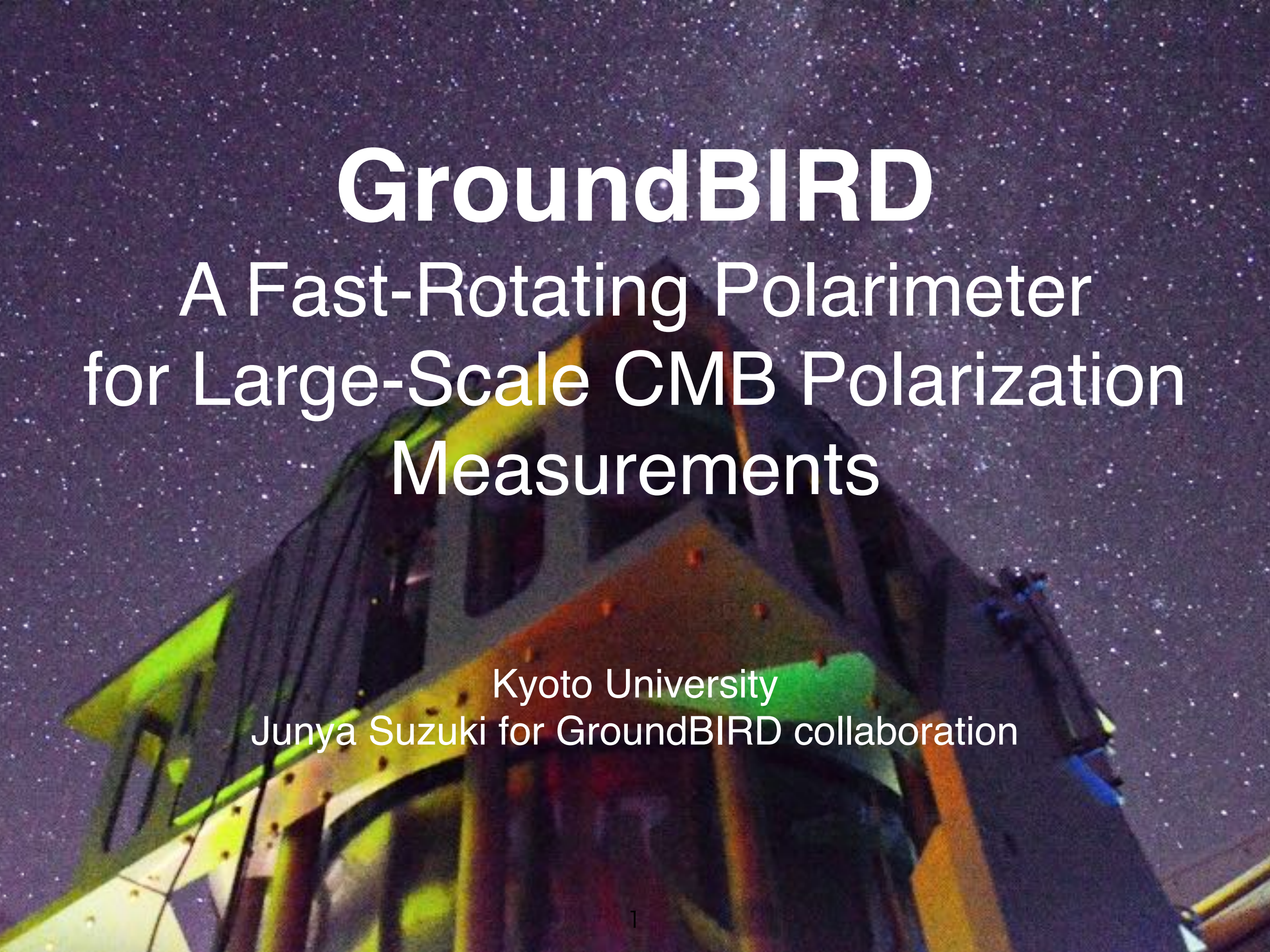




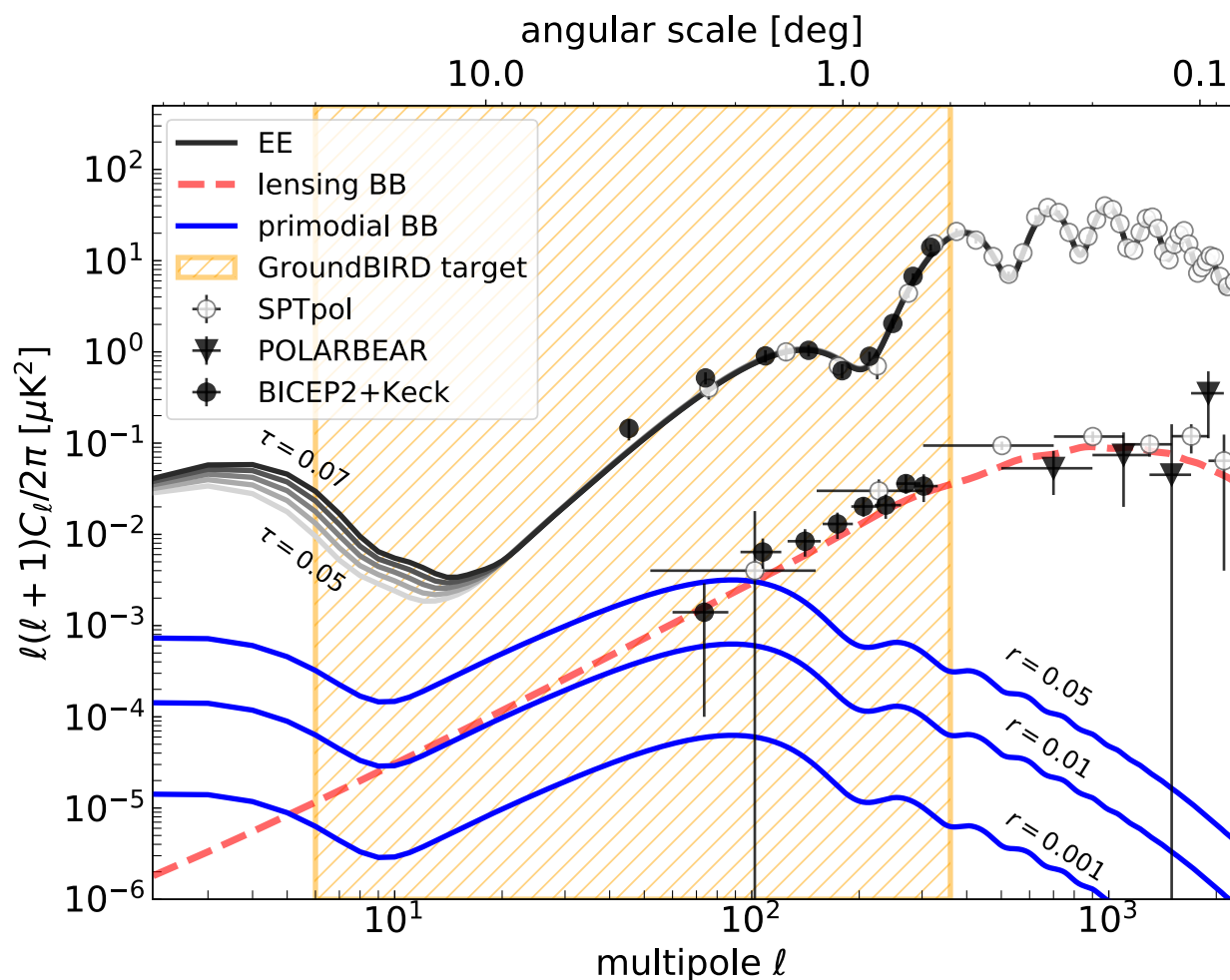
GroundBIRD

A Fast-Rotating Polarimeter
for Large-Scale CMB Polarization
Measurements

Kyoto University
Junya Suzuki for GroundBIRD collaboration

GroundBIRD

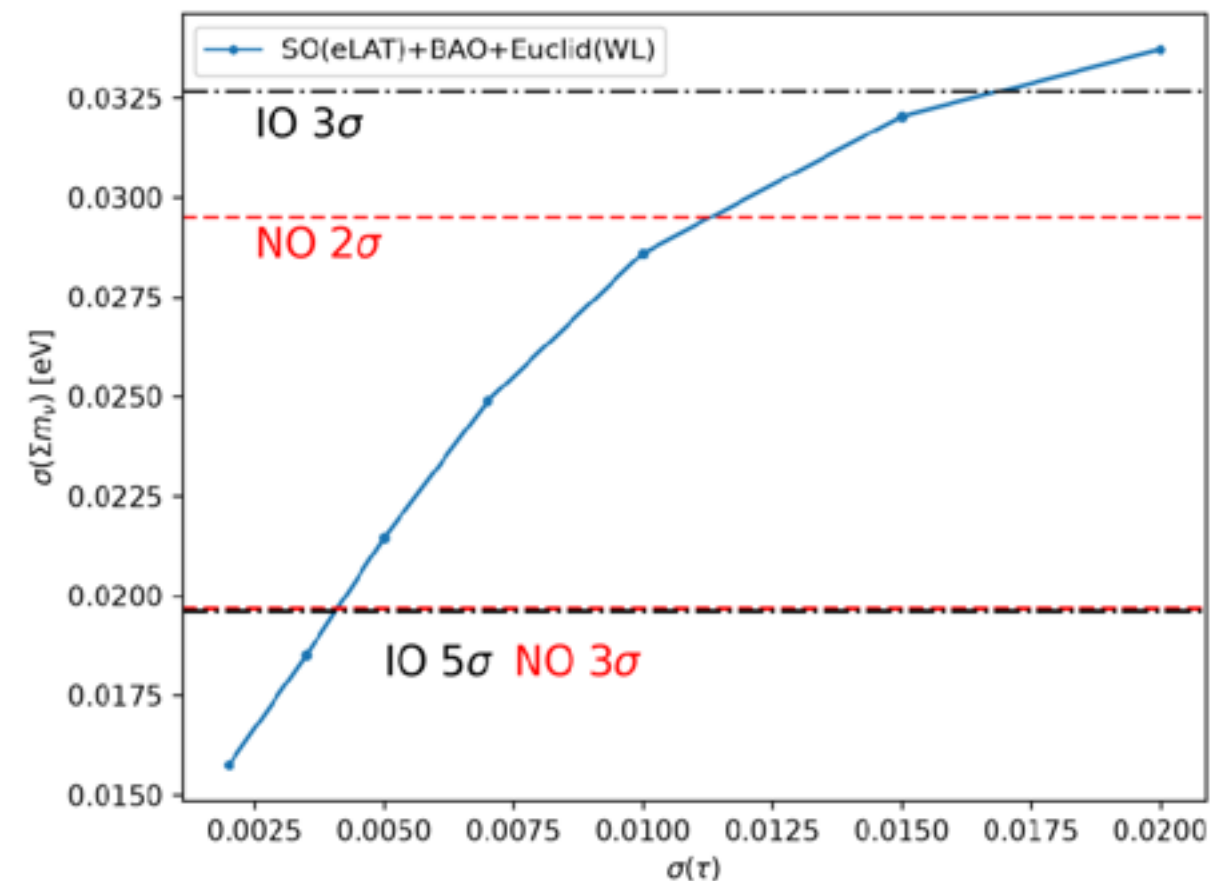
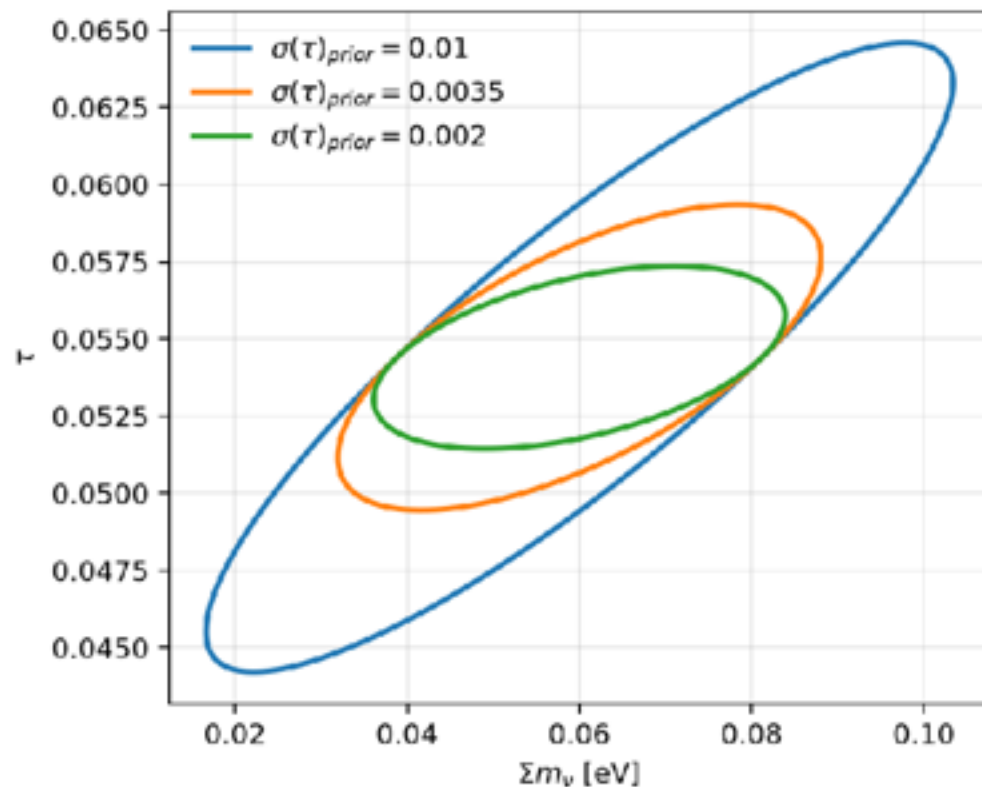
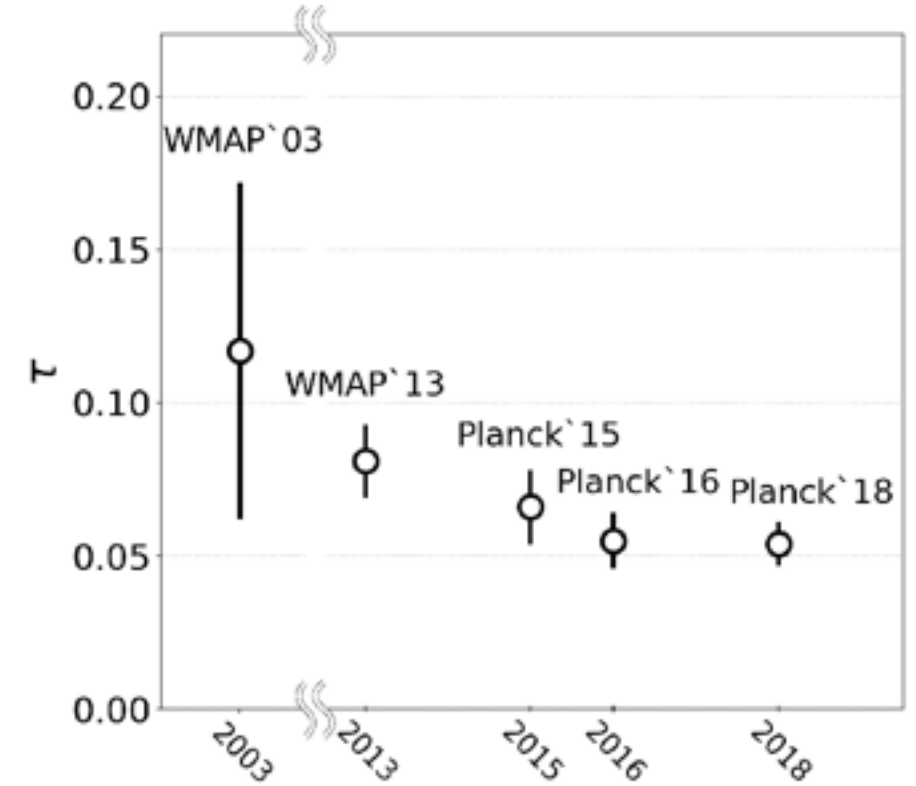
- CMB polarimeter for low- ℓ observation
 - High-speed rotation scan modulation to mitigate atmospheric fluctuation



