

DARK MATTER, NEW PHYSICS AND COSMOLOGY IN THE LIGHT OF LHC

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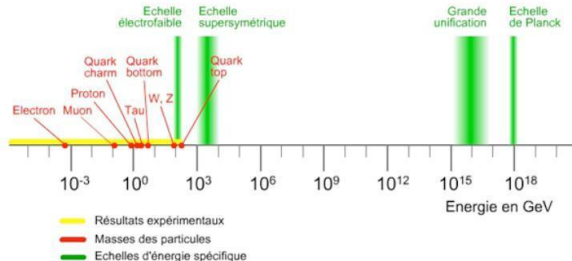
ANGERS, November 2010

Outline

- 1 Supersymmetry
- 2 MSSM & SUSY Breaking
- 3 Astrophysics, Cosmology & Dark Matter
- 4 Programs & Tools
- 5 Constraints of mSUGRA parameter space
- 6 DM of different models AMSB, MM-AMSB & HC-AMSB

Introduction

- The energy scales



- SUSY** : extension of **SM** to the TeV scale.
- LHC** : Machine at the TeV scale which describes new physics.

Definition

- **SUSY** : symmetry relates bosons and fermions.
- Q generator of **SUSY** algebra :
 $\bar{Q}|boson\rangle = |fermion\rangle$ & $Q|fermion\rangle = |boson\rangle$
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- generators create a new state called **superpartner**, it corresponds to a supersymmetric particle.
- **SUSY** algebra contains anti-commutators because of fermionic generators. (super Poincaré algebra)

$$\{Q_\alpha, \bar{Q}_{\dot{\alpha}}\} = 2\sigma_{\alpha,\dot{\alpha}}^\mu P_\mu$$

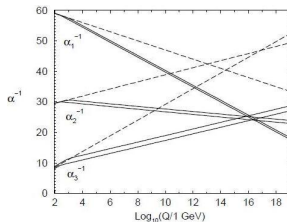
$$\{Q_\alpha, Q_\beta\} = \{\bar{Q}_{\dot{\alpha}}, \bar{Q}_{\dot{\beta}}\} = 0$$

$$[Q_\alpha, P^\mu] = [\bar{Q}_{\dot{\beta}}, P^\mu] = 0$$

$$[Q_\alpha, M_{\mu\nu}] = [\bar{Q}_{\dot{\beta}}, M_{\mu\nu}] = 0$$

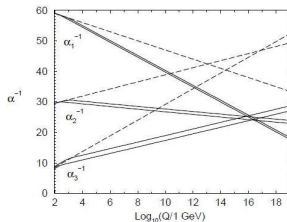
Motivations in particle physique

- Unification of gauge couplings :

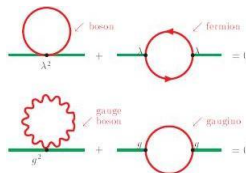


Motivations in particle physique

- Unification of gauge couplings :



- Solution of the hierarchy problem :



Description of gravity

- The graviton, hypothetic particle mediates the gravity, have a spin 2 and the others bosons (photon, gluons, W et Z) have a spin 1 \Rightarrow they correspond to different representations of Poincaré Algebra.
- **Supersymmetric Transformations** \Rightarrow
spin 2 \rightarrow spin 3/2 \rightarrow spin 1 \rightarrow spin 1/2 \rightarrow spin 0.

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spin 2 \rightarrow spin 3/2 \rightarrow spin 1 \rightarrow spin 1/2 \rightarrow spin 0.
- Infinitesimal Transformations $\delta_\epsilon = \epsilon^\alpha Q_\alpha$, $\bar{\delta}_{\bar{\epsilon}} = \bar{Q}_{\dot{\alpha}} \bar{\epsilon}^{\dot{\alpha}}$, \Rightarrow
 $\{\delta_\epsilon, \bar{\delta}_{\bar{\epsilon}}\} = 2(\epsilon\sigma^\mu\bar{\epsilon})P_\mu$
- Local coordinate transformation $\epsilon = \epsilon(x) \Rightarrow$ **Supergravity**
(*local version of SUSY*).

