

CP-violation from Non-Unitary Leptonic Mixing

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S. Antusch, C. Biggio, E. Fernández-Martínez and M.B. Gavela
[JHEP 0610:084,2006](#)

E. Fdez-Martinez, M.B. Gavela and O. Yasuda
[Phys.Lett.B 649:427-435,2007](#)

Motivations

- Neutrino masses and mixing → evidence of Physics Beyond the SM
- Typical explanations involve New Physics at higher energies
- This NP often induces deviations from **unitarity** of the **PMNS** at low energy



We will analyze the present constraints on the mixing matrix
without assuming **unitarity**

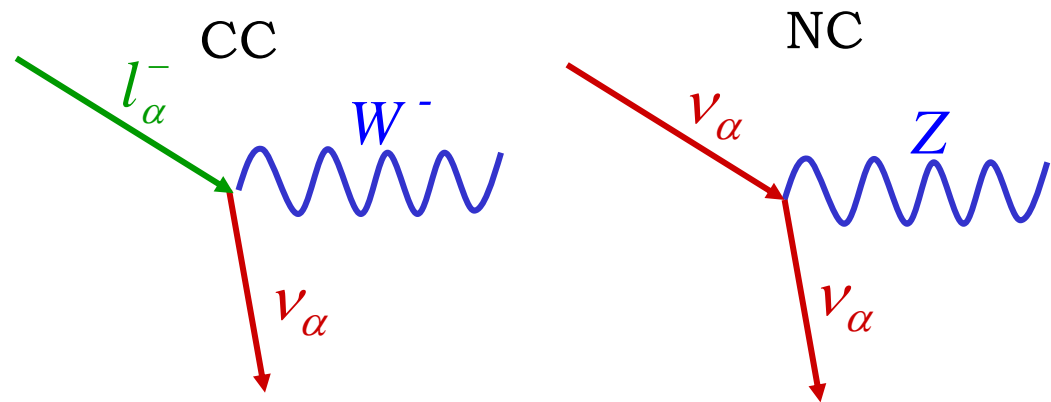
Study of **CP violation** in the context of **non-unitarity**

Usual Analysis

$$L = \frac{1}{2} \left(i\bar{\nu}_\alpha \partial \nu_\alpha - \bar{\nu}_\alpha M_{\alpha\beta} \nu_\beta + h.c. \right) - \underbrace{\frac{g}{\sqrt{2}} \left(W_\mu^+ \bar{l}_\alpha \gamma^\mu P_L \nu_\alpha + h.c. \right)}_{\text{CC}} - \underbrace{\frac{g}{\cos \theta_W} \left(Z_\mu \bar{\nu}_\alpha \gamma^\mu P_L \nu_\alpha + h.c. \right)}_{\text{NC}} + \dots$$



Diagonalizing

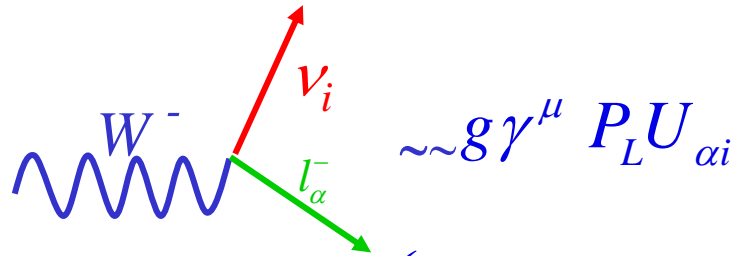


$$L = \frac{1}{2} \left(i\bar{\nu}_i \partial \nu_i - \bar{\nu}_i m_i \nu_i \right) - \frac{g}{\sqrt{2}} \left(W_\mu^+ \bar{l}_\alpha \gamma^\mu P_L \underbrace{U_{\alpha i}} \nu_i \right) - \frac{g}{\cos \theta_W} \left(Z_\mu \bar{\nu}_i \gamma^\mu P_L \nu_i \right) + h.c. + \dots$$

$$\nu_\alpha = U_{\alpha i} \nu_i$$

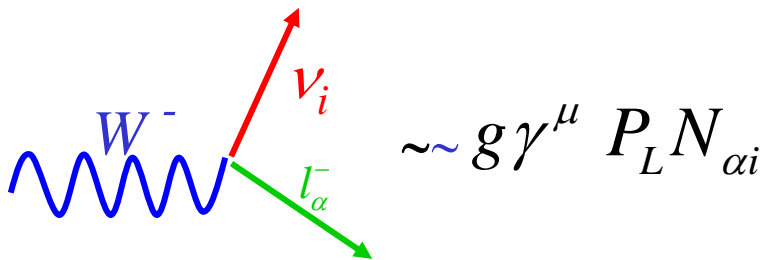
The mixing matrix U is unitary in the usual analysis with 3 light ν

The general idea



$$\sim g \gamma^\mu P_L U_{\alpha i}$$

$$U = \begin{pmatrix} c_{13}c_{12} & s_{12}c_{13} & s_{13}e^{i\delta} \\ -s_{12}c_{23} - s_{23}s_{13}c_{12}e^{-i\delta} & c_{12}c_{23} - s_{23}s_{13}s_{12}e^{-i\delta} & s_{23}c_{13} \\ s_{23}s_{12} - c_{23}s_{13}c_{12}e^{-i\delta} & -s_{23}c_{12} - c_{23}s_{13}s_{12}e^{-i\delta} & c_{23}c_{13} \end{pmatrix} \begin{pmatrix} e^{i\alpha} \\ e^{i\beta} \\ 1 \end{pmatrix}$$




$$\sim g \gamma^\mu P_L N_{\alpha i}$$

$$N = \begin{pmatrix} N_{e1} & N_{e2} & N_{e3} \\ N_{\mu1} & N_{\mu2} & N_{\mu3} \\ N_{\tau1} & N_{\tau2} & N_{\tau3} \end{pmatrix}$$

Effective Lagrangian allowing non-unitarity


- 3 light ν
- Assume NP at $\Lambda \gg \Lambda_{EW}$. In the flavor basis,

$$L = \frac{1}{2} (i\bar{\nu}_\alpha \partial K_{\alpha\beta} \nu_\beta - \bar{\nu}_\alpha^c M_{\alpha\beta} \nu_\beta + h.c.) - \frac{g}{\sqrt{2}} (W_\mu^+ \bar{l}_\alpha \gamma^\mu P_L \nu_\alpha + h.c.) - \frac{g}{\cos \theta_W} (Z_\mu \bar{\nu}_\alpha \gamma^\mu P_L \nu_\alpha + h.c.) + \dots$$



$$\nu_\alpha = N_{\alpha i} \nu_i$$

$$L = \underbrace{\frac{1}{2} (i\bar{\nu}_i \partial \nu_i - \bar{\nu}_i m_{ii} \nu_i)}_{\text{unchanged}} - \frac{g}{\sqrt{2}} (W_\mu^+ \bar{l}_\alpha \gamma^\mu P_L \underbrace{N_{\alpha i} \nu_i}_{\substack{\downarrow \\ N \text{ non-unitary}}}) - \frac{g}{\cos \theta_W} (Z_\mu \bar{\nu}_i \gamma^\mu P_L (N^\dagger N)_{ij} \nu_j) + h.c. + \dots$$



$$\langle \nu_i | \nu_j \rangle = \delta_{ij}$$

The effects of non-unitarity appear in the interactions



We can bound unitarity with weak decays

$$|\nu_\alpha\rangle = \frac{1}{\sqrt{\sum_i |N_{\alpha i}|^2}} \sum_i N_{\alpha i}^* |\nu_i\rangle \quad \longrightarrow \quad \langle \nu_\beta | \nu_\alpha \rangle \neq \delta_{\alpha\beta}$$

$$P_{\alpha\beta} = \left| \langle \nu_\beta | \nu_\alpha(L) \rangle \right|^2 = \frac{\left| \sum_i N_{\beta i} N_{\alpha i}^* e^{-i \frac{m_i^2}{2E} L} \right|^2}{\sum_i \left| \sum_{\alpha} N_{\alpha i} \right|^2 \sum_i \left| \sum_{\beta} N_{\beta i} \right|^2}$$

Zero Distance Effect

Constraints on Unitarity from weak decays

$W \rightarrow l_\alpha \nu$ universality tests

$Z \rightarrow$ invisible rare lepton decays

$$|NN^\dagger| \approx \begin{pmatrix} 0.994 \pm 0.005 & < 7.1 \cdot 10^{-5} & < 1.6 \cdot 10^{-2} \\ < 7.1 \cdot 10^{-5} & 0.995 \pm 0.005 & < 1.0 \cdot 10^{-2} \\ < 1.6 \cdot 10^{-2} & < 1.0 \cdot 10^{-2} & 0.995 \pm 0.005 \end{pmatrix} \quad 90\% \text{ C.L.}$$

↳ N is unitary at % level

In the future...

