

Updated Constraints from B Physics on the MSSM and the NMSSM.¹

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¹F. Domingo and U. Ellwanger, arXiv:0710.3714 [hep-ph].

Recent Experimental and SM Achievements on B Physics Observables

$BR(B \rightarrow X_s \gamma)$

- Status until 2005:

$$BR(B \rightarrow X_s \gamma)|_{E_\gamma > 1.6 \text{ GeV}}^{exp.} = (3.23 \pm 0.42) \cdot 10^{-4} \quad \text{Exp. World Average, 2001}$$

$$BR(B \rightarrow X_s \gamma)|_{E_\gamma > 1.6 \text{ GeV}}^{NLO} = (3.61 \pm 0.5) \cdot 10^{-4} \quad \text{SM NLO prediction, [Hurth et al., 2005]}$$

- Present situation:

$$BR(B \rightarrow X_s \gamma)|_{E_\gamma > 1.6 \text{ GeV}}^{exp.} = (3.55 \pm 0.24_{-0.10}^{+0.09} \pm 0.03) \cdot 10^{-4} \quad \text{Exp. World Average, 2005}$$

$$BR(B \rightarrow X_s \gamma)|_{E_\gamma > 1.6 \text{ GeV}}^{NNLO} = (3.15 \pm 0.23) \cdot 10^{-4} \quad \text{SM NNLO, [Misiak et al., 2006]}$$

$$BR(B \rightarrow X_s \gamma)|_{E_\gamma > 1.6 \text{ GeV}}^{NNLO} = (2.98 \pm 0.26) \cdot 10^{-4} \quad \text{SM NNLO, [Becher, Neubert, 2006]}$$

(using an improved treatment of the photon energy cutoff)

$$\Delta M_d = M_{B_d} - M_{\bar{B}_d}, \Delta M_s = M_{B_s} - M_{\bar{B}_s}$$

- ΔM_d :

$$\Delta M_d^{exp} = (0.507 \pm 0.004) .ps^{-1} \quad \text{HFAG, 2005}$$

$$\Delta M_d^{SM} = (0.59 \pm 0.19) .ps^{-1} \quad \text{SM prediction}$$

- ΔM_s :

$$\Delta M_s^{exp} = (17.77 \pm 0.12) .ps^{-1} \quad \text{CDF, 2006}$$

$$\Delta M_s^{SM} = (20.5 \pm 3.1) .ps^{-1} \quad \text{SM prediction}$$

$$BR(\bar{B}_s \rightarrow \mu^+ \mu^-)$$

$$BR(\bar{B}_s \rightarrow \mu^+ \mu^-)|_{exp} < 5.8 \times 10^{-8} \quad (95\% C.L.) \quad \text{CDF, 2007}$$

$$BR(\bar{B}_s \rightarrow \mu^+ \mu^-)|_{SM} = (3.8 \pm 0.1) \times 10^{-9} \quad [\text{Dedes et al., 2002}]$$

$$BR(\bar{B}^+ \rightarrow \tau^+ \nu_\tau)$$

$$BR(\bar{B}^+ \rightarrow \tau^+ \nu_\tau)|_{exp} = (1.32 \pm 0.49) .10^{-4} \quad \text{HFAG, 2005}$$

$$BR(\bar{B}^+ \rightarrow \tau^+ \nu_\tau)|_{SM} \text{ from } (0.85 \pm 0.13) .10^{-4} \quad [\text{Carena et al., 2007}] \quad |V_{ub}|_{excl.} \sim 3.7 \times 10^{-3}$$

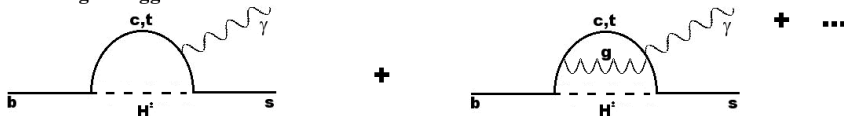
$$\text{to } (1.59 \pm 0.40) .10^{-4} \quad [\text{Isidori, Paradisi, 2006}] \quad |V_{ub}|_{incl.} \sim 4.4 \times 10^{-3}$$

