

QCD corrections to the electroweak sphaleron rate

Dietrich Bödeker (Bielefeld University)

Work with Philipp Klose 2510.20594

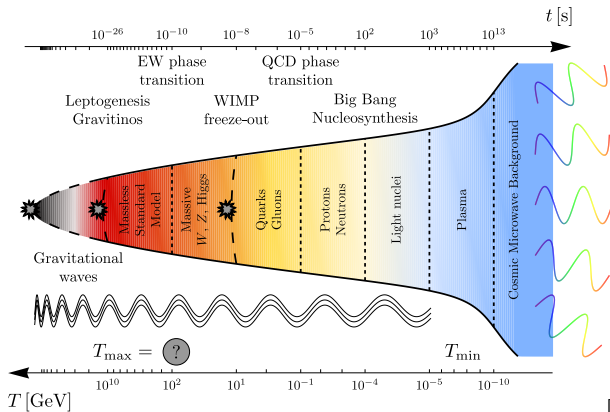


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- baryon number violation in the early Universe
- electroweak sphalerons
- non-abelian plasma physics
- weak-isospin conductivity
- QCD interactions
- effect on sphaleron rate
- summary and outlook

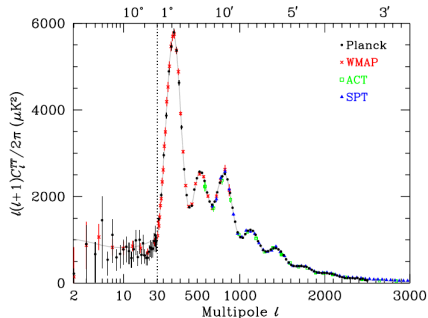
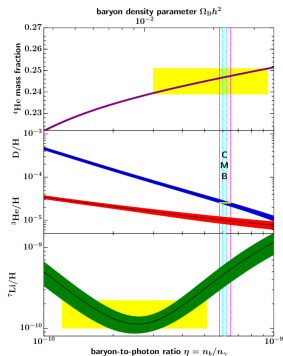
The early Universe



- beautiful connection particle physics $\leftrightarrow$ early universe
- different stages of evolution: inflation, matter-antimatter asymmetry, dark matter, dark energy
- how are these phases related to extensions of the Standard Model?

Matter-antimatter asymmetry

astrophysics $\Rightarrow$ no antimatter in the universe



[PDG 2019]

abundance of matter:

$$\frac{n_B}{n_\gamma} = [0.03 \text{ (visible)} \quad 5.8 - 6.6 \text{ (BBN)} \quad 6.09 \pm 0.06 \text{ (CMB)}] \times 10^{-10}$$

Independent measurements at very different times!

initial condition after inflation: $B = 0$

what is the origin of matter?

$$\eta_B = \frac{n_B}{n_\gamma} = (6.09 \pm 0.06) \times 10^{-10}$$

present value has to be explained!

necessary conditions for generating a matter-antimatter asymmetry [Sakharov]

1. baryon number violation
2. C and CP violation
3. deviation from thermal equilibrium

baryon and lepton number not conserved in Standard Model

divergence given by triangle anomaly

$$\begin{aligned}\partial_\mu J_B^\mu &= \partial_\mu J_L^\mu \\ &= \frac{n_f}{32\pi^2} \left(g^2 F_{\mu\nu}^a \tilde{F}^{a\mu\nu} - g'^2 B_{\mu\nu} \tilde{B}^{\mu\nu} \right)\end{aligned}$$

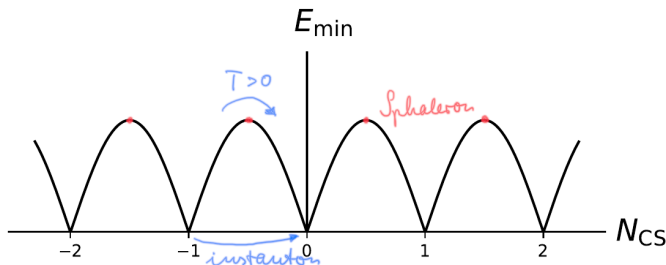
n_f : number of families,

$F_{\mu\nu}^a$, $B_{\mu\nu}$ SU(2), U(1) field strengths, gauge couplings g and g' .