TESTS OF UNITARITY

RARE LEPTON DECAYS

- $\mu \rightarrow e \gamma$ $(NN^\dagger)_{e\mu} < 10^{-6}$ (MEG)
- $\tau \rightarrow e \gamma$ $(NN^\dagger)_{e\tau} < 1.6 \cdot 10^{-2}$ (BABAR)
- $\tau \rightarrow \mu \gamma$ $(NN^\dagger)_{\mu\tau} < 1.0 \cdot 10^{-2}$ (Belle)

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ZERO-DISTANCE EFFECT @ NF

- $\nu_e \rightarrow \nu_\mu$ $(NN^\dagger)_{e\mu} < 2.3 \cdot 10^{-4}$
- $\nu_e \rightarrow \nu_\tau$ $(NN^\dagger)_{e\tau} < 2.9 \cdot 10^{-3}$
- $\nu_\mu \rightarrow \nu_\tau$ $(NN^\dagger)_{\mu\tau} < 2.6 \cdot 10^{-3}$

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PHASES

- **phases** \rightarrow *appearance* experiments: NUFACTs, *b*-beams

Can we measure the phases of N ?

E. Fdez-Martinez, M.B. Gavela, J. López-Pavón and O. Yasuda [hep-ph/0703098](#)

If we parametrize $N = (1 + \eta) \cdot U$ where $U \approx U_{PMNS}$ and $\eta = \eta^\dagger$

$$\eta = \begin{pmatrix} \eta_{ee} & \eta_{e\mu} & \eta_{e\tau} \\ \eta_{e\mu}^* & \eta_{\mu\mu} & \eta_{\mu\tau} \\ \eta_{e\tau}^* & \eta_{\mu\tau}^* & \eta_{\tau\tau} \end{pmatrix}$$

If $\frac{\Delta m^2 L}{2E} \equiv \Delta \ll 1$ \Rightarrow $\langle \nu_\alpha | \nu_\beta(L) \rangle = A_{\alpha\beta}^{\text{SM}} + 2\eta_{\alpha\beta}^* + \mathcal{O}(\eta^2)$

$P_{\alpha\beta}$ only depends on $\eta_{\alpha\beta}$

Can we measure the phases of N ?

E. Fdez-Martinez, M.B. Gavela, J. López-Pavón and O. Yasuda [hep-ph/0703098](#)

For instance, for two families:

$$P_{\alpha\beta} = \underbrace{\sin^2(2\theta) \sin^2\left(\frac{\Delta}{2}\right)}_{\text{SM}} - \underbrace{4|\eta_{\alpha\beta}| \sin(\delta_{\alpha\beta}) \sin(2\theta) \sin\Delta}_{\text{CP violating interference}} + \underbrace{4|\eta_{\alpha\beta}|^2}_{\text{Zero distance effect}}$$

SM

CP violating
interference

Zero distance
effect

New CP-violation signals

$$\hookrightarrow \eta_{\alpha\beta} \equiv |\eta_{\alpha\beta}| e^{-\delta_{\alpha\beta}}$$

$$\hookrightarrow \Delta_{ij} = \Delta m_{ij}^2 L / 2E \ll 1$$

Which is the best channel?

$$\left\{ \begin{array}{l} \nu_e \leftrightarrow \nu_\mu ? \end{array} \right.$$

$$|\eta_{e\mu}| < 3.5 \cdot 10^{-5}$$

$$\nu_e \leftrightarrow \nu_\tau ?$$

$$|\eta_{e\tau}| < 8.0 \cdot 10^{-3}$$

$$\nu_\mu \leftrightarrow \nu_\tau ?$$

$$|\eta_{\mu\tau}| < 5.1 \cdot 10^{-3}$$

$$P_{\mu\tau} \approx \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta_{31}}{2} \right) - 4 |\eta_{\mu\tau}| \sin \delta_{\mu\tau} \sin 2\theta_{23} \sin(\Delta_{31}) + 4 |\eta_{\mu\tau}|^2$$

This is the best channel to measure new CP phases:

$$P_{\mu\tau} - P_{\bar{\mu}\bar{\tau}} \approx 8 |\eta_{\mu\tau}| \sin \delta_{\mu\tau} \sin(2\theta_{23}) \sin(\Delta_{13})$$

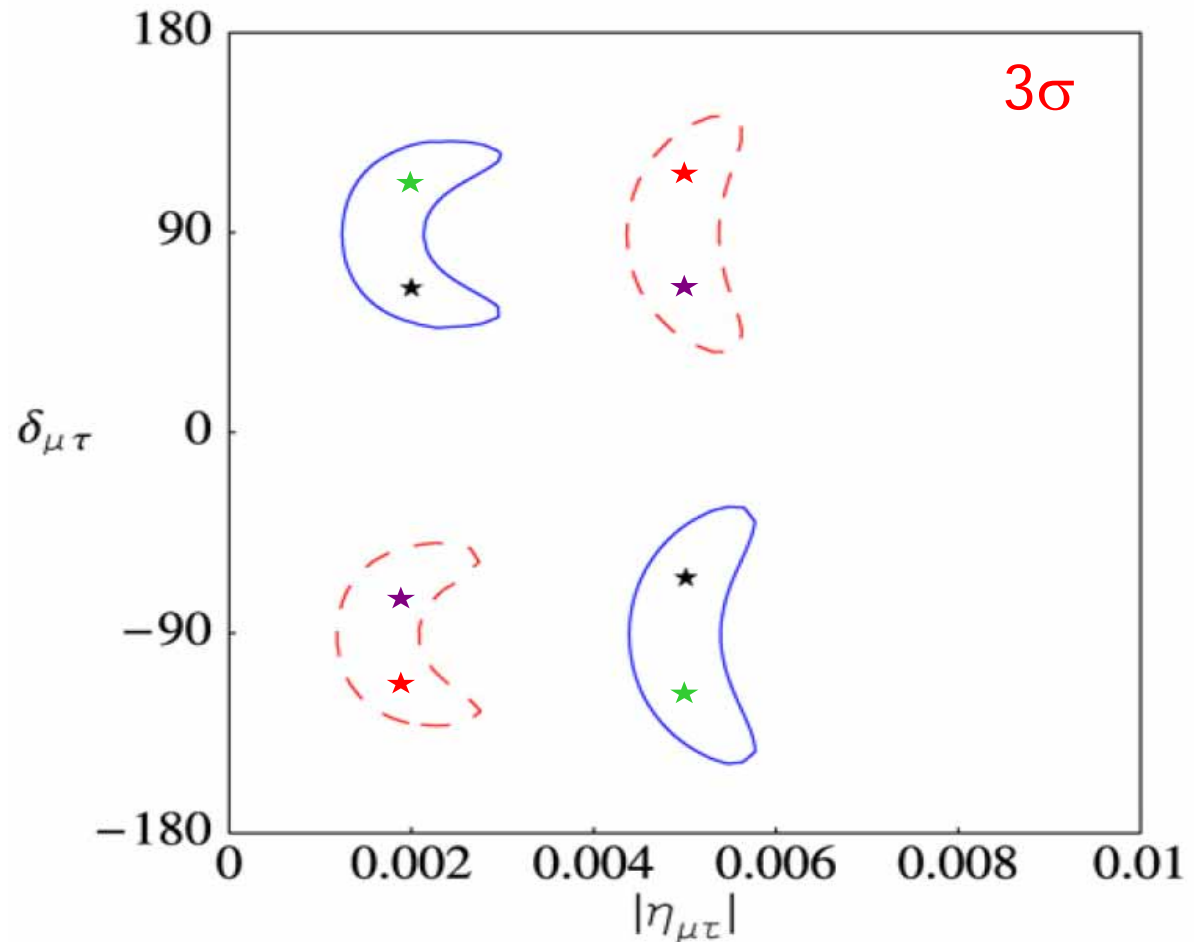
$\nu_\mu \leftrightarrow \nu_\tau$ The CP phase $\delta_{\mu\tau}$ can be probed

$$N_{\mu\tau} - N_{\bar{\mu}\bar{\tau}} \propto 8 |\eta_{\mu\tau}| \sin \delta_{\mu\tau} \sin(2\theta_{23}) \sin(\Delta_{13})$$

