

Latest Results from the LUX-Zeplin Experiment

Björn Penning
University of Zürich

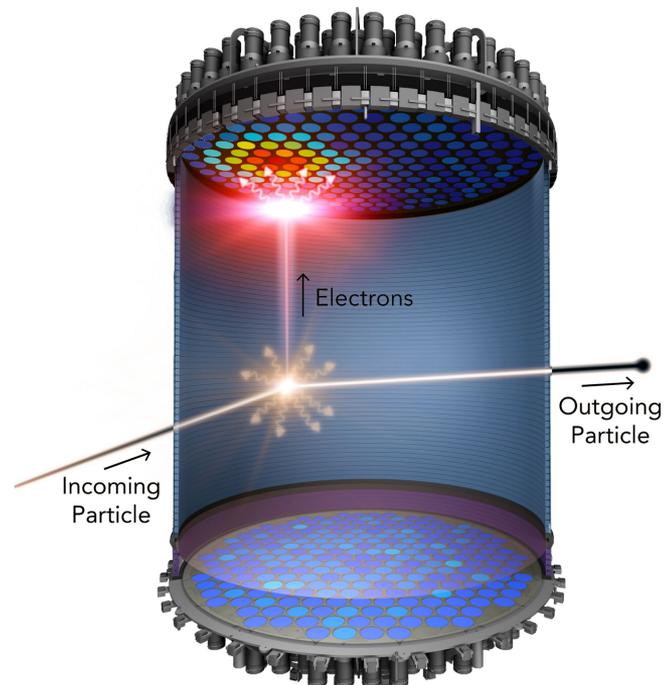
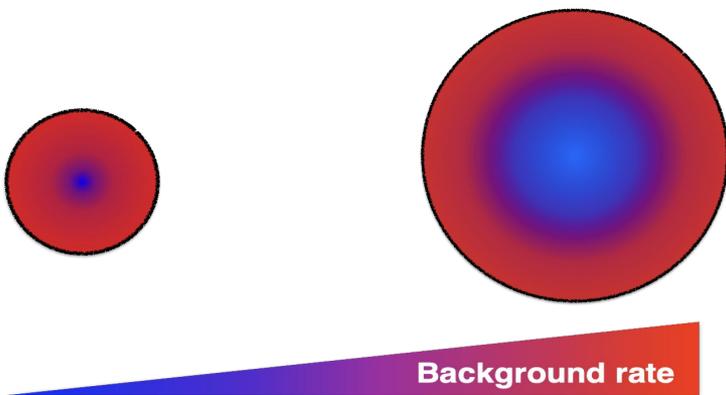
for the LZ Collaboration



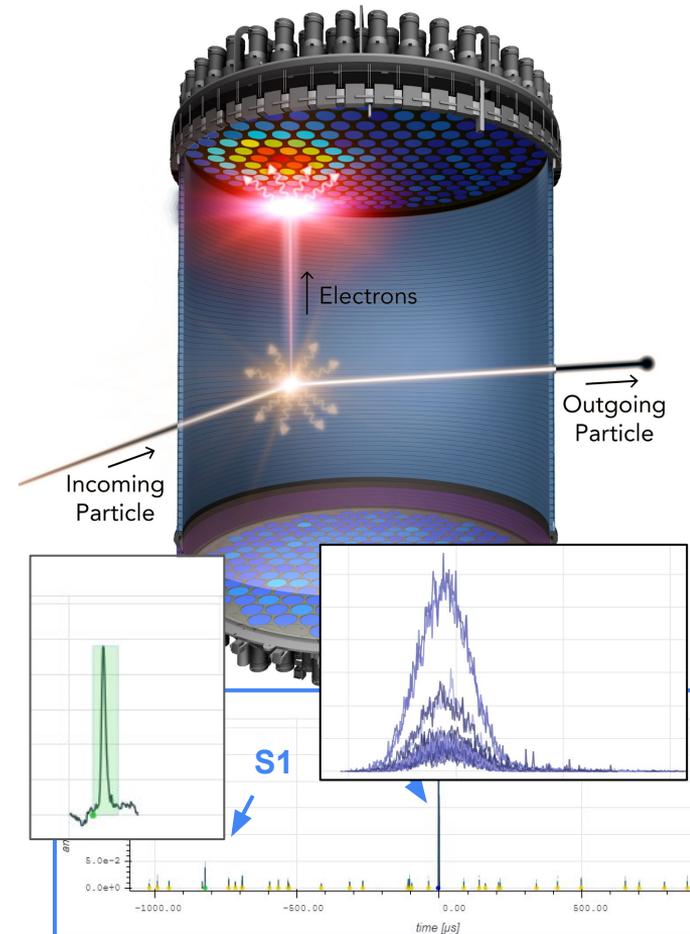
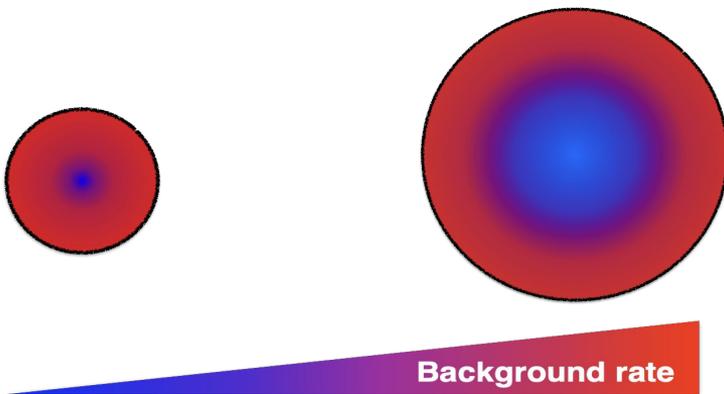
**University of
Zurich**^{UZH}



- **Dual phase TPC**, two signals
 - Prompt scintillation light (**S1**)
 - Delayed proportional charge signal amplified in gas (**S2**)
- **Depth** (z) from time difference between S1/S2 and light pattern provides (x, y) **position**
- LXe is dense and **self-shielding** → **fiducial volume** (FV)



- **Dual phase TPC**, two signals
 - Prompt scintillation light (**S1**)
 - Delayed proportional charge signal amplified in gas (**S2**)
- **Depth** (z) from time difference between S1/S2 and light pattern provides (x, y) **position**
- LXe is dense and **self-shielding** → **fiducial volume** (FV)
- LZ located 1.6 km deep at the Sanford Underground Research Facility (**SURF**)





Water shield:

- 230 t DI water

Outer Detector:

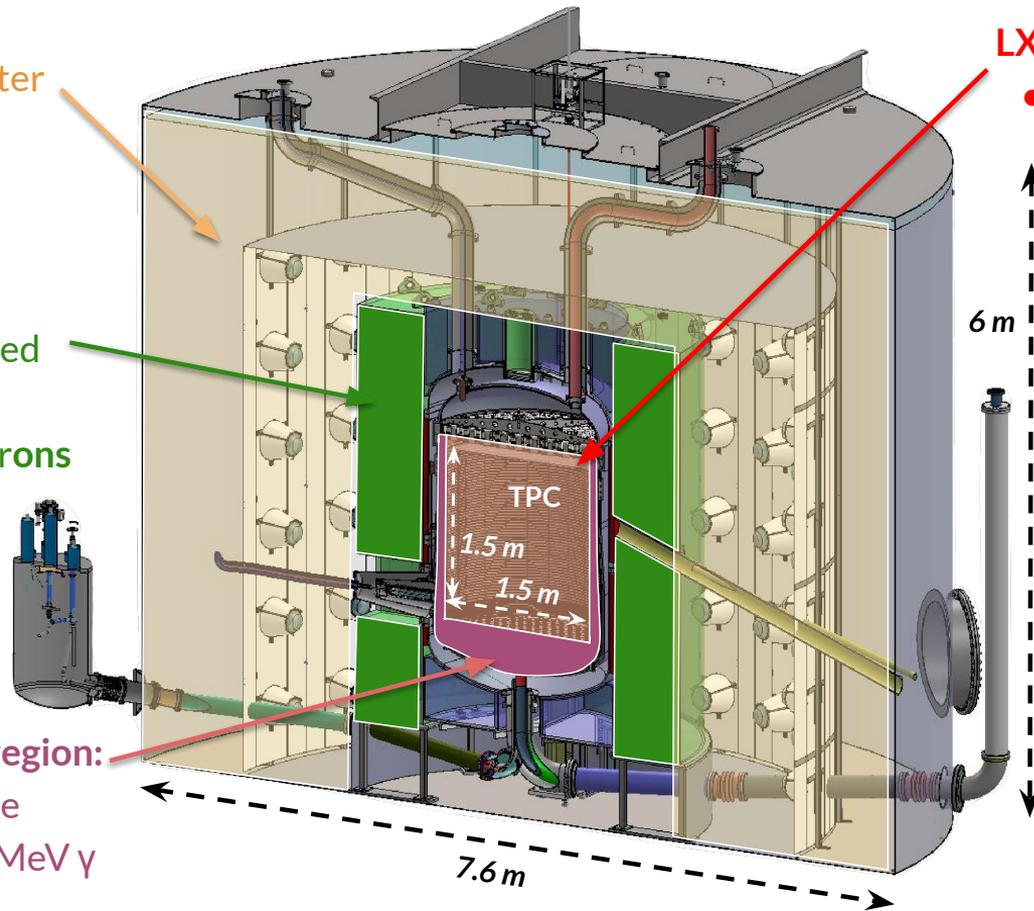
- 17 t Gd loaded scintillator
- Detect neutrons

LXe "skin" region:

- 2 t LXe
- Veto MeV γ

LXe target:

- 7 t LXe
TPC



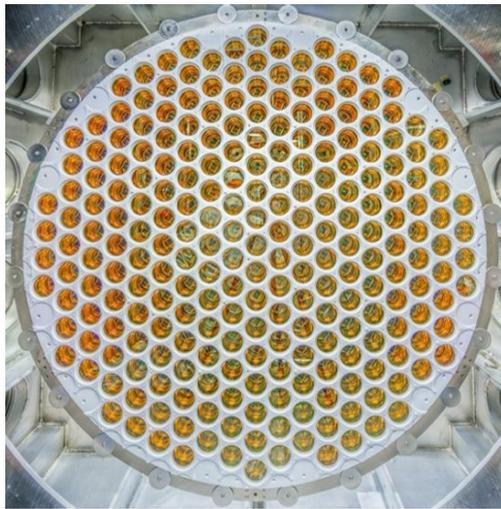


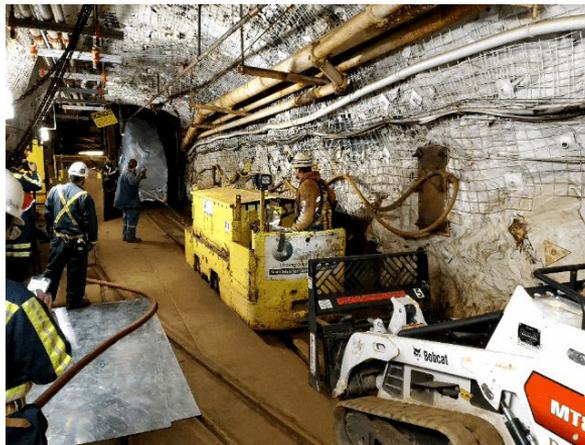
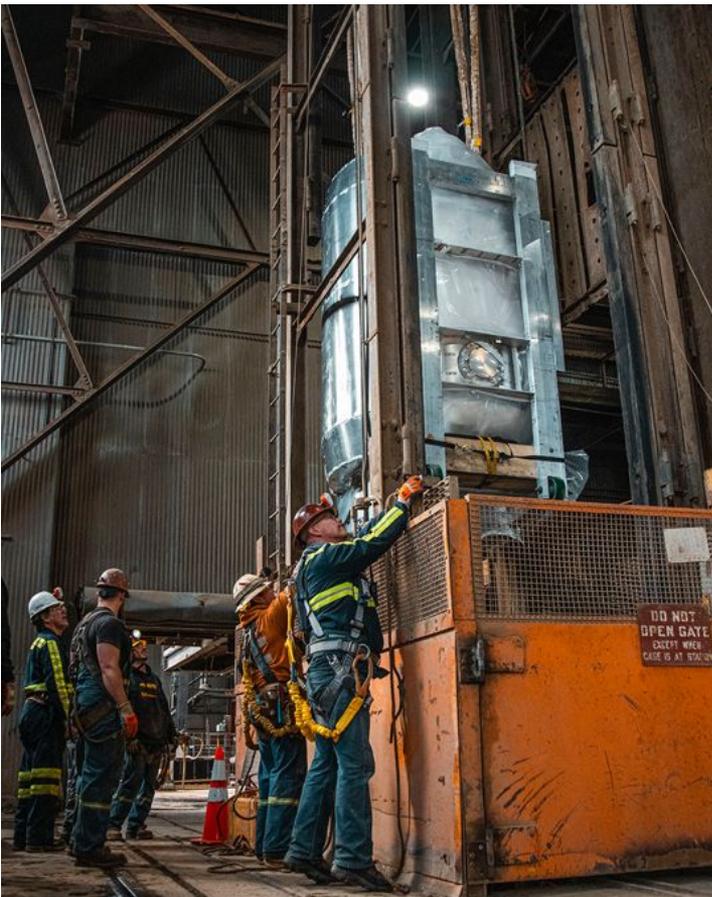
1 Banana = 15 Bq

- Bananas are actually somewhat radioactive due to potassium
 - **15 Bq/Banana** (100 g)
- Our target activity in the Xe: 2 $\mu\text{Bq/kg}$ - **1/75M Bananas**
- Cleaning, cleaning, cleaning, cleaning!

- **Need also to avoid all type of internal contaminants**
 - Use purest materials obtainable, screen all materials
 - Build everything in clean room, reduce dust on surfaces to **O(ng/cm^2)**
 - Keep circulating and purifying target material: aim xenon contaminants to **O(0.015 ppt)**

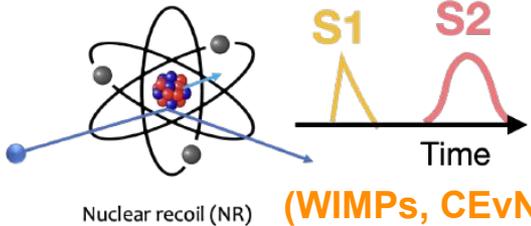
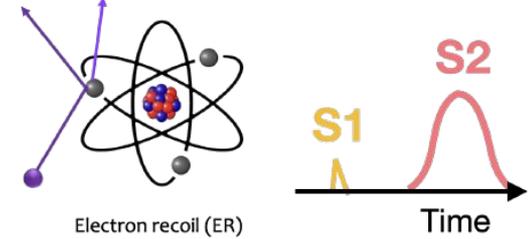




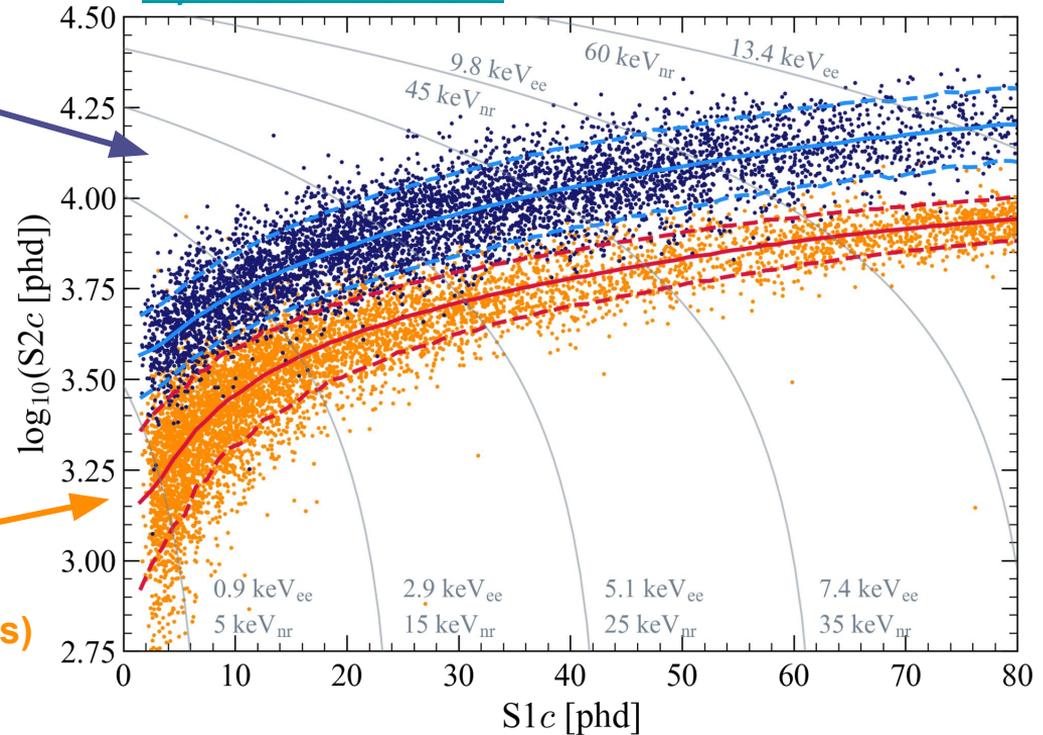


[Phys. Rev. Lett. 131, 041002](#)

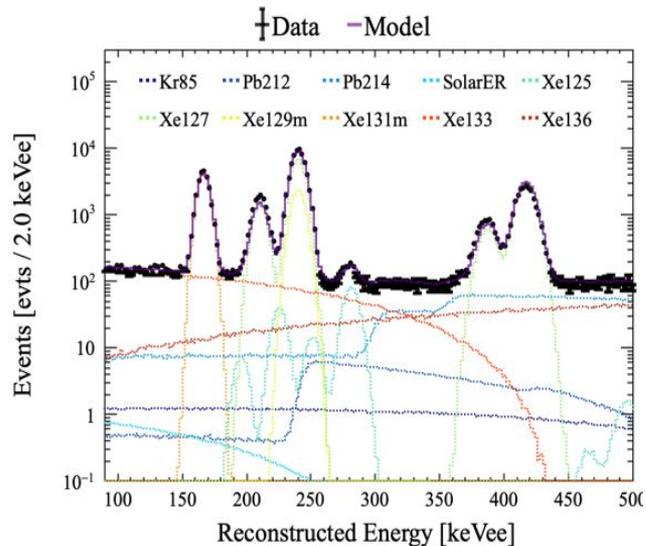
(β decays, γ rays, ν -e $^-$)



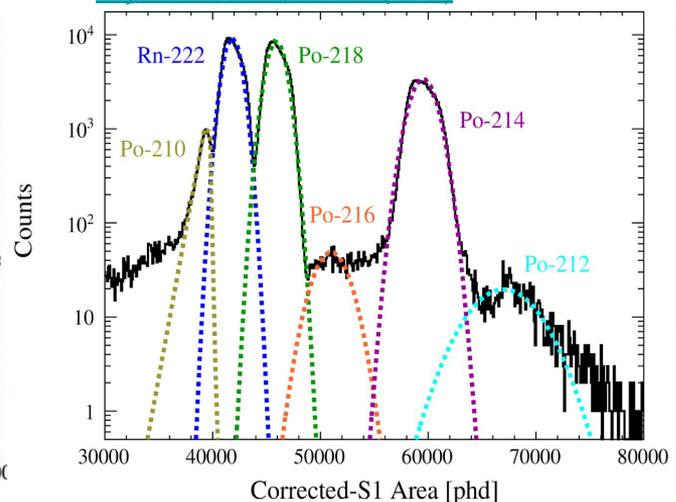
(WIMPs, CEvNS, neutrons)



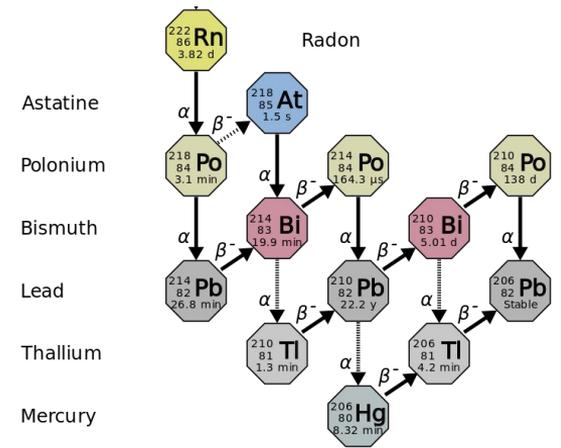
- Backgrounds predominantly electron recoils (ER), WIMPs produce nuclear recoils (NR)
 - **ER band:** Tritiated methane (CH₃T) injection, spatially homogeneous β source
 - **NR band:** DD neutron generator (NR band), Monoenergetic 2.45 MeV neutrons

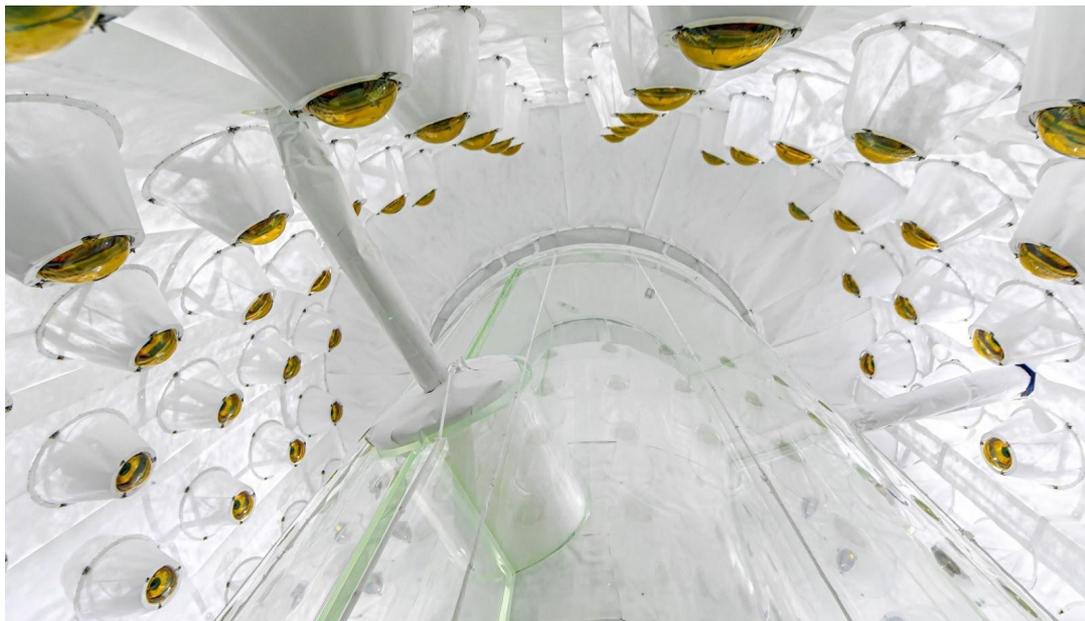


[Phys. Rev. D 108, 012010 \(2023\)](#)