baryon and lepton number are related to the topology of the weak non-abelian gauge fields:

$$\begin{aligned}
 B(t) - B(0) &= n_f \int_0^t dt' \int d^3x \frac{g^2}{32\pi^2} \underbrace{F_{\mu\nu}^a \tilde{F}^{a\mu\nu}}_{=\mathbf{E}^a \cdot \mathbf{B}^a} \\
 &= n_f [N_{\text{CS}}(t) - N_{\text{CS}}(0)]
 \end{aligned}$$

N_{CS} : Chern-Simons number



in pure gauge theory

$$\left\langle [N_{\text{CS}}(t) - N_{\text{CS}}(0)]^2 \right\rangle = \Gamma_{\text{sph}} V t$$

Γ_{sph} : CS-diffusion rate, sphaleron rate

baryon number dissipation

$$\frac{d}{dt} B = -\gamma_{B+L} B$$

fluctuation $\leftrightarrow$ dissipation

$$\gamma_{B+L} = \frac{n_f^2 V}{2 \langle B^2 \rangle} \Gamma_{\text{sph}}$$

analogous process in QCD: 'strong sphalerons'
 $\rightarrow$ anomalous chirality violation

only relevant in early Universe:

- low T : $\Gamma_{\text{sph}} \sim e^{-4\pi/\alpha} = O(10^{-154})$
- near EW scale: $\Gamma_{\text{sph}} \simeq \kappa \frac{m_W^7}{(\alpha_W T)^3} e^{-E_{\text{sph}}/T}$ $E_{\text{sph}} \simeq 8\pi v(T)/g$
- $T \gtrsim T_{\text{EW}}$: $v(T) = 0$, EW symmetry unbroken
 $\Gamma_{\text{sph}} = \kappa T^4$ unsuppressed

Γ_{sph} enters Boltzmann equations for baryogenesis

most popular scenarios:

- electroweak baryogenesis in first-order phase transition
[Kuzmin, Rubakov, Shaposhnikov]
- leptogenesis: create lepton asymmetry through sterile-neutrino decay,
partially converted to baryon asymmetry through sphalerons

[Fukugita, Yanagida]

$\Delta N_{\text{CS}} \sim 1$ requires

$$g^2 \int d^4x \underbrace{\mathbf{E} \cdot \mathbf{B}}_{\sim \dot{\mathbf{A}} \cdot \nabla \times \mathbf{A}} \sim g^2 k^{-2} \mathbf{A}^2 \stackrel{!}{\sim} 1$$

avoid Boltzmann suppression:

$$\text{energy} \sim \int d^3x \mathbf{B}^2 \sim k^{-1} \mathbf{A}^2 \stackrel{!}{\lesssim} T$$

$$\Rightarrow k \gtrsim g^2 T, \quad k \sim g \mathbf{A}$$

magnetic screening length $\sim (g^2 T)^{-1}$

$\Rightarrow k \sim g^2 T$ non-perturbative [Linde]

$\rightarrow$ lattice

occupancy $\frac{1}{e^{k/T} - 1} \simeq \frac{T}{k} \gg 1$ when $k \ll T$

classical field limit

solve classical field equation for magnetic scale ($\equiv$ ultrasoft) gauge fields

most field modes are hard ($k \sim T$), not classical

include current carried by hard particles

$$D_0 \mathbf{E} + \mathbf{D} \times \mathbf{B} = \mathbf{J}$$

mean free path for small-angle scattering $\sim \frac{1}{g^2 T \ln(1/g)}$

small angle scattering

- momenta almost unchanged
- weak gauge boson exchange changes weak isospin

logarithmic approximation: $\frac{1}{g^2 T \ln(1/g)} \ll \frac{1}{g^2 T} \Rightarrow \mathbf{J} = \sigma \mathbf{E}$

σ : weak-isospin conductivity

NB: not a transport coefficient!

