superbeam) are affected differently by NSI  
 $\Rightarrow$  Naive standard oscillation fits respond differently

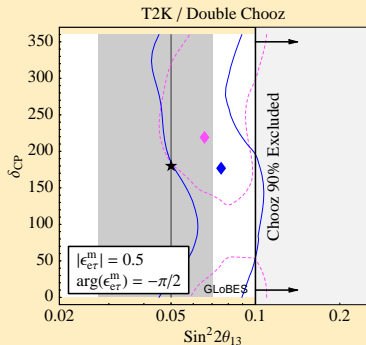
## Mismatch



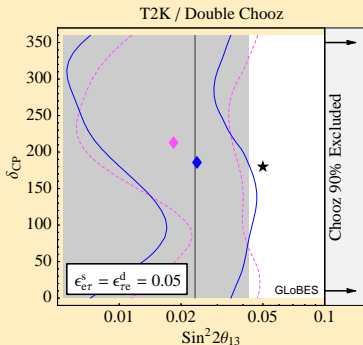
# NSI-induced mismatches and offsets

$P_{\bar{\nu}_e^s \rightarrow \bar{\nu}_e^d}$  (reactor) and  $P_{\nu_\mu^s \rightarrow \nu_e^d}$  (superbeam) are affected differently by NSI  
 $\Rightarrow$  Naive standard oscillation fits respond differently

## Mismatch



## Offset



# A more systematic approach

- Choose random values for the NSI parameters  $\varepsilon_{\alpha\beta}^s$ ,  $\varepsilon_{\alpha\beta}^d$ ,  $\varepsilon_{\alpha\beta}^m$   
( $|\varepsilon|$  distributed logarithmically,  $\arg(\varepsilon)$  distributed linearly)



# A more systematic approach

- Choose random values for the NSI parameters  $\varepsilon_{\alpha\beta}^s$ ,  $\varepsilon_{\alpha\beta}^d$ ,  $\varepsilon_{\alpha\beta}^m$   
 ( $|\varepsilon|$  distributed logarithmically,  $\arg(\varepsilon)$  distributed linearly)
- Take into account specific constraints  $\varepsilon_{e\alpha}^s = \varepsilon_{\alpha e}^{d*}$  and  $|\varepsilon_{\mu\alpha}^s| \gtrsim |\varepsilon_{\alpha\mu}^d|$

# A more systematic approach

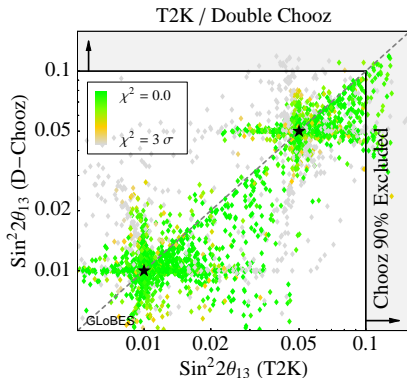
- Choose random values for the NSI parameters  $\varepsilon_{\alpha\beta}^s$ ,  $\varepsilon_{\alpha\beta}^d$ ,  $\varepsilon_{\alpha\beta}^m$   
 ( $|\varepsilon|$  distributed logarithmically,  $\arg(\varepsilon)$  distributed linearly)
- Take into account specific constraints  $\varepsilon_{e\alpha}^s = \varepsilon_{\alpha e}^{d*}$  and  $|\varepsilon_{\mu\alpha}^s| \gtrsim |\varepsilon_{\alpha\mu}^d|$
- Simulate event spectra