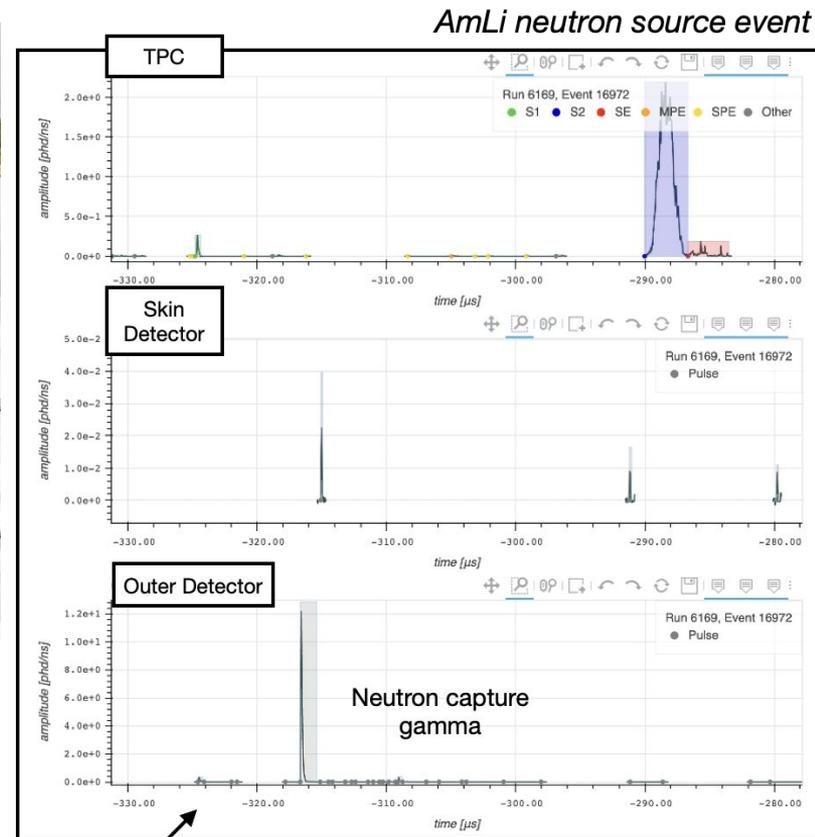


- Calibrations via a series of γ - and neutron sources (^{228}Th , $^{83\text{m}}\text{Kr}$, ^{57}Co , ^{22}Na , ^{54}Mn , DD, AmLi, AmBe,)
- Peaks from α -emitters bracket β -emitters that are fitted





- Neutrons are **main WIMP background**
- OD and Skin serve as active vetoes to reject neutrons
- Total neutron veto efficiency: $92 \pm 4 \%$

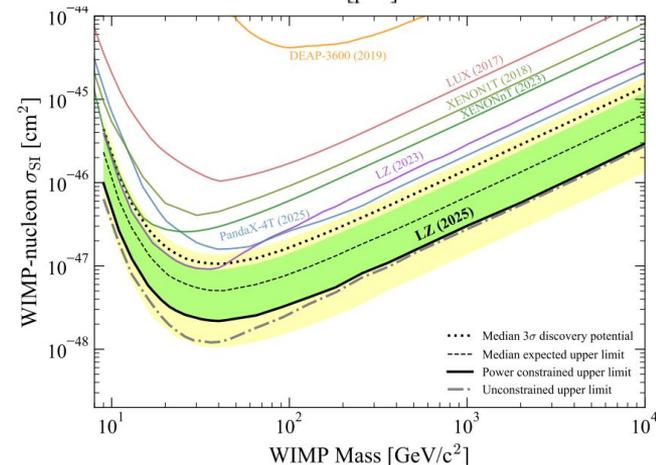
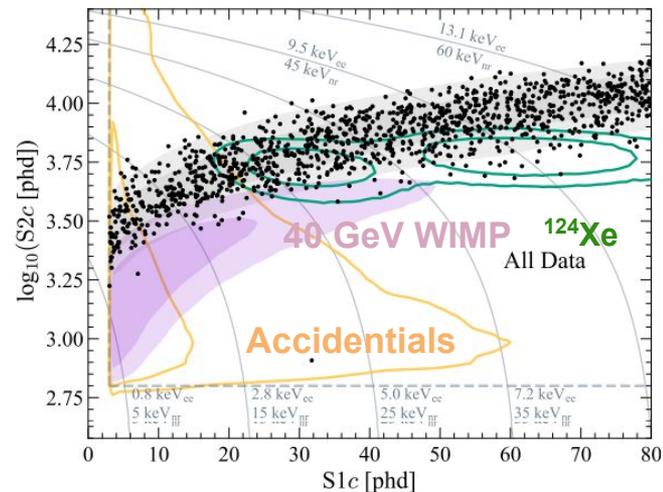


Proton recoil (LAB)

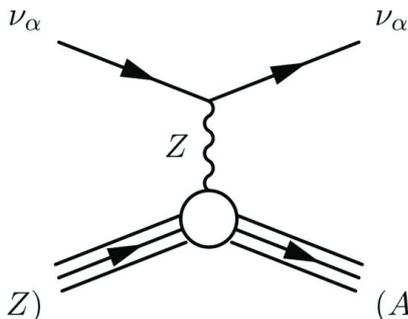
- Total **WS2024 exposure**: 280 days (4.2 tonne-years)
 - 220 days of new data
 - Combined with the WS2022 result (60 days)
- Base/WIMP analysis selections
 - $3 \text{ phd} \leq S1c < 80 \text{ phd}$, $S1 \text{ coinc.} \geq 3$,
 - $S2 \geq 645 \text{ phd}$ ($14.5e^-$), $S2c \leq 10^5 \text{ phd}$

- Profile likelihood ratio (PLR) analysis
- World's **best sensitivity for SI WIMP-type DM**
- Observed SI exclusion:

$$\sigma_{\text{DM}}^{\text{SI}} = 2.2 \times 10^{-48} \text{ cm}^2 \text{ at } 40 \text{ GeV}/c^2$$

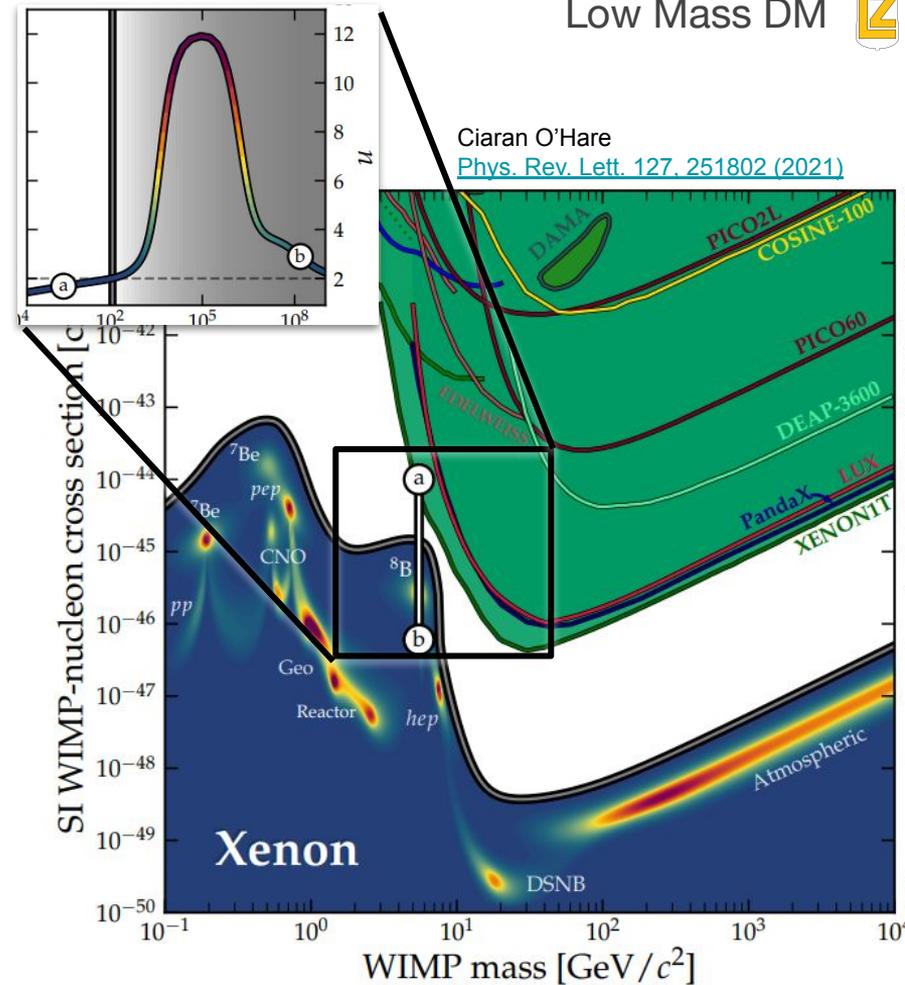


- **DM** might also be at **lower masses**
- **Neutrino fog also much higher (CEvNS from ^8B solar neutrinos)** at lower recoil energies
- Not a hard floor, a slowly **rising fog**
 - Energy and position dependent
 - Imagine: Going from $e^+e^- \rightarrow pp$
- First observation & novel SM measurements
- **First indications** by [PandaX-4T](#) & [XENONnT](#) ($< 3\sigma$)



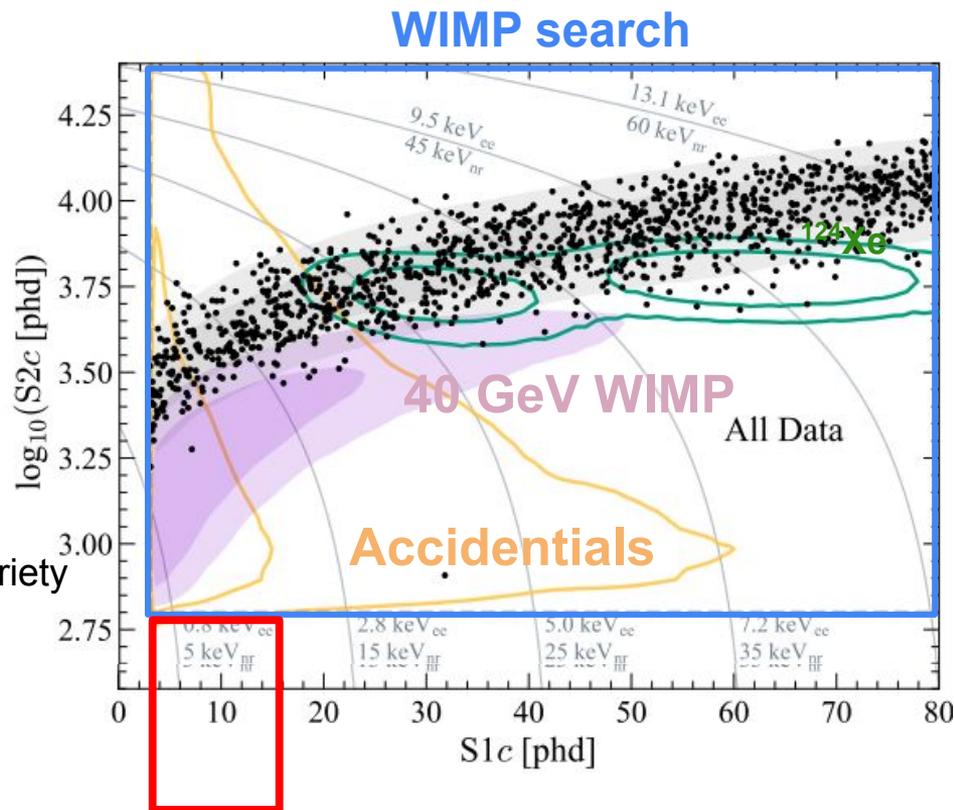
(A, Z)