Science target

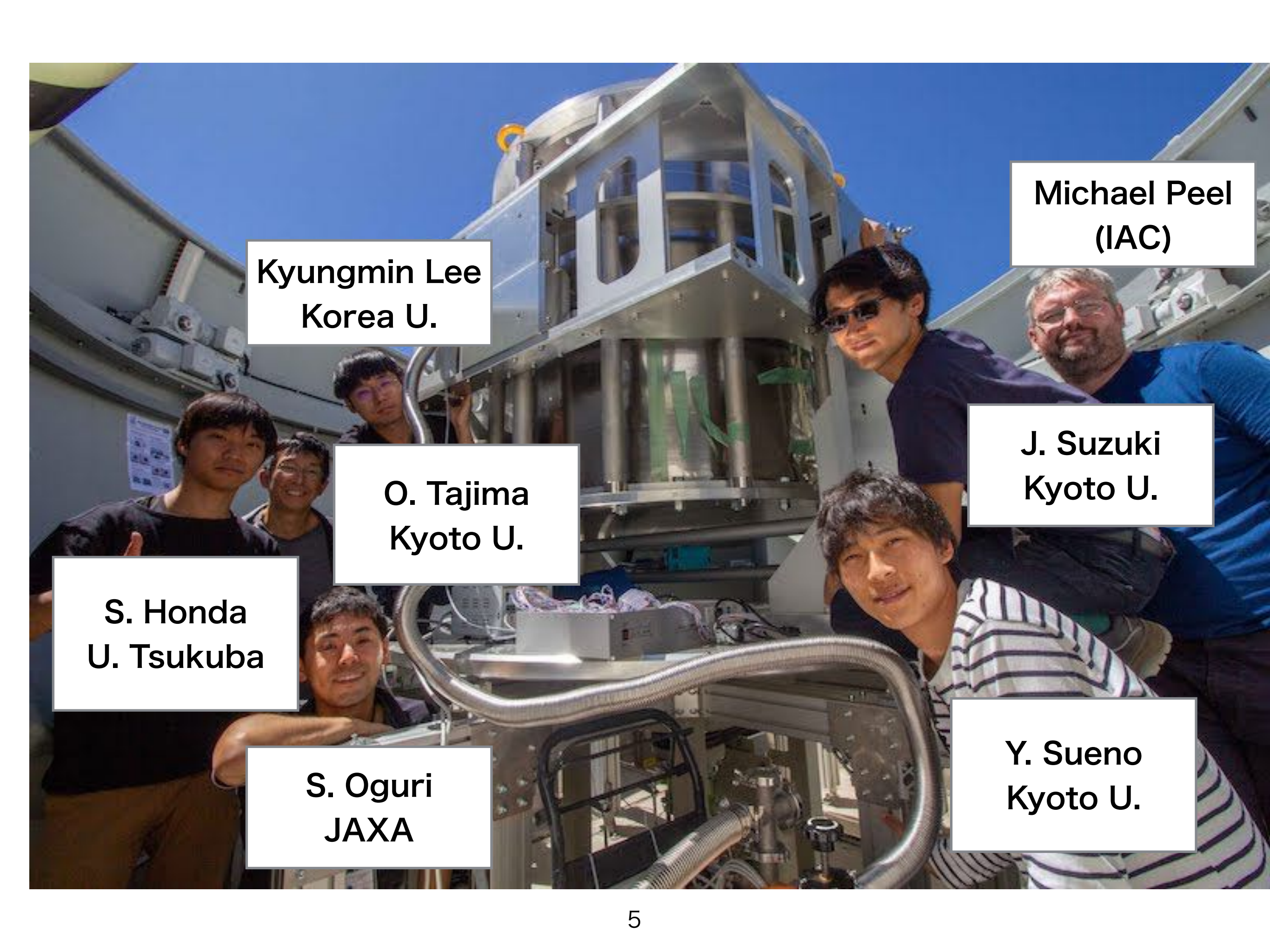
- Determination of Σm_ν requires τ measurement
- GroundBIRD + QUIJOTE expects $\sigma(\tau) \sim 0.01$ level independent measurement from the ground

Previous τ measurements



GroundBIRD



A group of researchers are gathered around a large, complex scientific instrument, likely a radio telescope component, mounted on a ship's deck. The instrument is made of metal and has a large, curved, corrugated metal tube in the foreground. The background shows the ship's structure and a clear blue sky. The researchers are of various ages and are looking towards the camera. The text labels are overlaid on the image in white boxes with black text.

Kyungmin Lee
Korea U.

Michael Peel
(IAC)

J. Suzuki
Kyoto U.

O. Tajima
Kyoto U.

S. Honda
U. Tsukuba

S. Oguri
JAXA

Y. Sueno
Kyoto U.



KEK large exp hall

Short history

- GroundBIRD was prepared and partially evaluated at KEK and shipped to Tenerife in 2019.
- After the COVID outbreak, full KIDs array was installed in 2023.