Contributions to $BR(b \rightarrow s \gamma)$

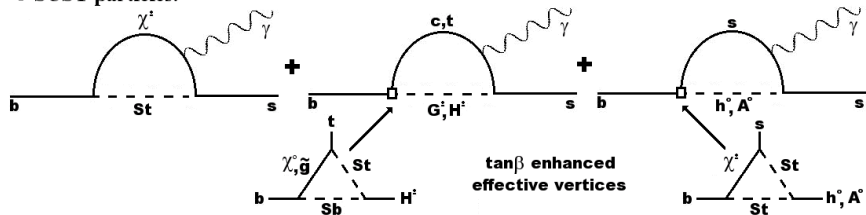
● SM:



● Charged Higgs:



● SUSY particles:



$BR(\bar{B} \rightarrow X_s \gamma)$: Formalism [Gambino, Misiak, 2001]; [Hurth, Lunghi, Porod, 2005]

Formula derived in a low-energy Effective Theory (matching scale $\mu_o \equiv m_t$):

$$BR(\bar{B} \rightarrow X_s \gamma) \Big|_{E_\gamma > E_o} = \frac{6\alpha_{em}}{\pi C} \left| \frac{V_{ts}^* V_{tb}}{V_{cb}} \right|^2 BR(\bar{B} \rightarrow X_c e \bar{\nu}) \times \left[K_c + \frac{m_b(m_t)}{m_b^{1S}} (K_t + K_{BSM}) + \epsilon_{ew} \right]^2 + B(E_o) + N$$

Quantities Involved:

- $K_c + \frac{m_b(m_t)}{m_b^{1S}} (K_t + K_{BSM})$: NLO QCD partonic amplitude for $b \rightarrow s\gamma$. Ambiguous dependence at NLO on $\frac{m_c}{m_b} = 0.23_{-0.05}^{+0.08}$; SM NNLO results reproduced for $\frac{m_c}{m_b} \simeq 0.307$;
- $BR(\bar{B} \rightarrow X_c e \bar{\nu})$: measured experimentally, $\simeq 0.1061$;
- $C = \left| \frac{V_{ub}}{V_{cb}} \right|^2 \frac{\Gamma(\bar{B} \rightarrow X_c e \bar{\nu})}{\Gamma(\bar{B} \rightarrow X_u e \bar{\nu})}$, calculable $\simeq 0.580$;
- N : Non-Perturbative corrections (Heavy Quark Effective Theory) $\sim \frac{\Lambda_{QCD}^2}{m_c^2}$ (+ higher orders);
- $B(E_o)$: (gluon) Bremsstrahlung corrections, depends on the lower limit E_o on the photonic energy E_γ . Here: $E_o = 1.6 GeV$;
- ϵ_{ew} : electroweak radiative corrections;

Calculation of ΔM_q , $q = d, s$

Formula for the Mass difference, [Buras *et al.*, 2003]

$$\Delta M_q = \frac{G_F^2 M_W^2}{6\pi^2} M_{B_q} \eta_B f_{B_q}^2 \hat{B}_{B_q} |V_{tq}^* V_{tb}|^2 |F_{tt}^q|$$

with:

$$F_{tt}^q = S_0(x_t) + \frac{1}{4r} C_{new}^{VLL} + \bar{P}_1^{SLL} (C_1^{SLL} + C_1^{SRR}) + \bar{P}_2^{LR} C_2^{LR} + \dots$$

Parameters

- Hadronic parameters (lattice QCD):

$$f_{B_s} \sqrt{\hat{B}_{B_s}} = (0.281 \pm 0.021) \text{ GeV} \quad [\text{Dalgic et al., 2007}]$$

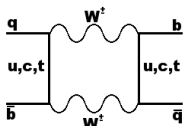
$$f_{B_s} \sqrt{\hat{B}_{B_s}} / f_{B_d} \sqrt{\hat{B}_{B_d}} = 1.216 \pm 0.041 \quad [\text{Okamoto, 2006}]$$

- CKM, from Tree Level measurements [Ball, Fleisher, 2006]:

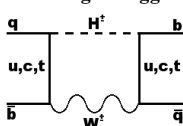
$$|V_{td}^* V_{tb}| = (8.6 \pm 1.4) \cdot 10^{-3}$$

$$|V_{ts}^* V_{tb}| = (41.3 \pm 0.7) \cdot 10^{-3}$$

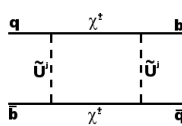
- SM



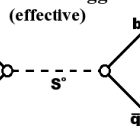
- Charged Higgs



- SUSY



-Neutral Higgs



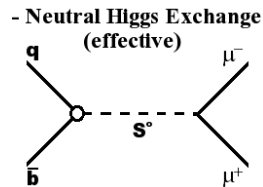
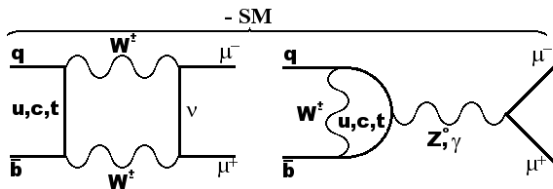
Branching Ratio $\bar{B}_s \rightarrow \mu^+ \mu^-$

Formula [Bobeth et al., 2002]

$$BR(\bar{B}_s \rightarrow \mu^+ \mu^-) = \frac{G_F^2 \alpha^2 M_{B_s}^5 f_{B_s}^2 \tau_{B_s}}{64 \pi^3 \sin^4 \theta_W} |V_{tb} V_{ts}^*|^2 \sqrt{1 - 4 \frac{m_\mu^2}{M_{B_s}^2}} \left[\frac{1 - 4 \frac{m_\mu^2}{M_{B_s}^2}}{\left(1 + \frac{m_s}{m_b}\right)^2} |c_S|^2 + \left| \frac{c_P}{1 + \frac{m_s}{m_b}} + \frac{2m_\mu}{M_{B_s}^2} c_A \right|^2 \right]$$

Effective coefficients

- SM contribution in c_A : 1 order of magnitude below the sensitivity of experiments;
- Effective Neutral Higgs contributions in c_S , c_P : enhanced for light scalars/large $\tan\beta$.



Branching ratio $\bar{B}^+ \rightarrow \tau^+ \nu_\tau$

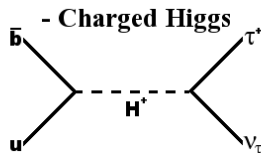
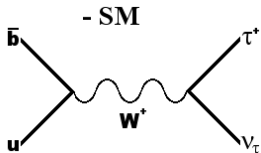
Formula [Akeroyd, Recksiegel, 2003]

$$BR(\bar{B}^+ \rightarrow \tau^+ \nu_\tau) = \frac{G_F^2 M_B m_\tau^2}{8\pi} \left(1 - \frac{m_\tau^2}{M_B^2}\right)^2 f_B^2 |V_{ub}|^2 \tau_B r_H$$

$$r_H = \left[1 - \left(\frac{M_B}{m_{H^\pm}}\right)^2 \frac{\tan^2 \beta}{1 + \tilde{\epsilon}_0 \tan \beta}\right]^2$$

Parameters

- Hadronic parameter: $f_B = (0.216 \pm 0.022) \text{ GeV}$, HPQCD, 2005;
- CKM: large uncertainty; we take $|V_{ub}| = (4.0 \pm 0.35) \cdot 10^{-3}$.