(A, Z) Moriond Electroweak 2026



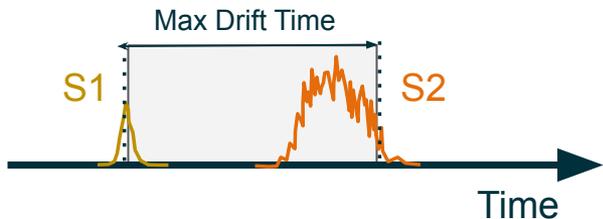
Ciaran O'Hare
[Phys. Rev. Lett. 127, 251802 \(2021\)](#)

- **Low energy signals** are very challenging
- Single scatters at **very low energies** (ROI about $1\text{-}6\text{ keV}_{\text{NR}}$)
- Different background composition:
 - Betas, gammas (**ER**) are **negligible**
 - Neutron backgrounds are $\ll 1$ event
 - Dominated by **accidental coincidence** events (very hard!)
- **Accidentals:** S1s and S2s mostly from a large variety of instrumental effects, e.g.
 - PMT dark counts
 - Charge loss near walls
 - Grid emissions
 - Photon/electron ‘trains’ after large energy deposits

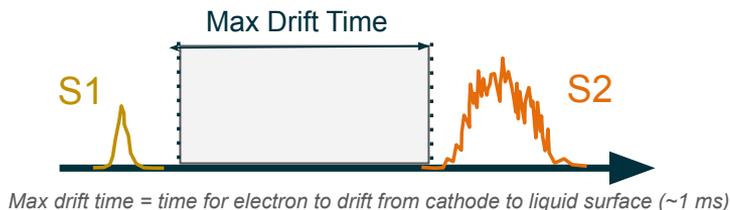


**Low Mass DM Search
(WS2025)**

Possible Physical Event



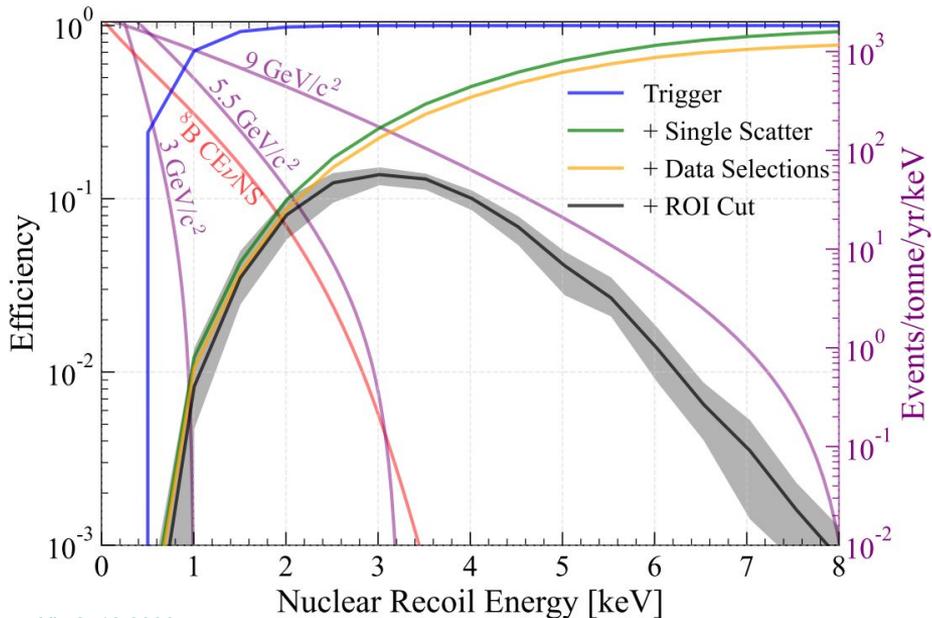
Definite Accidental Event



- Selection criteria developed on non-WIMP ROI background & calibration data
- Rejection of live time with detector instabilities, high TPC pulse rates
- Apply data quality, FV, base selections

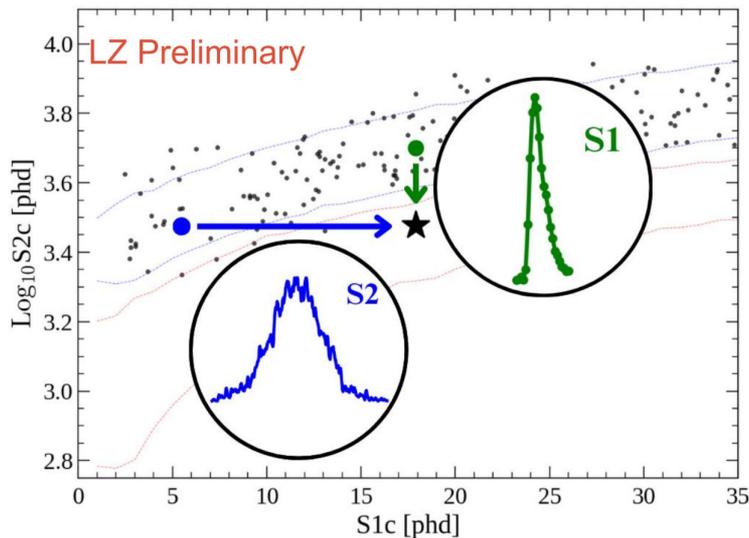
Event selections:

- Primary selection: 3-fold S1 coincidence
 $2 \leq S1c \leq 15$ phd, $3.5 \leq S2c \leq 14.5$ e⁻ (44.5 phd/e⁻)
- 417 live days, 5.1 T LXe FV

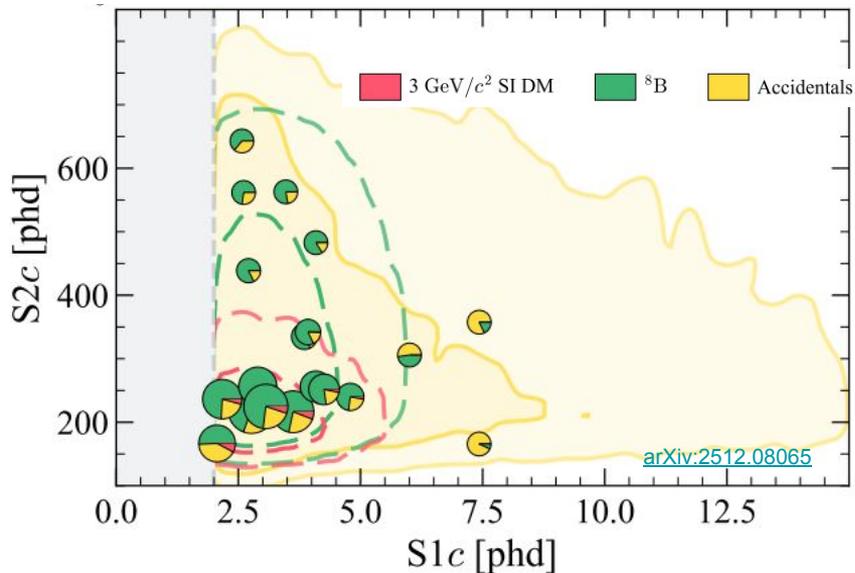


[arXiv:2512.08065](https://arxiv.org/abs/2512.08065)

- Slightly tighter 5.1 tonne fiducial volume for lowest background and uniform signals
- ‘Blind’ the data by injecting data drive **fake events**



Final data events, each pie chart represents the relative fit contributions and weight



• **Expected yields:**

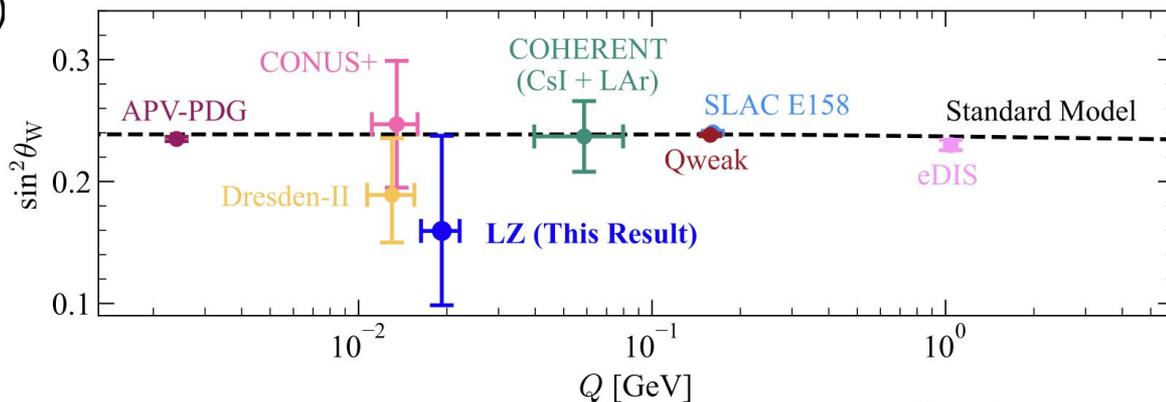
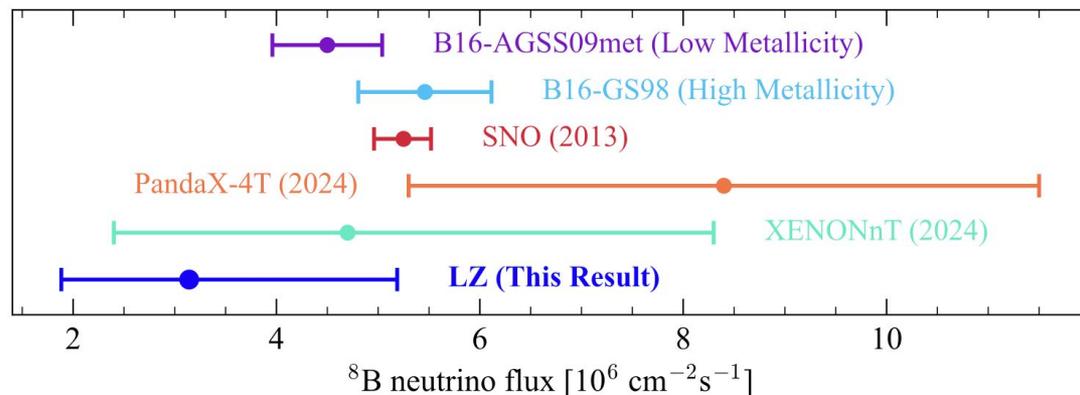
- **Bkgds:** $\sim 6.6 \pm 0.4$ events
- **⁸B CEvNS:** $20.6^{+8.9}_{-6.8}$ events