Test at IAC clean room



Reboot after COVID in 2023

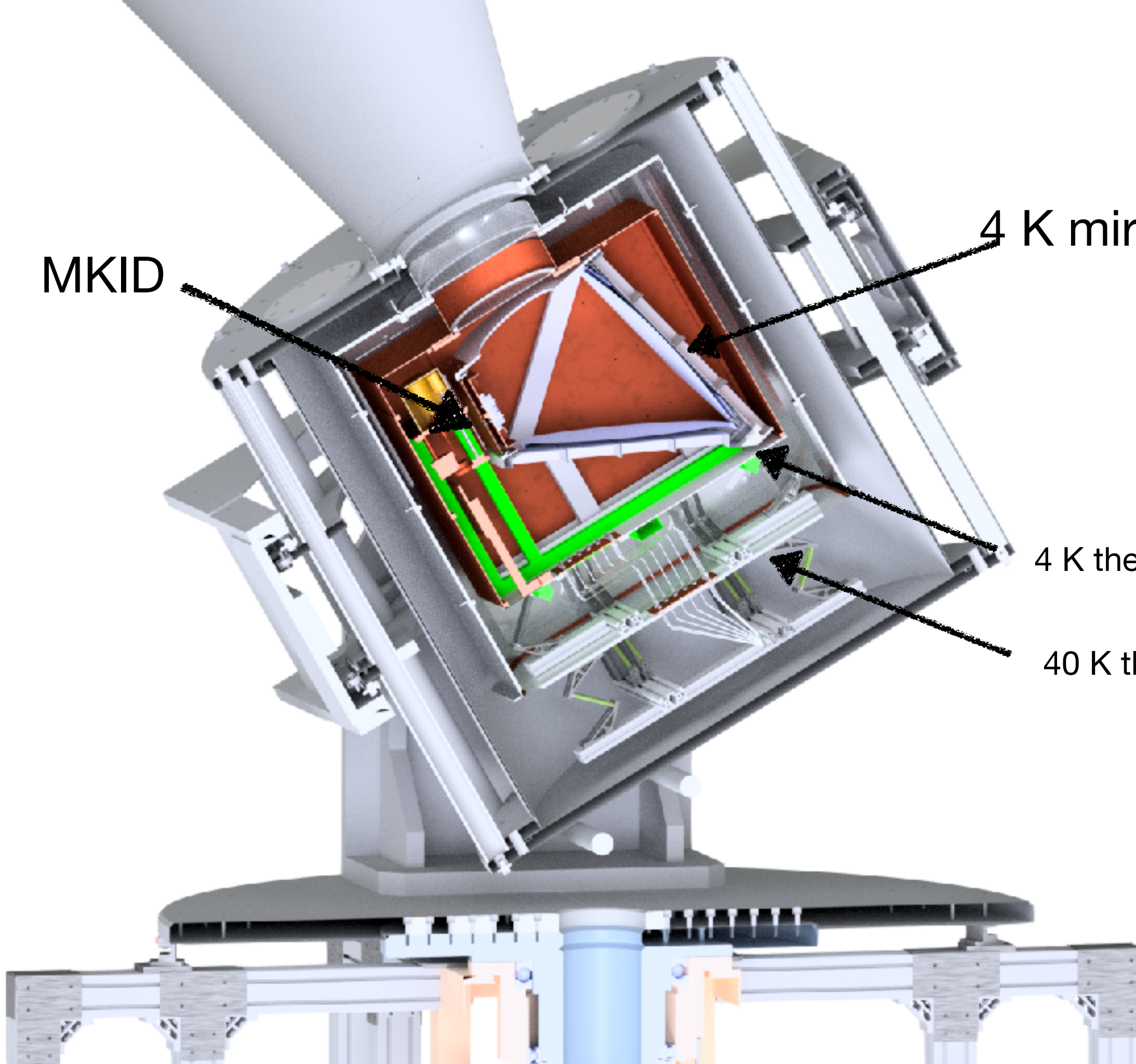


MKID

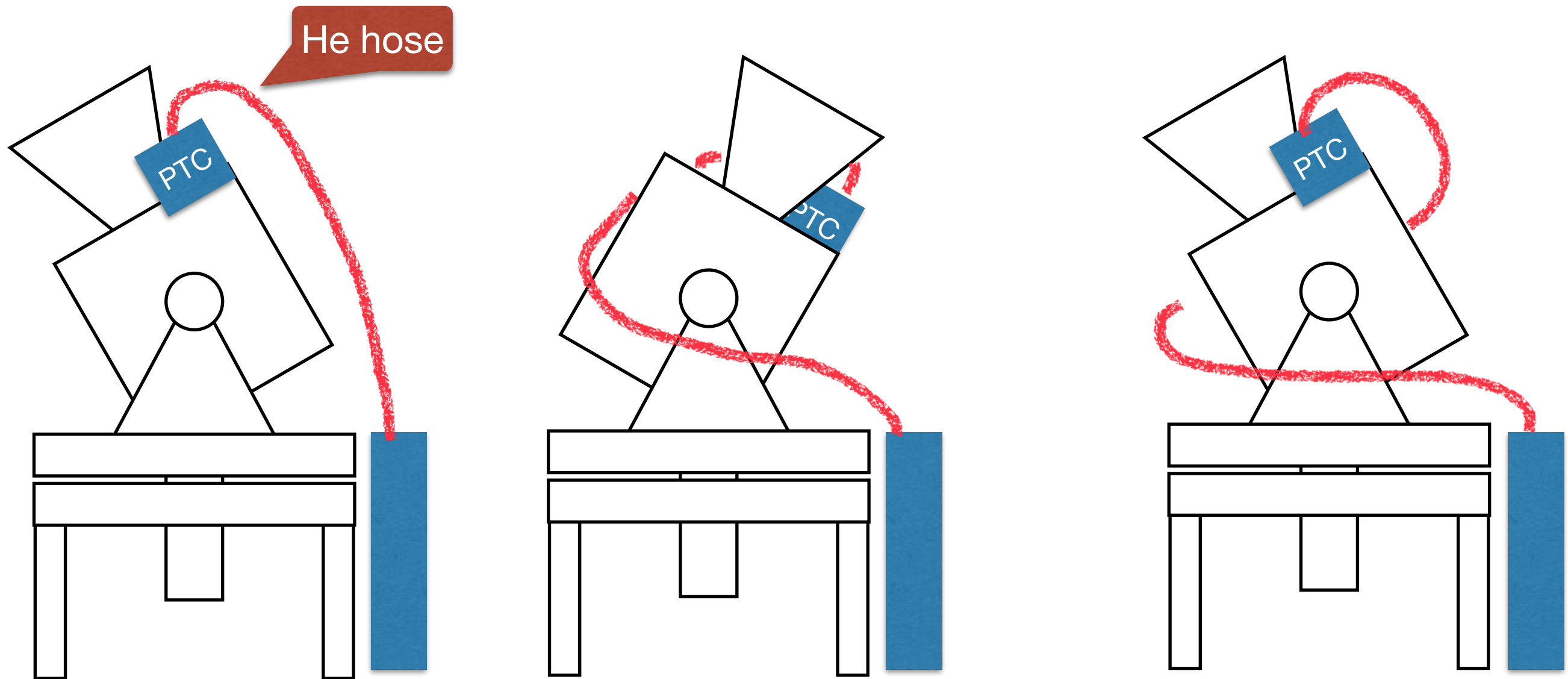
4 K mirror

4 K thermal shield

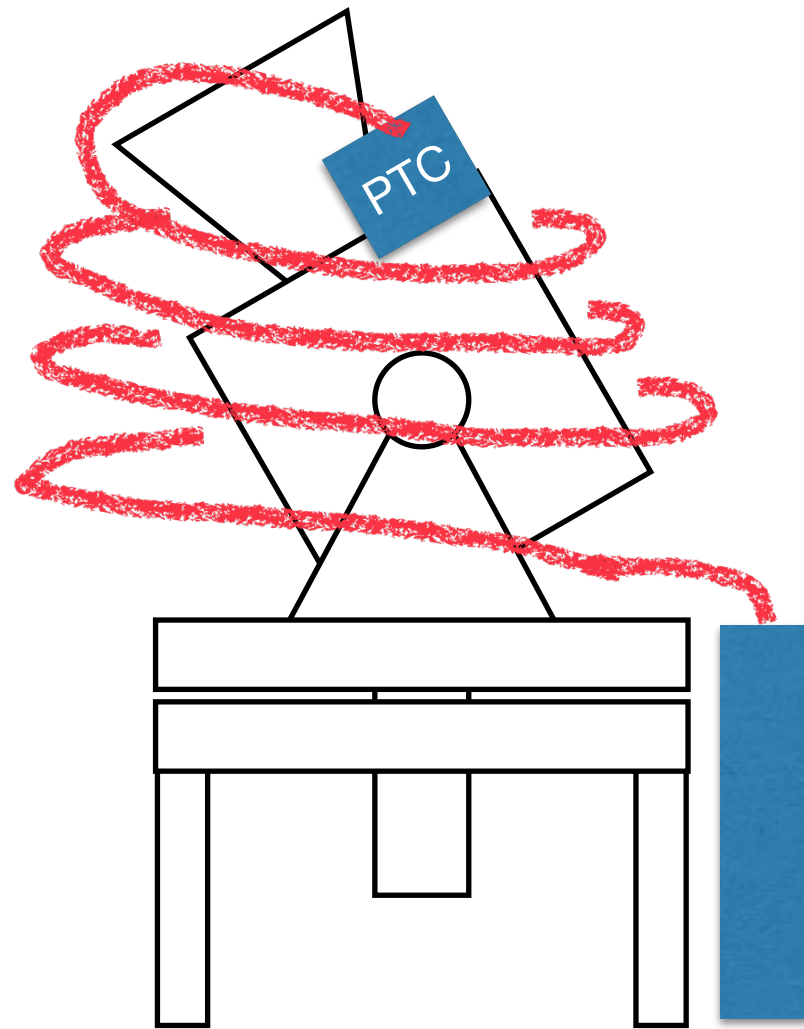
40 K thermal shield



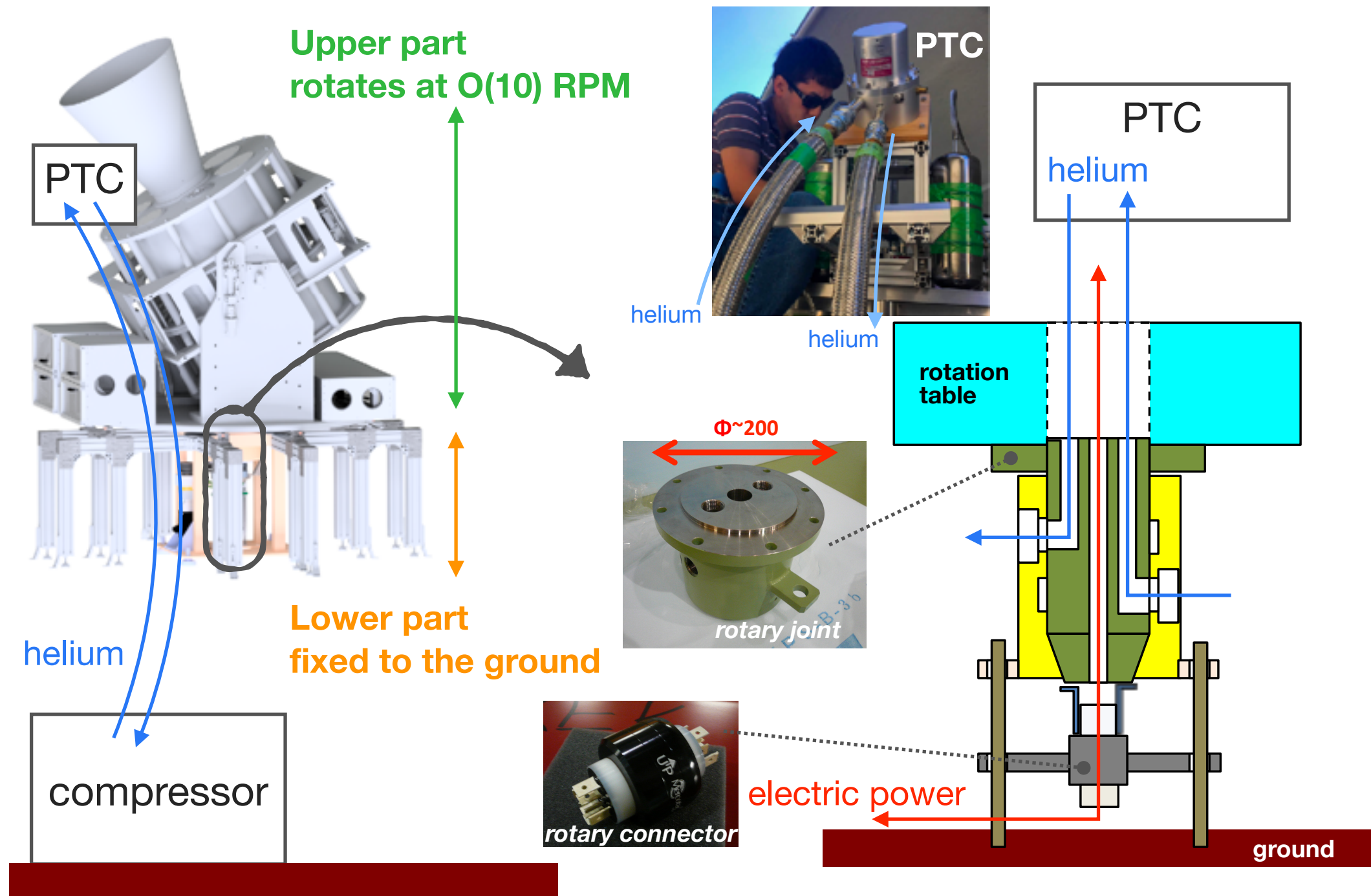
Rotation scan

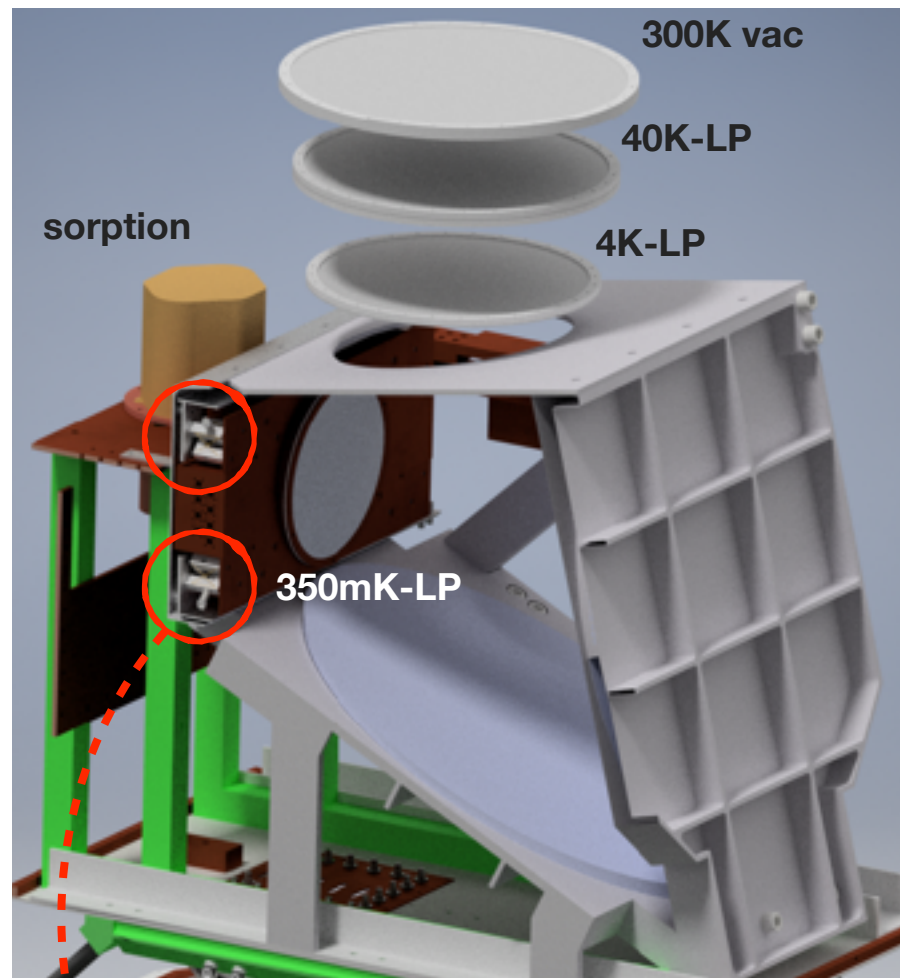


Rotation scan

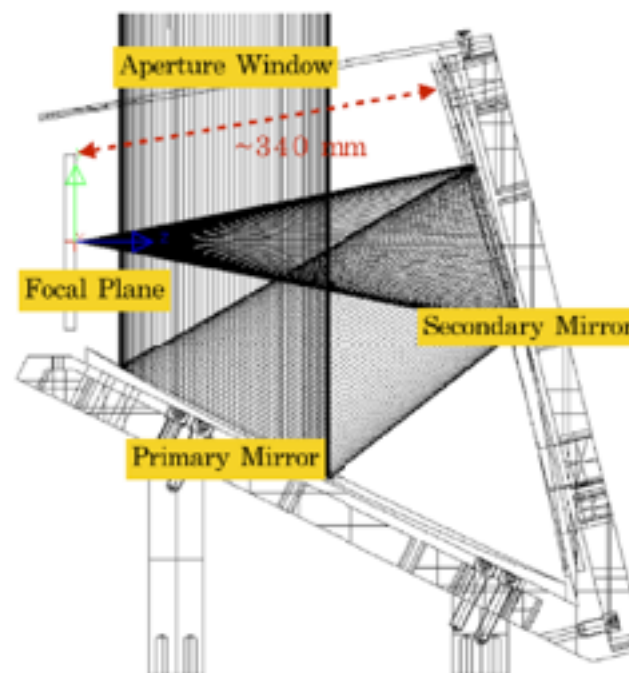


Solution: He rotary joint





mirror design

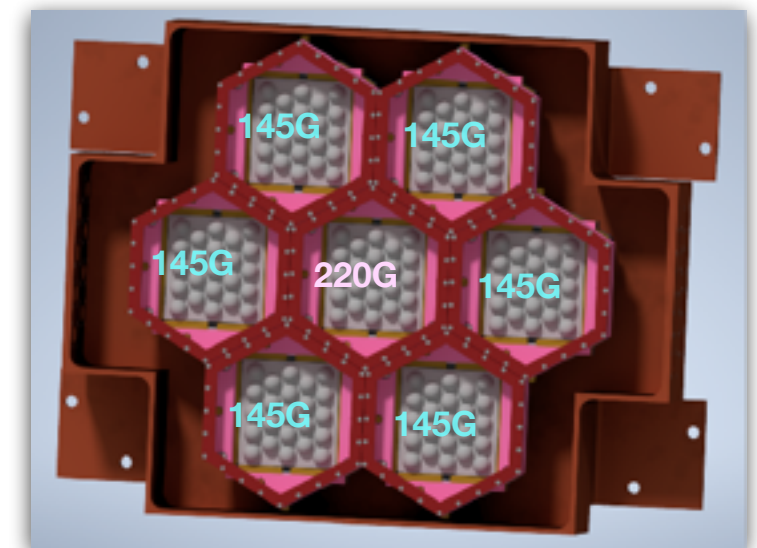


7 wafers on the focal plane

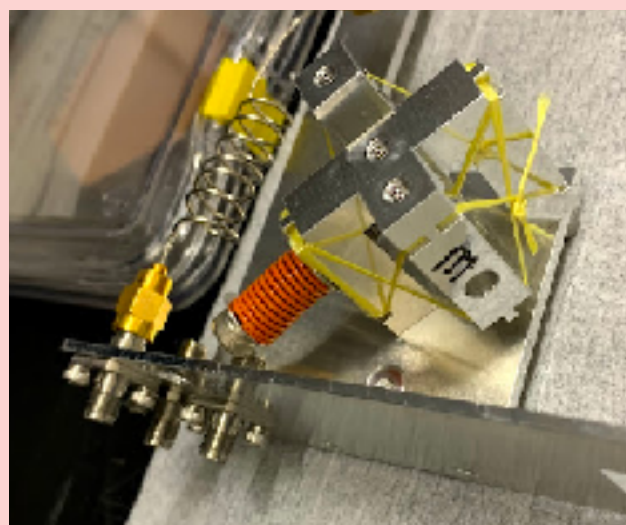
23pix. for each wafer

▶ 138 pix. with 145GHz for CMB

▶ 23 pix. with 220GHz for dust

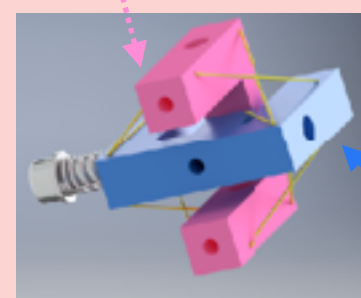


Kevlar small jigs for thermal isolation



Thermal isolation well achieved
by tensions of Kevlar wires

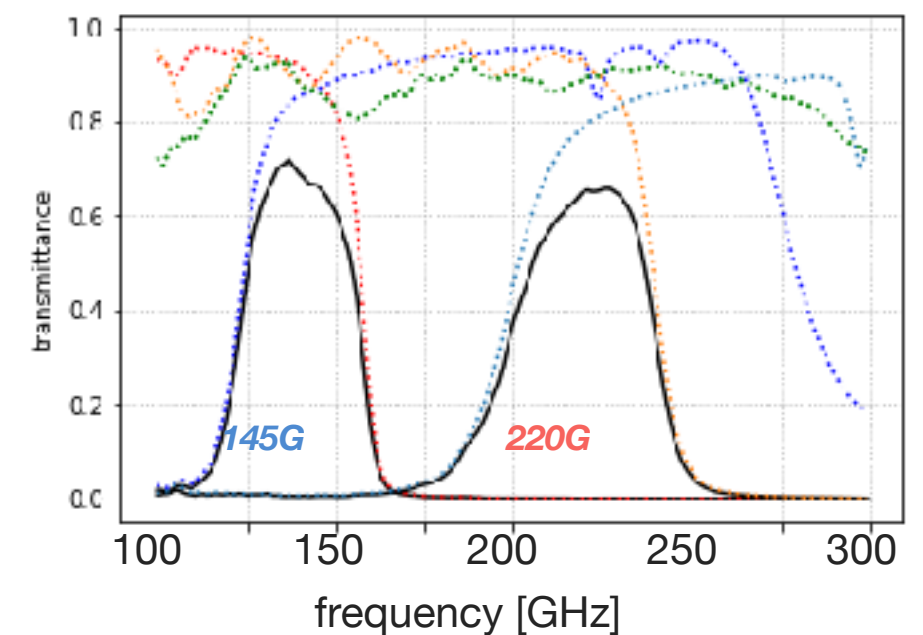
temp. 1



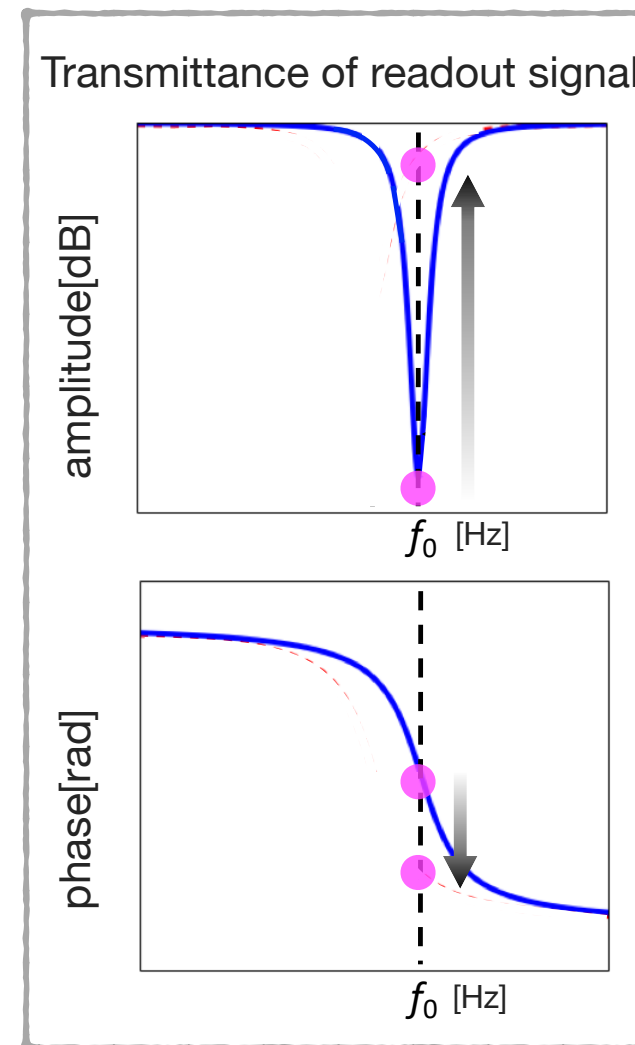
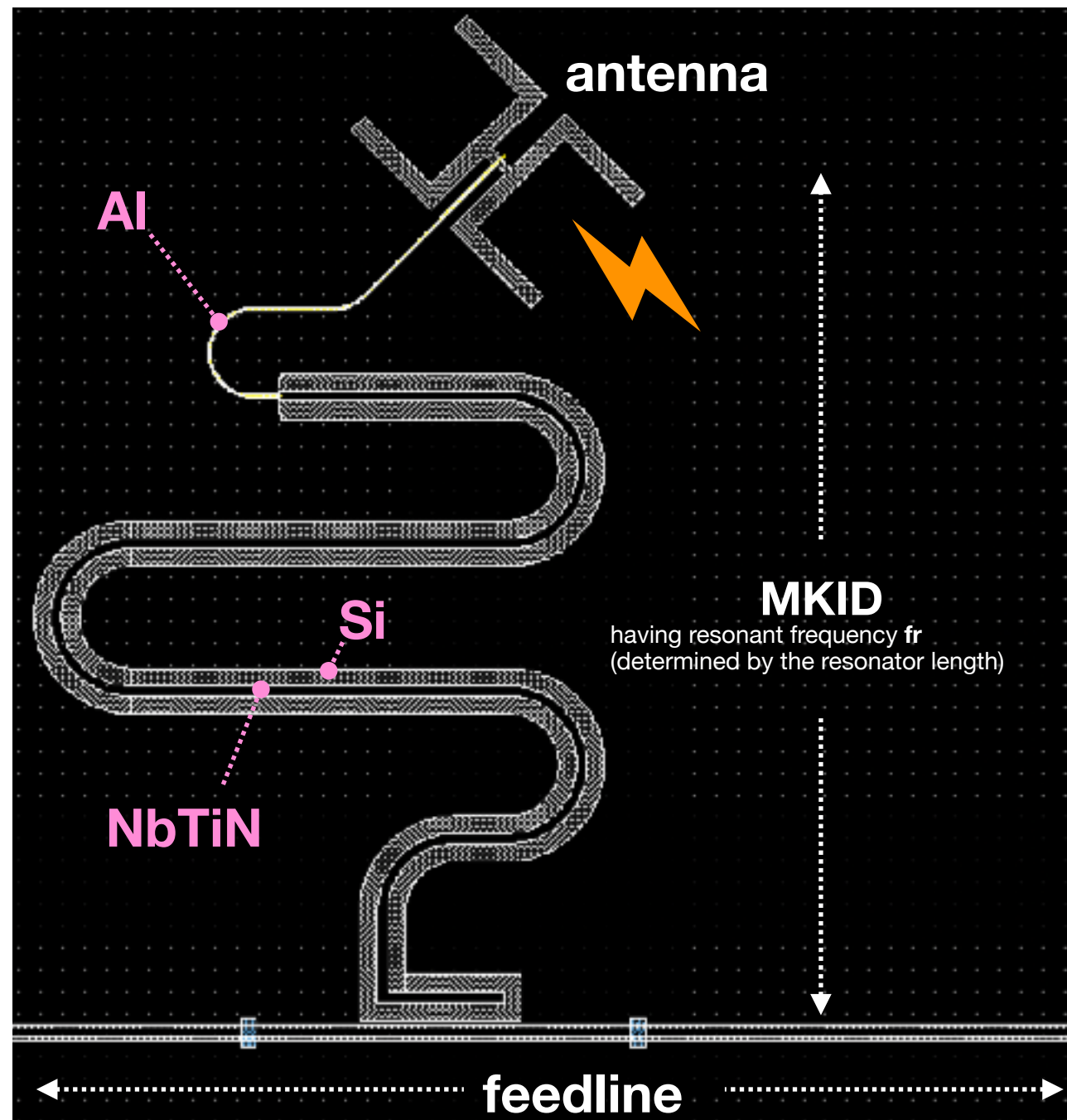
thermal conductivity
~ 0.0064 [mW/K]