Field content

- The content of **MSSM**

Superfield	Bosons	Fermions	$SU_c(3)$	$SU_L(2)$	$U_Y(1)$		
<i>Gauge</i>							
G^a	gluon g^a	gluino \tilde{g}^a	8	1	0		
V^k	Weak $W^k(W^\pm, Z)$	wino, zino $\tilde{w}^k(\tilde{w}^\pm, \tilde{z})$	1	3	0		
V'	Hypercharge $B(\gamma)$	bino $\tilde{b}(\tilde{\gamma})$	1	1	0		
<i>Matter</i>							
L_i	sleptons	$\tilde{L}_i = (\tilde{\nu}, \tilde{e})_L$	leptons	$L_i = (\nu, e)_L$	1	2	-1
E_i				$E_i = e_R$	1	1	2
Q_i	squarks	$\tilde{Q}_i = (\tilde{u}, \tilde{d})_L$	quarks	$Q_i = (u, d)_L$	3	2	1/3
U_i				$U_i = u_R^c$	3*	1	-4/3
D_i				$D_i = d_R^c$	3*	1	2/3
<i>Higgs</i>							
H_1	Higgses	H_1	higgsinos	\tilde{H}_1	1	2	-1
H_2				H_2	\tilde{H}_2	1	2

Lagrangian of MSSM



$$\mathcal{L} = \mathcal{L}_{SUSY} + \mathcal{L}_{Breaking} = \mathcal{L}_{Gauge} + \mathcal{L}_{Yukawa} + \mathcal{L}_{Breaking}$$

- **superpotentiel** which reproduces Yukawa interaction in SM

$$W = \epsilon_{ij}(y_{ab}^U Q_a^j U_b^c H_2^i + y_{ab}^D Q_a^j D_b^c H_1^i + y_{ab}^L L_a^j E_b^c H_1^i + \mu H_1^i H_2^j)$$

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- such as **SM** but : fields \rightarrow superfields.
Difference : term which describes the Higgs mixing (absent in SM because we have 1 Higgs doublet).

R-parity

- R-parity symmetry

$$R = (-1)^{3(B-L)+2S}$$

R-parity

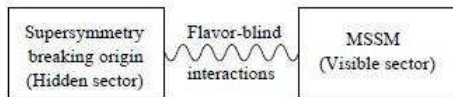
- R-parity symmetry

$$R = (-1)^{3(B-L)+2S}$$

- Consequences of R-parity conservation :
 - 1 Superparticles are created by pairs.
 - 2 The **LSP** is absolutely stable.
 - 3 The **LSP** is an excellent candidate for dark matter, and it is in general a *neutralino*.

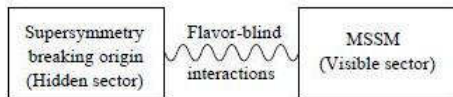
Breaking mechanisms

- The breaking occurs in a **hidden sector** and the ordinary matter will be in the **visible sector**



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- 4 important mechanisms for generating the SUSY breaking :
 - Gravity mediation (SUGRA);
 - Gauge mediation ;
 - Anomaly mediation ;
 - Gaugino mediation.

Soft terms for breaking SUSY

- **universality hypothesis** \Rightarrow only 5 parameters which define the mass scale :

$$\mu, m_0, m_{1/2}, A \text{ et } B \leftrightarrow \tan\beta = \frac{v_2}{v_1}$$

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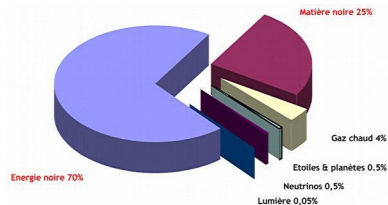
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$$\begin{aligned}
 -\mathcal{L}_{\text{Breaking}} = & m_0^2 \sum_i |\varphi_i|^2 \\
 & + \left(\frac{1}{2} m_{1/2} \sum_\alpha \tilde{\lambda}_\alpha \tilde{\lambda}_\alpha + B[\mu H_1 H_2] \right. \\
 & \left. + A[y_{ab}^U \tilde{Q}_a \tilde{U}_b^c H_2 + y_{ab}^D \tilde{Q}_a \tilde{D}_b^c H_1 + y_{ab}^L \tilde{L}_a \tilde{E}_b^c H_1] + h.c. \right)
 \end{aligned}$$