Obs: 19

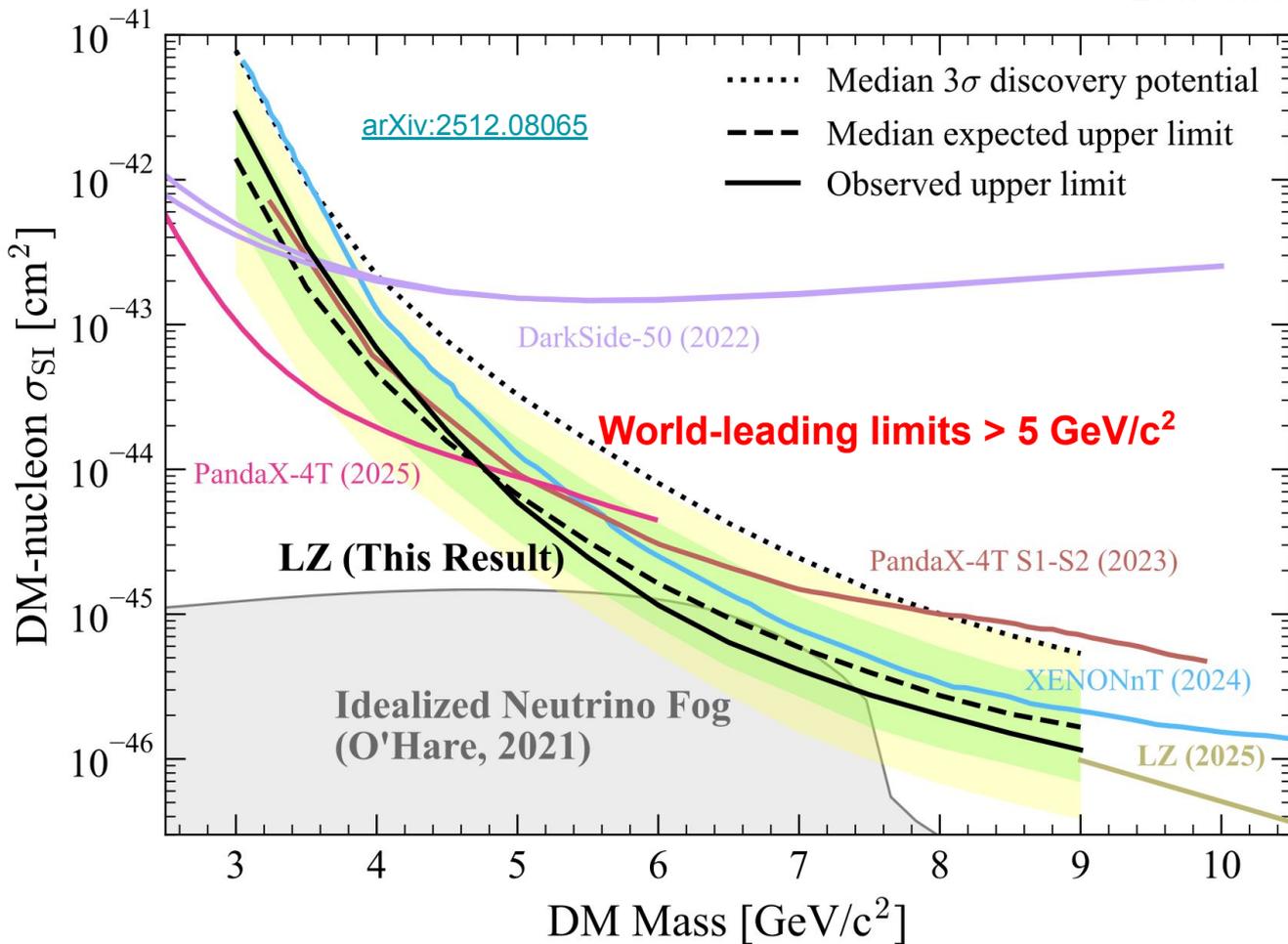
Fit: $18.9^{+7.0}_{-5.5}$

4.50 σ ⁸B observation!

- Measurement of ^8B CE ν NS can be used for ...
 - Flavor-independent measurement of the ^8B solar neutrino flux
 - Measurement of the weak-mixing angle ($\sin^2\theta_W$) at low momentum transfer Q
- Precision will improve with increased statistics and precision

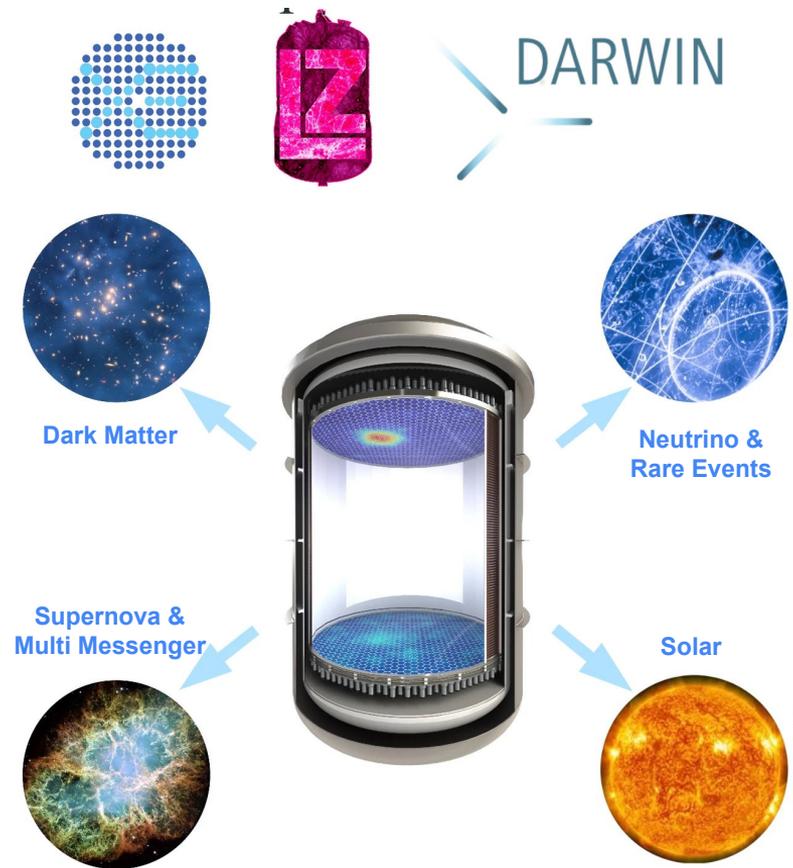
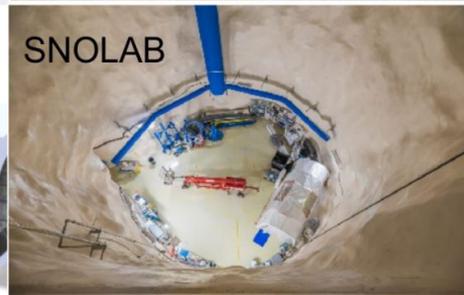
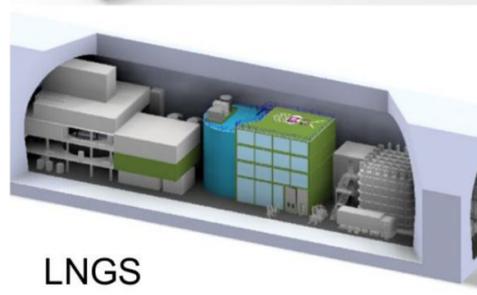


[arXiv:2512.08065](https://arxiv.org/abs/2512.08065)



- LZ is the largest & most sensitive WIMP liquid xenon TPC ([Phys.Rev.Lett. 135 \(2025\) 1, 011802](#))
- Other physics searches being performed in LZ...
 - Effective field theory coupling of DM to xenon nuclei ([Phys.Rev.D 109 \(2024\) 9, 092003](#), [Phys.Rev.Lett. 133 \(2024\) 22, 221801](#))
 - Solar axions, axion-like particles, dark photons ([Phys.Rev.D 108 \(2023\) 7, 072006](#), [arXiv:2511.17350](#))
 - Ultra heavy dark matter ([Phys.Rev.D 109 \(2024\) 11, 112010](#))
 - Cosmic ray boosted DM ([Phys.Rev.Lett. 134 \(2025\) 24, 241801](#))
 - Atmospheric millicharged particles ([Phys.Rev.Lett. 134 \(2025\) 24, 241802](#))
 - Muon Flux measurement ([arXiv.org:2602.16799](#))
 - ^{136}Xe $2\nu\beta\beta$ and $0\nu\beta\beta$, also DEC
- And many more...

- **XLZD** = XENON + LUX-ZEPLIN + DARWIN for a future 60-80 tonnes DM and neutrino experiment ([Eur. Phys. J. C \(2025\) 85: 1192](#))
- 17 countries, 76 institutions, 440+ members
- Four candidate sites:



- LZ has the **world leading sensitivity** for WIMP masses $> 5 \text{ GeV}/c^2$
- First **^8B CEvNS observation** with 4.50σ (exp. 6.70σ)
- LZ continues towards **1000 live days** of data and many more searches...
 - Investigating **extension beyond 2028**
- Beyond LZ we are planning a next generation experiment with the **XLZD collaboration**



- Black Hills State University
 - Brookhaven National Laboratory
 - Brown University
 - Center for Underground Physics
 - Edinburgh University
 - Fermi National Accelerator Lab.
 - Imperial College London
 - King's College London
 - Lawrence Berkeley National Lab.
 - Lawrence Livermore National Lab.
 - LIP Coimbra
 - Northwestern University
 - Pennsylvania State University
 - Royal Holloway University of London
 - SLAC National Accelerator Lab.
 - South Dakota School of Mines & Tech
 - South Dakota Science & Technology Authority
 - STFC Rutherford Appleton Lab.
 - Texas A&M University
 - University of Albany, SUNY
 - University of Alabama
 - University of Bristol
 - University College London
 - University of California Berkeley
 - University of California Davis
 - University of California Los Angeles
 - University of California Santa Barbara
 - University of Liverpool
 - University of Maryland
 - University of Massachusetts, Amherst
 - University of Michigan
 - University of Oxford
 - University of Rochester
 - University of Sheffield
 - University of Sydney
 - University of Texas at Austin
 - University of Wisconsin, Madison
 - University of Zürich
- US Europe Asia Oceania



LZ Collaboration Meeting at UCLA, March 2025

<https://lz.lbl.gov/>



US DOE



Science and Technology Facilities Council



Fundação para a Ciência e a Tecnologia



SANFORD UNDERGROUND RESEARCH FACILITY



Swiss National Science Foundation



Institute for Basic Science

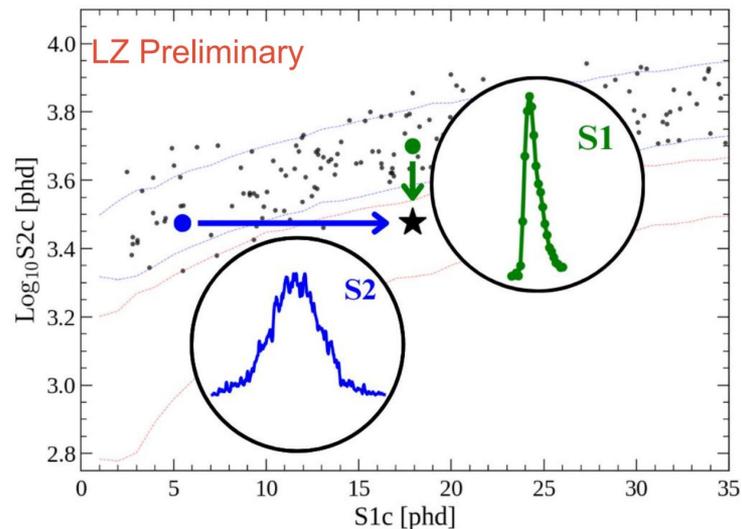
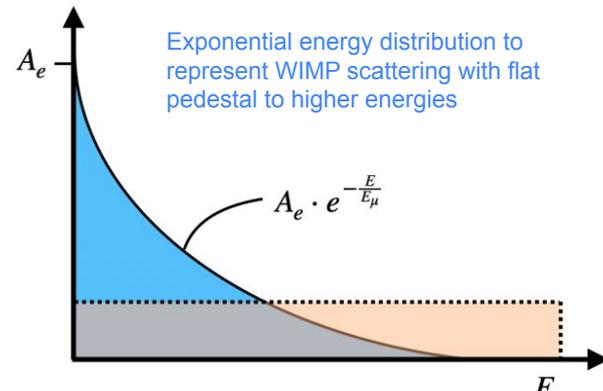