temp. 2

spec. of optical filters at focal plane

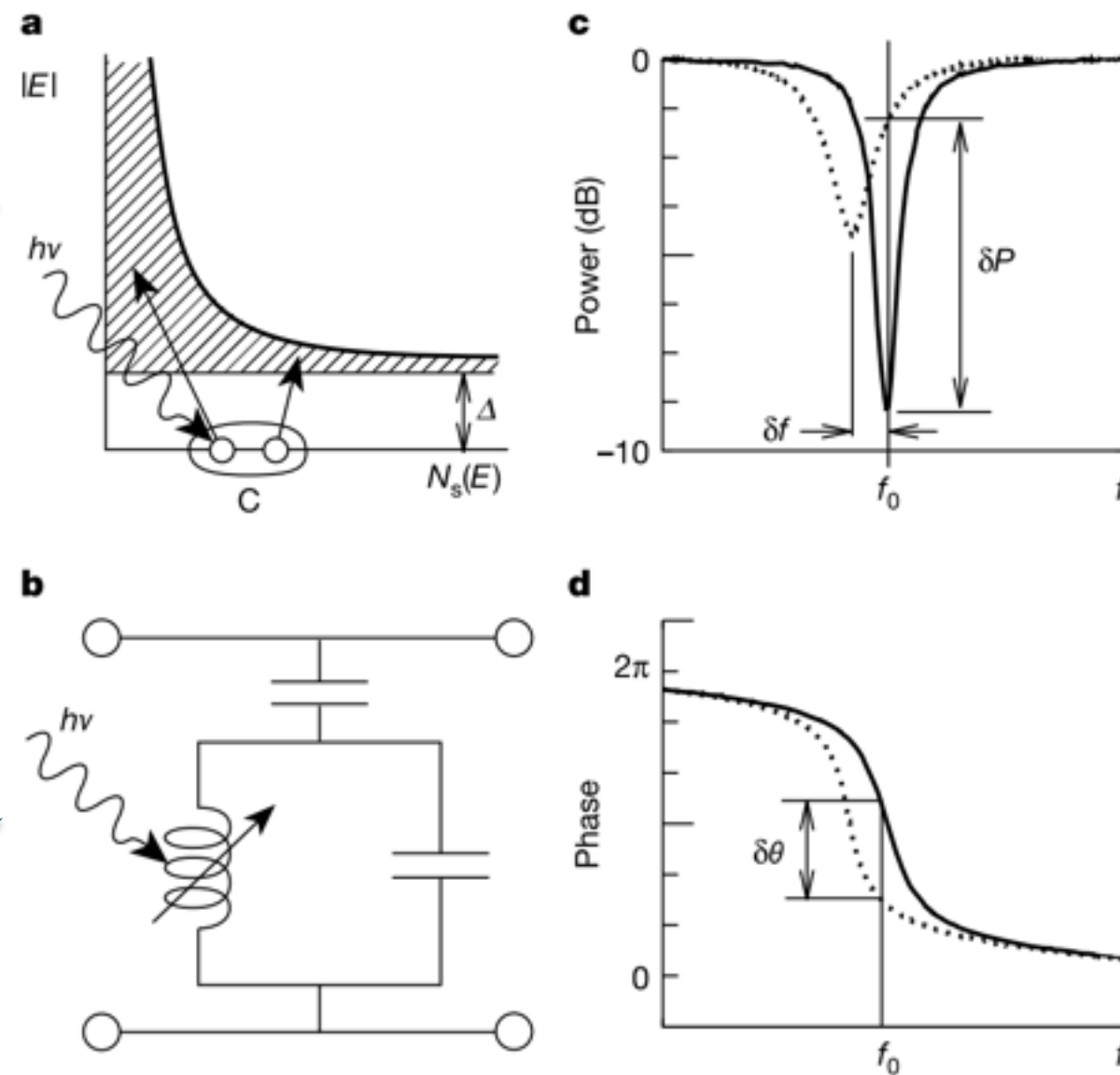


Microwave Kinetic Inductance Detector



MKID

Cooper-pair
breaking by incident
photon



Resonator
frequency/depth
change

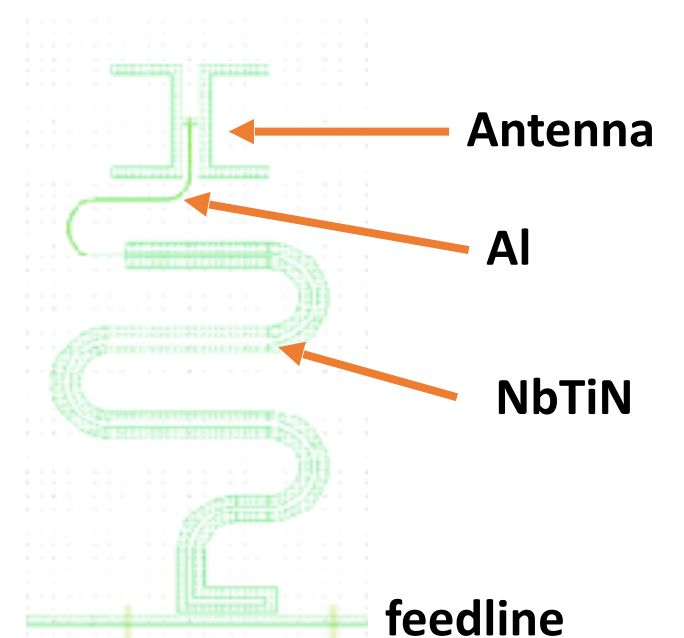
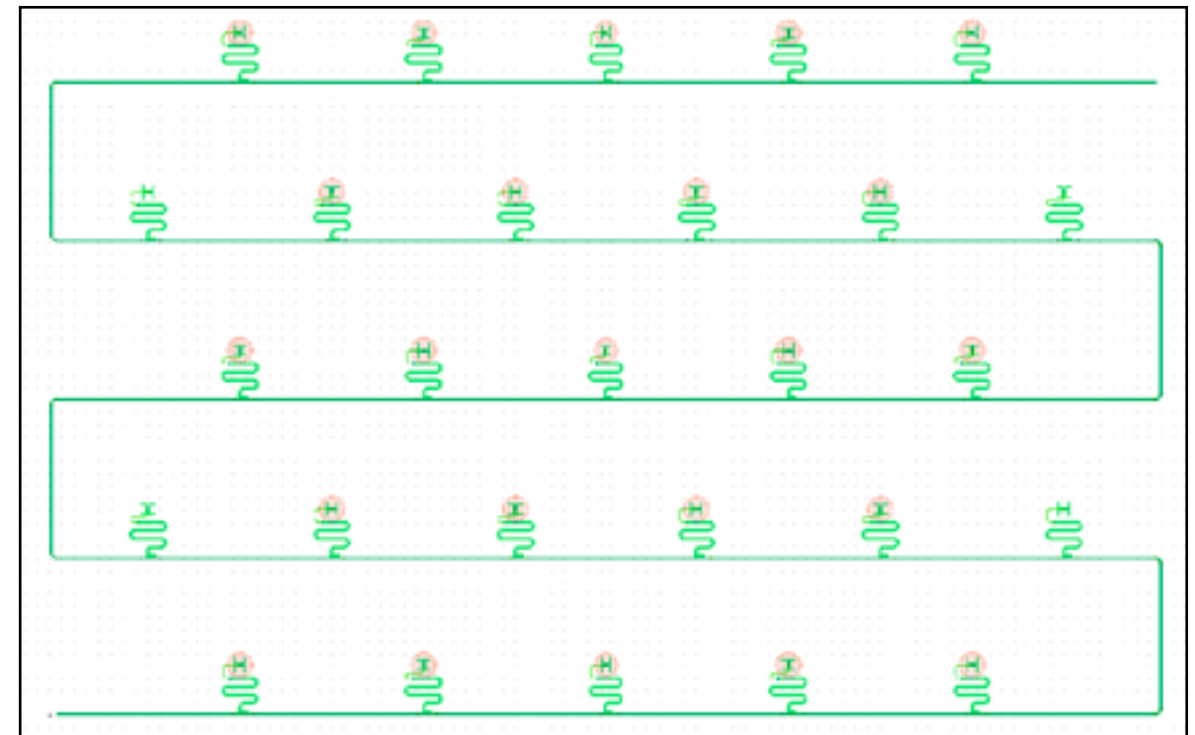
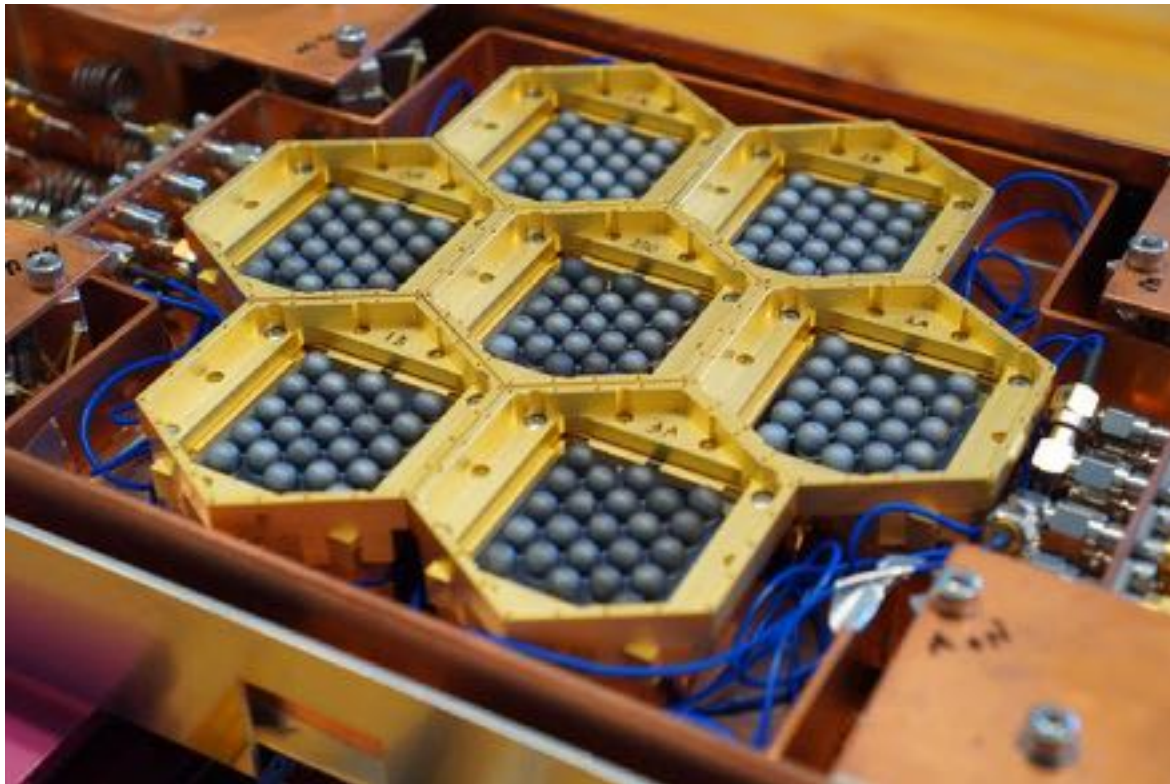
Inductance change

Day, Peter K., et al., Nature 425.6960 (2003): 817-821.

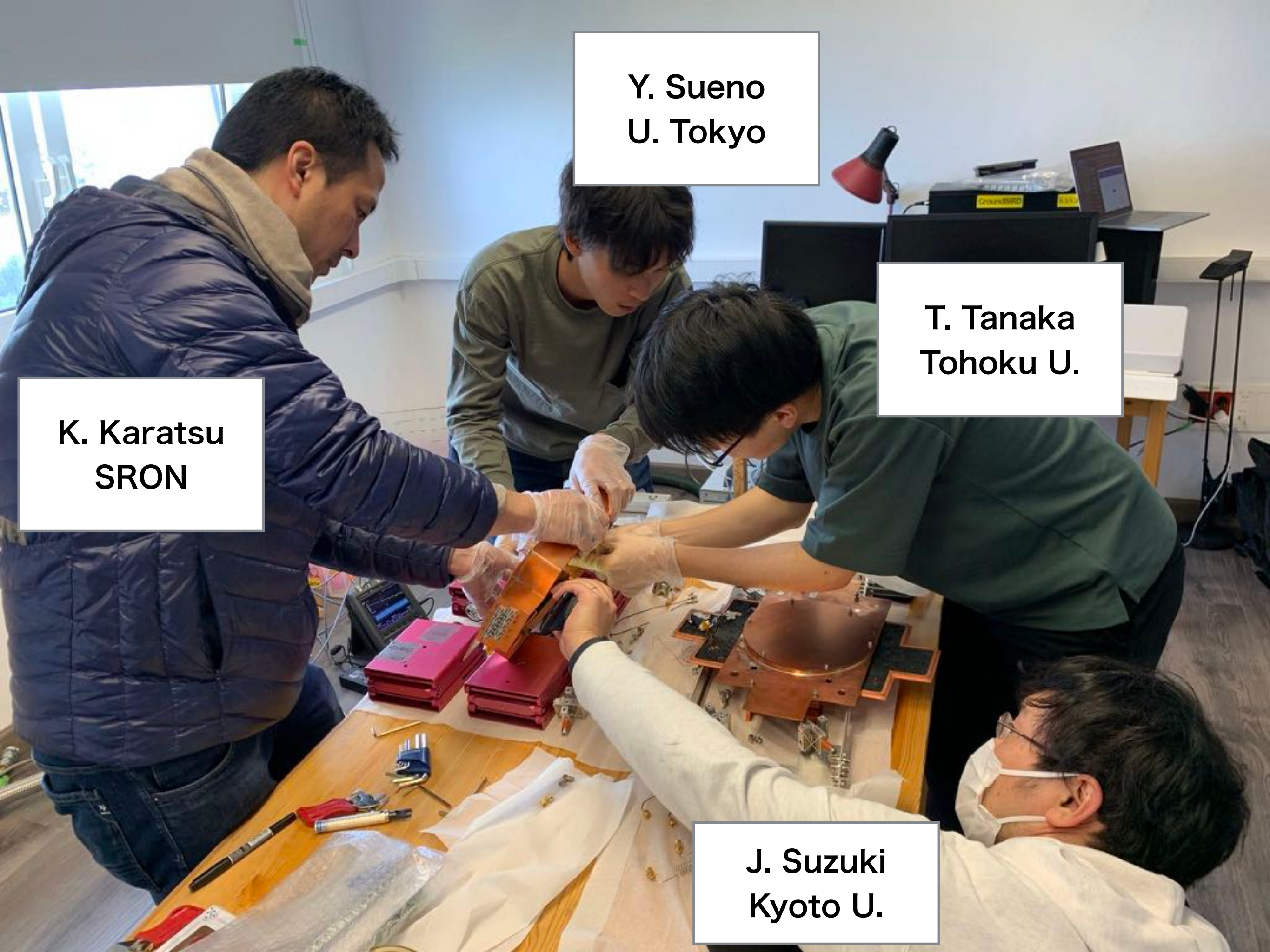
MKIDs for GroundBIRD

Collaboration with
SRON / TUDelft

- 23 pix / wafer
- Al/NbTiN Hybrid



MKID design

A photograph showing four researchers in a laboratory setting. They are gathered around a wooden table, working on a complex electronic assembly. The assembly consists of several red rectangular modules, copper-colored metal plates, and various electronic components. One person is wearing gloves. In the background, there are computer monitors, a desk lamp, and other laboratory equipment.