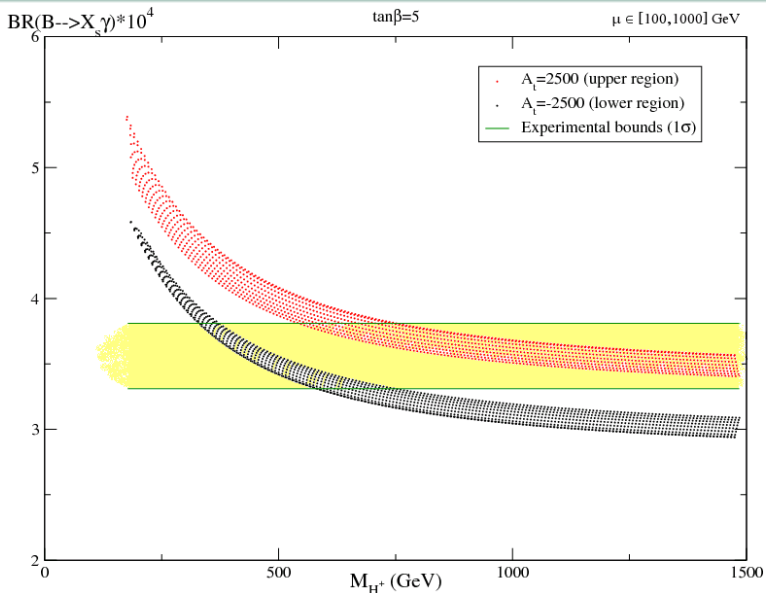
NMSSM specific effects relative to B constraints with respect to the MSSM...

Differences of the Model

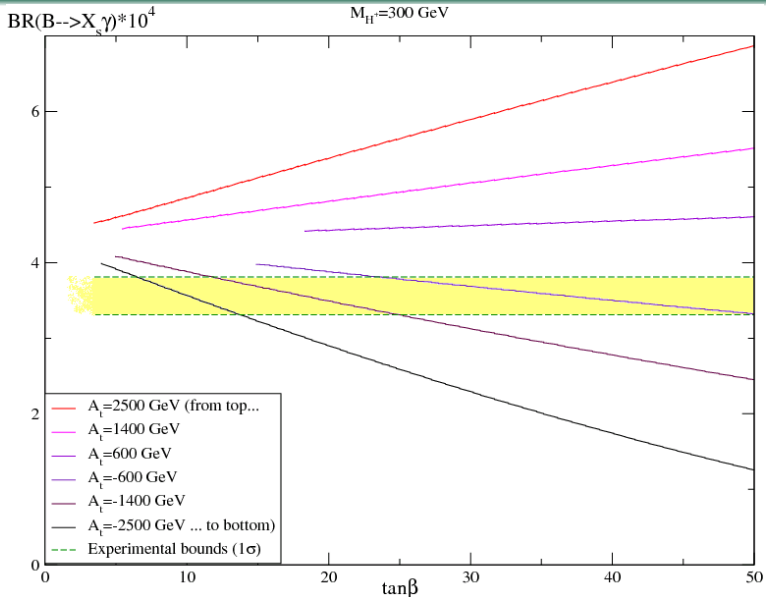
- Superfield content: $H_d, H_u, Q_L, U_R^c, D_R^c, L_L, E_R^c$ + 1 singlet S ;
- Superpotential: $W = \lambda S H_u \cdot H_d + \frac{\kappa}{3} S^3 + Y_u Q_L \cdot H_u U_R^c - Y_d Q_L \cdot H_d D_R^c - Y_e L_L \cdot H_d E_R^c$;
- MSSM " μ " term generated when $\langle S \rangle \neq 0$; possible solution of the " μ -problem".