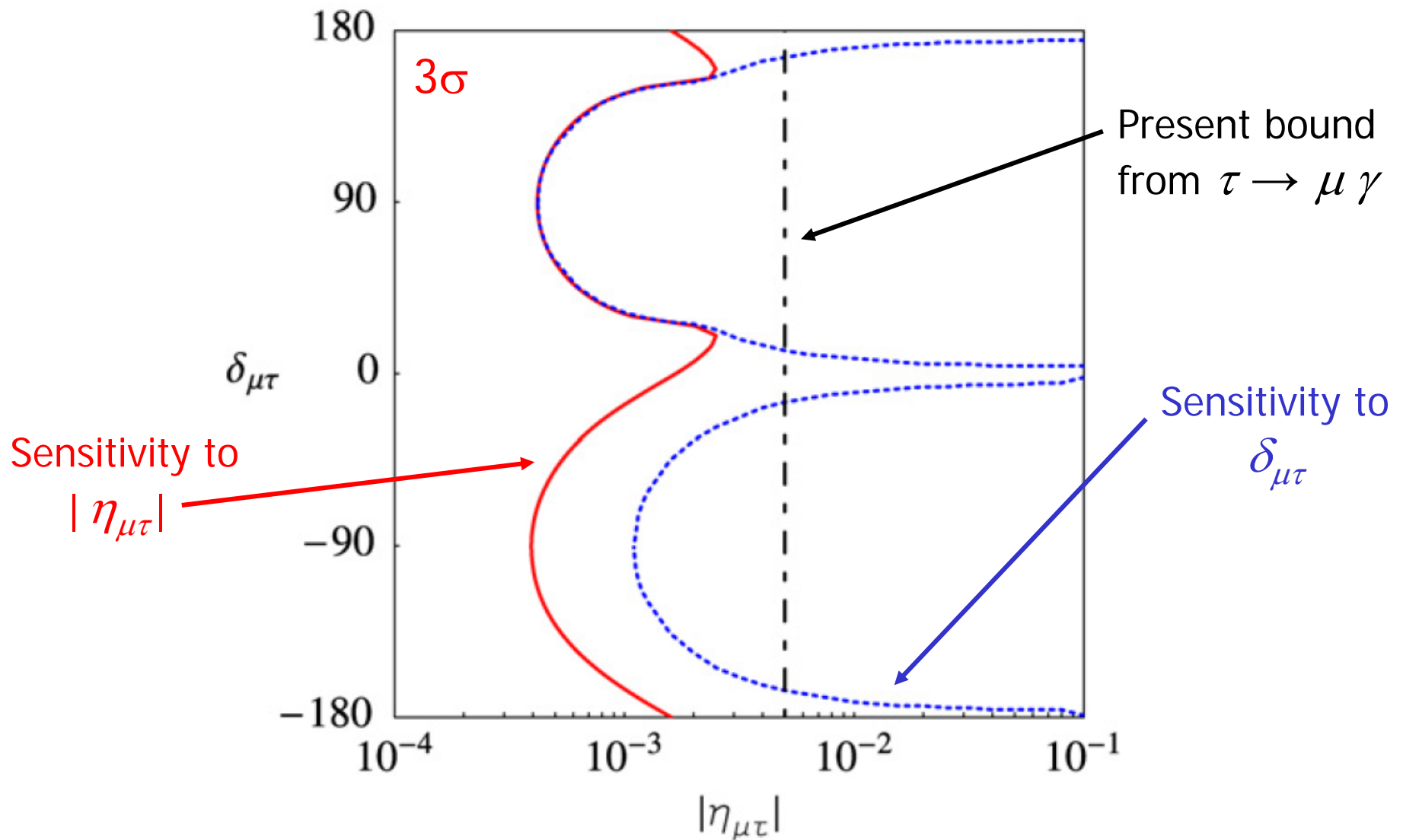
1. Confusion between
 $sign(\delta_{\mu\tau}) \leftrightarrow sign(\Delta_{13})$

2. Degeneracy
 $180^\circ - \delta_{\mu\tau} \leftrightarrow \delta_{\mu\tau}$

@ a Neutrino Factory
with $L = 130$ Km



$\nu_\mu \leftrightarrow \nu_\tau$ The CP phase $\delta_{\mu\tau}$ can be probed



For non-trivial $\delta_{\mu\tau}$, one order of magnitude improvement for $|\eta_{\mu\tau}|$

Conclusions

- Analyze neutrino data **without assuming unitarity**. We started the first analyses and conclude that, at present:

- EW decays confirms **unitarity at % level**

- Present bounds on unitarity are strong enough to match the unitarity analysis

- **$V_{\mu\tau}$** CP-asymmetry is a clean probe of the new phases.

- Non-unitary effects in typical models are too small to be detected at present experiments. They could be sizable in extensions/others models with $M \sim \text{TeV}, \dots$).

- > **keep tracking them in the future.**

- They are excellent signals of new physics.**

Back-up slides

Constraints on unitarity

$$|\eta| \approx \left(\begin{array}{ccc} |\eta_{ee}| < 5.5 \cdot 10^{-3} & |\eta_{e\mu}| < 3.5 \cdot 10^{-5} & |\eta_{e\tau}| < 8.0 \cdot 10^{-3} \\ |\eta_{e\mu}| < 3.5 \cdot 10^{-5} & |\eta_{\mu\mu}| < 5.0 \cdot 10^{-3} & |\eta_{\mu\tau}| < 5.1 \cdot 10^{-3} \\ |\eta_{e\tau}| < 8.0 \cdot 10^{-3} & |\eta_{\mu\tau}| < 5.1 \cdot 10^{-3} & |\eta_{\tau\tau}| < 5.0 \cdot 10^{-3} \end{array} \right)$$

Leptonic Mixing Matrix Elements From Oscillations

$$N = \begin{pmatrix} N_{e1} & N_{e2} & N_{e3} \\ N_{\mu1} & N_{\mu2} & N_{\mu3} \\ N_{\tau1} & N_{\tau2} & N_{\tau3} \end{pmatrix}$$

N elements from oscillations: e -row

CHOOZ: $P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \cong \left(|N_{e1}|^2 + |N_{e2}|^2 \right)^2 + |N_{e3}|^4 + 2 \left(|N_{e1}|^2 + |N_{e2}|^2 \right) |N_{e3}|^2 \cos(\Delta_{13})$

KamLAND: $P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \cong |N_{e1}|^4 + |N_{e2}|^4 + |N_{e3}|^4 + 2 |N_{e1}|^2 |N_{e2}|^2 \cos(\Delta_{12})$

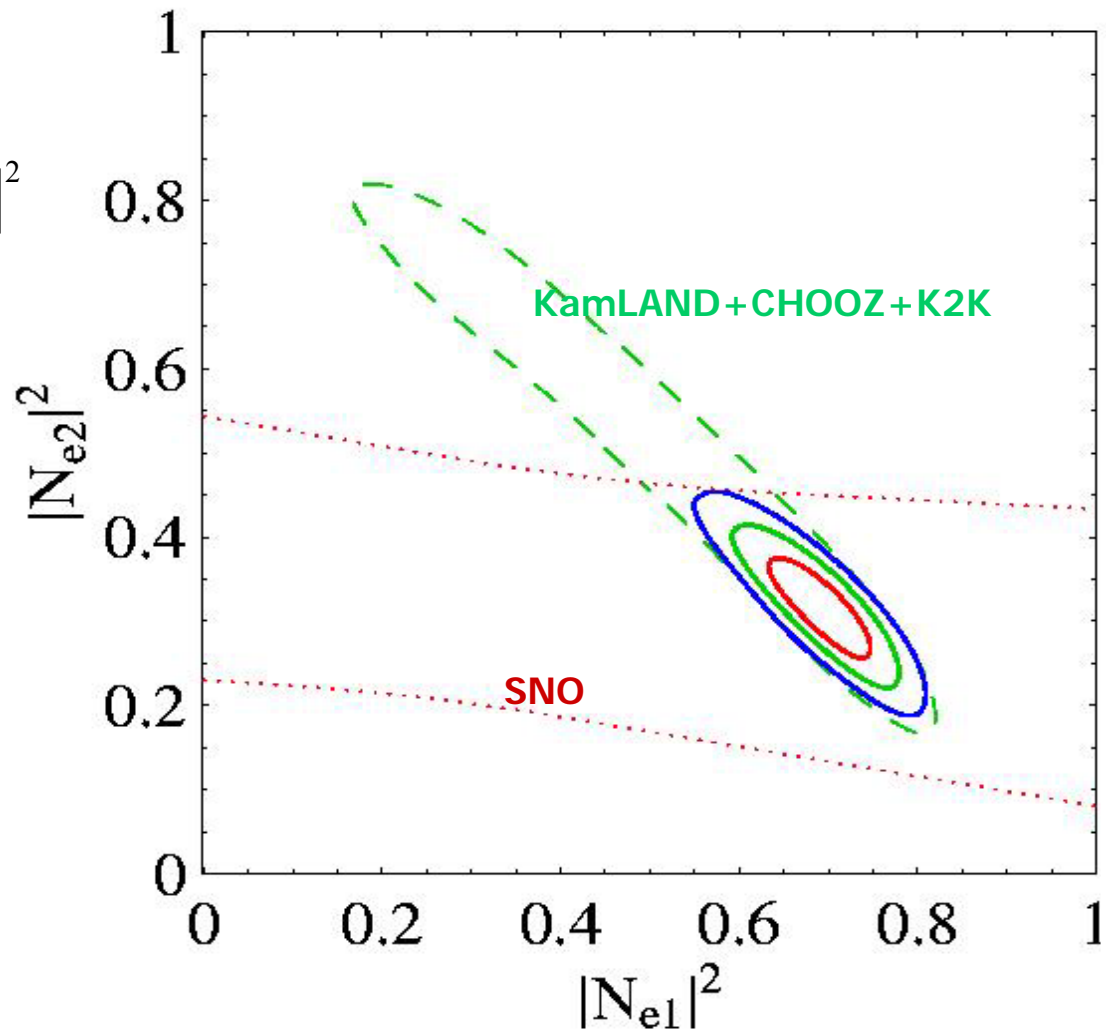
Δ_{13} from K2K

SNO:

$$P(\nu_e \rightarrow \nu_e) \cong 0.1 |N_{e1}|^2 + 0.9 |N_{e2}|^2$$

$$\Delta_{ij} = \Delta m_{ij}^2 L / 2E$$

→ all $|N_{ei}|^2$ determined



N elements from oscillations: μ -row

Atmospheric + K2K: $\Delta_{12} \approx 0$

$$\Delta_{ij} = \Delta m_{ij}^2 L / 2E$$

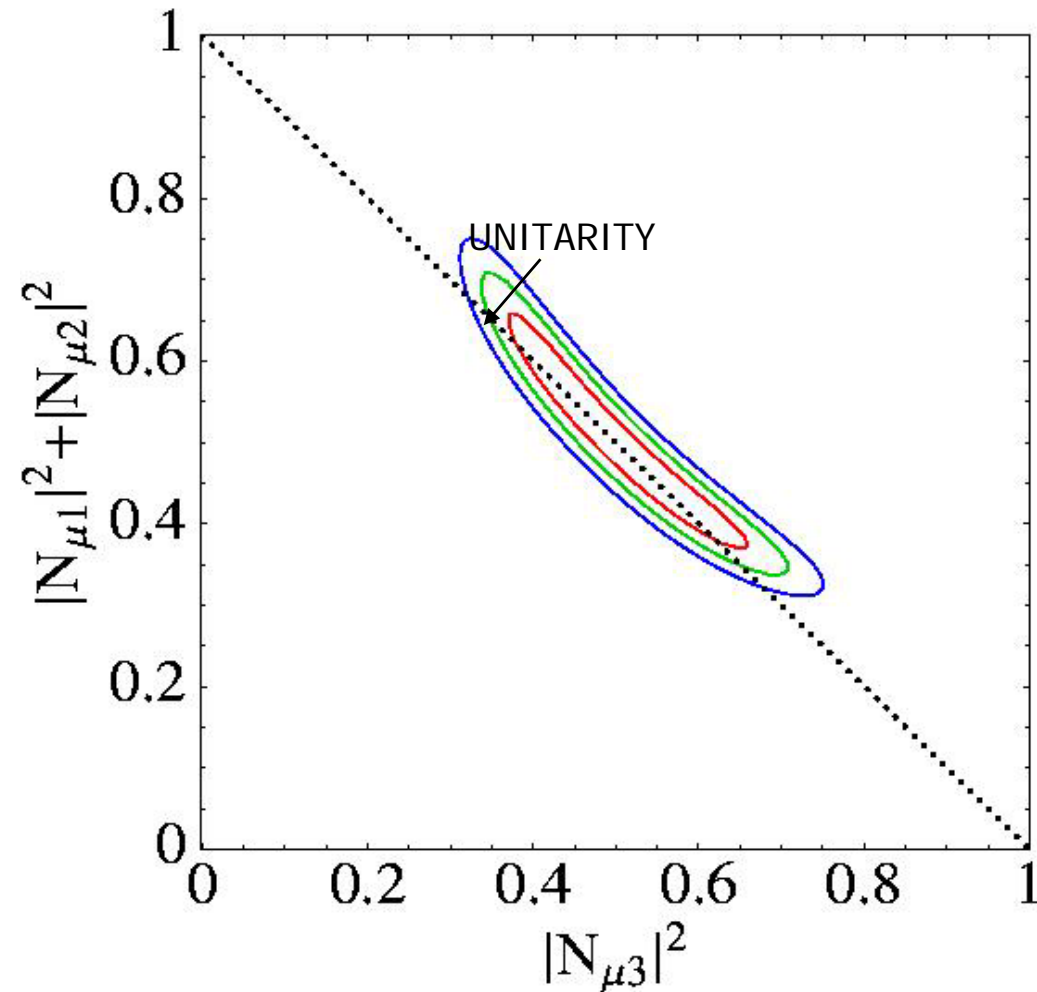
$$\hat{P}(v_\mu \rightarrow v_\mu) \cong \left(|N_{\mu 1}|^2 + |N_{\mu 2}|^2 \right)^2 + |N_{\mu 3}|^4 + 2 \left(|N_{\mu 1}|^2 + |N_{\mu 2}|^2 \right) |N_{\mu 3}|^2 \cos(\Delta_{23})$$

1. Degeneracy

$$|N_{\mu 1}|^2 + |N_{\mu 2}|^2 \leftrightarrow |N_{\mu 3}|^2$$

2. $|N_{\mu 1}|^2$, $|N_{\mu 2}|^2$

cannot be disentangled



N elements from oscillations only

without unitarity
OSCILLATIONS

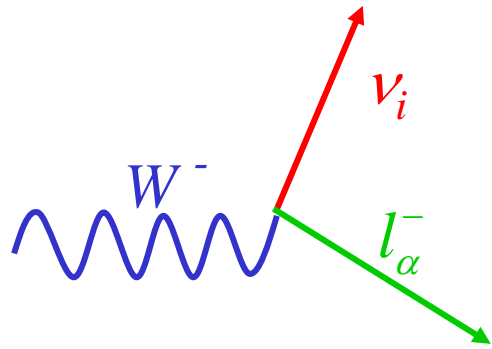
$$|N| = \begin{pmatrix} 0.75 - 0.89 & 0.45 - 0.66 & < 0.34 \\ \left[\left(|N_{\mu 1}|^2 + |N_{\mu 2}|^2 \right)^{1/2} = 0.57 - 0.86 \right] & 0.57 - 0.86 & \\ ? & ? & ? \end{pmatrix}$$

3σ

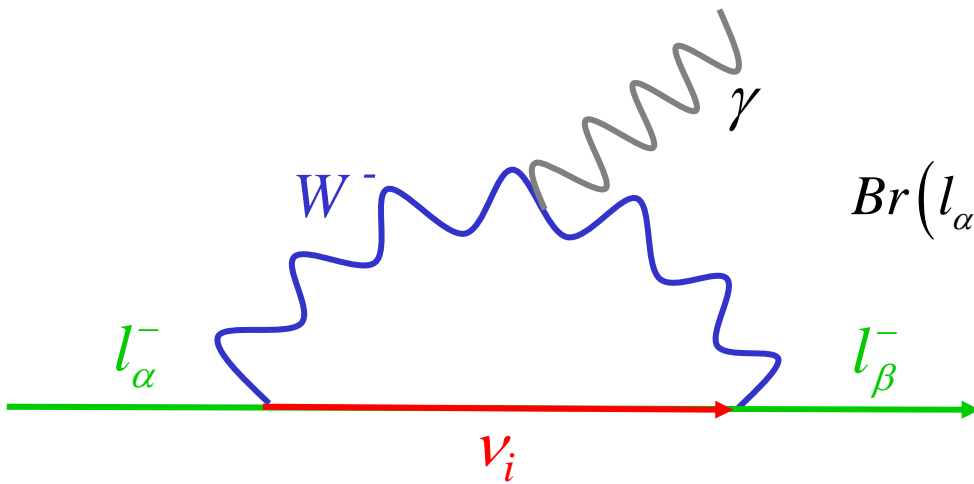
with unitarity
OSCILLATIONS

$$|U| = \begin{pmatrix} 0.79 - 0.86 & 0.50 - 0.61 & < 0.20 \\ 0.25 - 0.53 & 0.47 - 0.73 & 0.56 - 0.79 \\ 0.21 - 0.51 & 0.42 - 0.69 & 0.61 - 0.83 \end{pmatrix}$$

Decays



$$\Gamma(W \rightarrow l_\alpha \nu) \propto \frac{(NN^\dagger)_{\alpha\alpha}}{\sqrt{(NN^\dagger)_{ee} (NN^\dagger)_{\mu\mu}}}$$



$$Br(l_\alpha \rightarrow l_\beta \gamma) \propto \frac{|(NN^\dagger)_{\beta\alpha}|^2}{(NN^\dagger)_{\alpha\alpha} (NN^\dagger)_{\beta\beta}} + \mathcal{O}\left(\frac{m_\nu^2}{M_W^2}\right)$$

- $Z \rightarrow \text{invisible} \propto \frac{\sum_{i,j} (N^\dagger N)_{ij}}{\sqrt{(NN^\dagger)_{ee} (NN^\dagger)_{\mu\mu}}}$

- universality tests $\propto \frac{(NN^\dagger)_{\alpha\alpha}}{(NN^\dagger)_{\beta\beta}}$

Can we measure the phases of N ?