(1)

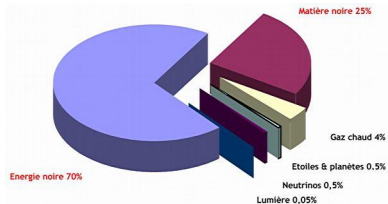
General Description

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- Types of non-baryonic DM : Cold DM formed by WIMP, Hot DM formed by relativistic light particles (neutrinos) and Warm DM sterile neutrino is a good candidate.

CDM & WIMP

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- For **WIMP** : no candidates in **SM**, SUSY offers an excellent candidate : the **neutralino (LSP)**.
- The evolution of density number $n_\chi(t)$ is described by the Boltzmann equation :

$$\frac{dn_\chi}{dt} + 3Hn_\chi = - \langle \sigma_a v \rangle [(n_\chi)^2 - (n_\chi^{eq})^2]$$

- The relic density is proportionally inverse to the cross section $\Omega_\chi h^2 \sim \frac{1}{\langle \sigma_a v \rangle}$.

LSP

- **Gravity mediation** : LSP is the lightest *neutralino* $\tilde{\chi}_1^0$.
- **Gauge mediation** : LSP is the *gravitino* \tilde{G} that is produced by the decay from **NLSP**. ($\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}, h \tilde{G}, Z \tilde{G}$ or $\tilde{l}_R \rightarrow \tau \tilde{G}$).
- **Anomaly mediation** : LSP is $\tilde{\chi}_1^0$ (decaying from NLSP) or $\tilde{\nu}_L$ (decaying from chargino $\tilde{\chi}^+ \rightarrow \tilde{\nu} l$).

SUPERISO

- **SuperIso** is a public C program that is :
 - dedicated to the flavour physics observable calculations
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 - dedicated to the flavour physics observable calculations
 - aimed to provide everyone the possibility to do the calculations in different models
 - based on the most precise calculations publicly available in the literature
- **Models** : SM, 2HDM, MSSM (mSUGRA, AMSB, NUHM, GMSB), NextMSSM.

SUPERISO RELIC

- SuperIso Relic = SuperIso (flavour physics calculations) + relic density calculation

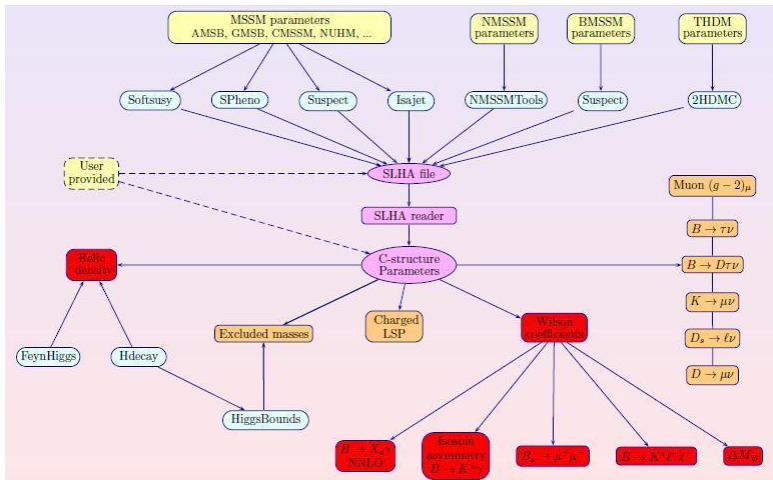
SUPERISO RELIC

- SuperIso Relic = SuperIso (flavour physics calculations) + relic density calculation
- Concept of the code consists of :
 - Automatized computation of flavour observables and relic density in SUSY
 - Flexible particle physics model implementation (mSUGRA, NUHM, AMSB, ...)
 - Flexible cosmological model implementation (dark energy, dark entropy, ...)
 - Publicly available on <http://superiso.in2p3.fr/relic>