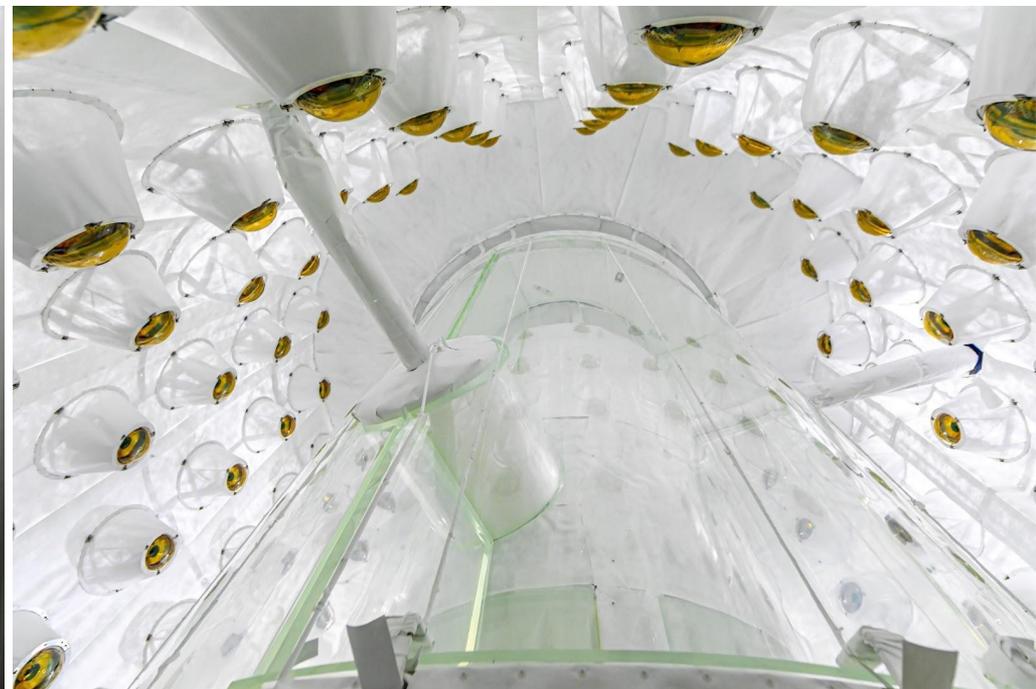
38 institutions, 250 scientists, engineers, and technical staff

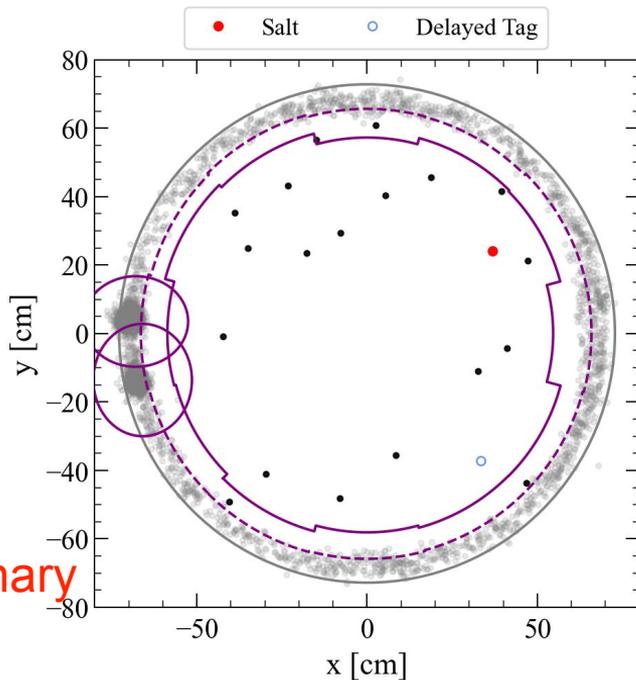


Backup

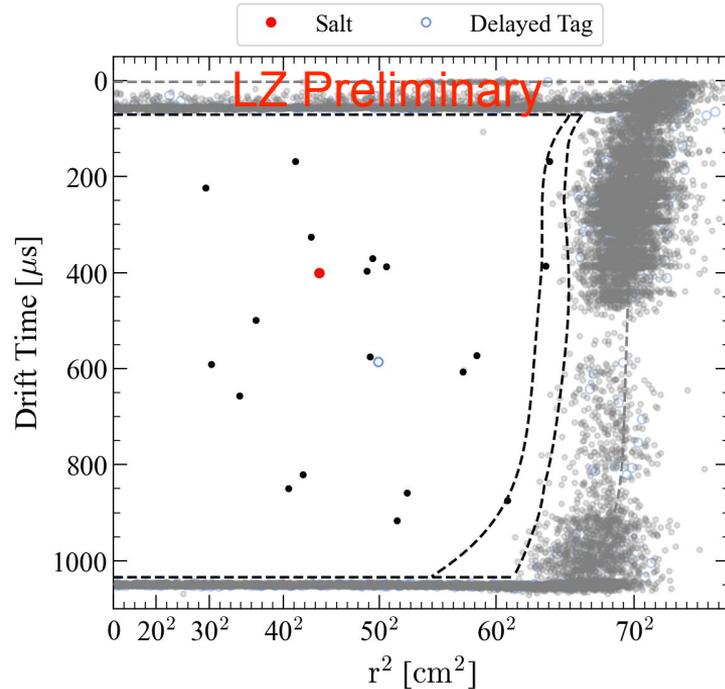
- **Salting:** ‘Blind’ the data by injecting fake events
- Generated by **combining real S1 & S2 pulses** from **calibration data** using exponential + flat spectrum
 - Events follow exponential+flat spectrum (exact parameters randomly generated, kept hidden)
 - Covers WIMP and higher-energy NR regions of
 - Injected salt rate chosen to match Poisson uncertainty on expected ${}^8\text{B}$ CE ν NS rate based on WS2024 NR model
- Identity of salt events revealed after analysis is finalised



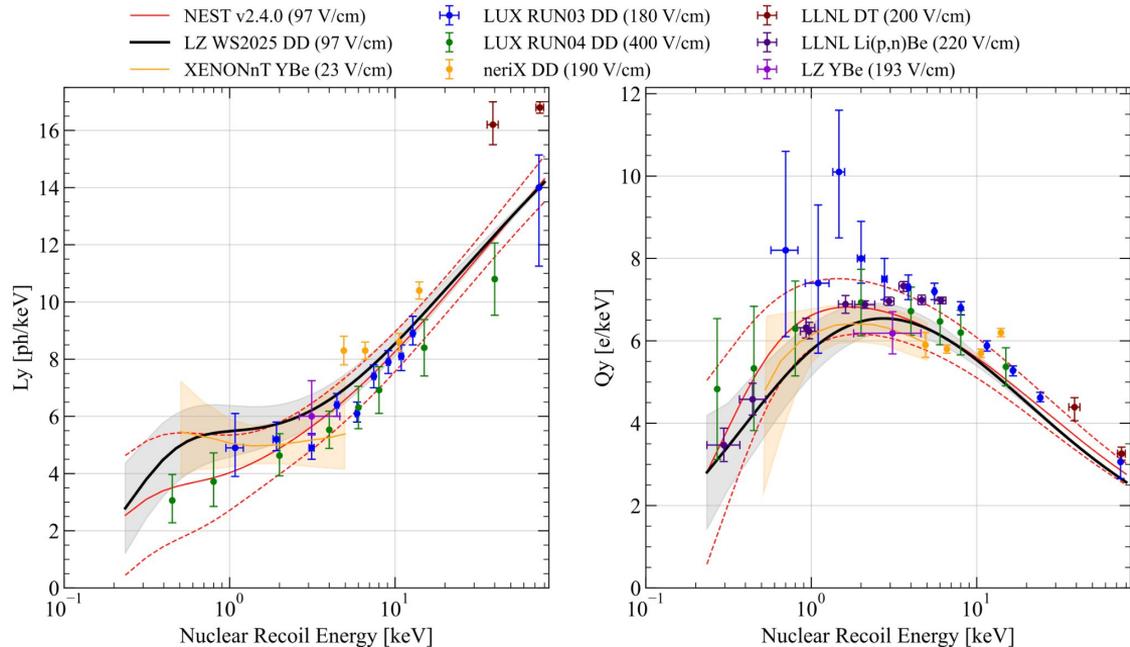
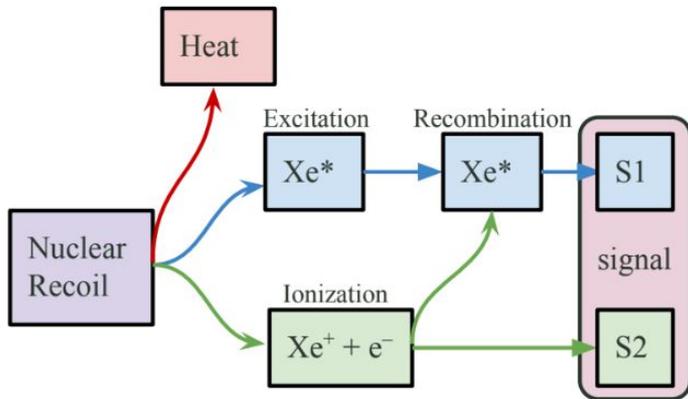




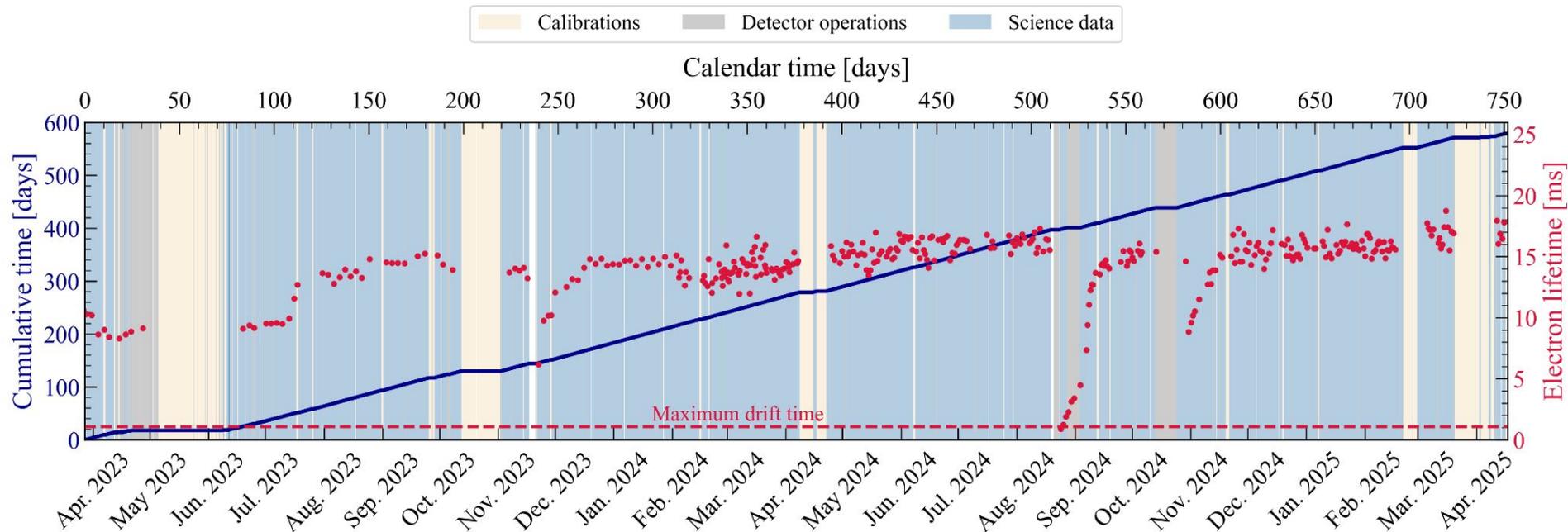
Z Preliminary



- Gray: TPC edge
- Purple: FVr extents at the top and bottom of the FVz
- Purple circles: Resistor cutout