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If we parametrize $N = (1 + \eta) \cdot U$ where $U \approx U_{PMNS}$ and $\eta = \eta^\dagger$

$$\eta = \begin{pmatrix} \eta_{ee} & \eta_{e\mu} & \eta_{e\tau} \\ \eta_{e\mu}^* & \eta_{\mu\mu} & \eta_{\mu\tau} \\ \eta_{e\tau}^* & \eta_{\mu\tau}^* & \eta_{\tau\tau} \end{pmatrix}$$

Normalization factors

$$\langle \nu_\beta | \nu_\alpha(L) \rangle = A_{\alpha\beta}^{\text{SM}}(L) (1 - \eta_{\alpha\alpha} - \eta_{\beta\beta}) + \sum_\gamma (\eta_{\alpha\gamma}^* A_{\gamma\beta}^{\text{SM}}(L) + \eta_{\beta\gamma} A_{\alpha\gamma}^{\text{SM}}(L)) + O(\eta^2)$$

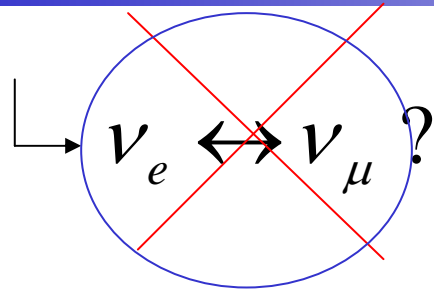
Standard amplitude

If $\frac{\Delta m^2 L}{2E} \equiv \Delta \ll 1 \Rightarrow \begin{cases} A_{\alpha\alpha}^{\text{SM}} \approx 1 \\ A_{\alpha\beta}^{\text{SM}} \sim |\eta_{\alpha\beta}| \ll 1 \end{cases}$

$$\langle \nu_\alpha | \nu_\beta(L) \rangle = A_{\alpha\beta}^{\text{SM}} + 2\eta_{\alpha\beta}^* + O(\eta^2)$$

$P_{\alpha\beta}$ only depends on $\eta_{\alpha\beta}$

Which is the best channel?

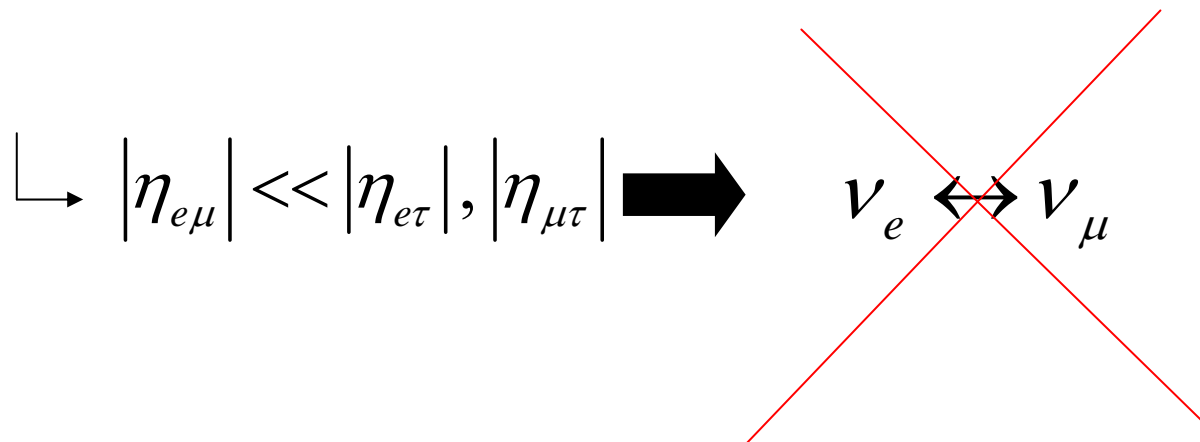


$$\nu_e \leftrightarrow \nu_\tau ?$$

$$\nu_\mu \leftrightarrow \nu_\tau ?$$

Constraints on unitarity:

$$|\eta| \approx \begin{pmatrix} |\eta_{ee}| < 5.5 \cdot 10^{-3} & |\eta_{e\mu}| < 3.5 \cdot 10^{-5} & |\eta_{e\tau}| < 8.0 \cdot 10^{-3} \\ |\eta_{e\mu}| < 3.5 \cdot 10^{-5} & |\eta_{\mu\mu}| < 5.0 \cdot 10^{-3} & |\eta_{\mu\tau}| < 5.1 \cdot 10^{-3} \\ |\eta_{e\tau}| < 8.0 \cdot 10^{-3} & |\eta_{\mu\tau}| < 5.1 \cdot 10^{-3} & |\eta_{\tau\tau}| < 5.0 \cdot 10^{-3} \end{pmatrix} \quad 90\% \text{ C.L.}$$



Which is the best channel?

$$\left\langle \begin{array}{l} \nu_e \\ \nu_\mu \end{array} \right\rangle \leftrightarrow ?$$

$$\left\langle \begin{array}{l} \nu_e \\ \nu_\tau \end{array} \right\rangle \leftrightarrow ?$$

$$\left\langle \begin{array}{l} \nu_\mu \\ \nu_\tau \end{array} \right\rangle \leftrightarrow ?$$

$$P_{e\tau} \approx P_{e\tau}^{SM} +$$

$$-2 |\eta_{e\tau}| c_{23} \sin(\delta + \delta_{e\tau}) \sin 2\theta_{13} (\Delta_{31}) + 2 |\eta_{e\tau}| s_{23} c_{13} \sin 2\theta_{12} \sin \delta_{e\tau} (\Delta_{21})$$

$$+ 4 |\eta_{e\tau}|^2$$

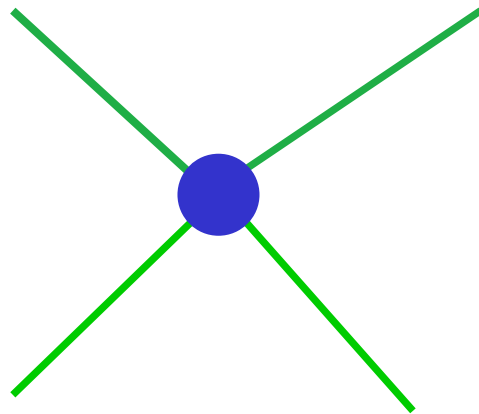
$$\Delta_{ij} = \Delta m_{ij}^2 L / 2E \ll 1$$

$\nu_e \leftrightarrow \nu_\tau$ is suppressed by the small parameters Δ_{21} and $\sin 2\theta_{13} \Delta_{31}$

Our analysis will also apply to “non-standard” or “exotic” neutrino interactions.

(Grossman, Gonzalez-Garcia et al., Huber et al., Kitazawa et al., Davidson et al. Blennow et al...)

They add 4-fermion exotic operators which affect production
or detection
or propagation in matter

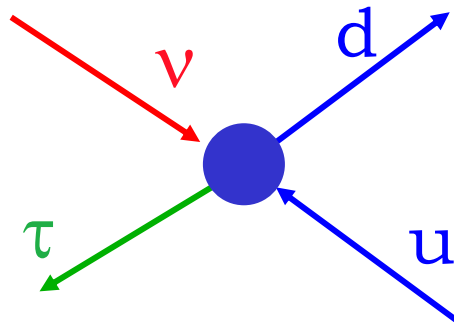


$$\bar{\Psi} \Psi \bar{\Psi} \Psi$$

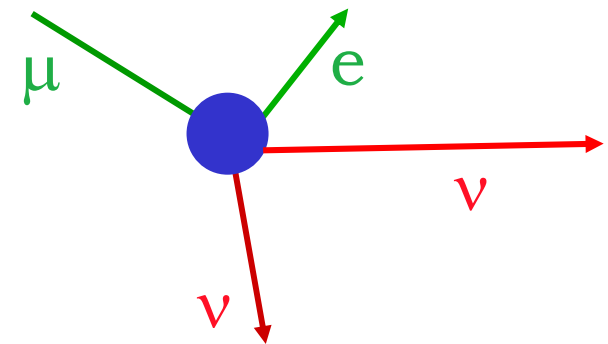
"NSI" neutrino interactions vs NonUnitarity

- SM

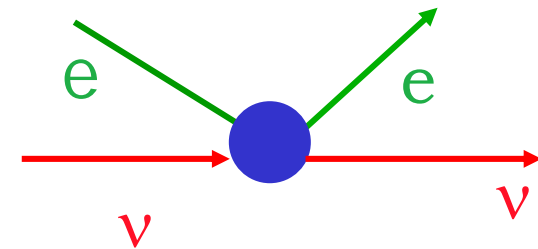
$$-\mathcal{L}^{\text{CC}} = \frac{4G_F}{\sqrt{2}} \left[\underbrace{\sum_{\alpha\beta} (\bar{\nu}_\alpha^{\text{SM}} \gamma_\mu P_L l_\alpha) (\bar{d}_\beta \gamma_\mu P_L u_\beta)}_{\text{Detection @ NF: } \bar{\nu}_\alpha d \rightarrow l_\alpha^- u} + \underbrace{(\bar{\nu}_\alpha^{\text{SM}} \gamma_\mu P_L l_\alpha) (\bar{l}_\beta \gamma_\mu P_L \nu_\beta^{\text{SM}})}_{\text{Production @ NF: } \mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu} + \text{h.c.} \right]$$