analogous consideration in QCD $\rightarrow$ color conductivity

$$\sigma \mathbf{E} = -\sigma \dot{\mathbf{A}} + \dots \quad \text{damping term}$$

$$\sigma \sim g^2 \times (\text{mean free path}) \times T^2 \sim \frac{T}{\log(1/g)}$$

$$D_t \mathbf{E} \ll \sigma \mathbf{E} \quad \text{over-damping}$$

effective classical equation of motion [DB]

$$\mathbf{D} \times \mathbf{B} = \sigma \mathbf{E} + \boldsymbol{\zeta}$$

$\boldsymbol{\zeta}$: Gaussian white noise

$$\Gamma_{\text{sph}} = q \frac{2\pi T}{3\sigma} \alpha_W^5 T^4, \quad \alpha_W \equiv \frac{g^2}{4\pi}$$

lattice: $q = 10.0 \pm 0.3$ [Moore; Annala, Rummukainen]

so far: only EW interactions included in σ

almost half of the weak-isospin current carried by quarks

quarks scatter through strong interaction

$$\text{mean free path} \sim \frac{1}{g_S^4 \log(g_S^{-1})T}$$

for $g_S^4 \ll g^2$ this would be much larger than the mean free path for EW small-angle scattering

however: at m_Z $g_S^4 = 2.2$, $g^2 = 0.44$

linearized Vlasov equation

electroweak interactions only:

$$(D_t + \mathbf{v} \cdot \mathbf{D}) \delta f_{\pm} \pm g \mathbf{v} \cdot \mathbf{E} f'_{\text{eq}} = \hat{C}_W \delta f_{\pm}$$

non-abelian $\mathbf{E} = \mathbf{E}^a \tau^a / 2$ (adjoint rep. of SU(2))

contribution to weak-isospin current:

$$\mathbf{J}^a = \frac{g}{2} N_{\text{fam}} N_c \int \frac{d^3 p}{(2\pi)^3} [\delta f_+^a - \delta f_-^a] \mathbf{v}$$

$$\underbrace{(D_t + \mathbf{v} \cdot \mathbf{D})}_{g^2 T} \delta f_{\pm} \pm g \mathbf{v} \cdot \mathbf{E} f'_F = \underbrace{\hat{C}_W}_{g^2 \log(1/g) T} \delta f_{\pm}$$

leading log approximation: neglect convective term

→ solution

$$\delta f_{\pm} = \frac{\mp g \mathbf{v} \cdot \mathbf{E}}{c_W} f'_F$$

insert into current →

$$\sigma = \frac{m_W^2}{3 c_W}$$

we may choose $\mathbf{E}$ diagonal in flavor space

$$\mathbf{E} = \frac{1}{2}\mathbf{E}^3 \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

$\Rightarrow \delta f_{\pm}$ diagonal

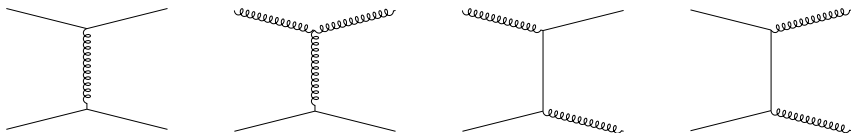
for left-handed (LH) quarks:

$$\delta f_{+} = f_{\text{F}}(1 - f_{\text{F}}) \begin{pmatrix} \eta_{uL} & 0 \\ 0 & \eta_{dL} \end{pmatrix}, \quad \eta_{dL} = -\eta_{uL}$$

RH quarks, LH antiquarks, and gluons remain in equilibrium

$$\pm g \mathbf{v} \cdot \mathbf{E} f'_{\text{F}} = \left[\hat{C}_W + \hat{C}_S \right] \delta f_{\pm}$$

leading log contributions to $\hat{C}_S$:



diffusion approximation [Arnold, Moore, Yaffe; Ghiglieri, Moore, Teaney]

$$\left[\hat{C}_S \delta f_+ \right]_{\alpha, \text{gl}} = \frac{g_S^2 \log(g_S^{-1})}{6\pi} m_G^2 T \partial_{p^i} \left\{ f_F(p_0) [1 - f_F(p_0)] \partial_{p^i} \eta_\alpha \right\} + \text{gain terms}$$

gain terms cancel in sum over α

collision term commutes with rotations $\Rightarrow \delta f_{\pm}$ and η_{α} have same angular dependence as force term