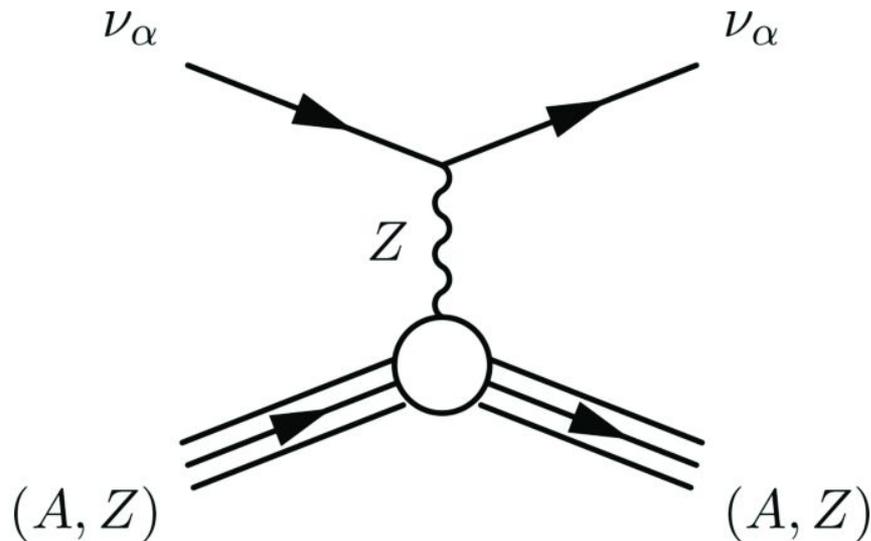


- Carefully calibrate low energy NR response and quantify systematic uncertainties
- Use mono-energetic 2.45 MeV DD neutron source, additional AmBe and AmLi calibrations
- Quantified systematic uncertainty associated with yields, fluctuations and detection efficiencies



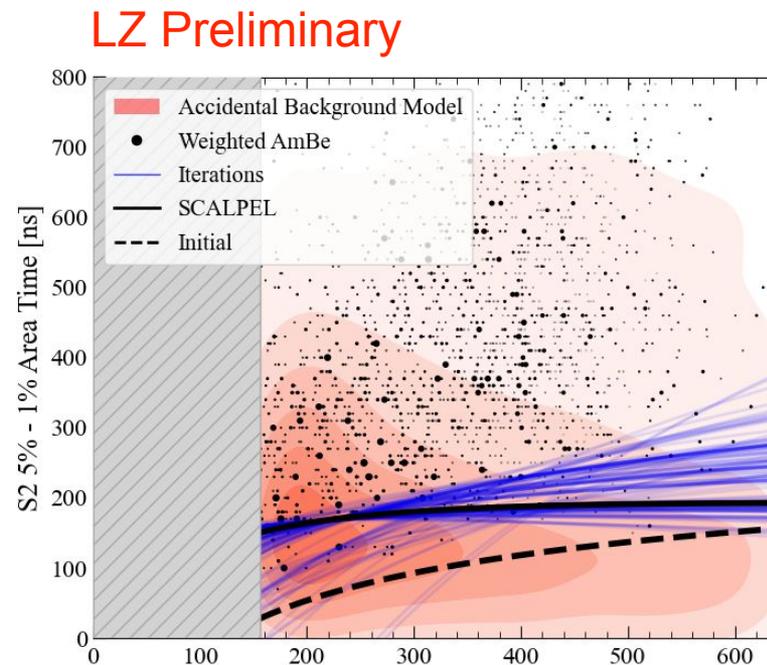
- Data taken from April 2023 - March 2025: 417 days
- Drift field of 97 V/cm, Extraction field of 3.4 kV/cm
- Dedicated AmBe and DD neutron calibrations
- Calibrated for uniformity in S1 and S2 response in time and space

- Well-predicted neutral current Standard Model process proposed in the 1970s where neutrino “sees” whole nucleus
 - Cross-section scales with neutrons squared
- First measured by the COHERENT experiment in 2017 using the Spallation Neutron Source
[Science 357, 6356 \(2017\)](#)
- Additional measurements with reactor anti-neutrinos, anti-neutrinos by Dresden-II
[PRL 129, 211802 \(2022\)](#), CONUS+ [Nature 643, 1229-1233 \(2025\)](#)

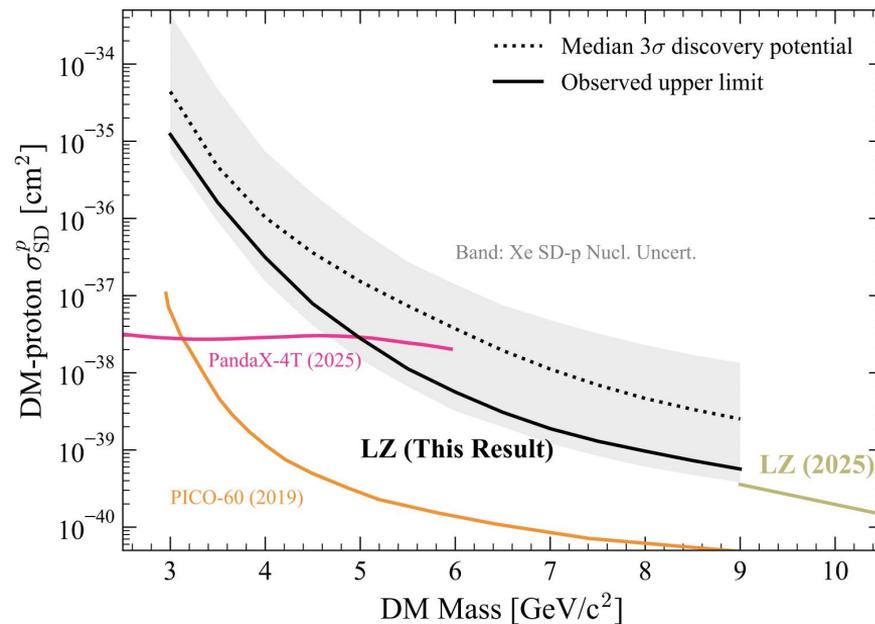
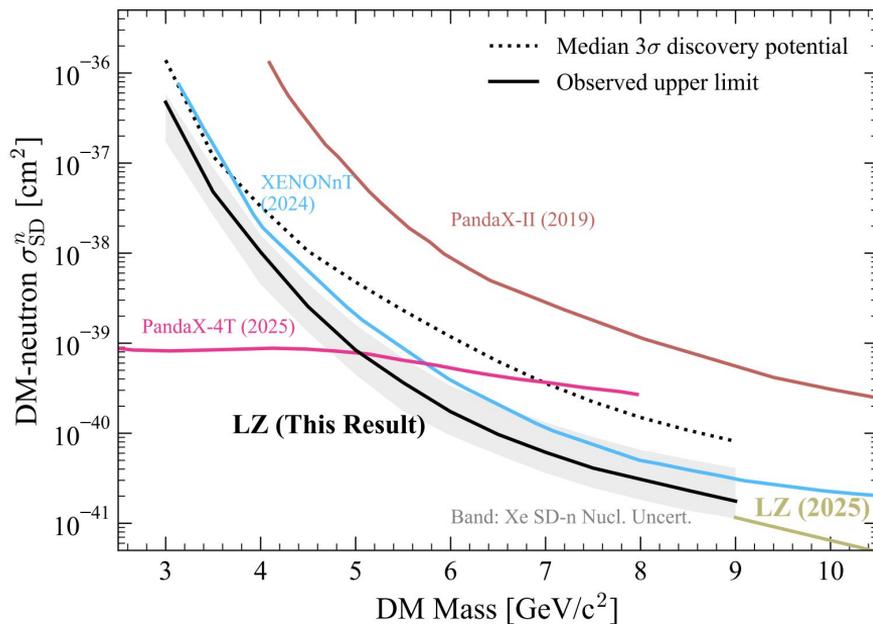


- First indications ($< 3\sigma$) of CE ν NS with solar neutrinos from PandaX-4T [PRL 133, 191001 \(2024\)](#), XENONnT [PRL 133, 191002 \(2024\)](#)

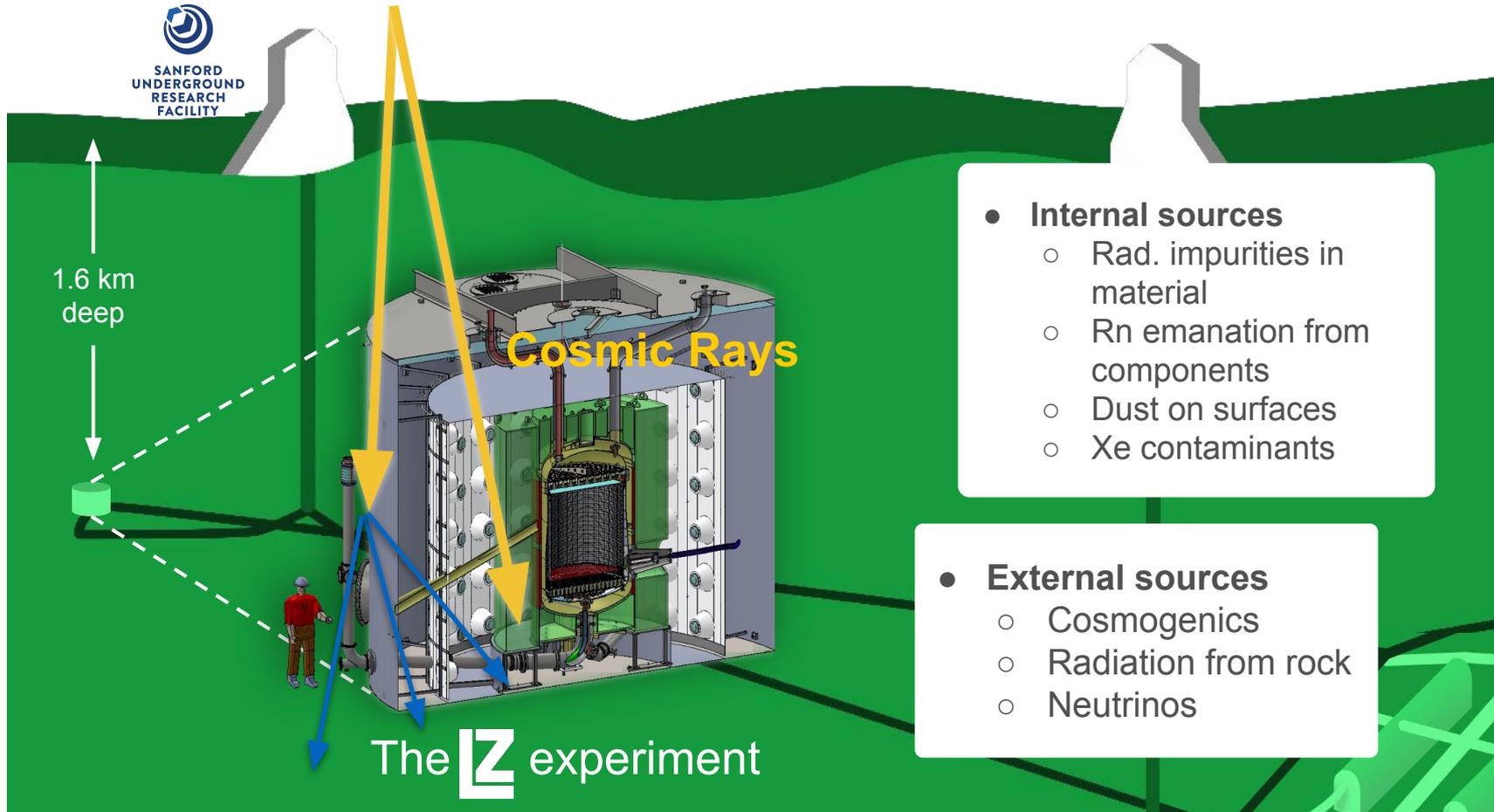
- Isolated S1s and S2s can **pile-up to fake a scatter**, where the S1 and S2 do not come from the same energy deposition
- Sources of S1s and S2s largely from a wide variety of instrumental effects, e.g.:
 - PMT dark counts
 - Charge loss near the TPC walls which degrades the S2 signal
 - Large photon/electron rates after big energy depositions
- Selections are designed and optimized using neutron (AmBe) calibration and fabricated background datasets
- Optimization technique (CMA-ES¹) used to simultaneously tune several selections across multiple parameter spaces
- Powerful, and not a black box!