SUPERISO RELIC

- Structure of the code consists of :
 - Generation of a SLHA file with Isajet, Softsusy, Spheeno or Suspect
 - Initialization of the variables using the SLHA file
 - Generation of additional Higgs sector variables with FeynHiggs or Hdecay
 - Calculation of W_{eff} with Fortran functions
 - Calculation of $\langle \sigma_{eff} v \rangle$ with C functions
 - Solving of the Boltzmann equation with C functions
 - Computation of the other SuperIso observables

How it works ?



Observables

- 1 CHARGED LSP
- 2 EXCLUDED MASS
- 3 DARK MATTER
- 4 BRANCHING RATIO $\bar{B} \rightarrow X_s \gamma$
- 5 ISOSPIN ASYMMETRY $B \rightarrow K^* \gamma$
- 6 MUON ANOMALOUS MAGNETIC MOMENT
- 7 BRANCHING RATIO $B_s \rightarrow \mu^+ \mu^-$
- 8 BRANCHING RATIO $B_u \rightarrow \tau \nu_\tau$
- 9 BRANCHING RATIO $B \rightarrow D \tau \nu_\tau$

mSUGRA

- In **mSUGRA**, the breaking of **SUSY** is propagated by the gravity due to the existence of a hidden sector.

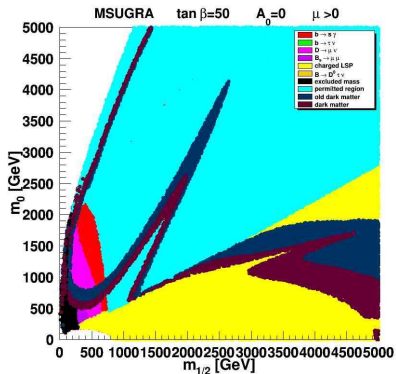
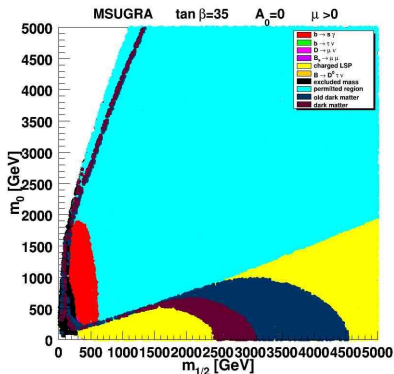
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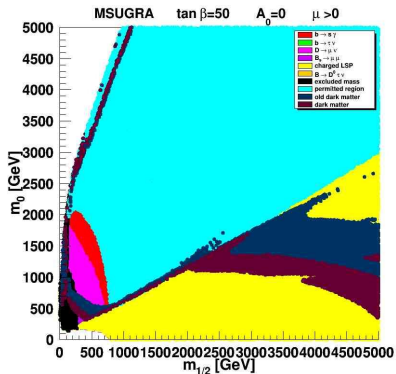
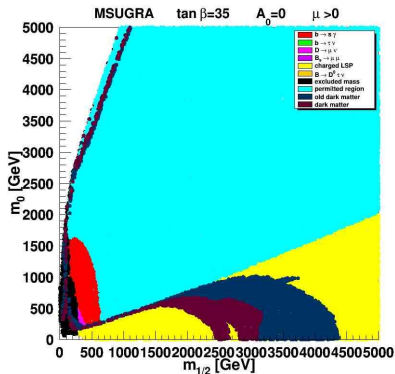
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- ISAJET, SOFTSUSY, SPHENO and SUSPECT are 4 generators of spectra \Rightarrow numerical calculation of RGE, deduction of particle masses and mixing matrices \Rightarrow calculation of decay widths, branching ratio, cross section of production...