Production @ conventional-beams
super-beams



Matter effects: $\alpha = \beta = e$

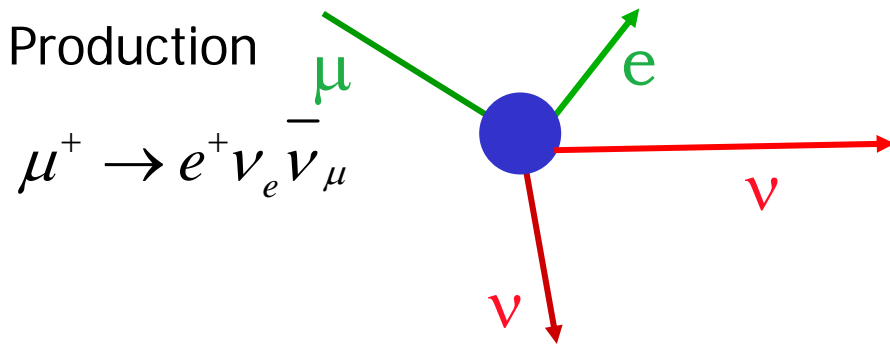


"NSI" neutrino interactions vs NonUnitarity

- Exotic Interactions

Add new effective 4-fermion operators which affects to:

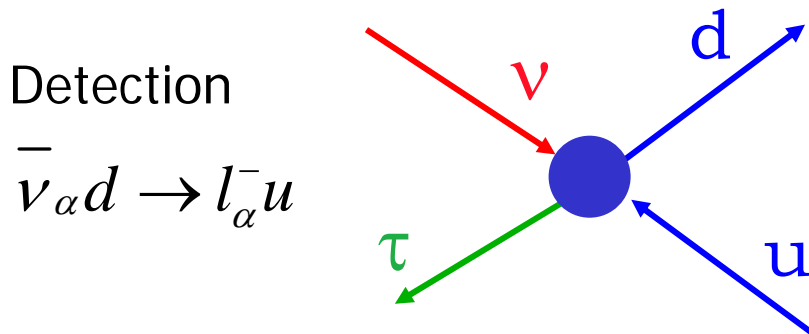
Production



$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$$

$$\frac{-4}{\sqrt{2}} \sum_{\alpha \neq e} G_{e\alpha}^P (\bar{\mu} \gamma_\mu P_L \nu_\mu^{\text{SM}}) (\bar{\nu}_\alpha^{\text{SM}} \gamma_\mu P_L e) + \text{h.c.}$$

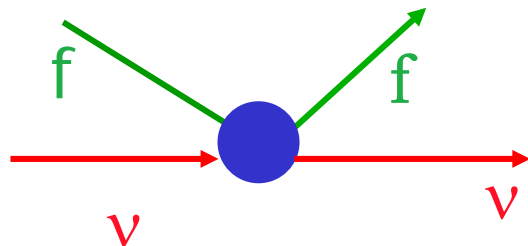
Detection



$$\bar{\nu}_\alpha d \rightarrow l_\alpha^- u$$

$$\frac{-4}{\sqrt{2}} \sum_{\beta \neq \mu} G_{\mu\beta}^D (\bar{\nu}_\beta^{\text{SM}} \gamma_\mu P_L \mu) (\bar{d} \gamma_\mu P_L u) + \text{h.c.}$$

Matter effects



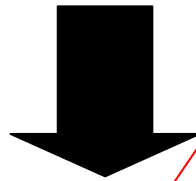
$$\sim (v_\alpha^{\text{SM}} \gamma_\mu P_L v_\beta^{\text{SM}}) (\bar{f} \gamma_\mu f)$$

"NSI" neutrino interactions vs NonUnitarity

- SM + Exotic Interactions

$$-\mathcal{L}^{\text{int}} = \frac{4G_F}{\sqrt{2}} \sum_{\alpha} \left[\left(\delta_{\alpha\mu} + \frac{G_{\mu\alpha}^D}{G_F} \right) (\bar{\nu}_{\alpha}^{\text{SM}} \gamma_{\mu} P_L \mu) (\bar{d} \gamma_{\mu} P_L u) + \text{h.c.} \right. \\ \left. + \left(\delta_{\alpha e} + \frac{G_{e\alpha}^P}{G_F} \right) (\bar{\nu}_{\alpha}^{\text{SM}} \gamma_{\mu} P_L e) (\bar{\mu} \gamma_{\mu} P_L \nu_{\mu}^{\text{SM}}) + \text{h.c.} \right]$$

$$\nu_{\alpha}^{\text{SM}} \equiv U_{\alpha i} \nu_i$$



$$\left. \begin{aligned} | \nu_e^P \rangle &= (1 + \epsilon^{*P})_{e\beta} U_{\beta i}^* | \nu_i \rangle, \\ | \nu_{\mu}^d \rangle &= (1 + \epsilon^{*d})_{\mu\beta} U_{\beta i}^* | \nu_i \rangle \end{aligned} \right\}$$

$$| \nu_{\alpha} \rangle = (1 + \eta^*)_{\alpha\beta} U_{\beta i}^* | \nu_i \rangle$$

Non-Unitarity

Exotic Interactions

"NSI" neutrino interactions vs NonUnitarity

- Exotic Interactions

$$\left. \begin{aligned} |v_e^p\rangle &= (1 + \epsilon^{*p})_{e\beta} U_{\beta i}^* |v_i\rangle, \\ |v_\mu^d\rangle &= (1 + \epsilon^{*d})_{\mu\beta} U_{\beta i}^* |v_i\rangle \end{aligned} \right\}$$

No relationship between $\epsilon_{\alpha\beta}^p$ and $\epsilon_{\beta\alpha}^d$

- Non-Unitarity

$$|v_\alpha\rangle = (1 + \eta^*)_{\alpha\beta} U_{\beta i}^* |v_i\rangle$$

$$\eta_{\alpha\beta} = \eta_{\beta\alpha}^*$$

If $\epsilon_{\alpha\beta}^p = \epsilon_{\beta\alpha}^{d*}$ \longrightarrow $P_{\mu\tau}^{EXOTIC} = P_{\mu\tau}^{NON-UNITARITY}$

If not \longrightarrow Idem , barring extreme fine tuned cancelations

Our constraints apply to Exotic Interactions !

Oscillation Probability in Matter

$$-\mathbf{L}_{eff}^{int} = \underbrace{\sqrt{2}G_F n_e \bar{\nu}_e \gamma^0 P_L \nu_e}_{V_{CC}} - \underbrace{\frac{1}{\sqrt{2}}G_F n_n \sum_{\alpha} \bar{\nu}_{\alpha} \gamma^0 P_L \nu_{\alpha}}_{V_{NC}}$$

$$i \frac{d}{dt} \begin{pmatrix} |\nu_e\rangle \\ |\nu_{\mu}\rangle \end{pmatrix} = \left[\tilde{N}^* \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix} (\tilde{N}^*)^{-1} + \begin{pmatrix} (V_{CC} - V_{NC})(NN^{\dagger})_{ee} & -V_{NC} \sqrt{\frac{(NN^{\dagger})_{\mu\mu}}{(NN^{\dagger})_{ee}}} (NN^{\dagger})_{\mu e} \\ (V_{CC} - V_{NC}) \sqrt{\frac{(NN^{\dagger})_{ee}}{(NN^{\dagger})_{\mu\mu}}} (NN^{\dagger})_{e\mu} & -V_{NC} (NN^{\dagger})_{\mu\mu} \end{pmatrix} \right] \begin{pmatrix} |\nu_e\rangle \\ |\nu_{\mu}\rangle \end{pmatrix}$$

Effective Evolution

1. non-diagonal elements

2. NC effects do not disappear

Matter effects in $\nu_\mu \leftrightarrow \nu_\tau$

Keeping only terms to second order in Δ_{21} , Δ_{31} , AL and $\sin^2 2\theta_{13}$ and first in $|\eta_{\alpha\beta}|$ (setting $\eta_{e\mu} = 0$):

$$\begin{aligned}
 P_{\mu\tau} = & \sin^2 2\theta_{23}(\Delta_{31}L/2)^2 - 2|\eta_{\mu\tau}| \sin \delta_{\mu\tau} \sin 2\theta_{23}(\Delta_{31}L) + 4|\eta_{\mu\tau}|^2 \\
 & - (1/2)c_{12}^2 \sin^2 2\theta_{23}(\Delta_{31}L)(\Delta_{21}L) - |\eta_{\mu\tau}| \sin 2\theta_{23} \cos \delta_{\mu\tau} (AL)(\Delta_{31}L) \\
 & + (1/4)c_{12}^4 \sin^2 2\theta_{23}(\Delta_{21}L)^2 + 2|\eta_{\mu\tau}|c_{12}^2 \sin \delta_{\mu\tau} \sin 2\theta_{23}(\Delta_{21}L) \\
 & - (1/4)s_{13} \sin 4\theta_{23} \sin 2\theta_{12} \cos \delta (\Delta_{31}L)(\Delta_{21}L).
 \end{aligned}$$