# A more systematic approach

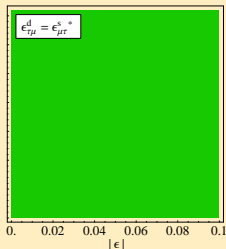
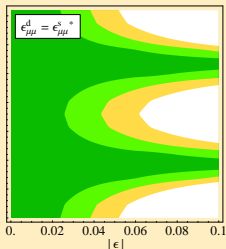
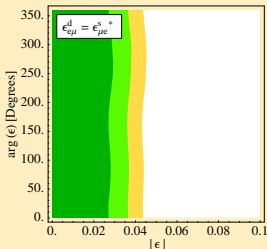
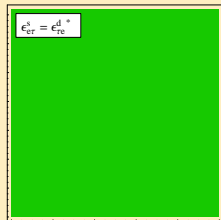
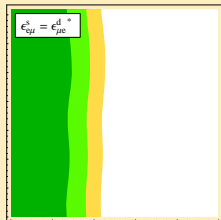
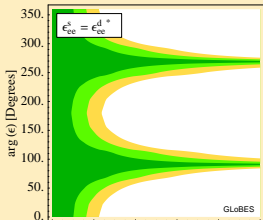
- Choose random values for the NSI parameters  $\varepsilon_{\alpha\beta}^s$ ,  $\varepsilon_{\alpha\beta}^d$ ,  $\varepsilon_{\alpha\beta}^m$   
 ( $|\varepsilon|$  distributed logarithmically,  $\arg(\varepsilon)$  distributed linearly)
- Take into account specific constraints  $\varepsilon_{e\alpha}^s = \varepsilon_{\alpha e}^{d*}$  and  $|\varepsilon_{\mu\alpha}^s| \gtrsim |\varepsilon_{\alpha\mu}^d|$
- Simulate event spectra
- Perform standard oscillation fit

# A more systematic approach

- Choose random values for the NSI parameters  $\varepsilon_{\alpha\beta}^s$ ,  $\varepsilon_{\alpha\beta}^d$ ,  $\varepsilon_{\alpha\beta}^m$  ( $|\varepsilon|$  distributed logarithmically,  $\arg(\varepsilon)$  distributed linearly)
- Take into account specific constraints  $\varepsilon_{e\alpha}^s = \varepsilon_{\alpha e}^{d*}$  and  $|\varepsilon_{\mu\alpha}^s| \gtrsim |\varepsilon_{\alpha\mu}^d|$
- Simulate event spectra
- Perform standard oscillation fit

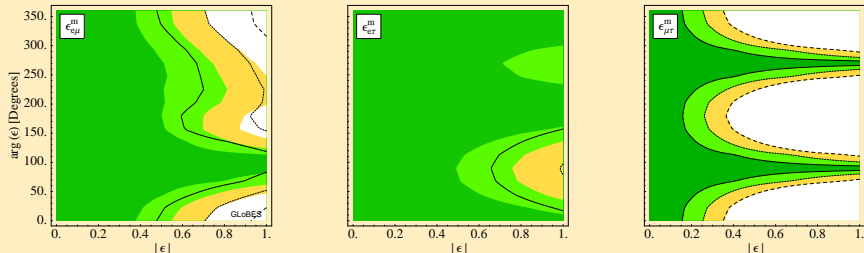


## Discovery reach of combined T2K/D-Chooz analysis

NSI in source and detector (current bound  $\sim 0.1$ )

## Discovery reach of combined T2K/D-Chooz analysis

## NSI in propagation

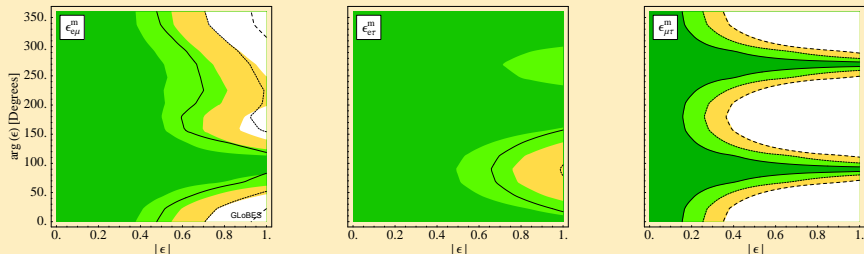


- Current bounds:  $|\varepsilon_{e\mu}^m| \lesssim 5 \times 10^{-4}$ ,  $|\varepsilon_{e\tau}^m| \lesssim 0.7$ ,  $|\varepsilon_{\mu\tau}^m| \lesssim 0.1$

Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011

## Discovery reach of combined T2K/D-Chooz analysis

## NSI in propagation



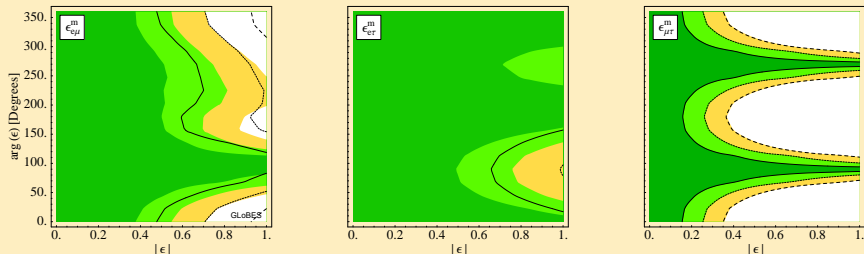
- Current bounds:  $|\varepsilon_{e\mu}^m| \lesssim 5 \times 10^{-4}$ ,  $|\varepsilon_{e\tau}^m| \lesssim 0.7$ ,  $|\varepsilon_{\mu\tau}^m| \lesssim 0.1$

Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011

- Note: Discovery reach depends strongly on  $\arg \varepsilon$   
 $\leftrightarrow P_{\nu_\alpha^s \rightarrow \nu_\beta^d}$  depends strongly on  $\arg \varepsilon$

## Discovery reach of combined T2K/D-Chooz analysis

## NSI in propagation



- Current bounds:  $|\varepsilon_{e\mu}^m| \lesssim 5 \times 10^{-4}$ ,  $|\varepsilon_{e\tau}^m| \lesssim 0.7$ ,  $|\varepsilon_{\mu\tau}^m| \lesssim 0.1$

Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011

- Note: Discovery reach depends strongly on  $\arg \varepsilon$   
 $\leftrightarrow P_{\nu_\alpha^s \rightarrow \nu_\beta^d}$  depends strongly on  $\arg \varepsilon$
- Optimal discovery reach requires reactor, superbeam, and near detectors



# Outline

- 1 Introduction to non-standard interactions
- 2 Sensitivity of reactor and superbeam experiments
- 3 Discovery reach of a neutrino factory**
- 4 Summary and conclusions

# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.

# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.
- Very challenging due to strong constraint  $\varepsilon_{e\mu}^m \lesssim 5 \times 10^{-4}$  (90% C.L.) from charged lepton flavour violation. Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011

# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.
- Very challenging due to strong constraint  $\varepsilon_{e\mu}^m \lesssim 5 \times 10^{-4}$  (90% C.L.) from charged lepton flavour violation. Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011
- Neutrino factory setup

# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.
- Very challenging due to strong constraint  $\varepsilon_{e\mu}^m \lesssim 5 \times 10^{-4}$  (90% C.L.) from charged lepton flavour violation. Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011
- Neutrino factory setup
  - Baseline 3000 km (1 detector only)

# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.
- Very challenging due to strong constraint  $\varepsilon_{e\mu}^m \lesssim 5 \times 10^{-4}$  (90% C.L.) from charged lepton flavour violation. Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011
- Neutrino factory setup
  - Baseline 3000 km (1 detector only)
  - Detector: 50 kt magnetized iron calorimeter

# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.
- Very challenging due to strong constraint  $\varepsilon_{e\mu}^m \lesssim 5 \times 10^{-4}$  (90% C.L.) from charged lepton flavour violation. Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011
- Neutrino factory setup
  - Baseline 3000 km (1 detector only)
  - Detector: 50 kt magnetized iron calorimeter
  - Parent muon energy: 50 GeV

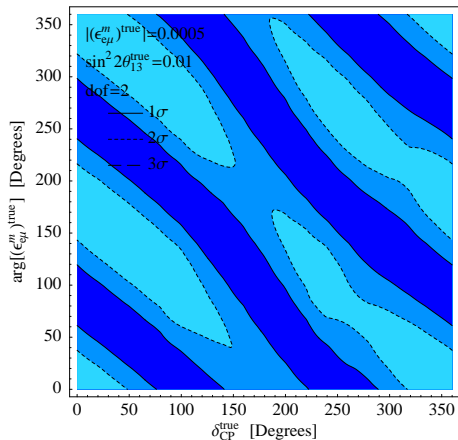
# Discovery reach for $\varepsilon_{e\mu}^m$ in a neutrino factory