Y. Sueno
U. Tokyo

T. Tanaka
Tohoku U.

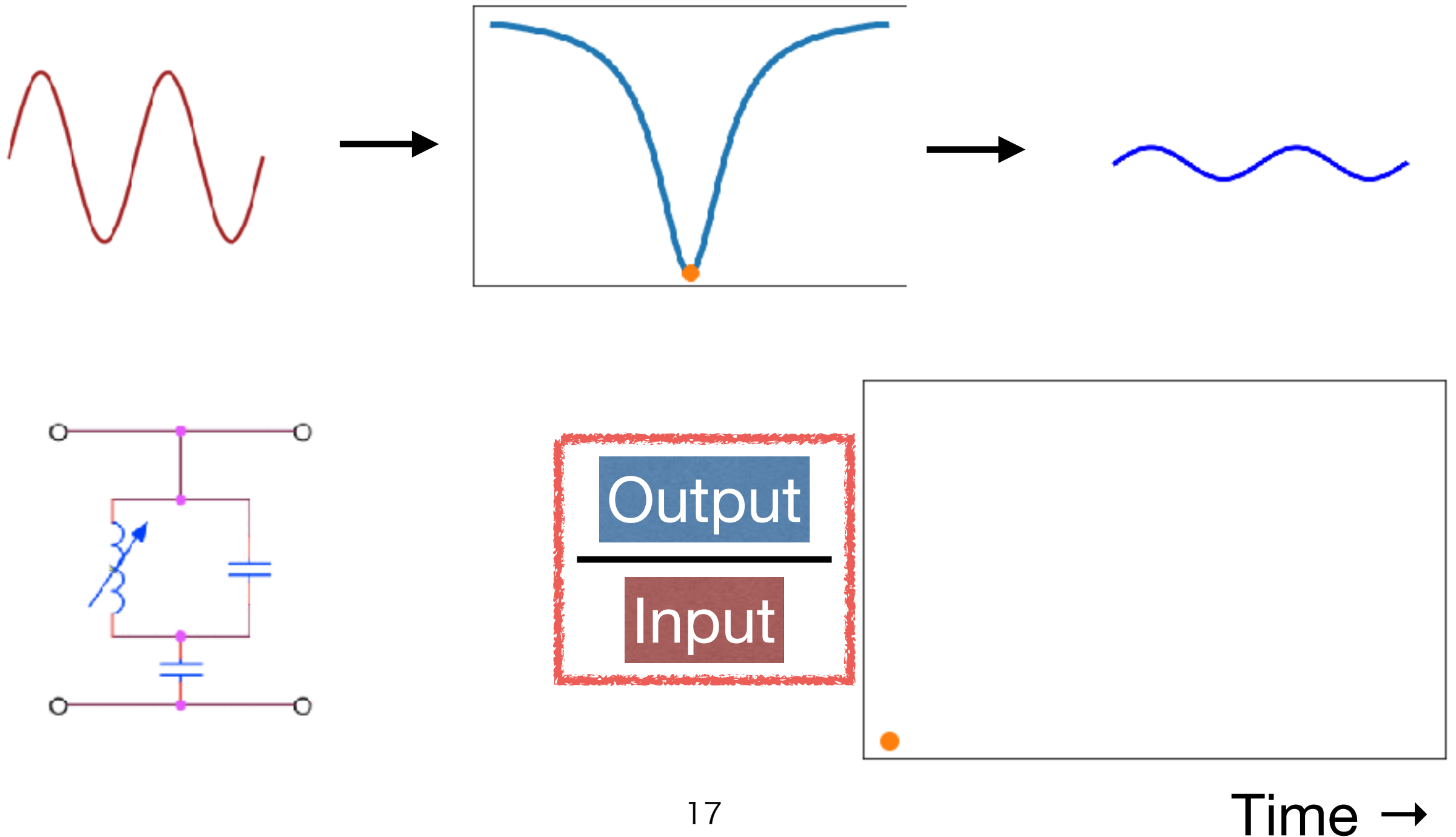
K. Karatsu
SRON

J. Suzuki
Kyoto U.

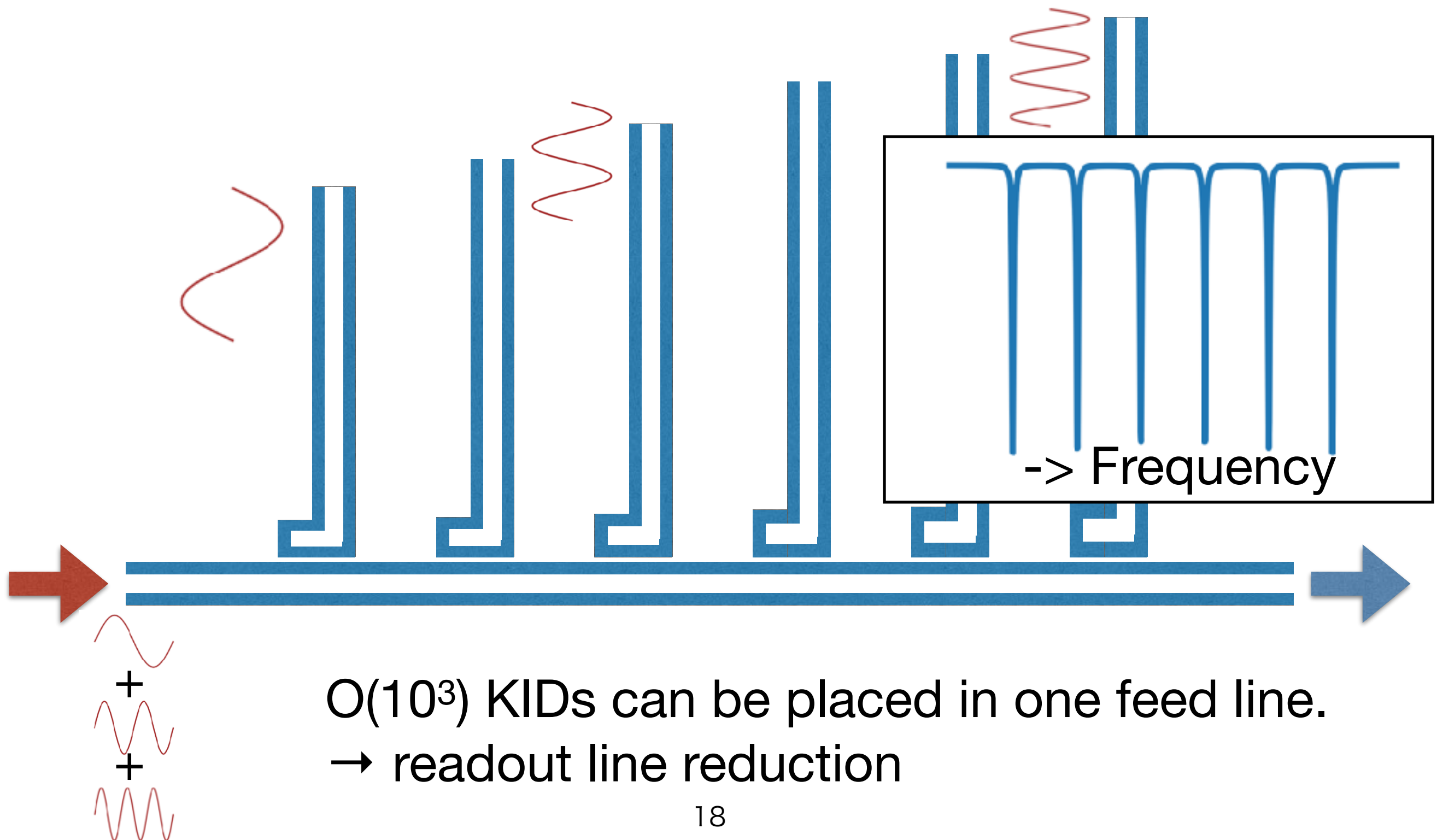
Basics of KIDs readout

Input

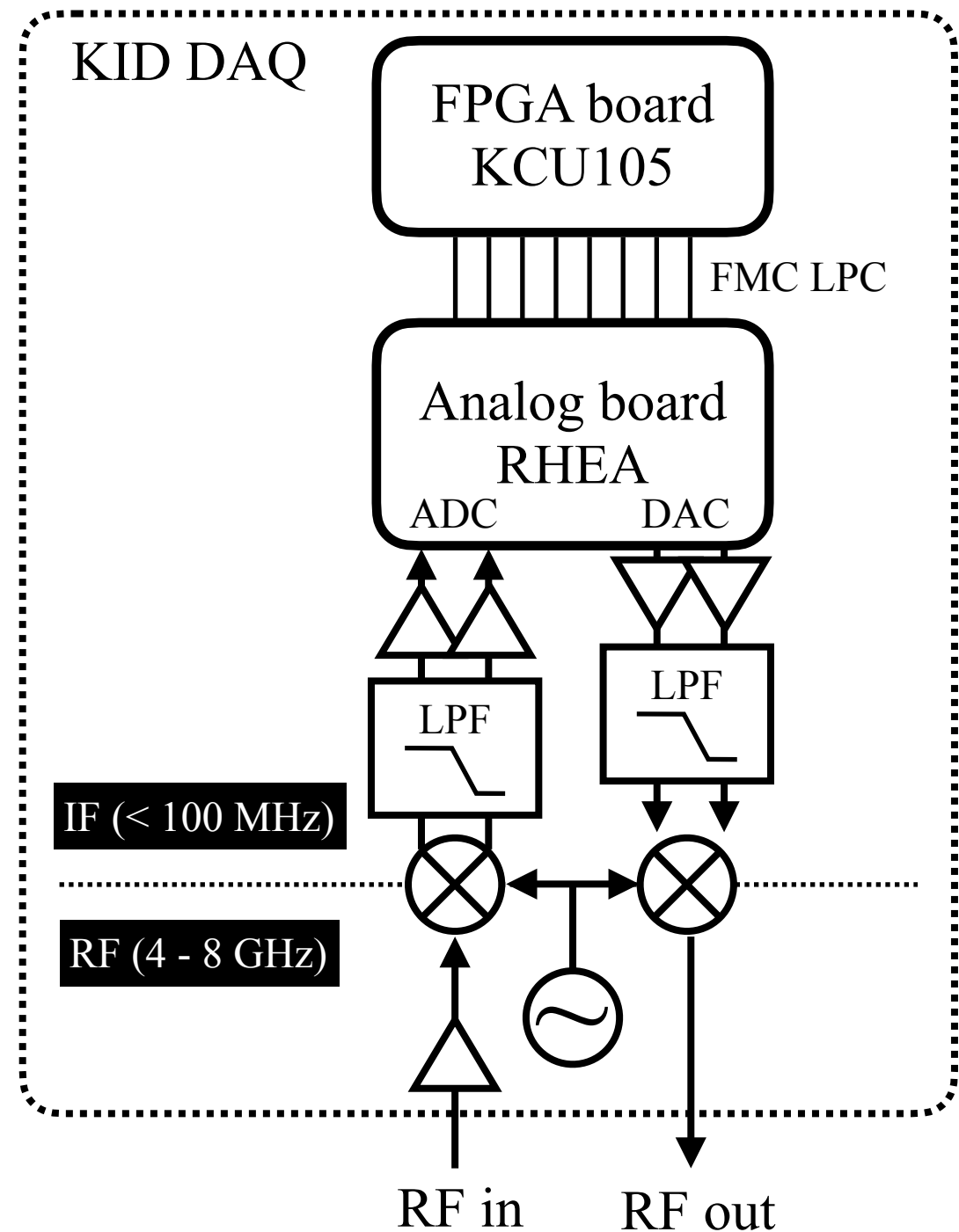
Output



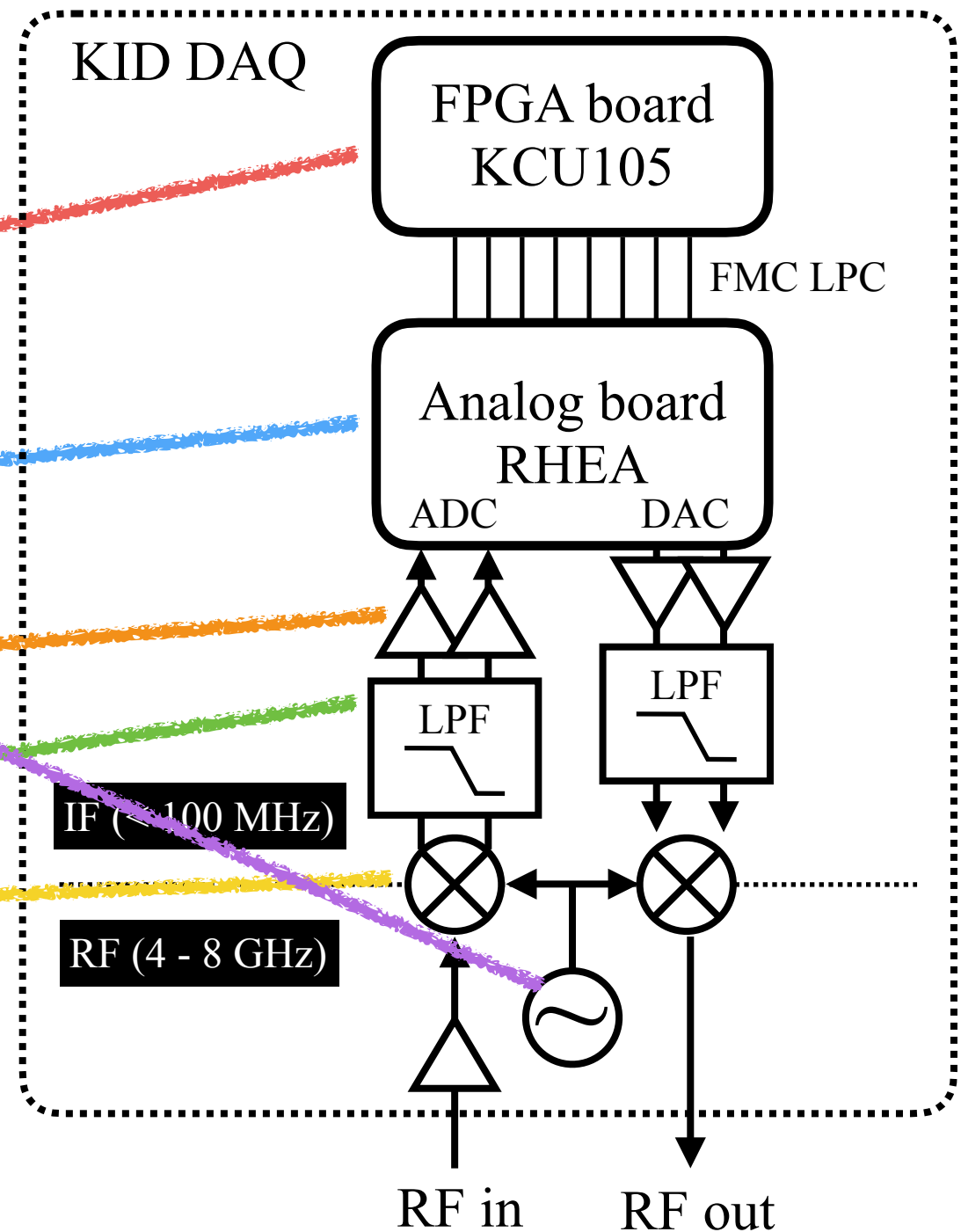
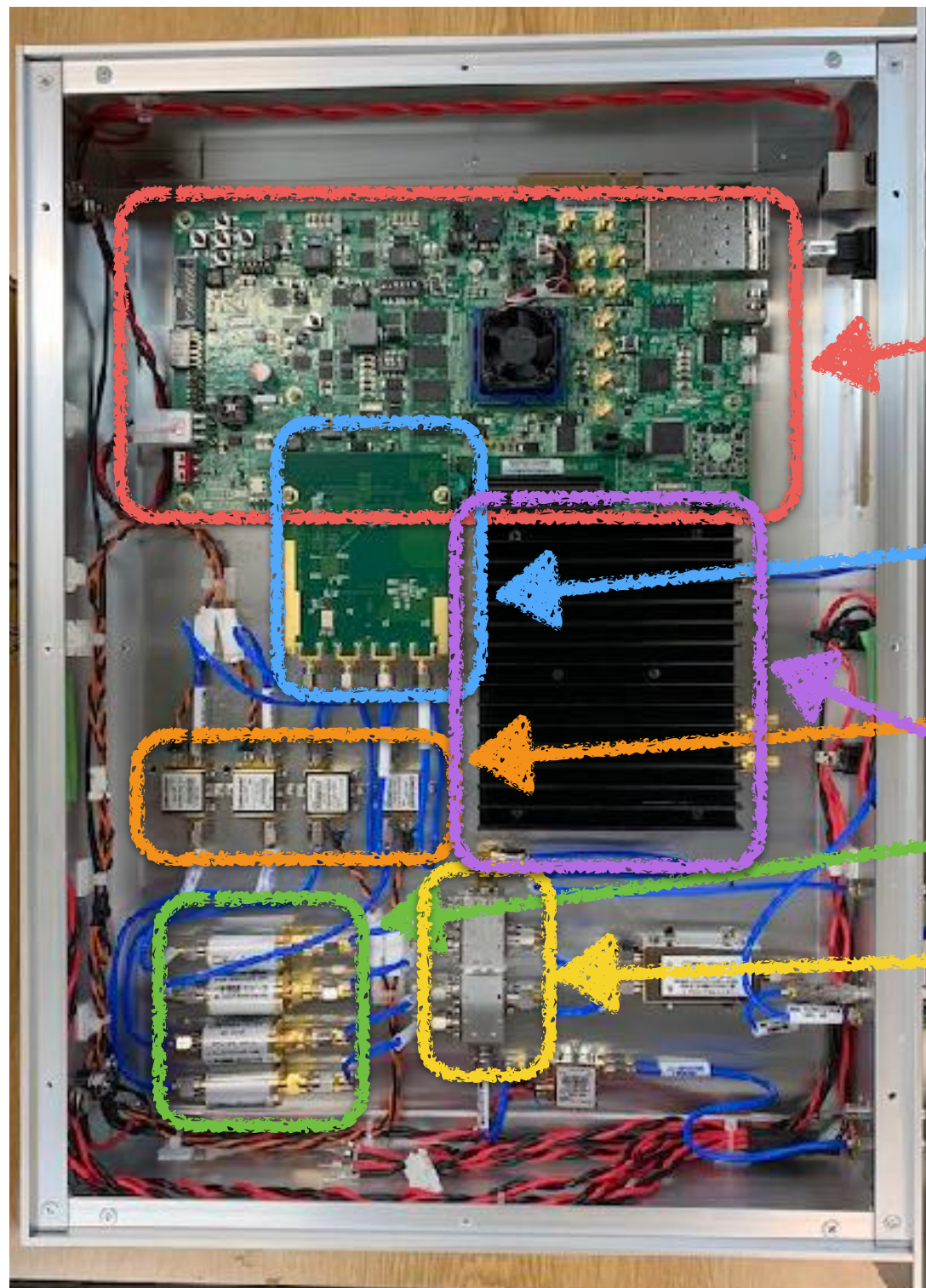
Multiplexing