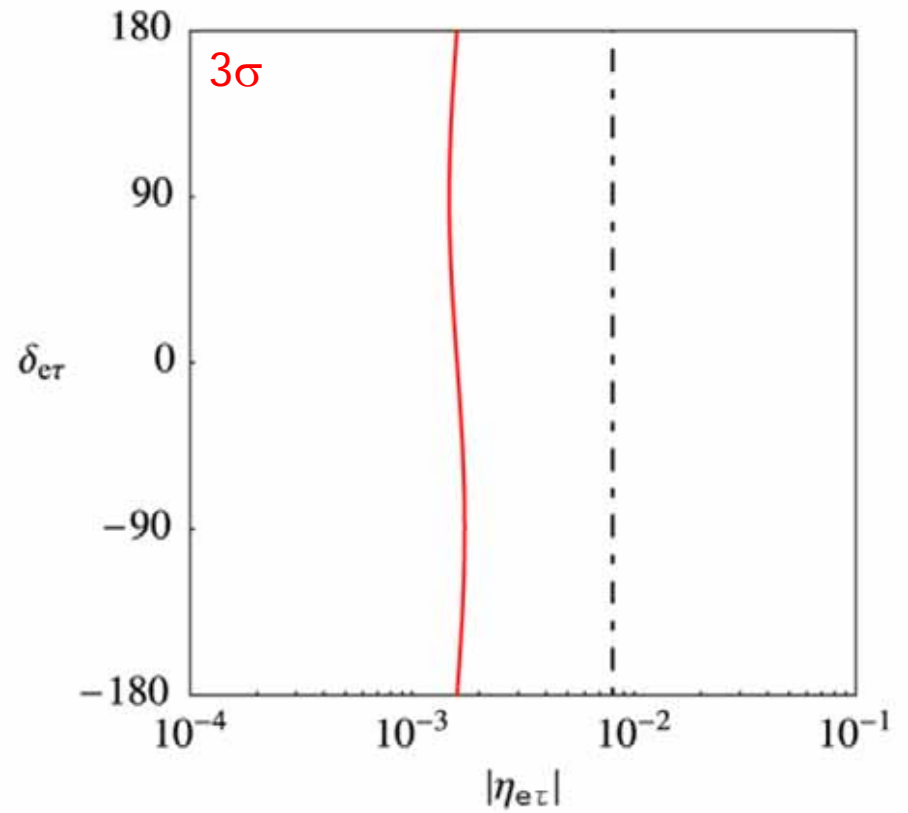
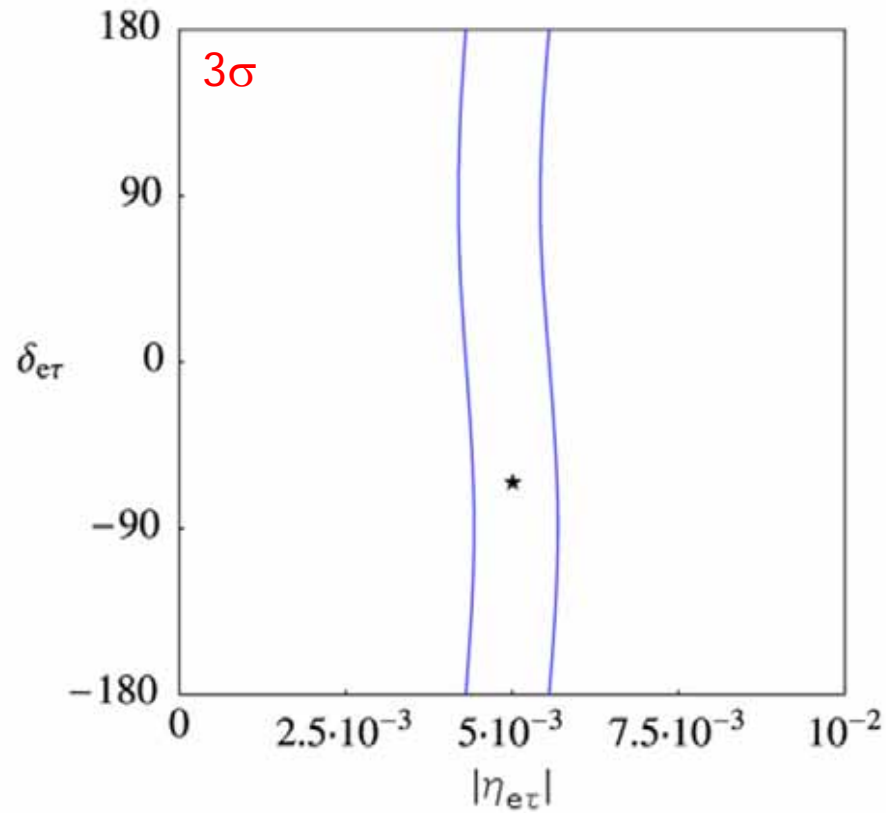
$$\sin^2 2\theta_{13} \lesssim 10^{-1} \quad \Delta_{31} \square AL \square 10^{-2} \quad \Delta_{21} \square 10^{-3.5} \quad |\eta_{\alpha\beta}| \lesssim 10^{-2}$$

Aproximate expression for $P_{e\tau}$

$$\begin{aligned} P_{e\tau} \approx & c_{23}^2 \sin^2 2\theta_{13} \left(\frac{\Delta_{31}}{2} \right)^2 + s_{23}^2 \sin^2 2\theta_{12} \left(\frac{\Delta_{21}}{2} \right)^2 - \\ & c_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \delta \left(\frac{\Delta_{21}}{2} \right) \left(\frac{\Delta_{31}}{2} \right) \\ & - 2 |\eta_{e\tau}| c_{23} \sin 2\theta_{13} \sin(\delta + \delta_{e\tau}) (\Delta_{31}) + 2 |\eta_{e\tau}| s_{23} c_{13} \sin 2\theta_{12} \sin \delta_{e\tau} (\Delta_{21}) \\ & + 4 |\eta_{e\tau}|^2 \end{aligned}$$

$$P_{e\mu} = P_{e\tau} \left(s_{23} \rightarrow -c_{23}; c_{23} \rightarrow s_{23}; \eta_{e\tau} \rightarrow \eta_{e\mu} \right)$$

Analysis of $\nu_e \leftrightarrow \nu_\tau$



The zero distance effect dominates over CP violating interference term

- No information on $\delta_{e\tau}$

- Sensitivities to $|\eta_{e\tau}|$ around 10^{-3}

Some details about the experiment $\nu_\mu \leftrightarrow \nu_\tau$ @ NF

We study a NF beam resulting from the decay of 50 GeV muons:

- Assuming 2×10^{20} useful decays/year
- 5 years running with each polarity
- 5 Kt Opera-like detector
- $L=130$ km
- $\nu_\mu \leftrightarrow \nu_\tau$ sensitivities and backgrounds = $5 \times (\nu_e \leftrightarrow \nu_\tau$ sensitivities and backgrounds)


hep-ph/0305185

(NN^\dagger) and $(N^\dagger N)$ from decays

$$|NN^\dagger| \approx \begin{pmatrix} 1.002 \pm 0.005 & < 7.2 \cdot 10^{-5} & < 1.6 \cdot 10^{-2} \\ < 7.2 \cdot 10^{-5} & 1.003 \pm 0.005 & < 1.3 \cdot 10^{-2} \\ < 1.6 \cdot 10^{-2} & < 1.3 \cdot 10^{-2} & 1.003 \pm 0.005 \end{pmatrix} \quad 90\% \text{ C.L.}$$

$$N = HV \quad NN^\dagger = H^2 = 1 + \varepsilon \quad \text{with } \varepsilon = \varepsilon^\dagger$$

$$N^\dagger N = 1 + V^\dagger \varepsilon V = 1 + \varepsilon'$$

$$|\varepsilon'_{ij}| \leq \sqrt{\sum_{\alpha\beta} |\varepsilon_{\alpha\beta}|^2} \approx 0.03$$

$$|N^\dagger N| \approx \begin{pmatrix} 1.00 \pm 0.032 & < 0.032 & < 0.032 \\ < 0.032 & 1.00 \pm 0.032 & < 0.032 \\ < 0.032 & < 0.032 & 1.00 \pm 0.032 \end{pmatrix}$$

→ N is unitary at % level

Low-energy theory

After EWSB:

$$L = \frac{1}{2} \left(i \bar{\nu}_\alpha \partial \mathbf{K}_{\alpha\beta} \nu_\beta - \bar{\nu}_\alpha^c \mathbf{M}_{\alpha\beta} \nu_\beta + h.c. \right) + \\ - \frac{g}{\sqrt{2}} \left(W_\mu^+ \bar{l}_\alpha \gamma^\mu P_L \nu_\alpha + h.c. \right)$$

$M_{\alpha\beta} \rightarrow$ diagonalized \rightarrow unitary transformation

$K_{\alpha\beta} \rightarrow$ diagonalized and normalized \rightarrow unitary transf. + **rescaling**

$$L = \frac{1}{2} \left(i \bar{\nu}_i \partial \nu_i - \bar{\nu}_i^c m_{ii} \nu_i \right) - \frac{g}{\sqrt{2}} \left(W_\mu^+ \bar{l}_\alpha \gamma^\mu P_L N_{\alpha i} \nu_i \right).$$