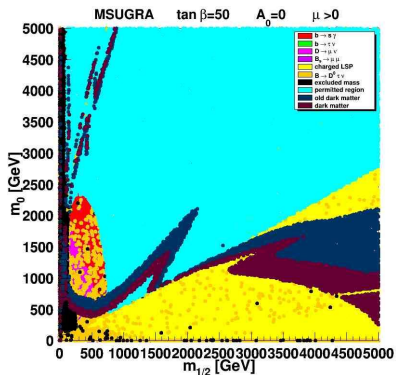
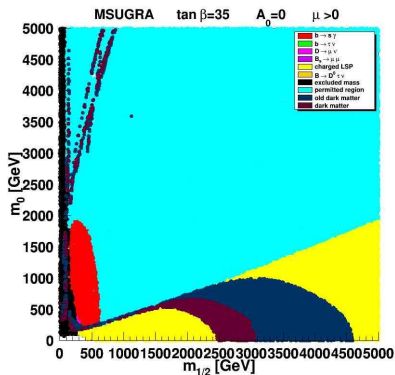
mSUGRA results (SOFTSUSY)



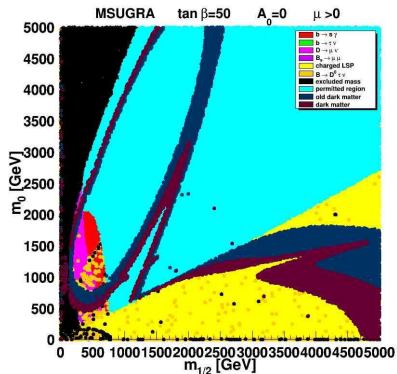
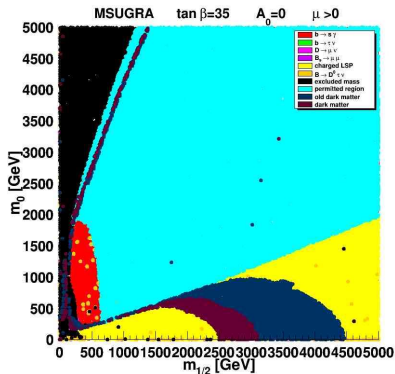
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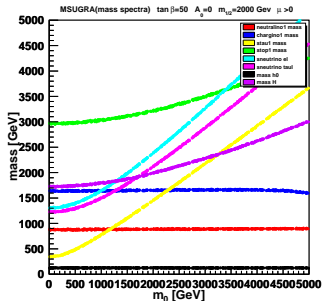
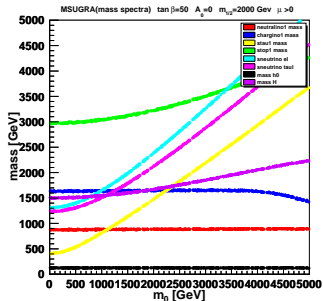
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mSUGRA results (SUSPECT)



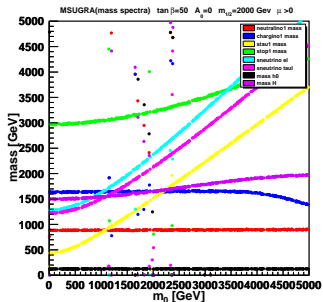
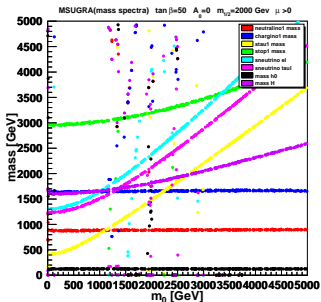
Mass spectra in mSUGRA



The mass spectra of mSUGRA model : (at left) SOFTSUSY, (at right) ISAJET

- The neutralino is the candidate of dark matter in mSUGRA.