$$\eta_{uL} = \frac{g}{2T} \mathbf{v} \cdot \mathbf{E}^3 \chi(p_0) .$$

$\rightarrow$ ordinary differential equation for χ

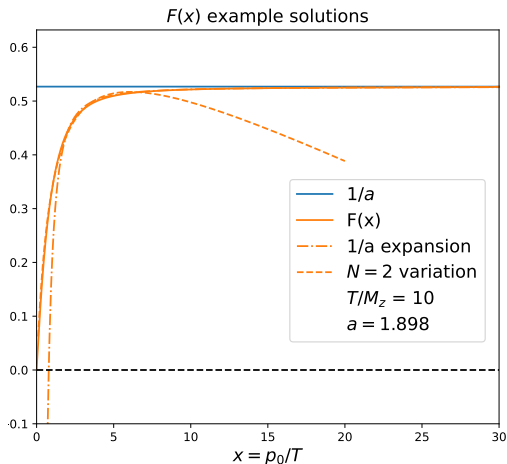
boundary value problem

follows from variational problem $\delta Q[F] = 0$ [Heiselberg; Arnold, Moore, Yaffe]

ansatz for F in terms of N basis functions

already $N = 2$ gives good approximation

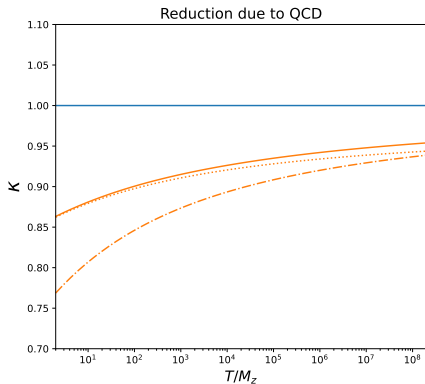
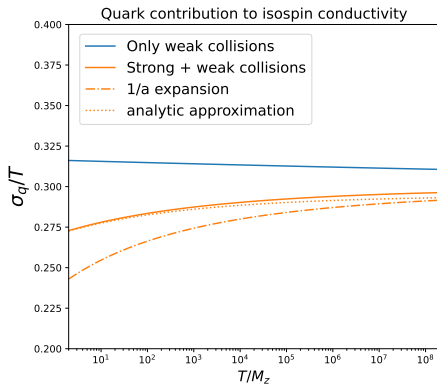
Results: variational approach



$$a = \frac{3\pi c_W T}{g_s^2 m_G^2 \log(\mu_*/m_G)}$$

quark contribution to weak-isospin conductivity

$$\sigma_q = \frac{m_{Wq}^2}{3 c_W} \kappa$$



Results: electroweak sphaleron rate

QCD reduces σ_q by 10 to 15%

→ electroweak sphaleron rate increases by 6%

modest increase, considering the size of g_S

- baryon number B is rapidly violated in the symmetric phase of the Standard Model
- rate for B violation enters computation of baryogenesis in the early Universe
- rate determined by non-perturbative W boson dynamics
- rate computed with effective classical theory
- overdamped due to large weak-isospin conductivity σ
- we included QCD collisions in LL approximation, decrease σ
- sphaleron rate decreased by 6%

- go beyond LL approximation
- full leading order in QCD scattering
- order g_S corrections
- include masses near EW crossover
- beyond LL in electroweak interaction: new operators in effective equation of motion

$$c_W = \gamma + \frac{g^2 T}{2\pi} K, \quad K = 3.0410,$$

where γ is defined implicitly as a solution of the equation

$$\gamma = \frac{g^2 T}{2\pi} \log \left(\frac{m_W}{\gamma} \right)$$

$$m_W^2 = \left(\frac{2}{3} + \frac{N_s}{6} + \frac{(1 + N_c)N_{\text{fam}}}{12} \right) g^2 T^2$$

maximize the integral

$$Q[F] = \int_0^{\infty} \frac{x^2 e^x}{(e^x + 1)^2} \left\{ F - \frac{1}{4} F'^2 - \left(\frac{a}{2} + \frac{b}{4x} \frac{e^x + 1}{e^x - 1} + \frac{1}{2x^2} \right) F^2 \right\}$$

$$b = \frac{m_{\infty}^2}{m_G^2} = \frac{1}{6}, \quad x = \frac{p_0}{T}, \quad F(x) = \frac{c_W \chi(p_0)}{a}. \quad (1)$$

$$a = \frac{3\pi c_W T}{g_s^2 m_G^2 \log(\mu_*/m_G)}, \quad b = \frac{m_{\infty}^2}{m_G^2} = \frac{1}{6}, \quad x = \frac{p_0}{T}, \quad F(x) = \frac{c_W \chi(p_0)}{a}. \quad (2)$$