¹Covariance Matrix Adaptation Evolution Strategy



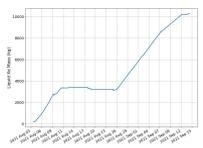
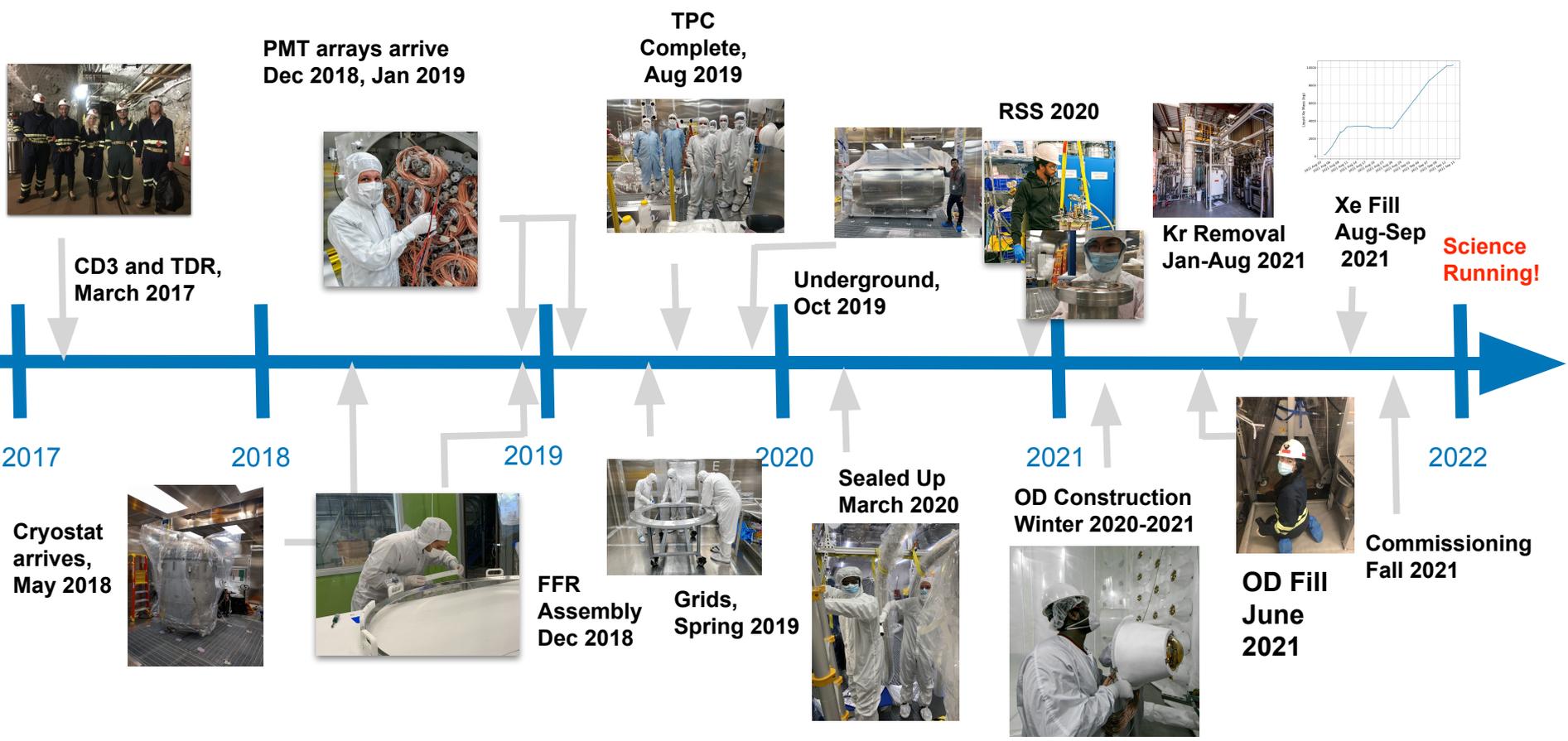
Constraining neutron-only or proton-only coupling to ^{129}Xe (26%) and ^{131}Xe (21%) which have unpaired neutrons leading to non-zero nuclear spin



- **Internal sources**
 - Rad. impurities in material
 - Rn emanation from components
 - Dust on surfaces
 - Xe contaminants

- **External sources**
 - Cosmogenics
 - Radiation from rock
 - Neutrinos

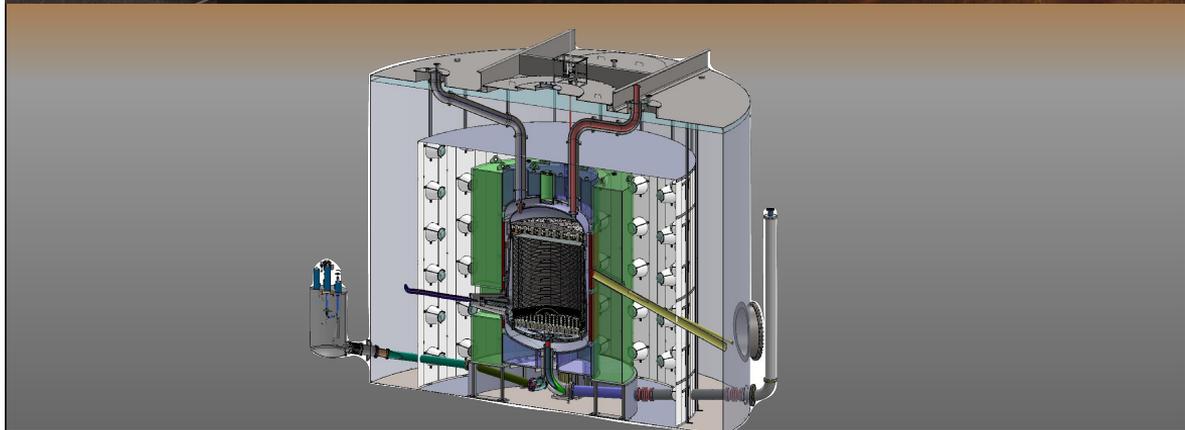
The **Z** experiment





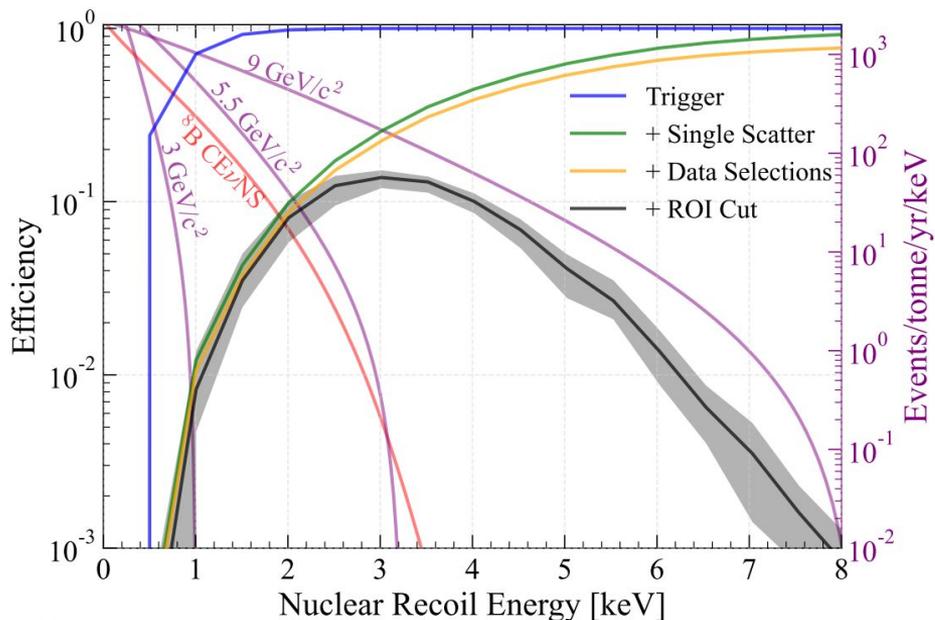
NO
YOND
WHILE
OR CAGE

CAUTION
NO ONE BEYOND
THIS POINT WHILE
WAITING FOR CAGE



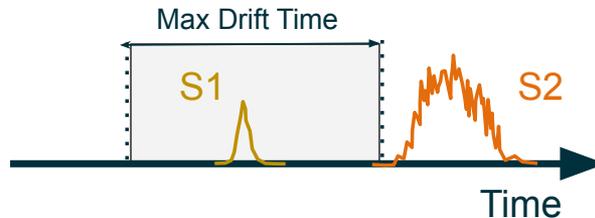
● **Event selections:**

- Primary selection: 3-fold S1 coincidence
 $2 \leq S1c \leq 15$ phd, $3.5 \leq S2c \leq 14.5$ e⁻ (44.5 phd/e⁻)
- Apply data quality, FV, base selections

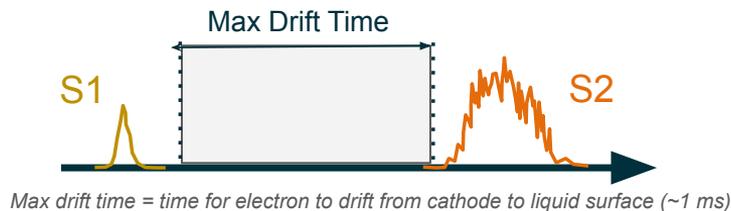


[arXiv:2512.08065](https://arxiv.org/abs/2512.08065)

Possible Physical Event



Definite Accidental Event



- Selection criteria developed on non-WIMP ROI background & calibration data
- Rejection of live time with detector instabilities, high TPC pulse rates
- **Key numbers:**
 - 417 live days, 5.1 T LXe FV