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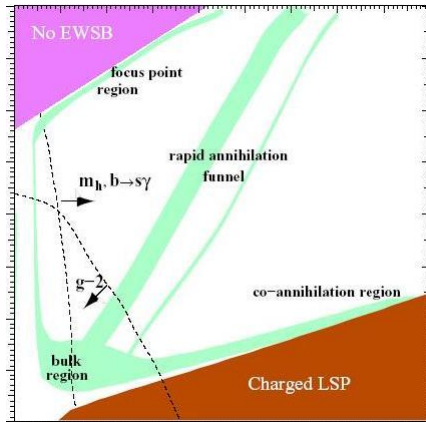


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Favoured regions

- 4 favoured regions of parameters space :



Discussion of results

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- A-funnel region : resonance situation $m_{A^0} \approx m_{H^0} \approx 2m_{\chi_1^0}$,
 s-channel Higgs exchange $\chi_1^0 \chi_1^0 \rightarrow A^0 \rightarrow b\bar{b}$ or $\tau\bar{\tau}$.

AMSB

- **AMSB** (Anomaly Mediated Supersymmetry Breaking) is a predictive framework for SUSY breaking in which the breaking of scale invariance mediates between hidden and visible sectors, and the sparticules acquire their masses due to this mediation.

AMSB

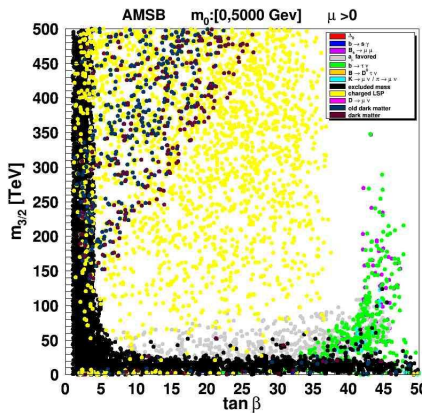
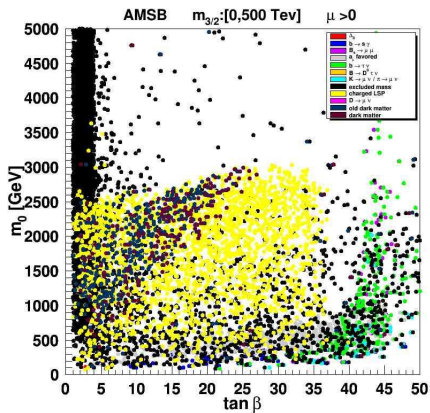
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- The parameters of this model are :

$$m_0, m_{3/2}, \tan\beta, \text{sign}(\mu).$$

AMSB results



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- There is no permitted for **DM**.
- \Rightarrow study extended models from AMSB that combine gravity with **conformal anomaly** in order to obtain some regions that propose candidates to DM.

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- \Rightarrow study extended models from AMSB that combine gravity with **conformal anomaly** in order to obtain some regions that propose candidates to DM.
- \Rightarrow **MM-AMSB** and **HC-AMSB**.

MM-AMSB

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$$m_{3/2}, \alpha, \tan\beta, \text{sign}(\mu), n_i, l_\alpha.$$

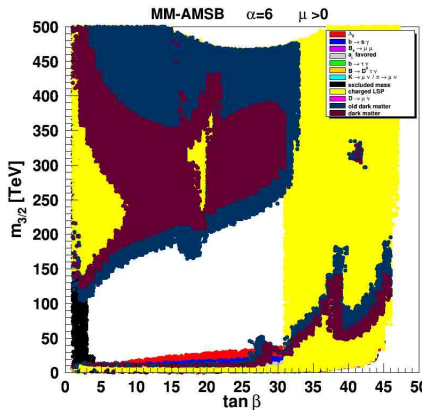
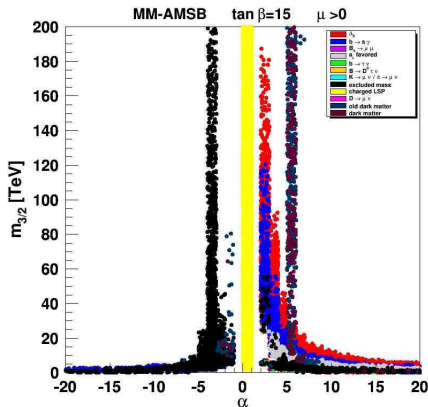
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MM-AMSB