N non-unitary

Some numbers

- CHOOZ: systematic error
hep-ex\0301017
}

 Normalization flux $\sim 2.7\%$
 Energy $\sim 1.1\%$

- SNO: $\frac{n_{CC}}{n_{NC}} \square 0.9|N_{e1}|^2 + 0.1|N_{e2}|^2$
└──────────┬──────────┘
└──────────────────────────▶ we let it vary 2%

- OPERA (NF): background: c-decays / charged current decays $\sim 10^{-6}$
systematic error $\sim 5\%$
efficiency $\sim 20\%$
hep-ph\0305185

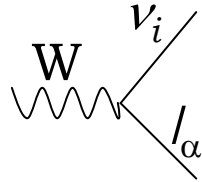
- K2K: Statistics
hep-ex\0411038

ZERO-DISTANCE EFFECT
 40Kt Iron calorimeter near NUFACT

- $\nu_e \rightarrow \nu_\mu$ $(NN^\dagger)_{e\mu} < 2.3 \cdot 10^{-4}$
4Kt OPERA-like near NUFACT
- $\nu_e \rightarrow \nu_\tau$ $(NN^\dagger)_{e\tau} < 2.9 \cdot 10^{-3}$
- $\nu_\mu \rightarrow \nu_\tau$ $(NN^\dagger)_{\mu\tau} < 2.6 \cdot 10^{-3}$

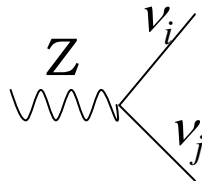
(NN^\dagger) from decays: G_F

- W decays



$$\rightarrow \frac{(NN^\dagger)_{\alpha\alpha}}{\sqrt{(NN^\dagger)_{ee}} \sqrt{(NN^\dagger)_{\mu\mu}}}$$

- Invisible Z



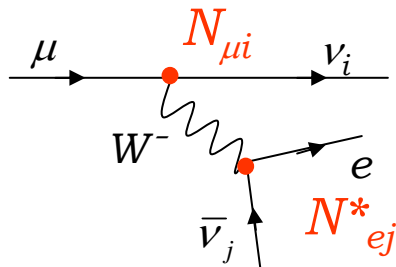
$$\rightarrow \frac{\sum_{ij} (N^+ N)_{ij}}{\sqrt{(NN^\dagger)_{ee}} \sqrt{(NN^\dagger)_{\mu\mu}}}$$

- Universality tests

$$\rightarrow \frac{(NN^\dagger)_{\alpha\alpha}}{(NN^\dagger)_{\beta\beta}}$$

Info on $(NN^\dagger)_{\alpha\alpha}$

G_F is measured in μ -decay



$$\Gamma = \frac{G_F^2 m_\mu^5}{192\pi^3} \sum_i |N_{\mu i}|^2 \sum_j |N_{ej}|^2$$

$$G_F^2 = \frac{G_{F,\text{exp}}^2}{\sum_i |N_{\mu i}|^2 \sum_j |N_{ej}|^2}$$

Oscillation Probability in Vacuum

$$i \frac{d}{dt} |\nu_i\rangle = \hat{H}_{free} |\nu_i\rangle = \sum_j E_j \delta_{ij} |\nu_j\rangle \quad \Leftrightarrow \quad \sum_i |\nu_i\rangle \langle \nu_i| = I$$

Effective Evolution

$$i \frac{d}{dt} \begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} = \tilde{N} \tilde{U}^* \begin{pmatrix} E_1 E_1 & 0 & 0 & 0 \\ 0 & E_2 E_2 & 0 & 0 \\ 0 & 0 & 0 & E_3 E_3 \end{pmatrix} (\tilde{N} \tilde{U})^{-1} \begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} \quad \Leftrightarrow \quad \sum_\alpha |\nu_\alpha\rangle \langle \nu_\alpha| \neq I$$

$$P_{\alpha\beta} = \left| \langle \nu_\beta | \nu_\alpha(L) \rangle \right|^2 = \frac{\left| \sum_i N_{\beta i} \sum_{\alpha i} N_{\alpha i}^* e^{-i \frac{m_i^2}{2E} L} \right|^2}{\sum_i |N_{\alpha i}|^2 \sum_i |N_{\beta i}|^2}$$

Zero
Distance
Effect

Non-unitarity from see-saw

$$\mathcal{L} = \mathcal{L}_{SM} + i \bar{N}_R \not{\partial} N_R - Y_\nu \bar{L} H N_R - M N_R N_R$$

Integrate out N_R $\mathcal{L}^{eff} = \mathcal{L}_{SM} + \frac{1}{M} \mathcal{L}^{d=5} + \frac{1}{M^2} \mathcal{L}^{d=6} + \dots$

$$YY^T/M (L L H H)$$

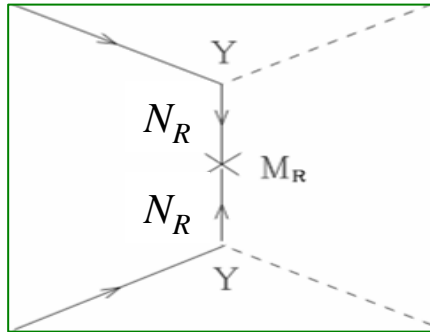
d=5 operator
it gives mass to ν

$$YY^+/M^2 (\bar{L} H) \not{\partial} (H L)$$

d=6 operator
it renormalises kinetic energy

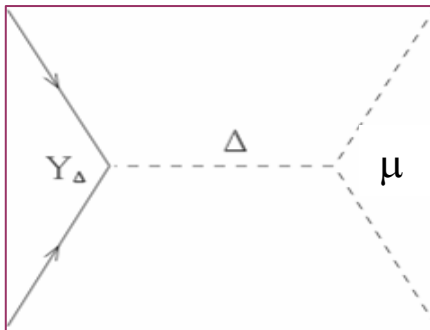
ν masses beyond the SM

★ Tree-level realizations



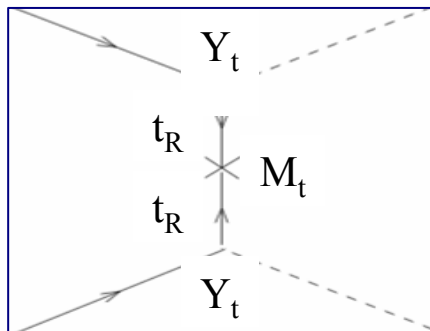
Heavy fermion singlet N_R (Seesaw I)

→ deviations from unitarity



Heavy scalar triplet Δ

→ no deviations from unitarity



Heavy fermion triplet t_R

→ deviations from unitarity

Abada,
Antusch,
Biggio,
Bonnet,
Hambye,
M.B.G.

...adding near detectors...

Test of zero-distance effect: $P_{\alpha\beta}(E,0) = |(NN^\dagger)_{\alpha\beta}|^2 \neq \delta_{\alpha\beta}$

- MINOS: $(NN^\dagger)_{\mu\mu} = 1 \pm 0.05$
- NOMAD: $(NN^\dagger)_{\mu\tau} < 0.09$ $(NN^\dagger)_{e\tau} < 0.013$
- BUGEY: $(NN^\dagger)_{ee} = 1 \pm 0.04$
- KARMEN: $(NN^\dagger)_{\mu e} < 0.05$

$$|N| = \begin{pmatrix} 0.75 - 0.89 & 0.45 - 0.66 & < 0.27 \\ 0.00 - 0.69 & 0.22 - 0.81 & 0.57 - 0.85 \\ ? & ? & ? \end{pmatrix}$$

→ also all $|N_{\mu i}|^2$ determined

Number of events

$$n_{ev} \sim \int dE \frac{d\Phi_{\alpha}(E)}{dE} P_{\alpha\beta}(E, L) \sigma_{\beta}(E) \varepsilon(E)$$

ν produced and detected in CC

$$\left\{ \begin{array}{l} \frac{d\Phi_{\alpha}}{dE} \sim \frac{d\Phi_{\alpha}^{SM}}{dE} (NN^{+})_{\alpha\alpha} \\ \sigma_{\beta} \sim \sigma_{\beta}^{SM} (NN^{+})_{\beta\beta} \end{array} \right.$$

$$n_{ev} \sim \int dE \frac{d\Phi_{\alpha}^{SM}(E)}{dE} (NN^{+})_{\alpha\alpha} P_{\alpha\beta}(E, L) (NN^{+})_{\beta\beta} \sigma_{\beta}^{SM}(E) \varepsilon(E)$$

$$\hat{P}_{\alpha\beta}(E, L) = \left| \sum_i N_{\alpha i}^* e^{iP_i L} N_{\beta i} \right|^2$$

Exceptions:

- measured flux
- leptonic production mechanism
- detection via NC

In the future...

TESTS OF UNITARITY

RARE LEPTON DECAYS

- $\mu \rightarrow e \gamma$ $(NN^\dagger)_{e\mu} < 7.1 \cdot 10^{-5}$ (MEGA)
- $\tau \rightarrow e \gamma$ $(NN^\dagger)_{e\tau} < 1.6 \cdot 10^{-2}$ (BABAR)
- $\tau \rightarrow \mu \gamma$ $(NN^\dagger)_{\mu\tau} < 1.0 \cdot 10^{-2}$ (Belle)