- Discovery reach for  $\varepsilon_{e\mu}^m$ : Minimum value of  $|\varepsilon_{e\mu}^m|$  which can no longer be fitted with  $\varepsilon_{e\mu}^m = 0$  at a given C.L.
- Very challenging due to strong constraint  $\varepsilon_{e\mu}^m \lesssim 5 \times 10^{-4}$  (90% C.L.) from charged lepton flavour violation. Davidson Pena-Garay Rius Santamaria JHEP **03** (2003) 011
- Neutrino factory setup
  - Baseline 3000 km (1 detector only)
  - Detector: 50 kt magnetized iron calorimeter
  - Parent muon energy: 50 GeV
  - Stored muons:  $4 \times 10^{21} \mu^+$ ,  $4 \times 10^{21} \mu^-$



Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 0.5 \times 10^{-3}$
- $|\varepsilon_{e\mu}^m|$  too small



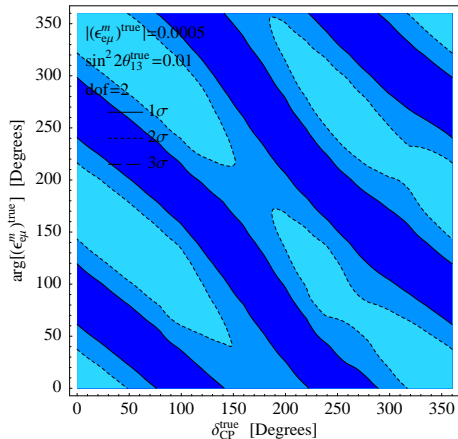
Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 0.5 \times 10^{-3}$

- $|\varepsilon_{e\mu}^m|$  too small

↓

$\chi^2$  of standard oscillation fit is below  $3\sigma$  for all true values of  $\delta_{\text{CP}}$  and  $\arg \varepsilon_{e\mu}^m$ .



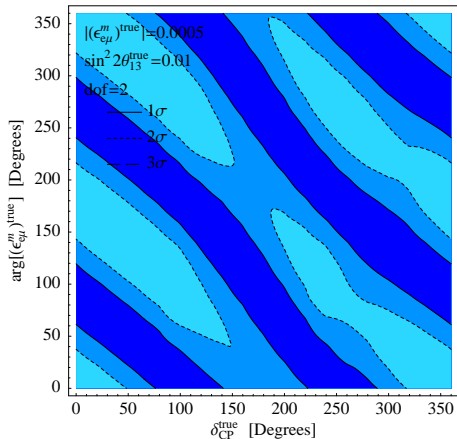
Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 0.5 \times 10^{-3}$

- $|\varepsilon_{e\mu}^m|$  too small

↓  
 $\chi^2$  of standard oscillation fit is  
 below  $3\sigma$  for all true values of  $\delta_{\text{CP}}$   
 and  $\arg \varepsilon_{e\mu}^m$ .

↓  
**No chance for discovery**

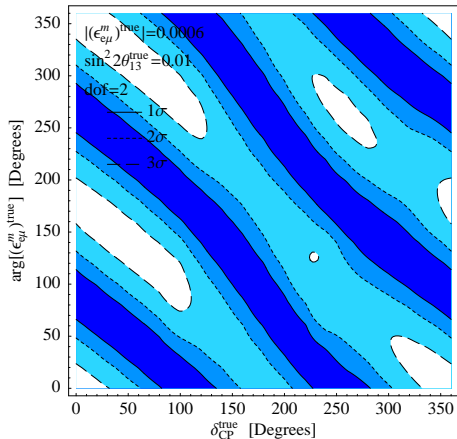


Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 0.6 \times 10^{-3}$
- $|\varepsilon_{e\mu}^m|$  becomes larger

↓  
 For some combinations of  $\delta_{\text{CP}}$  and  $\arg \varepsilon_{e\mu}^m$ , the standard oscillation fit becomes worse than  $3\sigma$  (white islands appear).

↓  
 Discovery possible for favorable phase combinations



Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 1 \times 10^{-3}$

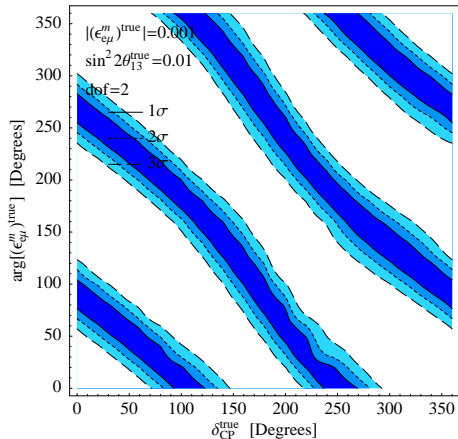
- $|\varepsilon_{e\mu}^m|$  becomes larger

↓

For some combinations of  $\delta_{\text{CP}}$  and  $\arg \varepsilon_{e\mu}^m$ , the standard oscillation fit becomes worse than  $3\sigma$  (white islands appear).