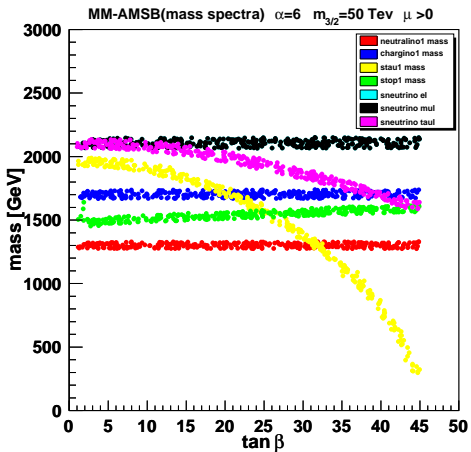
- The mass scale of SSB parameters is given by the gravitino mass $m_{3/2}$.
- α : relative contribution of anomaly mediation and gravity mediation (moduli) for soft terms.
- n_i : modular weights of visible sector of matter fields.
- l_a : the gauge kinetic function.
- I added this model in SuperIso taking $n_i=0$ and $l_a=1$.

MM-AMSB results



Parameters space in the MM-AMSB model with zero modular weights and $\mu > 0$, in the planes $(m_{3/2}, \alpha)$ at left and $(m_{3/2}, \tan\beta)$ at right

Mass spectra for MM-AMSB



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- The lightest **neutralino** is a good candidate for **DM** in MM-AMSB until $\tan\beta \sim 32$ with a mass of order 1.3 TeV.
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- Common point with **mSUGRA** model.
- LSP must be created by pairs (conservation of R-parity) \Rightarrow more difficult for detection at LHC in the near futur.
- Decay from NLSP which is the stop, $\tilde{t}_1 \rightarrow c\chi_1^0$ or $\tilde{t}_1 \rightarrow bW\chi_1^0$

Can axino be a candidate for DM ?!

- The stau can decay into an *axino* plus a particle from SM :

$$\tilde{\tau} \rightarrow \tilde{a} + \text{SM}, \quad n_{\tilde{\tau}} \sim n_{\tilde{a}} \quad \text{and} \quad \rho_{\tilde{\tau}} = m_{\tilde{\tau}} n_{\tilde{\tau}}. \quad \text{And,}$$

$$\rho_{\tilde{a}} = m_{\tilde{a}} n_{\tilde{a}} = m_{\tilde{a}} n_{\tilde{\tau}} = \frac{m_{\tilde{a}}}{m_{\tilde{\tau}}} \rho_{\tilde{\tau}}.$$

$$m_{\tilde{a}} \sim 0.1 \text{ GeV} \quad \text{and} \quad m_{\tilde{\tau}} \sim 1000 \text{ GeV} \quad \text{so} \quad \rho_{\tilde{a}} \sim 10^{-4} \rho_{\tilde{\tau}}. \quad \text{But}$$

$$\Omega_M = \frac{\rho_M^0}{\rho^c} \Rightarrow \Omega_{\tilde{a}} \sim 10^{-4} \Omega_{\tilde{\tau}}.$$

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- The stau can decay into an *axino* plus a particle from SM :
 $\tilde{\tau} \rightarrow \tilde{a} + \text{SM}$, $n_{\tilde{\tau}} \sim n_{\tilde{a}}$ and $\rho_{\tilde{\tau}} = m_{\tilde{\tau}} n_{\tilde{\tau}}$. And,
 $\rho_{\tilde{a}} = m_{\tilde{a}} n_{\tilde{a}} = m_{\tilde{a}} n_{\tilde{\tau}} = \frac{m_{\tilde{a}}}{m_{\tilde{\tau}}} \rho_{\tilde{\tau}}$.
 $m_{\tilde{a}} \sim 0.1 \text{ GeV}$ and $m_{\tilde{\tau}} \sim 1000 \text{ GeV}$ so $\rho_{\tilde{a}} \sim 10^{-4} \rho_{\tilde{\tau}}$. But
 $\Omega_M = \frac{\rho_M^0}{\rho^c} \Rightarrow \Omega_{\tilde{a}} \sim 10^{-4} \Omega_{\tilde{\tau}}$.
- This new value of $\Omega \Rightarrow$ (where axino is DM) is also excluded by other constraints of flavour physics !
- The **axino** can't be a candidate for DM in MM-AMSB.