KIDs readout system



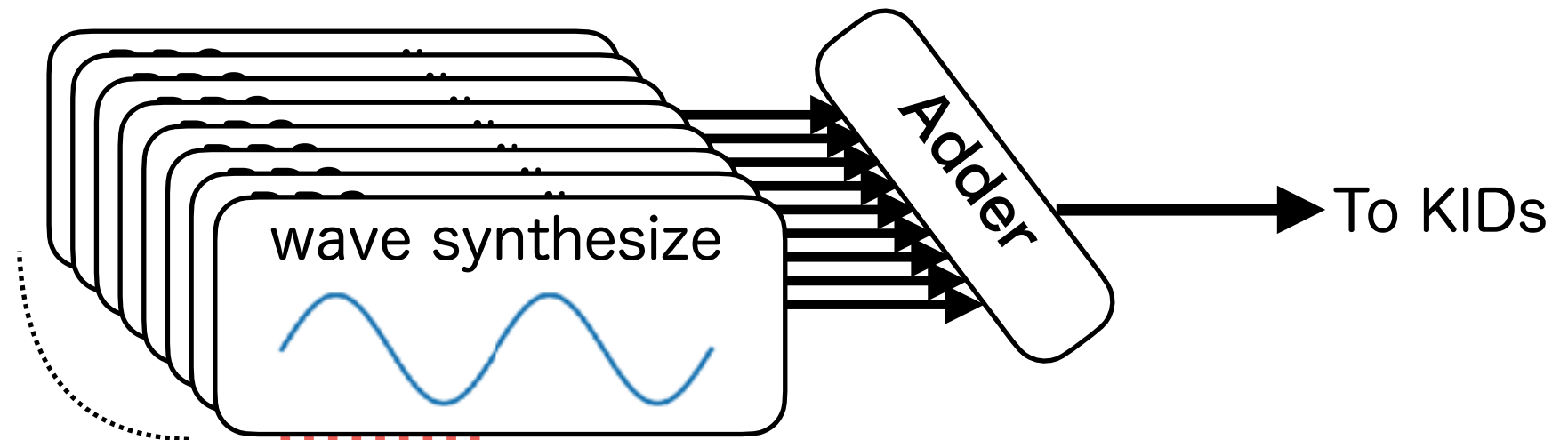
KIDs readout system



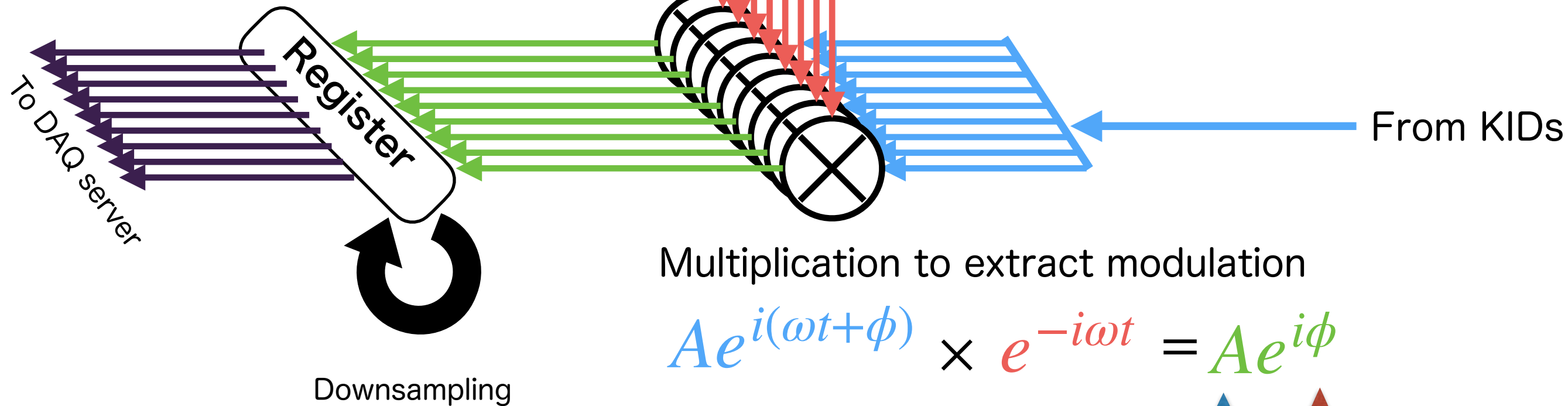
Digital signal processing

DDS (synthesis)

(# of KIDs) $\times$



DDC (decompose)

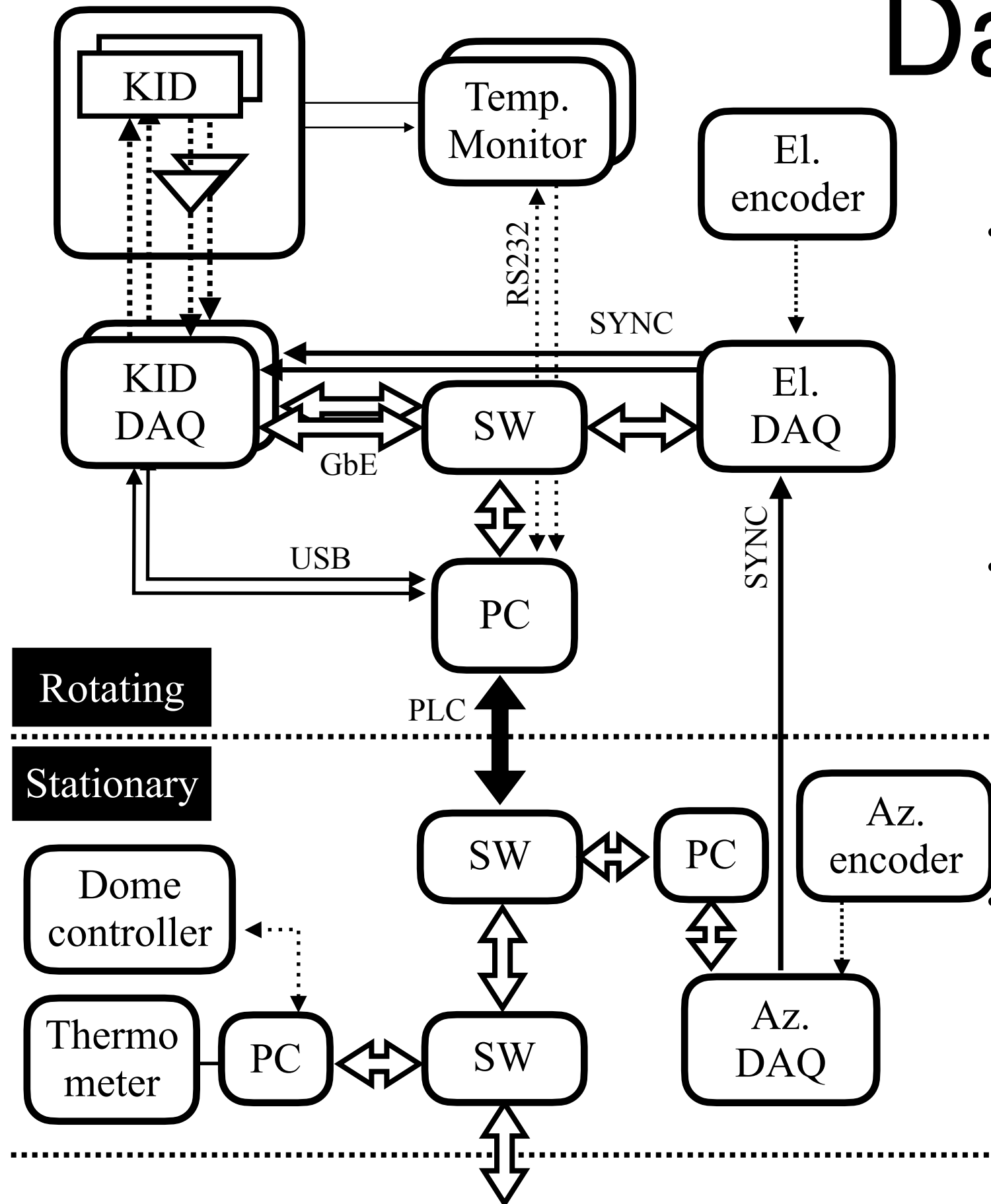


$$Ae^{i(\omega t + \phi)} \times e^{-i\omega t} = Ae^{i\phi}$$

Amplitude

Phase

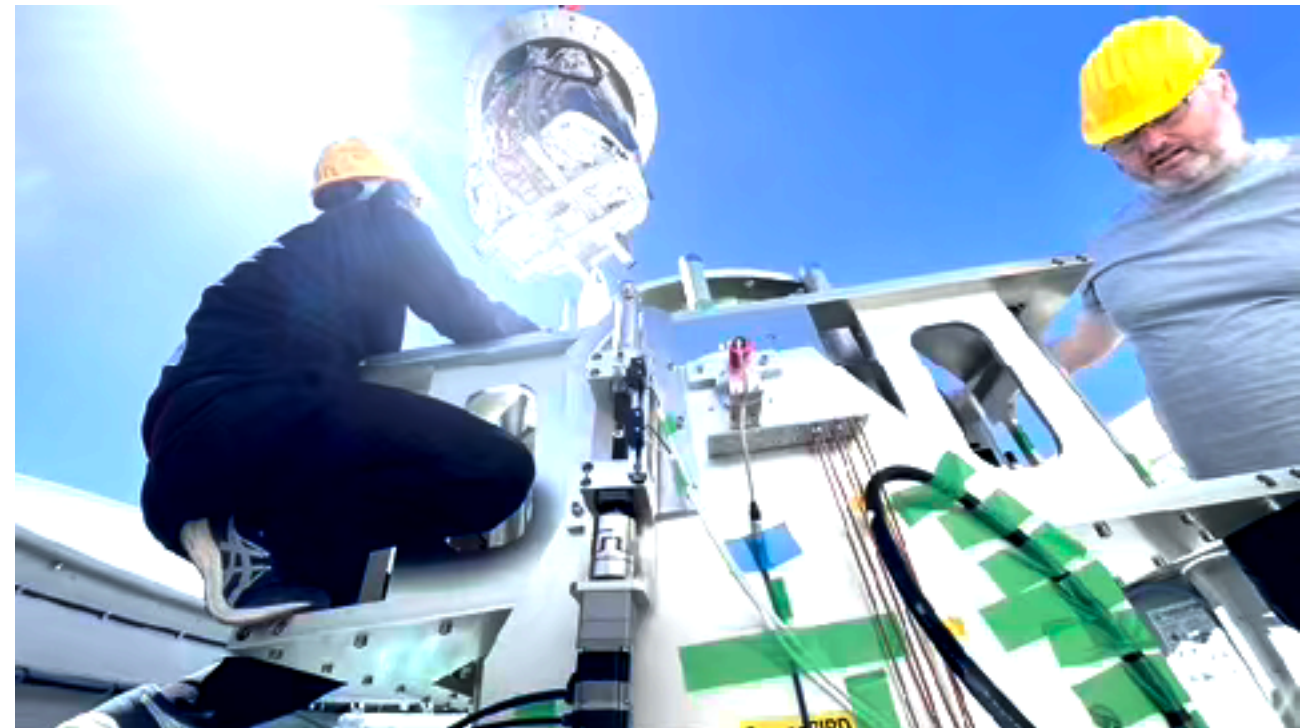
Data acquisition



- Rotating part and stationary part are connected through rotary joint of electricity (AC100V)
- Power Line Connection (PLC) is used for data transfer through the electricity line.
- Data acquisition is synced using a signal through rotary joint as well.

On-site work

- We use crane track to open the chamber
- Thermal shields are attached to the top of the OVC
- We can decompose the telescope by hands after taking off the top flange







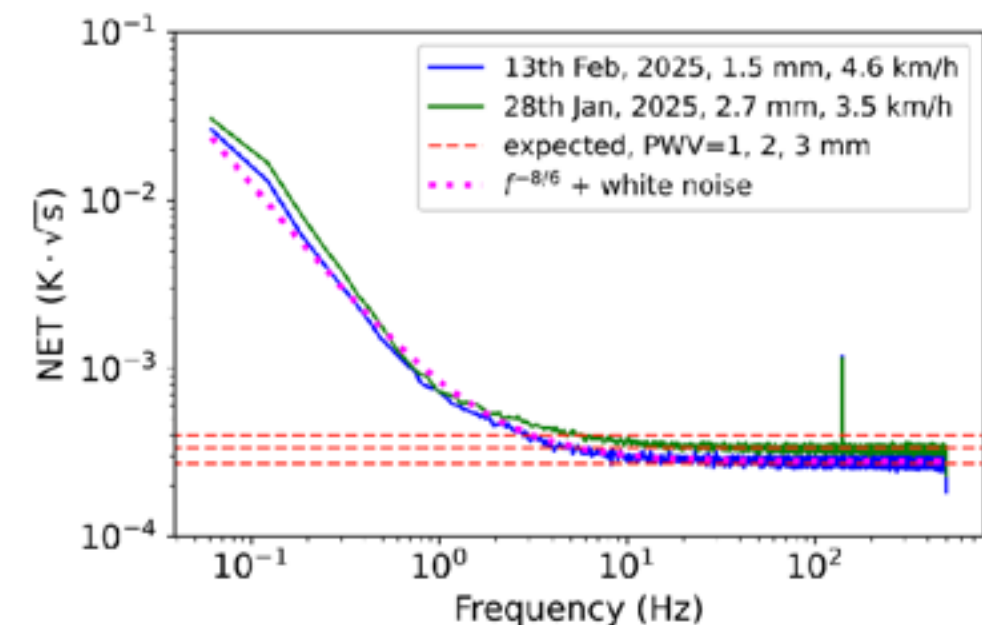
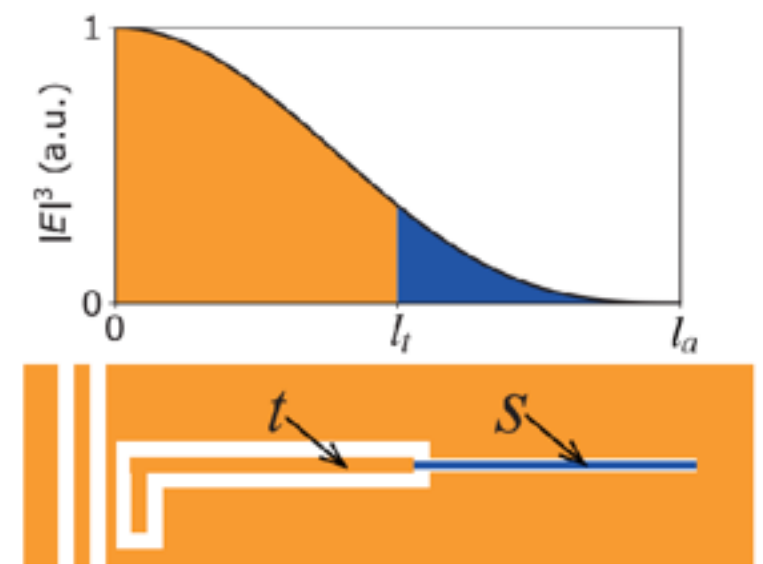
Performance evaluation / analysis

Detector design and performance eval.



T. Tanaka
Tohoku U.

- Detector design was done to reduce the effect of KID's intrinsic noise (TLS)
- The result of the on-site evaluation of the detector will be publish soon.