↓

Discovery possible for favorable phase combinations



Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 2 \times 10^{-3}$

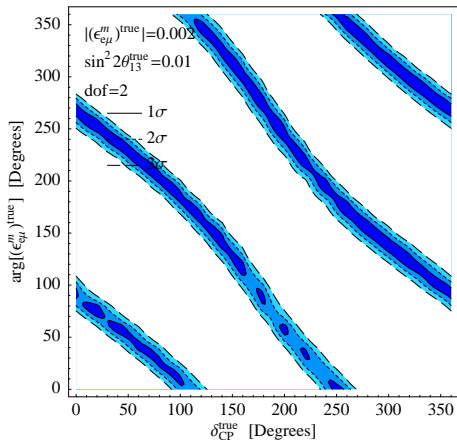
- $|\varepsilon_{e\mu}^m|$  becomes larger

↓

For some combinations of  $\delta_{\text{CP}}$  and  $\arg \varepsilon_{e\mu}^m$ , the standard oscillation fit becomes worse than  $3\sigma$  (white islands appear).

↓

Discovery possible for favorable phase combinations



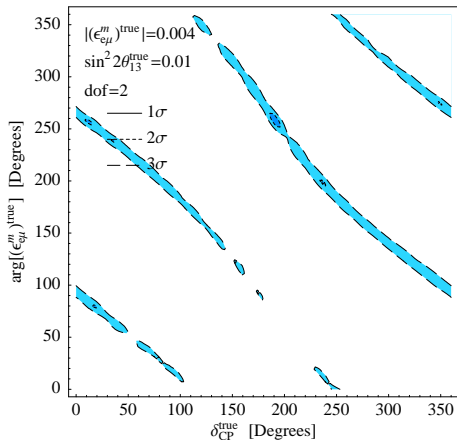
Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 4 \times 10^{-3}$

- $|\varepsilon_{e\mu}^m|$  becomes larger

↓  
 For some combinations of  $\delta_{\text{CP}}$  and  $\arg \varepsilon_{e\mu}^m$ , the standard oscillation fit becomes worse than  $3\sigma$  (white islands appear).

↓  
 Discovery possible for favorable phase combinations



Discovery reach for  $\varepsilon_{e\mu}^m$  in a neutrino factory

- $|\varepsilon_{e\mu}^m| = 5 \times 10^{-3}$

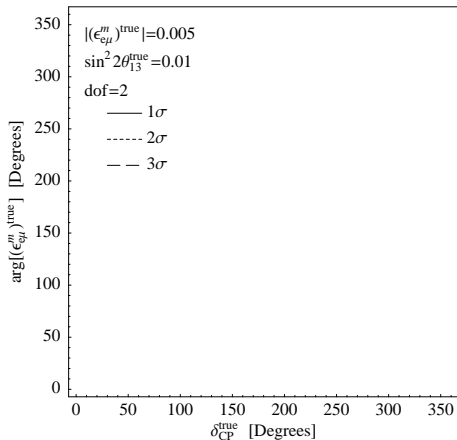
- $|\varepsilon_{e\mu}^m|$  is large enough

↓

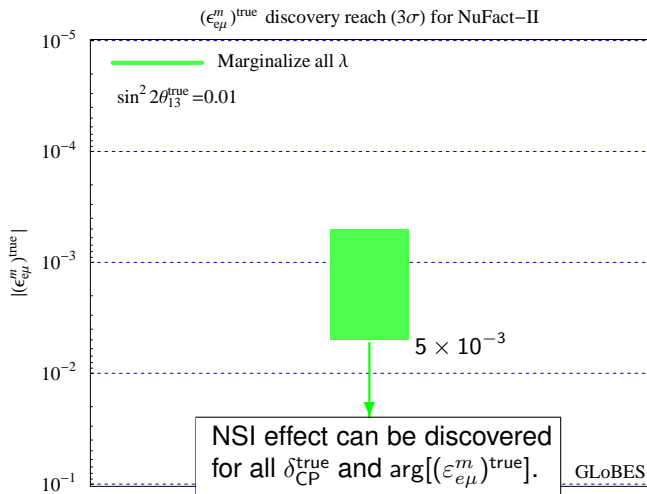
$\chi^2$  of standard oscillation fit exceeds  $3\sigma$  in the whole parameter plane.