Peculiarities concerning B processes

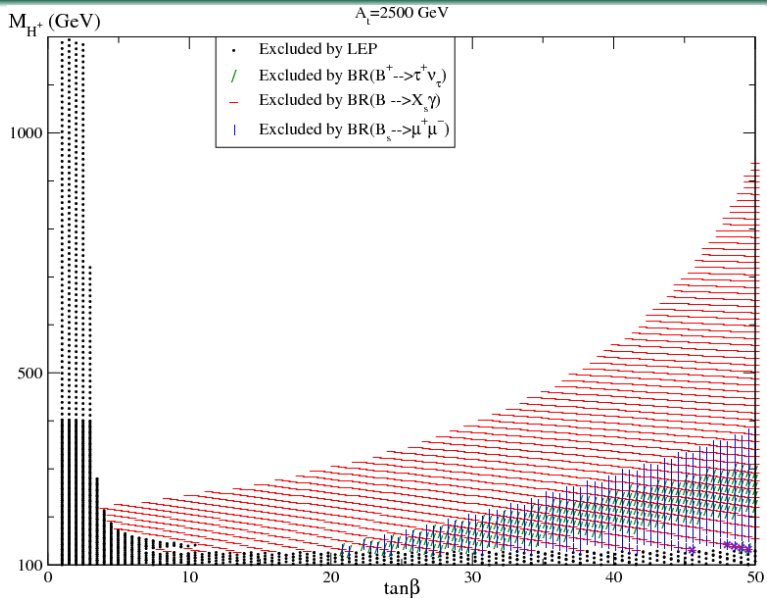
- Extended Unconstrained Parameter Space: in the NMSSM, low values of $\tan\beta$ (~ 1.5) are not excluded by LEP;
- Charged Higgs Mass: the NMSSM parameter λ gives a negative contribution to M_{H^\pm} , which allows for slight modulations on $\bar{B} \rightarrow X_s \gamma$;
- The effect of the extended neutralino sector is negligibly small;
- Light pseudoscalars (below 10 GeV) escape LEP constraints, but they are significantly constrained by ΔM_q and $BR(\bar{B}_s \rightarrow \mu^+ \mu^-)$.

$BR(\bar{B} \rightarrow X_s \gamma)$ as a function of M_{H^\pm} 

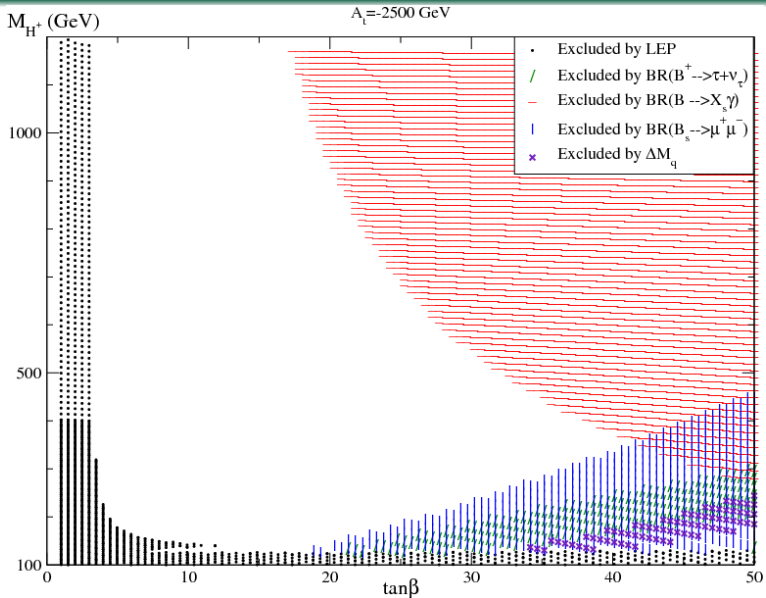
$BR(\bar{B} \rightarrow X_s \gamma)$ as a function of $\tan \beta$



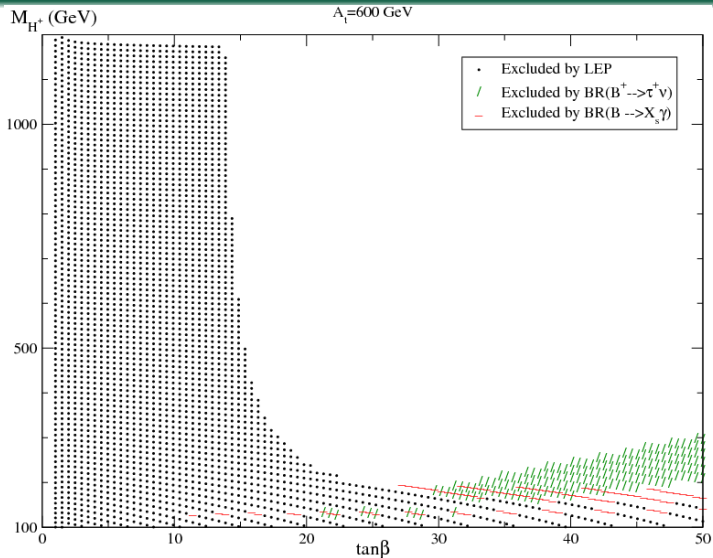
B constraints on the $(M_{H^+}, \tan\beta)$ plane