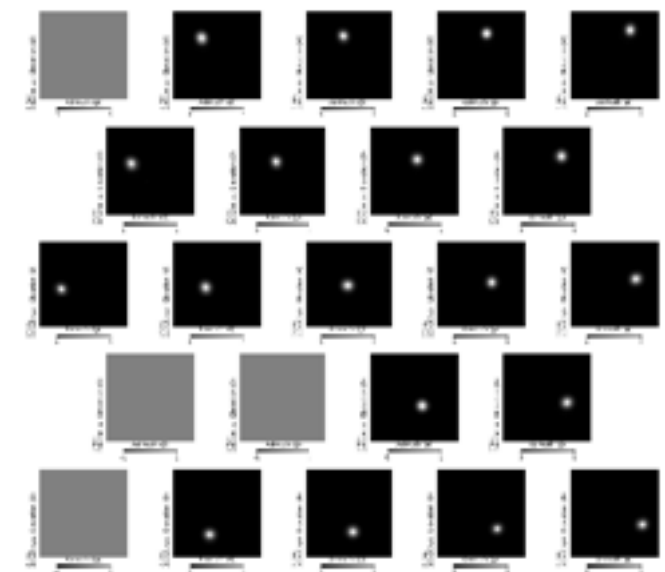
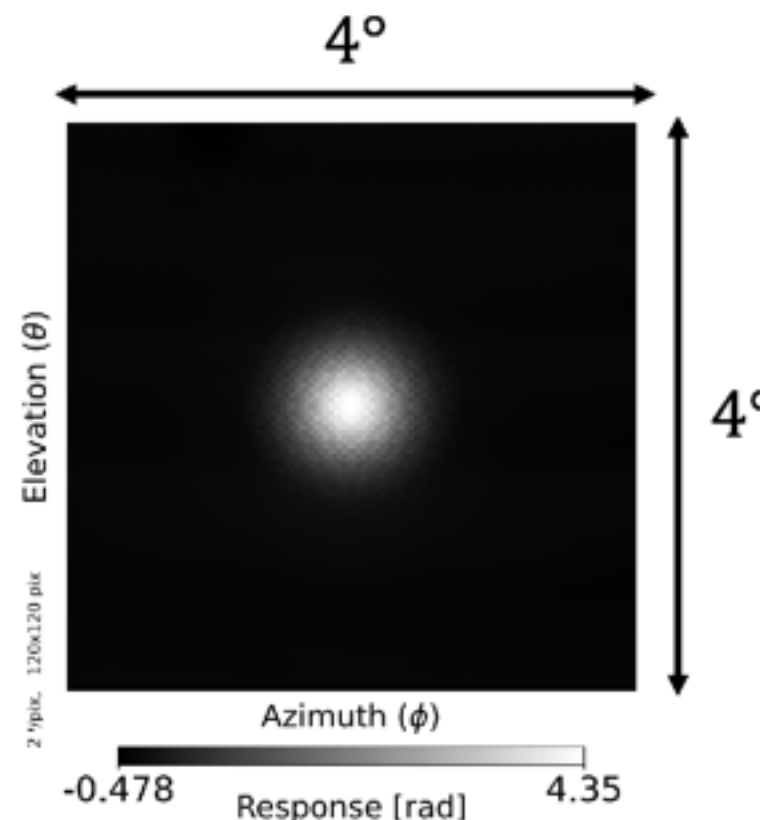
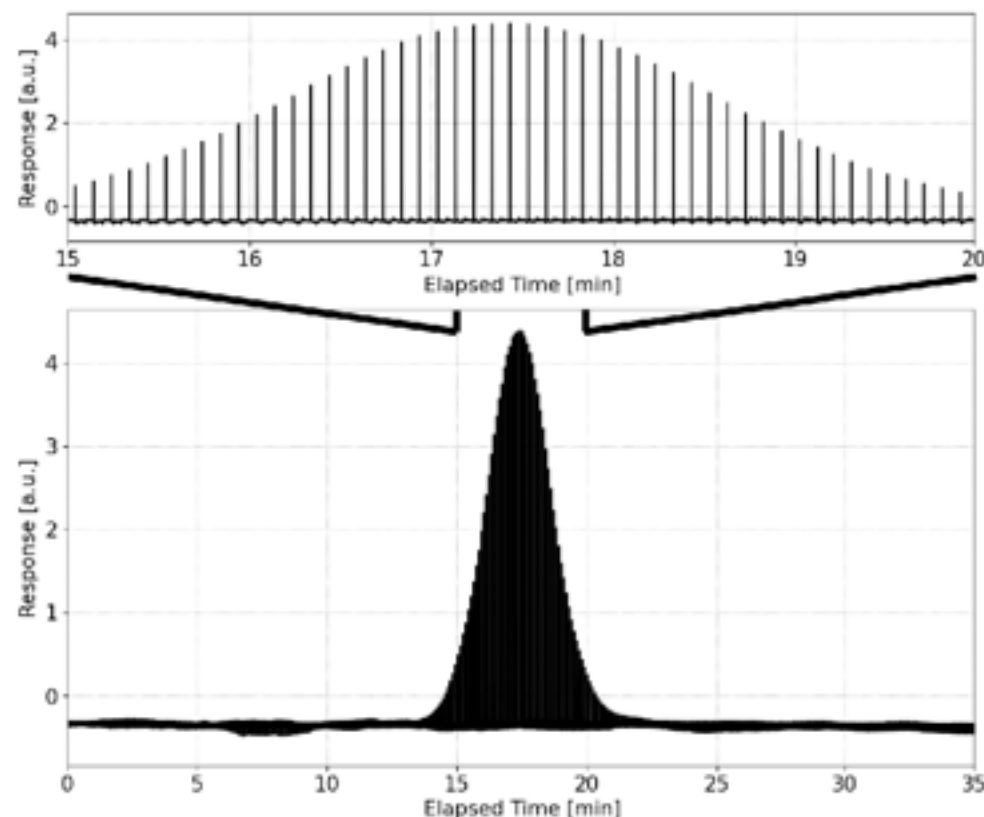


Pointing model from the moon observation

- Moon is used to get the pointing model of the focal plane detectors.
- Reconstruction of the image from rotation scan data $\rightarrow$ position calibration.

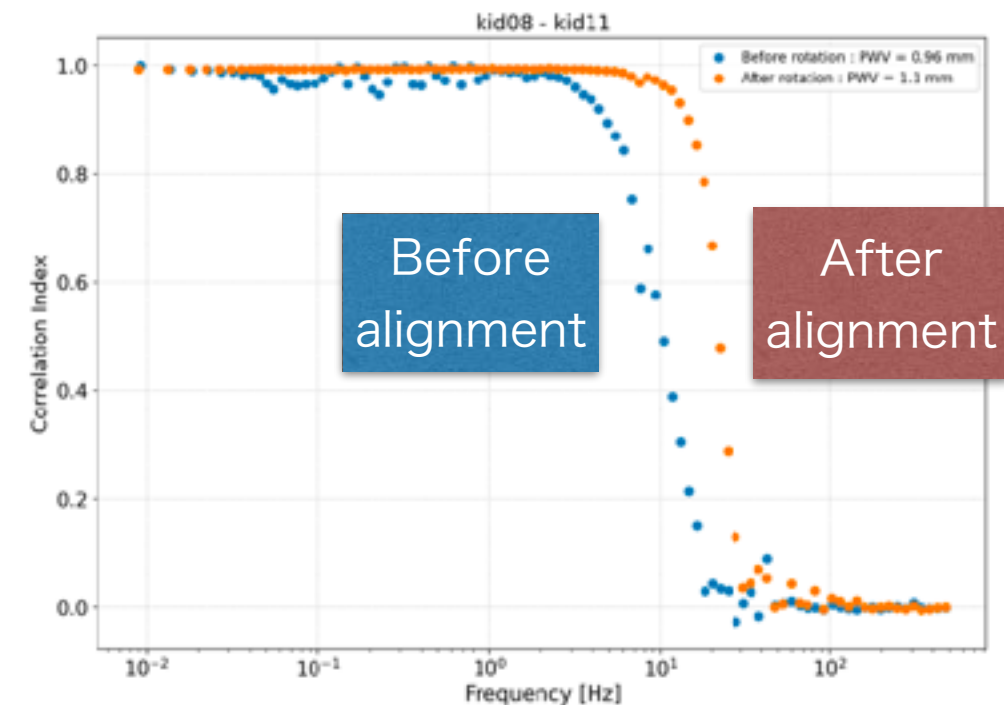
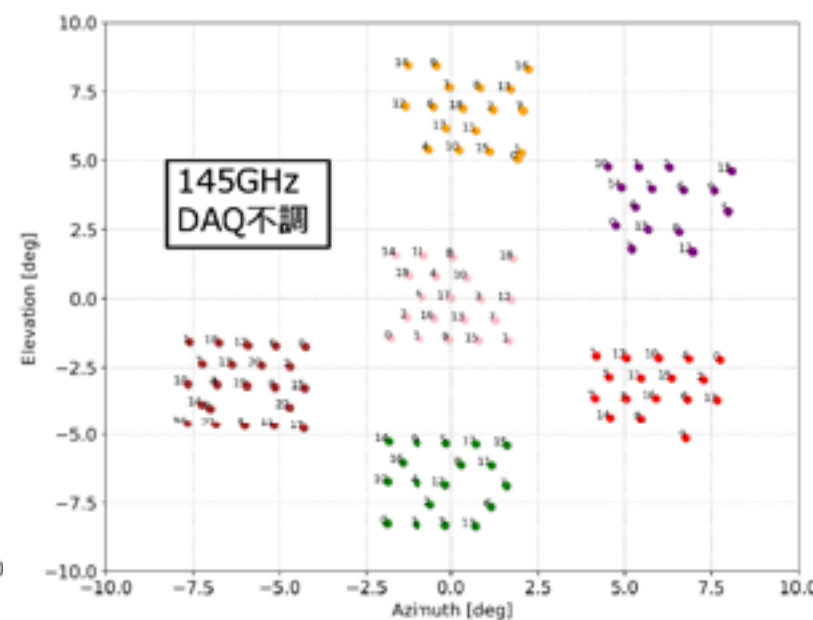
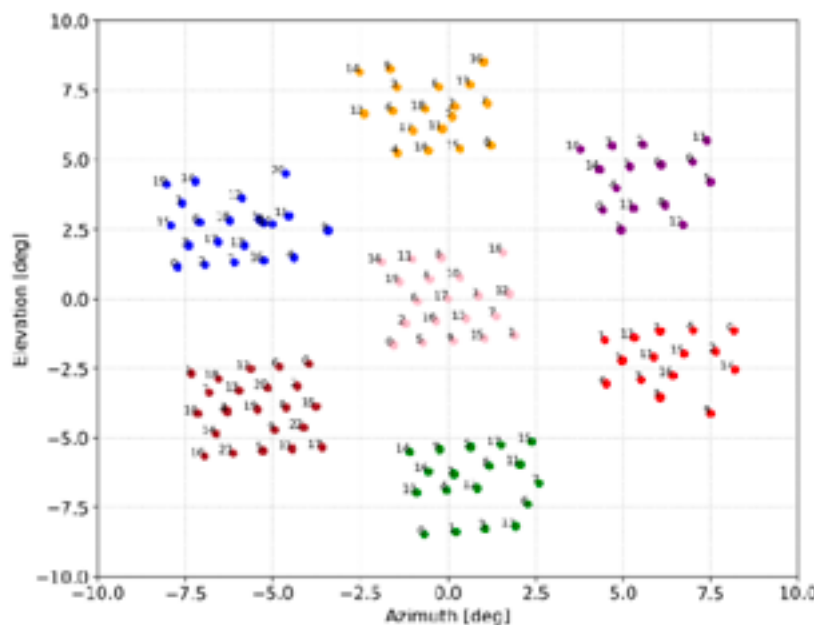
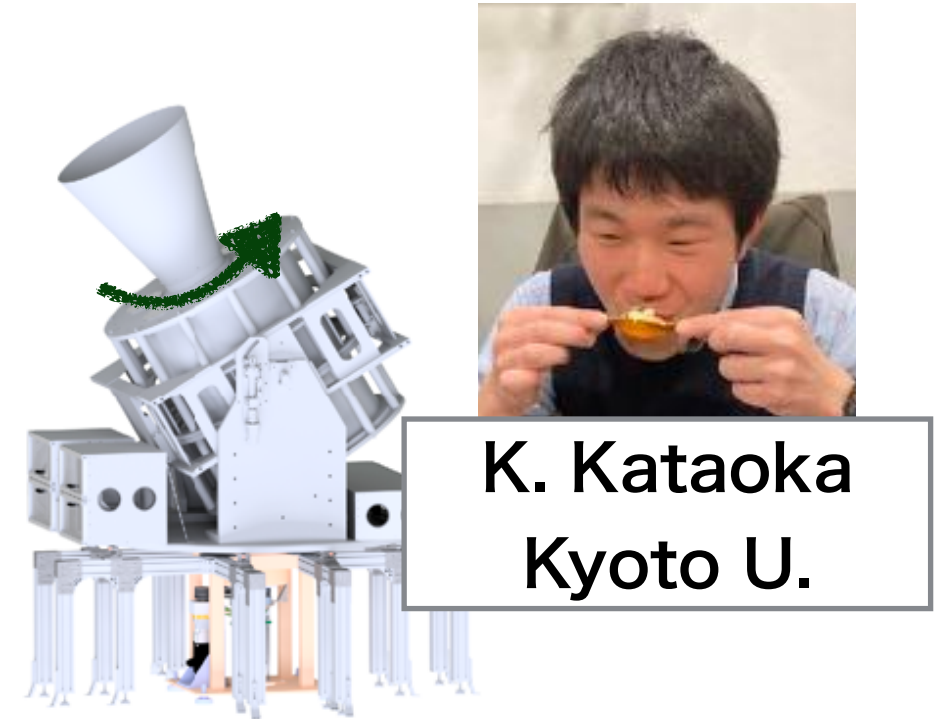


Y. Sueno
U. Tokyo

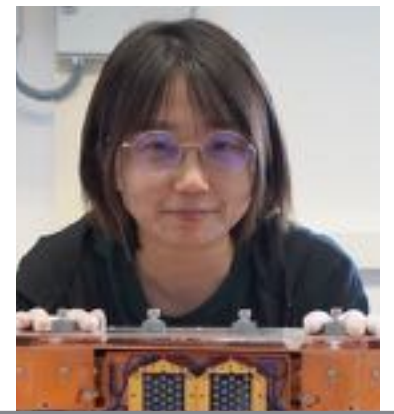


Boresight rotation

- GroundBIRD receiver is hold by outer frame and can be rotated along boresight.
- Pointing calibration was done with several boresight setup to build pointing model and align detectors on sky.
- Alignment effect was also seen on noise correlation between detectors.