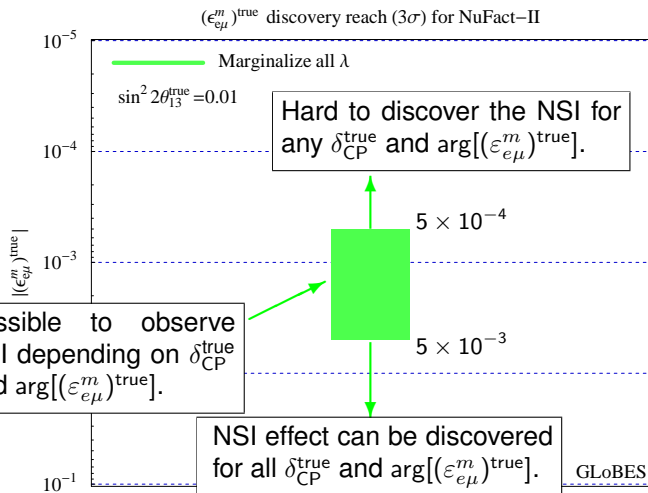
↓

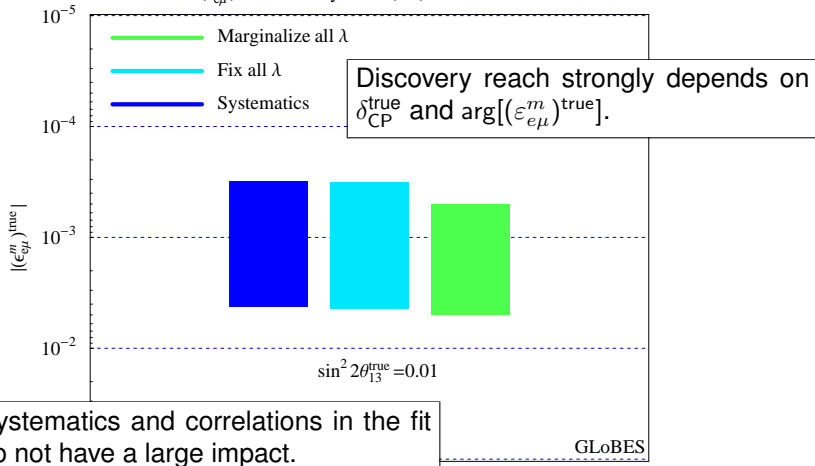
Discovery is possible for any phase combination