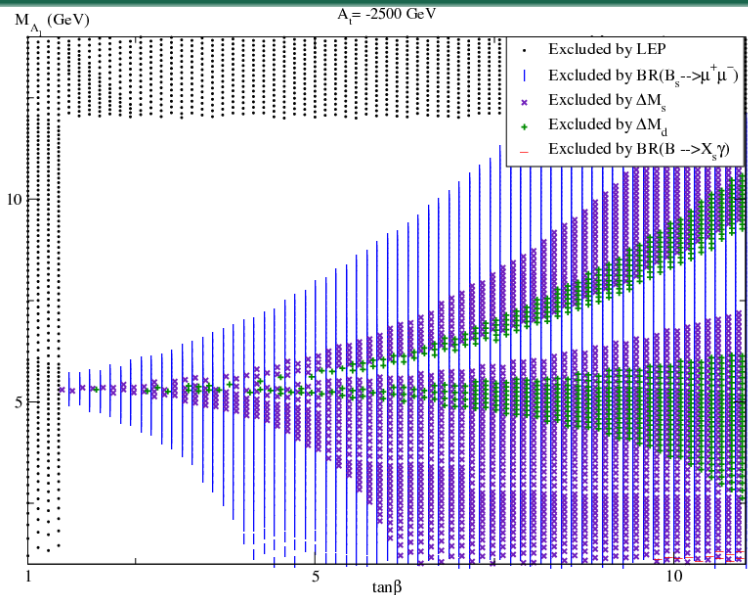
Playing SUSY against Charged Higgs Contribution

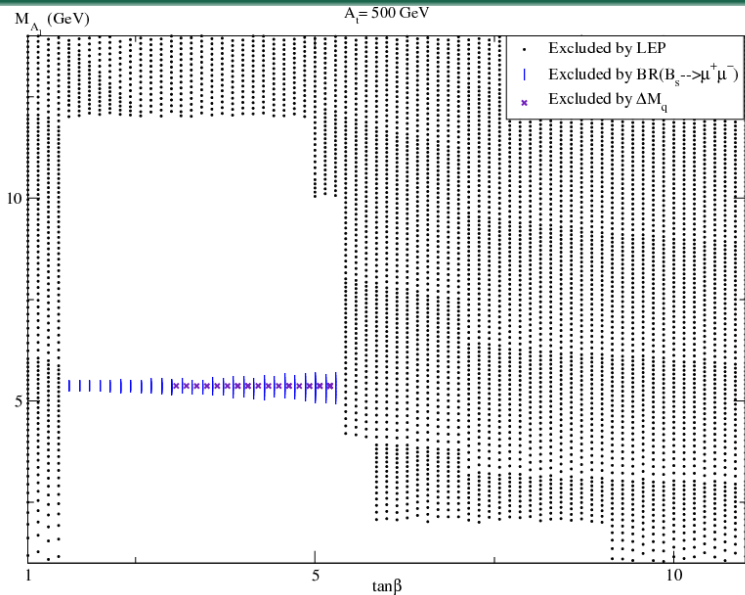


Low values of A_t : weaker B constraints vs enhanced LEP bounds



NMSSM Light Pseudoscalars



Low A_t 

Conclusions

- Constraints from $\bar{B} \rightarrow X_s \gamma$ are weaker than they used to be thanks to the recent improvements on the experimental side and the SM analysis.
- Still, very light ($\leq 200 \text{ GeV}$) charged Higgs lead to difficulties for low $\tan\beta$. Domains with both large A_t and large $\tan\beta$ are also strongly constrained.
- $BR(\bar{B}_s \rightarrow \mu^+ \mu^-)$ is the most sensitive observable depending on neutral scalar exchanges, provides us with significant constraints especially for light scalars.
- The Fortran code will be added to the NMSSMTools package (and could also be used for the MSSM...). Such codes for the MSSM: FeynHiggs, Suspect, MicrOmegas, Spheeno.