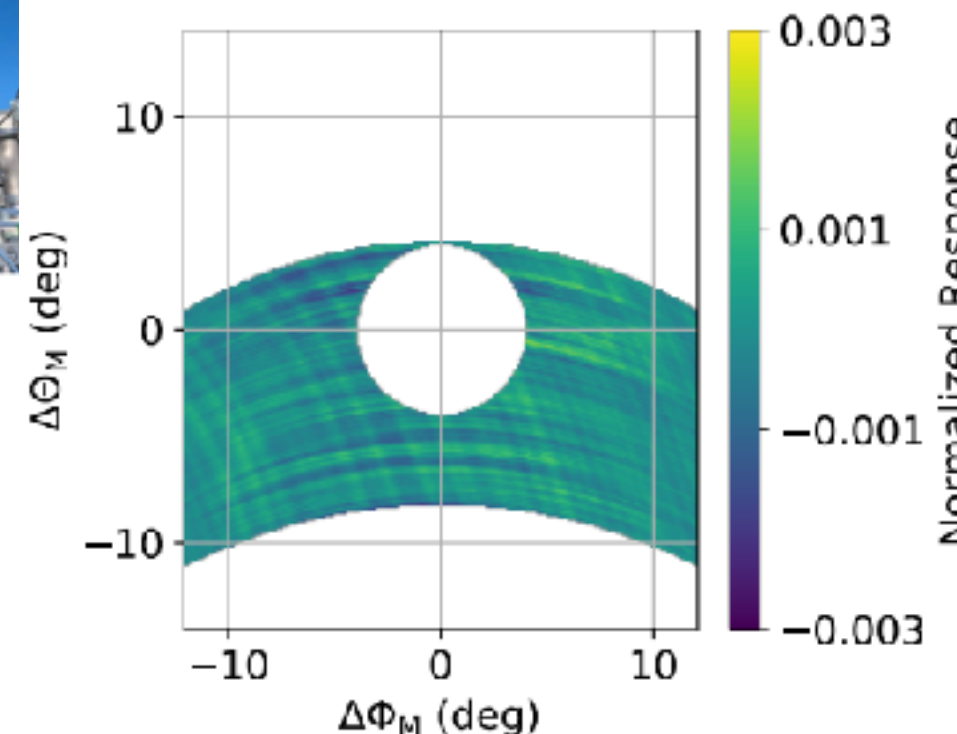
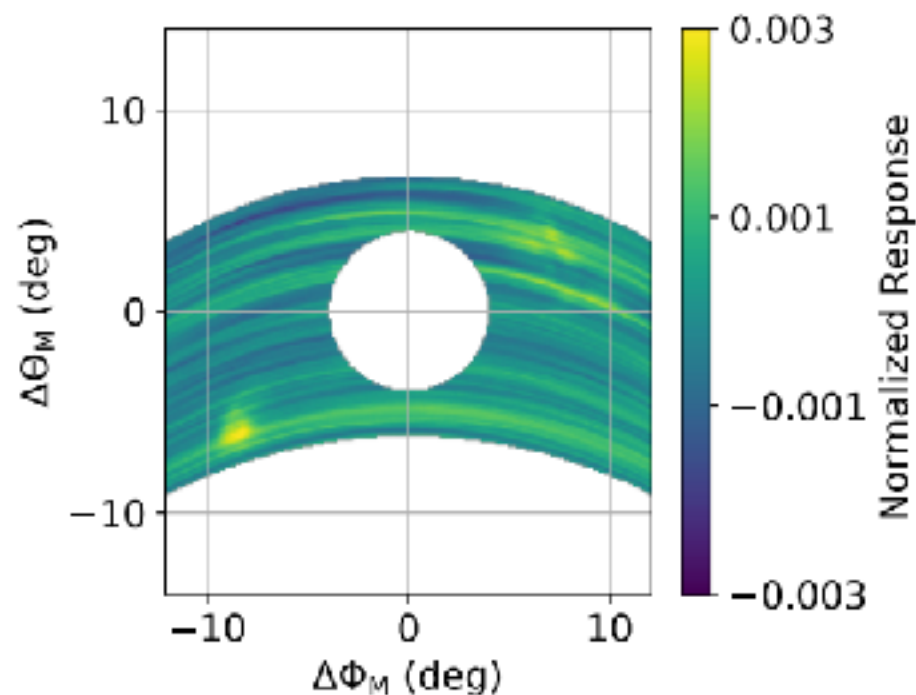


Forebaffle optimization and evaluation

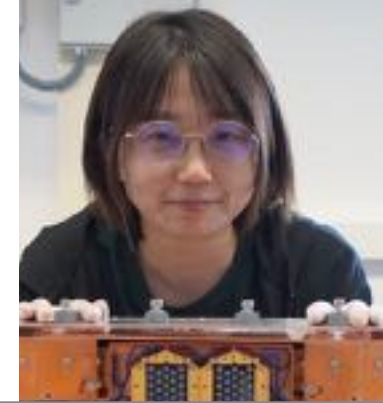


M. Tsujii
Tohoku U.

- Forebaffle was redesigned to mitigate stray light. Detailed simulation was performed to optimize the design.
- Analysis using Moon data showed clear improvement for stray light rejection.

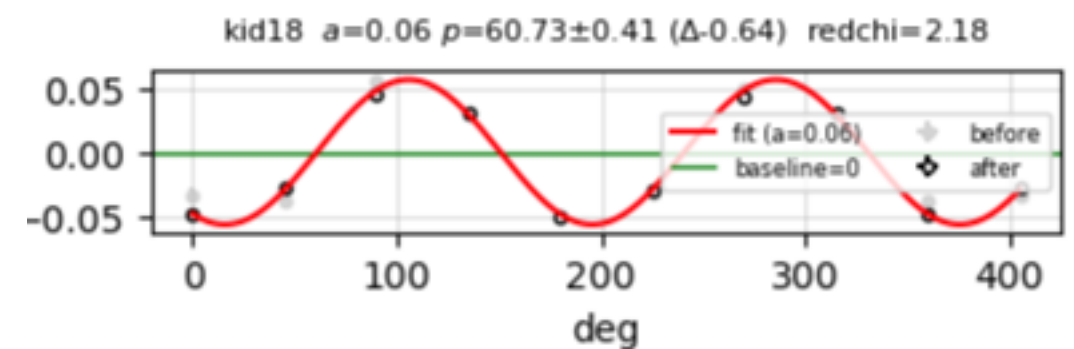
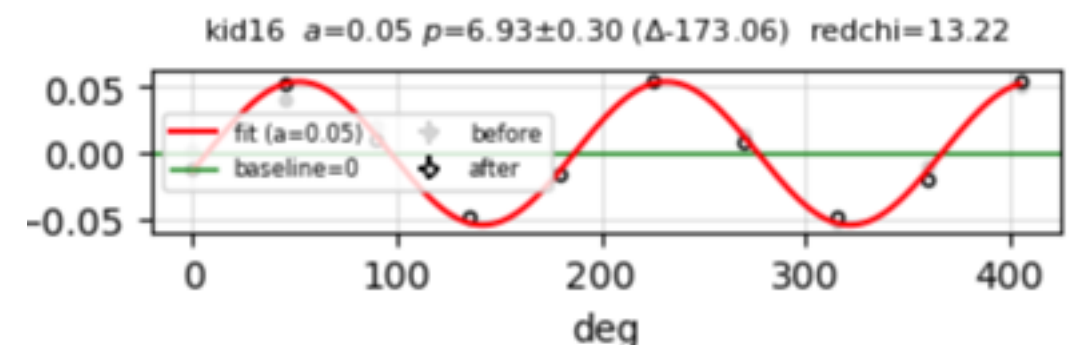
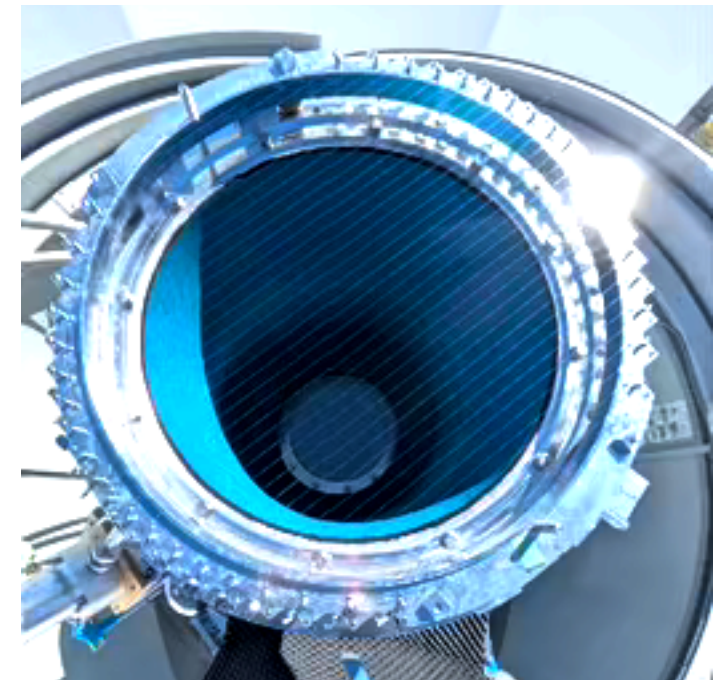


Polarization angle calibration



M. Tsujii
Tohoku U.

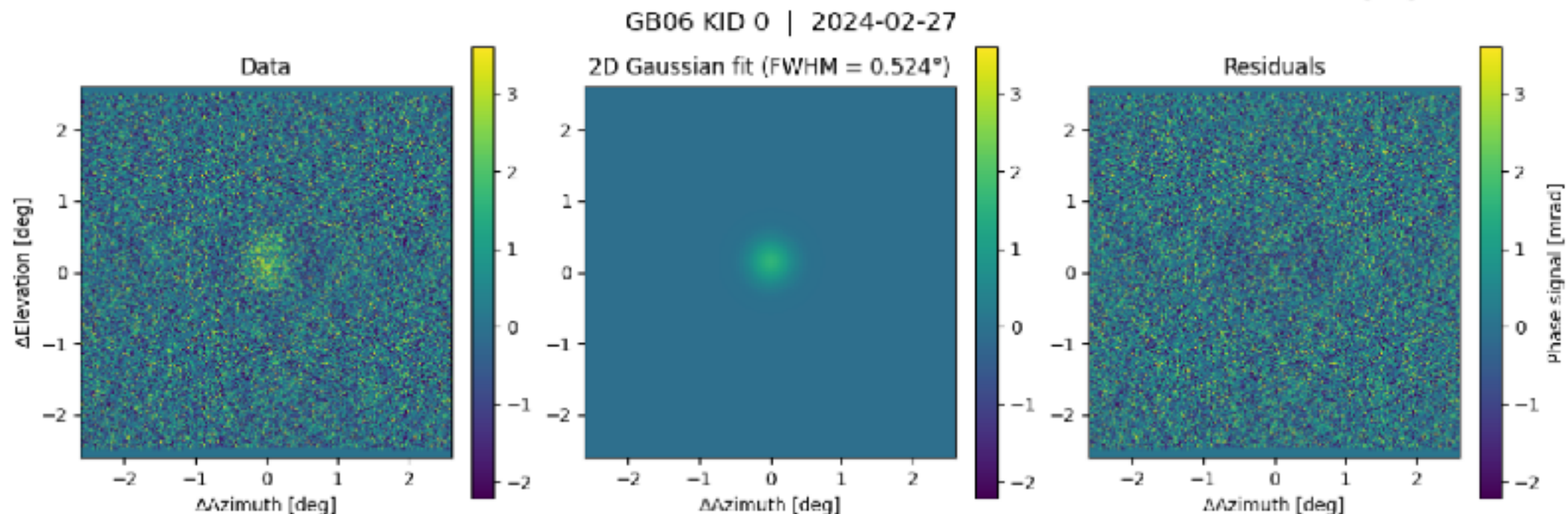
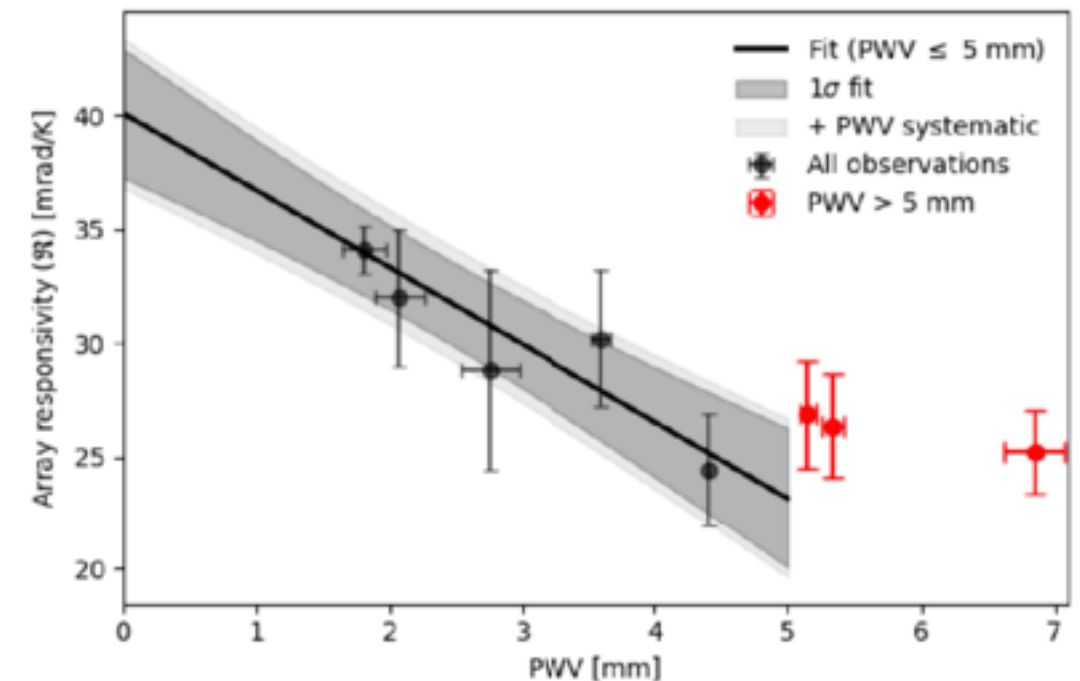
- Polarization angle calibration was performed using a sparse wire grid placed on top of the forebaffle.
- Miku will present about the status of the analysis in the following talk.



Gain calibration using a point source

A. Fasano
IAC

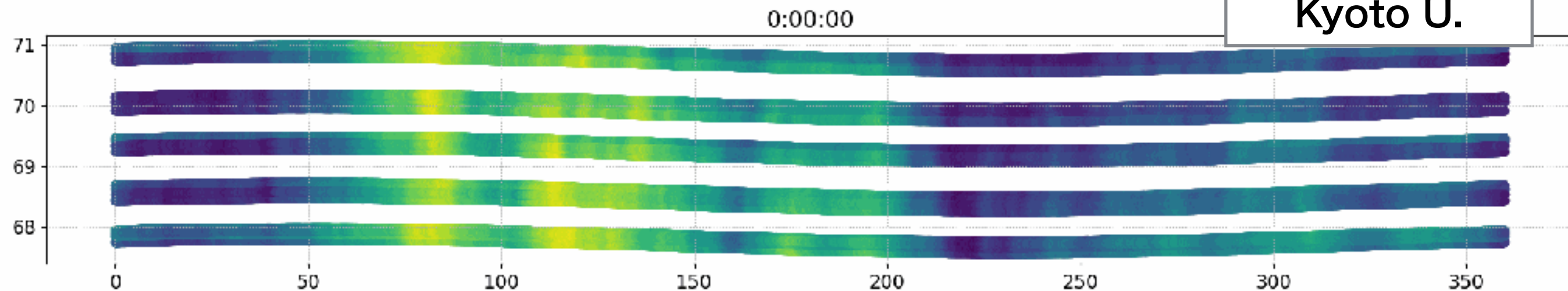
- Gain calibration and the beam stability were under study using Jupiter observation data.
- Result will be presented in the coming SPIE conference.



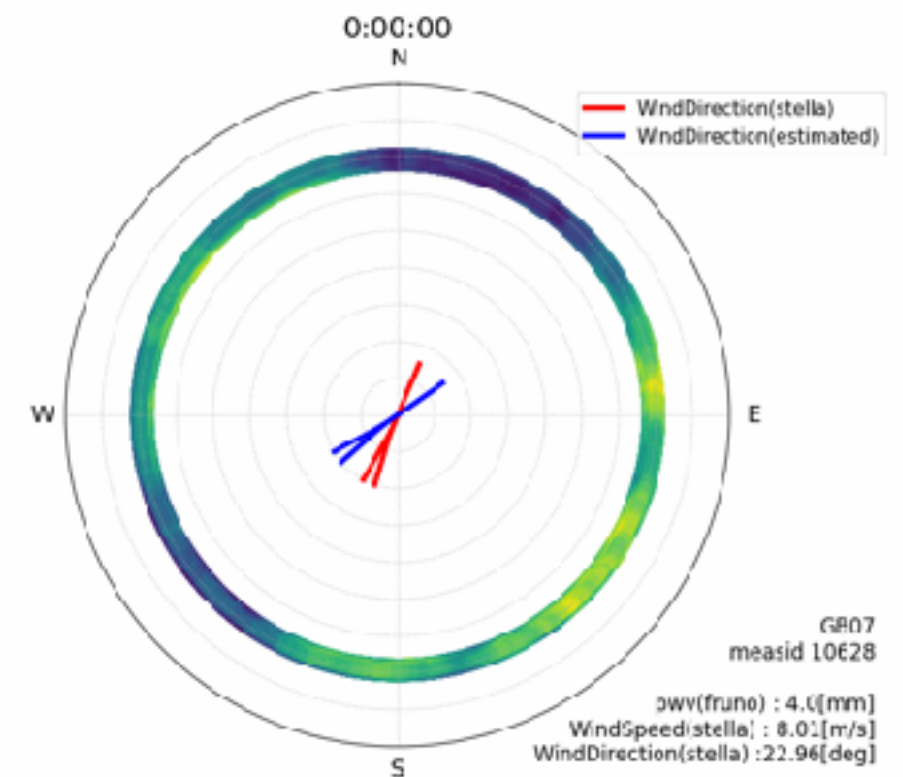
Atmospheric flow analysis



S. Takeichi
Kyoto U.



- We can trace atmospheric flow using fast-rotating feature of the GroundBIRD telescope
- Wind direction estimation roughly agrees with weather station data



Summary

- GroundBIRD is a fast-rotating polarimeter for large scale CMB polarization measurements
 - Primary target: optical depth to reionization τ
- GroundBIRD uses KIDs for its focal plane detectors
 - The result of lab / on-sky evaluation of the detector arrays will be published soon.
- Several instrumental / atmospheric analysis and pipeline construction works are now going on!
 - Pol angle, gain cal w point sources / el-nods, atmospheric characterization

Backup

Telescope	Frequency band [GHz]	Noise level ($\mu K \cdot arcmin$)
QUIJOTE	11	3600
QUIJOTE	13	3600
QUIJOTE	17	5100
QUIJOTE	19	5100
QUIJOTE	30	160
QUIJOTE	40	91
GroundBIRD	145	110
GroundBIRD	220	780