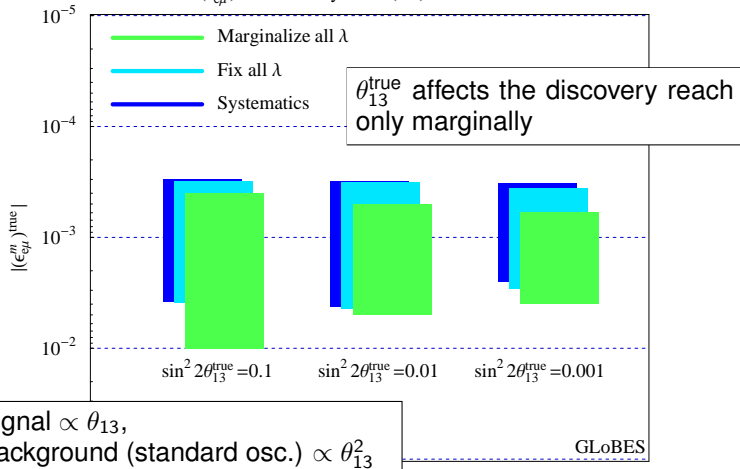


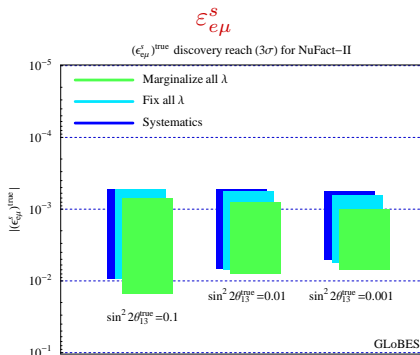
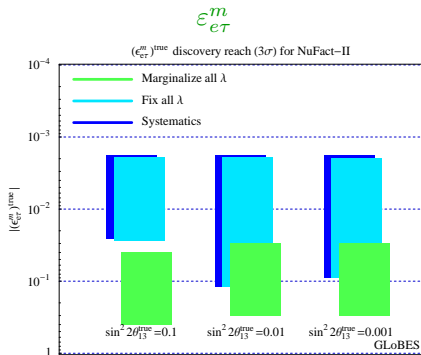


Summary of the discovery reach for  $\varepsilon_{e\mu}^m$ 

Summary of the discovery reach for  $\epsilon_{e\mu}^m$ 

Summary of the discovery reach for  $\epsilon_{e\mu}^m$  $(\epsilon_{e\mu}^m)^{\text{true}}$  discovery reach ( $3\sigma$ ) for NuFact-II

Summary of the discovery reach for  $\epsilon_{e\mu}^m$  $(\epsilon_{e\mu}^m)^{\text{true}}$  discovery reach ( $3\sigma$ ) for NuFact-II

Discovery reach for  $\epsilon_{e\tau}^m$  and  $\epsilon_{e\mu}^s$  in a neutrino factory

# Outline

- 1 Introduction to non-standard interactions
- 2 Sensitivity of reactor and superbeam experiments
- 3 Discovery reach of a neutrino factory
- 4 Summary and conclusions**

# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments

# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments
- In superbeam and reactor experiments: Sizeable effects only if  $\epsilon$  parameters are very large (i.e. only in very obscure models, since  $|\epsilon| \sim M_W^2/M_{\text{NSI}}^2$ ).



# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments
- In superbeam and reactor experiments: Sizeable effects only if  $\varepsilon$  parameters are very large (i.e. only in very obscure models, since  $|\varepsilon| \sim M_W^2/M_{\text{NSI}}^2$ ).
- Neutrino factory probes more interesting  $M_{\text{NSI}}$  range.

# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments
- In superbeam and reactor experiments: Sizeable effects only if  $\varepsilon$  parameters are very large (i.e. only in very obscure models, since  $|\varepsilon| \sim M_W^2/M_{\text{NSI}}^2$ ).
- Neutrino factory probes more interesting  $M_{\text{NSI}}$  range.
- Possible consequences of NSI are:

# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments
- In superbeam and reactor experiments: Sizeable effects only if  $\varepsilon$  parameters are very large (i.e. only in very obscure models, since  $|\varepsilon| \sim M_W^2/M_{\text{NSI}}^2$ ).
- Neutrino factory probes more interesting  $M_{\text{NSI}}$  range.
- Possible consequences of NSI are:
  - **Poor quality** of standard oscillation fit ( $\Rightarrow$  Detection of NSI possible)

# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments
- In superbeam and reactor experiments: Sizeable effects only if  $\varepsilon$  parameters are very large (i.e. only in very obscure models, since  $|\varepsilon| \sim M_W^2/M_{\text{NSI}}^2$ ).
- Neutrino factory probes more interesting  $M_{\text{NSI}}$  range.
- Possible consequences of NSI are:
  - **Poor quality** of standard oscillation fit ( $\Rightarrow$  Detection of NSI possible)
  - **Mismatch** between standard fits to different experiments

# Summary and conclusions

- Low-energy fingerprints of “new physics” can be seen in oscillation experiments
- In superbeam and reactor experiments: Sizeable effects only if  $\varepsilon$  parameters are very large (i.e. only in very obscure models, since  $|\varepsilon| \sim M_W^2/M_{\text{NSI}}^2$ ).
- Neutrino factory probes more interesting  $M_{\text{NSI}}$  range.
- Possible consequences of NSI are:
  - **Poor quality** of standard oscillation fit ( $\Rightarrow$  Detection of NSI possible)
  - **Mismatch** between standard fits to different experiments
  - **Offset**: Consistent, but wrong reconstruction of neutrino mixing parameters