HC-AMSB

- **HC-AMSB** (HyperCharged Anomaly Mediation) is a scenario motivated by the string theory which explain an extension of AMSB.
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- Depending on its size, the bino mass M_1 can lead to a small perturbation to the spectrum of anomaly mediation.

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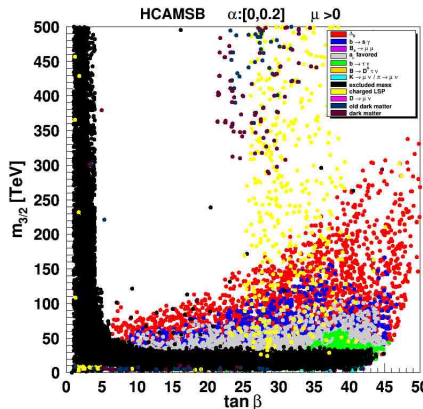
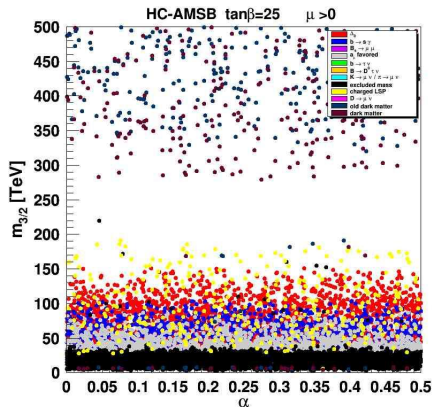
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- The parameters of this model are :

$$m_{3/2}, \alpha, \tan\beta, \text{sign}(\mu).$$

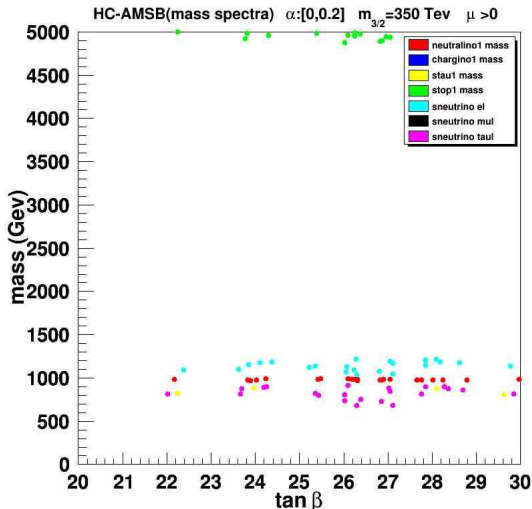
- α is a dimensionless quantity which represents the size of hypercharge contribution to soft terms relative to the AMSB contribution.

HC-AMSB results



The parameters of space in the HC-AMSB model with $\alpha \in [0, 0.2]$ and $\mu > 0$, in the planes $(m_{3/2}, \alpha)$ at left and $(m_{3/2}, \tan\beta)$ at right.

Mass spectra for HC-AMSB

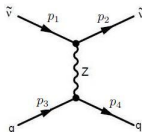


Mass spectra for HC-AMSB

- The **sneutrino tau** is a good candidate for **DM** in HC-AMSB with a mass of order of 700 GeV (this is different from mSUGRA model).
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- The diagram which corresponds to the interaction of sneutrinos with the nucleus via t-channel is :



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- mSUGRA : the neutralino is the LSP , while AMSB does'nt offer any allowed regions for the dark matter.
- MM-AMSB, as in mSUGRA, the neutralino is the LSP, but in HC-AMSB, it is the sneutrino tau that is the lightest and it is excluded by the direct detection of its interaction with the nucleus.

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- Addition of non-thermal production of relic particles.
- Implementation of an alternative entropy modification.
- Inclusion of a BBN code to test the cosmological modifications.